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# Methane and nitrous oxide from ground-based FTIR at Addis Ababa: observations, error analysis, and comparison with satellite data

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Abstract. A ground-based, high-spectral-resolution Fourier transform infrared (FTIR) spectrometer has been operational in Addis Ababa, Ethiopia (9.01° N latitude, 38.76° E longitude; 2443 m altitude above sea level), since May 2009 to obtain information on column abundances and profiles of various constituents in the atmosphere. Vertical profile and column abundances of methane and nitrous oxide are derived from solar absorption measurements taken by FTIR for a period that covers May 2009 to March 2013 using the retrieval code PROFFIT (V9.5). A detailed error analysis of CH<sub>4</sub> and N<sub>2</sub>O retrieval are performed. Averaging kernels of the target gases shows that the major contribution to the retrieved information comes from the measurement. Thus, average degrees of freedom for signals are found to be 2.1 and 3.4, from the retrieval of CH<sub>4</sub> and N<sub>2</sub>O for the total observed FTIR spectra. Methane and nitrous oxide volume mixing ratio (VMR) profiles and column amounts retrieved from FTIR spectra are compared with data from the reduced spectral resolution Institute of Meteorology and Climate Research/Instituto de Astrofísica de Andalucía (IMK/IAA) MI-PAS (Version V5R\_CH4\_224 and V5R\_N2O\_224), the Microwave Limb Sounder (MLS) (MLS v3.3 of N2O and CH4 derived from MLS v3.3 products of CO, N2O, and H2O), and the Atmospheric Infrared Sounder (AIRS) sensors on board satellites. The averaged mean relative difference between FTIR methane and the three correlative instruments MIPAS, MLS, and AIRS are 4.2 %, 5.8 %, and 5.3 % in the altitude ranges of 20 to 27 km, respectively. However, the biases below 20 km are negative, which indicates the profile of CH<sub>4</sub> from FTIR is less than the profiles derived from correlative instruments by -4.9%, -1.8%, and -2.8%. The averaged positive bias between FTIR nitrous oxide and correlative instrument, MIPAS, in the altitude range of 20 to 27 km is 7.8 %, and a negative bias of -4 % at altitudes below 20 km. An averaged positive bias of 9.3 % in the altitude range of 17 to 27 km is obtained for FTIR N<sub>2</sub>O with MLS. In all the comparisons of CH<sub>4</sub> from FTIR with data from MI-PAS, MLS, and AIRS, sensors on board satellites indicate a negative bias below 20 km and a positive bias above 20 km. The mean error between partial-column amounts of methane from MIPAS and the ground-based FTIR is -5.5%, with a standard deviation of 5% that shows very good agreement as exhibited by relative differences between vertical profiles. Thus, the retrieved CH<sub>4</sub> and N<sub>2</sub>O VMR and column amounts from Addis Ababa, tropical site, is found to exhibit very good agreement with all coincident satellite observations. Therefore, the bias obtained from the comparison is comparable to the precision of FTIR measurement, which allows the use of data in further scientific studies as it represents a unique environment of tropical Africa, a region poorly investigated in the past.

### 1 Introduction

Methane (CH<sub>4</sub>), nitrous oxide (N<sub>2</sub>O), and chlorofluorocarbons (CFCs) are tropospheric species, which are the main

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source gases for the chemical families  $NO_x$ ,  $CIO_x$ , and  $HO_x$  (Jacobson, 2005). The reaction of  $CH_4$  with hydroxyl radicals reduces ozone in the troposphere, and it influences the lifetime or production of other atmospheric constituents such as stratospheric water vapor and  $CO_2$  (Michelsen et al., 2000; Boucher et al., 2009), whereas the lifetime of  $N_2O$  is determined by its rate of UV photolysis or reaction with  $O(^1D)$  (Collins et al., 2010).

Methane retrievals from near-infrared spectra recorded by the SCIAMACHY instrument on board ENVISAT suggested an unexpectedly large tropical CH4 emissions and the impact of water spectroscopy on methane retrievals, with the largest impacts in the tropics (Frankenberg et al., 2008b). The recent increasing impact of CH<sub>4</sub> and N<sub>2</sub>O to global warming has also been assessed by the last AR4 IPCC report (IPCC, 2007; Sussmann et al., 2012). Nitrous oxide (N<sub>2</sub>O) becomes the dominant ozone-depleting substance emitted in the 21st century (Ravishankara et al., 2009). In 2007 and 2008, The Infrared Atmospheric Sounding Interferometer (IASI) on board METOP-A observed an increase in midtroposphere methane in the tropical region of  $9.5 \pm 2.8$  and  $6.3 \pm 1.7 \,\mathrm{ppbv}\,\mathrm{yr}^{-1}$ , respectively (Crevoisier et al., 2013). Long-lived compounds ascend in the tropics, across the tropical tropopause, and are subsequently redistributed by the Brewer–Dobson circulation (Holton, 2004). According to the World Meteorological Organization (WMO), the 2010 report (WMO, 2010), 96 % of the increase in radiative forcing is due to the five long-lived greenhouse gases: carbon dioxide, methane, nitrous oxide, CFC-12, and CFC-11. The sources and sinks of atmospheric methane (CH<sub>4</sub>) and its budget in the tropics are not yet well quantified and have large uncertainties due to the scarcity of measurements (e.g., Meirink et al., 2008).

The tropics is the location where two important exchange processes in the atmosphere are taking place, the interhemispheric exchange and the entry of tropospheric air mass into the stratosphere (Petersen et al., 2010; Fueglistaler et al., 2009). Moreover, composition of a tropical atmosphere also plays a critical role in the stratospheric chemistry (Solomon, 1999; IPCC, 2007). Measurements and interpretation of atmospheric trace gas composition of the tropics is vital for a better understanding of the budgets, sources, and sinks of trace gases in the atmosphere and their effects on atmospheric chemistry, greenhouse effect, and climate changes globally. Emissions within the tropics contribute substantially to the global budgets of many important trace gases (IPCC, 2007; Frankenberg et al., 2008a).

The ground-based FTIR measurement at the Addis Ababa site was launched in 2009, in collaboration with Karlsruhe Institute of Technology, Germany, to measure concentrations of various trace gases in the lower and middle atmosphere over Addis Ababa. Thus, Addis Ababa FTIR measurements of atmospheric trace gases and their importance to understand various lower and middle atmospheric processes have been reported in a number of previous studies (Takele Ke-

nea et al., 2013; Mengistu Tsidu et al., 2015; Schneider et al., 2015, 2016; Barthlott et al., 2017). H<sub>2</sub>O volume mixing ratio (VMR) profiles and integrated column amounts from ground-based FTIR measurements of the Addis Ababa site were also compared with the coincident satellite observations of the Tropospheric Emission Spectrometer (TES), Atmospheric Infrared Sounding (AIRS), and Modular Earth Submodel System (MESSy) model, and the result confirmed reasonably good agreement (Samuel, 2014). Laeng et al. (2015) found that the MIPAS CH<sub>4</sub> profiles V5R\_CH4\_222 below 20 to 25 km is biased high by +14%. For a later and improved data version, namely V5R\_CH4\_224, Plieninger et al. (2016) found a positive bias between 0.1 and 0.2 ppmv. For the MIPAS N<sub>2</sub>O data version V5R\_N2O\_224, Plieninger et al. (2016) determined the bias to be between 0 and +30 ppb.

In this study, previous work on the intercomparison of ozone (Takele Kenea et al., 2013) and water vapor (Samuel, 2014) are extended to source gases CH<sub>4</sub> and N<sub>2</sub>O from ground-based FTIR. Intercomparisons of vertical profiles and column amounts retrieved from solar spectra observed by the Fourier transform spectrometer at the Addis Ababa site with data from MIPAS, MLS, and AIRS sensors on board satellites were made to assess the quality of the data derived from FTIR. The observed differences between ground-based FTIR and satellite observation of CH<sub>4</sub> and N<sub>2</sub>O are analyzed using the statistical tools detailed in von Clarmann (2006). The measurement site and the FTIR spectrometer along with the retrieval approach will be introduced in Sect. 2, and the retrieved information content and spectral analysis will be discussed in Sect. 3. A short description of satellite measurement techniques followed by the detailed intercomparison with those products will be presented in Sects. 4 and 5, respectively. Finally, a summary and conclusions are given in Sect. 6.

# 2 Measurement site and instrumentation

#### 2.1 Measurement site

The ground-based FTIR at Addis Ababa was established to acquire high-quality, long-term measurements of trace gases to understand chemical and dynamical processes in the atmosphere and to validate models and satellite measurements of atmospheric constituents. The geographic position of the observatory is 9.01° N, 38.76° E, 2443 m a.s.l., and its suitability has been confirmed from the measurements of tropical stratospheric ozone, precipitable water vapor, and isotopic composition of water vapor (Takele Kenea et al., 2013; Mengistu Tsidu et al., 2015; Schneider et al., 2015, 2016; Barthlott et al., 2017). Addis Ababa is a tropical high-altitude observing site and as such is important to the understanding of processes near the tropical tropopause. Physical process in the tropics, mainly around tropopause layer, has a vital role in climate change and the general circulation of the tropi-

cal troposphere, which would control the transport of energy, water vapor, and trace gases in the climate system derived by the deep convection (Holton and Gettelman, 2001). Thus, the observed variation in the measurement of atmospheric trace gases would help us to understand the effects of tropical dynamics on the site. Besides, it fills a data gap due to the scarcity of ground-based measurements in tropics.

# 2.2 The FTIR spectrometer and retrieval

Fourier transform spectroscopy has been applied successfully to study trace gases in the atmosphere by examining atmospheric absorption lines in the infrared spectrum from solar radiation. Measurement of the sun's spectra at the Earth's surface provides information about atmospheric composition. This technique uses the sun as a light source in order to quantify molecular absorptions in the atmosphere and then retrieve trace gases abundance. The high-resolution FTIR spectrometer, Bruker IFS120M, upgraded with 125 M electronics, from Bruker Optik GmbH, in Germany, was installed in May 2009 at the Addis Ababa site. This interferometer is equipped with indium—antimonide (InSb) detector, which allows coverage of the 1500–4400 cm<sup>-1</sup> spectral interval. In this spectral range, a large number of species that reside in the atmosphere can be detected.

The measured spectra have been analyzed using an algorithm that simulates the spectra and Jacobians by the lineby-line radiative transfer model PRFFWD (PRoFit ForWarD model) to produce the synthesized spectra, and the vertical profiles of CH<sub>4</sub> and N<sub>2</sub>O would be derived by applying a retrieval code PROFFIT (Ver95) (Hase et al., 2004). It has been developed based on semiempirical implementation of the optimal estimation method (Rodgers, 2000) to derive the VMR profiles and column amounts of multiple species. Hence, CH<sub>4</sub> and N<sub>2</sub>O profiles from measured spectra in the microwindows that span a spectral range of 2400–2800 cm<sup>-1</sup> have been discussed in this paper. A Tikhonov-Phillips regularization method on a logarithmic scale was used to derive the profiles. Retrieved state vector  $\hat{x}$  is related to a priori  $(x_a)$ and true state vectors (x) by the following mathematical expression:

$$\hat{\mathbf{x}} = \mathbf{x}_{\mathbf{a}} + \hat{\mathbf{A}}(\mathbf{x} - \mathbf{x}_{\mathbf{a}}) + \varepsilon, \tag{1}$$

where  $\hat{\mathbf{A}}$  is averaging kernel matrix, and  $\varepsilon$  is the measurement error. Moreover, the actual averaging kernel matrix depends on several parameters including the solar zenith angle, the spectral resolution and signal-to-noise ratio, the choice of retrieval spectral microwindows, and the a priori covariance matrix  $\mathbf{S}_a$ . The elements of the averaging kernel for a given altitude gives the sensitivity of retrieved profiles at which the real profile is present, and its full width at half maximum is a measure of the vertical resolution of the retrieval at that altitude (Rodgers, 1990). Error estimation analysis is based on the analytical method suggested by Rodgers (2000) as fol-

lows:

$$\hat{\mathbf{x}} - \mathbf{x} = (\mathbf{A} - \mathbf{I})(\mathbf{x} - \mathbf{x}_x) + \mathbf{G}\mathbf{K}_b(\mathbf{b} - \mathbf{b}_a) + \mathbf{G}\varepsilon. \tag{2}$$

The averaging kernel matrix can be defined as  $\mathbf{A} = \mathbf{GK}$ ;  $\mathbf{I}$  is the identity matrix, and  $\mathbf{G}$  is the gain matrix that represents the sensitivity of retrieved parameters to the measurement.  $\mathbf{K}_b$  is the sensitivity matrix of the spectrum to the forward model parameters  $\mathbf{b}$ . Since we do not know the true state of the atmosphere, we can not specify the actual retrieval error; we can only make a statistical estimate of it, which is expressed in terms of a covariance matrix. The total error in the retrieved profile can be described as a combination of measurement error and forward model parameter error. It has been suggested by Rodgers (2000) to include smoothing error in the total error budget, but this concept has been revised by von Clarmann (2014).

### 3 Information content and error analysis

# 3.1 Spectroscopic data and a priori profiles

In our retrieval strategy, the profiles of CH<sub>4</sub> and N<sub>2</sub>O were retrieved, while the profiles of interfering species (see Table 1) were scaled. A prior  $x_a$  profiles for methane and the interfering species above Addis Ababa were taken from 40-year averages (1980-2020) of the Whole Atmosphere Community Climate Model (WACCM; Garcia et al., 2007). Similarly, the a priori profile for nitrous oxide has also been constructed from monthly average data available from WACCM (e.g., Tilmes et al., 2007), whereas the grid to be used for the Addis Ababa site is found with the WACCM mixing ratio profile data at ftp://ftp.acom.ucar.edu/user/jamesw/IRWG/2013/ WACCM/V6/Addis Ababa/ (last access: 21 July 2020), as recommended by the NDACC/IRWG (Network for the Detection of Atmospheric Composition Change/ Infrared Working Group). WACCM is a numerical model developed at the National Center for Atmospheric Research (NCAR). Daily profiles of pressure and temperature were taken from the National Centers for Environmental Prediction (NCEP) (http://www.cdc.noaa.gov/data/gridded/ data.ncep.reanalysis.html) (last access: 21 July 2020) reanalysis and are made available through the NASA Goddard Space Flight Center automailer. The spectroscopic parameters were taken from the HITRAN (high-resolution transmission molecular absorption) database, version 2008, for  $N_2O$ , and 2009 for  $H_2O$  (Rothmann et al., 2009), and the updated HITRAN 2012 for CO, CH<sub>4</sub>, and NO<sub>2</sub> (Rothmann et al., 2013) were used during retrieval of CH<sub>4</sub> and N<sub>2</sub>O.

Both methane (CH<sub>4</sub>) and nitrous oxide ( $N_2O$ ) are well mixed in the troposphere, and their VMR decreases with height and becomes negligible with no variation above 55 km. The vertical variability in  $N_2O$  and CH<sub>4</sub> in the lower stratosphere is characterized by a somewhat higher vertical gradient as compared to the other layers. Both

Gas	Microwindow (cm <sup>-1</sup> )	Interfering species	DOFs
CH <sub>4</sub>	(2599.8, 2600.5) (2614.87, 2615.4) (2650.8, 2651.29)	H <sub>2</sub> O, CO <sub>2</sub> , NO <sub>2</sub>	$2.045 \pm 0.18$
	(2760.6, 2761.23) (2778.22, 2778.55)		
N <sub>2</sub> O	(2464.2, 2465.57) (2486.55, 2488.18) (2491.86, 2492.9) (2522.95, 2524.1)	H <sub>2</sub> O, CO <sub>2</sub> , CH <sub>4</sub>	$3.38 \pm 0.15$

**Table 1.** Microwindows, interfering gases, and their DOFs listed in the table are used for the retrieval of VMR profiles and column amounts of CH<sub>4</sub> and N<sub>2</sub>O from FTIR spectra recorded at Addis Ababa.

profiles and columns of CH<sub>4</sub> and N<sub>2</sub>O over Addis Ababa have been obtained by fitting five and four selected spectral regions for CH<sub>4</sub> and N<sub>2</sub>O, respectively. Here, spectral microwindows used for the retrieval are selected such that they contain absorption features of the target species along with a minimal number of interfering absorption lines; and they have been adopted from different sources (Senten et al., 2008; Sussmann et al., 2011; Meier et al., 2004). Microwindows, target, and interfering species used in this paper are summarized in Table 1. However, the microwindows are somewhat modified for the Addis Ababa FTIR site from the windows recommended by NDACC, as mentioned in a result of work done within the EU UFTIR projects (http://projects.amap.no/project/uftirtime-series-of-upper-free-troposphere-observations-from -a-european-ground-based-ftir-network/, last access: 21 July 2020). The choice of these microwindows over those recommended by NDACC is due to their improved performance as indicated in the Supplement. The main criterion for selection of these microwindows is high sensitivity to methane and low interference from other gases. Our tests have shown that these windows are still appropriate for the Addis Ababa site. Methane and nitrous oxide vertical profiles over Addis Ababa have been obtained by fitting five and four microwindows, respectively. The retrieved state vector contains volume mixing ratios of the target gas defined in 41 layers of the tropical atmospheric conditions.

PROFFIT includes various retrieval options such as scaling of a priori profile, the Tikhonov–Phillips method (Phillips, 1962; Tikhonov, 1963), or the optimal estimation method (Rodgers, 2000). In this study, an optimized retrieval strategy for Addis Ababa has been established for CH<sub>4</sub> and  $N_2O$  by applying it first to single spectra, as test cases, and later routinely to the full set of measurements. Partly, the strategy to optimally retrieve the total columns of CH<sub>4</sub> and  $N_2O$  is to search for a set of spectra microwindows. A constraint, initial guess, and a priori profile are chosen in such a way that all the structures visible in the retrieved distributions originate from the measurements and are not artifacts

due to any constraints. At the Addis Ababa site, we did not use the a priori covariance matrix as an optimal estimation. However, the Tikhonov-type L<sub>1</sub> regularization method (Sussmann et al., 2009) on a logarithmic scale is used during the retrieval of CH<sub>4</sub> and N<sub>2</sub>O. The retrieval is performed on a fine vertical grid from 2.45 to 85 km and is stabilized by a first-order Tikhonov constraint,  $R = \alpha L^T L$ , where  $\alpha$  is the strength of the constraint, and L<sub>1</sub> is the first-order derivative (Borsdorff et al., 2014), which smooths the solution without biasing it towards the a priori profile. The parameter determines the weight of regularization, and it is also important to choose it to be appropriate to the problem. One way to fix this parameter is the L-curve method (Hansen, 1992). The regularization strength  $\alpha$ , is determined by finding a tradeoff between the number of degrees of freedom (a measure of the amount of information in methane and nitrous oxide retrieval), which is given by the trace of the averaging kernel matrix and the noise-induced error (Rodgers, 2000). A regularization strength  $\alpha$  of  $2.5 \times 10^4$  was found optimum for CH<sub>4</sub> retrieval.

The spectral fit and residual between measured and simulated spectra at five microwindows for CH<sub>4</sub> are shown in Fig. 1 for spectra recorded on 26 February 2013. Four microwindows are used for N<sub>2</sub>O and depicted in Fig. 2 for spectra recorded on 31 December 2009. The last column of Table 1 provides typical values for the degrees of freedom for signal (DOFs), and it indicates the possible independent pieces of information for the target gases distribution. The magnitude of residuals found from spectral fits span a range from a maximum of +0.25% to -0.64% for CH<sub>4</sub> and +0.34% to -0.34% for N<sub>2</sub>O. Hence, the residuals indicate systematic errors in the spectroscopic line data used to derive the concentration of CH<sub>4</sub> and N<sub>2</sub>O. Therefore, the fits are good with an averaged root-mean-square residual of 0.12% for the microwindows selected in the retrieval of CH<sub>4</sub>.

The quality of FTIR measurements during time period of May 2009–February 2011 for ozone has been revealed by Takele Kenea et al. (2013), whereas the measurements quality for CH<sub>4</sub> and N<sub>2</sub>O has also been assessed through the sen-

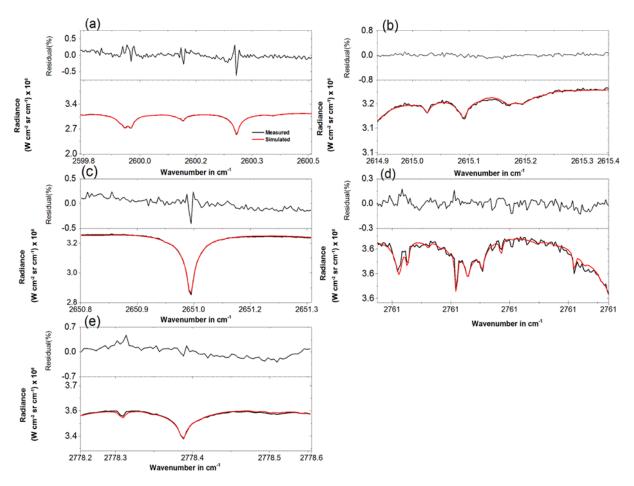


Figure 1. The five spectral microwindows used for retrieval of  $CH_4$ , with the measured spectrum in red, simulated spectrum in black, and residuals on top of the respective microwindow for spectrum recorded on 26 February 2013; time: 10 h 17 m 15 s; root mean square (rms) = 0.1189; solar zenith angle (SZA) =  $20.6^{\circ}$ ; optimal path difference (OPD) = 116.1; DOF = 2.23; field of view (FOV) = 2.27 mrad.

sitivity, DOFs, and the contribution of different error sources on measurements, in addition to the spectral residuals that indicate systematic errors in the spectroscopic line data.

# 3.2 Vertical resolution and sensitivity assessment

The spectral resolution of a measurement affects the amount of vertical information derived from the spectral line shape of a measured species (Livesey et al., 2008). Figure 3 shows averaging kernel matrices for the retrieval of the vertical profiles of  $CH_4$  and  $N_2O$  mixing ratios, respectively, from the FTIR measurements. The rows of the averaging kernel matrices at selected altitudes which indicate the sensitivity of retrieved  $CH_4$  and  $N_2O$  values at the level to true mixing ratios are also presented. The dotted line represents the sum of all the rows of the averaging kernel, which represents the overall sensitivity of the FTIR measurement to observe  $CH_4$  and  $N_2O$ .

Figure 3 shows a strong sensitivity in the altitude range of the troposphere and lower stratosphere, i.e., 2.45 up to 27 km for the retrieval of  $CH_4$  and  $N_2O$ . Thus, the sum of rows of  $\bf A$ 

for all the retrieval values of CH<sub>4</sub> and N<sub>2</sub>O is greater than 0.5 up to 27 km. The trace of the row-averaging kernel for CH<sub>4</sub>, which is 2.25 for the spectra recorded on 26 February 2013 and  $2.11 \pm 0.06$  for all the data, implies that partial columns representing two different altitude ranges in the atmosphere can be obtained from the observations of CH<sub>4</sub> in tropical atmospheric conditions. Similarly, the trace of the averaging kernel for N<sub>2</sub>O is  $3.38 \pm 0.15$  on for all the data.

The amplitude of the averaging kernels indicates the sensitivity of the retrieval, and the full widths at half maximum (FWHM) indicate the vertical resolution of the corresponding layer. We also ignore the altitude range where the resolution of the instrument is beyond 20 km, which has been computed using the reciprocal of the diagonal values of averaging kernels and multiplying by the intervals of the layers as reported in Rinsland et al. (2005). The vertical resolution is less than 20 km for altitudes below around 27 km (not shown).

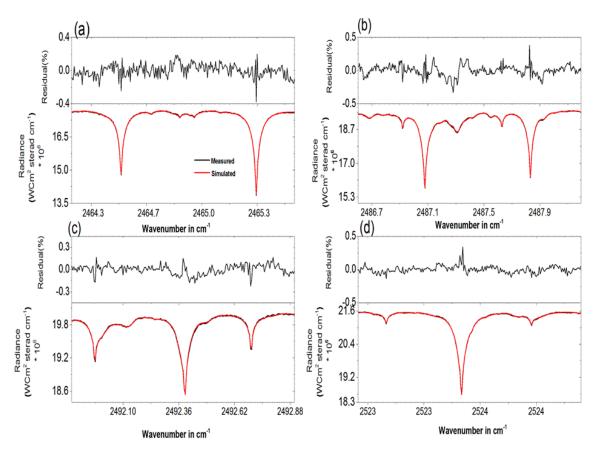
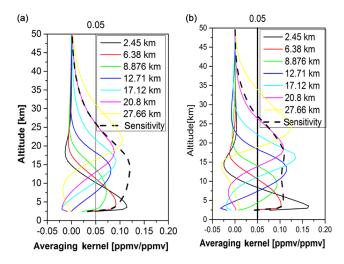


Figure 2. The four spectral microwindows used for retrieval of  $N_2O$ , with the measured spectrum in red, simulated spectrum in black, and residuals on top of the respective microwindow for spectrum recorded on 31 December 2009; time: 9 h 3 m 27 s; solar zenith angle  $(SZA) = 13.4^{\circ}$ ; optimal path difference (OPD) = 100; DOF = 3.35.



**Figure 3.** Sensitivity analysis of the retrieved profiles of (a)  $CH_4$  and (b)  $N_2O$  at Addis Ababa using the selected rows of the averaging kernels as a function of altitude. The dotted lines are the sum of the rows of averaging kernels for a spectrum measured on 26 February 2013 for  $CH_4$  and 31 December 2009 for  $N_2O$ .

#### 3.3 Error estimation

The error calculations conducted here are based on the error estimation package incorporated in the PROFFIT retrieval algorithm that was developed based on the analytical method suggested by Rodgers (2000). The quantified sources of errors are temperature, measurement noise, instrumental line shape, solar lines, line of sight, zero level baselines offset, and spectroscopy. It has been observed that baseline and atmospheric temperature uncertainties are the leading contribution to the total uncertainty. Details about the evaluation of individual contributions to the error budget are provided in Senten et al. (2008). Figure 4 shows the statistical (random) error, systematic error, and total fractional error (left to right) for CH<sub>4</sub> (top) and N<sub>2</sub>O (bottom) retrieval from a spectrum recorded on 26 February 2013 and 31 December 2009, respectively. It can be noted from Fig. 4 that the main systematic error source is the uncertainty in spectroscopic parameters, whereas the major statistical error source is the baseline. Random errors are dominated by the baseline offset uncertainty and the measurement noise in the troposphere. Total estimated random error due to parameter uncertainties is depicted as a dark-yellow line (see Fig. 4a). The total statisti-

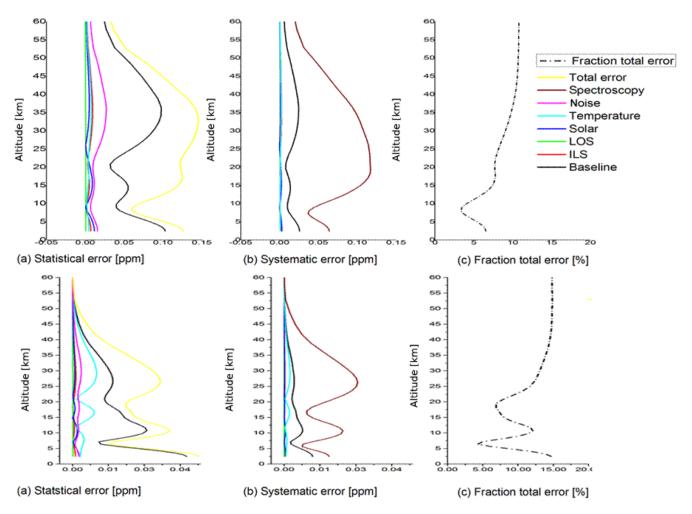


Figure 4. Estimated errors for the profiling retrieval of  $CH_4$  (top) and  $N_2O$  (bottom) over Addis Ababa: (a) statistical (random) errors (b) systematic errors of parameter listed in the legends, and (c) fractional total error (%).

cal error of CH<sub>4</sub> retrieval is about 0.07 ppmv  $(4.4\,\%)$  in the lower troposphere and about 0.04 ppmv  $(2.25\,\%)$  in the upper troposphere/lower stratosphere (UT/LS) region. Concerning systematic errors, spectroscopic parameters are the dominant uncertainty sources, and estimated total systematic error is about 0.05 ppmv  $(3.5\,\%)$  and 0.1 ppmv  $(7.2\,\%)$  for the lower troposphere and the UT/LS region, respectively.

Figure 4a–c (bottom panels) show the estimated random and systematic errors for the  $N_2O$  profile retrieved from FTIR. Random errors are dominated by the baseline offset uncertainty and temperature in the troposphere. The total statistical errors in the middle and upper troposphere are between 0.009 (3.5%) and 0.03 ppmv (9%), with its major contribution from the baseline. Spectroscopic parameters and baselines are the dominant uncertainty sources for systematic errors. The estimated total systematic error is less than 0.025 ppmv (8%) at altitudes below 22 km. The total fractional error of CH<sub>4</sub> and  $N_2O$  retrieved from ground-based FTIR has been shown in the last column of Fig. 4. Fractional error of CH<sub>4</sub> is less than 10% at altitudes below 27 km with

minimum fractional error of 4% at middle troposphere. On the other hand, the total fraction error of  $N_2O$  retrieval is less than 13% at altitudes below  $27 \,\mathrm{km}$ , with a minimum value of 4% at  $6 \,\mathrm{km}$  and 7.5% at  $17 \,\mathrm{km}$ .

# Time series partial-column amount

Concentrations of CH<sub>4</sub> and N<sub>2</sub>O were derived from 166 spectra of NDACC filter 3 recorded from May 2009 to March 2013. Figure 5 shows the time series of the retrieved total column amounts (in molecules cm<sup>-2</sup>) of CH<sub>4</sub> and N<sub>2</sub>O obtained from the Addis Ababa FTIR measurement site from 2009 to 2013. The mean total column amounts of CH<sub>4</sub> and N<sub>2</sub>O measured at Addis Ababa are  $2.9 \times 10^{19}$  molecules cm<sup>-2</sup>  $\pm 3.4\%$  and  $5.23 \times 10^{18}$  molecules cm<sup>-2</sup>  $\pm 6.93\%$ , respectively. Due to sensitivity of the observation in measuring CH<sub>4</sub> and N<sub>2</sub>O trace gases being limited to an altitude of around 27 km as explained using averaging kernel row of the measurement, the mean partial column of CH<sub>4</sub> and N<sub>2</sub>O

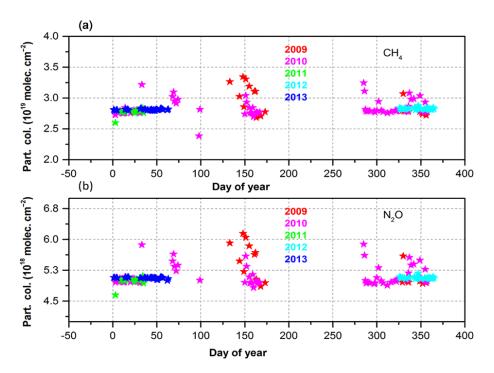


Figure 5. Partial columns of (a) CH<sub>4</sub> and (b) N<sub>2</sub>O gases over Addis Ababa in the altitude range of 2.45 to 27 km.

within the sensitivity range of the instrument are determined as  $2.85 \times 10^{19}$  molecules cm<sup>-2</sup>  $\pm 5.3\%$  and  $5.16 \times 10^{18}$  molecules cm<sup>-2</sup>  $\pm 6.95\%$ , respectively.

The sensitivity from the averaging kernel analysis is used to determine the upper altitude limit up to which CH<sub>4</sub> and  $N_2O$  data from ground-based FTIR can reasonably be used. The DOFs within these partial-column limits are about 1.03 and 1.27 for CH<sub>4</sub> and  $N_2O$ , respectively. Error analysis indicates that the statistical error accounts for 2.3 % in the total column amounts of CH<sub>4</sub> and 2.0 % in total columns of  $N_2O$ . Similarly, the systematic error accounts for 2.1 % in the total columns of CH<sub>4</sub> and 2.26 % in the total columns of  $N_2O$ . Generally, the overall contribution of both statistical and systematic errors to the total error during the retrieval of CH<sub>4</sub> and  $N_2O$  from ground-based FTIR are 3.1 % and 3 %, respectively.

# 4 Satellite measurements

# 4.1 Michelson Interferometer for Passive Atmospheric Sounding (MIPAS)

Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) is a Fourier transform spectrometer for the detection of limb emission spectra from the upper atmosphere to the lower thermosphere and is designed for global vertical profile measurement of many atmospheric trace constituents relevant to the atmospheric chemistry, dynamics, and radiation budget of the middle atmosphere. The verti-

cal resolution of MIPAS ranges from 2.5 to 7 km for CH<sub>4</sub> and from 2.5 to 6 km for  $N_2O$  in the reduced-resolution period (Plieninger et al., 2015). In this study, we have used the reduced spectral resolution Institute of Meteorology and Climate Research/Instituto de Astrofísica de Andalucía (IMK/IAA) MIPAS methane and nitrous oxide data products V5R\_CH4\_224 and V5R\_N2O\_224 (Plieninger et al., 2016, 2015). MIPAS profile points, with the diagonal element of the averaging kernels above 0.03 and the visibility flag of 1, have been used (Plieninger et al., 2016).

### 4.2 Microwave Limb Sounder (MLS)

The Earth Observing System (EOS) Microwave Limb Sounder (MLS) is one of four instruments on NASA's EOS Aura satellite, launched on 15 July 2004, into a near-polar, sun-synchronous orbit at 705 km altitude (Schoeberl et al., 2006). It measures N<sub>2</sub>O in spectral region, 640 GHz from the stratosphere into upper troposphere (Waters, 2006). Moreover, spatial coverage of this instrument is nearly global (-82° S to 82° N), and individual profiles are spaced horizontally by 1.5° or 165 km along the orbit track. Roughly, the satellite covers these latitudinal bands with 15 orbits per day or around 3500 profiles per day with vertical resolution of 4–6 km for N<sub>2</sub>O. This instrument ascends in the equatorial region at local time of around 13:45 LT.

MLS  $N_2O$  data set has been used to validate the ground-based FTIR measurements. However, methane (CH<sub>4</sub>) data are derived using coincident measurements of atmospheric water vapor (H<sub>2</sub>O), carbon monoxide (CO), and nitrous ox-

ide ( $N_2O$ ) from the EOS MLS instrument on the NASA Aura satellite, and details are given in Minschwaner and Manney (2015). Selection criteria were implemented as stated in Livesey et al. (2013). More details regarding the MLS experiment and data screening are provided in the above references in detail and at https://mls.jpl.nasa.gov/data/datadocs. php (last access: 21 July 2020). MLS  $N_2O$  v2.2 has been validated, and its precision and accuracy is in Lambert et al. (2007). The authors reported that MLS  $N_2O$  precision is 24–14 ppbv (9 %–41 %) and the accuracy is 70–3 ppbv (9 %–25 %) in the pressure range of 100–4.6 hPa.

# 4.3 Atmospheric Infrared Sounder (AIRS)

Operating in nadir sounding geometry, the Atmospheric Infrared Sounder (AIRS) on board the Aqua satellite launched into Earth orbit in May 2002 (Chahine et al., 2006). AIRS is a medium-resolution infrared grating spectroradiometer, and a diffraction grating disperses the incoming infrared radiation into 17 linear detector arrays comprising 2378 spectral samples. The satellite crosses the Equator at approximately 01:30 and 13:30 local time, resulting in near-global coverage twice a day. AIRS has 2378 channels that cover from 649 to 1136, 1217 to 1613, and 2169 to 2674 cm<sup>-1</sup>. It also measures trace gases such as O<sub>3</sub>, CO, and to some extent CO<sub>2</sub>. AIRS CH<sub>4</sub> and N2O retrievals have been characterized and validated by Xiong et al. (2008, 2014), respectively. Both AIRS and MLS data were obtained through the Goddard Earth Sciences Data and Information Services Center (https://daac.gsfc.nasa.gov/, last access: 21 July 2020).

# 5 Comparison of FTIR with MIPAS, MLS, and AIRS observations

### 5.1 Comparison methodology

The quality of FTIR CH<sub>4</sub> and N<sub>2</sub>O for a period that covers May 2009 to March 2013 is assessed through comparison with data from MIPAS (May 2009 to December 2010), MLS (May 2009 to March 2013), and AIRS (May 2009 to March 2013) sensors on board satellites. MIPAS, MLS, and AIRS retrievals were used after averaging data obtained within coincident criteria of  $\pm 2^{\circ}$  of latitude and  $\pm 10^{\circ}$  of longitude from the ground-based FTIR site in Addis Ababa and within time difference of  $\pm 24$  h. The more stringent latitudinal criterion has proven to be a good choice for all comparisons, since latitudinal variations are, in general, more pronounced than longitudinal ones (Takele Kenea et al., 2013). These criteria yielded 29, 77, and 118 d of coincident measurements between FTIR and MIPAS, MLS, and AIRS, respectively.

The ground-based FTIR measurements of  $CH_4$  and  $N_2O$  have been validated at different locations (e.g., Senten et al., 2008). The satellite data (MIPAS, MLS, and AIRS) have a considerably better vertical resolution than ground-based

FTIR profiles due to observation geometry, spectral windows, and measurement techniques. Thus, analysis of the comparison between volume mixing ratio values derived from FTIR and MIPAS were performed for the data sets collected on May 2009 to December 2010. Furthermore, the comparison of FTIR (CH<sub>4</sub> and N<sub>2</sub>O) with MLS (CH<sub>4</sub> and N<sub>2</sub>O) and AIRS (CH<sub>4</sub>) for the time period of May 2009 to February 2013 has also been applied to assess quality of the data derived from FTIR. Hence, the profiles from MIPAS, MLS, and AIRS have been smoothed to make a comparison with FTIR as satellite observations attain better vertical resolution. Therefore, the satellite measurement profiles are smoothed using the FTIR averaging kernels of individual species obtained from the ground-based FTIR retrieval by applying the procedures reported in Rodgers and Connor (2003) and given as

$$\mathbf{x}_{\mathrm{si}} = \mathbf{x}_{\mathrm{a}} + \mathbf{A}(\mathbf{x}_{i} - \mathbf{x}_{\mathrm{a}}),\tag{3}$$

where  $x_{si}$  is the smoothed profile,  $x_a$  and A represent the a priori and averaging kernel for CH<sub>4</sub> and N<sub>2</sub>O obtained from the ground-based FTIR instrument, respectively, and  $x_i$  is the retrieved profile obtained from satellite measurements after we interpolated it to the FTIR grid spacing. We also calculate the following error statistics that can characterize the features of the instruments and parameters to be observed, such as the bias between the instruments using the difference (absolute or relative) in the daily mean profile. The difference (absolute or relative) at each altitude layers of a pair of profile is calculated using

$$\delta_i(z) = [\text{FTIR}_i(z) - \mathbf{x}_{\text{si}}(z)]. \tag{4}$$

The mean squares error can be expressed as

$$MSE_{i}(z) = \sqrt{\frac{1}{N(z) - 1} \sum_{i=1}^{N(Z)} [\delta_{i}(z)]^{2}}.$$
 (5)

The mean difference (absolute or relative) for a complete set of coincident pair profiles obtained from the ground-based FTIR and the correlative satellites is expressed as

$$\Delta_{\text{rel}}(z) = 100(\%) \times \frac{1}{N(z)} \sum_{i=1}^{N(z)} \frac{[\text{FTIR}_i(z) - \mathbf{x}_{\text{si}}(z)]}{[\text{FTIR}_i(z) + \mathbf{x}_{\text{si}}(z)]/2}, \quad (6)$$

where  $\delta_i(z)$  is the difference (absolute or relative), N(z) is the number of coincidences at z, and  $FTIR_i(z)$  is the FTIR VMR at z and the corresponding  $x_{si}(z)$  volume mixing ratio derived from satellite instruments. The standard deviation from the mean differences (absolute or relative)  $\sigma_{diff}(z)$  is important to partially characterize the measurement error. As reported in von Clarmann (2006), some use debiased standard deviation, which measures the combined precision of the instruments instead of the standard deviation of the mean

differences.

$$\sigma_{\text{diff}}(z) = \sqrt{\frac{1}{N(z) - 1} \sum_{i=1}^{N(Z)} [\delta_i(z) - \triangle_{\text{abs}}(z)]^2},$$
 (7)

where  $\delta_i(z)$  is the difference (absolute or relative) for the ith coincident pair calculated using Eq.(4). The statistical uncertainty in the mean differences (absolute or relative), which is standard error of the mean (SEM), is the quantity used to judge the statistical significance of the estimated biases, and it can be expressed in terms of the standard deviation of the mean as follows:

$$SEM(Z) = \frac{\sigma(z)}{\sqrt{N(Z)}}.$$
(8)

One can also conduct the comparison of FTIR and MI-PAS using partial columns obtained from both FTIR and smoothed MIPAS CH<sub>4</sub> and N<sub>2</sub>O. Hence, the relative difference between ground-based FTIR and smoothed MIPAS partial columns of CH<sub>4</sub> and N<sub>2</sub>O by taking into account the lower altitude limit of MIPAS observations and upper limit of ground-based FTIR sensitivity has been calculated using

$$RDiff(\%) = 100 \times \left[ \frac{(PC_{FTIR}(z) - PC_{Sat}(z))}{(PC_{FTIR}(z) + PC_{Sat}(z))/2} \right], \tag{9}$$

where PC is a partial column of FTIR and the corresponding satellite measurements. Here in this paper coincidence and smoothing errors are not taken into account in the full error analysis of the comparisons between remotely sensed data sets (von Clarmann, 2006). Hence, we focus on the random uncertainties in each instrument (Combined random error) that has been used to evaluate the comparison uncertainty (standard deviation of the difference).

# 5.2 Comparison of FTIR CH<sub>4</sub>

In Fig. 6, mean profiles, mean differences, and estimated errors versus deviations of the difference between FTIR and MIPAS\_CH4\_224 mixing ratios are shown. The comparison has been made using 29 coincident data for a time period between November 2009 and December 2010. Figure 6b indicates a negative bias of -4.8% at around 16 km and 2% at 22 km. Between 23 and 27 km, the FTIR value is higher than MIPAS values. The difference increases with altitude from 23 to 27 km (4.6%) with a maximum at 27 km. A large negative bias in FTIR CH<sub>4</sub> is obtained; i.e., FTIR CH<sub>4</sub> values are lower by 0.07 (4.8%) to 0.04 ppmv (2.2%).

Figure 6a indicates that the standard deviation of the mean differences is larger than the combined random error of the two instruments throughout the altitude range. For instance, it is twice the combined standard deviation at altitudes above 20 km and less below 20 km, which indicates the underestimation of random errors from one or both of the instruments. In addition, the overestimation of standard deviation of the

difference may result from not taking all the error budget of MIPAS into account, and the spatial and temporal criteria sets used to collect the coincidence data of MIPAS can create a discrepancy as well. The natural variations in the methane have also contributed to the overestimation of a standard deviation of the difference, as biases vary with seasons as reported in Payan et al. (2009).

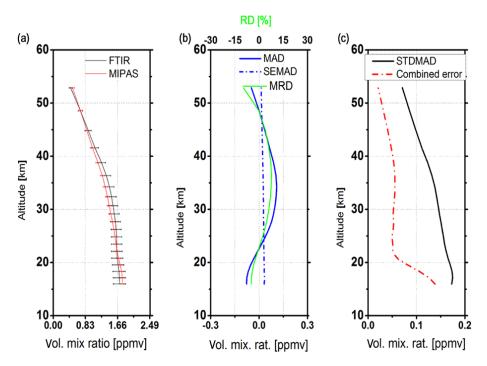
Figure 7b shows the comparison between FTIR CH<sub>4</sub> profiles and CH<sub>4</sub> derived from MLS measurements of atmospheric water vapor (H<sub>2</sub>O), carbon monoxide (CO), and nitrous oxide (N2O) and indicates that no significant bias in FTIR CH<sub>4</sub> data is present between 18 and 20 km. In the tropopause layer, the comparison indicates a negative bias of -1.7% at 17 km; i.e., the FTIR value is slightly high. FTIR CH<sub>4</sub> values are lower at altitudes between 20 and 27 km with a bias of below 11 %, which is maximum at 27 km or on average by 0.12 ppmv (6.7 %) between 20 and 27 km. The bias below 19 km and above 27 km can not be explained by the systematic errors of FTIR as the bias is larger than the systematic errors of FTIR. However, the latter, which is for altitudes above 27 km, is also out of the sensitivity ranges of FTIR. Furthermore, the standard deviation of the difference is larger than the combined random errors of the instruments. A bias in the altitude range of 20 to 27 km can be explained by the systematic error of FTIR.

In Fig. 8 mean profiles, differences, and estimated error versus deviation of the difference between FTIR and AIRS mixing ratios are shown. The largest negative bias is found at altitudes between 11 and 19 km, with a maximum difference of -0.08 ppmv at around 15 km. A negative bias of the AIRS mixing ratio of CH<sub>4</sub> is higher than that the FTIR as shown in Fig. 8. A positive bias existed at altitudes between 7 and 9 km, and similarly, it also has been shown at altitudes between 21 and 27 km, with a maximum value at around 27 km, and its bias is 0.14 ppmv (9 %). The standard deviation of the difference agrees with the combined random error at altitudes below 20 km, and it overestimates above 20 km.

In all the comparisons of FTIR CH<sub>4</sub> with data from MI-PAS, MLS, and AIRS, sensors on board satellites indicate a negative bias below 21 km and a positive bias above 21 km, with similar bias of not higher than 5.8% in the altitude range of 21-27 km (see Table 2.). The volume mixing ratios derived from the satellite are higher at altitudes lower than 21 km.

# 5.3 Comparison of FTIR N<sub>2</sub>O

FTIR  $N_2O$  mixing ratio MIPAS comparison results are shown in Fig. 9, where it represents the mean profiles, mean absolute difference, and standard deviation of the mean along with the combined errors of the two instruments. Mean profiles of FTIR show a maximum at around 23 km and decrease smoothly as altitude increases, and that of the MI-PAS\_N2O\_224 value starts to decline starting from the lowermost stratosphere.



**Figure 6.** Comparison of CH<sub>4</sub> from MIPAS reduced resolution (V5R\_CH4\_224) and FTIR. (a) Mean profiles of MIPAS (red) and FTIR (black) and their standard deviation (horizontal bars). (b) Mean difference FTIR minus MIPAS (MAD, solid blue), standard error of the difference (SEMAD; dotted blue), and mean relative differences FTIR minus MIPAS relative to their average (MRD; green, upper axis). (c) Combined mean estimated statistical error of the difference (combined error, dotted red, contains MIPAS instrument noise error, and FTIR random error budget) and standard deviation of the difference (STDMAD; solid black).

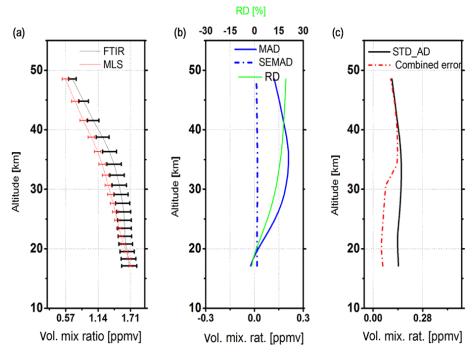


Figure 7. Comparison of CH<sub>4</sub> from MLS (V3.3) and FTIR. Details as in Fig. 11.

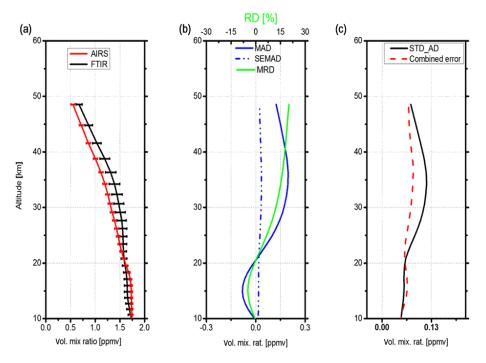


Figure 8. Comparison of CH<sub>4</sub> from AIRS and FTIR. Details as in Fig. 11.

**Table 2.** Averaged statistical means (M) and standard deviations (SD) of the relative differences  $100 \times \left\lfloor \frac{\text{FTIR-MIPAS}}{\text{FTIR+MIPAS}} \right\rfloor$  (%) defined in the altitude range of 17–20 and 21–27 km. The numbers of coincidences (N) within a spatiotemporal criterion of  $\pm 2^{\circ}$  of latitude and  $\pm 10^{\circ}$  of longitude and time difference of  $\pm 24$  h are selected for intercomparison. This is for FTIR CH<sub>4</sub> and N<sub>2</sub>O with the corresponding other instruments (stated in second column).

Gas	Instrument	Altitude range	$M\pm$ SD	Period	N
CH <sub>4</sub>	MIPAS MLS	17–20/21–27 17–19/20–27	$-4.8/4.2 \pm 5.2/5.5$ $-1.8/5.8 \pm 8/8.8$	May 2009–Dec 2010 Jun 2009–Feb 2013	29 77
	AIRS	17–20/21–27	$-2.8/5.3 \pm 3.5/5.4$	Jun 2009–Feb 2013	118

Comparison of FTIR N<sub>2</sub>O profiles **MIPAS** to (V5R N2O 224) measurements (see Fig. 9b) indicates that FTIR value is higher than the MIPAS above 20 km, and the maximum mean absolute difference in N<sub>2</sub>O is 15 % (0.04 ppmv) at around 24 km, while the FTIR value is less at altitudes below 20 km, with a maximum difference of -7%(-0.02 ppmv) at around 17 km. The bias at 19 km is not statistically significant, as the standard error of the mean is larger than the bias. In the remaining altitudes, standard error of the mean is smaller than the mean bias, and the biases are statistically significant. Since, the bias at altitudes between 20 and 27 km is smaller than the FTIR systematic errors, the bias could be explained in terms of systematic uncertainties in FTIR (see Fig. 4b, bottom panel). The standard deviation of the difference is larger than the combined error of the two instruments at altitudes above 20 km (see Fig. 9c), and the standard deviation of the difference agrees with the estimated combined random error in the altitude ranges

between 20 and 27 km. For the altitudes below 20 km, the estimated combined random error is overestimated.

Figure 10a represents the mean profiles of  $N_2O$  derived from the coincident pairs of FTIR and MLS. Throughout the whole altitude range, the value derived from FTIR is overestimated (relative to MLS). The FTIR values of  $N_2O$  are larger than the MLS value of  $N_2O$  by a factor of 1.2 and 1.1 at around 21 and 27 km, respectively. The mean relative difference in FTIR and MLS  $N_2O$  value increases as altitude increases; its value is less than 18.6 % at altitudes below 27 km, and its bias below 22 km is less than 8 %, which can be explained in terms of the systematic error of FTIR  $N_2O$ . Thus, the positive bias is statistically significant as the mean difference in the comparison is larger than the standard error of the mean.

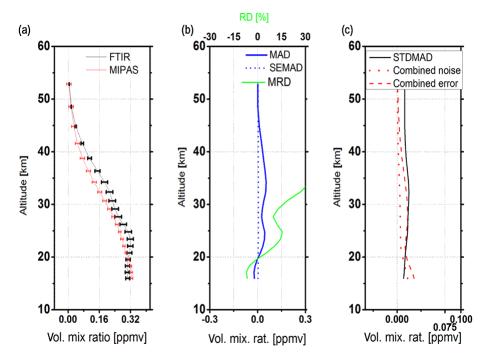


Figure 9. Comparison of N<sub>2</sub>O from MIPAS (V5R\_N2O\_224) and FTIR. Details as in Fig. 11.

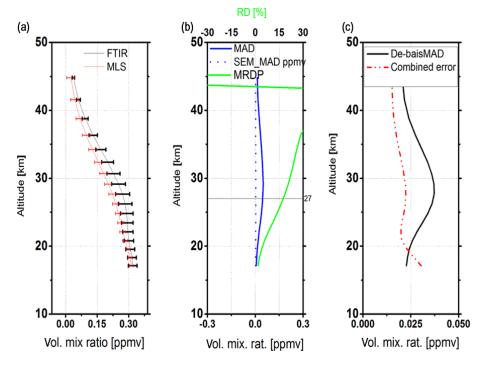
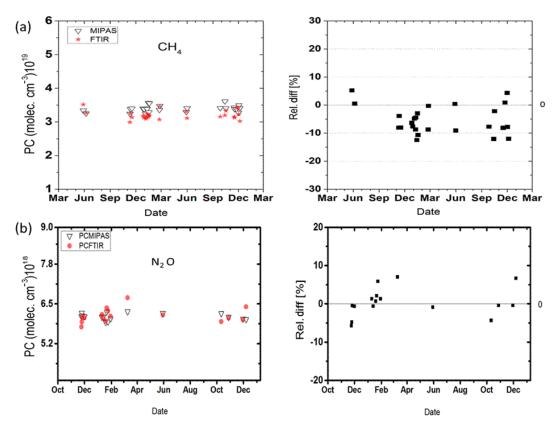


Figure 10. Comparison of N<sub>2</sub>O from MLS (V3.3) and FTIR. Details as in Fig. 11.



**Figure 11.** Time series of **(a)**  $CH_4$  and **(b)**  $N_2O$  partial-column comparisons for the altitude range of 15–27 km. Right panels: ground-based FTIR (stars) and MIPAS (V5R\_CH4\_224 and V5R\_N2O\_224) (inverted triangles) partial columns. Left panels: relative differences between ground-based FTIR and MIPAS (V5R\_CH4\_224 and V5R\_N2O\_224) partial columns.

# 5.4 Comparisons of partial columns

For the partial-column (PC) comparisons of FTIR with MIPAS, it is vital to take into account the lower altitude limit of MIPAS, which is 15 km for both target gases. The ground-based FTIR sensitivity is used to determine the upper altitude limit, which is reasonable up to  $\sim 27\,\mathrm{km}$  for CH<sub>4</sub> and N<sub>2</sub>O in the tropical atmospheric condition. Therefore, the PC that we use in the comparison is limited to the altitude range of 15–27 km. The DOFs within these partial columns' limit are about 1.0 and 1.2 for CH<sub>4</sub> and N<sub>2</sub>O, respectively.

Figure 11 shows the time series of the partial columns and relative differences of  $CH_4$  (a) and  $N_2O$  (b). The partial-column comparison of  $CH_4$  between values of FTIR and MI-PAS revealed a mean error of  $-5.5\,\%$ , mean squares error of  $7.4\,\%$ , and standard deviation from the mean error of  $5\,\%$ . Similarly,  $N_2O$  values between FTIR and MIPAS revealed a mean error of  $0.5\,\%$ , mean square error of  $3.7\,\%$ , and standard deviation from mean error of  $3.8\,\%$ ; in the latter case a significant positive bias is observed, and in  $CH_4$  negative bias was obtained.

# 6 Summary and conclusions

The vertical profiles and partial columns of CH<sub>4</sub> and N<sub>2</sub>O over Addis Ababa, Ethiopia, were derived from ground-based FTIR. The mean partial column of CH<sub>4</sub> and N<sub>2</sub>O within the sensitivity ranges of the instrument, which is from the surface to around 27 km, is determined as  $2.85 \times 10^{19}$  molecules cm<sup>-2</sup>  $\pm 5.3\%$  and  $5.16 \times 10^{18}$  molecules cm<sup>-2</sup>  $\pm 6.95\%$ , respectively. Furthermore, the overall contribution of both statistical and systematic errors, i.e., a total error of CH<sub>4</sub> and N<sub>2</sub>O from ground-based FTIR, is 3.1% and 3%, respectively.

From comparison of FTIR CH<sub>4</sub> and MIPAS\_CH4\_224 products, a statistically significant maximum negative bias of -4.8% at an altitude of 15 km that extends to 21 km and maximum positive bias of 4.6% at an altitude 27 km were obtained. The largest negative bias is found at altitudes between 11 and 19 km, with a maximum difference of -0.08 ppmv (-4.8%) at around 15 km, and a positive bias of less than 0.14 ppmv (9%) is found at altitudes between 21 and 27 km, with a maximum value at around 27 km in the FTIR CH<sub>4</sub> comparison with AIRS. On the other hand, from a comparison of CH<sub>4</sub> from ground-based FTIR and MLS version 3.3, we obtained a significant positive average bias of 0.12 ppmv

 $(6.7\,\%)$  in the altitude range of 20– $27\,\mathrm{km}$  and a negative bias of  $-1.7\,\%$  at  $17\,\mathrm{km}$ . In the case of FTIR  $N_2O$  and MI-PAS\_N2O\_224, a significant positive bias of less than  $15\,\%$  in the altitude range of 22– $27\,\mathrm{km}$  with a maximum value at around  $25\,\mathrm{km}$  and a negative bias of  $-7\,\%$  at  $17\,\mathrm{km}$  has been obtained. A positive bias of less than  $18.6\,\%$  for altitudes below  $27\,\mathrm{km}$  is noted for  $N_2O$  between FTIR and MLS, and its bias below  $22\,\mathrm{km}$  is less than  $8\,\%$ , which can be explained in terms of the systematic error of FTIR  $N_2O$ .

In general, the retrieved CH<sub>4</sub> and N<sub>2</sub>O VMR and column amounts from Addis Ababa, tropical site, exhibited very good agreement with all coincident satellite observations in the altitude ranges of 17-27 km with a positive mean relative difference within 20-27 km and negative below 20 km. In addition, the bias obtained from the comparison and precision of the FTIR measurements is also comparable. The intercomparisons of CH<sub>4</sub> and N<sub>2</sub>O VMR from ground-based FTIR with data from MIPAS, MLS, and AIRS sensors on board satellites reported in this work establish main features that characterize the FTIR instruments at Addis Ababa. The FTIR data can be used in further scientific studies as it represents a unique environment of tropical Africa, a region poorly investigated in the past. Furthermore, the results of this intercomparison for FTIR observations with the satellites can ensure that FTIR can now be used to validate satellite missions. Thus, for the FTIR data, it is anticipated that the use of the data in further scientific studies may provide some insight into the processes that govern chemical transport and chemistry in the atmosphere as well as sources of green gases in this part of the globe.

Data availability. The FTIR Data at Addis Ababa is available on request from the authors. All other data sets are publicly available at archives indicated in the paper.

*Supplement.* The supplement related to this article is available online at: https://doi.org/10.5194/amt-13-4079-2020-supplement.

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