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Ground-based water vapor Raman lidar measurements up to the upper troposphere and lower stratosphere – Part 1: Instrument development, optimization, and validation

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Abstract

Recognizing the importance of water vapor in the upper troposphere and lower stratosphere (UT/LS) and the scarcity of high-guality, long-term measurements, JPL began the development of a powerful Raman lidar in 2005 to try to meet these needs. This

- development was endorsed by the Network for the Detection of Atmospheric Compo-5 sition Change (NDACC) and the validation program for the EOS-Aura satellite. In this paper we review the stages in the instrumental development of the lidar and the conclusions from three validation campaigns: MOHAVE, MOHAVE-II, and MOHAVE 2009 (Measurements of Humidity in the Atmosphere and Validation Experiments). The data
- analysis, profile retrieval and calibration procedures, as well as additional results from 10 MOHAVE-2009 are presented in detail in a companion paper (Leblanc et al., 2011a). Ultimately the lidar has demonstrated capability to measure water vapor profiles from \sim 1 km above the ground to the lower stratosphere, reaching 14 km for 1-h integrated profiles and 21 km for 6-h integrated profiles, with a precision of 10% or better near 13 km and below, and an estimated accuracy of 5%.
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Introduction 1

Due to its radiative, chemical, and thermodynamic properties, water vapor has long been identified as a key constituent of the atmosphere. The water vapor molecule strongly absorbs infrared radiation and in the troposphere it is therefore a primary greenhouse gas. In the stratosphere it is produced by methane oxidation thus link-20 ing it to ozone chemistry. It was shown that a global increase in lower stratospheric H₂O mixing ratio similar to that observed locally since 1981 (Oltmans and Hofmann, 1995; Hurst et al., 2011), would contribute to a surface warming reaching 40% of that responsible from CO₂ increases over the same period (Forster and Shine, 1999).

The resulting lower stratospheric cooling would be of same order of magnitude as that 25 caused by changes in ozone concentrations. Therefore, to fully understand, quantify,





and predict future water vapor-related radiative and chemical processes impacting climate change, high accuracy water vapor measurements throughout the troposphere and stratosphere are required (typically 3–10%). Many instruments however currently cannot achieve such measurement accuracy without thorough calibration and valida-⁵ tion.

In order to address this issue, the Network for the Detection of Atmospheric Composition Change (NDACC, formerly known as NDSC) recently included water vapor Raman lidar in its suite of long-term monitoring techniques.

The fundamental Raman lidar measurement technique is relatively simple and easy
to implement (e.g. Melfi et al., 1969; Cooney, 1970; Vaughan et al., 1988; Sherlock et al., 1999a,b). A laser pulse is emitted into the atmosphere and scattered by the molecules and particles. A fraction of the laser light is collected back on the ground with a telescope, where it is spectrally separated, and sampled in time (i.e. distance). After typical corrections, the ratio of the lidar signals collected in the water vapor and nitrogen channels is proportional to water vapor mixing ratio. To obtain the absolute water vapor mixing ratio the lidar profile needs calibration, which for example, is done

using the mixing ratio measured by radiosonde in the lower troposphere. In the framework of NDACC, a high capability water vapor Raman lidar was recently built at the Jet Propulsion Laboratory (JPL) Table Mountain Facility (TMF) in California

(34.4° N, 117.5° W, elevation 2285 m), with the primary objective of reaching the upper troposphere and lower stratosphere (UTLS) with the best possible accuracy (5% or better). The instrument, referred to as "TMW" in the remainder of this paper, went through several configurations before routine measurements were finally obtained with the quality requirements initially set.

The TMW instrument technical description and achievements, including multiple upgrades over the period 2005–present, is presented here. A detailed description of the data analysis and calibration procedure is provided in a companion paper (Leblanc et al., 2011a). The initial lidar configuration will be described in Sect. 2. The various stages of instrument validation and optimization that led to the present configuration will





be described in Sect. 3. This section will be concluded by results from the MOHAVE-2009 campaign which characterize the present performance of the system.

2 Instrument initial setup

2.1 Transmitter

⁵ Throughout the development of the TMW lidar system, the laser transmitter, telescope, and the data acquisition system have remained essentially unchanged. The laser is a high pulse energy Nd:YAG laser (Continuum) operating at 355 nm at a repetition rate of 10 Hz. While specified to produce more than 900 mJ/pulse, in long-term operation it consistently provides around 650 mJ/pulse. The beam is expanded 7.5 × using a
 ¹⁰ refractive beam-expander (CVI) before being steered into the atmosphere. Alignment of the laser to the receiving telescope is performed automatically and the program and mechanics of this part of the system has been fully described by Aspey et al. (2008). Figure 1 of this reference also shows the basic arrangement of the lidar transmitter and receiver telescope.

15 2.2 Receiver

The primary telescope is a 91 cm-diameter Newtonian telescope for the far-range channels supplemented by three (later four) 7.5 cm-diameter refractive telescopes for the near range. Coupling of the large telescope to the receiver polychromator is one of the arrangements that was changed as the lidar evolved as discussed in Sect. 3. TMW makes use of the vibrational Raman shift at 2330 cm^{-1} for nitrogen and 3654 cm^{-1} for

makes use of the vibrational Raman shift at 2330 cm ⁻¹ for nitrogen and 3654 cm ⁻¹ for water vapor, which for an emission at 355 nm, translates to receiving at 387 nm and 407 nm respectively.

The small receivers are each designed to detect only a single wavelength and, as shown in Fig. 1, comprise a $7.5 \,\mathrm{cm}$ fused silica lens (L1), a variable diaphragm field



stop (FS), collimating and refocusing lenses sandwiching an interference filter (one each for 355 nm [F3], 387 nm [F1], and 407 nm [F2]). The received light is focused directly onto the photomultiplier (PMT) detector modules (P). Other than the addition of a fourth channel (very low intensity 355 nm for tropospheric aerosols information), the
 small receivers setup was basically unchanged since the lidar first light in 2005.

A diagram of the initial design of the channels coupled out of the large telescope is shown in Fig. 2. Light was sent from the telescope into the polychromator using a 1 mm diameter UV grade optical fiber (CeramOptec). The 0.22 numerical aperture of the fiber was matched to the f/2.7 focal ratio of the telescope and the diameter of the fiber also acted as a field-stop defining the field-of-view (fov) as 400 µrad. There were sev-

- ¹⁰ also acted as a field-stop defining the field-of-view (fov) as 400 µrad. There were several reasons for choosing to use fiber coupling. It was not clear at the outset whether it would be required to install a chopper to counter signal-induced-noise. Should this have proved necessary, it would have been much easier to implement with a fiber. Also, we have successfully used similar fibers in other lidar systems (e.g. McDermid)
- et al., 1999) and they minimize obscuration in the Newtonian configuration. Referring to Fig. 2, the light entering the polychromator from the fiber is collimated and then the first dichroic beamsplitter (B1) transmits 387 nm while reflecting 407 nm and 355 nm. An interference filter (F1) further selects only 387 nm and then the signal is divided using an anti-reflection coated fused silica plate (B3) into two parts with intensity ra-
- tio approximately 99:1. Using these 99:1 dual channels allows signal non-linearities, i.e. saturation, to be corrected and dynamic range to be extended. Similarly, beam-splitter (B2) reflects 407 nm and transmits the remaining 355 nm. As for the 387 nm channels, an interference filter (F2) is used to select only 407 nm and the signal is divided (B3) into high and medium intensity channels. Interference filter (F3) transmits only 355 nm onto a gated PMT.

In the original configuration, all 8 channels (5 channels from the large telescope and 3 from the small telescopes) utilized the same type PMT modules (Hamamatsu H5783P) except for the 355 nm high-intensity channel which used a bare R7400 PMT (Hamamatsu) in a home-built electronically gated base. This PMT was electronically





gated to turn on at about 20 km in order to suppress some of the saturation and signalinduced-noise caused by the very high intensity Rayleigh returns from lower altitudes. All PMT signals were input to a series of photon counting multi-channel-scalers (Licel).

2.3 Results from the initial configuration (2005–2006)

⁵ The configuration described above was used in 2005 and 2006 up to and including the first MOHAVE campaign (October 2006). Data were acquired typically during 2 h at the beginning of the night and simultaneously with the two other operating lidars on site (tropospheric ozone lidar, and stratospheric ozone/aerosols/temperature lidar). The 2-h long measurements were sliced into 5-min interval datasets, providing vertical profiles with high temporal resolution for altitude ranges with sufficient signal-to-noise ratios (typically, below 10 km).

The lidar signals from all three pairs of 387–407 nm channels were analyzed for the retrieval of water vapor (3–20 km) by our lidar analysis software LidAna (Leblanc, 2011a). Calibration was obtained using the mixing ratio values measured in the lower

- ¹⁵ troposphere by meteorological radiosondes launched from the lidar site. For that purpose the JPL lidar team uses its own Vaisala RS92 radiosonde station. One RS92 radiosonde was launched from TMF on each lidar measurement night. Typically the launch time is chosen to coincide with the first hour of the lidar measurement. A detailed description of the data analysis and calibration procedure is provided in our companion paper (Leblanc et al., 2011). More than 200 2-h-long lidar profiles were
 - analyzed during the June 2005–October 2006 period.

Figure 3 shows four examples of individual water vapor profiles obtained simultaneously by the lidar and RS92 radiosondes showing good agreement between the lidar and the radiosonde up to about 9–10 km and an increasing difference as we approach

and cross the tropopause, the RS92 being too dry and/or the lidar being too wet. The difference is clearly identified in the compiled statistics shown in Fig. 4. The excessive dryness of the RS92 in the upper troposphere has been known and well documented for many years (Miloshevich et al., 2004, 2009). However, the possible wet bias of the





lidar in the UT/LS could not be confirmed until the lidar profiles could be compared to accurate measurements in this region. These comparisons took place during the Measurements Of Humidity in the Atmosphere and Validation Experiments (MOHAVE) campaign held in October 2006 with the primary objective of validating the Raman water vapor lidar measurement in the UT/LS.

3 Instrument optimization and validation

3.1 The MOHAVE-I campaign (October 2006)

Fifty RS92 radiosondes and 10 Cryogenic Frost-point Hygrometers (CFH) (Vömel et al., 2007) were launched from TMF during the 15 nights of the campaign. Two mobile
lidars from the Goddard Space Flight Center (GSFC): Aerosol and Temperature Lidar (AT-Lidar) (McGee et al., 1991), and Scanning Raman Lidar (SRL) (Whiteman et al., 2006), were deployed at TMF for the campaign. The Naval Research Laboratory (NRL) Water Vapor Millimeter-wave Spectrometer (WVMS), one Global Positioning System (GPS) from JPL and one from GSFC also participated. Each of the three collocated
water vapor lidars acquired over 150 h of measurements (2 to 10 h per night). At least one balloon per night was launched, each payload including one or two RS92. Ten payloads also contained a CFH and an ozonesonde. The lidar data corresponding to the first hour after launch were systematically processed and compared with the balloon

- measurements. In these 1-h profiles, thin layered structures were well captured by all
 instruments but the lidar profiles begin to get noisy above 12 km. Figure 5 shows the average of four profiles acquired simultaneously by five different instruments/techniques during the campaign. If the CFH is taken as the reference profile, we find a dry bias in the upper troposphere for the RS92 and a wet bias for all three lidars, even though the lidars show excellent agreement with each other.
- ²⁵ After investigating the possible sources of the lidar wet bias and running additional test experiments, signal contamination by fluorescence in the lidar receivers was





identified, which led all three lidar teams to modify their instrument configuration. For the JPL lidar, a 355 nm blocking filter was temporarily installed in front of the fiber optic. The instrument acquired 3 profiles simultaneously with a CFH in this configuration. The difference between the CFH and lidar profiles during these three flights was compared to that before the modification. The results are plotted in Fig. 6 (left and center panels respectively). As anticipated, the wet bias completely disappeared after the modification (center panel). The nature of the fluorescence contamination was demonstrated by applying a physically-based empirical correction to the contaminated signals plotted in Fig. 6 (left panel). The profiles retrieved from the corrected signals are shown in Fig. 6

(right panel). The identification and removal of signal contamination by fluorescence was a major milestone in the TMF water vapor lidar project, and a major achievement of the MOHAVE campaign (Leblanc et al., 2008a)

3.2 Receiver second configuration (2007–2009)

After MOHAVE, the front-end of the lidar receiver was redesigned to permanently suppress the fluorescence identified in the fiber optic during the campaign. This modification required custom-designed mechanical and optical components and was fully implemented in July 2007. Between MOHAVE and July 2007, only minor modifications were made (including sensitivity tests to fluorescence of several fiber optics). Routine operation continued alternately with and without fluorescence contamination. Signals
²⁰ knowingly acquired with contamination were empirically corrected to remove the effect of the fluorescence. The new configuration is shown in Fig. 7 and was fully implemented in July 2007. The strong 355 nm Rayleigh signal was redirected out of the

- main optical path ahead of the focus of the large telescope using a long-wave pass beamsplitter (B4) that reflected 355 nm and transmitted 387 nm and 407 nm. An additional 355 nm blocking filter (F4) was placed ahead of the fiber connecting to the polychromator. A separate fiber was used to transmit the 355 nm to the gated PMT.
- This arrangement successfully eliminated fluorescence in the 387 nm and 407 nm signals (Leblanc et al., 2008a). However, the new and much stronger spectral selectivity





resulted in an overall loss of signal by a factor of two, causing a decrease of 3 km (typically from 17 km to 14 km for 1-h integration) of the uppermost altitude of the instrument.

3.3 The MOHAVE-II campaign (October 2007)

- To evaluate the effectiveness of the system modifications to remove fluorescence, a 5 second MOHAVE campaign was carried out in October 2007. Again, two mobile lidars from GSFC were deployed for MOHAVE-II. However, the SRL system having sustained major damage during transportation for MOHAVE-I, was replaced by the Atmospheric Laboratory for Validation, Interagency Collaboration, and Education (ALVICE) trailer.
- ALVICE contains an upgraded version of the Raman Airborne Spectroscopic Lidar 10 (RASL) adapted for ground-based measurements (Whiteman et al., 2007, 2010). The AT lidar had only minor modifications. MOHAVE-II was implemented following operational principles similar to that of MOHAVE-I (Leblanc et al., 2008b). Figure 8 shows the average of the 10 profiles measured simultaneously by all participating instruments
- and techniques during the campaign. All the lidars agreed well with the CFH up to 15 12 km. Though the lidar profiles appear noisier, partly due to the signal decrease mentioned above, no wet bias appears anymore between the JPL lidar (magenta curve) and the CFH (green). On average, a 1-h lidar integration reached \sim 15–16 km for the JPL system. A dry bias was still present on the uncorrected RS92 profile (red). A
- comparison of the results from the first two MOHAVE campaigns is shown in Fig. 9. 20 This figure clearly demonstrates that the measurement accuracy of the Raman lidar is better than that of the RS92 in the UT/LS. The CFH measurement remains the best quality among all instruments but its cost per profile is much higher (>\$ 3000 per CFH launch) than that of the lidar and clearly prohibitive for long-term, routine monitoring,
- i.e. multiple measurements per week. 25





3.4 Receiver current configuration (2009-present)

While the installation of the additional optics to divert the 355 nm returns into a separate fiber eliminated the fluorescence contamination and brought the lidar profiles into agreement with the CFH, the signal levels, and consequently the altitude range of the

- ⁵ measurement were reduced. Simulation of the expected lidar returns and comparison with other systems suggested that significant signal was being lost somewhere in our receiver; most likely coupling in and out of the fibers. To try to recover the signal level and potentially increase it, it was decided to eliminate the fibers completely and to reposition the receiver at the telescope Newtonian focus.
- Figure 10 is a schematic of the current receiver configuration (far-range channels). To make the system more compact a new 387/407 nm beamsplitter (B1') operating at 45° angle of incidence was acquired to replace the previous one that operated at 30°. All of the other optical components were the same as described previously except for the field stops (FS). The new field stops were calibrated, ring-activated iris diaphragms
- that enabled the field-of-view to be varied as opposed to the fixed fov defined by the fiber diameter in the earlier configurations. This flexibility provided better optimization of laser/telescope overlap and sky background rejection. Values of 600 µrad and 800 µrad were finally selected for the Raman and 355H channels respectively. The various components of the receiver were aligned on top of each other to minimize ob-
- scuration of the large telescope and a calibration lamp was permanently mounted to the side of the receiver module. The purpose of this lamp is presented in Leblanc and McDermid (2008) and is further discussed in our companion paper (Leblanc et al., 2011b).

One other change in this final configuration was to the 407H photomultiplier detector. The original Hamamatsu H5783P module was replaced with a H10682-110 photon counting head device. Within this module, the PMT high voltage supply and discriminator level are factory preset to optimum values. Also, the PMT in this device has a super-bialkali (SBA) photocathode which has higher quantum efficiency than standard





PMTs; 33 % at 407 nm compared to 25 % (eventually, when commercially available, ultra-biakali (UBA) photocathodes will further raise the QE to ~40 %). The net result of all these changes was an order of magnitude increase in signal levels. However, such major changes required re-evaluation/validation and therefore another campaign, MOHAVE-2009 was organized.

3.5 The MOHAVE-2009 campaign (October 2009)

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The modifications described above occurred in the summer of 2009 (July, August and September). No further changes occurred after that and the instrument remained untouched throughout the MOHAVE-2009 campaign with the exception of the calibration lamp (Leblanc et al., 2011b). This third MOHAVE campaign was by far the largest of all three MOHAVE campaigns with many participating instruments: two types of frost-point hygrometers (CFH and NOAA-Frost Point), two types of radiosonde (Vaisala and InterMet), two microwave radiometers (NRL and University of Bern), two Fourier-Transform spectrometers (the JPL MkIV and FTUVS), and two GPS receivers. The

- ¹⁵ AT lidar and ALVICE from GSFC also returned. An additional system from GSFC, the Stratospheric Ozone Lidar Trailer (STROZ) (McGee et al., 2011) participated for the first time and contains a system similar to that in the AT lidar. This was fortunate since a laser failure in the AT lidar precluded any measurements during MOHAVE-2009 from this system. Coincident measurements from several satellite instruments (Aura-
- MLS, Aqua-AIRS, Aura-TES, ENVISAT-MIPAS and ACE-FTS) were also compared to the balloon-borne and ground-based measurements. The campaign spanned approximately two weeks from 11 October to 27 October 2009. A review of the MOHAVE-2009 campaign operations and results is provided in Leblanc et al. (2011b). Documentation and an overview of the results can be found on the campaign website:
- http://tmf-lidar.jpl.nasa.gov/campaigns/mohave2009.htm. A total of 44 balloons were launched which included 58 RS92 radiosondes, 16 CFH, 4 NOAA-FPH, and 16 ECC ozonesondes. Approximately 300 h of water vapor lidar measurements were acquired. Results from MIPAS, the MkIV spectrometer, balloon measurements, and two lidar





instruments are presented in Stiller et al. (2011), Toon et al. (2011), Hurst et al. (2011), Whiteman et al. (2011), and McGee et al. (2011) respectively.

Figure 11a–d shows four examples of water vapor profiles measured simultaneously during the campaign. In all plots, the lidar profiles (TMW, ALVICE and STROZ) were

- ⁵ obtained from a 1-h integration starting at the corresponding balloon launch time. The CFH campaign-mean profile is provided on each plot for reference purposes only. 18 and 20 October (top-left and top-right panels) illustrate the very large natural variability in the troposphere from one night to another: very humid upper troposphere on 18 October and very dry on 20 October. Most of the very fine vertical structures
- (sometimes less than 1-km deep) are well captured by all balloon-borne and lidar instruments. A well-known RS92 dry bias is again systematically observed above 12 km (cyan curves). The RS92 profiles corrected following the method described by Miloshevich et al. (2009) show a better agreement with lidar and CFH above 13 km. The difference between corrected and uncorrected profiles can reach 35 % and is reviewed
- in Leblanc et al. (2011). The bottom-left and bottom-right panels of Fig. 11 show profiles measured on the night 27 October. Besides showing very large vertical changes (from 200 ppmv to 2000 ppmv in less than 2 km), the two figures illustrate the very high temporal variability on short timescales, namely a change at 4 km altitude from 200 ppmv to 4000 ppmv in a 5-h interval. The very fine and fast-changing structures
 observed on all plots of Fig. 11 confirm the need for extra care when producing com-
- parisons between the various instruments and techniques.

The campaign-mean profiles measured by the TMW and ALVICE lidars and their difference, is plotted in Fig. 12. Both uncorrected (blue) and fluorescence-corrected (green) ALVICE profiles are shown. During MOHAVE-2009, the ALVICE lidar signals experienced contamination by fluorescence unexpectedly induced in the instrument receiver. Using a source of external measurement in the UTLS (in this case CFH), this contamination can be characterized and the effect fluorescence can be mostly suppressed. This fluorescence correction is described in Whiteman et al. (2011). The ALVICE lidar wet bias due to fluorescence contamination remains small below 10 km,





but reaches 20 % above 13 km. Below 10 km, the lidars agree very well with each other with differences not exceeding 5 % (within the lidar uncertainty range).

Finally the campaign-mean profiles measured simultaneously by the TMW lidar and CFH, and their difference, are plotted in Fig. 13. Since the CFH is the most accurate

- instrument in the UTLS participating to the MOHAVE-2009 campaign, this figure is indeed the most important result of the campaign for the TMW lidar. The comparisons were made using two different coincidence criteria. Below 14 km, atmospheric variability is very high and only the strictly coincident profiles were used to compute the mean (i.e. 12 CFH flights and the 12 corresponding 1-h integrated lidar profiles). Using the 1-
- h integrated profiles prevents the lidar reaching altitudes above 14 km. For these higher altitudes, the lidar profiles integrated all-night were used (mean of 8 nights during which 9 CFH were launched). The TMW lidar and CFH profiles are in excellent agreement throughout the troposphere and the UTLS. The mean differences range from -5% in the lowermost troposphere (TMW drier) to +2% in the lower stratosphere (TMW wet ter). These values are well below the reported uncertainties of both instruments.
 - A close look at the green and pink dotted curves on the left panel of Fig. 13 reveals that the standard deviation of the lidar measurements above 16 km is two to three times larger than that of the CFH. This number provides a good indication of the number of additional lidar measurements needed to achieve the same precision as CFH. With
- the assumption that lidar data contains random noise following a Poisson distribution, using one to two CFH profiles yields a precision in the UTLS equivalent to integrating the lidar measurements for 5 to 8 full nights (i.e. 40 to 70 h). That is, one weekly CFH launch yields a precision in the UTLS similar to that of the integrated lidar measurement of 4 full nights per week. Similarly, one monthly CFH profile yields a precision in the
- ²⁵ UTLS equivalent to that of the integrated lidar measurement of one full night per week, or four 2-h-long measurements per week. These estimations therefore show that, with our current target to run TMW for two hours per night, 4 nights per week, we can achieve a precision in the UTLS equivalent to that achieved when launching one CFH per month.





4 Conclusions

A high capability water vapor Raman lidar was developed at the Jet Propulsion Laboratory Table Mountain Facility (California) with the objective of making accurate, routine, long-term water vapor measurements in the UTLS. The system, referred to as "TMW',

- ⁵ was built in 2005 and optimized over the past four years. To support and achieve a full optimization, the MOHAVE-I, MOHAVE-II, and MOHAVE-2009 validation campaigns took place in October 2006, 2007 and 2009 respectively. During these campaigns the lidar measurements were extensively compared to operational radiosonde (Vaisala RS92) and Research-grade instruments (CFH).
- ¹⁰ The MOHAVE-I campaign revealed the presence of fluorescence contaminating the lidar signal in its upper part (10–20 km). After a first major receiver configuration change in 2007, the MOHAVE-II campaign confirmed the removal of the contaminating fluorescence, but also revealed a substantial loss of signal resulting from this change. After a second major configuration change (summer 2009), the MOHAVE-2009 cam-
- paign finally showed that the TMW lidar was brought up to its initial performance expectations, i.e. capable of measuring water vapor mixing ratio in the UTLS with a precision of 10 % or better between 10 km and 16 km, and reaching 21–22 km with a degraded precision and vertical resolution.

The latest comparisons with CFH measurements (MOHAVE-2009) showed excellent agreement throughout the troposphere and UTLS, with mean biases well below the reported measurements uncertainties. The TMW lidar shows a 2 % mean dry bias with CFH in the lower troposphere and a 3 % mean wet bias in the UTLS.

A closer look at the lidar measurement standard deviations above 16 km during MOHAVE-2009 shows values two to three times larger than that of the CFH. At these

altitudes the lidar signals are noise limited and the standard deviation values reflect the precision of the measurements, which therefore provides a good indication of the number of additional lidar measurements needed to achieve the same precision as CFH. It was estimated that one monthly CFH profile yields a precision in the UTLS equivalent





to that of the integrated lidar measurement of one full night per week, or four 2-hlong measurements per week. These estimations show that, with our present target of routinely running TMW 2 h per night, 4 nights per week, we can achieve measurements with a precision in the UTLS equivalent to that achieved if launching one CFH per month. Though the TMW lidar is now fully optimized, its overall power-aperture

- product is well below most contemporary optimums. For example, commercial laser powers today can easily exceed 10 W at 355 nm as opposed to 6.5 W produced by our decade-old laser. Therefore it is not unlikely that a future laser upgrade, together with the upgrade of other components, will bring the precision of the TMW lidar closer
- to that of the CFH, thus requiring less laser running time to achieve similar levels of precision in the UTLS.

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Fig. 2. TMW lidar receiver initial configuration (2005–2007) for the high- and mid-intensity channels.

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Fig. 3. Example of four water vapor mixing ratio profiles measured by TMW in 2005 and 2006. Each profile was selected to represent a different season.







Fig. 4. Comparison of the climatological mixing ratio profiles measured over the period June 2005-October 2007 by the TMW lidar and simultaneous and co-located Vaisala RS92 radiosondes. Left panel: mean profiles. Right panel: mean difference, lidar measurement uncertainty and standard deviation.

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Fig. 6. Comparison of mean water vapor mixing ratio profiles measured simultaneously by the TMW lidar and CFH during MOHAVE-I. Left panel: lidar signals contaminated by fluorescence (no 355 nm block). Center panel: lidar signals free of fluorescence (355 nm block installed). Right panel: same as left panel, but after fluorescence correction is applied to lidar signals.



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Fig. 7. TMW lidar receiver second configuration (2007–2009) for the high- and mid-intensity channels.





Fig. 8. Averaged water vapor mixing ratio computed from 10 profiles measured simultaneously during MOHAVE-II by the TMW lidar, two other lidars, CFH, and Vaisala RS92 radiosondes.





Fig. 9. Mean water vapor mixing ratio computed from 8 profiles measured simultaneously by the TMW lidar, RS92 radiosondes, and CFH during MOHAVE-I (left panel) and computed from 9 profiles measured simultaneously during MOHAVE-II (right panel).



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Fig. 10. TMW lidar receiver final configuration (2009–present) for the high- and mid-intensity channels.

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Fig. 11. Example of four water vapor mixing ratio profiles measured simultaneously by TMW, two other lidars, CFH, Aura-MLS, and RS92 radiosondes during MOHAVE-2009.





Fig. 12. Comparison of mean water vapor mixing ratio (left panel) computed from 10 profiles measured simultaneously by the TMW lidar and the ALVICE lidar during MOHAVE-2009 and their difference (right panel). Both fluorescence-contaminated and fluorescence corrected profiles are shown for the ALVICE lidar.



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Fig. 13. Comparison of the MOHAVE-2009 campaign-averaged mixing ratio profiles (left panel) calculated from all coincident profiles measured by the TMW lidar and the CFH, and their difference (right panel). A loose (± 6 h) time coincidence criterion and longer lidar integration times were used in the UTLS, and a strict time coincidence criterion (± 1 h) and 60-min-integrated lidar profiles were used in the lower and mid-troposphere.



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