

1 **Ozone Comparison between Pandora #34, Dobson #061, OMI, and OMPS at Boulder**  
2 **Colorado for the period December 2013 – December 2016.**

3  
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14  
15 **Abstract**

16 A one-time calibrated (in December 2013) Pandora Spectrometer Instrument (Pan #034) has  
17 been compared to a periodically calibrated Dobson spectroradiometer (Dobson #061) co-located  
18 in Boulder, Colorado, and compared with two satellite instruments over a 3-year period  
19 (December 2013 – December 2016). The results show good agreement between Pan#034 and  
20 Dobson#061 within their statistical uncertainties. Both records are corrected for ozone retrieval  
21 sensitivity to stratospheric temperature variability obtained from the Global Modeling Initiative  
22 (GMI) and Modern-Era Retrospective analysis for Research and Applications (MERRA-2)  
23 model calculations. Pandora#034 and Dobson#061 differ by an average of  $2.1 \pm 3.2 \%$  when  
24 both instruments use their standard ozone absorption cross sections in the retrievals algorithms.  
25 The results show a relative drift ( $0.2 \pm 0.08\%$  per year) between Pandora observations against  
26 NOAA Dobson in Boulder, CO over a three-year period of continuous operation. Pandora drifts  
27 relative to the satellite Ozone Monitoring Instrument OMI and the Ozone Mapping Profiler  
28 OMPS are  $+0.18 \pm 0.2 \%$  per year and  $-0.18 \pm 0.2 \%$  per year, respectively, where the  
29 uncertainties are two standard deviations. The drift between Dobson #061 and OMPS for a 5.5-  
30 year period (January 2012 – June 2017) is  $-0.07 \pm 0.06 \%$  per year.

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38

39 **Introduction**

40 A Pandora Spectrometer Instrument #034 (PSI) located on top of the NOAA building in  
41 Boulder, Colorado has been operating since December 2013 with little maintenance and using  
42 the original calibration. The purpose of this paper is to present a comparison between two co-  
43 located ozone measuring instruments, Pandora #034 and Dobson #061 for the period December  
44 2013 to December 2016. Additional comparisons are made with satellite overpass data from  
45 OMI (Ozone Monitoring Instrument on board the AURA spacecraft) and OMPS (Ozone  
46 Mapping Profiler Suite on board the Suomi NPOESS satellite). This paper is an extension of a  
47 previously published paper (Herman et al., 2015) that presented just 1 year of data. The results  
48 demonstrate the accuracy and stability of both the Dobson and PSI for retrieval of total column  
49 ozone, and serves as a validation demonstration at one location for both the fairly new PSI and  
50 for satellite ozone data from OMI and OMPS. Part of the experiment comparing Pandora #034 to  
51 Dobson #061 was to see if Pandora #034 would perform well over a long period without  
52 additional calibration or adjustments. The only change made during the period 2014 to the  
53 present (August 2017) was to replace a broken motor on the suntracker that caused a data gap in  
54 early 2016.

55  
56 The characteristics of both the PSI and the Dobson Spectroradiometer are described in  
57 Herman et al. (2015). Briefly, the PSI consists of a small Avantes low stray light spectrometer  
58 (280 – 525 nm with 0.6 nm spectral resolution with 5 times oversampling) connected to an  
59 optical head by a 400 micron core diameter single strand fiber optic cable. The spectrometer is  
60 temperature stabilized at 20<sup>o</sup>C inside of a weather resistant container. The optical head consists  
61 of a collimator and lens giving rise to a 2.5<sup>o</sup> FOV (field of view) FWHM (Full Width Half  
62 Maximum) with light passing through two filter wheels containing diffusers, open hole, a UV340  
63 filter (blocks visible light), neutral density filters, and an opaque position (dark current  
64 measurement). The optical head is connected to a small suntracker capable of accurately  
65 following the sun’s center using a small computer-data logger contained in a weatherproof box  
66 along with the spectrometer. Pandora#034 is capable of obtaining NO<sub>2</sub> and Total Column Ozone  
67 TCO amounts sequentially over a period of 80 seconds. The integration time in bright sun is  
68 about 4 milli-seconds that is repeated and averaged for 30 seconds to obtain very high signal to  
69 noise and an ozone precision of less than 1 DU or 0.2% (1 DU = 2.69x10<sup>16</sup> molecules/cm<sup>2</sup>).

70  
71 The Dobson record in Boulder started in 1966 based on an improved design from the  
72 instrument first deployed in the 1920’s (Dobson, 1931). Dobson instrument uses a differential  
73 absorption method to derive total column ozone from direct–sun measurements using two UV  
74 wavelength pairs in the 300 – 340 nm range (see Herman et al., 2015). The extensive Dobson  
75 network uses the Bass-Paur (BP) ozone absorption cross sections (Bass and Paur, 1985) for  
76 operational data processing (Komhyr et al., 1993).

77  
78

79 All NOAA Dobson instruments are periodically calibrated against WMO world standard  
80 Dobson #083, which in turn uses Langley method calibrations at the Mauna Loa Observatory  
81 station (Komhyr et al., 1989). Standard lamps are used to check Dobson spectral registration  
82 stability. Recently, July 2017, intermediate calibrations from Dobson #083 were applied to the  
83 Dobson #061 ozone data record that improved its comparison with satellite data (the calibration  
84 updates were processed by one of the co-authors, Koji Miyagawa).

85  
86 The main sources of noise in the PSI measurement comes from the presence of clouds or  
87 haze in the FOV, which increases the exposure time needed to fill the CCD wells to 80% and  
88 reduces the number of measurements in 30 seconds. For this comparison study, data were  
89 selected for scenes that are clear-sky conditions as determined from the Dobson A-D pair direct-  
90 sun data record.

91  
92 Accuracy in the PSI spectral fitting retrieval is obtained using careful measurements of the  
93 spectrometer's slit function, wavelength calibration, and knowledge of the solar spectrum at the  
94 top of the atmosphere. The current operational PSI ozone retrieval algorithm used in this study is  
95 based on extraterrestrial solar flux from a combination of the Kurucz spectrum (wavelength  
96 resolution  $\lambda/1\lambda = 500\ 000$ ) radiometrically normalized to the lower-resolution shuttle Atlas-3  
97 SUSIM spectrum (Van Hoosier, 1996; Bernhard et al., 2004, 2005), BDM ozone cross sections  
98 (Brion et al. (1993, 1998) and Malicet et al. (1995)), corrections for stray light, and an effective  
99 ozone weighted temperature.

100  
101 The Dobson data used in this study contain the individual measurements (more than 1 per day  
102 between 09:00 and 15:00 local time with almost all of the data between 10:00 and 14:00) for  
103 clear-sky direct-sun observations using the quartz plate and A-D wavelength pairs for ozone  
104 retrieval (Dobson label ADDSGQP). These were made available by one of the co-authors (I.  
105 Petropavlovskikh, private communication, Table 2). The NOAA Dobson total ozone data are  
106 typically archived at WOUDC (World Ozone and Ultraviolet Radiation Data Centre) or NDACC  
107 (Network for the Detection of Atmospheric Composition Change) with one representative ozone  
108 value per day.

## 109 110 **1. Temperature Sensitivity**

111  
112 The PSI ozone retrieval algorithm is more sensitive to the effective ozone weighted  
113 average temperature than is the 4 wavelength Dobson retrieval (Redondas et al., 2014).  
114 Neglecting the temperature sensitivity creates a seasonal difference between the two instruments.  
115 To correct for this, we use an effective ozone temperature  $T_E$  based on daily ozone profile  
116 weighted altitude temperature averages (Redondas et al., 2014). The temperature and ozone  
117 profile data were obtained from the GMI (Global Modeling Initiative) model calculation for  
118 2012 to 2016. (<https://gmi.gsfc.nasa.gov/merra2hindcast/>). The GMI model provides

119 atmospheric composition hindcasts using MERRA-2 (Modern-Era Retrospective analysis for  
 120 Research and Applications, Version 2, meteorology (Strahan et al., 2013; Wargan and Coy,  
 121 2012) <https://gmao.gsfc.nasa.gov/reanalysis/MERRA-2/>). The simulation with 2 x 2.5 degree  
 122 resolution uses the CCMI (Chemistry–Climate Modelling Initiative, Morgenstern et al., 2017)  
 123 emissions and boundary conditions. MERRA-2 uses assimilation schemes based on  
 124 hyperspectral radiation, microwave observations and ozone satellite measurements. The resulting  
 125 seasonal cycle for  $T_E$  shows variations over the four year period, while day-to-day variability is  
 126 enhanced during winter and spring season (Fig. 1). An estimated 5<sup>th</sup> year (2017) has been added  
 127 (Fig. 1) by forming the average of the daily temperatures from the 2013 to 2016 period.

128  
 129 The  $T_E$  time series data are used for an ozone retrieval temperature correction  $TCO_{corr}$   
 130 coefficient per  $^{\circ}K$  given in the form  $TCO_{corr} = TCO (1 + C(T))$  and  $O_3(corr) = O_3 TCO_{corr}$   
 131 (Herman et al., 2015), where  $C(T_E)$  is given by eqns. 1 and 2.

$$132 \quad C_{\text{Pandora-BDM}}(T_E) = 0.00333(T_E - 225) \quad (\text{Herman et al., 2015}) \quad (1)$$

$$C_{\text{Dobson-BP}}(T_E) = -0.0013(T_E - 226.7) \quad (\text{Redondas et al., 2014}) \quad (2)$$

$$C_{\text{Dobson-BDM}}(T_E) = 0.00042(T_E - 226.7) \quad (\text{Redondas et al., 2014}) \quad (3)$$

133 As mentioned earlier, the Dobson TCO retrieval normally uses the Bass and Paur (BP)  
 134 ozone absorption coefficients, while Pandora uses the Brion-Daumont-Malicet (BDM)  
 135 coefficients. A change in  $T_E$  of  $+1^{\circ}K$  change leads to TCO changes for the Pandora(BDM),  
 136 Dobson(BP), and Dobson(BDM) instruments of  $+0.33\%$ ,  $-0.13\%$ , and  $0.042\%$ , respectively.  
 137 For a nominal TCO value of 325 DU, the change would be  $+1.1$  and  $-0.4$  DU, a net relative  
 138 change of 1.5 DU for a  $1^{\circ}K$  change between Pandora(BDM) and Dobson(BP).

139  
 140 While BDM cross sections are not currently recommended for use in standard Dobson  
 141 processing, their use yields slightly different values of TCO and a smaller sensitivity to  
 142 temperature. The basic Dobson algorithm, based on pairs of wavelengths, is intrinsically less  
 143 sensitive to  $T_E$  than Pandora's spectral fitting retrieval.

## 144 2. TCO Comparisons between Pandora, Dobson, OMI and OMPS

145  
 146 Comparing retrieved TCO from the PSI, Dobson, OMI and OMPS instruments show that  
 147 there are small, but significant differences between the PSI and Dobson instruments and between  
 148 the ground-based instruments and satellite derived values of TCO. The difference is calculated  
 149 using three-year estimates of secular change based on a linear least squares fit to the percent  
 150 differences between the instruments. The cloud-free direct-sun A-D pair Dobson ozone data are  
 151 selected for comparison with time-matched Pandora#034 retrieved ozone data (Herman et al.,  
 152 2015). The Pandora#034 retrieved ozone (every 80 seconds) are matched to the less frequent  
 153

154 Dobson#061 retrieval times that are obtained for mid-day solar zenith angles (SZAs) and  
155 averaged over  $\pm 8$  minutes (Fig. 2A).

156  
157 Each clear-sky PSI data point is an average of 2000 (early morning to evening SZAs) to  
158 4000 (mid-day SZAs) measurements obtained during 30 seconds. All data for this study were  
159 clear-sky within the instrument's field of view based on the Dobson criteria for A-D-pair direct-  
160 sun clear sky. In addition, the PSI data are averaged over a period of  $\pm 8$  minutes surrounding  
161 the Dobson time of measurement (2 to 3 times per day). Since PSI measurements are obtained  
162 every 80 seconds, there were an additional 10 PSI data points averaged together to compare to  
163 each Dobson, OMI, or OMPS measurement. The result is high signal to noise values for Pandora  
164 and high precision (0.1%). The same procedure using cloud-screened PSI data was used for  
165 comparisons with OMI and OMPS, where they measure once or twice per day over Boulder,  
166 Colorado. Some of the variations in the day to day ozone values are driven by changes in the  
167 local weather over Boulder, Colorado (see Fig. 14 in Herman et al., 2015), with weekly averages  
168 having much smaller variation.

169  
170 Figure 2B shows a Lowess(0.1) fits to the two time series in Fig. 2A that is approximately  
171 equivalent to a 3-month running average. The Lowess(f) procedure is based on local least  
172 squares fitting using low order polynomials applied to a specified fraction  $f$  of the data  
173 (Cleveland, 1979) that reduces the effect of outlier points from the mean. The smooth curves  
174 show a small variable difference between the Dobson and Pandora time series. Fig. 2C shows the  
175 percent difference PD between the time series in Fig. 2A and the residual seasonal variation in  
176 PD. Estimating the slope of the least squares fit to the percent difference can be sensitive to the  
177 selection of the end points of the time series. This effect can be minimized by removing the  
178 seasonal time dependence (Fig. 2C) using a low-pass filter function with zero slope derived from  
179 the Lowess(0.1) fit. The result is shown in Fig. 3A.

180  
181 Figure 3 shows the de-seasonalized percent differences PD(A,B) for six pairs between  
182 Pandora #034, Dobson #061, OMI, and OMPS for the 3-year period 2014 – 2016 (summarized in  
183 Table 1). The slightly curvy Lowess (0.1) lines about each linear fit show the residual seasonal  
184 cycles, which are too small to have an effect on slope determination. Error estimates (Fig. 3 and  
185 Table 1) for the linear least squares slopes and averages are one standard deviation (1-STD).  
186 Some of the error estimates are large enough to make the statistical significance of the slopes  
187 marginal (see Panel E OMPS vs Pandora;  $0.18 \pm 0.098$ ,  $p = 0.06$ ), while others are significant  
188 (see Panel D OMI vs Dobson:  $-0.18 \pm 0.08$ ,  $p = 0.03$ ) at the 2-STD level. The significance  
189 probability parameter  $p$  is given, where  $p$  is the probability (0 to 1) that the slope is statistically  
190 different from 0 relative to  $p = 0.05$ . Also shown are the numbers of data points in each time  
191 series.

192  
193

194 After removal of the residual seasonal variation in the calculated percent differences,  
195 there still is a statistically significant drift of 0.2% per year ( $p < 0.001$ ) between the Pandora#034  
196 and Dobson#061 (Panels A and B in Fig. 3) using either BP or BDM ozone cross sections for  
197 Dobson#061. The differences in the mean values (-2.1 and -2.8%) are not significant at the 2-  
198 STD level.

199  
200 The linear trend (Panel C,  $-0.09 \pm 0.08$  % per year,  $p = 0.3$ ) between the Dobson and  
201 OMPS is not significantly different from zero, while the drift with OMI (Panel D,  $-0.18 \pm 0.08$  %  
202 per year,  $p = 0.03$ ) is significant. This suggests that OMI ozone retrievals are drifting with  
203 respect to OMPS and the Dobson. Extending the period from 2012 to June 2017 gives a very  
204 small, but significant trend,  $-0.07 \pm 0.03$  % per year,  $p = 0.047$  for PD(OMPS,Dobson).

205  
206 Calculations for Pandora#034 (Panels E and F in Fig. 3) show marginally significant ( $p =$   
207 0.06) trends for Pandora#034 compared to OMPS (Panel E,  $-0.18 \pm 0.098$  % per year) and OMI  
208 (Panel F,  $+0.18 \pm 0.096$  % per year). If the Pandora#034 time series is extended into 2017 to  
209 minimize the effect of missing Pandora data in 2016, then the trends for Pandora compared to  
210 OMPS ( $-0.2 \pm 0.08$  % / Year  $p = 0.013$ ) and OMI ( $0.15 \pm 0.076$   $p=0.05$ ) are significant, but not  
211 different from the shorter 2014 – 2016 period. The secular trends for the difference between  
212 Pandora#034 and Dobson#061 ( $-0.2\%$  per year) are almost the same for both Dobson BP and  
213 BDM ozone absorption coefficients even though the temperature sensitivity using the Dobson  
214 BDM ozone absorption coefficients is small ( $0.042\%$  per  $^{\circ}\text{C}$ ). This suggests that the  
215 stratospheric effective ozone temperature change is not a source for the small differences  
216 between Pandora#034 and Dobson#061.

217  
218 Figure 4 shows that the TCO between Pandora#034 and Dobson#061 are highly  
219 correlated with 1:1 slope and the correlation coefficient  $r^2 = 0.97$  for the 3-year period 2014 to  
220 2016. Similar correlation plots (Fig. 5) for Pandora#034 and Dobson#061 with OMI and OMPS  
221 also show very high correlations. The correlations in TCO are obtained after only temperature  
222 corrections to Pandora#034 and Dobson#061 using  $T_E$  (TCO pairs similar to Fig. 2, panel A).

223  
224 The Pandora, OMI, and OMPS data used in this study are from the overpass files located  
225 on the public websites (Table 2).

## 226 227 **Summary**

228 Temperature corrected Pandora#034 and Dobson#061 differ by an average of  $2.1 \pm 3.2$  %  
229 with Pandora using its standard retrieval BDM ozone absorption cross sections and Dobson  
230 using the recommended BP ozone absorption cross sections. Pandora compared to Dobson  
231 shows a small, but significant drift ( $-0.2 \pm 0.08$  % per year,  $p < 0.001$ ) for the 2014 – 2016  
232 period. Comparisons of Pandora with OMI and OMPS are marginally significant drifts of  
233  $0.18 \pm 0.2$  and  $-0.18 \pm 0.2$   $p=0.06$  for 2014-2016, but are significant ( $0.15 \pm 0.15$  % per year,  
234  $p=0.05$  and  $-0.2 \pm 0.16$  % per year,  $p = 0.013$ , respectively), if the period is extended to mid-

235 2017 to minimize the effect of missing Pandora data during 2016. The small Pandora and  
236 Dobson trends compared to OMPS suggest that both instruments are stable. The conclusion is  
237 that the periodically calibrated Dobson#061 is able to detect smaller ozone trends than a Pandora  
238 instrument with no intermediate calibration during a 3-year period. The longer term trend for  
239 Dobson compared to OMPS for a 5.5-year period (2012 – June 2017) is  $-0.07 \pm 0.06$  % per year,  
240  $p = 0.047$ . All error estimates are two STD.

241  
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243 for supplying the atmospheric temperature data for Boulder, Colorado.

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317



318 **Tables**

319

Table 1 Percent Difference Summary of Linear Fit Slopes and Mean Differences in Fig. 3

Percent Diff(A,B)	Slope (% per Year)	Probability	Mean (%)	Points	Panel
Pan, Dob(BP)	-0.2 ± 0.04	P < 0.001	-2.1 ± 1.6	2020	A
Pan, Dob(BDM)	-0.2 ± 0.04	P < 0.001	-2.8 ± 1.6	2020	B
OMPS, Dob(BP)	-0.09 ± 0.08	P = 0.3	-1.4 ± 2.1	854	C
OMI, Dob(BP)	-0.18 ± 0.08	P = 0.03	-1.4 ± 1.9	654	D
OMPS, Pan	-0.18 ± 0.098	P = 0.06	0.96 ± 2.7	952	E
OMI, Pan	+0.18 ± 0.096	P = 0.06	1.1 ± 2.1	624	F

320

321

322

323

Table 2 Data Availability

**OMI:**

[https://avdc.gsfc.nasa.gov/index.php?site=1593048672&id=28/aura\\_omi\\_l2ovp\\_omto3\\_v8.5\\_boulder.co\\_067.txt](https://avdc.gsfc.nasa.gov/index.php?site=1593048672&id=28/aura_omi_l2ovp_omto3_v8.5_boulder.co_067.txt)

**OMPS:**

[ftp://toms.gsfc.nasa.gov/pub/omps\\_tc/overpass/suomi\\_npp\\_omps\\_l2ovp\\_nmto3\\_v02\\_boulder.co\\_067.txt](ftp://toms.gsfc.nasa.gov/pub/omps_tc/overpass/suomi_npp_omps_l2ovp_nmto3_v02_boulder.co_067.txt)

**Pandora34:**

<https://avdc.gsfc.nasa.gov/pub/DSCOVER/Pandora/DATA/Boulder/Pandora34/L3c/>

**Dobson061:**

[ftp://aftp.cmdl.noaa.gov/data/ozwv/Dobson/WinDobson/Pandora%20comparisons/Dobson61%20Boulder%20Ad-dsgqp%20120213-032717\\_w\\_Header.txt](ftp://aftp.cmdl.noaa.gov/data/ozwv/Dobson/WinDobson/Pandora%20comparisons/Dobson61%20Boulder%20Ad-dsgqp%20120213-032717_w_Header.txt)

324 **Figure Captions**

325 Fig. 1 Calculated  $T_E$  using model estimates of  $O_3$  and temperature profiles. The Trend is  
326 calculated from the difference of  $T_E$  from its 4-year daily mean that is also used for year 2017  
327 labelled Avg.

328  
329 Fig. 2 Panel A shows the retrieved ozone time series (December 2013 – June 2017) for Pandora  
330 (red) and Dobson (Black). Panel B shows Lowess(0.1) fit to the each time series. Panel C shows  
331 the percent difference, a linear least squares fit, and a Lowess(0.1) fit showing seasonal residuals.  
332

333 Fig. 3 Comparisons of Pandora(BDM) with Dobson(BP and BDM) retrieved ozone for Boulder,  
334 Colorado in percent differences of retrieved ozone and comparisons with OMI and OMPS. Slope  
335 = value of the linear least square fit,  $\pm N$  is 1 STD, and p is the probability (0 to 1) that the slope  
336 is statistically different from 0 relative to  $p = 0.05$ . The solid lines are a Lowess(0.1) fit and a  
337 linear least squares fit.

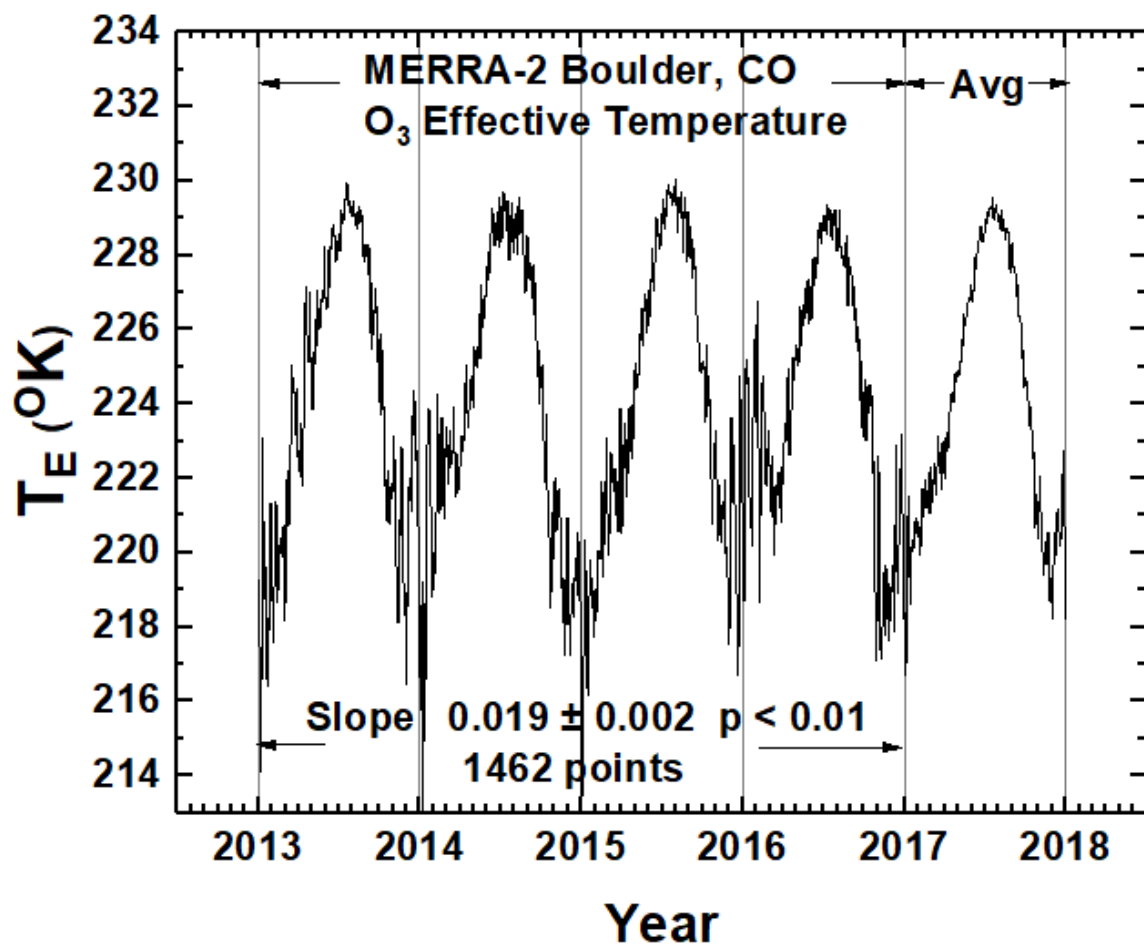
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339 Fig. 4 Correlation between Pandora #034 and Dobson #061: 2014 – 2016

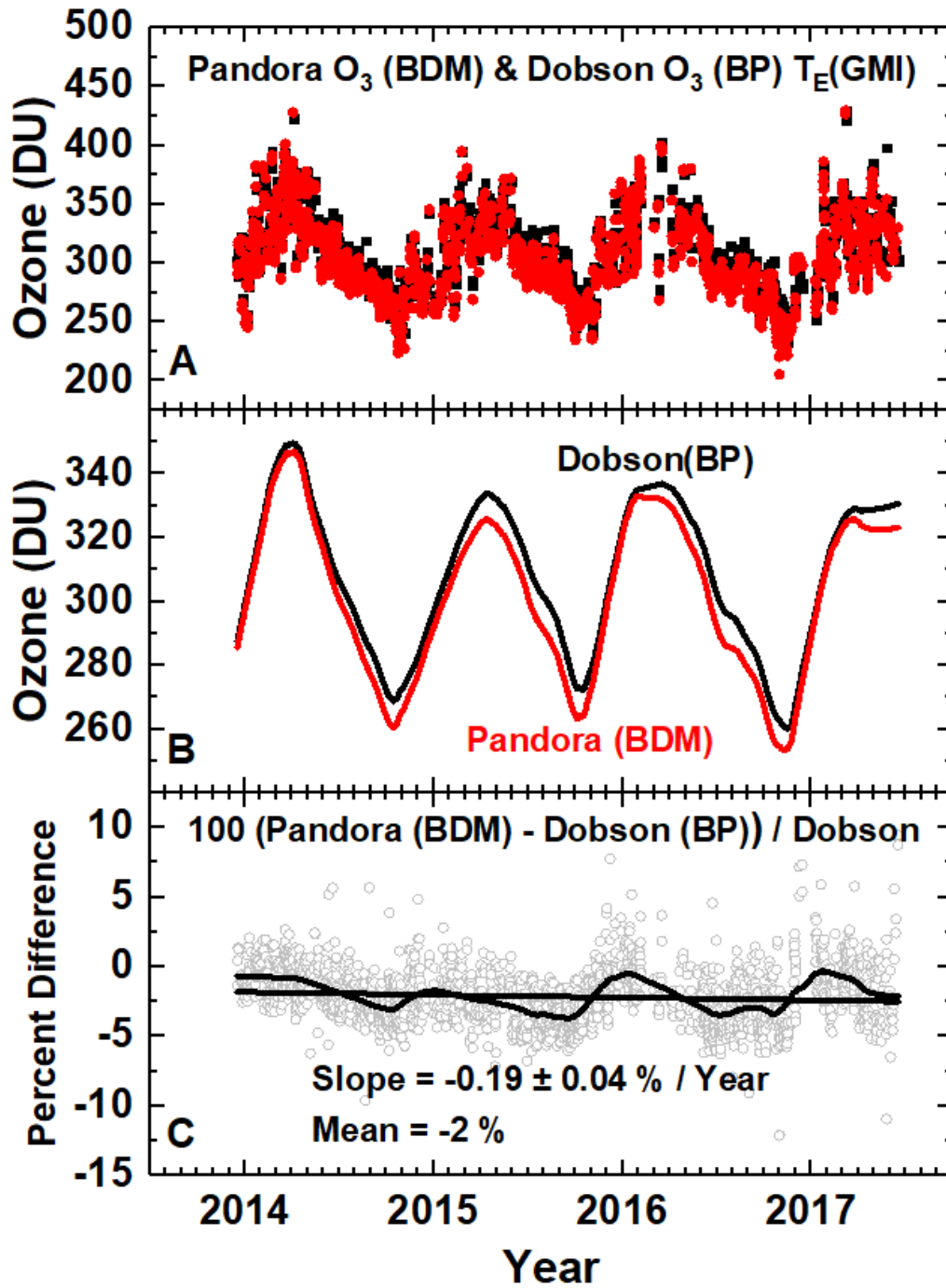
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341 Fig. 5 Correlation of Pandora#034 and Dobson#061 with OMI and OMPS: 2014 - 2016

342 **Figures**  
343

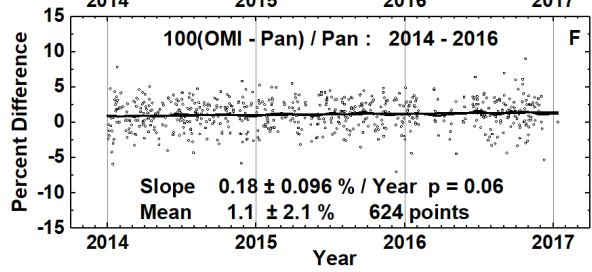
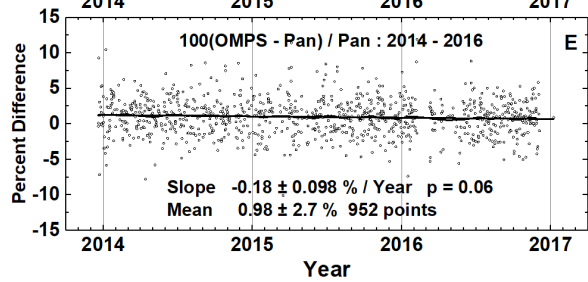
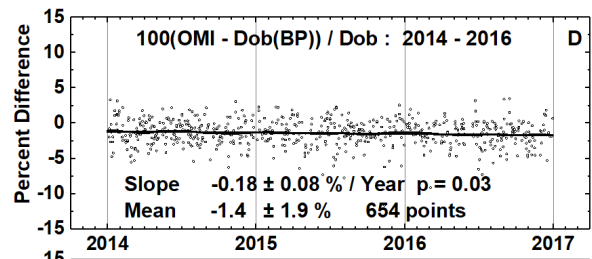
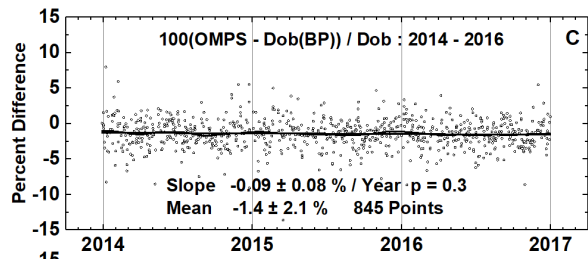
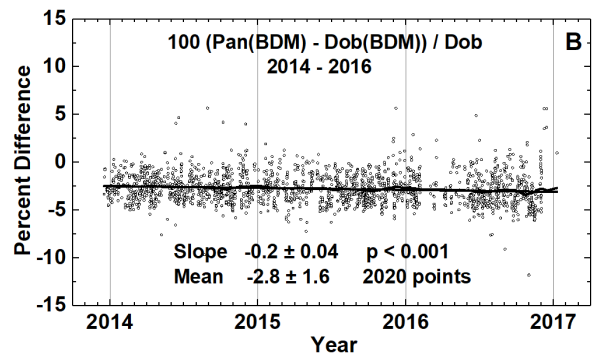
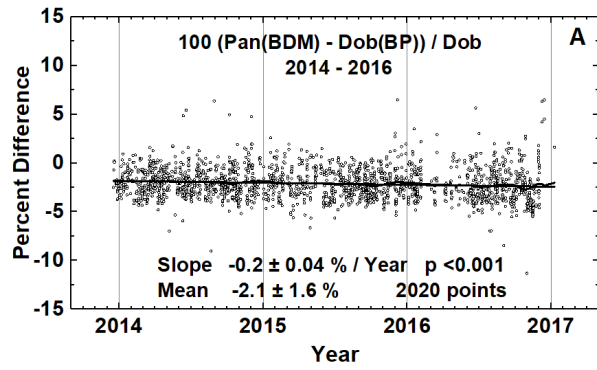


344  
345 **F1**



346 **F2**

347



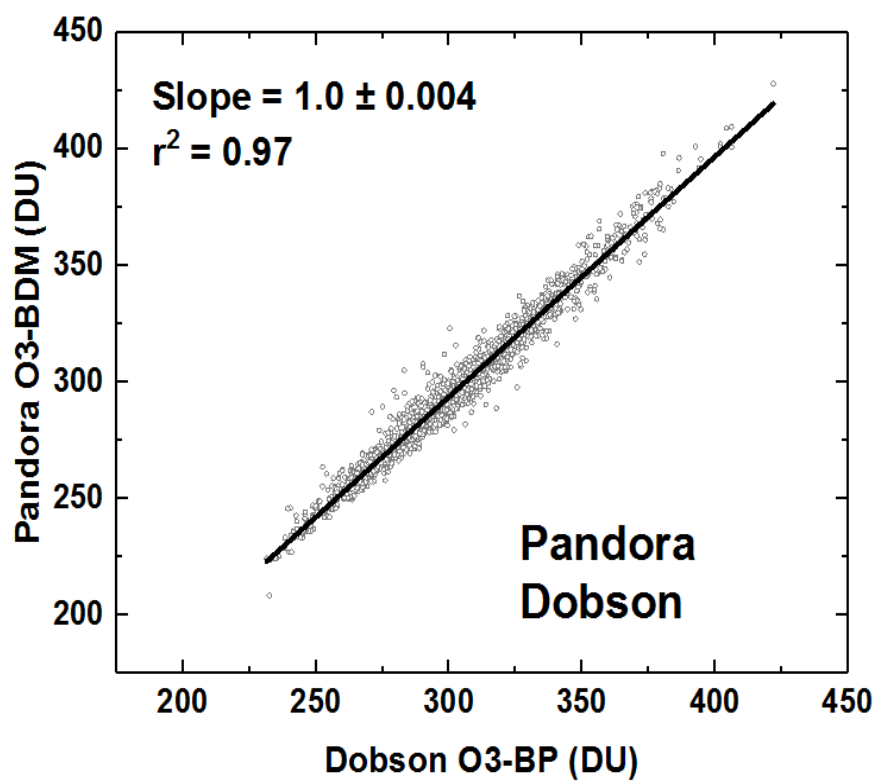
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350 **F3**

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355 **F4**

