



A new method for estimating UV fluxes at ground level in cloud-free conditions

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Abstract. A new method has been developed to estimate the global and direct solar irradiance in the UV-A and UV-B at ground level in cloud-free conditions. It is based on a resampling technique applied to the results of the k -distribution method and the correlated- k approximation of Kato et al. (1999) over the UV band. Its inputs are the aerosol properties and total column ozone that are produced by the Copernicus Atmosphere Monitoring Service (CAMS). The estimates from this new method have been compared to instantaneous measurements of global UV irradiances made in cloud-free conditions at five stations at high latitudes in various climates. For the UV-A irradiance, the bias ranges between -0.8 W m^{-2} (-3% of the mean of all data) and -0.2 W m^{-2} (-1%). The root mean square error (RMSE) ranges from 1.1 W m^{-2} (6%) to 1.9 W m^{-2} (9%). The coefficient of determination R^2 is greater than 0.98. The bias for UV-B is between -0.04 W m^{-2} (-4%) and 0.08 W m^{-2} ($+13 \%$) and the RMSE is 0.1 W m^{-2} (between 12 and 18 %). R^2 ranges between 0.97 and 0.99. This work demonstrates the quality of the proposed method combined with the CAMS products. Improvements, especially in the modeling of the reflectivity of the Earth's surface in the UV region, are necessary prior to its inclusion into an operational tool.

1 Introduction

Solar ultraviolet (UV) radiation at the Earth's surface has beneficial and adverse effects on human health (Juzeniene et al., 2011). For instance, UV radiation is a principal source of vitamin D, while the excess UV exposure is a risk factor for skin cancers, cataracts and immunosuppression. The wavelength dependence of these effects is typically characterized by action spectra. The most widely used one is the standardized action spectrum for erythema, which is also known as the CIE (Commission Internationale de l'Eclairage) spectrum (McKinlay and Diffey, 1987). There are also other action spectra related to skin cancer and melanoma (de Gruijl et al., 1993; Setlow et al., 1993). Emphasis has been placed mostly on the assessment of the solar UV erythema irradiance and the derived quantity, the UV index, which is a very popular quantity to inform the public about UV levels. The UV index is also used in campaigns promoting safe Sun exposure. While the UV-B [280, 320] nm band is the major contributor to erythema UV, interest is growing in the role of UV-A [320, 400] nm and UV [280, 400] nm on various diseases, such as viral infections (Norval, 2006), multiple sclerosis (Orton et al., 2011), Parkinson's disease (Kravietz et al., 2017), eye diseases (Delcourt et al., 2014), skin cancer (Coste et al., 2015; Fortes et al., 2016) or thyroid cancer (Mesrine et al., 2017), among many others (Juzeniene et al., 2011; Norval and Halliday, 2011).

Ground-based spectroradiometers are one of the means to monitor the intensity of solar UV fluxes. Such measurements are rare due to the high costs of the instruments, operations and maintenance. To overcome this scarcity, many researchers have looked for proxies and have studied the relationship between UV radiation and the surface downwelling solar radiation integrated from 280 to 2800 nm, called broadband radiation, since the latter is measured at a greater number of stations or can be estimated at any place from satellite images (Blanc et al., 2011; Lefèvre et al., 2014). Several empirical relationships have been published that relate, with the knowledge of the total ozone column (TOC), the broadband irradiance to the erythemal UV (den Outer et al., 2010; Calbó et al., 2005) or the UV-A, UV-B or UV irradiance (Aculinin et al., 2016; Canada et al., 2003; Foyo-Moreno et al., 1998).

An alternative way is the use of an appropriate radiative transfer model (RTM) together with accurate inputs describing the state of the atmosphere in cloud-free conditions and the properties of the ground surrounding the instrument, such as libRadtran (Emde et al., 2016; Mayer and Kylling, 2005). A comparison between 1200 measured UV spectra and estimates made with a previous version called uvspec – now part of libRadtran – with only ozone and aerosol optical properties as inputs yielded very good performance for simulating the UV irradiance under cloud-free conditions (Mayer et al., 1997). The relative biases ranged between -11 and $+2\%$ for wavelengths between 295 and 400 nm and solar zenithal angles (SZA) up to 80° . Using measurements from sites in Finland, Norway and Sweden, Lindfors et al. (2007, 2009) showed that erythemal UV and spectral UV irradiances can be accurately modeled using libRadtran and broadband radiation, TOC, the total water vapor column from the ERA-40 data set, the surface albedo as estimated from snow depth and the altitude of the location as input.

RTMs are usually computationally expensive; hundreds of spectral calculations are required to compute the UV irradiance in an RTM. Strategies have been built to reduce the amount of calculations. Among them are the k -distribution method and correlated- k approximation by Kato et al. (1999). The approach was originally designed for the calculation of the broadband solar irradiance. It consists in calculating the total solar irradiance, i.e., integrated between 240 and 4000 nm, with only 32 spectral calculations in the spectral range between 240 and 4606 nm. The operational McClear model estimating the total irradiance in cloud-free conditions accurately reproduces the irradiance computed by libRadtran based on the Kato et al. (1999) approach (Lefèvre et al., 2013). The McClear model uses several look up tables computed by libRadtran for selected values of inputs and provides the irradiance at each of the 32 spectral intervals. Hereafter, these 32 spectral intervals are named Kato bands and abbreviated KB with the number in subscript. Four KBs cover the whole UV range: KB₃ [283, 307] nm, KB₄ [307, 328] nm, KB₅ [328, 363] nm and KB₆ [363, 408] nm. In KB₁

and KB₂, atmospheric ozone attenuates the radiation before it reaches the ground.

Wandji Nyamsi et al. (2014) compared atmospheric transmissivities obtained by the Kato et al. (1999) approach against those obtained by spectrally resolved computations using two RTMs in each of the 32 KBs. These calculations were performed for a set of 200 000 realistic atmospheres and clouds. As for the UV band, the authors found that the Kato et al. (1999) approach offers very accurate estimates of irradiances in KB₅ and KB₆. On the contrary, a very large underestimation of the transmissivity was observed in KB₃ [283, 307] nm and KB₄ [307, 328] nm by respectively -93 and -16% in relative value and exhibits a relative root mean square error (RMSE) of 123 and 17 % in clear-sky conditions. Similar relative errors are observed for cloudy conditions. This is due to the assumption made by Kato et al. (1999) that in these bands a single ozone cross section at the central wavelength is sufficient to accurately represent the absorption by ozone over the whole interval. In a subsequent work, Wandji Nyamsi et al. (2015b) have proposed a novel parameterization using more than one single ozone cross sections which accurately represents the transmissivity due to ozone absorption. The novel parameterization of the transmissivity using more quadrature points yields maximum errors of respectively 0.0006 and 0.0143 for intervals KB₃ and KB₄. Version 2.0.1 of libRadtran (Emde et al., 2016) includes this correction and has been used in this study.

The KBs do not fit the UV spectral ranges exactly and a spectral resampling is necessary. This is the subject of the present article. The concept of the novel method is to determine several 1 nm spectral bands whose atmospheric transmissivities are correlated to those of the KB and are then used in a linear interpolation process to compute the UV irradiance. The method is empirically implemented by the means of libRadtran in cloud-free conditions. The concept has already been tested for photosynthetically active radiation simulated by libRadtran (Wandji Nyamsi et al., 2015a). Now, the concept is tested for actual UV fluxes. This work is part of a larger project whose overarching goal is to create an operational tool for estimating UV fluxes. In particular, it exploits the recent results on aerosol properties and TOC produced by the Copernicus Atmosphere Monitoring Service (CAMS) for any location and any time after 2003. The performance of the novel method is assessed by a comparison against high-quality measurements of UV fluxes performed in cloud-free conditions. Stations have been selected to fulfill two main constraints. The first one is that the measurement has to be carried out during a cloud-free instant, meaning that it either should be clearly marked or should use an algorithm for selecting cloud-free instants, which most of the time requires broadband measurements as inputs. The second one is that high-quality control and assurance should be applied to the measurement. Following these constraints, five stations which are located at high latitudes were selected.

2 Description of measurements used for comparison

Ground-based measurements were collected from three sites of the UV network of the National Science Foundation (NSF) of the USA and two sites of the Finnish Meteorological Institute (FMI). Table 1 reports the geographical coordinates of the stations, time period of data and their source, type of instruments, spectral interval and step of measurements.

The Barrow site is located approximately 6 km northeast of the Barrow city in Alaska, on the coast of the Chukchi Sea, part of the Arctic Ocean, usually covered by ice between November and July. The snow cover of the surroundings of the station extends from October until June. According to Bernhard et al. (2008), the effective UV albedo of the surface reaches its maximum of approximately 0.8 during March–April. It decreases to 0.05 in summer from August until September.

The Sodankylä site is approximately 6 km south of the village of Sodankylä. The site is located in the vicinity of the river Kitinen, and the surroundings are boreal pine forest and large peatland areas. A permanent snow cover is present during the winter and the annual number of snowy days is on average 190. The snow cover starts accumulating in October or November and melts away during May almost every year. The effective UV albedo follows a seasonal variation due to the snow cover. It ranges from very low values in boreal summer up to 0.65 in winter (Arola et al., 2003).

Jokioinen Observatory is located in a fairly flat rural area in the southwest of Finland surrounded by fields of agriculture and a boreal forest. The number of snowy days is typically 130. The snow conditions in Jokioinen vary from year to year and also within each winter. At the earliest, snow may appear at the end of October or early November, while it typically melts away in March or April. The effective UV albedo is highly variable; it may rise up to 0.58 in boreal winter but more typical values are 0.2 to 0.5 (Lindfors et al., 2007).

Palmer is situated on Anvers Island, which is on the western side of the Antarctic Peninsula. The ocean surrounding the island is frozen during austral winter and usually ice-free in summer. According to Bernhard et al. (2005), the effective UV albedo varies between 0.6 and 0.95 occurring from August until November and then decreases down to 0.3 to 0.5 after snowmelt. It is large even in austral summer because of the glaciers surrounding the site.

McMurdo is a coastal site located on Ross Island, a volcanic island of Antarctica surrounded by a persistence of the ice sheet. The surroundings of the station are mostly made of dark volcanic rocks. McMurdo has an annual cycle of change in effective UV albedo. It ranges between 0.54 (March) and 0.99 (October) (Bernhard et al., 2006).

WMO (2008) reports that uncertainties associated with the measurements in UV by spectrometers are difficult to estimate precisely. Beyond the technical specifics of the site itself, several errors may occur in the calibration of the instrument that include (i) the uncertainties associated

with irradiance transfer standards, (ii) the stability of instruments over time and (iii) imperfect directional response. The WMO guide estimates that a 5 % measurement uncertainty at 300 nm can be achieved only under the most rigorous conditions at the present time.

The data provided by the NSF are available online and can be downloaded freely. Only data of version 2, which have been corrected for the instruments' cosine error, have been selected to ensure higher accuracy. Data measured during clear-sky conditions are flagged (flag "CS"). The number of clear-sky instants is reported in Table 1. Integrated irradiances in the UV-A and UV-B range are available and have been downloaded from the website <http://uv.biospherical.com/Version2>.

The data for the two Finnish sites have been corrected for all known errors following the routine spectral UV data processing procedure of the FMI (Lakkala et al., 2008; Mäkelä et al., 2016). The irradiance scale of the FMI's Brewer spectrophotometers is traceable to that maintained by the Finnish National Standards Laboratory at VTT MIKES Metrology and maintained by a rigorous schedule for measurements of primary standard, secondary standard and working standard lamps (Heikkilä et al., 2016a). In this work, the measured UV spectra were first deconvoluted and then convoluted with a standard triangular slit function with a full width at half maximum of 1 nm and extrapolated using the SHICrvm software package (http://www.rivm.nl/en/Topics/U/UV_ozone_layer_and_climate/SHICrvm) (Slaper and Koskela, 1997; Slaper et al., 1995) to cover the full UV spectrum as explained in Heikkilä et al. (2016b). Spectral irradiances are integrated over the UV-B and UV-A. The uncertainties related to the extrapolation are less than 2 and 3 % in the integrated UV-A, for the Jokioinen and Sodankylä Brewers respectively when averaged over daily time window. However, for individual spectra the uncertainties are estimated to be somewhat higher, up to 5–6 % for Sodankylä Brewer with the highest measured wavelength at 325 nm (H. Slaper, personal communication, 2017).

In addition, at both Finnish stations, direct, diffuse and global broadband irradiances are measured every 1 min, with the global irradiance being the sum of the direct and diffuse irradiances on a horizontal plane. These series of data are exploited for selecting cloud-free instants by using the very restrictive algorithm proposed by Lefèvre et al. (2013). The latter is made of two successive filters. The first one is a constraint on the amount of diffuse irradiance with respect to the global irradiance since the direct irradiance is usually prominent in cloud-free conditions. The second filter analyzes the temporal variability in the global irradiance normalized by the irradiance received at the top of the atmosphere and by a typical air mass since this quantity should be steady for several hours in cloud-free conditions. We assume that a cloud-free instant detected by analyzing broadband irradiances is also cloud-free for the spectral measurements. It is possible that UV is affected by the presence of scattered cloudiness,

Table 1. Description of stations used for validation, ordered by decreasing latitude.

Station	Barrow	Sodankylä	Jokioinen	Palmer	McMurdo
Source	NSF	FMI	FMI	NSF	NSF
Latitude (°)	71.32	67.37	60.82	−64.77	−77.83
Longitude (°)	−156.68	26.63	23.50	−64.05	166.67
Altitude (m)	8	179	104	21	183
Instrument	SUV-100 spectro-radiometer	Brewer spectro-radiometer	Brewer spectro-radiometer	SUV-100 spectro-radiometer	SUV-100 spectro-radiometer
		MK-II no. 037	MK-III no. 107		
Average acquisition frequency	4 spectra per hour	between 1 and 2 spectra per hour	between 1 and 2 spectra per hour	4 spectra per hour	4 spectra per hour
Spectral range (nm)	280–600	290–325	286.5–365	280–600	280–600
Step (nm)	0.2 in UV-B to 1.0 in visible	0.5	0.5	0.2 in UV-B to 1.0 in visible	0.2 in UV-B to 1.0 in visible
Period	2005-01 to 2010-11	2007-01 to 2011-12	2007-01 to 2008-12	2005-01 to 2010-09	2005-01 to 2010-02
NCFI*	4293	2590	1140	1736	10 175

* NCFI: number of cloud-free instants.

which may go unnoticed in the broadband range, and that the retained series of cloud-free instants for broadband may comprise cloudy instants for UV. Given the high selectivity of the algorithm of Lefèvre et al. (2013), we believe that such cases are rare and that the conclusions will be unaffected as a whole.

3 Description of the new method

In brief, the method combines the fluxes estimated by libRadtran in the four KBs and performs a resampling of these fluxes for retrieving UV fluxes. For all the radiative transfer simulations, a plane-parallel atmosphere was assumed and the DISORT 2.0 (discrete ordinate technique) algorithm (Stamnes et al., 2000) with 16 streams was selected to solve the radiative transfer equation because several articles have demonstrated the accuracy of its results when compared to robust and more time-consuming solvers.

3.1 Inputs to libRadtran

In cloud-free conditions, UV irradiance at ground level depends mostly on the SZA (θ_s); the ground albedo; the total column content of ozone; the vertical profile of ozone, temperature, pressure, density and volume mixing ratio for gases as a function of altitude; aerosol optical depth (AOD); the Ångström coefficient; and aerosol type and the elevation of the ground above sea level. As the method shall be used operationally, the sources of these inputs have been selected to allow estimation of UV irradiance at any location and any time.

The Copernicus Atmosphere Monitoring Service of the European Commission provides aerosol properties together with physically consistent TOC for any place and any time after 2003. Along with TOC, the AOD at 550 nm, Ångström coefficient and aerosol type are collected from this source of data following exactly the path of the McClear model

(Lefèvre et al., 2013). θ_s is given by the SG2 algorithm for the Sun position and angles (Blanc and Wald, 2012). Ground elevation is extracted from the Shuttle Radar Topography Mission database and has been downloaded from the website <http://srtm.csi.cgiar.org/SELECTION/inputCoord.asp>.

The albedo is the ratio of the upwelling to downwelling flux at the surface and is the integral of the bidirectional reflectance distribution function (BRDF), which depends on the surface-type, its roughness and the wavelength of the impinging radiation. A few institutes provide BRDF products in the UV range or Lambertian equivalent reflectivity (e.g., Herman and Celarier, 1997) of the Earth's surface with a coarse resolution of 0.5 or 1°. In absence of the ideal solution – BRDF parameters in the UV range available worldwide with a grid cell of 0.05° or better – an approximate solution has been adopted by using the so-called shortwave [250, 5000] nm BRDF parameters proposed by Blanc et al. (2014). The US National Aeronautics and Space Administration (NASA) provides worldwide maps of the BRDF parameters that are derived from the MODIS (Moderate Resolution Imaging Spectroradiometer) instrument (Schaaf et al., 2002). Blanc et al. (2014) have created a series of maps of the MODIS BRDF parameters for each calendar month for the shortwave albedo with no missing values at a spatial resolution of 0.05°. In addition, these authors proposed a method for computing the albedo simultaneously for direct and global irradiances. These maps and this method are those used by the McClear model (Lefèvre et al., 2013). As a first approximation, the UV albedo is assumed to be spectrally constant and equal to the shortwave albedo. This assumption may result in biases depending on the surface. For example, in the case of snow surface, Varotsos et al. (2014) reported from many aircraft measurements that spectral albedo exhibits a tendency to decrease with increasing wavelength of about 0.7 in UV to about 0.4 in the near-infrared independently of the sky conditions. Therefore, the albedo integrated over the spectrum becomes less than 0.7, resulting in the un-

derestimation in UV albedo, hence in a lesser contribution to diffuse UV irradiance, and therefore in the underestimation of the global UV.

3.2 Resampling technique

Let λ be the wavelength, G_λ the global spectral irradiance at the surface and B_λ the direct normal spectral irradiance, i.e., the irradiance received from the direction of the Sun at the surface on a plane normal to the Sun's rays. The irradiance in a given interval $[\lambda_1, \lambda_2]$ is given by

$$G_{\lambda_1\lambda_2} = \int_{\lambda_1}^{\lambda_2} G_\lambda d\lambda. \quad (1)$$

Similar expressions may be obtained for irradiances in UV-A, UV-B, UV or over KB_j . For example, the UV irradiance G_{UV} and the direct normal irradiance B_{UV} or the irradiance G_{KB4} in KB_4 are given by

$$G_{UV} = \int_{280}^{400} G_\lambda d\lambda, \quad (2)$$

$$B_{UV} = \int_{280}^{400} B_\lambda d\lambda, \quad (3)$$

$$G_{KB4} = \int_{307}^{328} G_\lambda d\lambda, \quad (4)$$

where λ is expressed in nm.

The KBs do not fit the UV spectral ranges exactly. For example, the UV-B is covered by KB_3 and a part of KB_4 . One solution to estimate the irradiance in a UV interval is the use of weighted sums based on the overlap between KB_j and this interval. Another technique is adopted here whose concept is to determine several 1 nm spectral bands NB_i whose transmissivities are correlated to those of the KB_j and are then used in a linear interpolation process to compute the UV irradiance. A similar approach has been used by Wandji Nyamsi et al. (2015a) for the calculation of the photosynthetically active radiation with better results than a weighted sum.

If one assumes that the optical properties of the atmosphere do not change over a given NB_i , the integrals of Eqs. (2)–(4) may be replaced by Riemann sums over NB_i . For example, if λ_i denotes the central wavelength of NB_i , G_{UV} can be approximated by

$$G_{UV} = \sum_{i=1}^{120} G_{\lambda_i}. \quad (5)$$

If one defines the clearness index KT_i as

$$KT_i = \frac{G_{\lambda_i}}{E_{o_i} \cos(\Theta_s)}, \quad (6)$$

where E_{o_i} is the irradiance at the top of atmosphere on a plane normal to the Sun's rays for NB_i for a given instant t , Eq. (5) becomes

$$G_{UV} = \cos(\Theta_s) \sum_{i=1}^{120} E_{o_i} KT_i. \quad (7)$$

Similarly, the clearness index KT_{KB_j} for KB_j is given by

$$KT_{KB_j} = \frac{G_{KB_j}}{E_{o_{KB_j}} \cos(\Theta_s)}. \quad (8)$$

The solar spectrum of Gueymard (2004) available in libRadtran was combined with the algorithm SG2 for computing the Sun position (Blanc and Wald, 2012) to yield E_{o_i} and $E_{o_{KB_j}}$.

For the method presented here, we assume that simple and accurate relations, e.g., affine functions, may be found between each KT_{KB_j} and a subset of several KT_i , called KT_k hereafter. Then it may be possible to interpolate linearly between KT_k to obtain an estimate of KT_i for each λ_i . By summing the products of E_{o_i} and these interpolated KT_i 's over a given spectral interval, it is then possible to compute the corresponding irradiance. This is the principle of the resampling technique. The set of NB_k is selected at the beginning of the procedure and the same set is used for all processing. The number of NB_k should be as small as possible in order to decrease the amount of calculations but still large enough to allow a good accuracy.

The selection of NB_k is empirically determined by means of libRadtran. The current approach is empirical with no guarantee that the selected set of NB_k is the optimum. It could have been possible to use some mathematical optimization tools. This is not a straightforward process as the cost function should take into account that the number of NB_k is unknown a priori.

A set of 60 000 clear-sky atmospheric states was built by means of the Monte Carlo technique in order to select NB_k . Each state comprises the nine variables described above, and the value of each variable was randomly selected by taking into account their modeled marginal distribution. The distributions proposed by Lefèvre et al. (2013) and Oumbe et al. (2011) established from observations were adopted here (Table 2). More specially, the uniform distribution is chosen as a model for marginal probability for all variables except AOD, the Ångström exponent coefficient, and total column content of ozone. The chi-square law for AOD, the normal law for the Ångström exponent coefficient, and the beta law for TOC have been selected. The selection of these parametric probability density functions and their corresponding parameters have been empirically determined from the analyses of the observations made in the Aerosol Robotic Network for aerosol properties and from meteorological satellite-based ozone products (Lefèvre et al., 2013).

Each atmospheric state is input twice to libRadtran, (1) with the Kato et al. (1999) approach yielding KT_{KB3} ,

Table 2. Ranges and statistical distributions of values taken by the cosine of the solar zenith angle, the ground albedo and the 7 variables describing the clear atmosphere.

Variable	Value
Cosine of the SZA $\cos(\Theta_s)$	Uniform between 0 and 89 (degree) for Θ_s
Ground albedo ρ_g	Uniform between 0 and 0.9
Elevation of the ground above mean sea level	Equiprobable in the set $\{0, 1, 2, 3\}$ (km)
Total column content of ozone	Ozone content is $300 \times \beta + 200$, in Dobson unit. Beta distribution, with A parameter = 2 and B parameter = 2, to compute β
Atmospheric profiles (Air Force Geophysics Laboratory standards)	Equiprobable in the set $\{\text{midlatitude summer, midlatitude winter, subarctic summer, subarctic winter, tropical, US standard}\}$
Aerosol optical depth at 550 nm	Gamma distribution, with shape parameter = 2 and scale parameter = 0.13
Ångström exponent coefficient	Normal distribution, with mean = 1.3 and standard-deviation = 0.5
Aerosol type	Equiprobable in the set $\{\text{urban, rural, maritime, tropospheric, desert, continental, Antarctic}\}$

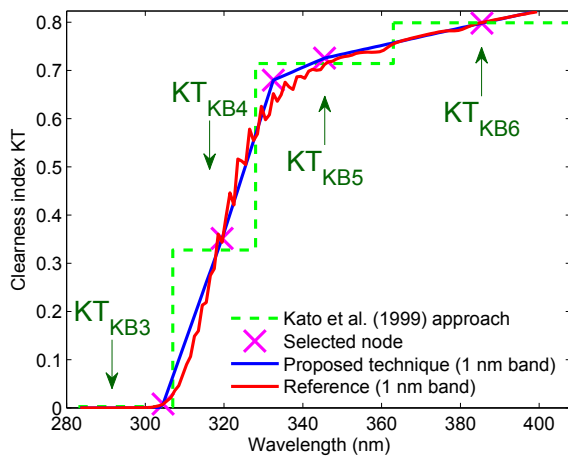


Figure 1. Illustration of the resampling technique.

KT_{KB4} , KT_{KB5} and KT_{KB6} and (2) with detailed spectral computations providing KT_i every 1 nm for the interval [283, 408] nm. Several plots were made superimposing KT_{KB3} , KT_{KB4} , KT_{KB5} and KT_{KB6} (in green), and KT_i (in red). Figure 1 is such a graph with the following inputs: θ_s of 53.76° , the midlatitude winter atmospheric profile, TOC of 470 DU, AOD of 0.78 at 1000 nm for a maritime polluted aerosol model with an Ångström exponent of 1.93, elevation of 0 m and surface albedo of 0.63. A visual inspection shows that KT_{KB_j} and KT_i are approximately equal for λ_j in the middle of KB_j , except for KB_5 . KT_i in this band exhibits a nonlinear behavior that cannot be accounted for with a single KT_k . If one selects 305, 320, 333, 346 and 386 nm as NB_k (magenta crosses), then the linear interpolation (in blue) provides a fairly accurate estimate of KT_i . The five NB_k 's were selected by a lengthy visual inspection of such plots and are reported in Table 3. The same NB_k 's apply for the global and direct irradiances. For each NB_k , the parameters of the affine function relating KT_{KB_j} and KT_k are determined by least-

squares fitting technique (Table 3):

$$KT_k = a_k KT_{KB_j} + b_k. \quad (9)$$

Another set of parameters is determined in the same way for the direct irradiance. In the operational mode, given an atmospheric state, a run of libRadtran, or a fast approximation of it, yields four KT_{KB_j} , from which the five KT_k 's are computed using the affine functions. Then, approximate KT_i^* 's are computed for each nm between 280 and 400 nm using a linear interpolation and extrapolation of KT_k . In cases where extrapolation provides negative values, KT_i^* is set to 0. Eventually, G_{UV} is obtained by

$$G_{UV} = \cos(\Theta_s) \sum_{i=1}^{120} Eo_i KT_i^*. \quad (10)$$

A similar process is performed for the direct normal irradiance B_{UV} as well as for the same quantities in UV-A and UV-B. As the method provides the spectrum KT_i^* , the equation may be extended to include any action spectrum $S(\lambda)$, for example,

$$G_{S(\lambda)} = \cos(\Theta_s) \sum_{\lambda=1}^{\lambda=2} Eo_i S(i) KT_i^*. \quad (11)$$

3.3 Numerical validation

In this section, results of the proposed technique are compared with results from the detailed spectral calculations made by libRadtran to assess the accuracy of the proposed technique for G_{UVA} , G_{UVB} , B_{UVA} and B_{UVB} for cloud-free conditions. The errors made by using the proposed technique for calculations of the UV-A and UV-B irradiances are presented. To that extent, an additional sample of 10 000 atmospheric states has been randomly constructed following the marginal distribution variables described in Table 2. The proposed technique was applied to the outputs of libRadtran using the Kato et al. (1999) approach and the estimates were

Table 3. KB covering the UV band and selected subintervals NB_k ; slopes and intercepts of the affine functions between the clearness indices in KB and subintervals NB_k .

KB	KB range, nm	Subinterval NB_k , nm (# k)	Global		Direct normal	
			Slope a_i	Intercept b_i	Slope c_i	Intercept d_i
3	283–307	304–305 (#1)	3.0900	0.0007	3.0852	0.0003
4	307–328	319–320 (#2)	1.1264	−0.0175	1.0886	−0.0007
5	328–363	332–333 (#3)	1.0247	−0.0519	0.8992	−0.0103
		345–346 (#4)	0.9946	0.0152	1.0112	−0.0004
6	363–408	385–386 (#5)	1.0030	−0.0032	0.9987	−0.0023

Table 4. Statistical indicators of the performances of the proposed technique for estimating UV fluxes.

UV fluxes	Mean ($W m^{-2}$)	Bias ($W m^{-2}$)	RMSE ($W m^{-2}$)	rBias (%)	rRMSE (%)	R^2
G_{UVA}	45.6	+0.1	0.1	+0.2	0.2	1.00
B_{UVA}	23.4	−0.1	0.2	−0.6	0.8	1.00
G_{UVB}	2.30	−0.04	0.14	−1.64	6.19	1.00
B_{UVB}	0.73	+0.07	0.15	+10.10	20.48	0.97

compared to the detailed calculations performed by libRadtran. Following the ISO standard (1995), the deviations were computed by subtracting measurements for each instant from the results of the method. They were summarized by the bias (mean error), the root mean square error, and their values rBias and rRMSE relative to the mean value of the measurements. In addition, the coefficient of determination (R^2) is computed.

Table 4 reports the statistical indicators for the global and direct normal UV-A and UV-B irradiances. For UV-A fluxes, the bias for the global irradiance and the direct irradiance is $+0.10 W m^{-2}$, i.e., $+0.2\%$ in relative value, and $−0.15 W m^{-2}$, i.e., $−0.7\%$ in relative value respectively. The RMSE is respectively $0.12 W m^{-2}$ (0.3%) and $0.18 W m^{-2}$ (0.8%). For UV-B fluxes, the bias for the global irradiance and the direct irradiance is $−0.04 W m^{-2}$, i.e., $−1.6\%$ in relative value, and $+0.07 W m^{-2}$, i.e., $+10.1\%$ in relative value respectively. The corresponding RMSE is respectively $0.14 W m^{-2}$ (6.2%) and $0.15 W m^{-2}$ (20.5%). The coefficient of determination R^2 is greater than 0.99 except for the direct normal UV-B irradiance which is 0.966. Expectedly, these indicators prove the good level of performance of the proposed technique.

4 Results

The results of the proposed method were compared to measurements of UV-A and UV-B irradiances at the surface for cloud-free conditions. Similar statistical indicators as those presented in the previous section are also computed to synthesize the errors.

4.1 Performance of the method for UV-A irradiance

Figure 2 exhibits the scatter density plot between ground-based instantaneous measurements made for each station in cloud-free conditions and estimates from the proposed method combined with inputs from CAMS. The station name is indicated at the top of each plot. Figure 2a exhibits the results for Barrow. All points are well located along the identity line. The slope of the fitting line is 0.995, i.e., very close to 1, showing a very good estimation of the measurements by the method. R^2 is 0.97, meaning that all the variability in the measurements is very well explained by the estimates. The bias is low with a value of $−0.2 W m^{-2}$, i.e., $−1\%$ of the mean value of the measurements, $20.5 W m^{-2}$. The RMSE is small with a value of $1.4 W m^{-2}$, which is around 7% of the mean. These statistical indicators for each station are reported in Table 5. The measurements are mostly between March and September. The shortwave albedo is 0.8 from the beginning of March until the middle of May. With the progressive snowmelt, this shortwave albedo decreases from mid-May down to 0.12 in mid-July. This variation corresponds well to the climatological evolution reported by Bernhard et al. (2008) and supports the choice of this approximation by the shortwave albedo.

Even if the points follows the perfect line (Fig. 2a) quite well, a set of points is seen where the method noticeably underestimates by more than 20% . These underestimations occurs between the end of May and mid-July. During that period, the shortwave albedo was less than the effective UV albedo by a factor 0.8. The effective UV albedo is part of the version 2 dataset and was derived by comparing measured clear-sky spectra with corresponding radiative transfer model results (Bernhard et al., 2007). As a smaller albedo means a smaller contribution to the diffuse part of the irradiance, the difference between the shortwave and effective UV albedo may explain these underestimations seen in Fig. 2a.

Results for Sodankylä are shown in Fig. 2b. Cloud-free conditions occur mostly between February and September. The points lie along the identity line with a slight overestimation by the method at low irradiance and an underestimation at large irradiance. R^2 is 0.98. The bias is low with a value of $−0.5 W m^{-2}$, i.e., $−2\%$ of the mean value of $20.8 W m^{-2}$.

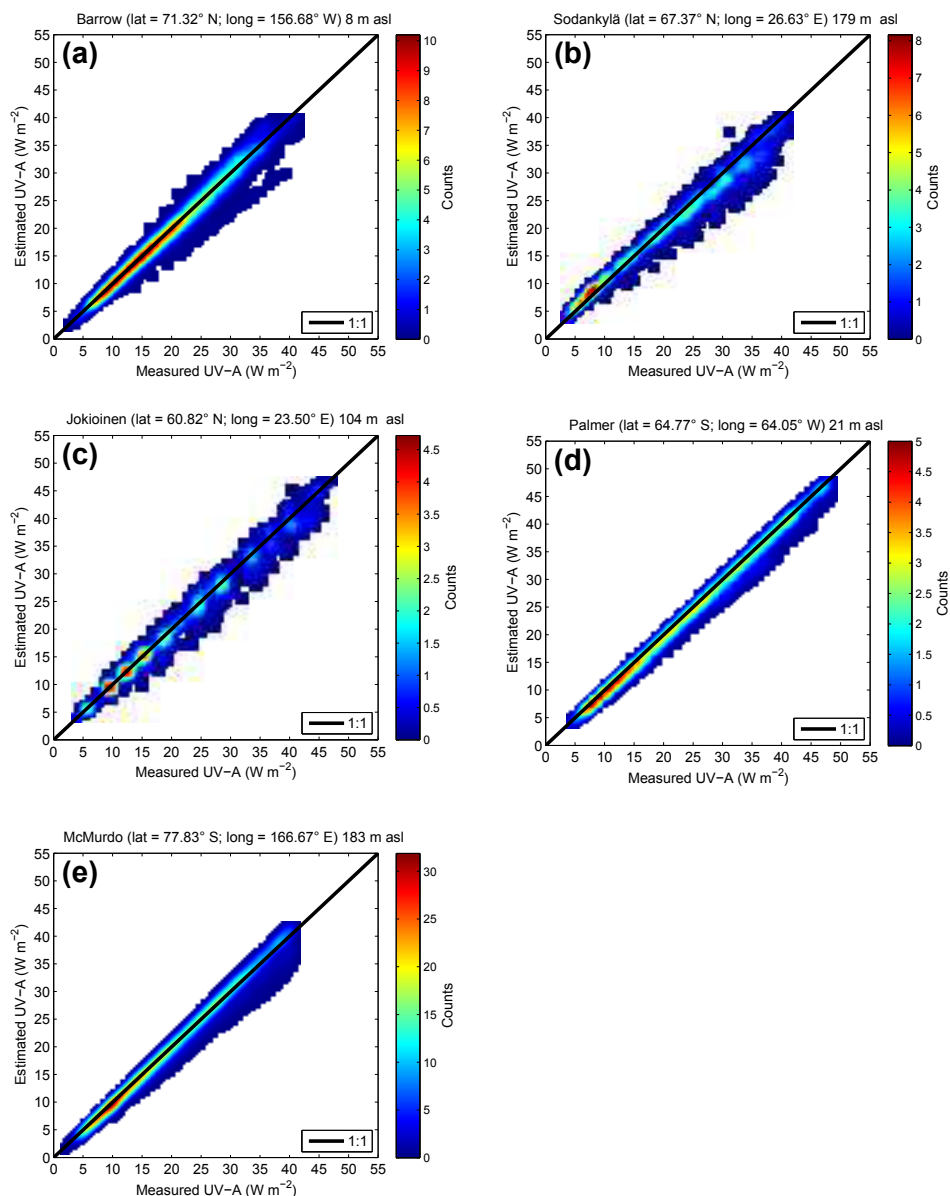


Figure 2. Scatter density plot between measurements of UV-A and estimates for each station with each station name at the top. The color bar indicates the number of points in the area within the interval $0.4 \text{ W m}^{-2} \times 0.4 \text{ W m}^{-2}$.

Table 5. Statistical indicators of the performances of the method for UV-A irradiance. N is the number of data points.

Station	N	Mean (W m^{-2})	Bias (W m^{-2})	RMSE (W m^{-2})	rBias (%)	rRMSE (%)	R^2
Barrow	4293	20.0	−0.2	1.4	−1.1	6.8	0.98
Sodankylä	2590	20.8	−0.5	1.9	−2.5	9.0	0.98
Jokioinen	1140	22.1	−0.5	1.6	−2.1	7.5	0.98
Palmer	1736	24.9	−0.8	1.2	−3.1	4.9	0.99
McMurdo	10175	20.3	−0.3	1.1	−1.6	5.6	0.99

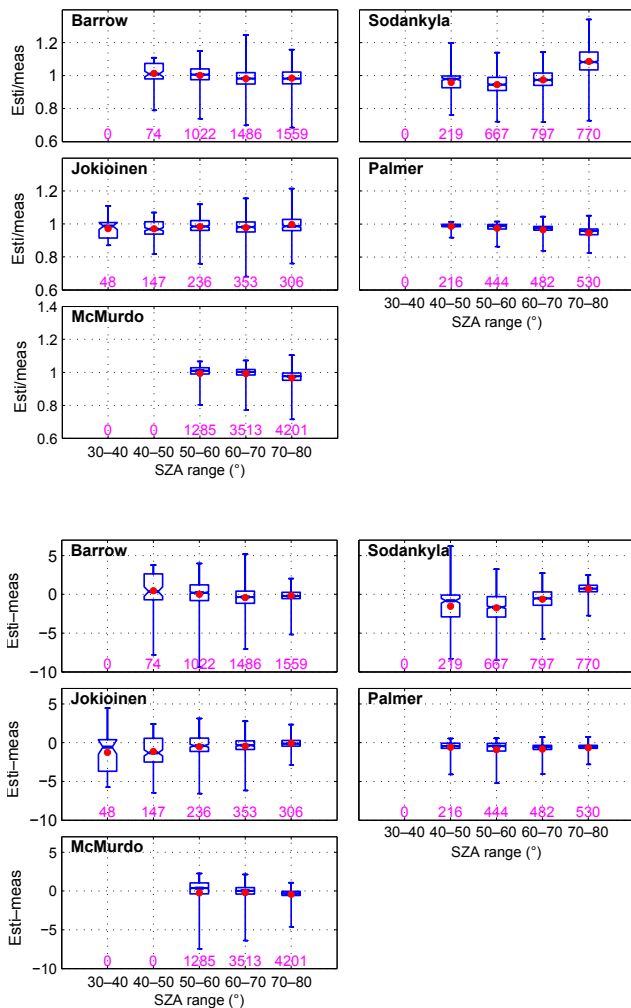


Figure 3. Dependence of ratio (top) of the estimated (esti) to the measured (meas) UV-A irradiances for each station and the difference between the estimated and measured (bottom) UV-A irradiances for each station as a function of SZA range. The red dots indicate the mean; the limits of the boxes are the first, second (median) and third quartiles. The lower whisker is the minimum and the upper one is the maximum. The pink number is the number of data in a single SZA range.

The RMSE is low with a value of 1.9 W m^{-2} (9 %). As for Jokioinen (Fig. 2c), all points are well located along the identity line. R^2 is 0.98. The bias is low with a value of -0.5 W m^{-2} , i.e., -2% of the mean value of the measurements, 22.1 W m^{-2} , as well as the RMSE with a value of 1.6 W m^{-2} (8 %).

Results for Palmer are shown in Fig. 2d. One may note that the points are well aligned with low scatter along a straight line whose slope is 0.99 with a slight underestimation by the method. The bias is -0.8 W m^{-2} (-3% of the mean value of 24.9 W m^{-2}). The RMSE is 1.2 W m^{-2} (5 %). R^2 is greater than 0.99. Cloud-free conditions occur mostly between Au-

gust and April. The shortwave albedo slightly increases from 0.28 to 0.32 between August and March and then decreases until April up to 0.20. These values are small and close to those of a ground free of snow or ice. The effective UV albedo is usually greater than 0.3 with peaks up to 0.8. This difference between the shortwave and effective UV albedo may explain the slight underestimation indicated in Fig. 5.

Results for McMurdo are shown in Fig. 2e. Cloud-free conditions occur mostly between April and September. The points are aligned along the identity line. R^2 is 0.99. The bias is low with a value of -0.3 W m^{-2} (-1% of the mean value of 21.0 W m^{-2}), as well as the RMSE with a value of 1.2 W m^{-2} (6 %). The shortwave albedo may reach 0.8 and there is no clear discrepancy between the shortwave and effective UV albedo. Nevertheless, the authors believe that the outliers may be explained by a difference in albedo.

The dependence of errors as a function of SZA was investigated. Figure 3 exhibits the ratio (top) and difference (bottom) as function of θ_s for each station for UV-A irradiances. For the ratio, the limits of the boxes are close from one quartile to another, meaning a very limited spread of the ratio. The deviations between maximum and minimum are approximately small. The median is similar to the mean. Regardless of the number of data (in pink color), the deviations are very close to 1 for all SZA ranges and stations except Sodankylä at high θ_s . As for the ratio, the similar observations are also seen in terms of the differences at the bottom of Fig. 3. The difference is very close to 0 for all the SZA ranges. The absolute value of the mean difference shows a tendency to decrease as θ_s increases, with the maximum being reached for low θ_s .

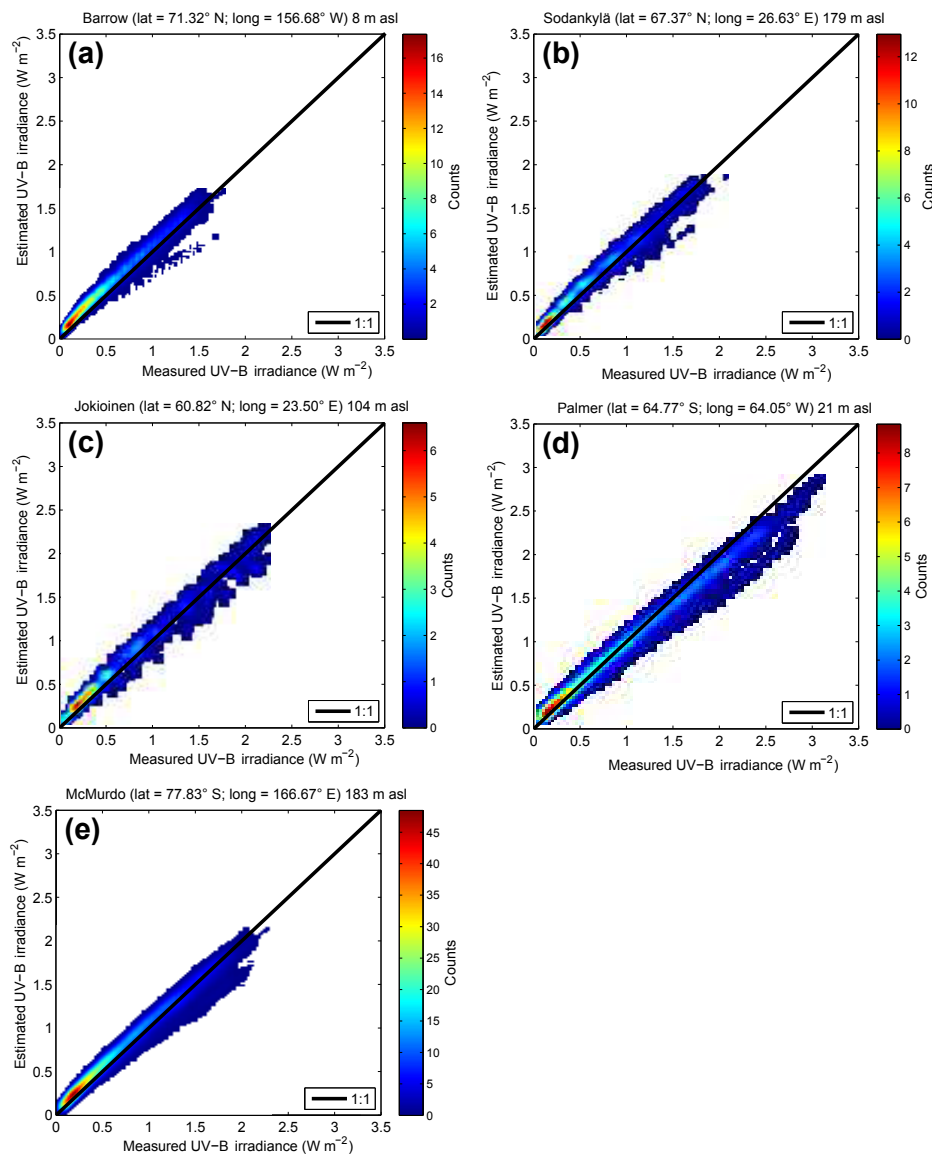
The dependence of errors as a function of TOC and albedo was also investigated (not shown). The results have revealed that there is no clear dependence of errors as a function of TOC or albedo for all stations. In addition, the absolute values of the bias (not shown) show a tendency to decrease as θ_s increases, with the maximum being reached for low θ_s , except McMurdo. In the opposite manner, the absolute values of the relative bias show a tendency to increase with θ_s . This could be related to the fact that low θ_s are reached in summer, with greater values in UV irradiance and lower values in effective UV albedo.

4.2 Performance of the method for UV-B irradiance

The UV-B band is the spectral region of UV irradiance where the ozone absorption is very strong. Figure 4 exhibits the scatter density plots between measurements of UV-B and estimates for each station. Table 6 reports the statistical indicators for UV-B. R^2 is greater than 0.97 for all stations, meaning that variability in UV-B is very well reproduced by the estimates. In general, the method overestimates the UV-B irradiances. Visually, one observes that the method clearly overestimates when the irradiance is low. For the wavelength less than 320 nm, in Fig. 1, the proposed method seems to

Table 6. Statistical indicators of the performances of the method for UV-B irradiance. N is the number of data points.

Station	N	Mean (W m^{-2})	Bias (W m^{-2})	RMSE (W m^{-2})	rBias (%)	rRMSE (%)	R^2
Barrow	4293	0.57	0.08	0.10	13.41	18.01	0.97
Sodankylä	2590	0.65	0.05	0.09	7.74	13.91	0.98
Jokioinen	1140	0.75	0.05	0.10	6.70	13.74	0.98
Palmer	1736	1.03	−0.04	0.12	−4.24	11.67	0.99
McMurdo	10 175	0.72	0.04	0.09	4.86	12.32	0.98

**Figure 4.** Same as Fig. 2, but for UV-B irradiance.

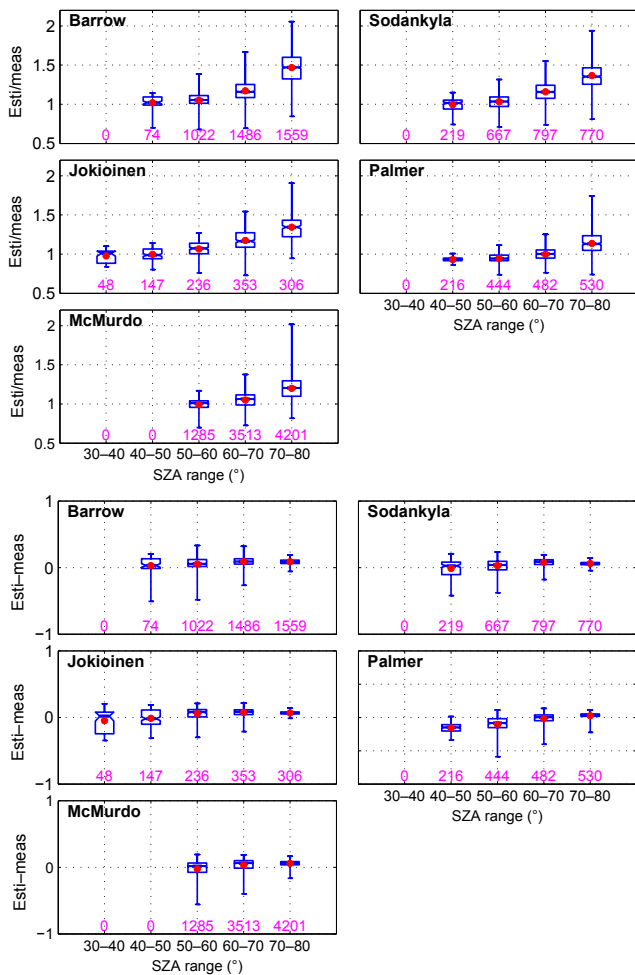


Figure 5. Same as Fig. 3, but for UV-B irradiance.

mostly overestimate when compared to the detailed spectral calculations serving as reference. This observation induces a systematic overestimation at low irradiance from the method. This mainly explains this previous observation. The absolute value of the bias is less than 0.1 W m^{-2} . The relative bias ranges between -4% (Palmer) and 13% (Barrow). The rRMSE ranges between 12% and 18% .

Figure 5 shows the change in ratio and relative difference as a function of θ_s . For both ratio and differences, the spread of limits of the boxes is more or less visible. Nevertheless, the median and the mean are close. The deviations show a tendency to increase with θ_s for all stations, probably meaning that a systematic bias from the method is more visible at high θ_s as mentioned above. These results explained the main sources of errors for the dependence of errors as a function of TOC and albedo.

In addition, one notices a tendency of the bias to reach a maximum between 65 and 75° (not shown); this appears in the form of a plateau around 65° for the RMSE which decreases as θ_s increases, i.e., as the irradiance decreases. The

relative bias increases with θ_s as well as the rRMSE. Both are closer in the high θ_s than in the small θ_s .

5 Discussion and conclusion

The comparison has demonstrated a reasonable agreement between the ground-based measurements of UV-A and UV-B and the estimates by the proposed method with CAMS products as inputs. The variability in UV fluxes is well reproduced by the method. A good level of accuracy is reached that is close to the uncertainty of the measurements themselves. The computations of the fluxes in the KBs can be performed quite fast with the use of precomputed look up tables as shown by the example of the McClear model. This model is an accurate approximation of libRadtran but 10^5 times faster. The proposed method extends the results of the McClear model to the UV range and can be used in future operational tools that are both accurate and fast.

Further improvements are needed. A major improvement would be the extension to all sky conditions. In this aspect, one may build on the work of Oumbe et al. (2014a, b), who demonstrated that, in the case of an infinite plane-parallel single- and double-layered cloud, the solar irradiance at ground level computed by a radiative transfer model can be approximated by the product of the irradiance under clear atmosphere and a modification factor that depends on cloud properties and ground albedo only as changes in clear-atmosphere properties have a negligible effect on this factor. Such an approximation has been exploited previously with limited justification by several authors in studies on broadband irradiance (Huang et al., 2011), UV or photosynthetically active radiation (see e.g., Calbo et al., 2005; den Outer et al., 2010; Krotkov et al., 2001).

Another improvement consists in the modeling of the surface albedo in the UV range. Maps of BRDF parameters in the UV range must be created with a satisfactory spatial resolution of 0.05° or better. The MODIS BRDF parameters may be a starting point as they are available at several wavelengths. It could be possible to apply the technique used by Blanc et al. (2014) to create BRDF maps for each wavelength and for each calendar month with no missing values. The smallest wavelength in the MODIS BRDF is approximately 470 nm , i.e., outside the UV range, and extrapolation towards small wavelengths will be necessary.

Data availability. UV data from Barrow, Palmer station and McMurdo station were provided by the NSF UV Monitoring Network operated by Biospherical Instruments Inc. and funded by the US National Science Foundation's Office of Polar Programs. Version 2 data used here are available from <http://uv.biospherical.com/Version2/Version2.asp>

FMI's spectral Brewer UV measurements are available through the European UV Database: <http://uv.fmi.fi/uvdb/>

Products from CAMS can be downloaded from the following website: <http://atmosphere.copernicus.eu/>

The BRDF maps by Blanc et al. (2014) may be downloaded from the following website: <http://www.oie.mines-paristech.fr/Valorisation/Outils/AlbedoSol/>

Competing interests. The authors declare that they have no conflict of interest.

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