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# Establishment of AIRS climate-level radiometric stability using radiance anomaly retrievals of minor gases and sea surface temperature

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Abstract. Temperature, H<sub>2</sub>O, and O<sub>3</sub> profiles, as well as CO<sub>2</sub>, N<sub>2</sub>O, CH<sub>4</sub>, chlorofluorocarbon-12 (CFC-12), and sea surface temperature (SST) scalar anomalies are computed using a clear subset of AIRS observations over ocean for the first 16 years of NASA's Earth-Observing Satellite (EOS) Aqua Atmospheric Infrared Sounder (AIRS) operation. The AIRS Level-1c radiances are averaged over 16d and 40 equal-area zonal bins and then converted to brightness temperature anomalies. Geophysical anomalies are retrieved from the brightness temperature anomalies using a relatively standard optimal estimation approach. The CO<sub>2</sub>, N<sub>2</sub>O, CH<sub>4</sub>, and CFC-12 anomalies are derived by applying a vertically uniform multiplicative shift to each gas in order to obtain an estimate for the gas mixing ratio. The minor-gas anomalies are compared to the National Oceanic and Atmospheric Administration (NOAA) Earth System Research Laboratory (ESRL) in situ values and used to estimate the radiometric stability of the AIRS radiances. Similarly, the retrieved SST anomalies are compared to the SST values used in the ERA-Interim reanalysis and to NOAA's Optimum Interpolation SST (OISST) product. These intercomparisons strongly suggest that many AIRS channels are stable to better than 0.02 to 0.03 K per decade, well below climate trend levels, indicating that the AIRS blackbody is not drifting. However, detailed examination of the anomaly retrieval residuals (observed - computed) shows various small unphysical shifts that correspond to AIRS hardware events (shutdowns, etc.). Some examples are given highlighting how the AIRS radiance stability could be improved, especially for channels sensitive to N<sub>2</sub>O and CH<sub>4</sub>. The AIRS shortwave channels exhibit larger drifts that make them unsuitable for climate trending, and they are avoided in this work. The AIRS Level 2 surface temperature retrievals only use shortwave channels. We summarize how these shortwave drifts impacts recently published comparisons of AIRS surface temperature trends to other surface climatologies.

# 1 Introduction

The Atmospheric Infrared Sounder (AIRS) on NASA's Aqua satellite platform (Aumann et al., 2003) measures 2378 high-spectral-resolution infrared radiances between 650 and 2665 cm<sup>-1</sup> with a resolving power ( $\lambda/\Delta\lambda$ ) of ~ 1200. Launched in 2002 into a Sun-synchronous polar orbit with a 13:30 UTC ascending node Equator crossing time, AIRS now has been operating almost continuously for 17+ years.

The long record of AIRS allows measurements of shortterm climate trends that are especially useful given its global coverage. Nominal decadal climate temperature trends are in the 0.1–0.2 K per decade range. For example, a recent Intergovernmental Panel on Climate Change (IPCC) report (Masson-Delmotte et al., 2018) suggests 20th century surface temperature trends (2000–2017) of about 0.17 K per decade. If AIRS is to contribute to climate-level trend measurements, uncertainty estimates for the time stability of the AIRS radiances are a prerequisite before using AIRS Level 2/3 products for climate-level trending. Estimating the level of any instrument-related trends, for a wide range of AIRS channels, is the subject of this work.

A recent study (Aumann et al., 2019) addressed the stability of a single AIRS channel by comparisons to sea surface temperatures (SSTs). Some limitations of that study are addressed below, but its major limitation is that it evaluates only one channel. AIRS retrievals use 400+ AIRS channels, and there is no guarantee that the AIRS stability in one channel applies to all channels, as acknowledged in Aumann et al. (2019).

AIRS is sensitive to a host of atmospheric and surface variables, including atmospheric temperature (via CO<sub>2</sub> emissions), humidity, surface temperature, O3, CH4, N2O, carbon monoxide, clouds, coarse-mode aerosols, and other minor gases. 1D-Var retrievals such as the AIRS Level 2 products (Susskind et al., 2014) attempt to retrieve all relevant atmospheric and surface variables in order to produce the most accurate temperature and H<sub>2</sub>O profiles. The atmospheric CO<sub>2</sub> concentration is especially important for AIRS retrievals since most of the radiance measured in the temperature sounding channels is due to CO<sub>2</sub> emission. However, it is difficult to separate the CO<sub>2</sub> concentration from variations in the temperature profile due to co-linearity of their Jacobians. Consequently, the AIRS Level 2 retrievals instead vary CO<sub>2</sub> in the forward model to account for CO<sub>2</sub> growth during the mission (John M. Blaisdell, personal communication, 2019).

The largest radiance trends seen by AIRS are due to the growth rate of  $CO_2$  in the atmosphere. Assuming a nominal growth rate of 2 ppm yr<sup>-1</sup>, and max sensitivity of AIRS channels of  $CO_2$  of  $0.03 \text{ K ppm}^{-1}$ , the brightness temperature (BT) shift in AIRS over 16 years is  $\sim 1 \text{ K}$ , or  $0.06 \text{ K yr}^{-1}$ . Concentrations of atmospheric carbon dioxide have been measured worldwide for many years with extremely high accuracy (Masarie and Tans, 1995; Tans and Keeling, 2019) by the National Oceanic and Atmospheric Administration (NOAA) Earth System Research Laboratory (ESRL). Averaged yearly, CO<sub>2</sub> concentrations are highly uniform globally, with little latitudinal variation in growth rates. Similarly, NOAA ESRL also provides a wide network of measurements of N<sub>2</sub>O and CH<sub>4</sub>, which are also relatively uniformly mixed over yearly time periods. Here, we use the high accuracy of the trends in these in situ measurements of minor gases to determine the stability of a large number of AIRS channels.

SST trends are also extremely well measured and generally referenced to the in situ Argo (Argo, 2019) buoy network but interpolated to a full grid using instruments such as the Advanced Very High Resolution Radiometer (AVHRR). Two SST products referenced to the buoy network are compared to AIRS trends here: (1) NOAA's Optimum Interpolation SST (version 2) (OISST) (Banzon et al., 2016), and (2) the Operational Sea Surface Temperature and Ice Analysis (OSTIA) (Stark et al., 2007), which has been used in the ERA-Interim reanalysis (ERA-I) since 2009 (Dee et al., 2011). Prior to February 2009, ERA-I used the National Centers for Environmental Prediction (NCEP) Real-Time Global (RTG) SST product, a precursor to OISST.

AIRS stability is referenced to trends in these minor gases and SST by performing 1D-Var retrievals of clear scene radiance anomalies averaged into 40 equal-area latitude bins and 16 d time periods. Comparisons of the retrieved gas concentrations and SST trends, combined with examination of the retrieval residuals, provides a number of powerful tests of AIRS radiometric stability as well as detailed information on AIRS performance changes due to several minor instrument shutdowns that took place occasionally over the mission.

After summarizing the characteristics of the AIRS instrument, and the data used in this work, the retrieval methodology is reviewed with a short discussion of the retrieved temperature profile time series. We follow with stability estimates derived from the anomaly spectra retrievals of  $CO_2$ , N<sub>2</sub>O, CH<sub>4</sub>, and SST. Although AIRS is most sensitive to the two best in situ data sets, CO<sub>2</sub> and SST, we also compare to retrievals of N<sub>2</sub>O and CH<sub>4</sub> since they are also relatively well measured and help test the AIRS performance in spectral regions not covered by CO<sub>2</sub> and SST. Finally we examine the time series of the anomaly retrieval residuals (BT observed – fit) time series since, together with the anomaly geophysical retrievals, they provide detailed information on AIRS radiances over time, especially the instrument response to various short shutdowns that occurred during the mission.

# 2 AIRS instrument and data

Several details of the AIRS instrument design are relevant to the processing performed here and are needed to understand some of the results. AIRS has 2378 spectral channels divided up into 17 different detector arrays. Appendix A gives the nominal wavenumber boundaries of these arrays. Arrays M-11 and M-12 are linear arrays of single photoconductive mercury-cadmium-telluride (HgCdTe) detectors. The other AIRS arrays are photovoltaic detectors, and each reported detector output is actually some linear combination of two detectors offset from each other in the vertical (not dispersive) direction. The photovoltaic detectors for each AIRS channel are labeled "A" and "B". The relative contributions of A and B detectors can be changed by command to the spacecraft. The majority of these detectors are wired for equal contributions by the A and B detectors, which we denote as A+B detectors. However, some detectors have always been inoperable, or their performance characteristics changed in orbit, so there are a number of A-only and B-only detectors.

The radiometric and spectral characteristics of the A versus B detectors can be slightly different. During the mission, good A+B detectors can suddenly exhibit greatly increased noise when one or the other of the two detectors fails or degrades. In many circumstances, the AIRS project has changed A+B detectors to be either A-only or B-only in order to recover that particular channel, albeit at slightly lower noise levels than if both detectors were working properly. Fortunately, many of the A-only and B-only detectors are in the window regions where AIRS has tremendous redundancy. Unfortunately, the M-10 array which covers the

tropospheric  $CO_2$  sounding channels also has a good number of A-only, B-only detectors.

Here, we avoid any photovoltaic channel that is not A+B, and any channel with a state change during the mission. Although A-only and B-only channels may perform well, many of these single detector channels exhibit drifts over the mission for colder scenes. This is especially apparent in time series of cold scene observations (deep convective clouds) by comparison to similar time series derived from the Infrared Atmospheric Sounding Interferometer (IASI) on MetOp-A. In addition, we avoid any channels with detector noise above 0.5 K noise-equivalent brightness temperature (NEDT) (for a 250 K scene). As discussed below in more detail, we also avoid all shortwave AIRS channels, meaning channels past  $2000 \text{ cm}^{-1}$  for our final trend measurements, since we find that the shortwave is drifting slightly.

#### 3 Radiance/brightness temperature anomalies

## 3.1 Clear selection

The new AIRS Level-1c (L1c) radiance product (Aumann et al., 2020) is used in this work rather than the standard L1b product. The L1c product provides single-footprint radiance estimates for channels in L1b that are not functional or are extremely noisy. Even high-quality L1b channels can sometimes "pop" or experience radiation hits that invalidate the measurement. In these extremely rare cases, the L1c algorithm substitutes an estimated radiance using a principalcomponent approach. These corrections are rare enough that they have no effect on the long-term trends under study in this work. L1c also includes some channels (between detector arrays) that do not exist.

More importantly for this work, the radiances in L1c have been corrected for small drifts in the channel center frequencies. These drifts are small but are large enough to have some minor impact on radiance trends. We emphasize that the channels selected for the anomaly retrievals are all valid L1b channels, and most have undergone no corrections other than adjusting the radiances back to a fixed frequency scale.

AIRS L1c clear scenes are primarily detected using a uniformity filter. (Throughout this paper, the term "scene" refers to a single AIRS nominal  $12 \times 12$  km footprint or field of view.) The BT of each AIRS ocean scene is subtracted from the BT of each of its eight neighbors for two window channels at 819.3 and 961.1 cm<sup>-1</sup>. A scene is initially deemed clear only if the absolute value of all of these differences, averaged over the two channels, is less than 0.4 K. The selected scenes are matched to ERA-I model fields and a simulated clear BT for the 961.1 cm<sup>-1</sup> channel is computed using a stand-alone version of the AIRS radiative transfer algorithm (Strow et al., 2003) called SARTA (StandAlone Rapid Transmittance Algorithm), implemented using highresolution transmission molecular absorption database (HI- TRAN) 2008 line parameters. If the difference between the observed and computed clear scene BT values is more than  $\pm 4$  K, the scene is discarded from the clear list. This test mostly removes colder scenes made up of very uniform marine boundary layer stratus clouds. The clear yield and mean zonal radiances are quite insensitive to the exact value of this threshold. The uniformity test is not performed on the first and last of the 135 along-track scans in each AIRS granule since they do not have eight neighbors and we wanted to avoid cross-granule processing. The total number of clear scenes is limited to  $\sim$  40000 daily clear scenes by randomly subsetting the detected clear scenes; however, this daily limit is almost never reached. In this work, we only use descending node observations in order to avoid solar and non-LTE contributions to the AIRS radiances in the shortwave. After subsetting for descending (ocean) only the total number of clear scenes detected is  $\sim 10000 \, \text{d}^{-1}$ .

The 4 K (observed – computed) BT test removes ~20% of the scenes detected with the uniformity filter. A map of these deleted scenes very clearly shows that they are almost all located along the west coasts of the Americas and Africa, where marine boundary layer stratus clouds commonly occur. The (observed – computed) BT values for the 1231 cm<sup>-1</sup> window channel have a nearly Gaussian distribution with a width of ~ 0.6 K. Note that this distribution of biases is well separated from the 4 K cutoff used to remove marine boundary layer stratus clouds.

Another important characteristic of this clear subset is the stability of the observing times. If the mean observing time changes during this 16-year time period, trends in the SST could be confused with the diurnal cycle of the SST. Due to the high stability of the Aqua orbit, this is not an issue. The short-term day-to-day variations in the mean clear subset times can vary by several hours. In addition, there is a seasonal variation of several hours in the clear subset. But these variation are extremely stable, and the total linear drift of the clear subset over the 16-year observing period, for any given latitude bin in the tropics, is  $\sim 20 \pm 40$  s ( $2\sigma$ ) per year, effectively zero.

All observing parameters, on a footprint basis, are saved, such as satellite-viewing zenith angle and noise (converted to BT units). In addition, the ERA-I model parameters (temperature,  $H_2O$ , and  $O_3$  profiles, and surface temperature with a spatial resolution of approximately 80 km on 60 levels in the vertical from the surface up to 0.1 hPa) are matched to each clear scene and saved along with their associated simulated L1c radiances. This allows our processing to use simulated rather than observed radiances for testing. The ERA-I profiles are also used to compute the anomaly Jacobians used in the retrievals and are discussed in detail in Sects. 4.3 and 5.

#### 3.2 Clear scene characteristics

Figure 1 illustrates the density and location of the clear ocean data set, averaged over 2012. Retrievals are only performed



Figure 1. Density of AIRS clear ocean scenes for the calendar year of 2012.

on zonally averaged data, which translate into  $\sim 44/25$  observations per day at  $-50/+50^{\circ}$  latitude, respectively, with a maximum of 200 observations per day at  $-0.5^{\circ}$  latitude. The non-uniform nature of this sampling should be kept in mind when examining temperature, H<sub>2</sub>O, or O<sub>3</sub> trends in that this data set is not necessarily representative of global/zonal climate trends. However, we do assume that the minor-gas anomaly trends we are retrieving are uniformly mixed over multi-year timescales. Our anomaly retrieval results show uniform mixing is generally quite accurate over even 16 d timescales.

Figure 2 illustrates the accuracy of ERA-I for this data set by plotting the (observed – ERA-I) BT bias for 28.4° N. The ERA-I simulated BT used our SARTA radiative transfer algorithm (RTA), which has a default value of 385 ppm for  $CO_2$ . This  $CO_2$  value is matched in the observations by comparing to AIRS observations for the time period centered around June 2008 when the nominal global CO<sub>2</sub> amount was 385 ppm. The window regions  $(800-1000 \text{ cm}^{-1})$  exhibit a bias of  $\sim -0.5$  K, which is quite small and likely some combination of instrument bias, evaporative cooling of the ocean surface relative to the ERA-I SST, incorrect ERA-I water vapor column affecting the H<sub>2</sub>O continuum, and some cloud contamination. Sampling errors may contribute to the larger biases in the water region beyond  $1300 \,\mathrm{cm}^{-1}$ . A zoom of the bias in Fig. 2b highlights the low bias in the  $700-750 \,\mathrm{cm}^{-1}$ region which is sensitive to tropospheric  $CO_2$ , with a mean of  $\sim 0.2$ –0.3 K.

The single-footprint standard deviation of the bias is also shown in Fig. 2b along with the average AIRS NEDT for these footprints. The ERA-I bias standard deviation is barely larger than the AIRS noise in this spectral region, indicating that ERA-I temperatures in the mid-troposphere track the AIRS observations very closely with a standard deviation considerably smaller than the AIRS noise. This makes a



**Figure 2.** (a) (AIRS – ERA-I simulated) BT bias for  $28.4^{\circ}$  N for a time period centered around June 2008 when the global CO<sub>2</sub> amount was ~ 385 ppm. (b) Zoom of panel (a), showing that the region near 700–760 cm<sup>-1</sup>, which is most sensitive to CO<sub>2</sub>, has a mean bias of ~ 0.2–0.3 K and a single-footprint standard deviation of ~ 0.3 K. Also shown is the AIRS NEDT, which is barely smaller than the bias standard deviation.

strong case for the accuracy of the BT Jacobians computed from ERA-I temperature fields.

Figure 3 shows the linear trend for the clear data set averaged over  $\pm 50^{\circ}$  latitude. These BT trends prominently exhibit the growth in  $CO_2$  in the tropospheric channels from 700 to  $750 \,\mathrm{cm}^{-1}$ , which results in a negative change in the observed BT since increasing CO2 shifts the emission to higher and therefore colder regions of the atmosphere. The growth in CO<sub>2</sub> in the stratospheric channels (a positive BT change) below  $700 \,\mathrm{cm}^{-1}$  is roughly canceled by cooling in the stratosphere. All window channels exhibit warming, with larger values in the shortwave past  $2450 \,\mathrm{cm}^{-1}$ . Spectral regions in Fig. 3 that exhibit trends smaller than the  $2\sigma$  uncertainty are often channels where the BT trends that are due to increasing CO<sub>2</sub>, CH<sub>4</sub>, and N<sub>2</sub>O are counterbalanced by changes of the opposite sign due to trends in either the atmospheric or surface temperature. These counterbalanced trends are all properly accounted for in the anomaly retrievals given the good agreement between the observed and in situ minorgas trends. The non-uniform spatial sampling of these clear scenes precludes any general statements about climate warming, although for these observations we clearly see surface warming in the  $800-1250 \text{ cm}^{-1}$  region, if the AIRS radiometry is stable. In addition, the effects of much stronger water vapor absorption in the longwave compared to the shortwave windows make definitive intercomparisons of the BT trends complicated, which is addressed below by doing retrievals on these data.



**Figure 3.** Mean BT trends ( $a_1$  in Eq. 1) averaged over  $\pm 50^{\circ}$  in  $\Delta$ BT yr<sup>-1</sup> units. The  $2\sigma$  uncertainty shown has been corrected for serial correlations in the BT time series. Channels used in the anomaly retrievals are denoted in red, and the BT trend uncertainty is in yellow.

# 3.3 Construction of anomalies

The clear scene radiance subset is sorted into 40 equivalentarea latitude bins that cover the full -90 to  $90^{\circ}$  latitude range and are averaged over every 16 d. This results in a data set for the first 16 years of AIRS that has a size of  $(40 \times 2645 \times 365)$ latitude bins, AIRS L1c channels, and the total number of 16 d averages. The following time series function was fit to these averaged radiances,  $r_{obs}(t)$ , for each latitude and AIRS L1c channel:

$$r_{\rm fit}(t) = r_o + a_1 t + \sum_{i=1}^4 c_i \sin(2\pi n t + \phi_i), \tag{1}$$

where t is AIRS mission times in years,  $r_o$  is a constant, the  $c_i$  values are the amplitudes of the season cycle and three harmonics, and the  $\phi_i$  is their associated phases. At 28° N, for example, the annual amplitude relative to the mean radiance,  $c_1/r_o$ , has a median value (taken over channel) of 4.2%. The median amplitudes of the three harmonics terms,  $c_2$ ,  $c_3$ , and  $c_4$ , relative to  $r_o$ , are 0.32%, 0.45%, and 0.23%, respectively, all with  $2\sigma$  uncertainties of  $\sim 0.05\%$ . The linear trends  $a_1$  are included in the anomaly time series fits for simple diagnostic purposes and are not used directly in the anomaly retrievals.

The radiance anomalies,  $r_a(t)$  are formed by removing the constant  $r_o$ , and the sinusoidal terms in Eq. (1), from the observed radiance time series  $r_{obs}(t)$ . This can be expressed as

$$r_{\rm a}(t) = r_{\rm obs} - \left(r_o + \sum_{i=1}^{4} c_i \sin(2\pi nt + \phi_i)\right).$$
(2)

The radiance anomalies  $r_a(t)$  were converted to brightness temperature units using



**Figure 4.** Mean and standard deviation of the AIRS BT anomalies for the zonal bin centered at 28.3° N.

$$\mathbf{y}(v) \equiv \mathrm{BT}_{\mathrm{a}}(v,t) = \frac{r_{\mathrm{a}}(v,t)}{\frac{\partial r(v)}{\partial \mathrm{BT}(v)}}.$$
(3)

The  $40 \times 2645 \times 365$  array of BT<sub>a</sub> vectors are the retrieval inputs **y** in the retrieval formulation discussed in Sect. 4.

The anomaly BT time series mean BT spectra and their standard deviations are shown in Fig. 4 for the 28.4° N latitude bin. The BT anomaly is set to zero at the mission start; therefore, the mean BT in the channels sensitive to tropospheric CO<sub>2</sub> between 700 and 750 cm<sup>-1</sup> is -0.5 K, which then increases by  $\sim 1$  K during the mission. The standard deviation indicates that the SST (and H<sub>2</sub>O continuum) vary by  $\sim 0.4$  K during this time period (window region channels from 800 to 1000 cm<sup>-1</sup>). Some of this is likely due to changes in sampling from day to day and ERA-I errors in SST and column H<sub>2</sub>O. Upper-tropospheric water vapor, which dominates the spectral region between 1350 and 1615 cm<sup>-1</sup>, has the highest variability, which is expected due to both the high temporal variability of water vapor and our non-uniform sampling.

An example radiance BT anomaly for the 710.141 cm<sup>-1</sup> channel is shown in Fig. 5, for the same latitude bin. This channel is heavily influenced by the CO<sub>2</sub> growth, so the AIRS observed trends are becoming more negative, although there is considerable noise, again due to weather and sampling. For comparison, we also plot the ERA-I simulated BT anomaly, which does not contain the CO<sub>2</sub> growth, since it is set to a fixed value of 385 ppm in the simulations. The difference between these two BT anomalies will primarily be due to CO<sub>2</sub> growth and is shown in black. Note that since the ERA-I tracks the atmospheric state quite accurately and most of the time series "noise" is removed. This helps lend credence to our use of the ERA-I model fields for Jacobian evaluation.



**Figure 5.** Sample AIRS observed and ERA-I simulated BT anomalies for the zonal bin centered at  $28.3^{\circ}$  N for the AIRS channel centered at 710.14 cm<sup>-1</sup>. The differences in the AIRS and ERA-I anomalies are plotted in black. Note that this difference anomaly is not used in the anomaly retrievals.

## 4 Retrieval methodology

#### 4.1 Approach

Geophysical retrievals are derived from the BT spectral anomalies y(v), defined in Eq. (3). Using standard retrieval notation, the atmospheric state x is derived from the observations y by minimizing the cost function J:

$$\mathbf{J} = (\mathbf{y} - F(\mathbf{x}))^T \mathbf{S}_{\epsilon}^{-1} (\mathbf{y} - F(\mathbf{x})) + (\mathbf{x} - \mathbf{x}_{a})^T \mathbf{R} (\mathbf{x} - \mathbf{x}_{a}), \quad (4)$$

where  $S_{\epsilon}$  is a diagonal observation error covariance matrix containing the square of the BT noise, **K** is the anomaly Jacobians, and **R** is a regularization matrix. The retrieved atmospheric state *x* (the geophysical anomalies) is given by

$$\boldsymbol{x} = \boldsymbol{x}_{a} + (\mathbf{K}^{T} \mathbf{S}_{\epsilon}^{-1} \mathbf{K} + \mathbf{R})^{-1} (\mathbf{K}^{T} \mathbf{S}_{\epsilon}^{-1} (\boldsymbol{y} - F(\boldsymbol{x}_{n})),$$
(5)

where

$$\mathbf{R} = \mathbf{S}_{a}^{-1} + \alpha \mathbf{L}^{\mathrm{T}} \mathbf{L}$$

 $S_a$  is the a priori covariance matrix, and  $\alpha L$  is an empirical regularization constraint using Tikhonov L1-type derivative smoothing. This retrieval approach is standard optimal estimation (OE) (Rodgers, 1976) enhanced to include both covariance and empirical Tikhonov regularization in **R** (Steck, 2002). Forward-model uncertainty is not included in the measurement error covariance. The mathematical approach is very similar to the author's single-footprint AIRS retrieval algorithm (DeSouza-Machado et al., 2018).

A priori estimates for  $x_a(t) \equiv 0$  for T(z),  $O_3(z)$ ,  $H_2O(z)$ , and  $T_{SST}$  were set to zero. Two approaches were used for the minor-gas a priori estimates. The first approach set  $x_a(t) =$  $x_a(t-1)$  where  $x_a(t=0) = 0$  for the minor gases, iteratively

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increasing the a priori gas amount in time based on the previous 16 d retrieval.

Another approach used the known growth rates in the minor gases (from ESRL) by setting  $x_a(t) = g \times (t - t_o)$ for the a priori minor-gas amount, where g is the nominal yearly growth rate for each gas from the NOAA ESRL atmospheric gas trends. For both approaches, we set the a prior icovariance to  $g \times 1$  year, the yearly variation in that gas. Nearly identical results are obtained if we increase the a prior icovariance to as much as  $g \times 5$  years. The iterative approach for setting the minor gas a priori produces noisier retrieval anomalies. However, if our retrievals are averaged over  $\pm 50^{\circ}$  latitude, both approaches produced identical differences compared to in situ measurements, including error uncertainties. The figures and trend results shown here use the a priori ramp from the ESRL data, although the figures for the iterative ramp are only distinguishable from what is shown for single zonal retrievals (such as the Mauna Loa and Cape Grim comparisons).

The temperature,  $H_2O$ , and  $O_3$  profile retrievals use 20 atmospheric layers, selected from the AIRS standard 100-layer pressure grid (Strow et al., 2003) by accumulating five of the standard AIRS layers at a time. The lowest layer is about 1.5 km thick, with increasingly wider layers as you go higher in the atmosphere. This layering scheme allows more layers than degrees of freedom (DOFs) although it does limit retrievals in the upper stratosphere. We wish to minimize our sensitivity to the upper stratosphere since our comparisons to in situ measurements are made in the troposphere. Consequently, we removed all channels peaking above 10 hPa.

Most of the regularization in the retrieval comes from the Tikhonov terms, since we do not want to invoke climatology too strongly for a climate-level measurement. Appendix B discusses the profile retrievals, and simulations of these retrievals, in more detail. In summary, after experimentation with Tikhonov regularization we added some a priori covariance uncertainties in temperature and water vapor of 2.5 K and 60 %, respectively. These are extremely large values for a priori uncertainties compared to the anomaly variations. For example, the retrieved 400 hPa temperature anomalies shown in Fig. B3 are all less then  $\pm 3$  K, indicating that the temperature a priori covariance uncertainty is providing very minimal regularization. This means that almost all of the retrieved temperature variability is coming from the data and is not damped by the a priori estimates, a desirable situation for the measurement of climate trends. These a priori covariance uncertainty terms did, however, improve the profile trends generated in simulation by a slight amount (3 %-10 %) and thus were retained in our retrieval.

The observation error covariances (noise) are the mean AIRS NEDT for each channel, averaged over 16 d, and then divided by the square root of N, the number of scenes averaged. Originally, a fixed value of 0.01 K observation noise was used, but we found that this noise value depressed the CO<sub>2</sub> anomaly retrievals as they grew in size over time. This

problem disappeared once we switched to the true measurement noise values, which are in the range of NEDT equal to 0.004 K for longwave CO<sub>2</sub> channels from 700 to  $750 \text{ cm}^{-1}$ , about 0.001 K in window regions between 800 and  $1250 \text{ cm}^{-1}$ , and 0.001 to 0.002 K in the water band that covers the  $1300-1615 \text{ cm}^{-1}$  spectra region. These are extremely low noise values, which help explain why the anomaly retrievals have a relatively high number of degrees of freedom.

As stated earlier, the profile Jacobians used the ERA-I profiles, which were converted to anomaly profiles for each pressure layer. The minor-gas Jacobians were computed using our pseudo line-by-line kCompressed Radiative Transfer Algorithm (kCARTA) (Strow et al., 1998; DeSouza-Machado et al., 2020). kCARTA allows for extremely accurate Jacobian calculations, including analytic trace gas and temperature Jacobians. Initial retrievals used a fixed value for the minor-gas Jacobians. However, given the large increase in the minor gases (10 % for CO<sub>2</sub>), we determined that the minorgas Jacobians need to be updated as the gas amounts increase. Therefore, we used finite-difference Jacobians, computed using the minor-gas amount retrieved from the previous time step during the anomaly retrievals (or from the gas amount estimated using NOAA ESRL in situ gas amount data). The minor-gas profiles used in the Jacobian calculations are from Anderson et al. (1986). The  $CO_2$  profile is essentially constant in ppm until you reach the highest atmospheric layer.

There exists a weak dependence of these retrievals on the ERA-I model fields since we use the ERA-I model fields for the temperature,  $H_2O$ , and  $O_3$  profiles in the profile Jacobians, **K**. While we could retrieve the atmospheric profiles from the full radiance at each time step and latitude zone, ERA-I is so accurate we do not believe this is needed. Section 5.4 discusses potential errors introduced by using ERA-I for Jacobian evaluation, where they are shown to be extremely small and unimportant.

The direct retrieval of anomalies from the BT anomaly spectra represents a very different approach than normally used in infrared remote sounding. Although the mathematical approach is the same as in single-footprint retrievals (DeSouza-Machado et al., 2018), the often troublesome problem of static measurement and RTA bias errors is largely removed here since instrument calibration and/or absolute RTA biases do not appear in the retrieval process.

## 4.2 Channel selection

As discussed in Sect. 2, only channels that remain A+B throughout the mission are used, noting that the designation A+B does not apply to detectors in the M-11 and M-12 long-wave detector arrays. Initial retrievals showed that the AIRS shortwave detectors are drifting slightly, so these channels are also excluded from the anomaly fits (except for demonstration tests as discussed below). Unfortunately, the use of

only A+B detectors greatly restricts the number of available channels in the important longwave  $CO_2$  temperature sounding region from 710 to  $780 \text{ cm}^{-1}$ , where many channels are either A only or B only. It is important to weight these channels relatively strongly in the retrieval minimization. Since we also wish to de-emphasize stratospheric contributions to the minor-gas rates only every fifth channel from 650 to  $720 \text{ cm}^{-1}$  was included in the retrieval. In addition, any channels in this range with Jacobians that peaked above 10 hPa were excluded.

All channels in the M-5 array were excluded since they have relatively poor radiometric stability (as will be shown later). Several window channels that are sensitive to Chlorofluorocarbon-11 (CFC-11) were excluded, although many channels sensitive to CFC-12 were included, and CFC-12 trends were retrieved. Many H<sub>2</sub>O channels were included, since they are mostly A+B and have been stable throughout the mission. After some experimentation, four channels sensitive to N<sub>2</sub>O were also excluded since they appear to be behaving significantly out of family. Three of these channels are located near the end of the M-4c array, which also exhibits some anomalous frequency shifting behavior (Aumann et al., 2020).

A total of 470 channels remained after this pruning process. These channels are nicely distributed throughout the AIRS spectrum and are easily sufficient for 1D-Var retrievals. The nominal number of DOFs for tropical scenes for this channel set are  $\sim 6$  ozone DOFs,  $\sim 8$  temperature DOFs, and 12 H<sub>2</sub>O DOFs. The larger number of H<sub>2</sub>O DOFs is likely due to the large number of H<sub>2</sub>O channels used (321 out of 470 channels).

The overall sensitivity of the anomaly retrievals to  $CO_2$  is shown is shown in Fig. 6 where the mean  $CO_2$  Jacobian, averaged over all channels, is plotted. The  $CO_2$  sensitivity peaks around 400 hPa, and drops to near zero at the surface. There is some dependence on stratospheric  $CO_2$ , but stratospheric  $CO_2$  trends, especially in the lower stratosphere, should track the tropospheric trends, albeit with growth rates that are slightly influenced by previous years due to age of air. This figure also shows the mean  $CO_2$  Jacobian if all channels below 700 cm<sup>-1</sup> are removed (all sensitive to the stratosphere). Retrieval tests using these restrictions are discussed later.

#### 4.3 Construction of Jacobians

The relatively high accuracy of ERA-I temperature fields was highlighted previously in Fig. 5, which plots the time dependence of the bias between the observed and simulated BT for this channel. This bias, in black, has very little variability (other than the smooth decrease due to increasing CO<sub>2</sub>) compared to either the observed or simulated BT values due to the high accuracy of the ERA-I temperature profiles. This is not unexpected in a reanalysis product that assimilates a wide range of in situ measurements (radiosondes) and satel-



Figure 6. Mean of  $CO_2$  Jacobians for all channels used in the anomaly retrievals, and the same if all channels below  $700 \text{ cm}^{-1}$  (stratospheric channels) are excluded.

lite measurements (microwave and infrared sounders, including AIRS). In principal, we could use the AIRS Level-2 atmospheric state for generating the Jacobians for the anomaly retrievals. However, for the large-scale averaging used in this work, errors introduced by the relatively large ERA-I spatial grid compared to AIRS are minimized.

Moreover, ERA-I is constrained by a large number of instruments and in situ measurements for the temperature profile. Monthly mean ERA-I observation – analysis differences for radiosonde temperatures are below 0.2 K throughout the troposphere, rising to 0.3 K in the lower stratosphere (Simmons et al., 2014). We note that the statistical accuracy of the AIRS Level-2 algorithm is mainly verified by intercomparisons with ECMWF forecast/analysis fields (Susskind et al., 2014), which are likely even more stable in a reanalysis product. The AIRS Level-2 retrieved temperature and H<sub>2</sub>O global biases relative to ECMWF are very small, well below 0.5 K for temperature and 5 % for water vapor.

In principle, we could have performed 1D-Var retrievals on each 16d averaged BT spectrum in each latitude zone, but given the relatively small biases between ERA-I and AIRS shown in Fig. 2, retrievals will produce minimal improvements to the ERA-I fields. Note that the ERA-I bias in the 700–750 cm<sup>-1</sup> region with the most sensitivity to tropospheric CO<sub>2</sub> is only in the 0–0.5 K range. Moreover, 1D-Var retrievals using AIRS will also be limited by uncertainties in the AIRS radiometric calibration, which is estimated to be in the 0.2 K range (Pagano and Broberg, 2016).

More importantly, since we are only retrieving anomalies, highly accurate Jacobians are unnecessary since the BT variations in the anomalies are so small, especially when applied to trends. A quantitative assessment of errors in our measured anomaly trends from using ERA-I for Jacobian evaluations is presented in Sect. 5.4 and 5.7.

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#### 4.4 Temperature and minor-gas Jacobian co-linearity

A non-standard "correction" is made to the minor-gas retrievals that attempts to correct for the co-linearity of the temperature and minor-gas Jacobians. We demonstrate that this new approach clearly removes unphysical variability in the CO<sub>2</sub> anomaly retrievals. Co-linearity of the temperature and minor-gas Jacobians makes it difficult for the retrieval to separate temperature profile variations from variations in CO<sub>2</sub>, CH<sub>4</sub>, and N<sub>2</sub>O. Usually this is managed by constraining the retrievals with accurate a priori estimates that have small enough covariances to allow some separation of T(z) and CO<sub>2</sub> variability. Kulawik et al. (2010) discuss this problem in the context of CO<sub>2</sub> retrievals using the NASA Earth-Observing Satellite (EOS) Aqua Tropospheric Emission Spectrometer (TES) instrument, where they describe the selection of constraints as a way to "determine the partitioning of shared degrees of freedom between CO2 and temperature".

Here, we take a different approach based on the fact that we have highly accurate simulated anomalies computed from the ERA-I model fields. The simulated anomalies are derived using our SARTA RTA and were generated using constant values for the minor gases throughout the 16-year time period. Except for the minor-gas signatures, the ERA-I spectral anomalies are very similar to the observed anomalies since both AIRS calibration errors and RTA errors are largely removed when forming the anomalies. Figure 5 shows the excellent agreement between the observed and simulated BT anomalies for the  $710.14 \text{ cm}^{-1}$  channel. The only major difference in these anomalies is the downward drift in the observations primarily due to the growth of CO<sub>2</sub>. Note that almost all the high-frequency variability in the observed and simulated anomalies is removed when taking their difference, shown in black in Fig. 5, indicating that the ERA-I temperature fields match the AIRS observations very closely.

Given that the ERA-I spectral anomalies are very similar to observed anomalies, we can largely determine the effect of the Jacobian co-linearities on the observed  $CO_2$  anomaly retrievals by retrieving a (fictitious, or non-existing)  $CO_2$ anomaly from the ERA-I simulated BT anomalies using an identical retrieval algorithm. Since the simulated anomalies have a constant value for each minor gas, the variations in the retrieved  $CO_2$  (or other minor gases) are a measure of the inability of the retrieval to separate the minor-gas anomalies from the temperature profile.

Figure 7 illustrates this process for (1) a single latitude bin near  $-55^{\circ}$  latitude (with a width of  $\sim 4^{\circ}$  latitude) in Fig. 7a and (2) the average of 30 latitude bins covering  $\pm 50^{\circ}$  latitude in Fig. 7b. The yellow curve (labeled "Simulated") is the retrieved CO<sub>2</sub> anomaly derived from the ERA-I simulated anomalies. Although close to zero in the mean, this retrieved CO<sub>2</sub> anomaly varies considerably by up to  $\pm 15$  ppm. The red curve (labeled "AIRS raw") shows the CO<sub>2</sub> anomaly retrieved from the AIRS observations, which has similar vari-



Figure 7. Illustration of "noise" removal in the CO<sub>2</sub> anomaly retrievals by subtracting the CO<sub>2</sub> retrieved from ERA-I simulations from the observed CO<sub>2</sub> retrievals: (a)  $-55^{\circ}$  latitude CO<sub>2</sub> retrieval; (b)  $\pm 50^{\circ}$  latitude average CO<sub>2</sub> retrieval.

ability superimposed on a linear ramp of  $\sim 2 \text{ ppm yr}^{-1}$ . The adjusted observed CO<sub>2</sub> anomaly is generated by subtracting the simulated from the observed CO<sub>2</sub>. This is shown in blue (labeled "AIRS adjusted"), showing that most of the "noise" has been removed resulting in a very smooth CO<sub>2</sub> anomaly curve.

Figure 7b shows similar results but using the average of all latitude bins between  $\pm 50^{\circ}$  latitude. The co-linearity of the temperature and CO<sub>2</sub> Jacobians apparently changes randomly enough with latitude that the simulated CO<sub>2</sub> has far less variability than in Fig. 7a. The utility of this approach is nicely illustrated by examining the dip of about 7 ppm in the "Simulated" CO<sub>2</sub> retrieval in early 2010 for the  $\pm 50^{\circ}$ latitude bin. This dip is also visible in the observed anomaly curve (AIRS raw). The "Simulated" CO<sub>2</sub> anomaly is subtracted from the "AIRS raw" curve to obtain the final observed CO<sub>2</sub> anomaly (AIRS adjusted) and it is quite evident that the dip in early 2010 has canceled out, as desired.

The above adjustments to the  $CO_2$  anomaly retrievals have little effect on estimates of AIRS stability over 16 years, as outlined later in Sect. 5.4, although it does increase the statistical uncertainty in the AIRS BT trends by a factor of 2.4. More importantly, the application of these adjustments greatly reduces the apparent noise in the derived  $CO_2$ trends, making the detection of instrument shifts in the AIRS BT time series much more sensitive.

#### 5 Anomaly retrievals

# 5.1 AIRS events

Evaluation of the anomaly retrievals requires some knowledge of the AIRS mission events. Table 1 summarizes the major events during the AIRS mission that had thermal consequences for either the spectrometer or the focal plane arrays. While most of these events were minor, recent measurements of the AIRS frequency shifts (Aumann et al., 2020)

**Table 1.** Summary of AIRS events that had a thermal impact on either the spectrometer, the focal plane, or both. Dates are in DD/MM/YY format.

Date	Event
29/10/03	Aqua shutdown lasting for several weeks (solar flare)
09/01/10	Single event upset, focal plane temperature cycling
28/03/14	Single event upset, small focal plane cooler variation
25/09/16	Single event upset, one cooler restart

highlight that these events are associated with small shifts in the AIRS frequency scale. These shifts are indicative of very small movements of the detectors relative to the instrument spectrometer axis and could, for example, slightly alter the detector's view of the blackbody and cold scene. Any small non-uniformities in these calibration looks could affect the absolute radiometry. We will refer to these events during discussions of the anomaly retrieval results.

#### 5.2 Truth anomalies

The retrieved minor-gas anomalies are compared to the NOAA ESRL monthly mean data derived from in situ measurements (Tans and Keeling, 2019) for the Mauna Loa and Cape Grim sites, and for the global mean data for CO<sub>2</sub>, N<sub>2</sub>O, and CH<sub>4</sub>. Monthly anomalies for these in situ data sets were computed using the same methods used to compute the BT anomalies for consistency. We focus mainly on the global CO<sub>2</sub> ESRL anomalies since they are derived from a wide geographical range and sites and carefully merged to avoid local sources. The N2O ESRL anomalies provide information on AIRS channels in the  $1250-1310 \text{ cm}^{-1}$  region that are distinct from the main  $CO_2$  channels below  $780 \,\mathrm{cm}^{-1}$ . (There are also strong N<sub>2</sub>O channels in the shortwave band of AIRS.) The CH<sub>4</sub> anomalies mostly probe AIRS channels from 1230 to  $1360 \,\mathrm{cm}^{-1}$ . There is some concern that  $\mathrm{CH}_4$ anomaly trends may have more spatial variability than CO<sub>2</sub>

and  $N_2O$ ; however, we find good overall agreement with the ESRL global CH<sub>4</sub> trends, and CH<sub>4</sub> provides some sensitivity to channels that overlap with  $N_2O$  but extend a bit further into the water band.

We focus mostly on the use of CO<sub>2</sub> for AIRS stability estimations since CO<sub>2</sub> is so well measured and has the largest BT signal in the AIRS spectrum (relative to N<sub>2</sub>O and CH<sub>4</sub>). In addition, the N<sub>2</sub>O and CH<sub>4</sub> spectra overlap strongly in the AIRS BT spectrum, possibly introducing some retrieval uncertainty relative to CO<sub>2</sub>. Absolute errors in the ESRL CO<sub>2</sub> data are estimated to be  $\sim 0.2$  ppm (https:// www.esrl.noaa.gov/gmd/ccl/ccl uncertainties co2.html, last access: December 2019), with yearly growth rate uncertainties of  $\sim 0.07 \text{ ppm yr}^{-1}$  (https://www.esrl.noaa.gov/ gmd/ccgg/trends/gl gr.html, last access: December 2019). Anomaly growth rate errors averaged over 16 years are likely much lower since yearly sampling errors should diminish over time. Moreover, most absolute errors will not be applicable to the  $CO_2$  anomaly, which is a relative measurement. Therefore, it is difficult to definitively estimate the ESRL anomaly trend uncertainty. If the yearly growth rate uncertainties of 0.07 ppm yr<sup>-1</sup> are random, then the average of 16 of these growth rates would be  $0.018 \text{ ppm yr}^{-1}$ , which corresponds to a percentage uncertainty of 0.8 % in the anomaly trend.

Estimates for N<sub>2</sub>O and CH<sub>4</sub> anomaly trend uncertainties using the ESRL stated uncertainties in yearly growth rates, and assuming these are random errors each year, are 3.5%and 2.4%. These larger uncertainties, and the smaller total impact of these two gases on the AIRS BT anomalies, suggest that the best estimates for AIRS stability are likely derived from the CO<sub>2</sub> anomalies.

#### 5.3 Shortwave trends

Most of the anomaly retrievals performed here only included AIRS channels located below  $1615 \text{ cm}^{-1}$ , avoiding the shortwave channels in the 2181 to  $2665 \text{ cm}^{-1}$  region. Early retrievals showed that the AIRS shortwave channels exhibit a positive trend compared to the longer wave channels. Moreover, anomaly fits to just the shortwave channels return SST trends that are significantly larger than both the longwave channels and both the ERA-I (OSTIA) and OISST SST products.

The behavior of the AIRS shortwave channel relative to the longwave is easily seen in the anomaly retrieval fit residuals. Figure 8 shows the mean value (taken over the 365 16 d time steps for  $\pm$  30° latitude) for the residuals. All AIRS L1c channels are plotted, which includes many bad channels, and channels that do not exist but are filled during L1c creation (Aumann et al., 2020). The channels selected for the anomaly fits (see Sect. 4.2) are shown in red circles. The fit residuals for channels used in these retrievals are almost all well below 0.02 K. However, the shortwave channels show anoma-



Figure 8. Anomaly fit residual, averaged over all 365 16d time steps for  $\pm 30^{\circ}$  latitude. The L1c fill channels have no L1b counterparts and are simulated in the production of L1c. Note the offset in the shortwave.

lies inconsistent with the longwave of up to  $\sim 0.07 \text{ K}$  in the window channels past 2450 cm<sup>-1</sup>.

The anomaly retrievals can respond to drifts/offsets in the AIRS radiances by retrieving geophysical variables (CO<sub>2</sub>, temperature, etc.) that vary incorrectly in time. Alternatively, unphysical changes in the radiances could also be reflected in larger non-zero fit residuals. This could happen when the forward-model Jacobians cannot model time-dependent radiance errors, especially for jumps in the radiometric calibration that happen due to AIRS events (shutdowns). One way to examine this possibility is to look for any remaining trends in the anomaly fit residuals, which are shown in Fig. 9. Most of the channels used in the anomaly fits have residual slopes below  $0.002 \text{ K yr}^{-1}$ , although careful examination of the residual time series for particular channels can exhibit jumps associated with AIRS shutdowns.

The main observation in Fig. 9 is a clear positive trend in the shortwave relative to the longer wave channels used in the retrievals. The (AIRS - ERA) SST trend plotted as a solid horizontal line in this figure (discussed in Sect. 5.7) shows that the AIRS shortwave trends are more different from the ERA-I SST trends than the longwave channels. Most of the shortwave channels, including those in the mid-troposphere, exhibit positive trends relative to the longwave, except for some channels that are peaking very high in the stratosphere, below 10 hPa, that are marked in gray.

Consequently, unless otherwise noted, all the remaining results presented here avoid shortwave channels and use the channel set (470 channels) denoted in these figures.

#### 5.4 CO<sub>2</sub> anomaly retrievals

Figure 10 shows the retrieved CO<sub>2</sub> anomalies averaged over  $\pm$  50° latitude in blue and the ESRL global anomaly product



**Figure 9.** Linear trends in the anomaly fit residuals, averaged over all 365 16 d time steps for  $\pm 30^{\circ}$  latitude. Note the linear trend in the shortwave in these fit residuals. Also shown is the trend difference (ERA-I SST – AIRS SST) for these data.

in red. The correspondence over time is excellent. The AIRS - ESRL anomaly differences are shown in yellow. In order to convert the variation in the gas anomalies to an equivalent AIRS BT anomaly temperature, we computed anomaly retrievals with the observed AIRS BT anomaly spectra modified by a 0.01 K yr<sup>-1</sup> ramp for all channels. This 0.01 K yr<sup>-1</sup> ramp is divided by the resulting changes in the CO<sub>2</sub> anomaly linear trends (ppm  $yr^{-1}$ ) to obtain the sensitivity of the retrieval to a trend in the AIRS radiances, in K ppm<sup>-1</sup>. For CO<sub>2</sub>, this sensitivity is -0.073 K ppm<sup>-1</sup>. This is about twice as large as the largest column Jacobians in the AIRS spectra, which have a value of  $\sim 0.030 \,\mathrm{K}\,\mathrm{ppm}^{-1}$ . This is not unexpected, since the CO<sub>2</sub> column measurement is partially a relative measurement, especially for weak CO<sub>2</sub> channels in the window region where the absolute BT errors are mostly accounted for by (incorrect) adjustments in the SST that minimize the effect of the  $0.01 \,\mathrm{K \, yr^{-1}}$  applied ramp. It is also possible that the temperature profile could also adjust to minimize sensitivity of the ramp on the CO<sub>2</sub> ppm values. In addition, this sensitivity estimate assumes all AIRS channels are drifting, which is clearly an approximation given the results shown here.

The magenta curve in Fig. 10 is the (AIRS – ESRL) anomaly differences converted to BT units using the  $-0.073 \text{ K ppm}^{-1}$  sensitivity factor. This curve has been slightly smoothed for clarity. The right-hand-side vertical axis shows the variations in this curve in BT units. Most of the BT variability is within  $\pm 0.05 \text{ K}$ ; however, a transition in BT in late 2003 is larger. This larger transition is likely due to the Nov 2003 shutdown of the Aqua spacecraft. The AIRS channel center frequencies were shifted due to this shutdown (Strow et al., 2006) and were subsequently corrected in the AIRS L1c product (Aumann et al., 2020; Manning et al., 2019). In addition, as reported in Strow et al. (2006), inter-



Figure 10. Retrieved  $CO_2$  anomalies compared to ESRL global in situ data. The  $CO_2$  anomaly difference between AIRS and ESRL is shown in yellow. The magenta curve is that difference converted into BT units.

ference fringes in the AIRS entrance filters shifted after the November 2003 Aqua shutdown because AIRS was restarted at a slightly different spectrometer temperature. The fringes change the AIRS spectral response functions, which has not yet been corrected in the AIRS L1c product radiances.

Figure 11 illustrates the differences between the AIRS and ESRL  $CO_2$  linear growth rates. The growth rates for both our  $CO_2$  retrievals and the ESRL  $CO_2$  time series were computed by reusing the fitting function in Eq. (1) but now applied to the retrieved  $CO_2$  anomalies, i.e.,

$$CO_2(t) = CO_2(t=0) + b_1 t + \sum_{i=1}^4 d_i \sin(2\pi nt + \phi_i), \qquad (6)$$

where  $b_1$  are the CO<sub>2</sub> trends in ppm yr<sup>-1</sup>. Later, this equation will be used to fit the N<sub>2</sub>O, CH<sub>4</sub>, and SST anomalies, instead of the CO<sub>2</sub> anomalies as shown here.

Figure 11 plots the fitted values for the AIRS growth rates (the  $b_1$  term in Eq. 6), computed as a function of latitude. The CO<sub>2</sub> growth rates are not completely uniform from year to year, so Eq. (6) cannot perfectly fit the trend data. However, it provides a convenient metric for intercomparing these two CO<sub>2</sub> anomalies. Note that the error bars shown for AIRS are slightly overestimated because of the fact that Eq. (6) does not perfectly fit the slightly non-linear anomaly curve. The error estimates are 95 % confidence intervals and they have been corrected for serial correlations in the anomaly time series using the popular lag-1 autocorrelation approach detailed in Santer et al. (2000).

The Mauna Loa and Cape Grim growth rates are also shown, also derived using Eq. (6), as is the ESRL global rate, indicated by the dark black horizontal line. If the 16-year in situ rates indeed have an estimated error of 0.018 ppm yr<sup>-1</sup> (assuming the 0.07 ppm yr<sup>-1</sup> uncertainties in the ESRL rates are random), then AIRS is in close agreement with ESRL averaged over latitude. The latitude dependence of the AIRS



**Figure 11.** Observed linear trend in the AIRS  $CO_2$  anomalies versus latitude, compared to NOAA ESRL Mauna Loa (MLO), ESRL Cape Grim (GCRIM), and the ESRL global  $CO_2$  product trends (black line).

**Table 2.** Slope of the (AIRS – ESRL)  $CO_2$  anomalies in ppm yr<sup>-1</sup> units.

Data set CO <sub>2</sub>	Mean trend difference $(ppm yr^{-1})$	Uncertainty in trend (ppm yr <sup>-1</sup> )
Global	0.032	0.012
Mauna Loa	0.033	0.023
Cape Grim	0.056	0.020

derived rates appear to have clear latitudinal dependencies, with lower rates near the Intertropical Convergence Zone (ITCZ) and higher rates in regions of descending air. We do not examine this latitude dependence in this work, not only is it small, it could also be related to small inaccuracies in our retrieval algorithm.

Since the  $CO_2$  linear growth rate measurements are not sensitive to year-to-year variability in the CO<sub>2</sub> anomaly, we instead use the (AIRS - ESRL) global anomaly differences shown in Fig. 10 to quantify the AIRS stability. Any linear-trend differences between the AIRS and ESRL CO2 in Fig. 10 are quantified by fitting the (AIRS – ESRL) CO<sub>2</sub> anomaly differences to Eq. (6). Table 2 summarizes any trend in AIRS relative to ESRL by tabulating the  $b_1$  terms from the fit for the ESRL global, Mauna Loa, and Cape Grim sites. The uncertainties are as before, 95 % confidence intervals corrected for lag-1 autocorrelations. As one might expect, the global trends agree the best, and Cape Grim the worst. The higher errors for Cape Grim may be related to our clear subset having fewer samples at  $-40^{\circ}$  latitude relative to the 20° latitude zone occupied by Mauna Loa. These mean differences are extremely small, corresponding, for global, to  $1.5 \pm 0.6$  % trend differences.

Table 3 shows the conversion of the  $CO_2$  ppm trend differences to equivalent BT differences using the

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 $-0.073 \,\mathrm{K\,ppm^{-1}}$  sensitivity conversion. The baseline entry (first line of the table) represents the final configuration for the anomaly retrievals and is our best estimate for the differences between the ESRL and AIRS CO<sub>2</sub> anomaly trends,  $-0.023 \pm 0.009$  K per decade. This is an exceedingly small trend difference. While suggesting that AIRS is extremely stable, for channels sensitive to  $CO_2$ and temperature, systematic errors may be larger than the differences reported here. Our estimate of the ESRL global anomaly trend uncertainty discussed in Sect. 5.2, 0.8%, is equivalent to  $0.017 \text{ ppm yr}^{-1}$ . The AIRS – ESRL global trend difference shown in Table 2 is about 2 times larger than this estimate for the ESRL uncertainty and slightly larger than the statistical uncertainty in this trend difference. In BT units, this potential uncertainty in the ESRL global  $CO_2$  anomaly trend is ~ 0.012 K per decade.

The sensitivity of these results to uncertainties in the Jacobians are derived from the second partial derivative of BT as follows:

$$M_{\rm unc} = \frac{\partial}{\partial X} \left( \frac{\partial BT}{\partial Y} \right) \times X_{\rm unc} \times Y_{\rm meas}$$
$$= \left( \frac{\partial^2 BT}{\partial X \partial Y} \right) \times X_{\rm unc} \times Y_{\rm meas}, \tag{7}$$

where *M* is the quantity being measured (here,  $CO_2$  anomalies and trends),  $X_{unc}$  is the uncertainty in the profile variables used to compute the Jacobians, and  $Y_{meas}$  is either the maximum anomaly or the mean trend measured for *Y*. These are quantified in Table 4.

The first entry accounts for errors in  $\partial BT/\partial CO_2$  due to uncertainties in the CO<sub>2</sub> spectroscopy. The HITRAN database (Gordon et al., 2017) reports uncertainties in the CO<sub>2</sub> line strengths of 1 %-2 %. These uncertainties would translate into the same percentage error in the Jacobians. In addition, atmospheric spectra are sensitive to line widths, line shape, and line mixing, often at temperatures that are not measured in laboratory spectra. Characterizing the combination of these errors is essentially impossible, so here we assume a 1 % uncertainty in the CO<sub>2</sub> Jacobians, using the line strength uncertainty only. The maximum CO<sub>2</sub> anomaly error occurs at the end of the time series when the CO<sub>2</sub> anomaly is highest (35 ppm). Therefore, the max anomaly error is  $1\% \times 35$  ppm = 0.35 ppm. Using the retrieval sensitivity of  $-0.073 \,\mathrm{K}\,\mathrm{ppm}^{-1}$ , this translates into an effect max error in the BT anomaly error of 0.026 K. Dividing this anomaly uncertainty by the 16-year time period under study gives a trend uncertainty due to CO<sub>2</sub> spectroscopy errors of 0.016 K per decade as shown in Table 4. This value is slightly larger than the statistical uncertainty in the baseline CO<sub>2</sub> trend shown in Table 3 and slightly smaller than the derived trend differences versus ESRL CO<sub>2</sub>.

The second entry in Table 4 lists estimated uncertainties in the  $CO_2$  anomalies and trends (converted to BT units) that could arise due to errors in the ERA-I temperature profile.

CO <sub>2</sub> test	Mean trend difference (K per decade)	Uncertainty in trend (K per decade)
Global		
Baseline	-0.023	0.009
Baseline (no CO <sub>2</sub> adjustment)	+0.019	0.022
No strat.	-0.034	0.008
No cov. reg.	-0.043	0.009
No $\nu$ cal.	-0.059	0.010
Shortwave only	+0.070	0.009
ERA-I $T(z)$	+0.060	0.035
Mauna Loa		
Baseline	-0.024	0.017
Cape Grim		
Baseline	-0.040	0.020

**Table 4.** Anomaly and trend error estimates for  $CO_2$  and SST due to uncertainties in BT Jacobians via their second derivatives with respect to possible ERA-I uncertainties. As noted, the maximum effect on the  $CO_2$  anomalies would be at the time of the largest anomaly, which is at the end of our time series in August 2019. See the text for details.

Jacobian	Sensitivity	Uncertainty	Max effect on anomaly	Effect on trend
$\frac{\partial BT}{\partial CO_2}$	$\frac{\partial^2 BT}{\partial CO_2^2}$	1 % CO <sub>2</sub> spectroscopy	0.026 K (in August 2019)	0.016 K per decade
	$\frac{\partial^2 \mathbf{BT}}{\partial T_{\mathrm{air}} \partial \mathbf{CO}_2}$	0.5 K T profile	0.0035 K (in August 2019)	$2.2 \times 10^{-3}$ K per decade
$\frac{\partial \mathbf{BT}}{\partial T_{\mathrm{SST}}}$	$\frac{\partial^2 \mathbf{BT}}{\partial T_{\mathrm{SST}}^2}$	0.5 К <i>Т</i> <sub>SST</sub>	$4.0 \times 10^{-4}$	$9.6 \times 10^{-5}$ K per decade
	$\frac{\partial^2 \mathbf{BT}}{\partial T_{\mathrm{air}} \partial T_{\mathrm{SST}}}$	0.5 K T profile	$8 \times 10^{-4} \mathrm{K}$	$1.8 \times 10^{-4}$ K per decade
_	$\frac{\partial^2 \mathrm{BT}}{\partial T_{\mathrm{H_2O}} \partial T_{\mathrm{SST}}}$	10% H <sub>2</sub> O column	0.02 K	$4.5 \times 10^{-3}$ K per decade

The second partial derivative was computed with finite differences using a fixed temperature offset for all levels and then summed over all levels, a worst case scenario. The mean of these second-order derivatives, taken over the retrieval channels in the ~ 700–750 cm<sup>-1</sup> spectral region that has high sensitive to CO<sub>2</sub>, represents an effective scalar value for  $\partial^2(BT)/(\partial X \partial Y)$  in Eq. (7). This term is multiplied by an assumed uncertainty in the ERA-I temperature of 0.5 K and by the maximum anomaly value of 35 ppm to obtain a maximum uncertainty of 0.0035 K in the CO<sub>2</sub> anomaly. The maximum effect on the CO<sub>2</sub> trend is again this value divided by 16 years, giving an uncertainty of  $2.2 \times 10^{-3}$  K per decade, an insignificant uncertainty. Note that our assumed uncertainty of 0.5 K is higher than ERA-I error estimates discussed in Sect. 4.3.

Clearly, the estimated 1% uncertainty in the CO<sub>2</sub> spectroscopy is the dominant source of error in our CO<sub>2</sub> re-

trievals. If the ESRL 0.8% uncertainty is combined in quadrature with the 1% HITRAN uncertainty, a total minimum expected uncertainty in the  $CO_2$  anomaly trends is 1.3%. This translates to a BT uncertainty of 0.02 K per decade, close to our derived mean trend difference between AIRS and ESRL based on the  $CO_2$  anomaly measurements. This may be a more accurate uncertainty estimate for this measurement rather than the 0.009 K per decade statistical uncertainty derived from fitting the AIRS – ESRL anomalies.

The second entry in Table 3 lists the mean trend difference and its uncertainty if the adjustment for co-linearity discussed in Sect. 4.4 is not applied, which leads to a larger trend uncertainty by a factor of 2.4. The resulting trend difference is somewhat smaller but with a different sign. The baseline retrievals with and without the co-linear CO<sub>2</sub> adjustments do not quite overlap within their respective  $2\sigma$  uncertainties, missing statistical agreement by 0.013 K per decade, which is relatively small. However, based on the discussion in Sect. 4.4, we believe that the application of the co-linear CO<sub>2</sub> adjustment improves the accuracy of the AIRS CO<sub>2</sub> anomaly.

Table 3 also shows the results of a number of fit testing the sensitivity of the retrievals to various retrieval alternatives. The "no strat." entry removed all channels that primarily sense the stratosphere by removing all channels below  $700 \,\mathrm{cm}^{-1}$ . Figure 6 shows how this modifies the mean CO<sub>2</sub> Jacobian used in the retrieval, essentially removing all sensitivity to CO<sub>2</sub> above 60 hPa. Unfortunately channels above  $700 \,\mathrm{cm}^{-1}$  have some residual sensitivity to  $\mathrm{CO}_2$  in the stratosphere, and removing channels below  $700 \,\mathrm{cm}^{-1}$  may make it more difficult to properly minimize the retrieval residuals for some channels above  $700 \text{ cm}^{-1}$ . If  $S_a$  is completely removed, removing a priori profile regularization, the CO<sub>2</sub> anomaly trend difference increases by a factor of 2. Removing the L1c frequency calibration adjustments increases the anomaly trend differences by nearly a factor of 3, and changes their sign. If only shortwave channels are fit (excluding channels that peak above 10 hPa, and some channels sensitive to both carbon monoxide), the mean trend differences are more than 3 times larger than the baseline, again with a sign change.

The last test, labeled "ERA-I T(z)", examines the need for performing simultaneous retrievals of temperature profiles while retrieving the CO<sub>2</sub> anomalies by using the ERA-I temperature profiles anomalies, instead of fitting for them from the observed anomalies. This test increased the anomaly differences between AIRS and ESRL by almost a factor of 3, with a significant increase in the uncertainty of the trend, giving 0.35 K per decade instead of close to 0.009 K per decade for the baseline.

Table 3 also shows the Mauna Loa anomaly difference, which is close to the global result, although accompanied by a higher uncertainty of 0.017 K per decade compared to the 0.009 K per decade for the global anomaly. Cape Grim anomaly differences are almost 2 times higher than the global trend differences, but this is not surprising given the much lower number of observations at that latitude.

The retrieved AIRS global CO<sub>2</sub> anomalies did exhibit a small seasonal pattern for latitudes above 40° N of with an amplitude of  $\sim 0.5$  ppm. This is due to the residual of the seasonal cycle of CO<sub>2</sub> that is not completely removed when constructing the BT anomalies.

Note that radiometric shifts or drifts in the AIRS BT time series could be either reflected in incorrect geophysical trends, or partially buried in the anomaly fit residuals. The high quality of the anomaly retrievals for  $CO_2$  and the small fit residuals for  $CO_2$  channels strongly suggest that the AIRS blackbody is extremely stable, at least for long and mid-wave A+B channels. The SST retrievals discussed later reinforce this conclusion. However, we do see evidence of radiometric shifts due to discrete AIRS events (especially for N<sub>2</sub>O and CH<sub>4</sub>) that might be amenable to correction. Future



Figure 12. Retrieved  $N_2O$  anomalies compared to ESRL global in situ data. The  $N_2O$  anomaly difference between AIRS and ESRL is shown in yellow. The magenta curve is that difference converted into BT units.

Table 5. Slope of the (AIRS – ESRL)  $N_2O$  anomalies in K per decade units.

Data set N <sub>2</sub> O	Mean trend difference (K per decade)	Uncertainty in trend (K per decade)
Global	-0.141	0.012
Mauna Loa	-0.200	0.030
Cape Grim	-0.080	0.033

work will include careful examination of both the anomaly retrievals and their residuals, likely in an iterative fashion, in order to determine what channels are responsible for unphysical shifts in the anomaly products.

# 5.5 N<sub>2</sub>O anomaly retrievals

The N<sub>2</sub>O retrieved anomaly time series is shown in Fig. 12 and primarily senses the  $1240-1325 \text{ cm}^{-1}$  spectral region. Clearly the observed N<sub>2</sub>O anomaly is growing slightly faster than the ESRL values. The N<sub>2</sub>O anomalies are converted to equivalent BT variations just as for CO<sub>2</sub> but with a derived sensitivity of 0.140 K ppb<sup>-1</sup>. Table 5 tabulates the derived trend for the (AIRS – ESRL) anomaly by fitting the difference to Eq. (6) and then converting to BT units.

The trend differences here are much larger than for  $CO_2$ . Examination of either the AIRS – ESRL anomalies in ppb, or their equivalent in BT units (left-hand y axis), suggests that two unphysical steps might be present in the time series: one in mid-2005 and another one in mid-to-late 2010. Unfortunately, these steps do not closely coincide with AIRS events, possibly appearing more than 1 year after the November 2003 event and slightly less than 1 year after the January 2010 event.

To illustrate the effect of these two discrete shifts on the anomaly trend differences, we empirically introduce a step



**Figure 13.** Retrieved  $CH_4$  anomalies compared to ESRL global in situ data. The  $CH_4$  anomaly difference between AIRS and ESRL is shown in yellow. The magenta curve is that difference converted into BT units.

in our retrieved N<sub>2</sub>O time series of -0.6 ppb on 1 July 2005 and another step on 18 January 2010 of -0.5 ppb. The trend difference between this empirically modified time series and ESRL, in BT units, becomes  $-0.022 \pm 0.009$  K per decade, very similar to the CO<sub>2</sub> trend differences. The main point of this exercise is to illustrate that just two small discrete radiometric shifts could be responsible for the higher trend differences between AIRS and ESRL for N<sub>2</sub>O. More work is needed to map these discrete non-physical events in the retrieved N<sub>2</sub>O anomaly time series back into steps in the AIRS BT time series. The hope is that careful examination of the anomaly time series residuals during this process would highlight specific channels (or cluster of channels) that are behaving non-physically.

# 5.6 CH<sub>4</sub> anomaly retrievals

The CH<sub>4</sub> retrieved anomalies have some similarities to the N<sub>2</sub>O anomalies, since the spectra of both gases occur in the same general spectral region. The CH<sub>4</sub> the region of sensitivity is  $\sim 1210-1380$  cm<sup>-1</sup>. Figure 13 shows the CH<sub>4</sub> results using the same approach as for CO<sub>2</sub> and N<sub>2</sub>O. The ppb to BT conversion for CH<sub>4</sub> was measured to be 0.023 K ppb<sup>-1</sup>, significantly lower than for CO<sub>2</sub> or N<sub>2</sub>O, although total BT trend due to CH<sub>4</sub> is only marginally lower than CO<sub>2</sub> and N<sub>2</sub>O.

The high variability of atmospheric CH<sub>4</sub> growth is well known, as can be seen in the ESRL curve in Fig. 13. The AIRS derived anomalies follow that variable growth rate quite nicely overall. It should be noted that the ESRL CH<sub>4</sub> curve is more variable than CO<sub>2</sub> and N<sub>2</sub>O, and may be less uniform globally, making CH<sub>4</sub> a less ideal gas for testing AIRS stability. However, the AIRS – ESRL anomaly differences are valuable in that they, like N<sub>2</sub>O, highlight discrete jumps that can often be identified with AIRS events, such as late 2003 (biggest jump), early 2010, and possibly in early

**Table 6.** Slope of the (AIRS – ESRL)  $CH_4$  anomalies in K per decade units.

Data set CH <sub>4</sub>	Mean trend difference (K per decade)	Uncertainty in trend (K per decade)
Global	-0.107	0.024
Mauna Loa	-0.062	0.039
Cape Grim	-0.100	0.037

2014. The positive jump in the  $CH_4$  anomaly difference near March 2014 also coincides with a jump in the N<sub>2</sub>O anomaly difference, both taking place after the March 2014 event. However, this apparent jump seems to fade within 1 year for both gases. We believe this might be caused by AIRS frequency shifts that occurred in the M-4a and M-4c detector modules after this event. Those frequency shifts appeared to disappear within 1 year, and at present they are not corrected for in the AIRS L1c product.

Table 6 lists the trend differences between AIRS and ESRL for  $CH_4$ , showing trends differences that similar to those for  $N_2O$ , presumably since both gases absorb in the same spectral region.

#### 5.7 SST retrievals

The SST anomaly retrievals are compared to the ERA-I supplied SST (mostly OSTIA) and to NOAA's OISST operational SST product. Although both of these SST products are tied to the Argo floating buoy network, they are gridded SST products using interpolation derived from satellite data such as AVHRR.

A recent study (Fiedler et al., 2019) compared various SST products to the buoy network and found differences for OS-TIA of  $1.1 \text{ mK yr}^{-1}$ , and  $7.8 \text{ mK yr}^{-1}$  for OISST. This establishes a rough estimate of the differences in these products when evaluating them relative to our retrieved SST anomalies.

Figure 14 plots time series of our retrieved SST anomaly and the co-located ERA-I SST (mostly OSTIA) anomaly, averaged over  $\pm 30^{\circ}$  latitude, where these products are expected to be most accurate since most buoys are located in the tropics. The AIRS SST trend derived from this time series is  $0.096\pm0.046$  K per decade. The AIRS and ERA-I 16 d averaged anomalies agree very closely; their difference is shown in black. A zoom of the AIRS – ERA-I SST anomaly is shown in Fig. 15 to highlight their differences. Steps in these differences are possibly evident near the end of 2003 and especially near the end of September 2016 when AIRS had a cooler restart.

Table 7 summarizes the AIRS – (ERA-I and OISST) anomaly trend differences, computed using Eq. (6). The trend differences are quite small for both SST products. The (AIRS – ERA-I) trend has the same magnitude as the trend derived using  $CO_2$  but with the opposite sign. Overlap of the  $CO_2$ 



**Figure 14.** Tropical  $(\pm 30^{\circ})$  SST anomalies retrieved from AIRS compared to the ERA-I anomalies. The black curve is the difference between the AIRS and ERA-I anomalies.



**Figure 15.** Zoom of Fig. 14 that highlights the shift in the AIRS – ERA-I SST anomaly presumably due to the AIRS 25 September 2016 cooler restart. A small shift is also seen at the date of the November 2003 Aqua shutdown.

and ERA-I SST within their stated uncertainty estimates is missed by 0.01K per decade, which is very small. The CO<sub>2</sub> and OISST trend estimates miss overlap by slightly more, 0.02 K per decade. However, this overlap difference is small compared to the differences between OISST and the buoy network reported by Fiedler et al. (2019). Overall, the excellent agreement of these two extremely independent assessments (CO<sub>2</sub> versus SST) to within 0.02 K per decade is very encouraging given the complexity of the CO<sub>2</sub> measurement and the uncertainties in the SST product trends.

Comparisons between AIRS-derived SST and ERA-I or OISST products will contain biases due to time aliasing between the AIRS observations and daily means used in the SST products. Although these time-dependent biases can have random and seasonal variations of several hours, the observed linear drift in the AIRS local observing time over the 16-year observation period was less than 1 min yr<sup>-1</sup>, which is

**Table 7.** Slope of the (AIRS – (ERA/OISST)) SST anomaly differences.

Data set	Mean trend difference (K per decade)	Uncertainty in trend (K per decade)
(AIRS – ERA-I)	0.022	0.012
(AIRS – OISST)	0.034	0.021

far too small to introduce any drifts in the AIRS SST relative to the ERA-I or OISST daily averages.

Uncertainties in the SST anomaly retrievals due to our use of ERA-I fields for the evaluation of the SST Jacobians were estimated using the same approach for the CO<sub>2</sub> anomaly retrievals. The BT Jacobians (dBT/dSST) for channels sensitive to SST depend on accurate values for the SST itself, the air temperature profile, and most importantly the H<sub>2</sub>O profile, especially in the lower troposphere. We computed the partial derivatives of the BT Jacobians with respect to all three of these variables, again using finite differences and a constant offset for the air temperature profile and constant percentage offsets for the H<sub>2</sub>O profile. The partial derivatives were averaged for all AIRS channels used in our retrievals in the  $800-1235 \text{ cm}^{-1}$  region that is sensitive to surface temperature. The uncertainties assumed in the ERA model fields  $(X_{unc} \text{ in Eq. 7})$  are listed in column three of Table 4 and are likely higher than the estimated uncertainties summarized in Sect. 4.3. The uncertainties in the BT Jacobians are then multiplied by  $Y_{\text{meas}}$  in Eq. (7) which is either 0.4 K (the maximum SST anomaly; see Fig. 14) or  $0.0096 \,\mathrm{K} \,\mathrm{yr}^{-1}$  (our retrieved trend in SST).

The results shown in columns four and five of Table 4 clearly indicate that using ERA profile fields for estimated BT surface temperature Jacobian is extremely accurate. The highest uncertainties are due to  $H_2O$ , but even these are far below the statistical uncertainties shown in Table 7.

Aumann et al. (2019) recently compared the  $1231 \text{ cm}^{-1}$ AIRS channel trends to RTG SST, a precursor to OISST. This study used a statistical approach to remove trends in water vapor that affect the 1231 cm<sup>-1</sup> channel radiances, which Aumann et al. concede could introduce artifacts if there is a shift in the mean vertical distribution of water vapor. Our approach does not contain this limitation in principle, although we have not carefully examined the retrieved water vapor trends, mainly because there is no truth for comparison. An intercomparison of our results to Aumann et al.'s are not strictly possible since we used different SST products for truth and our SST anomalies used many channels. However, the trend of the 1231 cm<sup>-1</sup> channel in our retrievals can be derived by adding the slope of our fit residual for the  $1231 \text{ cm}^{-1}$  channel (-0.7 mK yr<sup>-1</sup>) to our derived SST trends for ERA-I and OISST. Using Aumann et al.'s units of mK yr<sup>-1</sup>, the result is a trend of 1.5 and 2.7 mK yr<sup>-1</sup> for ERA-I and OISST, respectively, with respective uncertainties of 1.2 and 2.1 mK yr<sup>-1</sup>. These two trends compare fa-



Figure 16. Latitude dependence of the linear trend in the AIRS retrieved SST, OISST, and ERA-I SST. Also shown are the SST trends when only the AIRS shortwave channels are used to compute the anomalies.

vorably with Aumann et al.'s night trend for  $1231 \text{ cm}^{-1}$  of  $+2.9\pm0.4 \text{ mK yr}^{-1}$ . It is interesting that our OISST trend differences agrees more closely with Aumann et al.'s RTG SST trend difference since these two data sets have similar heritage. Of course, the extremely low statistical errors reported by Aumann et al. do not allow overlap of these two results, but that is not necessarily expected since we use different SST products. Agreement for AIRS radiometric trends at the several mK yr<sup>-1</sup> level for at least a single channel should be considered quite remarkable.

We also derived AIRS – (ERA-I, OISST) SST trend differences using AIRS shortwave-only anomaly retrievals. For tropical latitudes,  $\pm 30^{\circ}$ , the (AIRS – ERA-I) trend is  $0.078\pm$ 0.040 and  $0.065\pm0.09$  K per decade for OISST. These represent significantly higher trend than observed using longwave and mid-wave channels only. The trend difference between (AIRS longwave – AIRS shortwave) anomaly fits is  $-0.058\pm0.026$  K per decade, clearly indicating the shortwave positive drift relative to the longwave.

The latitude dependence of the AIRS derived SST trends versus ERA-I and OISST may eventually help determine the source of some of these differences. Figure 16 shows these trends between  $\pm 60^{\circ}$  latitude. The uncertainties in these trends are  $\sim 0.005 \text{ K yr}^{-1}$  but are not shown since these uncertainties are primarily geophysical in nature (how linear is the SST trend) and affect each SST product identically. Agreement is quite good among all products in the Northern Hemisphere, while OISST is systematically lower than AIRS and ERA-I in the Southern Hemisphere. Also shown are the AIRS SST trends using only the shortwave channels (gray curve), which are always higher than the longwave AIRS trends except at the highest latitudes and near the Equator.

Unfortunately, the AIRS Level 2 retrieval algorithm only uses shortwave channels for surface temperature retrievals (Susskind et al., 2014). A recent intercomparison of surface temperature trends from the AIRS Level 2 retrievals to three established surface temperature climate products (Susskind et al., 2019) concluded that the AIRS surface temperature trends were 0.24 K per decade, slightly higher than Goddard Institute for Space Studies (GISS) Surface Temperature Analysis (GISTEMP)'s (Hansen et al., 2010) value of 0.22 K per decade, and significantly higher than the Had-CRUT4 (Morice et al., 2012) and Cowtan et al. (2015) values of 0.17 and 0.19 K per decade, respectively.

The results presented here conclude that the AIRS shortwave channels are drifting positive by about 0.058 K per decade relative to the longwave channels, which appear to be in extremely good agreement with established SST climate products as discussed above. If we subtract this 0.058 K per decade AIRS shortwave drift from the AIRS 0.24 K per decade trend presented in Susskind et al. (2019), we obtain a corrected AIRS trend of 0.18 K per decade, much more in line with the HadCRUT4 and C+W values. In this case, GIS-TEMP is now the only outlier. A more straightforward way to validate the reported AIRS Level 2 surface trends reported by Susskind et al. (2019) would be to directly compare them to other SST products such as OISST, but unfortunately this was not part of the Susskind et al. (2019) analysis.

# 5.8 CFC-12 retrieval

All anomaly retrievals presented here included CFC-12 retrievals. Although these are not used for quantitative assessments of AIRS radiometric stability, the retrieved CFC-12 anomaly is shown in Fig. 17 for completeness. Excellent agreement is found between the AIRS observed CFC-12 and the ESRL Northern Hemisphere measurements (ESRL, 2019). The linear trends derived from these two curves are  $-2.94 \pm 0.04$  ppt yr<sup>-1</sup> for AIRS, and  $-2.93 \pm 0.02$  ppt yr<sup>-1</sup> for ESRL, nearly perfect agreement. These results give us confidence that the SST retrievals have not been compromised by CFC-12 contamination, since there are a number of channels sensitive to both. Note that the trend of ~ 40 ppt of CFC-12 derived here from AIRS is equivalent to only ~ 0.11 K in BT!

#### 6 Retrieval BT breakouts and residuals

The anomaly fit residuals provide a wealth of information on the behavior of each AIRS channel versus time. As stated earlier, unphysical shifts in the AIRS radiance time series can be reflected in either the retrieved geophysical anomalies or in the fit residuals. Jumps in the fit residuals will generally take place when the shifted radiances cannot be "adjusted away" by the BT Jacobians, which require a reasonably accurate physical response to radiance jumps. We believe that the anomaly retrieval approach presented here will allow objective corrections to AIRS radiances, especially for radiance jumps that can be tied to instrument events. The excellent



**Figure 17.** AIRS CFC-12 retrieved anomaly compared to the NOAA ESRL Northern Hemisphere anomaly. Note that a 40 ppt trend in CFC-12 corresponds to about 0.11 K in brightness temperature for the channel with the highest CFC-12 Jacobian.

agreement between the  $CO_2$  and SST anomalies and in situ data strongly suggests that the AIRS blackbody is very stable, which is key to climate-level trend measurements.

There are several likely causes for some of the differences seen here between our observed anomalies and the N2O and CH<sub>4</sub> truth anomalies from ESRL. Shifts in the frequency calibration of AIRS (Strow et al., 2006; Manning et al., 2019) have largely been removed in the AIRS L1c product, although some transient shifts in the AIRS M-4a and M-4c arrays (that cover N<sub>2</sub>O and CH<sub>4</sub> channels) have not yet been corrected in L1c (see Aumann et al., 2020). The AIRS frequency shifts imply that detector views of the blackbody and cold scene targets have also shifted during the mission. While these shifts are very small, radiometric drifts/shifts could arise from these focal plane movements if the blackbody and cold scene targets are not perfectly uniform. As mentioned in Sect. 5.4, shifts of interference fringes in some of the AIRS entrance filters when Aqua was restarted in November 2003 may also contribute to the observed anomaly shifts. These fringe shifts have been modeled by the authors and future work may include modification of AIRS radiances before November 2003 to remove the effects of these small shifts in the instrument spectral response function.

Here, we present several views of the AIRS anomaly fits and their residuals as examples on how future work might proceed to potentially correct the AIRS radiances for small remaining radiometric drifts/shifts.

# 6.1 Retrieved anomalies in BT units

First, to provide some context, Fig. 18 shows the contribution of the various geophysical trends to the observed BT anomalies for channels sensitive to different geophysical variables. This is done by multiplying the BT Jacobian for some particular geophysical variable by its retrieved anomaly over time.



**Figure 18.** Contribution to the observed BT anomalies caused by the retrieved geophysical anomalies. These are simply the BT Jacobian multiplied by the time-dependent retrieved geophysical anomalies. The BT anomalies in panel (**b**) are multiplied by the sum, over all layers, of the retrieved profile anomalies.

For illustration purposes, we averaged the trends over the latitude bins from  $\pm 50^{\circ}$  latitude.

Figure 18a shows that the retrieved  $CO_2$  anomaly translates into a BT trend for the 722.1 cm<sup>-1</sup> channel of more than -1 K. Channels very sensitive to the retrieved CH<sub>4</sub> and N<sub>2</sub>O anomalies have BT trends that are lower than CO<sub>2</sub>. The anomaly for a channel sensitive to SST in Fig. 18a has an upward trend due to increasing SST values, but these are quite small compared to the minor-gas trends.

Figure 18b plots the BT anomalies due to the retrieved temperature, H<sub>2</sub>O, and O<sub>3</sub> anomalies. The profile anomalies have been summed over all levels for this figure. The same channel chosen to illustrate the BT anomaly due to the  $CO_2$  anomaly, 722.1 cm<sup>-1</sup>, is also used to illustrate the contribution of the temperature anomaly. The BT trend for the  $722.1\,\mathrm{cm}^{-1}$  channel due to the temperature anomaly is far smaller than for CO<sub>2</sub>, is slightly noisier, and has a small positive trend that mostly occurs after 2014. This would be expected since there is also a positive trend for SST with the same general time dependence. The BT trend due to the retrieved H<sub>2</sub>O anomaly is plotted for the 1418.6 cm<sup>-1</sup> channel sensitive to mid-tropospheric H<sub>2</sub>O. This BT anomaly moves in the opposite direction to the BT anomaly due to temperature, which is expected since on a large scale increasing temperatures raise H<sub>2</sub>O amounts, which leads to lower BT values.

Spectra illustrating how the various geophysical anomalies contribute to the BT anomalies are constructed by multiplying the BT Jacobians by the 16-year mean of the retrieved geophysical anomalies. Since these are computed quantities, all channels can be included. These are plotted in Fig. 19, where we separate the geophysical contributions just as in



**Figure 19.** Contribution to the observed BT anomalies caused by the retrieved geophysical anomalies. These are simply the BT Jacobian multiplied by the mean, over time, of the 16-year record of geophysical anomalies. The BT anomalies in panel (**b**) are multiplied by the sum, over all layers, of the retrieved profile anomalies.

Fig. 18. If the trends are linear in time, the 16-year mean anomalies represent the anomalies for year 8. Dividing these by 8 gives the nominal BT trend in K yr<sup>-1</sup>.

Figure 19a clearly shows that CO<sub>2</sub> dominates the changes in most of the longwave region, as expected. The N<sub>2</sub>O and CH<sub>4</sub> BT anomalies are concentrated in the 1230–1400 cm<sup>-1</sup> region with significant overlap, which is largely separable in the retrieval. On this scale, the BT changes due to SST are quite small. The temperature, H<sub>2</sub>O, and O<sub>3</sub> BT anomaly trends shown in Fig. 19b are derived from the sum of the profile Jacobians over all layers. In many regions of the spectrum the temperature and H<sub>2</sub>O BT anomaly trends are dominant, an indication that our anomaly retrievals successfully accounted for variability in those parameters. BT trends in the channels sensitive to tropospheric temperature (700– 750 cm<sup>-1</sup>) are in the range of 0.01–0.02 K yr<sup>-1</sup> (after dividing the plotted mean anomaly by 8), nominally consistent with global warming during this period.

The H<sub>2</sub>O greenhouse effect is clearly seen in Fig. 19b. The increased emission in the water band  $(1200-1615 \text{ cm}^{-1})$  due to higher atmospheric temperatures is largely negated by the decrease in emission due to increasing amounts of H<sub>2</sub>O, which shifts the emission in any given channel to higher altitudes where the temperature is lower.

Also note that channels sensitive to stratospheric temperatures in the  $650-690 \text{ cm}^{-1}$  region have a negative trend, indicating stratospheric cooling. This is also an expected result for global warming, but great care should be taken in using this data set for general conclusions since the sampling is non-uniform, and the air temperature trend standard deviation (Fig. 20) is about 80 % larger than the air temperature trend shown in Fig. 19.



**Figure 20.** Standard deviation for the contribution to the observed BT anomalies caused by the mean retrieved geophysical anomalies shown in Fig. 19.

The uncertainties in the mean spectral BT anomalies shown in Fig. 19 can be estimated from the mean differences between the observed and computed BT anomalies per channel shown in Fig. 8. An average over all fitted channels gives a mean residual of  $-0.0021 \text{ K} \pm 0.03 \text{ K}$ . This excellent fit, combined with the good agreement between the observed and in situ truth data for the CO<sub>2</sub>, CH<sub>4</sub>, and N<sub>2</sub>O anomalies indicates that the anomalies shown in Fig. 19a for a few sample channels are likely accurate to the anomaly fit  $2\sigma$  uncertainty of level 0.03 K.

For completeness, the standard deviation of the nominal linear anomaly trends shown in Fig. 19 are plotted in Fig. 20 using the same breakouts of geophysical anomalies. The CO<sub>2</sub> BT anomaly trend maximum standard deviation of  $\sim 0.35$  K near  $730 \,\mathrm{cm}^{-1}$  is nearly equal to the standard deviation expected if it was solely due to the linear trend in CO<sub>2</sub>. The air temperature stratospheric standard deviation is large, as previously noted, presumably due to the effects of the quasibiennial oscillation (QBO) and possibly El Niño-Southern Oscillation (ENSO) variability. The variability due to air temperature and H<sub>2</sub>O produces standard deviations in the water region  $(1250-1615 \text{ cm}^{-1})$  that are generally larger than variability due to trends in CH<sub>4</sub> and N<sub>2</sub>O, but apparently our retrieval successfully removes those interferences. Note the relatively high O<sub>3</sub> variability, which we do retrieve but have not examined carefully. It is important to remember that these are anomaly standard deviations, so they do not include seasonal variability.

# 6.2 Anomaly BT residuals

The anomaly fits shown above are summed and then subtracted from the observed BT anomalies to obtain the BT anomaly fit residuals. Any trends in these residuals can also



**Figure 21.** Slope of the AIRS anomaly residuals separated by A+B (fit channels), A-only, and B-only. This illustrates trends in the A-only and B-only channels relative to A+B channels in some modules. The A-only and B-only channels were not used in the fitting, so they are not strictly residuals but observed – computed differences.

be examined to search for channels that changed characteristics during the 16-year time period.

Figure 21 shows the BT anomaly fit residual slopes for A+B, A-only, and B-only channels separately. Most of the A+B channels shown, all of which were used in the anomaly retrievals, are within  $\pm 0.004 \,\mathrm{K \, yr^{-1}}$  of zero. While a large number of A-only and B-only channels are in agreement, there are a number of cases where they exhibit significant slopes (trends) that are not in agreement with the A+B channels. Module M-05 channels near  $1100 \,\mathrm{cm}^{-1}$  are clearly drifting differently than the other channels (we did not use any A+B M-05 channels in the retrievals since they are also in error). Module M-08 channels near  $851 \text{ cm}^{-1}$  show a clear separation between A+B channels and A-only, B-only. Clearly, the opposite signs of the A-only versus B-only drifts are largely canceled when A+B channels are used. Since the SST retrievals are quite good, and because the surface channels near  $1200 \,\mathrm{cm}^{-1}$  agree with the A+B channels, we conclude that the A-only and B-only drifts are real, and possibly due to drifts, or offsets, in the exact part of the blackbody and/or cold target scenes observed by these detectors.

Since the N<sub>2</sub>O retrieved anomalies exhibit some small unphysical behaviors, we examine the fit residuals for the 24 channels (used in the retrievals) that are most sensitive to N<sub>2</sub>O. Visual inspection of these channels' residual time series clearly indicated that 12 of them had easily identifiable features due to AIRS events. Figure 22 shows three different averages of these residual time series: (a) 12 good channels, with no strong evidence of AIRS events, (b) 12 bad channels which clearly exhibit jumps at the time of AIRS events, and (c) the mean time series for all 24 channels used in the anomaly fits. We see that the good channel mean (blue) is very flat, with a slight indication of a jump near the November 2003 event. The bad channel curve (red) shows a large



Figure 22. Anomaly fit residual time series for various combinations of 24 channels sensitive to  $N_2O$  in the longwave. The bad  $N_2O$  channels have easily visible jumps at times corresponding to AIRS hardware events.

jump near November 2003, possibly some longer-term drifts, and a feature in March 2014 that seems to last for 1 to  $\sim 1.5$  years. This last feature can change sign depending on which bad channel is observed, making it very likely that this is due to the M-4a/M-4c frequency calibration shift that is not yet corrected in the L1c product.

A new set of anomaly retrievals was produced but with the 12 bad N<sub>2</sub>O channels removed. When compared with the ESRL N<sub>2</sub>O anomalies, this change produced slightly better agreement with ESRL after November 2013. The slope of the (AIRS - ESRL) anomaly difference curve was reduced from -0.141 K per decade (as reported in Table 5) to -0.113 K per decade, a slight improvement. This drift relative to ESRL reduces to -0.069 K per decade if anomaly data before November 2013 are ignored. This illustrates that improvements to the AIRS products can be achieved by removing channels with residuals that have non-physical jumps. If the November 2013 radiometric jumps can be removed (whether due to frequency shifts, fringe shifts, or pure radiometric jumps), even higher stability is possible. However, one could presently begin the AIRS time series, say on 1 January 2004, and retain a stability approximately 2 times better than climate trends.

These results illustrate a simple case for how the anomaly fit residuals can be used to improve AIRS trend products. In this work, we have not looked for non-physical jumps in the retrieved temperature,  $H_2O$ , and  $O_3$  profile anomalies. These products likely exhibit some of these behaviors and need to be included in any comprehensive study to further improve the AIRS radiance stability. Some sort of iterative approach will likely be needed in order to ensure that these small remaining radiometric jumps become undetectable in both the retrieved anomalies and in the anomaly residuals.

#### 7 Conclusions

A framework for establishing stability of the AIRS radiances has been introduced that uses retrievals of minor-gas and SST trends from BT anomaly spectra. Extremely good agreement between retrieved CO<sub>2</sub> trends (or anomalies) and in situ trends from NOAA ESRL to  $-0.023 \pm 0.009$  K per decade illustrates that a large fraction of AIRS channels are extremely stable, well below climate trends. The SST anomaly retrievals also compare favorably to the ERA-I reanalysis and to NOAA's OISST SST product, with differences of less than 0.022 K per decade, and slightly higher values for comparisons to OISST. Such good agreement for a wide range of detectors strongly suggests that the AIRS blackbody is very stable.

Unphysical radiometric jumps are observed in all the retrieved anomaly time series but especially for  $N_2O$  and  $CH_4$ . These jumps can largely be related to AIRS events, and we illustrate how the anomaly fit residuals, combined with intercomparisons to truth anomaly trends such as  $N_2O$ , may provide a way to correct small remaining jumps in some AIRS channels. This work emphasizes that users of AIRS radiances (both L1b and L1c) for climate applications must pay careful attention to channel selection, since certain detector arrays and channels are presently not suitable for climate trending, including all of the AIRS shortwave channels. However, establishment of such a high level of stability for so many remote sensing observations/channels is unusual, and should lead to a high level of trust in AIRS climate trends that pay careful attention to only using validated climate-level channels.

## Appendix A: AIRS detector array wavenumbers

Table A1 shows the wavenumber ranges covered by each of the 17 AIRS arrays.

**Table A1.** The wavenumber ranges covered by each of the 17 AIRS arrays.

Array name	Start $\nu$ (cm <sup>-1</sup> )	End $v$ (cm <sup>-1</sup> )
1a	2552	2677
2a	2432	2555
1b	2309	2434
2b	2169	2312
4a	1540	1614
4b	1460	1527
3	1337	1443
4c	1283	1339
4d	1216	1273
5	1055	1136
6	973	1046
7	910	974
8	851	904
9	788	852
10	727	782
11	687	729
12	649	682

## Appendix B: Anomaly and profile trend retrievals

A complete simulated BT anomaly data set was generated using ERA-I model fields, by matching each AIRS clear observation to ERA-I and generating a simulated radiance. This simulated data set was used to set the regularization parameters for the profile inversions. The measurement of anomalies largely removes systematic errors in both the radiance observations (radiometric accuracy) and in the RTA (spectroscopy errors). We believe that these two factors helped make the retrieval inversions quite stable, requiring only minimal regularization.

Since our interest is mainly in the minor-gas profile offsets, we used 20 atmospheric layers for the retrievals (20 each for temperature,  $H_2O$ , and  $O_3$ ), created by concatenating layers from the 100-layer atmospheric profile model in Strow et al. (2003). This choice, coupled with our regularization, provided more layers than degrees of freedom, as desired. We found that the low noise of the AIRS zonally averaged 16 d anomalies (see Sect. 4.1) coupled with low bias errors in the measurement covariances permitted the use of only minimal regularization.

Retrieval trials started with Tikhonov-only first-derivative (L1-type) regularization, which removes obvious outliers, mostly in the higher latitudes in the stratosphere. This gave averaged linear-trend accuracies in the simulations of



**Figure B1.** Temperature kernels for the anomaly retrievals. These are taken from a random day for the zonal bin centered at 28.3° N.



**Figure B2.**  $H_2O$  kernels for the anomaly retrievals. These are taken from a random day for the zonal bin centered at  $28.3^{\circ}$  N.

 $-0.03 \pm 0.07$  K yr<sup>-1</sup> compared to the ERA-I model field trends used to generate the anomaly data set. (This degrades to  $-0.05 \pm 0.08$  K yr<sup>-1</sup> if the regularization is lowered by a factor of 10.) A reasonable goal is to achieve trends in simulation accurate to 0.01 K yr<sup>-1</sup>, averaged over the troposphere. A priori uncertainties were then introduced for the temperature and H<sub>2</sub>O profiles of 2.5 K and 60 %, respectively, which are roughly the maximum variation in these quantities over time for  $\pm 50^{\circ}$  latitude. These covariances are not very restrictive given that measurement uncertainties are so low. It appears that their main impact is again for high latitudes under conditions where we have higher noise due to a low number of clear samples.

The temperature and water vapor retrieval kernels are shown in Figs. B1, B2. They exhibit a very regular spacing in the troposphere with roughly 12 well-separated kernels.



**Figure B3.** Retrieved 400 hPa temperature anomalies versus latitude. (a) Our retrievals from the AIRS observations. (b) ERA-I anomalies.

Figure B3 illustrates the 400 hPa temperatures retrieved from the AIRS data (Fig. B3a) along with the ERA-I anomalies computed directly from the model fields. We do not expect these two data sets to compare perfectly, since, for example, the ERA-I anomalies are from relatively large gridded data and the AIRS measurement are from a nominal  $15 \times 15$  km field of view. Given the non-uniform sampling of this data set, we do not think detailed examination of the observed versus ERA-I anomalies is warranted for scientific purposes. However, note that there are many similarities in time and latitude that give some measure of validation to our profile retrievals. Similar results are seen with water vapor profiles.

Figure B4 summarizes the temperature trend simulations and comparisons between ERA-I trends, our anomaly retrievals from the ERA-I generated radiances, and those observed with the AIRS clear subset. The trends are computed from the anomaly retrievals (or model fields) using Eq. (6), where the input is the layer temperature instead of a  $CO_2$ amount.

These results have been slightly smoothed to make visual intercomparisons easier. Figure B4a shows the vertical trends versus latitude directly computed from the ERA-I temperature fields. Figure B4b shows our simulated temperature trend retrievals. These simulations agree quite well with the ERA-I model fields: the largest differences are seen in the lower troposphere at the higher latitudes and near the boundary layer in the tropics. The simulated retrievals are also placing the tropopause too high, which is not surprising given the lack of sensitivity of the infrared to the tropopause height and our limited number of vertical layers. Figure B4c shows the temperature anomaly trends retrieved from the AIRS observed anomalies. Clearly, there are significant differences between the ERA-I temperature profile trends and those we retrieved from AIRS, although the basic structure is relatively similar. Note that the uncertainties in these trends are quite high in the stratosphere (not shown) due to variations in the QBO, especially in the tropics, with errors larger than the observed trends in the vicinity of the tropopause. However, these uncertainties are largely present in both ERA-I and the AIRS observations.

The AIRS observed anomalies may also be impacted by errors in the BT Jacobians. Figure B4b used similar RTAs for both simulations and the retrieval. The version of SARTA used for the radiance simulations is based on HITRAN2008, while the Jacobians used in the retrieval used kCARTA which is based on HITRAN2016 and a slightly modified version of CO<sub>2</sub> line mixing. We expect that these spectroscopy differences have little impact since the CO<sub>2</sub> line strengths for the strong 15 µm bands have not changed between HITRAN versions. In addition, no noise was added to the simulated anomalies.

We believe that these results show that the anomaly retrievals used for measuring minor-gas trends exhibit realistic behavior and given our simulation testing this retrieval approach is likely to give accurate minor-gas trends. The impacts of some of the regularization choices are discussed in Sect. 5.4.



**Figure B4.** Temperature trends from the 16-year data period studied here. (a) ERA-I trends derived directly from the model temperature fields. (b) Simulated retrievals of the ERA-I trends using radiance anomalies created from the ERA-I fields and our SARTA RTA. (c) Temperature profile trends retrieved from the AIRS observed anomalies. The panel (b) simulation assumes that RTA is perfectly accurate.

*Data availability.* The AIRS L1c radiances are available for download from the NASA GES DISC at https://airs11.gesdisc.eosdis.nasa.gov/data/Aqua\_AIRS\_Level1/AIRICRAD.6.7/ (last access: 28 August 2020; Aumann et al., 2020).

The subset of 470 AIRS channels used for the anomaly retrievals presented in this article is available from Zenodo at https://doi.org/10.5281/zenodo.3878741 (Strow and DeSouza-Machado, 2020).

Author contributions. LLS led the study and made the comparisons between the anomaly fits and in situ data. SDM developed the anomaly retrieval algorithm. LLS and SDM together optimized the anomaly retrieval regularization.

*Competing interests.* The authors declare that they have no conflict of interest.

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