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# Thermal infrared remote sensing of mineral dust over land and ocean: a spectral SVD based retrieval approach for IASI

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Discussion Paper

Discussion Paper

Discussion Pape

## **AMTD**

4, 461–509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.



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Discussion Paper

Discussion Paper

Printer-friendly Version

Interactive Discussion



# **Abstract**

From the high spectral resolution thermal infrared observations of the Infrared Atmospheric Sounding Interferometer (IASI) mineral dust AOD (transferred from thermal infrared to 0.5 µm) is retrieved using a Singular Vector Decomposition of brightness temperature spectra. As infrared retrieval based on 8–12 µm observations, dust observation with IASI is independent from solar illumination. Through the linear combinations of suitable independent singular vectors weighted by their contribution to the observed signal, and a projection of different a-priori dust spectra on the resulting signal the dust can be well distinguished from the influence of surface emissivity and gas absorption. In contrast to lookup-table based single-channel retrievals this method takes advantage of the spectral shape of dust extinction and surface and atmosphere influence over the total 8-12 µm window band. Using different a-priori spectra for dust extinction allows also for an estimation of dust particle size in terms of effective radius based on the respective dust model size distributions. These dust models are also used for the transfer of infrared AOD to 0.5 µm.

Four months of IASI observations covering Northern Africa and Arabia are used for evaluation. Two large scale dust events, one covering the Arabian Peninsula and adjacent parts of the Indian Ocean, the other over the Atlantic Ocean off the coast of West-Africa, are analysed and compared with other satellite images. They also show the good suitability of IASI data for dust observation at day and night. Monthly means derived from IASI observations represent well the known seasonal cycles of dust activity over Northern Africa and Arabia. IASI Dust AOD<sub>0.5 um</sub> and AERONET coarse mode  $AOD_{0.5\, \mu m}$  are reasonably well (linearly) correlated with  $\rho$ =0.655. Moreover, comparison of time series of AERONET and IASI observations 25 shows that the evolution of dust events is very well covered by the IASI observations. Rank correlation between dust effective radius and AERONET Angström exponent is -0.614 indicating the general capability of (qualitative) dust particle size information being provided by this method.

**AMTD** 

4, 461-509, 2011

## SVD based dust retrieval for IASI

L. Klüser et al.

Title Page Introduction **Abstract** Conclusions References **Figures** Tables  $\triangleright$ Back Close Full Screen / Esc

Discussion

Paper

Printer-friendly Version

Interactive Discussion



#### Introduction

Airborne mineral dust plays an important role in the climate system as it has several effects which directly or indirectly impact on the atmospheric state and climate. It reduces surface insolation by extinction of solar radiation (King et al., 1999; Slingo et al., 2006), which is of importance not only for climate and atmospheric processes, but also for solar energy applications, for which thus mineral dust information is important (Breitkreuz et al., 2007). By the same means, airborne dust can heat the atmospheric layer due to absorption of solar radiation (Shenk and Curran, 1974; King et al., 1999; Dunion and Velden, 2004). Moreover airborne mineral dust has effects on cloud microphysics (Mahowald and Kiehl, 2003; Ansmann et al., 2008; Klüser and Holzer-Popp, 2010) and precipitation (N'Tchavi Mbourou et al., 1997; Rosenfeld et al., 2001; Hui et al., 2008). When being transported over the Atlantic ocean, airborne dust from Northern Africa can also interact with tropical storm strengths (Dunion and Velden, 2004: Evan et al., 2006a: Jenkins et al., 2008) and is a major source of iron input into the ocean (Jickells et al., 2005; Mahowald et al., 2005) and for the Amazonia rainforest (Koren et al., 2006). Airborne dust activity follows diurnal and seasonal cycles and has a very high spatial variability over Northern Africa and Arabia (Engelstaedter et al., 2006).

Satellite remote sensing of dust is quite well established over dark surfaces such as ocean areas (King et al., 1999; Evan et al., 2006b) or vegetated surfaces (Kokhanovski and de Leeuw, 2009), where also aerosol type speciation is generally possible (Holzer-Popp et al., 2008). But over desert regions with bright reflecting surfaces many traditional aerosol retrieval algorithms cannot detect airborne dust. Efforts have been made to improve remote sensing capabilities also over bright surfaces, e.g., the Absorbing Aerosol Index AAI (Torres et al., 1998) or the "Deep Blue" Algorithm for the MODerate resolution Imaging Spectro-radiometer MODIS (Hsu et al., 2004, 2006). The Multi-angle Imaging SpectroRadiometer (MISR) provides the potential to use multiangle observations for aerosol retrievals, which also can be used to determine surface

**AMTD** 

4, 461-509, 2011

## SVD based dust retrieval for IASI

L. Klüser et al.

Title Page Introduction **Abstract** Conclusions References **Figures** Tables Back Close Full Screen / Esc

reflectance and thus to retrieve aerosol properties also over desert surfaces (Diner et al., 2005; Kahn et al., 2005; Kalashnikova et al., 2005). These algorithms can detect aerosol also over deserts, but they are not capable of a reliable aerosol type separation.

Due to large particle sizes and rather high absorption mineral dust can also be detected in the thermal infrared (TIR) range (Shenk and Curran, 1974; Ackerman, 1997). Different methods based either on temporal comparison of brightness temperatures (Legrand et al., 2001; Miller et al., 2008), on brightness temperature differences (Dunion and Velden, 2004; Chaboureau et al., 2007; Schepanski et al., 2007) or combinations of both (Li et al., 2007; Klüser and Schepanski, 2009) have been expolited successfully for dust remote sensing from geostationary satellites TIR instruments. For polar orbiting satellites a brightness temperature difference approach has been applied to MODIS observations over Asia by Zhang et al. (2006) for dust storm classification. Moreover the Atmospheric InfraRed Sounder AIRS provides a high number of spectral bands which is suitable for remote sensing of dust Aerosol Optical Depth (AOD), particle size and height information over ocean (Pierangelo et al., 2004, 2005; DeSouza-Machado et al., 2010; Peyridieu et al., 2010).

The Infrared Atmospheric Sounding Interferometer IASI flown on the METOP satellites has an even better spectral resolution than AIRS and also a better spatial resolution of observations. Thus this instrument is very well suited for remote sensing of airborne dust from thermal infrared observations.

Over land surface emissivity with its high variability for desert soils is one critical parameter for dust remote sensing with thermal infrared methods (Shenk and Curran, 1974; Ackerman, 1997; Ogawa et al., 2003; Pierangelo et al., 2004; Klüser and Schepanski, 2009). Thus some effort is needed to overcome the dependency on a (not yet existing) complete database of global surface emissivity spectra. Any approach of direct radiative transfer modelling for AOD retrieval in the TIR spectrum over land requires sufficient information (or assumptions) on surface emissivity, as otherwise the modelling of radiances will imply very high uncertainty due to background errors.

# **AMTD**

4, 461–509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ◀ ▶I

■ Back Close

Full Screen / Esc



One possibility of avoiding the problem of surface emissivity is using a Singular Vector Decomposition (SVD, e.g., Rodgers, 2000) of observed spectra in order extract useful independent pieces of information. Such an approach is used here for detecting the mineral dust signal within observed IASI spectra.

After a short overview over the basic theory behind singular vector decomposition in Sect. 2, the retrieval method is described in Sect. 3 whereas Sect. 4 presents results and evaluation. After a discussion of strengths and limitations of the retrieval method in Sect. 5 conclusions are drawn in Sect. 6.

## 2 Basic principles of singular vector decomposition

Principle component analysis (PCA) is a general term for methods, which aim at separating data fields into different uncorrelated signals, the principle components of the system. They can be used for reducing the dimensionality of a problem and for finding different uncorrelated modes (components) which describe the data (Rodgers, 2000).

PCA is a classical tool from statistical methods to explore the extraction of the uncorrelated and independent variables and concentrate attention only on main (principal) components which should be retrieved. Such an approach has advantages (in comparison with, for example, Factor-analysis) and is useful since it will give in an unbiased way both the number of parameters that can be retrieved and their interpretation.

Good overviews over the basic theory of PCA with respect to radiance spectra (as e.g., from IASI) are given in Matricardi (2010) as well as in Rodgers (2000). Classical principle component theory consists of solving the Eigenvalue problem of the covariance matrix of any data  $n \times m$  matrix  $\mathbf{X}$  formed by n data vectors (e.g., observations)  $\mathbf{x}$  of dimension m. If  $\mathbf{A}$  denotes the  $m \times m$  matrix formed by the m Eigenvectors (sorted in ascending order with respect to the magnitude of the corresponding Eigenvalues), the vector of principle components (or PC weights),  $\mathbf{y}$ 

**AMTD** 

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ✓ ▶I

Back Close

Printer-friendly Version

Full Screen / Esc



$$y = \mathbf{A}x \tag{1}$$

(Matricardi, 2010). From the knowledge of  $\bf A$  and the respective  $\bf y$  any observation  $\bf x$  can be reconstructed by

$$5 \quad \mathbf{X} = \mathbf{A}^{\mathsf{T}} \mathbf{y} \tag{2}$$

If the elements of x are correlated, a number k < m of Eigenvectors contained in A is sufficient to reconstruct x, hence a reduction of dimensionality is achieved (Matricardi, 2010). The possible magnitude of dimension reduction depends on the desired level of accuracy of the reconstructed field.

As  $\mathbf{A}x$  has a dimension different from that of  $\mathbf{y}$  in the case of a non-symmetric field, a Singular Vector Decomposition (SVD) can be regarded mathematically as solving an equivalent to the Eigenvalue problem of non-symmetric fields (Rodgers, 2000). Thus SVD can be seen as principle component analysis of the original data matrix instead of the covariance matrix. For a non-symmetric  $n \times m$  data matrix  $\mathbf{X}$ , a singular vector decomposition is solving the equation

$$\mathbf{X} = \mathbf{U} \mathbf{\Lambda} \mathbf{V}^{\mathsf{T}} \tag{3}$$

where  $\Lambda$  is the diagonal  $n \times n$  array of singular values  $\Lambda_i$ . **U** is the  $m \times m$  matrix of left-singular vectors and **V** is the  $n \times n$  matrix of right-singular vectors. The left- and the right-singular vectors are basis vectors for the observation space **M**, which includes, in particular, the observation data matrix **X**, respectively. The observation matrix **X** is normalized such that the euclidian norm of the singular vectors is unity and singular values are also normalized.

SVD is applied to the observation field directly and the corresponding Singular Vectors (SV) form a basis of the observation space – without any statistical transformation (like correlation or covariance) being necessary. Thus weights of the singular vectors (similar to the Eigenvectors in classical PCA) can directly be used to reconstruct the observation (spectrum) as a linear combination of the singular vectors.

AMTD

Discussion Paper

Discussion Paper

Discussion Paper

Discussion Paper

4, 461–509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I 

I 

Back Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



## Pre-processing of IASI data

Mineral dust is characterised by high extinction around 9.5 µm and much lower extinction at wavelengths around 8 µm and around 12 µm (e.g., Volz, 1973; Ackerman, 1997; DeSouza-Machado et al., 2010). As the spectral signal of dust extinction (largely due to absorption: Volz, 1973; Takashima and Masuda, 1987; DeSouza-Machado et al., 2010) is very characteristic in the 8–12 µm region (TIR window), the high spectral sampling of IASI can be exploited to avoid gas absorption lines in this region and to obtain mineral dust properties from the observed radiance spectra. In order to reach this goal, brightness temperature spectra are calculated from the IASI radiance measurements. The brightness temperatures between 8 µm and 12 µm are collected into 42 bins (equidistant in wavenumber space). The brightness temperature used for each of the 42 bins is represented by the maximum brightness temperature of the 10 respective channel values observed within each bin. This maximum approach is the first step to avoid narrow gas absorption lines which disturb the dust retrieval. The strongest signal of such a spectrum is assumed to be the temperature of the emitting body, e.g., the surface or clouds. Many gas absorption features can be eliminated by this way, but of course broad band features such as ozone absorption around 9.6 µm and continous water vapour absorption cannot be avoided and remain part of the brightness temperature spectra.

From the spectrum of binned maximum brightness temperatures the "baseline" temperature  $(T_{\text{base}})$  is determined as the maximum brightness temperature of the spectrum. As a first guess, this baseline temperature is assumed to represent the emission temperature of the emitting body (surface) in case of a constant emissivity spectrum (thus in the case of a blackbody  $T_{\text{base}}$  would equal the true temperature of the body) for low thermal infrared optical depths. In cases with high TIR optical depth this assumption is wrong as dust emission becomes a major contribution to the brightness temperature spectrum and the radiance extinction causes significant

**AMTD** 

# SVD based dust retrieval for IASI

4, 461-509, 2011

L. Klüser et al.

Title Page

**Abstract** Conclusions

Tables

References **Figures** 

Introduction













Printer-friendly Version

Interactive Discussion



467

Discussion Paper

Discussion Paper

Discussion Paper







underestimation of emission temperature of the surface. Thus the estimation of the emission temperature (baseline temperature) is updated with the TIR optical depth of the dust through one iteration cycle of the retrieval.

Prior to the SVD and retrieval some filtering of IASI observations is applied, reducing the number of IASI Fields-of-View (FOV) or pixels to be processed without a-priori dust flagging (as e.g., done in the AIRS dust retrieval described in DeSouza-Machado et al., 2010). This filtering selects observations which possibly contain dust information while observations very likely to show no dust are not treated further (see also Klüser and Schepanski, 2009 for a similar approach of data filtering). As first step a temperature threshold is applied for the baseline temperature  $T_{\text{base}}$ . Observations with  $T_{\text{base}}$  lower than 240 K are assumed to represent high level clouds and thus being inappropriate for dust retrieval. This threshold, being inspired by the ice cloud threshold of 238 K reported in Chylek et al. (2006), is somewhat arbitrary but visual inspection throughout the analysis period used here shows that at least in the Sahara-Arabia domain it is sufficiently low to avoid false masking of intense dust events. The Integrated Imaging Subsystem (IIS) of the IASI instrument is exploited to obtain the spatial homogeneity within one IASI pixel. In contrast to e.g., cumulus clouds, mineral dust is characterised by rather high spatial homogeneity (Evan et al., 2006b), moreover thermally very inhomogeneous conditions within one FOV can largely influence the radiance spectrum observed by the IASI instrument. Only if the variance of IIS brightness temperatures  $(\sigma^2(IIS))$  for the respective IASI spectrometer observation is sufficiently low (<4 K<sup>2</sup>, obtained from visual inspection of a large number of dusty and cloudy scenes), the pixel is passed on to the dust retrieval. Thus the initial filtering conditions are:

$$T_{\text{base}} \ge 240 \,\text{K}$$
 (4)

$$_{25} \quad \sigma^2(IIS) \le 4 \,\mathsf{K}^2 \tag{5}$$

In order to further reduce the numbers of FOVs to be processed without any likelihood of mineral dust, only scenes showing the typical spectral behaviour of mineral dust in the 10-12 µm region are used. An overview over a rather simple score test scheme

# **AMTD**

4, 461-509, 2011

## SVD based dust retrieval for IASI

L. Klüser et al.

Title Page **Abstract** Introduction Conclusions References **Figures** Tables  $\triangleright$ Back Close Full Screen / Esc

Printer-friendly Version

Interactive Discussion



applied here is provided in Table 1. The tests are directly motivated from the shape of the spectral extinction profile of airborne dust (Volz, 1973; Hess et al., 1998) and are weak enough to allow for additional surface emissivity effects (Ogawa et al., 2003) to also pass the tests. If the score (sum of all positive tests) is larger than 250 (very conservative), the FOV is passed on to further processing. In contrast to e.g., Evan et al. (2006b) or DeSouza-Machado et al. (2010) the scoring scheme does not represent a dust flag, i.e., FOVs with scores higher than 250 can also be dust free. This second filtering step is only to be regarded as a test which filters out FOVs which certainly do not contain dust signals.

Brightness temperature spectra having passed the initial filtering tests are transformed into a spectrum of a quantity  $\tau_{\rm eqv}(\lambda)$ , which can be regarded as an "equivalent optical depth". The calculation of  $\tau_{\rm eqv}$  follows the very simple approach of

$$L_{\text{obs}}(\lambda) = e^{-\tau_{\text{eqv}}(\lambda)/\cos(\Theta_{v})} B_{\lambda}(T_{\text{base}})$$
(6)

where  $L_{\rm obs}(\lambda)$  denotes the radiance obtained from the binned maximum brightness temperature spectrum and  $\Theta_{\rm v}$  is the viewing zenith angle (see e.g., Ackerman, 1997; Thomas et al., 2009).  $\tau_{\rm eqv}$  should not be interpreted as a real optical depth of the atmosphere, as it still contains significant pieces of information about surface emissivity, which is not spectrally constant in real nature. Also e.g., the ozone optical depth peaking around 9.6  $\mu$ m and the influence of the water wapour continuum contribute to  $\tau_{\rm eqv}$ . Equation (6) directly motivates the denomination of  $\tau_{\rm eqv}$  as "equivalent optical depth". In contrast to brightness temperature spectra the emission temperature is no longer the predominant signal in the analysed spectra, but these are still the result of a huge amount of influencing factors. Remaining clouds and atmospheric effects and surface emissivity variability for different surface types affect the observed spectra. The surface emissivity of bare soils in desert regions has a strong influence on the spectral variation, as can be seen e.g., from the results of Ogawa et al. (2003) and Hulley et al. (2009). Especially in the traditional "split-window"

AMTD

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

l≼ ≻l

Back Close

Full Screen / Esc

Printer-friendly Version



region between 10  $\mu$ m and 12  $\mu$ m the signals of desert soil emissivity and mineral dust extinction are quite similar.

There are several ways possible to solve the problem of separating the signals of dust, emissivity and other atmospheric constituents. The presented algorithm uses singular vector decomposition to extract different pieces of information from the observed spectra.

The SVD based dust retrieval does not rely on external cloud screening or dust flagging prior to the retrieval (which are both strong error sources, see e.g., discussions in Holzer-Popp et al., 2008; Klüser and Schepanski, 2009; DeSouza-Machado et al., 2010). As clouds are rather opaque in the TIR and also have high emissivity, no distinguishing between surface and clouds is necessary if the dust layer is above the clouds. Thus, in contrast to solar wavelength retrievals, dust observation above clouds is generally possible with this method.

# 3.2 Dust optical properties

AOD spectra for three different dust mixtures from the Optical Properties of Aerosol and Clouds (OPAC) database (Hess et al., 1998) between 8 μm and 12 μm are used for dust characterisation in this method. Only few other databases of TIR optical properties for mineral dust are available, e.g., from Volz (1973), Fouquart et al. (1987), Dufresne et al. (2002) or Sokolik et al. (1998), which moreover are (partly) based on each other. The general spectral shape of airborne mineral dust extinction in the TIR is rather similar between all of them. The OPAC dust mixtures used here are MINM, MIAM and MITR representing nucleation, accumulation and transportation modes of dust size distributions. The different spectral characteristics are the result of Mie calculations with identical refractive indices of mineral aerosol components after shifting the particle size distributions (Hess et al., 1998).

The OPAC dust type "mineral coarse mode" (MICM) is not used for retrieval here as the spectral shape especially in the  $10-12\,\mu m$  range is quite similar to the extinction

**AMTD** 

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ◀ ▶I

Full Screen / Esc

Close

Back

Interactive Discussion

Printer-friendly Version



Discussion Paper

Printer-friendly Version

Interactive Discussion



spectrum of water vapour in this region and thus the retrieval has shown to be sensitive to water vapour when including the MICM type. Also within the Sahara and the Arabian deserts there are regions (mainly sandbeds and dune fields, Pye and Tsoar, 2009) which have emissivity spectra very similar to the transmissivity calculated with the MICM dust type. Figure 1 (left) shows AOD spectra of MIAM, MINM, MITR and MICM from OPAC (all scaled to AOD<sub>10 um</sub>=0.25). It is evident that MITR, MIAM and MINM spectra all have spectral shapes with a major peak around 9.5-10 µm and minima at lower and higher wavelengths. This spectral shape of dust AOD is also reported reported in the spectral database of Volz (1973). MICM does not have such a strong characteristical dust peak, generally the spectral shape of MICM is much more different from the other dust types (and also from the observations) then those to each other. The more or less linear shape of the MICM AOD spectrum between 10 µm and 12 µm also explains the similarity to water vapour absorption and surface emissivity. Moreover in the spectral range between 8 µm and 12 µm no clear primary AOD peak is found which characterises the other three dust models as well as observations (Thomas et al., 2009) and other databases. Dufresne et al. (2002) show extinction spectra of mineral dust for different effective radii. Dust extinction spectra with several effective radii up to 3 µm have similar characteristics as the three OPAC models used here (significant primary extinction peaks around 10 µm) while the spectrum with effective radius 10 µm is similar to the MICM spectrum. Such large dust particle effective radii may be found in very thick dust layers near source regions. But the spectral shape of the extinction coefficient is not suitable for use in this kind of dust remote sensing application. Thus the MICM type, although potentially representing a major part of the coarse mode dust fraction, has been excluded from the IASI retrieval. Nevertheless although the MICM model is called "mineral coarse mode", also the other aerosol models include coarse mode contributions to the particle size distribution (e.g., Hess et al., 1998; Dufresne et al., 2002).

The shapes of the MIAM and MITR spectra are very similar (but have different gradients, thus provide different useful information) whereas the MINM spectrum

# **AMTD**

4, 461-509, 2011

## SVD based dust retrieval for IASI

L. Klüser et al.

Title Page Introduction **Abstract** Conclusions References **Figures** Tables Back Close Full Screen / Esc

is characterised by a significant shift of the primary AOD peak towards smaller wavelengths. Thus the differences between MINM, MIAM and MITR may lead to information about the particle size in the projection of the observed spectra. Unfortunately all OPAC aerosol classes assume monomodal lognormal-type number size distributions, whereas dust aerosol in nature often shows bi-modal size distributions (see e.g., Chen et al., 2010). Moreover the Mie calculations as used in OPAC may be insufficient in TIR due to non-shericity effects. Brindley and Ignatov (2006) report strong effects of non-sphericity for scattering of solar reflection by coarse mode dust, Meland et al. (2010) show the effect of non-sphericity on TIR optical properties. A fixed mineralogy is used for all dust models in OPAC, while Highwood et al. (2003) report a strong dependency of spectral refractive indices on mineralogy and a minor influence of the dust particle distribution. Nevertheless OPAC provides a widely used available database which represents well the observations of e.g., Thomas et al. (2009) and is also consistent with other TIR dust models.

Generally it is still quite difficult to obtain reliable datasets of (infrared) refractive indices of mineral dust due to the wide range of mineralogies, particle sizes and shapes (including nonspheric particles), external and internal aerosol mixtures, vertical structure of aerosol layers during campaigns and last but not least different methods to reconstruct refractive indices (Sokolik et al., 1993; Highwood et al., 2003; Meland et al., 2010).

Another general problem of available dust extinction models is e.g., described by Kalashnikova and Sokolik (2004) and Meland et al. (2010): serious errors in IR extinction as well as in light scattering may arise from shape distribution uncertainties. Thus also the transfer between solar wavelengths and TIR optical depth of the dust may be strongly sensitive to non-sphericity of the dust particles.

Within the European Space Agency's Climate Change Initiative – Aerosol (Aerosol CCI) project it is planned to provide dust refractive indices accounting for non-sphericity by T-matrix approaches for the near future. Given that newer (better) dust models including non-sphericity effects, other mineralogies or bi-modal particle size

**AMTD** 

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Full Screen / Esc

Close

Back

Interactive Discussion

Printer-friendly Version



distributions will become available in the future, they can easily be implemented in the retrieval method – either as additional dust types or replacing the OPAC model(s).

# 3.3 Singular vector decomposition of $au_{\sf eqv}$ -spectra

A number of days (11–17 March 2009) with sufficient dust activity over Northern Africa, the Arabian Peninsula and also the adjacent ocean regions (0° N–50° N and 30° W–70° E) has been selected for the singular vector decomposition of the IASI spectra, which have passed the initial filtering. Although it is not straightforward that such a temporal subset is generally transferable to other seasons, the large domain should provide some generality with respect to the covered region. Nevertheless tests with different subsets from different periods throughout the year 2009 showed no large variation of the characteristical features of the obtained singular vectors and resulting AOD, thus it is assumed that the singular vectors obtained from this period are generally applicable.

For the singular vector decomposition all suitable IASI spectra (having passed the filtering) are stored in the observation matrix  $\mathbf{X}$  and the singular vector decomposition (Eq. 3) leads to a set of n=42 singular vectors  $\mathbf{v}_{i \le n}$  (right singular vectors in Eq. 3, forming a basis of the spectral space).

No differentiation is made between land and ocean observations in order to obtain absolutely the same retrieval over both surfaces. Also no discrimination between daylight and night observations is done, as radiance in the thermal infrared window is (almost) independent of solar illumination. The leading five basis vectors are depicted in Fig. 1 (right). Not surprisingly none of them can be attributed to the dust extinction spectra alone but all include partial impacts of surface emissivity and atmospheric effects.

## **AMTD**

4, 461–509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I 

I 

I 

Back Close

Full Screen / Esc



Printer-friendly Version

Discussion Paper

Discussion Paper

Discussion Paper

Discussion

# SVD based dust retrieval for IASI

**AMTD** 

L. Klüser et al.

where *i* denotes the number of the corresponding singular value.

Dust AOD at 10  $\mu m$  is now retrieved by projection of appropriate singular vector representation of  $\tau_{eqv}$  onto the three OPAC based dust types. The strongest impacts on the observed brightness temperature spectra are those of surface emissivity and gas absorption (by ozone and water vapour). As the first two singular vectors represent mainly those signals (with only small contributions of other atmospheric constituents as e.g., mineral dust), they are not used for dust retrieval. The dust signal is mainly carried by the singular vectors three to five. Only spectral ranges between 8  $\mu$ m and 9  $\mu$ m and between 10  $\mu$ m and 12  $\mu$ m are used for the weight calculation and AOD projection to avoid the spectral regions of strongest interference of ozone and surface.

avoid the spectral regions of strongest interference of ozone and surface.

The dependence of traditional TIR retrievals on dust layer height is mainly a result of different spectral sensitivity to atmospheric layers rather than to emission by the dust (see e.g., Pierangelo et al., 2004; DeSouza-Machado et al., 2010). It can be assumed that the spectral dependence of IASI observations on dust layer height is already part of the singular vectors, thus no height-dependent correction factors are necessary as e.g., applied by DeSouza-Machado et al. (2010).

Although this approach without sophisticated radiative transfer modelling is very simple, it leads to reasonable results as shown in Sect. 4. The main advantage is avoiding the necessity of determining the surface emissivity spectrum. As well Pierangelo et al. (2004) as also DeSouza-Machado et al. (2010) state that the main

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ■ ▶I

■ ▶I

Back Close

Full Screen / Esc

(C) (I)

Interactive Discussion

reason for the low applicability of those dust AOD retrievals over (arid) land is the huge problem with surface emissivity spectra characterisation.

In the first retrieval step the singular vector (SV) three to five based representation of the dust related equivalent optical depth spectrum is

$$\tau_{sv}(\lambda) = \sum_{j=3}^{5} w_j \cdot v_j(\lambda)$$
 (8)

and is used for dust AOD calculation.

For each of the OPAC types a projection of  $\tau_{\rm sv}(\lambda)$  onto a respective AOD spectrum with AOD<sub>10 µm</sub>=1 is calculated as the scalar product between the singular vector based spectrum and the corresponding OPAC spectrum:

10 
$$AOD_{type} = \sum_{\lambda} [AOD_{opac,type}(\lambda) \cdot \tau_{sv}(\lambda)]$$
 (9)

The inverse weight  $\varphi_{\mathrm{type}}$  of each OPAC type in the overall retrieval result is obtained by

$$\varphi_{\text{type}}^{-1} = \text{acos}\left(\frac{\text{AOD}_{\text{type}}}{\sqrt{\sum_{\lambda} \left(\text{AOD}_{\text{opac,type}}(\lambda)\right)^2} \sqrt{\sum_{\lambda} (\tau_{\text{sv}}(\lambda))^2}}\right)$$
(10)

which in the classical representation of the scalar product in Euclidian space would correspond to the angle between two vectors. Thus a higher value of  $\varphi_{\text{type}}$  results in a better representation of the observed aerosol by the respective dust model. The weights are normalised by the sum of all three weights leading to the fraction of the respective type within the dust representation.

As only the singular vectors three to five are used in this step, the vast majority of the strong influence of surface emissivity, contained in singular vectors one and two, is eliminated. Nevertheless the two first singular vectors also contain dust signals which are needed for a complete estimation of the dust AOD in the 8–12 µm band and thus

#### **AMTD**

4, 461-509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ← ►I

← ► Back Close

Full Screen / Esc

Printer-friendly Version



The OPAC dust spectrum (AOD<sub>opac</sub>) and the representation of the first two singular vectors ( $\tau_{[1,2]}$ ) are calculated as

$$_{5} \quad \mathsf{AOD}_{\mathsf{opac}}(\lambda) = \sum_{\mathsf{type}} \left[ \varphi_{\mathsf{type}} \cdot \mathsf{AOD}_{\mathsf{opac},\mathsf{type}}(\lambda) \right] \tag{11}$$

$$\tau_{[1,2]}(\lambda) = w_1 v_1(\lambda) + w_2 v_2(\lambda) \tag{12}$$

Given the standardized (unweighted) projection of  $\{v_3, v_4, v_5\}$  onto  $\{v_1, v_2\}$ , denoted as  $P_w$ , and the signed squared linear correlation  $s_{\text{var}}$  (explained variance) between  $AOD_{\text{opac}}$  and  $\tau_{[1,2]}$ , the impact of the first two singular vectors on the TIR AOD retrieval based on singular vectors three to five is calculated as

$$c_{[1,2]} = s_{\text{var}} \cdot \sum_{\text{type}} \left( \varphi_{\text{type}} \cdot P_w \cdot \text{AOD}_{\text{type}} \right)$$
 (13)

This correction term (being mainly positive over oceans and negative over desert soils) is added to the resulting 10  $\mu$ m preliminary optical depth AOD<sub>10  $\mu$ m</sub> calculated with

$$AOD_{10 \,\mu\text{m}} = \sum_{\text{type}} \left[ \varphi_{\text{type}} \cdot AOD_{\text{type}} \right] + c_{[1,2]} \tag{14}$$

For each observation where Eq. (14) leads to a positive result, from AOD  $_{10\,\mu m}$  the respective OPAC based spectral AOD variation

$$AOD_{\lambda} = \sum_{\text{type}} AOD_{10 \, \mu\text{m}} \cdot \varphi_{\text{type}} AOD_{\text{opac,type}}(\lambda)$$
 (15)

is calculated for all  $\lambda$  (as the OPAC spectra are scaled to  $AOD_{10\,\mu m}=1$ ).  $AOD_{\lambda}$  is then used for retrieving the equivalent emission brightness temperature of the dust (being a rough indicator for dust layer effective temperature respective height) and for temperature correction of the AOD retrieval. First the dust layer equivalent brightness

4, 461–509, 2011

# SVD based dust retrieval for IASI

**AMTD** 

L. Klüser et al.

Title Page

Introduction

.

Conclusions References

**Abstract** 

Tables Figures





Full Screen / Esc

Printer-friendly Version

Interactive Discussion



scussion Paper

Discussion Paper

Back

Interactive Discussion

temperature is calculated by inversion of Eq. (6) with a simple dust emission extension (Takashima and Masuda, 1987; Ackerman, 1997; Zhang et al., 2006):

$$L_{\uparrow}(\lambda) = e^{-AOD_{\lambda}}B_{\lambda}(T_{\text{surface}}) + \left(1 - e^{-AOD_{\lambda}}\right)\epsilon_{\text{dust}}B_{\lambda}(T_{\text{dust}})$$
(16)

As  $T_{\text{surface}}$  remains unknown,  $T_{\text{base}}$  is used instead as initial guess. In one iteration  $_{5}$  of the retrieval  $T_{\rm base}$  is recalculated with the respective retrieved value of AOD<sub>10 um</sub> leading to a more realistic effective surface emission temperature. With dust emissivity values  $e_{dust}$  as reported by Takashima and Masuda (1987) for different size distributions in the 10.5-11.5 µm band, this leads to

$$B_{\lambda}(T_{\text{dust}}) = \frac{1}{\epsilon_{\text{dust}}} B_{\lambda}(T_{\text{base}}) \frac{e^{-\tau_{\text{eqv}}(\lambda)} - e^{-\text{AOD}_{\lambda}}}{1 - e^{-\text{AOD}_{\lambda}}}$$
(17)

This inversion is done for seven spectral bins around 10.5 µm, with the dust layer effective emission temperature being the mean value of the obtained results. After retrieval of  $T_{dust}$  the observed radiance is "corrected" for thermal emission of the dust (with Eqs. 6, 7 and 16) and a second iteration of Eqs. (6) to (17) is performed leading to the final values for AOD<sub>10  $\mu$ m</sub> and  $T_{dust}$ .

With the weighted OPAC dust representations also a transfer coefficient between 10  $\mu m$  and 0.5  $\mu m$  ( $\gamma_{type}$ ) is calculated from OPAC for each observation, thus (coarse mode) dust contribution to AOD<sub>0.5 µm</sub> can also be provided as a retrieval result

$$AOD_{0.5 \,\mu\text{m}} = \left(\sum_{\text{type}} \left[\varphi_{\text{type}} \cdot \gamma_{\text{type}}\right]\right) \cdot AOD_{10 \,\mu\text{m}} \tag{18}$$

The weights of the different OPAC dust types and the OPAC based dust particle size distribution  $(n_r(type))$  lead to the estimated size distribution

$$n_r = \sum_{\text{type}} [\varphi_{\text{type}} \cdot n_r(\text{type})] \tag{19}$$

# **AMTD**

4, 461-509, 2011

## SVD based dust retrieval for IASI

L. Klüser et al.

Title Page **Abstract** Introduction Conclusions References **Figures** 

Close

From this size distribution the effective radius of the dust particles,

$$R_{\text{eff}} = \frac{\int n_r(r)r^3 dr}{\int n_r(r)r^2 dr}$$
 (20)

can directly be obtained.

The intrinsic uncertainty, which can be characterised theoretically, is composed of three parts. One part is the uncertainty which is introduced by the selection of three different fixed OPAC types and the respective weighting. This uncertainty,  $\delta_{\text{type}}$ , is calculated as the spread between maximum and minimum retrieved AOD<sub>type</sub> for the three different OPAC models, in percent of AOD<sub>10 um</sub>. The second internal error source is the temperature correction. This error  $(\delta_{\text{temp}})$  is characterised by the product of the difference (absolute value) between AOD<sub>10 um</sub> before and after the correction and the standard deviation of the seven different values for  $T_{
m dust}$  obtained from seven wavelengths. Also the correction of the 10 µm AOD for the contribution of the first two singular vectors is an error source ( $\delta_{corr}$ ). It is characterised by the fraction of the absolute value of the correction term and the sum of original prior retrieval result and correction term.

The intrinsic retrieval uncertainty then is estimated in percent as

$$\delta AOD_{10 \, \mu m} = 100\% \cdot \frac{\sqrt{(\delta_{type}^2 + \delta_{temp}^2 + \delta_{corr}^2)}}{AOD_{10 \, \mu m}}$$
 (21)

This value gives the percentage of result uncertainty due to the constraints of the retrieval mechanism (OPAC dust models and temperature correction) because of the assumptions being made. It is not to be regarded as the absolute retrieval uncertainty as obtained from validation with independent data. Nevertheless it is an important quality measure intrinsic to the retrieval scheme which allows observations with too high internal retrieval uncertainty ( $\delta AOD_{10 \, \mu m} > 50\%$ ) to be rejected.

**AMTD** 

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page Introduction **Abstract** 

Conclusions References

> **Figures** Tables

 $\triangleright$ 

Back Close

Full Screen / Esc

Printer-friendly Version



Discussion Paper

Discussion Paper

Discussion Paper

# **AMTD**

4, 461-509, 2011

# SVD based dust

L. Klüser et al.

# retrieval for IASI

# Title Page Introduction **Abstract** Conclusions References **Figures** Tables Back Close Full Screen / Esc

Printer-friendly Version

Interactive Discussion



#### Monitoring of large scale dust events

The evolution of large scale dust events in the Sahel-Soudan belt and over Arabia is presented in Fig. 2 for four consecutive days in February 2009. Retrieved Dust 5 AOD<sub>0.5 µm</sub> is plotted in colours whereas the greyscale background represents mean IIS temperature of the respective IASI observation (range from 220 K to 320 K. inverted).

The observations show morning overpasses of 11-14 February 2009. 11 February dust uplift behind a cold front over Kuwait and Central Saudi Arabia can be observed as well as some dust activation and transport in the Sahel. The Arabian dust plume propagates further South and reaches the Jemen coast line on 12 February as well as the airborne dust covers large parts of the Persian Gulf. The dust plume has decoupled from the cold front with a Southward propagation which is faster than that of the cloud band (for decoupling of dust fronts from synoptic disturbances see e.g., Miller et al., 2008). The Sahelian dust pattern has weakened to some degree and dust transport towards the Atlantic Ocean is observed. On 13 February the dust front over Arabia has reached the Indian ocean and is re-strengthened along the Red Sea coast of Saudi Arabia. The Arabian inland dust uplift breaks down between 13 and 14 February. The Sahelian dust plume is still active over Mali and the Atlantic Ocean.

Figure 3 shows RGB composites of the Advanced Very High Resolution Radiometer (AVHRR) operated on the same spacecraft ([R,G,B]=channels [1.6 μm, 0.8 μm, 0.6µm]) and IASI dust effective radius for 12 and 13 February daytime overpasses over Arabia and Northern Africa. On 12 February dust is visible in the AVHRR images especially over the Southern part of the Arabian Peninsula, over the Persian Gulf and in the Southern Sahel belt North of the deep convective cloud cells of the ITCZ, which very well compares to the respective AOD patterns observed with IASI. The dust plume over the Atlantic ocean has rather low AOD (Fig. 2) and thus is not really evident from the AVHRR imagery. Also the dust over the Sahara cannot be detected in the AVHRR

Discussion Paper

Printer-friendly Version

Interactive Discussion



image due to the bright underlying surface. The retrieved effective radius is rather large over the Arabian Peninsula and also slightly increased over the Persion Gulf. Over West-Africa high effective radius is observed in the Sahel near to the Sahara, while it is much lower further away from the source regions and over the Atlantic ocean (where also AOD is much lower).

On 13 February the South-Eastward propagation of the dust plume over the Indian Ocean is clearly visible from AVHRR, the pattern of higher reflectance over the ocean well fits the pattern of dust observed from IASI. Also the band of high dust AOD obtained from IASI over Western Saudi Arabia can be detected as high reflectance in the AVHRR composite image as well as the dust transport over the Southern part of the Red Sea. The dust pattern in the Sahel can be seen in the AVHRR images only very weakly, which is not very surprising as the AOD obtained from IASI are smaller than on the day before. Highest effective radii are observed in Western Arabia and over the Red Sea close to the Arabian coast, while effective radius is much lower over the Indian Ocean. Over the Sahel-Sahara source region the effective radius is slightly reduced compared to the day before and again the effective radius decreases with increasing distance to the dust source region. Lowest effective radii are observed over the Atlantic Ocean together with rather low AOD.

Figure 4 shows the evolution of a large-scale dust plume emerging from the Western Sahara towards the Atlantic Ocean from 14 through 16 September 2010 with twice daily IASI observations. The dust is emitted in the Western Sahara dust source region in the afternoon of 14 September and is first seen in IASI imagery in the ascending orbits of the same day. Near the source regions Dust AOD reaches very high values of more than 1.5 and the dust first reaches the Atlantic Ocean at the Mauretanian coast between 20°N and 25°N (14 September, ascending orbits). On 15 September the plume propagates further West and also expands towards the South (the main plume now crosses the coast south of 20° N). During the course of the day the dust plume also reaches the outer cloud bands of the hurricane "Julia" being active on the Eastern tropical North Atlantic Ocean. Inland dust transport towards the Atlantic Ocean still

## **AMTD**

4, 461-509, 2011

## SVD based dust retrieval for IASI

L. Klüser et al.

Title Page Introduction **Abstract** Conclusions References **Figures** Tables Back Close Full Screen / Esc

continues throughout 15 September. In the morning hours of 16 September, the dust plume over the Ocean has detached from the inland dust and travelled further West, while another dust plume is being formed near the Western Sahara source region. The Atlantic Ocean dust plume can still be seen to be in close contact with the cloud systems of hurricane "Julia".

The evolution of the dust plume can also be followed well in the AVHRR day-time composite images in Fig. 5. It is evident from the comparison of IASI AOD and AVHRR images that dust observation over (shallow cumulus) clouds is generally possible with this method. Only partly no dust above clouds is retrieved which is the result of the spatial homogeneity test not being passed if the cloud field is too inhomogeneous. Moreover the effective radius retrieved by IASI is depicted exemplarily for 15 September descending (daylight) and ascending (night time) orbits. From the effective radius retrieval it is evident that largest dust particle sizes are detected inland very close to the source region, while over the Atlantic Ocean particle sizes get much smaller.

## 4.2 Evaluation against AERONET and sensitivity analysis

Coarse mode AOD observations from the AErosol RObotic NETwork (AERONET, Holben et al., 1998) processed with the Spectral Deconvolution Algoritm (SDA, O'Neill et al., 2003) have been used for comparison with IASI derived dust AOD. One should be aware that in the SDA algorithm coarse mode AOD is somewhat sensitive to the presence of thin clouds and thus may be overestimated (O'Neill et al., 2003). All available AERONET observations over Northern hemispheric Africa, Southern Europe and the Arabian Peninsula (SDA available in 2009 for 23 stations, 250 coincidences over four months representing different seasons) are used for overpasses within a spatial distance of the observations of 50 km (to the center of the 12 km IASI FOV point-spread-function). All respective AERONET observations within half an hour deviation of the overpass are averaged and their temporal standard deviation is calculated.

AMTD

4, 461–509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ✓ ▶I

✓ ▶I

Printer-friendly Version

Full Screen / Esc

Close

Back



Figure 6 (left) presents the resulting scatter plot between IASI derived Dust AOD<sub>0.5 um</sub> and corresponding coarse mode AOD observed from AERONET. Colour coding of the symbols represents the Ångström exponent  $\alpha$  of the AERONET observations (total AOD, not only coarse mode), defined by

$$\frac{\text{AOD}_{0.44 \,\mu\text{m}}}{\text{AOD}_{0.87 \,\mu\text{m}}} = \left(\frac{0.44 \,\mu\text{m}}{0.87 \,\mu\text{m}}\right)^{\alpha} \tag{22}$$

 $\alpha$  is used quite often as indicator of aerosol particle size for mono-modal size distributions and for dust separation (e.g., Dubovik et al., 2002; Klüser and Schepanski, 2009); small  $\alpha$  indicates large particle sizes. It is obvious that for most observations with co-located IASI dust observations  $\alpha$  is rather small, thus in the majority of observations the aerosol is represented mainly by the coarse mode. Nevertheless there are also some observations with rather high  $\alpha$ , for which both IASI and AERONET coarse mode AOD are rather low.

The linear correlation between AERONET observations and IASI retrieved dust AOD is  $\rho$ =0.655. There is a negative bias of -0.15, which could be expected as TIR observations are mainly sensitive to coarse mode particles (Pierangelo et al., 2005) and dust particle number size distributions normally are bi-modal in contrast to the mono-modal assumption used here (Chen et al., 2010). 68% of all IASI derived AOD values are within a band of coarse mode AOD<sub>AFRONET</sub>±0.2. The Root Mean Square Deviation (RMSD) is 0.28. Although overall performance is quite reasonable compared with AERONET (given the differences in wavelengths and observation methods as well as a potential thin cloud contamination of AERONET coarse mode AOD in some cases), it is obvious that for very high AOD the IASI retrieval underestimates AOD at 0.5 µm, which largely explains the rather high RMSD. These underestimations are mainly from the stations Kuwait University (in May, no data available in February), Dhadnah and Eilat.

The comparison between IASI  $R_{\rm eff}$  and  $\alpha_{\rm AFRONET}$  (Fig. 6, right) indicates that high effective radii observed by IASI correspond to rather small Angström exponents

**AMTD** 

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page **Abstract** Introduction Conclusions References **Figures** Tables

Back Close

 $\triangleright$ 

Full Screen / Esc

Printer-friendly Version



observed by AERONET. For smaller  $R_{\rm eff}$  the range of AERONET Ångström exponents is much higher, but generally these also are higher, indicating smaller particle sizes. The linear correlation between both is  $\rho$ =-0.54, but as the comparison clearly indicates (and could be expected from theory), the relationship between  $R_{\rm eff}$  and  $\alpha$  is not linear. The (non-linear, non-Gaussian) rank correlation coefficient (Spearman, 1908) is  $\rho_{\rm rank}$ =-0.61, thus has slightly higher magnitude.

Figure 7 shows time series of AERONET and IASI AOD at the stations of Capo Verde (Cape-Verde Islands), Solar Village (Saudi Arabia), Tamanrassset (Algeria) and Dhadnah (United Arabian Emirates) for the months February, May and August 2010. Although also November 2010 has been processed and used here, none of the four presented stations had coincidences of AERONET and IASI dust observations in this month due to the generally very low dust activity in November, thus the respective November time series are not shown. It is evident that in most cases the transition of dust events is well represented by the IASI observations while in Solar Village and partly in Dhadnah AOD is generally underestimated by IASI in cases with high AOD observed by AERONET. In Tamanrasset two events with very high AOD in May are missed in the IASI observations, partly due to overlying cirrus clouds at the time of the IASI observations (not shown). In Dhadnah and Tamanrasset cases with rather low AERONET coarse mode AOD and overestimation by IASI are also present. From the analysis it is clearly evident that in cases where elevated coarse mode AOD (assumed to represent mineral dust at the selected stations) is observed by AERONET the IASI retrieval also detects mineral dust. Only in Capo Verde, an island location, there are much more coarse mode observations from AERONET than from IASI, which might be the result of airborne sea salt aerosol (which is also coarse mode aerosol). Time series of the other stations show very similar results but are not shown here.

For a sensitivity analysis the deviation of the IASI dust retrieval from AERONET observations

 $\delta AOD_{(IASI,AERONET)} = AOD_{IASI} - AOD_{AERONET}$  (23)

AMTD

4, 461-509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I 

I 

I 

Back Close

Full Screen / Esc

Printer-friendly Version



is used. Its correlation to the variable to be tested is calculated with both the linear and rank correlation. The results are given in Table 2, where  $\Delta d$  denotes the spatial distance of the FOV center to the AERONET station in km while  $\Delta t$  is the temporal deviation (in seconds) between the IASI observation and the median (with respect to time) of the AERONET observations. The anti-correlation between  $\delta AOD_{(IASI AERONET)}$ and IASI AOD (with magnitude of about 0.25) again indicates the underestimation of high AOD by the IASI retrieval. The highest correlation magnitude (with negative correlation coefficients) is observed for the retrieved effective radius, indicating that higher errors are related to lower  $R_{\rm eff}$  (as also seen in Fig. 6, right). Temporal and spatial deviations have hardly any influence on the retrieval error as is indicated by correlations very close to zero. The same is observed for the IIS temperature variance (as long as the IIS variance threshold is applied), the retrieved dust layer emission temperature and the viewing zenith angle. The weak correlations between relative error and  $T_{dust}$  can be regarded as being indicative that the dust layer height does not systematically affect the AOD retrieval (see e.g., the discussion in DeSouza-Machado et al., 2010 for height correction methodology in other TIR sounder retrievals).

# 4.3 Sampling the seasonal cycle of airborne dust

Monthly means of IASI daytime dust AOD are presented in Fig. 8 for February, May, August and November 2009, exemplarily. In February especially the high dust loads over Iraq, Kuwait and the Persian Gulf as well as over large parts of the Arabian Peninsula and the neighbouring Indian Ocean are evident. They are related to the high dust emissions connected with the Northerly Shamal wind (Laity, 2008) of the winter season. The Shamal is also responsible for dust transport towards the Indian Ocean and across the Red Sea, which is is evident from rather high AOD over these regions. Sahelian dust and transport towards the Atlantic Ocean ist the second main pattern evident from IASI AOD observations in February 2009. These are connected to the North-Easterly Harmattan wind (Laity, 2008; Pye and Tsoar, 2009; Klüser and Holzer-Popp, 2010) which activates dust throughout the Sahel belt. One clearly evident

AMTD

4, 461–509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

Back Close
Full Screen / Esc

Printer-friendly Version



dust hot spot is the Bodélé Depression in Northern Chad, often referred to as "the most dusty place in the world" (Koren et al., 2006), and neighbouring downstream regions.

In May 2009 high dust activity can still be observed in the Bodélé Depression, but much less airborne dust is found in downstream Sahel, compared to February. Being the transition period between winter (dry season) and summer (monsoon season), also the respective dust sources have transition activity, thus the weakening of the Sahelian dust activity while in the Western Sahara the dust source regions already start being active. The Atlantic Ocean dust plume becomes stronger and is found further North than in February, while dust activity over the Arabian Peninsula and the Red Sea weakens and is more confined to some hot spot areas.

August monthly mean dust AOD is also high over the Persian gulf, the Red Sea, parts of the Arabian Peninsula and the adjacent Indian Ocean. The dust uplift and transport in this part of the world in summer is connected to the Persian Gulf trough and the tightly connected Etesian winds, which promote a cyclonic circulation around the Persian Gulf and a Westerly transport onto the Indian Ocean (Pye and Tsoar, 2009). The second well known dust hot spot in August is the large Western Saharan dust source region (e.g., Engelstaedter et al., 2006; Klüser and Schepanski, 2009; Knippertz and Todd, 2010) and the strong advection towards the Atlantic Ocean.

In November, after the retreat of the West-African Monsoon (e.g., Klüser and Holzer-Popp, 2010), dust activity is very low throughout the whole observation area, some larger scale dust activity can be observed in the Sahel and North of the Persian Gulf.

#### 5 Discussion

The retrieval method for IASI presented here is based on Singular Vector Decomposition of equivalent optical depth spectra rather than on radiative transfer modelling and look-up tables (LUTs). This approach has some advantages but of course also a couple of disadvantages. The high variability of spectral surface emissivity of desert soils together with the evident lack of knowledge about its

AMTD

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I4 ►I

Back Close

Full Screen / Esc

Printer-friendly Version



spatial distributions (no reliable spectral emissivity maps available) complicate every radiative transfer modelling based approach of dust remote sensing in the TIR over deserts. Either one tries to retrieve emissivity and dust properties simultaneously (computationally very expensive and not always possible), or one has to use other approaches than radiative transfer modelling, if reliable results are desired also for arid land regions. One such approach is the singular vector decomposition and the accompanying projection scheme used here.

Another advantage of this approach is that it is possible to avoid cloud masking prior to the dust retrieval. This is not only computationally costly, but it is also a large source of errors especially in the TIR, as the separation between dust and stratocumulus clouds or fog is not always straightforward. Moreover the SVD based retrieval is also capable of retrieving dust information above low clouds beneath the dust layer. This would not be possible with reliable a-priori cloud screening.

Another advantage of thermal infrared observations in the  $8-12\,\mu m$  is the independence from sunlight, so that dust observation is possible twice daily with the SVD based approach from IASI. All three events presented here strongly indicate that there is high consistancy of morning and afternoon observations which allows to twice daily sample the evolution of airborne dust.

Dust AOD retrieved from IASI correlates quite well with AERONET coarse mode AOD given the asumptions and constraints of the retrieval and the fact that IASI is only sensitive to the coarse mode dust particles being responsible for considerable extinction in the  $8-12\,\mu m$  band.

The time series analysis of the four AERONET stations show exemplarily that the transition of dust events is well captured by IASI observations throughout all seasons. Although IASI observations tend to underestimate  $\mathsf{AOD}_{0.5\,\mu m}$  in scenes with high AERONET AOD, the ability of IASI dust detection in cases of elevated coarse mode  $\mathsf{AOD}_{0.5\,\mu m}$  is clearly evident. The  $\mathsf{AOD}_{0.5\,\mu m}$  underestimation is higher in the Solar Village time series than e.g., in Tamarasset or Cabo Verde. The highest underestimation is found for Kuwait University in May.

**AMTD** 

4, 461–509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ◀ ▶I

■ Back Close

Full Screen / Esc



Printer-friendly Version

The stations with most significant  $AOD_{0.5\,\mu m}$  underestimations (Kuwait University, Dhadnah, Eilat) are close to the sea, the Kuwait University station moreover is located directly inside Kuwait City. Thus coarse mode AOD contributions of other aerosols than dust, especially from near-surface particulate matter or sea-salt, cannot be excluded. Moreover all three stations are in environments of very inhomogeneous surface types (urban, desert, ocean), which might also impact on the strength of the dust signal in the IASI observations. Thus for these stations it is not very surprising that within a 50 km circle partly large deviations from AERONET observations occur.

Moreover, also AERONET coarse mode AOD retrieved with the SDA algorithm has some limitations and weaknesses. O'Neill et al. (2003) state that for coarse mode aerosol there is some sensitivity to thin homogeneous cloud contamination. While they report that given this sensitivity the fine mode AOD is not much affected by this problem, this is not true for the coarse mode AOD which has been used here. Although it is quite unlikely that cirrus is present in IASI observations due to the strong thermal signal of these clouds, they may well be present in AERONET observations within the 50 km distance and 0.5 h deviation from the IASI observations. The underestimation of AOD by IASI compared to AERONET coarse mode AOD might also point into this direction.

Although no explicit radiative transfer modelling is performed and the retrieval is not based on surface emissivity assumptions or atmospheric temperature and humidity profiles, some a-priori information is needed which has to be assumed to be generally valid. One of the a-priori assumptions is that the singular vectors derived from a set of days during March 2009 is transferable to other regions, seasons or years. As the singular vector decomposition has been tested with different periods and no large differences between singular vectors has been found, it can be concluded that the region covered by the singular vector decomposition performed here is sufficiently large, including sufficiently many different atmosphere and surface states, that the transferability is really present. It cannot be concluded from these results that this would be also the case for a transfer to different regions like e.g., East-Asia.

AMTD

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ◀ ▶I

Full Screen / Esc

Close

Back

Printer-friendly Version



Paper

Back

Interactive Discussion



Another a-priori assumption of the retrieval mechanism is the validity of OPAC extinction spectra. There is evidence that AOD spectra observed in the Sahara compares quite well with OPAC in the TIR, but as OPAC assumed log-normal size distributions it is not clear if the transfer to AOD<sub>0.5 um</sub> can be done more precisely with a better dust aerosol representation. The same might be the case for including non-sphericity effects in the dust model. As so far no suitable dust representation including bi-modal size distributions and non-sphericity is available for remote sensing applications, the choice of three (out of four possible) different OPAC models remains the preferred dust representation. Although IR methods are mainly sensitive to the coarse mode portion of the total aerosol burden (Pierangelo et al., 2004), the coarse mode mineral dust class of OPAC is not used as its spectral variability in 8-12 µm is very similar to signals of water vapour and surface emissivity. Thus using that class would increase the "false alarm rate" drastically. The comparison to AERONET and the possibility to follow the evolution of the dust plumes in the IASI observations shows that the OPAC assumption, while not being perfect, performs reasonably well. Lacking the coarse mode dust size distribution representation, the three OPAC dust types used here may not represent the full range of airborne mineral dust and the retrieval of the effective radius might suffer from that limitation. If the OPAC models also perform reasonably well in terms of particle size will be subject to future work focussing on evaluation of the effective radius estimation within the IASI dust retrieval.

In some cases rather high Ångström exponent  $\alpha$  is observed, indicating that the majority of the total aerosol burden of these observations is not of the coarse mode. For these observations both, IASI AOD and AERONET coarse mode AOD are rather small. Thus there is evidence that it is really only the dust fraction of the total aerosol burden, which is observed by IASI with the presented method. This conclusion is supported by the results of a comparison of IASI AOD with AERONET total AOD (not shown). For total AOD the evaluation results get much worse with a linear correlation coefficient  $\rho$ =0.584, RMSD=0.469 and bias=-0.338. Thus compared to total AERONET AOD there is a very strong underestimation of AOD by IASI, which can easily be attributed to

AMTD

4, 461–509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Close

Discussion

Printer-friendly Version

Interactive Discussion



the fine mode fraction of the aerosol. Moreover O'Neill et al. (2003) report a sensitivity of the AERONET coarse mode AOD retrieved with the SDA algorithm to thin clouds, which could result in an overestimation of coarse mode AOD. Thus this could be another error source being responsible for the large differences between IASI Dust 5 AOD and AERONET coarse mode AOD.

The AOD retrieval is done as a statistical weighting of the projections of linear combinations of suitable singular vectors on different dust models (from OPAC). The main difference of these dust models is their particle size distribution (Hess et al., Thus this approach automatically leads to an estimation of dust particle size, expressed in terms of effective radius in this retrieval. But as no evaluation of the effective radius has been presented (and can be presented) here and as the general validity of mono-modal size distributions as e.g., OPAC is based on has to be questioned, it cannot be concluded how good or bad the effective radius retrieval really is. Nevertheless, the AERONET analysis reveals that there is some correlation between AERONET Anström exponent and retrieved dust effective radius. Thus, although it is no quantitative validation of the effective radius, this observable seems to provide at least some qualitative information about the size distribution of airborne dust particles. Generally the spatial distributions of effective radius with larger  $R_{\rm eff}$  near source regions and lower  $R_{\text{off}}$  over the oceans seem to look quite reasonable and thus support the finding of  $R_{\rm eff}$  providing at least qualitative information.

For  $T_{dust}$ , no evaluation has been performed so far, thus also no  $T_{dust}$  results are shown here. Future work will focus on the potential of retrieving dust layer height information from  $T_{dust}$  quantitatively.

The error correlation analysis with AERONET data shows that the underestimation of high AOD and small effective radii representing conditions with rather small dust particles being airborne are the largest error sources. Small effective radii are connected with a rather wide range of AERONET Angström exponents, which indicates that in those cases the particle size distribution assumption of OPAC seems to be invalid. This invalidity might be the result of bi-modal particle size distributions

# **AMTD**

4, 461-509, 2011

## SVD based dust retrieval for IASI

L. Klüser et al.

Title Page Introduction **Abstract** Conclusions References **Figures** Tables  $\triangleright$ Back Close Full Screen / Esc

instead of the mono-modal lognormal distributions assumed in the OPAC database, as discussed above. For larger effective radii there seems to be a quite reasonable relationship between Ångström exponent from AERONET and estimated effective radius.

Both, observations during day and night conditions have been presented here showing the potential of TIR sounder observations for dust retrieval regardless of solar illumination. Nevertheless due to the lack of corresponding independent nighttime data almost all validation steps are performed during daytime. Evidence for the nighttime retrievals being as good as the daytime retrieval is given (only) by the general agreement of day- and nighttime observations of the specific events (plausible evolutions of the dust plumes).

The monthly means of dust AOD presented for February, May, August and November 2009 represent well the seasonal cycle of mineral dust as e.g., described in N'Tchayi Mbourou et al. (1997); Engelstaedter et al. (2006); Laity (2008); Klüser and Holzer-Popp (2010); and Knippertz and Todd (2010). Especially in the Sahelian dry season (mainly February) the IASI dust AOD reveals the possibility of separating dust from undetected biomass burning AOD, which has very strong influence especially on the aerosol plume over the Atlantic Ocean (rather weak in the IASI observations).

#### 6 Conclusions and outlook

A new approach of retrieving desert dust optical depth from IASI observations has been presented. The retrieval is based on a singular vector decomposition of IASI spectra in the 8–12 µm band rather than on direct simulation of radiative transfer. Avoiding radiative transfer modelling and making use of the decomposition of the IASI spectra in uncorrelated (orthogonal) principal components allows for a separation of surface emissivity and dust signals without the need for a-priori assumptions of emissivity spectra. Moreover the method does not rely on a-priori assumptions of atmospheric temperature and humidity profiles as maximum brightness temperature per bin is used instead of single-channel radiance. A set of singular vectors is determined

AMTD

4, 461–509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I4 ►I

Back Close

Full Screen / Esc

Printer-friendly Version



from several days of IASI observations over the Sahara-Arabia domain.

Two different cases of dust events over Northern Africa and Arabia show good performance in dust detection with the IASI method. Being a thermal infrared method (8–12  $\mu$ m band), the IASI retrieval provides observations independent from solar illumination conditions and thus is capable of dust observations twice daily. Good correspondence between day and night observations is found. The use of different OPAC dust types with different size distributions offers the possibility to retrieve not only dust AOD but also to estimate effective radius by means of contributions of the different OPAC types to total dust AOD. Due to a lack of suitable comparison datasets the effective radius retrieval could not be evaluated, but the results arising from the general evaluation indicate, that there is some useful information in the  $R_{\rm eff}$  results. Equivalent optical depth spectra used in the retrieval are corrected for dust thermal emission in one iteration of the retrieval algorithm. Thus as a by-product dust emission temperature is retrieved.

The set-up of the retrieval gives the advantage that not only AOD, effective radius and dust emission temperature can be derived, but also the intrinsic uncertainty of AOD (and dust temperature) are direct output variables of the retrieval. Thus every single retrieved AOD is accompanied by an intrinsic uncertainty value based on dust type (which means particle size distribution), dust temperature and singular vector projection errors. The delivery of pixel-wise uncertainty estimations is a great advantage in many applications as selections of subsets of observations can be made based on the desired degree of reliability.

Dust related research such as e.g., dust-cloud interaction studies will clearly benefit from the increased sampling rate (twice daily). Especially the potential of dust observation after sunset, when deep convective systems over land are much more pronounced than in the morning hours, is a great opportunity for such studies. Moreover results largely depend on the aerosol type, while a proper discrimination of dust contribution from other aerosol types has been shown to be of great importance (Klüser and Holzer-Popp, 2010).

**AMTD** 

4, 461–509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ✓ ▶I

Back Close

Full Screen / Esc



Another possible application is a revision of dust source area activation frequency maps used e.g., for numerical modelling of mineral dust (Tegen, 2003; Schepanski et al., 2007). The sampling rate of observations twice daily can also offer some help to understand the diurnal cycle of dust activity near source regions (Schepanski et al., 2007) and to revise applications based on the diurnal cycle of dust activity (Klüser and Schepanski, 2009; Klüser and Holzer-Popp, 2010).

Assimilation of aerosol optical depth into numerical modelling systems ideally would require assumptions about the type of aerosol, their size distribution and their vertical profile together with pixel-wise error estimates of these variables. The IASI retrieval can provide AOD of mineral dust clearly separated from other aerosol contributions. It moreover provides an estimate of the dust particle size distribution and, together with the temperature profile of the model system, possibly some height information. Moreover each observation is accompanied by an independent uncertainty estimation. Thus the IASI retrieval meets several of the key requirements of data assimilation and might prove being very useful in this context.

MODIS Deep Blue validation against AERONET shows that "Deep Blue" AOD values are generally within 20% to 30% of AERONET observations, also over deserts. No correlation coefficients or RMSD values are reported for the "Deep Blue" validation (Hsu et al., 2004, 2006). MISR AOD reaches correlation coefficients larger than 0.7 over desert regions and larger than 0.9 over oceans. Both products are total AOD retrievals based on reflected solar radiation. Thus, although performing somewhat weaker in the AERONET comparison (correlation coefficient of about 0.66), the IASI method presented here still performs reasonable well taking into account that only Dust AOD is observed (and that it is transferred from TIR to 0.5 μm) and that it is a thermal infrared retrieval.

The weaker performance in the AERONET evaluation (compared to the solar retrievals) is at least balanced by the advantages of the IASI retrieval. In contrast to solar retrievals it is sensitive to the mineral dust fraction of the aerosol only, thus a separation of aerosol types required by several applications is possible. As an

**AMTD** 

4, 461–509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I 

I 

I 

Back Close

Printer-friendly Version

Full Screen / Esc

Interactive Discussion



Discussion Paper

SVD based dust retrieval for IASI

**AMTD** 

4, 461-509, 2011

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I 

I 

I 

Back Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

thermal infrared method dust observation is possible twice daily leading to a much better sampling. The use of singular vectors instead of radiative transfer modelling avoids the necessity of knowledge about surface emissivity (or generally surface properties) and leads to applicability of the same method over desert, vegetated land and ocean. Generally the method is also suitable to detect dust above clouds, if the cloud field is not too inhomogeneous. The independence of solar illumination enables a twice daily sampling, allowing not only for much larger sample sizes in applications, but also for (partly) assessment of the diurnal cycle of dust activity near source regions. Besides the AOD also some information on dust layer temperature (height) and dust particle side is estimated in the IASI retrieval and can be provided simultaneously. Quantitative assessment of the retrieval of these observables from IASI, including validation, will be the subject of future work.

The IASI dust observations will soon become available in the World Data Center for Remote Sensing of the ATmosphere (WDC-RSAT), mandated by ICSU and WMO and hosted by the German Remotesensing Datacenter (http://wdc.dlr.de).

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**AMTD** 

4, 461–509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I 

I 

Back Close

Full Screen / Esc

Printer-friendly Version



Interactive Discussion

**AMTD** 

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

- Title Page

  Abstract Introduction

  Conclusions References

  Tables Figures

  I 

  ▶ I

  Back Close

  Full Screen / Esc
  - Printer-friendly Version



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**AMTD** 

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page **Abstract** Introduction Conclusions References

> **Figures** Tables

Close

Full Screen / Esc

4, 461-509, 2011

## SVD based dust retrieval for IASI

L. Klüser et al.

- Title Page Introduction **Abstract** Conclusions References **Figures** Tables Back Close Full Screen / Esc
- - Printer-friendly Version
  - Interactive Discussion

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  - 4, 461-509, 2011

### SVD based dust retrieval for IASI

**AMTD** 

L. Klüser et al.

Title Page **Abstract** Introduction Conclusions References **Figures** Tables Back Close Full Screen / Esc Printer-friendly Version

Interactive Discussion

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464, 469

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4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I 

I 

I 

Back Close

Full Screen / Esc

Printer-friendly Version

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**Table 1.** FOV dust likelihood score test, the final score is the sum of the scores of all positive tests. Only FOVs with scores higher than 250 are processed further.  $\lambda_{\chi}$  denotes the wavelength corresponding to the maximum brightness temperature ( $T_{\rm base}$ ).

Test no.	Test condition	Score
1	$\tau_{\text{eqv}}$ (10.00 μm)/ $\tau_{\text{eqv}}$ (10.63 μm)>2	128
2	$\tau_{\rm eqv} (10.10  \mu {\rm m})/\tau_{\rm eqv} (10.87  \mu {\rm m}) > 1$	64
3	$\tau_{\text{eqv}} (10.10  \mu\text{m}) / \tau_{\text{eqv}} (11.49  \mu\text{m}) > 1$	64
3	$(\lambda_x$ >11 μm or $\lambda_x$ <9 μm) and $\tau_{\rm eqv}$ (10.10 μm)/ $\tau_{\rm eqv}$ (11.49 μm)>2	64

4, 461-509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I 

I 

Back Close

Full Screen / Esc

Printer-friendly Version



Table 2. Linear and rank correlations between IASI error (with respect to AERONET) and

several influencing parameters.

	AOD	Δd	$\Delta t$	$\sigma^2$ (IIS)	$R_{ m eff}$	$T_{ m dust}$	Θν
$ ho_{ m linear}$	-0.265	+0.076	-0.136	-0.089	-0.354	-0.142	+0.142
$ ho_{rank}$	-0.234	+0.112	-0.148	-0.162	-0.323	-0.198	+0.101

4, 461-509, 2011

# SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

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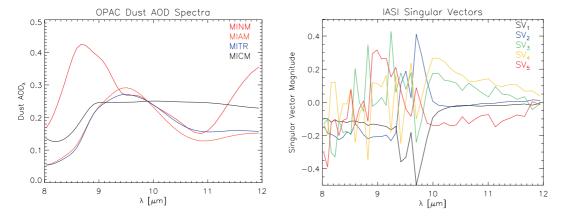
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Back Close

Full Screen / Esc

Printer-friendly Version





**Fig. 1.** Dust AOD spectra for four dust types from the OPAC database (left) and leading five singular vectors obtained from IASI observations (right).

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I 

I 

I 

Back Close

Full Screen / Esc



Printer-friendly Version

Interactive Discussion



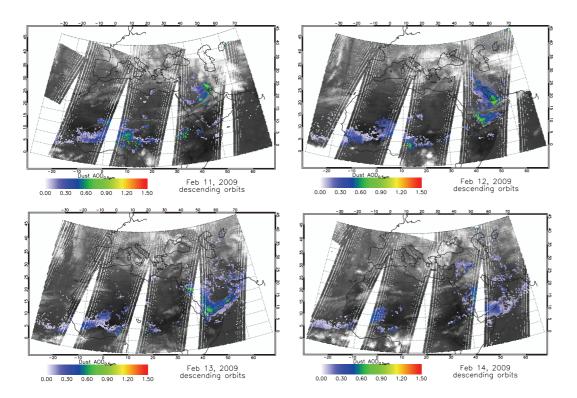


Fig. 2. IASI morning orbit observations of Dust AOD  $_{0.5\,\mu m}$  (colours) and IIS temperature (greyscale) for 11–14 February 2009 (top left to bottom right).

**AMTD** 

4, 461-509, 2011

#### SVD based dust retrieval for IASI

L. Klüser et al.

Title Page **Abstract** Introduction

Conclusions References

> **Tables Figures**

Close Back

Printer-friendly Version

Printer-friendly Version

Interactive Discussion



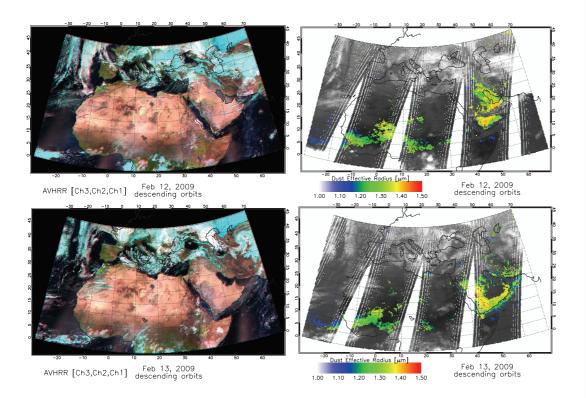


Fig. 3. AVHRR RGB composites (left) and IASI dust effective radius (right) for 12 (top) and 13 February (bottom).

**AMTD** 

4, 461-509, 2011

#### SVD based dust retrieval for IASI

L. Klüser et al.

Title Page Introduction **Abstract** Conclusions References

> **Figures Tables**

Close Back





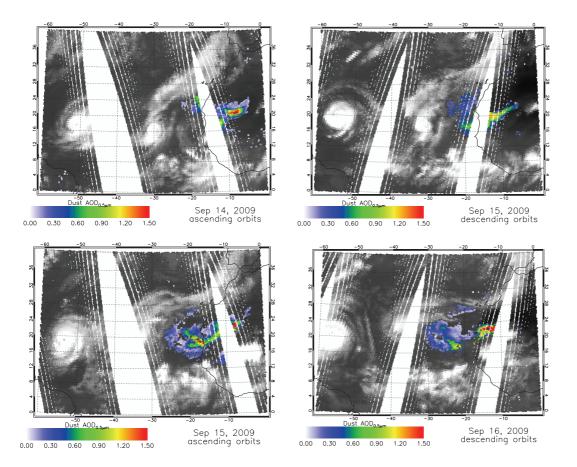


Fig. 4. IASI observations of Dust  $AOD_{0.5\, \mu m}$  (colours) and IIS temperature (greyscale) from 14 September afternoon through morning 16 September 2010.

**AMTD** 

4, 461-509, 2011

#### SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

**Abstract** 

Introduction

Conclusions

References

**Tables** 

**Figures** 









Close





Printer-friendly Version

Interactive Discussion



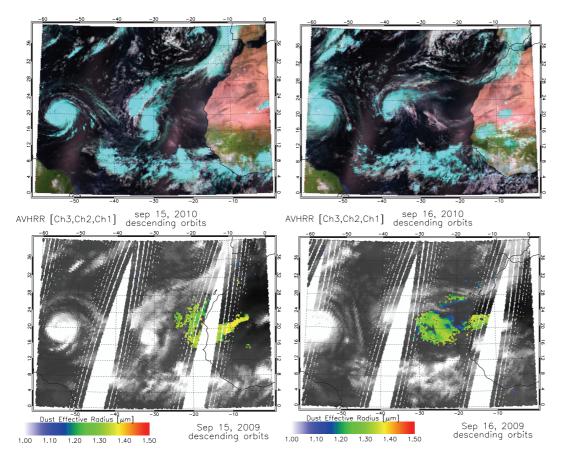


Fig. 5. AVHRR RGB composites (top) and IASI dust effective radius (bottom) for 15 (left) and 16 September (right) descending orbits.

**AMTD** 

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

**Abstract** 

Introduction

Conclusions

References

**Tables** 

**Figures** 







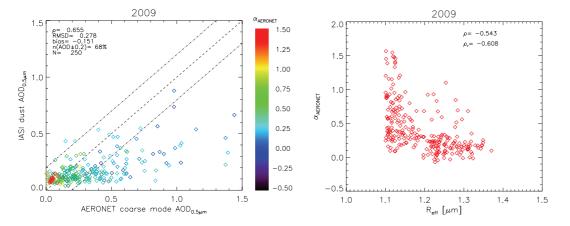


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**Fig. 6.** Left: Evaluation of IASI Dust AOD<sub>0.5 μm</sub> against AERONET coarse mode AOD<sub>0.5 μm</sub>. Symbol colours represent AERONET Ångström exponent  $\alpha$ . Right: AERONET Ångström exponent plotted over IASI retrieved dust effective radius.

4, 461-509, 2011

SVD based dust retrieval for IASI

L. Klüser et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I ✓ ▶I

Back Close

Full Screen / Esc





Full Screen / Esc

Printer-friendly Version



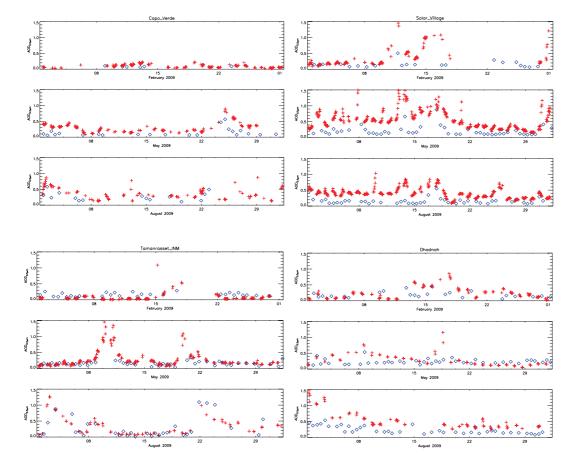


Fig. 7. Time series of AERONET (red) and IASI (blue) derived AOD at the AERONET stations Capo Verde (top left), Solar Village (top right), Tamarasset (bottom left) and Dhadnah (bottom right) for the months February, May and August 2009.

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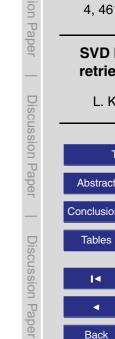
4, 461-509, 2011

#### SVD based dust retrieval for IASI

L. Klüser et al.

Title Page **Abstract** Introduction Conclusions References **Figures Tables** 

> Close Back



Printer-friendly Version

Interactive Discussion



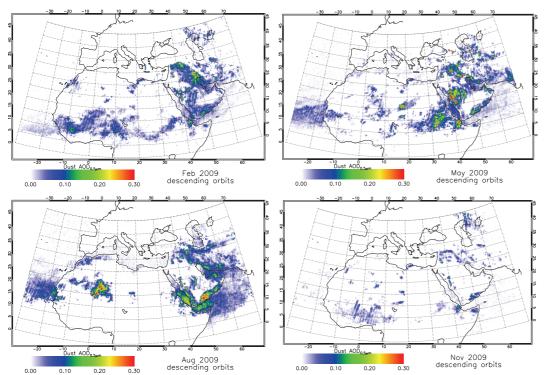


Fig. 8. IASI Dust  $AOD_{0.5\,\mu m}$  for February (top left), May (top right), August (bottom left) and November (bottom right) 2009.

**AMTD** 

4, 461-509, 2011

## SVD based dust retrieval for IASI

L. Klüser et al.

Title Page **Abstract** Introduction Conclusions References **Figures** Close Back Full Screen / Esc