Atmos. Meas. Tech. Discuss., doi:10.5194/amt-2017-269-AC1, 2017 © Author(s) 2017. This work is distributed under the Creative Commons Attribution 4.0 License.



Interactive comment on "Retrieval of $O_2(^1\Sigma)$ and $O_2(^1\Delta)$ volume emission rates in the mesosphere and lower thermosphere using SCIAMACHY MLT limb scans" by Amirmahdi Zarboo et al.

Amirmahdi Zarboo et al.

amirmahdi.zarboo@kit.edu

Received and published: 16 November 2017

[ACPD,manuscript]copernicus

amsmath

We would like to thank the referee for the careful review of the paper.

1. Lines 10–12 on page 1. It is mentioned that " $O2(^1\Sigma)$ shows secondary maxima during winter and spring, which are related to the downwelling of atomic oxygen after large

C1

sudden stratospheric warmings (SSW)." What does the "secondary maxima" mean? It was not pointed out in the description in the text.

Response: We mean "Additional polar high values of $O_2(^1\Sigma)$ ".

The manuscript is changed as follows:

" $O_2(^1\Sigma)$ emissions show additional high values at polar latitudes during winter and spring. These additional high values are presumably related to the downwelling of atomic oxygen after large sudden stratospheric warmings (SSW)."

2. Line 2 on page 2. It is better to add a reference for the sentence "has been a matter of dispute for some years".

Response: In our response we take into account the comments from both of the reviewers. The text now reads as follows:

 O_2^* represents any of the seven states below the first dissociation limit. Bates and others argue that the population distribution between these states can best be approximated statistically, in which the $^5\Pi_g$ state is produced in almost 40% of the collisions (Smith, 1984; Bates, 1992; Wraight, 1982). Most of the O_2^* derived from recombination is found in the $A^3\Sigma_u^+$ state (Slanger and Copeland, 2003), and in a recent review Huestis concludes that all of the recombining atoms pass through the Herzberg states $c^1\Sigma_u^-$, $A'^3\Delta_u$, and $A^3\Sigma_u^+$ (Huestis, 2013). Stegman and Murtagh (1991) provide the quenching parameters resulted from analysing the measurements of the near-ultraviolet portion of the nightglow to fit the synthetic spectra of the Herzberg bands of O_2 . These parameters set an upper limit of 10% production efficiency on the generation of $O_2(c^1\Sigma_u^-)$ in the atomic oxygen association reaction . Admittedly, proper accounting of the correct products of (R1) can be complex. Recent research has investigated this issue, e.g. Kirillov (2012, 2014). Therefore we assume

the production of a surrogate "hybrid" state O₂ in the photochemical model.

3. Line 4 on page 5. It is better to add "based on $O_2(^1\Sigma)$ emission " after "performed remote sensing measurements of upper atmosphere winds and temperatures."

This phrase was added to the corresponding place.

4. Line 20–21. "Daylight measurements by SCIAMACHY during the Envisat orbit begin with limb measurements of the twilit atmosphere (Bovensmann et al., 1999) which are located in the northern polar region (above 75 °N)." I can not catch the meaning of this sentence.

Response: The text was changed to make the meaning more clear. The manuscript is changed as follows:

"A typical orbit starts with a limb measurement of the twilit atmosphere, followed by the solar occultation measurement during sunrise over the North Pole and an optimized limb-nadir sequence. (Bovensmann et al., 1999)."

5. Lines 27–28 on page 9. Figure 4a just shows emission intensities at some altitudes; especially the intensity at 83 km is not shown in this Figure. How to draw the conclusion from Figure 4a that the largest SNR of the emission intensities for daytime $O_2(^1\Sigma)$ are located around 83–99 km and the largest SNR are ovserved at the edges of the altitude range?

Response: Indeed, this result corresponds to a draft version of the manuscript in which, the emission intensities have been shown for all of the measurement altitudes.

СЗ

Whereas in this version, only some of the measurement altitudes are shown for the purpose of improving the visibility of the emission intensities inside the figure. Yes, it should be stated like this:

"Evaluating all altitudes, not shown here but indicated in Figure 4a, we observe the strongest signal of daytime $O_2(^1\Sigma)$ around 83–99 km."

6. Line 2 on page 10. "The largest SNRs at the lowest altitudes (Figure 4b),". Does the "lowest altitudes" means 54 km? If 54 km is the lowest altitude in the retrieval of $O_2(^1\Delta)$ emission intensity in this paper? Is the intensity below 54 km not significant or can it not be observed?

Response: In accordance with the answer to the previous question, not all of the measurement altitudes are included in Figure 4. Yes, the 54 km is the lowest observable altitude in our measurements. We changed the manuscript as follows:

"The daytime $O_2(^1\Delta)$ emission intensities are strongest at the lowest observable altitudes, i.e., 54 km (Figure 4b). The strongest twilight $O_2(^1\Delta)$ emissions are located in the 83–96 km altitude range (Figure 4d shows only a selection of altitudes)."

7. Line 4 on page10. Similar to question 5, the intensity at 83 km is not given in Figure

Response: See reply to point #6 and the updated text therein.

8. It would be better to add a simple introduction about Figure 5 before that about Figure 6, for example in line 7 on page 10.

Thanks for pointing this out. A short introduction about Figure 5 has been added before the introduction about Figure 6. Change in manuscript: "Volume emission rate profiles for one sample satellite orbit (41455 on 03/02/2010) for daytime $O_2(^1\Sigma)$ and $O_2(^1\Delta)$ are shown in Figure 5a and Figure 5b respectively. Examples of the volume emission rate latitude-altitude distributions for the same orbit for daytime $O_2(^1\Sigma)$ are shown in Figure 6a and for daytime $O_2(^1\Delta)$ in Figure 6b."

9. Line 10 on page 10. Figure 5b gives the altitude distributions above 50 km (may be 54 km). And some profiles have not reached their maxima at the lowest altitude given in the figure. The altitude distributions of $O_2(^1\Delta)$ dayglow observed by the TIMED/SABER satellite often show its maximum around 50 km (usually lower than 50 km). The distribution of $O_2(^1\Delta)$ dayglow were given by Mlynczak et al. [2007,JGR, 112, D15306]. It is suggested to compare the altitude distributions of $O_2(^1\Delta)$ dayglow with that given by Mlynczak et al.

Response: An increase in the dayglow $O_2(^1\Delta)$ VER profiles is observed with decreasing altitude, but no turning point or maximum is observed.

We changed the manuscript as follows:

"The volume emission rate profile of dayglow $O_2(^1\Delta)$ observed by TIMED/SABER, often has its maximum around 50 km altitude, as shown for example in Figure 1 of Mlynczak et. al. (2007). Figure 5b shows that the SCIAMACHY MLT volume emission rate profiles are largest at the bottom of the observed altitude range, around 54 km. These VER profiles sometimes show secondary maxima in the range 80–90 km, which are at least one order of magnitude smaller than the largest SCIAMACHY VER. This secondary maximum occurs especially around equinox times."

C5

10. "around 80° N latitude" is mentioned in both lines 1 and 3 on page 12. Are both the VERs significant at only this latitude and not significant at other latitudes?

Response: There are only twilight measurements at these latitudes. The significance relates to the altitude range. Therefore, since mentioning 80°N latitude, and 'significant data' is misleading, we removed these terms from this part of the text.

It should be noted that we make an improvement to the geometry of the retrieval and we get the smoother results with smaller variations. This is true for all of our results and we make the figures once again. So we get less noise and therefore the signal to noise ratio of the data is improved. As it can be seen, we have data with signal to noise ratio of greater than one in most of the altitudes and latitudes.

Change in the manuscript: "To assess the signal to noise ratio for the daytime VERs, Figure 6e shows the daily mean $O_2(^1\Sigma)$ VERs signal to noise ratios. We observe the strongest signal of daytime $O_2(^1\Sigma)$ in the 70–130 km altitude range. The strongest signal of the twilight $O_2(^1\Sigma)$ is observed between 84 km and 95 km (not shown here). Figure 6f shows that the stronger signal of daytime $O_2(^1\Delta)$ is observed below 105 km, with the strongest around 70 km. The largest signal of twilight $O_2(^1\Delta)$ is observed in the altitude range of 83–97 km (not shown). "

11. Lines 13–14 on page 14. "The altitudes of the peak values of $O_2(^1\Sigma)$ roughly follows the maximum intensity of solar radiance". This can not be seen clearly from Figure 8c. What is the meaning of the maximum intensity of solar radiance?

We refer to the light blue and yellow areas of the Figure 8c. We refer to the maximum intensity of solar radiance as the latitudes and times in which solar zenith angles have their lowest values in.

Change in the manuscript: "The altitudes of the peak values of $O_2(^1\Sigma)$ roughly follows

the maximum intensity of solar radiance, but shows highest values at low-to middle latitudes. We refer to the maximum intensity of solar radiance as the latitudes and times in which solar zenith angles have their lowest values in."

12. Lines 19–20 on page 14. The "maximal value" and "peak altitude" are derived from the profile between 85–100 km. Therefore, when the second peak does not appear or the second peak is not obvious, the derived maximal value must be the emission at 85 km and the peak altitude must be 85 km. In the situation, the peak altitude given here is not the altitude for a real peak. This means the peak values and peak altitudes at some latitudes or some time shown in Figure 8 are not for real peaks but for 85 km; however, they are for the real peaks when the second peaks are obvious. In fact, the lowest peak altitudes (85 km) occur not only in summer at high latitudes but also in the tropics in Figure 8d. It is better to solve this problem. At least, it should be explained in the paper in case some readers could think that all the maximum values and peak altitudes given in Figure 8 are for real second peaks.

Response: Indeed, with changing Figure 8b, 8d, and 8f such that only the regions where the secondary maxima of $O_2(^1\Delta)$ exist, we see that secondary maxima are confined to winter at mid-to high latitudes. That means they definitely do not follow the position of the sun - i.e. the highest photolysis rates of O_3 . There are two possibilities for this - this is the region where the secondary ozone maximum is strongest, also atomic oxygen densities might be strongest due to enhanced mixing with the lower thermosphere.

Changes in the manuscript:

On page 14 line 11: "For this, we derive the maximum values from the daily mean VERs for $O_2(^1\Sigma)$ and between 85–100 km altitude for $O_2(^1\Delta)$, which are shown in Figures 8a and 8b respectively. Only those regions are shown in Figures 8b, 8d, and 8f where the secondary maxima of $O_2(^1\Delta)$ exist."

C7

On page 14, line 19: " The secondary maxima of the $O_2(^1\Delta)$ VER are confined to winter at mid-to-high latitudes. "

On page 15, line 4: "There are two possibilities for secondary maxima of $O_2(^1\Delta)$. They happen in the region where the secondary ozone maximum is strongest. Also, atomic oxygen densities might be strongest due to enhanced mixing with the lower thermosphere. Detailed study of the processes which result in the formation of the secondary maxima of $O_2(^1\Delta)$ is beyond our work."

On page 15, line 13: "In the regions where the secondary maxima of $O_2(^1\Delta)$ happen, the peak altitudes occur in \sim 84–89 km altitude range."

On page 15, line 25: "The $O_2(^1\Delta)$ secondary maximums occur in winter at high latitudes. The values of the $O_2(^1\Delta)$ h_{CA}rangebetween~88 km and ~89 km altitude."

On page 20, line 2: "The time series of $O_2(^1\Delta)$ VER is two orders of magnitude larger than $O_2(^1\Sigma)$ VER at its maximal values which are located below the observation altitude (<60 km), but it shows some secondary maxima with about one order of magnitude smaller than the primary maxima at 84-89 km. This happens in winter at high latitudes."

13. Equation (6). Please give the altitude range.

Response: It "ranges from $\sim\!50$ km to $\sim\!150$ km for $O_2(^1\Sigma)$ and from $\sim\!85$ km to $\sim\!100$ km for $O_2(^1\Delta)$."

The above is added to the text.

14. Line 5 on Page 15. "hemispheric wintertimes" Please check if it should be "hemispheric summertimes".

Response: This is the case only for $O_2(^1\Sigma)$ in Figure 8e, and is not the case for $O_2(^1\Delta)$ in Figure 8f. It is removed from the manuscript.

15. Lines 2–3 on page 16. Please explain in brief why you suspect that these abrupt changes are related to a change in the altitude sequence of the satellite measurements during that time.

Response: In our response we take into account the comments from both of the reviewers. The text now reads as follows:

"Figures 8c, 8d, and 8f, show a decrease in the altitude of the maximum $O_2(^1\Sigma)$, altitude of the maximum $O_2(^1\Delta)$, and $O_2(^1\Delta)$ h_{CA}, respectively between November 2010 and February 2011. This is due to a change in the limbs eq

16. Lines 12-13 on page 18. The larger intensities could begin earlier (than three weeks late

Response: We observe larger intensities on the next measurement day of SCIA-MACHY, i.e. 10 February 2009 which is about three weeks later than the previous measurement day of SCIAMACHY, i.e. 22 January 2009. To be more specific the following is inserted in the text:

Change in the manuscript: "Based on the temporal evolution of mesospheric temperature during the SSW event, Gao et al. (2011) divided the response in the mesosphere into three stages: prior to day 15 is considered the normal stage; days 15–22 correspond to the cooling stage; days post 22 correspond to the recovery stage. According to this, they reported that the O_2 nightglow brightness decreased by about a factor of 10 during the cooling stage and then increased by about a factor of 3 during the recovery stage relative to the normal stage. Figures 10a and 10b show the time series from 12/2008 to 04/2009 of the daytime $O_2(^1\Sigma)$ and $O_2(^1\Delta)$ daily mean VER, averaged from 60°N to 70°N. We observe that on the last day of the cooling stage, i.e., 22 January 2009, the daily zonal mean $O_2(^1\Sigma)$ and $O_2(^1\Delta)$ VERs show a reduced maximum intensity in the 82–87 km altitude range. We observe larger intensities on the next measurement day of SCIAMACHY about three weeks later, i.e., on 10 February 2009. This is expected from a decrease of atomic oxygen due to horizontal

C9

mixing and upwelling during the cooling stage, and then downward extension of the MLT region with large mixing ratio of O during the recovery stage of the SSW (Gao et al., 2011)."

17. Line 13 on page 18. It is better to add a citation for the sentence "as expected from enhanced mixing with oxygen—rich thermospheric air after the SSW".

Response: The manuscript is changed as follows:

"This is expected from downward extension of the MLT region with large mixing ratio of O during the recovery stage of the SSW (Gao et al., 2011)."

18. Line 14 on page 18. It is mentioned that "After the recovery phase, the $O_2(^1\Delta)$ signal is less prominent compared to $O_2(^1\Sigma)$ ". It is better of you can indicate the time when the recovery phase ended. In addition, it seems that the $O_2(^1\Delta)$ signal is less prominent compared to $O_2(^1\Sigma)$ after the recovery phase in Figure 10. However, the color bar in Figure 10a is different from that in Figure 10b. The $O_2(^1\Delta)$ emission after the recovery phase is evidently weaker than that during the recovery phase; however it is still further stronger than the $O_2(^1\Sigma)$ emission after the recovery phase. So is the SNR

Response: The manuscript is changed as "On the measurement day of SCIAMACHY after the recovery phase, i.e. 23 March 2009, the relative difference in the $O_2(^1\Delta)$ signal is less prominent compared to the relative difference in $O_2(^1\Sigma)$."

19. Line 2 on page 19. Similar to question 18.

Response: The manuscript is changed to read as follows "although the relative difference in $O_2(^1\Delta)$ band signal is less prominent compared to the relative difference in $O_2(^1\Sigma)$."

20. Figure 5. Please revise the label in Figure 5.

Response: It is revised from "Volume emission rate ($h\nu.cm^{-3}.s^{-1}$)" to "Volume Emission Rate (photon. $cm^{-3}.s^{-1}$)"

21. Figure 6. Please explain in brief the blank regions in Figure 6 in the text

Response: The blank regions correspond to the regions with SNR less than 1.

Change in the manuscript: "The blank regions represent areas with signal to noise ratios of less than one." is added to page 10, lines 16–17.

22. Figure 7. Please revise the label (yy-Mmm) in Figure 7.

Response: It has been changed as suggested.

23. Figure 8. Please revise the labels (yy-Mmm and yy-mm-dd) in Figure 8. In addition, it is better to change the ranges of the color bars to show some characteristics described in the text more clearly; especially for Figure 8d, it is better to decrease the upper limit value of the color bar.

Response: It is done as suggested.

24. Labels in Figure 9. Add unit of "Altitude". It is better to use $O_2(^1\Sigma)$ to replace $O_2(b)$ and use $O_2(^1\Delta)$ to replace $O_2(a)$. It is better to use (photons cm $^{-3}$ s $^{-1}$) to replace [photons cm $^{-3}$ s $^{-1}$].

It is applied in the Figure 9.

C11

References

- Bates, D.: Nightglow emissions from oxygen in the lower thermosphere, Planetary and Space Science, 40, 211 221, doi:10.1016/0032-0633(92)90059-W, 1992.
- Bovensmann, H., Burrows, J. P., Buchwitz, M., Frerick, J., Noël, S., Rozanov, V. V., Chance, K. V., and Goede, a. P. H.: SCIAMACHY: Mission Objectives and Measurement Modes, Journal of the Atmospheric Sciences, 56, 127–150, 1999.
- Gao, H., Xu, J., Ward, W., and Smith, A. K.: Temporal evolution of nightglow emission responses to SSW events observed by TIMED/SABER, Journal of Geophysical Research: Atmospheres, 116, 2011.
- Huestis, D. L.: Current Laboratory Experiments for Planetary Aeronomy, pp. 245–258, American Geophysical Union, doi:10.1029/130GM16, 2013.
- Kirillov, A.: The calculation of quenching rate coefficients of O2 Herzberg states in collisions with CO2, CO, N2, O2 molecules, Chemical Physics Letters, 592, 103 108, doi:10.1016/j. cplett.2013.12.009, 2014.
- Kirillov, A. S.: Model of vibrational level populations of Herzberg states of oxygen molecules at heights of the lower thermosphere and mesosphere, Geomagnetism and Aeronomy, 52, 242–247, doi:10.1134/S0016793212020077, 2012.
- Slanger, T. G. and Copeland, R. A.: Energetic Oxygen in the Upper Atmosphere and the Laboratory, Chemical Reviews, 103, 4731–4765, doi:10.1021/cr0205311, 2003.
- Smith, I. W. M.: The role of electronically excited states in recombination reactions, International Journal of Chemical Kinetics, 16, 423–443, doi:10.1002/kin.550160411, 1984.
- Stegman, J. and Murtagh, D.: The molecular oxygen band systems in the UV nightglow: measured and modelled, Planetary and Space Science, 39, 595–609, 1991.
- Wraight, P.: Association of atomic oxygen and airglow excitation mechanisms, Planetary and Space Science, 30, 251 259, doi:10.1016/0032-0633(82)90003-4, 1982.