



# 1 Title: Evaluation And Attribution Of OCO-2 XCO<sub>2</sub> Uncertainties

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# 11 Abstract

- 12 Evaluating and attributing uncertainties in total column atmospheric CO<sub>2</sub>
- 13 measurements (*XCO*<sub>2</sub>) from the OCO-2 instrument is critical for testing hypotheses
- related to the underlying processes controlling *XCO*<sub>2</sub> and for developing quality flags
- 15 needed to choose those measurements that are usable for carbon cycle science.
- 16 Here we test the reported uncertainties of Version 7 OCO-2 *XCO*<sub>2</sub> measurements by
- 17 examining variations of the *XCO*<sub>2</sub> measurements and their calculated uncertainties
- 18 within small regions ( $\sim$ 100 km x 10.5 km) in which CO<sub>2</sub> variability is expected to be
- 19 small relative to variations imparted by noise or interferences. Over 39,000 of
- 20 these "small neighborhoods" comprised of approximately 190 observations per
- 21 neighborhood are used for this analysis. We find that a typical ocean measurement
- should have a precision and accuracy of 0.35 and 0.24 ppm respectively for
- 23 calculated precisions larger than  $\sim$ 0.25 ppm. These values are approximately
- 24 consistent with the calculated errors of 0.33 and 0.14 ppm for the noise and
- 25 interference error (assuming that the accuracy is bounded by the calculated
- 26 interference error). The actual precision for ocean data becomes worse as the
- 27 signal-to-noise increases or the calculated precision decreases below 0.25 ppm for
- 28 reasons that not well understood. A typical land measurement (both nadir and
- 29 glint) is found to have a precision and accuracy of approximately 0.75 ppm and 0.65
- 30 ppm respectively as compared to the calculated precision and accuracy of





- 1 approximately 0.36 ppm and 0.2 ppm. However, this precision includes the effects of
- 2 synoptic variability in the total column that could be as high as 0.5 ppm during the
- 3 summer drawdown period. The accuracy is likely related to interferences such as
- 4 aerosols or surface albedo and is a lower bound as it is evaluated by comparing
- 5 gradients in OCO-2 estimates of *XCO*<sub>2</sub> to expected gradients across the region and
- 6 not by direct comparison to well-calibrated XCO<sub>2</sub> measurements from the ground
- 7 network.
- 8

# 9 **1.0 Introduction**

10

Variations of total column CO<sub>2</sub> (XCO<sub>2</sub>) resulting from photosynthesis and 11 respiration in tropical forests (e.g. Parazoo et al. 2013), urban emissions (e.g. Kort et 12 13 al., 2012) or tropical fires (e.g. Bloom et al., 2016) range from 2 – 5 ppm. 14 Consequently, in order to use space-based measurements of XCO<sub>2</sub> to infer fluxes or 15 properties of the processes controlling these variations, uncertainties in  $XCO_2$ should ideally be much much smaller than this variability (Miller et al. 2007). The 16 17 Orbiting Carbon Observatory-2 (OCO-2) was launched in July 2014, to measure the 18 atmospheric column averaged carbon dioxide  $(CO_2)$  dry air mole fraction, XCO<sub>2</sub> with 19 the precision, accuracy, and coverage needed to quantify variations on regional 20 scales at monthly intervals. These measurements are being used to investigate the underlying carbon cycle processes controlling atmospheric CO<sub>2</sub>. The radiative 21 22 transfer and XCO<sub>2</sub> estimation (or retrieval) algorithms (Boesch et al. 2006; 2011; Connor et al. 2008; O'Dell et al., 2012) were developed and tested using observed 23 24 radiances from the Japanese TANSO GOSAT instrument (Kuze et al. 2009; Yoshida et 25 al. 2011), which measured similar spectral regions as the OCO-2 mission. These 26 algorithms also allowed extensive evaluation of quality flags and metrics needed to 27 reject estimated XCO<sub>2</sub> values which were clearly spurious, likely because of poorly estimate values for aerosols, clouds, surface albedo or surface pressure (Crisp et al., 28 29 2012; Mandrake et al., 2013). In this paper we evaluate the calculated uncertainties 30 due to noise and interferences in the OCO-2 data product (Version 7).





Our approach follows the methodology described in Boxe et al. [2010] and Kuai 1 2 et al. [2013] in which variations of the observed trace gas over a small "area" are 3 compared to the calculated errors. Figure 1 shows the distribution of latitudinal 4 gradients in XCO<sub>2</sub> over the ocean and over North America based on the "high 5 resolution" Carbon Tracker model (e.g. Peters et al., 2007) with ~100 km spatial 6 resolution. This distribution is calculated by differencing XCO<sub>2</sub> from adjacent model 7 grid points, as a function of latitude, using all modeled XCO<sub>2</sub> values in July 2015. We 8 find that the root-mean-square (RMS) value of these gradients is approximately 0.3 9 ppm/100 km during the summer and ~0.1 ppm/100 km during November. Keppel-Aleks [2011, 2012] also found North American summertime gradients in *XCO*<sub>2</sub> 10 between 0.1 ppm/100 km to 0.3 ppm/100 km using ground based total column data 11 and measured wind speeds. In addition, these studies found synoptic variability 12 13 could change  $XCO_2$  values by up to 0.5 ppm over the study time period in a random 14 manner (Figure 5 in Keppel-Aleks [2011]). In contrast, Figures 1a and 1b show that typical variations in the gradients over the ocean should be less than that of land. 15 between  $\sim 0.1$  ppm/100 km to 0.2 ppm/100 km. While in situ measurements [e.g. 16 17 Wofsy *et al.*, 2011] and model data do show variations in  $XCO_2$  that are sometimes 18 larger than 0.2 ppm/100 km we would expect that these variations do not represent 19 typical  $XCO_2$  gradients, especially since the total column of  $CO_2$  integrates the effects 20 of many sources and sinks from hundreds to thousands of kilometers away from the observation [e.g. Keppel-Aleks *et al.*, 2011]. Because the expected variability in 21 22 XCO<sub>2</sub> from models, ground-based data, and in situ measurements are comparable or less than the calculated OCO-2 uncertainties, we can compare the observed 23 24 variability of XCO<sub>2</sub> from OCO-2 data within a small region, covering an orbit track 25 that spans 100 km in latitude, to evaluate the magnitude and character of their 26 corresponding calculated uncertainties. 27 28 2.0 Overview of OCO-2 data

29





1	The OCO-2 instrument measures radiances in the molecular oxygen ( $O_2$ ) A-band		
2	(0.765 microns), the "weak" $\mathrm{CO}_2$ band at 1.61 microns and the "strong" $\mathrm{CO}_2$ band at		
3	2.06 microns. The OCO-2 instrument is an imaging spectrometer that collects with 8		
4	samples, or "spatial footprints" across a narrow (0.8-degree) swath track observes		
5	near the "glint spot" where sunlight is specularly reflected by the surface.		
6	Observations are taken in three different modes, (1) "Nadir", where the space-craft		
7	points the instrument's aperture at the ground directly downward along the orbit		
8	track, (2) "Glint," where the space craft points instrument's aperture near the "glint		
9	spot" where sunlight is specularly reflected by the surface, near the specular		
10	reflection point for sunlight, and (3) Target, where the space-craft points the		
11	instrument aperture at a stationary surface target, such as a validation site or city.		
12	Nadir observations usually return useful measurements only over land. Glint		
13	observations return useful data over both land and ocean. Here, we discriminate		
14	land-glint and ocean glint observations because they have different error statistics.		
15	We do not evaluate Target data in this analysis due to spurious statistics that are		
16	observed with the Target data.		
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- 1 data with higher warn levels likely or are too strongly affected by interfering effects.
- 2 The warn levels are primarily evaluated empirically; for these reasons we
- 3 conservatively use only data with warn levels of 10 or smaller to ensure that the
- 4 corresponding errors are likely well characterized:
- 5 <u>http://disc.sci.gsfc.nasa.gov/OCO-2/documentation/oco-2-</u>
- 6 v7/0C02 XC02 Lite Files and Bias Correction 0915 sm.pdf.
- 7
- 8

# 9 **3.0 Evaluation of Uncertainties**

- 10
- 11 *3.1 Overview of Error Analysis and Methodology*
- We evaluate the uncertainties of the *XCO*<sub>2</sub> observations by examining the 12 13 variations of  $_{XC02}$  within small neighborhoods of approximately 10.5 km by 100 km 14 in size. After warn level filtering, this "small neighborhood" test set is composed of 15 approximately 1.5 million Land-Nadir soundings, 1.0 million Land-Glint soundings, and 5.0 million Ocean-Glint soundings. Each neighborhood contains at least 50 16 17 soundings, with roughly 190 soundings per neighborhood on average, and 18 approximately 39,000 small neighborhoods in total across the three modes. 19 stretching from approximately 30S to 30N. The strict filtering used in this analysis 20 (Warn Levels <= 10), and the need for at least 50 measurements per bin limits this analysis to latitudes between 30S to 30N, primarily over drier, sub-tropical regions 21 22 over land but no obvious preferential distribution over the ocean (not shown). As discussed in [O'Dell et al., 2012], a CO<sub>2</sub> profile is simultaneously estimated 23 24 with all other geophysical parameters that affect the observed radiance such as aerosols, albedo, and surface pressure. The "column-averaged dry air mole fraction" 25 26 of CO<sub>2</sub> or XCO<sub>2</sub> is then calculated by applying the column operator [e.g. Connor et al., 27 2008; Worden et al., 2015] to the estimated  $CO_2$  profile. As discussed in Rodgers
- 28 [2000], Worden *et al.* [2004], Connor [2008], and Bowman *et al.* [2006], when this
- $\label{eq:29} non-linear retrieval converges to a solution, the estimated XCO_2 can be written as:$
- 30





 $\hat{X} = X_a + h^T \mathbf{A}_{xx}(x - x_a) + h^T \mathbf{A}_{xy}(y - y_a) + h^T \mathbf{G} \mathbf{n} + h^T \mathbf{G} \sum_{i} \mathbf{K}_i \delta_i \quad (1)$ 1 2 where  $\hat{X}$  is the estimated total column for CO<sub>2</sub>,  $\hat{X}_{a}$  is the *a priori* value used to help 3 regularize the retrieval, the vector  $\mathbf{x}$  is the "true" CO<sub>2</sub> profile in units of volume 4 mixing ratio (VMR), discretized onto the forward model atmospheric pressure grid 5 6 used to calculate the transfer of radiation needed to model the observed radiance. 7 The  $x_a$  is the *a priori* for the CO<sub>2</sub> profile. The vector "y" contains all the other 8 parameters that are simultaneously estimated with **x** such as aerosol properties, 9 surface albedo, surface pressure. The vector "n" is the actual noise in the radiance. The quantities **x**, **y**, and **n** are not known exactly, only their statistical properties can 10 be estimated. The vector "**h**" is the column operator which maps a profile on the 11 12 pressure grid defined by "*x*" into a dry air total column. The averaging kernel matrix 13 A describes the sensitivity of the estimate to each retrieved parameter [Rodgers, 2000]. In equation 1 the averaging kernel matrix is composed of two parts,  $A_{xx}$  and 14 15 A<sub>xy</sub>, described by: 16

17 
$$\mathbf{A} = \begin{bmatrix} \mathbf{A}_{xx} & \mathbf{A}_{xy} \\ \mathbf{A}_{yx} & \mathbf{A}_{yy} \end{bmatrix}$$
(2)

18

For example  $A_{xx}$  describes the sensitivity (or  $\frac{\partial \hat{x}}{\partial x}$ ) of the estimated CO<sub>2</sub> on each level, 19 20 x, to its true value, whereas  $A_{xy}$  describes the sensitivity of the estimated  $CO_2$  on 21 each level, x, to all other simultaneously estimated parameters, e.g., aerosols, etc. The matrix, " $G_{1}$ " is the gain matrix, , which is the derivative of the estimated CO<sub>2</sub> on 22 each level, x, to the observed radiance, "L" (or  $G = \frac{\partial \hat{x}}{\partial t}$ ). The matrix, "K," is the 23 Jacobian, or sensitivity of the observed radiance to a parameter (e.g. K =  $\frac{\partial L}{\partial x}$ ). The last 24 term,  $\delta$ , describes the error in all parameters that are not estimated for this 25 retrieval, but are assumed constant, such as absorption coefficients or instrument 26 27 functions (e.g. Connor et al., 2008). The mean CO<sub>2</sub> column is written as:





1	$\hat{X}_{mean} = X_a + h^T \frac{1}{N} \sum_{j=1}^{N} \mathbf{A}_j (\mathbf{x}_j - \mathbf{x}_a) + \frac{1}{N} h^T \sum_{j=1}^{N} \mathbf{A}_{xy}^j (\mathbf{y}_j - \mathbf{y}_a) + \frac{1}{N} h^T \sum_{j=1}^{N} \mathbf{A}_{xy}^j (\mathbf{y}_j - \mathbf{y}_a) + \frac{1}{N} h^T \sum_{j=1}^{N} \mathbf{A}_{yj}^j (\mathbf{y}_j - \mathbf{y}_a) + \frac{1}{N} h^T \sum_{j=1}^{N} h^T \sum$
2	$\frac{1}{N}\sum_{j=1}^{N} \boldsymbol{h}^{T} \mathbf{G}_{j}(\boldsymbol{n}_{j} + \sum_{i,j} \mathbf{K}_{i,j} \boldsymbol{\delta}_{i,j}) $ (3)
3	
4	where N is the number of observations within the small area and for simplicity we
5	assume the column operator $m{h}$ is constant across the domain.
6	
7	For the next three sections, we test the following hypotheses regarding the observed
8	distributions within the collection of "small neighborhoods" and their calculated
9	uncertainties:
10	
11	H1: Uncertainties within a small area are primarily due to random noise
12	H2: Uncertainties are correlated
13	H3: Uncertainties within a small area are described by a slowly varying bias
14	(consistent with the expected effects of interference error).
15	
16	We look at the variability with respect to the neighborhood mean in two ways: (1)
17	for small neighborhoods; the predicted errors for a neighborhood are averaged
18	from the observations that comprise that neighborhood, making the statistics
19	technically a sum of Gaussians, and (2) the variability with respect to the
20	neighborhood mean, sorted by predicted error and aggregated over many
21	neighborhoods; the statistics in this case should be Gaussian, however the locality of
22	the analysis is somewhat reduced.
23	
24	3.2 H1: Error due to noise
25	
26	To evaluate whether measurement noise in the radiances is the primary
27	factor driving variability within a small area we assume that the terms $A_{xy}(y_j - y_a)$
28	and systematic errors $\mathbf{K}_{\mathbf{i},\mathbf{j}} oldsymbol{\delta}_{\mathbf{i},\mathbf{j}}$ do not vary. Based upon these approximations, the
29	difference between an observation and its mean is given by:

7





1  
2 
$$\hat{X}_{obs} - \hat{X}_{mean} = \delta_{obs} = \delta_{XCO2} + G_{obs}n_{obs} - \frac{1}{n} \sum_{j}^{N} G_{j}n_{j}$$
 (4)  
3  
4 where  $\delta_{XCO2} = h^{T} A(x_{obs} - x_{mean})$  and is the difference between the individual  
5 "true" *XCO*<sub>2</sub> and the mean of the "true" *XCO*<sub>2</sub> values within the neighborhood.  
6 Assuming the measurement noise is spatially uncorrelated, the variance within the  
7 small neighborhood is [e.g. Bowman *et al.*, 2006] is:  
9  $Var \left| \left| \hat{X}_{obs} - \hat{X}_{mean} \right| \right| = \sigma_{obs}^{2} = \sigma_{XCO2}^{2} + \sigma_{noise}^{2} + \frac{1}{n^{2}} \sum_{j=1}^{N} \sigma_{j}^{2} - \frac{2}{n} \sigma_{k}^{2}$  (5)  
10 where  $\sigma_{noise} = G_{K}S_{k}G_{K}^{T}$  is the measurement uncertainty due to noise. The  $\sigma_{XCO2}$  is  
11 where  $\sigma_{noise} = G_{K}S_{k}G_{K}^{T}$  is the measurement uncertainty due to noise. The  $\sigma_{xCO2}$  is  
12 the variability of the true *XCO*<sub>2</sub> within the small neighborhood. The  $S_{k}$  is the  
13 spectral instrumental noise covariance and is calculated during calibration of the  
14 instrument. The individual  $\sigma_{noise}$  values are provided for each measurement in the  
15 OCO-2 product files. For large N, Equation 5 is approximately equal to:  
16  $\sigma_{XCO2}^{2} + \sigma_{noise}^{2}$ .  
17 We next evaluate these uncertainties using two approaches. In the first  
18 approach we gather all observations that have approximately the same calculated  
19 measurements uncertainty,  $\sigma_{noise}$ , (to within 0.01 ppm) as provided in the OCO-2  
20 product files and compare to the actual variability of these observations. The steps  
10 Calculate the  $\delta_{obs}$  or difference between an observation and its mean  
21 within a small neighborhood as shown in Equation 4.  
22 Collect all of the  $\delta_{obs}$  values from all neighborhoods used in this analysis  
23 whose corresponding  $\sigma_{noise}$  values (measurement uncertainty) are the  
24 same to within 0.01 ppm and bin them as a function of  $\sigma_{noise}$ . There are  
23 typically about 1000 observations per  $\sigma_{noise}$  bin.





1	3) Compare the standard deviation of the collection of $\delta_{obs}$ values within			
2	each bin to the expected standard deviation due to noise or, $\sigma_{noise}$ . Based			
3	on Equation 5 we should expect to get a linear, one-to-one relationship if			
4	the dominant parameter affecting the variability within a small			
5	neighborhood is noise.			
6				
7	The results of these comparisons for land-nadir, land-glint, and ocean-glint			
8	observations are shown in the upper left panels of Figures 2, 3, and 4 respectively.			
9	These results show the calculated measurement error has skill, i.e. there is a linear			
10	relationship between calculated and actual error. However, over land the observed			
11	random variability is approximately 0.4 ppm larger than the variability expected			
12	from noise. Synoptic variations in XCO <sub>2</sub> could potentially explain much of this extra			
13	0.4 ppm however other sources of variability could be due to the strong non-			
14	linearities in the retrieval [e.g. Kulawik et al., 2008] or local variability between the			
15	true and <i>a priori</i> in the interferences, or non-retrieved parameters. Over the ocean			
16	there appears to be an even stronger one-to-one relationship between the			
17	calculated uncertainty and the actual uncertainty except for calculated uncertainties			
18	less than approximately 0.25 ppm which show a strong inverse relationship. We			
19	find that these observations (not shown) tend to occur in the tropics in cloudy			
20	regions and that the observations tend to have very high signal-to-noise ratios.			
21	We next test whether the calculated measurement noise is a useful value for			
22	predicting the expected distribution of observations within a neighborhood.			
23	Because each $\delta_{obs}$ is drawn from a distribution with a different variance, we treat			
24	the sample of each set of observations, $[\delta_1,\delta_2,\delta_N]$ , as being drawn from an			
25	uncorrelated distribution with individual variances $\sigma^2_{obs}$ . Accordingly, the variance			
26	of this sample should be the average of the individual variances $\sigma^2_{obs}$ :			
27				
28	$Var [\widehat{X}_{obs} - \widehat{X}_{mean}]   = \left  \left  [\delta_1, \delta_2, \dots \delta_N] \right  \right  = \frac{1}{N} \sum_{i}^{N} \sigma_i^2 $ (6)			
29				





1	The top right panel of Figure 1 shows a comparison of the observed variance of the				
2	$XCO_2$ distributions (using the left side of Equation 6) within each neighborhood				
3	(black circles) versus the expected variance in XCO <sub>2</sub> using the right side of Equation				
4	6. Each black symbol represents a single neighborhood. In contrast to the top left				
5	panel of Figure 1, this result suggests that the measurement error has no skill in				
6	predicting the observed variance of <i>XCO</i> <sub>2</sub> within a neighborhood.				
7	We next test whether the observed variance, versus that due to measurement				
8	noise or sampling, explains the upper right panel of Figures 2, 3, and 4. To perform				
9	this test, we perform the following steps:				
10					
11	1) Within each neighborhood, replace the calculated measurement error with				
12	the "actual" measurement error as shown by the solid red line in the upper				
13	left panel of Figures 2, 3, and 4, for each observation.				
14	2) Create a simulated distribution of observations based on this new				
15	uncertainty.				
16	3) Randomly sample (or take) one of these observations $\rightarrow$ label this the				
17	"modeled" observation.				
18	4) Repeat steps 1-3 for all observations in the neighborhood.				
19	5) Calculate the variance of this "modeled" set of observations for each				
20	neighborhood.				
21					
22	The red dots in Figures 2b, 3b, and 4b show the modeled distributions using the				
23	steps discussed above. The modeled distribution is more consistent with the mean				
24	of the observed distribution relative to the one-to-one line. However, it is clear from				
25	this simulation that errors due to random noise and sampling do not explain the				
26	observed variance for each neighborhood although the distribution of variances for				
27	the ocean show much better agreement relative to the land distributions.				
28					
29	3.3 H2: Uncertainties are correlated				
30					





- 1 We next test whether observed correlations in the data could explain the
- 2 distributions of the data within a neighborhood. Figures 5 shows the joint
- 3 distribution of the XCO<sub>2</sub> anomaly and a 0.3 second lagged anomaly in a
- 4 neighborhood. If the data were uncorrelated then the joint distribution should be
- 5 circular; the asymmetric distribution therefore implies that the errors, as
- 6 empirically described by the differences, are correlated. Figures 6a and 6b show
- 7 that autocorrelation is observed both in time for measurements made on the order
- 8 of 1 second of each other, and with respect to the spatially adjacent "footprints," the
- 9 8 simultaneous measurements made by the OCO-2 instrument at each time. The
- 10 range of correlations for the different observation types, land nadir, land glint, and
- ocean glint are 0.45, 0.43, and 0.28 as a function of footprint and 0.31, 0.34, and 0.24
  as a function of time.
- 12 as a function of time.
- 13 In order to test whether these observed correlations could explain the
- 14 distributions shown in Figures 2, 3, and 4, we conservatively use a correlation
- 15 coefficient of 0.7 for all observations (an extreme case). We then use the following
- 16 procedure, building on the steps described in the previous section.
- Within each neighborhood replace the calculated measurement error with
   the "actual" measurement error as shown in the upper left panels of Figures
   2, 3, and 4 for an observation
- 2) Starting with the first observation (in time) within a neighborhood for
  21 Footprint #1, sample a value for the observation from the distribution of
- 22 "actual" measurement errors. Label this the "modeled" observation.
- 3) For all subsequent observations in time for Footprint #1, sample each
  "modeled" observation from a distribution that is correlated with the
  modeled observation at the previous time step and has a variance
  corresponding to the "actual" measurement error.
- 4) For observations in Footprints #2-8, sampling each modeled observation
  from a distribution correlated with the modeled observation at the same
  time step in the previous (adjacent) footprint, again with a variance
- 30 corresponding to the "actual" error.





1	5) Calculate variance of this "modeled" set of observations, for each
2	neighborhood.
3	
4	As can be seen in the lower left panels of Figures 2, 3, and 4, adding correlations to
5	the data makes the comparison worse because the modeled distributions become
6	much narrower relative to the modeled distributions in the upper right panels of
7	these figures. Our conservative choice of a 0.7 correlation between observations at
8	adjacent times and footprints illustrates this effect clearly. We therefore conclude
9	that while correlations are empirically observed in the data, they cannot completely
10	explain the observed distributions within the small neighborhoods.
11	
12	3.4 H3: Uncertainties within a small area are characterized as a slowly varying bias.
13	
14	We next examine whether "non-random" uncertainties could explain the
15	observed distributions in the upper right panels of Figures 2,3, and 4. For example,
16	as shown in Equation (1), the jointly retrieved parameters $(y - y_a)$ might remain
17	constant across a neighborhood but the Averaging kernel associated with this term,
18	which is given by $A_{xy} = \frac{\partial x}{\partial L} \frac{\partial L}{\partial y} = GK_y$ , can vary across a neighborhood as the pointing
19	angle varies. The effect of non-retrieved parameters such as instrument effects or
20	spectroscopy on the estimate can vary for the same reason.
21	Figure 7 shows the variation of $XCO_2$ across one of the ocean neighborhoods
22	for all 8 OCO-2 footprints (denoted by "FP"). The right panel shows the observed
23	distribution in black relative to the mean $XCO_2$ of the neighborhood. For reference,
24	the red dashed line in the right panel indicates the expected distribution if only
25	random noise explained the variability. The slope shown in Figure 7 represents an
26	extreme case but demonstrates that observations can pass the set of quality flags
27	but still show this unlikely behavior over the ocean. Figure 8 shows the distribution
28	of all slopes across all land-nadir neighborhoods used in this study and different fits
29	(Gaussian, Lorentz, Laplace) to the distribution. The Laplace distribution provides





1	the best overall fit so we use its functional form as a simple, convenient description			
2	of the shape of the sharply peaked slope distribution. More complex models such as			
3	Gaussian mixtures might also describe the shape of this distribution of slopes as			
4	drawn from several distinct "populations" of neighborhoods, but we leave such an			
5	analysis to future work. The RMS of the distribution is approximately 1.28. which is			
6	much larger than expected variations in XCO <sub>2</sub> (e.g. Figure 1 and Keppel-Aleks <i>et al.</i>			
7	[2012]).			
8	For land-nadir, land-glint, and ocean-glint data the variance of the slopes is given			
9	by $1.28~\mathrm{ppm}/100~\mathrm{km}$ , $1.12~\mathrm{ppm}$ / $100~\mathrm{km}$ , and $0.48$ / $100\mathrm{km}$ respectively. To test			
10	whether these slowly varying changes explains the distribution of $XCO_2$ within small			
11	neight	porhoods we follow the same steps described in Section 3.2 and 3.3 but now		
12	add ar	other:		
13				
14	1)	Within each neighborhood replace the calculated measurement error with		
15		the "actual" measurement error as shown in the upper left panels of Figures		
16		2, 3, and 4 for an observation		
17	2)	Starting with the first observation (in time) within a neighborhood for		
18		Footprint #1, sample a value for the observation from the distribution of		
19		"actual" measurement errors. Label this the "modeled" observation.		
20	3)	For all subsequent observations in time for Footprint #1, sample each		
21		"modeled" observation from a distribution that is correlated with the		
22		modeled observation at the previous time step and has a variance		
23		corresponding to the "actual" measurement error.		
24	4)	For observations in Footprints #2-8, sampling each modeled observation		
25		from a distribution correlated with the modeled observation at the same		
26		time step in the previous (adjacent) footprint, again with a variance		
27		corresponding to the "actual" error.		
28	5)	Adjust each modeled observation with a linear function where the slope of		
29		the linear function is randomly chosen from the fitted Laplace distribution to		
30		the slopes (e.g., the Laplace function shown in Figure 8)		





1 6) Calculate variance of this "modeled" set of observations, for each

- neighborhood.
- 2 3

Figures 2, 3, and 4 (lower right panels) show the best overall agreement
between modeled distributions of XCO<sub>2</sub> relative to the mean and the expected
distributions based on observations, demonstrating that a slowly varying bias is
needed to best explain the observed distributions within a grid of approximately
100 km x 10 km.
The expected "true" variability across a typical 100 km neighborhood is ~0.1 to

10 ~0.3 ppm (e.g. Figure 1). Each typical observation has a random error related to

11 noise and a systematic error that is in principal bounded by the calculated

12 interference error (e.g. Boxe *et al.*, 2010) and is approximately 0.2 ppm. The 100 km

13 x 10.5 sizes for the small neighborhoods used for this analysis is a fortuitous size

14 because the expected latitudinal variability is approximately the same or smaller as

15 the mean interference error (Figure 1). Within a typical grid box an OCO-2 observed

16 measurement over land is within 1.28 / 2, or  $\sim$ 0.65 ppm of the mean XCO<sub>2</sub> value.

17 For these reasons, and we expect that a typical observation over land has at least a

systematic error of at least 0.65 ppm, about 2 to 3 times larger than the calculatedinterference error.

20 In contrast, the observed distributions of slopes and (mean slope of 0.48 ppm / 21 100 km or mean error of 0.24 ppm) for the ocean data is only 70% larger than the 22 mean calculated interference error of 0.14 ppm. Because the distribution of ocean data within "bins" (Figure 4, upper left panel) is also well described by the 23 24 calculated random error, we conclude that the ocean glint data is reasonably well 25 characterized by its calculated uncertainties for this size of a grid box, except for 26 calculated noise (or precision) uncertainties that are less than ~0.25 ppm. 27 We find no relationship between the distribution of slopes for a neighborhood 28 and the corresponding mean of the calculated interference error suggesting that the 29 calculated interference error does not explain the observed slope within a

30 neighborhood, in contrast to the measurement error. However, there is a





- 1 correlation between the slope and the estimated magnitude of interferences, such as
- 2 aerosol optical depth, surface albedo, and surface pressure. For example, the
- 3 correlation between the slopes of land-glint data with the mean uncertainty in the
- 4 interferences is 0.06 whereas the correlation between the observed slopes in XCO<sub>2</sub>
- 5 and similarly calculated observed slopes in aerosol optical depth is 0.37. This
- 6 correlation suggests that the observed slow variations in XCO<sub>2</sub> across a
- 7 neighborhood could be related to how interferences affect the XCO<sub>2</sub> estimate as
- 8 OCO-2 takes observations across a neighborhood.
- 9

#### 10 4.0 **Summary**

- 11
- 12 The analysis described in this paper uses the observed XCO<sub>2</sub> variability across
- 13 small neighborhoods, in comparison to expected variations, to evaluate the
- 14 precision and accuracy of the XCO<sub>2</sub> data. We find that the precision and accuracy of a
- 15 typical ocean measurement is approximately 0.35 and 0.2 ppm respectively,
- 16 consistent with the calculated errors (assuming that the accuracy is bounded by the
- 17 calculated interference error and does not include smoothing error). The precision
- 18 and accuracy of a typical land measurement (both nadir and glint) is approximately
- 19 0.75 ppm and 0.65 ppm. These values can be compared to the calculated
- 20 measurement and interference errors of approximately 0.36 ppm and 0.2 ppm.
- 21 Much of the difference between the observed precision and calculated measurement
- 22 error could be due to natural synoptic variability in XCO<sub>2</sub> but is also likely due to
- 23 non-linearities in the retrieval or random components of interference error. The
- 24 accuracy is estimated from observed gradients in XCO<sub>2</sub> of approximately 1.28 ppm /
- 25 100 km across the small neighborhoods used in this analysis. Natural variability can
- 26 likely explain at most about 0.1 to 0.3 ppm of this of 1.28 ppm. The accuracy is
- estimated as being at least half the value of this slope or  $\sim 0.65$  ppm.
- 28 This 0.65 ppm estimate for the accuracy of the land data could be a lower bound
- 29 because it is based on observed gradients across a region and not direct
- 30 comparisons against TCCON, although the OCO-2 data are bias corrected using





- 1 TCCON data (Wunch *et al.* 2011). We find a relationship between these gradients
- 2 and interferences such as aerosol optical depth and surface albedo suggesting that
- 3 these interferences are the cause of the gradients.
- 4 The analysis discussed in this paper can be applied to future versions of the
- 5 OCO-2 data in which more accurate calculations of the interferences are included or
- 6 additional data quality flags are used to remove spurious individual observations or
- 7 sets of observations. For example, another set of data quality flags could be
- 8 developed to remove observations that vary too much over a region. In addition,
- 9 Connor *et al.* (2016, submitted) finds that other instrumental and spectroscopic
- 10 uncertainties need to be included in the error analysis and that these additional
- 11 components will likely have a random and systematic component, thus possibly
- 12 explaining the discrepancy between calculated and actual uncertainties discussed
- 13 here. A future study in which the calculated uncertainties discussed in Connor *et al.*
- 14 (submitted) repeats the steps shown in this paper could be of great value for
- 15 explaining the observed variations across the small neighborhoods used in our
- 16 analysis.
- 17

# 18 Acknowledgements

- 19 Part of this research was carried out at the Jet Propulsion Laboratory, California Institute
- 20 of Technology, under a contract with the National Aeronautics and Space Administration.
- Funding for Susan Kulawik provided by NASA Roses NMO710771/NNN13D771T,
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- 25

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10

11 Figure 1: Distribution of latitudinal XCO<sub>2</sub> gradients as calculated by the high resolution, "Real

Time", Carbon Tracker model for November 2015 (left panel) and July 2015 (right panel) over
 North America and the nearby oceans. The latitude grid is 1 degree or ~110 km. The gradients

are re-scaled to 100 km for comparison to the XCO<sub>2</sub> gradients discussed in this paper.







4

5 Figure 2: Calculated, observed, and modeled uncertainties for Land-Nadir observations. Black

6 circles are the observed distributions and red circles are modeled distributions assuming 7 sampling and random error (upper right), correlated errors (bottom left) and correlated plus

- 8 trend in error (bottom right).
- 9







1 2 3







Figure 4: Observed and modeled distributions for Sea-Glint data.









6 "small neighborhood" used in this analysis.

- 7
- 8







5



Pearson's Correlation

7 neighborhood. (Bottom) correlation between observations for a single pixel.







- representative of one of the OCO-2 observations. The right panel shows the observed
  distribution (actual) and one calculated if the distributions were representative of the
- 7 calculated random error.
- 8

1 2 3

4

- 9
- 10







- 2 3 4 neighborhoods corresponding to Land Nadir observations.
- 5
- 6