



**Solar channel
inter-calibration
using SEVIRI**

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This discussion paper is/has been under review for the journal Atmospheric Measurement Techniques (AMT). Please refer to the corresponding final paper in AMT if available.

Inter-calibration of polar imager solar channels using SEVIRI

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Received: 6 February 2013 – Accepted: 26 March 2013 – Published: 3 April 2013

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Published by Copernicus Publications on behalf of the European Geosciences Union.

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Abstract

Accurate calibration of satellite imagers is a prerequisite for using their measurements in climate applications. Here we present a method for the inter-calibration of geostationary and polar-orbiting imager solar channels based on regressions of collocated near-nadir radiances. Specific attention is paid to correcting for differences in spectral response between instruments. The method is used to calibrate the solar channels of the Spinning Enhanced Visible and Infrared Imager (SEVIRI) on the geostationary Meteosat satellite with corresponding channels of the Moderate Resolution Imaging Spectroradiometer (MODIS) on the polar-orbiting Aqua satellite. The SEVIRI operational calibration is found to be stable during the years 2004 to 2009 but off by -8 , -6 , and $+3.5$ % for channels 1 ($0.6 \mu\text{m}$), 2 ($0.8 \mu\text{m}$), and 3 ($1.6 \mu\text{m}$), respectively. These results are robust for a range of choices that can be made regarding data collocation and selection, as long as the viewing and illumination geometries of the two instruments are matched. Uncertainties in the inter-calibration method are estimated to be 1 % for channel 1 and 1.5 % for channels 2 and 3. A specific application of the method is the inter-calibration of polar imagers using SEVIRI as a transfer instrument. This offers an alternative to direct inter-calibration, which in general has to rely on high-latitude collocations. Using this method we have tied MODIS-Terra and Advanced Very High Resolution Radiometer (AVHRR) instruments on National Oceanic and Atmospheric Administration (NOAA) satellites 17 and 18 to MODIS-Aqua for the years 2007 to 2009. While reflectances of the two MODIS instruments differ less than 2 % for all channels considered, deviations of an existing AVHRR calibration from MODIS-Aqua reach -3.5 and $+2.5$ % for the 0.8 and $1.6 \mu\text{m}$ channels, respectively.

1 Introduction

Over the past decades many different meteorological imaging instruments have been put in space on board of polar and geostationary satellites. Initially intended for

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qualitative imagery, these instruments have increasingly also been used for the quantitative retrieval of atmospheric and surface properties. With the latter application and the demand for stable time series, accurate calibration of the measured radiances has become essential. This holds in particular for the solar channels, which often lack on-board calibration facilities.

In general, two in-flight calibration approaches can be distinguished: (1) using carefully selected targets in combination with radiative transfer modelling to simulate the imager radiances or (2) comparing with radiances measured by other satellite or aircraft instruments. In the first approach a wide range of targets, including deserts, ice caps, deep convective clouds, and the moon, can be used as a reference (e.g. Vermote and Kaufman, 1995; Smith et al., 2002; Heidinger et al., 2003; Govaerts et al., 2004; Doelling et al., 2013a). While this is a powerful method, and the only option for the early imager era, it does have limitations, notably the assumed stability of the reflectivity of the targets and the accuracy of radiative transfer modelling. Furthermore, no single target type can be applied over the full range of satellite instruments (geostationary and polar orbiting) and solar channels.

The second approach, inter-calibration of satellite instruments, is normally based on simultaneous nadir overpasses (SNOs) for which near-nadir, collocated radiances are collected, without the need for specific target selection. SNOs have been applied both to pairs of polar imagers (e.g. Minnis et al., 2002, 2008; Heidinger et al., 2002, 2010; Cao et al., 2008) and polar-geostationary instrument combinations (e.g. Minnis et al., 2002; Roebeling et al., 2006; Ham and Sohn, 2010). Although this method is not hampered by some of the limitations of the “target” approach, it does require consensus on a well-calibrated reference instrument. The Moderate Resolution Imaging Spectroradiometer (MODIS) instruments onboard the Terra and Aqua platforms appear to be good candidates given their on-board calibration devices (Xiong and Barnes, 2006).

In this paper, an SNO method for polar-geostationary imager solar channel inter-calibration is presented. An important requirement for such a method is that differences in spectral response functions (SRFs) between the imagers are taken into account.

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Recently, Doelling et al. (2013b) employed Scanning Imaging Absorption Spectrometer for Atmospheric Cartography (SCIAMACHY) hyperspectral radiances for this purpose. Here, we use radiative transfer calculations to estimate pixel-by-pixel SRF corrections. The method is applied to calibrate the three solar channels of the geostationary Spinning Enhanced Visible and Infrared Imager (SEVIRI) on Meteosat with MODIS-Aqua. The sensitivity of the results to assumptions underlying the method, including collocation, sampling, and viewing/illumination geometry selection criteria, is assessed.

Direct inter-calibration of polar-orbiting sensors usually has to rely on SNOs at high latitudes. Potential drawbacks of these SNOs are a limited dynamic range of measured radiances and problems in characterizing scenes over bright snow- or ice-covered surfaces. Similar to Wang et al. (2011), who applied double differencing to inter-calibrate infrared radiances of two polar-orbiting sounders, we explore the use of the SEVIRI instrument to transfer calibrations from one polar imager to another. In this way, MODIS-Terra as well as two Advanced Very High Resolution Radiometers (AVHRRs) are tied to MODIS-Aqua, providing a means to verify existing calibrations based on direct SNOs.

The work described here is closely linked to the Global Space-based Inter-Calibration System (GSICS) effort (Goldberg et al., 2011), which aims to monitor, improve and harmonize the quality of observations from operational weather and environmental satellites.

The setup of this paper is as follows. The measurements are introduced in Sect. 2. Then the calibration approach is outlined in Sect. 3. Results of the inter-calibration of SEVIRI and MODIS are presented in Sect. 4, followed by a sensitivity analysis in Sect. 5. Section 6 describes the use of the method for polar imager inter-calibration. Conclusions are finally drawn in Sect. 7.

2 Measurements

This section starts with some definitions and notation of solar channel measurements, after which information is provided on the three satellite instruments that are considered in this paper: SEVIRI, MODIS, and AVHRR.

2.1 Terminology

The calibration of satellite imager solar channels centres around the determination of the calibration slope S and dark count D in the relation

$$I = S(C - D), \quad (1)$$

where C is the measured count and I is the radiance in the direction of the satellite in a specific channel. The units of I and S used here are $\text{Wm}^{-2}\text{sr}^{-1}\mu\text{m}^{-1}$. I is obtained by integrating the spectral radiance $I_s(\lambda)$ weighted by the instrument spectral response function (SRF), $\xi(\lambda)$, over wavelength λ :

$$I = \frac{\int \xi(\lambda) I_s(\lambda) d\lambda}{\int \xi(\lambda) d\lambda}. \quad (2)$$

The reflectance R is then defined as:

$$R = \frac{\pi I}{E_0 \cos \theta_0}, \quad (3)$$

where E_0 is the extra-terrestrial solar irradiance perpendicular to the solar beam contained within the channel's SRF in $\text{Wm}^{-2}\mu\text{m}^{-1}$, and θ_0 is the solar zenith angle. The calibration results in this paper will be based on comparing measured reflectances, but we will also use the sun-normalized radiance R_n in a few occasions:

$$R_n = R \cos \theta_0. \quad (4)$$

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2.2 Instruments

SEVIRI is a 12-channel imager on the Meteosat Second Generation (MSG) geostationary satellites operated by the European Organization for the Exploitation of Meteorological Satellites (EUMETSAT). Three of a planned total of four MSG satellites, Meteosat-8, -9, and -10, have been launched to date. Apart from one high-resolution visible channel, SEVIRI carries 11 channels between 0.6 and 14 μm with a resolution of $3 \times 3 \text{ km}^2$ at nadir and a 15 min repeat cycle. This study focuses on the short-wave channels 1, 2, and 3, for which spectral information is provided in Table 1. SEVIRI is not equipped with an on-board calibration device for these channels. An operational vicarious calibration algorithm relying on desert and ocean targets (Govaerts and Clerici, 2004; Govaerts et al., 2004) was put in place by EUMETSAT for the Meteosat Visible and Infrared Imager (MIVIRI) originally, and has been carried over to SEVIRI. However, this operational calibration was shown to considerably deviate from MODIS for channel 1 (Doelling et al., 2004; Ham and Sohn, 2010).

MODIS is a cross-track-scanning imager on the polar-orbiting National Aeronautics and Space Administration's (NASA) Earth Observing System (EOS) platforms Terra and Aqua. It measures radiation from 0.4 to 14.5 μm in 36 spectral channels. Considerable efforts have been devoted to the calibration and characterization of the MODIS instrument, taking advantage of the on-board calibration facilities (Xiong and Barnes, 2006). From the two MODIS instruments the one on Aqua has proven to be the most stable, as the MODIS-Terra solar diffuser door has been kept permanently in an open position since July 2003 (Wu et al., 2013). Thus, MODIS-Aqua is considered here as the reference for the other sensors. The uncertainty of MODIS reflectance for the solar channels under consideration is estimated to be less than 2% (Xiong et al., 2005). In this study, MODIS radiances at 1-km spatial resolution, (the MOD021KM and MYD021KM collection 5 products) have been used. Of interest are channels 1, 2, and 6 (see Table 1), as well as channels 17 and 18 near 0.94 μm to characterize atmospheric water vapour (see Sect. 3).

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AVHRR is an imager flown on the Polar-orbiting Operational Environmental Satellite (POES) series of National Oceanic and Atmospheric Administration (NOAA) satellites since 1978, as well as on the EUMETSAT Metop series since 2006. With continued planned extension until at least 2018, the AVHRR record will span over 40 yr and is thus very valuable for climate studies. In this study, the AVHRR/3 instrument comprising 6 channels between 0.6 and 12 μm is considered. The inter-calibration is done for channels 1, 2, and 3a, the latter of which was only active on some of the NOAA satellites. AVHRR observations are taken at 1 km spatial resolution, but aggregated to $3 \times 5 \text{ km}^2$ for global area coverage (GAC) transmittance to the ground. AVHRR lacks an on-board calibration device, but extensive efforts have been dedicated to post-launch calibration. Heidinger et al. (2010) developed a calibration approach based on SNOs as well as Antarctic and desert targets, which was radiometrically tied to MODIS. That approach includes a conversion of the measured AVHRR/3 dual-gain count to what a single-gain count would have been.

The starting point of this paper are existing reference calibrations, here termed “nominal” calibrations, of the three instruments. For MODIS the nominal calibration is provided in the MOD02/MYD02 files. For AVHRR we use the Heidinger et al. (2010) calibration. For SEVIRI the operational calibration by EUMETSAT is used. Since this shows slight step-wise changes over time, we have chosen the calibration coefficients of January 2007 as the reference for both MSG satellites. These are $S = 0.5736$, 0.4531, and $0.08783 \text{ W m}^{-2} \text{ sr}^{-1} \mu\text{m}^{-1}$ for Meteosat-8 channels 1, 2, and 3, respectively, and $S = 0.4993$, 0.3951, and $0.08276 \text{ W m}^{-2} \text{ sr}^{-1} \mu\text{m}^{-1}$ for Meteosat-9 channels 1, 2, and 3, respectively. In addition, values of E_0 as provided by EUMETSAT are used to convert SEVIRI radiances to reflectances (see http://www.eumetsat.int/groups/ops/documents/document/pdf_msg_seviri_rad2refl.pdf).

The calibration of SEVIRI (SEV) with respect to a polar-orbiting imager (P) will be expressed in terms of a re-calibration slope s :

$$s_{\text{SEV/P}} = \left\langle \frac{R_{\text{SEV}}}{R_{\text{P}}} \right\rangle, \quad (5)$$

where the mean reflectance ratio is in practice determined by a linear fit. In practice, MODIS-Aqua will be taken as the reference instrument P . The re-calibration slope can then be used to calculate a corrected calibration slope S^* for SEVIRI:

$$S_{SEV}^* = S_{SEV} / s_{SEV/P}. \quad (6)$$

Instrument dark counts D are assumed to be correct as provided. This assumption will be verified by performing both free linear calibration fits as well as fits forced through the origin.

In this study, six years (2004 to 2009) of SEVIRI-Meteosat and MODIS-Aqua data are considered. In addition, three years (2007 to 2009) of data from other polar imagers are evaluated. These time periods are long enough to make statements about the stability of the calibration of the respective satellite instruments, and in some cases sufficient to calculate significant trends in calibration coefficients.

3 Method

Our calibration method consists of the following steps:

1. Collect SEVIRI and polar imager near-nadir reflectances near the SEVIRI sub-satellite point and with the smallest possible time difference. Given the SEVIRI repeat cycle of 15 min, the time difference is about 7.5 min at most.
2. Convert SEVIRI reflectances to the spectral response of the polar imager by correcting for within-band atmospheric transmissivity.
3. Aggregate reflectances to a common regular latitude–longitude grid. A $0.15^\circ \times 0.15^\circ$ grid was used for the MODIS–SEVIRI comparisons, while for the AVHRR–SEVIRI comparisons a somewhat larger ($0.25^\circ \times 0.25^\circ$) grid was chosen because of the larger AVHRR pixel size.

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4. Generate a subset of the reflectance pairs based on viewing/illumination geometry and possibly other selection criteria.
5. Perform a linear regression of SEVIRI against polar imager reflectances. Since there is no true independent variable we apply orthogonal fits, i.e. least squares fits based on the orthogonal distance of reflectance pairs to the fit line. The regression yields a re-calibration slope.

We will first outline the SRF correction approach applied in step 2. Then the method is illustrated with a concrete example.

3.1 Correcting for differences in spectral response

The spectral response functions of the relevant SEVIRI, MODIS, and AVHRR channels around 0.6, 0.8, and 1.6 μm are shown in Fig. 1, together with the atmospheric transmissivity spectrum. The MODIS channels are generally narrower and less affected by atmospheric gas absorption than the SEVIRI and AVHRR channels. The AVHRR channel 2 is much broader than the others. These differences need to be taken into account before measured reflectances can be compared, which is done by the following procedure.

We used the Moderate Resolution Atmospheric Transmission code (MODTRAN4 Version 2; Anderson et al., 2001) to simulate top-of-atmosphere reflectance spectra over the satellite instrument wavelength bands, including Rayleigh scattering and absorption by atmospheric gases, for a Lambertian surface with fixed surface albedo. Clouds were not explicitly taken into account but were modelled by placing the surface at the height of the cloud top, H_c . Apart from H_c the reflectance depends on the solar zenith angle (θ_0), the satellite zenith angle (θ), and the above-cloud partial column amounts of atmospheric trace gases. While some of the relevant absorbing gases, O_2 , CO_2 , and CH_4 , are long-lived and therefore well-mixed in the atmosphere, water vapour and – to a lesser extent – ozone are characterized by significant total column variations. The illumination and viewing geometry were combined in the geometrical

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air-mass factor ($AMF = 1/\cos\theta_0 + 1/\cos\theta$), which is proportional to the length of the light path. Reflectances R^{sim} were then simulated with MODTRAN for the various instrument channel SRFs following Eqs. (2)–(3), and stored in a look-up table as a function of four variables: H_c , AMF, total column ozone (TCO) and total column water vapour (TCWV).

H_c was retrieved from 11 μm SEVIRI radiances using the approach outlined by Roebeling and Holleman (2009). This H_c can be regarded as the effective infrared radiating height rather than the physical cloud-top height, and is thus consistent with the placement of a Lambertian surface at H_c in the MODTRAN simulations. TCO was taken from a monthly mean climatology at $1^\circ \times 1.5^\circ$ (lat–lon) derived from the Multi Sensor Reanalysis (MSR) dataset by Van der A et al. (2010). It was assumed that all ozone resides above the clouds. TCWV was taken from a monthly-mean climatology at $1^\circ \times 1^\circ$ derived from the ERA-Interim reanalysis (Dee et al., 2011). A fixed vertical profile (McClatchey et al., 1971) was used to determine the amount of water vapour above the cloud. These assumptions constitute potential sources of error, especially in channel 2 which is characterized by strong water vapour absorption. This will be further discussed below. In our spectral response correction procedure, for each SEVIRI pixel with measured reflectance R_{SEVIRI} , the reflectance R'_{SEVIRI} that would have been observed given the SRF of a polar imager (MODIS or AVHRR) is estimated as:

$$R'_{\text{SEVIRI}} = R_{\text{SEVIRI}} \frac{R_{\text{polar}}^{\text{sim}}}{R_{\text{SEVIRI}}^{\text{sim}}}, \quad (7)$$

where $R_{\text{SEVIRI}}^{\text{sim}}$ and $R_{\text{polar}}^{\text{sim}}$ are the MODTRAN-simulated reflectances for the SEVIRI and polar-orbiter channels, respectively. More details on the atmospheric correction method can be found in Meirink et al. (2009).

The SRF correction procedure has various sources of uncertainty, e.g. from potential errors in the H_c and trace-gas amount input used, and in the accounting for below-cloud absorption in broken and semi-transparent cloud scenes. Uncertainties are largest

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2004), is reached by requiring the differences in θ_0 , θ , and scattering angle Θ between collocated pixels to be smaller than a specific limit, here chosen as 10° . Note that for θ_0 this requirement is automatically fulfilled by the spatio-temporal collocation between the instruments. After application of these criteria only a fraction of the data is retained, but these typically yield a significantly higher correlation as well as a much improved consistency between the two fit types (see lower right panel of Fig. 3).

Since the number of pixels selected for the fits varies a lot from granule to granule, the granule-based fit results are relatively noisy. Therefore, fits are performed for the collection of all data in a month. We collected all MODIS-Aqua granules near (0° W, 0° N), the Meteosat-9 sub-satellite point, usually one or two per day. Scatter plots of ray-matched reflectance pairs for channels 1 to 3 from all granules in a selected month (September 2007) are shown in Fig. 4. The correlations are generally very high and the two types of linear fits are consistent. The highest correlations are obtained for channel 1 ($0.6\ \mu\text{m}$), probably because there is very little absorption by trace gases in this channel, so that the SRF correction is almost negligible. In addition, the dynamic range is high, enabling more robust fits. For channel 2 ($0.8\ \mu\text{m}$) a complicating feature is observed (Fig. 4b): the MODIS channel saturates around a value of the sun-normalized radiance of about 0.7–0.75, because it has been optimized for land-surface applications. As a consequence, if all data points are taken into account for the fit, the slope is overestimated. To take care of this issue, only R_n pairs with an average value less than 0.6 are considered for the regressions further on (see Fig. 4c), giving somewhat lower regression slopes. Figure 4d shows what happens if the reflectances are not corrected for SRF differences. Since the SEVIRI channel 2 is much more affected by trace gas absorption (in particular of water vapour) than the MODIS channel 2, the SRF correction leads to an increase in the slope of about 7%. SRF correction typically also yields a slightly improved correlation between the reflectance pairs. Fit results for channel 3 ($1.6\ \mu\text{m}$) with and without SRF correction are presented in Fig. 4e and f, respectively. For this channel the effect of SRF correction is about 2%. A significant number of $1.6\ \mu\text{m}$ detectors on MODIS-Aqua are defect, causing stripes in the 1 km

level-1b images. Despite of this, we have applied the usual aggregation procedure, still finding sufficient valid pixels in the $0.15^\circ \times 0.15^\circ$ grid cells. The inter-calibration does not appear to suffer from the missing MODIS pixels.

In the remainder a “standard” pixel selection configuration will be used to generate calibration regression statistics. This configuration includes all pixels with $\Delta\theta_0 < 10^\circ$, $\Delta\theta < 10^\circ$, and $\Delta\Theta < 10^\circ$, as well as an R_n limit of 0.6 for channel 2 as explained above. Sensitivities to other selection configurations will be assessed in Sect. 5.

4 Inter-calibration of SEVIRI and MODIS

The method has been applied to three years (April 2004 to March 2007) of collocated MODIS-Aqua and SEVIRI-Meteosat-8 reflectances and three years (January 2007 to December 2009) of collocated MODIS-Aqua and SEVIRI-Meteosat-9 reflectances. Time series of monthly re-calibration slopes for the three solar channels are displayed in Fig. 5. These are the slopes of the linear fits through the origin, both for observed and SRF-corrected SEVIRI reflectances, and based on the standard geometric selection criteria.

For channel 1, re-calibration slopes around 0.92 are found for both Meteosats. Since MODIS-Aqua is considered to be the reference instrument, this means that the SEVIRI nominal reflectance is about 8% too low. This result is consistent with earlier findings by Doelling et al. (2004) and Ham and Sohn (2010). The calibration method gives very stable month-to-month slopes, with a standard deviation relative to the trend line of about 0.004. The correlation between SEVIRI and MODIS is high for all months and shows very little temporal variation (not shown). The impact of the spectral response correction, quantified by the difference between red and black symbols/curves, is negligible for this channel, because there is little absorption by trace gases. In contrast, for channel 2 the difference in SRFs between the two instruments has a much larger effect: SEVIRI is more affected by water vapour absorption than MODIS. The impact of this absorption is different for every individual scene that is observed. On average,

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over all the scenes analysed here, the respective SRFs lead to about 7 % difference in reflectance. Thus, the re-calibration slope is enhanced from about 0.87 to 0.94 after SRF correction. The final result is that SEVIRI reflectances in channel 2 are about 6 % too low as compared to MODIS-Aqua. It is encouraging to see that the standard deviation of monthly slopes is reduced after SRF correction, although it is still larger than for channel 1. This is another indication that the correction method works well. A seasonal cycle in derived re-calibration slopes is present for channel 2. This may point to imperfections in the accounting for water vapour variability. However, since a slight seasonal cycle is also observed for channel 1, other factors, probably related to sampling varying parts of the scene Bidirectional Reflection Distribution Function (BRDF), must play a role as well. For channel 3 the magnitude of the SRF correction is both smaller and less variable over time as compared to channel 2. The lower variability is explained by the fact that the main absorbing gases in channel 3 (CO₂ and CH₄) have nearly constant concentrations. For this channel the SEVIRI reflectance is found to be 3.5 % too high.

The re-calibration slopes for Meteosat-8 and Meteosat-9 are nearly identical for all channels, demonstrating that the operational vicarious calibration performed at EU-METSAT is consistent between both satellites. In Fig. 5 the trends of the re-calibration slopes are also shown. The annual trend is in all cases much smaller than 1 % yr⁻¹. We have calculated the uncertainty of the trends following Weatherhead et al. (1998). Only the trends for channel 1 (both Meteosats) are significant at the 95 % level, indicating a slow degradation of this SEVIRI channel compared to MODIS-Aqua. Overall, the results demonstrate that the SEVIRI instrument behaves very stable over time.

5 Sensitivity analysis of pixel selection

In this section, we elaborate on the robustness of the regression statistics to various underlying choices regarding selection and sampling of pixels. Fit statistics for monthly SEVIRI–MODIS regressions for the years 2007–2009 are shown in Table 2. As outlined

before, the “standard” configuration contains all pixels with $\Delta\theta_0 < 10^\circ$, $\Delta\theta < 10^\circ$, and $\Delta\Theta < 10^\circ$, as well as an R_n upper limit of 0.6 for channel 2. Results are shown for channel 1 but are similar for the other channels, except for a few cases which are included in Table 2.

5 First, the sensitivity to the matching of viewing angles is evaluated. If no limit to the viewing angle difference is set, poor regressions are obtained. The correlation coefficients are lower than for the standard case (on average 0.967 vs. 0.991), and there is a considerable difference between linear fits forced through the origin and free linear fits. The resulting mean re-calibration slope is 3% smaller than with the standard selection criteria. Subsequent analyses with looser and stricter matching criteria for the viewing angle difference ($\Delta\theta < 20^\circ$ and $\Delta\theta < 5^\circ$, respectively) indicate that the results are not critically dependent on the exact limit. The matching of scattering angles is found to be less important. If the $\Delta\Theta$ criterion is not applied, similar slopes are obtained but their monthly variability is higher and the mean correlation coefficients is lower than in the standard configuration.

10 In addition to ray-matching criteria, it might be important to avoid geometries for which reflectance is a fast-varying function of the scattering angle for particular scenes. We tested the following selection criteria to avoid

- the cloud bow: $\Theta < 135^\circ$ or $\Theta > 145^\circ$.
- the backscatter peak: $\Theta < 170^\circ$.
- sunglint: Θ more than 25° away from the glint angle.

15 It turns out that application of none of these criteria makes a difference to the mean re-calibration slope, while a positive impact on the correlation coefficients and on the standard deviation of monthly slopes is observed. However, the drawback is that the number of grid points for the regressions shows a larger month-to-month variation (not shown). For satellites with different overpass times these criteria lead to an even more variable distribution of data points through the year, and deteriorated the fits. Therefore, these criteria have not been included in the standard configuration.

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The impact of the horizontal resolution at which the reflectances of both instruments are aggregated is found to have no influence on the re-calibration slopes. This is a good result, since the choice of resolution is somewhat arbitrary. Different resolutions do yield slightly different correlation coefficients, with – as expected – consistently lower correlations at higher resolutions.

Next, the impact of the amount of data is evaluated by keeping data from only one out two, four, or eight days in the regressions. It turns out that even with approximately weekly data the same mean re-calibration coefficient is obtained, although the standard deviation of monthly slopes becomes larger.

We also investigated what happens if radiances below or above a certain limit are neglected. It turns out that neither setting a lower limit to R_n of 0.1 nor an upper limit of 0.5 has significant impact on the channel-1 calibration slopes. Nevertheless, a drop in the correlation coefficients is observed, which might be attributed to the decreased dynamic range of radiances. Table 2 also shows the regression statistics for applying the upper R_n limit to channels 2 and 3. For channel 2 the results are similar to the standard settings, which demonstrates that the exact value of the upper R_n limit, needed to account for the saturation of MODIS channel 2, is not very critical. Channel 3 sun-normalized radiances rarely exceed 0.5, and therefore the number of data points and thus the fit statistics are hardly affected by applying an upper limit to R_n .

In our calibration method, the albedo of observed scenes is implicitly assumed to be spectrally uniform over the various SRFs for a particular channel. While this assumption is well satisfied over ocean surfaces and clouds, land surfaces are often characterized by a spectrally varying albedo. To test the impact of this variability, we repeat the regressions for an ocean-only pixel selection. The fit statistics for channel 1 are virtually the same for the ocean-only selection compared to the standard configuration. For channel 2 this is different: the re-calibration slope increases by 1.2 %. At the same time the correlation coefficient decreases somewhat and becomes more variable. For channel 3 the selection of ocean pixels has no impact on the re-calibration slope, but yields a considerably higher standard deviation of the slopes as well as a lower and more variable

correlation coefficient. This suggests that the ocean-only fits are somehow less robust than the standard fits including land and ocean pixels.

In summary, the sensitivity study shows that results of the inter-calibration regressions are very robust and not critically dependent on the chosen selection and sampling strategy, although it is clear that certain limits to the viewing zenith angle and scattering angle differences need to be applied. The largest sensitivity is found for the restriction to scenes over ocean for channels 2. Based on these results we can try to come up with uncertainty estimates of the derived re-calibration slopes based on the sensitivity to pixel selection and sampling criteria as well as the standard deviation of monthly slopes. This yields conservative estimates of 1 % for channel 1, and 1.5 % for channels 2 and 3. It needs to be emphasized that there may be systematic errors in the calibration procedure, e.g. related to SRF correction, which are not included in these estimates.

6 Polar imager inter-calibration

Our method for the inter-calibration of MODIS-Aqua and SEVIRI can be applied to any other polar-orbiting imager than MODIS-Aqua with shortwave channels corresponding to those of SEVIRI. This is not of direct interest, but allows indirectly to inter-calibrate these other polar-orbiting imagers with MODIS-Aqua, as will be demonstrated below.

First we show, as an example, inter-calibration results between AVHRR onboard NOAA-17 and SEVIRI for the period 2007 to 2009 in Fig. 6. For AVHRR the nominal calibration coefficients are taken from Heidinger et al. (2010). SEVIRI channel-1 reflectances are about 6 % lower than AVHRR, and the SRF correction has very little effect. In contrast, for channel 2 the impact of SRF correction is very large (almost 7 %), because the AVHRR channel is much broader than its SEVIRI counterpart, and is much more affected by atmospheric water vapour absorption. The final result is that SEVIRI channel-2 reflectances are 2.5 % lower than AVHRR. It is encouraging to see that the standard deviation of re-calibration slopes is more than halved by the SRF

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correction. The 1.6 μm channels, although they are quite different on AVHRR and SEVIRI, are on average similarly impacted by trace gas absorption. Hence, SRF correction has little effect. Averaged over the years 2007 to 2009, SEVIRI and AVHRR channel-3 reflectances are almost at the same level, but there is a large and significant trend of about 1.5 % yr^{-1} in the re-calibration slopes.

When two polar imagers (P1 and P2) have been compared to SEVIRI, it is straightforward to inter-calibrate these polar imagers by taking the ratio of the respective re-calibration slopes:

$$S_{P1/P2} = \frac{S_{SEV/P2}}{S_{SEV/P1}}. \quad (8)$$

A similar strategy was applied by Wang et al. (2011), who inter-calibrated infrared radiances from the Atmospheric Infrared Sounder (AIRS) and the Infrared Atmospheric Sounding Interferometer (IASI) using the Geostationary Operational Environmental Satellite (GOES) as a transfer instrument. With their double differencing approach they were able to verify an inter-calibration based on direct AIRS–IASI SNOs.

Using SEVIRI as a transfer instrument, we inter-calibrated MODIS-Terra and the AVHRRs on NOAA-17 and NOAA-18 with respect to MODIS-Aqua for the period 2007 to 2009. Time series of resulting re-calibration slopes are presented in Fig. 7. The two MODIS instruments are found to be very consistent. For channel 1 a difference of about 2 % is obtained, Terra having lower reflectances than Aqua. A similar difference was reported by Minnis et al. (2008) and Wu et al. (2013). The 0.8 and 1.6 μm MODIS channels on Aqua and Terra differ less than 1 %. A significant negative trend is calculated for Terra channel 2. There is a clear hint, though, for a break in the time series around May–June 2009, when Terra/Aqua re-calibration slopes appear to decrease by around 1 % for all channels. A drop in MODIS-Terra reflectances in 2009 was also identified by Wu et al. (2013) as being due to a one-time calibration correction.

The AVHRR channel-1 reflectances are on average about 2 and 1 % lower than MODIS-Aqua for NOAA-17 and NOAA-18, respectively. In addition, negative trends of around 0.9 % yr^{-1} are observed, which is slightly larger than the 95-% significance

level. AVHRR Channel-2 reflectances are found to be 3.5% too low compared to MODIS-Aqua. The re-calibration slopes are also more variable than for other channels. For channel 3, which is only active on NOAA-17, AVHRR is found to overestimate reflectances by 2.5% on average, with a significant negative trend of $1.6\% \text{ yr}^{-1}$.

The AVHRR nominal calibration relies on direct SNOs with MODIS, and its accuracy was estimated to be 2 and 3% for channels 1 and 2, respectively (Heidinger et al., 2010). Thus, although the deviations obtained here are on average close to that accuracy, this is not the case during the full 2007–2009 time period. The average offsets and temporal trends in Fig. 7 point to inconsistencies between both calibration methods. These may result from different approaches to SRF correction in combination with different atmospheric composition and scene types at the sampling locations, i.e. high latitudes for the direct SNOs versus low latitudes in the present study. However, the SRF correction is negligible for channel 1 and small for channel 3, so the differences cannot be solely explained by the SRF correction approach. In addition, such differences should not lead to temporal trends. The negative trend of, in particular, the AVHRR 1.6 μm channel is puzzling, and requires further research to be elucidated.

7 Conclusions

In this study we have calibrated the solar channels of the geostationary SEVIRI instrument on MSG with MODIS on Aqua. The inter-calibration was based on regressions of collocated reflectances. Differences in spectral response between the instrument channels were corrected for by tabulated radiative transfer calculations of the effect of trace gas absorption on top-of-atmosphere reflectance. SEVIRI operational reflectances were found to be quite stable during the years 2004–2009, but considerably biased with respect to MODIS: -8 , -6 , and $+3.5\%$ for channels 1 ($0.6 \mu\text{m}$), 2 ($0.8 \mu\text{m}$), and 3 ($1.6 \mu\text{m}$), respectively. A sensitivity study revealed that the derived re-calibration coefficients do not critically depend on various choices regarding pixel selection, as long as the viewing and illumination geometry of the two instruments is

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matched. Based on the sensitivity study we estimated uncertainties in the re-calibration slopes to be 1% for channel 1 and 1.5% for channels 2 and 3. Of course, the total uncertainty includes also potential systematic errors in the calibration procedure (e.g. related to SRF correction) and the uncertainty of the Aqua-MODIS calibration. Interestingly, the channel-3 inter-calibration did not seem to be affected by the significant number of defect detectors in the corresponding MODIS-Aqua 1.6 μm channel.

Using double differencing, the calibration method was applied for the inter-calibration of other polar imagers with Aqua-MODIS for the years 2007–2009. Here, SEVIRI served as a calibration transfer instrument. This analysis demonstrated that Terra-MODIS channel 1 has a bias of about -2% compared to Aqua, while other channels are unbiased. The AVHRR instruments on NOAA-17 and NOAA-18 were also investigated using the direct AVHRR–MODIS SNO based calibration by Heidinger et al. (2010) as a nominal reference. We found reasonable agreement for channel 1, with a mean AVHRR–MODIS reflectance difference below 2% and a just-significant negative trend. For channels 2 and 3 the differences were larger (-3.5 and $+2.5\%$, respectively) and AVHRR channel-3 reflectances had a negative trend of more than $1.5\% \text{yr}^{-1}$. Overall, this indicates that the methods used by Heidinger et al. (2010) and in the present study yield reasonably consistent results but there are differences that require further investigation.

The approach for inter-calibration of polar-orbiting imagers using the geostationary SEVIRI as a transfer instrument has the advantage of daily availability of collocated reflectances. Since both clear and cloudy pixels are used, a large dynamic range of reflectances is obtained and robust statistics can be developed. The double difference method allows verifying calibrations based on polar–polar SNOs, which typically occur at high latitudes only. An obvious limitation of the method is that a geostationary sensor with the appropriate solar channels is required, which reduces the applicability backward in time. With the SEVIRI record already spanning almost ten years and a similar time period to come, and with the upcoming Meteosat Third Generation as well as other new geostationary imagers, the applicability of the method is guaranteed



in the far future, for even more channels than in this study. The work presented here will be extended to longer time series in the future, making use of the latest (Collection 6) MODIS calibration.

Acknowledgements. This work was funded by EUMETSAT's Satellite Application Facility on Climate Monitoring (CM-SAF) and ESA's Climate Change Initiative on Clouds (Cloud.CCI).

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Table 1. Channel numbers and central wavelengths λ_c of the solar channels considered in this study together with the approximate wavelengths used for reference.

Satellite	ch1, 0.6 μm		ch2, 0.8 μm		ch3, 1.6 μm	
	channel	λ_c (μm)	channel	λ_c (μm)	channel	λ_c (μm)
SEVIRI	1	0.635	2	0.810	3	1.640
MODIS	1	0.645	2	0.858	6	1.640
AVHRR	1	0.630	2	0.865	3a	1.610

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Table 2. Sensitivity of regression statistics^a to pixel selection criteria and sampling strategies. Results are for SEVIRI-Meteosat-9 to MODIS-Aqua fits for the years 2007–2009.

		\bar{N}	\bar{s}	$\overline{s_{2p}}$	$\sigma(s)$	\bar{r}	$100 \times \sigma(r)$
channel 1							
	standard settings	54 637	0.920	0.916	0.0040	0.991	0.23
geom. criteria	no $\Delta\theta$ criterion	161 863	0.888	0.935	0.0124	0.967	0.75
	$\Delta\theta < 20^\circ$	30 682	0.919	0.915	0.0038	0.991	0.24
	$\Delta\theta < 5^\circ$	74 292	0.918	0.915	0.0040	0.991	0.18
	no $\Delta\Theta$ criterion	92 003	0.920	0.917	0.0077	0.986	1.00
	add sunglint criterion	34 744	0.921	0.912	0.0028	0.993	0.12
	add rainbow criterion	36 147	0.920	0.913	0.0024	0.993	0.16
	add glory criterion	53 329	0.920	0.916	0.0038	0.991	0.27
spatial resol.	low resolution (0.3°)	13 785	0.920	0.916	0.0040	0.994	0.23
	high resolution (0.1°)	122 603	0.919	0.913	0.0040	0.987	0.25
temp. sampling	1 of 2 days selection	27 718	0.921	0.916	0.0043	0.992	0.23
	1 of 4 days selection	14 140	0.921	0.916	0.0053	0.991	0.36
	1 of 8 days selection	7197	0.921	0.916	0.0070	0.991	0.51
R_n limits	$R_n > 0.1$	34 614	0.920	0.914	0.0038	0.988	0.24
	$R_n < 0.5$	51 098	0.923	0.919	0.0057	0.983	0.56
spatial domain	ocean only ^b	29 371	0.920	0.917	0.0066	0.991	0.43
channel 2							
	standard settings	52 392	0.940	0.932	0.0059	0.986	0.41
R_n limits	$R_n < 0.5$	50 112	0.937	0.923	0.0062	0.984	0.45
spatial domain	ocean only ^b	28 589	0.952	0.942	0.0079	0.983	0.97
channel 3							
	standard settings	54 640	1.032	1.028	0.0069	0.988	0.53
R_n limits	$R_n < 0.5$	54 304	1.032	1.028	0.0072	0.987	0.60
spatial domain	ocean only ^b	29 377	1.033	1.032	0.0160	0.979	1.99

^a \bar{N} is the average number of pixels per month; \bar{s} and $\overline{s_{2p}}$ are the average of monthly re-calibration slopes from linear fits forced through the origin and free linear fits, respectively; $\sigma(s)$ is the standard deviation of monthly re-calibration slopes with respect to the trend line; \bar{r} and $\sigma(r)$ are the average and standard deviation of monthly Pearson correlation coefficients, respectively.

^b The ocean-only domain includes all pixels south of 4° N and west of 8° E.

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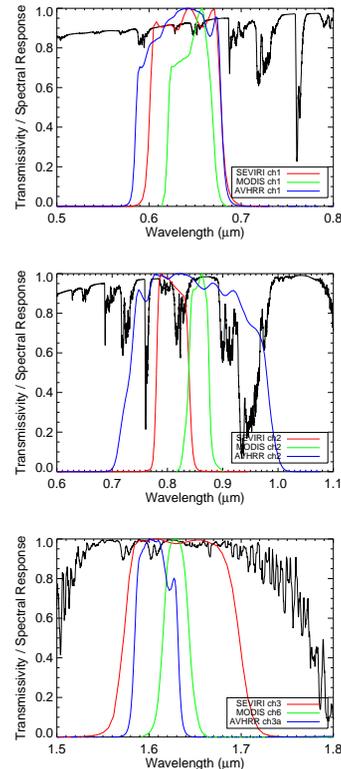


Fig. 1. Spectral response functions of SEVIRI, MODIS and AVHRR for the three solar channels considered in this paper, as obtained from <http://www.eumetsat.int/Home/Main/Satellites/MeteosatSecondGeneration/Instruments/index.htm>, <ftp://mcst.hbsss-sigma.com/pub/permanent/MCST>, and <http://www2.ncdc.noaa.gov/docs/klm/html/d/app-d.htm>, respectively. A nadir transmissivity spectrum calculated with MODTRAN and based on a tropical atmosphere (McClatchey et al., 1971) is shown in black.

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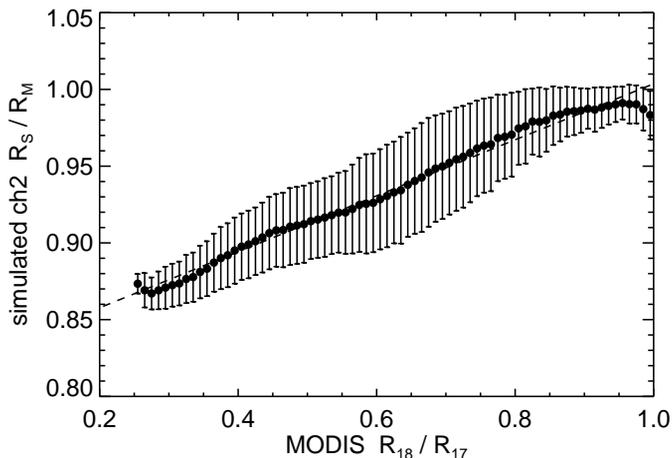


Fig. 2. Relation between simulated SEVIRI to MODIS reflectance ratio in channel 2 ($0.8\ \mu\text{m}$) and MODIS ch18/ch17 reflectance ratio, demonstrating the water vapour correction capability in our SRF correction procedure. Circles show the mean and vertical bars the standard deviation of simulated ch2 reflectance ratio in a MODIS ch18/ch17 reflectance ratio bin. The figure is based on about 200 000 data points from the year 2008. The correlation coefficient of the underlying individual data points is 0.77.

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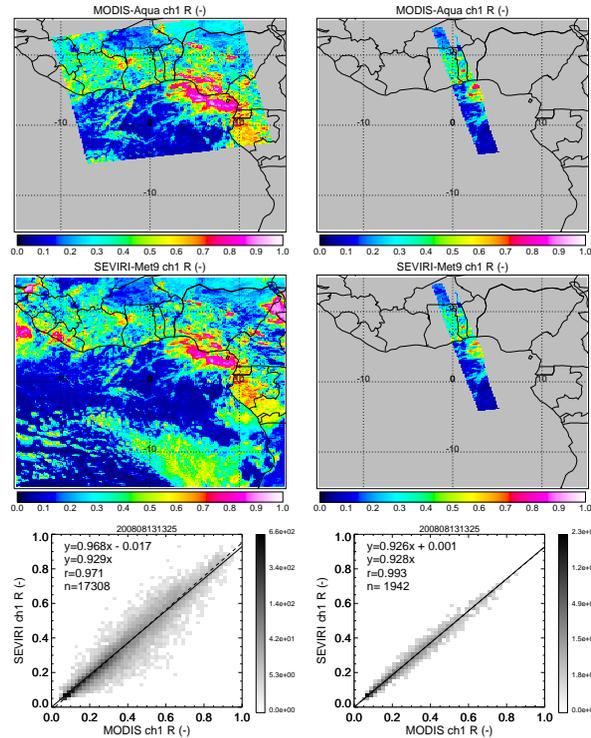


Fig. 3. Illustration of the calibration approach for a MODIS-Aqua granule on 13 August 2008 between 13:25 and 13:30 UTC and a SEVIRI image around 13:21 UTC. The top row shows MODIS ch1 reflectance, and the middle row shows SEVIRI ch1 reflectance corrected to the MODIS SRF using Eq. (7). The bottom row shows scatter density plots between SEVIRI and MODIS R in which the free linear fit (dashed) and the linear fit through the origin (solid) are indicated. A cubic color scale, as indicated by the vertical bars, has been used to make the low densities visible. The left panels are for all available data, while the right panels are restricted to those grid cells satisfying the standard geometric selection criteria.

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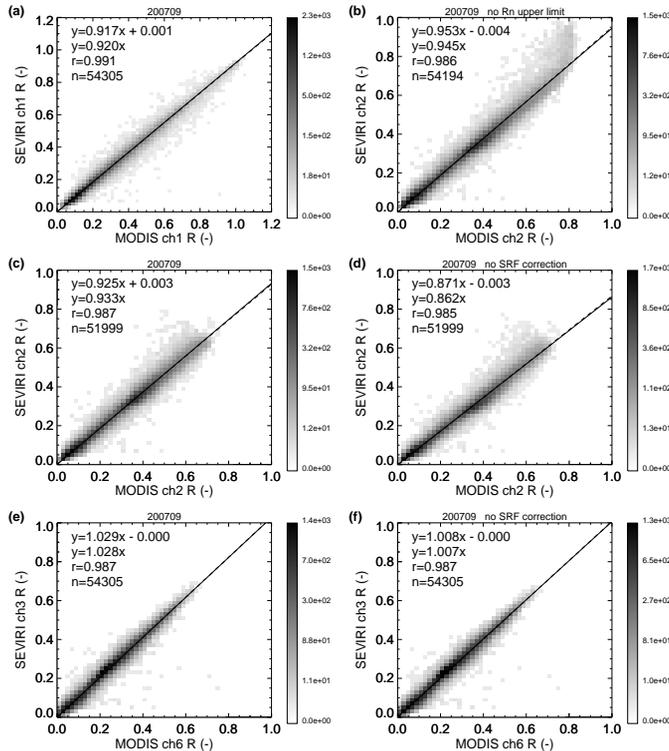


Fig. 4. Monthly (September 2007) fits of SEVIRI-Meteosat-9 against MODIS-Aqua reflectances: **(a)** channel 1, **(b)** channel 2 without upper limit to R_n , **(c)** channel 2 with upper limit to R_n (this is the standard setting for channel 2; see text), **(d)** channel 2 without SRF correction, **(e)** channel 3, **(f)** channel 3 without SRF correction. For all fits the standard geometric selection criteria have been applied. A cubic color scale, as indicated by the vertical bars, has been used to make the low densities visible. The free linear fits (dashed) and linear fits through the origin (solid) are indicated in the plots.

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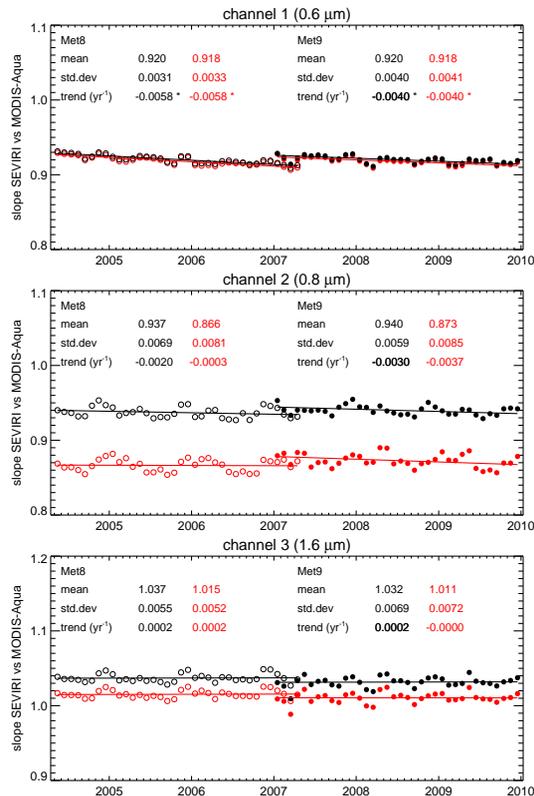


Fig. 5. Time series of re-calibration slopes for SEVIRI-Meteosat against MODIS-Aqua for three solar channels. The open circles are the monthly slopes for Meteosat-8, while the filled circles are for Meteosat-9. The solid lines are linear fits through those monthly slopes. Black symbols and lines correspond to data that have been corrected for SRF differences, while red symbols and lines correspond to data that have not been corrected for SRF differences. Mean, standard deviation, and trend of the re-calibration slopes are indicated in the plots. Trends that are significant at the 95 % level are marked by an asterisk.

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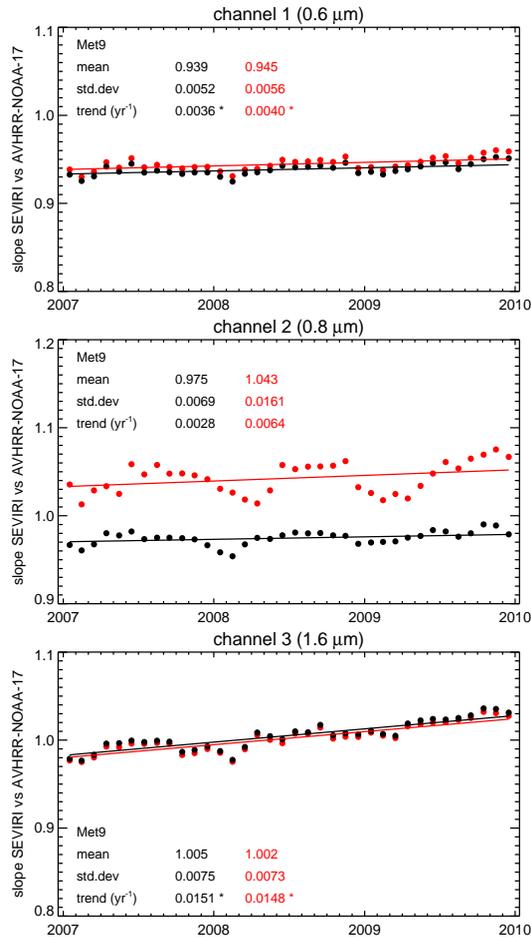


Fig. 6. Same as Fig. 5 but for AVHRR-NOAA17 instead of MODIS-Aqua, and for Meteosat-9 only.

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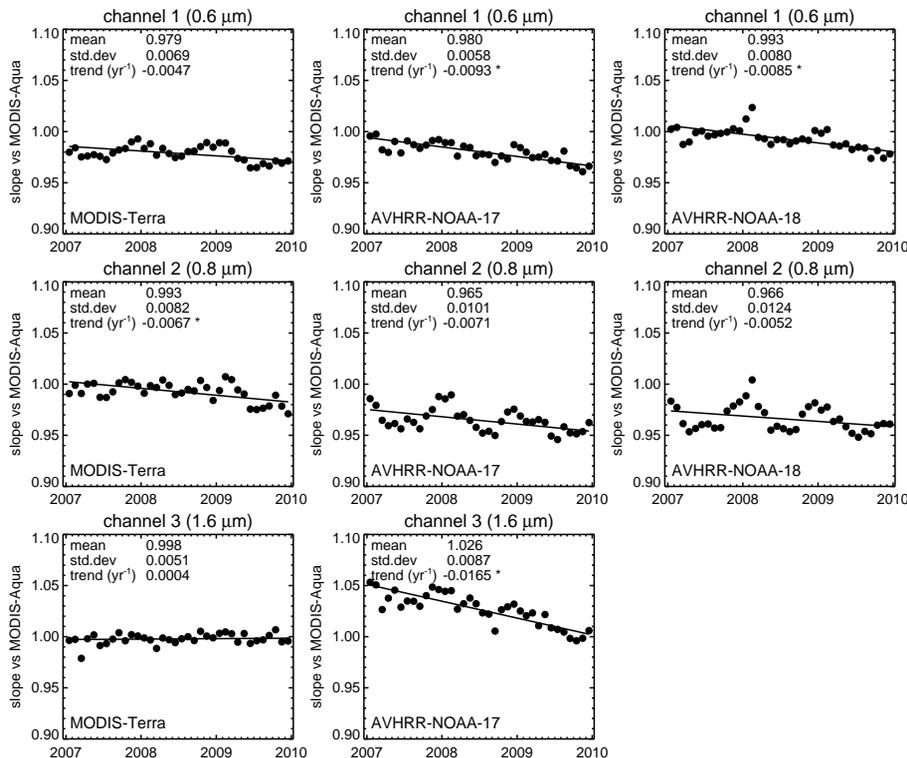


Fig. 7. Time series of re-calibration slopes of polar-orbiting imagers against MODIS-Aqua using SEVIRI-Meteosat-9 as a transfer instrument. Results are shown for MODIS-Terra (left column), AVHRR-NOAA-17 (middle column) and AVHRR-NOAA-18 (right column) and channels around 0.6 μm (top row), 0.8 μm (middle row) and 1.6 μm (bottom row). Note that the 1.6 μm channel was not active on AVHRR-NOAA-18. Mean, standard deviation, and trend of the re-calibration slopes are indicated in the plots. Trends that are significant at the 95 % level are marked by asterisks.