A comparative study of aerosol microphysical properties retrieved from ground-based remote sensing and aircraft in-situ measurements during a Saharan dust event

M. J. Granados-Muñoz^{1,2,*}, J.A. Bravo-Aranda^{1,2}, D. Baumgardner³, J.L. Guerrero-Rascado^{1,2} D. Pérez-

Ramírez^{4,5}, F. Navas-Guzmán⁶, I. Veselovskii⁷, H. Lyamani^{1,2}, A. Valenzuela^{1,2}, F.J. Olmo^{1,2}, G. Titos.^{1,2}, J. Andrey^{8,**}, A. Chaikovsky⁹, O. Dubovik¹⁰, M. Gil-Ojeda⁸ and L. Alados-Arboledas^{1,2}

¹Andalusian Institute for Earth System Research (IISTA-CEAMA), Avd. del Mediterráneo, 18006, Granada, Spain

²Dpt. Applied Physics, University of Granada, Fuentenueva s/n, 18071, Granada, Spain

³Droplet Measurement Technologies, Boulder, CO 80301

⁴ Mesoscale Atmospheric Processes Laboratory, NASA Goddard Space Flight Center, 20771, Greenbelt, Maryland, United States.

⁵Universities Space Research Association, 21044, Columbia, Maryland, United States. ⁶Institute of Applied Physics (IAP), University of Bern, Bern, Switzerland

⁷Physics Instrumentation Center of General Physics Institute, Troitsk, Moscow Region, 142190, Russia. ⁸Instituto Nacional de Técnica Aeroespacial (INTA), Ctra. Ajalvir km. 4, 28850 Torrejón de Ardoz, Spain.

⁹Institute of Physics, National Academy of Science, Minsk, Belarus

¹⁰Laboratoire d'Optique Atmospherique, CNRS Universite de Lille 1, Bat P5 Cite scientifique, 59655 Villeneuve d'Ascq Cedex, France

*Currently at Table Mountain Facility, NASA/Jet Propulsion Laboratory, California Institute of Technology, Wrightwood, California, USA.

**Currently at CNRM-GAME, Météo-France.

Corresponding author: María José Granados Muñoz. Departamento de Física Aplicada, Universidad de Granada, Granada, Spain.

Phone: +34 958241000 (31174)

E-mail: mjgranados@ugr.es

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ABSTRACT

In this work we present an analysis of aerosol microphysical properties during a mineral dust event taking advantage of the combination of different state-of-the-art retrieval techniques applied to active and passive remote sensing measurements and the evalution of some of those techniques using independent data acquired from in-situ aircraft measurements. Data were collected in a field campaign performed during a mineral dust outbreak at the Granada, Spain experimental site (37.16° N, 3.61° W, 680 m a.s.l.) on June 27, 2011. Column-integrated properties are provided by sun- and star-photometry which allows a continuous evaluation of the mineral dust optical properties during both day and night-time. Both the Linear Estimation and AERONET (Aerosol Robotic Network) inversion algorithms are applied for the retrieval of the column-integrated microphysical particle properties. In addition, vertically-resolved microphysical properties are obtained from a multi-wavelength Raman lidar system included in EARLINET (European Aerosol Research Lidar Network), by using both LIRIC (Lidar Radiometer Inversion Code) algorithm during daytime and an algorithm applied to the Raman measurements based on the regularization technique during night-time. LIRIC retrievals reveal the presence of dust layers between 3 and 5 km a.s.l with volume concentrations of the coarse spheroid mode up to 60 μ m³/cm³. The combined use of the regularization and LIRIC methods reveals the night-to-day evolution of the vertical structure of the mineral dust microphysical properties and offers complementary information to that from column-integrated variables retrieved from passive remote sensing. Additionally, lidar depolarization profiles and LIRIC retrieved volume concentration are compared with aircraft in-situ measurements. This study presents for the first time a comparison of the total volume concentration retrieved with LIRIC with independent in situ measurements, obtaining differences within the estimated uncertainties for both methods and quite good agreement for the vertical distribution of the aerosol layers. Regarding the depolarization, the first published dataset of the CAS-POL sonde for depolarization ratios is presented here and qualitatively compare with the lidar technique. Results indicate reasonable agreement within the estimated uncertainties.

1. Introduction

1 Mineral dust is estimated to be the most abundant aerosol type in the atmosphere 2 (~half of the total global aerosol burden) (e.g. Textor et al., 2007; Choobari et al., 2014), with global emission between 1000 and 3000 Mt yr^{-1} (Zender et al., 2003; Zender et al., 3 2004; Shao et al., 2011). Mineral dust directly scatters and absorbs solar and infrared 4 5 radiation (Miller and Tegen, 1998), and impacts the optical properties of clouds (Ferek et 6 al., 2000; Rosenfeld et al., 2001; Creamen et al., 2013). In addition, mineral dust particles 7 can act as cloud condensation and ice nuclei (Twohy et al., 2009; Ansmann et al., 2009a; 8 DeMott and Prenni, 2010) and affect air quality (Fairlie et al., 2010). The high temporal and 9 spatial variability of dust particles and the complexity in their microphysical and optical 10 properties present a significant challenge to our understanding of how these particles impact the 11 environment.

12 Numerous field campaigns have been conducted to better characterize mineral dust 13 properties, e.g. the Saharan Mineral Dust Experiments SAMUM-1 and SAMUM-2 (Ansmann 14 et al., 2009b, 2011a and references therein) and the Saharan Aerosol Long-Range TRansport 15 and Aerosol Cloud interaction experiment SALTRACE (http://www.pa.op.dlr.de/saltrace/), 16 among others. However, the information on mineral dust properties is still quite scarce 17 (Formenti et al., 2011), even though many measurements worldwide have been made using 18 different approaches. Satellites are providing global coverage but the retrievals of particle 19 properties are still affected by large uncertainties (Levy et al., 2013). Moreover, the 20 interaction of dust particles with solar and terrestrial radiation is complex due to their 21 irregular shapes and variable refractive indices (Mishchenko et al., 1997). Because of this, 22 in the past years it has been difficult to develop accurate algorithms for the retrieval of dust 23 microphysical properties from optical measurements. Dubovik et al. (2006), one of the first 24 studies that addressed this problem, developed an algorithm that took into account the 25 scattering patterns of non-spherical particles and implemented an inversion method for column-integrated radiometric measurements in the AERONET (Aerosol Robotic Network) 26 27 network (http://aeronet.gsfc.nasa.gov/, Holben et al., 1998). Other approximations using 28 non-spherical models have also been proposed, e.g. Olmo et al. (2006), Valenzuela et al. 29 (2012a). However, due to the inherent characteristics of the sun-photometer only daytime 30 retrievals are possible. The Linear Estimation (LE) algorithm, introduced in Veselovskii et

31 al., (2012), allows retrieving column-integrated aerosol microphysical properties with high 32 temporal resolution (Kazadzis et al., 2014) during both day and nighttime taking advantage 33 of sun- and star-photometry measurements (Perez-Ramirez et al., 2015).

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Currently, information on the vertical distribution of the mineral dust can also be 35 retrieved by different techniques. Information of mineral dust on the vertical coordinate is 36 essential for understanding particle transport processes from regional to intercontinental 37 scales, to improve radiative forcing calculations and to analyze the influence of mineral 38 dust on cloud formation (Ansmann et al., 2008,2009,2011a; Seifert et al., 2010). Therefore, 39 advanced methods to characterize dust microphysical properties profiles from remote 40 sensors like lidar systems are crucial. In this framework, Chaikovsky et al. (2012, 2015), 41 developed the Lidar Radiometer Inversion Code (LIRIC) that can provide vertically-42 resolved profiles of aerosol microphysical properties during daytime by combining elastic 43 lidar measurements and column-integrated sun-photometer microphysical properties 44 retrieved from AERONET. Even though multiwavelength lidar measurements are required 45 at least in three wavelengths, the widespread use of multiwavelength elastic backscattered 46 lidar systems in networks such as EARLINET (European Aerosol Research Lidar Network, 47 Pappalardo et al., 2014) and LALINET (Latin American Lidar Network, Guerrero-Rascado 48 et al., 2014) provides enough global coverage. The availability of collocated simultaneous 49 AERONET measurements in most of these lidar stations widely expands the applicability 50 of LIRIC.

51 As in the case of the column-integrated microphysical properties, LIRIC retrievals are 52 also limited to daytime because of the use of the sun-photometer measurements. To 53 overcome this difficulty, the regularization technique implemented by Veselovskii et al., 54 (2010) can be used to retrieved nighttime aerosol microphysical profiles. The method by 55 Veselovskii et al. (2010) implemented the kernel functions of Dubovik et al. (2002) in a 56 regularization technique (Muller et al., 1999; Veselovskii et al., 2002) to obtain vertically-57 resolved dust properties. This approach was first implemented using data of the SAMUM 58 field campaign (Müller et al., 2013) and from measurements of long-range transport of dust 59 over Europe (e.g. Veselovskii et al., 2010; Papayannis et al., 2012). This technique has the 60 limitation that for the input data Raman or HSRL measurements are required. Because of the low signal-to-noise ratio in these systems, their use is mostly limited to nighttime. 61

62 Additionally, acquiring and maintaining Raman and HSRL systems is quite costly. Hence, 63 their implementation is not as widespread as elastic backscatter systems.

64 During the summer of 2011, a measurement campaign was launched at Granada 65 experimental station during a dust event in order to contribute to better characterize mineral dust properties. One of the main goals of this campaign (previously presented in Bravo-66 67 Aranda et al., 2015) was to use the synergy between different remote sensing techniques to 68 retrieve the evolution of microphysical properties of mineral dust during day and nighttime 69 at different levels. Results obtained for both column-integrated and vertical profiles of 70 mineral dust microphysical properties are presented in this study.

71 Measurements performed at Granada were coordinated with a flight of the CASA C-72 212-200 research aircraft above the station (Andrey et al., 2014; Bravo-Aranda et al., 2015) 73 on June 27, 2011. The aircraft was equipped with in-situ instrumentation that allowed us to 74 compare for the first time, up to the best of our knowledge, the microphysical properties 75 profiles retrieved with LIRIC with well-established simultaneous in-situ measurements. On 76 the other hand, the aircraft was equipped with a CAS-POL sonde measuring in-situ 77 depolarization data. With the measurements acquired by this sonde, we were seeking to 78 make for the very first time a qualitative comparison of an in-situ measurement of 79 depolarization ratio with that measured remotely by the lidar depolarization technique.

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2. Experimental site and instrumentation

81 2.1 Experimental site

82 The data from ground based-instrumentation were acquired at the Andalusian 83 Institute for Earth System Research (IISTA-CEAMA) located in the city of Granada 84 (37.16° N, 3.61° W, 680 m a.s.l.), (Lyamani et al., 2010; Titos et al., 2012; Valenzuela et 85 al., 2012b). Granada is a medium-size city in the South-East of Spain located in a natural 86 basin, delimited on the East by mountains with peaks up to 3000 m a.s.l. Air masses 87 affecting the area arrive mainly from the Atlantic Ocean, Central Europe, the 88 Mediterranean Basin and North-Africa (Valenzuela et al., 2012a; Perez-Ramirez et al., 89 2012a). The number of mineral dust events at the Granada station originating in North 90 Africa is quite high, especially during summer, with an occurrence of 45% of the days in

91 June, July and August (Valenzuela et al., 2012b). These events can transport particles at 92 altitudes of up to 5500 m a.s.l, not always affecting the surface level (Guerrero-Rascado et 93 al., 2008; Guerrero-Rascado et al., 2009; Navas-Guzmán et al., 2013). The experimental 94 site is also impacted by anthropogenic particles from local and regional aerosol sources 95 (Lyamani et al., 2008, 2010, 2012; Titos et al., 2014).

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2.2 Ground-based instrumentation

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2.2.1. Lidar system MULHACEN

98 The Raman-lidar MULHACEN (based on LR331D400, Raymetrics, Greece) used 99 for aerosol vertical-profiling is described in detail by Guerrero-Rascado et al., (2008; 2009) 100 and Navas-Guzmán et al., (2011). It employees a Nd:YAG laser that emits at three 101 different wavelengths (355, 532 and 1064 nm). The receiving system consists of detectors 102 that split the radiation according to the three elastic channels (355, 532 and 1064 nm), two 103 nitrogen Raman channels (387, 607 nm) and a water vapour Raman channel (408 nm). 104 These Raman measurements have sufficient signal-to-noise ratio only for night-time 105 detection. The system also measures depolarization of the returned signal at 532 nm (532-106 cross and 532-parallel detection channels) (Bravo-Aranda et al., 2013) for retrieving vertical profiles of the particle linear depolarization ratio (δ_{λ}^{P}) . The aerosol optical 107 108 properties profiles presented in this study were obtained by means of the Klett-Fernald 109 (Fernald et al., 1972; Fernald, 1984; Klett, 1981) algorithm during daytime and using the 110 Raman technique (Ansmann et al., 1990) at night. The depolarization profiles were retrieved according to the methodology described in Bravo-Aranda et al., 2013 and 111 112 Freudentaler et al., 2009. More details on the retrieval of the aerosol optical properties 113 profiles from the lidar data presented in this study can be found in Bravo-Aranda et al., 114 2015. The estimated uncertainties associated with the lidar signals are between $\pm 15\%$ and 20% for the aerosol particle backscatter coefficient, β^{aer}_{λ} , and $\pm 20\%$ for the aerosol particle 115 extinction coefficient, α^{aer}_{λ} . These estimates are based on the statistical uncertainties 116 retrieved with Monte Carlo techniques according to Pappalardo et al. (2004) and Guerrero-117 118 Rascado et al. (2008). The procedure described by Wandinger and Ansmann et al. (2002) to 119 correct the incomplete overlap of the system is applied to our data. The use of this overlap correction allows to obtain reliable β^{aer}_{λ} profiles at 355 and 532 nm down to 320 m above 120

121 the station (Navas-Guzmán et al., 2011); however, reliable data are obtained only from 122 ~1000 m above the station for α^{aer}_{λ} . The Raman lidar system is part of EARLINET and 123 currently is included in the ACTRIS2 (Aerosols, Clouds, and Trace gases Research 124 InfraStructure Network 2) European project (http://www.actris.net/).

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2.2.2. Sun-photometer CIMEL CE-318

126 Sun photometric measurements obtained at Granada are used to obtain column-127 integrated aerosol properties using a CIMEL CE-318. This instrument is included in the AERONET-RIMA network (Iberian Network for Aerosol Measurements, federated to 128 129 AERONET) (http://www.rima.uva.es/index.php/en/) since 2002 and is calibrated following the AERONET protocols. Details about the CIMEL sun-photometer can be found in 130 131 Holben et al., (1998), however a brief description is presented here. This instrument makes 132 direct sun measurements at 340, 380, 440, 500, 670, 870, 940 and 1020 nm and sky 133 radiance measurements at 440, 670, 870 and 1020 nm. The direct sun measurements are used to retrieve aerosol optical depth (τ_{λ}) at 340, 380, 440, 500, 675, 870 and 1020 nm. The 134 τ_{λ} uncertainties provided by AERONET are ± 0.02 for $\lambda < 400$ nm and ± 0.01 for $\lambda > 400$ 135 136 nm. Additionally, the spectral dependency of the τ_{λ} has been considered through the 137 Ångström exponent, AE(440–870), calculated in the range 440–870 nm. Also included in 138 the analysis are aerosol optical depths at 500 nm for fine mode (τ_{fine}) and for coarse mode (τ_{coarse}) as well as the fine mode fraction (η) (ratio of τ_{fine} to τ), determined using the spectral 139 140 de-convolution algorithm method developed by O'Neill et al. (2003). In addition, columnintegrated aerosol microphysical properties (size distribution, refractive index, volume 141 142 concentration, etc.) provided by the AERONET code are also used (Dubovik and King, 143 2000; Dubovik et al., 2002,2006). For the retrieval of the aerosol microphysical properties 144 both the direct sun and the sky radiance measurements are used. The reported size distribution retrieval uncertainties are $\pm 10-35\%$, for the size range 0.1 μ m < r < 7 μ m, and 145 146 outside this range they are as large as $\pm 80-100\%$. All the data used here are Level 1.5 data 147 obtained using the AERONET Version 2 algorithm. Only a small number AERONET 148 Level 2.0 were available due to the restrictions imposed by AERONET code ($\tau_{440nm} > 0.4$ 149 and solar zenith angle >50°). Therefore, AERONET Level 1.5 (cloud screened data with 150 pre- and post-calibrations applied) were used in this study, considering only those data that 151 fulfil the following conditions to assure their quality: $\tau_{440nm} > 0.2$ and solar zenith angle >50°. 152

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2.2.3. Star-photometer EXCALIBUR

154 EXCALIBUR provides τ_{λ} measurements from the stars' direct irradiance using 155 interferential filters at 380, 436, 500, 670, 880 and 1020 nm (nominal wavelengths) using a 156 CCD camera as detector during night-time. More details can be found in Perez-Ramirez et 157 al., (2008a, b). A regular calibration of the instrument is performed once a year at a high mountain location. Errors in the τ_{λ} are ± 0.02 for $\lambda < 800$ nm and ± 0.01 for $\lambda > 800$ nm 158 (Perez-Ramirez et al., 2011). Data are cloud-screened and quality assured over 30 minute 159 160 intervals (Perez-Ramirez et al., 2012b).

161 2.2.4. Aerodynamic Particle Sizer APS-3321

162 Auxiliary measurements of the particle size distribution and concentration at the surface 163 were performed with an aerodynamic particle sizer (APS-3321; TSI). This instrument is an optical particle counter that measures particle diameter and aerosol concentration, in real 164 165 time, in 52 nominal size bins in the aerodynamic diameter range 0.50-20 µm by determining the time-of-flight of individual particles in an accelerating flow field. The APS 166 can measure concentrations up to 1000 particles cm-3 at 0.5 and 10 µm, with coincidence 167 errors inferior to 5% and 10%, respectively. The minimum and maximum concentrations 168 169 that can be measured are 0.001 and 10 000 particles cm-3, respectively. For solid particles, counting efficiencies range from 85% to 99% (Volcken and Peters, 2003). The APS was 170 171 operated at flow rate of 5 l min-1 and with data averaging time of 5 min.

172 Air sampling for APS instrument was obtained from the top of a stainless steel tube, 20-cm diameter and 5-m length (Lyamani et al., 2008). The inlet was located about 15 m above 173 174 the ground surface. The measurements were performed without aerosol size cut-off and no 175 heating was applied to the sampled air.

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2.3 Aircraft in-situ instrumentation

Meteorological and aerosol particle measurements were made with instruments 177 178 mounted on the CASA C-212-200 research aircraft, operated by the Spanish National Institute of Aerospace Technology (INTA). Details on aircraft measurements can be found
in Andrey et al. (2014) and Bravo-Aranda et al. (2015). Here we give a short overview of
the instruments used for our study: the Cloud and Aerosol Spectrometer with Polarization
detection (CAS-POL) and the the Passive Cavity Aerosol Spectrometer (PCASP-100X).

183 The CAS-POL measures the light scattered by individual particles passing through a 184 focused, 680 nm polarized laser beam (Baumgardner et al., 2001). The equivalent optical 185 diameter (EOD), over a nominal size range from 0.6 to 50 µm, is derived from the light that 186 is collected over a 4-12° solid angle in the forward direction assuming that the particle is 187 spherical with a refractive index of 1.54 -i0.0 (McConnell et al., 2010). The backscattered 188 light is collected over the solid angle 168-176° and then separated with a beam splitter into 189 two components. One component is directed to a detector with a polarized filter to transmit 190 only scattered light that is perpendicular to the plane of polarization of the incident light. 191 This component is referred to as the "S" polarization. The other component is directed to a 192 detector with no filter so it measures the total backscattered light. In order to get the 193 backscattered light that is polarized in the same plane as the incident light, the parallel "P" polarization, the signal from the "S" polarization detector is subtracted from the total 194 195 backscattering detector.

The size derivation is well calibrated, using reference particles of polystyrene latex spheres (PSL) and glass beads (crown glass and boric acid) of monodispersed size and known refractive index at the wavelength of the CAS-POL. The total backscatter signals are also related to particle size with the same reference particles.

For the depolarization data, there are currently no reference particles that produce a known intensity of polarized scattered light. Hence, the "S" and derived "P" signals are presently used only qualitatively and the depolarization ratio, defined as the quotient "S" to "P", cannot be quantitatively compared with the same ratio derived from the lidar measurements. The procedure for setting the gains of the "S" and "P" detectors is explained in greater detail in appendix A. As will be amplified in the following sections, the CAS-POL depolarization ratios will be compared with those from the lidar vertical profiles in a qualitative way to evaluate the performance and potential of the in-situ depolarizationmeasurements with the CAS-POL, not absolute values are evaluated.

The measurement uncertainties are associated with the accuracy of determining the sample volume and the derivation of the EOD. The estimated uncertainty in the sample volume is $\pm 20\%$ and the sizing uncertainty ranges from 20-50% depending on the variation in the refractive index and the asphericity of the particles. The degree of asphericity and the orientation of the particle in the beam leads to variations of the depolarization ratio of 20-100%, depending on the aspect ratio of the dust (see Appendix A and Figure 9 for more detail).

The errors in the sample volume and size are propagated using the root sum squared (RSS) approach to estimate the uncertainty in the derived bulk parameters of number and volume concentrations and the median volume diameter. These estimated uncertainties are $\pm 20\%$, $\pm 90\%$ and $\pm 50\%$, respectively. A detailed description of the CAS-POL and associated uncertainties is given in Appendix A

The other instrument on the aircraft that measured aerosol size distributions was the Passive Cavity Aerosol Spectrometer (PCASP-100X) that provides aerosol size distributions in the 0.1- 3 μ m diameter range in 15 different bins. The measurements from the PCASP increased the size range provided by the CAS-POL measurements down to 0.1 μ m in our study. The measurement principle and uncertainties (20% for the size distributions) have been described in detail by Cai et al. (2013), Rosenberg et al. (2012) and references therein.

228 **3. Methodology: Retrievals of aerosol particles microphysical properties**

3.1. LIRIC code

The LIRIC algorithm is applied to lidar and sun-photometer data to retrieve aerosol particle microphysical properties profiles during daytime (Chaikovsky et al., 2008, 2012, 2015; Wagner et al., 2013; Granados-Muñoz et al., 2014, 2015; Binietoglou et al., 2015). This algorithm requires as input data the column-integrated optical and microphysical properties retrieved from AERONET code (Dubovik et al., 2002; Dubovik et al., 2006) and 235 measured lidar elastic backscattered signals at least at three different wavelengths (355, 236 532, and 1064 nm). The depolarization information from lidar data can optionally be used 237 in LIRIC. To perform the retrieval, an aerosol model, based on the AERONET code, which 238 assumes a mixture of randomly oriented spheroid and spherical particles defined by the 239 column-integrated volume concentrations of each mode, is used (Dubovik and King, 2000; 240 Dubovik et al., 2006). Subsequently, an iterative procedure based on the Levenberg-241 Marquardt method is applied. The combined lidar and sun-photometer information provides 242 volume concentration profiles for the fine and coarse modes, distinguishing also between 243 coarse spherical and coarse spheroid particles if depolarization information is considered. 244 The separation between the fine and the coarse modes is made by searching the radius 245 located at the minimum of the AERONET bimodal size distribution in the radii range 246 between 0.194 and 0.576 µm. Uncertainties of the volume concentration profiles related to 247 the different user-defined parameters used within LIRIC algorithm are usually below 15% 248 (Granados-Muñoz et al., 2014), proving LIRIC to be quite stable. A detailed analysis on the 249 uncertainty due to the input lidar and sun-photometer data uncertainties is still lacking, but 250 it can be estimated for our case. The uncertainty in the retrieved volume concentration is 251 related to the size distribution that is retrieved from AERONET. As discussed above, this 252 uncertainty can range from 10-35% if the particle radii are less than 7 µm but can be as 253 much as 100% if there is significant volume in particles larger than this (Dubovik et al., 254 2000). As will be illustrated below, measurements of the volume size distribution with the 255 CAS-POL show that there are particles larger than 7 μ m radius and in the dust layer they 256 can be larger than 30 µm. Hence, a reasonable estimate of the retrieved volume uncertainty 257 in our case is approximately 50%, less than the 90% uncertainty estimated for the volume 258 measured by the CAS-POL (Appendix A).

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3.2. Inversion of Raman lidar measurements to retrieve microphysical properties

The Raman lidar system provides α^{aer}_{λ} profiles at 355 and 532 nm, β^{aer}_{λ} at 355, 532, 260 1064 nm and the δ^{P}_{λ} at 532 nm. This $3\beta + 2\alpha + 1\delta$ data set was inverted to retrieve the aerosol 261 particle microphysical properties using the regularization approach, described in detail by 262

- Müller et al. (1999) and Veselovskii et al. (2002, 2004). 263

264 To account for mineral dust particles non-sphericity the model of randomly oriented 265 spheroids was used, as described in Veselovskii et al., (2010). Following this approach, the 266 minimum and maximum particle size are determined during the retrieval. A large number of inversion windows are tested by using the look up tables introduced in Veselovskii et el. 267 268 (2010), which comprise radii between 0.05 and 25 μm . The real and imaginary part of the refractive indices are varied in the range 1.45-1.55 for the real part and $5 \cdot 10^{-4}$ -0.01 for the 269 270 imaginary part. The particle volume concentration and effective radius in our case were 271 estimated with an uncertainty of about 50% and 25%, respectively. The real part of the 272 refractive index, m_r , was also estimated, with an uncertainty of ± 0.05 . A detailed 273 description on the procedure to calculate the uncertainties is included in Veselovskii et al., 274 (2010).

275 **3.3.** Inversion of τ_{λ} spectral measurements

276 In the present study, we used two different techniques to retrieve the columnintegrated aerosol microphysical properties. Firstly, data provided by AERONET 277 operational algorithm (version 2) were used to retrieve the aerosol microphysics during 278 279 daytime by inversion of τ_{λ} and sky radiances obtained from the sun-photometer. In addition to these AERONET aerosol retrievals, we inverted τ_{λ} spectral values with the LE algorithm 280 281 to get high-temporally resolved aerosol properties such as the integrated volume and 282 effective radius (Kazadzis et al., 2014). The LE algorithm is described in Veselovskii et al., 283 (2012) and more details are provided by Perez-Ramirez et al. (2015), where correction 284 functions to get accurate aerosol data close to that of the operational AERONET code were 285 introduced. The LE algorithm is applied in our study to both sun- and star-photometer data, 286 allowing us to retrieve column-integrated aerosol microphysical properties during day and 287 night.

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4 **Results and discussion**

Different models such as NAAPS (Navy Aerosol Analysis and Prediction System) (Christensen, 1997) and BSC-DREAM8b (Pérez et al., 2006a, b; Basart et al., 2012), forecast mineral dust above Granada station on the June 27, 2011. Based on these forecasts, a measurement campaign was launched using all the available instrumentation at the IISTA-CEAMA experimental station coordinated with a simultaneous flight of the CASA C-212-200 research aircraft (Bravo-Aranda et al., 2015). The flight took place around 10:30
UTC on 27 June whereas ground bases measurements started the night before to cover the
dust event during night and daytime. The aircraft performed a pseudo spiral track flying
from 1200 to 5200 m a.s.l. at a distance of approximately 8 km from the IISTA-CEAMA
station.

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4.1. Column-integrated aerosol microphysical properties

300 Columnar aerosol optical and microphysical properties from the sun and star photometry 301 measurements are shown in Figure 2. Figure 2a illustrates the temporal trends in the τ_{500nm} , 302 and the corresponding contributions of fine (τ_{fine}) and coarse (τ_{coarse}) modes using the 303 Spectral Deconvolution Algorithm (O'Neill et al., 2001a,b; 2003). The τ_{i} -related Ångström 304 exponent (AE(λ_1 - λ_2)) computed using the wavelengths between 440 and 870 nm (436 and 305 880 nm for star photometry) and the fine mode fraction (η) are shown in Figure 2b. As can be observed, there was a smooth temporal-evolution of the aerosol properties with small 306 307 variations. τ_{500nm} time series indicated that the aerosol load slightly varied during the 308 analysed period, with values ranging between 0.27 and 0.37. τ_{coarse} was significantly larger 309 than τ_{fine} during the end of the night and day-time measurements whereas τ_{fine} was almost 310 constant around 0.1 during the study period. The simultaneous increase of τ_{500nm} and τ_{coarse} 311 from midnight suggests an increase of the incoming mineral dust in the atmospheric column. However, during the first hours of the night, the fine mode had a relevant 312 313 contribution to the total τ_{500nm} . The AE ranged between 0.80 at night-time and 0.4 during 314 day-time, suggesting a clear predominance of coarse particles that can be associated with the transport of dust particles. However, the larger values obtained at the beginning of the 315 316 night period are higher than those acquired from AERONET measurements in the presence 317 of mineral dust (e.g. Dubovik et al., 2002), and thus the contribution of fine particles to the 318 aerosol mixture is also considerable.

The evolution of the effective radius, r_{eff} and the column volume concentration, V, obtained both by LE and AERONET retrievals during day-time and by LE during nighttime for the 27th of June 2011 is shown in Figure 2c and Figure 2d. Good agreement between AERONET and LE retrievals was obtained during daytime corroborating the good performance of the LE method. In addition, continuity between the star- and sunphotometer measurements was observed, which shows the utility of combining both 325 instruments to perform a continuous monitoring of aerosol properties. As observed in Figures 2c-d, both r_{eff} and V were slightly lower during the night compared to daytime. The 326 r_{eff} values were around 0.3 -0.4 μ m during night-time and almost constant around 0.5 μ m 327 during daytime. These values suggest a predominance of coarse particles in the atmospheric 328 column. The V values were increasing from 0.15 $\mu m^3/\mu m^2$ during the night up to 0.25 329 $\mu m^3/\mu m^2$ in the early morning and then they decreased again down to 0.1 $\mu m^3/\mu m^2$ at noon. 330 331 The changes observed in aerosol properties suggest a mixture of different aerosol types 332 changing from night to day.

- 333
- 334

[Figure 2]

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4.2. Mineral dust microphysical properties profiles

337 4.2.1. Daytime retrieval of mineral dust microphysical properties profiles with 338 LIRIC

A comprehensive analysis of the optical properties profiles retrieved during daytime by Klett-Fernald (Fernald et al., 1972; Fernald, 1984; Klett, 1981) method and during nighttime by Raman method (Ansmann et al., 1990) was previously presented by Bravo-Aranda et al. (2015) and is out of the scope of our study. However, some of the optical profiles (Figure 3) and a brief discussion are included here to contextualize the mineral dust event in a way that we have enough information to properly discuss and understand the evolution of the microphysical properties profiles.

346 Figure 3 shows β_{532nm}^{aer} , β -AE(355-532nm), which is the backscatter related Ångström exponent between 355 and 532nm and δ^{P}_{532nm} , obtained during day-time from lidar elastic 347 348 measurements. According to this figure and Bravo-Aranda et al. (2015) two different layers 349 were clearly distinguished. The first layer was in the height range between 3000 and 5000 m a.s.l., where β_{532nm}^{aer} values were decreasing during the morning and β -AE(355-532nm) 350 profiles suggest the predominance of coarse particles with values close to zero during the 351 352 analysed period. δ_{532nm}^{P} values ranged between 0.23 and 0.28, suggesting an important contribution of non-spherical particles (Gross et al., 2011). However, below 2000 m a.s.l., 353

 β -AE(355-532nm) was decreasing from 2 in the early morning down to ~0.5 around midday. At the same time, δ^{P}_{532nm} was increasing from 0.08 up to 0.20, suggesting the influence of anthropogenic particles from local origin that were mixed with the mineral dust due to convective processes within the planetary mixing layer (Bravo-Aranda et al., 2015).

359

[Figure 3]

360 LIRIC retrievals obtained for the morning of the June 27, 2011 are shown in Figure 4. 361 A clear predominance of the coarse spheroid mode is observed from the surface up to 5000 m a.s.l., as expected for mineral dust events. In addition, a decrease of the total volume 362 concentration values occurred throughout the morning, as expected from the observed 363 decrease in β_{532nm}^{aer} (Figure 3) and the decrease in the integrated volume concentration V 364 (Figure 2d). A maximum peak in the volume concentration of the coarse spheroid mode 365 366 was observed between 4000 and 4500 m a.s.l., in coincidence with the maximum in β_{532nm}^{aer} profiles, indicating the presence of a distinguished aerosol layer at this height. It is worth 367 368 noting that fine particles were also observed during the different analysis periods as indicated by the volume concentration profiles, but in low concentrations (~6 μ m³/cm³). 369 370 The fact that the profiles of the fine mode volume concentration show a maximum peak at 371 the same height as the coarse spheroid mode suggests that these fine particles could have been advected with the mineral dust. Therefore, they might correspond to the fine mode of 372 373 mineral dust (Mamouri and Ansmann, 2014) or even anthropogenic aerosol advected from 374 the industrial areas of North of Africa together with the mineral dust (Rodríguez et al., 2011; Valenzuela et al., 2012c), as indicated by the backward trajectories analysis 375 376 performed with HYSPLIT model using GDAS meteorological database 377 (http://ready.arl.noaa.gov/HYSPLIT traj.php, not shown). However, Granados-Muñoz et al., (2014) suggested that it might be an artifact introduced by LIRIC due to problems with 378 379 the incomplete overlap or in cases of non-homogeneous layering distributing the 380 contribution of local pollution at high altitudes. Additional datasets would be required to 381 clarify the origin of these fine particles in our study.

The coarse spherical mode volume concentration slightly increased during the morning reaching its maximum values (~ 7 μ m³/cm³) between 11:15 and 11:45 UTC. This slight increase in the volume concentration of the coarse spherical mode (from 0 up to 5 μ m³/cm³) around 4500 m a.s.l. was in agreement with a slight decrease in δ^{P}_{532nm} values (from 0.28 to 0.23) in the same height range, indicating a larger contribution of spherical particles and corroborating once again the coherence between the aerosol optical and microphysical properties profiles obtained.

389 Even though layering was observed in the optical properties profiles (Figure 3), with 390 mineral dust in the upper layer between 3000 and 5000 m a.s.l. and anthropogenic particles 391 in the lower part of the troposphere, LIRIC results showed much more homogeneous 392 profiles with no such layering. As explained in previous studies (Granados-Muñoz et al., 393 2014) the assumption of height-independent AERONET retrieved properties (i. e., 394 refractive index, size distribution, sphericity, etc.) in LIRIC leads to much more vertically 395 homogeneous volume concentration profiles regarding the distribution of the modes than 396 those observed on the lidar optical properties. The layering observed in the optical 397 properties profiles (Figure 3) indicated the presence of mineral dust in the upper layer 398 between 3000 and 5000 m a.s.l. and the presence of anthropogenic particles in the lower 399 part of the troposphere that was not detected by LIRIC. This can be explained by the fact that LIRIC assumes several AERONET retrieved properties (i. e., refractive index, size 400 401 distribution, sphericity, etc.) as height-independent and therefore LIRIC results are much 402 more vertically homogeneous regarding the distribution of the modes than those retrieved 403 from the lidar data with Klett-Fernald algorithm. A combined inversion of lidar and sun-404 photometer data without such assumptions, as the one proposed in GARRLIC (Generalized 405 Aerosol Retrieval from Radiometer and Lidar Combined data) (Lopatin et al., 2013), might 406 provide more accurate results regarding the vertical distribution of the aerosol properties, although GARRLIC is out of the scope of this work. Meanwhile, the interpretation and 407 408 analysis of the LIRIC-retrieved volume concentration profiles needs to be carefully made 409 when the atmospheric aerosol layers comprise different aerosol types, as shown in 410 Granados-Muñoz et al. (2014).

411 [Figure 4]

412 4.2.2. Night-time retrieval of mineral dust microphysical properties profiles with the 413 regularization technique

Figure 5 shows the aerosol optical and microphysical properties profiles retrieved for the first period of lidar measurements (00:00-01:00 UTC). The Raman $3\beta+2\alpha+\delta$ profiles (Figures 5a, b and c) were inverted to retrieve vertical profiles of volume concentration, r_{eff} and the real part of the refractive index (Figures 5d and e) using the regularization approach by Veselovskii et al., (2010).

During the night period, values of β^{aer}_{532nm} (Figure 5b) were similar to those observed 419 420 at daytime (Figure 3), especially in the early morning (06:15-06:45 UTC). As during the 421 morning, two differentiated regions with different aerosol types were observed at night. Namely there were mineral dust particles above 2250 m a.s.l. ($\delta^{P}_{532nm} \sim 0.25$) and a mixing 422 423 of local anthropogenic aerosol with mineral dust below this height. Even though, similar 424 layers were observed during both analysed periods, the vertical structure of the aerosol 425 layer was different, indicating an evolution of the aerosol vertical structure during the night. 426 As observed when comparing Figure 5 and Figure 3, two peaks were detected during 427 nighttime around 3200 and 4200 m a.s.l. whereas a more homogeneous structure was 428 observed during daytime, with only one maximum at 4100 m a.s.l. Besides, a significant 429 increase of the anthropogenic pollution in the lower layer was observed during the morning.

430 Mineral dust microphysical properties profiles were obtained with the regularization 431 technique using the optical data, averaged in 250-m layers for the heights between 2000 432 and 3600 m a.s.l. Above that height averaging was done in 1000-m layers. The overlap effects prevented using extinction data below 2000 m and therefore inversion of 433 434 microphysical properties was not possible below this height. The retrieved profiles are shown in Figures 5e, 5e and 5f. As we can observe in Figure 5d, r_{eff} values varied in the 435 436 different layers. The highest values, ~1.76 µm, were found in the layer between 2500 and 437 3500 m a.s.l., which is the mineral dust layer. The smallest values, $\sim 0.53 \mu m$, were found in 438 the layer where the anthropogenic pollution was mixed with the mineral dust, below 2300 439 m a.s.l. layer . Values around 4200 m a.s.l. (~ 1.1 μ m) were still larger than those in the 440 lower layer, but lower than in the maximum around 3200 m a.s.l. This decrease in the 441 radius with height can be due to an artifact of the algorithm because of the larger averaging

442 height interval used in the upper part of the profiles and not necessarily related to a 443 decrease in the size of the particles. Volume concentration values were also larger for the 444 dust layer than for the one with mixed particles. This fact can be explained by both the volume concentration dependence on particle radius and also the larger aerosol load in this 445 layer (as observed in the layering structure of β_{λ}^{aer} in Figure 5b). Two regions were also 446 447 clearly distinguished in the m_r profile obtained with the regularization technique. The m_r 448 values were larger in the upper part of the profile, corresponding to the mineral dust layer 449 and lower values were obtained in the lower part of the profile, due to the presence of the mixture of mineral dust and anthropogenic particles. The mean value of m_r in the profile 450 451 was 1.55±0.05. A summary with the values of the different aerosol properties in both 452 detected layers is shown in Table 1.

453 Finally, figure 5f shows the volume size distributions for different altitudes. At the 454 surface level, the APS was used to obtain the hourly average volume size distribution in the 455 aerodynamic radius range 0.25-10 µm for the period 00:00 to 01:00. The volume size 456 distributions at higher altitudes were obtained by the regularization technique. In addition, the closer in time column-integrated AERONET size distribution (June 26, 2011 at 18:15 457 458 UTC) was included for comparison. From both APS and lidar measurements, it was observed a clear increase in the coarse mode radius with height, as the location of the 459 maximum is displaced towards larger radii, in agreement with r_{eff} profiles. There was also 460 461 an increase of the aerosol load since the maximum in the volume concentration strongly 462 increased with height. Fine mode was almost insignificant in the different layers. However, 463 the AERONET column-integrated distribution showed a small contribution of the fine 464 mode aerosol particles and also the coarse mode radius shifted to smaller values. The low 465 column-integrated values of r_{eff} obtained with the LE algorithm in Figure 2 together with 466 the AERONET distribution suggest an important contribution of fine particles in the 467 region below 2 km a.s.l. during the analysed period, not observed by the lidar because of 468 the incomplete overlap. This contribution of fine particles also explains the difference in the 469 real part of the refractive index between the closest in time AERONET retrieval, which was 470 1.45, and the value obtained with the regularization technique (1.55). The limitations of the 471 APS to measure fine mode particles at the surface do not allow to confirm this, but in-situ 472 measurements presented by Bravo-Aranda et al. (2015) also pointed in this direction.

[Figure 5]

474

475 4.3. Comparison of airborne in-situ measurements and ground-based retrieved 476 profiles

477 **4.3.1.** Comparison of the volume concentration profiles

478 Figure 6a shows the vertical profiles of the total volume concentration profiles obtained 479 with LIRIC algorithm and those obtained from the aircraft in-situ measurements. The 480 horizontal bars on the LIRIC profile (blue curve) are the estimated uncertainties of $\pm 50\%$. 481 The horizontal bars on the aircraft profile (black curve) are the standard deviation about the 482 mean value. For the retrieval of the volume concentration profiles from the in-situ aircraft 483 measurements, a refractive index of 1.54-i0.0 (McConnell et al., 2010) was assumed and 484 Mie theory was applied assuming spherical aerosol particles as explained by Andrey et al. 485 (2014) and discussed in Appendix A. This refractive index value is very similar to the one 486 obtained with the regularization technique for the night-time measurements, which was 487 1.55±0.05. Therefore, the assumption of the refractive index is not expected to introduce 488 large uncertainties. The agreement in the vertical structure between the aircraft and LIRIC 489 volume concentration profiles, with respect to where the peak values are located, is quite 490 good considering that the aircraft profiles are horizontally displaced from those by the lidar 491 by approximately 8 km. Similar layering was detected with both LIRIC and the airborne 492 data, distinguishing two maximum peaks around 3500 and 4200 km a.s.l. The geometrical 493 thickness of the different layers observed was also very similar for both LIRIC and the 494 aircraft data. The slight difference between the height of the first maximum obtained by 495 aircraft in-situ measurements (3450 m a.s.l.) and LIRIC algorithm (3250 m a.s.l.) can easily 496 be explained taking into account the differences between both techniques and the horizontal 497 distance between the lidar and the aircraft measurements.

498 Regarding the volume concentration values, the differences are within the expected 499 uncertainties and natural variability, marked by the horizontal bars in Figure 6a, revealing 500 quite promising results. In general, average in situ values exceed those from LIRIC by less 501 than 20 μ m³/cm³, except for the concentration maximum between 3200 and 3500 m

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502 a.s.l.The difference between the profiles reveals a likely underestimation of the volume 503 concentration profiles retrieved by LIRIC by some fraction that depends on the amount of 504 particle volume in sizes larger than 15 µm (in radius) not considered in AERONET. However, this underestimation does not completely explain the observed differences. The 505 506 intrinsic differences between both techniques and the differences in sampling can also lead 507 to discrepancies between the profiles. This comparison analysis raises the question as to 508 which technique is providing the more accurate value. Both techniques rely on assumptions 509 that could potentially bias the results. In the discussion of the CAS-POL uncertainties in 510 Appendix A, the two most important assumptions are related to the refractive index and 511 shape of the particles. If the refractive index of the particles differ from what is assumed, 512 this can bias the sizing either smaller or larger, but no more than 20%. The assumption that 513 the particles are spherical will bias the sizes, and hence the volumes, perhaps even as much 514 as a factor of two if the dust particles are very aspherical. This could explain the 515 discrepancies between the two techniques over most of the profiles where the differences 516 are within a factor of two. Countering that argument, however, is the comparison of the size 517 distributions measured with the combined PCASP-100X and the CAS-POL and those 518 derived from AERONET. Figure 7 shows average volume concentration size distributions, 519 measured with the PCASP and CAS-POL, in the five altitude ranges labelled (1) through 520 (5) in Figure 6a. The size distribution provided by AERONET at 10:19 UTC is also included in the plot (multiplied by a factor of 10^3) on order to facilitate the comparison 521 between the aircraft and the column-integrated data. When compared, the aircraft and 522 523 AERONET show a very similar shape for the size distribution, e.g. the peak volume falls at 524 9 µm EOD. This suggests that the larger volume reported from the CAS-POL is not a result 525 of oversizing.

526

[Figure 6]

[Figure 7]

527

528 **4.3.2.** Comparison of the depolarization ratio profiles

529 The interest of this comparison relies on the fact that in situ measurements of 530 depolarization ratios from aircraft have never been published before, so the results 531 presented here are the first opportunity to compare profiles from aircraft with those derived 532 from remote sensors and analyse the potential of aircraft depolarization measurements. The 533 CAS-POL averaged depolarization ratio at a wavelength of 680 nm is compared with the 534 δ^{P}_{532nm} profile retrieved at 10:30 UTC from the lidar data (Figure 6b). The horizontal bars on the lidar profile (blue curve) show the estimated uncertainties while the bars on the 535 536 CAS-POL values (black curve) are the standard deviations about the average. The 537 polarization ratios from the CAS-POL are calculated as the summed perpendicular-538 polarization values divided by the summed parallel-polarization values. In general, we 539 observe reasonable agreement in the vertical distribution of the depolarization derived from lidar δ^{P}_{532nn} profiles and CAS-POL. Given that the polarization measurements from the 540 541 CAS-POL are not calibrated, as noted previously, we are interested in comparing if both 542 techniques show similar, relative changes in the depolarization ratio with altitude. In order 543 to better compare these trends, in Figure 6c we have normalized the depolarization ratios at 544 each level by their maximum ratios, 0.26 and 0.34 for the lidar and CAS-POL techniques, 545 respectively. Both methods are able to clearly distinguish between the dust layer in the 546 upper part and the mixed dust with anthropogenic pollution in the lower part of the profiles. 547 Some discrepancies are still observed in the lower layer of the profiles and in the region 548 around 4.75 km, but to explain them it is worthy to consider that the CAS-POL is 549 measuring instantaneous values whereas the lidar data correspond to 30-min averaged 550 profiles. Taking into account the large variability observed at these regions throughout the 551 morning in Figure 3 and the depolarization uncertainties for both datasets, differences are 552 within the expected values. According to these results, calibrated profiles provided by the 553 CAS-POL using laboratory measurements will provide very valuable information on the 554 aerosol vertical structure regarding depolarization. Depolarization information retrieved 555 from aircraft measurements is currently very limited and the use of the CAS-POL offers 556 promising results at this respect in the near future.

557 5. Summary

558 An exhaustive measurement campaign was performed June 27, 2011 at Granada 559 station during a Saharan dust outbreak. One of the main goals of the campaign was to 560 analyse the evolution of mineral dust microphysical properties by using the synergy of different techniques. The campaign was coordinated with an aircraft research flight to beable to perform some comparisons based on completely independent datasets.

The different methods used for the retrieval of the microphysical properties were LIRIC and the regularization method, to retrieve vertical profiles with lidar data during both day and nighttime, and the AERONET code and the LE technique, to retrieve columnintegrated properties from sun- and star-photometry. To our knowledge this is the first time that these techniques are processed and compared together to retrieve continuous microphysical properties for day and night during a mineral dust event.

569 Results indicated that the night-to-day evolution of the aerosol optical depth at 500 570 nm and the Ångström exponent retrieved using star and sun-photometer data followed a 571 smooth behaviour explained by the natural variability of the aerosol particle population. 572 The aerosol optical depth at 500 nm ranged between 0.27 and 0.37 during the analysis 573 period, while the Ångström exponent varied from 0.80 at night-time to 0.4 during the 574 daytime. Column-integrated microphysical properties retrieved with the operational 575 AERONET and Linear Estimation codes were in good agreement. Moreover, the Linear 576 Estimation allowed the retrieval of the dust microphysical properties during night-time 577 using the star-photometer data, which was not possible with the AERONET code. A 578 smooth temporal-evolution was observed, with effective radius ranging between 0.3 and 0.5 579 μm. The derived Ångström exponent and effective radius suggest a predominance of coarse 580 particles but with some contribution by the fine mode particles. The fine contribution was quite significant during some periods, especially during the night period between 00:00 and 581 582 01:00 UTC. For this period, lidar measurements were performed and combined vertical and 583 column-integrated information revealed that most of the aerosol fine load was located 584 below 2 km a.s.l., not affecting the mineral dust layer.

In the vertical coordinate, two aerosol layers were observed during both day and night: a lower layer corresponding to a mixture of anthropogenic aerosol with mineral dust and a higher layer corresponding to pure mineral dust. As expected in presence of mineral dust, the LIRIC analysis during daytime indicated a clear predominance of the coarse spheroid mode, with average volume concentrations around 30 μ m³/cm³, suggesting a medium intensity dust event. Fine particles were observed in lower concentrations (~6 μ m³/cm³) at high altitudes that could be associated to anthropogenic particles originated, 592 along with the dust particles, in the industrial areas in the North of Africa or they might 593 correspond to the mineral dust fine mode. However, they could also be related to a LIRIC 594 artifact that distributes the local anthropogenic aerosol at higher altitudes. Results obtained 595 with LIRIC are in agreement with the results obtained from the analysis of the optical 596 properties retrieved from the lidar data in the upper layer corresponding to mineral dust. 597 However, there is some disagreement in the lowermost part of the profiles, below 2000 m 598 a.s.l., which corresponds to the mixture of mineral dust with anthropogenic aerosol. These 599 discrepancies are mainly related to the assumption of height-independent aerosol properties 600 in LIRIC and indicate that LIRIC profiles need to be carefully interpreted in cases of non-601 homogeneous aerosol layers.

602 The retrieval of the total volume concentration profile at night (with the regularization technique) indicated a decrease in the volume concentration from maximum 603 604 values of 75 μ m³/cm³ to 50 μ m³/cm³ and changes in the aerosol vertical structure when compared to LIRIC retrieved volume concentration profiles in the early morning. 605 606 Nonetheless, column integrated values showed small changes in the aerosol properties, 607 remaining almost constant during the night. These results underscore the need for vertically 608 resolved measurements to adequately monitor the evolution of the aerosol properties. The 609 discrepancies between regularization and LIRIC results are mainly explained by the natural 610 variability of the aerosol during the night. Because of this temporal variability, results 611 obtained from the two different approaches are not comparable in absolute terms. However, 612 from our analysis we can conclude that the combined use of LIRIC and the regularization 613 technique improves our capability for evaluating the evolution of microphysical properties 614 profiles during night- and day-time.

For a direct intercomparison of LIRIC with the regularization technique more simultaneous datasets are required. Advances in vibrational Raman will allow measurement of the extinction coefficients during the daytime (e.g. Brocard et al., 2013) and measurements from HSRL systems could also help in this aspect. On the other hand, advances in techniques to retrieve microphysical properties from star photometry will allow the use of LIRIC during night-. Improvements in the application of the Linear Estimation method also could be addressed in this way. Such approaches are essential both to study the temporal evolution of aerosol microphysical profiles and also to define strategies forevaluating satellite products.

624 In the last part of the study, profiles of linear particle depolarization retrieved with the 625 lidar data and volume concentration retrieved with LIRIC were compared with CASA C-626 212-200 aircraft measurements. The vertical profiles of the volume concentration, retrieved 627 from the combined PCASP-100X and CAS-POL size distributions and LIRIC retrievals 628 showed good agreement regarding the vertical distribution of the aerosol, with maxima 629 3200-3500 and 4100-4400 m identified by both techniques and discrepancies between the 630 volume concentration values were within the expected uncertainties. Discrepancies were 631 generally below 20 μ m³/cm³ Results from the comparison suggest that the volume 632 concentrations from the CAS-POL are likely overestimates due to the asphericity of the 633 dust particles but the LIRIC derived values are underestimates because of the presence of 634 particles with equivalent optical diameters larger than 20 µm.

The first-ever published dataset of depolarization data retrieve with the CAS-POL instrument on board a research aircraft was also presented in this study and qualitatively compared with the lidar retrieved depolarization profiles. The two separate techniques showed roughly comparable results with respect to the depolarization ratios in the dust layer compared to the layers below and above.

640

6. Conclusions and future outlook

641 The synergy of instruments and different remote-sensing techniques presented in this 642 study reveals the current existing potential to obtain a complete characterization of the 643 aerosol properties using a combination of different measurements obtaining promising 644 results. However, the comparison between the different ground-based remote sensing 645 techniques and with independent aircraft in-situ data also points out to the limitations of the 646 different approaches used here and the need for reducing the uncertainty of the measured 647 and retrieved aerosol properties in order to obtain more reliable and accurate aerosol 648 properties databases. Future efforts should aim in this direction and validation campaigns and studies using independent datasets are crucial to achieve this goal. 649

650

Appendix A: CAS-POL Measurement Principles and Uncertainties

651 The CAS-POL measures the light that is scattered by individual particles that pass 652 through a focused, 680 nm, polarized laser beam (Baumgardner et al., 2001). Optical 653 components are positioned to collect some of the light that is scattered into a forward cone from 4-12° and a backward cone from 168-176°. The backward scattered light is divided 654 into two components: one that is measured with a detector that is behind an optical filter 655 656 that passes only the scattered light with polarization that is perpendicular to the polarization 657 of the incident light and another detector with no filter (Glen and Brooks, 2013). The 658 equivalent optical diameter (EOD) of each particle in the nominal size range from 0.6 to 50 µm is derived from the peak intensity of the collected forward scattered light using Mie 659 660 theory (Mie, 1908) and an assumption of particle sphericity at a known refractive index. 661 The terminology, EOD, is used here to underscore that ambient aerosol particles are only 662 spherical if they are liquid or have a liquid coating and the refractive index can vary over a 663 wide range. Hence, the EOD refers to the size of a spherical particle with known refractive 664 index that would have scattered the equivalent intensity of light.

665 The three signals, forward scattering (FS), backward scattering (BS) and polarized 666 (POL) provide three pieces of information that, from a relative perspective, can 667 differentiate regions of air masses that have different particle characteristics. For the current 668 study, FS is used to derive the EOD, and BS and POL are used to calculate the average 669 depolarization ratios defined in the main body of the manuscript. The interaction of the 670 linearly polarized laser radiation with a particle leads to scattered light with some of the 671 polarization in the same plane as the incident light and some fraction at a perpendicular plane (Nicolet et al., 2007, 2010, 2012; Schnaiter et al., 2012; Baumgardner et al., 2014). In 672 673 laboratory studies, Glen and Brooks (2013) showed that the relationships between the FS, 674 BS and POL signals were related to the type of dust, i.e. hematite, quartz and zeolite could 675 be clearly differentiated by comparing the three signals. In the analysis of CAS-POL 676 measurements in the current study, we will only report the depolarization ratio (defined 677 below).

The measurement uncertainties are associated with the accuracy of determining the sample volume, the derivation of the EOD and the depolarization ratio. The estimated uncertainty in the sample volume is $\pm 20\%$ primarily due to the optical technique used to 681 qualify particles within the beam (Baumgardner et al., 2001). The sizing uncertainty ranges 682 from 20-50% and depends on the variation in the refractive index and the asphericity of the 683 particles. Figure 8a shows the theoretical scattering cross section (Mie, 1908) of spherical 684 particles as a function of diameter for three refractive indices and the collection angles of 685 the CAS-POL. Only a small size range is shown in order to illustrate the potential magnitude of uncertainty. We can see that particles with different sizes have the same 686 687 scattering cross section. The scattering intensity is directly proportional to the particle 688 optical cross section; hence, when we measure scattering of this magnitude with the 689 instrument, we do not know if the particle was a water droplet with 1.25, 2.0 or 2.5 µm 690 EOD, a salt particle (NaCl) with a 1.75 or 2.5 µm EOD or a sulfate particle (refractive 691 index of 1.44) with an EOD of 2.5 µm. The average uncertainty due to deriving an EOD of 692 unknown refractive index was estimated by analyzing the variation in size when assuming a 693 particle had a refractive index of 1.48 then finding what particle size would have the same 694 scattering cross section if it actually had a different refractive index. The calculation was carried out over a range of refractive indices from 1.33 (water) to 1.60 (some types of 695 696 organics). Figure 8b presents the results of this evaluation where the ordinate is the average 697 derived value for the given size at a refractive index of 1.48 and the vertical bars are one 698 standard deviation around the mean. The red lines are the one-to-one (solid) and $\pm 20\%$ 699 (dotted) around the one to one. From this figure we see that the uncertainty falls within 700 $\pm 20\%$ of the assumed values except for the EOD between 2 and 4 μ m. Hence, the 701 uncertainty due to the refractive index variations is on average $\pm 20\%$.

702 The uncertainties that are related to the asphericity of a particle are more difficult to 703 estimate since they are dependent on the complexity of the morphology, the degree of 704 asphericity and the orientation of the particle when it passes through the beam. Borrmann et 705 al. (2000) applied T-matrix theory to estimate the amount of undersizing of prolate and oblate spheroids dependent on their aspect ratios and demonstrated that the Forward 706 707 Scattering Spectrometer Probe (FSSP), the predecessor to the CAS, could undersize ice 708 crystals as much as 50%. Single particle light scattering probes like the FSSP and CAS 709 measure the optical size of a particle. Converting this size to a physical volume requires an 710 assumption of the shape, refractive index and density of the particle. Given the complexity 711 of the morphology of dust, it is beyond the scope of this paper to evaluate the relationship

between light scattering and particle morphology in order to assess the uncertainty in sizing

- related to the properties of dust. Based on the simple analysis conducted by Borrmann et al.
- (2000), we use $\pm 50\%$ as a conservative estimate for the uncertainty in sizing due to shape.

The errors in the sample volume and size are propagated using the root sum squared (RSS) approach to estimate the uncertainty in the derived bulk parameters of number and volume concentrations and the median volume diameter. These estimated uncertainties are $\pm 20\%$, $\pm 90\%$ and $\pm 50\%$, respectively.

719 The change in polarization caused by a particle is also dependent on aspect ratio and 720 orientation. Volten et al. (2001) measured the light scattering by Sahara dust particles, 721 among others, at angles ranging from 5-173°. In the angle range of the CAS-POL the 722 depolarization ratio varied from 0.2 to 0.5. In recent laboratory measurements were made 723 with a CAS-POL (not the one used in the study reported in this paper) of various spherical 724 and aspherical particles. The results from those measurements are summarized in Figure 9 where the average depolarization ratio is shown as a function of the sum of the S and P 725 726 polarization signals. The horizontal and vertical bars show the standard deviations about the 727 average values. From these very preliminary results we can see that the depolarization ratio 728 can vary as much as a factor of two, or 100%. As seen in the results of measurements in the 729 dust layer (Figure 6), the variation is much less than this.

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Figure captions:

Figure 1. Time series of the RCS for the period 00:00-01:00 UTC and 06:45-12:10 UTC on the 27th of June 2011.

Figure 2: Night-to-day temporal evolution of a) aerosol optical depth (τ_{λ}) including also its separation between fine (τ_{fine}) and coarse (τ_{coarse}) mode at 500 nm, b) Ångström exponent (AE) between 440 and 870 nm (436-880 nm for star photometry) and fine mode fraction η c) effective radius (r_{eff}) and d) column integrated volume concentration (V) on the 27th of June 2011.

Figure 3: Profiles of a) aerosol particle backscatter at 532 nm (β_{532nm}^{aer}) b) backscatterrelated Ångström exponent between 355 and 532 nm (β -AE(355-532 nm)) and c) linear particle depolarization ratio at 532 nm (δ_{532nm}^{P}) retrieved from lidar elastic measurements at different hours between 07:30 and 11:30 UTC on the morning of the 27th of June 2011.

Figure 4. Volume concentration profiles of the fine, coarse spherical and coarse spheroid modes obtained with LIRIC from 30-min averaged lidar data for different periods on the 27th of June 2011. The error bars represent the uncertainty associated to the selection of the user-defined input parameters in LIRIC, obtained as indicated in Granados-Muñoz et al., (2014)

Figure 5. a) β_{532nm}^{aer} and α_{532nm}^{aer} retrieved with Raman technique and b) derived β -AE(355-532nm) and δ_{532nm}^{P} at 00:00-01:00 UTC on the 27th of June 2011. c) Total volume concentration, r_{eff} and d) m_r retrieved for the same period with the regularization technique applied to the $3\beta + 2\alpha + 1\delta$ lidar data and e) Volume size distributions retrieved at different height levels using the APS (surface level) and the regularization technique applied to the $3\beta + 2\alpha + 1\delta$ lidar data at 00:00-01:00 UTC on the 27th of June 2011. Column-integrated AERONET size distribution corresponding to the 26th of June 2011 at 18:15 UTC is included (bottom).

Figure 6. a) 30-min averaged volume concentrations retrieved with LIRIC (blue line) centered at 10:30 UTC and the combined PCASP-100X and CAS-POL concentrations (black line) retrieved during the flight ascent on the 27th of June 2011. The blue horizontal bars are the estimated uncertainy in the LIRIC values and the black horizontal bars are the

standard deviation about the average aircraft values b) Depolarization ratio retrieved from the CAS-POL measurements (black line) and δ^{P}_{532nm} retrieved from the lidar data (blue line) between 10:30 and 11:00 UTC Horizontal bars show the estimated uncertainties from the lidar (blue) and the satandard deviation about the average CAS-POL values (black). c) Depolarization ratios normalized to the maximum values derived from the lidar (blue) and CAS-POL (black).

Figure 7. Size distributions of volume concentrations derived from the combined PCAP-100X and CAS-POL measurements averaged over the five altitude ranges labeled in Figure 6a.

Figure 8. a) theoretical scattering cross section (Mie, 1908) of spherical particles as a function of diameter for three refractive indices and the collection angles of the CAS-POL. Only a small size range is shown in order to illustrate the potential magnitude of uncertainty. The dashed lines show that particles with four different sizes have the same scattering cross section. b) Average retrieved diameter for a given size at a refractive index of 1.48. Vertical bars represent one standard deviation around the mean. The red lines are the one-to-one (solid) and $\pm 20\%$ (dotted) around the one to one.

Figure 9. The average depolarization ratio as a function of the total backscattering (S pus P polarization) derived from CAS-POL measurements in the laboratory of a number of different types of particles: calibration particles, labeled by their nominal size, nebulized water droplets, Arizona Test Dust (ATD), volcanic ash, ice crystals and pollen.

<u>Tables</u>

	β^{aer}_{532nm} (Mm ⁻¹ ·sr ⁻¹)	δ^{P}_{532nm}	β-AE(355-532nm)	r _{eff} (µm)	V(µm ³ ·cm ⁻³)
Region 1	38	0.15	0.18	0.61	44
Region 2	82	0.25	0.05	1.23	54

Table 1. Aerosol properties on region 1, corresponding to the mixture of anthropogenic aerosol and mineral dust below 2250 m a.s.l. and region 2, corresponding to the mineral dust particles located above 2250 m a.s.l.