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# 1 km fog and low stratus detection using pan-sharpened MSG SEVIRI data

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# Abstract

In this paper a new technique for the detection of fog and low stratus in 1 km resolution from MSG SEVIRI data is presented. The method relies on the pan-sharpening of 3 km narrow-band channels using the 1 km high-resolution visible (HRV) channel. As solar and thermal channels had to be sharpened for the technique, a new approach based on an existing pan-sharpening method was developed using local regressions. A fog and low stratus detection scheme originally developed for 3 km SEVIRI data was used as the basis to derive 1 km resolution fog and low stratus masks from the sharpened channels. The sharpened channels and the fog and low stratus masks based on them were evaluated visually and by various statistical measures. The sharpened channels deviate only slightly from reference images regarding their pixel values as well as spatial features. The 1 km fog and low stratus masks are therefore deemed of high quality. They contain many details, especially where fog is restricted by complex terrain in its extent, that cannot be detected in the 3 km resolution.

## 1 Introduction

Fog and low stratus (FLS) have operationally been detected from AVHRR and MODIS data for quite some time (e.g. Eyre et al., 1984; Kudoh and Noguchi, 1991; Bendix et al., 2006). Also for geostationary satellite systems, which have the advantage of a high temporal resolution, many reliable approaches for FLS detection exist (e.g. Lee et al., 1997; comprehensive overview in Gultepe et al., 2007). For Meteosat Second Generation (MSG) Spinning-Enhanced Visible and Infrared Imager (SEVIRI) data the Satellite-based Operational Fog Observation Scheme (SOFOS) developed by Cermak (2006) is a recent and reliable algorithm for FLS detection. It has been extensively validated and proven its suitability for operational deployment (Cermak and Bendix, 2008), which is why it was used as the underlying FLS detection scheme in this study.



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However, for the detection of several fog types such as small-valley fog in a lower mountain range topography as found in many parts of Central Europe, the current scheme has some disadvantages, particularly because the nominal spatial resolution of 3 km at the sub-satellite point of the SEVIRI instrument's multispectral bands is too coarse to detect these small-scale FLS phenomena. On the other hand, a panchromatic high resolution visible (HRV) channel with a nominal resolution of 1 km per pixel is available on MSG, which could generally help to overcome this resolution problem. The high potential of the HRV channel in FLS detection was already highlighted by Bugliaro and Mayer (2004) using two different approaches. Since their first approach, the simple application of radiance thresholds in the HRV channel, is not transferrable to a complex multi-channel classification scheme such as SOFOS, the more promising method for 1 km FLS detection is their second approach; this uses the HRV channel's high-resolution information to sharpen the SEVIRI 3 km channels. However, this procedure has major limitations regarding the quality of the resulting FLS masks and needs comprehensive manual corrections so that it cannot be used for operational applications.

The main aim of the current paper is to develop an automatic method that can be used operationally without manual post-processing. The technique is based on a specific pan-sharpening approach, also including the SEVIRI thermal bands. This is an innovation in comparison to other common pan-sharpening algorithms (see Strait et al., 2008, for an overview). The newly developed technique is based on a local regression approach presented by Hill et al. (1999). Sharpened solar and thermal channels are then used to detect FLS by using the SOFOS approach.

Data and techniques used for this study are described in Sect. 2. The results are shown and discussed in Sect. 3. In Sect. 4, a conclusion is drawn and a short outlook given.

## 2 Data and methods

### 2.1 SEVIRI data

SEVIRI raw data distributed via EUMETCast (EUMETSAT, 2012) is operationally received at the Marburg Satellite Station. The data is decoded, unpacked and further processed by the FMet package (Cermak et al., 2008). Radiances in  $\text{Wm}^{-2}\text{sr}^{-1}$ , reflectance and blackbody temperatures (BBTs) in K are calculated for a section of the SEVIRI full disk showing Europe. Several products, such as the SOFOS 3 km FLS mask, are computed operationally from these data. In this study 144 day-time scenes from 17 November 2008, 10 December 2008, 22 December 2008 and 17 January 2011 showing Europe were incorporated.

### 2.2 METAR data

For validation purposes 6799 METereological Aerodrome Reports (METAR) from 21 points in time on the same days, obtained from a freely available online archive (Berge, 2012) were used. Geographic coordinates of the 982 European stations (Fig. 1) that have published these METARs were taken from Thompson (2011).

### 2.3 FLS identification via SOFOS

1 km FLS masks were created from pan-sharpened channels using SOFOS. SOFOS generally checks every pixel of a scene for FLS cloud properties using a range of tests (Cermak and Bendix, 2008, 2011). Most of these tests are conducted on a per-pixel basis:

- cloud identification;
- snow pixel elimination;
- test for liquid phase;

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- test for small droplet sizes.

Pixels which do not pass all of these tests are rejected as obviously non-FLS. Further tests checking for spatial parameters do not consider single pixels' properties but properties of spatially coherent entities of cloud. For this purpose spatially connected  
 5 pixels that were not rejected as obviously non-FLS are grouped together in entities. The following tests are performed on these entities:

- test for stratiformity;
- test for low clouds (cloud top height < 1000 m).

Entities which do not pass both tests are excluded from the resulting mask. Altogether  
 10 SOFOS utilizes reflectance 0.6, 0.8 and 1.6  $\mu\text{m}$  and BBTs at 3.9, 8.7, 10.8 and 12.0  $\mu\text{m}$  to derive FLS masks from SEVIRI data, which now has to be provided in 1 km spatial resolution by the pan-sharpening algorithm.

## 2.4 The pan-sharpening-algorithm

For global regression-based pan-sharpening as used by Bugliaro and Mayer (2004) the  
 15 HRV panchromatic image is degraded to match the resolution of the SEVIRI narrow-band channels. Afterwards the pixel values of the degraded panchromatic channel can be directly related to any other SEVIRI channel as shown in Fig. 2 for a solar and a thermal channel. A regression function computed on this basis can be used to calculate approximate narrow-band spectral values from the degraded broad-band HRV  
 20 channel or to calculate a high-resolution version of a given narrow-band channel from the original HRV channel as done by global regression pan-sharpening. The better the regression line fits the point cloud, the better the resulting image quality.

Obviously  $R^2$  is high (0.96530) for the displayed solar narrow-band channel but very  
 25 low (0.08098) for the thermal channel. This was to be expected, as the HRV is a solar channel. This is the reason why most pan-sharpening techniques are not applicable in the thermal regions. However, local features such as the contrast between cloud

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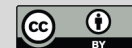
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and clear-sky regions register in both the solar and thermal regions of the spectrum. The fringe area of the clouds in the northwest of the scene shown in Fig. 3 is an example of this. Here, the radiance in the HRV channel fades from high values in pixels completely covered by thick clouds to low values in cloud free pixels. The pixels with a high fractional cloud cover have low BBTs in the  $10.8\text{ }\mu\text{m}$  channel while pixels with less and thinner clouds have higher BBTs. Thus, a negative correlation can be assumed between the  $10.8\text{ }\mu\text{m}$  channel and the HRV for this area. Therefore, local regressions based only on small areas of an image should be suitable for the pan-sharpening of thermal SEVIRI channels (Except SEVIRI's water vapor channels).

A window-based pan-sharpening technique using linear local regressions is described by Hill et al. (1999). The idea of the method is not to calculate one regression function for a whole scene but one regression function for the calculation of every single 3 km pixel. Each of these is based on the degraded HRV and narrow-band pixel values in a window of  $5 \times 5$  surrounding pixels. Similar to global regression pan-sharpening a 1 km narrow-band channel is generated from the HRV channel on this basis. In contrast to global regression pan-sharpening, however, not the same regression function is applied on all HRV pixels. For every HRV pixel the regression function that was calculated for the corresponding 3 km pixel is used instead (cf. Fig. 4).

Since this algorithm was originally developed for high-resolution remotely sensed data in the solar spectrum, originating from airborne sensors and polar orbiting satellites, an extension of the technique concerning the sharpening of thermal channels and the adaptation to SEVIRI's relatively low resolution was necessary. The following improvements to the pan-sharpening technique were realized, described in more detail below:

- (i) application of a potential regression;
- (ii) distance weighting;
- (iii) different combinations of window sizes and shapes.

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(i) The first improvement is related to the type of regression used by the pan-sharpening algorithm. In Fig. 5 the pixel values of solar and thermal channels were plotted against the degraded HRV channel for different single windows to determine the nature of the relationship between these channels. Some relationships (a and b) can be described well by linear as well as potential regressions. For other windows (c and d) they can much better be described by potential regressions and in a few windows there is no relationship between both channels at all (e). The latter case is reflected in locally deteriorated sharpening quality (see Sect. 3). The other findings suggest improving the algorithm by using a potential regression in the form of

$$y = a \cdot x^b \quad (1)$$

with

$$b = \frac{n \cdot \sum_{i=1}^n (\ln(x_i) \cdot \ln(y_i)) - \sum_{i=1}^n \ln(x_i) \cdot \sum_{i=1}^n \ln(y_i)}{n \cdot \sum_{i=1}^n \ln(x_i)^2 - \left( \sum_{i=1}^n \ln(x_i) \right)^2} \quad (2)$$

and

$$a = \frac{\sum_{i=1}^n \ln(y_i) - b \cdot \sum_{i=1}^n \ln(x_i)}{n} \quad (3)$$

where  $x_i$  is the degraded HRV value and  $y_i$  is the narrow-band channel value for a window pixel  $i$ .  $n$  is the number of pixels in the current window (some windows next to an image's border are cut off and therefore consist of fewer pixels).

(ii) The second change to Hill et al.'s (1999) algorithm was implemented to account for Tobler's (1970) first law of geography saying that near things are more related than distant things. In this context this means that a window's regression should particularly

be affected by pixels close to the window center. Therefore we introduced a weighting factor  $w_i$  for each window pixel  $i$  that depends on the Euclidean distance  $d_i$  (measured in pixels) to a window's central pixel:

$$w_i = \begin{cases} \frac{1}{d_i} & \text{if } d_i > 0 \\ \frac{1}{0.5} & \text{if } d_i = 0 \end{cases} \quad (4)$$

- 5 Since  $d_i$  is 0 for the central pixel and therefore this pixel would be weighted infinitely,  $d_i$  is set to 0.5 for these pixels. To account for  $w_i$  in such a way that window pixels which are nearer to the window's center have a stronger influence on the central pixel Eqs. (2) and (3) were altered, such that:

$$b = \frac{\sum_{i=1}^n w_i \cdot \sum_{i=1}^n (\ln(x_i) \cdot \ln(y_i) \cdot w_i) - \sum_{i=1}^n (\ln(x_i) \cdot w_i) \cdot \sum_{i=1}^n (\ln(y_i) \cdot w_i)}{\sum_{i=1}^n w_i \cdot \sum_{i=1}^n (\ln(x_i)^2 \cdot w_i) - \left( \sum_{i=1}^n (\ln(x_i) \cdot w_i) \right)^2} \quad (5)$$

10 and

$$a = \frac{\sum_{i=1}^n (\ln(y_i) \cdot w_i) - b \cdot \sum_{i=1}^n (\ln(x_i) \cdot w_i)}{\sum_{i=1}^n w_i} \quad (6)$$

(iii) Since relationships between the degraded HRV channel and narrow-band channels are spatially limited, windows of only  $3 \times 3$  pixels in size, and windows of a roughly round shape were tested (round windows cannot be achieved with only a few pixels).

- 15 Besides the use of a potential regression and the changes regarding the window shape and size, the basic idea of the pan-sharpening-algorithm described by Hill et al. (1999) and shown in Fig. 4 was not altered.

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## 2.5 Methodology of validation

Several variations of the pan-sharpening approach described in Sect. 2.4 are possible since the different window sizes and shapes, distance weighting and the two types of regression can be combined in various ways. To find the best-suited version for 1 km FLS detection, several steps of validation were necessary. First, the spectral as well as the spatial quality of pan-sharpened channels were measured for each possible combination (see Sect. 2.5.1). The most promising variations of the pan-sharpening technique were used to sharpen SEVIRI channels for FLS masking. These masks were compared to reference masks and METARs (see Sect. 2.5.2). The additional METAR validation was necessary to account for some special characteristics of SOFOS regarding its reaction to an increased resolution of its input channels.

For the statistical validation of the pan-sharpened channels and the comparison of FLS masks with reference masks 144 daylight scenes under various illumination conditions from (randomly chosen) FLS appearances on 17 November 2008, 10 December 2008, 22 December 2008 and 17 January 2011 were utilized. For the METAR validation the number of scenes was reduced to 21 to avoid overload of the used online archive.

Besides the validation using statistical measures, the quality of sharpened channels and FLS masks based on them was additionally assessed visually for individual scenes. FLS masks based on channels the resolution of which was increased by using a nearest-neighbour interpolation were included in this visual validation to be able to judge to what degree differences between 3 km and 1 km channels were caused by SOFOS' above-mentioned reaction to high-resolution input channels.

The validation study on the performance of the pan-sharpening was conducted including all seven SEVIRI channels used as input for SOFOS (see Sect. 2.3).

### 2.5.1 Validation of pan-sharpened channels

In order to validate the spectral quality of the sharpened channels the average root mean square error (RMSE, see Strait et al., 2008) was calculated for all 144 scenes.

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Since it is a measure for the average difference between a data series and a reference data series, it can be used for measuring the difference between the pixel values of an image and the corresponding pixel values in a reference image. As it is in the nature of pan-sharpening that images in a resolution in which the corresponding channels did not exist before are created, the RMSE cannot be calculated in the SEVIRI 1 km resolution for lack of a reference image. To overcome this problem a method suggested by Thomas and Wald (2004) was utilized. According to their approach (hereinafter referred to as “validation approach A”) the narrow-band channels as well as the HRV channel are downsampled by the factor 3. The 3 km HRV channel obtained in this way is used to sharpen the degraded 9 km narrow-band channels. The resulting 3 km narrow-band channels can be compared to the original ones as reference. The area on the Earth’s surface covered by a window of a certain pixel size is dependent on the resolution of the image. As spatial aspects are taken into account by the new pan-sharpening algorithm in many ways, the validation via approach A, in which the pan-sharpening algorithm is applied on channels in a reduced resolution, may lead to wrong results regarding the window shape and size as well as the distance weighting. Therefore in a second validation approach (further referred to as “validation approach B”) the narrow-band channels were sharpened to the HRV channel’s resolution and afterwards degraded to their original size, in which they were compared to the unaltered channels as reference images. Validation approach B has the disadvantage that information from individual pixels is lost when the sharpened images are degraded to the original resolution. Thus, errors affecting only isolated pixels cannot be accounted for by this approach. For this reason both approaches were used.

The spatial quality, a measure of the degree of similarity between the geometry of a sharpened channel and a reference image, for the different variations of the pan-sharpening algorithm was calculated using a slightly modified variation of a method developed by Zhou et al. (1998). By application of a Laplacian filter using

the convolution mask:

$$D_{xy} = \begin{bmatrix} -1 & -1 & -1 \\ -1 & 8 & -1 \\ -1 & -1 & -1 \end{bmatrix} \quad (7)$$

the high-frequency component of each sharpened channel was extracted and compared to the high-frequency component of a reference image using the Pearson product-moment correlation coefficient of the pixel values of both high frequency images. Since Zhou et al. (1998) apply their method on sharpened solar channels only, they use the panchromatic channel as the reference image. This method is not suited for thermal channels as their high frequency components differ from those of solar channels. So, analogous to validation approach A, the high frequency component was calculated for channels that were first degraded and sharpened afterwards, and compared to the high frequency component of the unaltered channels. The average value of the product-moment correlation was calculated for all 144 scenes.

## 2.5.2 Validation of FLS masks

FLS masks based on sharpened channels were validated in two different ways. First, channels treated according to validation approaches A and B were used to create 3 km FLS masks via SOFOS. These masks were compared with reference masks calculated by SOFOS from unaltered 3 km channels. Confusion matrices (cf. Table 1) based on each pixel and each of the 144 scenes used were calculated. From these tables the following statistical measures (Wilks, 1995; Joliffe and Stephenson, 2003) were calculated (see Appendix A for formulas):

- Proportion Correct (PC);
- Bias;
- Probability of Detection (POD);

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- Probability of False Detection (POFD);
- False Alarm Rate (FAR);
- Hanssen-Kuipers Discriminant (HKD);
- additionally a new measure, the Edge Precision (EP), was developed. It describes how well the positions of edge pixels of FLS entities in the validated masks fit the positions of edge pixels of the corresponding entities in the reference mask. For this purpose the convolution mask  $D_{xy}$  was applied on each mask and each reference mask, in order to gain images showing only the edges of each mask's FLS entities. For these edge images and reference edge images contingency tables were created and the POD, which in this context describes the fraction of edges pixels that are on the correct position, was calculated.

For purposes of comparison these measures were also calculated for channels the resolution of which was increased after degradation by using a simple nearest-neighbour interpolation instead of pan-sharpening. This was done for validation approach A only since approach B would gain perfect results for interpolated channels.

The second approach to validate the FLS masks was necessary as SOFOS shows a tendency to identify different (mostly larger) areas as FLS when the input channels' resolution is increased. This tendency is not caused by spectral errors that appear during sharpening as it can also be observed for channels the resolution of which was increased by simple nearest-neighbour interpolation (see Sect. 3). To check whether this behaviour affects the FLS masks negatively METAR data was utilized. A contingency table was calculated for FLS masks based on 1 km pan-sharpened as well as on unaltered 3 km channels using the METARs as reference for all scenes used. From these tables the same statistical measures as used for the first approach for FLS validation were calculated (except the Edge Precision, as it does not make sense for the comparison with point data).

A principal problem when validating a satellite based cloud product by making use of METAR data is mentioned by Cermak and Bendix (2007): as the instruments at

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ground-based stations publishing METARs are below the clouds they have another perspective on the clouds than satellites. To eliminate this problem as far as possible, not only the cloud base height but also the fractional cloud cover, the range of sight and (if included) the cloud type were utilized. Assuming a typical cloud thickness of 200 m, FLS METARs with cloud base heights of less than 800 m were interpreted as showing FLS presence if no cumulonimbus or cumulus congestus with a large vertical extent had been reported. Additionally, METARs reporting a line of sight of less than 1000 m were interpreted as indicating ground fog. Reports mentioning additional cloud layers above the 800 m threshold were interpreted as FLS-free since the additional layers would block the view from the satellite on underlying layers. Finally METARs with a fractional cloud coverage of less than 3/8 were interpreted as free of FLS since it is very unlikely for such a low coverage that the respective pixel would be cloud-covered.

Due to the small sample size the METAR validation here should not be seen as a general validation of SOFOS' sharpening quality. For a detailed METAR validation of SOFOS based on 1030 scenes see Cermak's (2006) original work.

## 3 Results and discussion

### 3.1 The quality of sharpened channels

The validation of the sharpened channels showed that the use of a potential regression as well as distance weighting significantly improve the pan-sharpening quality, regardless of the window size and shape. The validation results for the original algorithm by Hill et al. (1999) and the most promising variations of the new algorithm are given in Table 2. The spectral quality is the highest if potential regression and distance weighting are used in combination with a  $3 \times 3$  pixels "round" window (further referred to as 3r). The algorithm version using a  $5 \times 5$  pixels square window (further referred to as 5s), which yields a lower spectral quality (but is still better than that of the unaltered algorithm by Hill et al. (1999)), yields the highest spatial quality. The spatial as

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well as the spectral quality of the other algorithm versions using potential regression and distance weighting ( $5 \times 5$  pixels round window and  $3 \times 3$  pixels square window, not shown in Table 2) lies between the quality of 3r and 5s. These findings are not surprising: the spectral quality improves for smaller windows consistent with the consideration that relationships between the HRV channel and the narrow-band channels are mostly spatially limited. As the sample size of each regression gets smaller for smaller windows, the regression calculated a 3 km pixels is of low significance in some cases and cannot be used for every corresponding 1 km pixel. The resulting failures occur rarely and only affect single pixels, which is why the overall spectral quality is hardly affected. Since these single pixel errors have impact on the high frequency component of an image, they strongly affect the spatial quality. Due to the trade-off between the spatial and the spectral quality the window size and shape has to be chosen depending on the intended use. In order to check whether the spatial or spectral quality of SOFOS input channels is more important for the quality of the FLS detection the validation of the resulting masks was necessary to choose the best version of the pan-sharpening algorithm.

Figure 6 shows the quality of the new algorithm using 5s and 3r for the sharpening of each of the SOFOS input channels. For the 3r version the mentioned errors in individual pixels can be clearly identified. Since the lowered spectral quality of the 5s variation cannot be recognized by the human eye, a  $5 \times 5$  pixels window seems to be a suitable choice for visualization purposes. The scene shows fog in the Po Valley in the center and the snow-covered Alps in the north. North of the fog in the plain Lake Iseo can be recognized. For each channel these different surfaces were sharpened well. The fog margin is very sharp and even small valleys, along which the fog reaches into the northern parts of the Apennine Mountains, can be recognized. In the south east of the scene there are high clouds above the FLS, which are shown in dark shades in the thermal channels (low temperature). In the thermal spectrum the spatial quality of the scene is lowered for these clouds by some weakly pronounced rectangular artifacts in the original 3 km resolution. These artifacts are caused by a situation as shown in

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Fig. 5e. As no relationship between the narrow-band channel and the degraded HRV channel can be established, a more or less horizontal regression function is assumed, which is why for these areas quite similar values are assigned to all 1 km pixels covered by the same 3 km pixel. These flaws do not affect the spectral quality. Further they are rare and can mostly be observed for high clouds, especially for cumuliform clouds where the brightness of the HRV channel is strongly influenced by the illumination geometry. Therefore they should only minimally affect the quality of FLS masks based on sharpened channels. In addition to Fig. 6 the pan-sharpening quality of 5s is demonstrated for a thermal channel in Fig. 7 showing an FLS entity covering Lake Constance and adjacent valleys.

## 3.2 The quality of FLS masks based on sharpened channels

The validation results for FLS masks based on sharpened SEVIRI channels are given in Table 3 for the original sharpening technique by Hill et al. (1999) as well as for 3r and 5s. Additionally the quality measures according to validation approach A (see Sect. 2.5.2) are given for the masks based on interpolated channels. The PC is almost 1 for each pan-sharpening technique as well as for each validation approach. This is no surprise since even the simple nearest-neighbour interpolation achieves such high values. The reason for that is that the majority of the pixels of each mask are far away from the boundaries of any FLS entity, but the resolution enhancement is only important to classify pixels near boundaries correctly. Since the interpolation does not cause any spectral errors, the mask quality inside and outside of the FLS entities is very good, while it is low on boundary pixels. All in all this causes a high PC, which is slightly better than that of the original algorithm by Hill et al. (1999) and the 5s variation. The same can be observed for the bias, the POD, the POFD, the FAR and the HKD. Although Figs. 6 and 7 show that the spatial quality of the 5s method is much higher than that of a simple interpolation, even the Edge Precision (EP) of sharpened masks is lower than that of masks based on interpolated channels. All of these findings are most probably caused by an insufficient spectral quality of the original algorithm and the 5s version.



This hypothesis is supported by the fact that only the 3r variation, which has the best spectral quality, outperforms the interpolation for all statistical measures. In addition this variant delivers a spatial quality (cf. Fig. 6) that is much higher than that of a simple interpolation and thus gains the best Edge Precision. Therefore it is the best choice for the sharpening of SOFOS input channels.

### 3.3 Comparison of 3 km and 1 km FLS masks

In order to determine whether the mentioned differences between the FLS area in masks on base of 3 km channel and on base of 1 km channels are altogether advantageous or disadvantageous, the METAR validation was performed. The results are shown in Table 4. As the bias is low for both resolutions, the area of FLS is apparently underestimated by SOFOS in general. The enhanced overall quality (PC and HKD) and a higher probability of detection in combination with a bias closer to 1 imply that the masks' quality is enhanced due to a lowered degree of underestimation. The POD, on the other hand, is a little bit higher (but still on a low level) for the 1 km masks, which means that the area of FLS is overestimated in some additional cases. Despite this finding the FAR is slightly better for the 1 km resolution. So, all in all, the FLS masks do strongly benefit from SOFOS' reaction to resolution enhancement.

Additional insights into the degree of enhancement of FLS masks via pan-sharpening can be obtained from Figs. 8–12. Figure 8 shows fog on 17 January 2011 covering Lake Constance in the north east and Lake Geneva and Lake Neuchâtel in the south west. The outlines of FLS masks based on 3 km SEVIRI channels as well as on channels that were sharpened by the 3r method are shown. In both resolutions the FLS has been detected very well. Nevertheless, the quality of the 1 km mask is much higher, since fine details that cannot be displayed at the 3 km resolution can be found along the borders of the FLS entities. For example the fog-covered Thur Valley (white arrow) south of Lake Constance can hardly be recognized in the 3 km resolution, while its shape can be identified precisely in the 1 km mask. Figure 9 shows a fog occurrence on the same date in Central Germany. Again, many details (white arrows), especially

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if fog and snow or other clouds are situated directly side by side, they fell into common windows and therefore the algorithm tries to describe the dependency of a narrow-band channel from the pan channel by the same regressions for both materials. Since this is not reasonable in many cases a situation similar to that shown in Fig. 5e could occur and therefore drawbacks in the spatial quality of the sharpened channels and in the resulting mask should be possible. 3 km artifacts in the 1 km FLS mask caused by this could not be observed for the margins of snow-covered areas or for any entity of ground fog. In the margin area of FLS without ground contact, however, this kind of artifacts could be observed in rare cases (cf. Fig. 12).

A general disadvantage of every algorithm using local regressions is the fact that only those pixel types can be sharpened that are accounted for by the regression used to sharpen them. This means that a 1 km narrow-band pixel can only be calculated from the corresponding HRV pixel by using a regression that was calculated on the basis of at least one 3 km pixel the value of which is linked to the HRV pixel values in the same way as the 1 km narrow-band pixel value. Thus, FLS entities that are not visible in the 3 km resolution and not in the same window as bigger entities cannot be detected by this approach.

## 4 Conclusion and outlook

A method for the creation of 1 km fog and low stratus (FLS) masks from MSG SEVIRI data was presented. The method uses a newly developed pan-sharpening algorithm suitable to enhance the spatial quality of solar as well as thermal 3 km SEVIRI channels to 1 km by using the instrument's panchromatic HRV channel. The resulting 1 km synthetic narrow-band channels are used as input for the well-validated FLS detection scheme SOFOS in order to produce 1 km FLS masks. The pan-sharpening approach strongly improves the spatial quality of the resulting masks. As the shape of many small valleys cannot be distinguished at a 3 km spatial resolution, this is especially important for the proper mapping of FLS entities delimited by topography. The new approach

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combines a relatively high spatial resolution with the high temporal resolution of a geo-stationary instrument. Therefore it offers new possibilities for short-range forecasting and the creation of fog climatologies. Errors in the high-resolution FLS masks caused by a low spectral quality of the sharpened input channels could not be found. The only flaws which can be traced back to the pan-sharpening algorithm are local reductions of the spatial quality reflected in rectangular artifacts in the mask border. These artifacts are rare and could only be observed for the fuzzy border of FLS without ground contact. As they never reduce the spatial quality of the mask to a level lower than that of 3 km masks, a sharpening of SOFOS input channels using the improved pan-sharpening algorithm seems to be a suitable method to gain high quality 1 km FLS masks.

Meteosat Second Generation is going to be replaced by the Meteosat Third Generation in 2015 (Bensi et al., 2004). Several adaptations of the pan-sharpening algorithm to the MTG's main instrument Flexible Combined Imager (FCI) are conceivable. As the FCI will have 5 panchromatic channels, a variation of the algorithm using multiple regressions would be useful to improve the sharpening quality. Additionally multiple panchromatic channels could be utilized to classify pixels by their spectral signature. Therefore not only spatially adjacent pixels but also pixels that are similar to each other in their spectral signatures could be grouped together for each regression. For each spectral signature in each window an individual regression could be calculated. When the regressions are used to transfer panchromatic channel information into the high resolution version of a narrow-band channel, there would be several regressions to be chosen from depending on the spectral signature of each 1 km pixel.

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## Appendix A

### Formulas utilized for the validation of FLS masks

The following formulas were used for calculating of the statistical measure used in Sect. 2.5.2:

$$5 \quad PC = \frac{n_{11} + n_{00}}{n_{11} + n_{10} + n_{01} + n_{00}} \quad (A1)$$

$$\text{Bias} = \frac{n_{11} + n_{01}}{n_{11} + n_{10}} \quad (A2)$$

$$\text{POD} = \frac{n_{11}}{n_{11} + n_{10}} \quad (A3)$$

$$\text{POFD} = \frac{n_{01}}{n_{01} + n_{00}} \quad (A4)$$

$$\text{FAR} = \frac{n_{01}}{n_{11} + n_{01}} \quad (A5)$$

$$10 \quad \text{HKD} = \text{POD} - \text{POFD} \quad (A6)$$

*Acknowledgements.* This research was funded by a grant from the German Research Council DFG (BE 1780/14-1).

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**Table 1.** Layout of the confusion matrices used to calculate the different statistical measures.

FLS in reference mask	FLS in validated mask	
	True/1	False/0
True/1	$n_{11}$	$n_{10}$
False/0	$n_{01}$	$n_{00}$

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**Table 2.** Validation results for the spectral and spatial quality of different variations of the pan-sharpening approach. The RMSE is given in absolute values (reflectance for solar channels, blackbody temperature for thermal channels) as well as in % of the mean pixel value of each validated scene.

Method of validation	0.6 $\mu\text{m}$	0.8 $\mu\text{m}$	1.6 $\mu\text{m}$	3.9 $\mu\text{m}$	8.7 $\mu\text{m}$	10.8 $\mu\text{m}$	12.0 $\mu\text{m}$
RMSE, approach A							
Hill et al. (1999)	0.038/11.147 %	0.046/14.476 %	0.047/16.457 %	4.576/1.677 %	2.905/1.094 %	3.096/1.160 %	3.134/1.176 %
5s	0.037/10.749 %	0.045/12.152 %	0.044/15.518 %	4.268/1.564 %	2.471/0.931 %	2.620/0.981 %	2.645/0.993 %
3r	0.037/10.719 %	0.045/12.093 %	0.042/14.887 %	4.025/1.475 %	2.134/0.803 %	2.243/0.840 %	2.254/0.846 %
RMSE, approach B							
Hill et al. (1999)	0.029/8.422 %	0.036/9.652 %	0.032/11.383 %	2.814/1.031 %	1.437/0.541 %	1.493/0.559 %	1.497/0.562 %
5s	0.025/7.324 %	0.031/8.394 %	0.028/9.846 %	2.546/0.934 %	1.193/0.415 %	1.141/0.427 %	1.142/0.428 %
3r	0.019/5.339 %	0.023/6.096 %	0.021/7.383 %	2.118/0.777 %	0.570/0.214 %	0.580/0.217 %	0.580/0.217 %
Spatial quality							
Hill et al. (1999)	0.689	0.665	0.635	0.185	0.381	0.389	0.368
5s	0.693	0.668	0.644	0.217	0.403	0.413	0.396
3r	0.683	0.657	0.618	0.197	0.361	0.369	0.354

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**Table 3.** Statistical measures describing the quality of FLS masks based on channels that were sharpened by different variations of the pan-sharpening algorithm as well as on interpolated channels.

Method of validation	PC	Bias	POD	POFD	FAR	HKD	EP
Validation approach A							
Hill et al. (1999)	0.933	0.987	0.746	0.038	0.244	0.708	0.336
5s	0.936	1.045	0.787	0.040	0.246	0.747	0.356
3r	0.950	1.0002	0.816	0.029	0.184	0.787	0.425
Interpolation	0.948	1.001	0.807	0.030	0.194	0.777	0.372
Validation approach B							
Hill et al. (1999)	0.956	1.011	0.842	0.026	0.167	0.816	0.502
5s	0.966	1.013	0.880	0.021	0.132	0.859	0.612
3r	0.974	1.002	0.904	0.015	0.098	0.888	0.703
Best Possible Result	1	1	1	0	0	1	1

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Statistical measure	3 km masks	1 km masks
PC	0.574	0.600
Bias	0.197	0.378
POD	0.654	0.658
POFD	0.057	0.107
FAR	0.346	0.342
HKD	0.072	0.141



**Fig. 1.** Positions of the METARs stations used in this study.

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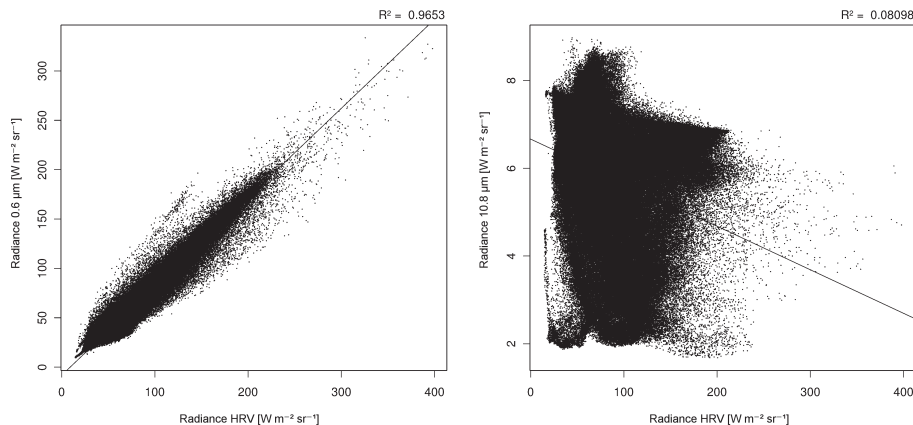
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**Fig. 2.** Relationship between the HRV radiance and the radiance at 0.6  $\mu\text{m}$  (left), and 10.8  $\mu\text{m}$  (right) for each pixel of a section of the SEVIRI scene of 17 January 2011 showing Europe.

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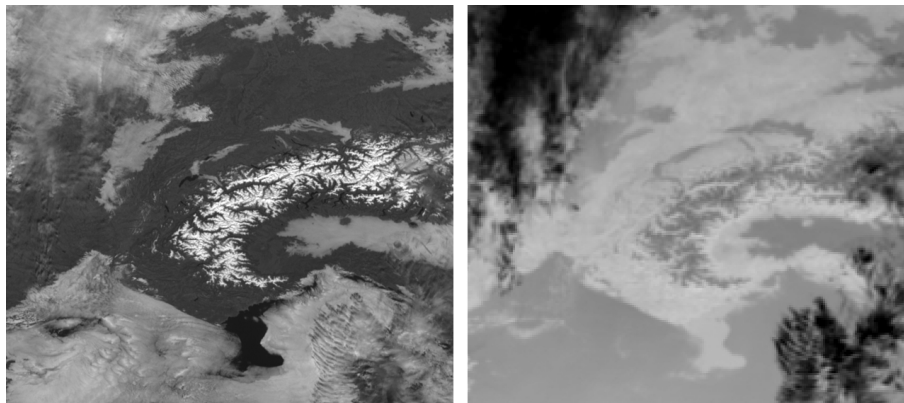
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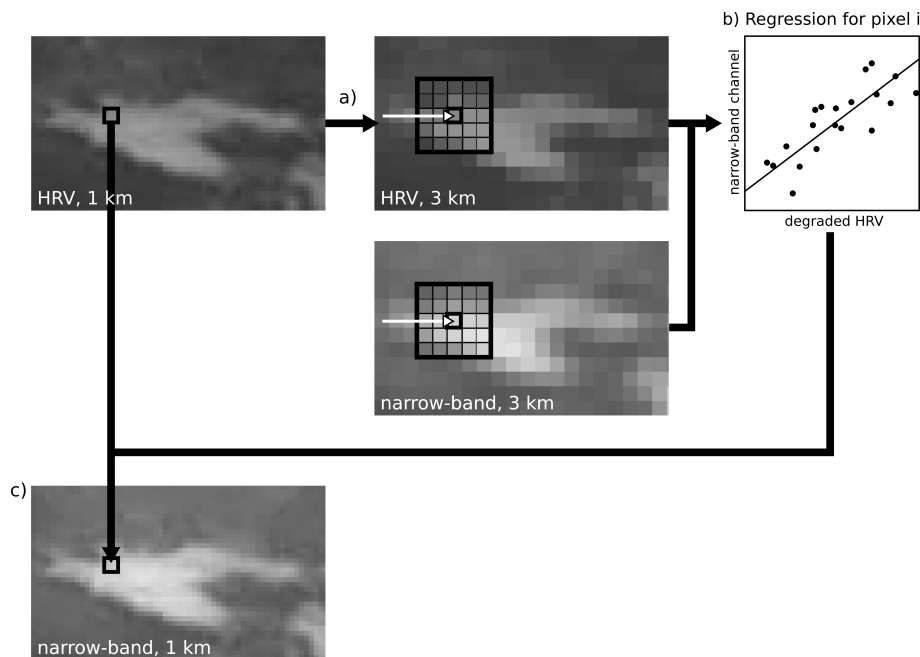


**Fig. 3.** HRV channel (left) and  $10.8\,\mu\text{m}$  channel (right) for a section of the SEVIRI scene from 17 January 2011 showing the Alps in the center of the image and Mediterranean Sea in the south.

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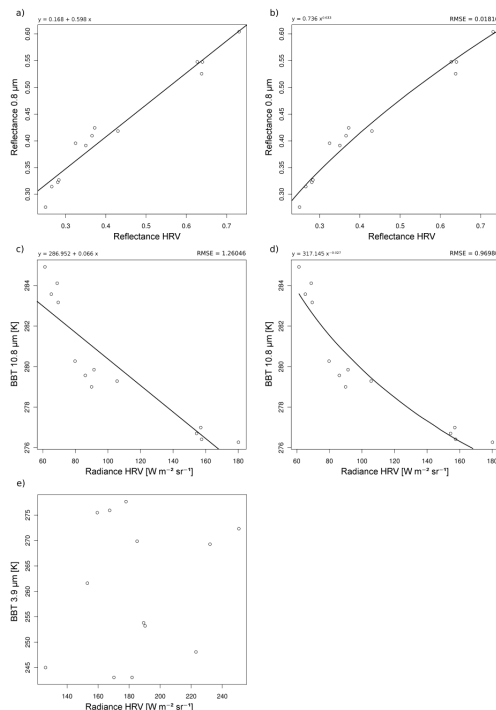


**Fig. 4.** The pan-sharpening algorithm adapted from Hill et al. (1999). After degradation of the HRV channel **(a)** a regression for each 3 km pixel is calculated from the values of narrow-band pixels and degraded HRV pixels covered by a  $5 \times 5$  pixels window **(b)**. Each regression is used to calculate the values for  $3 \times 3$  1 km narrow-band pixels (corresponding to one 3 km pixel) from  $3 \times 3$  HRV pixels **(c)**.

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**Fig. 5.** Types of relationships between different narrow-band channels and the HRV channel for single windows in a SEVIRI scene from 30 September 2003. The relationship shown in the upper row can be described well by a linear **(a)** and a potential regression **(b)**. The relationship shown in the middle row is better described by a potential regression **(d)** than by a linear regression **(c)**, while the relationship shown in **(e)** cannot be explained by the HRV values at all. Please note that (as reflectance channels were sharpened using the HRV reflectance while the HRV radiance was used to sharpen BBT channels in our study) the narrow-band reflectance values are plotted against the HRV reflectance and the narrow-band BBT values are plotted against HRV radiance values.

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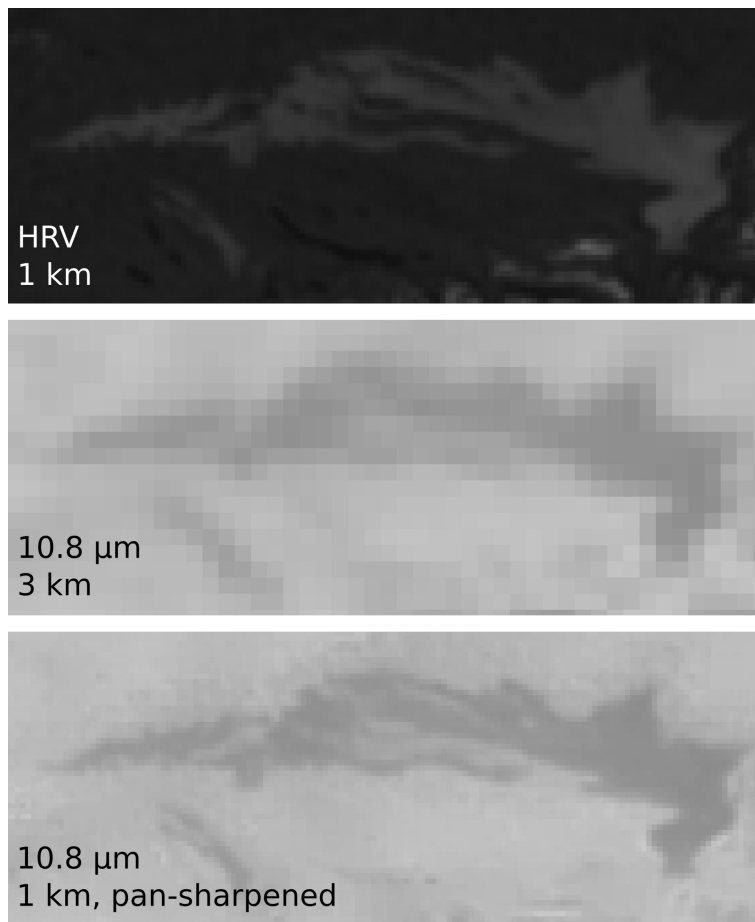
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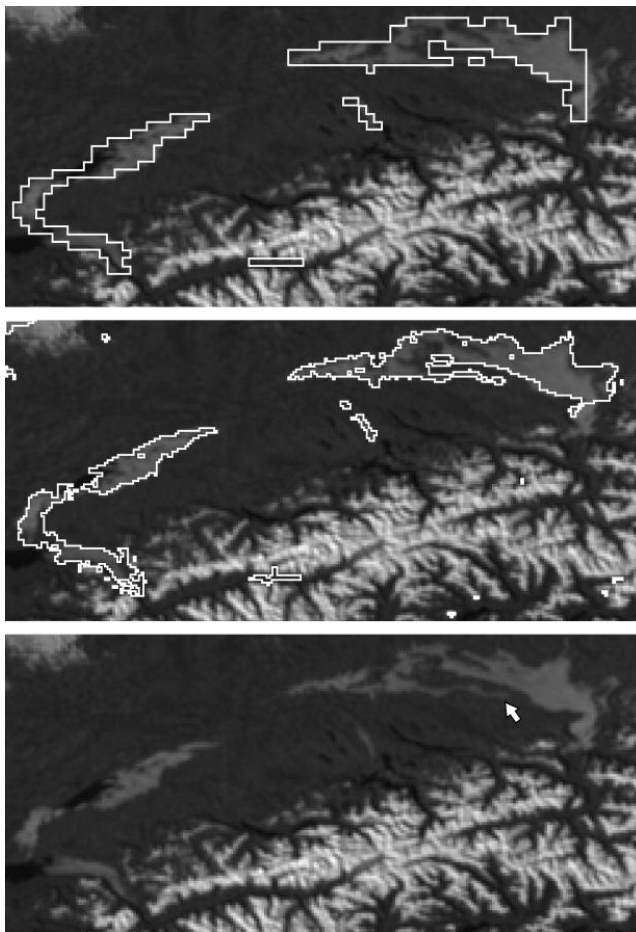
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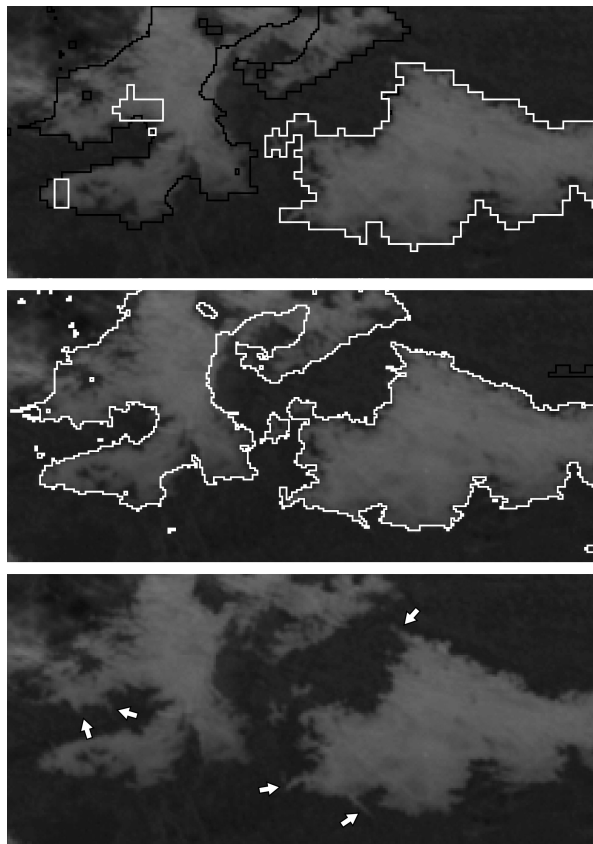




**Fig. 7.** Fog covering Lake Constance and adjacent valleys on 17 January 2011. In order to achieve better comparability the 3 km channels were interpolated to the size of the 1 km images via nearest-neighbour interpolation.



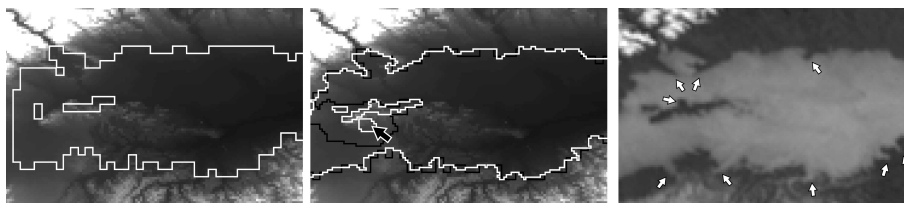
**Fig. 8.** 3 km (top) and 1 km (middle) FLS mask for fog in Southern Germany and Switzerland on 17 January 2011. White line = boundary of the masks. White arrow = detail that strongly benefits from the resolution enhancement.



**Fig. 9.** 3 km (top) and 1 km (center) FLS for fog in Central Germany on 17 January 2011. White line = boundary of the 3 km and 1 km masks. Black line = boundary of an additional FLS mask based on channels interpolated to 1 km resolution (notice that this mask matches the 3 km mask perfectly for the eastern FLS entity, which is why the black outline is not visible here). White arrows = details that strongly benefit from the resolution enhancement.

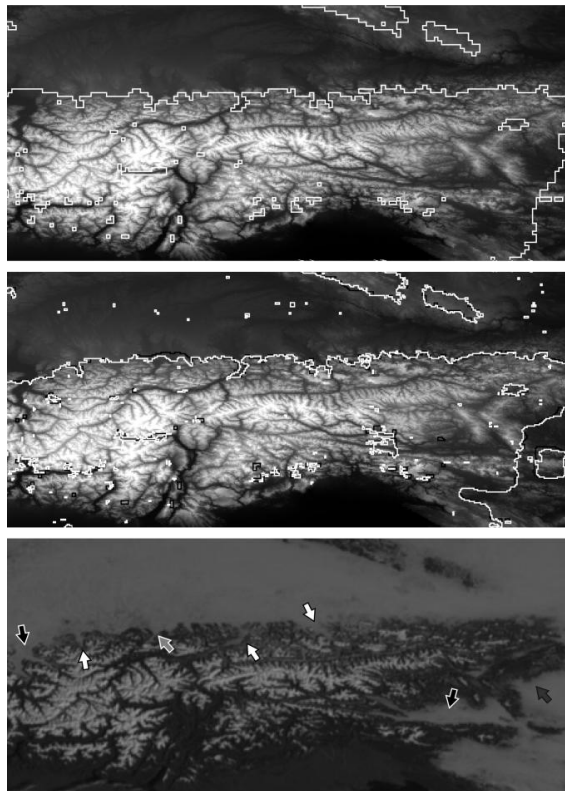
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**Fig. 10.** FLS masks (indicated by the white outline) in 3 km (left) and 1 km (center) resolution for fog in the Po Valley on 22 December 2008. The masks' outlines are underlaid by a DEM. White line = boundary of the 3 km and 1 km masks. Black line = boundary of an additional FLS mask based on channels interpolated to 1 km resolution. White arrows = details that strongly benefit from the resolution enhancement. Black arrow = area where FLS was not detected.

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**Fig. 11.** 3 km (top) and 1 km (center) FLS masks (indicated by the white outline) for fog north of the Alps on 2 February 2011. The masks' outlines are underlaid by a DEM. White line = boundary of the 3 km and 1 km masks. Black line = boundary of an additional FLS mask based on channels interpolated to 1 km resolution. White arrows = details that strongly benefit from the resolution enhancement. Black arrows = Areas that were not identified as FLS regardless of the resolution. Light grey arrow = an area that was only detected as FLS in 3 km resolution. Dark grey arrow = an area that was only detected as FLS in 1 km resolution.

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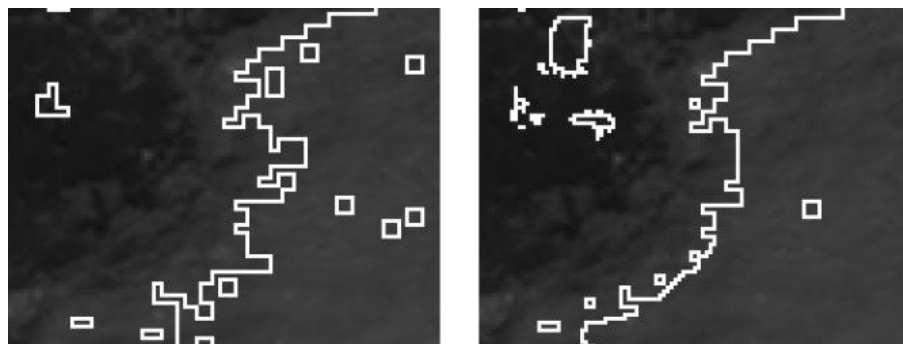
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**Fig. 12.** FLS masks (indicated by the white outline) for low stratus without ground contact over Central France on 10 December 2008 based on 3 km (left) and 1 km (right) SEVIRI channels. The spatial quality of the 1 km mask is reduced by rectangular artifacts in the 3 km resolution.