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**Long-term evolution
and seasonal
modulation of
methanol**

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



Abstract

Methanol (CH_3OH) is the second most abundant organic compound in the Earth's atmosphere after methane. In this work, we present the first long-term time series of methanol total, lower tropospheric and upper tropospheric-lower stratospheric partial columns derived from the analysis of high resolution Fourier transform infrared solar spectra recorded at the Jungfraujoch station (46.5°N , 3580 m a.s.l.). The retrieval of methanol is very challenging due to strong absorptions of ozone in the region of the selected ν_8 band of CH_3OH . Two wide spectral intervals have been defined and adjusted in order to maximize the information content. Methanol does not exhibit a significant trend over the 1995–2012 time period, but a strong seasonal modulation characterized by maximum values and variability in June–July, minimum columns in winter and a peak-to-peak amplitude of 130 %. In situ measurements performed at the Jungfraujoch and ACE-FTS occultations give similar results for the methanol seasonal variation. The total and lower tropospheric columns are also compared with IMAGESv2 model simulations. There is no systematic bias between the observations and IMAGESv2 but the model underestimates the peak-to-peak amplitude of the seasonal modulations.

1 Introduction

Methanol (CH_3OH) is the second most abundant organic molecule in the atmosphere after methane with concentrations between 1 (Singh et al., 2001) and 20 ppbv (Heikes et al., 2002), despite a lifetime that has been estimated to lie between 4.7 days (Millet et al., 2008) and 12 days (Atkinson et al., 2006). Plant growth is the largest source of methanol with a 65–80 % contribution to its emissions (Galbally and Kirstine, 2002; Jacob et al., 2005). The atmospheric production of CH_3OH through peroxy radical reactions represents up to 15–23 % of its sources (Madronich and Calvert, 1990; Tyndall et al., 2001). Other sources of methanol are plant matter decaying (Warneke, 1999),

AMTD

7, 4659–4692, 2014

Long-term evolution and seasonal modulation of methanol

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

biomass burning (Dufour et al., 2006; Paton-Walsh et al., 2008), fossil fuel combustion, vehicular emissions, solvents and industrial activities.

Methanol influences the oxidizing capacity of the atmosphere through reaction with the hydroxyl radical (Jimenez et al., 2003), its main sink, leading to the formation of water vapour and either CH_3O or CH_2OH radicals which both react with O_2 to give HO_2 and formaldehyde (H_2CO) (Millet et al., 2006). The photo-oxidation of formaldehyde, a key intermediate in the oxidation of numerous volatile organic compounds, leads to the formation of HO_2 radicals and carbon monoxide (CO). As a consequence, CH_3OH is considered as a source of CO with a yield close to 1 (Duncan et al., 2007). The main sources and sink of methanol are characterized by significant seasonal modulations. This results in a strong signal for CH_3OH , with maximum and minimum abundances observed in the Northern Hemisphere at the beginning of July and in December, respectively (Rinsland et al., 2009; Stavrakou et al., 2011; Wells et al., 2012; Cady-Pereira et al., 2012), reflecting the seasonality of biogenic sources.

In the last decade, ground-based (Schade and Goldstein, 2001, 2006; Karl et al., 2003; Carpenter et al., 2004), ship (Warneke et al., 2004) and aircraft (Singh et al., 2006; Fehsenfeld et al., 2006) in situ measurements combined to space-based detections including the Infrared Atmospheric Sounding Interferometer (IASI) onboard the MetOp-A satellite (Razavi et al., 2011), the TES (Tropospheric Emission Spectrometer) nadir-viewing Fourier transform spectrometer (FTS), on board the Aura satellite (Beer et al., 2008), and the solar occultations recorded by the Atmospheric Chemistry Experiment-FTS (ACE-FTS, Bernath et al., 2005; Dufour et al., 2006, 2007) have supplied numerous observations of CH_3OH which have provided valuable insights on the distribution and budget of methanol at the global scale. However, there still remain large uncertainties in our knowledge of the methanol global sources and sinks in the atmosphere, as indicated by the large discrepancies existing between different measurement-based estimates of the total sources (Galbally and Kirstine, 2002; Tie et al., 2003; von Kuhlmann et al., 2003a, b; Jacob et al., 2005; Millet et al., 2008; Stavrakou et al., 2011). Previous studies have reported the measurement of

Long-term evolution and seasonal modulation of methanol

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

⏪

⏩

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



methanol from ground-based infrared solar absorption observations performed at Kitt Peak (31.9° N, 111.6° W, 2090 ma.s.l.; Rinsland et al., 2009) and at Saint-Denis (Reunion Island, 21° S, 55° E, 50 ma.s.l.; Stavrakou et al., 2011; Vigouroux et al., 2012).

In this paper, we report the first methanol time series derived from ground-based high-resolution infrared spectra recorded with a Fourier Transform InfraRed (FTIR) spectrometer operated under clear sky conditions at the high-altitude International Scientific Station of the Jungfraujoch (ISSJ, Swiss Alps, 46.5° N, 8.0° E, 3580 ma.s.l.; Zander et al., 2008). Most of the available spectra have been recorded within the framework of the Network for Detection of Atmospheric Composition Change monitoring activities (NDACC; see <http://www.ndacc.org>). A detailed analysis was conducted to optimize the retrieval strategy of atmospheric methanol in order to minimize the fitting residuals while maximizing the information content. A thorough discussion of the retrieval strategy, data characterization (information content and error budget), long-term trend and seasonal cycle of total and partial columns of methanol above Jungfraujoch is presented here. This paper is organized as follows. A detailed description of the optimized retrieval strategy is given in Sect. 2. The characterisation of our data by their eigenvectors and error budget is discussed in Sect. 3. Finally, in Sect. 4, we present and discuss the results, focusing on the intra-annual and intra-day variability of methanol at ISSJ along with comparisons with in situ measurements, satellite occultations and model calculations.

2 Retrieval strategy

Regular FTIR observations have been carried out at the ISSJ with a homemade spectrometer since 1984, complemented in the early 1990s and definitely replaced in 2000 by a commercial Bruker IFS120HR instrument (Zander et al., 2008). This spectrometer is equipped with HgCdTe and InSb cooled detectors, allowing us to cover the 650 to 4500 cm^{-1} region of the electromagnetic spectrum. Since 1991, the FTIR instruments are affiliated to the NDACC network.

Long-term evolution and seasonal modulation of methanol

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



3 Data characterization and error budget

Information content has been carefully evaluated and typical results are displayed on Fig. 2. The information content is significantly improved, with a typical Degree Of Freedom for Signal (DOFS) of 1.82, in comparison with DOFS of about 1 in previous studies (e.g., Rinsland et al., 2009; Vigouroux et al., 2012). In Fig. 2, the first eigenvector and eigenvalue (see left panel, in orange) show that the corresponding information is mainly coming from the retrieval (99%). The increase of information content allows us to retrieve a tropospheric column (Tropo, from 3.58 to 10.72 km) with only 1% of a priori dependence as well as two partial columns with less than 30% of a priori dependence (second eigenvector) i.e. a low-tropospheric (LT, from 3.58 to 7.18 km) and an upper troposphere-lower stratosphere (UTLS, from 7.18 to 14.84 km).

The error budget is calculated following the formalism of Rodgers (2000), and can be divided into three different error sources: the smoothing error expressing the uncertainty due to finite vertical resolution of the remote sounding system, the forward model parameters error, and the measurement noise error. The right panel of Fig. 2 gives the corresponding error budget, with identification of the main error components, together with the assumed variability. Error contributions for total and all three partial columns are reported in Table 1.

Through a perturbation method, we also accounted for other error sources: systematic errors such as the spectroscopic line parameters and the misalignment of the instrument while uncertainty on the temperature and on the solar tracking is considered to be source of random error. Table 1 provides an error budget resulting from major instrumental and analytical uncertainties. For the spectroscopic line parameters, we included in our error budget the uncertainty on line intensities provided by the HITRAN database. As methanol line intensities matter, a rough idea of the accuracy of the intensities can be obtained from Table 8 of Xu et al. (2004) work as it reports an RMS deviation of 7%. It should be noted that the uncertainty on ozone and its isotopologues lines, according to HITRAN-08 parameters, amounts to between 5 and 10% (Rothman

measurements. The second one, “IASI”, uses emissions constrained by IASI vertical column data in an inverse modelling framework based on the adjoint of IMAGESv2.

4.2 Time series and long term trend

In order to produce the first long-term time series of atmospheric methanol above Jungfraujoch, three criteria were used to reject noisy measurements or weak absorption: (i) when negative methanol mixing ratios are retrieved, (ii) when rms (root mean square, difference between calculated and observed absorption) was out of the interval defined by the 95 % level of confidence ($2\text{-}\sigma$), (iii) when the number of iterations reached the fixed maximum. After implementation of these criteria, the total number of valid measurements is 4271 obtained on 1476 days of measurements between 1995 and 2012. For the trend calculations, we used the statistical tool developed by Gardiner et al. (2008) that employs a bootstrap resampling method. The function fitted to the time series is a combination of a linear component and a 3rd order Fourier series, i.e.:

$$F(t, b) = c_0 + c(t - t_0) + b_1 \cos 2\pi(t - t_0) + b_2 \sin 2\pi(t - t_0) + b_3 \cos 4\pi(t - t_0) + b_4 \sin 4\pi(t - t_0) + b_5 \cos 6\pi(t - t_0) + b_6 \sin 6\pi(t - t_0) \quad (1)$$

where c_0 is the abundance at the reference time t_0 for the linear component (seasonalized data), and c is the annual trend. Figure 3 shows the whole times series of daily mean methanol total columns above Junfraujoch. We evaluated the trend of methanol total columns over the 1995–2012 time period and found a yearly negative trend of $(-1.34 \pm 2.71) \times 10^{13}$ molecules cm^{-2} or $-0.18 \pm 0.36 \%$ ($2\text{-}\sigma$), i.e. a non-significant trend at this level of confidence which is consistent with the trend computed by Rinsland et al. (2009). A non-significant trend has been computed also for both partial column subsets. Hence the results indicate a long-term trend which is not statistically significant and a strong seasonal variation.

Long-term evolution and seasonal modulation of methanol

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



4.3 Methanol seasonal modulation

As the results for the full time series do not indicate a statistically significant trend, we illustrate in Fig. 4 the daily mean total columns over a 1 year time base. The strong seasonal modulation of methanol is characterized by minimum values and variability in December to February and maximum columns in June–July. The methanol maximum in summer indicated by our results is consistent with the maximum observed for free tropospheric methanol above Kitt Peak (Rinsland et al., 2009) and the analysis of IASI tropospheric measurements over Europe (Razavi et al., 2011). The mean peak-to-peak amplitude of a seasonal cycle computed by Gardiner’s tool and expressed as a percentage of the corresponding CH₃OH yearly mean column amounts to $130.1 \pm 1.6\%$ ($2\text{-}\sigma$) while the seasonal modulation above Kitt Peak amounts to $64.6 \pm 0.1\%$ showing a similar amplitude with the IASI measurements (Razavi et al., 2011) for subtropical regions.

The IMAGESv2 model estimates a seasonal modulation of methanol in phase with the one we measured but underestimate the peak-to-peak amplitude with $88.6 \pm 1.3\%$ and $70.4 \pm 1.2\%$ for “IASI” and “MEGAN” respectively. The MEGAN emission fluxes being dependent on temperature, visible radiation fluxes, leaf area index and leaf age, they show a pronounced seasonal variation at mid-latitudes, with peak values in early summer. The IASI-derived emissions peak somewhat earlier than in the MEGAN inventory, a result consistent with modeling studies using TES methanol data (Wells et al., 2012; Cady-Pereira et al., 2012) as well as with other studies based on in situ concentration measurements (Jacob et al., 2005) or on flux measurements (Laffineur et al., 2012) which concluded to substantially higher methanol emission rates by young leaves compared to mature or senescent leaves.

No systematic bias is observed on the whole time series, but a seasonal bias is characterized (see Fig. 4): the maximum fractional difference $[(\text{IMAGES-FTIR})/((\text{IMAGES+FTIR})/2)]$ between monthly mean results from FTIR measurements and both “IASI” and “MEGAN” simulations is found to occur in July, with

Long-term evolution and seasonal modulation of methanol

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

⏪

⏩

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

**Long-term evolution
and seasonal
modulation of
methanol**

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

from FTS spectra (Dufour et al., 2007; Rinsland et al., 2009; Vigouroux et al., 2012), we have significantly improved the information content. With a typical DOFS of 1.82, a total column and two partial columns time series are available, i.e. a lower-tropospheric (LT, 3.58–7.18 km) and an upper tropospheric-lower stratospheric one (UTLS, 7.18–14.84 km). Both random and systematic error sources have been identified and characterized using the spectra recorded in the year 2010, and are found to be respectively 5 and 7 % for the total column.

The analysis of the time series does not reveal a significant long-term trend but shows a high peak-to-peak amplitude of the seasonal cycle of $129.4 \pm 5.5\%$ (2σ) for total columns. Methanol total and partial columns are characterized by a strong seasonal modulation with minimum values and variability in December to February and maximum columns in June–July. First analysis of methanol diurnal variation shows an increase of methanol in the morning and a decrease during the afternoon for all seasons but summer.

Comparisons with methanol measurements obtained with other techniques (in situ and satellite) give satisfactory results. Although the seasonal amplitude is larger in the FTIR lower tropospheric data compared to in situ measurements, a good agreement is generally found regarding the data dispersion. Concerning the UTLS partial columns, there is a close to statistical agreement with ACE-FTS occultations despite higher ACE columns of methanol in March and May.

The IMAGESv2 simulations underestimate the peak-to-peak amplitude for total and lower-tropospheric columns. Despite the absence of a systematic bias between our results and the IMAGESv2 simulations, comparisons show seasonal differences with an overestimation of winter methanol and an underestimation during summertime, which might be explained by an overestimation of the vertical gradient of methanol mixing ratios by the model. Regarding UTLS columns, the peak-to-peak amplitude and timing of the maximum (June–July) in both IMAGESv2 simulations are in very good agreement with the FTIR results.

**Long-term evolution
and seasonal
modulation of
methanol**

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



Even though the role of plant growth in methanol budget is confirmed by its season-
ality, large uncertainties remain on the methanol budget. Thanks to the improvement of
the information content of our retrieval and therefore our vertical resolution, our partial
column time series should contribute to better constraints for model simulations and
therefore may lead to a better understanding of methanol budget.

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**Long-term evolution
and seasonal
modulation of
methanol**

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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**Long-term evolution
and seasonal
modulation of
methanol**

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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- 30

**Long-term evolution
and seasonal
modulation of
methanol**

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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Long-term evolution and seasonal modulation of methanol

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

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**Long-term evolution
and seasonal
modulation of
methanol**

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

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**Long-term evolution
and seasonal
modulation of
methanol**

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

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Long-term evolution and seasonal modulation of methanol

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

⏪

⏩

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



Long-term evolution and seasonal modulation of methanol

W. Bader et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

Table 1. Error budget for total and all three partial columns. TC: total column, Tropo: tropospheric column, LT: lower tropospheric layer, UTLS: upper troposphere/lower stratosphere.

Error Sources	Max. Error (%)				Comments
	TC	Tropo	LT	UTLS	
Variability	46	50	57	48	
Systematic Errors (%)					
	TC	Tropo	LT	UTLS	
Line intensity CH ₃ OH	7.02	7.11	6.39	9.22	Xu et al. (2004)
Line Intensity interfering gases	1.00	1.73	3.96	0.91	Rothman et al. (2009)
ILS	0.41	0.33	1.19	2.39	±10% misalignment
Forward model	1	< 1	< 1	< 1	Retrieval algorithm-related
Total	7.17	7.39	7.68	9.62	
Random Errors (%)					
	TC	Tropo	LT	UTLS	
P-T profiles	1.2	2.3	11.3	8.6	From NCEP
SZA	0.2	0.4	3.1	1.4	0.2°
Smoothing	0.4	4.4	16.1	15.2	Barret et al. (2002)
Measurement noise	5.2	19.4	35.9	37.5	
Model parameters	0.7	0.6	0.5	1.2	
Total	5.37	20.04	40.18	41.43	
Relative Standard Deviation	6.60	8.34	22.59	21.11	

Long-term evolution and seasonal modulation of methanol

W. Bader et al.

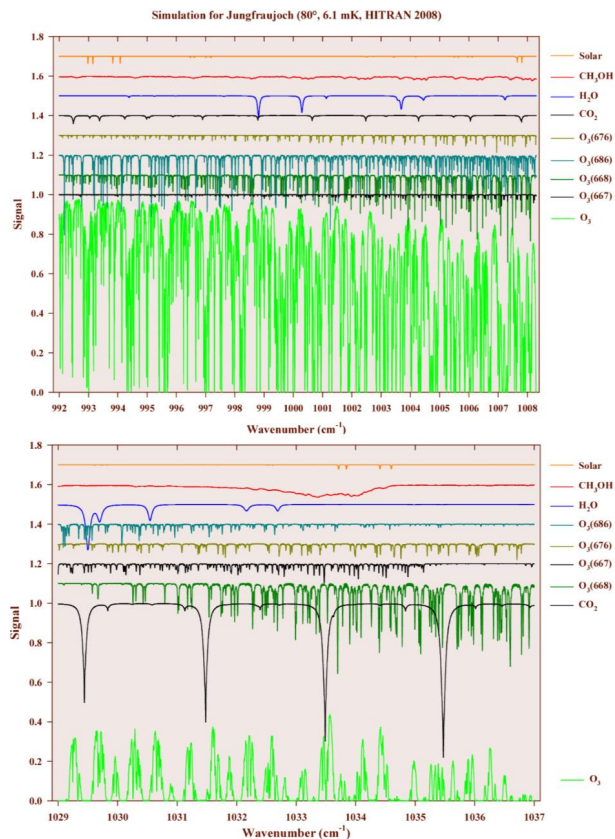


Fig. 1. Simulation for Jungfraujoch, 80° zenith angle, 6.1 mK. For both windows, we display the synthetic spectra for individual contributors (see color codes). HITRAN 2008 and averaged mixing ratio profiles based on the WACCM model climatology have been used for the simulations, except for CH₃OH for which our a priori was used (see text). For clarity, the contributions of each species have been vertically shifted.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

Long-term evolution and seasonal modulation of methanol

W. Bader et al.

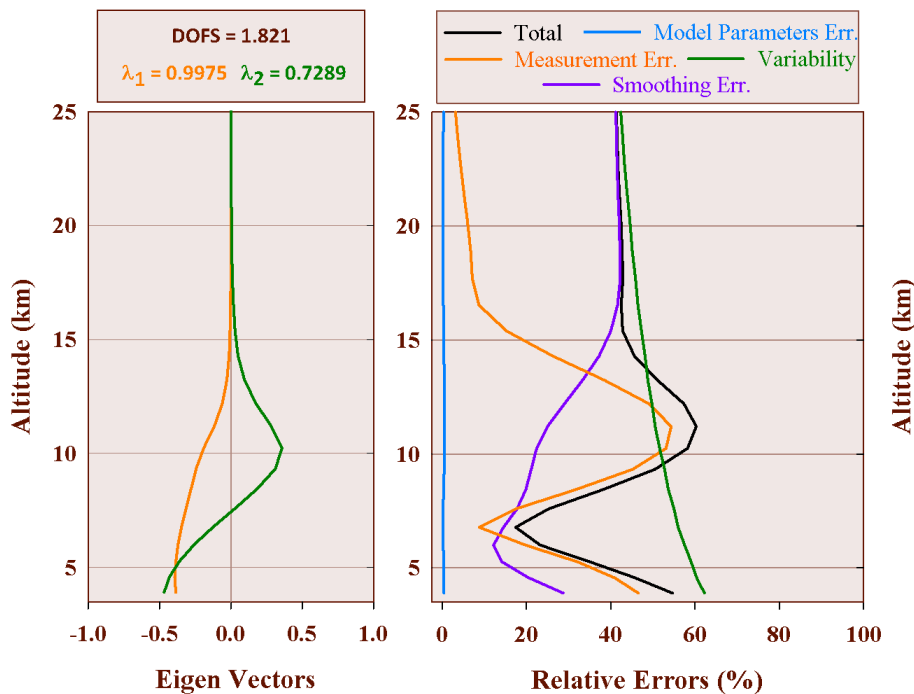


Fig. 2. Typical results for information content and error budget. Left frame: first eigen vectors and corresponding eigenvalues. Right frame: error budget, with identification of the main error components, together with the assumed variability (see color codes and Table 1 for additional information).

[Title Page](#)
[Abstract](#)
[Introduction](#)
[Conclusions](#)
[References](#)
[Tables](#)
[Figures](#)
[⏪](#)
[⏩](#)
[◀](#)
[▶](#)
[Back](#)
[Close](#)
[Full Screen / Esc](#)
[Printer-friendly Version](#)
[Interactive Discussion](#)

Long-term evolution and seasonal modulation of methanol

W. Bader et al.

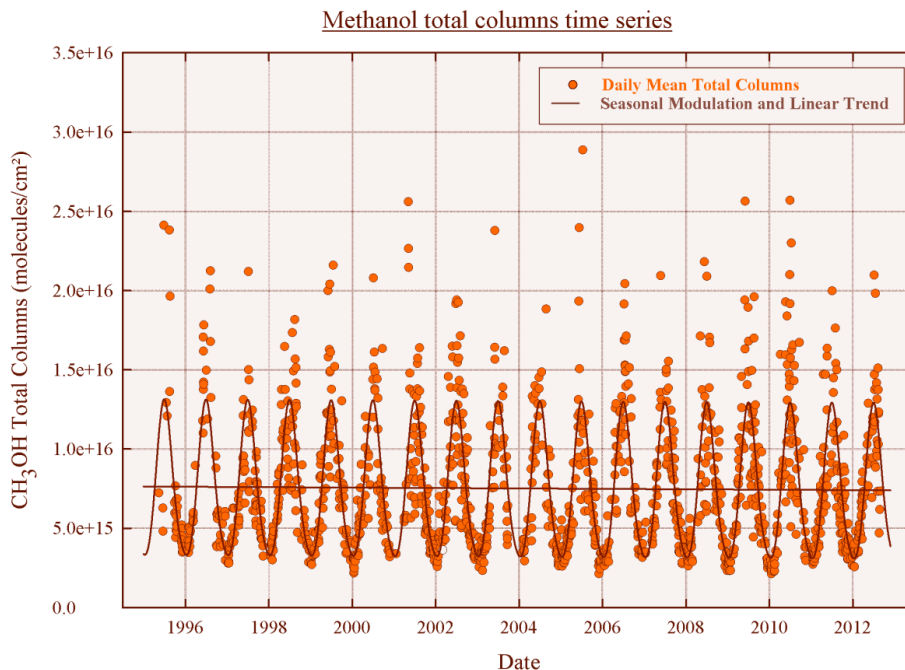


Fig. 3. Daily mean total (orange circles) column time series of CH₃OH above Jungfraujoch. Brown curves show the linear and seasonal trend components computed with the bootstrap resampling method (Gardiner et al., 2008).

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

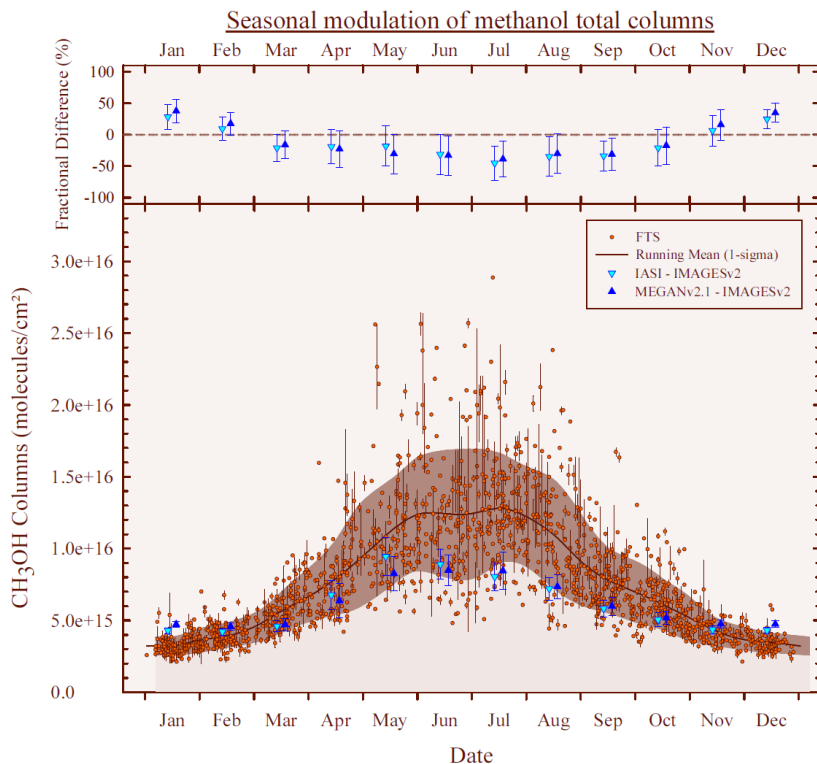


Fig. 4. Seasonal modulation of methanol total columns. Dots with vertical lines represent the daily mean total columns over a 1 year time base and their associated standard deviation. The brown curve corresponds to a running mean fit to all data points, with a 15 day step and a 2 month wide integration time. The area corresponds to the 1- σ standard deviation associated to the running mean curve. Up and down blue triangles are monthly means of the model IMAGESv2 simulations for MEGAN and IASI respectively. Upper frame shows monthly fractional difference between FTIR results and IMAGESv2 simulations.

**Long-term evolution
and seasonal
modulation of
methanol**

W. Bader et al.

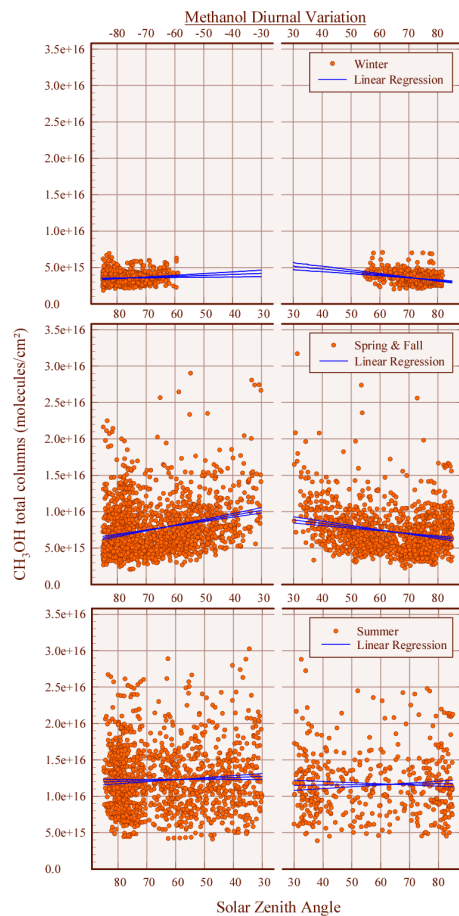


Fig. 5. Methanol diurnal variation. Total columns vs. the solar zenith angle for winter, summer and the rest of the year. Blue lines represent linear regressions.

Long-term evolution and seasonal modulation of methanol

W. Bader et al.

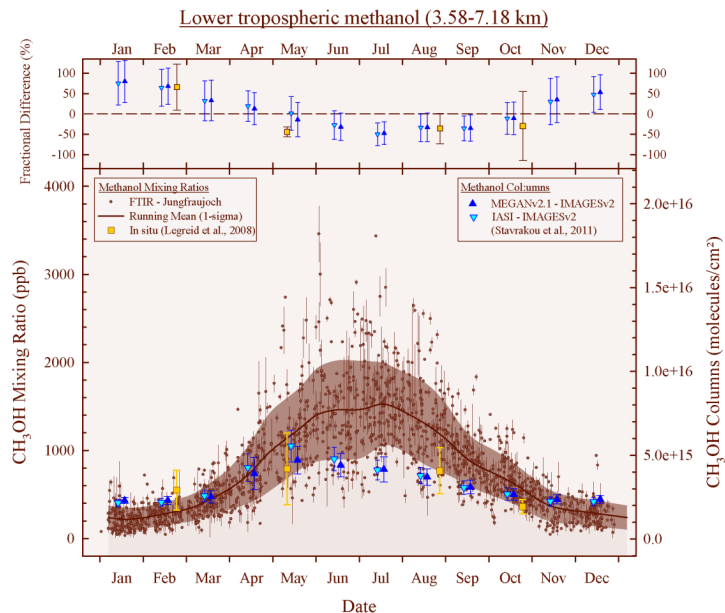


Fig. 6. Lower-tropospheric methanol (3.58–7.18 km). Dots with vertical lines represent the daily mean lower-tropospheric columns over a 1 year time base and their associated standard deviation. The brown curve corresponds to a running mean fit to all data points, with a 15 day step and a 2 month wide integration time. The area corresponds to the 1- σ standard deviation associated to the running mean curve. Up and down blue triangles are monthly means of the model IMAGESv2 simulations for MEGAN and IASI respectively (Stavrakou et al., 2011). Yellow squares are seasonal means of methanol in situ measurements (Legreid et al., 2008). The upper panel shows monthly fractional difference between FTIR results and IMAGESv2 simulations and seasonal fraction difference with in situ measurements.

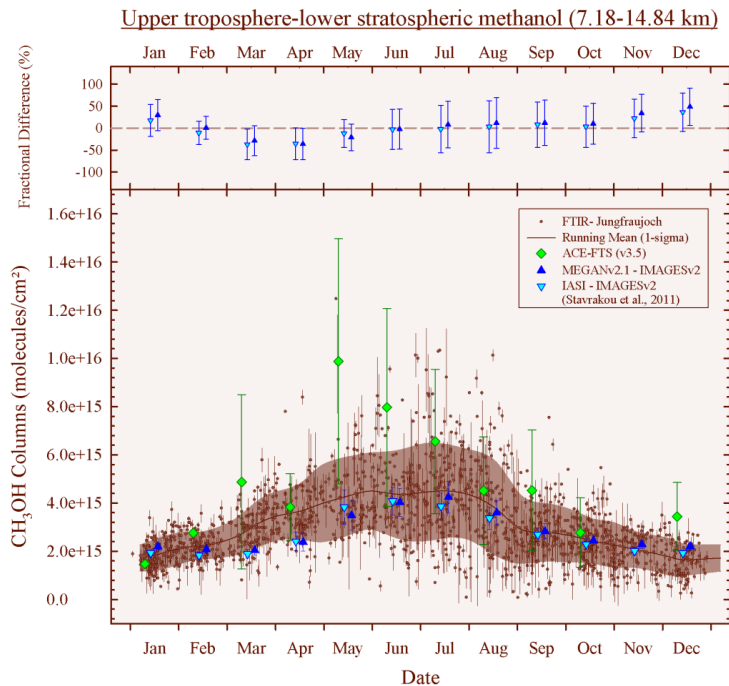


Fig. 7. Upper troposphere-lower stratospheric methanol (7.18–14.84 km). Dots with vertical lines representing daily mean lower-tropospheric columns over a 1 year time base and their associated standard deviation. The brown curve corresponds to a running mean fit to all data points, with a 15 day step and a 2 month wide integration time. The area corresponds to the 1- σ standard deviation associated to the running mean curve. Up and down blue triangles are monthly means of the model IMAGESv2 simulations for MEGAN and IASI respectively (Stavrakou et al., 2011). Green diamonds are monthly means of methanol retrieved from ACE-FTS occultations with the error bars representing the standard deviation (2- σ). Upper frame show monthly fractional difference between FTIR results and IMAGESv2 simulations and ACE-FTS results.