



A new zenith-looking narrow-band radiometer-based system (ZEN) for dust aerosol optical depth monitoring

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Abstract. A new zenith-looking narrow-band radiometer based system (ZEN), conceived for dust aerosol optical depth (AOD) monitoring, is presented in this paper. The ZEN system comprises a new radiometer (ZEN-R41) and a methodology for AOD retrieval (ZEN-LUT). ZEN-R41 has been designed to be stand alone and without moving parts, making it a low-cost and robust instrument with low maintenance, appropriate for deployment in remote and unpopulated desert areas. The ZEN-LUT method is based on the comparison of the measured zenith sky radiance (ZSR) with a look-up table (LUT) of computed ZSRs. The LUT is generated with the LibRadtran radiative transfer code. The sensitivity study proved that the ZEN-LUT method is appropriate for inferring AOD from ZSR measurements with an AOD standard uncertainty up to 0.06 for $\text{AOD}_{500\text{nm}} \sim 0.5$ and up to 0.15 for $\text{AOD}_{500\text{nm}} \sim 1.0$, considering instrumental errors of 5%. The validation of the ZEN-LUT technique was performed using data from AEROSOL ROBOTIC NETWORK (AERONET) Cimel Electronique 318 photometers (CE318). A comparison between AOD obtained by applying the ZEN-LUT method on ZSRs (inferred from CE318 diffuse-sky measurements) and AOD provided by AERONET (derived from CE318 direct-sun measurements) was carried out at three sites characterized by a regular presence of desert mineral dust aerosols: Izaña and Santa Cruz in the Canary Is-

lands and Tamanrasset in Algeria. The results show a coefficient of determination (R^2) ranging from 0.99 to 0.97, and root mean square errors (RMSE) ranging from 0.010 at Izaña to 0.032 at Tamanrasset. The comparison of ZSR values from ZEN-R41 and the CE318 showed absolute relative mean bias (RMB) < 10%. ZEN-R41 AOD values inferred from ZEN-LUT methodology were compared with AOD provided by AERONET, showing a fairly good agreement in all wavelengths, with mean absolute AOD differences < 0.030 and R^2 higher than 0.97.

1 Introduction

Atmospheric aerosols play an important role in the environment, especially affecting air quality and climate. Concerning the Earth's climate, atmospheric aerosols are one of the main drivers of climate change and the most uncertain of them (Stocker et al., 2013). These atmospheric constituents can affect the Earth's radiative balance by scattering and absorbing the incoming solar radiation and the outgoing terrestrial radiation (aerosol–radiation interactions) but also by influencing cloud formation and reflectivity (aerosol–cloud interactions; Klein et al., 2010; Hoose and Möhler, 2012; Stocker et al., 2013). Furthermore, aerosols are not

uniformly distributed in the atmosphere, but show a high spatial and temporal variability. For these reasons, a good knowledge of the microphysical and optical properties of atmospheric aerosols is necessary all over the world. In this respect, aerosols remote sensing by means of satellite and ground-based devices is normally used to routinely monitor the aerosols columnar optical and microphysical properties, the latter obtained by inversion techniques (i.e. Dubovik and King, 2000). Regarding the aerosol optical properties, the aerosol optical depth (AOD) and its spectral dependence are the most common and important parameters for aerosol characterization, widely used in models and satellite sensors.

Remote sensing from satellite platforms has proved to be an effective tool for global and long-term monitoring of aerosols in the atmosphere. Early satellite sensors, such as the Advanced Very High Resolution Radiometer (NOAA/AVHRR Stowe et al., 1997) or TOMS (Total Ozone mapping Spectrometer; Herman et al., 1997; Torres et al., 1998), permitted the AOD estimation spanning from 1979 to the present. In recent years, a more advanced generation of active and passive satellite sensors for aerosol detection have been used to estimate aerosol radiative forcing on a global scale, such as MODIS (Moderate Resolution Imaging Spectrometer), MISR (Multi-angle Imaging Spectro-Radiometer), both in operation since 2000 (Tanré et al., 1997; Zhang and Reid, 2006, 2010), and CALIPSO (Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observations; Winker et al., 2010). However, as Li et al. (2009) suggested, satellite AOD retrievals are subject to important uncertainties due to radiometric calibration, a priori assumed aerosol properties, cloud contamination and correction of the surface effect. In particular, a reliable determination of the ground reflectance is mandatory, because the signal measured by the on-board sensor is the sum of the top of atmosphere (TOA) reflection and the surface reflection. A reliable determination is especially important over bright land surfaces like deserts, where the main dust sources are located. Moreover, the temporal resolution of satellite-borne sensors over a specific point is quite limited (one observation per day as maximum) and represents a high limitation for dust storm monitoring since they develop very fast in time periods of few hours.

There are currently extensive networks of ground-based sun photometers all over the world dedicated to aerosol monitoring. These networks are integrated by very precise instruments, close to the 0.02 AOD accuracy suggested as a goal by the World Meteorological Organization (WMO; Schmid et al., 1999; WMO, 1993). Among others, the WMO GAW-PFR network (Wehrli, 2005), SKYNET (Takamura and Nakajima, 2004), CARSONET (Che et al., 2015) and AERONET (Holben et al., 1998) are the most important. AERONET (AErosol RObotic NETwork) is the most widespread network using the Cimel Electronique 318 photometer as the standard instrument (hereinafter CE318). At present, this network provides unique long-term open-access

data with aerosol optical, microphysical and radiative properties, which constitute an important source of information for climate and environmental sciences, since their results have been widely tested under quite different conditions (Eck et al., 1999, 2001, 2003a, b, 2008, 2009, 2010; Holben et al., 2001). However, in the Northern Hemisphere there is a lack of stations in desert regions, which are the most important source of mineral dust, leading to an unsatisfactory description of dust cycles. Although AERONET sun photometers operate automatically, they need to be checked on a daily basis and frequently maintained by trained staff to assure data quality. These instruments operate using a sun tracker, which is a moving element and the origin of most of the operational problems. In addition, the relatively high cost of sun photometers is a constraint for many developing countries located in desert regions.

For all the above reasons, in this paper we present a new system (ZEN), composed of the radiometer ZEN-R41 and a new methodology especially conceived to estimate AOD from downwelling zenith sky radiance (ZSR) observations for desert dust conditions. This system has been jointly developed by SIELTEC Canarias S.L. company (SIELTEC) and the Izaña Atmospheric Research Center (IARC) from the State Meteorological Agency of Spain (AEMET). The use of ZSR as the direct measurable magnitude simplifies the design of the radiometer ZEN-R41, avoiding the use of sun tracker, making it more robust, automated and available at a lower cost than classical sun photometers. However, the AOD calculation is not as straightforward and precise as using direct-sun measurements, and some a priori assumptions must be made about aerosol properties, and significant additional efforts must be done in the modelling part. As made in the techniques employed by some on-board satellites sensors (Kaufman et al., 1997; Tanré et al., 1997; Stowe et al., 1997; Torres et al., 1998), and some previous work with ground based instrumentation (Lindfors et al., 2013), we have developed a look-up table (LUT)-based methodology (ZEN-LUT) specifically designed for desert aerosols to estimate AOD. The LUTs are composed of a set of simulated ZSR and AOD values obtained with the radiative transfer code LibRadtran (Mayer and Kylling, 2005). The AOD is then inferred by minimizing a function which depends on the difference between simulated and measured ZSR at all used wavelengths. In this case, the surface albedo does not have as much influence as in the upwelling radiance measured by satellite sensors. These characteristics make the ZEN a suitable system for aerosol monitoring in remote desert areas, filling the current observational gaps and being useful in the validation of satellite sensors and models dust products.

The test sites and ancillary observations used in this study are described in Sect. 2. In Sect. 3 a description of the new ZEN-R41 radiometer and a performance comparison with CE318 are given. The AOD retrieval method, including the results of a sensitivity study performed to examine the impact of key input parameters as well as to assess the influence of



Figure 1. Meteosat/TERRA image showing a Saharan dust outbreak over the study area on 12 January 2015. The Izaña, Santa Cruz and Tamanrasset sites are indicated with yellow stars.

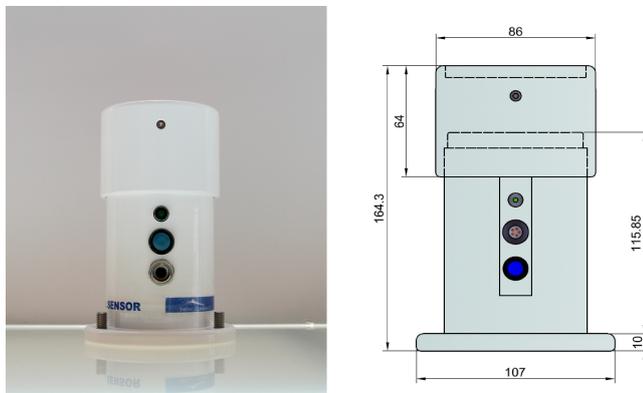


Figure 2. ZEN-R41 scheme and dimensions.

the instrumental errors, is examined in Sect. 4. Section 5 describes the AOD comparisons between ZEN-R41 and collocated CE318. Finally, the main conclusions of this work are presented in Sect. 6.

2 Test sites and ancillary information

2.1 Test sites

The ZEN-LUT methodology and the ZEN-R41 instrument have been tested in three different sites impacted by the presence of desert mineral dust (Fig. 1).

The Izaña station (Izaña, Canary Islands, Spain; 28.3° N, 16.5° W, 2373 m a.s.l.) is a high mountain subtropical station which represents atmospheric background conditions most of the time as a consequence of its location over a strong temperature inversion layer as a result of general subsidence processes and the presence of cool trade winds at lower levels. However, the proximity to the Saharan desert introduces an important influence of mineral dust on its aerosol climatology. As Basart et al. (2009) showed, there is an enhancement of dust transport from the Sahara to Izaña altitudes during summer ($\text{AOD}_{675\text{ nm}} > 0.15$ and large particles with

Angström Exponent, $\alpha_{440-870\text{ nm}} < 0.25$), with $\text{AOD}_{675\text{ nm}}$ values < 0.15 prevailing in the rest of the year, especially in winter, which represent $\sim 85\%$ of the overall conditions. Guirado (2014) also performed a detailed aerosol characterization at Izaña. This author found predominant dust conditions associated to $\text{AOD}_{500\text{ nm}} > 0.10$ and large particles with $\alpha < 0.60$. Izaña Observatory is part of the WMO Global Atmosphere Watch programme (GAW) and the Network for the Detection of Atmospheric Composition Change (NDACC). Izaña is an absolute sun calibration site of AERONET and a World Radiation Center (WRC) Global Atmospheric Watch Precision Filter Radiometer (GAW/PFR) station and calibration site. The Izaña station is also a WMO CIMO (Commission for Instruments and Methods of Observation) Testbed for Aerosols and Water Vapour Remote Sensing Instruments (WMO, 2014).

Santa Cruz de Tenerife station (SCO, Canary Islands, Spain; 28.5° N, 16.2° W; 52 m a.s.l.) is an urban station located at sea level. The aerosol climatology at this station is dominated by the well-mixed combination of fine fraction of pollution aerosols and coarse model marine particles (prevailing $\text{AOD}_{675\text{ nm}} > 0.15$) with mineral dust influence from spring to autumn (Basart et al., 2009), increasing AOD and reducing α values. Following the work developed by Guirado et al. (2014), those situations with $\text{AOD}_{500\text{ nm}} > 0.15$ and $\alpha < 0.5$ can be considered prevalent dust conditions at this station.

Tamanrasset station (TAM, Algeria; 22.8° N, 5.5° E, 1377 m a.s.l.) is located in southern Algeria in a key location near the most important dust sources of Mali, Algeria, Lybia and Chad, where there is little impact from industrial activities. Guirado et al. (2014) performed a thorough study to characterize aerosols at this Saharan station. They found that desert mineral dust is the predominant aerosol type in this station, where the dry-cool (winter) season is characterized by prevailing clear-sky conditions ($\text{AOD}_{440\text{ nm}} \sim 0.09$ and $\alpha \sim 0.62$), while high turbidity events with coarse dust particles are frequent during the wet-hot (summer) season, associated to an $\text{AOD}_{440\text{ nm}}$ modal value of 0.15 and $\alpha \sim 0.4$.

2.2 Ancillary information

2.2.1 CE318 sun photometer and AERONET network

In this work we have used AOD data provided by AERONET (Holben et al., 1998) to validate the results obtained with the ZEN-R41 radiometer and the ZEN-LUT technique. AERONET also provides information about microphysical and optical parameters, such as particle size distribution, refractive indices, single scattering albedo (SSA) or phase function using the inversion algorithm developed by Dubovik and King (2000); Dubovik et al. (2002, 2006).

The standard instrument used in AERONET is the CE318 sun photometer, which performs direct-sun and diffuse-sky measurements. Direct-sun measurements performed at 340,

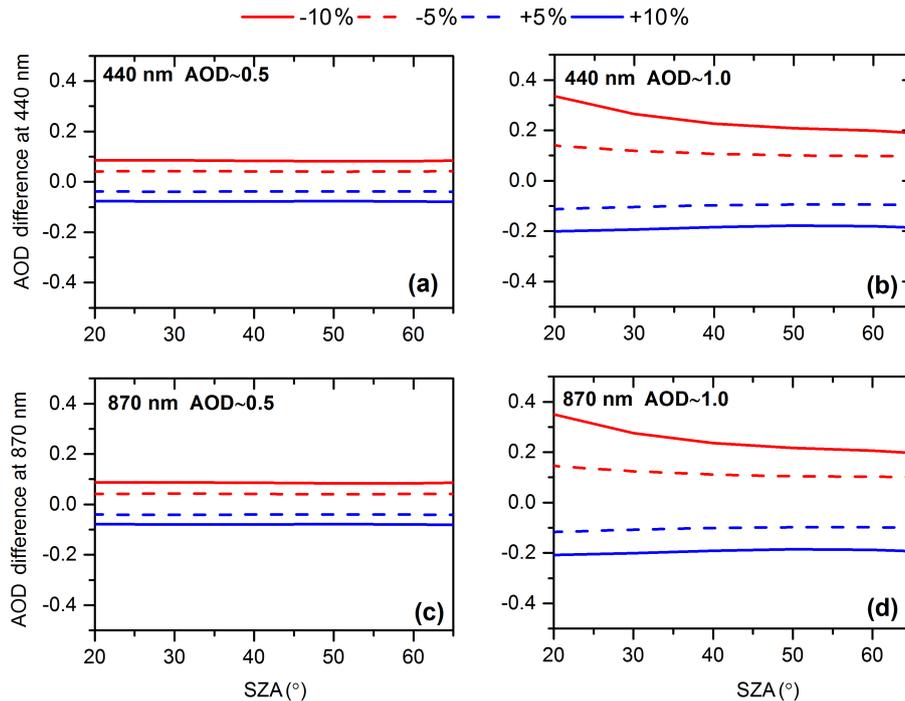


Figure 3. AOD difference versus SZA for different AOD conditions ($AOD_{500\text{nm}} \sim 0.5$ and 1.0) and ZSR perturbed $\pm 5\%$ and $\pm 10\%$, for 440 nm (a, b) and 870 nm (c, d).

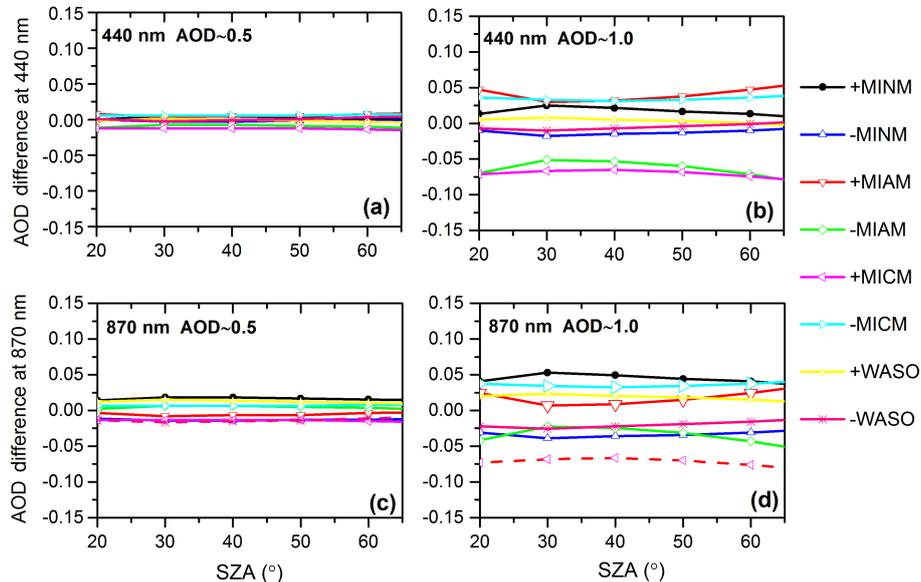


Figure 4. AOD difference versus SZA for different AOD conditions ($AOD_{500\text{nm}} \sim 0.5$ and 1.0) and 440 nm (a, b) and 870 nm (c, d) wavelengths. We have included different concentrations of four different aerosol components, MINM, MIAM and MICM perturbed in $\pm 5\%$ and WASO concentration perturbed in $\pm 50\%$.

380, 440, 500, 675, 870, 940 and 1020 nm are used to derive accurate AOD and precipitable water vapour with typical AOD uncertainties for field instruments ranging between ± 0.01 and ± 0.02 , with the higher errors in the UV spectral range (Eck et al., 1999). Diffuse-sky measurements with two

different routines are performed to infer the aerosol optical and microphysical properties: the almucantar (ALM) and the principal plane (PPL).

In the ALM routine, the azimuth angle is varied while the zenith angle is kept constant (equals to the solar zenith an-

Table 1. Coefficient of determination (R^2), relative root mean square error (RRMSE) in %, relative mean bias (RMB) in % and number of coincident data (N) for the ZSR comparisons between CE318 and ZEN-R41 measurements at four different spectral bands (440, 500, 675 and 870 nm) performed at Izaña in 2015.

Wavelength	R^2	RRMSE	RMB	N
870	0.99	4.7	-6.3	616
675	0.99	7.2	6.7	616
500	0.99	4.3	9.2	616
440	0.99	3.4	5.1	616

gle). On the other hand, in the PPL routine, the zenith is varied while the azimuth angle is kept constant (equals to the solar azimuth angle). The ALM measurements are performed in two wings: right (azimuth angle displaced towards the right of sun position) and left (azimuth angle displaced towards the left of sun position). In a homogeneous atmosphere, the signal measured at both wings should be equal, so this fact can be used to detect wrong data, such as cloud-contaminated measurements. Contrary to ALM, PPL is not symmetric but wrong data can also be detected by checking the smoothness of the PPL curve (Holben et al., 1998). In the present study, the CE318 ZSR values are obtained by linear interpolation of the PPL data to the zenith position (it is not possible to do it with ALM measurements). To detect and remove the presence of clouds in data, we have visually checked hemispheric images in those places where an all-sky camera was available. In case an all-sky camera was not available at the station, a smoothness criterion was applied on the PPL curve to detect clouds. This smoothness criterion is based on the analysis of the second derivative of the PPL radiances with respect to the scattering angle. The data are considered cloud contaminated if the second derivative is negative at any scattering angle between 2 and 90°. The threshold value for this smoothness criterion is not 0 but -1×10^{-7} , as determined empirically.

2.2.2 SONA all-sky camera

In the present study we have checked cloud interference by means of independent measurements from all-sky Automatic Cloud Observation System SONA cameras developed by SIELTEC (González et al., 2012). The SONA cameras have been used to detect and remove CE318 and ZEN-R41 cloud-contaminated ZSR data at Izaña and Santa Cruz stations. Cloud cover detection was performed by visual inspection, by analysing each individual hemispheric image from the total sky cameras and identifying the presence of visible clouds around the zenith.

3 The ZEN-R41 narrow-band radiometer

The ZEN-R41 is a prototype radiometer jointly developed by SIELTEC and IARC and designed to obtain AOD from downwelling zenith sky radiation at different wavelengths. This prototype incorporates collimating lenses and internal baffles to achieve a $\sim 3^\circ$ field of view. ZEN-R41 is equipped with four silicon detectors (350–1100 nm) and four optical filters of 10 nm FWHM with nominal wavelengths centred in 440, 500, 675 and 870 nm. These filters are hard coated to prevent ageing of their optical properties. The measurements, made simultaneously in the four channels, are amplified and acquired with 16 bit resolution (65.536 counts per each level of amplification). Inside the instrument there are sensors for internal humidity and temperature monitoring, as well as a fan for temperature homogenization and electronic components protection. Using this internal information, ZEN-R41 signal is temperature corrected, allowing the minimization of the temperature dependence of the silicon detectors. It is equipped with a small aluminum weatherproof (IP67 grade) case (white powder coated), with no moving parts and is protected by a thick borosilicate bk7 window (see Fig. 2), preventing it from damages such as scratches or direct impacts. The device also has a plate base which allows the instrument to be levelled and fixed on flat and rigid surfaces. As a result, the instrument's design is very robust and operates in a wide temperature range, between -40 and 85°C . An external blower attached to the external case is foreseen in order to facilitate the removal of dirt, dust or any other element that may affect measurements. In the case of remote locations, it is possible to use solar panels as a power source.

The on-board electronics comprise an internal 16 bits data logger and a 4 MB internal memory with Ethernet communications for data acquisition, display, download, set-up and diagnosis of the instrument. On-board processing is possible through its microcontroller board, with 16 MIPS of CPU speed and 96 KB of RAM memory, which result in a fast user-oriented data delivery (no additional parts or components are required) and a friendly user interface. The power and communications are performed together through a PoE port.

The instrument can work with fixed IP or in DHCP mode. Data can be downloaded manually or automatically and can also be sent via UDP or requested via TCP/IP protocol. It allows ZEN-R41 to form part of a ZEN-1 network for aerosol monitoring. This instrument can be also used independently or as a dependent sensor part of a weather station or other instrument which requires AOD as input.

The web interface shows alarms when these internal variables are out of the normal ranges. All the settings and calibration factors can be entered via a web graphical user interface.

ZEN-R41 was calibrated using an integrating sphere at the IARC facilities in order to convert the output signal into radiance units ($\text{W m}^{-2} \text{sr nm}^{-1}$). We performed a set of 10 mea-

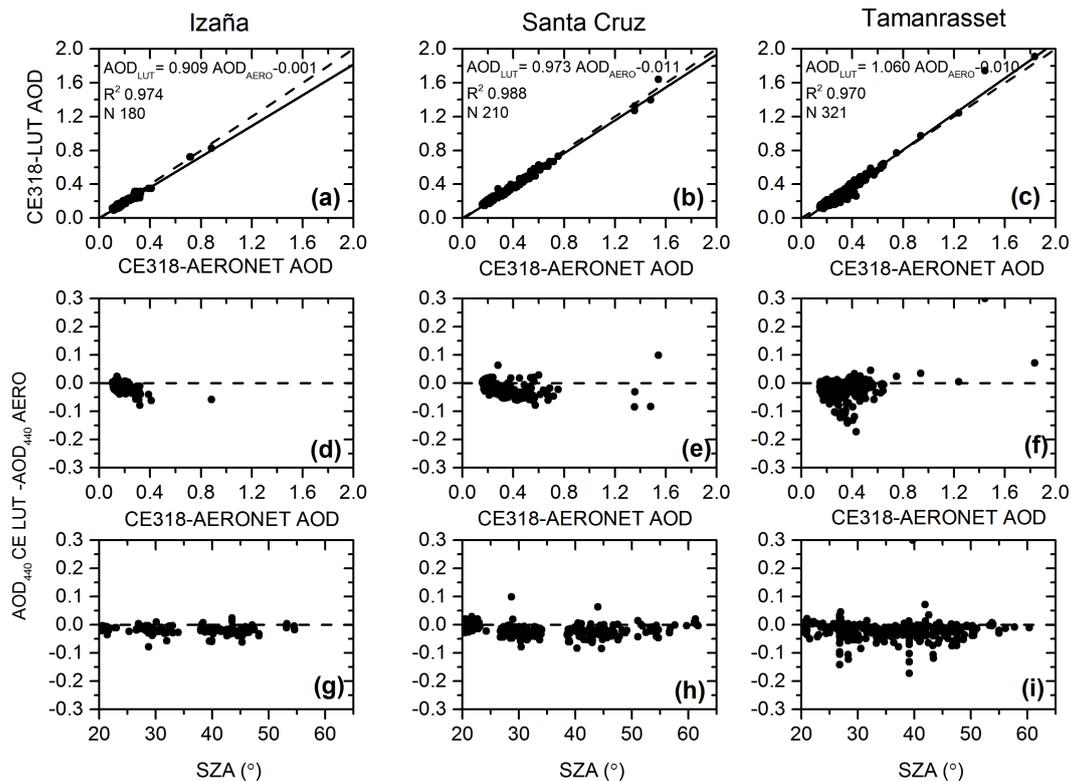


Figure 5. AOD scatter plot at 440 nm between CE318-AERONET and CE318-LUT for Izaña (a), Santa Cruz (b) and Tamanrasset (c) stations in 2013 for $20^\circ < \text{SZA} < 65^\circ$ (a, b, c). The black solid lines are the least-square fits, and the dashed lines are the diagonals ($y = x$). The least-square fit parameters are shown in the legend (slope, intercept, correlation coefficient R and number of data N). AOD differences in 440 nm between CE318-AERONET and CE318-LUT with respect to CE318-AERONET (d, e, f) and SZA ($^\circ$) (g, h, i).

measurements of the sphere's radiance, in which we observed a very low variability in ZEN-R41 digital counts ($\sim 2\%$). Since the uncertainty involved in this calibration procedure was established in $\sim 5\%$ by Walker et al. (1991) and, assuming the sphere has a perfect precision (Holben et al., 1998), we can use this information to estimate the ZEN-R41 radiance uncertainty in a $\sim 5\%$.

3.1 ZEN-R41 and CE318 radiometric comparison at Izaña

In order to check the goodness of the radiometric measurements of the ZEN-R41, we have compared ZSR observations from this device with those provided by CE318 instruments, derived from PPL measurements. The ZSR measurements performed in Izaña station in 2015 were cloud screened using ancillary SONA all-sky images. In this comparison we have not included ZSR data for $\text{SZA} < 20^\circ$ and $\text{SZA} > 65^\circ$ as we have detected higher discrepancies in ZSR for such SZA ranges. In the case of $\text{SZA} < 20^\circ$, the signal measured by ZEN-R41 device is larger than that obtained with CE318, which might be attributed to the lower stray light rejection and/or larger field of view of the ZEN-R41 radiometer. For $\text{SZA} > 65^\circ$, we attribute the observed discrepancies to lower

signals measured, which reduces the signal to noise ratio, increasing the instrument's uncertainty. We present a basic statistics of ZSR intercomparison for the four coincident spectral bands (440, 500, 675 and 870 nm) in Table 1. ZSR measured by both instruments are highly correlated, with coefficients of determination (R^2) ~ 0.99 for the four spectral ranges, discarding possible nonlinearities of ZEN-R41 in the selected measurement range. However, the relative mean bias results (RMB) showed that ZEN-R41 ZSR slightly overestimates at 675, 500 and 440 nm, but underestimates at 870 nm. RMB also showed a moderate variability in the measured ZSR, with values ranging from 6.3 % at 870 nm to 9.2 % at 500 nm, which is within the combined uncertainty of both instruments ($\sim 5\%$).

4 ZEN-LUT method

For estimating the AOD from downwelling ZSR measurements under cloud-free conditions, we used the ZEN-LUT method. This method is based on the comparison of measured ZSRs, corrected from earth–sun distance, with a set of simulated ZSR values at the four wavelengths available in the ZEN-R41 radiometer (440, 500, 675 and 870 nm). The

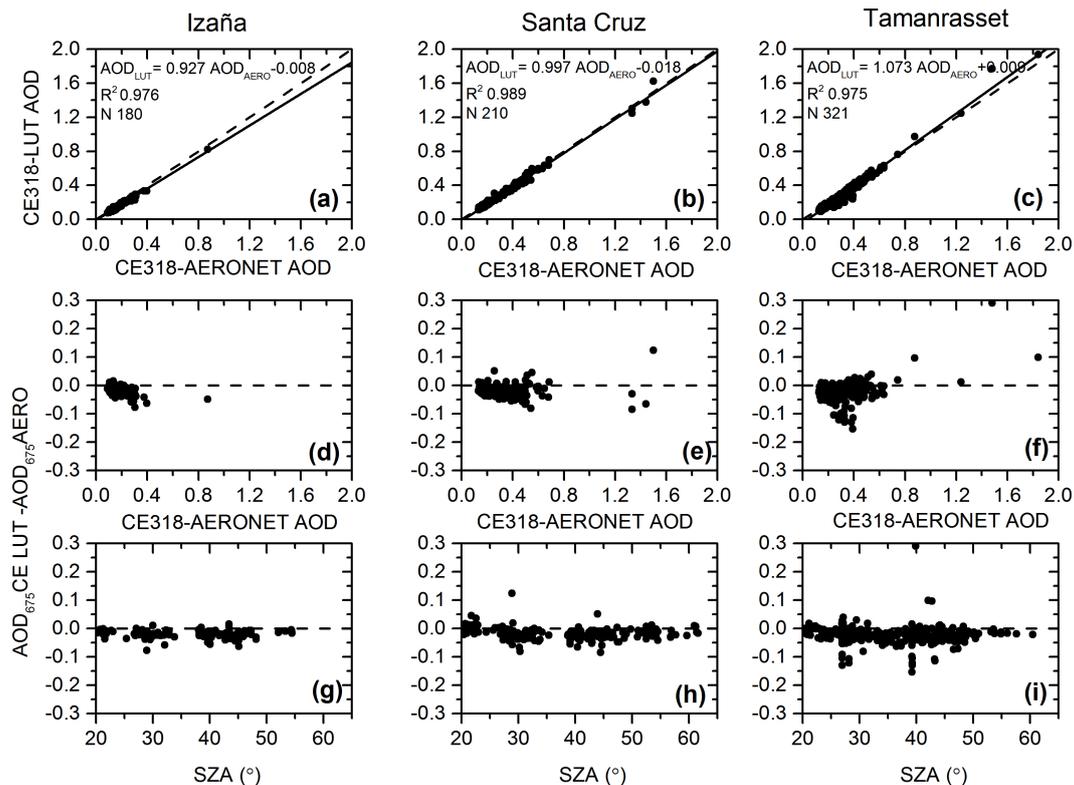


Figure 6. AOD scatter plot at 675 nm between CE318-AERONET and CE318-LUT for Izaña (a), Santa Cruz (b) and Tamanrasset (c) stations in 2013 for $20^\circ < \text{SZA} < 65^\circ$ (a, b, c). The black solid lines are the least-square fits, and the dashed lines are the diagonals ($y = x$). The least-square fit parameters are shown in the legend (slope, intercept, correlation coefficient R and number of data N). AOD difference in 675 nm between CE318-AERONET and CE318-LUT with respect to CE318-AERONET (d, e, f) and SZA ($^\circ$) (g, h, i).

LUT was determined using the LibRadtran radiative transfer model (RTM). This model is available in a complete software package containing a suite of tools for radiative transfer calculations in the Earth's atmosphere (freely available from <http://www.libradtran.org>; Mayer and Kylling, 2005; Emde et al., 2016). The LibRadtran structures the atmosphere as multi-layers, considering the vertical profiles of temperature, pressure and atmospheric components, such as gases and aerosols. A complete treatment of the absorption scattering processes offers hundreds of options and input parameters to handle all the structure that has a detailed RTM. It also includes several libraries which help to describe the atmosphere and the ground surface contribution on the simulated radiation field. Concerning the aerosol contribution, we have used the OPAC (Optical Properties of Aerosols and Clouds) library (Hess et al., 1998). The aerosols in OPAC can be defined through 10 basic components. These are water insoluble (INSO), water soluble (WASO), soot (SOOT), two sea salt components (sea salt accumulation mode or SSAM and sea salt coarse mode or SSCM), four mineral dust components (mineral nucleus mode, MINM, mineral accumulation mode, MIAM, mineral coarse mode, MICM, and mineral transported, MITR) and the sulfate component

(SUSO). Additionally, LibRadtran includes four spheroid components (MINM, MIAM, MITR and MICM spheroids) to define the mineral dust aerosols. The effects of relative humidity are taken into account for those components affected by water. Every component is defined by their microphysical properties, that is, the refractive index, mostly taken from d'Almeida et al. (1991) and a log-normal size distribution (Deepak and Gerber, 1983). Then, through Mie-scattering calculation (Wiscombe, 1980; Bohren and Huffman, 1983), or the T-matrix method (Mishchenko and Travis, 1998) in the case of spheroids, the optical properties are calculated for every component and normalized to 1 particle cm^{-3} .

The ZSR values were calculated by using the radiative transfer equation (RTE) solver CDISORT (Buras et al., 2011). The extraterrestrial solar flux was selected from Kurucz (1992) with a spectral resolution of 1 nm. The mid-latitude summer standard model for the atmosphere profile and the molecular absorption, parameterized with the LOWTRAN band model (Pierluissi and Peng, 1985), as adopted from the SBDART code (Ricchiazzi et al., 1998), were used. We have also used a common normalized Gaussian filter function for the four considered spectral bands centred in the nominal wavelength (440, 500, 675 and 870 nm). To

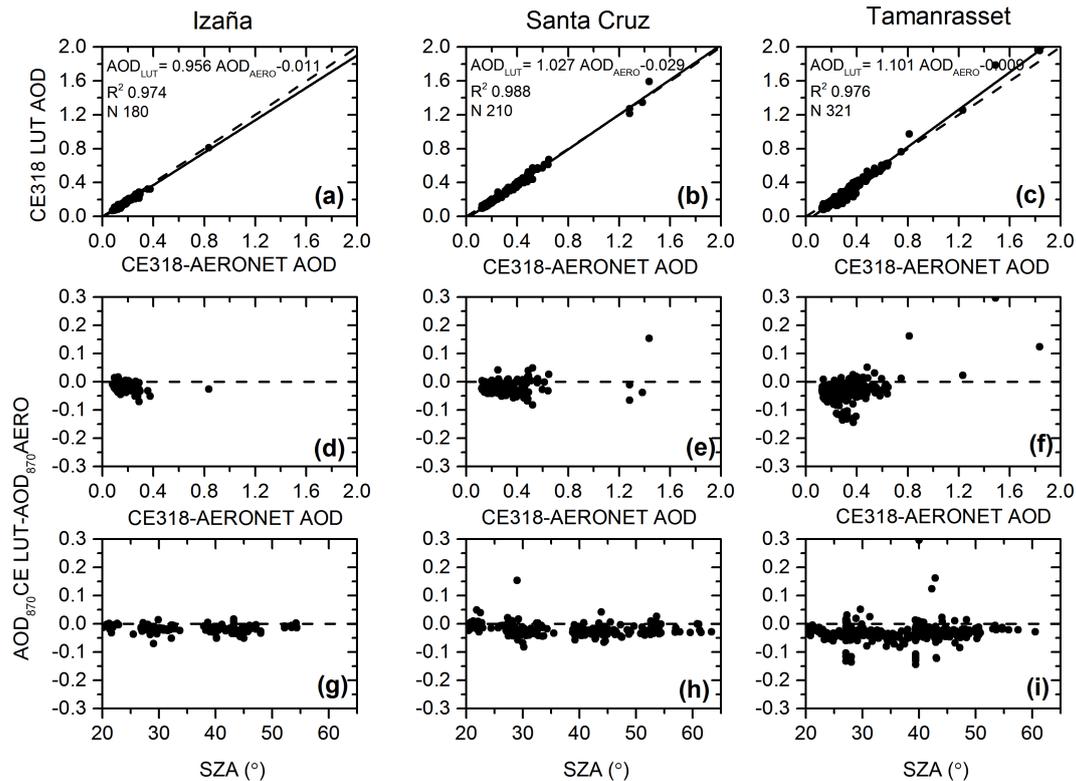


Figure 7. AOD scatter plot at 870 nm between CE318-AERONET and CE318-LUT for Izaña (a), Santa Cruz (b) and Tamanrasset (c) stations in 2013 for $20^\circ < SZA < 65^\circ$ (a, b, c). The black solid lines are the least-square fits and the dashed lines are the diagonals ($y = x$). The least-square fit parameters are shown in the legend (slope, intercept, correlation coefficient R and number of data N). AOD difference in 870 nm between CE318-AERONET and CE318-LUT respect to CE318-AERONET (d, e, f) and SZA ($^\circ$) (g, h, i).

set the surface reflectance contribution, we have considered the albedo definitions given in the IGBP (International Geosphere Biosphere Programme) library, which originates from the NASA CERES/SARB Surface Properties Project (Belward and Loveland, 1996).

Concerning the solar position, only the solar zenith angle (SZA) is given as input to the model, since the computed radiance in the zenith direction for a cloud-free sky is invariable with the solar azimuth angle. The earth–sun distance correction is directly applied in the measured ZSR, so the ZEN-LUT ZSRs are only simulated for the mean earth–sun distance.

The aerosol optical properties were obtained with the OPAC library (Hess et al., 1998) by defining the height profiles of every aerosol component present at every layer. The height of all layers and the mix of aerosol components present at every layer except the boundary layer were set up following the indications given in Hess et al. (1998). In the case of the boundary layer, we decided to set the mix of aerosol components dynamically, adopting the suggestions given in the Global Aerosol Data Set (GADS) report (Koepke et al., 1997) for desert regions. They proposed a mix of four different components, three mineral dust components (MINM, MIAM and MICM), plus a certain fixed quantity of

water soluble component (WASO). The ratio of three mineral dust components present in the mixture is variable and depends on the total mineral dust particle density (N_{mineral}). The three mineral dust components are related to N_{mineral} through the following expressions (Koepke et al., 1997):

$$\ln(N_{\text{MINM}}) \sim 0.104 + 0.963 \cdot \ln(N_{\text{mineral}}) \quad (1)$$

$$\ln(N_{\text{MIAM}}) \sim -3.94 + 1.29 \cdot \ln(N_{\text{mineral}}) \quad (2)$$

$$\ln(N_{\text{MICM}}) \sim -13.7 + 2.06 \cdot \ln(N_{\text{mineral}}). \quad (3)$$

In our calculations we have used the spheroid LibRadtran definitions.

It is important to note that the empirical relationships presented on these equations were derived for average conditions in desert areas, therefore they seem to be appropriate to describe the actual aerosols in Tamanrasset station, located in the middle of the Sahara desert, but not for Izaña and Santa Cruz stations, which are located hundreds of kilometres away from the Sahara desert, over the North Atlantic. In the latter, the presence of large mineral dust particles is considerably reduced due to a faster deposition during its transport out of dust source regions and to the predominance of other aerosols in the absence of Saharan intrusions. Be-

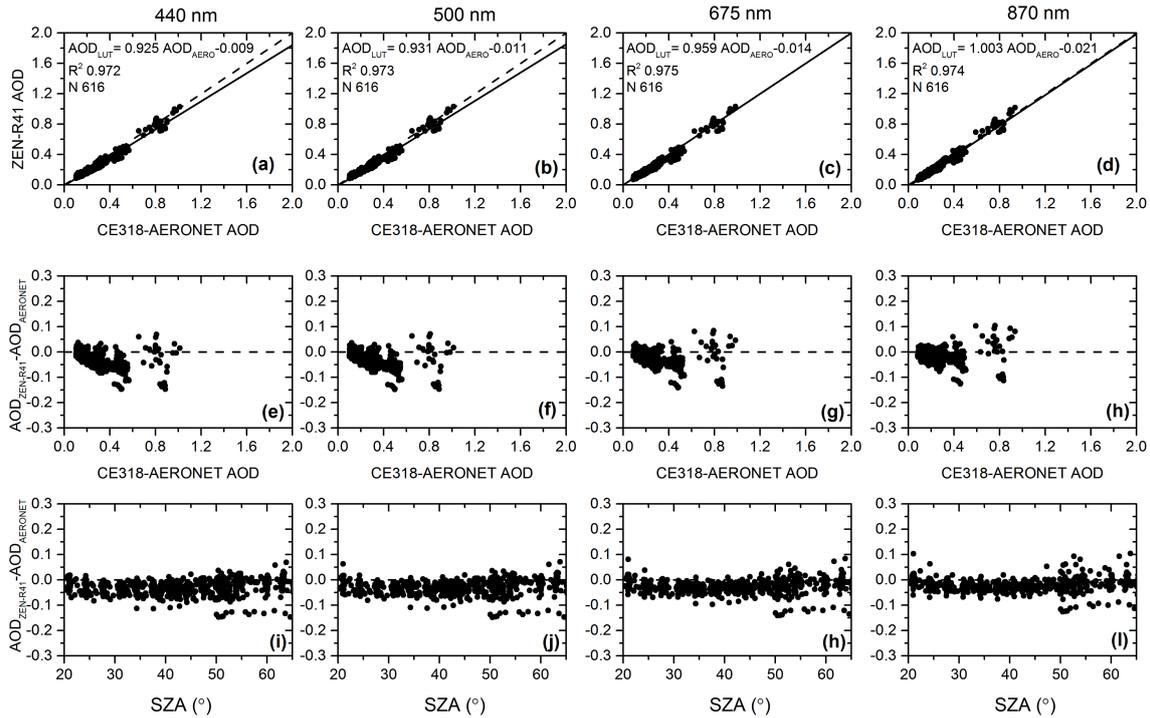


Figure 8. AOD comparisons between CE318-AERONET and ZEN-R41 for four different spectral bands (440, 500, 675 and 870 nm) performed at Izaña station in 2015. In the upper panel (a–d) AOD scatter plots AERONET/ZEN41 are presented. The middle panel (e–h) shows the AOD differences versus AERONET AOD. AOD differences versus solar zenith angle (SAZ in °) are shown in the lower panel (i–l).

cause there is not a comprehensive definition of aerosols in terms of OPAC components, we propose a different mixture of components for these sites. This mixture involves the same WASO and MINM contribution as desert-type aerosols, but the MIAM component is replaced by MITR, keeping the same relationship with N_{mineral} as MIAM, and the MICM contribution is discarded.

We have generated the LUT using a set of n values for N_{mineral} , ranging from 0 to 4000 particles cm^{-3} , in order to calculate n height profiles, each one composed of all aerosol components present in every layer ($N_i(h)$). Then, with these profiles, we simulate a set of n ZSR values for each SAZ, in addition to the corresponding AOD in the ZEN-LUT.

Finally, to find out the right value for N_{mineral} and thus the right AOD from the LUT, we applied an adapted version of the method described in Tanré et al. (1997), originally designed to select the right aerosol model from satellite radiance measurements. In this case, the selection is performed by finding the minimum value of the following expression:

$$\epsilon_l = \sqrt{\frac{1}{N_\lambda} \sum_{\lambda=1}^{N_\lambda} \left(\frac{L_\lambda^m(\theta_v=0, \theta_s) - L_{\lambda,l}^c(\theta_v=0, \theta_s)}{L_\lambda^m(\theta_v=0, \theta_s)} \right)^2}, \quad (4)$$

where L_λ^m and $L_{\lambda,l}^c$ are the measured (earth–sun distance corrected) and the computed radiances in wavelength λ , N_λ is the total number of wavelengths and l is the index indicating

a value defined in the array N_{mineral} . θ_v is the viewing zenith angle which is equal to zero and θ_s the solar zenith angle. Then the value of the index l , which minimizes the quantity ϵ_l , indicates the right value for N_{mineral} and AOD.

4.1 ZEN-LUT method sensitivity study

The precision and accuracy of the solution given by the presented ZEN-LUT method are related to random and systematic errors. In the present work we have focused on performing a sensitivity study on the main systematic errors, which can be classified in two categories: instrumental errors, and errors made in the a priori considered RTM inputs, causing underestimation or overestimation in the measured and simulated ZSR, which affects the retrieved AOD. In the first category, radiance calibration error, levelling error, misalignments of the optical parts, the effect of the finite field of view and the stray light contribution are included. The assumed values for surface albedo and the a priori considered mixture of aerosol components are included in the second category.

We have performed this sensitivity study using the inputs previously defined in Sect. 4.

This type of analysis is commonly used to identify the dominant contributors to the output variability. However, if the effect of each source of error is conveniently quantified, it is possible to estimate the contribution of each parameter to the uncertainty in the final AOD retrieval.

4.1.1 Instrumental error sensitivity

Taking into account previous studies in the literature (Eck et al., 1999; Basart et al., 2009; Guirado et al., 2014; Cesnu-lyte et al., 2014; Cuevas et al., 2015) and AERONET climatology tables (https://aeronet.gsfc.nasa.gov/cgi-bin/climo_menu_v2_new), AOD desert stations such as Tamanrasset (Algeria), Ouarzazate (Morocco), Dakar (Senegal) and Solar Village (Saudi Arabia) show monthly and seasonal values typically ranging between 0.04 (Ouarzazate in January) and 0.67 (Dakar in June), with sporadic AOD maxima > 1 associated with strong desert dust outbreaks or local dust resuspension. Therefore, we have focused our sensitivity study in AOD conditions varying between 0.5, which corresponds to hazy conditions by dust, and 1, indicating strong dust intrusions.

We have assumed the ZEN-R41 calibration uncertainty to be $\sim 5\%$ considering the comparison analysis with an integrating sphere presented in Sect. 3. In order to evaluate the influence of the overall instrumental errors on the inferred AOD we have perturbed the computed ZSR ± 5 and $\pm 10\%$ for two aerosol loads ($\text{AOD}_{500\text{nm}} \sim 0.5$ and 1.0) at all the available wavelengths and SZAs ranging from 20 to 65° in 10° steps. Although this perturbed range should be statistically estimated by comparing several ZEN-R41 instruments, we have selected these values according to the relative mean bias results obtained in the ZSR comparison between CE318 and ZEN-R41 presented in Sect. 3.1 and Table 1, in which we considered the CE318 as the reference instrument. Differences in AOD obtained from unperturbed and perturbed ZSR values are presented in Fig. 3 for 440 and 870 nm wavelengths. In case of $\text{AOD}_{500\text{nm}} \sim 0.5$, we obtained absolute differences < 0.05 and < 0.1 for perturbed radiances of ± 5 and $\pm 10\%$, which are not dependent on SZA and wavelength. For higher aerosol load conditions ($\text{AOD}_{500\text{nm}} \sim 1.0$), we have found nearly constant AOD differences of approximately -0.1 and -0.2 with SZA for negative radiance perturbations of -5 and -10% respectively. However, in the case of positive ZSR perturbations, we observed a SZA dependence with higher AOD differences as SZA decreases, up to ~ 0.35 for a positive radiance perturbation of 10% and SZA of 20° .

Table 2. Sensitivity study for an albedo perturbation of $\pm 15\%$. AOD differences between perturbed and unperturbed situations (ΔAOD) for 440 and 870 nm spectral bands are shown for two AOD values ($\text{AOD}_{500\text{nm}} \sim 0.5$ and 1.0) and different solar zenith angles (SZA) ranging from 20 to 65° .

AOD	SZA	ΔAOD_{440}	ΔAOD_{870}
0.5	20	$-0.006/0.006$	$-0.006/0.007$
	30	$-0.010/0.010$	$-0.011/0.011$
	40	$-0.013/0.013$	$-0.013/0.013$
	50	$-0.015/0.016$	$-0.015/0.017$
	60	$-0.017/0.017$	$-0.017/0.017$
	65	$-0.016/0.017$	$-0.016/0.017$
1.0	20	$-0.016/0.017$	$-0.016/0.018$
	30	$-0.023/0.024$	$-0.025/0.026$
	40	$-0.028/0.030$	$-0.031/0.033$
	50	$-0.032/0.033$	$-0.035/0.038$
	60	$-0.033/0.036$	$-0.037/0.040$
	65	$-0.033/0.034$	$-0.037/0.040$

4.1.2 RTM input sensitivity

Surface albedo sensitivity

The effect of the albedo uncertainty on the inferred AOD was also assessed for the same AOD and solar zenith angle range. We have performed the study for albedo values $\pm 15\%$ of those values used in the generated look-up table. We have considered this variation range, taking into account the results obtained by Tsvetsinskaya et al. (2006), who studied the spatial and temporal variability of surface albedo using MODIS data. They found very stable albedo values with a temporal variability that are essentially negligible in cases of surfaces with high reflectivity, such as deserts, and spatial variations between 14 ($\lambda < 700\text{nm}$) and 9% ($\lambda > 700\text{nm}$). Consequently, a variability in surface albedo of 15% seems reasonable for our sensitivity study. The computed ZSR with the modified albedo values are used as inputs to retrieve the AOD and the result is compared with the actual AOD. The AOD differences for 440 and 870 nm, and several SZA and $\text{AOD}_{500\text{nm}}$ values are shown in Table 2. This table shows that the surface reflectance effect on AOD is relatively low, with AOD differences somewhat higher with SZA, ranging from -0.016 to 0.017 for an $\text{AOD}_{500\text{nm}} \sim 0.5$, and from -0.037 to 0.040 for an $\text{AOD}_{500\text{nm}} \sim 1.0$. The wavelength dependence is also almost negligible.

Mix of aerosol components sensitivity

In order to test the influence of the mixture of aerosol components in the inferred AOD, we have modified the assumed mixing ratio of the reference LUT for mineral dust and WASO components present in the boundary layer. Mineral dust components were perturbed in $\pm 5\%$ of the slope co-

efficient in Eqs. (1), (2) and (3), and WASO components, assumed as a fixed value, were perturbed assuming a variation in concentration of $\pm 50\%$. AOD differences considering perturbed and unperturbed values for the mix of aerosol components are presented in Fig. 4 for two AOD conditions ($\text{AOD}_{500\text{nm}} \sim 0.5$ and 1.0) and two spectral ranges (440 and 870 nm). We have found negligible influence for $\text{AOD}_{500\text{nm}} \sim 0.5$, with AOD differences ranging from ± 0.02 for all aerosol components, although a slight dependence with SZA and wavelength can be appreciated. For higher aerosol content ($\text{AOD}_{500\text{nm}} \sim 1.0$), we observed little dependence with SZA and wavelength on AOD difference. We found AOD differences ranging from -0.1 to 0.05 .

This sensitivity analysis showed that ZEN AOD combined standard uncertainty (determined by means of summation in quadrature of each term analysed in the sensitivity study) is up to 0.06 (for $\text{AOD}_{500\text{nm}} \sim 0.5$) and 0.15 (for $\text{AOD}_{500\text{nm}} \sim 1.0$) when an instrumental error of 5% is considered. Higher AOD uncertainty is expected if instrumental error is $> 5\%$, up to 0.23 in the 10% limit of instrumental error (0.37 in case of very low SZAs).

5 Results

5.1 ZEN-LUT method validation at three different AERONET sites

The performance of the ZEN-LUT methodology was tested in three different sites in which mineral dust plays an important role in their respective aerosol climatologies (Izaña, Santa Cruz and Tamanrasset) using a CE318 instrument as a reference, since it has been already widely tested under quite different conditions (Eck et al., 1999, 2001, 2003a, b, 2010; Holben et al., 2001, 2006; Dubovik et al., 2000). The CE318-LUT AOD was inferred after applying the ZEN-LUT methodology on ZSR data derived from PPL data provided by AERONET.

AOD data from Izaña and Santa Cruz stations were cloud screened using information from the SONA all-sky imager. Since any all-sky camera was available at Tamanrasset station, the cloud screening was performed by applying the PPL smoothness criteria described in Sect. 2.2.1.

We have restricted this analysis to desert dust events. We have identified the desert aerosol conditions following the criteria given in Guirado (2014), and Guirado et al. (2014) (Sect. 2.1). We have used AERONET data for the year 2013, as it was the most recent period of available AERONET 2.0 level AOD data in the three sites. The comparison analysis performed at the three stations at every available wavelength (440, 675 and 870 nm) is shown in Figs. 5, 6 and 7 and Table 3. We present the AOD CE318-LUT/CE318-AERONET scatter plot in Figs. 5a–c, 6a–c and 7a–c, which shows a good agreement for all channels in the three locations, with high correlations ($R^2 > 0.97$). We have also found low RMSE val-

Table 3. Coefficient of determination (R^2), root mean square error (RMSE), mean bias (MB) and number of coincident data (N) for the AOD comparison between CE318-AERONET and CE318-LUT at three different spectral bands (440, 675 and 870 nm) performed at Izaña (IZO), Santa Cruz (SCO) and Tamanrasset (TAM) stations in 2013.

Station	Wavelength (nm)	R^2	RMSE	MB	N
IZO	870	0.97	0.011	−0.018	180
	675	0.98	0.010	−0.020	180
	440	0.97	0.011	−0.018	180
SCO	870	0.99	0.021	−0.021	210
	675	0.99	0.021	−0.019	210
	440	0.99	0.021	−0.020	210
TAM	870	0.98	0.030	−0.031	321
	675	0.98	0.030	−0.022	321
	440	0.97	0.032	−0.023	321

Table 4. Coefficient of determination (R^2), root mean square error (RMSE), mean bias (MB) and number of coincident data (N) for the AOD comparisons between CE318-AERONET and ZEN-R41 at four different spectral bands (440, 500, 675 and 870 nm) performed at Izaña station in 2015.

Wavelength (nm)	R^2	RMSE	MB	N
870	0.97	0.026	−0.020	616
675	0.97	0.026	−0.025	616
500	0.97	0.026	−0.029	616
440	0.97	0.027	−0.030	616

ues for Izaña and Santa Cruz (up to 0.011 and 0.021) but higher values for Tamanrasset (up to 0.032). The linear dependence between the two AOD data sets is different in the considered channels, with the lower slope having the shorter wavelength at all sites, and more evidently for Izaña. With regard to the AOD differences plotted in Figs. 5d–i, 6d–i and 7d–i, it can be said that AOD is mostly underestimated by the ZEN-LUT method, with negligible dependence with SZA for the SZA range considered in this study. We have found maximum AOD differences up to -0.07 for Izaña and between -0.08 and 0.12 for Santa Cruz. These values are in agreement with the expected uncertainty involved in the ZEN-LUT methodology of AOD determination presented in Sect. 4.1. It also confirms the low instrumental errors affecting the CE318 ($\sim 5\%$). The considerably higher differences found for Tamanrasset might be explained by possible dust contamination of the CE318 lenses.

5.2 ZEN-R41 and CE318 AERONET AOD comparison at Izaña

We have performed an intercomparison of AOD from ZEN-R41 and CE318 at Izaña station during 2015. ZEN-R41 AOD, retrieved by applying the ZEN-LUT method to cloud-screened ZSR data, was compared with AERONET level 1.5 AOD for the 20–65° SZA range. The results of the intercomparison for the four spectral bands are presented in Fig. 8 and Table 4. In Fig. 8a–d we present a scatter plot between ZEN-R41 and CE318-AERONET AOD. A good correlation can be observed between both AODs, with $R^2 = 0.97$ and RMSEs ~ 0.026 – 0.027 for all channels, indicating that the ZEN-R41 instrument and ZEN-LUT method together are adequate for AOD estimation. However, as we pointed out in the previous subsection, the linear dependence between the two AOD data sets is different in the four channels. The AOD differences plotted in Fig. 8e–l shows a similar behaviour to those presented in the former subsection for Izaña, showing that AOD is mostly underestimated; although in this case the differences are larger as we have found maximum observed AOD differences up to 0.15 and RMSE values up to 0.030 (Table 4), but within the uncertainties estimated in Sect. 4.1.

6 Conclusions

In this study we have presented the new ZEN system for dust AOD monitoring from ZSR observations. The ZEN system comprises the development of a new and robust radiometer (ZEN-R41) and a methodology to retrieve AOD from ZSR measurements through a look-up table (LUT) method (ZEN-LUT), especially designed for desert aerosols. This methodology, inspired by previous methodologies commonly applied to on-board satellite sensors, uses the radiative transfer code LibRadtran and its packages to simulate ZSRs and the associated AODs. Then, AOD is inferred by minimizing a function which depends on the differences between simulated and measured ZSR at the corresponding solar zenith angle (SZA) in all the available wavelengths.

The main conclusions of this study are as follows.

1. The comparison of ZSR from ZEN-R41 and CE318 showed a high coefficient of determination (R^2) for all wavelengths (0.99), although we observed relative mean bias (RMB) as high as -6.3% at 870 nm and 9.2% at 500 nm.
2. The sensitivity analysis, performed to identify the systematic errors (instrumental and radiative transfer inputs, or RTM) exerting the most influence on the final AOD, showed the instrumental errors and the aerosol model as the most important contributors to the final AOD uncertainty. This analysis estimated an AOD combined standard uncertainty in the ZEN system up to 0.06 in the case of $\text{AOD}_{500\text{nm}} \sim 0.5$ and 0.15 for

$\text{AOD}_{500\text{nm}} \sim 1.0$, provided instrumental errors are minimized ($\sim 5\%$).

3. We have compared AOD from AERONET with AOD retrieved from CE318 ZSR by means of the ZEN-LUT method (CE318-LUT) in a common period in which AERONET level 2.0 is available at three stations (Izana, Santa Cruz and Tamanrasset). A good AOD agreement (R^2 from 0.99 to 0.97) and RMSE values from 0.011 (at Izaña) to 0.032 (at Tamanrasset) have been obtained. The relatively bad results observed at Tamanrasset might be explained by possible contamination of lenses by dust. We observed maximum AOD differences up to -0.07 at Izaña and between 0.08 and 0.12 for Santa Cruz, in agreement with the expected uncertainty involved in the ZEN-LUT methodology.
4. The AOD comparison at Izaña showed a good agreement between ZEN-R41 and AERONET (R^2 of 0.97), with observed AOD differences up to 0.15, and ZEN-R41 AOD systematically underestimated (mean bias ranging from -0.020 to -0.030). These results are also in agreement with the expected uncertainty for the ZEN system.

The results of this preliminary study indicate that the ZEN-LUT method is appropriate to infer dust AOD from ZSR measurements from the ZEN-R41, with an expected uncertainty between 0.06 and 0.15 in the AOD range between 0.5 and 1.0, which seems reasonable for most Saharan and Middle Eastern sites affected by dust. However, a thorough validation with a higher number of ZEN-R41 radiometers installed in stations located in quite different environments affected by desert dust will be carried out in the near future to confirm and complement the results presented in this paper. The study of the impact caused by other aerosols is an important issue to be addressed in order to adapt this instrument to other environments, far from dust sources. Although clouds are not a major problem in desert regions for much of the year, autonomous cloud-screening system is being implemented into the ZEN-R41 radiometer in order to discriminate cloud-contaminated ZSR data.

With the ZEN system we do not intend at all to replace accurate AOD measurements performed by sun photometer networks (as AERONET) but to complement these observations with the aim of improving mineral dust monitoring in remote locations, where it is difficult to deploy sun photometers for logistical reasons and poor infrastructure. The ZEN system could be used individually with autonomous data processing, to create networks with centralized data processing or ultimately be incorporated into automatic weather stations in desert regions in an inexpensive and simple way. As a consequence, this instrument could play a key role in dust model data assimilation near dust source regions, in satellite validation and in early warning within the WMO Sand

and Dust Storm Warning Advisory and Assessment System (SDS-WAS).

7 Data availability

The reference aerosol optical depth and zenith sky radiance data at the three sites (Izaña, Santa Cruz and Tamanrasset) were obtained from Cimel Electronique 318 photometers and they are available from the AERONET website (<https://aeronet.gsfc.nasa.gov/>). All the look-up tables as well as the sensitivity study data were generated by means of radiative transfer simulations carried out with LibRadtran version 1.7 code, available at <http://www.libradtran.org/>. All input files for performing these radiative transfer simulations are available upon request. All ZEN-R41 data are available upon request. The SONA all sky camera images for Izaña and Santa Cruz station are available upon request.

Competing interests. The authors declare that they have no conflict of interest.

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