Atmospheric Measurement Techniques



A review of sources of systematic errors and uncertainties in

² observations and simulations at 183GHz

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27 Abstract

Several recent studies have observed biases between measurements in the 183.31GHz water vapour 28 line by space-borne sounders and calculations using radiative transfer models with inputs from either 29 radiosondes (RAOBS) or short range forecasts by Numerical Weather Prediction (NWP) models. 30 This paper discusses all the relevant categories of observation-based or model-based data, 31 quantification of their uncertainties, and separation of biases that could be common to all causes from 32 33 those attributable to a particular cause. Reference observations from radiosondes, Global Navigation Satellite Systems (GNSS), Differential Absorption Lidar (DIAL) and Raman lidar are thus 34 overviewed. Biases arising from their procedures of calibration, NWP models and data assimilation, 35 36 instrument biases and radiative transfer models (both the models themselves and the underlying 37 spectroscopy) are presented and discussed. Although no single process in the comparisons seems 38 capable of explaining the observed structure of bias, recommendations are made in order to better understand the causes. 39

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41 **1. Introduction**

Recent cross-comparisons between the existing satellite microwave sounders of the tropospheric 42 43 humidity in the 183.31GHz line, SAPHIR (Sondeur Atmospheric du Profil d'Humidité Intertropical 44 par Radiométrie, on Megha-Tropiques), ATMS (Advanced Technology Microwave Sounder, on Suomi-NPP), SSMI/S (Special Sensor Microwave Imager/Sounder, on DMSP-F17 and F18) and 45 MHS (Microwave Humidity Sounder on MetOp-A and B and NOAA-18 and 19), show very good 46 agreement between them, in a 0.3-0.7K range of mean difference, well within the radiometric noises 47 of the instruments [Wilheit et al., 2013; Moradi et al, 2015]. However, when the measurements are 48 compared to radiative transfer model (RTM) calculations using profiles of temperature and humidity 49 either from radiosondes (RAOBS) or from Numerical Weather Prediction (NWP) models, a channel-50 dependent bias, which increases from the center to the wings of the line, is observed. It was only with 51 the arrival of ATMS and SAPHIR, both launched in October 2011, that the spectral shape of the bias 52 became clear [Clain et al., 2015; Moradi et al., 2015]. Indeed these two instruments sample the 53 183.31GHz line 5 and 6 times respectively between the line center (providing humidity information 54 for the upper troposphere, above 300hPa) and line wings (for a deeper sounding of the atmosphere) 55 compared to only 3 times for SSMI/S, MHS and AMSU-B (Advanced Microwave Sounding Unit-B 56 57 that precedeed MHS onboard NOAA 15, 16 and 17). The observed minus calculated brightness temperatures (BTs) are shown in Figure 1, for SAPHIR, ATMS, MHS, SSMI/S against RAOBS or 58





NWP systems (Météo-France and European Centre for Medium-range Weather Forecasts) as input 59 for the RTTOV.v11 RTM (Radiative Transfer for the Television InfraRed Observation Satellite 60 (TIROS) Operational Vertical Sounder, [Eyre, 1991; Saunders et al., 1999; Matricardi et al., 2004; 61 Saunders et al., 2013]). It shows a consistent spectrally-dependent bias which is worse further from 62 the line center. Figure 2 presents two maps of the differences between the observed ATMS BTs and 63 64 the calculated BTs from the all-sky first-guess fields of the ECMWF NWP system for channels 22 (183±1GHz) and 18 (183±7GHz) for a 2-month period. While great differences in areas of snow of 65 66 persistent thick ice clouds are produced by the RTM's omission of their effect in the simulation of such scenes, there is a clear background difference, between 60°N and 60°S, that increases towards 67 the line wings and that cannot be related to a particular region or atmospheric state. 68

The importance of knowledge of absolute errors for climate applications (e.g. re-analysis) is clear 69 70 because these need to detect small changes in the mean state. For real-time NWP, knowing absolute values is less important because observations are bias-corrected, with state-of-the-art NWP systems 71 72 updating the corrections frequently. Bias corrected observations at 183.31GHz are used widely in NWP models and have important impacts [Geer et al., 2014]. However, the absolute errors in the 73 74 short-range forecast are also not well known. If sources of systematic errors can be reduced, the need for bias correction would also be reduced, and perhaps one day could be eliminated completely, 75 allowing 183.31GHz observations to become a reference for humidity. This would enable us both to 76 77 understand better biases in other humidity observations and to improve separation of biases arising 78 from the model (such as cloud cover) from others (such as humidity fields). Bias correction has proven a successful and pragmatic strategy for handling unknown biases, but it does mean we can 79 only correct and characterize the random component of error in model short-range forecasts. 80

The attribution of the biases requires discussion of all the relevant observational data and model results, quantification of their uncertainties, and separation of biases that could be common to all approaches from those attributable to some particular methodology. Thus we need reference observations from RAOBS, Global Navigation Satellite Systems (GNSS), Differential Absorption Lidar (DIAL) and Raman lidar. We can then attribute biases arising from their calibration procedures, NWP models and data assimilation, instrument biases and RTMs (both the models themselves and the underlying spectroscopy).

If we can successfully attribute the biases then it becomes possible to work on a solution. To illustrate this consider the case of a new observation type for humidity. If we believe that there is a calibration error and we can characterize this error accurately, it may become clear to instrument engineers why the bias exists. If the bias is the sum of ones from a number of sources, this process





92 can't happen. This is also the motivation behind projects such as GAIA-CLIM (Gap Analysis for
93 Integrated Atmospheric ECV - Essential Climate Variable - Climate Monitoring, http://www.gaia-

94 clim.eu/) and FIDUCEO (Fidelity and Uncertainty in Climate Data record from Earth Observations,

95 http://www.fiduceo.eu/), which are also attempting to provide a more accurate assessment of 96 systematic biases in observations.

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2. State-of-the-art in observation and modelling

99 2.1 Reference water vapor measurements

Radiosondes: Uncertainties in relative humidity (RH) measurements by in-situ probes arise mainly 100 101 from pre-launch calibration procedures, calibration corrections, time-lags and (for some probes) solar 102 radiation heating of the sensor. The GCOS Reference Upper-Air Network (GRUAN, http://www.gruan.org) aims to establish a network of temperature and humidity measurements with 103 traceability to SI standards, and best possible characterization of uncertainties [Dirksen et al., 2014; 104 105 Bodeker et al., 2015]. For the Vaisala RS92-SGP (hereafter RS92), one of the most common RAOBS probes used operationally and during field campaigns, the GRUAN-characterized uncertainty in RH 106 107 measurements after correcting known biases is overall below 6%RH, and the only uncorrected bias is a dry bias of $\sim 5\%$ RH at night for temperatures colder than -40°C [Dirksen et al., 2014]. The 108 109 reliability of the RS92 product in the troposphere has been verified by comparisons with frostpoint 110 hygrometer (FPH) measurements, which are highly accurate balloon-borne humidity measurements, in both tropical and extra-tropical regions. The GRUAN RS92 profile for humidity does not vary 111 112 greatly from Vaisala's default processing below the upper troposphere (pressures below 200hPa, [Yu 113 et al., 2015]). Further, the most recent intercomparison campaign [Nash et al., 2011] held in Yangjiang (a tropical site in South China) showed good agreement between RS92 and most other 114 115 operational RAOBS up to the mid-to upper troposphere (pressure above 500hPa), and also with FPH 116 measurements [Vömel et al., 2007]. In the lower to mid-troposphere there is therefore robust agreement and evidence that RAOBS biases could at most be a few percent, with somewhat broader 117 random sampling uncertainties on individual profiles. Also, a given fractional change in humidity 118 causes a considerably small fractional change of BTs. Therefore RAOBS errors could only explain 119 120 substantial biases near the line center, where the opacity is highest and thus sensitive to water vapor 121 of the upper part of the troposphere, whereas the 183.31GHz biases are the largest towards the wings.





123 GNSS receivers: GNSS estimates of the atmospheric precipitable water (PW) rely on the measurement of the zenith total delay, the sum of terms due to water vapor (the "wet" delay) and to 124 the dry gases of the troposphere (the "dry" delay). The GNSS-estimated PW has an uncertainty of 125 126 $\sim 2\%$ in the mean (<1.0mm) [Ning et al., 2015]. However, a recent analysis of the upper-air sounding network deployed during the CINDY/DYNAMO/AMIE field campaign has revealed an unclear, and 127 128 statistically significant, dry bias in the GNSS values (~2.0mm in PW) in moist conditions for some sites of the campaign [Ciesielski et al., 2014]. Understanding of this dry bias might contribute to the 129 130 current discussion on the 183GHz bias.

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Lidar systems: Two types of lidars can be used to measure water vapor profiles [Behrendt et al., 132 2007; Bhawar et al., 2011]. Differential absorption lidars (DIAL) measure the water vapor number 133 134 density with two backscatter signals (at adjacent near-IR wavelengths with high -online- and low -135 offline- absorption) yielding a self-calibrating system. It only relies on the difference of the water vapor absorption cross-sections σ_{on} and σ_{off} at these two wavelengths (thus eliminating any bias 136 137 common to both cross-sections). Raman lidars are based on the inelastic scattering by water vapor 138 molecules, and require one instrument-dependent calibration factor for all heights to obtain the 139 mixing ratio. All lidar systems provide data in the cloud-free atmosphere, or until the laser beam reaches an optically thick cloud. Performance simulations as well as intercomparisons have 140 141 confirmed accuracies < 5% in the troposphere for both systems [Wulfmeyer et al., 2015]. 142 Comparisons between water vapor profiles measured by lidar and by radiosondes further support the 143 accuracy estimates for the global radiosonde data in the troposphere, quoted above.

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145 **2.2 Radiative transfer modeling and spectroscopy**

Behavior of radiative transfer models: Many cross-comparisons of microwave RTMs have been 146 performed over the years. For instance, Garand et al. (2001) compared different reference and fast 147 models (including RTTOV.v5 and v6 [Saunders et al., 1999]) from the operational NWP community 148 for the 183±1GHz channel. At the time, they found an overall agreement better than 0.5K in BT, 149 which is consistent with the results of Melsheimer et al. (2005), a study dedicated to line-by-line 150 RTMs, with differences within roughly 0.5, 1.5, and 2.5K, respectively for $183\pm1, \pm3$, and ±7 GHz 151 152 channels (channels 3 to 5 of MHS). The differences were mainly attributed to the differences in 153 spectroscopy (line parameters and continua) and not the RTMs themselves. This was confirmed by 154 Buehler et al. (2006), who showed that, on average over the globe, the difference between two





155 independent models, RTTOV.v7 and the line-by-line model ARTS (Atmospheric Radiative Transfer Simulator [Buehler et al., 2005a; Eriksson et al., 2011]) is between 0.01K for the 183±1GHz channel 156 and 0.2K for the 183±7GHz channel, except for a few extreme situations. More recently, Chen et al. 157 158 (2010) have shown that, on a sample of diverse atmospheric profiles, fast and line-by-line RTMs agree within 0.1K for MHS channels, and that this result was highly stable under humid conditions. 159 160 In their evaluation of SAPHIR using RAOBS, Clain et al. (2015) have shown that the three RTMs RTTOV.v10, ARTS and MonoRTM (Monochromatic Radiative Transfer Model [Clough et al., 2005; 161 162 Payne et al., 2011]) provide fairly consistent BTs on a common set of tropical profiles, the differences being in the range -1.50K / 0.78K, with the largest differences observed for the central 163 channel (183.31±0.2GHz). 164 Patterns in the IR are similar in terms of dependence on distance from the line center. Indeed, 165

- 166 sensitivity studies on a sample of channels in the 6.3µm water vapor band of the IASI (Infrared 167 Atmospheric Sounding Interferometer) instrument carried out using RTTOV.v11 showed that the difference between calculated and observed BTs increases as the peak altitude of the weighting 168 functions shifts downwards. The same qualitative behavior is observed irrespective of the 169 170 atmospheric state used in the simulations (i.e. either RAOBS or NWP data). It was also found that, on a purely empirical basis, the pattern of increasing bias can be removed by applying adjustments to 171 the humidity fields input to the simulations (3 to 10% increase below 500hPa) and/or to the strength 172 of the continuum absorption (30% increase in foreign continuum plus a 20% increase of self-173 174 continuum). This does not necessarily mean that the same mechanisms must be responsible for the 175 biases observed in the MW radiances, although for the three RTMs mentioned here (RTTOV, ARTS 176 and MonoRTM), the absorption due to the water vapor continuum is parameterized using the same 177 semi-empirical model MT CKD (MlawerTobin CloughKneisysDavies, [Mlawer et al., 2012]).
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179 Spectroscopy status: One of the considerations for the accuracy of the line-by-line RT models is the 180 spectroscopic input for the modeling of molecular absorption. The main contributions to the 181 molecular absorption in the MW region of the spectrum are from H₂O, O₂ and N₂, with some minor contributions from O_3 and N_2O . For instance John and Buehler (2004) found that absorption by O_3 182 183 corresponds to a decrease of up to 0.5K in the 183±1GHz channel of MHS using ARTS, reaching 184 even smaller values for the other channels. Additional sensitivity tests performed by Clain et al. 185 (2015) using MonoRTM showed that the impact is also marginal in the 183±11GHz channel of SAPHIR (0.05K on average). The details of how the molecular absorption is modeled may vary 186





between different models, but the overall absorption is most commonly calculated for both thecontribution near the line centers and the smoothly-varying continuum.

Line parameters (line position and strength, the air-broadened half-width, the self-broadened half-189 190 width, the temperature exponent of the width and the pressure shift) may be obtained from laboratory experiments or from theoretical calculations and are collected in databases such as the widely used 191 192 High Resolution Transmission compilation (HITRAN, [Rothman et al., 2013]). The current edition of HITRAN provides the parameters for Voigt profiles, but future editions of the database will allow for 193 194 inclusion of additional parameters for more sophisticated line shape models [Tennyson et al., 2014]. Sensitivity tests on Voigt parameters, performed using MonoRTM, have shown that illustrative 195 uncertainties on the foreign $(\pm 3\%)$ and self-broadened $(\pm 15\%)$ half widths, on the temperature 196 197 exponent (maximum of 15%) and the pressure shift (maximum of 20%) are certainly too small to explain the observed bias (Payne et al., 2008) and the spectroscopic community (lab and modellers) 198 believes that confidence limits on these parameters are lower than those assumed above. The 199 200 uncertainty of the dry air absorption including dry continuum and resonance absorption by O_2 , O_3 , 201 N₂O, NO, CO and other minor atmospheric constituents, as well as uncertainty related to wings of 202 neighboring water lines is not thought to be large enough to account for the observed model-203 measurement bias.

The physical origin and properties of the water vapor continuum have been debated and probed with 204 205 measurements for decades. In the current version of the MT CKD continuum model water vapor 206 contributions are modeled as monomer absorption and the spectral variation of the continuum is assumed to be extremely smooth (sampled every 300GHz in MT CKD). Figure 3 summarizes 207 comparisons of the continuum coefficients obtained from known laboratory and field (~1km path 208 209 along the surface, insensitive to vertical distribution of absorbers) measurements (shown by symbols) 210 against the continuum parameters (dotted and dashed lines) that would provide agreement with 211 ground-to-sky radiometric data in MT CKD. It shows discrepancies (offsets in a log scale) that are 212 not yet understood. There is thus an inconsistency between two large sets of experimental data, 213 namely laboratory together with surface path measurements and radiometric measurements within currently accepted modeling of spectroscopy including both modeling of the continuum and 214 modeling of the atmosphere. This is confirmed by Payne et al. (2011) who concluded that for 215 216 atmospheric path lengths (looking up) the combination of MPM (Millimeter-wave Propagation 217 Model, [Liebe 1989; Rosenkranz, 1998]) foreign and self-continuum (solid lines in Figure 3) is inconsistent with the radiometric measurements at high column water vapor amounts. 218





219 Finally recent laboratory studies have resulted in unambiguous detection of H₂O dimer absorption in the millimeter-wave range [Tretyakov et al., 2013; Serov et al., 2014] and to the development of a 220 221 model to describe it [Odintsova et al., 2014]. This absorption shows spectral variation on scales that 222 are not accounted for in the current version of MT CKD or in Liebe-based models. Odintsova et al. (2014) indicate that the inclusion of dimer absorption can result in small-scale spectral (~1GHz) 223 224 variation of 0.5 to 1K in up-looking (ground-based) spectra. The impact of accounting for dimer absorption on RT modeling for the 183.31GHz satellite radiometer channels has yet to be 225 226 determined.

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228 2.3 Water vapor analysis

229 In NWP models, the main observation types influencing humidity analyses are in-situ data like RAOBS and remote sensing observations from IR and MW sensors [Andersson et al., 2005]. Cai and 230 231 Kalnay (2005) have illustrated how a balance can arise between a forecast model (subject to a 232 inherent bias with respect to the truth) and observations (with different bias characteristics). 233 Variational Bias Correction (VarBC) techniques have been developed to adaptively estimate, as part 234 of the constrained optimization that also provides the best estimate of the latest atmospheric state, a bias correction for each of the various assimilated observations [Dee, 2005; Auligné et al., 2007]. In 235 236 order to anchor the bias corrections and the final humidity analysis, RAOBS are not bias corrected with VarBC (although corrections are applied to standardize them to night-time RS92 observations: 237 238 Agusti-Panareda et al. 2009). Hence it is possible that humidity analyses share similar biases to radiosondes, which might explain some of the consistency between the channel-dependent bias found 239 240 with both in-situ measurements and NWP simulations. More investigations would be needed to test 241 the impact of this anchoring, although it is expected from the discussion in Section 2.1 that the effect 242 would be more pronounced in the upper-tropospheric channels (close to the line center) than in the 243 lower tropospheric channels (in the line wings).

An issue that affects most comparisons between 183.31GHz observations and a reference is cloud detection. Indeed clouds and precipitation are not usually included in MW radiative transfer simulations, while they naturally reduce 183.31GHz BTs, particularly in the lower-peaking channels (see for instance the negative bias in the Inter-Tropical Convergence Zone in Figure 2), either by scattering or by absorption, which shifts upwards the altitude of the weighting function. There are major sources of uncertainty in modeling the effects of cloud and precipitation around the 183.31GHz line, coming from the difficulty in specifying shape, density, and particle size





distribution of solid precipitating particles [Burns et al., 1997; Doherty et al., 2007; Geer and Baordo,

252 2014]. However, if the model cloud and precipitation fields were unbiased compared to reality, a

negative first guess departure in a 183.31GHz channel would indicate that the observation selection

is too cloudy (i.e. cold). This would be consistent with a cloud detection that is missing some cloud-

255 affected scenes.

256 Different cloud screening methods are applied depending on the sensor and the channels available, leading to different accuracies of the cloud screening. For instance, ATMS offers a large suite of 257 channels that can complement each other to separate clear and cloudy scenes [Bormann et al., 2013], 258 259 while SAPHIR provides only channels in the 183.31GHz line, giving fewer possibilities for cloud detection [Chambon et al., 2015]. Several filtering techniques using only 183.31GHz channels exist 260 261 and rely on the strong reduction in BT induced by scattering and/or absorption, which can reach up to 20K in regions of intense convection [Greenwald and Christopher, 2002; Hong et al., 2005; Buehler 262 263 et al., 2007]. However, it is difficult to screen all clouds, so residual biases may be present. The allsky first guess departures from assimilation at ECMWF [e.g. Geer et al., 2014] are the only routinely 264 computed differences in the 183.31GHz line which attempt to take into account the effects of cloud 265 and precipitation in the radiative transfer modeling. Compared to those differences computed using 266 clear-sky radiative transfer and cloud-screening (which are shown in Figure 1 marked "ECMWF"), 267 the all-sky ATMS biases are smaller by 0.4K in the ± 7 GHz channel, suggesting some but not nearly 268 all of the bias in the lower peaking channels can be explained by residual cloud effects. 269

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271 2.4 Space-borne microwave radiometers

272 Intercomparison of instruments: Ways to compare measurements by different satellite instruments 273 include the simultaneous-nadir-overpasses technique (limiting the comparisons to the highest 274 latitudes, [John et al., 2012]) and the use of "natural targets" that have very little variability or the averaging over a lot of scenes [John et al., 2013a]. Another technique is the double difference 275 technique that uses NWP fields input to a RTM as a transfer function between two radiometers that 276 277 have few common overpasses. Using this technique, recent comparisons of the 183.31GHz channel calibrations have been performed by the Global Precipitation Mission (GPM) intercalibration 278 279 working group, or (XCAL team). Mean differences between the GMI (GPM Microwave Imager) BTs 280 and the four operational MHS sensors onboard MetOp-A, MetOp-B, NOAA-18, and NOAA-19 as 281 well as the Suomi NPP ATMS and Megha-Tropiques SAPHIR instruments are given in Table 2.





Calibration issues: Since GPM is in a precessing orbit with an inclination of 65°, it frequently 283 crosses the orbits of the other sounders providing near coincident observations several times each day 284 in the 65°N/65°S belt. Post launch, a series of GPM calibration maneuvers was performed, and the 285 resulting data were used to develop corrections for magnetic-induced biases, cross-track biases and 286 updates to the pre-launch spillover corrections, as well as to verify the channel polarizations. The 287 resulting GMI calibration is based on the data from these calibration maneuvers and does not depend 288 289 on RTM. The GMI calibration is also completely independent of the calibration of the MHS, ATMS, and SAPHIR instruments, thus providing a useful measure of the absolute calibration accuracy of the 290 291 183.31GHz channels for these sensors. The differences in Table 2 show consistent results, with 292 values within 1K for all channels and all sensors. Errors in the calibration of the SSMI/S 183.31GHz channels are substantially larger due to substantial biases caused by reflector emission and solar 293 294 intrusion issues [Berg and Sapiano, 2013].

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Scan asymmetry: Cross-track scanners sometimes show asymmetries across the swath, indicating 296 297 issues related to the imperfectly known antenna pattern, which often evolve with the age of the 298 instruments. Far from nadir, channels sounding deep into the troposphere, located on the wings of the 299 183.31GHz line, might be more affected by antenna issues than higher peaking channels, located in the center of the line, due to radiation measured by the side-lobes. Such asymmetries were found for 300 AMSU-B and MHS [Buehler et al., 2005b; John et al., 2013b] but so far the monitoring of SAPHIR 301 302 has not shown any scan-asymmetry. For ATMS, comparisons of Temperature Data Records (TDR, calibrated antenna temperatures) and Sensor Data Records (SDR, BT after further applying beam 303 304 efficiency and scan position dependent bias corrections) for the 183.31GHz channels show the same behavior meaning that the TDR to SRD conversion is not responsible for the bias, although it seems 305 306 to introduce some dependence on the viewing angle that warrants further investigation. For other sensors, in particular AMSU-B and MHS this has so far not been thoroughly investigated, partly due 307 308 to a lack of publicly available pre-launch measurements of antenna pattern.

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310 **3. Recommendations**

311 <u>On the reference measurements:</u> Research on better characterization of uncertainties of the Vaisala 312 RS92 measurements, and on offline corrections of biases, indicate that the remaining RAOBS biases 313 could only explain discrepancies in the center of the 183.31GHz line, not in its wings. Nevertheless 314 cross-comparisons of water vapor measured by lidars and RAOBS and simulated by NWP models





are strongly encouraged in order to confirm the spectral pattern of the 183.31GHz bias. Finally the
 next WMO upper-air intercomparison campaign is strongly encouraged to include the new Vaisala

RS41 probes, which are likely to be even more accurate than the RS92.

On the radiative transfer and the spectroscopy: For the purposes of atmospheric remote sensing, 318 319 consistency with in-situ atmospheric radiometric measurements is key. Currently, the cause of the apparent discrepancy between the laboratory measurements and the atmospheric results remains an 320 open question. Continuation and augmentation of laboratory measurements are strongly encouraged 321 to check the uncertainty levels for the main spectroscopic parameters, and to explore new line shape 322 323 parametrizations. Also, the use of ground based 183.31GHz instruments can help to better constrain the parametrizations, because the surface does not contaminate the measurements. In particular 324 measurements at high values of precipitable water (> 3cm) are required to constrain the self-325 326 broadened continuum. For instance, recent opacity measurements performed with the radio occultation active spectrometer ATOMMS (Active Temperature Ozone Moisture Microwave 327 328 Spectrometer, [Kursinski et al., 2012]) have shown two spectral discrepancies. The first discrepancy is a poor match between the Liebe MPM93 model and the measured line shape within 4 GHz of the 329 330 183.31GHz line center. In this interval, the HITRAN-based AM6.2 model of Scott Paine (Harvard-331 Smithonian Center for Astrophysics) matches the measurements very well, to 0.3%. However, there 332 is a significant spectral discrepancy in the wing of the line with respect to the AM6.2 modeled 333 opacity, which is apparently lower than the measured opacity. Viewed from space, this would 334 translate into a modeled BT that is higher than measured (the modeled radiation coming from deeper 335 in the atmosphere). Detailed understanding requires additional measurements and more quantitative 336 examination. In particular it is recommended that ATOMMS measurements be made from aircraft at 337 a range of pressures in order to determine the true line shape variation with pressure.

Better coordination between instrument and calibration experts and RT modelers would also help to ensure that RT simulations are consistent with the most recent spectroscopy measurements.

340 <u>On the water vapor analysis:</u> The two main approaches to handling clouds (avoiding observations 341 affected by cloud - the « clear-sky » approach; or using model cloud-fields and analysing all data -342 the « all-sky » approach) give similar but not identical biases. Accordingly, the method for handling 343 clouds can only explain part of the bias, but further investigation is needed to determine more 344 precisely the impact on the size and characteristics of the bias. Testing the impact of the RAOBS





constraint in the bias correction schemes of NWP systems with respect to the various observingsystems (MW, IR, GNSS) should be also performed.

347 On the space-borne sounders: The instrument spectral response functions (SRF) are too often averaged over the pass bands and assumed to be rectangular when unknown. Therefore the pre-348 349 launch recording and permanent availability of digital data and metadata of SRF and antenna patterns are strongly encouraged for future instruments. Moreover, in the future, it is considered likely that 350 Radio Frequency Interference (RFI) may become a threat at 183 GHz. It is therefore important both 351 to measure the bandpasses of the instrument accurately, and also to ensure there is no sensitivity to 352 353 bands outside the protected frequency. This point reinforces the need to have (an easy) access to the SRF of each instrument, as well as the conversion procedures between counts/radiances/BT. Finally, 354 biases arising from asymmetric effects for channels with widely spaced double-side bands, as 355 356 SAPHIR does, may also have an impact on the spectral characteristics of the bias. The size and 357 nature of this asymmetry needs to be accurately documented.

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562 Tables

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Table 1: List of satellites, sensors and their channels located in the f_0 = 183.31 GHz absorption line

Satellite	Instrument	Scanning geometry, maximum satellite zenith angle (θ_{zen} , in °), swath width (km) and number of pixels/samples per scan line	Channels around f ₀ = 183.31 GHz (GHz)	Bandwidth (MHz)	In-flight NeDT (K)
Megha- Tropiques	SAPHIR	cross-track $\theta_{zen} = \pm 50.7^{\circ}$ swath width = 1700 km 130 contiguous pixels	$f_0 \pm 0.2$	200	1.44
			$f_0 \pm 1.1$	350	1.05
			$f_0 \pm 2.8$	500	0.91
			$f_0\!\pm 4.2$	700	0.77
			$f_0\pm 6.8$	1200	0.63
			$f_0 \pm 11.0$	2000	0.54
Suomi-NPP	ATMS	cross-track $\theta_{zen} = \pm 64^{\circ}$ swath width = 2503 km 96 samples	$f_0 \pm 1.0$	500	0.9
			$f_0 \pm 1.1$	1000	0.8
			$f_0\!\pm 2.8$	1000	0.8
			$f_0 \pm 4.2$	2000	0.8
			$f_0\pm7.0$	2000	0.8
MetOp-A / B NOAA-18 / 19	MHS	cross-track $\theta_{zen} = \pm 50^{\circ}$ swath width = 1920 km 90 contiguous pixels	$f_0 \pm 1.0$	500	1.06
			$f_0\pm 3.0$	1000	0.70
			$f_0 + 7.0$	2200	0.84
DMSP F17 / F18	SSMI/S	conical $\theta_{zen} = 53.1^{\circ}$ swath width = 1707 km 180 samples	$f_0 \pm 1.0$	513	0.81
			$f_0\pm 3.0$	1019	0.67
			$f_0\pm 6.6$	1526	0.97

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- 572 **Table 2:** Mean calibration differences versus the GPM GMI radiometer for the cross-track sounders
- 573 listed in Table 1 (GMI sounder) for channels near the f_0 =183.31 GHz absorption line, computed by
- 574 the GPM XCAL intercalibration team.

				575
Satellite	Instrument	High frequency channels around f_0 = 183.31 GHz (GHz)	Calibration difference (K)	576
Megha- Tropiques	SAPHIR	$f_0 \pm 0.2$	0.18	
		$f_0 \pm 1.1$	-0.56	
		$f_0\pm 2.8$	-0.42	
		$f_0\pm 4.2$	-0.66	
		$f_0 \pm 6.8$	-0.32	
		$f_0 \pm 11.0$	-0.41	
Suomi-NPP	ATMS	$f_0 \pm 1.0$	0.40	
		$f_0 \pm 1.8$	-0.30	
		$f_0\pm 3.0$	-0.97	
		$f_0\pm 4.5$	-0.82	
		$f_0\pm 7.0$	-0.78	
MetOp-A	MHS	$f_0 \pm 1.0$	0.66	
		$f_0\pm 3.0$	-0.17	
		$f_0 + 7.0$	-0.14	
MetOp-B	MHS	$f_0 \pm 1.0$	0.60	
		$f_0\pm 3.0$	0.00	
		$f_0 + 7.0$	0.09	
NOAA-18	MHS	$f_0 \pm 1.0$	0.24	
		$f_0\pm 3.0$	-0.12	
		$f_0 + 7.0$	-0.06	
NOAA-19		$f_0 \pm 1.0$	0.94	
	MHS	$f_0 \pm 3.0$	0.03	
		$f_0 + 7.0$	-0.05	







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Figure 1: Mean observed minus calculated BT. All the calculated BTs are from RTTOVv11 run on 579 580 RAOBS measurements (triangles) or Météo-France NWP profiles (MF, circles) or European Centre for Medium-range Weather Forecasts NWP profiles (ECMWF, squares). Each color refers to a 581 specific sensor, as in the legend and the horizontal gray bars indicate the width of the band passes. 582 The RAOB measurements were collected during the Cooperative Indian Ocean Experiment on 583 584 Intraseasonal Variability (CINDY)/Dynamics of the Madden-Julian Oscillation (DYNAMO)/ARM 585 Madden-Julian Oscillation (MJO) Investigation Experiment (AMIE) field campaign, winter 2011-586 2012. The inset is a scaled representation of the 183GHz line, assuming a Van Vleck-Weisskopf 587 shape.

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Figure 2: First-guess departures (observed BTs - those calculated from the forecast background) for
 Suomi-NPP/ATMS (a) channel 18 (183±7GHz) and (b) channel 22 (183±1GHz) and the current
 ECMWF NWP system. The maps show the mean, with the global minima, maxima and means in the

595 captions.





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599 Figure 3: Self-broadened (pure water vapor or quadratic with humidity) and foreign-broadened (mixture with air or linear with humidity) continuum coefficients C (in $(dB/km)/(GHz kPa)^2$), as in 600 Rosenkranz (1998, 1999). Symbols are field and laboratory data [Bauer et al., 1993 & 1995; Kuhn et 601 al., 2002; Godon et al., 1992; Katkov et al., 1995; Liebe et al. 1984 & 1987; Koshelev et al., 2011]. 602 603 Statistical uncertainty of points in each series approximately equal or less than size of points. Solid lines are continuum coefficients derived by Rosenkranz (1998) for Millimeter-wave Propagation 604 Model (MPM). Dotted and dashed blue lines correspond to scaling of these coefficients on the basis 605 of radiometric data suggested by Turner (2009) and Payne et al. (2011). 606

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