



Measurements of diurnal variations and Eddy Covariance fluxes of glyoxal

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Measurements of diurnal variations and Eddy Covariance (EC) fluxes of glyoxal in the tropical marine boundary layer: description of the Fast LED-CE-DOAS instrument

S. Coburn^{1,2}, I. Ortega^{1,2}, R. Thalman^{1,2,*}, B. Blomquist³, C. W. Fairall⁴, and R. Volkamer^{1,2}

¹Department of Chemistry and Biochemistry, University of Colorado, Boulder, CO, USA

²CIRES – Cooperative Institute for Research in Environmental Sciences, Boulder, CO, USA

³Department of Oceanography, University of Hawaii, Honolulu, HI, USA

⁴Earth System Research Laboratory, NOAA, Boulder, CO, USA

* now at: Brookhaven National Laboratory, Upton, NY, USA

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Correspondence to: R. Volkamer (rainer.volkamer@colorado.edu)

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during the day. After sunset the ocean was a source for glyoxal to the atmosphere (positive flux) in the SH; this primary ocean source was operative throughout the night. In the NH, the nighttime flux was positive only shortly after sunset, and negative during most of the night. Positive EC fluxes of soluble glyoxal over oceans indicate the presence of an ocean surface organic microlayer (SML), and locate a glyoxal source within the SML. The origin of atmospheric glyoxal, and possibly other oxygenated hydrocarbons over tropical oceans warrants further investigation.

1 Introduction

Eddy covariance (EC) fluxes are a well-established and widely used technique to measure surface-atmosphere gas exchange. The EC flux method provides insight into sources and sinks of atmospheric parameters (physical, chemical state variables) suitable to test our process level understanding (Baldocchi et al., 2001). EC fluxes are defined as the time average covariance between deviations from the mean of vertical wind velocity and deviations from the mean in the parameter of interest, e.g. here, the mixing ratio of a trace gas:

$$F_c = \overline{w'c'} = \int_0^{f_n} C_{wc}(f)df \quad (1)$$

where F is the flux, w' is the vertical wind velocity component, c' is the mixing ratio of the trace gas component, the prime denotes the instantaneous deviation from the mean, f_n is the Nyquist frequency of the measurements, and C_{wc} is the cospectrum.

A requirement of the EC flux technique is that measurements of both vertical wind velocities and the trace gas of interest are performed at high sampling frequencies, f , (typically a minimum of several Hz), sufficient to capture a majority of those frequencies that contribute to the overall flux. Balancing this requirement with preserving sufficient

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sensitivity in the measurements is one of the major challenges with developing chemical sensors suitable for EC flux applications. For mobile deployments, a portable and robust sensor is needed. Further, additional measurements of platform motion need to be performed, and corrections on the wind velocity data are needed. A description of the system deployed in this study and the method of correction is described by Fairall et al. (1997) and Edson et al. (1998), respectively. A particular challenge arises for EC flux measurements of short-lived species in the marine boundary layer (MBL), for which concentrations often do not exceed 10s to 100s of parts per trillion (1 pptv = 10^{-12} volume mixing ratio (VMR) = 2.46×10^7 molec cm $^{-3}$ at 298 K temperature and 1013 mbar pressure). As a result of these challenges, ship based EC flux measurements have today only been reported for the seven trace molecules: dimethyl sulfide (DMS) (Huebert et al., 2004; Blomquist et al., 2006, 2010; Marandino et al., 2007, 2008, 2009; Miller et al., 2009; Edson et al., 2011; Bell et al., 2013), Carbon dioxide (CO $_2$) (Fairall et al., 2000; McGillis et al., 2001, 2004; Kondo and Tsukamoto 2007; Miller et al., 2009, 2010; Taddei et al., 2009; Norman et al., 2012), Ozone (O $_3$) (Bariteau et al., 2010; Helmig et al., 2012), carbon monoxide (CO) (Blomquist et al., 2012), acetone (Marandino et al., 2005; Taddei et al., 2009; Yang et al., 2014), acetaldehyde (Yang et al., 2014), and methanol (Yang et al., 2013). Table 1 lists typical concentrations for these molecules in the MBL, and compares them with glyoxal in terms of their Henry's Law constants (K_H , at 298 K), and typical atmospheric lifetimes. Notably, glyoxal is the molecule with the shortest atmospheric lifetime, and is present in the lowest abundance. The short lifetime of glyoxal limits the spatial scale over which it can be transported in the atmosphere to few 10 km. Further, glyoxal is the most soluble molecule in Table 1, i.e. its Henry's Law constant is 2000, 13 860, and 30 000 times larger than that of the other oxygenated hydrocarbons (OVOC) methanol, acetone and acetaldehyde, respectively. The differences in the physical and chemical properties have fundamental implications for the air-sea exchange of glyoxal. For example, while it is possible to supersaturate the surface ocean with acetaldehyde (Zhou and Mopper, 1990; Kieber et al., 1990; Millet et al., 2010; Yang et al., 2014) it is impossible

is present. The water measurement was used to monitor ambient conditions, NO₂ was used as a tracer for sampling the ship stack plume, and the O₄ measurement was used to correct the DOAS data for sampling time lag and inlet characterization (discussed in Sects. 3.1.1 and 3.1.2).

The primary measurement of the DOAS technique is Slant Column Density (SCD) which is the integrated concentration of the measured species along all light paths. It is easily converted using Lambert–Beer’s Law to concentration if the light path length within the cavity is known. Two different methods were utilized to experimentally determine the cavity light path: (1) comparison of measured O₄ SCDs and the calculated concentration of O₄ within the cavity; and (2) using the ratio of the signal measured in two different pure gases whose Rayleigh scattering cross-sections are well known (Thalman and Volkamer, 2010). For this study, method 2 was employed and N₂ and He were used for this process (referred to as mirror curves from this point forward). Mirror curves were taken on a near daily basis which enabled the continuous monitoring of the cavity performance. Additionally, an inherent consistency check exists from the comparison of O₄ SCD measurements with those calculated from the mirror curve, the Rayleigh scattering cross section of air, and known temperature and pressure (Thalman and Volkamer, 2010). For the duration of the cruise, the peak mirror reflectivity was maintained between 99.9967–99.9973 %, translating into routine cavity path lengths of 18–20 km at 455 nm.

In order to accelerate the data acquisition of the instrument to rates sufficient to accommodate EC fluxes, software was developed to simultaneously eliminate shutter movements and decrease readout time (through binning of CCD rows). The final instrument measurement frequency of ~ 2 Hz strikes a balance between time resolution, and the duty cycle dedicated to collecting photons (as compared to read-out time of the CCD). The measurement detection limit with CE-DOAS measurements is typically photon shot-noise limited. We assess the instrument performance by investigating the root mean square (RMS) of the optical density of the residual remaining after the non-linear least squares fitting routine, and comparing it with the theoretical photon shot

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noise (Coburn et al., 2011). Individual spectra were summed, and analyzed to improve the signal to noise ratio of the measurements. In an ideal instrument (i.e., completely limited by photo shot noise), the RMS of the fitting routine should follow photon counting statistics, where the theoretical RMS is inversely proportional to the square root of the number of photons collected.

$$\text{RMS} \equiv \frac{1}{\sqrt{N}} \quad (2)$$

where N is the number of photons collected.

The measured RMS of the Fast LED-CE-DOAS instrument field deployment is compared to the theoretical RMS, and plotted as a function of the number of photons in Fig. 3. The grey points are raw data at different levels of averaging and the colored squares represent the median values for each set: light blue is the raw 400 ms data; dark blue is the sum of 5 spectra (~ 2 s); purple is the sum of 20 spectra (~ 8 s); red is the sum of 100 spectra (~ 40 s); and the green is the sum of 1000 spectra (~ 8 min). As can be seen, the RMS during this campaign fairly closely follows counting statistics for the measured spectra, as well as for different levels of binning. Shown on the right axis is the corresponding 1σ precision for glyoxal.

Appropriate quality assurance filters were applied to the raw CE-DOAS measurements prior to calculating glyoxal fluxes in order to exclude the use of any stack contamination, or otherwise questionable data. These filters removed periods of elevated NO_2 (contaminated by the ship stack plume: values greater than ~ 30 pptv), instability in the cavity (O_4 and internal cavity pressure measurements: acceptable pressure range 470–500 torr), and any spectra where the DOAS fitting resulted in RMS values larger than 5×10^{-3} .

2.2 TORERO field campaign

While the cruise started on 25 January 2012, only data taken 2–28 February 2012 will be considered for this study. The inlet for the cavity was mounted near the top of a 10 m

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jackstaff (18 m a.s.l.) on the bow along with the inlets for the CO₂ flux system (Blomquist et al., 2014) and the in situ O₃ monitor, the sonic anemometer, and a motion system. The sampling line between the inlet and the instrument was ~ 65 m long, and consisted of 3/8" ID coated aluminum tubing (Eaton SynFlex Type 1300). Additionally, an aerosol filter (changed every other day) was included after the inlet in order to prevent collection of sea salt in the sampling line, and keep the air reaching the CE-DOAS system aerosol free. The filter was regularly changed to avoid aerosol accumulation, and experiments of glyoxal transfer through Teflon filters showed no visible attenuation (Thalman and Volkamer, 2010). Previous experiments to characterize the effect of In order to maintain turbulent flow throughout the sampling line, a high flow pump maintained a flow of ~ 120 L min⁻¹ (Lenschow and Raupach, 1991). From this main flow, a sample flow of ~ 9 L min⁻¹ was pulled through the cavity. These flow conditions resulted in an operating pressure inside the cavity of ~ 470–500 torr. This sub-ambient cavity pressure had to be actively addressed due to the sensitivity of optical cavities to fluctuations in pressure (which can de-align the mirrors). This was accomplished by the addition of stabilizing mounts for the mirrors to prevent movement during measurements. Figure 4 contains a plumbing diagram for the CE-DOAS system with arrows indicating the direction of air flow at various points along the sampling line. Two pumps and three Mass Flow Controllers (MFCs) were used in this system, the main flow through the sampling line was set at ~ 120 L min⁻¹ (controlled by MFC 1), the smaller sample flow through the cavity was set at ~ 9 L min⁻¹ (controlled by MFC 2), and the calibration gases for the Fast-LED-CE-DOAS system (used for monitoring cavity performance and determining cavity path length) were controlled by MFC 3. Photographs of the inlet, operational cavity, and instrument rack containing all controlling electronics and spectrometer can be found in the Supplement Fig. S1.

3 Results

3.1 Instrument characterization

The following sections will describe the characterization of instrument properties pertinent to the measurement of fluxes via the EC technique.

3.1.1 Phase correction (N₂ pulse)

Wind sensor data was collected at 10 Hz and in order to calculate the glyoxal fluxes the CE-DOAS measurements needed to be synchronized to this data. Rather than degrading the high resolution wind data, the CE-DOAS measurements were first interpolated from 2 Hz to 10 Hz. Since the trace gas is drawn through an inlet, there is a finite time difference between the instantaneous wind velocity measurements and those of the trace gas measurements. The flux system deployed here includes a method for experimentally determining this correction. The method is described in detail in Bariteau et al. (2010), so only a brief overview will be given here: the inlet is equipped with a fast-switching solenoid valve that injects pure nitrogen (supplied from a compressed air cylinder) into the sample flow. The valve is triggered for 3–5 s at the beginning of every hour and the signal used as the trigger is recorded on the same timestamp as the anemometer. This data is used in conjunction with the accompanied drop in the trace gas signal (recorded on a different timestamp) to continuously monitor, and apply a correction to the time stamps prior to correlating both sensors. In the cavity, the measurement of O₄ was used for this correction. Figure 5 contains a plot showing an example of the corrected O₄ signal overlaid on the nitrogen pulse signal (black trace), also shown is the fit of the step response function from method 1 (blue trace, see below). The raw O₄ measurements are shown as black circles, and the interpolated data are the smaller red circles. Two methods were used to determine the phase correction based on the drop in the O₄ signal: (1) fitting of a first order step response function; (2) manual determination. Method 2 involved using O₄ data averages to identify when

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the N_2 was attenuating the O_4 signal, and from there determining the time at which the signal actually started dropping. Each analysis was performed on hourly data files; 626 files were analyzed and 50 of these files did not meet basic criteria to enable the pulse matching and so were rejected; the total number of usable hours for the flux data was 576. The average difference found between the two phase correction methods was 0.11 s and the statistics associated with each analysis can be found in Table 2.

3.1.2 Response time

The pulse of nitrogen described in the previous section was also used to characterize the response time of the instrument. Introducing pure N_2 gas into the sample flow created a drop in the O_4 signal which was exploited to determine the response time of the instrument. The same two methods employed for the phase correction were used to calculate the instrument response time, which also gave an average difference between methods of 0.11 s (statistics in Table 2). The instrument response is best determined experimentally, since high frequency flux attenuations can be caused by drawing the sample through the aerosol filter and long sampling line. Here, a low-pass filter function was chosen to represent the attenuation.

$$H(f) = \frac{1}{1 + (2\pi f \tau_c)^2} \quad (3)$$

where τ_c is the instrument response time.

Using the measured response time and the filter function, the instrument cut-off frequency (f_c) (the frequency at which the signal fluctuations drop by $1/\sqrt{2}$) was calculated, which corresponds to a drop in the signal to 0.5.

$$f_c = \frac{1}{2\pi\tau_c} \quad (4)$$

Using the average values of the response time of 0.283 s and 0.282 s for the first-order step response function and the manual determination, respectively, the calculated cut-off frequency is 0.56 Hz. The application of the filter function for this system and the

effect of the response time on the high frequency attenuation will be discussed in Sect. 3.3.1. These small differences in response time determined from the two methods add certainty about the correction of the flux measurements, as is discussed in more detail in Sect. 3.3.3.

3.1.3 Fast measurements

The variance spectra for glyoxal as a function of frequency for a 6 h time period on 4 February 2012 from 15:00–21:00 UTC are shown in Fig. 6. Data from both 10 min (purple) and 30 min (light blue) averaging periods (see Sect. 3.3.2) are included in this plot.

The constant variance per Hz in the frequency range sampled by the instrument demonstrates that the Fast-LED-CE-DOAS system is indeed a white noise sensor. The horizontal black line represents the integral of the data in the frequency range 6×10^{-4} to 1 Hz of $\sim 1585 \text{ pptv}^2$. The solid vertical black line depicts the cut-off frequency of the instrument calculated from the average response time of the instrument, and the dashed vertical black lines represent ± 1 standard deviation of this data.

3.2 Diurnal cycle measurements

Analyzing data created from summing 1000 spectra (~ 8 min total integration time) enabled the measurement of a diurnal cycle of glyoxal between 2–28 February 2012. A time series of these measurements can be found in Fig. 7 (top panel, left axis). Summing 1000 spectra allowed the realization of an average RMS value of $(1.0 \pm 0.1) \times 10^{-4}$, which translates into an average detection limit of 5.9 pptv; lower detection limits are possible from further averaging of the data. Included in Fig. 7 are time traces for in situ O_3 , solar zenith angle (SZA), NO_2 , RH, ambient air temperature, ambient pressure, wind speed (from the sonic anemometer), and a flag indicating periods that were suitable for EC fluxes.

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3.3 Ambient flux measurements

3.3.1 Signal attenuation

As introduced in Sect. 3.1.2, a low-pass filter function was used to assess the high frequency flux attenuation due to the aerosol filter and sampling line length; Eq. (1) can be re-written as Eq. (5):

$$F_c = \overline{w'c'_m} = \int_0^{f_n} C_{wc}(f)[H(f)]^{\frac{1}{2}} df = \int_0^{f_n} C_{wcm}(f) df \quad (5)$$

where the subscript m represents the measured values, see Eq. (1) for other variables (note that the square root appears in the modified equation because only the signal of the trace gas is attenuated).

This relationship can then be used to assess the effect of attenuation on the overall flux by applying the filter function using the Kaimal model neutral-stability cospectrum (Kaimal et al., 1972), derived via Eqs. (6a) and (6b)

$$\frac{fC_{wc}(f)}{F_c} = \frac{11n}{(1 + 13.3n)^{1.75}}, \quad n \leq 1.0 \quad (6a)$$

$$\frac{fC_{wc}(f)}{F_c} = \frac{11n}{(1 + 3.8n)^{2.4}}, \quad n \geq 1.0 \quad (6b)$$

where the surface normalized frequency $n = fz/\overline{u_r}$, z is the measurement height, and $\overline{u_r}$ is the average relative wind speed. Using the calculated “true” and “measured”

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proved to be noisy, further binning of the 30 min flux measurements reveals trends in the data. Figure 8 contains example cospectra from this data, where the average of all data (green trace) scatters around zero. The examples of both positive (red trace) and negative (blue trace) cospectra were created by binning data: from 16 February 21:15–17 February 00:45 LT (positive cospectrum); and from 2 February 14:45–15:45 LT (negative cospectrum).

3.3.3 Error sources

The potential sources of error in this data are: (1) inaccuracies in determining the phase shift of the CE-DOAS measurements; (2) high frequency signal loss due to sampling line attenuation; and (3) uncertainty surrounding the noisy raw glyoxal measurements. Phase shift determination was deemed to be rather robust (through the comparison of the values determined using the two different methods), and any small inaccuracies would have negligible effect on the flux data (as assessed by comparing the results from the 4 methods previously mentioned). The high frequency flux loss due to signal attenuation was calculated as being, at most, 10% from the characterization in the instrument response time. Based on the cospectra (Fig. 8), it seems that glyoxal efficiently transferred through the sample lines and using the O_4 measurements to characterize sample transfer gives reasonably good agreement.

4 Discussion and conclusions

The Fast-LED-CE-DOAS instrument is a multispectral sensor suitable to measure eddy covariance (EC) fluxes of glyoxal in the remote marine boundary layer (MBL). The measurement frequency of ~ 2 Hz is sufficient to capture ~ 90 % of the glyoxal flux. Inlet and sampling line attenuation was determined using the measured response time of the instrument (0.28 ± 0.14 s, based on O_4 measurements) and accounts for a correction of < 10 %. Multiple gases are selectively detected simultaneously with glyoxal, and are

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glyoxal inside the MBL to explain satellite VCDs; this is particularly true over the NH tropical Eastern Pacific Ocean (by a factor 2 to 6). Furthermore, our in situ data show glyoxal is more abundant in the SH tropical MBL. By contrast, both satellites find $\sim 25\text{--}42\%$ lower glyoxal VCDs in the SH compared to the NH. The campaign average VMR during mornings in the SH (47 ± 7 pptv glyoxal) corresponds to $\sim 1.2 \times 10^{14}$ molec cm^{-2} glyoxal VCD over the SH cruise segment, which is 2–3 times lower than long-time average VCD observed from space. The reason for this apparent mismatch in glyoxal amounts, and reversed hemispheric gradient is currently not understood. A particularly interesting development to investigate the diurnal variation of glyoxal over oceans consists in the Tropospheric Emissions: Monitoring of Pollution (TEMPO) satellite mission (planned to launch in 2019), which will provide first time-resolved glyoxal VCD observations from geostationary orbit. Our diurnal profiles show further that glyoxal concentrations change by 30% over the course of the day. With the caveat that changes in MBL VMRs may not be indicative of VCD changes, this also implies that $\sim 15\%$ lower VCDs are expected at the time of the OMI satellite overpass (13:45 LT at equator). The differences between satellite and in situ measurements are as of yet difficult to reconcile. Notably, a direct comparison between in situ observations and column data is complicated due to the lack of vertical profile measurements of glyoxal at tropical latitudes, different averaging times, and spatial scales probed by in situ and column observations, as well as uncertain assumptions about a priori profiles, cloud screening, and other factors that influence air mass factor calculations. Some of these factors were further investigated from collocated measurements of glyoxal from aircraft during the TORERO project. The fast in situ LED-CE-DOAS instrument holds great potential for future deployments on research aircraft.

The diurnal variations in our glyoxal flux measurements qualitatively reflect the variations in the VMRs seen in Fig. 9. The maximum fluxes are seen at night (SH: $(5.3 \pm 3.3) \times 10^{-2}$ pptv m s^{-1} ; NH: $(2.3 \pm 3.1) \times 10^{-2}$ pptv m s^{-1}) and minimum fluxes during the daytime (SH: $(-1.6 \pm 3.8) \times 10^{-2}$ pptv m s^{-1} ; NH: $(-5.6 \pm 4.1) \times 10^{-2}$ pptv m s^{-1}). All nighttime fluxes in the SH are significantly greater

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Table 1. Overview of Eddy Covariance flux measurements from ships.

Molecule	MBL Concentration (pptv)	K_H ($M \text{ atm}^{-1}$)	Lifetime* (days)	Reference flux measurement in MBL
CO ₂	380–400 ($\times 10^6$)	0.035	$> 3 \times 10^5$	Fairall et al. (2000)
CO	60–150 ($\times 10^3$)	1×10^{-3}	16	Blomquist et al. (2012)
Acetone	700–900	30.3	10	Yang et al. (2014)
O ₃	10–30 ($\times 10^3$)	0.011	6	Bariteau et al. (2010)
Methanol	300–900	222	4	Marandino et al. (2005)
DMS	20–1500	0.485	0.8	Hubert et al. (2004)
Acetaldehyde	200–300	14.1	0.2	Yang et al. (2014)
Glyoxal	25–80	4.2×10^5	9×10^{-2}	This work

* Lifetimes calculated against reaction with OH (assuming $[\text{OH}] = 3 \times 10^6 \text{ molec cm}^{-3}$), and photolysis rates calculated for aerosol free, noon time at equator conditions.

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Table 2. The average phase correction and time response of the Fast-LED-CE-DOAS instrument (with standard deviation) for the two different methods employed in this study. See text for details.

	Method 1		Method 2	
	Phase Correction (s)	Time Response (s)	Phase Correction (s)	Time Response (s)
Average	−2.54	0.28	−2.61	0.28
Standard Deviation	0.23	0.16	0.24	0.14

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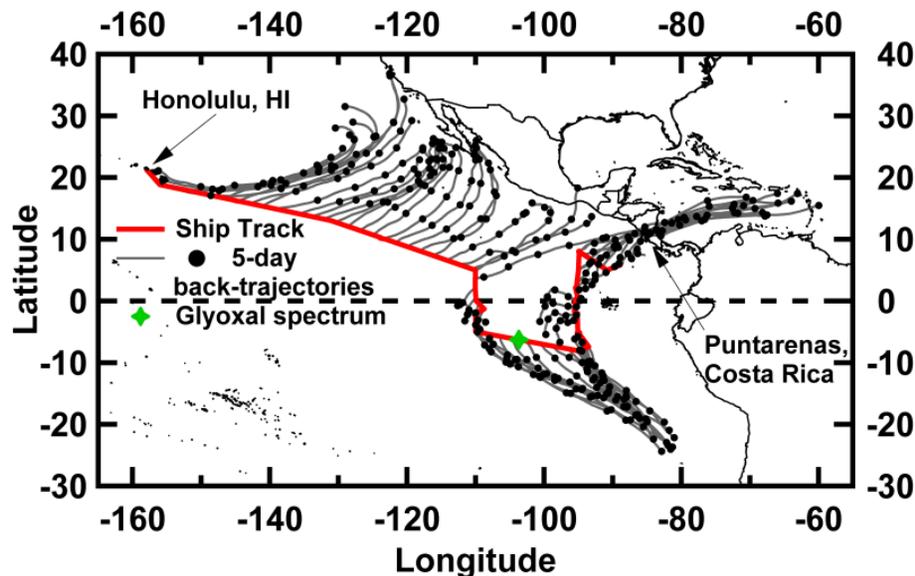


Figure 1. Cruise track of the NOAA RV *Ka'imimoana* during the TORERO 2012 field experiment (red trace). The ship set sail from Honolulu, HI on 25 January 2012 and made final port in Puntarenas, Costa Rica on 1 March 2012 (37 days at sea). Shown along the ship track are HYSPLIT 5-day back trajectories (initiated at 00:00 and 12:00 LT every day; solid grey lines). The black circles along the trajectories are spaced by 1 day. Air sampled in the Northern Hemisphere had been over the ocean for at least 2 days prior to reaching the ship, and often did not experience land influences for at least 5 days. Air sampled in the Southern Hemisphere had been over the ocean for more than 5 days without obvious land/pollution influences. The location of the example glyoxal spectrum is marked by the green star.

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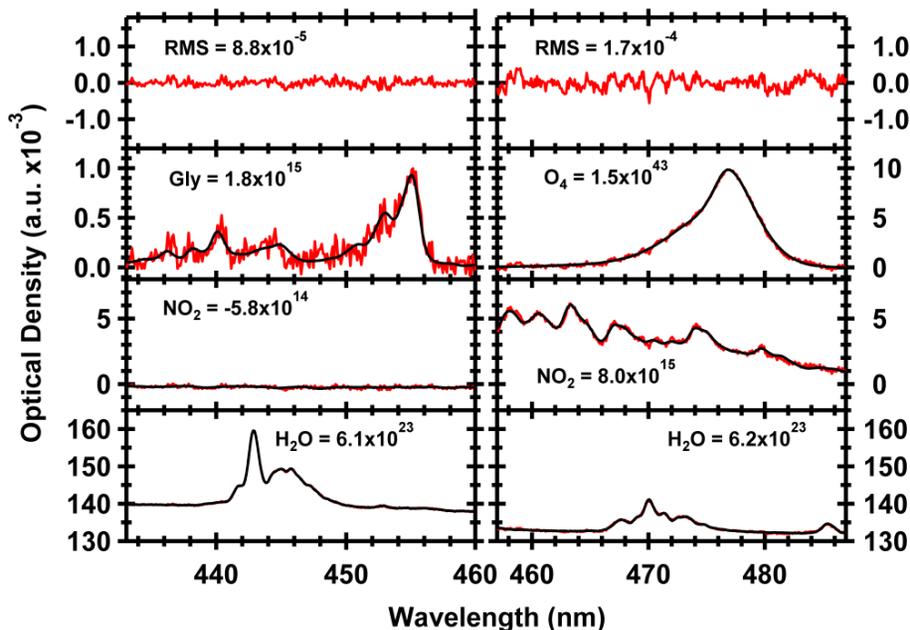


Figure 2. Example spectra of molecules measured by the Fast LED-CE-DOAS instrument. The DOAS fits are shown for glyoxal (left panel, 433–460 nm fit window), O_4 and NO_2 (right panel, 457–487 nm), and water vapor (both windows). The RMS residual for each fit is shown in the top row. The spectra shown here were recorded on 14 February 2012 at $\sim 06:20$ LT (glyoxal), and 11 February 2012 at $\sim 13:40$ LT (O_4).

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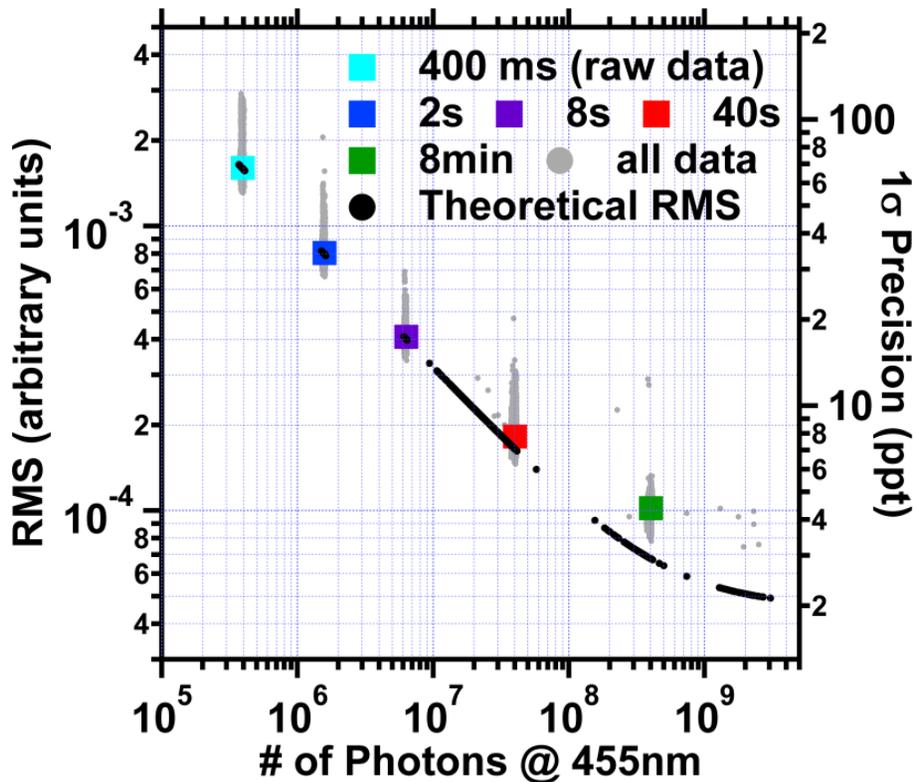


Figure 3. Fast LED-CE-DOAS instrument performance: sensitivity. The residual noise (RMS) from the DOAS analysis is shown as a function of the number of photons corresponding to different averaging of the data. Grey points represent all data, while colored squares represent their respective mean value; black circles represent the theoretical RMS value determined from photon counting statistics (Coburn et al., 2011). The corresponding 1σ precision of glyoxal is plotted on the right axis.

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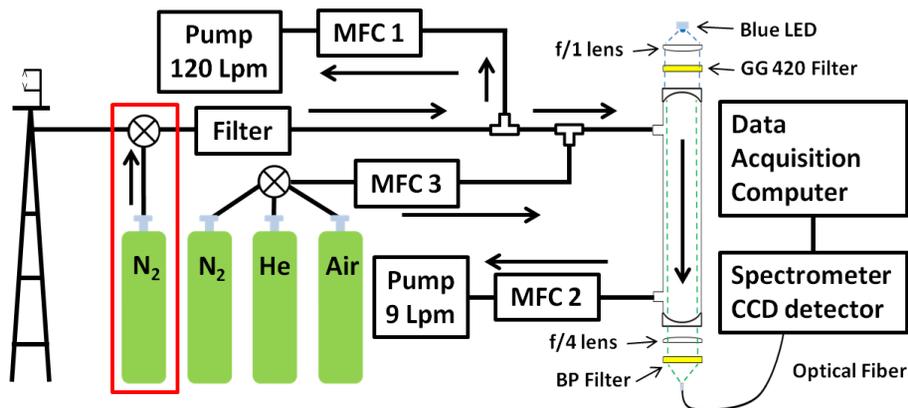


Figure 4. Sketch of the Fast-LED-CE-DOAS setup and plumbing diagram for sampling during TORERO 2012. The N_2 “puff” system is indicated by the red box. Arrows show the direction of flow through various portions of the system. Photographs of this set up can be found in Supplement Fig. S1.

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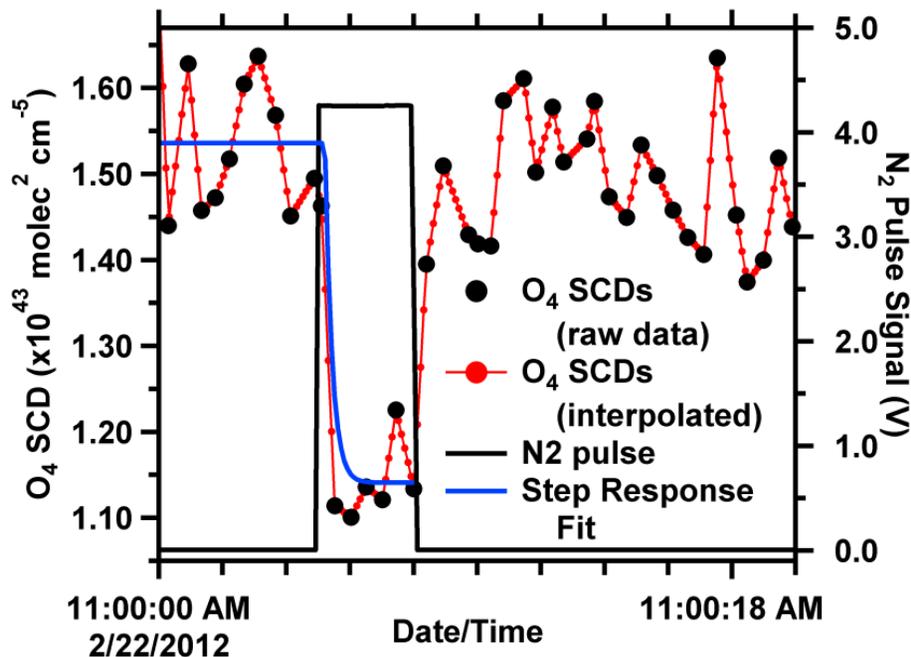


Figure 5. Illustration of the phase-correction and time response using O_4 . Individual CE-DOAS O_4 measurements (black dots) were interpolated onto the timestamp of the wind sensor (red dots). The N_2 pulse signal (solid black line) is visible as the drop in O_4 SCDs; the data has already been time-shifted to match this N_2 trigger. Also shown is the fit of a step response function (solid blue line) to the drop in the O_4 signal, from which an instrument time response can be determined.

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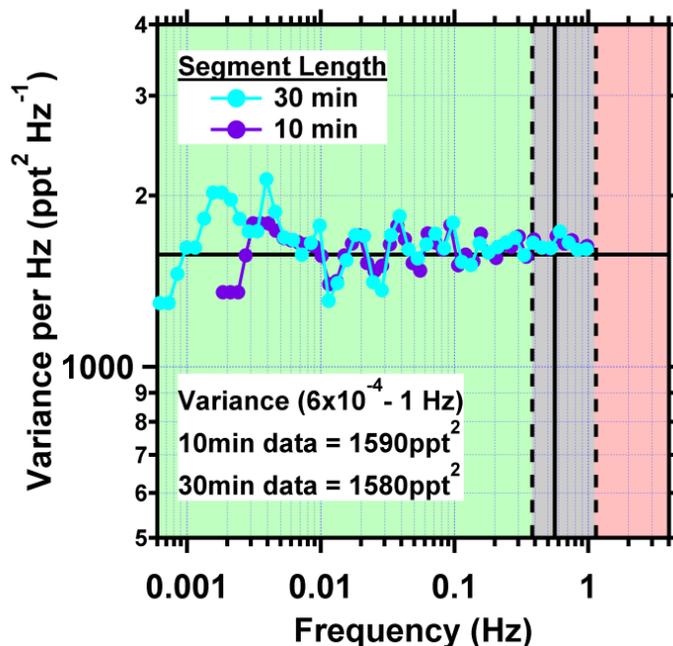


Figure 6. Fast LED-CE-DOAS instrument performance: frequency response for glyoxal. The glyoxal variance distribution per frequency bin is for a 1 h section of data (representative of the global average) on 22 February 2012 from 04:00–05:00 LT. The horizontal line represents the integral variance ($\sim 1600 \text{ ppt}^2 \text{ Hz}^{-1}$) at frequencies measured by the instrument (green shading). The solid vertical line represents the cut-off frequency determined from the average time response of the instrument; the dashed vertical lines represent the standard deviation of the time response data (grey background shading); higher frequencies were not measured by our setup (red shading). The Nyquist frequency of our setup is 1 Hz.

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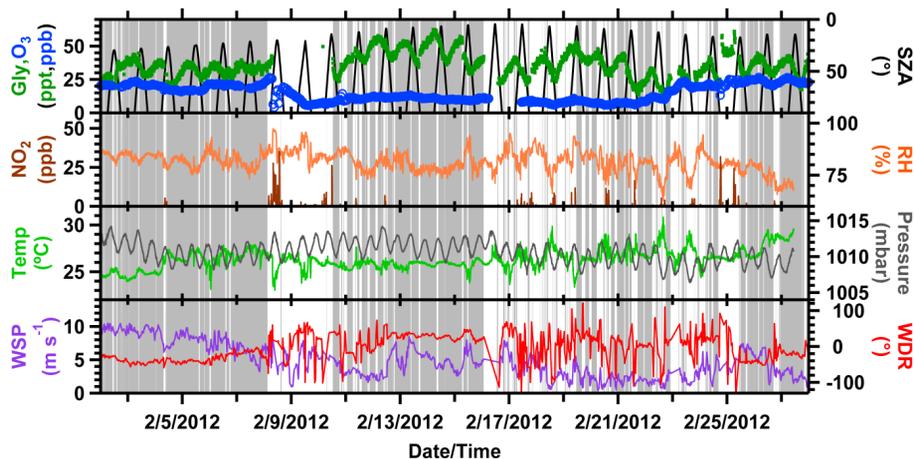


Figure 7. Time series of glyoxal, O_3 and NO_2 , as well as meteorological parameters. Grey shaded background represents times suitable for flux calculations; filters included NO_2 , cavity pressure, wind direction, wind direction standard deviation, ship heading range, $dG/y/dt$, and the horizontal glyoxal flux components. See text and Supplement Fig. S2 for details.

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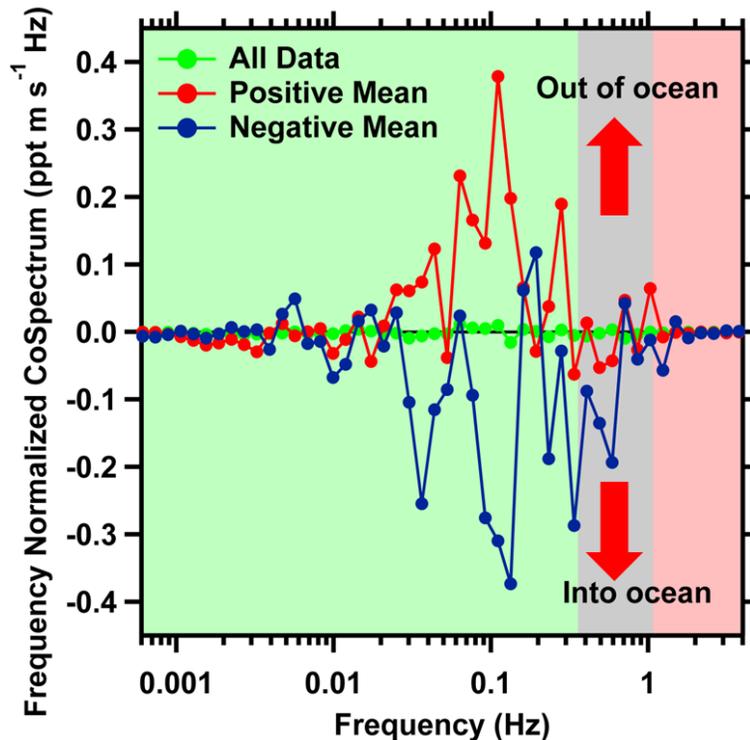


Figure 8. Cospectra of glyoxal and vertical wind from the flux calculations. The green trace represents the average of all cospectra that passed the quality assurance filters. The positive cospectrum (red) represents data averaged over a period of ~ 3.5 h from 16 February 2012 21:15LT to 17 February 2012 00:45LT. The negative cospectrum (blue) represents a ~ 1 h average on 2 February 2012 from 14:45–15:45LT. The background color shading is identical to that in Fig. 6.

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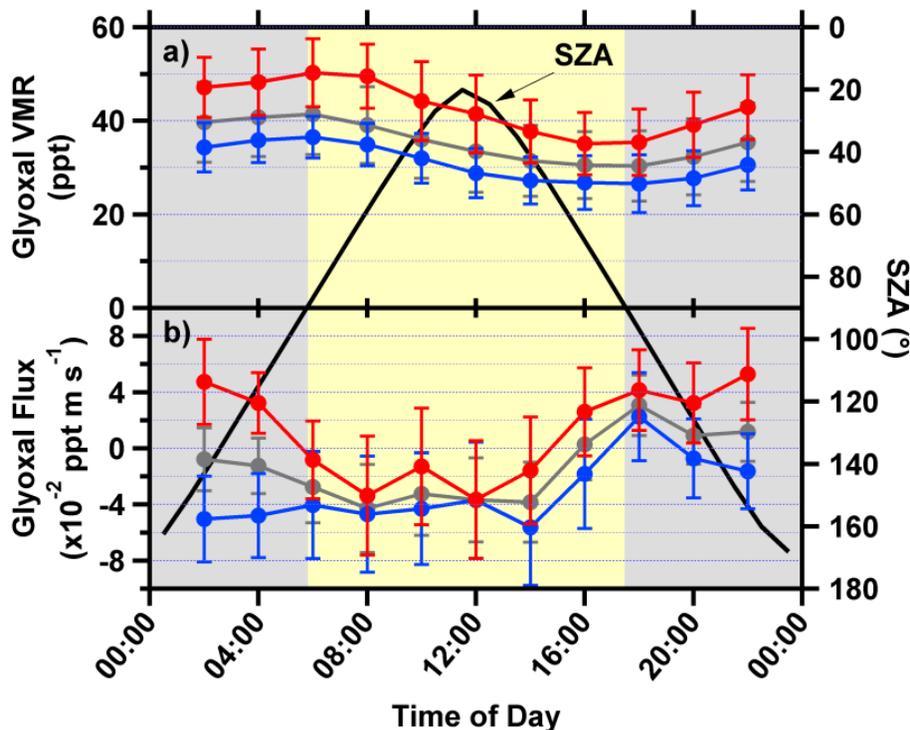


Figure 9. Diurnal variation in the glyoxal mixing ratio (a) and the glyoxal flux (b) in the Northern (blue) and Southern Hemisphere (red). Only data that qualifies for flux calculations has been averaged. Yellow shading indicates daytime, while grey indicates nighttime; the SZA is also shown on the right axis.

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