



Water vapor  
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# Water vapor observations up to the lower stratosphere through the Raman lidar during the MAïdo Lidar Calibration Campaign

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## Abstract

A new lidar system devoted to tropospheric and lower stratospheric water vapor measurements has been installed at the Maïdo altitude station facility of La Reunion Island, in the southern subtropics.

The main objectives of the MAïdo Lidar Calibration Campaign (MALICCA), performed in April 2013, were to validate the system, to set up a calibration methodology, to compare the acquired water profiles with radiosonde measurements and to evaluate its performances and capabilities with a particular focus on the UTLS measurements.

Varying the characteristics of the transmitter and the receiver components, different system configuration scenarios were tested and possible parasite signals (fluorescent contamination, rejection) were investigated. A hybrid calibration methodology has been set up and validated to insure optimal lidar calibration stability with time. In particular, the receiver transmittance is monitored through the calibration lamp method that, at the moment, can detect transmittance variations greater than 10–15%. Calibration coefficients are then calculated through the hourly values of IWV provided by the co-located GPS. The comparison between the constants derived by GPS and Vaisala RS92 radiosondes launched at Maïdo during MALICCA, points out an acceptable agreement in terms of accuracy of the mean calibration value (with a difference of approximately 2–3%), but a significant difference in terms of variability (14 vs. 7–9%, for GPS and RS92 calibration procedures, respectively).

We obtained a relatively good agreement between the lidar measurements and 15 co-located and simultaneous RS92 radiosondes. A relative difference below 10% is measured in low and middle troposphere (2–10 km). The upper troposphere (up to 15 km) is characterized by a larger spread (approximately 20%), because of the increasing distance between the two sensors.

To measure water vapor in the UTLS region, nighttime and monthly water vapor profiles are presented and compared. The good agreement between the lidar monthly profile and the mean WVMR profile measured by satellite MLS has been used as

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a quality control procedure of the lidar product, attesting the absence of significant wet biases and validating the calibration procedure.

Thanks to its performance and location, the MAIDO H<sub>2</sub>O lidar is devoted to become a reference instrument in the southern subtropics, allowing to insure the long-term survey of the vertical distribution of water vapor, and to document scientific themes such as stratosphere–troposphere exchange, tropospheric dynamics in the subtropics, links between cirrus clouds and water vapor.

## 1 Introduction

Water vapor is a crucial climate variable involved in many processes, widely determining the energy budget of our planet. It is the dominant greenhouse gas in Earth's atmosphere and its condensed forms (liquid and ice) exert a profound influence on both incoming solar and outgoing infrared radiation. The water vapor distribution in the upper troposphere (UT) and lower stratosphere (LS) is of central importance in several ways: it plays a major role in the balance of planetary radiation; it influences and responds to atmospheric motions; and it plays a key role in many aspects of UT/LS chemistry. In fact, it strongly contributes to the stratospheric radiative balance via its greenhouse effect (e.g. Kiehl and Trenberth, 1997), and is the main precursor of HO<sub>x</sub> radicals contributing to the catalytic destruction of ozone in the lower stratosphere (e.g. Wennberg et al., 1994; Osterman et al., 1997). Furthermore, the presence of cirrus clouds in the upper troposphere, highly dependent on the concentration of water vapor and the local temperature, also strongly impacts the radiative balance (Jensen et al., 1994).

Although methane oxidation is a major source of water in the stratosphere, the question of the mechanism controlling the amount of water vapor in the stratosphere still remains (Sherwood and Dessler, 2000; Kley et al., 2000; Oltmans et al., 2000). This can be partly explained through the lack of reliable water vapor observations in the tropical UTLS, limited to a few balloon, high altitude aircraft measurements, and remote

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## 2 Theory and instruments

### 2.1 Raman lidar WV profile retrieval

Raman-scattering-based lidar for atmospheric water vapor measurements has been amply described in the literature (Melfi, 1972; Sherlock et al., 1999b; Leblanc et al., 2012). However, to discuss the technical solutions adopted in the system configuration of RMR-H<sub>2</sub>O, it is useful to report the equation relating the water vapor mixing ratio (WVMR,  $w$  in the equation) to the recorded Raman signals:

$$w(z) = \frac{O_N \xi_N \Gamma_N F_N [T(z)] \frac{d\sigma_N/d\Omega}{d\Omega} N_H(z)}{O_H \xi_H \Gamma_H F_H [T(z)] \frac{d\sigma_H/d\Omega}{d\Omega} N_N(z)} \quad (1)$$

In the following, the notation  $x$  stands for the Raman wavelength of the considered atmospheric component (N<sub>2</sub> or H<sub>2</sub>O,  $N$  and  $H$  in the equation, respectively);  $k$  is the ratio between the molecular weight of water vapor and dry air multiplied by 0.781 (the factor expressing the constant fraction of the nitrogen molecule in dry air in the homosphere);  $O_x$  is the overlap function of the lidar channel;  $\xi_x$  is the total lidar receiver optical efficiency;  $F_x [T(z)]$  is the temperature dependent term;  $d\sigma_x/d\Omega$  is the Raman differential backscattering cross section;  $N_x = S_x - B_x$  is the recorded signal  $S_x$  at the Raman wavelength of the atmospheric component  $x$ , subtracted by the associated background  $B_x$ , which is computed by averaging the signal return from above 100–150 km;  $\Gamma_x(z) = \Gamma_x^m \Gamma_x^p$  is the total extinction coefficient term that is usually separated into the molecular ( $\Gamma_x^m$ ) and the particulate ( $\Gamma_x^a$ ) contribution.

Depending on the lidar instrument setup each multiplicative term in the Eq. (1) can have a varying impact on the WVMR measurement.

### 2.2 Instrument characteristics

Whereas the previous Raman water vapor lidar system (Baray et al., 2006; Hoareau et al., 2012), installed at the Saint Denis, near the sea level, was an instrumental

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wavelength-dedicated spherical mirrors relatively to the operational configuration (visible or UV). Pure simultaneous comparisons using both wavelengths were not possible. All other optics are coated to be  $R_{\max}$  at both wavelength. The wavelength swift in the emitter configuration takes 10 min thanks to an easy access to this mirror.

5 A coaxial geometry for emission and reception has been implemented to avoid parallax effects, to extend measurement down to few meters from the ground and to facilitate the alignment. This configuration as well as the global system design has been schematically represented in the Fig. 14 of Hoareau et al. (2013). The primary mirror is a 1.2 m diameter telescope that was previously used at Biscarosse for Rayleigh and Raman measurements (Hauchecorne et al., 1991) and that was refurbished in 10 2011. Light coming from this element is reflected by a secondary flat mirror, tilted at  $45^\circ$  in order to direct the light in one of the side of the telescope where an adjustable diaphragm field stop, located in the focal plane, defines the variable field of view of the system (3.0–0.5 mrad). This element is placed at the entrance of the optical box unit used to separate the Raman and Rayleigh backscattered signals. Thus, the current system uses a set of lenses and mirrors instead of optical fibers to transfer the backscattered signals to the optical ensemble. This configuration, despite a possible increase of optical losses, permits to avoid a systematic bias in water vapor measurements due to fluorescence in fiber-optic cables.

20 The spectral separation of the light is firstly realized by a dichroic beam splitter (BS1) that reflects the visible component of the backscattered radiation toward the visible separation unit (VSU) and transmits the UV component to UV separation unit (USU). These permanently-installed units have the purpose to split the Raman from the Rayleigh–Mie signals and have the same configuration in terms of optical path and equivalent optic elements.

25 Considering the USU, the filtered beam is split by another dichroic beam splitter (BS2) that reflects its 355 nm component toward a band pass interference filter (BP-IFF, bandwidth = 1 nm, maximum transmittance of 55.3 %) and, subsequently, a beam splitter ( $R : T = 92 : 8$ ) that separates the 355 nm beam to the Rayleigh–Mie channels (low









The lidars operated, on average, 3 to 4 h per night, with the exception of the 8 h continued lidar sessions taken during the nights of 9, 10, 11 (during new moon) and 22 April.

### 3.1 Characterization of the system configurations

To enhance the SNR, besides the large collecting surface of the telescope, the RMR-H<sub>2</sub>O lidar can assume several configuration scenarios. As described in Sect. 2, it is possible to double the emitted power by synchronizing the two lasers, to change the wavelength emission from UV to visible and to lower the background noise by reducing the receiver field of view (FOV). Considering that the intensity of the Raman H<sub>2</sub>O channel depends mainly to the highly variable concentration of atmospheric water vapor, the different lidar setups have been evaluated by estimating some representative parameters of the Nitrogen Raman channel. In particular for 30 min time integration lidar sessions, we calculated the maximum altitudes at which the SNR on the nitrogen signal is lower than 0.1, 0.3 ( $z_{\text{err}10}$  and  $z_{\text{err}30}$ , respectively), the signal detectability ( $\text{dtb} = [(S_x - B_x)/B_x]$ ) is higher or equal to 0.1 ( $z_{\text{dtb}}$ ), the altitude,  $z_{\text{ov}}$ , of the full overlap between the emitter and the receiver (i.e. the overlap function  $f_{z_{\text{ov}}} = 1$ ), the background noise of both Raman channels ( $B_N$  and  $B_H$ ) and the correction of the signal linearity. Table 3 reports the values of those parameters for each of the tested measurement scenarios during the nighttime lidar acquisition of MALICCA.

#### UV and visible emission

The opportunity of using the emitting wavelength at 355 and 532 nm (see Sect. 2) allowed a direct comparison of the UV and visible system capabilities that are difficult to determine theoretically, depending on several factors such as the Raman backscattering cross-section, laser source availability and power, and detectors' efficiency.

The lidar sessions acquired with the visible configuration during the first experimental period (September–November 2012) have been compared with the UV lidar sessions of MALICCA. In particular, the first two column of Table 3 resume the results for the

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back-scattering signal and of the background noise, affecting the SNR, the detectability and, in the case of very high-count rates, the linearity response of the PMTs. To find a compromise between these constraints, the effects of several field apertures have been tested during MALICCA. Table 3 reports only the results for the diaphragm aperture of 2 and 2.5 mm (i.e. a FOV of 0.55 and 0.69 mrad, respectively) that optimize the above listed parameters. The two configurations have similar values in terms of SNR and detectability, with the narrower (broader) FOV that optimize the detectability (SNR) of the system and that raises (lowers) the full overlap altitude ( $z_{ov}$ ) due to the defocusing effect that enlarges the spotlight on the diaphragm aperture decreasing the signal intensity at low range.

### Signal linearity correction

Another element, which has to be considered for the choice of the FOV and of the emitter set up, is the saturation of PMT that, in case of a too high number of received photons, causes a nonlinear response of the detector. This phenomenon is corrected using the following exponential law (Singh, 1996):

$$N_c = N_r \exp\left(-\frac{N_r}{N_{\max}}\right) \quad (2)$$

where  $N_r$  are the received photons,  $N_c$  the number of counted photons, and  $N_{\max}$  the number maximum photons that can be counted by the PMT (system). Due to the coaxial emission-reception geometry, the nitrogen Raman channel of the RMR-H<sub>2</sub>O is subjected to saturation. To evaluate and correct this effect using Eq. (2), the value of  $N_{\max}$  for each PMTs of the system has been measured (saturating on purpose the N<sub>2</sub> Raman channel) and then a recursive method to resolve the equation has been applied.

The linearity correction (i.e. the ratio  $N_c / N_r$  in percentage) for the adopted FOVs are reported in Table 3 as the maximum value of the ratio applied in the Nitrogen vertical profile. As expected the saturation effect is higher in case of two-laser emission and with a broader FOV. In conclusion, the FOV of 0.69 mrad will be adopted.

## 3.2 Rejection of the residual signals

To optimize the Raman lidar technique to water vapor measurements, it is necessary to quantify the systematic biases affecting the technique. In particular, several studies (Sherlock et al., 1999; Ferrare et al., 2004; Whiteman et al., 2006; Leblanc et al., 2012) have highlighted that many lidar systems experienced an excess amount of water vapor (wet bias) in the mid-upper troposphere lidar profile, significantly impacting their measurements. The recent work of Whiteman et al. (2012) identified three general causes for this effect: (1) instrumental effects, (2) data processing, (3) atmospheric constituents.

The RMR MAIDO lidar system has been conceived to prevent the wet bias effect. During MALICCA several tests were performed to verify the correct rejection in the water vapor Raman channel system of residual signals due to fluorescence and to Rayleigh, Mie or Raman signal leakage.

### 3.2.1 Excess signal due to fluorescence

As stated by the study of Sherlock et al. (1999a), the weak Raman backscattering signal due to water vapor molecules is susceptible to contamination by fluorescence processes, which can cause systematic errors in Raman Stokes measurements. To reduce this bias, one of the technical solutions adopted for the RMR MAIDO has been to avoid the using of an optical fiber to transfer the backscattered signals to the optical ensemble. This element has been proved to be one major source of fluorescence, causing a contamination signal on the water vapor Raman channel.

However fluorescence processes could arise in any optical component of the lidar system. Thus, to verify the possible presence of such contamination, during the night of 4 April, the interference filter on the water vapor channel has been replaced by one 10 nm band-pass cavity interference filter centered at 432 nm. Since a significant backscatter contribution from atmospheric constituents is not present in this spectral region, any observed signal may be due to the fluorescence.

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random errors that increase to 30 %, more than 50 and 100 % around 15, 16 and 17 km, respectively.

To lower further the statistical error in the UTLS region, lidar data have to be integrated over one or more nighttime sessions (see Sect. 5).

## 4 Calibration

### 4.1 Long-term calibration strategy

The characteristics of the system and of the location, together with the possibility of calculating molecular extinction with the profiles of air number density derived by models, climatological data or measurements (as well as the density of the atmospheric absorbers), permit, in first approximation, to formulate the Eq. (1) of the WVMR measured by the RMR-H<sub>2</sub>O MAIDO Lidar in a simplified form:

$$WVMR(z) = k \frac{\xi_N d\sigma_N/d\Omega}{\xi_H d\sigma_H/d\Omega} \times \frac{S_H(z) - B_H}{S_N(z) - B_N} \times \Gamma_{\Delta}^a \quad (4)$$

where  $C$  is the calibration coefficient of the measurements, namely the factor that converts the measured profiles of backscattered radiation into a useful geophysical variable (i.e. mixing ratio), while  $\Gamma_{\Delta}^a$  is the particulate differential extinction term for the Raman wavelengths of nitrogen and water vapor.

The estimation of the calibration coefficient represents a well-known issue that can still limit a systematic and operational employment of this technique. For this reason, during the last two decades, several efforts have been made to develop a methodology relatively simple, repeatable, stable, and that can be fully characterized in terms of accuracy and associated uncertainties (Ferrare et al., 1995; Whiteman et al., 2003b). In the frame of the NDACC, these requirements are fundamental to ensure the proper long-term monitoring of the (UTLS) water vapor mixing ratio.

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In particular, for the MAIDO RMR-H<sub>2</sub>O lidar two methods have been foreseen: the calibration lamp (CL), and the passive daytime observations (PDO). As highlighted by the works of Leblanc and McDermid (2008) and of Whiteman et al. (2011a) for CL and by the work of Hoareau et al. (2009) for PDO, it is noteworthy to specify that these methods cannot be used to provide an accurate quantification of the system optical efficiency, but only to identify ISPs.

Both the methods are based on collecting the ratio of the collected signals in the water vapor and the nitrogen channels that represents the ratio of the transmittance functions of the two Raman channels ( $TF_{387}/TF_{407}$ ). Previous works show that even if the lamp emission can vary with time the ratio will remain the same.

An ORIEL model 6251NS 75 W Xenon lamp has been mounted on a removable support on the top of the primary telescope to directly illuminate its surface. The CL monitoring procedure consists of acquiring the signals coming exclusively from the illumination by the lamp and then deriving  $TF_{387}/TF_{407}$ . This procedure, which lasts 10 min before the beginning of each water vapor lidar acquisition, has been tested for 11 lidar sessions between 1 and 24 April.

The time series of  $TF_{387}/TF_{407}$ , calculated as the mean of 1 min ratio, are shown in Fig. 4. Because of the high background noise registered in the nitrogen Raman channel during the first days of MALICCA, the 3 April we provided to substitute the PMT on this channel. This instrumental change is well detected by the doubling of the  $TF_{387}/TF_{407}$  mean (horizontal black dashed line) calculated for the lidar session before and after 3 April respectively. A residual variability (mean  $\pm$  SD, blue light regions in the plot) of approximately 9 and 7% characterizes the two identified periods. This is due to the fact that the optical arrangement of the lamp allows lightening only a portion of the telescope surface, causing a not uniform illumination of all of the receiver components. Furthermore, this arrangement has been subjected to small variations. The right side of Fig. 5 depicts the partial illumination of the mirror by the optical arrangement of the CL, while the left side schematically represents the effect on the  $R_{\text{ff}}$  values caused by illuminating four different parts of the RMR-H<sub>2</sub>O telescope surface. A similar range of



values (7 %) was obtained by Whiteman et al. (2008) using a calibration lamp scanned over the full aperture of the Howard University Raman Lidar.

A light trend of approximately 1–2 % is also recorded during each CL session, probably due to an insufficient heating (warm up) of the lamp.

Passive daytime observations to identify ISPs were also tested during MALICCA. The technique consists of measuring the daytime sky background radiation at a given time, changing with season to keep the same solar zenith angle, on the two Raman channels. The main limitation of the method is that clear-sky conditions must be fixed for every measurement because the effect of aerosol and clouds has a strong impact on the  $TF_{387}/TF_{407}$  retrieved values. This requirement limits the employment of the technique. In fact, contrary to the nighttime, the observatory, during daytime, is characterized by a predominance of cloudy conditions. This fact is pointed out by the Fig. 6, where the PDOs performed on 2 and 5 April are depicted in the left and the right plot, respectively. The measurements, both starting at 08:20 UTC (corresponding to a zenith angle of approximately  $63^\circ$ ), last 30 min. The PMT change is still noticeable (the mean  $TF_{387}/TF_{407}$  value is 0.55 and 1.6 approximately for 2 and 5 April, respectively), but the  $TF_{387}/TF_{407}$  of 5 April are strongly affected by a rapid transit of several small clouds (a typical condition at Maïdo site during daytime convection), causing a variation of  $TF_{387}/TF_{407}$  values of even 20 %. Furthermore, the hypothesis that the system behaves similarly during nighttime and daytime has to be verified.

Given these results, major instrumental changes (i.e. variations of  $TF_{387}/TF_{407}$  greater than 10–15 %) of the RMR- $H_2O$  lidar system will be monitored through the implementation of CL method. However in the future, to gain on lamp stability and ameliorate the method sensitivity, it is planned to wait ten minutes before starting such a measure and to fix that the lamp arrangement so that it will not be subjected to any variation.

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only 15 m above the station. Furthermore the co-located GPS, described in Sect. 2.3, can provide every hour a reference value of IWV. For these reasons a calibration strategy based on GPS IWV was tested during MALICCA.

The RMR-H<sub>2</sub>O IWV is calculated using the lidar water vapor profile completed adding a surface point derived by the humidity measurement of the co-located COMET T7310 automatic weather station and an upward extension (above 16 km) based on the European Center for Medium-Range Weather Forecast (ECMWF) operational water vapor profiles. It must be noted that while the ground point can affect the RMR-H<sub>2</sub>O IWV value even for 1 %, the ECMWF data, re-sampled on a latitude-longitude resolution grid of 1.125° and converted to water vapor mixing ratio by means of the empirical saturation vapor pressure over liquid water formulas of Hyland and Wexler (1983) has an impact of less than 0.1 %.

The calibration procedure consists of integrating only the lidar profiles acquired 30 min before and after the hourly IWV values retrieved by the GPS, calculating the corresponding un-calibrated RMR-H<sub>2</sub>O IWV value and scaling it to the IWV GPS coincident value.

The time-series of the IWV GPS calibration coefficients associated to their errors (black vertical bars) are displayed in Fig. 7 for the period 1–24 April. The horizontal black dashed lines depict the median calibration factors for the two ISPs identified by the calibration lamp. The N<sub>2</sub> PMT substitution causes a jump of the calibration median coefficient by a factor more than 5, with a variability (i.e. the normalized pseudo-SD) of approximately 13–14 % for the two periods. To validate the procedure, the calibration coefficients have been also estimated through 11 of the 15 RS92 launched during the campaign. In particular two methods were performed: radiosonde-lidar comparison of water vapor profiles and of water vapor columns (PROF RS92 and IWV RS92, respectively). For the former, the raw lidar signals are integrated for 60 min starting at the radiosonde launching time ( $t = t_0$ ). The calibration coefficient is computed through the median of the ratio of all radiosonde-lidar matching pairs, in the altitude range between 3 and 11 km. The upper limit is fixed to keep the lidar signal to noise ratio (SNR) higher

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of different altitudes are presented in Table 6, where the errors obtained by a standard 120 min integration are also shown. The  $\Delta\text{WVMR}$  has been estimated using the formula obtained by combining Eqs. (1) and (3) and following Whiteman et al. (2003b):

$$\frac{\Delta\text{WVMR}}{\text{WVMR}} = \sqrt{\left(\frac{dw}{w}\right)^2 + \left(\frac{dC}{C}\right)^2 + \left(\frac{d\Gamma_{\Delta}}{\Gamma_{\Delta}}\right)^2}, \quad (5)$$

where  $\Gamma_{\Delta}$  is the ratio between the total extinction coefficient terms at nitrogen and water vapor Raman wavelengths.

Neglecting in a first approximation the contribution of the extinction term and assuming as  $dC$  the pseudo-SD calculated for the calibration method in Sect. 4.3, it is possible to fully quantify the  $\Delta\text{WVMR}$  of the Maïdo H<sub>2</sub>O lidar in the UTLS during MAL-ICCA. The results for the single day integrations (i.e. 120 and 240 min) are the mean values calculated over the 8 sessions that have been also used to simulate the monthly lidar profile.

For a two-hours integration, the  $\Delta\text{WVMR}$  is more than 4 ppmv above 15 km, which corresponds of a total relative error of 65 and of more than 100 % at 16 km, confirming the impossibility, with this temporal resolution, of covering the whole troposphere.

On the contrary, the daily integration gives a  $\Delta\text{WVMR}$  that ranges between 1.5 and 2 ppmv (i.e. a relative error of up to 50 %) in the upper troposphere (from 15 to 17/18 km), a region where a recent research (Whiteman et al., 2011b) indicated that random uncertainties of 50 % are acceptable for trend detection purposes if regular and frequent (e.g. every three or four days) measurements are taken. Thus, this temporal integration seems to be a good compromise, in terms of accuracy and timescale variability, to study the upper tropospheric water vapor.

The latter approach allows extending the water vapor measurements in the LS. In fact, as illustrated in Table 6, the integration of eight 4 h lidar sessions (i.e. the number of sessions that would be acquired during a month of regular observations) could lower the  $\Delta\text{WVMR}$  to less than 1 ppmv at 20 km, with a relative error kept below 25 %. This type of integration could be addressed to the LS, which is characterized by a less

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natural water vapor variability (Hurst et al., 2011), but more sensitive to additional measurement noise than the upper troposphere. Furthermore, this monthly lower stratospheric water vapor profile might also be useful for the quality control of the data. In fact, as shown by the work of Whiteman et al. (2012), the monthly average water vapor mixing ratios measured by the Microwave Limb Sounder (MLS) can be used to quality control Raman water vapor lidar data. This sounder installed on the AURA satellite observes thermal microwave – far infrared emissions from the Earth’s atmosphere in 5-spectral regions. The water vapor profiles are retrieved from 183 GHz H<sub>2</sub>O rotational line spectrum measurements and their precision and accuracy in LS are well documented in literature (Lambert et al., 2007; Vömel et al., 2007b; Livesey et al., 2013).

In our case, the comparison between the campaign-integrated lidar profile and the MLS (version 3.3) mean WVMR profile, derived by the selection of 7 AURA-MLS passages over a 2° × 3° grid box centered on Reunion Island during MALICCA, is depicted in Fig. 12 together with their relative difference (i.e. (MLS-Maïdo)/Maïdo).

Below 16.5–17 km (100 hPa), MLS shows a significant dry bias (30–40 %). This feature could be caused both by the different instrumental sampling and by the MLS systematic bias in the upper troposphere due to its poor resolution in the very fast transition from dry stratosphere to wet troposphere (Leblanc et al., 2012).

On the contrary, a good agreement is observed in LS between 17 and 20 km (i.e. 90 and 55 hPa) with a relative difference of less than 10 % and the lidar profile that falls inside the MLS mean  $\pm 2\sigma$  values. This result confirms the absence of wet biases in the UTLS water vapor lidar profile and validates the value of the calibration coefficient. Above 21 km (50 hPa), due to the increase of the  $\Delta$ WVMR, the lidar water vapor profile is unreliable.

## 6 Summary and conclusions

A new RMR-H<sub>2</sub>O lidar has been installed at the Maïdo altitude station facility of La Réunion. The system, designed to ameliorate the critical drawbacks of the previous WV Raman prototype located at Saint Denis near the sea level, will be devoted to the long-term survey of water vapor in Upper Troposphere Lower Stratosphere.

The objectives of the MALICCA campaign, held in April 2013, were to validate the water vapor measurements of the new lidar, to set up a calibration methodology, and to evaluate its performances and capabilities with a particular focus on the UTLS domain.

The validation of the RMR-H<sub>2</sub>O measurements passed through three phases:

- a. Testing the different system configuration scenarios. Regarding the transmitter, the UV emission mode is preferable to the visible one in terms of the maximum heights reached by the SNR and the detectability, while doubling the emitted power (i.e. coupling two lasers) increases the SNR, but also the background noise and the saturation effect of the PMT in the nitrogen Raman channel. For the receiver, the fields of view of 0.55 and 0.69 mrad are those that better satisfy the constraints of the SNR and the linearity response of the PMTs.
- b. Verifying the presence of possible parasite signals. The absence of a distinguishable fluorescent contamination in the Raman water vapor channel has been verified measuring the signal at 432 nm, a spectral region where there is a negligible backscatter contribution from atmospheric constituents. Additionally, the nominal OD of the Raman H<sub>2</sub>O channel (15 and 9, at 355 and 387 nm, respectively) seems to guarantee a correct rejection to the signal contamination due to the Rayleigh, Mie or Raman nitrogen signals. This has been confirmed by comparing the water vapor mixing ratio profiles measured by the lidar and the co-located RS92 radiosondes in correspondence of a cirrus layer. No evidence of signal leakage into the water vapor Raman channel has been detected.

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c. Determining the height dependence of the lidar statistical error. The lidar performances measured during MALICCA have been compared to those simulated by Hoareau et al. (2012). The mean altitudes above the sea level where the H<sub>2</sub>O measurements have a relative statistical error per bin within 15 and 30 % are 12.3 and 13.4 km respectively, 1.7 and 2.9 km lower than those estimated by the lidar simulation. Applying an height dependent sliding average to the lidar raw data, with a temporal integration of 30 and 120 min, limits the statistical error to less than 10 % below 13 km, maintaining a high vertical resolution in the lower and the middle troposphere. Above 13 km the vertical resolution gradually degrades with random errors equal to more than 50 % at 16 km.

Since one of the overall goals of the RMR-H<sub>2</sub>O Maïdo lidar is to provide long-term monitoring, a hybrid calibration methodology has been set up and validated to insure optimal lidar calibration stability with time. The receiver transmittance is monitored through the calibration lamp method that, at the moment, can detect transmittance variations greater than 10–15 %. The calibration coefficients are then calculated through the hourly values of IWV provided by the co-located GPS. The comparison between the calibration constants derived by the GPS and the Vaisala RS92 radiosondes launched at Maïdo during MALICCA, points out an acceptable agreement in terms of accuracy of the mean calibration value (with a difference of approximately 2–3 %), but a significant difference in terms of variability (14 vs. 7–9 %, for GPS and RS92 calibration procedures, respectively). Further studies are needed to characterize these dissimilarities, which can be partly explained by the sampling difference of the considered instruments (i.e. lidar, GPS and radiosonde) that is stressed by the high and local variation of water vapor regimes in La Reunion Island. However, the higher variability of IWV GPS strategy is balanced by the possibility of having a greater number of samples during a lidar session.

During MALICCA, the lidar measurements have been compared to 15 co-located and simultaneous RS92 radiosondes. A relatively good agreement between the instruments (i.e. relative difference below 10 %) is measured in the low and the middle troposphere

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(2–10 km). The upper troposphere (up to 15 km) is characterized by a larger spread (approximately 20%), which lowers below 10% by excluding from the statistics the nights of 10 and 11 April. This result confirms that, at high altitudes and depending on the water vapor spatial distribution, the distance of the two sensors can significantly affect the comparison between lidar and radiosoundings.

To measure the water vapor in the UTLS region two different integration methodologies have been adopted: nighttime integration and monthly integration. The former, which consists of a temporal integration of 240 min, allows measuring the WVMR in the UT (up to 17/18 km) with an absolute error of 2 ppmv. The latter, obtained simulating a month of regular measurements (240 min  $\times$  8 lidar sessions), allows extending the measurements in the lower stratosphere, lowering the absolute error to 1 ppmv at 20 km.

Finally, the comparison between the lidar monthly profile and the mean WVMR profile measured by MLS can be used as a quality control procedure of the lidar product. Following Whiteman et al. (2012), the good agreement observed in the lower stratosphere (from 17 to 20 km) could attest the absence of significant wet biases and validate the calibration procedure.

In conclusion, the design and the performance of this new lidar system permit the covering of a large altitude range from the ground up to the lower stratosphere (19–20 km). In particular the obtained results show the capabilities of the H<sub>2</sub>O lidar to measure water vapor in UTLS down to few ppmv with random errors around 50 and 25% accordingly to the adopted integration scheme. The achievement of this objective opens up new opportunities for the characterization of the water vapor in this atmospheric region, in terms of long-term monitoring, process investigation and instrumental inter-comparison and satellite validation. Within this frame, further tests are planned to optimize the calibration procedure, with the goal of increasing the accuracy and stability of the method. In the next future, to use the MAIDO H<sub>2</sub>O lidar as a reference instrument in the southern subtropics, it will be crucial to improve the data quality tests,

implementing operational procedures to characterize the measurements and minimize the influence of systematic errors.

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**Table 1.** Pass band interference filter characteristics of the Raman channels.

	N <sub>2</sub> Vis	H <sub>2</sub> O Vis	N <sub>2</sub> UV	H <sub>2</sub> O UV
Central wavelength (nm)	606.9	660.0	386.7	407.44
Passband width, FWHM (nm)	1.0	1.0	3.0	0.98
Peak transmittance (%)	66	72	63	68

**Table 2.** Transmitter and receiver characteristics of the MAIDO-H<sub>2</sub>O lidar system.

MAIDO H <sub>2</sub> O (41.8° S, 12.6° E, 2168 m a.s.l.)	
<b>Transmitter</b>	
Laser Type	Nd:Yag
Wavelength	532 and/or 355 nm
Energy per pulse	800–400 mJ
Pulse repetition rate	30 Hz
Power	11–22 or 24–48 W
beam diameter	200 mm (with a 5x beam expander)
beam divergence	0.1 mrad
Emission-reception geometry	Coaxial
<b>Receiver</b>	
Type of telescope	Newtonian
Diameter, focal length	1200, 3007 mm
Field of view (mrad)	0.1–2
Optic fiber	no
<b>Data acquisition</b>	
Raman channels N <sub>2</sub> (nm)	387, 607
H <sub>2</sub> O (nm)	407, 660
Elastic channels (nm)	355_a, 355_b, 532_a, 532_b
Sounding range (km)	2–25 (Raman) 7–100 (elastic)
Time resolution (sec)	60
Vertical resolution (m)	15

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**Table 3.** Lidar performance parameters of the Nitrogen Raman channel of the MAIDO-H<sub>2</sub>O lidar for different tested configurations. A temporal integration of 30 min has been applied to the raw lidar data and no vertical integration. The percentage errors of the linearity correction values (last row) are given in parentheses.

Day	23 Oct 2012	21 Apr 2013	21 Apr 2013	4 Apr 2013	4 Apr 2013
Moon	1st qrt + 3	1st qrt + 3	1st qrt + 3	3st qrt + 1	3st qrt + 1
Aerosol	clear sky	clear sky	clear sky	clear sky	clear sky
Configuration:					
Laser	1 (532 nm)	1 (355 nm)	2 (355 nm)	1 (355 nm)	1 (355 nm)
FOV	0.55 mrad	0.55 mrad	0.55 mrad	0.55 mrad	0.69 mrad
$z_{\text{err}10}$ [km]	23.3	28.9	30.4	28.3	29.1
$z_{\text{err}30}$ [km]	28.1	37.5	38.5	36.7	37.3
$B_N$ [no of photons]	17.13	0.72	1.32	0.41	0.60
$B_H$ [no of photons]	0.34	0.42	0.72	0.14	0.25
$z_{\text{dtb}}$ [km]	33.1	47.7	46.7	45.2	44.7
$z_{\text{ov}}$ [km]	5.5	9.0	9.2	7.4	7.1
Linearity correction	0.07 (3%)	0.06 (3%)	0.08 (3%)	0.10 (3%)	0.19 (3%)

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**Table 4.** Comparison of the MAIDO-H<sub>2</sub>O capabilities estimated by the numerical simulation of Hoareau et al. (2012) and calculated as the mean of ten nighttime measurements acquired during MALICCA. The minimum and the maximum values (in the brackets) of the measured parameters are also reported.

	H <sub>2</sub> O Simulation	H <sub>2</sub> O MALICCA (10 session)
Lidar configuration	1 (355 nm) 0.25 mrad	1 (355 nm) 0.55 mrad
Data integration	30 min–150 m	30 min–150 m
$B_H$ [no of pht]	4.8	0.7 (0.5–1.8)
$Z_{H_2O15\%}$ [km]	14.6	12.3 (9.7–13.0)
$Z_{H_2O30\%}$ [km]	16.3	13.4 (11.0–14.3)

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**Table 5.** Principal results (median, pseudo-SD, standard error values and the number of points) for the three calibration methods tested during MALICCA.

	$C_{\text{med}}$	PSTD/ $C_{\text{med}}$ (%)	Se/ $C_{\text{med}}$ (%)	# points
Lid-GPS IWV	214	13	2	55
Lid-RDS prof	221	9	3	11
Lid-RDS IWV	220	7	2	11

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**Table 6.** Total absolute ( $\Delta$ WVMR) and relative errors of the calibrated Raman lidar water vapor measurements in the UTLS (between 13 and 20 km) for three different data products: two – hours, daily (240 min) and monthly integration.

Alt. (km)	120 min integration	Nighttime integration	Monthly integration	Vert. Resol. (km)
	$\Delta$ WVMR (ppmv, %)	$\Delta$ WVMR (ppmv, %)	$\Delta$ WVMR (ppmv, %)	
13	3.1 (10 %)	2.6 (8 %)	2.5 (8 %)	0.435
14	4.0 (23 %)	2.1 (12 %)	1.5 (8 %)	0.585
15	2.5 (27 %)	1.5 (16 %)	0.9 (9 %)	1.005
16	4.1 (65 %)	1.5 (24 %)	0.7 (10 %)	2.055
17	4.5 (110 %)	2.0 (50 %)	0.6 (16 %)	4.065
18	–	2.0 (55 %)	0.5 (16 %)	5.265
19	–	3.1 (75 %)	0.9 (25 %)	6.015
20	–	3.2 (75 %)	1.0 (25 %)	6.765

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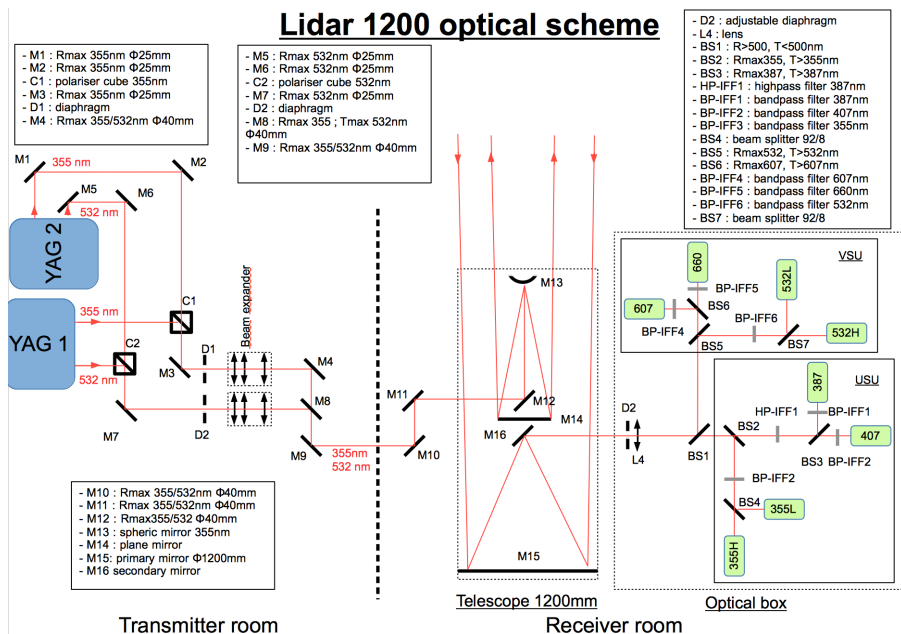
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**Figure 1.** Optical scheme of the Mäido lidar. The optical components of the visible separation unit (VSU) and the UV separation unit (USU) are described in the text.

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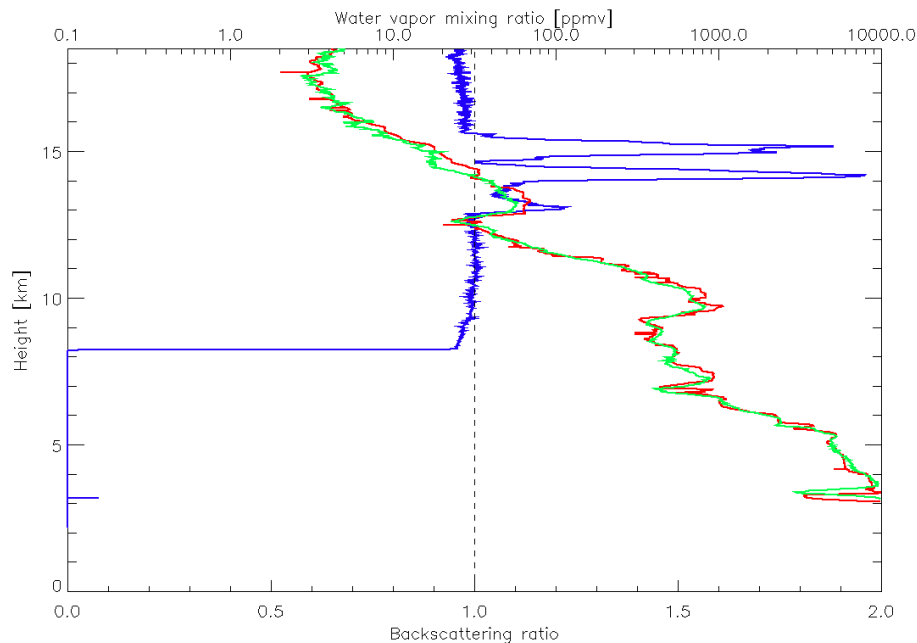
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**Figure 2.** Backscattering ratio (i.e. the ratio between the Rayleigh and the Raman channels at 355 and 387 nm, respectively) and WVMR profiles (blue and green curves, respectively) observed during the night of 8 April 2013, together with the WVMR measured by the co-located RS92 radiosonde (red curve). Both lidar profiles are integrated for 60 min starting at the radiosonde launching time (i.e. 20:50 UT).

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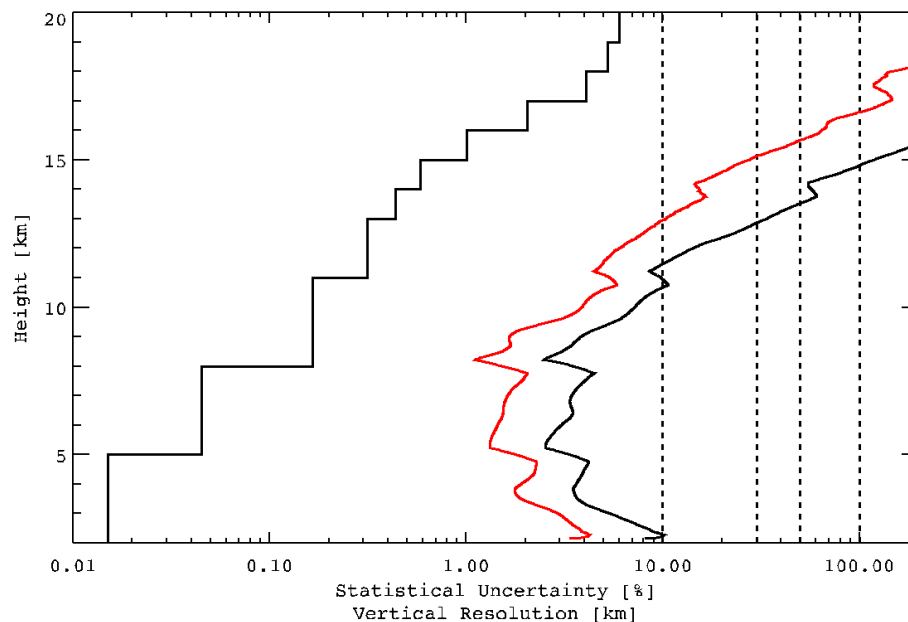
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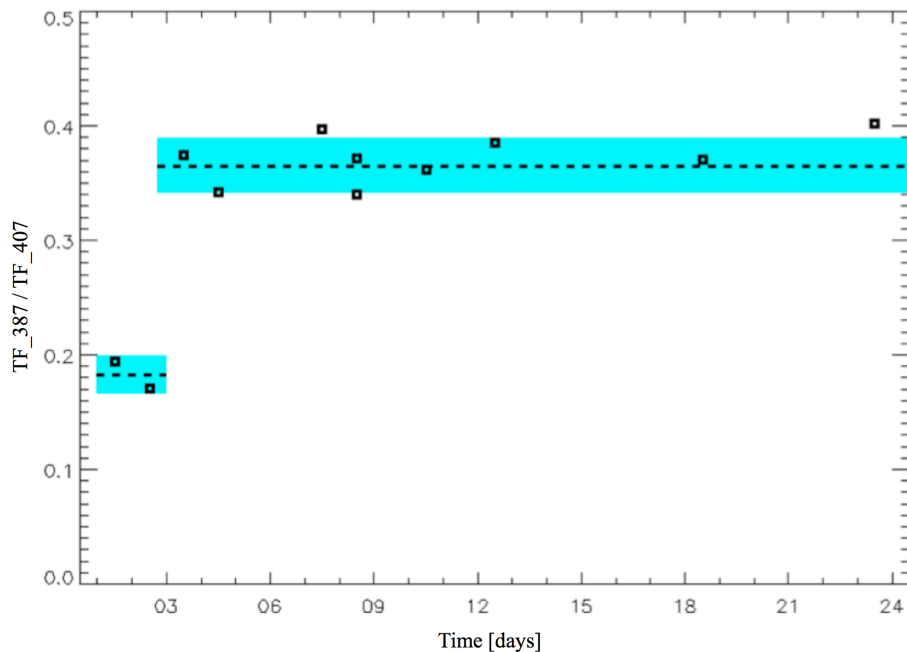
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**Figure 3.** Mean statistical uncertainty (%) after the vertical filtering scheme calculated for ten nighttime measurements with the same lidar configuration. Data are temporally integrated for 30 and 120 min (black and red curves, respectively). The step black curve represents the corresponding vertical resolution (km).



**Figure 4.** Temporal evolution of the transmittance functions of the two Raman channels measured through the lamp method during the MALICCA campaign. Dashed horizontal lines represent the median values, while the blue light region defines the residual variability (mean  $\pm$  SD).

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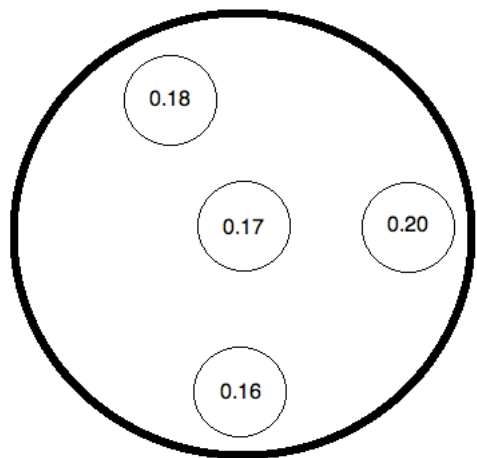
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**Figure 5.** Left: representation scheme of the  $R_{ff}$  values in functions of the illuminated portions of the RMR-H<sub>2</sub>O telescope surface. Right: example of the partial illumination of the mirror by the optical arrangement of the calibration lamp.

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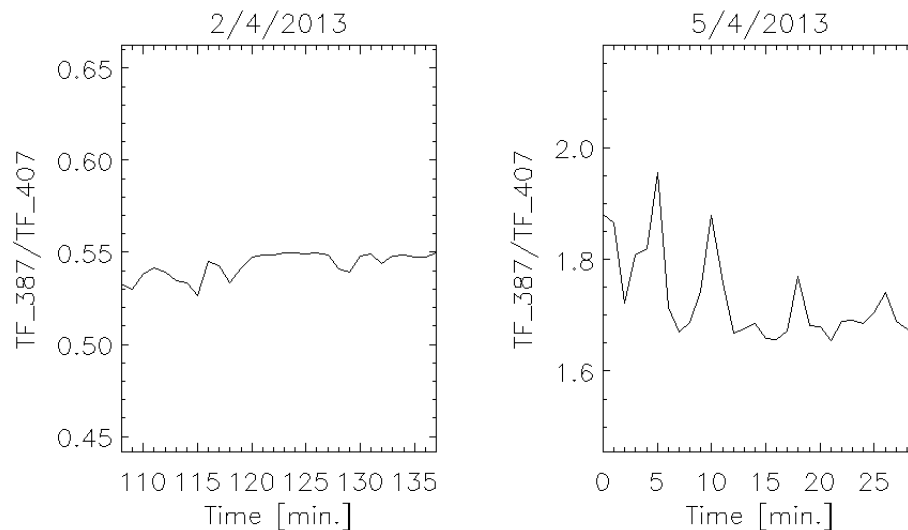
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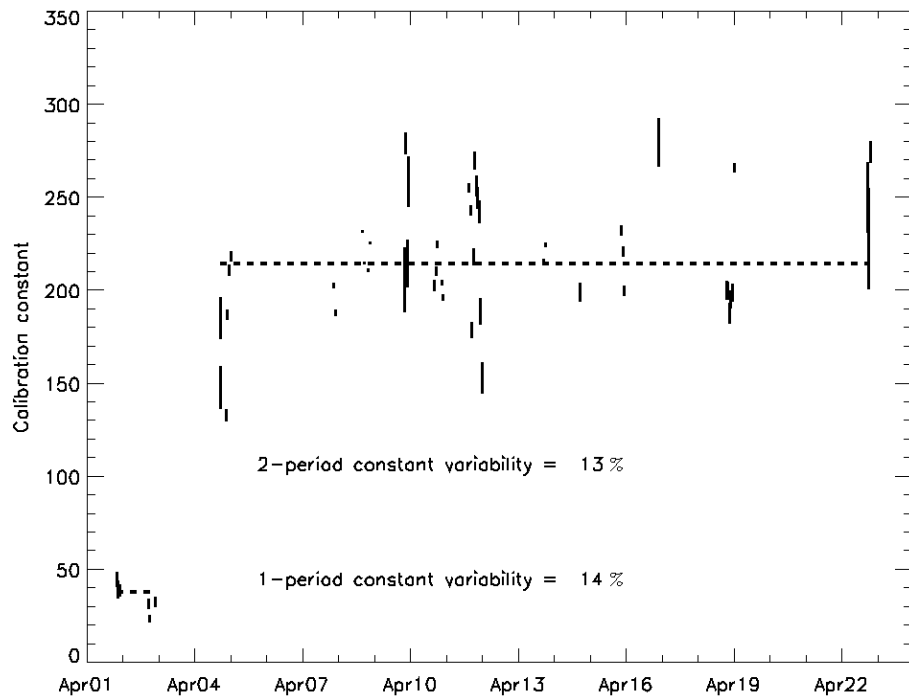
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**Figure 6.**  $TF_{387}/TF_{407}$  determined by 30 min of passive daytime observations at approximately  $63^\circ$  of the solar zenith angles for 2 and 5 April (left and right plot, respectively).

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**Figure 7.** Time series of lidar calibration factors with their associated errors for the RMR-H<sub>2</sub>O lidar determined by IWV GPS during the MALICCA campaign. Dashed horizontal lines indicate the median values for the ISPs identified by the CL monitoring method.

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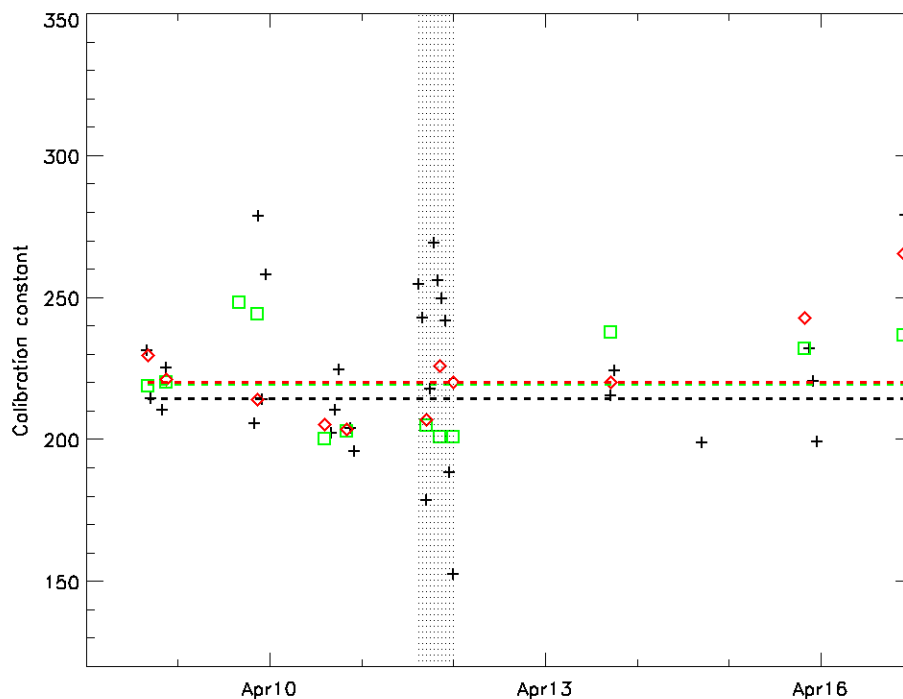
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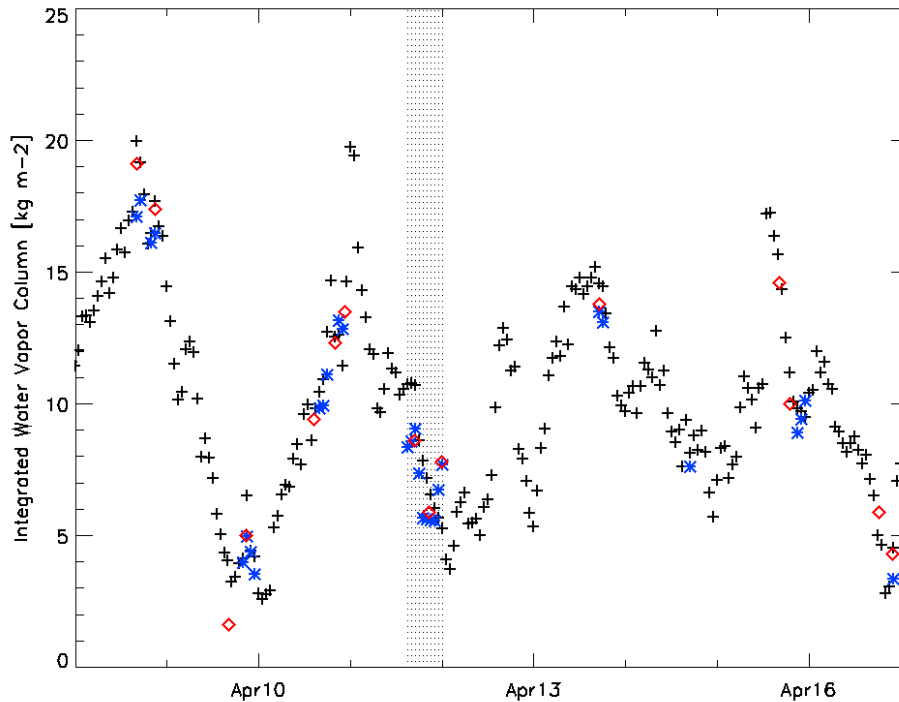
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**Figure 8.** Time series of the lidar calibration factors for the RMR-H<sub>2</sub>O lidar determined by IUV GPS, IUV RS92 and PROF RS92 approaches (black crosses, red diamonds and green squares, respectively) for the period 8–16 April. Dashed horizontal lines indicate the median values for each method. Dotted vertical lines highlight the lidar measurement session acquired between 15:00 UTC of the 11 April and the 00:00 UTC of 12 April 2013.

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**Figure 9.** Time series of IWV estimated by GPS, RS92 and RMR-H<sub>2</sub>O lidar calibrated through the GPS procedure (black crosses, red diamonds and blue stars, respectively) for the period 8–16 April. Dotted vertical lines highlight the period between 15:00 UTC of the 11 April and the 00:00 UT of 12 April 2013.

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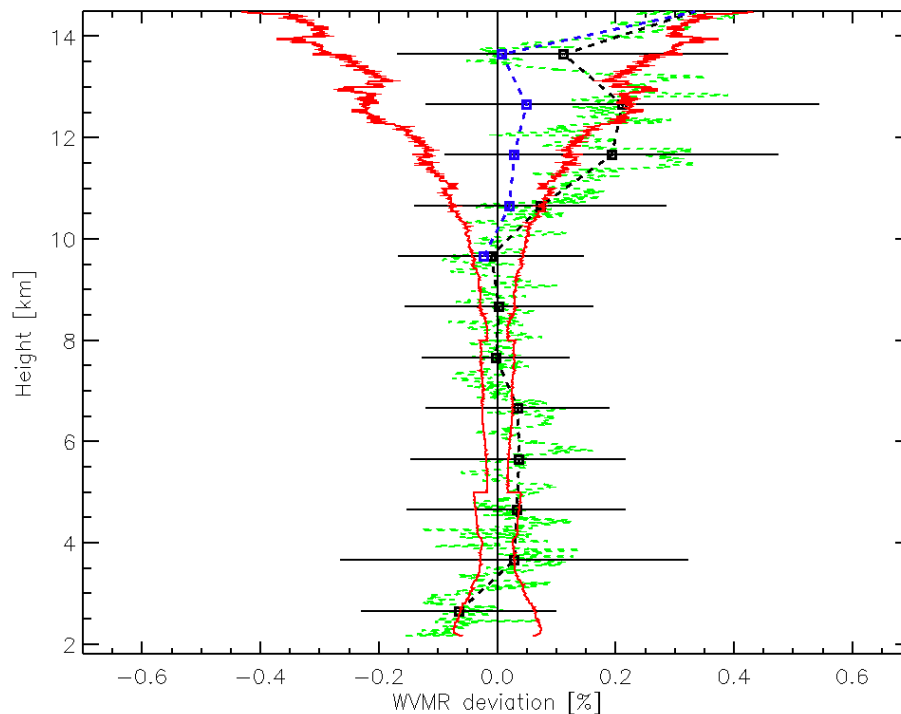
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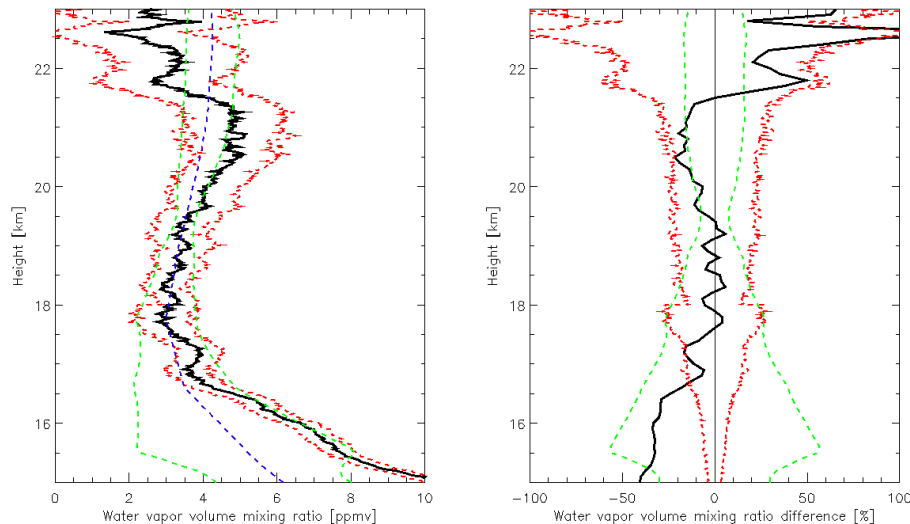
**Figure 11.** WVMR (lidar-rds)/rds relative difference (green dashed curve) between 12 RS92 flights and the 12 corresponding 1 h integrated lidar profiles acquired during MALICCA. Black squares and horizontal bars depict the relative difference averaged on 1 km thick layer and its related SD, while the blue squares represent the WVMR relative deviation excluding the lidar-RS92 comparisons of 10 and 11 April. Red curves are the mean lidar statistical error.

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**Figure 12.** Left plot: UTLS water vapor measurements derived by the lidar campaign-integrated profile (black line) and by the MLS average profile calculated during MALICCA (blue dashed line). Red dotted curves are the associated total lidar error, while green dashed lines represent the mean  $\pm 2\text{-}\sigma$  of the MLS profile. Right plot: relative difference,  $100 \times (\text{MLS-Maïdo})/\text{Maïdo}$ , between the lidar and the MLS UTLS water vapor measurement (black line), together with the associated lidar uncertainty and the  $2\text{-}\sigma$  MLS profiles (red dotted and green dashed curves, respectively).

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