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# A two-channel, tunable diode laser-based hygrometer for measurement of water vapor and cirrus cloud ice water content in the upper troposphere and lower stratosphere

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# Abstract

The recently developed NOAA Water instrument is a two-channel, closed-path, tunable diode laser absorption spectrometer designed for the measurement of water vapor and enhanced total water (vapor + inertially enhanced condensed-phase) in the upper

- <sup>5</sup> troposphere/lower stratosphere from the NASA Global Hawk unmanned aircraft system (UAS) or other high-altitude research aircraft. The instrument utilizes wavelengthmodulated spectroscopy with second harmonic detection near 2694 nm to achieve high precision with a 79 cm double-pass optical path. The detection cells are operated under constant temperature, pressure and flow conditions to maintain a constant sensitivity to
- <sup>10</sup> H<sub>2</sub>O independent of the ambient sampling environment. An on-board calibration system is used to perform periodic in situ calibrations to verify the stability of the instrument sensitivity during flight. For the water vapor channel, ambient air is sampled perpendicular to the flow past the aircraft in order to reject cloud particles, while the total water channel uses a heated, forward-facing inlet to sample both water vapor and cloud
- <sup>15</sup> particles. The total water inlet operates subisokinetically, thereby inertially enhancing cloud particle number in the sample flow and affording increased cloud water content sensitivity. The NOAA Water instrument was flown for the first time during the second deployment of the Airborne Tropical TRopopause EXperiment (ATTREX) in February– March 2013 on board the Global Hawk UAS. The instrument demonstrated a typical
- <sup>20</sup> in-flight precision (1 s, 1 $\sigma$ ) of better than 0.17 parts per million (ppm, 10<sup>-6</sup> mol mol<sup>-1</sup>), with an overall H<sub>2</sub>O vapor measurement uncertainty of 5 % ± 0.23 ppm. The inertial enhancement for cirrus cloud particle sampling under ATTREX flight conditions ranged from 33–48 for ice particles larger than 8 µm in diameter, depending primarily on aircraft altitude. The resulting ice water content detection limit (2 $\sigma$ ) was 0.023–0.013 ppm, cor-<sup>25</sup> responding to approximately 2 µg m<sup>-3</sup>, with an estimated overall uncertainty of 20 %.



# 1 Introduction

Water in the upper troposphere and lower stratosphere (UT/LS) plays an important role in Earth's climate system through aspects of radiative transfer, cirrus cloud formation and stratospheric ozone chemistry (Kirk-Davidoff et al., 1999; Forster and Shine, 2002; Solomon et al., 2010). The dominant transport of  $H_2O$  into the stratosphere occurs in the tropics and is controlled to first order by the cold temperatures of the tropical tropopause (Liu et al., 2010; Schoeberl and Dessler, 2011; Randel and Jensen, 2013). However, uncertainty remains in both the microphysics of dehydration in the tropical tropopause layer (TTL) and the potential contributions of other processes (e.g.

- overshooting convection, transport from the extra-tropics) to the water budget of the lower stratosphere (Sherwood and Dessler, 2001; Fueglistaler et al., 2009; Randel and Jensen, 2013). Cirrus clouds occur with high frequency and large spatial extent in the TTL, and those occurring near the thermal tropopause facilitate the final dehydration of stratosphere-bound air parcels. The net role of TTL cirrus clouds in the climate system
- and how cloud effects may change as the climate changes remains uncertain (Dessler, 2010; Zelinka and Hartmann, 2011; Randel and Jensen, 2013). Further accurate measurements of water vapor in the tropical UT/LS are needed to improve understanding of the dynamical and microphysical processes that control the dehydration of air transported into the stratosphere. In contrast to water vapor, fewer in situ measurements of cirrus ice water content (IWC) exist for the UT/LS. IWC is an important metric for estimating both the radiative impact of cirrus and their role in the dehydration of tropo-
- estimating both the radiative impact of cirrus and their role in the dehydration of tropospheric air in the TTL.

A number of different measurement techniques have been used to measure water vapor in the UT/LS. These include chilled mirror hygrometry (Brewer et al., 1948; <sup>25</sup> Mastenbrook and Oltmans, 1983; Vömel et al., 2007), Lyman- $\alpha$  photofragment fluorescence (Kley and Stone, 1978; Weinstock et al., 1994; Zöger et al., 1999), and both open- and closed-path tunable diode laser (TDL) absorption spectroscopy (May, 1998; Diskin et al., 2002; Buchholz et al., 2012; Sargent et al., 2013).



Persistent disagreements among collocated measurements of H<sub>2</sub>O in the UT/LS have produced uncertainties in our understanding of the cirrus cloud microphysics that control dehydration of tropospheric air in the tropical tropopause region and determine the amount of H<sub>2</sub>O that reaches the lower stratosphere (Kley et al., 2000; Jensen et al.,

- <sup>5</sup> 2005; Peter et al., 2006; Weinstock et al., 2009). These disagreements among collocated in situ measurements have prompted community efforts to determine the sources of error in the H<sub>2</sub>O measurements, including the chamber-based AquaVIT-1 intercomparison in 2007 (Fahey et al., 2014), and a recent in situ intercomparison during the Mid-latitude Airborne Cloud Properties EXperiment (MACPEX) in 2011 (Rollins et al.,
- <sup>10</sup> 2014). In the AquaVIT-1 intercomparison, a number of the primary UT/LS in situ H<sub>2</sub>O instruments reported values that generally agreed within 20% at mixing ratios down to 1 part per million (ppm,  $10^{-6}$  mol mol<sup>-1</sup>) under static conditions similar to those in the TTL. This was a significant improvement over the 30–100% differences observed at low mixing ratios in previous UT/LS field measurements (Kley et al., 2000; Jensen
- et al., 2005; Weinstock et al., 2009), suggesting that additional factors affect instrument performance when operating on aircraft in field campaigns. During the MACPEX campaign, the observed differences among instruments at low mixing ratios were typically within 20%, similar to the AquVIT-1 results. A careful analysis of the data found that all of the instruments were potentially affected by variable background artifacts that be-
- <sup>20</sup> came significant at mixing ratios below 10 ppm (Rollins et al., 2014). Variations at this level in long-term datasets are sufficient to preclude the early identification of trends in UT/LS H<sub>2</sub>O that are predicted to occur with changes in Earth's climate (Kley et al., 2000; Fueglistaler et al., 2013).

Measurement of TTL cirrus IWC is even more challenging because it combines the issues of water vapor measurement at low mixing ratios with the need to characterize the sampling efficiency of cloud particles. A number of different instruments have been developed and deployed for cirrus IWC determination, using both TDL (Davis et al., 2007a; Dorsi et al., 2014) and Lyman- $\alpha$  (Brown and Francis, 1995; Weinstock et al., 2006a; Schiller et al., 2008) detection. Typically these instruments measure the sum of



water vapor and evaporated condensed-phase water, and rely on a second instrument to measure only water vapor in order to derive a value for the IWC. The counterflow virtual impactor instrument described by Twohy et al. (1997) eliminates this need by inertially separating the cloud particles into a dry synthetic air stream and measuring

- the evaporated water vapor. This method produces a size cut-off for sampled particles and therefore generally will not afford detection of the IWC of sufficiently small ice crystals. Total water sampling inlets are operated either isokinetically (Weinstock et al., 2006a, b) to simplify the characterization of the particle sampling, or subisokinetically (Davis et al., 2007a; Schiller et al., 2008) to enhance the particle concentration sampled
- into the instrument for improved IWC sensitivity. Davis et al. (2007b) report on a comparison of three in situ IWC measurements made during the NASA Midlatitude Cirrus Experiment in 2004. The measurements demonstrated reasonable agreement at high IWC loadings, but the differences were significant at values below 5 mg m<sup>-3</sup>; a value more than an order of magnitude larger than typically found in TTL cirrus (Davis et al., 2007b; Schiller et al., 2008).

To take advantage of the TTL sampling capabilities of the NASA Global Hawk UAS platform, we have developed a new, compact, two-channel instrument for the simultaneous measurement of water vapor and total water in the UT/LS, and integrated it onboard the Global Hawk during the recent NASA Airborne Tropical TRopopause EXperiment (ATTREX) mission. Here we describe details of the instrument design and validation, and present instrument performance from recent measurements in the UT/LS

obtained during ATTREX.

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# 2 The NOAA water instrument

# 2.1 General description

<sup>25</sup> The NOAA Water instrument is a two-channel, closed-path, TDL absorption spectrometer for the measurement of UT/LS water vapor (WV) and enhanced total water (eTW,



vapor + inertially enhanced condensed-phase) from the NASA Global Hawk UAS or other high-altitude research aircraft. The instrument utilizes wavelength modulation spectroscopy with second harmonic (2*f*) detection at a wavelength near 2694 nm to achieve high sensitivity and precision at low ppm mixing ratios. The instrument is cal-

- <sup>5</sup> ibrated on the ground by comparison to a reference chilled-mirror hygrometer (Model 373LX, MBW Calibration, Ltd., Wettingen, Switzerland) and the stability of the calibration during flight is assessed periodically by addition of water vapor to one of the channels. The absorption cells are operated under conditions of constant pressure, temperature, and flow to maintain constant sensitivity independent of pressure and temper-
- ature changes in the payload compartment or in the ambient sampling environment. This approach eliminates the need for highly accurate knowledge of the pressure- and temperature-dependent spectral line parameters and consequently complex data reduction in exchange for a reduced absolute absorption signal that results from operating the cells at pressures below the lowest ambient pressure encountered in flight.
- <sup>15</sup> A summary of the instrument physical and performance specifications appears in Table 1. The main instrument enclosure (Fig. 1) measures 49 cm × 43 cm × 36 cm and houses the optical absorption cells, calibration system, custom electronics, and data acquisition and control computer. The inlet pylon is attached to an interface flange at the fuselage surface and extends outside the fuselage perpendicular to the free stream
- flow. Two small scroll pumps, their associated controllers, and a 24VDC power supply are located external to the main instrument enclosure. The total instrument weight, including pumps and inlet pylon, is 40 kg. During flight, the typical power consumption is 70 W of DC (28 V) and 400 W of AC (120 V, 400 Hz). The AC power draw varies depending on the power required by the heaters used to maintain instrument component temperatures.

## 2.2 Sample flow system

A schematic of the instrument flow system is shown in Fig. 2. Sample flow through the instrument is produced by a pair of small scroll pumps (Model V12H20N2.5, Air



Squared, Broomfield, CO, USA). The sample mass flow through each channel is controlled using a custom butterfly valve (BV) based on the design described by Gao et al. (1999) and similar to those used previously with the NOAA CIMS H<sub>2</sub>O instrument (Thornberry et al., 2013). The BV is located in the inlet line upstream of the absorption cell and is servo-controlled by a mass flow meter (MFM, Model D6F-03A3, Omron Corp., Kyoto, Japan) located in the exhaust line from the absorption cell. The mass flow through each cell was maintained at 600 standard (273.15 K, 1013.25 hPa) cm<sup>3</sup> min<sup>-1</sup> (sccm). The flow value was chosen to be as large as possible, within the capacity of the vacuum pumps, in order to reduce the potential for hysteresis arising from the adsorption/desorption of water to inlet tubing and cell surfaces. The pressure in each

- <sup>10</sup> sorption/desorption of water to inlet tubing and cell surfaces. The pressure in each absorption cell is controlled using a second BV located in the exhaust line from the cell, immediately downstream of the sample flow MFM, which is servo-controlled by a pressure sensor (Model 19C015PA4K, Honeywell Sensing and Control, Minneapolis, MN, USA) located at the laser/detector end of the absorption cell. The absorption cell
- pressures were set to 45 hPa and 50 hPa in the WV and TW channels, respectively, in order for the pressures to be sufficiently lower than the inlet sampling pressures at the highest flight altitudes for Global Hawk operation to enable cell pressures and flows to be held constant at all aircraft flight levels. The control pressure difference between the two channels is due to a combination of the additional ram pressure at the opening of the feruard foring TW inlet and the reduced pressure at the side facing appring of the
- the forward-facing TW inlet and the reduced pressure at the side-facing opening of the WV inlet.

## 2.3 Absorption cells

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The absorption cells for the NOAA Water instrument were designed and manufactured by Port City Instruments (Wilmington, NC, USA). A schematic of an optical absorption cell is shown in Fig. 3. Each cell consists of a 35 cm long, 1.9 cm ID (3.2 cm OD) Ni-plated aluminum cylinder with a Ni-plated aluminum block housing the inlet port and a gold mirror on one end and another block housing the outlet port, laser, detector, and pressure port on the other. This single-reflection arrangement produces



a 78.6 cm round-trip optical absorption path length. The internal cell volume is approximately 120 cm<sup>3</sup>, yielding a sample residence time of less than 0.52 s at a flow rate of 600 sccm, cell pressure of 50 hPa and cell temperature of 40 °C. The inlet, outlet and pressure ports were designed for o-ring face-seal fittings (VCO, Swagelok, Inc., Solon, 5 OH, USA) to minimize potential for leaks into the cells.

In order to thermally stabilize the optical cavity and minimize the migration of interference fringes through the spectrum, the absorption cells are housed in a temperaturecontrolled sub-enclosure within the overall instrument enclosure. The temperature of the sub-enclosure is stabilized by controlling the air temperature within the subenclosure by means of forced circulation of air over temperature-controlled heat sinks and by separately controlling the temperatures of the individual aluminum panels comprising the sub-enclosure. Multiple thermistors are used to monitor the thermal stability

of the absorption cells.
 The laser and detector are mounted directly into the block at the outlet end of the
 absorption cell to minimize the potential for any volume in the optical path that is not actively swept by the sample flow through the cell. A lens mounted in an open structure in front of the laser is used to collimate the beam. The gold mirror is sealed against the inlet end of the absorption cell using an o-ring face seal. Four small screws in the mirror assembly that compress the mirror against the o-ring allow adjustment of the laser of the laser beam.

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## 2.4 Lasers and detectors

The lasers used in the instrument are distributed feedback lasers (DFB, nanoplus GmbH, Gerbrunn, Germany) mounted on thermoelectric coolers (TEC) and sealed in TO5 packages with antireflection-coated, wedged windows. The laser packages are evacuated and backfilled with  $N_2$  at 1 bar pressure and contain less than 30 ppm residual  $H_2O$ . The sealed optical path between the diode facet and TO5 package window is 0.5 mm, resulting in a maximum optical depth due to residual  $H_2O$  in the laser packages equivalent to 0.375 ppm in the sample cell at 50 hPa. Lasing threshold current



and tuning characteristics are inherently variable with DFB lasers; however the useful tunable range accessible in a single current scan for the lasers used here was typically  $2-3 \text{ cm}^{-1}$ .

- The InAs detectors (J12TE1-37S-R01M, Teledyne Judson Technologies, Montgomeryville, PA, USA) used are mounted on TECs in sealed 37s packages. These packages were evacuated, baked, and backfilled with N<sub>2</sub> during fabrication to 50 hPa with a maximum of 5 ppm residual H<sub>2</sub>O. The optical path between the photodiode and window is approximately 2 mm, resulting in a signal from trapped water equivalent to less than 0.014 ppm water in the detection cell.
- For measurement of H<sub>2</sub>O, a pair of absorption lines near 2694 nm was identi-10 fied from the HITRAN 2008 database (Rothman et al., 2009). To access the selected H<sub>2</sub>O absorption features, the laser TEC temperatures are set to values that vield output near the target wavelength for each individual laser and the laser drive currents are ramped to scan the lasers across a spectral range of approximately  $1 \text{ cm}^{-1}$ . The primary absorption line used to detect low mixing ratios of H<sub>2</sub>O is from 15 the asymmetric stretch mode  $((0,0,1) \leftarrow (0,0,0))$  at 3712.20 cm<sup>-1</sup> (2693.82 nm) with a line strength of  $1.49 \times 10^{-19}$  cm<sup>-1</sup> cm<sup>2</sup>. This line is approximately an order of magnitude stronger than strong lines in the  $H_2O$  absorption band near 1400 nm that have typically been used for UT/LS H<sub>2</sub>O measurement. A weaker line at 3711.88 cm<sup>-1</sup> (2694.06 nm) due to the symmetric stretching mode  $((1,0,0) \leftarrow (0,0,0))$  with a line 20 strength of  $4.41 \times 10^{-21}$  cm<sup>-1</sup> cm<sup>-2</sup> is used at higher H<sub>2</sub>O mixing ratios to provide a linear response up to and exceeding 2000 ppm. Two prominent CO<sub>2</sub> absorption lines at 3711.62 cm<sup>-1</sup> (2694.24 nm) and 3711.75 cm<sup>-1</sup> (2694.15 nm) also fall within the scan window and could potentially be used for additional instrument diagnostics.
- <sup>25</sup> The 1 cm<sup>-1</sup> spectral sweep is performed at 10 Hz and the spectra are co-averaged to produce 1 s data. Improved precision is achieved by use of 2*f* detection, which is employed by adding a 30 kHz sine wave modulation to the current ramp, producing a 0.08 cm<sup>-1</sup> modulation depth of the scan, and demodulating the signal from the photodiode detector at 60 kHz. Figure 4 presents the spectral region scanned showing



calculated transmission through the optical cell using line parameters from the HITRAN 2008 database along with the measured direct absorption and 2f spectra of ambient air containing 100 ppm H<sub>2</sub>O. The asymmetry observed in the peaks in the 2f spectrum is due to the increasing laser power with current across the scan. To account for

<sup>5</sup> variations in laser power that would affect the instrument sensitivity, the 2*f* amplitude is normalized to the laser power at each spectral feature (N2f = 2f peak-to-peak amplitude/laser power at line center) before calculating water mixing ratios. Laser power at the center of each line is determined from a polynomial fit to the direct absorption baseline.

#### 10 2.5 Inlets

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The inlets for the two channels of the NOAA Water instrument are housed in an airfoil shaped pylon assembly that extends 27 cm perpendicular to the skin of the aircraft in order to sample in the free stream flow around the aircraft. The pylon is 14 cm wide (fore-aft) and 3.5 cm thick. A cross-section of the inlet pylon is shown in Fig. 5. Gas flows between the inlet pylon and the instrument pass through custom o-ring face seal fittings at the interface flange. This design provides a good, reproducible seal while allowing the inlet pylon to be easily installed and removed when necessary for accessing the instrument.

#### 2.5.1 Water vapor inlet

The sampling point for the WV channel is located at the end of the inlet pylon, oriented perpendicular to the flow past the aircraft (Fig. 5). This sampling arrangement, combined with the thin plate at the end of the pylon, allows for sampling of ambient water vapor while rejecting cloud particles. The design is similar to that described in Popp et al. (2004) for sampling gas-phase nitric acid while excluding nitric acid trihydrate particles, and that described by Perring et al. (2013) for sampling submicron aerosol while rejecting cloud particles. A similar inlet was used with the NOAA CIMS



 $H_2O$  instrument (Thornberry et al., 2013) for water vapor measurements during the MACPEX campaign, where the measurements showed no detectable contamination from cloud water during extensive cirrus sampling. The WV sample line is constructed of 0.46 cm ID electropolished stainless steel tubing (WinTech 10, Winter Technologies,

Pacific, MO, USA), which was found in our previous laboratory experiments to produce little hysteresis to rapid changes in water vapor at mixing ratios between 1 and 50 ppm (Thornberry et al., 2013). A section of the WV inlet line closest to the tip is heated to 50 °C to prevent condensation in the inlet.

## 2.5.2 Total water inlet

<sup>10</sup> The sampling point for the TW channel is located at the end of a forward facing 0.46 cm ID tube that affords sampling of both water vapor and condensed-phase water. The inlet tube is oriented parallel to the local flow at a distance of 22.5 cm from the aircraft surface and extends 6.5 cm forward from the leading edge of the pylon. Inside the pylon, 10.5 cm from the tip, the inlet tube makes a sharp (mitered) 90° turn to facilitate impaction of ice particles on the heated inlet wall. 9 cm of the forward facing sample line is enclosed in a 0.2 cm thick copper sheath extending from a copper block at the bend that is heated to 180°C. After the bend, the sample line passes through the catalyst heater block, also controlled to 180°C, before entering the instrument.

The controlled inlet flow of 600 sccm causes the sampling to be subisokinetic, i.e. the sample air velocity inside the inlet (*U*) is lower than the free stream velocity ( $U_0$ ) at sampling altitudes and typical high altitude research aircraft speeds (140–200 m s<sup>-1</sup>). This subisokinetic sampling leads to an inertially enhanced aspiration efficiency for particles into the inlet (Belyaev and Levin, 1974; Krämer and Afchine, 2004; Eddy et al., 2006; Davis et al., 2007a). As a result, cirrus particle number densities in the sample flow are enhanced by a factor, EF, defined as the ratio of particle number concentra-

tion entering the inlet tube to that in the free stream. The EF is a function of the inlet geometry, aircraft attitude, the ratio of the free stream flow to the inlet flow  $(U_0/U)$  and the Stokes numbers (Stk) of the particles. Since the mass flow through the TW inlet is



constant, the volumetric flow rate and therefore the sample flow velocity, U, is a function of the ambient pressure and temperature and varies significantly with altitude. The parameterization developed by Eddy et al. (2006, Eq. 15) based on computational fluid dynamic simulations of various inlet geometries was used to calculate the particle size-

- dependent EF for the TW inlet. Figure 6 shows the calculated EF curves as a function of particle diameter for typical TTL operating conditions of the Global Hawk UAS. A density of 0.7 g cm<sup>-3</sup> was used for calculations based on the reported effective density of small (< 50 µm) cirrus particles by Cotton et al. (2013). As expected, the value of EF is near unity for small particles (< 0.2 µm) and increases with particle size. For particles larger than approximately 10 µm in diameter, typical of cirrus particles, the EF for the</li>
- TW inlet approaches the maximum value of  $U_0/U$ , which is in the range of 30 to 50. This enhancement results in a significantly increased sensitivity to cirrus IWC.

After cloud particles are sampled into the inlet, the condensed-phase water must be vaporized in order for it to be measured by the instrument. The heated section of the

- <sup>15</sup> TW inlet consists of 9 cm of the forward-facing tube and an additional 18.5 cm length following the 90° bend. Calculations of the evaporation of ice particles (assumed to be solid spheres with density  $0.7 \text{ g cm}^{-3}$ ) in the inlet at a pressure of 85 hPa indicate that particles less than 10 µm diameter will vaporize before reaching the bend, and particles less than 20 µm diameter would vaporize in the 27.5 cm heated region even with-
- $_{20}$  out contacting the hot walls. Particles larger than 20  $\mu m$  initial diameter have sufficient residual inertia to impact the tubing at the 90° bend, greatly increasing heat transfer and potentially producing shattering of the particle, accelerating subsequent vaporization. Calculations at 150 hPa indicate that particles up to 26  $\mu m$  would evaporate in the heated region, while particles with an initial diameter > 24  $\mu m$  will impact the tubing at
- the bend. Based on these calculations, and the additional 25 cm of warm (20–50 °C) inlet tubing through which the sample flow passes prior to reaching the absorption cell, it is reasonable to conclude that all condensed-phase water entering the TW inlet is vaporized prior to measurement.



The direct measurement of the water mixing ratio in the TW channel is defined as enhanced total water (eTW), which is composed of ambient water vapor ( $H_2O_{vap}$ ) and evaporated, inertially enhanced IWC (eIWC) as:

$$_{5}$$
 eTW = H<sub>2</sub>O<sub>vap</sub> + EF · IWC = H<sub>2</sub>O<sub>vap</sub> + eIWC

The actual IWC of a sampled cirrus cloud is then determined as

 $IWC = (eTW - H_2O_{vap})/EF$ 

where H<sub>2</sub>O<sub>vap</sub> is measured in the WV channel, and the overall EF is determined by convolving the calculated EF curve for the immediate sampling conditions with an assumed cirrus ice particle size distribution or that provided by a co-measuring cloud probe instrument.

#### 2.6 Calibration system

The NOAA Water instrument uses an on-board calibration system to evaluate the sta-<sup>15</sup> bility of the sensitivity and background of the WV channel during flight. The calibration system is similar to the one used previously in the NOAA CIMS H<sub>2</sub>O instrument (Thornberry et al., 2013) and is based on the catalytic oxidation of H<sub>2</sub> standards to produce known H<sub>2</sub>O mixing ratios (Rollins et al., 2011). The current system (see Fig. 2) consists of one dry synthetic air (ZA) cylinder (M22A, Luxfer Gas Cylinders, Riverside,

- <sup>20</sup> CA, USA) and two cylinders (M09B, Luxfer) containing different mixing ratios of H<sub>2</sub> in air. The ZA is passed through a 1 cm ID by 25 cm long stainless steel tube containing a molecular sieve desiccant (BMSR-1, Agilent Technologies, Inc., Santa Clara, CA, USA), which reduces the residual water vapor mixing ratio in the ZA. This moisture trap is located downstream of the ZA pressure regulator (Series 4300-N, Premier In-
- <sup>25</sup> dustries, Blaine, MN, USA) to capture water from as many potential leaks as possible before the ZA is directed to the inlets. The catalyst assembly, located in the inlet pylon, consists of a 14 cm long, 0.21 cm ID, electropolished stainless steel tube containing



(1)

(2)

two 2.5 cm × 5 cm pieces of tightly rolled Pt mesh (part no. 298093-1.7G, Sigma-Aldrich Co., St. Louis, MO, USA). The catalyst is maintained at a temperature of 180 °C, which was demonstrated to yield complete conversion of  $H_2$  to  $H_2O$  for a similar geometry and flow rates to those used in the NOAA Water instrument (Rollins et al., 2011).

<sup>5</sup> Mass flow controllers (MC-Series, Alicat Scientific, Inc., Tucson, AZ, USA), which were tested and found to be insensitive to changes in ambient enclosure pressure, are used to control the ZA flows to the inlets and the ZA and H<sub>2</sub>/air flows through the Pt catalyst.

To perform a calibration, the ZA flow to the WV inlet is set to 1000 sccm, overflowing the inlet with ZA. The H<sub>2</sub>/air flows through the catalyst are adjusted in a series of steps between 2 and 20 sccm and the output of the catalyst is added to the ZA in the sample flow to produce a series of known H<sub>2</sub>O mixing ratios, which are used periodically to determine the sensitivity of the WV channel. Due to space and weight constraints and the desire to use a minimum number of fittings (potential leaks), only the WV channel is calibrated in flight. The stability of the sensitivity of the TW channel during flight is assessed by comparing its signal with that of the WV channel when sampling in cloudfree air.

To mitigate the potential effects of parasitic water from infinitesimal leaks into the calibration system, 10 sccm flows of ZA are used to continuously flush the ZA and calibration system tubing out to the inlet addition points and back through return flow controllers (see Figs. 2 and 5). The return flow controllers are set to additionally pull small (10–15 sccm) flows of sample air from the inlets to prevent ZA from reaching the sample flow during ambient sampling.

## 2.7 Data system

Instrument control, data acquisition, and communication with the aircraft are accomplished using a custom LabView program running on a CompactRIO system (National Instruments, Inc., Austin, TX, USA). Custom Arduino-based electronics boards are



used for temperature control of instrument components and operation of the flow- and pressure-control butterfly valves.

Instrument temperatures, pressures and flows (housekeeping) along with laser parameters, N2f signals and calculated  $H_2O_{vap}$  and eTW mixing ratios are recorded at

<sup>5</sup> 1 Hz (~4 MB h<sup>-1</sup>) for post-flight data processing and instrument performance evaluation. A subset of this data can be transmitted to the ground for real-time reporting of mixing ratios and monitoring of instrument performance during flight if communication is available. 1024-point direct absorption and 2*f* spectra for each channel are acquired at 10 Hz and coaveraged and saved at 1 Hz (29.5 MB h<sup>-1</sup>). The instrument can oper <sup>10</sup> ate entirely autonomously or with user interaction, depending on the communications capability of the aircraft platform.

#### 3 Instrument performance

The first field deployment of the NOAA Water instrument occurred in the second phase of the ATTREX mission in January–March 2013, during which the instrument was integrated on the NASA Global Hawk UAS. During this mission, the Global Hawk conducted six research flights flight with durations of 24 (±0.3) hours, allowing for transit south from NASA Dryden Flight Research Center in California to the central and eastern tropical Pacific with substantial time at low latitudes to conduct extensive sampling in the TTL. Measured water vapor mixing ratios in the TTL reached values as low as
1.5 ppm and cirrus clouds with a wide range of IWC were encountered.

Operational procedures for the NOAA Water instrument during ATTREX involved calibrating the instrument on the ground between flights, as well as extensively in the laboratory before and after the mission. These calibrations were conducted with sample flows containing H<sub>2</sub>O mixing ratios between 0.5 and 2500 ppm as quantified by the MBW frost point hygrometer. The calibration flows were generated by an external H<sub>2</sub>/Pt

<sup>25</sup> MBW frost point hygrometer. The calibration flows were generated by an external  $H_2/Pt$  catalytic  $H_2O$  source (Rollins et al., 2011) and distributed through a common manifold. Figure 7 shows results from a laboratory calibration over the range 2–2140 ppm. The



non-linearity in the N2*f* signal that arises from high optical depth is apparent at mixing ratios above 100 ppm in the strong absorption lines, while electronic noise and optical fringing produce non-linearity in the WV and TW weak absorption line signals below 40 ppm and 60 ppm, respectively. Repeated calibrations performed with many hours of

- <sup>5</sup> continuous operation at mixing ratios less than 20 ppm were used to verify the linearity of the strong absorption line N2*f* signals at these low mixing ratios, and to quantify the background water signal in each channel (Fig. 8). The background equivalent water value in the WV channel was determined to be  $0.3 \pm 0.1$  ppm, while that in the TW channel was  $0.3 \pm 0.2$  ppm.
- <sup>10</sup> During the ATTREX flights, H<sub>2</sub>O vapor and eTW mixing ratios were calculated in real time on a 1 Hz basis from the N2*f* signals in the WV and TW channels. Because data reduction is performed by empirically calibrating the N2*f* signal against a H<sub>2</sub>O reference standard (MBW), the measurement accuracy does not depend on highly accurate knowledge of the laser modulation amplitude. Further, uncertainties due to changes in
- <sup>15</sup> spectral line shape with changes in gas temperature and pressure do not contribute to measurement uncertainty as long as these parameters do not change between groundbased calibrations and in-flight measurements. Mixing ratios were calculated from the N2*f* signals in the low optical depth regime for each line, where the non-linearity in the instrument response is less than 0.5%. Only the TW channel measured values high enough to require use of the weak absorption line, which was used for mixing ratios
- <sup>20</sup> enough to require use of the weak absorption line, which was used for mixing ratios above 80 ppm.

During flights, the detection cell pressures were maintained to within  $\pm 0.7$ % of the pressure at which the channel was calibrated in the laboratory, and the absorption cell temperatures were maintained within a  $\pm 0.03$ °C range after the optics sub-enclosure

reached its temperature set point. Therefore variations in the instrument sensitivity due to changes in these parameters are negligible. To directly verify the stability of the instrument response during flight operation, the ambient flow in the WV channel was periodically replaced with calibration flows produced using the Pt/H<sub>2</sub>/ZA system. Figure 9 shows results from the in-flight calibrations during the flight on 21–22 February 2013.



The measured instrument response (sensitivity) for both  $H_2O$  spectral lines in individual calibrations during the flight agreed with those measured in the laboratory to better than 4 %, which is within the uncertainty expected of the onboard calibration system. These results are typical of the in-flight calibrations on all of the ATTREX flights.

- <sup>5</sup> In-flight measurement precision was evaluated by calculating the distribution of 1 s changes in the WV channel measurements ( $\Delta WV = WV[t] - WV[t - 1]$ ). Figure 10 shows a histogram of  $\Delta WV$  from the 21–22 February 2013 flight with a Gaussian fit indicating that the noise is normally distributed with a width of 0.34 ppm, corresponding to a 1 $\sigma$  precision of 0.17 ppm. Noise in the TW channel was typically slightly higher than in the WV channel (0.25 for the 21–22 February flight). The observed in-flight pre-
- than in the WV channel (0.25 for the 21–22 February flight). The observed in-flight precision is independent of signal amplitude and compares well, given a contribution from real atmospheric variability, with the typical value of 0.1 ppm observed while sampling stable mixing ratios in the laboratory.
- Also shown in Fig. 10 is the histogram of the difference between 1 s TW and WV the channel signals during cloud-free sampling. This difference was normally distributed and centered near zero (-0.012 ppm). The detection limit for eIWC is defined as twice the standard deviation ( $2\sigma$ ) in the clear-sky difference of the signal. For the 21–22 February flight this results in an eIWC detection limit of 0.641 ppm. Assuming the cirrus particles are greater than 8 µm effective diameter, the calculated EFs of 33–48 yield
- an IWC mixing ratio detection limit from 0.023 to 0.013 ppm, depending on altitude. Converting these numbers to mass, the IWC detection limit was approximately 2 µg m<sup>-3</sup> across the sampling altitude range. Figure 11 shows a sample time series of WV and eTW mixing ratios during a vertical profile through the TTL from the 21–22 February flight. Cirrus encounters are clearly visible on both the ascent and descent. This data
- sample illustrates the typical instrumental precision in the measurements, as well as the good agreement between the WV and TW channels.



#### 3.1 Water vapor measurement uncertainty

The principal sources of uncertainty in the determination of ambient  $H_2O$  mixing ratios in the NOAA Water instrument are (1) the accuracy of the MBW reference frost point hygrometer, (2) uniformity of the  $H_2O$  mixing ratio in the external calibration flows to the MBW and the NOAA Water instrument (analytical split), (3) stability of the absorption call pressures (4) linearity of the calibration fit, and (5) variation of antical fringes in the

cell pressures, (4) linearity of the calibration fit, and (5) variation of optical fringes in the wavelength scan.

The NIST-traceable accuracy of the MBW frost point determination is 0.1 °C, which at 850 hPa (typical ambient pressure in Boulder, CO) corresponds to an uncertainty of 1.6 % at 1 ppm decreasing to approximately 1 % at 1000 ppm. The uncertainty aris-

- <sup>10</sup> of 1.6 % at 1 ppm decreasing to approximately 1 % at 1000 ppm. The uncertainty arising from changes to the sensitivity (H<sub>2</sub>O absorption) with changes in cell pressure are similar in magnitude to the variations in the cell pressure for small changes, and were typically less than 0.7 %. The uncertainty due to changes to the N2*f* signal caused by variations in the optical fringes in the spectrum that arise from small changes in the
- optical path with fluctuations in cell temperature was assessed by analyzing extended (up to 12 h) sampling of constant mixing ratios. The observed variations in the signal were equivalent to less than 0.07 ppm H<sub>2</sub>O for temperature fluctuations similar to those observed in flight (±0.03 °C). Potential issues with analytical split of the external calibration between the MBW and the instrument were addressed by examining the
- <sup>20</sup> repeatability of the laboratory calibrations with a number of different plumbing configurations and found to be negligible. During the 2013 ATTREX deployment, including preand post-mission calibrations, the goodness of the individual linear fits to the calibration data were better than 4 %, and the average difference between individual calibrations of the WV channel is described by the linear function 0.1 + 0.005  $\cdot$  H<sub>2</sub>O. Combining the
- <sup>25</sup> uncertainty arising from these factors yields an accuracy of  $5\% \pm 0.15$  ppm for the WV channel. Including the typical precision value of 0.17 ppm, the overall uncertainty in the H<sub>2</sub>O mixing ratio in the WV cell is  $5\% \pm 0.23$  ppm for a 1 s measurement. For the eTW H<sub>2</sub>O measurement in the TW channel, the overall uncertainty was slightly higher



at  $6\% \pm 0.32$  ppm due to contributions from larger optical fringes in the spectrum and slightly higher electronic noise (precision). The overall good agreement between TW and WV channels during periods of cloud-free sampling in all of the ATTREX flights is shown in the correlation plot in Fig. 12. An orthogonal linear regression fit to all of the data yielded a slope of 0.96 and the correlation had an  $r^2$  value of 0.97. Slopes of the regressions for individual flights ranged from 0.95 to 0.98, indicating a small variation in

sensitivity occurred over time. This variation appeared to be primarily associated with shifts in the optical fringes in the TW channel. No consistent offset between the two channels in flight was observed.

#### 10 3.2 Ice water content measurement uncertainty

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A number of additional sources of uncertainty affect the IWC measurement. These include (1) the particle aspiration efficiency of the inlet, (2) size distribution of the ice particles, (3) completeness of ice particle vaporization within the inlet, and (4) disturbance of the ambient cirrus ice particle distribution at the sampling point by the presence of the aircraft. Conservative calculations of ice particle vaporization indicate that an assumption of complete vaporization of all ice particles within the TW inlet is well justified and therefore contributes negligibly to the overall IWC measurement uncertainty. Calculating the TW inlet EF requires knowledge of a number of atmospheric (pressure, temperature) and aircraft (true air speed, pitch) and instrument (inlet flow)

- <sup>20</sup> parameters that all contribute some uncertainty. For ATTREX, the atmospheric and aircraft parameters were measured by the Meteorological Measurement System (NASA Ames Research Center, T. V. Bui, PI), which reports uncertainties of  $\pm 0.3$  hPa,  $\pm 0.3$  K and  $\pm 1.0$  m s<sup>-1</sup> for temperature, pressure and wind vector, respectively. These uncertainties are sufficiently small to contribute only about 1 % uncertainty to the calculated
- value  $U_0$ . The uncertainty in inlet flow measured by the mass flow meter is 3%. The typical pitch of Global Hawk was 1.4° (±0.6°), so the correction for reduction in geometric inlet sampling area is negligible. Eddy et al. (2007) report that the EF correlation presented (Eq. 15) agreed with their numerical simulations within 20% for most Stk



values. The uncertainty in IWC determination arising from the size dependence of the EF depends on the fraction of ice mass in crystals with small diameters, where the EF decreases from the asymptotic value of  $U/U_0$ . If all ice mass were in particles of 5 µm diameter, assuming the  $U_0/U$  value for the EF would result in an underestimate of IWC

- <sup>5</sup> by less than 10%, while if the particles were all 3μm, the underestimate would rise to 33%. The magnitude of this uncertainty can be reduced by using a co-measured ice particle size distribution, if one is available. Combining these uncertainty factors results in an overall uncertainty for IWC determination of 15–30%, depending on the mass-weighted cloud particle size distribution.
- Estimating the additional IWC uncertainty associated with potential sampling artifacts caused by the presence of the aircraft is difficult. During the ATTREX mission, the NOAA Water instrument was mounted in a forward fuselage bay on the right side of the aircraft, 3 m behind the nose. Potential ice crystal bounce and deflection effects on sampling have not been investigated for the GH using computational fluid dynamic simulations, although in the future they could potentially yield some insight into these
- Isinulations, altrough in the future they could potentially yield some insight into these issues. In the comparison of total water measurements during the MidCiX campaign reported by Davis et al. (2007b), the reasonable agreement between three instruments sampling from different locations on the NASA WB-57F suggest that the effect may be small.

#### 20 4 Summary

The NOAA Water instrument was developed for accurate in situ measurements of water vapor and cirrus IWC in the TTL and elsewhere in the UT/LS, which are important for our understanding of the dynamical and microphysical processes that regulate the  $H_2O$  distributions in this climatically important region. The instrument is based on TDL

absorption spectroscopy of  $H_2O$  in two single-reflection absorption cells with an absorption path length of 78 cm. Wavelength modulation spectroscopy with a strong  $H_2O$  absorption line near 2694 nm is used to achieve high sensitivity and good measurement



precision at low-ppm H<sub>2</sub>O mixing ratios. A second, weaker absorption line is used to provide dynamic range for measurement of mixing ratios up to 2500 ppm. Data reduction is simplified by operating the absorption cells at constant pressure, temperature, and flow, independent of changes in the ambient sampling environment. The instru <sup>5</sup> ment is calibrated against a reference hygrometer on the ground and utilizes an on-board calibration system to verify the stability of the measurement sensitivity during flight.

The dual-cell design of the instrument allows simultaneous measurement of water vapor and enhanced total water for the determination of cirrus IWC. Air for the measurement of  $H_2O$  vapor is sampled through an inlet oriented perpendicular to the free stream flow past the aircraft in order to reject ice crystals, while a forward-facing, subisokinetic inlet produces inertially enhanced sampling of cirrus particles to yield high sensitivity to IWC.

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The instrument was successfully deployed in multiple flights in the tropical UT/LS on <sup>15</sup> board the NASA Global Hawk UAS during the second deployment of the ATTREX mission. Based on flight data, the overall uncertainty in the measurement of water vapor (1 s, 1 $\sigma$ ) is 5 % ± 0.23 ppm, while uncertainty in the determination of cirrus IWC is estimated to be approximately 20 % (15–30 %). The enhancement factor for ice particles larger than 8 µm was calculated to be in the range of 33–48 depending on aircraft alti-<sup>20</sup> tude, which results in an IWC detection limit of approximately 2 µg m<sup>-3</sup>. Measurements of H<sub>2</sub>O mixing ratio and cirrus IWC in the Pacific TTL during the ATTREX mission will be used to investigate the dynamic processes and microphysics related to the dehydration of air as it is transported through this region.

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 Table 1. NOAA Water instrument specifications.

Size	49 cm × 43 cm × 36 cm	
Weight	40 kg	
Power	70 W DC, 400 W AC	
Data rate	1 Hz	
	WV	TW
Accuracy	5 % ± 0.15 ppm	6 % ± 0.20 ppm
Precision (1 s, 1 $\sigma$ )	0.17 ppm	0.25 ppm
Uncertainty (1 s)	5 % ± 0.23 ppm	6 % ± 0.32 ppm
IWC Uncertainty (1 s)		15–30 %

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**Figure 1.** Schematic of the NOAA Water instrument showing the inlet pylon and some of the principal instrument components. The enclosure structure and pylon are fabricated from aircraft-grade aluminum.





**Figure 2.** Schematic of the NOAA Water instrument showing the components of the dualchannel flow system. The custom butterfly valves located upstream and downstream of each cell are used to control the mass flow (upstream) and pressure (downstream) in the optical cells to constant values independent of variations in sampling pressure as the aircraft changes altitude. The numbers in the mass flow controller symbols indicate the flow value or range used during flight operation.





**Figure 3.** Schematic of a NOAA Water instrument optical absorption cell. Arrows indicate the direction of flow through the cell. "P" indicates the pressure transducer. The laser and detector are visible in the detail on the left (outlet) and the mirror assembly in the detail on the right (inlet).

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**Figure 4.** Top: HITRAN calculation of transmission through the detection cell with 100 ppm  $H_2O$  (blue) and 370 ppm  $CO_2$  (red). Middle: 1-s coaveraged direct absorption signal in the WV channel from the final descent on 27 February 2013 at an ambient  $H_2O$  mixing ratio of approximately 100 ppm. The green dashed line is the polynomial fit to the baseline for determination of the  $I_0$  values at the strong and weak  $H_2O$  absorption lines. Bottom: corresponding 1-s coaveraged 2*f* spectrum. Note, the scan is plotted as it is recorded in time and the wavenumber scale decreases from left to right.





**Figure 5.** Schematic of the NOAA Water instrument inlet pylon showing the WV and TW sample inlet tubes, the ZA delivery lines (blue) to each inlet, the Pt catalyst and calibration delivery line (cyan) to the WV inlet, and the ZA and calibration return lines (magenta).





**Figure 6.** TW inlet particle aspiration enhancement factor (EF) as a function of ice particle diameter calculated using the formulation of Eq. (15) in Eddy et al. (2006) for the TW inlet geometry and an assumed particle density of  $0.7 \,\mathrm{g \, cm^{-3}}$ . The curves represent conditions near the tropical tropopause (approximately 17.6 km, blue) and near the bottom (approximately 14.4 km, green) of the typical NASA Global Hawk operational altitude range during the ATTREX mission.





**Figure 7.** Laboratory calibration curves showing the N2*f* signals from the strong and weak absorption lines (2693.82 and 2694.06 nm, respectively) in the WV and TW channels as a function of the H<sub>2</sub>O mixing ratio in the sample flow measured by the MBW reference hygrometer. Each data point represents an average of 100–300 s of data at a constant mixing ratio. The error bars are the standard deviation of the N2*f* signal in each interval. The black lines are the weighted orthogonal linear regression fits to the data, which are used to determine instrument sensitivity for each absorption line. The differences in N2*f* values between channels at the same MBW value are due to intrinsic performance and configuration differences between the channels.





**Figure 8.** Laboratory calibration of the WV and TW channels at low  $H_2O$  mixing ratios. Top: time series of water vapor values in the sample flow as measured by the MBW reference hygrometer. Middle: corresponding time series of the strong absorption line N2*f* signals in each channel. Bottom: Linearity of the N2*f* signals at low mixing ratios. Repeated long-duration calibrations at low mixing ratios were used to quantify and verify the stability of the background water signal from residual water in the laser and detector packages.





**Figure 9.** WV channel sensitivity measured during in-flight calibrations from a flight on 21–22 February 2013. Top panel: water vapor mixing ratios and aircraft altitude during a series of profiles through the TTL. Middle panel: calibration slopes (sensitivity) for the weaker absorption line  $(3711.9 \text{ cm}^{-1})$ . Bottom panel: calibration slopes for the strong H<sub>2</sub>O absorption line  $(3712.2 \text{ cm}^{-1})$ . Each data point represents a calibration procedure in which ZA displaced the ambient flow for 8.7 min and a series of 8 water vapor mixing ratios between 2 and 400 ppm were added in time intervals of 40–60 s. In the middle and bottom panels, the dash-dot horizontal lines indicate the sensitivity measured in the laboratory prior to the flight. The dotted lines indicate the ±5% range surrounding the pre-flight laboratory value. The blue solid lines simply connect the data points as an aid to the reader.





**Figure 10.** Left: Histogram of 1 s changes in the WV channel (strong line) during the 21–22 February 2013 flight. The width of the Gaussian fit to the distribution is 0.34, indicating the signal noise is normally distributed with  $\sigma = 0.17$  ppm. Right: Histogram of the difference between 1 s measurements in the TW and WV channels during cloud-free sampling in the 21–22 February 2013 flight. The difference is normally distributed around 0 with a width of 0.45 ppm. This scatter yields a detection threshold ( $2\sigma$ ) for eIWC of 0.641 ppm.











**Figure 12.** Correlation of  $H_2O$  measured in the TW channel with that measured in the WV channel for cloud-free sampling conditions during ATTREX flights. The slope of the correlation calculated using linear orthogonal distance regression was 0.96. Individual flight correlation slopes ranged from 0.95 to 0.98 and the  $r^2$  values of the correlations ranged from 0.92 to 0.99.

