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Aerosol Single Scattering Albedo retrieved from ground-based measurements in the UV-visible

V. Buchard, C. Brogniez, F. Auriol, and B. Bonnel

Laboratoire d'Optique Atmosphérique, Université Lille 1 Sciences et Technologies, UMR8518, Villeneuve d'Ascq, France

Received: 16 July 2010 – Accepted: 21 July 2010 – Published: 30 July 2010

Correspondence to: V. Buchard (virginie.j.buchard-marchant@nasa.gov)

Published by Copernicus Publications on behalf of the European Geosciences Union.

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Abstract

Estimates of Aerosol Single Scattering Albedo (SSA) from ground-based spectral measurements in the UV-visible are conducted at Villeneuve d'Ascq (VdA) in France. In order to estimate this parameter, measurements of global and diffuse UV-visible solar irradiances performed under cloud-free conditions since 2003 with a spectroradiometer operated by the Laboratoire d'Optique Atmosphérique (LOA) are used. The technique consists in comparing the measured irradiance values to modelled irradiances computed for various SSA. The retrieval is restricted to the 330–450 nm range to avoid ozone influence.

For validation purpose, the retrieved values of SSA at 440 nm are compared to the ones obtained from sunphotometer measurements of the AERONET/PHOTONS network available on the LOA site. The results are rather satisfying: in 2003 and 2005–2006 the Root Mean Square (RMS) of the differences are about 0.05, these values are within the uncertainty domain of retrieval of both products. Distinction between days characterized by different aerosol content, by means of the aerosol optical thickness (AOT) retrieved from ground-based measurements at the same wavelength, shows that the comparisons between both products are better when AOT are higher. Indeed in case AOT are greater than 0.2, the RMS is 0.027 in 2003 and 0.035 in 2005–2006. The SSA estimated at 340 and 380 nm from ground-based spectra are also studied, though no validation can be carried out with sunphotometer data (440 nm is the shortest wavelength at which the SSA is provided by the network). The good comparisons observed at 440 nm can let assume that the SSA retrieved from spectroradiometer measurements at the two other wavelengths are also obtained with a good confidence level. Thus these values in the UV range can be used to complete aerosol data provided by AERONET/PHOTONS at VdA. Moreover they can be used for a best knowledge of the aerosol absorption that is necessary to quantify the error on surface UV irradiances estimated from satellite.

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1 Introduction

Beside geographical and temporal factors, several atmospheric parameters other than ozone influence the quantity of UV radiation reaching the Earth's surface. The most important are clouds and aerosols. Since several years aerosols receive a particular attention regarding their influence on climate, through both direct and indirect effects (Forster et al., 2007). The influence of atmospheric aerosols on radiation is not fully understood due to their high spatial and temporal variability. The aerosol studies related to aerosol forcing have been conducted mainly over the visible spectrum. However, some progress have been made in recent years to understand the effects of aerosols on ultraviolet UV radiation (WMO, 2007). In addition to the Aerosol Optical Thickness (AOT), a detailed knowledge of the Single Scattering Albedo (SSA), which provides information about the absorbing properties of the aerosols, is necessary to quantify the impact of aerosols on climate. Moreover a good knowledge of the regional characteristics of the SSA would allow decreasing the errors in the surface UV irradiance derived from space-borne measurements (Arola et al., 2009; Ialongo et al., 2010).

The AERONET/PHOTONS network (Holben et al., 1998) has been developed to determine aerosol optical properties but the shortest wavelength at which the SSA is restituted is 440 nm, i.e. in the visible. The SSA parameter is difficult to estimate in the UV range and even in the visible close to the UV due to rather large uncertainties existing on ground-based and space measurements in this wavelength range (Petters et al., 2003; Bais et al., 2005; Torres et al., 1998). Nevertheless, several methodologies using ground-based spectral measurements associated with radiative transfer model calculation have been developed (Petters et al., 2003; Bais et al., 2005; Krotkov et al., 2005) to determine SSA, and we have conducted such a retrieval.

The spectroradiometer performing spectral UV irradiance measurements in VdA used in this study is described in Sect. 2 along with the methodology used to derive the SSA from these ground-based measurements.

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Section 3 reports comparisons between the previous product at 440 nm and the SSA derived at the same wavelength from measurements of AERONET/PHOTONS sunphotometers located close to the spectroradiometer. SSA values retrieved from the spectra at shorter wavelengths are also presented, extending the spectral variations of the SSA provided by AERONET/PHOTONS towards the UV. Sect. 4 reports the conclusions.

2 Instruments and methodology

2.1 Ground-based instrument

The instrument used in this study is a Jobin-Yvon HD10 spectroradiometer located in Villeneuve d'Ascq (VdA) on the roof of the laboratory building (50.61° N, 3.14° E, 70 m a.s.l.). This site is typical of a flat region near a city in the north of France.

The spectroradiometer scans the wavelength range from 280 to 450 nm with a sampling step of 0.5 nm and its spectral resolution is about 0.7 nm. A single spectral scan takes about 6 min. Calibration is performed every 3 months with two standard lamps traceable to NIST (National Institute of Standard lamp and Technology) and NPL (National Physical Laboratory), and each spectrum is corrected of the wavelength misalignment via a software tool developed at LOA (Houët, 2003). The irradiance expanded uncertainty (coverage factor $k=2$) is about 7% at 320 nm and about 5% at 400 nm for a high irradiance level and 9% and 7%, respectively for a low irradiance level (Houët, 2003). The instrument has been intercompared with the QASUME spectroradiometer (Quality Assurance of Spectral Ultraviolet Measurements in Europe) in September 2004 (Gröbner et al., 2006).

Since the implementation of a shadow disc in 2003, the spectroradiometer performs alternately scans of the global and of the diffuse irradiance every 15 min. The difference between the global and diffuse spectral irradiances enables to derive the direct spectral irradiance. A careful analysis has shown that the use of the shadower adds a negligible

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uncertainty on the direct irradiance data, reaching only 1% for SZA > 60° (Brogniez et al., 2008). On clear sky conditions, the analysis of the direct irradiance is used to derive the spectral aerosol optical thickness (AOT) in the 330–450 nm range with uncertainties varying from about 0.03 at 440 nm for SZA=70° to 0.05 at 340 nm for SZA=40° (Brogniez et al., 2008).

2.2 Retrieval of the Single Scattering Albedo

The SSA is estimated using global and diffuse irradiance measurements as well as direct-to-diffuse irradiance ratio (dr/df). The technique consists in comparing the measured quantities with modeled ones as proposed by Petters et al. (2003) and Bais et al. (2005).

The modeled irradiances are computed using a radiative transfer (RT) code (DIS-ORT, DIScret Ordinates Radiative Transfer, Stamnes et al., 1988) for various SSA values. The SSA values range from 0.60 to 1, with a step of 0.01. Following the study of Bais et al. (2005), we select the SSA for which the modelled and the measured quantities agree to within 1%. Generally, for the three quantities, more than one SSA value can satisfy this condition, leading to a SSA range.

The other input parameters used in the RT model are the mid-latitude temperature and pressure profiles and the aerosol parameters. The spectral AOT is derived from the spectroradiometer measurements, the asymmetry parameter (g) is supposed to be spectrally constant. We choose 0.7 following the studies of Madronich (1993), who found g varying between 0.6 and 0.8 in the UV wavelength range, and of Bais et al. (2005), who found g around 0.7 for urban aerosols. Moreover this value is consistent with an analysis of AERONET/PHOTONS data (level 1.5) showing a mean value of 0.70 ± 0.02 at 440 nm in VdA. The surface albedo is supposed equal to 0.02 for all wavelengths (Feister and Grewe, 1995). The accuracy on the retrieved SSA results from the uncertainties existing on these input parameters and on the measured irradiances. Concerning g and the surface albedo, a variation of these parameters leads to a slight variation of the SSA value, about 0.01. SSA uncertainties have been estimated

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for two SZA cases using modelling with two AOT values. The results at 380 nm for the three quantities (global, diffuse and ratio dr/df) are reported in Table 1. The accuracy on the retrieved SSA depends on the amount of aerosol and for the global quantity, on the SZA, and it appears that the three quantities are less sensitive to changes in SSA at small AOT and for the global quantity also at small SZA. Moreover, the dr/df ratio provides the better accuracy in the estimation of SSA, $\Delta(SSA) \approx 0.05$. The same conclusions were obtained at 340 and 440 nm. These results are consistent with those retrieved by Bais et al. (2005) and Petters et al. (2003).

As discussed in Brogniez et al. (2008), important high frequency variations appear in the AOT spectrum deduced from the spectroradiometer measurements, especially near Fraunhofer lines. To remove these rapid oscillations a triangular smoothing over 4 nm is performed (Brogniez et al., 2008). Few oscillations remain, more or less pronounced depending on the extraterrestrial spectrum used. To reduce the error on the retrieved SSA linked to these AOT oscillations, a SSA average is computed over 5 nm around each wavelength, (i.e. 2.5 nm before and after each wavelength). Figure 1 shows an example of diurnal variations of the SSA obtained on a cloud-free day around 340 nm (Fig. 1a) and 440 nm (Fig. 1b), the errors bars are the standard deviations of the SSA. As expected from the results in Table 1, the dispersion is larger when global or diffuse measurements are used, meaning that the SSA is worse estimated when using these two measurements rather than the dr/df ratio. Thus, in the following we have only considered the SSA derived from this later quantity.

Moreover, Brogniez et al. (2008), have noticed that the AOT at 440 nm obtained with the Thuillier et al. (2003) spectrum (referred to as TRS in the following) agree better with the AOT obtained from AERONET/PHOTONS measurements, but that the oscillations are larger than when using the reference spectrum provided in the SCHICrivism software (Slaper et al., 1995), (referred to as SRS in the following). Thus, given the sensitivity of SSA to the AOT we have constructed a “new spectrum” in the spectral range 400–450 nm with the SRS resolution and the TRS radiometry. For that purpose we multiplied the SRS spectrum by 0.95 in order to reach the TRS radiometry. The coefficient 0.95

is a mean value over 400–450 nm of the ratio of both extraterrestrial spectra. The three extraterrestrial spectra are presented in Fig. 2.

The effect of the extraterrestrial spectrum on SSA retrieval at 440 nm is shown in Figs. 3 and 4, which present two examples of the retrieved results for two days characterized by different aerosol content. The daily mean AOT is 0.12 on the first day while it is about 0.26 on the second day. The left-hand side panels in both figures are obtained with the “new spectrum” and the right-hand side panels are obtained with the TRS. For the first day the AOT variations obtained with both spectra are very close, there is about 0.01 difference on average (Fig. 3b and d). The SSA variations (Fig. 3a and c) exhibit the same behaviour but the values are different, about 0.04 difference on average, and the dispersion is larger with the TRS. For the other day, the AOT values differ also by about 0.01 on average (Fig. 4b and d) but the SSA values are less different, about 0.02 on average (Fig. 4a and c) and the dispersion is very similar for both reference spectra. So we can conclude that the choice of the extraterrestrial spectrum is important in case of low aerosol content due to the sensitivity of the SSA retrieval to the aerosol load.

3 Applications

The shortest wavelength at which SSA is deduced from sunphotometer measurements of the AERONET/PHOTONS network is 440 nm. The inversion algorithm provides SSA using the sky radiance measurements along the solar principal plane and along the solar almucantar (Dubovik et al., 2000; Dubovik and King, 2000). According to these authors, the uncertainty in the retrieved SSA varies between 0.03 and 0.07 depending on the aerosol type and loading. For urban aerosol type, the SSA retrieval accuracy is between 0.05 and 0.07 when the AOT is lower or equal to 0.2 and 0.03 when the AOT is greater than 0.2. Due to the weakness of the AOT in VdA, the AERONET/PHOTONS data used here are level 1.5 values, which are cloud-screened.

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Figure 5a–b shows scatter plots of SSA retrieved with the spectroradiometer and with the sunphotometer on clear sky conditions, during 2003 and during 2005–2006 at 440 nm. The number of sunphotometer's inversion in a day is generally small because of the required criteria for a good inversion (see above), which explains the small number of comparison points (109 and 51 for 2003 and 2005–2006, respectively).

Accounting for the uncertainties on both products, the comparison is rather satisfying, the correlation coefficient are $r=0.84$ in 2003 and $r=0.71$ in 2005–2006, and the RMS are 0.050 and 0.048 in 2003 and 2005–2006, respectively. We can notice that points characterized by high absorption are retrieved by both photometer and spectroradiometer. If we consider only days with an AOT greater than 0.2 (red crosses), the comparison is better than for all points, the RMS is equal to 0.027 in 2003 and 0.035 in 2005–2006. When looking at the days with an AOT lower than 0.2, differences between SSA obtained from the spectroradiometer and from photometers are in general larger, RMS are 0.060 for both periods. This result confirms the difficulty to retrieve a SSA when the AOT is low.

If we consider all points for the two periods (Fig. 5a and b), we notice that on average the SSA retrieved from the spectroradiometer are lower than those retrieved from the sunphotometers, mean differences (spectroradiometer – photometer) being negative and equal to -0.03 for 2003 and -0.02 for 2005–2006.

As stated in Sect. 2.2, the SSA is also retrieved at shorter wavelengths from the spectroradiometer measurements. In that case no validation via the sunphotometers can be carried out, nevertheless since the AOT inferred at 380 and 340 nm have been well validated against AERONET/PHOTONS (Brogniez et al., 2008), the corresponding SSA are assumed to be reliable. Figure 6a–b presents an example of the spectral dependence of the retrieved-SSA from the spectroradiometer at 340, 380 and 440 nm combined with the SSA deduced from AERONET/PHOTONS measurements at 440, 675, 870 and 1020 nm with their corresponding uncertainties on 29 January 2006 at 11:00 UTC and 11 September 2006 at 13:30 UTC. For the spectroradiometer they are estimated following the Table 1 and for sunphotometers they are estimated to 0.05.

First, one can note that the spectroradiometer's SSA values at 440 nm are very close to sunphotometers' ones. The sunphotometer SSA values are well within the spectroradiometer uncertainty bars. One can notice also that the aerosols are more absorbent as wavelength increases on 29 January 2006 whereas the opposite is observed on 11 September 2006. Thus the aerosol spectral absorption seems to depend on the period of the year, but it depends also on the year, as confirmed in the following.

Figure 7a–b presents the monthly means of the SSA at 340, 380 and 440 nm for the spectroradiometer and at 440, 675, 870 and 1020 nm for sunphotometers in 2003 and 2006. The same days are considered to compute the monthly mean and dispersion around the mean at each wavelength, the number of days is indicated on top of Fig. 7a–b. The SSA monthly mean values at 440 nm from the spectroradiometer are well within the sunphotometer dispersion bars for most of the months.

The SSA spectral dependence is more or less pronounced depending on the period of the year. For example, in January 2006, the SSA mean value at 440 nm for the spectroradiometer and for the sunphotometer are equal (0.84 ± 0.05), and the value at 340 nm is 0.92 ± 0.03 compared with the value at 1020 nm equal to 0.72 ± 0.10 , showing large spectral variations over the whole wavelength range. Nevertheless it must be noted that for this month the dispersions in the near infrared domain are rather large, indicating a strong variability of the 13 observations. Moreover, depending on the year, for a same month, the SSA means at a given wavelength are quite different and the spectral dependence of SSA may vary. Indeed, if we compare the months of January and September in 2003 (Fig. 7a) and in 2006 (Fig. 7b), the spectral dependence of the mean SSA are very different (January) and even opposite (September). Thus, the SSA derived from the spectroradiometer measurements in the UV spectral range adds complementary information to that provided by AERONET/PHOTONS in the visible and near infrared, that could be use to identify the aerosol type in VDA.

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4 Conclusions

The global and diffuse spectral irradiance measurements performed by a spectroradiometer have been used with model calculations in order to determine the SSA in the UV-visible spectral range in VDA. The AOT obtained from ground-based spectroradiometer measurements are used as input data of the DISORT radiative transfer code to compute irradiances. Sensitivity tests show that the direct-to-diffuse irradiance ratio is the best quantity to estimate the SSA with the best accuracy. Under cloudless conditions, the SSA retrieved from this quantity at 440 nm have been compared with SSA obtained at the same wavelength with AERONET/PHOTONS sunphotometers operating close to the spectroradiometer. The comparisons show good agreement, especially when the aerosol amount is high, i.e. when AOT are greater than 0.2. SSA retrievals at 340, 380 and 440 nm derived from the spectroradiometer are combined with those retrieved by sunphotometers at 440, 675, 870 and 1020 nm in order to exhibit the spectral variations of the SSA over a large spectral range. It appears that the SSA value retrieved in VdA in the UV can be different from the one retrieved in the visible. Thus, estimating the SSA in the UV from spectroradiometer measurements leads to complementary information on the aerosol absorption already provided in the visible, that is useful for the monitoring of the aerosol in UV.

Acknowledgements. We thank L. Blarel and T. Podvin from the PHOTONS team for their help in selecting the data. The site is supported by CNES within the french program TOSCA. The figures were drawn using the Mgraph package developed at LOA by L. Gonzalez and C. Deroo (<http://www-loa.univ-lille1.fr/Mgraph>).



The publication of this article is financed by CNRS-INSU.

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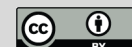
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Table 1. Estimated uncertainty on the SSA retrieved from spectral global and diffuse irradiances and from the direct-to-diffuse irradiance ratio (dr/df) at 380 nm for two AOT, at two SZA.

Quantity	Global irradiance		Diffuse irradiance		Ratio dr/df	
AOT	0.2	0.4	0.2	0.4	0.2	0.4
Δ (SSA)						
SZA=30°	0.16	0.09	0.08	0.05	0.06	0.04
SZA=60°	0.13	0.07	0.08	0.05	0.06	0.04

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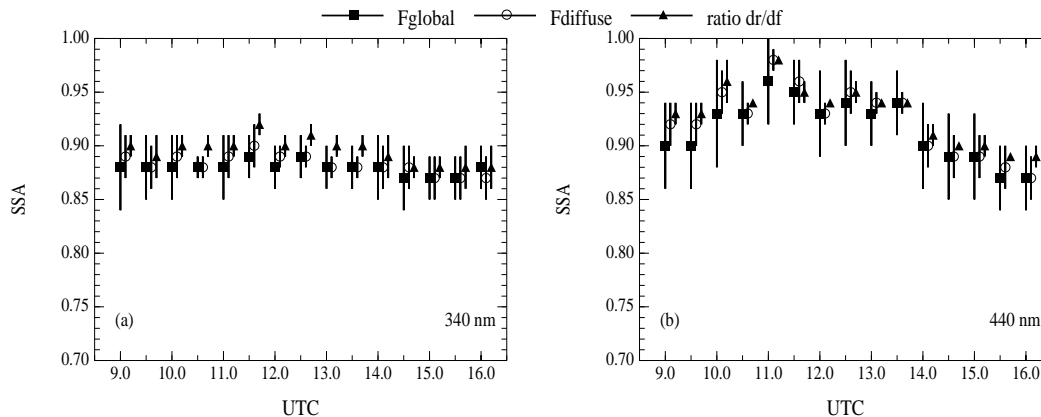


Fig. 1. Daily variations of SSA retrieved from global (square), diffuse (circle) and the direct-to-diffuse ratio (triangle) measured by the spectroradiometer on 11 September 2006: **(a)** at 340 nm; **(b)** at 440 nm. Vertical bars correspond to the standard deviation around the mean SSA computed over 5 nm around the central wavelength (see text).

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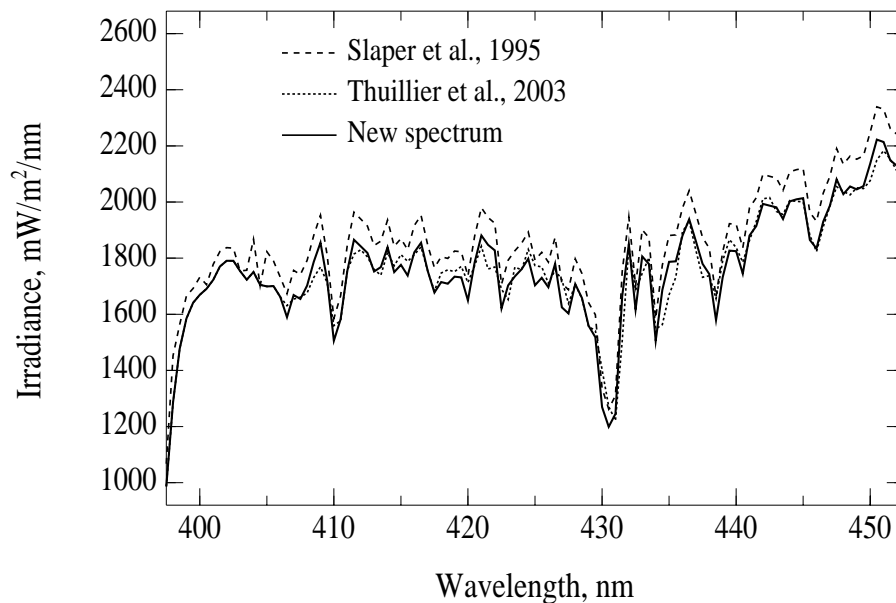


Fig. 2. Extraterrestrial spectra between 400 and 450 nm: Slaper et al., 1995 in dash-line; Thuillier et al., 2003 in microdash-line and the “new spectrum” constructed from both in solid-line.

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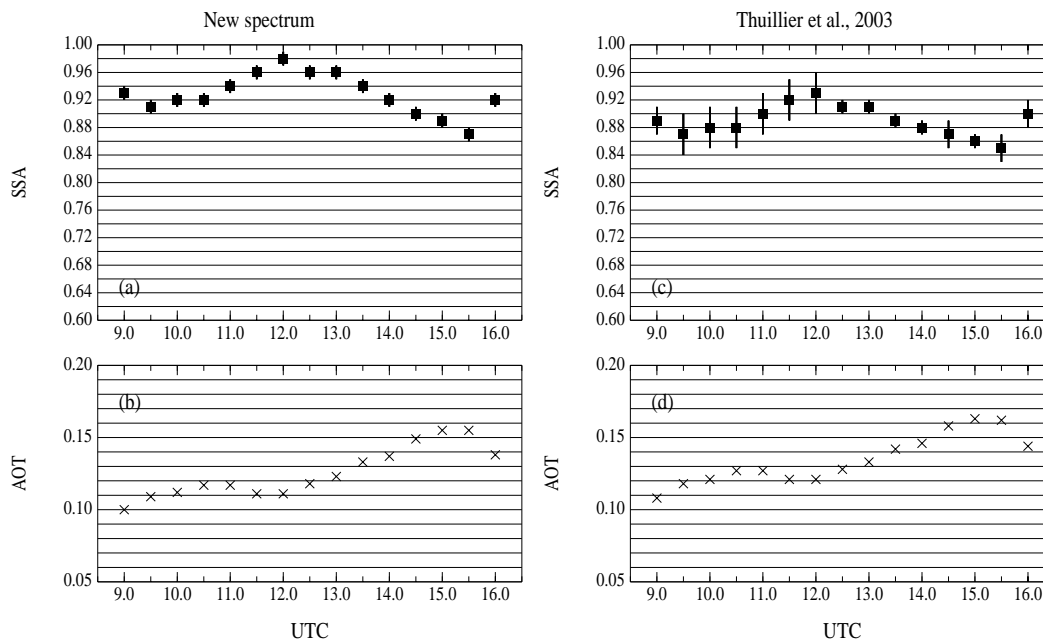


Fig. 3. Daily variations of AOT and SSA at 440 nm retrieved from spectroradiometer’s measurements on a day characterized by a low aerosol content (12 June 2006): **(a)** and **(b)** for retrieval with the “new spectrum”; **(c)** and **(d)** for retrieval with Thuillier et al., 2003 spectrum. Vertical bars in (a) and (c) are the standard deviations around the mean SSA computed over 5 nm interval.

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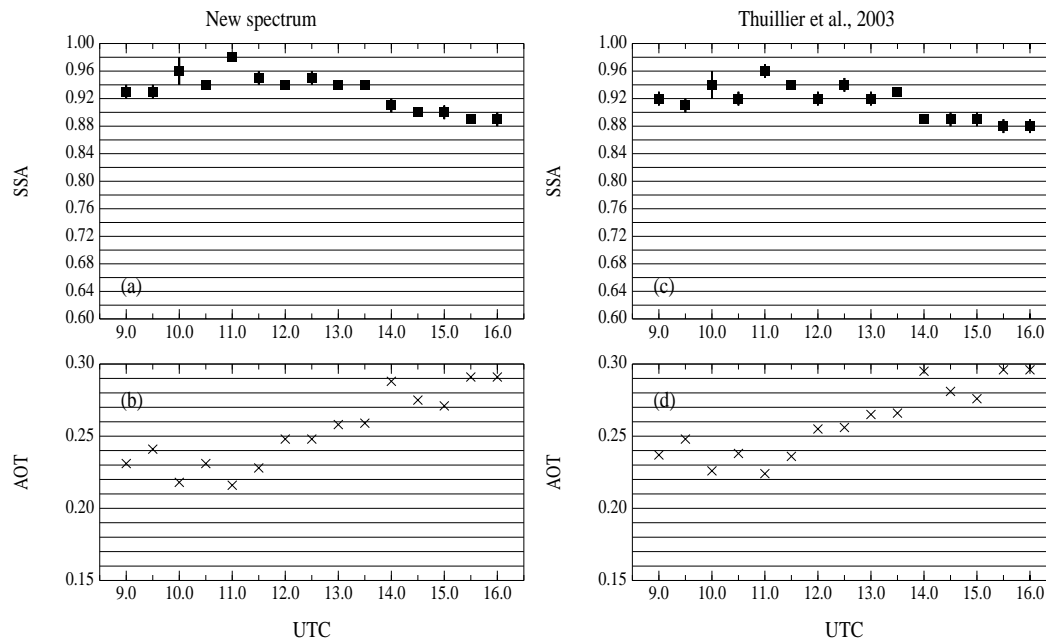


Fig. 4. Same as Fig. 3 for a day characterized by a high aerosol content (11 September 2006).

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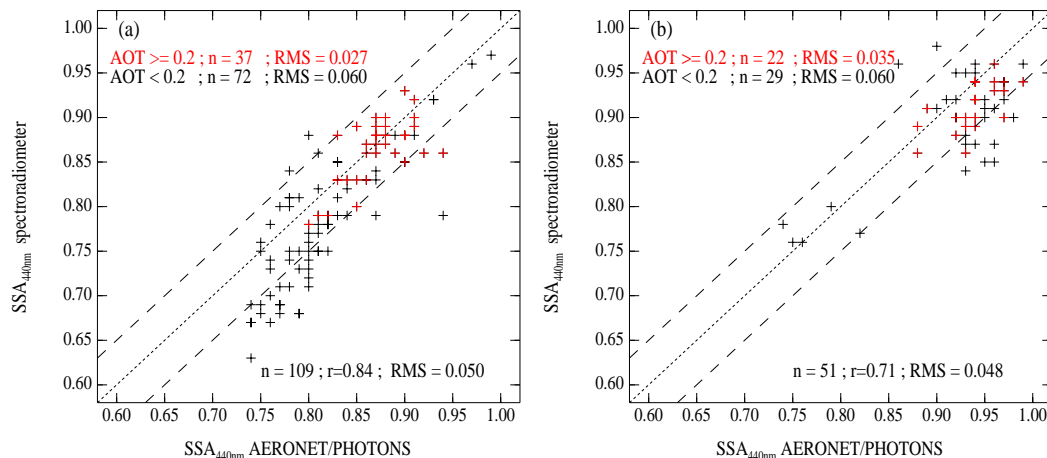


Fig. 5. Scatter plots of SSA retrieved at 440 nm from the spectroradiometer versus SSA from AERONET/PHOTONS: **(a)** in 2003; **(b)** in 2005–2006. Red crosses for AOT greater than or equal 0.2, black crosses for AOT lower than 0.2. The correlation coefficient is indicated, the micro-dash line is the first bisector and dash lines are located at ± 0.05 from the first bisector. n is the number of points and RMS is the Root Mean Square of the difference (Spectro-AERONET).

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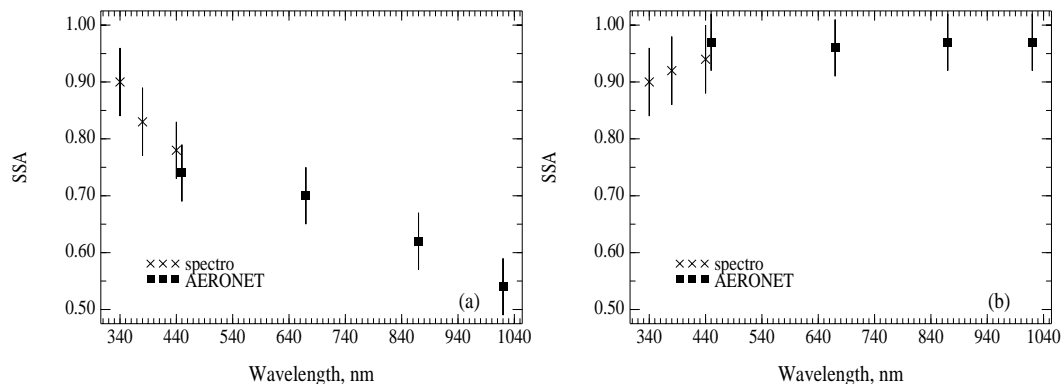


Fig. 6. Spectral variations of SSA retrieved from spectroradiometer's measurements at 340, 380 and 440 nm (crosses) and from AERONET/PHOTONS at 440, 675, 870 and 1020 nm (squares): **(a)** 29 January 2006 at 11:00 UTC; **(b)** 11 September 2006 at 13:30 UTC. Vertical bars correspond to SSA uncertainties.

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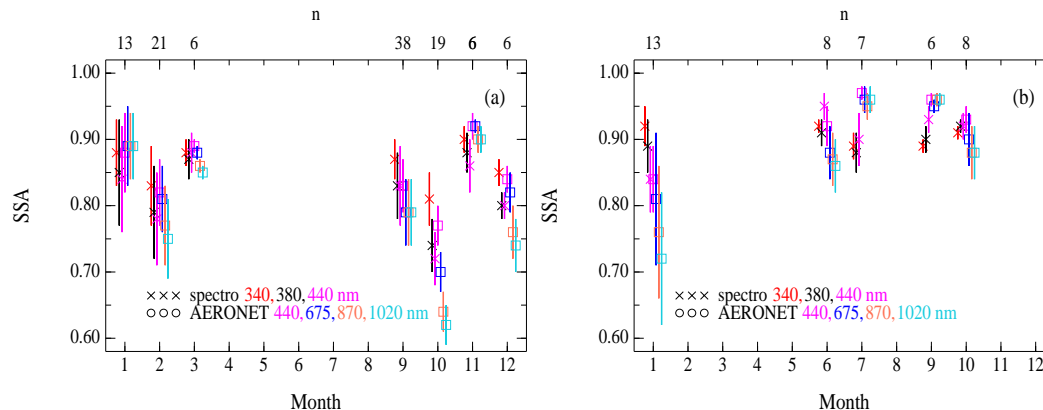


Fig. 7. Monthly means of SSA retrieved from spectroradiometer's measurements (crosses) at 340 (red), 380 (black) and 440 nm (magenta) and from AERONET/PHOTONS (circles) at 440 (magenta), 675 (blue), 875 (sienna) and 1020 nm (turquoise): **(a)** in 2003; **(b)** in 2006. The vertical bars represent the dispersions around the means and n , on the top, is the number of observations used to determinate the mean. For clarity the abscissa values of the various wavelengths have been slightly shifted around the month number.

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