

1 **A New Algorithm for Detecting Cloud Height using OMPS/LP Measurements**

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10

11 **Abstract**

12 The Ozone Mapping and Profiler Suite Limb Profiler (OMPS/LP) ozone product requires
13 the determination of cloud height for each event to establish the lower boundary of the
14 profile for the retrieval algorithm. We have created a revised cloud detection algorithm
15 for LP measurements that uses the spectral dependence of the vertical gradient in
16 radiance between two wavelengths in the visible and near-IR spectral regions. This
17 approach provides better discrimination between clouds and aerosols than results
18 obtained using a single wavelength. Observed LP cloud height values show good
19 agreement with coincident Cloud-Aerosol Lidar and Infrared Pathfinder Satellite
20 Observation (CALIPSO) measurements.

21

22 **1. Introduction**

23 The Ozone Mapping and Profiler Suite Limb Profiler (OMPS/LP) is one of three OMPS
24 instruments onboard the Suomi National Polar-orbiting Partnership (S-NPP) satellite
25 (Flynn et al., 2007). S-NPP was launched in October 2011, into a sun-synchronous polar
26 orbit. The local time of the ascending node of the S-NPP orbit is 13:30. The LP
27 instrument collects limb scattered radiance data and solar irradiance data on a 2-D charge
28 coupled device (CCD) array over a wide spectral range (290-1000 nm) and a wide
29 vertical range (0-80 km) through three parallel vertical slits. Each slit provides a 1.85°
30 vertical field of view (FOV) corresponding to 112 km vertical extent at the tangent point.
31 The FOV of each slit is separated horizontally by 250 km in the cross-track direction. The
32 OMPS/LP produces three ozone profiles every 19 seconds along the orbit track, which
33 corresponds to a sampling distance of about 150 km (approximately 1° latitude).

34 OMPS/LP has been operating continuously since April 2012, collecting approximately
35 160-180 measurements (events) per orbit for each of the three slits and each of the 14-15
36 orbits per day. Jaross et al. (2013) provides more details about the OMPS/LP instrument
37 design and capabilities.

38

39 Retrieval of ozone profiles from limb scattering measurements becomes extremely
40 difficult in the presence of tropospheric clouds, because these clouds shield the signal
41 from the lower atmosphere, and also reflect a part of the incoming radiation back to
42 space. Due to the potential bias in the retrieved profiles from clouds, the OMPS/LP
43 retrievals are based on a cloud-free assumption. Thus, the current ozone retrieval
44 algorithm applied to LP measurements is designed to identify cloud height (if present) for
45 each event, and terminate the retrieval 1 km above this height.

46

47 Several techniques to retrieve cloud information from remote sensing measurements have
48 been developed. Most of them use changes in the oxygen A-band where the absorption of
49 oxygen is sensitive to cloud top height for retrieving cloud information (Kuze and
50 | Chance, 1994; Koelemeijer et al., 2001; Rozanov and Kokhanovsky, 2004; [Bourassa et](#)
51 | [al., 2005](#); [Eichmann et al., 2005](#); Kokhanovsky et al., 2005; [von Savigny et al., 2005](#);
52 | Loyola et al., 2007; van Diedenhoven et al., 2007; Loyola et al., 2010; Schuessler et al.,
53 | 2014; [Eichmann et al., 2015](#)). All these algorithms need a forward model with necessary
54 | assumptions to solve the radiative transfer equation in a multi-layer, multiple- scattering
55 | and absorbing atmosphere. In view of the OMPS/LP sensor relatively coarse spectral
56 | resolution (10 nm in visible region), Rault and Loughman (2013) determine cloud height

57 based on the identification of a sharp change in the vertical gradient of visible or near-
58 infrared radiances. Clouds appear as either faint or sharp discontinuities in the reflected
59 sunlight radiance vertical profiles. However, aerosol layers can also cause relatively
60 abrupt changes in the radiance profile at visible and IR wavelengths, so this approach
61 cannot always differentiate between tropospheric cloud and aerosols.

62

63 This paper describes a revised approach to cloud-top height detection using OMPS/LP
64 measurements, based on the spectral dependence of the vertical gradient in radiance
65 between two wavelengths. The approach is simple to implement. It is capable of
66 distinguishing between aerosols and water cirrus clouds in many cases. We show that the
67 performance of this approach is consistent with Cloud-Aerosol Lidar and Infrared
68 Pathfinder Satellite Observation (CALIPSO) results for quasi-coincident orbits in
69 individual cases, as well as for a larger statistical comparison.

70

71 **2. Algorithm design**

72 The new gradient-based LP cloud detection algorithm assumes that clouds produce a
73 sharper slope of the vertical gradient in radiances than aerosols. Because of the different
74 size distributions between aerosol particles and cloud hydrometeors, their scattering of
75 incoming solar radiation shows a different behavior. At UV and shorter visible
76 wavelengths, Rayleigh scattering reduces the contrast between cloudy and clear pixels.
77 This contrast increases with longer visible and near-IR wavelengths. Since aerosol
78 particles are smaller, their increase in brightness is less pronounced for the same change
79 in wavelength, so the increase in contrast for aerosols is not as large as for clouds.

80

81 We define the vertical gradient of observed radiances as the rate of change in radiances
82 with tangent height:

83

$$84 \quad G(\lambda, z) = \partial \ln I(\lambda, z) / \partial z \quad (1)$$

85

86 where $I(\lambda, z)$ is the limb radiance as a function of wavelength λ and tangent height z . The
87 variation of the radiance gradient (i.e., the height derivative) with wavelength between
88 500 nm and 900 nm for various targets is shown in [Figure- 1](#). Clear sky scenes show
89 Rayleigh scattering with almost no wavelength dependence, as expected. Note that the
90 tropospheric cloud at 14.5 km shows a steeper wavelength dependence than the aerosol
91 layer at 25.5 km. At visible and near-IR wavelengths longer than $\lambda_{\text{short}} = 674$ nm, where
92 the absorption of light by ozone can be neglected, the dependence of the radiance
93 gradient on wavelength can be parameterized using a linear relationship.

94

$$95 \quad G(\lambda, z) \approx \alpha(z)(\lambda - \lambda_s) + k(z); \quad \lambda \geq \lambda_s. \quad (2)$$

96

97 where α and k , the slope and intercept respectively, are a function of z and independent of
98 wavelength λ . Thus, the absolute value of α is the measure of the strength of the spectral
99 variation in radiance gradient. The slope α [in Eq. \(2\)](#) can be determined by choosing a
100 longer wavelength (λ_{long}) ~~in Eq. (2)~~:

101

$$102 \quad \alpha(z) = [G(\lambda_s, z) - G(\lambda_l, z)] / (\lambda_s - \lambda_l). \quad (3)$$

103

104 | We choose the two wavelengths λ_{short} and λ_{long} within the LP measurement range to
105 | maximize the cloud signature. 674 nm is chosen as the shorter wavelength λ_{short} to avoid
106 | Chappuis band ozone absorption. Data rate limitations on the S-NPP spacecraft mean that
107 | not all possible wavelengths and altitudes measured by LP can be downloaded during
108 | regular operations. Although LP measurements extend to ~1015 nm, changes in spectral
109 | coverage during the S-NPP mission mean that 868 nm represents the longest wavelength
110 | λ_{long} which is used in ~~e~~Equation (3) available with full temporal coverage.

111

112 | Calculating the slope values for the cases shown in ~~Figure-~~ 1, we find that
113 | $\alpha(6.5\text{km}\sim 10.5\text{km}) \approx 0$ is consistent with the spectrally independent gradient expected for
114 | clear sky, $\alpha(25.5 \text{ km}) = -0.00027$ represents the weaker spectral dependence of radiance
115 | gradient for an aerosol, and $\alpha(14.5 \text{ km}) = -0.0013$ corresponds to the strongest spectral
116 | dependence of radiance gradient for a cloud. We note that since the slope values are
117 | typically negative, we can rewrite Eq. (3) to define the gradient difference $\ln R(z)$:
118 | ~~Considering the sign of the slope, Eq. (3) can be rewritten as:~~

119

$$120 \quad \ln R(z) = [G(\lambda_s, z) - G(\lambda_l, z)] = \alpha(z)(\lambda_s - \lambda_l). \quad (4)$$

121

122 | Identifying the largest values of the gradient difference $\ln R$ in a measured profile should
123 | therefore provide a sensitive indicator for the presence of clouds.

124

125 | **3. Results**

126 **3.1 Threshold determination**

127 The method described in Sect. 2 has been used to determine cloud-top height from
128 OMPS/LP measurements. We assign a positive cloud detection if the value of $\ln R$ in Eq.
129 (4) meets a threshold value F at some altitude in the radiance profile. To determine the
130 cloud detection threshold, we use CALIPSO 532 nm backscattering daytime data
131 (Winker et al., 2003) and the corresponding CALIPSO Vertical Feature Mask (VFM)
132 version 3 data product (Vaughan et al., 2004; Kacenelenbogen et al., 2011) on selected
133 days where the satellite tracks of Suomi NPP and CALIPSO most closely overlap. Figure
134 2 provides an example of the determination of F during S-NPP orbit 4163 on August 16,
135 2012 for three events with clouds as well as one event without a cloud. These events
136 show distinctly different signatures in their $\ln R$ profiles. The sharpest vertical gradient,
137 with a maximum value of $\ln R = 0.33$ at 23.5 km, is observed for a polar stratospheric
138 cloud (PSC). For clouds at lower altitudes, the maximum values of $\ln R$ fall between 0.18
139 and 0.20. However, for the clear-sky event, the maximum value of $\ln R$ is very small (less
140 than 0.05). Further comparisons with CALIPSO observations indicate that $F = 0.15$ is a
141 reasonable threshold for positive cloud detection in LP data.

142

143 **3.2 Distinction between clouds and aerosols**

144 Confirming the presence of a cloud at any altitude requires an ability to discriminate
145 between cloud and aerosol signals. We define a quantity called aerosol scattering index
146 (ASI) at 674 nm for detecting aerosols in LP measurements:

147

148
$$ASI = (I_m - I_{c0})/I_{c0}. \tag{5}$$

149

150 where I_m is the measured radiance and I_{c_0} is the calculated radiance using a forward
151 model (Herman et al., 1995) for a Rayleigh atmosphere. Both I_m and I_{c_0} are normalized
152 at 45 km, assuming that there is no aerosol at that altitude. Figure 3a shows aerosols at
153 20-22 km at tropical latitudes, identified using ASI values for a single orbit on June 19,
154 2014. Figure 3b shows clouds at 10-15 km identified by CALIPSO data for the same
155 event. In the CALIPSO image, the red-gray-white colored features indicate clouds
156 between 10-15 km detected by lidar data, and the red dots represent LP cloud height
157 values detected by our new algorithm for the same orbit. Note that the LP cloud locations
158 are consistently at the top of the CALIPSO cloud regions. Figure 3c illustrates the LP
159 radiance gradient profiles for a single event at 3°S, identified by the dashed line in Figure-
160 3a and 3b. Note that $G(868 \text{ nm})$ shows peaks of comparable magnitude at 12.5 km
161 (tropospheric cloud) and 21.5 km (aerosol), whereas $G(674 \text{ nm})$ has a similar magnitude
162 peak at 21.5 km but a smaller peak at 12.5 km. Thus, the gradient difference $\ln R$ clearly
163 identifies the maximum cloud altitude using the threshold specified in Sect. 3.1, and does
164 not select the aerosol layer.

165

166 The OMPS/LP viewing geometry produces high single scattering angle (SSA) values for
167 Southern Hemisphere measurements (up to 160°), and low SSA values for Northern
168 Hemisphere measurements (down to 20°). This relationship leads to large variations in
169 ASI values over an orbit due to Mie scattering phase function effects. Figure 4 shows the
170 variation of $\ln R$ and ASI as a function of SSA at 14.5 km and 20.5 km for the same orbit
171 presented in Figure 3. ASI values increase rapidly for $SSA < 80^\circ$ at both altitudes. In

172 | contrast, lnR values are essentially constant throughout the orbit, and are well below our
173 | cloud detection threshold except for the tropical region that is consistent with CALIPSO
174 | cloud detections. We therefore use a constant cloud detection threshold to evaluate all
175 | LP measurements.

177 | **3.3 Comparison with LP Version 2 results**

178 | The cloud detection algorithm used in the OMPS/LP Version 2 ozone product (which is
179 | available at <https://ozoneaq.gsfc.nasa.gov/data/omps>) is based on the identification of
180 | sharp radiance profile changes at selected individual wavelengths (Rault and Loughman,
181 | 2013). Figure 54 shows cloud-top heights derived by our new radiance gradient
182 | algorithm and the LP Version 2 algorithm for a single orbit on August 16, 2012, with
183 | comparisons to the CALIPSO 532 nm backscatter coefficient for the same orbit. The
184 | Version 2 algorithm identifies many clouds that are not seen by CALIPSO, while the
185 | radiance gradient method only finds a few higher clouds between 33°-46°N at locations
186 | where CALIPSO also shows such clouds. This suggests that the Version 2 algorithm
187 | may misidentify aerosols as clouds. To further illustrate this result, we focus on two
188 | selected events at 36.5°N (event A) and 55.8°N (event B). The Version 2 algorithm finds
189 | sufficiently sharp changes in 892 nm radiance profiles to identify clouds at 14.5 km for
190 | both events (Fig. 54b). In contrast, the radiance gradient algorithm finds a clear cloud
191 | signature in lnR values for event A, but a much weaker signature that falls below the
192 | detection threshold for event B (Fig. 54c). These results give us confidence that the
193 | radiance gradient algorithm is not creating “false positive” cloud identifications.

194 Removing the incorrect cloud detections will also provide increased sampling in the
195 upper troposphere for LP retrieval products.

196

197 **4. Validation of LP cloud height product**

198 In order to quantify the accuracy of LP cloud-top height values derived by the new LP
199 radiance gradient algorithm, we evaluate our results against the CALIPSO VFM daytime
200 product. The similarity in orbits between the S-NPP and CALIPSO spacecraft makes it
201 possible to select many events with reasonably tight coincidence criteria [Δ latitude $<$
202 $\pm 0.15^\circ$, Δ longitude $< \pm 3.25^\circ$, Δ time $< \pm 1$ hr]. Only the CALIPSO measurements within
203 the footprint of the S-NPP orbit have been considered. These requirements yielded
204 approximately 439,000 cases spread over 70 sample days between April 2012 and
205 February 2015. We do not consider LP cloud detections below 5 km because our
206 approach is not effective at such low altitudes.

207

208 | Figure [65](#) shows the latitude distribution of cloud-top heights from these coincidence data
209 | sets in 5° zonal mean latitude bands. The cloud-top heights derived from the LP
210 | algorithm agree quite well with CALIPSO data in the tropics and mid-latitudes (up to
211 | approximately 50°). The cloud altitudes derived from both data sets decrease towards the
212 | poles due to the general decrease of the tropopause height. The LP cloud height values
213 | are higher in polar regions because our data set consistently includes polar stratospheric
214 | clouds (PSCs), which are identified at 15-30 km in winter and spring months (see
215 | example in Fig. 2). [LP measurements may also detect clouds that are located at different](#)

216 | positions along the line of sight, which will give higher derived cloud heights than if the
217 | same cloud is located at the tangent point position.

218

219 | Figure 76 shows a histogram of cloud height differences between the LP and CALIPSO
220 | data sets. The difference values are calculated as the LP cloud-top height minus the
221 | collocated CALIPSO value. The histogram has been constructed using bins of 1 km, the
222 | vertical sampling of the LP measurements. The most common difference values occur
223 | between -1 km and +4 km, with a median difference of $\Delta z_{\text{cloud}} = 1.8$ km. A Gaussian fit
224 | to these data yields a similar median difference value (2.0 km). We note that the LP
225 | cloud detection algorithm identifies the upper edge of a cloud, so it is not surprising to
226 | find a high bias in reported heights relative to CALIPSO cloud height values based on
227 | nadir-viewing lidar measurements. In addition, the LP vertical resolution is ~1.6-1.8 km,
228 | whereas CALIPSO data have much finer vertical sampling and resolution. The extended
229 | tail of this distribution towards large negative values corresponds to scattered high cloud
230 | values ($z_{\text{cloud}} > 20$ km) in the CALIPSO data set.

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232 | Figure 87 shows two histograms of cloud-top heights in the tropics as detected from the
233 | LP algorithm and from CALIPSO data. These distributions have very similar shapes, and
234 | the distributions are roughly Gaussian. The maximum cloud height occurrence frequency
235 | is observed between 14 and 16 km for both instruments. We note that the CALIPSO data
236 | shows some clouds up to 25 km height, which confirms previous studies that CALIPSO
237 | can sometimes misidentify aerosols as clouds (Chen et al., 2010; Chen et al., 2012).
238 | However, the LP data set does identify a population of clouds at 20-22 km, which are

239 clearly above the tropopause when individual orbits are inspected. The presence of these
240 unusually high clouds in the tropics is connected with the Kelut volcano which erupted in
241 February 2014. Remember that the $\ln R$ value calculation presented in Sect. 2 determines
242 the slope of the radiance gradient. Larger aerosol particles, such as those found in fresh
243 volcanic plumes, will increase the slope of the radiance gradient, and makes these events
244 more difficult to distinguish from “normal” clouds. In addition, patchy clouds in the near
245 and far sides of the tangent point may also cause high biased estimates of cloud height.
246 This potential error source was investigated in detail by Kent et al. (1997).
247

248 **5. Summary and conclusions**

249 We have developed a revised cloud detection algorithm for use with OMPS/LP
250 measurements. This algorithm uses the spectral dependence of the vertical gradient of
251 radiance at 674 nm and 868 nm to identify clouds and distinguish them from aerosols.
252 Comparison of cloud detection results for individual events with CALIPSO data confirms
253 the success of this approach. The revised LP cloud detection algorithm is also more
254 effective than the LP Version 2 algorithm in identifying only valid clouds. Our cloud
255 detection results are consistent with CALIPSO observations in terms of latitude
256 dependence and frequency distribution of altitudes. The offset in absolute cloud height
257 for coincident measurements is consistent with differences between the detection
258 methods. The LP cloud detection algorithm also consistently identifies polar stratospheric
259 clouds in both hemispheres, which may be useful for directly examining the impact of
260 PSCs on LP ozone retrievals. We do not attempt to retrieve cloud heights below 5 km
261 with this algorithm. Aerosol layers with larger particles, such as fresh volcanic plumes,

262 are more likely to be classified as clouds. Further theoretical studies of spectral
263 properties and scattering effects are needed to fully understand the applicability range and
264 limitations of this method. The new cloud detection algorithm will be implemented for
265 the forthcoming LP Version 3 ozone and aerosol retrieval algorithms, and the LP cloud
266 height values will also be distributed as a public data product.

267

268 **6. Data availability**

269 The OMPS LP Level 1 gridded radiance product (LP-L1G-EV) used to create the cloud
270 height product described in this paper can be obtained at

271 <https://ozoneaq.gsfc.nasa.gov/data/ozone/>.

272

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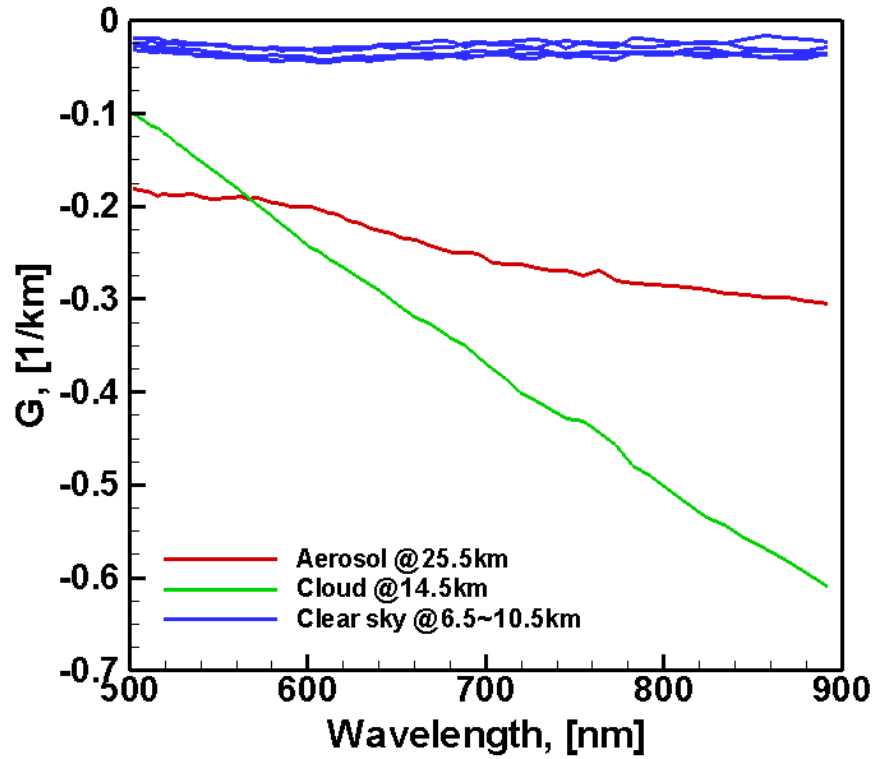
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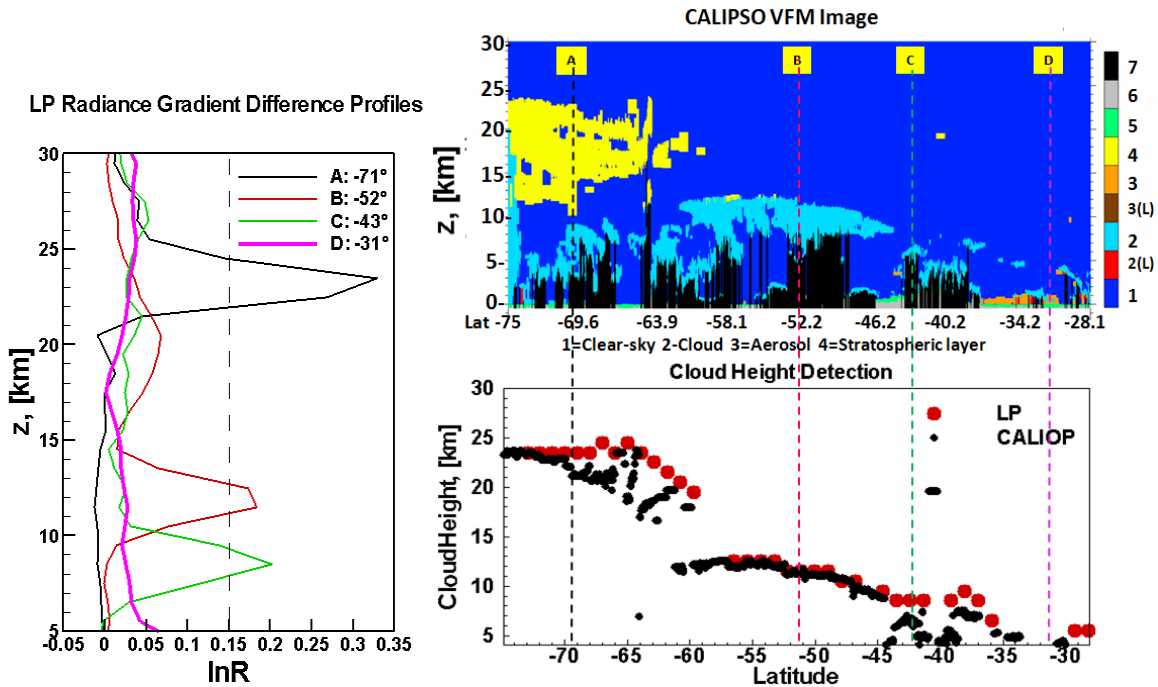
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358 **Fig. 1.** Variations in the radiance gradient $G(\lambda, z)$ from OMPS LP data at 5°S during orbit
 359 | 16754 on January 21, 2015. *Blue* = clear sky, *green* = ~~cloud~~aerosol, *red* = aerosol~~cloud~~.

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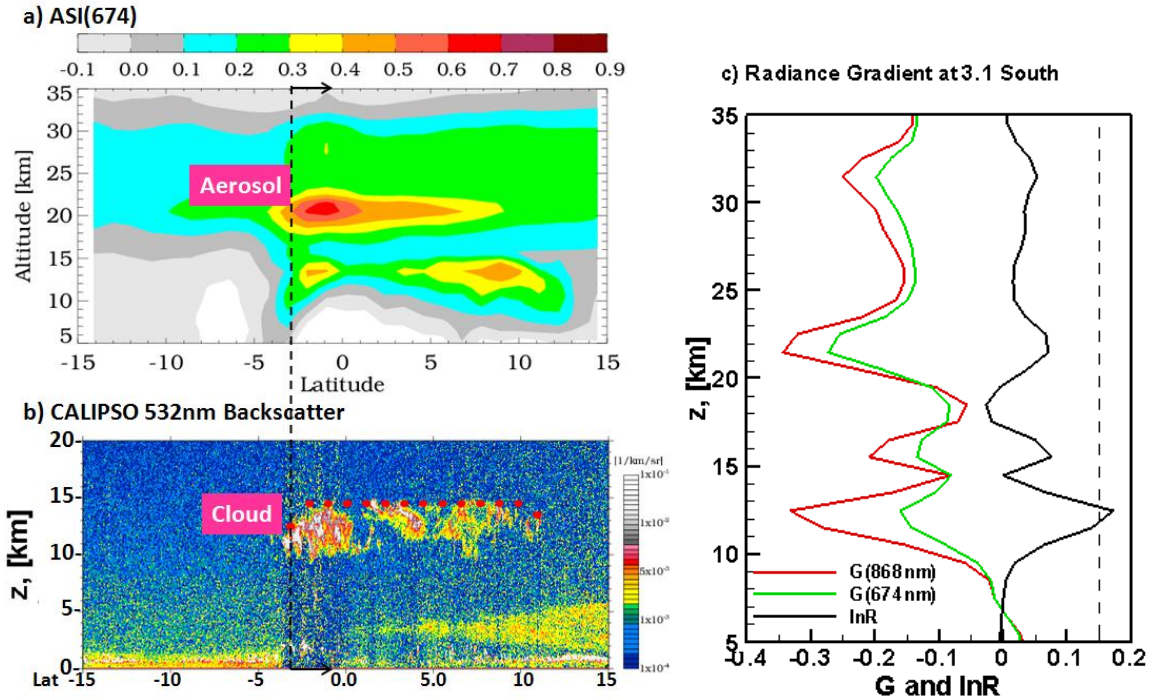
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Fig. 2. Radiance gradient and cloud detection results for four Southern Hemisphere events from a single orbit on August 16, 2012, using OMPS/LP measurements and CALIPSO Vertical Feature Mask (VFM) daytime data.

Left panel: Vertical profiles of the LP radiance gradient $\ln R$ for each event. The dashed black line represents the cloud detection threshold, which identifies clouds in events A, B, and C.

Top right panel: CALIPSO VFM image for the same orbit. Yellow features indicate polar stratospheric clouds (PSCs), and light blue patches represent clouds.

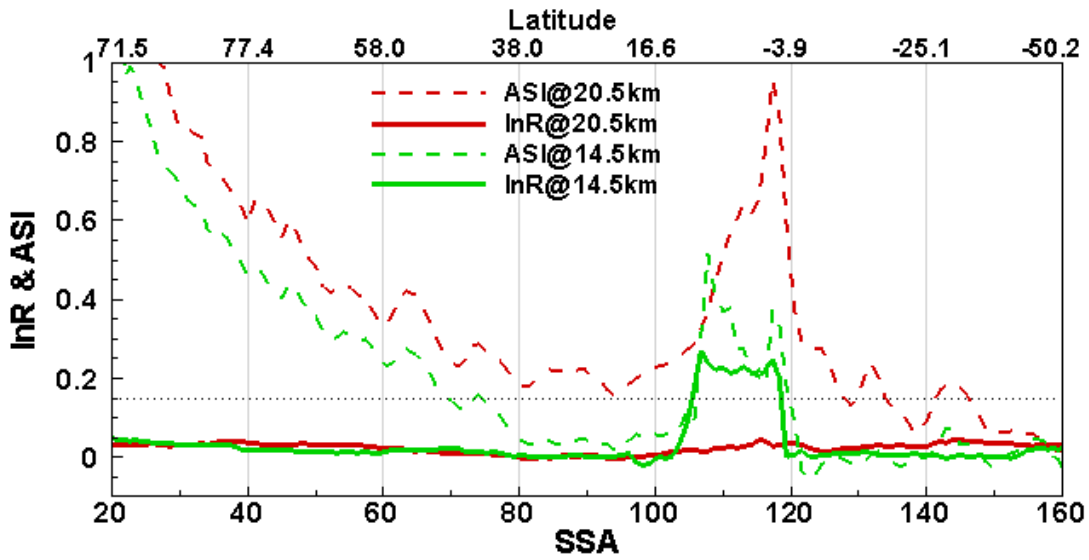
Bottom right panel: Cloud height values detected by the LP algorithm (red dots) and from CALIPSO VFM data (black dots) in the same orbit. The four colored dotted lines in the right panel indicate the four events labeled as A, B, C and D respectively.



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Fig. 3. Example of discrimination between clouds and aerosols, using OMPS/LP observations taken on June 19, 2014.
 (a) Aerosol layer at 20-22 km in tropics identified using OMPS/LP aerosol scattering index (ASI).
 (b) Tropospheric clouds at 10-15 km identified in CALIPSO data for the same orbit. The red dots represent LP cloud-top height values derived from the radiance gradient algorithm.
 (c) LP radiance gradient profiles (red = 868 nm, green = 674 nm) for a single event at 3.1°S, identified by the dashed line in panels (a) and (b). The difference between profiles (lnR) is shown as the black line.

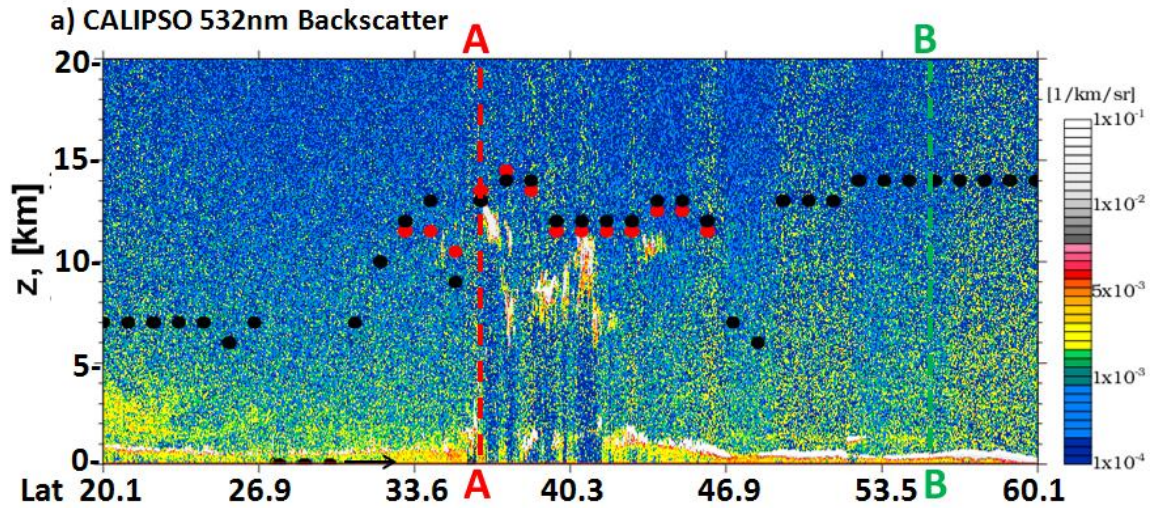
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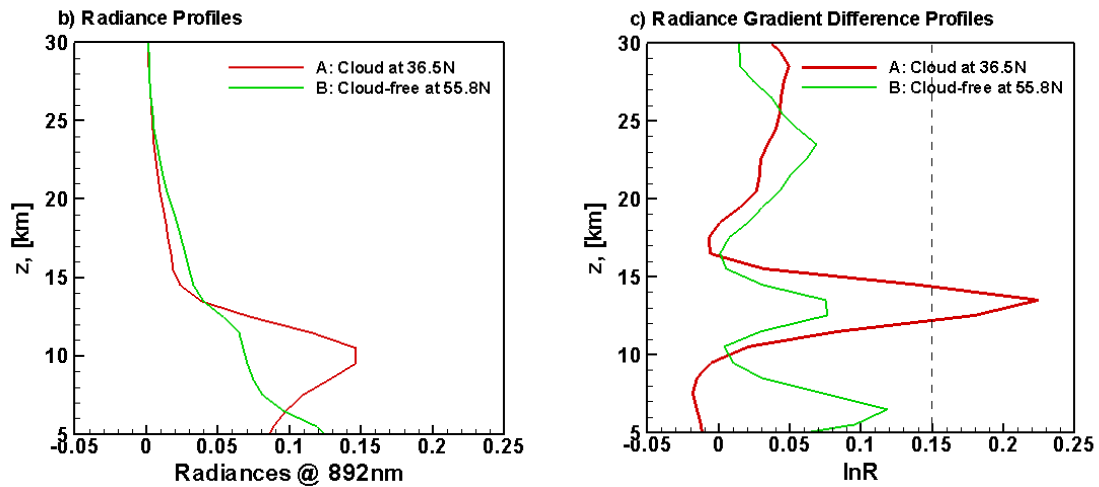
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Fig. 4. InR (solid line) and ASI (dashed line) plotted as a function of single scattering angle (SSA) at 14.5 km (green) and 20.5 km (red). This figure uses OMPS/LP observations taken on 19 June 2014, from the same orbit shown in Figure 3. The dotted black line represents the value of the cloud detection threshold for the InR curves.

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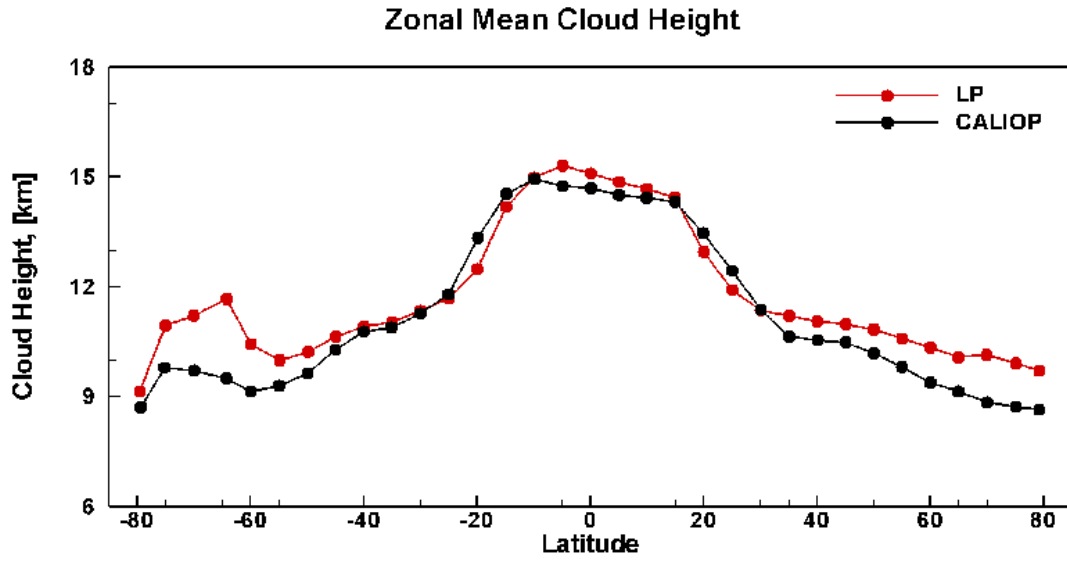


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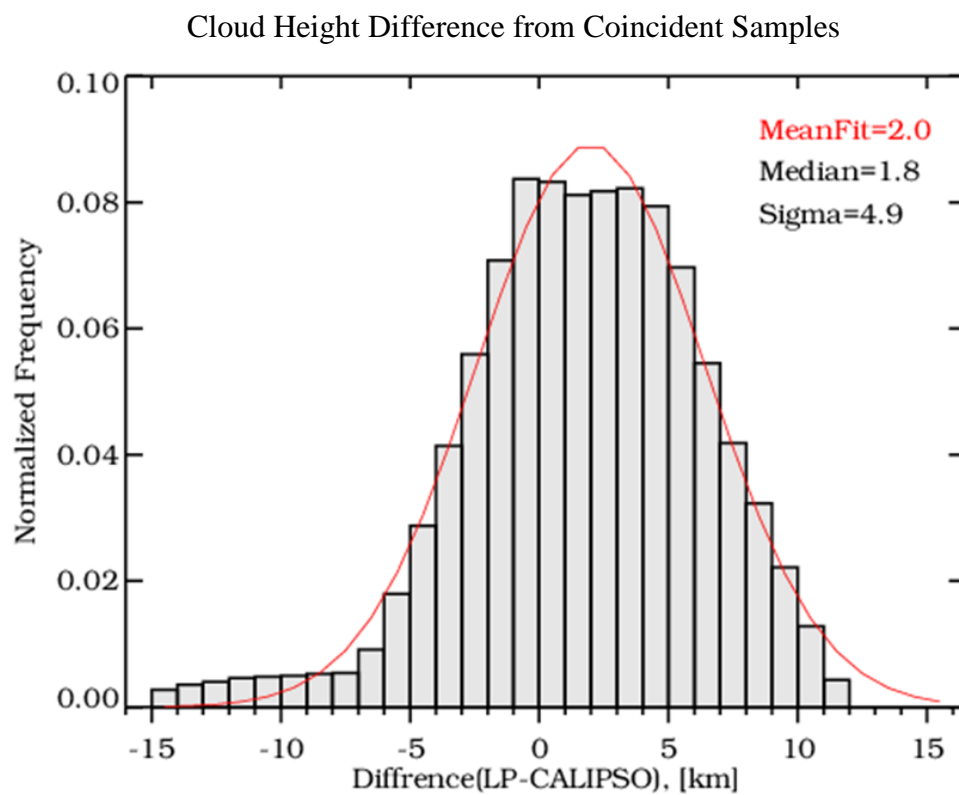
Fig. 54. Comparison of LP cloud detection results for a single orbit on August 16, 2012. (a) CALIPSO 532 nm daytime backscatter data for the same orbit. The red-grey-white regions in the image denote the cloud layers. The red and black dots in the image represent cloud-top heights derived from the LP radiance gradient algorithm and the LP Version 2 algorithm, respectively. Lines A and B indicate OMPS/LP measurements at 36.5°N and 55.8°N . (b) Radiance profiles at 892 nm used as the basis for Version 2 algorithm cloud identification. *Red* = event A, *green* = event B. (c) Radiance gradient difference profiles used for new LP algorithm cloud identification. *Red* = event A, *green* = event B. The dashed line represents the value of the cloud detection threshold.



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Fig. 65. Zonal mean cloud height calculated from LP cloud detection algorithm results (red line) and collocated CALIPSO data (black line) in 5° latitude bands. Results averaged over 70 sample days between April 2012 and February 2015.

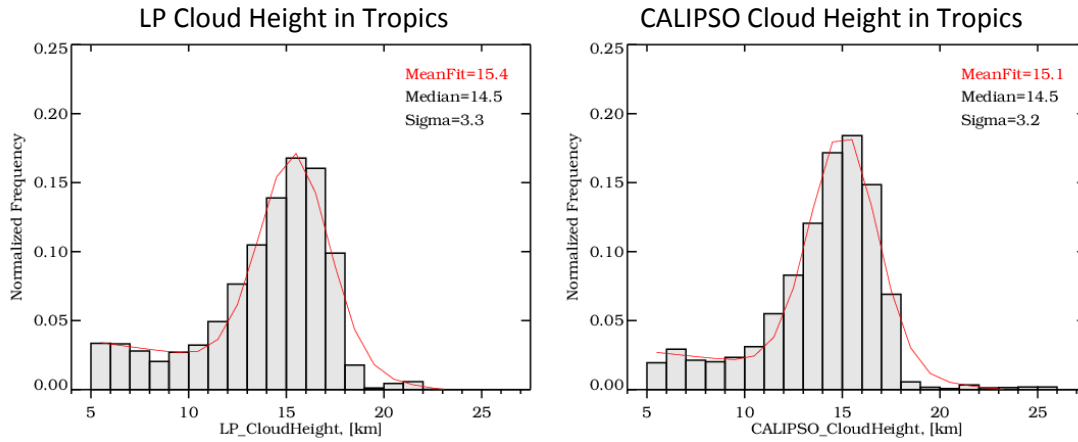
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Fig. 76. Normalized frequency histogram of all cloud height differences (LP-CALIPSO) from coincidence data sets in $\Delta z_{\text{cloud}} = 1$ km intervals. The red curve represents a Gaussian fit to the data.

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478 | **Fig. 87.** Normalized frequency histograms of all cloud height values from LP cloud
479 | detection results (*left panel*) and the collocated CALIPSO data (*right panel*) in the tropics
480 | (latitude $< \pm 30^\circ$) in $\Delta z_{\text{cloud}} = 1$ km intervals. The red curves represent linear combination
481 | of a Gaussian and quadratic function fit to each data set.

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