



- 1 CFC-11, CFC-12 and HCFC-22 ground-based remote sensing FTIR measurements at Reunion  
2 Island and comparisons with MIPAS/ENVISAT data  
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## 14 Abstract

15 Profiles of CFC-11 (CCl<sub>3</sub>F), CFC-12 (CCl<sub>2</sub>F<sub>2</sub>) and HCFC-22 (CHF<sub>2</sub>Cl) have been obtained  
16 from Fourier transform infrared (FTIR) solar absorption measurements above the Saint-Denis (St  
17 Denis) and Maïlo sites at Reunion Island (21 °S, 55 °E) with low vertical resolution. FTIR profile  
18 retrievals are performed by the SFIT4 program and the detail retrieval strategies along with the  
19 systematic/random uncertainties of CFC-11, CFC-12, and HCFC-22 are discussed in this study.  
20 The FTIR data of all three species are sensitive to the whole troposphere and the lowermost  
21 stratosphere, with the peak sensitivity between 5 and 10 km. The trends derived from the  
22 combined St Denis and Maïlo FTIR time-series are  $-0.86 \pm 0.12\%$  and  $2.75 \pm 0.12\%$  for CFC-11 and  
23 HCFC-22, respectively, for the period 2004 to 2016, and  $-0.76 \pm 0.05\%$  for CFC-12 for 2009 to  
24 2016. These measurements are consistent with the trends observed by the National Oceanic and  
25 Atmospheric Administration (NOAA) Global Monitoring Division's (GMD) Halocarbons & other  
26 Atmospheric Trace Species Group (HATS) measurements at Samoa (14.2 °S, 170.5 °W) for  
27 CFC-11 ( $-0.87 \pm 0.04\%$ ), but slightly weaker for HCFC-22 ( $3.46 \pm 0.05\%$ ) and stronger for CFC-12  
28 ( $-0.60 \pm 0.02\%$ ).

29 The ground-based FTIR data have also been compared with the collocated Michelson  
30 Interferometer for Passive Atmospheric Sounding (MIPAS/ENVISAT) data, and found to be in  
31 good agreement: the observed mean relative biases and standard deviations of the differences  
32 between the smoothed MIPAS and FTIR partial columns (6-30 km) are ( $-4.3\%$  and  $4.4\%$ ), ( $-2.9\%$   
33 and  $4.6\%$ ) and ( $-0.7\%$  and  $6.0\%$ ) for CFC-11, CFC-12, and HCFC-22, respectively, which are  
34 within the combined error budgets from both measurements.

## 35 1. Introduction

36 CFC-11 (CCl<sub>3</sub>F), CFC-12 (CCl<sub>2</sub>F<sub>2</sub>) and HCFC-22 (CHF<sub>2</sub>Cl) are the major sources of chlorine  
37 in the stratosphere after photolytic decomposition, and therefore play an important role in  
38 stratospheric ozone depletion (Molina and Rowland, 1974). In addition, these gases absorb  
39 thermal infrared radiation and contribute significantly to the greenhouse effect (Lashof and Ahuja,  
40 1990). Due to the long lifetime of these gases (CFC-11 : ~60 years ; CFC-12: ~120 years;



41 HCFC-22: ~12 years (Ko et al., 2013)), they are good tracers to study transport and mixing  
42 processes in the upper troposphere and lower stratosphere region (Hoffmann and Riese, 2004).

43 Because of the vital importance of these gases, the Advanced Global Atmospheric Gases  
44 Experiment (AGAGE) in-situ network has been measuring CFC-11 and CFC-12 continuously  
45 since 1978 and HCFC-22 since the 1990s (Cunnold et al., 1997; Dunse et al., 2005). NOAA's  
46 Halocarbons & other Atmospheric Trace Species Group (HATS) sampling network started  
47 monitoring CFCs from flask grab samples in 1977 and via online in-situ techniques starting in  
48 1977 (Elkins et al., 1993). HCFC-22 was added to the NOAA/HATS measurements in 1992.  
49 Because of the use of chlorofluorocarbon (CFCs) as propellant and refrigerant in the 1980s, the  
50 in-situ measurements show the rapid rise of CFC-11 and CFC-12 at that time. To reduce  
51 substances that deplete the ozone layer, amongst others CFCs, 27 nations around the world signed  
52 a global environmental treaty, the Montreal Protocol, on September, 1987 (Murdoch and Sandler,  
53 1997). The hydrochlorofluorocarbons (HCFCs) were applied to replace the CFCs after the  
54 Montreal Protocol, since they react with tropospheric hydroxyl (OH), resulting in a shorter  
55 lifetime compared with CFCs. As a result, accelerated increases are observed for HCFCs since  
56 2004 in the global atmosphere (Montzka et al., 2009). The tropospheric concentrations of CFC-11  
57 and CFC-12 reached their maximums in 1992 and 2003 respectively, and a decline has been  
58 observed since then (Elkins et al., 1993; Montzka et al., 1996; Walker et al., 2000).

59 Apart from the in-situ measurements, observations of CFCs and HCFCs abundances have  
60 also been made using remote sensing infrared spectroscopy techniques. Space-based observations  
61 provide the global distributions and trends of CFCs and HCFCs; examples are the measurements  
62 of CFC-11 and CFC-12 from ILAS (Improved Limb Atmospheric Spectrometer), of CFC-11,  
63 CFC-12, CFC-113, HCFC-22, HCFC-142a and HCFC-142b from ACE-FTS (Atmospheric  
64 Chemistry Experiment - Fourier Transform Spectrometer) and of CFC-11, CFC-12 and HCFC-22  
65 from MIPAS (Michelson Interferometer for Passive Atmospheric Sounding) (Khosrawi et al.,  
66 2004; Hoffmann et al., 2008; Mahieu et al., 2008). Also ground-based FTIR measurements are  
67 able to monitor the CFCs and HCFCs (Notholt, 1994), especially at the Swiss Jungfraujoch station  
68 (Zander et al., 2005; Mahieu et al., 2010; Mahieu et al., 2013), where comparisons with the  
69 ACE-FTS measurements show a good agreement (Mahieu et al., 2015). They can provide long  
70 time-series of CFC-11, CFC-12, HCFC-22 total columns and are therefore very good candidates  
71 for supporting the evaluation of satellite and model data and for the evaluation of trends: the  
72 Jungfraujoch CFC-11, CFC-12, and HCFC-22 time series and trends have been included in the  
73 most recent Scientific Assessment of Ozone Depletion (Carpenter et al., 2014).

74 In this study, we provide the first ground-based FTIR time series of CFC-11, CFC-12 and  
75 HCFC-22 in the Southern Hemisphere, namely at two stations, both located at Reunion Island  
76 (21 S, 55 E): Saint-Denis (St Denis) and Ma'ïlo, and we compare them to MIPAS/ ENVISAT  
77 collocated data. Section 2 describes the FTIR experiments at Reunion Island as well as the  
78 retrieval strategy for each target CFC along with the uncertainty analysis. In addition, we provide  
79 the three species' trends derived from our FTIR time-series and compare them to the trends  
80 observed at American Samoa (SMO) from in-situ measurements for CFC-11 and CFC-12 and  
81 flask samplings for HCFC-22, which is one of NOAA's baseline observatories (14.2°S, 170.5°W,  
82 77m a.s.l.) at a similar latitude as Reunion Island. Vertical profile and partial column comparisons  
83 of the FTIR measurements with the MIPAS data are discussed in Sect. 3. Conclusions are drawn  
84 in Sect. 4.



## 85 2. Reunion Island FTIR data

### 86 2.1 FTIR experiments at Reunion Island

87 As explained in Baray et al., (2013), the atmospheric observations at Reunion Island are  
88 carried out at two sites, namely St Denis (20.90 °S, 55.48 °E, 85 m a.s.l.), close to the coast, and  
89 the Maido mountain site (21.07 S, 55.38 °E, 2155 m a.s.l.). At present, both sites are equipped  
90 with a Bruker 125HR FTIR instrument. These FTIR instruments contribute to two important  
91 networks: NDACC (Network for the Detection of Atmospheric Composition Change) and  
92 TCCON (Total Carbon Column Observing Network) dedicated to greenhouse gas observations.  
93 Each network requires a particular spectral coverage (mid-infrared (600-4500 cm<sup>-1</sup>) in NDACC  
94 and near-infrared (4000-8000 cm<sup>-1</sup>) in TCCON) and therefore a corresponding instrumental  
95 configuration (optical filters, beamsplitters and detectors) and operation mode (including spectral  
96 resolution). Since March 2013, when the FTIR spectrometer at Maido became operational, the  
97 Maido FTIR has been dedicated to NDACC and the St Denis FTIR to TCCON.

#### 98 2.1.1 La Reunion - St Denis

99 The Royal Belgian Institute for Space Aeronomy (BIRA-IASB) started the FTIR solar  
100 absorption experiments at La Reunion in St Denis in 2002, with a Bruker 120M FTIR  
101 spectrometer, first on a campaign basis with campaigns in 2002 (October), 2004 (August to  
102 October) and 2007 (May to November), and then in continuous mode since June 2009 (Senten et  
103 al., 2008; Vigouroux et al., 2009; Duflot et al., 2010; Vigouroux et al., 2012; Baray et al., 2013). In  
104 September 2011, BIRA-IASB started the replacement of the Bruker 120M by a Bruker 125HR:  
105 the Bruker 125HR was installed next to the Bruker 120M and both instruments were set up to  
106 make collocated measurements until November 2011, when BIRA-IASB disassembled the Bruker  
107 120M. Since then, at St Denis, BIRA-IASB operates only the Bruker 125HR.

108 Since October 2013 the instrument is primarily dedicated to TCCON measurements and no  
109 more NDACC measurements have been made. Because the CFC retrieval windows requires a KBr  
110 beamsplitter and a MCT detector (600 to 1400 cm<sup>-1</sup>), the CFC data presented in this work at  
111 Saint-Denis and requiring the NDACC observational configuration cover the August 2004 -  
112 November 2011 period.

#### 113 2.1.2 La Reunion – Ma ïlo

114 BIRA-IASB started operating a second Bruker 125HR FTIR spectrometer at the Ma ïlo  
115 observatory in March 2013 and dedicated it primarily to NDACC measurements with MCT and  
116 InSb detectors. As such, our CFC time series at Ma ïlo cover the March 2013 – present time  
117 period.

### 118 2.2 FTIR retrieval

119 The NDACC ground-based FTIR experiment observes the absorption of the direct solar  
120 radiation with high spectral resolution (0.0035-0.0110 cm<sup>-1</sup>) and uses the pressure broadening  
121 effect of absorption lines to retrieve volume mixing ratio (vmr) profiles of target gases. In this  
122 study, the FTIR retrievals are based on an optimal estimation method (Rodgers, 2000), carried out  
123 with the SFIT4 algorithm (<https://wiki.ucar.edu/display/sfit4>), which is an open source code,  
124 jointly developed at the NASA Langley Research Center, the National Center for Atmospheric  
125 Research (NCAR) and the National Institute of Water and Atmosphere Research (NIWA). The



126 National Centers for Environmental Prediction (NCEP) provide the 6-hourly pressure and  
 127 temperature profiles. The difference between NCEP and the balloon sounding measurements  
 128 above Reunion Island is used to create the systematic and random error covariance matrices of the  
 129 water vapor and temperature profiles. HBr cell measurements performed on a daily basis to verify  
 130 the alignment of the instrument and to obtain the instrument line shape (ILS) using the  
 131 LINEFIT14.5 program (Hase et al., 1999); the ILS is provided as an input parameter in the  
 132 forward model of SFIT4.

### 133 2.2.1 Retrieval strategy

134 CFC-11, CFC-12 and HCFC-22 have weak absorptions in the infrared spectral range,  
 135 requiring careful selection of the retrieval spectral windows in order to minimize the interfering  
 136 absorptions from other species. The microwindows (see Table 1) are the same as in the work of  
 137 Mahieu et al. (2010), except for CFC-12: the 922.50-923.60  $\text{cm}^{-1}$  window is however not  
 138 appropriate for the humid site of Reunion Island because of the strong water vapor lines present at  
 139 the edge of the window (922.13  $\text{cm}^{-1}$ ). To avoid such interferences from water vapor, we prefer to  
 140 use the 1160.2-1161.4  $\text{cm}^{-1}$  for our CFC-12 retrievals at Reunion Island. The left panels in Fig. 1  
 141 show the typical transmittances along with the absorption lines of the target and interfering species,  
 142 and the fitting residuals for the CFC-11, CFC-12 and HCFC-22 retrievals at St Denis. The  
 143 interfering gases are also listed in Table 1: as indicated in the Table, either a full profile retrieval is  
 144 performed or only a scaling of the a priori profile (column retrieval).

145 In each microwindow, the background transmittance  $\beta$  describes the shape caused by the  
 146 optics in the instrument (especially the bandpass filter) as a second order polynomial of the  
 147 wavenumber:

$$148 \quad \beta = [1 + a(\mathbf{w} - w_0) + b(\mathbf{w} - w_0)^2] / (1 + z_0) \quad (1).$$

149 In Equation (1),  $a$  is the slope coefficient,  $b$  is the curvature coefficient,  $w_0$  is the first  
 150 wavenumber in the microwindow ( $\text{cm}^{-1}$ ),  $\mathbf{w}$  is the vector of all wavenumber in the microwindow,  
 151 and  $z_0$  is the zero level offset (zshift). The parameters  $a$ ,  $b$ , and  $z_0$  can be fitted in SFIT4  
 152 in addition to the target gases profiles and interfering species profiles or columns. The calculated  
 153 transmittance  $y_c$  is the result of bringing the background, absorptions and zshift together:

$$154 \quad y_c = \beta \cdot (\zeta(\mathbf{t}(\mathbf{w})) + z_0) \quad (2),$$

155 Where  $\mathbf{t}(\mathbf{w})$  is the calculated transmittance (after the absorption from each species and solar  
 156 lines) and  $\zeta(\mathbf{t}(\mathbf{w}))$  is the transmittance after convolution with the ILS. If necessary, a beam  
 157 correction can also be applied to fitting the baseline of the transmittance in each microwindow. It  
 158 creates a zshift-like parameter  $z_b$  for the interferogram perturbation (IP) model:

$$159 \quad y_c = \beta (\zeta(\mathbf{t}) + z_0 + z_b) \quad (3),$$

160 SFIT4 uses 4 parameters ( $A$ : amplitude;  $T$ : period;  $\varphi$ : phase and  $\tau$ : slope for the amplitude)  
 161 to retrieve each beam (the maximum number of beams is 20):

$$162 \quad \mathcal{G} = A(1 + \tau(\mathbf{w} - w_0))e^{i(2\pi/T(\mathbf{w} - w_0))} \quad (4),$$

$$163 \quad z_b = \zeta(\mathcal{G}) \quad (5),$$



164 Table 1 lists the parameters used for fitting the background in the CFC-11, CFC-12 and  
165 HCFC-22 retrievals. Since the retrieval windows of CFC-12 and HCFC-22 are narrow, linear fit is  
166 enough to characterize the spectral background ( $b=0$ ). However, the retrieval window of CFC-11  
167 ( $830\text{--}860\text{ cm}^{-1}$ ) is very wide and contains several saturated  $\text{H}_2\text{O}$  absorption lines, therefore zshift,  
168 slope, and curvature were retrieved together to fit the oscillating shape of the background in the  
169 CFC-11 microwindow, and for the CFC-11 retrieval at Ma ïlo, it turned out necessary to retrieve  
170 also one IP-type beam. Together with one IP-type beam, the retrieved CFC-11 total columns at  
171 Ma ïlo show better agreements with MIPAS data and the ground-based HATS SMO in-situ  
172 measurements along with a much smaller fitting residual in comparison with without-beam  
173 retrievals. Fig. 2 shows the average residual transmittance for the CFC-11 retrievals at Ma ïlo,  
174 with and without fitting a beam parameter. The spikes mainly result from the strong absorption  
175 lines of  $\text{H}_2\text{O}$ . The a priori values for the IP beam parameters were obtained by fitting the mean  
176 residuals of all without-beam retrievals. It is clear from Fig. 2 that adding one IP-type beam was  
177 useful to remove the background oscillation of the residuals at Ma ïlo. As such oscillation was not  
178 found for the St Denis CFC-11 residuals, the beam parameters retrieval was only applied for  
179 Ma ïlo.

180 We use the empirical pseudo-line-lists (PLL) created by G. Toon (details see  
181 <http://mark4sun.jpl.nasa.gov/pseudo.html>) for the CFC-11, CFC-12, HCFC-12 and  $\text{COCl}_2$   
182 spectroscopy, and HITRAN 2012 (Rothman et al., 2013) for the remaining species (see Table 1).  
183 The a priori profiles of interfering gases, except  $\text{H}_2\text{O}$  and  $\text{O}_3$ , are the mean of 1980-2020 monthly  
184 data from Whole Atmosphere Community Climate Model (WACCM, version 6,  
185 <ftp://acd.ucar.edu/user/jamesw/IRWG/2013/WACCM/V6/>). In order to reduce the influence of  $\text{O}_3$   
186 and  $\text{H}_2\text{O}$  uncertainties, preliminarily retrieved profiles of  $\text{O}_3$  and  $\text{H}_2\text{O}$  obtained with the settings of  
187 Vigouroux et al. (2015) are used, as input for the CFC-11, CFC-12 and HCFC-22 retrievals. The a  
188 priori profiles of target species are the mean of 2004-2016 monthly data from WACCM after  
189 scaling to the annual mean of ground-based NOAA/HATS SMO flask grab samples and in-situ  
190 measurements of 2009 for St Denis and 2014 for Ma ïlo. As such, the a priori profiles of CFC-11  
191 and CFC-12 for St Denis are a little larger than the ones for Ma ïlo, while the a priori profile of  
192 HCFC-22 for St Denis is a little smaller than that for Ma ïlo (see Fig. 2). However, all the  
193 retrievals at St Denis or at Ma ïlo use the same a priori profiles. The profiles of the three gases  
194 decrease rapidly above 20 km and become close to zero vmr values at 30 km for CFC-11, 40 km  
195 for CFC-12 and 100 km for HCFC-22.

196 The a priori covariance matrix (regularization matrix) is another important input parameter in  
197 the optimal estimation method. Ideally, the diagonal values of covariance matrix represent the  
198 natural variability of the gas concentration around the a priori profile. Therefore, in our study,  
199 2004-2016 monthly data from WACCM are used to provide the variability for the FTIR retrieval,  
200 which agrees with the a priori profile ensemble. The variabilities of CFC-11, CFC-12 and  
201 HCFC-22 are then 5%, 2% and 15%, respectively. The gas profile correlation width is set to 4 km  
202 from 0 to 100 km in the SFIT4 retrieval, and the retrieved profiles for CFC-11, CFC-12 and  
203 HCFC-22 are shown in Fig. 2.

204 Table 1 lists the DOFS of the total columns of CFC-11, CFC-12 and HCFC-22, along with  
205 the standard deviation ( $1\sigma$ ); they are  $1.1\pm 0.1$ ,  $1.5\pm 0.1$ ,  $0.9\pm 0.1$  respectively at St Denis and  $1.1\pm 0.1$ ,  
206  $1.6\pm 0.1$ ,  $1.1\pm 0.1$  respectively at Ma ïlo. The right panels of Fig. 1 shows the typical averaging  
207 kernels of the CFC-11, CFC-12 and HCFC-22 retrievals at St Denis; they represent the vertical



208 sensitivity of the measurement as a function of altitude. The FTIR retrievals of all three species are  
 209 sensible to the whole troposphere and the lowermost stratosphere, with the peak sensitivity around  
 210 5-10 km. We have to keep in mind that the retrieved profiles of CFC-11, CFC-12 and HCFC-22  
 211 have very poor vertical resolution: the DOFS range from 0.9 to 1.6 and the  
 212 full widths at half maximum of the averaging kernels are very wide (~8 km).

### 213 2.2.2 Error budget

214 According to the optimal estimation method (Rodgers, 2000), the final state  $\hat{x}$  satisfies

$$215 \hat{x} = x_a + \hat{G}(y - F(\hat{x}, b) + \hat{K}(\hat{x} - x_a)) \quad (6),$$

$$216 \hat{G} = (S_a^{-1} + \hat{K}^T S_\epsilon^{-1} \hat{K})^{-1} \hat{K}^T S_\epsilon^{-1} \quad (7),$$

217 where  $x_a$  is the a priori state vector;  $\hat{G}$  is the contribution function, indicating the  
 218 sensitivity of the retrieval to the measurements;  $\hat{K}$  is the weighing function, representing the  
 219 sensitivity of the measurements to the state vector;  $y$  is the observed spectrum and  $F(\hat{x}, b)$  is  
 220 the forward model (with model parameters represented by  $b$ ) evaluated in the final state;  $S_a$   
 221 is the a priori covariance matrix and  $S_\epsilon$  is the measurement error covariance matrix. If we consider  
 222 the different uncertainty components, formula (6) can be approximated as,

$$223 \hat{x} = x_a + \hat{G}(F(x_t, b) + \epsilon_f + \epsilon_y + K_b b - F(\hat{x}, b) + \hat{K}(\hat{x} - x_a)) \quad (8),$$

$$= x_a + \hat{G}(\hat{K}(x_t - \hat{x}) + \hat{K}(\hat{x} - x_a)) + \hat{G}(\epsilon_f + \epsilon_y + K_b b)$$

224 in which,  $x_t$  is the true state of the atmosphere;  $\epsilon_f$  is the forward model error,  $\epsilon_y$  is the  
 225 measurement noise;  $K_b$  is the sensitivity of the measurements to the forward model parameters  
 226  $K_b = \partial F(x, b) / \partial b$ . It is worth noting that this equation is approximated using a Newton  
 227 iterative algorithm, which is also subject to an error, but the convergence criterion guarantees that  
 228 the iteration error gets smaller than the noise error on the spectrum. So it is ignored.

229 We can rewrite formula (8) as

$$230 \hat{x} - x_t = (\hat{A} - I_n)(x_t - x_a) + \hat{G}(\epsilon_f + \epsilon_y + K_b b) \quad (9),$$

231 where  $\hat{A} = \hat{G} \hat{K}$  is the averaging kernel.

232 The first term in the right side of the equation (9) is the smoothing error, the second term  
 233 contains three parts: the forward model error; the measurement error and the forward model  
 234 parameters error. The forward model parameters error comes from the atmospheric (temperature, a  
 235 priori profiles, pressure ...), spectroscopic, geometrical and instrumental parameters, which are  
 236 not included in the state vector, but do have an impact on the forward model calculation. Each  
 237 error contains both a systematic and a random part.

238 Tables 2 and 3 list the different contributions to the total average retrieval uncertainty, at St  
 239 Denis and Maïlo, resp., including smoothing, measurement noise, retrieval parameters (slope;  
 240 curvature; wavenumber shift; zero-level offset (zshift); beam parameters; solar line shift; simple  
 241 phase correction), interfering species, temperature profile, solar zenith angle (SZA), spectroscopic  
 242 parameters (line intensity, air-broadened half-width, temperature dependence of the air-broadened  
 243 half-width). We assume that the measurement and retrieval parameters have very small systematic  
 244 uncertainties (set to zero in our case) and that the spectroscopic parameters have negligible  
 245 random errors. Because zshift is in the state vector for CFC-11, the zshift uncertainty is included in  
 246 retrieval parameters uncertainties but in the model parameters uncertainties for CFC-12 and



247 HCFC-22. The total average systematic/random uncertainties associated with the retrieved  
 248 columns for CFC-11, CFC-12, and HCFC-22 are 7.0%/2.0%, 1.8%/1.1%, 4.4%/4.5%, respectively,  
 249 at St Denis and 6.7%/1.6%, 1.8%/1.1%, 4.5%/4.1% respectively at Ma ĩlo. The systematic  
 250 uncertainties originate mainly in the uncertainties on the spectroscopic parameters, as well as in  
 251 the temperature uncertainty. The random uncertainty is dominated by the smoothing error, the  
 252 uncertainty on the SZA and the measurement noise; especially for HCFC-22, the measurement  
 253 noise error is very significant due to the narrow and weak absorption of HCFC-22 (see the left  
 254 bottom panel in Fig. 1).

### 255 2.2.3 Total column time series

256 Fig. 4 shows the time series of retrieved total columns of CFC-11, CFC-12 and HCFC-22 at  
 257 St Denis and Ma ĩlo, together with their uncertainties (in unit of molecules/cm<sup>2</sup>). Although the  
 258 retrievals use the same database of spectra, the numbers of successful retrievals for CFC-11,  
 259 CFC-12 and HCFC-22 are different. Fig. 4 indicates that the time series of CFC-12 has the largest  
 260 data density, because the 1160.2-1161.4 cm<sup>-1</sup> microwindow falls in the middle of the spectrum  
 261 with high signal to noise ratio (SNR), while the microwindows 830.0-860.0 cm<sup>-1</sup> and 828.75-829.4  
 262 cm<sup>-1</sup> lie in the edge of the spectrum, with lower SNR due to the optical filter shape. Fig. 4 also  
 263 shows that there is an offset between the total columns of all three species at St Denis and Ma ĩlo,  
 264 since the altitude of St Denis (85 m a.s.l.) is much lower than that of Ma ĩlo (2155 m a.s.l.),

### 265 2.2.4 Trend analysis and seasonal cycle

266 Because of the different altitudes of St Denis and Ma ĩlo, it is difficult to compare the total  
 267 columns of St Denis and Ma ĩlo directly. Therefore, for the trend analysis we extract from the  
 268 profiles at St Denis the partial columns (2.155-100 km) of CFC-11, CFC-12 and HCFC-22. We  
 269 plot in Fig.5 these partial columns at St Denis (light coral) together with the total columns at  
 270 Ma ĩlo, and we can see that they are in good agreement. Also shown for comparison in Fig.5, are  
 271 the in-situ and flask daily mean measurements at SMO. We use the Chromatograph for  
 272 Atmospheric Trace Species (CATS) in-situ daily mean data for CFC-11 and CFC-12  
 273 (<http://www.esrl.noaa.gov/gmd/hats/insitu/cats/>) and the flask data for HCFC-22 (Montzka et al.,  
 274 1993). The precision of the in-situ and flask measurements is about a few of tenths of a ppt for  
 275 CFC-11, CFC-12 and HCFC-22. Our FTIR measurements could capture the main trends of these  
 276 species very well, but the scatters of the FTIR retrievals are much larger than the in-situ  
 277 measurements mainly due to the larger retrieval uncertainties of the FTIR measurements, and the  
 278 FTIR columns is also associated with the scatter on the air column (e.g., due to tropopause shifts  
 279 and P-variations).

280 To derive the secular trends from the FTIR and in-situ measurements daily means  $Y(t)$ , with  $t$   
 281 the time in fractional year, we use a regression model that includes a Fourier series (3<sup>rd</sup> order) to  
 282 describe the seasonal cycle:

$$283 \quad Y(t) = A_0 + A_1 \cdot t + \sum_{k=1}^3 A_{2k} \cos(2k\pi t) + A_{2k+1} \sin(2k\pi t) + \varepsilon(t) \quad (10),$$

284 where  $A_0$  is the intercept at  $t=0$ ,  $A_1$  is the secular (annual) trend,  $A_2$  to  $A_7$  are the seasonal cycle  
 285 parameters, and  $\varepsilon(t)$  are the residuals between the observations and the model. The  
 286 auto-correlation in the residuals must be taken into account to avoid the underestimation of trend  
 287 uncertainties. We follow the approach of (Santer et al., 2000), which by combining their Eqs. 3 to  
 288 6, leads us to the following corrected uncertainty  $\sigma_c$  on the regression parameters:



$$289 \quad \sigma_c = \sigma_d * \frac{(n-2)}{(n_e-2)} \quad (11),$$

290 with  $\sigma_d$  the uncertainty directly provided by the regression model,  $n$  the number of daily means in  
 291 the  $Y(t)$  time-series, and  $n_e$  the effective sampling size:

$$292 \quad n_e = n * \frac{1-r}{1+r} \quad (12),$$

293 with  $r$  the auto-correlation, with a time-lag of 1, in the residuals.

294 Table 4 gives the annual percent changes and their uncertainties of CFC-11, CFC-12 and  
 295 HCFC-22 (%) from both FTIR at Reunion Island and from the in-situ and flask measurements at  
 296 SMO. The trends from MIPAS, ACE, and Jungfraujoch FTIR measurements in Carpenter et al.  
 297 (2014) are also listed in Table 4 for comparison. The ACE-FTS (2004-2010) trends were  
 298 determined by averaging the mixing ratios in molecule-dependent altitude ranges within tropical  
 299 occultation (30°N-30°S), the MIPAS (2004-2010) trends were calculated by 10-15 km partial  
 300 column for the 20°N-20°S, and Jungfraujoch is the total columns of FTIR measurements. The  
 301 annual percent change, in this study, is defined as the ratio of the annual change to the mean of all  
 302 the measurements that are used to do the trend analysis. Since the time range of Maïlo  
 303 measurements only covers about 3 years, we do not perform trend analysis on Maïlo data only.  
 304 For CFC-11, the total column annual change at St Denis (2004-2011) is  $-0.69 \pm 0.15\%$ , which is  
 305 slightly weaker than the one derived from SMO measurements ( $-0.89 \pm 0.01\%$ ). It is also weaker  
 306 than the trends reported by MIPAS, ACE, and Jungfraujoch FTIR measurements. However, the  
 307 annual change of the combined FTIR partial columns at St Denis and total columns at Maïlo  
 308 (2004-2016) is very close to the SMO measurements ( $-0.86 \pm 0.12\%$  vs.  $-0.87 \pm 0.04\%$ ). For CFC-12,  
 309 the total column annual change at St Denis (2004-2011) is  $-0.26 \pm 0.10\%$ , which is also slightly  
 310 weaker than the one derived from SMO measurements ( $-0.37 \pm 0.08\%$ ), but in agreement within the  
 311 estimated uncertainties. It is also in agreement with the ACE-FTS and Jungfraujoch reported  
 312 trends. Fig. 4 shows that the concentration of CFC-12 has a significant trend change around 2004  
 313 (increasing before 2004 and decreasing after), therefore, it is better to select the data after 2009 to  
 314 do the trend analysis. The annual change of the combined partial columns at St Denis and total  
 315 columns at Maïlo (2009-2016) is stronger than that derived from the SMO measurements  
 316 ( $-0.76 \pm 0.05\%$  vs.  $-0.60 \pm 0.02\%$ ). For HCFC-22, the annual change of St Denis partial columns  
 317 ( $3.14 \pm 0.43\%$ ) is close to that reported from ACF-FTS data and Jungfraujoch measurements for  
 318 approximately the same period, but slightly smaller than that derived from the SMO  
 319 measurements ( $4.04 \pm 0.06\%$ ). The trend of the combined partial columns at St Denis and total  
 320 columns at Maïlo (2004-2016) is also smaller than that of the SMO measurements ( $2.75 \pm 0.12\%$   
 321 vs.  $3.46 \pm 0.05\%$ ).

322 Fig. 6 shows the seasonal cycles of CFC-11, CFC-12 and HCFC-22. The red lines represent  
 323 the modeled seasonal cycle obtained by Eq. 10 for the 2004-2016 St Denis -Maïlo time-series,  
 324 and blue lines represent the mean of FTIR measurements for each month during the 2004-2016  
 325 period, after subtraction of the trend (removing the systematic retrieval error), together with the  
 326 standard deviation  $\sigma$  on the mean (thin error bars). The standard deviation  $\sigma$  represents the random  
 327 error of the FTIR retrievals, and the  $2\sigma$  error on the mean ( $2\sigma / \sqrt{n}$ ;  $n$  being the number of  
 328 measurements for each month) is also shown with thick blue lines. For CFC-11 and HCFC-22,  
 329 there is no obvious seasonal variation, since the uncertainty is very large. However, there is a  
 330 significant seasonal variation for CFC-12, for which the concentration is highest in March-May



331 and lowest in September-November.

### 332 **3. Comparison with ENVISAT/MIPAS data**

#### 333 **3.1 MIPAS introduction**

334 ENVISAT was successfully launched into space on March 1, 2002 carrying several sensors,  
335 including MIPAS, a cryogenic limb emission Fourier transform spectrometer (FTS) which  
336 observes many trace gases from a wide spectrum covering 865-2410  $\text{cm}^{-1}$  (Fischer and Oelhaf,  
337 1996). The mission ended on 08 April 2012. From July 2002 to March 2004, MIPAS was operated  
338 in full spectral resolution (FR) mode (spectral resolution: 0.05  $\text{cm}^{-1}$ ), covering the altitude range  
339 from 6 km to 68 km. Due to the failure of one of the interferometer slides, MIPAS was operated  
340 with a reduced spectral resolution (0.121  $\text{cm}^{-1}$ ), the so-called RR mode, starting January 2005. The  
341 RR mode covers the altitude range from 6 km to 70 km (Fischer et al., 2008). In this paper, we use  
342 the Institute of Meteorology and Climate Research (IMK) generated MIPAS/ENVISAT products,  
343 taken from <https://www.imk-asf.kit.edu/english/308.php>. The retrieval windows of MIPAS are  
344 831.0-853.0  $\text{cm}^{-1}$  for CFC-11; 915.0-925.0  $\text{cm}^{-1}$  for CFC-12; 803.50-804.75  $\text{cm}^{-1}$ , 808.25-809.75  
345  $\text{cm}^{-1}$ , 820.50-821.12  $\text{cm}^{-1}$  and 828.75-829.50  $\text{cm}^{-1}$  for HCFC-22. A dedicated spectroscopic  
346 database was applied for MIPAS retrieval (Flaud and Teffo, 2003). The detailed MIPAS CFCs  
347 retrieval strategies can be found in previous publications (Hoffmann et al., 2005; Kellmann et al.,  
348 2012; Chirkov et al., 2016); all the products have been validated to some degree by comparison  
349 with other space experiments, air-borne in-situ instruments, ground-based measurements, or  
350 independent ENVISAT MIPAS analyses (Hoffmann et al., 2008).

#### 351 **3.2 Vertical profile comparison**

352 There is no temporal overlap between MIPAS data and Maïlo measurements, so the MIPAS  
353 footprints within  $\pm 2$  latitude  $\pm 5$  longitude around St Denis are selected to compare with the St  
354 Denis FTIR measurements. The overpass times of MIPAS above Reunion Island are around 6:30  
355 and 18:30 UTC, due to the sun-synchronous orbit of ENVISAT. As the FTIR measurements are  
356 recorded only during daytime, the MIPAS data around 6:30 UTC are chosen in the following  
357 analysis.

358 Fig. 7 shows the comparison of averaged profiles between FTIR measurements and MIPAS  
359 data. The individual FTIR-MIPAS data pair was selected when the FTIR measurement and the  
360 MIPAS observation were collocated within  $\pm 3$  hours around 6:30 UTC on the same day. If more  
361 than one MIPAS datapoint was found on a given day, the closest (in geodetic distance) MIPAS  
362 data point was taken. If more than one FTIR measurement exists on a given day, each FTIR  
363 measurement together with the closest MIPAS datapoint will be taken as one individual data pair.  
364 In total, there are 60, 86 and 42 FTIR-MIPAS data pairs for CFC-11, CFC-12 and HCFC-22,  
365 respectively. It is worth noting that, to account for the sensitivity of the retrieval to the true  
366 profiles and to take into account the low vertical resolution of the FTIR retrieved profiles, the  
367 MIPAS profile  $\hat{\mathbf{x}}_{\text{MIPAS}}$  is smoothed by the FTIR averaging kernel (AK)  $\mathbf{A}_{\text{FTIR}}$  (Rodgers and  
368 Connor, 2003):

$$369 \quad \hat{\mathbf{x}}'_{\text{MIPAS}} = \mathbf{x}_{\text{FTIR}}^{\text{a priori}} + \mathbf{A}_{\text{FTIR}}(\hat{\mathbf{x}}_{\text{MIPAS}} - \mathbf{x}_{\text{FTIR}}^{\text{a priori}}) \quad (13),$$

370 Where  $\mathbf{x}_{\text{FTIR}}^{\text{a priori}}$  is the FTIR a priori profile and  $\hat{\mathbf{x}}'_{\text{MIPAS}}$  is the MIPAS profile after



371 smoothing correction is applied. The MIPAS profile was interpolated onto the FTIR retrieval grids  
372 (keeping the total column unchanged). Fig. 7 mainly focuses on the vertical range from 6 to 30 km,  
373 because there aren't any MIPAS measurements below 6 km (Fischer et al., 2008) and the FTIR  
374 sensitivity is very weak above 30 km (see Fig. 1). In addition, the low FTIR sensitivity above 30  
375 km leads to a very small relative difference (less than 1%) between the smoothed MIPAS and  
376 FTIR  $((\text{MIPAS-FTIR})/\text{FTIR} \times 100\%)$  above 30 km for all three species (see right panels in Fig. 7).  
377 Since the FTIR retrievals have very poor vertical resolution, the “oscillation” of the profiles of the  
378 relative difference between FTIR and MIPAS could be caused by the FTIR retrievals. Anyhow, for  
379 CFC-11, the FTIR concentration is larger than the smoothed MIPAS concentration value between  
380 6 km and 30 km: the largest difference is of order -7% around 15 km. For CFC-12, the FTIR  
381 retrieval is larger below 14 km and smaller above 14 km than the smoothed MIPAS data. The  
382 peaks are around 6 km (-8%) and 18 km (2%). For HCFC-22, the FTIR retrieval is very close to  
383 the smoothed MIPAS data: the relative difference is within  $\pm 5\%$  between 6 and 30 km.

### 384 3.3 Partial column comparison

385 In this section, we compare the MIPAS and St Denis FTIR partial columns (PC) from 6 to 30  
386 km, for the same collocated pairs as in Sect. 3.2. The DoFS of the partial columns of CFC-11,  
387 CFC-12, HCFC-22 are  $0.6 \pm 0.1$ ,  $0.9 \pm 0.1$ ,  $0.6 \pm 0.1$ . Table 5 exhibits the statistical values of the  
388 comparison: relative bias and standard deviation of the difference between the MIPAS (raw and  
389 smoothed) and FTIR, together with the partial column uncertainties from both data sets. The  
390 largest mean relative bias is found for CFC-11 (-4.3%), showing that the FTIR partial columns are  
391 larger than the MIPAS ones, which is probably caused by the large systematic error of FTIR  
392 CFC-11 retrievals (10.5%). For CFC-12, the bias is -2.7% which is also within the uncertainty  
393 budget of combined data sets. The lowest relative bias is found for HCFC-22 (-0.7%) but the  
394 standard deviation is large (6.0%) because of the large FTIR retrieval errors (7.8%/7.1%) and  
395 MIPAS retrieval error (5.0%). Overall, the biases and standard deviations between the two data  
396 sets lie within the uncertainty budgets for the three species.

397 Fig. 8 shows the time series of the monthly means of partial columns of CFC-11, CFC-12 and  
398 HCFC-22 FTIR measurements at St Denis (grey) along with the raw MIPAS data (red). The  
399 smoothed MIPAS data are not shown here, because more than half of MIPAS data do not  
400 correspond with an individual FTIR measurement within one day or even one week, and the  
401 differences between partial columns of smoothed and unsmoothed data are within 1.0%. Note that  
402 the bias between the raw MIPAS and FTIR data also contains the smoothing error, but the bias  
403 already lies within the uncertainty budget even without smoothing error (see Table 5). Fig. 8  
404 shows that the monthly means of MIPAS and FTIR data are in a good agreement.

## 405 4. Summary

406 CFC-11, CFC-12 and HCFC-22 mixing ratio profiles were retrieved at Reunion Island from  
407 St Denis and Ma'ïlo ground-based solar absorption FTIR measurements between 2004 and 2016.  
408 The retrieval microwindows are carefully selected to minimize the interfering absorptions from  
409 other species. The averaging kernels of CFC-11, CFC-12 and HCFC-22 are very similar, and the  
410 retrieved information comes mainly from the troposphere and lower stratosphere with low vertical  
411 resolution. As expected as a response to the Montreal Protocol, negative trends of total columns of  
412 CFC-11 and CFC-12 and a positive trend of HCFC-22 were observed at St Denis and Ma'ïlo,



413 which is in good agreement with the in-situ surface data and other remote sensing results (E.g.,  
414 SMO in-situ and flask measurements and Jungfraujoch FTIR data, resp.). The observed FTIR total  
415 column trends above St Denis between 2004 and 2011 are  $-0.69 \pm 0.15\%/yr^{-1}$  for CFC-11,  
416  $-0.26 \pm 0.10\%/yr^{-1}$  for CFC-12 and  $3.14 \pm 0.43\%/yr^{-1}$  for HCFC-22. The trends of combined FTIR  
417 partial columns (2.155-100 km) at St Denis and total columns at Maïlo are  $-0.86 \pm 0.12\%$  for  
418 CFC-11 and  $2.75 \pm 0.12\%$  for HCFC-22 between 2004 and 2016, and  $-0.76 \pm 0.05\%$  for CFC-12  
419 between 2009 and 2016. These trends are consistent with the ones observed at SMO for CFC-11  
420 ( $-0.87 \pm 0.04\%$ ), but slightly smaller for HCFC-22 ( $3.46 \pm 0.05\%$ ) and larger for CFC-12  
421 ( $-0.60 \pm 0.02\%$ ).

422 The FTIR measurements were also compared with collocated MIPAS/ENVISAT data around  
423 St Denis. There are 60, 86 and 42 FTIR-MIPAS collocated data pairs for CFC-11, CFC-12 and  
424 HCFC-22 within  $\pm 2^\circ$  latitude,  $\pm 5^\circ$  longitude and  $\pm 3$  hours around 6:30 UTC. The differences  
425 between FTIR and smoothed MIPAS profiles from 6 to 30 km altitude are within  $\pm 10\%$  for  
426 CFC-11 and CFC-12, and  $\pm 5\%$  for HCFC-22. The relative biases and standard deviations of the  
427 differences between the partial columns (6-30 km) of smoothed MIPAS and FTIR are  $-4.3\%$   
428  $\pm 4.4\%$ ,  $-2.9\% \pm 4.6\%$  and  $-0.7\% \pm 6.0\%$  for CFC-11, CFC-12, and HCFC-22, respectively, which lie  
429 within the error budgets from both data sets. Overall, the time series of MIPAS monthly partial  
430 columns show a good agreement with the St Denis FTIR partial column data.

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578



579 Table 1. Microwindows, interfering gases, spectroscopic database, a priori profile and background  
 580 parameters (slope, curvature, zshift and beam as discussed in Section 2.2.1) used for the SFIT4  
 581 retrievals of CFC-11, CFC-12 and HCFC-22. The degree of freedom (DOF, mean and standard  
 582 deviation) of retrievals at St Denis and Ma ĩo.

Target gas	CFC-11 (CCl <sub>3</sub> F)	CFC-12 (CCl <sub>2</sub> F <sub>2</sub> )	HCFC-22 (CHF <sub>2</sub> Cl)
Microwindows (cm <sup>-1</sup> )	830.0-860.0	1160.2-1161.4	828.75-829.4
Profile retrieval	CFC-11, H <sub>2</sub> O	CFC-12, N <sub>2</sub> O	HCFC-22
Column retrieval	HNO <sub>3</sub> O <sub>3</sub> COCl <sub>2</sub> CO <sub>2</sub>	O <sub>3</sub> CH <sub>4</sub> H <sub>2</sub> O	CO <sub>2</sub> H <sub>2</sub> O O <sub>3</sub>
Spectroscopy	PLL, HITRAN2012	PLL, HITRAN2012	PLL, HITRAN2012
A priori profile	WACCM	WACCM	WACCM
St Denis background	slope, curvature, zshift	slope	slope
Ma ĩo background	slope, curvature, zshift, beam	slope	slope
DOFS (St Denis)	1.1±0.1	1.5±0.1	0.9±0.1
DOFS (Ma ĩo)	1.0±0.1	1.6±0.1	1.1±0.1

583



584 Table 2. Systematic and random uncertainties (in %) for CFC-11, CFC-12, and HCFC-22 at St  
 585 Denis. Sb represents the relative uncertainties of the non-retrieved parameters. For temperature,  
 586 the systematic/random Sb matrix was created by the mean/standard deviation of the differences  
 587 between NCEP and the balloon observations. For the spectroscopic parameters, 0.07, 0.01 and  
 588 0.05 are the relative uncertainties of CFC-11, CFC-12 and HCFC-22, which are according to the  
 589 PLL database, respectively. When a relative uncertainty is smaller than 0.01%, it is considered  
 590 negligible and represented as “-”. For zshift, the same uncertainty is adopted for the systematic  
 591 and random error; zshifts included in the retrieval parameters for CFC-11, but not for CFC-12 and  
 592 HCFC-22. For SZA, the systematic uncertainty is 0.001 and the random uncertainty is 0.002 (in the  
 593 bracket).

Error	Sb	CFC-11		CFC-12		HCFC-22	
		Systematic	Random	Systematic	Random	Systematic	Random
Smoothing		0.18	1.03	0.18	0.79	0.47	0.51
Measurement		-	0.80	-	0.21	-	4.05
Retrieval parameters		-	-	-	-	-	0.51
Interfering species		0.12	0.84	-	0.10	0.02	0.13
Temperature		1.35	0.30	0.87	0.15	1.00	0.18
SZA	0.001(0.002)	0.21	0.42	0.36	0.72	0.63	1.26
Line intensity	0.07/0.01/0.05	6.59	-	1.00	-	3.96	-
T-dependence of line width	0.07/0.01/0.05	-	-	-	-	0.15	-
Air-broadening of line width	0.07/0.01/0.05	0.45	-	0.07	-	1.53	-
szshift	0.01	-	-	0.12	0.09	0.10	0.12
Total		7.0	2.0	1.8	1.1	4.4	4.5

594



595 Table 3. Same as Table 2, but for Ma řlo.

Error	Sb	CFC-11		CFC-12		HCFC-22	
		Systematic	Random	Systematic	Random	Systematic	Random
Smoothing		0.23	0.90	0.02	0.67	0.40	0.58
Measurement		-	0.70	-	0.20	-	3.36
Retrieval parameters		-	0.26	-	-	-	0.12
Interfering species		0.08	0.20	-	-	0.03	0.03
Temperature		1.82	0.76	1.02	0.19	0.84	0.16
SZA	0.001(0.002)	0.25	0.51	0.22	0.44	0.22	0.44
Line intensity	0.07/0.01/0.05	6.22	-	0.98	-	4.22	-
T-dependence of line width	0.07/0.01/0.05	0.05	-	-	-	0.05	-
Air-broadening of line width	0.07/0.01/0.05	1.54	-	0.05	-	1.48	-
zshift	0.01	-	-	0.12	0.07	0.10	0.05
Total		6.7	1.6	1.8	1.1	4.5	3.6

596



597 Table 4. The annual percent changes (in %/year, relative to the mean of data used in the trend  
 598 analysis) and uncertainties of FTIR total columns of CFC-11, CFC-12 and HCF-22 at St Denis  
 599 (2004-2011) and of the combined partial columns (2.155-100 km) at St Denis along with the total  
 600 columns at Ma ĩlo (2004-2016 for CFC-11 and HCFC-12, and 2009-2016 for CFC-12). The trends  
 601 from in-situ and flask measurements at SMO are also given for the same time periods. The trends  
 602 observed by MIPAS (2004-2010), ACE-FTS (2004-2010) and by the ground-based FTIR at  
 603 Jungfraujoch (2004-2010) % in %/yr relative to the 2007 annual mean are taken from Carpenter et  
 604 al. (2014) as described in the text.

Dataset	Time range	CFC-11	CFC-12	HCFC-22	Reference
St Denis (TC)	2004-2011	-0.69±0.15	-0.26±0.10	3.14±0.43	
SMO	2004-2011	-0.89±0.01	-0.37±0.08	4.04±0.06	
MIPAS	2004-2010	-1.03±0.09	-0.51±0.09	-	Kellmann et al., 2012
ACE-FTS	2004-2010	-0.9±0.1	-0.4±0.1	3.7±0.1	Brown et al., 2011
Jungfraujoch	2004-2010	-0.99±0.10	-0.38±0.07	3.52±0.08	Zander et al., 2008
St Denis(PC) + Ma ĩlo(TC)	2004-2016	-0.86±0.12	-	2.75±0.12	
	2009-2016	-	-0.76±0.05	-	
SMO	2004-2016	-0.87±0.04	-	3.46±0.05	
	2009-2016	-	-0.60±0.02	-	

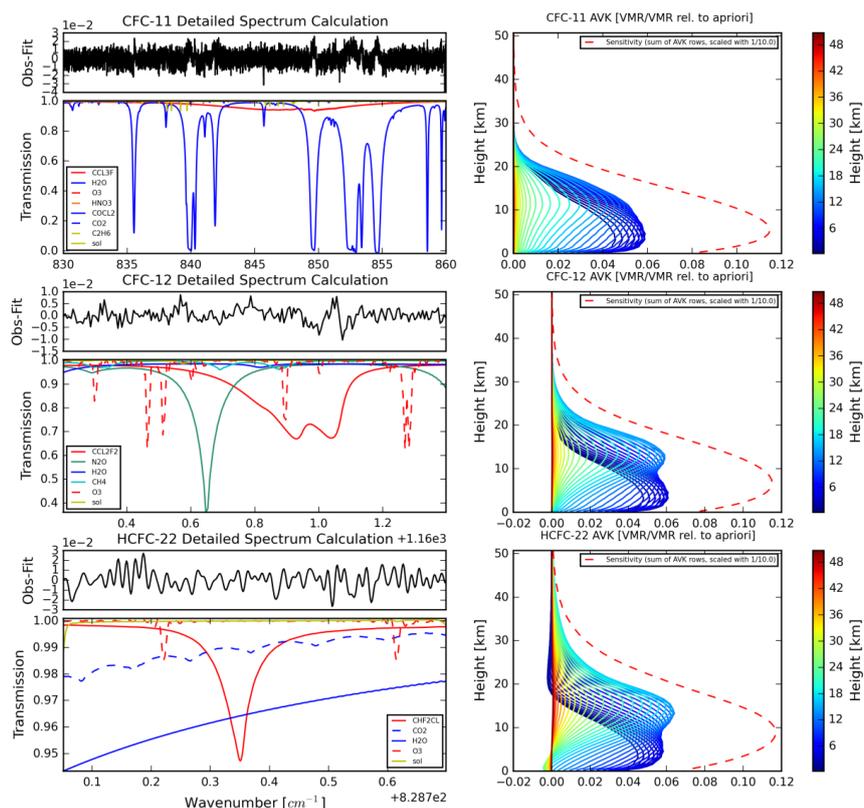
605



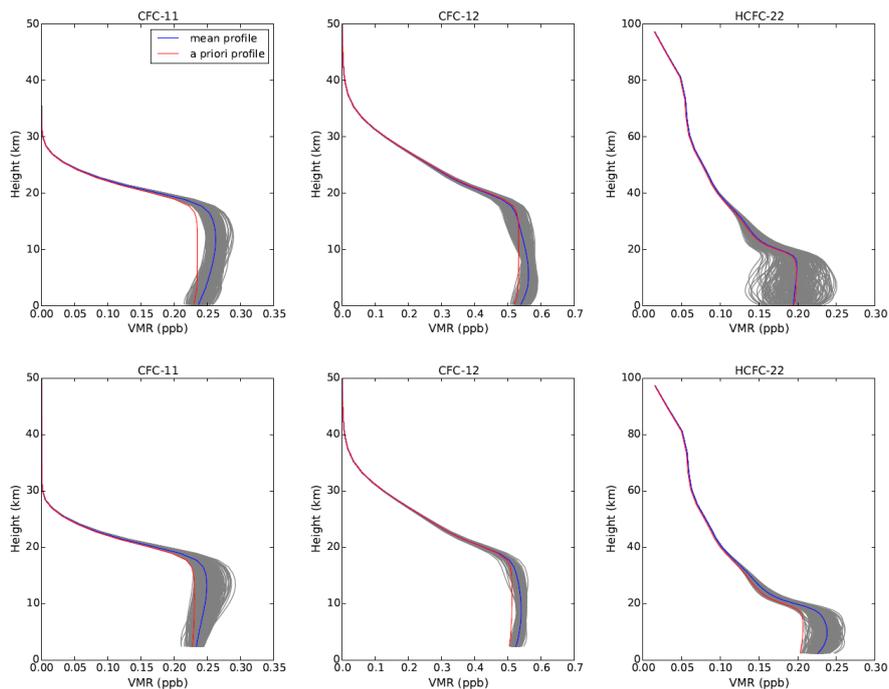
606 Table 5. The number of collocated MIPAS-FTIR pairs, bias and standard deviation (std) of the  
 607 relative differences  $((\text{MIPAS-FTIR})/\text{FTIR} \times 100\%)$  between the partial columns (6-30 km) of  
 608 MIPAS (both the raw and smoothed data) and FTIR, together with the mean random and  
 609 systematic uncertainties of the FTIR partial columns and the retrieval error of the MIPAS data  
 610 (in %).

	CFC-11	CFC-12	HCFC-22
Number of data pairs	60	86	42
Bias_raw	-4.4	-3.3	0.2
Std_raw	4.7	4.5	6.3
Bias_smoothed	-4.3	-2.9	-0.7
Std_smoothed	4.4	4.6	6.0
FTIR PC random error	4.2	3.5	7.8
FTIR PC systematic error	10.5	2.6	7.1
MIPAS PC retrieval error	4.1	4.3	5.0

611

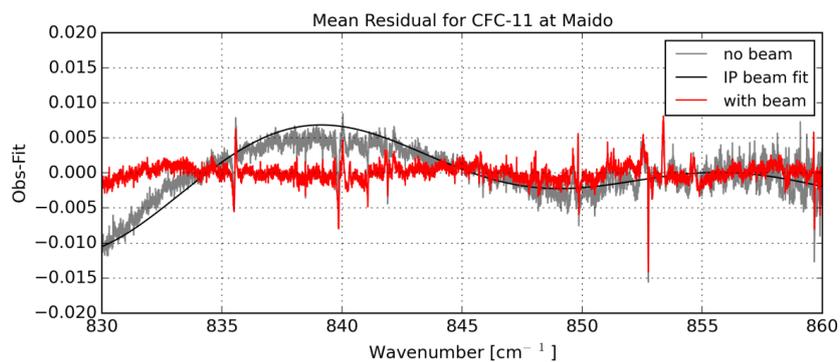


612  
 613 Figure 1. The typical spectrum and averaging kernels of CFC-11 (upper), CFC-12 (middle) and  
 614 HCFC-22 (bottom) at St Denis. The left panels show the transmission and residual (observed –  
 615 calculated spectrum) for the three retrieval microwindows, along with the absorption contribution  
 616 from each specie. Right panels show the averaging kernels for the target gases (no sensitivity  
 617 above 50 km). The solid lines represent the sensitivities at specific altitudes. The red dashed line is  
 618 the sum of the row of averaging kernel scaled with 0.1, indicating the vertical sensitivity.

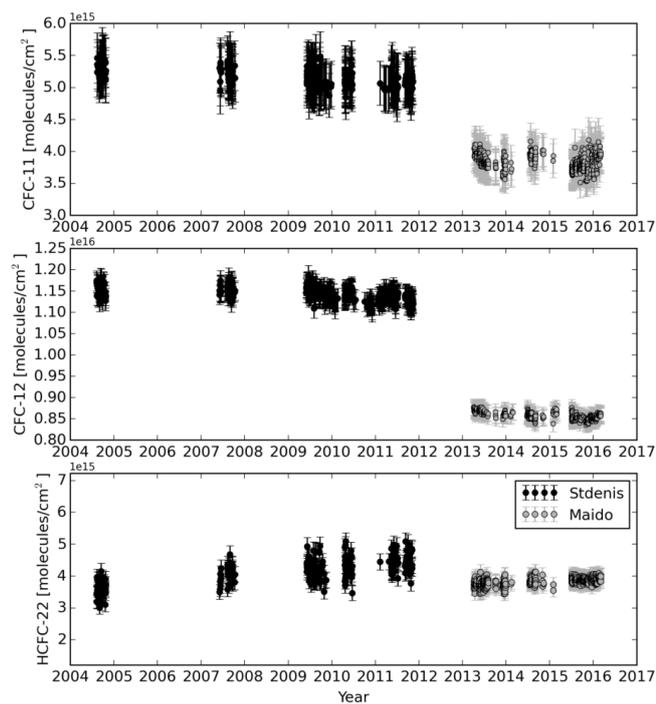


619

620 Figure 2. The a priori (red line), retrieved profiles (grey lines) and the mean retrieved profile (blue  
621 line) of CFC-11, CFC-12 and HCFC-22 at St Denis (upper panels) and Ma'ilo (bottom panels).



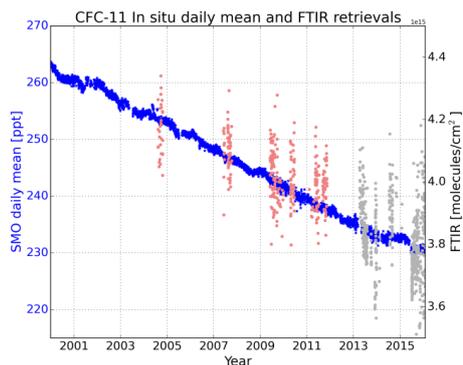
622  
623 Figure 3. The mean residual transmittance (observed – calculated) of the CFC-11 retrievals with  
624 and without beam parameters at Mađo. The IP beam fit line is used as a priori IP beam  
625 parameters.



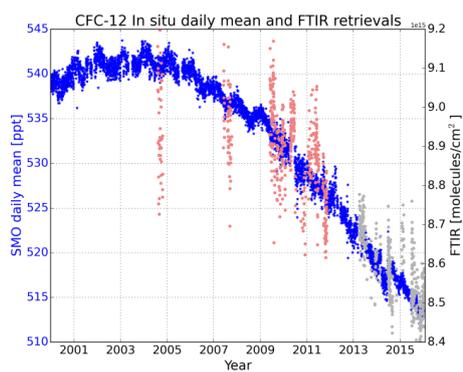
626  
627 Figure 4. The time series of the total columns and total uncertainties of CFC-11, CFC-12 and  
628 HCFC-22 at St Denis (black) and Maïdo (grey). The error bar contains both systematic and  
629 random uncertainties from SFIT4 retrieval ( $\sqrt{\varepsilon_s^2 + \varepsilon_r^2}$ ).



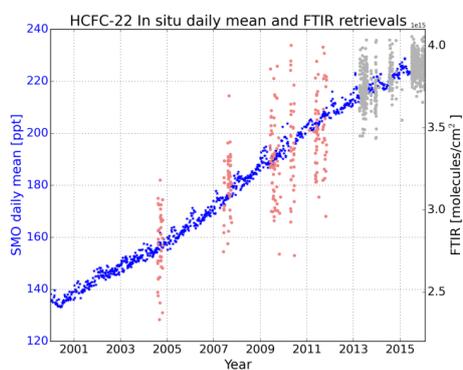
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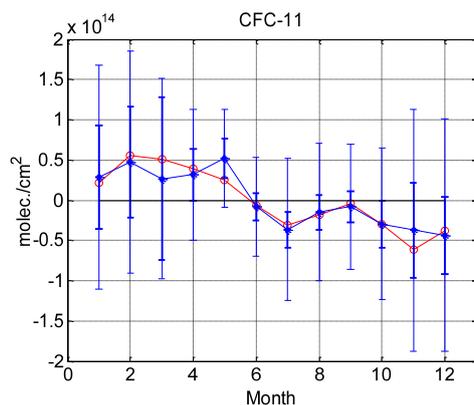
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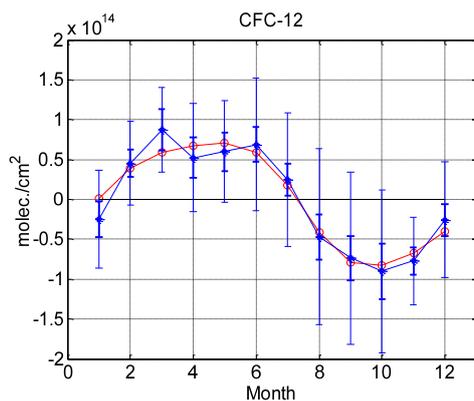
633 Figure 5. In-situ daily mean (CFC-11 and CFC-12) and flask pair measurements (HCFC-22) at  
634 SMO site (blue) and individual FTIR partial column at St Denis (2.155-100 km; light coral) and  
635 total column at Ma ño (2.155-100 km; grey). Upper: CFC-11; middle: CFC-12; bottom: HCFC-22.



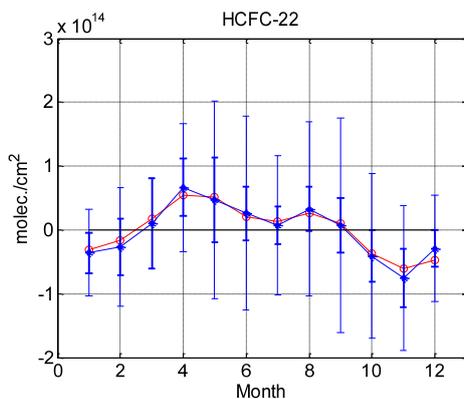
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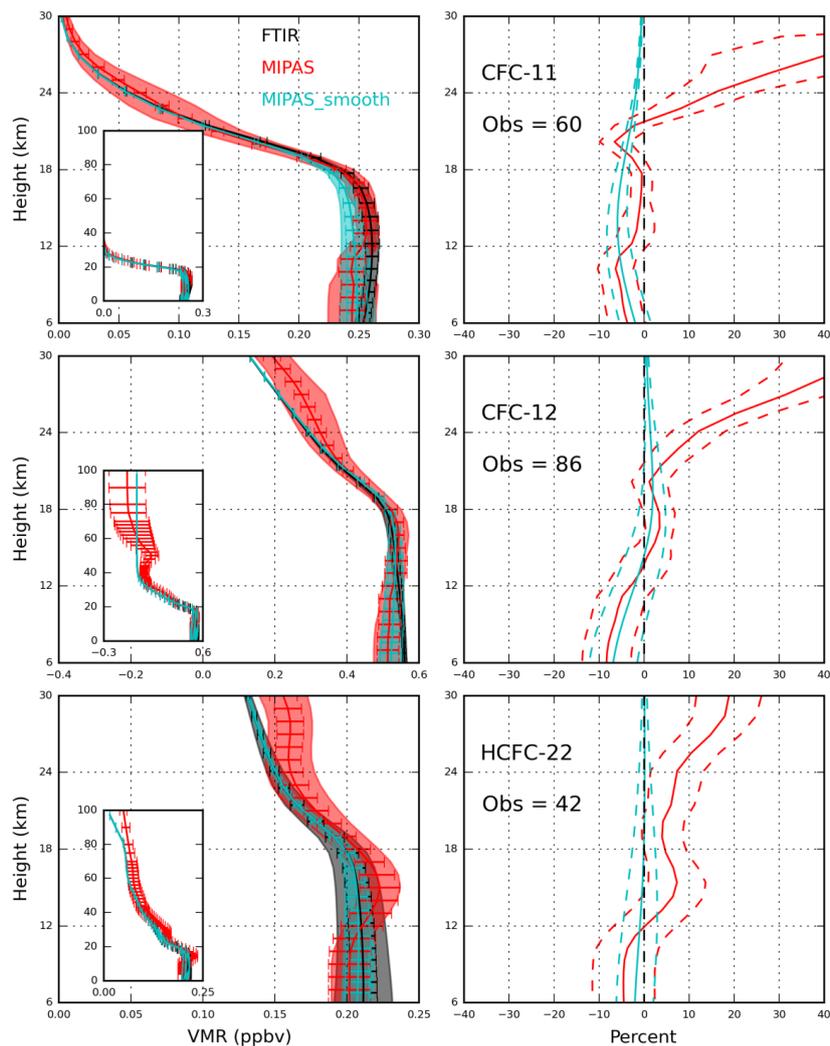
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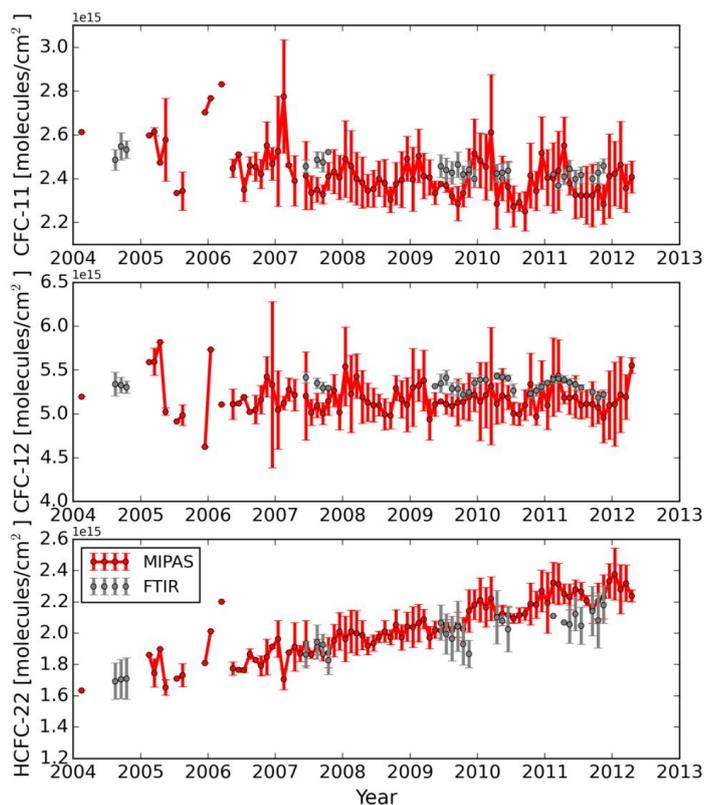
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639 Figure 6: Seasonal cycles of CFC-11, CFC-12 and HCFC-22. The modelled seasonal cycle  
640 obtained by Eq. 10 for the 2004-2016 St Denis -Ma ão time-series is shown in red. In blue, the  
641 mean of FTIR measurements for each month during the 2004-2016 period, after subtraction of the  
642 trend, is shown, together with the standard deviation  $\sigma$  of the mean (thin error bars). The  $2\sigma$  error  
643 on the mean ( $2\sigma / \sqrt{n}$ ;  $n$  being the number of measurements for each month) is also shown with  
644 thick blue lines.



645  
 646 Figure 7. Left panel, for each target species (from top to bottom: CFC-11, CFC-12 and HCFC-22):  
 647 averaged target species mixing ratio profile, random uncertainty (error bar) and the standard  
 648 deviation of all the co-existing data pairs (shade area) for FTIR (in black) and for MIPAS (in red :  
 649 raw data; in sky blue: after smoothing with the corresponding FTIR averaging kernel). The  
 650 profiles (from 0 to 100 km) are also manifested in the left panels. Right panel, for each target  
 651 species, averaged relative difference between MIPAS and FTIR  $((\text{MIPAS} - \text{FTIR})/\text{FTIR} \times 100\%)$   
 652 (solid lines), along with the standard deviation (dash lines).



653  
654 Figure 8. The time series of the monthly means of partial columns (6 to 30 km) of CFC-11,  
655 CFC-12 and HCFC-22 from St Denis FTIR measurements (grey) and raw MIPAS data (red). Error  
656 bars represent the standard deviations of the monthly means.