### **Revised manuscript**

Dear Dr. Michel Van Roozendael,

we would like to take the opportunity to thank you for your efforts and that you accepted the editorship of our manuscript "GOME-2A retrievals of tropospheric  $NO_2$  in different spectral ranges - influence of penetration depth". Furthermore, we would like to thank you for the extension of the review period for the manuscript.

Please find enclosed a revised version of our manuscript where we implemented all comments by the referees. We revised the original manuscript according to their suggestions and provided also additional information as requested by the referees. In particular, we added an additional fitting window for the green spectral range and a section about stratospheric NO<sub>2</sub>. More details are provided on both the retrieval errors and temperature dependency of the NO<sub>2</sub> cross section. Supporting Figures have been compiled into an extensive supplement to our manuscript. Furthermore, we changed again the title of section from "2.6 Removing stratospheric NO<sub>2</sub>" to "2.6 Conversion to tropospheric NO<sub>2</sub> columns".

Below the author comments are provided that have already been uploaded to the AMT web page on 06 March 2018. We also provide here a version of the revised manuscript in which changes in comparison to the initial version are highlighted. We hope that with the submission of the author comments and the revision of the manuscript, our article will be accepted for publication in AMT.

Yours sincerely, Lisa Behrens (on behalf of the co-authors)

List of Attachments

- Author comments to Referee #1
- Author comments to Referee #2
- Revised manuscript with highlighted changes
- Revised manuscript
- Supplementary information

## Author reply to Referee #1

Lisa K. Behrens et al.

March 6, 2018

We thank Referee #1 for carefully reading our manuscript and for the helpful comments which will improve the quality of our manuscript. We will reply to the comments point by point.

Legend:

- referee comments

- authors comments

- changed text in the manuscript

The paper "Investigation of NO2 vertical distribution using two DOAS retrievals for GOME-2A measurements in the UV and vis spectral range" by Behrens et al. presents a retrieval of NO2 vertical columns in the UV spectral range and discusses how far the comparison to the "standard" retrieval in the blue spectral range provides information on the vertical distribution of NO2 within the troposphere. The manuscript is well written, clearly structured and overall convincing in its conclusions. It matches well in the scope of AMT.

Thank you very much for the positive comments.

#### I have two fundamental comments:

1. The goal of investigating the vertical distribution of NO2, as stated in the title, seems quite ambitious to me; one measured spectrum just provides information on the integrated column (SCD), without any information on the vertical distribution. It is not clear to me why the authors seem to have expected to get direct quantitative information on the vertical distribution by just adding one further piece of information (the SCD at a different wavelength). ...

We agree with the referee that the investigation of the vertical distribution is ambitious and the title can be confusing, because we are not able to give a real profile shape. Therefore, we will change the title to: "GOME-2A retrievals of tropospheric NO<sub>2</sub> in different spectral ranges - influence of penetration depth". Furthermore, we will add maps with retrieved top-altitudes of NO<sub>2</sub> layer height (Fig. 1) in the revised manuscript in Sect. 3.1. For the altitude retrieval simple box profiles are assumed for tropospheric  $NO_2$ . A seasonal dependency of the retrieved altitude can be clearly observed in the global maps.



Figure 1: Monthly mean top-altitudes retrieved from the ratio between the UV and blue spectral range. To retrieve the altitude, box profiles are assumed for the tropospheric  $NO_2$ . The light grey coloured values indicate values which are below the threshold defined for the ratio (see Fig. 8 manuscript; will be added in a revised version)

... So the aim of deriving vertical information in the troposphere seems to be rather a second step; as first step, I would have expected improved stratospheric estimates. The separation of stratospheric and tropospheric column is still a fundamental challenge for the retrieval of tropospheric NO2 VCDs. While it is often no problem over highly polluted regions, it is still a crucial prerequisite for accurate emission estimates of large-scale sources such as soil emissions. Thus I miss some discussion about how far a UV retrieval might improve stratospheric estimates. Empirical approaches such as the RSM and modifications (allowing for longitudinal gradients) as well as assimilation approaches use the measured total column for the estimation of the stratospheric fraction. This requires some kind of a-priori knowledge on how large the tropospheric contribution is expected to be. For a UV retrieval, the sensitivity to the tropospheric column is weaker. Thus, a modified RSM approach might benefit from UV measurements since it can include more measurements over weakly polluted regions. This topic should be discussed in the manuscript.

We agree with the referee that the derivation of the vertical column is a second step and that it would be nice to improve the separation between troposphere and stratosphere which was indeed our original intention. Unfortunately, it was not possible to improve the stratospheric estimation with a UV NO<sub>2</sub> retrieval as this suffers not only from noise but also from systematic biases. However as shown by the BAMFs (Fig. 6 (Fig. 2, manuscript, will be replaced in a revised version)), the UV spectral range has still a small sensitivity to the lower troposphere and therefore, the tropospheric contribution can also be observed in the UV spectral range. This tropospheric pollution can be observed in the total SCDs in both spectral ranges (Fig. 2, will be added in a revised version) as well as an additional fitting window in the green spectral region.

BAMFs show the largest differences between the UV and blue spectral range close to the ground ( $\sim$  factor 3). In higher altitudes of the troposphere, the differences are clearly reduced. As shown by Delon et al. (2008) and Stewart et al. (2008), also soil emissions (large-scale sources) are partly located in elevated layers which increases the visibility in the UV spectral range, and therefore, reduces the differences between the spectral ranges. A key issue here is the seasonal dependency between the spectral ranges (Fig. 3, Fig. 5; will be added in a revised version). In January, the  $vis_{blue}$  NO<sub>2</sub> values are higher than the UV  $NO_2$  values especially above polluted areas (Fig. 3) whereas in July, both spectral ranges are similar or the UV  $NO_2$  values are higher than the vis<sub>blue</sub>  $NO_2$  values. Similar offsets can be found between the blue and the green fitting window. Therefore by introducing an UV fitting window, improving the separation of stratosphere and troposphere is not possible and further investigations are needed for this special point. In our case, we are using tropospheric columns by subtracting the reference sector, and therefore, the differences are cancelled out to a large extent. We will add the following paragraph (between Sect. 2.4 and Sect. 2.5 in the old manuscript (2.6 in the new manuscript)):

#### **2.5 Stratospheric** $NO_2$

Figure 5 (will be added in a revised version) shows the latitudinal and seasonal dependency for the three  $NO_2$  fitting windows. The seasonal dependency clearly differs between the three fitting windows also over regions dominated by stratospheric  $NO_2$ , especially for the green wavelength range strong interferences are observable. Although the differences are smaller between the UV and blue spectral range, they are clearly observable, for example at the equator (Figure 5 b; will be added in a revised version). In northern hemispheric summer the UV  $NO_2$  values are higher than the  $NO_2$  values derived from the blue spectral range whereas in northern hemispheric winter the  $NO_2$ values from the blue spectral range are slightly higher. Therefore, it is currently not possible to improve the stratospheric  $NO_2$  retrieval by using different wavelength ranges.



Figure 2: Monthly mean total NO<sub>2</sub> SCD for (a, b, c) January and (d, e, f) July 2008 for the (a, d) UV, (b, e) blue, and (c, f) green spectral range. (will be added in a revised version, supplement)



Figure 3: Relative difference between monthly mean total  $NO_2$  SCDs in the blue and UV spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no  $NO_2$  values available. (will be added in a revised version, supplement)



Figure 4: Relative difference between monthly mean total NO<sub>2</sub> SCDs for the blue and green spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no NO<sub>2</sub> values available. (will be added in a revised version, supplement)



Figure 5: Time series of NO<sub>2</sub> total VCDs of 2008 for 30° S (averaged: 28° S – 32° S), 0° N (averaged: 2° S – 2° N), and 30° N (averaged: 28° N – 32° N) above the area of the reference sector for cloud free pixels. (will be added in a revised version)

2. The authors use the established blue spectral range plus a fit window in the UV. But what about the green spectral range? As shown in Fig. 1, the NO2 cross section still shows absorption bands above 550 nm. This would be another complementary piece of information, even more sensitive towards the lower troposphere than the standard retrieval. Of course the noise will be higher, and the choice of the fitting window and the water vapour cross section might be challenging. I understand that such a "green" retrieval would require major data processing; but still the authors should discuss the potential of adding additional wavelengths for the goal of assessing profile information, and I would appreciate if they could provide some sensitivity studies.

We calculated BAMF for three different wavelength 352 nm, 438 nm, and 515 nm. As can be seen in Fig. 6, the sensitivity for larger wavelength increases towards the surface compared to the standard fitting window. Therefore, in principle further information about the vertical profile can be gained from this spectral range as suggested by the reviewer.



Figure 6: BAMF for three wavelengths and the ratio of BAMFs to the mean wavelength of the standard fitting window calculated with the radiative transfer model SCIATRAN. BAMFs converge at higher altitude (not shown). The BAMFs are calculated for 352 nm (UV), 438 nm (vis), and 515 nm at (a, b)  $30^{\circ}$  (c, d)  $50^{\circ}$  and at (e, f)  $70^{\circ}$  solar zenith angle (SZA). A surface spectral reflectance of soil is assumed. (Fig. 2, manuscript, will be replaced in a revised version)

An additional fitting window for the green spectral range was already presented in Richter and Burrows (2000). They used a fitting window from 490-540 nm for the GOME instrument. In combination with a UV fitting window, they tried to separate stratospheric and tropospheric NO<sub>2</sub> VCDs. They concluded that for GOME measurements it is not possible to retrieve global maps of NO<sub>2</sub> VCDs, because of the smaller signal to noise ratio and strong interferences with other absorbers which leads to systematic errors.

Here, we found similar results for GOME-2A. Large areas with liquid water absorption can be clearly identified. In our fit, we do not use a cross section for liquid water, because this introduces further interferences over land where no liquid water absorption is expected. Furthermore, in the current fitting window, areas with clear water (see for example north of Australia, Fig. 8 (b); will be added in a revised version) and interferences with the surface in Africa can be observed (Fig. 7, 8; will be added in a revised version). These effects can also be observed in the ratio between the blue and green spectral range (Fig. 10; will be added in a revised version).

Compared with the fitting windows in the UV and blue spectral range, the fitting windows shows a similar distribution of total VCDs above the Pacific Ocean as the UV spectral range (Fig. 11 (Fig. 4, manuscript, will be replaced in a revised version)) with a small shift to negative values. Nevertheless, the NO<sub>2</sub> retrieval works quite well above highly polluted areas. Therefore, we will include a case study for China for the green spectral range and an additional section about the green fitting window in our revised manuscript.



Figure 7: Monthly mean tropospheric  $NO_2$  SCD for (a, c) January and (b, d) July 2008 for the (a, b) green and (c, d) blue spectral range. (will be added in a revised version, supplement)



Figure 8: Absolute difference between monthly mean tropospheric NO<sub>2</sub> SCDs in the green and blue spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no NO<sub>2</sub> values available. (will be added in a revised version, supplement)



Figure 9: Relative difference between monthly mean tropospheric NO<sub>2</sub> SCDs in the green and blue spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no NO<sub>2</sub> values available. Light grey coloured values indicate values where the vis NO<sub>2</sub> is close to zero, which have been filtered out. (will be added in a revised version, supplement)



Figure 10: Ratio between monthly mean tropospheric SCDs of NO<sub>2</sub> in the blue and green spectral range. (a) January 2008 and (b) July 2008. Dark gray shaded area: no NO<sub>2</sub> values available. Light grey coloured values indicate values where the vis NO<sub>2</sub> is close to zero, which have been filtered out. (will be added in a revised version, supplement)



Figure 11: Distribution of total NO<sub>2</sub> VCDs over a clean region (equatorial Pacific:  $5^{\circ} S - 5^{\circ} N$  and  $150^{\circ} E - 210^{\circ} E$ ) for the UV and vis spectral range for January 2008. Curves are normalised to unit area and centered on zero. (Fig. 4, manuscript, will be replaced in a revised version)



Figure 12: Time series 2007-2015 of tropospheric NO<sub>2</sub> SCDs for the UV (red line) and blue (blue line) spectral range for different regions. Additionally for China, the green (green line) spectral range is shown. Note the different y-axes. The vertical lines indicate January of the individual years. (a-c) biomass burning regions and (d-f) regions with high anthropogenic air pollution. (Fig. 9, manuscript, will be replaced in a revised version)



Figure 13: Monthly mean tropospheric  $NO_2$  VCD for (a, c) January and (b, d) July 2008 for the (a, b) green and (c, d) blue spectral range. (will be added in a revised version, supplement)



Figure 14: Absolute difference between monthly mean tropospheric NO<sub>2</sub> VCDs in the green and blue spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no NO<sub>2</sub> values available. (will be added in a revised version, supplement)



Figure 15: Relative difference between monthly mean tropospheric NO<sub>2</sub> VCDs in the green and blue spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no NO<sub>2</sub> values available. Light grey coloured values indicate values where the vis NO<sub>2</sub> is close to zero, which have been filtered out. (will be added in a revised version, supplement)



Figure 16: Time series 2007-2015 of tropospheric NO<sub>2</sub> VCDs for the UV (red line) and blue (blue line) spectral range as well as the TM5-model VCDs (gray line) for different regions. Additionally for China, the green (green line) spectral range is shown. Note the different y-axes, the same as in Fig. 12. The vertical lines indicate January of the individual years. (a-c) biomass burning regions and (d-f) regions with anthropogenic air pollution. (Fig. 13, manuscript, will be replaced in a revised version)

#### Minor comments:

Page 3 line 32: "can provide information about the accuracy": I would suggest a different formulation such as "allows to infere the validity of a-priori profiles".

Changed as suggested.

Page 4 line 12: should be "summary and conclusions".

Changed as suggested.

Section 2.1: Please add Munro et al. (2016) also as reference for GOME-2 A.

Done.

Page 4 line 30: add "before launch" or similar.

Done.

Section 2.2: The details of the cross sections and respective references might be shifted to Table 1.

Done — we changed the paragraph as follows:

For this study, we developed a NO<sub>2</sub> DOAS retrieval for the GOME-2A instrument in the UV which uses a fitting window between 342 and 361.5 nm, and a polynomial degree of four. In this retrieval, one NO<sub>2</sub> cross section measured with the GOME-2A instrument before launch as well as two O<sub>3</sub> cross sections are used with an additional I<sub>0</sub> correction of  $10^{20}$  molec cm<sup>-2</sup> (Platt et al., 1997; Richter, 1997). Additionally, cross sections for O<sub>4</sub>, BrO, HCHO, the Ring effect and the instrumental cross section Zeta are included in the fitting procedure (see Tab. 1).

Table 1: Fit settings for the  $NO_2$  retrievals in the UV and the vis spectral range.

|                | UV spectral range                     | vis spectral range                                |
|----------------|---------------------------------------|---|
| fitting window | $342 - 361.5 \mathrm{nm}$             | $425 - 450 \mathrm{nm}$                           |
| polyn. degree  | 4                                     | 4   |
| cross sections |                                       |   |
| $NO_2$         | 223 K; Gür et al. (2005)              | 243 K; Gür et al. (2005)                          |
| $O_3$          | $223\mathrm{K}$ and $243\mathrm{K}$ ; | $223\mathrm{K};\mathrm{G\ddot{u}r}$ et al. (2005) |
|                | Serdyuchenko et al. (2014)            |   |
| $O_4$          | Greenblatt et al. (1990)              | Greenblatt et al. (1990)                          |
| $H_2O$         | _                                     | Rothman et al. (2010)                             |
| BrO            | Wilmouth et al. (1999)                | _   |
| HCHO           | Meller and Moortgat (2000)            | _   |
| Ring           | calculated with SCIATRAN,             | Vountas et al. (1998)                             |
|                | Rozanov et al. (2014)                 |   |
| instr. func.   | Zeta; EUMETSAT (2011)                 | _   |

Page 5 line 6: Which were the criteria for identifying "best results"?

Best results means smallest RMS — we changed the sentence as follows: Among many different wavelength windows we tested, the selected window from 342-361.5 nm provided the smallest root mean squared error (RMS).

Page 5 line 11: For a focus on stratospheric patterns, this would probably be crucial.

We agree with the reviewer but as shown above a stratospheric correction is also difficult without the introduced offset. We will add an additional section about the stratospheric  $NO_2$ .

Page 5 line 12: So daily Earth is the alternative, but what is the default?

Yes, it is an alternative. For  $NO_2$  the daily solar reference is default. Daily Earth spectra are often used in the UV and for weak absorbers as CHOCHO and HCHO.

Page 5 line 20: Details and references for O4 (vis) and H2O are missing (might also be added to Table 1).

Done — see Table 1.

Page 6 line 21: What does "upper atmospheric" mean?

"upper atmospheric" is removed.

Page 6 line 30: What is meant by "this correction method"? The RSM? Then write it. Or any correction of the RSM close to the polar vortex? Then provide further details.

RSM is meant — we changed the sentence as follows: For the reference sector method, the same cloud screening as for the data selection is used (see Sect. 2.7).

Page 7 line 2: "This correction"  $\rightarrow$  RSM?

Changed as suggested.

Page 7 line 5: "no negative values are expected": if the model is correct!

Done — we changed the sentence as follows: Therefore, no significantly negative values are expected for the VCDs assuming the model is correct.

Page 7 line 9: "... using a linear approach"  $\rightarrow$  "the NO2 VCDs are scaled by a correction factor linear in T in order to correct for the temperature dependency of the NO2 cross section, as suggested in Boersma..."

Changed as suggested.

Page 12 line 24: avoid "believe"; you have provided several arguments for this conclusion.

Done.

Additionally as suggested by Referee #2, we will change four main points in the revised manuscript:

- 1. We will discuss errors in more detail. Therefore, a table with error sources is added to the revised manuscript.
- 2. We will include a discussion about the temperature dependency of the NO<sub>2</sub> SCDs which affects the UV and vis NO<sub>2</sub> retrievals different.
- 3. The figure with the absolute differences will be replaced by figures with relative differences which makes the differences between the SCDs and the VCDs for the two spectral ranges more clear.
- 4. We will discuss in more detail how CTM profiles should be changed to reconcile VCDs of the different spectral ranges as well as the model VCDs. Here, our observations suggest that compared to the real NO<sub>2</sub> profiles, in the TM5 model the NO<sub>2</sub> is higher in the atmosphere with lower surface concentrations.

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## Author reply to Referee #2

Lisa K. Behrens et al.

March 6, 2018

We thank Referee #2 for carefully reading our manuscript and for the helpful comments which will improve the quality of our manuscript. We will reply to the comments point by point.

Legend:

- referee comments

- authors comments

- changed text in the manuscript

Behrens et al. present a study showing the relationship between UV and VIS tropospheric NO2 VCDs and how it can relate to information on vertical distribution of NO2 in the troposphere. Although the paper is well written, it appears to be rather deceiving. I find the paper too qualitative and the reader could expect an attempt to effectively derive some information on the vertical distribution of NO2 from the combination of UV and vis NO2 measurements. The authors provide a number of possible reasons for the differences between UV and vis NO2 VCDs (a-priori profiles, effect of clouds/aerosols, etc) that are all plausible (and speculative) but there is no clear way forward. They almost conclude on the current impossibility to derive profile information. A weak point is that it is difficult to separate possible errors in the retrievals (in the UV spectral fits) from real effects.

We agree with the referee that the title may suggest that combined UV and vis NO<sub>2</sub> measurements would be provided information about the NO<sub>2</sub> profiles in this manuscript which may lead to confusion. Therefore, we will change the title to: "GOME-2A retrievals of tropospheric NO<sub>2</sub> in different spectral ranges - influence of penetration depth". Furthermore, we will add maps with retrieved top-altitudes of NO<sub>2</sub> layer height (Fig. 1) in the revised manuscript in Sect. 3.1. For the altitude retrieval simple box profiles are assumed for tropospheric NO<sub>2</sub>. A seasonal dependency of the retrieved altitude can be clearly observed in the global maps.



Figure 1: Monthly mean top-altitudes retrieved from the ratio between the UV and blue spectral range. To retrieve the altitude, box profiles are assumed for the tropospheric  $NO_2$ . The light grey coloured values indicate values which are below the threshold defined for the ratio (see Fig. 8 manuscript; will be added in a revised version)

Furthermore, we agree that it is difficult to separate possible errors in the retrieval from real effects. With the help of the referee's comments we will make the manuscript more quantitative, by including an error discussion and the discussion of the temperature dependency of SCDs.

# To be published in AMT, the following points need to be addressed:

-a comprehensive error analysis on the UV retrievals needs to be undertaken.

Done — we will include the following table in Sect. 2.7 (2.8 in the revised manuscript):

|                       | 1  |  |
|-----------------------|--|--|
|                       | errors UV  | errors vis   |
| total SCDs            | $1.8 \times 10^{15} \mathrm{molec}\mathrm{cm}^{-2}$                      | $0.6 \times 10^{15} \mathrm{molec}\mathrm{cm}^{-2}$                  |
|                       | (calculated above the Pacific Ocean, see Fig. 4 (manuscript))            |  |
| strato. VCDs          | $7.4 \times 10^{14} \mathrm{molec}\mathrm{cm}^{-2}$                      | $2.1 \times 10^{14} \mathrm{molec}\mathrm{cm}^{-2}$                  |
|                       | (see Fig. 4 (manuscript))  |  |
| SSR                   | $40\%$ at $320\mathrm{nm}$   | $5\%$ at $500\mathrm{nm}$  |
|                       | (Kleipool et al., 2008)  | (Kleipool et al., 2008)  |
| 0.05, increase $0.01$ | BAMF increases 9%  | BAMF increases $11\%$  |
|                       | (338  nm, Lorente et al., 2017)  | $(440 \mathrm{nm}, \mathrm{Lorente} \mathrm{et} \mathrm{al.}, 2017)$ |
| $\operatorname{AMF}$  |  |  |
| cloud fraction        | 0-30%, Boersma et al., 2004  |  |
| cloud height          | $<\!10\%,$ Boersma et al., 2004  |  |
| aerosols              | not included in our calculations $\rightarrow$ 15%, Boersma et al., 2004 |  |
| profile shape         | $<\!15\%$ (regions with little NO_2: $>\!50\%$ ), Boersma et al., 2004   |  |

Table 1: Errors for the UV and vis spectral range.

-section 2.2: the effect of T on Uv+vis NO2 retrievals is not well discussed. NO2 cross sections are varying with T but only one T cross-section is included in the fits (both for UV and VIS DOAS fits). What is the impact on the results and conclusions of this study?

The NO<sub>2</sub> cross sections have a temperature dependency, which differs for the UV and vis spectral range as shown in Fig. 2 and Fig. 3 exemplarily for the two spectral ranges. This temperature dependency can further increase the differences between the two spectral ranges. Figure 4 and Fig. 5 show the altitude dependent sensitivity of the scaling coefficients for the UV and vis spectral ranges. The temperature dependency influences the tropospheric as well as the stratospheric NO<sub>2</sub> measurements. For the stratospheric NO<sub>2</sub> measurements, the sensitivity of the UV NO<sub>2</sub> spectral range is approximately 10% higher than for the vis spectral range. Close to the surface, the difference in temperature sensitivity is up to 10% stronger in the vis spectral

range. For VCDs, the temperature dependency of the  $NO_2$  cross section is scaled by a linear correction factor for both spectral ranges, as suggested in Boersma et al., 2004. In the revised manuscript, we will include this issue in our discussion, but we will move this point to Sect. 2.4 (p. 5, l. 27, manuscript):

[...] Furthermore, the measurement sensitivity for  $NO_2$  decreases towards the surface. This can be clearly observed in the BAMF (Fig. 2, manuscript). This effect is enhanced by the temperature dependency of the  $NO_2$  cross section. The temperature dependency influences the tropospheric as well as the stratospheric  $NO_2$ measurements (see Fig. 4 and Fig. 5). For the stratospheric  $NO_2$ retrieval, the sensitivity in the UV spectral range is up to 10% higher than for the vis spectral range. Close to the surface, the temperature sensitivity is up to 10% stronger in the vis spectral range compared to the UV spectral range. The temperature vertical sensitivity introduces a seasonal and a latitudinal dependency. This effect is stronger in the tropics than for higher / lower latitudes and in the mid-latitudes it is more pronounced in summer and less in winter.

[In combination with ...]



Figure 2:  $NO_2$  cross section and their temperature dependency for the UV spectral range. The scaling coefficient of the temperature dependency is calculated. (will be added in a revised manuscript, supplement)



Figure 3:  $NO_2$  cross section and their temperature dependency for the vis spectral range. The scaling coefficient of the temperature dependency is calculated. (will be added in a revised manuscript, supplement)



Figure 4: Altitude dependency of the  $NO_2$  scaling coefficient for China. The profiles are calculated for model data simulated with the TM5 model for 2008. (will be added in a revised manuscript, supplement)



Figure 5: Altitude dependency of the  $NO_2$  scaling coefficient for ASE. The profiles are calculated for model data simulated with the TM5 model for 2008. (will be added in a revised manuscript, supplement)

-section 2.6: what is the uncertainty due to errors in SSR?

This issue has been addressed in the new Tab. 1.

-section 3.3, p14, l20: it is written that VCD differences are small but it is hard to judge as only absolute values for the differences are shown. It would be better to incorporate relative differences as well as proper error calculation (see comment above).

In response to the suggestion of the reviewer, we will show the relative differences in the revised manuscript (see Fig. 6 and Fig. 7). The figures with the absolute differences will be moved to the supplement.



Figure 6: Relative difference between monthly mean tropospheric NO<sub>2</sub> SCDs in the vis and UV spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no NO<sub>2</sub> values available. Light grey coloured values indicate values where the vis NO<sub>2</sub> is close to zero, which have been filtered out.



Figure 7: Relative difference between monthly mean tropospheric  $NO_2$  VCDs in the UV and vis spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no  $NO_2$  values available. Light grey coloured values indicate values which are filter out. The same filter as for Fig. 6 is used.

-p15: it is not clear how the CTM profiles should be changed to reconcile the vis, uv and modeled VCDs.



Figure 8: SCD, VCD, and AMF for different NO<sub>2</sub> profiles, calculated for one scenario in January over China with a SZA of  $66^{\circ}$ . Blue: true profile, the SCDs are calculated for this profile. Other color: changed input profiles for AMF calculation leading to changes in the retrieved VCDs.

As shown in Fig. 8, the profile shape has an influence on the retrieved VCDs, if the assumed and true profile shape do not agree. Exemplarily for one scenario over China in January with a SZA of 66°, the VCDs are calculated. To avoid possible measurement errors in the retrieved SCDs, the SCDs are also calculated with the radiative transfer model SCIATRAN. The "true"  $NO_2$  profile is the blue curve in Fig. 8a. Additionally, the  $NO_2$  profiles were changed (colour coded) and AMF (Fig. 8b) were calculated for these changed profiles which introduce differences in the retrieved VCDs (Fig. 8 d). The two profiles with lower  $NO_2$  values than the true  $NO_2$  profile close to the surface and with a slightly higher PBL and a smoother decrease of NO<sub>2</sub> values (red and green) for lead to a similar situation as observed in our study. The observed VCDs are higher for both spectral ranges than model VCDs and the VCDs for the UV spectral range are lower than for the vis spectral range. For a scenario with a much higher PBL, a constant mixing in the PBL (yellow), and a sharp decrease above, the model values are higher than the retrieved values which we have not observed in our study. For lower PBL with higher  $NO_2$  values (cyan), the differences between the spectral ranges and the model VCDs are less pronounced. Therefore, our observations suggest that compared to the real  $NO_2$  profiles, in the TM5 model the  $NO_2$  is higher in the atmosphere with lower surface concentration values.

We will include Fig. 8 in the supplement and discuss this issue in more detail in the revised manuscript.

-p16, l25-26: the reading nearly suggests that it could be better to use UV retrievals but it is quite unlikely given the larger uncertainties.

We will change the sentence as follows:

For ASE and ANE during biomass burning season, the UV  $NO_2$  VCDs and the model VCDs agree quite well, while the vis  $NO_2$  values are higher. It should however be kept in mind that in the vis retrieval, both the uncertainties and the contribution of the a priori are lower, making these results more reliable.

-p18, l35: It is stated that 'concurrent measurement of the same air mass from different view geometries could yield insight on aerosols and vertical distribution' but these measurements do not exist.

We gave an example for developments for possible space-borne measurements in the future. Furthermore, there are already concurrent measurements of the same air mass. Therefore, there might be already the possibility to retrieve these kind of information from GOME-2 measurements using the instruments on the MetOp-A and MetOp-B. For example on the 01 January 2013 (before reduction of spatial resolution of GOME-2A) for a pixel of South Korea, the GOME-2A instrument has an overpass time of 2:04 UTC with a SZA of 61.86°, a LOS of -43.26°, and a RAA of -46.66°, whereas GOME-2B has an overpass time of 1:17 with a SZA of 66.84°, a LOS of 18.41°, and a RAA 137.35°.

We will change the sentence as follows:

Future developments in space-based observation of the Earth's atmosphere could however increase the capability to retrieve vertical  $NO_2$  distribution. For example, concurrent measurements of the same air mass from different viewing geometries could yield insight on aerosol types and vertical profiles which might be also possible for GOME-2A and GOME-2B observations.

#### Minor comments:

-p2, l34-35: I find misleading that the examples on past studies on vertical profiling are mostly unrelated to NO2 retrievals as given in the present paper (which considers optically thin atmosphere as for NO2).

The reviewer is right that we gave examples which are unrelated to  $NO_2$ . However, a similar method has so far not been used before for an optically thin gas. Therefore, we used examples for ozone which is an optically thick gas. We will point out this differences more clearly and change the paragraph as follows in the revised manuscript:

[...] In consideration of this fact, knowledge of the vertical distribution of  $NO_2$  can be gained by combining measurements at different wavelengths. The idea of using the penetration depth in the UV to determine vertical profiles of ozone was first proposed by Singer and Wentworth (1957). The use of the temperature dependence of the Huggins absorption bands coupled with penetration depth was proposed to retrieve information about the vertical profile of ozone in the troposphere (Chance et al., 1997). Here, will use a similar method for the optically thin trace gas  $NO_2$ .

#### -p3, 113: ...path within the NO2 layer relative to the vertical path.

Done.

#### -p3, l15: environmental effects is too vague.

Done — we will change the sentence as follows in the revised manuscript: AMFs are calculated by radiative transfer models, which take into account the viewing geometry and environmental effects, e.g., scattering processes in the Earth's atmosphere, SSR and the vertical distribution of trace gases (Platt and Stutz, 2008).

-throughout the manuscript, the author often use the word 'visibility' to express the idea that NO2 signal is more clear 'visible' in some spectral range. It is ambiguous as it might be interpreted by 'visible wavelength'.

In revised the manuscript, we will replace the word 'visible'. Then, it will be only used for the visible spectral range. However we think, it is not necessary to replace the word 'visibility'.

#### -p3: the example of bAMFs should be presented here.

We will move Fig. 2 to the introduction as suggested and refer to this figure in the later chapters in the revised manuscript.

# -p3: the effect of BRF on NO2 retrievals is not developed enough as possible reason between UV and vis

Done — we will change the paragraph as follows in the revised manuscript: .... For wavelengths in the UV, the BAMF in layers close to the ground is considerably smaller than for the vis spectral range (Fig. 1, revised manuscript; Fig. 2 manuscript). This effect is even less pronounced for longer wavelengths. In general, BAMFs for longer wavelengths have a smaller dependency on altitude compared with BAMFs for shorter wavelengths (Burrows et al., 2011). The altitude of highest sensitivity further depends on the solar zenith angle (SZA). For increasing SZAs, the altitude of highest sensitivity moves upwards to the stratosphere. Furthermore, the surface spectral reflectance (SSR) depends on the wavelength, and therefore, the SCDs are influenced by the SSR (Burrows et al., 2011). Generally, for the UV and vis spectral range the SSR is quite low between 2 and 30% depending on the surface type except for snow or ice (Burrows et al., 2011). For these kind of surface types, the SSR is lower in the UV than in the vis spectral range. For smaller SSR, the UV shows a stronger decrease and for larger SSR (e.g. snow) the UV shows stronger increase towards the surface thus the UV SCDs decrease or increase compared to the vis SCDs. Therefore, the SSR can strengthen the effect of the Rayleigh scattering which can be further increased by higher SZAs. Additionally, aerosols influence the measurements and also the visibility of  $NO_2$  is influenced by the presence of aerosols (Burrows et al., 2011). Depending on the type and the optical thickness of aerosols the influences on the measurement differs.

-p3, l34: sofar  $\rightarrow$  so far.

Done.

-p5, l 15: "strong absorption lines" -> "strong differential absorption lines"

Changed as suggested.

-section 2.5: for the SCDs stratospheric correction, are the averages performed for both UV and VIS data separately?

Yes, they are calculated separately for both spectral ranges — we will change the sentence as follows:

For the SCDs, we use the "reference sector method" (Richter and Burrows, 2002; Martin et al., 2002) for both spectral ranges separately, in which a monthly average of SCDs measured over a presumably clean area above the Pacific ( $180^{\circ}$  E to  $210^{\circ}$  E) is subtracted from all measurements per latitude band.

-p10, l16-17: this is a bit contradictory. If it is below the detection, then how meaningful is the 0.6 SCD ratio?

Yes, that is true — we will change the sentence as follows: Finally, Fig. 8 (manuscript) shows SCD ratios over the well known shipping lane leading from South India to the Strait of Malacca.

-p13, l8, first word: India–>China?

No, the paper from Hilboll et al. (2017) is about air pollution in India.

-p13, l18: SZA is lower-> SZA is higher?

Yes — Done.

Additionally as suggested by Referee #1, we will change two main points in the revised manuscript:

- 1. We will add discussion about stratospheric  $NO_2$  and show that it is so far not possible to improve the stratospheric  $NO_2$  retrieval by using different wavelength ranges.
- 2. We will discuss the possibility of an additional fitting window in the green spectral range. The fitting window in the green spectral range

has a higher sensitivity to the lower troposphere. However, in the green spectral range interferences with the surface are clearly visible and therefore in the revised manuscript, we will include only a case study for China.

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## **Investigation** GOME-2A retrievals of tropospheric NO<sub>2</sub> vertical distribution using two DOAS retrievals for **GOME-2A** measurements in the UV and vis different spectral range - influence of penetration depth

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Abstract. In this study, we present a novel  $NO_2$  DOAS retrieval in the ultraviolet (UV) spectral range for satellite observations from the Global Ozone Monitoring Instrument 2 on board EUMETSAT's MetOp-A (GOME-2A) satellite. We compare the results to those from an established NO<sub>2</sub> retrieval in the visible (vis) spectral range from the same instrument and infer information about investigate, how differences between the two are linked to the  $NO_2$ 

vertical profile shape in the troposphere. 5

As expected, radiative transfer calculations for satellite geometries show that the sensitivity close to the ground is higher in the vis than in the UV spectral range. Consequently, NO<sub>2</sub> slant column densities (SCDs) in the vis are usually higher than in the UV, if the NO<sub>2</sub> is close to the surface. Therefore, these differences in NO<sub>2</sub> SCDs between the two spectral ranges contain information on the vertical distribution of  $NO_2$  in the troposphere. We combine these

results with radiative transfer calculations and simulated NO<sub>2</sub> fields from the  $\frac{TM5}{TM5}$  chemistry transport 10 model to evaluate the simulated  $NO_2$  vertical distribution.

We investigate regions representative for of both anthropogenic and biomass burning  $NO_2$  pollution. Anthropogenic air pollution is mostly located in the boundary layer close to the surface, which is reflected by the large differences between UV and vis SCDs of  $\sim 60$  %. Biomass burning NO<sub>2</sub> in contrast is often uplifted into elevated layers above

the boundary layer. This is best seen in tropical Africa south of the equator, where the biomass burning  $NO_2$  is 15 well observed in the UV, and the <u>SCDs</u> difference between the two spectral ranges is only  $\sim 36$  %. In tropical Africa north of the equator, however, the biomass burning  $NO_2$  is located closer to the ground, reducing its visibility in the UV.

While not enabling a full retrieval of the vertical  $NO_2$  profile shape in the troposphere, our results can help to constrain the vertical profile of  $NO_2$  in the lower troposphere and, when analyzed analyzed together with simulated 20  $NO_2$  fields, can help to better interpret the model output.
# 1 Introduction

Nitrogen dioxide  $(NO_2)$  is an important indicator for natural phenomena and anthropogenic air pollution, as it is produced in the troposphere by, e.g., biomass burning or combustion of fossil fuels (Lee et al., 1997). Additionally,  $NO_2$  is produced by lightning (Lee et al., 1997; Beirle et al., 2004) and microbiological soil emissions activity in soils

- 5 (Williams et al., 1992; Bertram et al., 2005). The relevance of the individual NO<sub>2</sub> sources depends on the region of the Earth (van der A et al., 2008). Biomass burning is important in equatorial regions like central Africa, whereas anthropogenic air pollution is emissions are mostly important in industrialised areas the industrialised mid-latitudes like China or Europe. The distribution of NO<sub>2</sub> is of major interest because it is harmful to human health, adds to local radiative forcing, catalyses surface ozone production during summer smog, and causes acid rain (Finlayson-Pitts
- 10 and Pitts, 1999).

Using spectrometers, solar radiation scattered upwelling at the top of in the Earth's atmosphere can be measured and the amount of trace gases inverted mathematically from the depth of molecular absorption bands. Such measurements have been made from satellite (e.g., Burrows et al., 1999), airborne (e.g., Heue et al., 2005; Wang et al., 2005), and ground-based (Noxon, 1975) platforms.

- 15 Since hyperspectral satellite observations began in the mid-1990s (Burrows et al., 1999), the horizontal distribution of the column amounts of many trace gases is well known, e.g., for NO<sub>2</sub>. Using sun-synchronous orbits, a nearly daily global coverage at similar local times can be obtained. The global coverage as well as the spatial resolution of the measurements depends on the instrument (see for example: Callies et al., 2000). However, there is only limited knowledge of the vertical distribution from satellite data. In contrast, ground based Compared to satellite
- 20 <u>observations, ground-based</u> Multi AXis Differential Optical Absorption Spectroscopy (MAX-DOAS) measurements have high temporal resolution and can provide vertical profiles of trace gases up to several kilometres altitude (Wittrock et al., 2004), but can only measure at one particular location and are thus limited in resolving horizontal gradients.
- Irrespective of the instrument's viewing geometry, the measured spectra can be analysed using the Differential 25 Optical Absorption Spectroscopy (DOAS; Platt and Stutz, 2008; Burrows et al., 2011) technique which is a well established method based on Lambert Beer's law discribing describing the spectral reduction of the initial intensity of light due to absorption. The main result is the integrated concentration of trace gases  $\rho(s)$  along the effective light path s which is called the total slant column density (SCD; Platt and Stutz, 2008; Burrows et al., 2011):

$$SCD = \int \rho(s)ds.$$
 (1)

30 As Rayleigh scattering in the atmosphere depends strongly on wavelength ( $\lambda$ ; the scattering cross-section is proportional to  $\lambda^{-4}$ ), sun light sunlight having longer wavelengths has a larger penetration depth in can penetrate deeper into the atmosphere compared to shorter wavelengths. This results in a wavelength-dependency of the SCDs (Burrows et al., 2011). Consequently, if NO<sub>2</sub> is located close to the ground, SCDs retrieved from satellite measurements at shorter wavelengths are in theory smaller than SCDs retrieved from longer wavelengths. In consideration of this fact, knowledge of the vertical distribution of  $NO_2$  can be gained by combining measurements at different wavelengths. The idea of using the penetration depth in the UV to determine vertical profiles of ozone-trace gas vertical profiles was first proposed for ozone by Singer and Wentworth (1957). The use of the temperature dependence of the Huggins absorption bands coupled with penetration depth was proposed to retrieve information about the vertical

5 profile of ozone in the troposphere (Chance et al., 1997). Here, will use a similar method for the optically thin gas  $NO_{2}$ .

In the presence of clouds, the described behaviour changes. For cloudy scenes only smaller differences are expected. Due to the clouds the sensitivity towards the surface is decreased, and therefore, the  $NO_2$  below the clouds is partly invisible for to the instrument. For the  $NO_2$  above the clouds, the sensitivity is similar or partly increased for the UV spectral range (Burrows et al., 2011). In previous studies, clouds at different altitudes are have been used to

10 UV

obtain information about  $NO_2$  profiles (Choi et al., 2014; Belmonte Rivas et al., 2015).

The vertical column densities (VCDs) are the integral of the trace gas concentration from the surface to the top of the atmosphere along the altitude z (Platt and Stutz, 2008; Burrows et al., 2011):

$$VCD = \int \rho(z)dz.$$
 (2)

15 They can be calculated using air mass factors (AMFs), which are defined as the ratio of SCDs and VCDs (Platt and Stutz, 2008; Burrows et al., 2011), and are an indicator of the measurement sensitivity or the length of the light path within the NO<sub>2</sub> layer relative to the vertical path:

$$AMF(\lambda) = \frac{SCD(\lambda)}{VCD}.$$
(3)

AMFs are calculated by radiative transfer models, which take into account the viewing geometry and environmen-20 tal effects (Platt and Stutz, 2008). The sensitivity of the measurement to an absorber varies with altitude; this is expressed by the so-called box air mass factor (BAMF; Burrows et al., 2011), which is defined as  $BAMF_i = SCD_i/VCD_i$  for an atmospheric layer *i*. For wavelengths in the UV, the BAMF in layers close to the ground is considerably smaller than for vis wavelengths. the vis spectral range (Fig. 1). This reduction in sensitivity to lower atmospheric layers is further reduced for longer wavelengths. Figure 1 shows an example for surface type soil which

25 results in a stronger altitude dependency compared to other surface types e.g., vegetation. In general, BAMFs for longer wavelengths have a smaller dependency on altitude compared with BAMFs for shorter wavelengths (Burrows et al., 2011).

Furthermore, the surface spectral reflectance (SSR) depends on the wavelength, and therefore, the SCDs are wavelength dependence of the SCDs is influenced by the SSR (Burrows et al., 2011). Generally, for the UV and

30 vis spectral range the SSR is quite low between 2 and 30% depending on the surface type except for snow or ice (Burrows et al., 2011). For this kind of surface types, the SSR is lower in the UV than in the vis spectral range. For smaller SSR, the UV sensitivity shows a stronger decrease and for larger SSR the UV shows (e.g. snow) a stronger increase towards the surface. Therefore, the SSR can strengthen the effect of the Rayleigh scattering.

Generally, the SSR is for the UV smaller than for the vis spectral range, resulting in corresponding SCD decreases or increases compared to the vis SCDs. Furthermore, the altitude of highest sensitivity depends on the SZA. For increasing SZAs, the altitude of highest sensitivity moves upwards to the stratosphere. Additionally, aerosols influence the measurements and also the visibility of  $NO_2$  is influenced by the presence of aerosols (Burrows et al., 2011). Depending on the type and the optical thickness of aerosols the influences on the measurement differs.

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**Figure 1.** BAMF for UV (red line), green (green line), and blue (blue line) spectral range and the ratio of UV and green BAMFs calculated with the radiative transfer model SCIATRAN. BAMFs converge at higher altitude (not shown). The BAMFs are calculated for 352 nm (UV), for 438 nm (vis<sub>blue</sub>), and for 515 nm (vis<sub>green</sub>) at (a, b) 30° (c, d) 50° and at (e, f) 70° SZA. A surface spectral reflectance of 0.04 (352 nm), 0.06 (438 nm), and 0.09 (515 nm) representative for bare soil is assumed.

Another parameter needed for the AMF calculation is an a-priori  $NO_2$  profile (Burrows et al., 2011), as the total AMF is calculated as the average of the BAMFs of all atmospheric layers, weighted by the absorber concentration. The retrieved VCDs therefore , depend on the a-priori  $NO_2$  profiles, and differences between the a-priori and actual  $NO_2$  profiles can introduce systematic errors in the VCDs. As in principle, the final VCDs should not depend on the wavelength, analysing the differences in the VCDs retrieved from different wavelength regions can provide information about the accuracy of the used allows to infer the validity of a-priori  $NO_2$  profiles.

Typically, NO<sub>2</sub> DOAS fitting windows used in satellite data analysis are in the vis-blue spectral range (see for example: Richter et al., 2011 and van Geffen et al., 2015). The only One satellite NO<sub>2</sub> retrieval sofar-using UV

wavelengths has been developed for the Ozone Mapper and Profiler Suite (OMPS) on board of the Suomi National Polar-orbiting Partnership (SNPP) satellite, employing a DOAS-like method to derive  $NO_2$  VCDs (Yang et al., 2014). Compared to VCDs retrieved in the vis-blue spectrum with DOAS from the Ozone Monitoring Instrument (OMI) on board the Aura satellite, Yang et al. found good agreement between the two retrievals.

- 5 Richter and Burrows (2000) introduced also a  $NO_2$  fitting window in the UV as well as in the green spectral range to retrieve tropospheric VCDs from GOME measurements. They conclude that it is possible to retrieve  $NO_2$ from both spectral ranges and that stratosphere and troposphere can be separated by combining the two retrievals. However, the green spectral range is strongly influenced by liquid water absorption and interferences with vegetation and therefore, a reliable global retrieval of  $NO_2$  was therefore not possible from the green spectrum.
- In the present study, we present introduce a DOAS retrieval for NO<sub>2</sub> in the UV spectral range and use the results and compare this compare the results with a DOAS retrieval in the vis-blue spectral range, which is our standard approach, to investigate the vertical sensitivity of satellite-based NO<sub>2</sub> measurements. Furthermore, we present a case study for the green spectral range for China. In Sect. 2, the NO<sub>2</sub> DOAS retrieval in the UV spectral range will be introduced for measurements from the Global Ozone Monitoring Experiment 2 (GOME-2) on board MetOp-A. The
- 15 UV NO<sub>2</sub> retrieval settings will be discussed in detail as well-and the dataset from the vis-blue and green spectral range will be presented. In Sect. 3, results for the two/three NO<sub>2</sub> retrievals in the UVand vis-, green, and blue spectral ranges will be compared, and their implications for the NO<sub>2</sub> vertical distribution will be discussed. The manuscript ends with conclusions and a summary a summary and conclusions in Sect. 4.

#### 2 Datasets and methods

### 20 2.1 The GOME-2A instrument

Different types of remote sensing instrumentation on satellite platforms yield spectral radiance at the top of the atmosphere, which <u>yield contains information on</u> the amounts and distribution of  $NO_2$  in the Earth's atmosphere. Here, we use the GOME-2 instrument (Callies et al., 2000) flies (Callies et al., 2000; Munro et al., 2016) on board the MetOp-A satellite (hereafter GOME-2A), which has a swath width of 1920 km, resulting in nearly global coverage

- 25 every day. MetOp-A was launched in October 2006 into a sun-synchronous orbit with an equator crossing time of 09:30 local time (LT) in descending node. The GOME-2A instrument is a nadir viewing instrument spectrometer with a ground pixel size of 80 x 40 km<sup>2</sup> (Callies et al., 2000). In July 2013, the ground pixel size was reduced to 40 x 40 km<sup>2</sup>, when the identical GOME-2 instrument on board MetOp-B (hereafter GOME-2B), launched in September 2012, took over operational measurements (Munro et al., 2016). The spectrometer is separated into four channels covering
- 30 wavelengths from 240 to 790 nm (Callies et al., 2000). Channel 2 provides data in the UV spectral range from 311 to 403 nm with a spectral resolution of 0.26 to 0.28 nm, while channel 3 provides data in the vis spectral range from 401 to 600 nm with a spectral resolution of 0.44 to 0.53 nm. These two channels are used for the following analysis.

# 2.2 NO<sub>2</sub> DOAS retrieval in the UV spectral range

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For this study, we developed a NO<sub>2</sub> DOAS retrieval for the GOME-2A instrument in the UV which uses a fitting window between 342 and 361.5 nm, and a polynomial degree of of degree four. In this retrieval, one NO<sub>2</sub> cross section at 243K measured with the GOME-2A instrument (Gür et al., 2005) before launch, as well as two O<sub>3</sub> cross sections are used (at 223K and 243K, Serdyuchenko et al., 2014) with an additional I<sub>0</sub> correction of  $10^{20}$  molec cm<sup>-2</sup> (Platt et al., 1997; Richter, 1997). Additionally, cross sections for O<sub>4</sub>(Greenblatt et al., 1990), BrO(Wilmouth et al., 1999), HCHO(Meller and Moortgat, 2000), the Ring effect (calculated with SCIATRAN,

**Bro(winnouth et al., 1999)**, **HCHO(Meher and Moortgat, 2000)**, the King effect (calculated with SCIATRAN, Rozanov et al., 2014) and the instrumental cross section Zeta (EUMETSAT, 2011) are included in the fitting procedure (see Tab. 1).

|                | UV spectral range  | $vis_{blue}$ spectral range   | $vis_{green}$ spectral range  |  |  |
|----------------|--|---|---|--|--|
| fitting window | $342 - 361.5{ m nm}$   | $425-450\mathrm{nm}$  | $490-540\mathrm{nm}$  |  |  |
| polyn. degree  | 4  | 4   | 7   |  |  |
| cross sections |  |   |   |  |  |
| $NO_2$         | $223 \mathrm{K}; \mathrm{Gür} \mathrm{et} \mathrm{al.} (2005)$ | $243 \mathrm{K}; \mathrm{G\ddot{u}r} \mathrm{et} \mathrm{al.} (2005)$ | $243 \mathrm{K}; \mathrm{G\ddot{u}r} \mathrm{et} \mathrm{al.} (2005)$ |  |  |
| O <sub>3</sub> | $223 \mathrm{K}$ and $243 \mathrm{K}$ ;                        |   | $223\mathrm{K};$  |  |  |
|                | Serdyuchenko et al. (2014)                                     | 223  K;  Gur et al.  (2005)   | Serdyuchenko et al. (2014)  |  |  |
| $O_4$          | Greenblatt et al. (1990)                                       | Greenblatt et al. (1990)  | Greenblatt et al. (1990)  |  |  |
| $H_2O$         | _  | Rothman et al. $(2010)$   | Rothman et al. $(2010)$   |  |  |
| BrO            | Wilmouth et al. (1999)   | _   | _   |  |  |
| HCHO           | Meller and Moortgat (2000)                                     | _   | _   |  |  |
| Ring           | calculated with SCIATRAN, Vountas et al. (1998)                |   |   |  |  |
| instr. func.   | Zeta; EUMETSAT (2011)  | -   | _   |  |  |

Table 1. Fit settings for the  $NO_2$  retrievals in the UV and the vis spectral range.

10 These settings are the result of a number of careful sensitivity tests. Among many different wavelength windows we tested, the selected window from 342-361.5 nm provided the best resultssmallest root mean squared error (RMS) of the fit residual. The instrumental correction function Eta (EUMETSAT, 2011) only had minor influence on the results when included in the fit; we therefore chose not to include it in the fitting procedure use it to keep the number of fit parameters small. Including additional O<sub>3</sub> cross sections to correct the O<sub>3</sub> non-linearity (Pukīte et al., 2010) improved the fit in polar latitudes, where color angles and thus O<sub>4</sub> abcomptions are large. However, this also

15 2010) improved the fit in polar latitudes, where solar angles and thus  $O_3$  absorptions are large. However, this also

introduced an offset in the data, and as we are mainly interested in polluted and not in polar areasmid-latitudes, we chose not to use it here. Finally, we investigated the effect of using a daily Earth reflectance spectrum as reference. However, the differences between instead of a daily solar reference. The differences between results using the two reference spectra was were only minor so that we use the daily solar reference in order to preserve consistency with the fit in the vis-blue wavelength region (see Sect. 2.3).

### 2.3 NO<sub>2</sub> DOAS retrieval in the visible spectral range

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Due to the considerably stronger <u>differential</u> absorption lines in the <u>vis\_blue</u> compared to the UV spectral range (Fig. 2), NO<sub>2</sub> DOAS retrievals normally use measurements in the <u>vis\_blue</u> spectral range, see for example Richter et al. (2011) or van Geffen et al. (2015). Here for the <u>vis\_blue</u> spectral range, a retrieval from 425-450 nm with a polynomial

10 degree of four is used (in the following vis<sub>blue</sub>; introduced in Burrows et al., 1999 for the GOME instrument, applied to GOME-2A measurements in Richter et al., 2011). The fit settings are summerized summarized in Tab. 1. Cross sections for NO<sub>2</sub> at 243K and O<sub>3</sub> at 223K are used for this retrieval. Additional cross section are O<sub>4</sub>, H<sub>2</sub>O, and the Ring effect.



Figure 2.  $NO_2$  absorption cross section measured at 243 K with the GOME-2 instrument now deployed on MetOp-A. The fitting windows used in this study are shaded by colorcolour. GreenRed: new  $NO_2$  fitting window in the UV spectral range. Green:  $NO_2$  fitting window in the green spectral range. Blue:  $NO_2$  fitting window in the vis-blue spectral range (see Tab. 1).

Furthermore, a second fitting window in the vis spectral range from 490-540 nm with a polynomial degree of
seven is used (in the following vis<sub>green</sub>; introduced in Richter and Burrows, 2000 for the GOME instrument). The cross sections used in this fit are NO<sub>2</sub>, O<sub>3</sub>, O<sub>4</sub>, H<sub>2</sub>O, and the Ring effect. The fit settings are shown in Tab. 1. The high polynomial degree of seven is chosen in order to correct for interferences with the surface and liquid water absorption. Including a liquid water cross section for the green spectral range introduces errors over land; therefore, no liquid water cross section is used in the NO<sub>2</sub> fit for the green spectral range.

# 20 2.4 Comparison of the NO<sub>2</sub> DOAS retrievals in the UV and visible spectral range

In the UV and green spectral range, the differential NO<sub>2</sub> absorption lines are weaker than in the vis-blue spectral range. In Fig. 2 the NO<sub>2</sub> cross section (Gür et al., 2005) at 243 K measured with the GOME-2A instrument is

shown. Coloured shaded areas indicate the fitting windows in the UV (green) and vis (blue) spectral range. The wavelength dependency of the differential absorption strength is clearly observed. Furthermore, the measurement sensitivity for  $NO_2$  decreases towards the surface. This can be clearly observed in the BAMF (Fig. 1). In-This effect is enhanced by the temperature dependency of the  $NO_2$  cross section which is not taken into account in the

- 5 DOAS fit. The temperature dependency influences the tropospheric as well as the stratospheric NO<sub>2</sub> measurements (see Fig. S4 and Fig. S5). For the stratospheric NO<sub>2</sub> the combined radiative transfer and temperature dependency leads to a 10% higher sensitivity in the UV spectral range than in the blue spectral range. Close to the surface, the temperature sensitivity is up to 10% stronger in the blue spectral range compared to the UV spectral range. For the green spectral range, the differences are less pronounced compared to the blue spectral range (Fig. S6 and
- 10 Fig. S7). The green spectral range has a slightly reduced temperature sensitivity (up to 3%) close to the ground and an enhanced sensitivity (up to 9%) in the stratosphere compared to the blue spectral range. Additionally, the temperature vertical sensitivity introduces a seasonal and a latitudinal dependency. This effect is stronger in the tropics than for higher/lower latitudes and in the mid-latitudes it is more pronounced in summer and less in winter.
- 15 The smaller absorption lines in the UV in combination with reduced intensity from the sun , this and the weaker signal from the surface leads to a considerably higher noise level in the NO<sub>2</sub> differential optical depths in the UV spectral range compared to the visblue. Similarly, in the green spectral range the absorption lines are smaller than in the blue spectral range and therefore, also the green spectral range has a higher noise level than the blue spectral range. To illustrate this, Fig. 3 shows an example for the UV (a) and vis (b) fit results of the fit results from different
- 20 <u>wavelength regions</u> for one measurement above Teheran( $35.38^{\circ}N$ ,  $51.47^{\circ}E$ ) on 22 January 2008. The Figure shows the reference (soild lines) and reference plus the unexplained part of the signal, called *residual* (dashed lines). As expected, the retrieved NO<sub>2</sub> SCD is often smaller in the UV spectral range compared to the vis(for the measurement shown in Fig. 3, the retrieved SCDs are  $6.31 \times 10^{16}$  moleccm<sup>-2</sup> in the UV and  $9.33 \times 10^{16}$  moleccm<sup>-2</sup> in the vis spectral range).
- 25 BAMF for UV (green line) and vis (blue line) spectral range and the ratio of both BAMFs calculated with the radiative transfer model SCIATRAN. BAMFs converge with higher altitude (not shown). The BAMFs are calculated for 352nm (UV) and for 438nm (vis) at (a, b) 30°(c, d) 50° and at (e, f) 70° solar zenith angle (SZA). A surface spectral reflectance of 0.04 (UV) and 0.06 (vis) is assumed.

Due to the noisier differential optical depth signal, the UV NO<sub>2</sub> fit has larger uncertainty compared to the vis<sub>blue</sub> 30 or the vis<sub>green</sub> NO<sub>2</sub> fit. The UV NO<sub>2</sub> fit of the example has a random error of 4.3 %, whereas the vis<sub>blue</sub> NO<sub>2</sub> fit has a smaller error of 0.8 %. This is visible in Fig. 4 which shows the distribution of total The NO<sub>2</sub> VCDs over a presumably clean area fit in the green spectral range has only a slightly higher error than the blue fit with 0.9 %. Generally, the errors for the spectral ranges differ (see Tab. 2). For the estimation of the errors for the total SCDs over an assumed clean region (equatorial Pacificfrom : 5°S to 5°N and from 150°E to 210°E) are calculated. In the blue spectral

35 range, the standard deviation is smallest with  $0.6 \times 10^{15}$  molec cm<sup>-2</sup> followed by the green spectral range with



Figure 3. The NO<sub>2</sub> reference (differential NO<sub>2</sub> cross section multiplied with retrieved NO<sub>2</sub> SCD; solid line) and reference plus residual (dashed line) for the (a) UVand. (b) vis blue, and (c) green spectral range for one pixel above Teheran (35.38°N, 51.47°E) on 22 January 2008. The SCD for this pixel for the UV spectral range is  $6.31 \times 10^{16}$  molec cm<sup>-2</sup> with a fit error of 4.3%. The SCD for this pixel, for the vis blue spectral range it is  $9.33 \times 10^{16}$  molec cm<sup>-2</sup> with a fit error of 0.8%, and for the green spectral range is  $9.88 \times 10^{16}$  molec cm<sup>-2</sup> with a fit error of 0.9%. Note the different y-axes.

 $1.4 \times 10^{15}$  molec cm<sup>-2</sup>. The highest scatter is found for the UV spectral range with  $1.8 \times 10^{15}$  molec cm<sup>-2</sup>. This can also be seen in Fig. 4 which shows the distribution of total NO<sub>2</sub> VCDs over the selected region. Here, a stratospheric AMF was applied to calculate the NO<sub>2</sub> VCDs, as the common assumption is that no or very little tropospheric NO<sub>2</sub> is present in this area (Richter and Burrows, 2002; Martin et al., 2002; Peters et al., 2012). Due to the lower fit quality in the UV, the NO<sub>2</sub> columns retrieved there have a larger standard deviation of  $7.4 \times 10^{14}$  molec cm<sup>-2</sup> compared to the vis<sub>blue</sub> with a standard deviation of  $2.1 \times 10^{14}$  molec cm<sup>-2</sup>. For the green spectral range, a standard deviation of  $5.8 \times 10^{14}$  molec cm<sup>-2</sup> is found. The green fitting window shows a bias to slightly more negative values over the Pacific Ocean which is related to liquid water absorption (not shown; see Sect. 2.3).



Figure 4. Distribution of total NO<sub>2</sub> VCDs over a clean region (equatorial Pacific:  $5^{\circ}S - 5^{\circ}N$  and  $150^{\circ}E - 210^{\circ}E$ ) for three fitting windows in the UV, green, and vis blue spectral range for January 2008. Curves are normalised normalized to unit area and centered centered on zero.

Table 2. Errors for the UV and vis spectral range.

|                       | errors UV  | errors vis <sub>blue</sub>  | $ m errors vis_{green}$                             |  |  |
|-----------------------|--|---|---|--|--|
| total SCDs            | $1.8\times10^{15}\mathrm{molec}\mathrm{cm}^{-2}$                             | $0.6 \times 10^{15} \mathrm{molec} \mathrm{cm}^{-2}$ $1.4 \times 10^{15} \mathrm{molec} \mathrm{cm}^{-2}$ |   |  |  |
|                       | (STD calculated above the Pacific Ocean, see Fig. 4)                         |   |   |  |  |
| total VCDs            | $7.4\times10^{14}\mathrm{moleccm^{-2}}$                                      | $2.1 \times 10^{14} \mathrm{molec}\mathrm{cm}^{-2}$   | $5.8 \times 10^{15} \mathrm{molec}\mathrm{cm}^{-2}$ |  |  |
|                       |  | (see Fig. $4$ )   |   |  |  |
| SSR                   | $40\%$ at $320\mathrm{nm},$  | $5\%$ at $500\mathrm{nm},$  |   |  |  |
|                       | Kleipool et al., 2008  | Kleipool et al., 2008   |   |  |  |
| 0.05, increase $0.01$ | BAMF increases $9\%$   | BAMF increases $11\%$   | _   |  |  |
|                       | $(338\mathrm{nm},$   | $(440\mathrm{nm},$  |   |  |  |
|                       | Lorente et al., 2017 $)$   | Lorente et al., 2017 )  |   |  |  |
| AMF                   |  |   |   |  |  |
| cloud fraction        | 0-30%, Boersma et al., 2004  |   |   |  |  |
| cloud height          | <10%,Boersma et al., 2004  |   |   |  |  |
| aerosols              | not included in our calculations $\rightarrow$ 15%, Boersma et al., 2004     |   |   |  |  |
| profile shape         | < 15% (regions with little NO <sub>2</sub> : $> 50%$ ), Boersma et al., 2004 |   |   |  |  |

Generally, both Furthermore, for VCD calculation several input parameters are needed (see also Sect. 2.7) which have further uncertainties. Several previous studies calculated errors for DOAS retrievals, as for example Boersma et al. (2004) and Lorente et al. (2017). Their results are summarized in Tab. 2. Additionally, the errors of the SSR differ strongly between the UV (~40%) and vis spectral range (~5%) which increases the uncertainty in

5 the UV stronger (Kleipool et al., 2008). Overall, the uncertainties for the NO<sub>2</sub> concentration due to the SSR can be large, but in total it is smaller than 50% (Boersma et al., 2004)

<u>Generally, all three</u> NO<sub>2</sub> retrievals lead to similar SCD <u>spatial</u> patterns. In Fig. 5, NO<sub>2</sub> SCDs for the UV<del>and vis</del>, <u>blue and green</u> spectral range are shown for one orbit passing above East China on 03 January 2008. The larger uncertainty of the UV and vis<sub>green</sub> NO<sub>2</sub> retrieval is reflected in this figure by the larger spread of the slant columns

10 and the existence of unphysical negative values. Nevertheless, the NO<sub>2</sub> columns in the UV and vis spectral range spectral range and both vis spectral ranges are comparable for background regions and show a similar dependency on latitude. Over highly polluted regions, as for example China (25° N – 50° N), the vis NO<sub>2</sub> values are larger, especially the green spectral range, compared to the NO<sub>2</sub> values in the UV spectral range.



Figure 5. Total NO<sub>2</sub> SCDs of one orbit (above East China on 03 January 2008) for three fitting windows in the UVund vis , blue, and green spectral range. Only data with SZA smaller 70° are shown.

# 2.5 Removing stratospheric and upper atmospheric Stratospheric NO<sub>2</sub>

Stratospheric  $NO_2$  shows large differences between the three spectral ranges. Figure 6 shows the latitudinal and seasonal dependency for the three  $NO_2$  fitting windows. The seasonal dependency clearly differs between the three fitting windows also over regions dominated by stratospheric  $NO_2$ , especially for the green wavelength range, strong

- 5 interferences are observable. Although the differences are smaller between the UV and blue spectral range, they are clearly observable, for example at the equator (Figure 6 b). In northern hemispheric summer the UV NO<sub>2</sub> values are higher than the NO<sub>2</sub> values derived from the blue spectral range whereas in northern hemispheric winter the NO<sub>2</sub> values from the blue spectral range are slightly higher. Consequently, the differences between the UV and blue spectral range have a clear seasonal dependency and therefore, it is currently not possible to consistently improve
- 10 the stratospheric NO<sub>2</sub> retrieval by using different wavelength ranges as suggested in Richter and Burrows (2000).



Figure 6. Time series of NO<sub>2</sub> total VCDs of 2008 for 30°S (averaged:  $28^{\circ}S - 32^{\circ}S$ ), 0°N (averaged:  $2^{\circ}S - 2^{\circ}N$ ), and 30°N (averaged:  $28^{\circ}N - 32^{\circ}N$ ) above the area of the reference sector for cloud free pixels.

# 2.6 Conversion to tropospheric NO<sub>2</sub> columns

For the comparison of tropospheric SCDs and VCDs, a correction for the impact of stratospheric  $NO_2$  on the measured SCDs is needed. For the SCDs, we use the "reference sector method" (Richter and Burrows, 2002; Martin et al., 2002) for the three spectral ranges separately, in which a monthly average of SCDs measured over a presumably

- 5 clean area above the Pacific (180° E to 210° E) is subtracted from all measurements per latitude band. The underlying assumption for this simple correction method is that no NO<sub>2</sub> is present over the reference sector in the troposphere, and that stratospheric NO<sub>2</sub> is zonally homogeneously distributed (Richter and Burrows, 2002; Martin et al., 2002). However, as this assumption is not always valid (Richter and Burrows, 2002; Boersma et al., 2004; Hilboll et al., 2013b), the reference sector method leads to areas with negative tropospheric NO<sub>2</sub> values, which are related to
- 10 zonal inhomogeneities in stratospheric NO<sub>2</sub>, for example close to the polar vortex (Dirksen et al., 2011). For this correction the reference sector method, the same cloud screening as for the data selection is used (see Sect. 2.8). As the sensitivity to the stratosphere is nearly similar for both similar in the three spectral ranges (see Sect. 2.7), this correction can be used the reference sector method is not expected to introduce systematic differences.
- For the VCDs, we use a more sophisticated correction method, in which stratospheric VCDs from the Bremen-3D chemistry transfer model (B3dCTM, see Hilboll et al., 2013b and references therein) are used for the stratospheric correction. Therefore, no <u>significantly</u> negative values are expected for the VCDs <u>assuming the model is correct</u>. A Lambertian surface is assumed and the Lambert-equivalent <u>reflectly reflectivity</u> is taken as the SSR. For the AMF, a cloud correction is applied using the independent pixel approximation and cloud radiance fractions derived from the Fast REtrieval Scheme for Clouds from the Oxygen A-band (FRESCO+, version 6; Wang et al., 2008) dataset.
- 20 Furthermore, the temperature dependency of the cross section is included by using a linear approach as suggested in Boersma et al., 2004.

### 2.7 Radiative transfer simulations

The vertical sensitivity of the measurements to NO<sub>2</sub> has been investigated using box air mass factors simulated for both BAMFs simulated for the UV (352 nm) and the vis, the blue (438 nm), and the green (515 nm) fitting windows

- by the radiative transfer model SCIATRAN (version 3.6.5, see Rozanov et al., 2014). To calculate the radiance, SCIATRAN requires knowledge about the measurement scenario, e.g., viewing geometry, solar position, atmospheric absorbers, spectral surface reflectance SSR, surface elevation, aerosols, and clouds as input parameters. Here, SSR and surface elevation have been taken from the OMLER\_V003 (5-year climatology 2005-2009; wavelength: 354 nmand, 442 nm, and 499 nm; Kleipool et al., 2008) and GMTED2010 (Danielson and Gesch, 2011) datasets, respectively.
- 30 Exemplarily, results for 30°, 50° and 70° solar zenith angle (SZA) and SSR of 0.04 (UV)and 352 nm), 0.06 (vis438 nm), and 0.09 (515 nm) are shown in Fig. 1. The SSR differs between the two spectral ranges; the two SSR values are chosen to be representative of soil which shows the strongest differences in BAMF close to the ground. For 50° SZA, the sensitivity at the surface in the vis blue is about three to four times larger than in the UV. The sensitivity difference

in the green spectral range at this SZA is eight times higher than in the UV spectral range. In this wavelength range, the sensitivity at the surface depends stronger on SZA (for increasing SZA the differences between the two spectral ranges increase). The sensitivity difference between the wavelength regions decreases with increasing altitude, until at up to about 9 km, where they are identical. Above 9 km the UV shows a slightly (up to 4%) higher sensitivity

- 5 to NO<sub>2</sub> compared to the vis-blue spectral range. The sensitivity of the green spectral range is slightly smaller (up to 3%) than for the blue spectral range. With higher altitude the BAMFs converge, up to at around 30 km, from which onwards there are no significant differences between the UV and the vis spectral range spectral ranges (not shown). As shown in Fig. 1, the point of identical sensitivity increases slowly moves upward with increasing SZA, which also influences the sensitivity in the atmosphere below the point of identical visibility. However, close to the
- 10 ground, the differences between to two wavelength ranges the UV and blue spectral range stay nearly constant. In order to calculate the effective AMF for one measurement scene, the BAMF profile needs to be combined with an a-priori vertical profileNO<sub>2</sub> concentration. This a-priori profile has a strong impact on the AMFs, as they provide it provides the basis for the effective visibility of the NO<sub>2</sub> in the measurement scene. In this study, we
- used vertical NO<sub>2</sub> profiles from the TM5-MP model, described in Williams et al., 2017. For the European project
  QA4ECV (http://www.qa4ecv.eu) a 15 year run (2002-2016) was performed with the TM5-MP model version of July 2016. The model was run with a resolution of 1° × 1° on 34 levels, using ERA-Interim reanalysis meteorology from ECMWF. Output was provided for , and with a temporal sampling every of 2 hours. This dataset has been used within the QA4ECV project as a-priori for the retrieval of e.g. formaldehyde, and for other projects such as the one described here. In particular, biomass burning emissions are taken from the monthly estimates provided by the
- 20 GFEDv3 inventory (van der Werf et al., 2010), and latitude-dependent injection heights and a tropical burning cycle are implemented following Huijnen et al., 2010. See Williams et al., 2017 for other model details. In order to correct for the temperature dependency of the NO<sub>2</sub> cross-section, the NO<sub>2</sub> concentrations in each altitude are scaled by a correction factor linear in temperature for all spectral ranges, as suggested in Boersma et al., 2004. Temperatures were taken from the model; the temperature dependencies are illustrated in the supplement (Fig. S1–Fig. S7).

#### 25 2.8 Data selection and post-processing

Both the UV and vis dataset. The datasets for the three spectral ranges were gridded to a  $0.25^{\circ} \times 0.25^{\circ}$  grid and monthly means were calculated. Because of the fit quality and the reduced intensity in the UV spectral range, only measurements at solar zenith angles smaller than 70° are included in the following analyses. As a result of the fitting problems in the green spectral range (Sect. 2.3 and Sect. 2.4), in the following only time series for China are

30 shown where the NO<sub>2</sub> concentrations are high and therefore, further influences are less pronounced (global maps are shown Fig. S11–Fig. S14). Only measurements with a geometric cloud fraction smaller than 0.2 are included, unless otherwise noted. Cloud filtering was performed using the FRESCO+ (version 6) dataset (Wang et al., 2008). Furthermore, only fits with a  $\chi^2$  (describing the fit residual) smaller than 0.005 and 0.001 for the UV and vis spectral range, respectively, were used. While for the first years of the GOME-2A measurements, a consistent  $\chi^2$ -limit could be used for both all spectral ranges, instrumental degradation severely impacts the instrument's performance in later years (Dikty and Richter, 2011). Since channel 2 is more strongly affected by this degradation than channel 3, a larger  $\chi^2$ -limit for the UV spectral range is needed for a statistically meaningful comparison.

# 3 Results and Discussion

# 5 3.1 Spatial distribution of NO<sub>2</sub> slant columns

Slant column densities (SCDs) are retrieved from the DOAS fit; they do not depend on any a-priori assumptions on the state of the atmosphere. Therefore, a comparison of the spatial distribution of the SCDs from the two UV and vis<sub>blue</sub> fit windows provides a first opportunity to assess the NO<sub>2</sub>. However The SCDs depend on the VCDs and on the AMFs (see Eq. 3). Furthermore, the SCDs depend on the measurement geometry, SSRand-wavelength, and

- 10 wavelength, and the AMFs additionally depend on the a-priori vertical NO<sub>2</sub> profile. Consequently, maps of SCDs show not only the altitude dependency, but also other changes of the AMF. For naturally polluted areas, the AMFs for the vis and the UV spectral ranges differ by a factor of ~0.9–1.1 (not shown) whereas the changes in SCDs differ between a factor of ~1.1 (northern hemispheric, NH, summer in Africa north of the equator; ANE) and ~2 (see for example Fig. 11). For anthropogenically polluted areas, the AMFs differ by a factor of ~0.9–1.3 whereas the
- 15 SCDs differ by a factor of  $\sim 2$  3 or moreone particular measurement, the geometry is the same irrespective of the wavelength window. Thus, differences in the SCDs are introduced by the different sensitivities and therefore, large differences between the two spectral ranges are expected when the NO<sub>2</sub> is located close to the surface; when the NO<sub>2</sub> located higher in the atmosphere, the differences between the two spectral ranges are reduced. Consequently, the part of the SCD differences can be partly explained by the different AMFsmeasurement geometries and SSR, which
- 20 <u>change during the year</u>, but other factors like injection height and relative vertical distribution have to contribute as well.

Figure 7 shows monthly averages of tropospheric  $NO_2$  SCDs for January (upper panels) and July (lower panels) and July 2008. Tropospheric SCDs are shown for both, UV (a, c) and vis (b, d) fits. In both spectral ranges, similar spatial patterns are found and anthropogenic as well as natural air pollution can be detected. Especially over

- 25 the respective winter hemisphere, anthropogenic source regions are clearly observed. In January, for example, the highest NO<sub>2</sub> columns are located over China (a, b), whereas in July, there are high NO<sub>2</sub> values over the Highveld Plateau region in South Africa (HPSA; c, d). In addition, biomass burning regions are also clearly visible observable in the data. For instance, in July over Africa south of the equator (ASE), enhanced values are detected in both spectral ranges, consistent with van der A et al. (2008) and Schreier et al. (2014), who found a similar seasonal
- 30 pattern. Finally, artefacts originating from the simplified stratospheric correction are observed in Fig. 7 (a, b), when enhanced NO<sub>2</sub> values over the Pacific reference area lead to large areas with too low NO<sub>2</sub> values in the subtropics of the northern hemisphere (see Sect. 2.6) in both spectral ranges.



Figure 7. Monthly mean tropospheric NO<sub>2</sub> SCDs for (a, b) January and (c, d) July 2008. (a, c) UV spectral range and (b, d) vis blue spectral range.

In order to better compare the two NO<sub>2</sub> retrievals, Fig. 8 shows the absolute relative differences between these results (absolute differences are shown in Fig. S15). Over highly polluted areas, NO<sub>2</sub> values in the UV spectral range are generally lower than in the vis-blue spectral range (see Sect. 2.4). These differences are introduced by the wavelength dependency of the penetration depth (see Sect. 1), which leads to better visibility of lower tropospheric pollution to the vis<sub>blue</sub> retrieval. Larger differences between the two NO<sub>2</sub> retrievals can be found in the respective winter hemisphere, e.g., in January over China and the east coast of the USA or in July over HPSA (see also Fig. 7). These large differences are related to first, both the larger tropospheric NO<sub>2</sub> SCDs , and second, and to the larger SZA (see Fig. 1). In addition, large differences are observed in the western South Atlantic, where the retrieval noise in both spectral ranges is strongly enhanced due to the South Atlantic Anomaly (Richter et al., 2011).

5

10 To further analyse the differences between the two spectral windows, we use the ratio of UV to vis<sub>blue</sub> tropospheric NO<sub>2</sub> SCDs which is a first approximation of an altitude dependency; Fig. 9 shows these for January (a) and July 2008-January and (b). To visualise the ratio of the two retrievals, a threshold value was used to filter out pixels



Figure 8. Absolute Relative difference between monthly mean tropospheric  $NO_2$  SCDs in the vis and UV and blue spectral range. Differences for (a) January and (b) July 2008. Dark gray grey shaded area: no  $NO_2$  values available. Light grey coloured values indicate values where the vis<sub>blue</sub>  $NO_2$  is close to zero, which have been filtered out. This threshold was defined as a smoothed latitude-dependent mean over the reference sector area. The Pacific threshold is one standard deviation of the gridded  $NO_2$  values for both retrievals. To smooth the latitude-dependent threshold, a 5°-running-mean was used. The regions defined in Tab. 3 are depicted in the right figure (b).

where a too low denominator (due to noise) would lead to unrealistically high ratios. The threshold was defined as a smoothed latitude-dependent mean over the reference sector area. The Pacific threshold is one standard deviation of the gridded values for both retrievals. To smooth the latitude-depend threshold, a 5°-running-mean is used.

- July 2008. Highly polluted areas are clearly visibleobservable, with ratios as low as ~0.2, indicating the NO<sub>2</sub>
  being located close to the surface. As before, differences between the two retrievals are larger in the respective winter hemisphere, showing as lower ratios. The most obvious example is East China in January, where large areas with very low ratios can be detected. However, all strong anthropogenic source regions, e.g., India, northern America, HPSA, and the Middle East, but also individual large cities like Madrid, Moscow, and Mexico City, show low ratios between the two retrievals. Here, it should be noted that due to the rather strict selection criterion on the SZA (see Sect. 2.8), no data are available for the strong source regions in central Europe in January.
- In addition to anthropogenic air pollution close to the ground, NO<sub>2</sub> pollution from biomass burning can be detected in the retrieval ratios, e.g., in ANE and ASE (Fig. 9). In July (southern hemispheric, SH, winter), values of ~0.6 are found for the  $UV/visvis_{blue}$  ratio over the ASE region. These low values (as well as the absence of a signal in January) correspond to the seasonal pattern of biomass burning in this region.
- 15 In the ANE region, the vis<sub>blue</sub> fit results show clearly enhanced NO<sub>2</sub> SCDs in January (NH winter, coinciding with the annual biomass burning peak in this region) but not in July (see Fig. 7). However, no significant differences



Figure 9. Ratio between monthly mean tropospheric SCDs of NO<sub>2</sub> in the UV and vis-blue spectral range. (a) January 2008 and (b) July 2008. Dark gray grey shaded area: no NO<sub>2</sub> values available. Light grey coloured values indicate values where the vis<sub>blue</sub> NO<sub>2</sub> is close to zero, which have been filtered out (Fig. 8).

between the two months are observed in the UV fit results, where slightly enhanced values can be seen throughout the year (not shown). These findings will be further discussed in Sect. 3.2.

North of the ANE region at approx.  $10^{\circ}-20^{\circ}$  N, SCD ratios of ~0.7 can be found in July (see Fig. 9). This is an indicator for NO<sub>2</sub> enhancement most probably related to soil emissions (Jaeglé et al., 2004; Zörner et al., 2016).

- 5 While this NO<sub>2</sub> is more clearly observed in the vis-blue spectral region, also the UV results show enhanced SCDs (see Fig. 7). As shown by Delon et al. (2008) and Stewart et al. (2008),  $NO_x$  ( $NO_x = NO + NO_2$ ) from soil emissions is usually well mixed and thus not only located close to the ground, but also in elevated layers, enhancing visibility in the UV.
- Finally, Fig. 9 shows SCD ratios of ~0.6 over the well known shipping lane leading from South India to the Strait of Malacca. The shipping signal in this region was first identified in satellite NO<sub>2</sub> data by (Beirle et al., 2004; Richter et al., 2004) (Beirle et al., (2004) and Richter et al. (2004). Here, while the shipping lanes are clearly visible observable in the vis<sub>blue</sub> SCDs, they cannot be identified in the UV NO<sub>2</sub> data. This is probably because the NO<sub>2</sub> shipping emissions are located close to the groundstay with the shallow marine boundary layer. Consequently, the shipping emissions when averaged over the ground scene are below the detection limit for the UV spectral range (see
- 15 Fig. 7). Higher horizontal resolution (e.g., in OMI and the upcoming S-5P, S4S-4, and S-5 missions) might improve detection of shipping lanes in the UV.

From the retrieved ratios of SCDs, a first approximation of the top-altitude of  $NO_2$  can be calculated. For the altitude retrieval, simple box profiles from the surface up to a top altitude are assumed for tropospheric  $NO_2$  and the ratios of these box profile SCDs are compared with the retrieve ratios of the measurements. The altitude of

ratio from the box profile SCDs which fits best to the ratio of the measurement is taken as the  $NO_2$  top-altitude. The retrieved top-altitudes for the individual pixels are shown in Fig. 10. A seasonal dependency of the retrieved altitude can be clearly observed in the global maps. Low altitudes are found in regions with low ratios e.g., China in January and HPSA in July. High altitudes are found over the Oceans and in the surroundings of highly polluted areas which is expected from transport events (Richter et al., 2014 ; Zien et al., 2014 ).



Figure 10. Monthly mean top-altitudes retrieved from the ratio between the UV and blue spectral range. To retrieve the altitude, box profiles are assumed for the tropospheric  $NO_2$ . Light grey coloured values indicate values where the vis<sub>blue</sub>  $NO_2$  is close to zero, which have been filtered out.

# 3.2 Temporal variability of regional NO<sub>2</sub> slant columns

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In this section, we investigate the temporal variability of tropospheric SCDs from the two<u>/three</u>NO<sub>2</sub> retrievals over six regions (see Tab. 3 and Fig. 8 for the regions' definitions) and make inferences about the vertical distribution of NO<sub>2</sub> in the atmosphere. Figure 11 shows monthly mean time series for three natural / biomass burning source
regions (a-c) and three anthropogenic source regions (d-f). The seasonal cycle observed in Fig. 11 a-c corresponds to the seasonal pattern of biomass burning activity in these regions (Schreier et al., 2014). Slightly negative SCDs in the UV fit in north Australia (NAUS; Fig. 11 c) and Riyadh (Kingdom of Saudi Arabia; Fig. 11 f) are artefacts caused by the stratospheric correction (see Sect. 2.6). Another artefact in the data is related to degradation of the GOME-2A instrument (see Sect. 2.8; Dikty and Richter, 2011), which is in some regions represented by a slightly decreasing linear trend (for example: in the UV in ANE and HPSA; Fig. 11 b and c).

In both ASE and NAUS, the seasonal cycles in the UV and vis-blue spectral range are similar (Fig. 11 a and c), but the tropospheric NO<sub>2</sub> SCDs in the UV are smaller than in the vis-blue spectral range. As can be seen in Fig. 12, this results in the strong correlation coefficients (0.87 in both cases, see Tab. 4) and slopes  $\leq 1$  of the regression

Table 3. Selected regions for  $NO_2$  time series with their abbreviation and their location.

| Region                           | Abbreviation | geographical location   |
|----------------------------------|--------------|---|
| Africa south of the equator      | ASE          | $5^{\circ}\mathrm{S} - 20^{\circ}\mathrm{S}$ and $10^{\circ}\mathrm{E} - 40^{\circ}\mathrm{E}$        |
| Africa north of the equator      | ANE          | $0^{\circ}\mathrm{N} - 10^{\circ}\mathrm{N}$ and $15^{\circ}\mathrm{W} - 40^{\circ}\mathrm{E}$        |
| North Australia                  | NAUS         | $15^{\circ}\mathrm{S} - 25^{\circ}\mathrm{S}$ and $120^{\circ}\mathrm{E} - 145^{\circ}\mathrm{E}$     |
| China                            |              | $30^{\circ}\mathrm{N}-40^{\circ}\mathrm{N}$ and $110^{\circ}\mathrm{E}-123^{\circ}\mathrm{E}$         |
| Highveld Plateau in South Africa | HPSA         | $22^{\circ}\mathrm{S} - 30^{\circ}\mathrm{S} \mathrm{and}27^{\circ}\mathrm{E} - 34^{\circ}\mathrm{E}$ |
| Riyadh                           |              | $23^{\circ}\mathrm{N} - 26^{\circ}\mathrm{N}$ and $45^{\circ}\mathrm{E} - 48^{\circ}\mathrm{E}$       |



Figure 11. Time series 2007-2015 of tropospheric NO<sub>2</sub> SCDs for the UV (green red line) and vis<sub>blue</sub> (blue line), and vis<sub>green</sub> (green line, China only) spectral range for different regions. Note the different y-axes. The vertical lines indicate January of the individual years. (a-c) biomass burning regions and (d-f) regions with high anthropogenic air pollution.

lines (see Tab. 5). When the seasons are considered individually, the slope differs between 0.65 (biomass burning season) and 1.04 (rainy season). Hence, for both regions in the rainy season the two spectral ranges differ only by an offset whereas in the biomass burning season large differences are observed. Furthermore, in both regions for the individual seasons the correlation coefficient is also high (0.50-0.99); see Tab. 4), especially in the biomass burning

season ( $\geq 0.92$ ), which is probably related to the higher signal to noise ratio for this season. In ANE, conversely, significant differences in the seasonal cycle between the two spectral ranges can be observed, as the NO<sub>2</sub> signal from the biomass burning peak in NH winter shows in the vis spectral range (Fig. 11b) blue spectral range only while in the UV, no interannual variability can be detected (Fig. 11b). The increased NO<sub>2</sub> load visible observed in the vis<sub>blue</sub> SCDs cannot be detected in the UV SCDs (nevertheless the correlation coefficient is only 0.75). In all the other seasons, the correlation coefficient is comparable to that observed in ASE and NAUS (see Fig. 12 b and Tab. 4).



Figure 12. Scatter plot of tropospheric UV NO<sub>2</sub> SCDs against tropospheric vis<u>blue</u> NO<sub>2</sub> SCDs for the period 2007-2015 for the six selected regions. Note the different axes. The seasons are <u>color colour</u> coded. The related correlation, slope and intercept of the regression line can be found in Tab. 4 and 5. December, January, February: DJF; March, April, May: MAM; June, July, August: JJA; September, October, November: SON.

Several effects could contribute to an explanation for the differences in visibility of  $NO_2$  in ANE in the two spectral ranges:

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- 1. Cloud influence on NO<sub>2</sub> measurements: When considering only those measurements flagged as cloudy (i.e., having a cloud coverage  $\geq 0.3$ ), only a weak seasonal cycle can be found in ANE (Fig. A1, App. A1), whereas ASE shows a similar seasonal cycle in the cloud covered case as in the cloud free case. This indicates that the biomass burning NO<sub>2</sub> in ASE might be located partially in elevated NO<sub>2</sub> layers above the clouds, whereas in ANE NO<sub>2</sub> is located closer to the ground, i.e., usually below and thus partly shielded by the clouds.
- 2. NO<sub>2</sub> layer altitude: The altitude of biomass burning NO<sub>2</sub> emissions influence their visibility (see Fig. 1). According to simulations with the TM5-model (Williams et al., 2017; for details, see App. A2), the layer

| Region | Annual | DJF  | MAM  | JJA  | SON  |
|--------|--------|------|------|------|------|
| ASE    | 0.87   | 0.50 | 0.69 | 0.92 | 0.80 |
| ANE    | 0.53   | 0.75 | 0.77 | 0.76 | 0.63 |
| NAUS   | 0.87   | 0.76 | 0.75 | 0.87 | 0.90 |
| China  | 0.98   | 0.98 | 0.99 | 0.83 | 0.98 |
| HPSA   | 0.82   | 0.74 | 0.80 | 0.90 | 0.87 |
| Riyadh | 0.95   | 0.73 | 0.86 | 0.70 | 0.95 |

Table 4. Correlation between UV  $NO_2$  SCDs and visblue  $NO_2$  SCDs for the six selected regions.

height of biomass burning  $NO_2$  over ANE is lower than over ASE (Fig. A2). This is consistent with the above result, which shows high and low visibility for the biomass burning  $NO_2$  in the UV for the ASE and ANE regions, respectively. For ASE, several previous studies could show that biomass burning plumes are regularly located at least partly above the boundary layer (Coheur et al., 2007; Rio et al., 2010; Gonzi and Palmer, 2010).

- 5
- 3. Stratospheric correction: For observations in the UV spectral range, total SCDs (i.e., the SCDs resulting from the DOAS fit, without the stratospheric correction described in Sect. 2.6) over ANE show two distinct peaks over the course of the year: one peak in <u>NH</u> winter and another peak in <u>NH</u> summer (Fig. A3, <u>App. A3</u>). The <u>NH</u> winter peak in Dec./JanJan. falls in the main biomass burning season and is thus expected to be of tropospheric origin. In contrast, the <u>NH</u> summer peak in May-Jul-July is of stratospheric origin, as stratospheric NO<sub>2</sub> has its maximum in summer. It is noteworthy that in the vis<sub>blue</sub> spectral range, this <u>NH</u> summer peak is smaller compared to the UV, which might be partly related to small differences in the stratospheric sensitivity between the two wavelength ranges. In the UV SCDs, the biomass burning NO<sub>2</sub> might thus be obscured by the stratospheric correction due to the similar magnitude of the stratospheric <u>NH</u> summer peak.
- 4. Cloud influence on data sampling: The data shown in Fig. 11 are filtered to include only measurements not strongly influenced by clouds. In satellite-based cloud retrievals, smoke is often misinterpreted as cloud (Boersma et al., 2004). Assuming that smoke and NO<sub>2</sub> from biomass burning is advected together, this could lead to large parts of the biomass burning NO<sub>2</sub> being filtered out due to apparent cloud contamination of the measurement. However, the vis<sub>blue</sub> NO<sub>2</sub> SCDs are clearly enhanced in the cloud-free measurements, showing that not all biomass burning NO<sub>2</sub> is filtered out. Furthermore, there are only small differences in the cloud coverage between ANE and ASE (not shown); in both regions, the number of pixels being filtered out due to cloud contamination is similar. The minimum fraction of cloud free pixels is about 30 %, observed during

the rainy season, while the maximum fraction of about 70-80% and 80% in ANE and ASE, respectively, is observed during the dry season, when the biomass burning occurs. It seems thus unlikely that the cloud filtering significantly impacts the visibility of the biomass burning NO<sub>2</sub> over ANE.

5. Aerosols: The NO<sub>2</sub> measurements could be influenced by aerosols. According to CALIPSO measurements, the main aerosol-type in ASE is smoke from biomass burning, while other aerosol types contribute only a smaller amount to the total aerosol concentration (<u>Omar et al., 2009</u>; Fig. A4, App. A4). In contrast, the dominant aerosol-types in ANE are dust and polluted dust (<u>Omar et al., 2009</u>). As in the UV spectral range, the single scattering albedo (SSA) of dust is smaller than that of biomass burning aerosols (Russell et al., 2010; Dubovik et al., 2001; Bergstrom et al., 2007), the biomass burning NO<sub>2</sub> signal in ANE could be shielded by the darker dust aerosol in the UV. However, this hypothesis seems unlikely, as the number of dust aerosols (per area) is even higher in Riyadh compared to ANE, while in Riyadh a clear NO<sub>2</sub> seasonal cycle can be identified in both spectral ranges.

In summary, we <u>believe conclude</u> that the absence of <u>a</u> biomass burning  $NO_2$  signal in the UV spectral range over ANE is predominantly formed by the lower injection height, causing the  $NO_2$  to be located lower in the atmosphere

- 15 compared to ASE. Due to the strong differences in sensitivity between UV and vis blue spectral ranges close to the ground, this could to a large extent explain the invisibility of UV  $NO_2$  in ANE. However, also further effects could influence the visibility. For example, interferences of the seasonal cycles of stratospheric and tropospheric  $NO_2$  over ANE might contribute, while we deem the effect caused by differing prevalent aerosol types in the three regions unlikely to be significant.
- Figures 11 d-f show three regions dominated by anthropogenic air pollution. Here, absolute differences between UV and  $vis_{blue}$ NO<sub>2</sub> values are often larger than in biomass burning source regions (see also Fig. 9). The seasonal cycle is mostly larger in the <u>vis-blue</u> spectral range compared to the UV spectral range, indicating a seasonal dependency of the reasons underlying the differences between the spectral regions. In general, the time series of the two spectral ranges show a similar behaviour and are highly correlated (0.82; Tab. 4).
- Over East China, the NO<sub>2</sub> SCDs in UVand vis vis<sub>green</sub>, and vis<sub>blue</sub> have similar shape and show a very high correlation of  $\geq 0.98$  (Fig. 11 d; see also Table S1). Until 2012, increasing NO<sub>2</sub> winter values in both-the three spectral ranges can be observed; these increases have been analysed in depth by Richter et al. (2005) and Hilboll et al. (2013a). In 2013, there are nearly no changes in NO<sub>2</sub> and afterwards, NO<sub>2</sub> SCDs are slightly decreasing, which was already reported by Richter et al. (2015) and Irie et al. (2016) and is consistent with findings reported by Hilboll
- 30 et al. (2017) for anthropogenic NO<sub>2</sub> pollution in India. Over the HPSA region, no NO<sub>2</sub> trend is observed (Fig. 11 e). A clear seasonal cycle is observed in both spectral ranges, with the highest correlation coefficient in SH winter (0.90; Tab. 4). The detected seasonal cycle in the vis-blue spectral range is similar to the cycle found by Noije et al. (2006) in the year 2000 in data from the Global Ozone Monitoring Experiment (GOME). In Riyadh (Fig. 11 f) a small trend is observed in maximum values. As shown by Hilboll et al. (2013a), during the whole period a slight increase can

be detected. The detected month with maximum NO<sub>2</sub> is in agreement with van der A et al. (2008). In this area the NO<sub>2</sub> values derived from the UV and vis blue spectral range show a similar seasonal cycle and a similar year-to-year variability with a correlation coefficient of  $\geq 0.95$ .

| Region | Annual           | DJF             | MAM               | JJA              | SON              |
|--------|------------------|-----------------|-------------------|------------------|------------------|
| ACE    | 0.66/ -0.19      | 0.98/ -0.45     | 0.96/ - $0.44$    | 0.83/ -0.61      | 0.65/ -0.10      |
| ASE    | $(0.03/ \ 0.05)$ | $(0.23/\ 0.20)$ | $(0.17 / \ 0.16)$ | $(0.07/ \ 0.11)$ | $(0.09/ \ 0.15)$ |
| ANE    | 0.40/ 0.26       | 1.08/ -0.93     | 0.84/ -0.23       | 0.92/ -0.23      | 0.67/ $0.01$     |
| ANE    | $(0.05/ \ 0.06)$ | $(0.16/\ 0.26)$ | $(0.12/\ 0.15)$   | $(0.14/ \ 0.14)$ | $(0.13/ \ 0.14)$ |
| NAUS   | 0.74/ -0.32      | 0.89/ -0.41     | 1.04/ -0.50       | 0.78/ -0.30      | 0.71/ -0.34      |
|        | $(0.04/\ 0.03)$  | $(0.13/\ 0.11)$ | $(0.16/\ 0\ 09)$  | $(0.08/ \ 0.04)$ | $(0.06/ \ 0.06)$ |
| China  | $0.37/\ 1.04$    | 0.42/ -0.07     | 0.46/ $0.25$      | 0.49/ $0.52$     | 0.37/ $0.89$     |
|        | $(0.01/\ 0.08)$  | $(0.02/\ 0.30)$ | $(0.02/\ 0.14)$   | $(0.06/ \ 0.36)$ | $(0.02/ \ 0.16)$ |
| HPSA   | 0.41/ $0.24$     | 0.74/ -0.37     | 0.46/ $0.21$      | 0.68/ -1.03      | 0.60/ - $0.32$   |
|        | $(0.03/ \ 0.09)$ | $(0.12/\ 0.26)$ | $(0.06/\ 0.21)$   | $(0.06/ \ 0.27)$ | $(0.06/ \ 0.19)$ |
| Riyadh | 0.74/ -1.26      | 0.63/ -1.15     | 0.64/ -0.97       | 0.46/ $0.06$     | 0.71/ -1.10      |
|        | $(0.02/\ 0.08)$  | $(0.10/\ 0.26)$ | $(0.07/ \ 0.19)$  | $(0.08/ \ 0.40)$ | $(0.04/ \ 0.17)$ |

Table 5. Slope and intercept between UV NO<sub>2</sub> SCDs and  $vis_{blue}$  NO<sub>2</sub> SCDs for the six selected regions. Intercept in 1e15 molec cm<sup>-2</sup>. In the brackets the standard errors for both values are shown.

In anthropogenically polluted areas, the differences in the NO<sub>2</sub> signal strength between the two-UV and blue 5 spectral ranges are a result of the larger sensitivity to the lower troposphere in the vis-blue spectral range. In winter the air pollution is mostly stronger, compared to summer, the SZA is lowerhigher, the boundary layer is more shallow (von Engel and Teixeira, 2013), and even within this boundary layer also a strong gradient is expected over high-emission areas (Dieudonné et al., 2013). All these effects increase the observed differences between the two spectral ranges. Due to the higher profile sensitivity close to the ground and the resulting larger differences, a 10 lower regression slope (see Fig. 12) is observed for the anthropogenically polluted areas than for the biomass burning regions (Tab. 5) in all seasons. For example, in NAUS and ASE the regression line between all UV and vis<sub>blue</sub> SCDs has a slope of 0.66 and 0.74, respectively, whereas for China and HPSA, a considerably smaller slope (~0.4) is found. One has to note that the regions ANE and Riyadh do not follow this pattern — ANE is attributed the missing of the due to the missing biomass burning NO<sub>2</sub> signal in the UV, and Riyadh probably results from because of

15 differences in boundary layer height and the maximum in  $NO_2$  seasonal cycle (von Engel and Teixeira, 2013). In

Riyadh, the seasonal cycle of  $NO_2$  emission and the cycle of the boundary layer height are in phase, as it is the case for naturally polluted areas, whereas for China and HPSA, the  $NO_2$  maximum can be found in the respective winter season, where a lower boundary layer height is observed (Fig. S16). Consequently, in Riyadh the UV  $NO_2$  signal is stronger and the slope (0.74) is larger than usually expected for observed in other anthropogenic source regions.

- 5 Finally, a seasonal dependency of the regression line slope may result from the combination of the boundary layer height and the surface concentration of NO<sub>2</sub>. For ASE and NAUS, the slope shows a seasonal pattern, which might be influenced by the small signal to noise ratio (see Tab. 5). Slopes of  $\sim 1.0$  can be found in these areas for the rainy season, with the largest uncertainties for these areas, possibly related to high cloud cover. The differences are thus most likely introduced by an offset. During biomass burning season, the slope is about 0.83 and 0.71 for ASE
- 10 and NAUS, respectively. Since the biomass burning NO<sub>2</sub> signal cannot be seen in UV SCDs over ANE, no clear seasonal pattern is visible can be observed there. In East China and the HPSA region, a larger slope can be found in NH summer, when the boundary layer is higher; spring, autumn and winter show a smaller slope (for China this can also be found for the green spectral range, Table S1). This is expected due to the increased sensitivity (see Fig. 1) of the UV measurements to NO<sub>2</sub> in a thick compared to a shallow boundary layer. In Riyadh however, this
- 15 effect is not observed, possibly due to a decreased signal to noise ratio caused by the smaller area, entailing fewer measurementsor a stronger stratification of the boudary layer.

## 3.3 Spatial distribution of NO<sub>2</sub> vertical columns

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The NO<sub>2</sub> vertical column densities (VCDs) are the final result of the DOAS retrieval procedure, as they are a physically meaningful and universally comparable quantity. In this section, we compare NO<sub>2</sub> tropospheric VCDs retrieved from the UV and the vis blue spectral range. For the calculation of the VCDs, NO<sub>2</sub> vertical profiles from

- the TM5-model (Williams et al., 2017) are used as a-priori (see Sect. 2.7). Assuming perfect measurements, radiative transfer simulations, and a-priori profiles, the AMF would equalise all differences between the two spectral ranges, leading to <u>idential-identical</u> VCDs for UV and vis<sub>blue</sub> fits. Figure 13 shows maps of monthly <u>means-mean</u> NO<sub>2</sub> VCDs for January (<u>upper panels</u>) and July (lower panels) 2008, retrieved from the UV (a, c) and vis fits (b, d). and July
- 25 <u>2008.</u> The spatial patterns in both datasets are very similar and agree well with those visible shown in SCD maps (Fig. 7).

However, also the VCDs of both spectral ranges do not perfectly match, as illustrated in Fig. 14. Compared to the SCD differences relative differences of the SCDs shown in Fig. 8, the VCD differences are smaller relative differences of the VCDs are smaller (absolute differences are shown in Fig. S17). Especially over areas without significant NO<sub>2</sub>

30 sources, but also over the biomass burning regions, the differences between the two retrievals are reduced compared to the SCDs. Over anthropogenically polluted areas however, e.g., over China in January, the VCDs retrieved from the two spectral regions still show significant differences. The most probable reason for this is that the simulated NO<sub>2</sub> profiles used as a-priori do not represent the actual NO<sub>2</sub> vertical distribution; furthermore, aerosols and the SSR might influence the calculated VCDs, as will be discussed in Sect. 3.4.



Figure 13. Monthly mean tropospheric NO<sub>2</sub> VCDs for (a, b) January and (c, d) July 2008. (a, c) UV spectral range and (b, d) vis blue spectral range.

# 3.4 Temporal variability of regional NO<sub>2</sub> vertical columns

To better understand the reasons for the differences between the tropospheric VCDs of the two/three\_spectral regions, we again investigate the temporal variability over the six regions shown before (see Sect. 3.2, Tab. 3, Fig. 8 for the region definitions). The general shape of the seasonal cycles, which was described for SCDs in Sect. 3.2, can also be found in the VCD time series (Fig. 15). Similar to the SCD time series (see Sect. 3.2), the NO<sub>2</sub> VCDs are still mostly larger in the vis-blue spectral range. Compared to the SCD differences, however, the VCD differences between the two spectral ranges UV and the blue spectral range are reduced in all six regions (Fig. 15) which is also visible in the higher can also be seen in the larger slope of the regression line (Tab. 6). For nearly all regions and seasons, the slope of the regression line (Tab. 7) is still below one, indicating that the calculated AMFs are not representative for the actual state of the atmosphere. The SSR and the relative vertical profile as well as aerosols

10 representative for the actual state of the atmosphere. The SSR and the relative vertical profile as well as aerosol may contribute to the differences of VCDs between the two spectral ranges UV and the blue spectral range.



**Figure 14.** Absolute Relative difference between monthly mean tropospheric NO<sub>2</sub> VCDs in the UV and vis-blue spectral range. Differences for (a) January and (b) July 2008. Dark gray grey shaded area: no NO<sub>2</sub> values available. Light grey coloured values indicate values which are filter out. The same filter as for Fig. 8 is used.

In the ASE region, for example, the seasonal cycle in the VCDs is smoother than in the SCDs, especially for the UV data (Fig. 15 a). The seasonal cycles of the VCDs in the two spectral regions UV and the blue spectral range are more parallel, leading to a high correlation coefficient of 0.96-0.97 between the two data sets (see Tab. 6). However, rather than being separated by a simple offset, the VCDs from the two spectral ranges seem to be linearly related as indicated by the regression slopes differing from one (see Tab. 7 and Fig. 16). Compared to SCDs, the slope for the VCDs increases slightly in nearly every season; for the annual values, the slope of VCDs is 0.70-0.77 compared

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to 0.66 for SCDs. In the ANE region, the VCDs from the UV and vis spectral ranges blue spectral range agree considerably better than for the respective SCDs (Fig. 15b). The UV NO<sub>2</sub> VCDs show a clear seasonal cycle, of similar shape but

- 10 reduced amplitude compared to the vis<sub>blue</sub> VCDs, leading to reduced differences between the two spectral ranges. Comparing to the SCDs in the ANE region, this indicates that the biomass burning NO<sub>2</sub> seems to be located too close to the ground to be detected by the UV retrieval; however, as the NO<sub>2</sub> emitted by fires is included in the simulation from which the a-priori profiles are taken, the AMFs introduce the observed seasonal cycle in the UV VCDs. This is reflected by the considerably improved correlation coefficients of 0.91 for all VCDs (see Tab. 6) compared to
- 15 0.53 for all SCDs; similar improvements can be seen for all seasons individually. Interestingly, when the seasons are considered separately, the regression line has higher slopes between 0.72-0.79 (SON) and 0.98-1.06 (JJA), showing that in the rainy season, the UV and vis<sub>blue</sub> VCDs only differ by an offset. Nevertheless, discrepancies between the two spectral ranges are still observed, especially during biomass burning season.



Figure 15. Time series 2007 - 2015 of tropospheric NO<sub>2</sub> VCDs for the UV (green red line) and vis<sub>plue</sub> (blue line), and vis<sub>green</sub> (green line, China only) spectral range as well as the TM5-model VCDs (gray-grey line) for different regions. Note the different y-axes, the same as in Fig. 11. The vertical lines indicate January of the individual years. (a-c) biomass burning regions and (d-f) regions with anthropogenic air pollution.

Table 6. Correlation between UV  $NO_2$  VCDs and vis<sub>blue</sub>  $NO_2$  VCDs for the six selected regions.

| Region | Annual | DJF  | MAM  | JJA  | SON  |
|--------|--------|------|------|------|------|
| ASE    | 0.97   | 0.66 | 0.89 | 0.93 | 0.95 |
| ANE    | 0.91   | 0.75 | 0.92 | 0.98 | 0.87 |
| NAUS   | 0.95   | 0.92 | 0.78 | 0.93 | 0.94 |
| China  | 0.98   | 0.96 | 0.98 | 0.96 | 0.97 |
| HPSA   | 0.96   | 0.79 | 0.97 | 0.90 | 0.93 |
| Riyadh | 0.92   | 0.83 | 0.89 | 0.70 | 0.96 |

For NAUS, the correlation coefficient between UV and  $vis_{blue}$  VCDs is higher (0.940.95) than for the SCDs (0.78). Also in this area, the slope of the VCD regression lines is constant or increased compared to the SCDs (Tab. 7; Fig. 16). During rainy season, the observed differences between UV and vis<sub>blue</sub> VCDs are mostly related to an offset, whereas during biomass burning season the slope is smaller than one.

For VCD time series of over East China (Fig. 15d), the upward trend which was visible observed in the winter NO<sub>2</sub> SCDs (see Fig. 11 d) is not present in the VCDs. However, also here the UV NO<sub>2</sub> values are still smaller than

- the vis<sub>blue</sub>NO<sub>2</sub> values. In spite of these discrepancies, the correlation coefficient is highest for this region ( $\geq 0.96$ 5 in all seasons). During all seasons the correlation coefficients are increased or similar compared to those of the SCDs. While the slope of the regression line is still smaller than one (0.540.59 - 0.690.76; Tab. 7), it is still increased compared to SCDs. The green and blue spectral range show a good agreement with a correlation coefficient of 0.98 (see also Table S2). Furthermore, the slope (1.07) agrees well which shows that there are only smaller differences between these spectral ranges.
- 10

The Highveld (HSPA, Fig. 15e) region differs from all other regions: a high correlation coefficient (0.90) can be found in SH winter, when also the slope of the regression line (0.921.00) as well as the offset is high. In SH summer, however, the smallest correlation coefficient is observed for the region (0.780.79) and also the slope is the smallest with 0.690.77.

15 In Riyadh, large relative differences between the two-UV and blue spectral ranges can be found compared to the other regions (Fig. 15 f). This is reflected by the slope of the regression line which is  $\sim 0.650.75$ , except for NH winter when the slope is 1.09 - 1.19 with a large offset (Fig. 16 f).

The observed discrepancies between the two-UV and blue spectral ranges can result from different reasons: they can be related to the a-priori profiles used for VCD calculations, or they can be related to other influencing factors

- 20 like aerosols and surface properties. The importance of the a-priori profile shape for the resulting VCDs was shown by Martin et al. (2002) and Boersma et al. (2004); the resulting uncertainty for this parameter is assumed to be smaller than 15% (Boersma et al., 2004). A potential error in the a-priori profiles from the TM5-model in a given altitude layer will lead to different errors in the VCDs calculated for the two spectral ranges UV and the blue spectral range, due to the spectral dependence of the measurement sensitivity (i.e., the a-priori error is weighted
- with different sensitivities in the two spectral ranges). This ; see Fig. S18). Higher NO<sub>2</sub> concentrations close to 25 the surface combined with a lower injection height would reduce the differences between the UV and blue spectral range. The a-priori profile is especially relevant in biomass burning season, when the quality of the a-priori profiles is probably lower than usual, since the modeling modelling of biomass burning emissions entails large uncertainties due to spatial resolution, emission uncertainties, and assumptions made on the plume injection height (see, e.g.,
- Archer-Nicholls et al., 2015 and references therein). Similarly, since our retrieval does not explicitly account for the 30 effect of aerosols, an aerosol layer at a given altitude will influence the visibility of the  $NO_2$  both above and below the aerosol differently, enhancing the difference in the VCDs retrieved from the two spectral ranges. The same is true for potential errors in the assumed SSR. In order to bring UV and vis retrievals into better agreement, either the UV SSR would have to be lower than assumed or the vis SSR larger than assumed or a combination of both.



Figure 16. Scatter plot of tropospheric UV NO<sub>2</sub> VCDs against tropospheric vis<sub>blue</sub> NO<sub>2</sub> VCDs for the period 2007–2015 for the six selected regions. Note the different axes. The seasons are <u>color\_colour</u> coded. The related correlation, slope and intercept of the regression line can be found in Tab. 6 and 7. December, January, February: DJF; March, April, May: MAM; June, July, August: JJA; September, October, November: SON.

As shown by Boersma et al. (2004), in March the uncertainties of NO<sub>2</sub> concentration are largest in polluted areas with low SSR, but they are generally below 50% (see also Tab. 2).

Tropospheric model VCDs are plotted in Fig. 15, showing a regional dependency of the discrepancies between the model VCDs and the observed VCDs. For ASE and ANE during biomass burning season, the UV NO<sub>2</sub> VCDs and

5 the model VCDs agree quite well, while the vis<sub>blue</sub> NO<sub>2</sub> values are higher. It should however be kept in mind that in the vis<sub>blue</sub> retrieval, both the uncertainties and the contribution of the a priori are lower, making these results more reliable. In rainy season, especially in ANE, there are larger discrepancies between the UV and vis<sub>blue</sub> NO<sub>2</sub> and the model VCDs. The model values are lower though the UV and vis blue spectral ranges agree quite well. This confirms the above finding that the differences between the spectral ranges are caused by an offset. The differences 10 during biomass burning season might be related to the aerosol treatment or wrong SSR in AMF calculations.

Also for NAUS the simulated and observed values agree quite well in the rainy season, whereas in the biomass burning season large differences between the three values are observed. Therefore, in this regions a combination of the different factors, described above, will most likely provide the best explanation.

For China and Riyadh, in the less polluted season the model and UV  $NO_2$  values agree quite well whereas the 15 vis<sub>blue</sub>  $NO_2$  values are larger. In the season with high air pollution, in both regions large discrepancies between the

| Region | Annual           | DJF              | MAM              | JJA              | SON               |
|--------|------------------|------------------|------------------|------------------|-------------------|
| ASE    | 0.77/ 0.04       | 1.62/ -0.73      | 0.91/ -0.15      | 0.94/ -0.37      | 0.76/ $0.01$      |
|        | $(0.02/ \ 0.03)$ | $(0.30/ \ 0.25)$ | $(0.09/\ 0.08)$  | $(0.07/ \ 0.12)$ | $(0.05/ \ 0.08)$  |
| ANE    | 0.69/ $0.15$     | 0.90/ -0.29      | 1.04/ -0.24      | 1.06/ -0.21      | 0.79/ $0.06$      |
|        | $(0.03/ \ 0.04)$ | $(0.14/\ 0.23)$  | $(0.08/\ 0.11)$  | $(0.04/ \ 0.04)$ | $(0.08/ \ 0.09)$  |
| NAUS   | 0.81/ -0.07      | 0.98/ -0.24      | 1.12/ -0.30      | 0.94/ -0.12      | 0.78/ -0.08       |
|        | $(0.03/\ 0.02)$  | $(0.08/\ 0.09)$  | $(0.16/\ 0.11)$  | $(0.07/ \ 0.05)$ | $(0.06/ \ 0.07)$  |
| China  | 0.69/ -0.15      | 0.75/ -0.85      | 0.72/ -0.11      | 0.76/ -0.12      | 0.59/ $0.79$      |
|        | $(0.01/\ 0.16)$  | $(0.04/\ 0.78)$  | $(0.03/ \ 0.32)$ | $(0.04/ \ 0.29)$ | $(0.03/ \ 0.34)$  |
| HPSA   | 0.94/ -0.54      | $0.77/\ 0.02$    | 0.93/ -0.29      | 1.00/ -0.89      | 0.87/ -0.49       |
|        | $(0.03/ \ 0.11)$ | $(0.11/\ 0.31)$  | $(0.05/ \ 0.19)$ | $(0.09/ \ 0.48)$ | $(0.07 / \ 0.26)$ |
| Riyadh | 0.74/ -0.37      | 1.19/ -1.43      | 0.79/ -0.51      | 0.67/ -0.19      | 0.74/ -0.37       |
|        | $(0.03/ \ 0.09)$ | $(0.15/\ 0.36)$  | $(0.08/\ 0.20)$  | $(0.12/ \ 0.42)$ | $(0.04/ \ 0.12)$  |

Table 7. Slope and intercept between UV NO<sub>2</sub> VCDs and vis<sub>blue</sub> NO<sub>2</sub> VCDs for the six selected regions. Intercept in 1e15 molec cm<sup>-2</sup>. In the brackets the standard errors for both values are shown.

three values are visible observed. For anthropogenically polluted areas, the discrepancies between the modelled and the measured  $NO_2$  values are well known from previous studies (e.g., Huijnen et al., 2010). For these regions, also aerosols play an important role and can further influence the differences between the two spectral ranges. As shown by Russell et al. (2010), especially dust reduces the visibility in the UV spectral range which could partly explain the differences in Riyadh. Additionally, as shown by Boersma et al. (2004), for these anthropogenically polluted areas

5 differences in Riyadh. Additionally, as shown by Boersma et al. (2004), for these anthropogenically polluted areas also the SSR can introduce large differences. In China only small differences between the blue and green spectral range were found. The reason for the different behaviour between the two vis spectral ranges might be that the green spectral range is less sensitive to the a-priori profile compared to the UV and blue spectral range (see Fig. S18).

In HPSA, an offset between the model and the observed values can clearly be observed, which could partly 10 explain the differences in SH winter values (Tab. 7). However, also here the SSRs and aerosols might influence the measurements, and therefore, the calculation of AMF.

### 4 Summary and Conclusion

In this study, we present a new NO<sub>2</sub> DOAS retrieval for measurements by the GOME-2A instrument, using the ultraviolet (UV) spectral range. Having NO<sub>2</sub> measurements from this wavelength range in addition to the usually exploited visible (vis) blue spectral region gives rise to conclusions about insights into the vertical distribution of

5 tropospheric NO<sub>2</sub>, as the vertical sensitivity of space-borne NO<sub>2</sub> DOAS retrievals strongly depends on the wavelength. NO<sub>2</sub> slant column densities (SCDs) can in principle be retrieved from the UV spectral range, as shown by Yang et al. (2014). However, the UV retrievals have larger uncertainty compared to the vis-blue spectral range, caused by the smaller differential absorption and the reduced intensity in the UV spectral range, leading to inferior fit qualityand, and also due to the larger AMF uncertainties. Nonetheless, the spatial distribution of both retrievals

10 agrees very well.

We focused our analysis on six regions: three biomass burning regions (Africa south of the equator / ASE, Africa north of the equator / ANE, and northern Australia / NAUS), and as three anthropogenic source regions (Eastern China, the Highveld Plateau in South Africa / HPSA, and the surroundings of Riyadh).

- The differences between the UV and vis<sub>blue</sub> fits are lower over the biomass burning areas ASE, ANE, and NAUS compared to the regions dominated by anthropogenic pollution. The NO<sub>2</sub> from biomass burning is in some regions and periods located in elevated layers, leading to an effectively smaller sensitivity difference between UV and vis <u>blue</u> spectral ranges. This is reflected by the slope of the regression lines between UV and vis<sub>blue</sub> SCDs, which is 0.66 and 0.74 for the ASE and NAUS regions, respectively. Anthropogenic air pollution on the other hand is mostly located in the boundary layer close to the ground (see Stull, 1988). As the sensitivity of the UV retrieval drops sharply towards the surface, this leads to considerably lower regression line slopes of 0.37 and 0.41 for East China
- and HPSA, respectively. Nonetheless, SCDs in the UV and  $vis_{blue}$  show high correlation coefficients  $\geq 0.82$  in both cases.

In ANE, the biomass burning NO<sub>2</sub> in NH winter is not visible cannot be observed in the UV SCDs. Our analysis of UV NO<sub>2</sub> SCDs over cloudy scenes and of NO<sub>2</sub> profiles simulated by the  $\frac{\text{TM5-MP}}{\text{TM5-MP}}$  model both

- suggest that in ANE the NO<sub>2</sub> is located close to the ground, where sensitivity in the UV is limited. The invisibility non-detectability could be increased by enhanced the stratospheric NO<sub>2</sub> over the Pacific which counteracts with seasonal cycle, which counteracts the biomass burning NO<sub>2</sub> over ANE and therefore balance enhanced SCDssignal over ANE. Moreover, the ANE biomass burning NO<sub>2</sub> could be shielded by the predominantly dust aerosols, which are dark in the UV and thus lead to decreased sensitivity compared to other biomass burning regions. In ASE, for
- 30

example, the predominant aerosol type is smoke, which is brighter than dust in the UV spectral range; combined with the higher altitude of the  $NO_2$  this could lead leads to increased visibility in ASE.

In Riyadh, the seasonal cycles of boundary layer height and tropospheric  $NO_2$  are in phase, contrary to other anthropogenically polluted areas, where  $NO_2$  columns are highest in winter, when the boundary layer is <u>lowshallow</u>. Therefore, the maximum  $NO_2$  values coincide with high boundary layers and thus better visibility in the UV fit, leading to better agreement between SCDs from the two spectral regions, expressed by a higher slope of the regression line of 0.74.

Even though in theory, given perfect measurements, radiative transfer modelling, and a-priori information, vertical column densities (VCDs) retrieved by the DOAS technique should be independent on the wavelength, VCDs retrieved

- 5 from the UV and vis fits do still show considerable differences. If the a-priori profile does not capture the actual vertical distribution of NO<sub>2</sub> in the troposphere, the vertical measurement sensitivity will be incorrectly aggregated in the air mass factor calculation, which will lead to wavelength-dependent VCDs. One, admittedly unlikely, explanation would be if the model over-estimated vertical mixing, leading to an underestimation of the AMF UV/vis ratio. Having said this, the observed VCD differences can also be introduced by further input parameters like surface spectral
- 10 reflectance (SSR; wavelength: 354 nm and 442 nm) and the lack of explicit aerosol treatment. These parameters can lead to large discrepancies between the two spectral ranges when assumed wrongly. Even though differences exist between the two-UV and blue spectral ranges, the correlation coefficient between the VCDs is high everywhere (≥91%0.91). Generally, the differences between the two spectral ranges indicate that the NO<sub>2</sub> in the atmosphere is located in lower atmospheric layers than in the model which has stronger inferences in the UV or SSR is too high in the vis-blue spectral range.
  - While  $vis_{blue}$  NO<sub>2</sub> VCDs are often larger than UV VCDs, the latter agree with the simulated VCDs in most seasons. For China and HPSA, the differences between the two-UV and blue spectral ranges are pronounced in the respective winter season, whereas for Riyadh large differences can be found throughout the whole year. However in HPSA, in SH winter the correlation coefficient and the slope of the regression line is largest with an high offset.
- 20 In the biomass burning regions, the agreement is better for the rainy seasons; during dry season, when a larger aerosol load is expected, the differences are larger. After conversion into VCDs, in ANE a clear seasonal cycle can be observed in both spectral ranges, which is mostly introduced by the a-priori assumptions in the AMF calculation.

An additional fitting window in the green spectral range also shows a higher uncertainty compared to the blue spectral range which is related to the smaller absorption lines as well as interferences with liquid water and vegetation,

- as already reported by Richter and Burrows (2000). Nevertheless, over highly anthropogenically polluted areas, NO<sub>2</sub> retrieved from the green spectral range is even more sensitive to the lower troposphere compared to the blue spectral range. This can be clearly observed in the higher vis<sub>green</sub> SCDs over China compared with the UV and blue spectral range. In general, the SCDs of the three spectral ranges agree well with correlation coefficients  $\geq 0.98$ . The VCDs for the green and blue spectral show also a good agreement (correlation coefficient of 0.98 and slope of 1.07) which
- 30 might be related to the smaller sensitivity to the a-priori profile.

It is currently not possible to retrieve vertical profiles of tropospheric  $NO_2$  from satellite measurements, due to the low information content of the measured spectra (see, e.g., Hilboll et al., 2014). Further developments in space-based observation of the Earth's atmosphere could however increase the capability to retrieve vertical  $NO_2$  distribution. For example, concurrent measurements of the same air mass from different viewing geometries could yield insight

35 on aerosol types and vertical profiles which might be also possible for GOME-2A and GOME-2B observations. This

in turn would increase the possibility to retrieve more quantitative information about the vertical distribution of  $NO_2$  in the troposphere. Improvements in signal to noise ratio would reduce the uncertainty in the UV retrievals and thus increase the value of the UV slant columns.

However, our study shows that even for the current generation of instrumentation in the global observing system,
the difference between vis<sub>blue</sub> and UV NO<sub>2</sub> DOAS retrievals can be used to constrain our knowledge of NO<sub>2</sub> vertical distribution in the lower troposphere. For example when coupled with data assimilation techniques and high resolution models the amount of boundary layer NO<sub>2</sub> and its trends can be inferred more accurately.

### Appendix A: Shielding of northern hemispheric winter values in Africa north of the equator

#### A1 Shielding by clouds

- 10 Figure A1 shows NO<sub>2</sub> time series for cloud free scenes ( $\leq 20\%$  cloud fraction; a and c) and for cloudy scenes ( $\geq 30\%$  cloud fraction; c and d). The decrease/increase of NO<sub>2</sub> values in recent years visible shown in Fig. A1 is probably related to instrumental degradation (Dikty and Richter, 2011); a similar decrease/increase cannot be observed in GOME-2B data (not shown). In ASE, the seasonal cycles for cloud covered and cloud free measurements are similar. In ANE, however, only a weak seasonal cycle can be found for the cloudy scenes in either spectral range, whereas
- 15 in the cloud free scenes a clear seasonal cycle can be found in the vis-blue but not in the UV spectral range. This implies that the NO<sub>2</sub> layer over ASE could be located also in higher altitudes, while over ANE the NO<sub>2</sub> could be located closer to the ground and thus less visible sensitive to the UV observations.

### A2 NO<sub>2</sub> profile shape

To compare the vertical distribution of NO<sub>2</sub> over the biomass burning areas ANE, ASE, and NAUS, simulated NO<sub>2</sub>
concentration fields from the TM5-model dataset TM5-MP (Williams et al., 2017) for 2008 at 09:00 LT were used. We assume that the small time difference between the satellite measurements and the model output can be neglected. Figure A2 (a) shows profiles of NO<sub>2</sub> concentration for the regions ANE and ASE for four months. Apparently, NO<sub>2</sub> is located at higher altitudes in ASE, which leads to differences in the visibility of the NO<sub>2</sub> signal between the two spectral ranges. This is partly related to the surface altitude which is on average ~830 m above sea level in ASE and only ~440 m in ANE. This could explain why the UV fit shows enhanced NO<sub>2</sub> over ASE and not over ANE.

# A3 Seasonality of stratospheric NO<sub>2</sub>

Figure A3 shows time series of detrended total SCDs for the regions ANE and ASE. Detrending is necessary because of the degradation of the instrument (Dikty and Richter, 2011), and is implemented as subtraction of a linear trend. Here, a shortened time series from 2008 to 2012 is used, since changes in the instrument configuration (EUMETSAT, 2015) might introduce an offset after 2012



Figure A1. Tropospheric NO<sub>2</sub> SCD time series for ASE (a, c) and ANE (b, d) for cloud free scenes ( $\leq 20 \%$  cloud fraction; a, b) and for cloudy scenes ( $\geq 30 \%$  cloud fraction; c, d). The vertical lines indicate January of the individual years. (a, b) are identical to Fig. 11 a, b and are replicated here to facilitate direct comparison.



Figure A2. Average profiles of normalised NO<sub>2</sub> concentration for the three regions ANE and ASE for 2008, derived from the TM5-model TM5-MP for 09:00 LT. The gray-grey dashed lines indicating the surface height of the two regions.

In ASE a clear seasonal cycle can be seen for both spectral ranges (Fig. A3 a), which agrees well with the seasonal cycle expected from biomass burning activity. However, comparison with Fig. A3 c shows that a similar seasonal cycle can also be observed in SCD data over the Pacific Ocean (averaged over  $180^{\circ}-210^{\circ}$  E), which is void of any NO<sub>x</sub> emissions and should therefore only have very low tropospheric NO<sub>2</sub> SCDs. This indicates that the seasonality over ASE is at least partly caused by stratospheric fluctuations.

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Over ANE, slightly different seasonal cycles can be observed in the UV and vis blue spectral range (see Fig. A3 b). The peak related to biomass burning in NH winter is clearly observed in the vis<sub>blue</sub> SCD data. In some years a small secondary peak in NH summer can be noticed. Similar seasonal variability can be seen in the UV spectral range, but both peaks have similar magnitude there. This double peak structure can also be found in NO<sub>2</sub> SCDs of from the Ozone Monitoring Instrument (OML not shown). Over the Pacific Ocean (Fig. A3d), the second peak in NH

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the Ozone Monitoring Instrument (OMI, not shown). Over the Pacific Ocean (Fig. A3d), the second peak in NH summer can clearly be observed in the UV data, corresponding to the known seasonal variability of stratospheric  $NO_2$ . As the seasonal cycles of biomass burning and stratospheric  $NO_2$  are in-phase and out-of-phase for ASE and ANE, respectively, the visibility of the biomass burning  $NO_2$  differs strongly between the two regions.



Figure A3. Time series of detrended total NO<sub>2</sub> SCDs for the UV and  $\frac{\text{vis-blue}}{\text{vis-blue}}$  spectral range for ASE (a) and ANE (b) as well as detrended time series over the Pacific reference sector area for the latitude of ASE (c) and ANE (d). The vertical lines indicate January of the individual years.

#### A4 Aerosol effect

The Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observations (CALIPSO) satellite is in a sun-synchronous polar orbit with an equator crossing time of 13:30 LT (Winker et al., 2007). It was launched in April 2006 and has a repeat cycle of 16 days. On CALIPSO, the Cloud-Aerosol Lidar with Orthogonal Polarisation (CALIOP) instrument

5 is operating since June 2006. CALIOP is a nadir-viewing two wavelength polarization-sensitive lidar, operating at 532 nm and 1064 nm. The horizontal and vertical resolution depends on the altitude. Closer to the ground the resolution increases in both cases (Winker et al., 2007; Winker et al., 2009).

Here, the CALIOP level 3 product "Aerosol Profile All Sky (daytime)" was used. The level 3 data are on a 2° latitude times 5° longitude grid (Atmosheric Science Data Center, 2009). For the calculation of regional aerosol contribution for smaller regions, the amount of aerosols in the grid box was weighted by the contribution of the

grid box to the region. This dataset provides the six aerosol types "dust", "polluted continental", "smoke", "clean marine", "clean continental", and "polluted dust", where the latter accounts for a mixture of dust and smoke or a mixture of dust and urban pollution (Omar et al., 2009). The aerosol-types were derived using aerosol models based on cluster analysis of an AERONET dataset and the aerosol extinction-to-backscatter ratio (Omar et al., 2009).
15 In the present study, the differences in equator crossing time can be ignored, because especially for ASE and ANE

larger areas are averaged.

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Figure A4 shows the sample number for the six aerosol types over the three regions ANE, ASE, and Riyadh. The dominant aerosol-type in ASE is smoke whereas dust dominates in ANE and Riyadh.



Figure A4. Time series for the number of samples of aerosol-types for three different regions (CALIPSO data). To calculate the time series for the individual areas, the amount of aerosol in the grid box was weighted by the contribution of the grid box to the region. Note the different y-axes. The vertical lines indicate January of the individual years.

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# GOME-2A retrievals of tropospheric $NO_2$ in different spectral ranges - influence of penetration depth

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Abstract. In this study, we present a novel NO<sub>2</sub> DOAS retrieval in the ultraviolet (UV) spectral range for observations from the Global Ozone Monitoring Instrument 2 on board EUMETSAT's MetOp-A (GOME-2A) satellite. We compare the results to those from an established NO<sub>2</sub> retrieval in the visible (vis) spectral range from the same instrument and investigate, how differences between the two are linked to the NO<sub>2</sub> vertical profile shape in the troposphere

5 troposphere.

As expected, radiative transfer calculations for satellite geometries show that the sensitivity close to the ground is higher in the vis than in the UV spectral range. Consequently,  $NO_2$  slant column densities (SCDs) in the vis are usually higher than in the UV, if the  $NO_2$  is close to the surface. Therefore, these differences in  $NO_2$  SCDs between the two spectral ranges contain information on the vertical distribution of  $NO_2$  in the troposphere. We combine

10 these results with radiative transfer calculations and simulated NO<sub>2</sub> fields from the TM5-MP chemistry transport model to evaluate the simulated NO<sub>2</sub> vertical distribution.

We investigate regions representative of both anthropogenic and biomass burning NO<sub>2</sub> pollution. Anthropogenic air pollution is mostly located in the boundary layer close to the surface, which is reflected by large differences between UV and vis SCDs of ~ 60 %. Biomass burning NO<sub>2</sub> in contrast is often uplifted into elevated layers above

15 the boundary layer. This is best seen in tropical Africa south of the equator, where the biomass burning NO<sub>2</sub> is well observed in the UV, and the SCDs difference between the two spectral ranges is only  $\sim 36$  %. In tropical Africa north of the equator, however, the biomass burning NO<sub>2</sub> is located closer to the ground, reducing its visibility in the UV.

While not enabling a full retrieval of the vertical NO<sub>2</sub> profile shape in the troposphere, our results can help to 20 constrain the vertical profile of NO<sub>2</sub> in the lower troposphere and, when analysed together with simulated NO<sub>2</sub> fields, can help to better interpret the model output.

# 1 Introduction

25

Nitrogen dioxide  $(NO_2)$  is an important indicator for natural phenomena and anthropogenic air pollution, as it is produced in the troposphere by, e.g., biomass burning or combustion of fossil fuels (Lee et al., 1997). Additionally,  $NO_2$  is produced by lightning (Lee et al., 1997; Beirle et al., 2004) and microbiological activity in soils (Williams

- 5 et al., 1992; Bertram et al., 2005). The relevance of the individual NO<sub>2</sub> sources depends on the region of the Earth (van der A et al., 2008). Biomass burning is important in equatorial regions like central Africa, whereas anthropogenic emissions are mostly important in the industrialised mid-latitudes like China or Europe. The distribution of NO<sub>2</sub> is of major interest because it is harmful to human health, adds to local radiative forcing, catalyses surface ozone production during summer smog, and causes acid rain (Finlayson-Pitts and Pitts, 1999).
- 10 Using spectrometers, solar radiation scattered in the Earth's atmosphere can be measured and the amount of trace gases inverted mathematically from the depth of molecular absorption bands. Such measurements have been made from satellite (e.g., Burrows et al., 1999), airborne (e.g., Heue et al., 2005; Wang et al., 2005), and ground-based (Noxon, 1975) platforms.
- Since hyperspectral satellite observations began in the mid-1990s (Burrows et al., 1999), the horizontal distribution of the column amounts of many trace gases is well known, e.g., for NO<sub>2</sub>. Using sun-synchronous orbits, a nearly daily global coverage at similar local times can be obtained. The global coverage as well as the spatial resolution of the measurements depends on the instrument (see for example: Callies et al., 2000). However, there is only limited knowledge of the vertical distribution from satellite data. Compared to satellite observations, ground-based Multi AXis Differential Optical Absorption Spectroscopy (MAX-DOAS) measurements have high temporal resolution and
- 20 can provide vertical profiles of trace gases up to several kilometres altitude (Wittrock et al., 2004), but can only measure at one particular location and are thus limited in resolving horizontal gradients.

Irrespective of the instrument's viewing geometry, the measured spectra can be analysed using the Differential Optical Absorption Spectroscopy (DOAS; Platt and Stutz, 2008; Burrows et al., 2011) technique which is a well established method based on Lambert Beer's law describing the spectral reduction of the initial intensity of light due to absorption. The main result is the integrated concentration of trace gases  $\rho(s)$  along the effective light path

s which is called the total slant column density (SCD; Platt and Stutz, 2008; Burrows et al., 2011):  

$$SCD = \int \rho(s) ds.$$
(1)

As Rayleigh scattering in the atmosphere depends strongly on wavelength (λ; the scattering cross-section is proportional to λ<sup>-4</sup>), sunlight having longer wavelengths can penetrate deeper into the atmosphere compared to
30 shorter wavelengths. This results in a wavelength-dependency of the SCDs (Burrows et al., 2011). Consequently, if NO<sub>2</sub> is located close to the ground, SCDs retrieved from satellite measurements at shorter wavelengths are

in theory smaller than SCDs retrieved from longer wavelengths. In consideration of this fact, knowledge of the vertical distribution of  $NO_2$  can be gained by combining measurements at different wavelengths. The idea of using the penetration depth in the UV to determine trace gas vertical profiles was first proposed for ozone by Singer and

Wentworth (1957). The use of the temperature dependence of the Huggins absorption bands coupled with penetration depth was proposed to retrieve information about the vertical profile of ozone in the troposphere (Chance et al., 1997). Here, will use a similar method for the optically thin gas  $NO_2$ .

In the presence of clouds, the described behaviour changes. For cloudy scenes only smaller differences are expected. 5 Due to the clouds the sensitivity towards the surface is decreased, and therefore, the NO<sub>2</sub> below the clouds is partly invisible to the instrument. For the NO<sub>2</sub> above the clouds, the sensitivity is similar or partly increased for the UV spectral range (Burrows et al., 2011). In previous studies, clouds at different altitudes have been used to obtain information about NO<sub>2</sub> profiles (Choi et al., 2014; Belmonte Rivas et al., 2015).

The vertical column densities (VCDs) are the integral of the trace gas concentration from the surface to the top 10 of the atmosphere along the altitude z (Platt and Stutz, 2008; Burrows et al., 2011):

$$VCD = \int \rho(z)dz.$$
(2)

They can be calculated using air mass factors (AMFs), which are defined as the ratio of SCDs and VCDs (Platt and Stutz, 2008; Burrows et al., 2011), and are an indicator of the measurement sensitivity or the length of the light path within the  $NO_2$  layer relative to the vertical path:

15 
$$AMF(\lambda) = \frac{SCD(\lambda)}{VCD}$$
. (3)

AMFs are calculated by radiative transfer models, which take into account the viewing geometry and environmental effects (Platt and Stutz, 2008). The sensitivity of the measurement to an absorber varies with altitude; this is expressed by the so-called box air mass factor (BAMF; Burrows et al., 2011), which is defined as BAMF<sub>i</sub> = SCD<sub>i</sub>/VCD<sub>i</sub> for an atmospheric layer *i*. For wavelengths in the UV, the BAMF in layers close to the ground is considerably smaller than for the vis spectral range (Fig. 1). This reduction in sensitivity to lower atmospheric layers is further reduced for longer wavelengths. Figure 1 shows an example for surface type soil which results in a stronger altitude dependency compared to other surface types e.g., vegetation. In general, BAMFs for longer wavelengths have a smaller dependency on altitude compared with BAMFs for shorter wavelengths (Burrows et al., 2011).

Furthermore, the surface spectral reflectance (SSR) depends on the wavelength, and therefore, the wavelength dependence of the SCDs is influenced by the SSR (Burrows et al., 2011). Generally, for the UV and vis spectral range the SSR is quite low between 2 and 30% depending on the surface type except for snow or ice (Burrows et al., 2011). For this kind of surface types, the SSR is lower in the UV than in the vis spectral range. For smaller SSR, the UV sensitivity shows a stronger decrease and for larger SSR (e.g. snow) a stronger increase towards the surface, resulting in corresponding SCD decreases or increases compared to the vis SCDs. Furthermore, the altitude

30 of highest sensitivity depends on the SZA. For increasing SZAs, the altitude of highest sensitivity moves upwards to the stratosphere. Additionally, aerosols influence the measurements and also the visibility of NO<sub>2</sub> is influenced by the presence of aerosols (Burrows et al., 2011). Depending on the type and the optical thickness of aerosols the influences on the measurement differs.



**Figure 1.** BAMF for UV (red line), green (green line), and blue (blue line) spectral range and the ratio of UV and green BAMFs calculated with the radiative transfer model SCIATRAN. BAMFs converge at higher altitude (not shown). The BAMFs are calculated for 352 nm (UV), for 438 nm (vis<sub>blue</sub>), and for 515 nm (vis<sub>green</sub>) at (a, b) 30° (c, d) 50° and at (e, f) 70° SZA. A surface spectral reflectance of 0.04 (352 nm), 0.06 (438 nm), and 0.09 (515 nm) representative for bare soil is assumed.

Another parameter needed for the AMF calculation is an a-priori  $NO_2$  profile (Burrows et al., 2011), as the total AMF is calculated as the average of the BAMFs of all atmospheric layers, weighted by the absorber concentration. The retrieved VCDs therefore depend on the a-priori  $NO_2$  profiles, and differences between the a-priori and actual  $NO_2$  profiles can introduce systematic errors in the VCDs. As in principle, the final VCDs should not depend on the wavelength, analysing the differences in the VCDs retrieved from different wavelength regions allows to infer the validity of a-priori  $NO_2$  profiles.

Typically,  $NO_2$  DOAS fitting windows used in satellite data analysis are in the blue spectral range (see for example: Richter et al., 2011 and van Geffen et al., 2015). One satellite  $NO_2$  retrieval using UV wavelengths has been developed for the Ozone Mapper and Profiler Suite (OMPS) on board of the Suomi National Polar-orbiting

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10 Partnership (SNPP) satellite, employing a DOAS-like method to derive NO<sub>2</sub> VCDs (Yang et al., 2014). Compared to VCDs retrieved in the blue spectrum with DOAS from the Ozone Monitoring Instrument (OMI) on board the Aura satellite, Yang et al. found good agreement between the two retrievals.

Richter and Burrows (2000) introduced also a  $NO_2$  fitting window in the UV as well as in the green spectral range to retrieve tropospheric VCDs from GOME measurements. They conclude that it is possible to retrieve  $NO_2$ from both spectral ranges and that stratosphere and troposphere can be separated by combining the two retrievals. However, the green spectral range is strongly influenced by liquid water absorption and interferences with vegetation and therefore, a reliable global retrieval of  $NO_2$  was therefore not possible from the green spectrum.

In the present study, we introduce a DOAS retrieval for  $NO_2$  in the UV spectral range and compare the results with a DOAS retrieval in the blue spectral range, which is our standard approach, to investigate the vertical sensitivity of satellite-based NO<sub>2</sub> measurements. Furthermore, we present a case study for the green spectral range for China. In Sect. 2, the NO<sub>2</sub> DOAS retrieval in the UV spectral range will be introduced for measurements from the Global

10 Ozone Monitoring Experiment 2 (GOME-2) on board MetOp-A. The UV NO<sub>2</sub> retrieval settings will be discussed in detail and the dataset from the blue and green spectral range will be presented. In Sect. 3, results for the two/three  $NO_2$  retrievals in the UV, green, and blue spectral ranges will be compared, and their implications for the  $NO_2$ vertical distribution will be discussed. The manuscript ends with a summary and conclusions in Sect. 4.

#### $\mathbf{2}$ Datasets and methods

#### The GOME-2A instrument 15 2.1

Different types of remote sensing instrumentation on satellite platforms yield spectral radiance at the top of the atmosphere, which contains information on the amounts and distribution of  $NO_2$  in the Earth's atmosphere. Here, we use the GOME-2 instrument (Callies et al., 2000; Munro et al., 2016) on board the MetOp-A satellite (hereafter GOME-2A), which has a swath width of 1920 km, resulting in nearly global coverage every day. MetOp-A was

- 20 launched in October 2006 into a sun-synchronous orbit with an equator crossing time of 09:30 local time (LT) in descending node. The GOME-2A instrument is a nadir viewing spectrometer with a ground pixel size of  $80 \times 40 \text{ km}^2$ (Callies et al., 2000). In July 2013, the ground pixel size was reduced to  $40 \times 40 \text{ km}^2$ , when the identical GOME-2 instrument on board MetOp-B (hereafter GOME-2B), launched in September 2012, took over operational measurements (Munro et al., 2016). The spectrometer is separated into four channels covering wavelengths from 240 to
- 25 790 nm (Callies et al., 2000). Channel 2 provides data in the UV spectral range from 311 to 403 nm with a spectral resolution of 0.26 to 0.28 nm, while channel 3 provides data in the vis spectral range from 401 to 600 nm with a spectral resolution of 0.44 to 0.53 nm. These two channels are used for the following analysis.

#### NO<sub>2</sub> DOAS retrieval in the UV spectral range $\mathbf{2.2}$

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For this study, we developed a NO<sub>2</sub> DOAS retrieval for the GOME-2A instrument in the UV which uses a fitting window between 342 and 361.5 nm, and a polynomial of degree four. In this retrieval, one NO<sub>2</sub> cross section measured 30 with the GOME-2A instrument before launch, as well as two  $O_3$  cross sections are used with an additional  $I_0$  correction of  $10^{20}$  molec cm<sup>-2</sup> (Platt et al., 1997; Richter, 1997). Additionally, cross sections for O<sub>4</sub>, BrO, HCHO, the Ring effect and the instrumental cross section Zeta are included in the fitting procedure (see Tab. 1).

| UV spectral range |   | $vis_{blue}$ spectral range   | $vis_{green}$ spectral range |  |  |
|-------------------|---|---|------------------------------|--|--|
| fitting window    | $342 - 361.5{ m nm}$  | $425-450\mathrm{nm}$  | $490 - 540  { m nm}$         |  |  |
| polyn. degree     | 4   | 4   | 7                            |  |  |
| cross sections    |   |   |                              |  |  |
| $NO_2$            | $223 \mathrm{K}; \mathrm{G\ddot{u}r} \mathrm{et} \mathrm{al.} (2005)$ | $243 \mathrm{K}; \mathrm{G\ddot{u}r} \mathrm{et} \mathrm{al.} (2005)$ | 243 K; Gür et al. (2005)     |  |  |
| 0                 | $223 \mathrm{K}$ and $243 \mathrm{K}$ ;                               | 222 K. Cür et al. (2005)  | $223\mathrm{K};$             |  |  |
| $O_3$             | Serdyuchenko et al. (2014)  | 225 K; Gur et al. (2005)  | Serdyuchenko et al. (2014)   |  |  |
| $O_4$             | Greenblatt et al. (1990)  | Greenblatt et al. (1990)  | Greenblatt et al. (1990)     |  |  |
| $H_2O$            | _   | Rothman et al. $(2010)$   | Rothman et al. (2010)        |  |  |
| BrO               | Wilmouth et al. (1999)  | _   | _                            |  |  |
| НСНО              | Meller and Moortgat (2000)  | _   | _                            |  |  |
| Ring              | calculated with SCIATRAN, Vountas et al. (1998)                       |   |                              |  |  |
| instr. func.      | Zeta; EUMETSAT (2011)   | _   | _                            |  |  |

Table 1. Fit settings for the  $NO_2$  retrievals in the UV and the vis spectral range.

These settings are the result of a number of careful sensitivity tests. Among many different wavelength windows we tested, the selected window from 342-361.5 nm provided the smallest root mean squared error (RMS) of the fit residual. The instrumental correction function Eta (EUMETSAT, 2011) only had minor influence on the results when included in the fit; we therefore chose not to use it to keep the number of fit parameters small. Including additional O<sub>3</sub> cross sections to correct the O<sub>3</sub> non-linearity (Pukīte et al., 2010) improved the fit in polar latitudes, where solar angles and thus O<sub>3</sub> absorptions are large. However, this also introduced an offset in the data, and as we are mainly interested in polluted mid-latitudes, we chose not to use it here. Finally, we investigated the effect

10 of using a daily Earth reflectance spectrum as reference instead of a daily solar reference. The differences between results using the two reference spectra were only minor so that we use the daily solar reference in order to preserve consistency with the fit in the blue wavelength region (see Sect. 2.3).

## 2.3 NO<sub>2</sub> DOAS retrieval in the visible spectral range

Due to the considerably stronger differential absorption lines in the blue compared to the UV spectral range (Fig. 2), NO<sub>2</sub> DOAS retrievals normally use measurements in the blue spectral range, see for example Richter et al. (2011) or van Geffen et al. (2015). Here for the blue spectral range, a retrieval from 425-450 nm with a polynomial degree

5 of four is used (in the following vis<sub>blue</sub>; introduced in Burrows et al., 1999 for the GOME instrument, applied to GOME-2A measurements in Richter et al., 2011). The fit settings are summarized in Tab. 1. Cross sections for  $NO_2$  and  $O_3$  are used for this retrieval. Additional cross section are  $O_4$ ,  $H_2O$ , and the Ring effect.



Figure 2.  $NO_2$  absorption cross section measured at 243 K with the GOME-2 instrument now deployed on MetOp-A. The fitting windows used in this study are shaded by colour. Red: new  $NO_2$  fitting window in the UV spectral range. Green:  $NO_2$  fitting window in the green spectral range. Blue:  $NO_2$  fitting window in the blue spectral range (see Tab. 1).

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Furthermore, a second fitting window in the vis spectral range from 490-540 nm with a polynomial degree of seven is used (in the following vis<sub>green</sub>; introduced in Richter and Burrows, 2000 for the GOME instrument). The cross sections used in this fit are NO<sub>2</sub>, O<sub>3</sub>, O<sub>4</sub>, H<sub>2</sub>O, and the Ring effect. The fit settings are shown in Tab. 1. The high polynomial degree of seven is chosen in order to correct for interferences with the surface and liquid water absorption. Including a liquid water cross section for the green spectral range introduces errors over land; therefore, no liquid water cross section is used in the NO<sub>2</sub> fit for the green spectral range.

# 2.4 Comparison of the NO<sub>2</sub> DOAS retrievals in the UV and visible spectral range

- 15 In the UV and green spectral range, the differential NO<sub>2</sub> absorption lines are weaker than in the blue spectral range. In Fig. 2 the NO<sub>2</sub> cross section (Gür et al., 2005) at 243 K measured with the GOME-2A instrument is shown. The wavelength dependency of the differential absorption strength is clearly observed. Furthermore, the measurement sensitivity for NO<sub>2</sub> decreases towards the surface. This can be clearly observed in the BAMF (Fig. 1). This effect is enhanced by the temperature dependency of the NO<sub>2</sub> cross section which is not taken into account in the DOAS
- 20 fit. The temperature dependency influences the tropospheric as well as the stratospheric NO<sub>2</sub> measurements (see Fig. S4 and Fig. S5). For the stratospheric NO<sub>2</sub> the combined radiative transfer and temperature dependency leads to a 10% higher sensitivity in the UV spectral range than in the blue spectral range. Close to the surface, the

temperature sensitivity is up to 10% stronger in the blue spectral range compared to the UV spectral range. For the green spectral range, the differences are less pronounced compared to the blue spectral range (Fig. S6 and Fig. S7). The green spectral range has a slightly reduced temperature sensitivity (up to 3%) close to the ground and an enhanced sensitivity (up to 9%) in the stratosphere compared to the blue spectral range. Additionally, the

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temperature vertical sensitivity introduces a seasonal and a latitudinal dependency. This effect is stronger in the tropics than for higher/lower latitudes and in the mid-latitudes it is more pronounced in summer and less in winter. The smaller absorption lines in the UV in combination with reduced intensity from the sun and the weaker signal from the surface leads to a considerably higher noise level in the NO<sub>2</sub> differential optical depths in the UV spectral range compared to the blue. Similarly, in the green spectral range the absorption lines are smaller than in the blue spectral range and therefore, also the green spectral range has a higher noise level than the blue spectral range.

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spectral range and therefore, also the green spectral range has a higher noise level than the blue spectral range. To illustrate this, Fig. 3 shows an example of the fit results from different wavelength regions for one measurement above Teheran. As expected, the retrieved  $NO_2$  SCD is smaller in the UV spectral range compared to the vis.



Figure 3. The NO<sub>2</sub> reference (differential NO<sub>2</sub> cross section multiplied with retrieved NO<sub>2</sub> SCD; solid line) and reference plus residual (dashed line) for the (a) UV, (b) blue, and (c) green spectral range for one pixel above Teheran (35.38°N, 51.47°E) on 22 January 2008. The SCD for this pixel for the UV spectral range is  $6.31 \times 10^{16}$  molec cm<sup>-2</sup> with a fit error of 4.3%, for the blue spectral range it is  $9.33 \times 10^{16}$  molec cm<sup>-2</sup> with a fit error of 0.8%, and for the green spectral range is  $9.88 \times 10^{16}$  molec cm<sup>-2</sup> with a fit error of 0.9%. Note the different y-axes.

Due to the noisier differential optical depth signal, the UV NO<sub>2</sub> fit has larger uncertainty compared to the vis<sub>blue</sub> or the vis<sub>green</sub> NO<sub>2</sub> fit. The UV NO<sub>2</sub> fit of the example has a random error of 4.3%, whereas the vis<sub>blue</sub> NO<sub>2</sub>
15 fit has a smaller error of 0.8%. The NO<sub>2</sub> fit in the green spectral range has only a slightly higher error than the blue fit with 0.9%. Generally, the errors for the spectral ranges differ (see Tab. 2). For the estimation of the errors for the total SCDs over an assumed clean region (equatorial Pacific: 5°S to 5°N and from 150°E to 210°E) are calculated. In the blue spectral range, the standard deviation is smallest with 0.6 × 10<sup>15</sup> molec cm<sup>-2</sup> followed by the green spectral range with 1.4 × 10<sup>15</sup> molec cm<sup>-2</sup>. The highest scatter is found for the UV spectral range with 1.8 × 10<sup>15</sup> molec cm<sup>-2</sup>. This can also be seen in Fig. 4 which shows the distribution of total NO<sub>2</sub> VCDs over the selected region. Here, a stratospheric AMF was applied to calculate the NO<sub>2</sub> VCDs, as the common assumption

is that no or very little tropospheric NO<sub>2</sub> is present in this area (Richter and Burrows, 2002; Martin et al., 2002; Peters et al., 2012). Due to the lower fit quality in the UV, the NO<sub>2</sub> columns retrieved there have a larger standard deviation of  $7.4 \times 10^{14}$  molec cm<sup>-2</sup> compared to the vis<sub>blue</sub> with a standard deviation of  $2.1 \times 10^{14}$  molec cm<sup>-2</sup>. For the green spectral range, a standard deviation of  $5.8 \times 10^{14}$  molec cm<sup>-2</sup> is found. The green fitting window shows a bias to slightly more negative values over the Pacific Ocean which is related to liquid water absorption (not shown; see Sect. 2.3).

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|                       | errors UV  | $\operatorname{errors}\operatorname{vis}_{\operatorname{blue}}$  | $\operatorname{errors}\operatorname{vis}_{\operatorname{green}}$ |  |  |
|-----------------------|--|--|--|--|--|
| total SCDs            | $1.8\times10^{15}\mathrm{molec}\mathrm{cm}^{-2}$                             | $0.6 \times 10^{15} \mathrm{molec} \mathrm{cm}^{-2} \qquad 1.4 \times 10^{15} \mathrm{molec} \mathrm{cm}^{-2}$ |  |  |  |
|                       | (STD calcu   | lated above the Pacific Ocean,   | see Fig. 4)  |  |  |
| total VCDs            | $7.4\times10^{14}\mathrm{moleccm^{-2}}$                                      | $2.1\times10^{14}\mathrm{moleccm^{-2}}$  | $5.8\times10^{15}\mathrm{moleccm^{-2}}$                          |  |  |
|                       | (see Fig. 4)   |  |  |  |  |
| SSR                   | $40\%$ at $320\mathrm{nm},$  | 5% at 500 nm,  |  |  |  |
|                       | Kleipool et al., 2008  | Kleipool et al., 2008  |  |  |  |
| 0.05, increase $0.01$ | BAMF increases $9\%$   | BAMF increases $11\%$  | _  |  |  |
|                       | $(338\mathrm{nm},\mathrm{Lorente}\mathrm{et}\mathrm{al.},$                   | $(440\mathrm{nm},\mathrm{Lorente}\mathrm{et}\mathrm{al.},$   |  |  |  |
|                       | 2017)  | 2017)  |  |  |  |
| AMF                   |  |  |  |  |  |
| cloud fraction        | 0-30%, Boersma et al., 2004  |  |  |  |  |
| cloud height          | <10%,Boersma et al., 2004  |  |  |  |  |
| aerosols              | not included in our calculations $\rightarrow$ 15%, Boersma et al., 2004     |  |  |  |  |
| profile shape         | < 15% (regions with little NO <sub>2</sub> : $> 50%$ ), Boersma et al., 2004 |  |  |  |  |

Table 2. Errors for the UV and vis spectral range.

Furthermore, for VCD calculation several input parameters are needed (see also Sect. 2.7) which have further uncertainties. Several previous studies calculated errors for DOAS retrievals, as for example Boersma et al. (2004) and Lorente et al. (2017). Their results are summarized in Tab. 2. Additionally, the errors of the SSR differ strongly between the UV ( $\sim 40\%$ ) and vis spectral range ( $\sim 5\%$ ) which increases the uncertainty in the UV stronger (Kleipool

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and Lorente et al. (2017). Their results are summarized in Tab. 2. Additionally, the errors of the SSR differ strongly between the UV ( $\sim 40\%$ ) and vis spectral range ( $\sim 5\%$ ) which increases the uncertainty in the UV stronger (Kleipool et al., 2008). Overall, the uncertainties for the NO<sub>2</sub> concentration due to the SSR can be large, but in total it is smaller than 50% (Boersma et al., 2004)



Figure 4. Distribution of total NO<sub>2</sub> VCDs over a clean region (equatorial Pacific:  $5^{\circ}S - 5^{\circ}N$  and  $150^{\circ}E - 210^{\circ}E$ ) for three fitting windows in the UV, green, and blue spectral range for January 2008. Curves are normalized to unit area and centred on zero.

Generally, all three NO<sub>2</sub> retrievals lead to similar SCD spatial patterns. In Fig. 5, NO<sub>2</sub> SCDs for the UV, blue and green spectral range are shown for one orbit passing above East China on 03 January 2008. The larger uncertainty of the UV and vis<sub>green</sub> NO<sub>2</sub> retrieval is reflected in this figure by the larger spread of the slant columns and the existence of unphysical negative values. Nevertheless, the NO<sub>2</sub> columns in the UV spectral range and both vis spectral ranges are comparable for background regions and show a similar dependency on latitude. Over highly polluted regions, as for example China ( $25^{\circ}N-50^{\circ}N$ ), the vis NO<sub>2</sub> values are larger, especially the green spectral range, compared to the NO<sub>2</sub> values in the UV spectral range.





Figure 5. Total NO<sub>2</sub> SCDs of one orbit (above East China on 03 January 2008) for three fitting windows in the UV, blue, and green spectral range. Only data with SZA smaller  $70^{\circ}$  are shown.

# 2.5 Stratospheric NO<sub>2</sub>

Stratospheric  $NO_2$  shows large differences between the three spectral ranges. Figure 6 shows the latitudinal and seasonal dependency for the three  $NO_2$  fitting windows. The seasonal dependency clearly differs between the three fitting windows also over regions dominated by stratospheric  $NO_2$ , especially for the green wavelength range, strong

5 interferences are observable. Although the differences are smaller between the UV and blue spectral range, they are clearly observable, for example at the equator (Figure 6 b). In northern hemispheric summer the UV NO<sub>2</sub> values are higher than the NO<sub>2</sub> values derived from the blue spectral range whereas in northern hemispheric winter the NO<sub>2</sub> values from the blue spectral range are slightly higher. Consequently, the differences between the UV and blue spectral range have a clear seasonal dependency and therefore, it is currently not possible to consistently improve





Figure 6. Time series of NO<sub>2</sub> total VCDs of 2008 for 30°S (averaged:  $28^{\circ}S - 32^{\circ}S$ ), 0°N (averaged:  $2^{\circ}S - 2^{\circ}N$ ), and 30°N (averaged:  $28^{\circ}N - 32^{\circ}N$ ) above the area of the reference sector for cloud free pixels.

# 2.6 Conversion to tropospheric NO<sub>2</sub> columns

For the comparison of tropospheric SCDs and VCDs, a correction for the impact of stratospheric NO<sub>2</sub> on the measured SCDs is needed. For the SCDs, we use the "reference sector method" (Richter and Burrows, 2002; Martin et al., 2002) for the three spectral ranges separately, in which a monthly average of SCDs measured over a presumably
clean area above the Pacific (180° E to 210° E) is subtracted from all measurements per latitude band. The underlying assumption for this simple correction method is that no NO<sub>2</sub> is present over the reference sector in the troposphere, and that stratospheric NO<sub>2</sub> is zonally homogeneously distributed (Richter and Burrows, 2002; Martin et al., 2002). However, as this assumption is not always valid (Richter and Burrows, 2002; Boersma et al., 2004; Hilboll et al., 2013b), the reference sector method leads to areas with negative tropospheric NO<sub>2</sub> values, which are related to zonal

20 inhomogeneities in stratospheric  $NO_2$ , for example close to the polar vortex (Dirksen et al., 2011). For the reference sector method, the same cloud screening as for the data selection is used (see Sect. 2.8). As the sensitivity to the

stratosphere is similar in the three spectral ranges (see Sect. 2.7), the reference sector method is not expected to introduce systematic differences.

For the VCDs, we use a more sophisticated correction method, in which stratospheric VCDs from the Bremen-3D chemistry transfer model (B3dCTM, see Hilboll et al., 2013b and references therein) are used for the stratospheric correction. Therefore, no significantly negative values are expected for the VCDs assuming the model is correct. A Lambertian surface is assumed and the Lambert-equivalent reflectivity is taken as the SSR. For the AMF, a cloud correction is applied using the independent pixel approximation and cloud radiance fractions derived from the Fast

REtrieval Scheme for Clouds from the Oxygen A-band (FRESCO+, version 6; Wang et al., 2008) dataset.

#### 2.7 Radiative transfer simulations

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- 10 The vertical sensitivity of the measurements to NO<sub>2</sub> has been investigated using BAMFs simulated for the UV (352 nm), the blue (438 nm), and the green (515 nm) fitting windows by the radiative transfer model SCIATRAN (version 3.6.5, see Rozanov et al., 2014). To calculate the radiance, SCIATRAN requires knowledge about the measurement scenario, e.g., viewing geometry, solar position, atmospheric absorbers, SSR, surface elevation, aerosols, and clouds as input parameters. Here, SSR and surface elevation have been taken from the OMLER\_V003 (5-year
- 15 climatology 2005-2009; wavelength: 354 nm, 442 nm, and 499 nm; Kleipool et al., 2008) and GMTED2010 (Danielson and Gesch, 2011) datasets, respectively. Exemplarily, results for 30°, 50° and 70° solar zenith angle (SZA) and SSR of 0.04 (352 nm), 0.06 (438 nm), and 0.09 (515 nm) are shown in Fig. 1. The SSR differs between the spectral ranges; the SSR values are chosen to be representative of soil which shows the strongest differences in BAMF close to the ground. For 50° SZA, the sensitivity at the surface in the blue is about three to four times larger than in the UV.
- 20 The sensitivity in the green spectral range at this SZA is eight times higher than in the UV spectral range. In this wavelength range, the sensitivity at the surface depends stronger on SZA (for increasing SZA the differences between the two spectral ranges increase). The sensitivity difference between the wavelength regions decreases with increasing altitude, up to about 9 km, where they are identical. Above 9 km the UV shows a slightly (up to 4%) higher sensitivity to NO<sub>2</sub> compared to the blue spectral range. The sensitivity of the green spectral range is slightly
- 25 smaller (up to 3%) than for the blue spectral range. With higher altitude the BAMFs converge, up to at around 30 km, from which onwards there are no significant differences between the spectral ranges (not shown). As shown in Fig. 1, the point of identical sensitivity moves upward with increasing SZA, which also influences the sensitivity in the atmosphere below the point of identical visibility. However, close to the ground, the differences between the UV and blue spectral range stay nearly constant.
- 30 In order to calculate the effective AMF for one measurement scene, the BAMF profile needs to be combined with an a-priori NO<sub>2</sub> concentration. This a-priori profile has a strong impact on the AMFs, as it provides the basis for the effective visibility of the NO<sub>2</sub> in the measurement scene. In this study, we used vertical NO<sub>2</sub> profiles from the TM5-MP model, described in Williams et al., 2017. For the European project QA4ECV (http://www.qa4ecv.eu) a 15 year run (2002–2016) was performed with the TM5-MP model version of July 2016. The model was run with a resolution

of  $1^{\circ} \times 1^{\circ}$  on 34 levels, using ERA-Interim reanalysis meteorology from ECMWF, with a temporal sampling of 2 hours. This dataset has been used within the QA4ECV project as a-priori for the retrieval of e.g. formaldehyde, and for other projects such as the one described here. In particular, biomass burning emissions are taken from the monthly estimates provided by the GFEDv3 inventory (van der Werf et al., 2010), and latitude-dependent injection heights

5 and a tropical burning cycle are implemented following Huijnen et al., 2010. See Williams et al., 2017 for other model details. In order to correct for the temperature dependency of the NO<sub>2</sub> cross-section, the NO<sub>2</sub> concentrations in each altitude are scaled by a correction factor linear in temperature for all spectral ranges, as suggested in Boersma et al., 2004. Temperatures were taken from the model; the temperature dependencies are illustrated in the supplement (Fig. S1-Fig. S7).

#### 10 2.8 Data selection and post-processing

The datasets for the three spectral ranges were gridded to a  $0.25^{\circ} \times 0.25^{\circ}$  grid and monthly means were calculated. Because of the fit quality and the reduced intensity in the UV spectral range, only measurements at solar zenith angles smaller than 70° are included in the following analyses. As a result of the fitting problems in the green spectral range (Sect. 2.3 and Sect. 2.4), in the following only time series for China are shown where the NO<sub>2</sub> concentrations

- 15 are high and therefore, further influences are less pronounced (global maps are shown Fig. S11-Fig. S14). Only measurements with a geometric cloud fraction smaller than 0.2 are included, unless otherwise noted. Cloud filtering was performed using the FRESCO+ (version 6) dataset (Wang et al., 2008). Furthermore, only fits with a χ<sup>2</sup> (describing the fit residual) smaller than 0.005 and 0.001 for the UV and vis spectral range, respectively, were used. While for the first years of the GOME-2A measurements, a consistent χ<sup>2</sup>-limit could be used for all spectral ranges,
- 20 instrumental degradation severely impacts the instrument's performance in later years (Dikty and Richter, 2011). Since channel 2 is more strongly affected by this degradation than channel 3, a larger  $\chi^2$ -limit for the UV spectral range is needed for a statistically meaningful comparison.

#### 3 Results and Discussion

#### 3.1 Spatial distribution of NO<sub>2</sub> slant columns

- 25 Slant column densities (SCDs) are retrieved from the DOAS fit; they do not depend on any a-priori assumptions on the state of the atmosphere. Therefore, a comparison of the spatial distribution of the SCDs from the two UV and vis<sub>blue</sub> fit windows provides a first opportunity to assess the NO<sub>2</sub>. The SCDs depend on the VCDs and on the AMFs (see Eq. 3). Furthermore, the SCDs depend on the measurement geometry, SSR, and wavelength, and the AMFs additionally depend on the a-priori vertical NO<sub>2</sub> profile. Consequently, maps of SCDs show not only the altitude dependency, but also other changes of the AMF. For one particular measurement, the geometry is the same
- irrespective of the wavelength window. Thus, differences in the SCDs are introduced by the different sensitivities

and therefore, large differences between the two spectral ranges are expected when the  $NO_2$  is located close to the surface; when the  $NO_2$  located higher in the atmosphere, the differences between the two spectral ranges are reduced. Consequently, the SCD differences can be partly explained by the different measurement geometries and SSR, which change during the year, but other factors like injection height and relative vertical distribution have to contribute as well.

Figure 7 shows monthly averages of tropospheric  $NO_2$  SCDs for January and July 2008. In both spectral ranges, similar spatial patterns are found and anthropogenic as well as natural air pollution can be detected. Especially over the respective winter hemisphere, anthropogenic source regions are clearly observed. In January, for example, the highest  $NO_2$  columns are located over China (a, b), whereas in July, there are high  $NO_2$  values over the Highveld

10 Plateau region in South Africa (HPSA; c, d). In addition, biomass burning regions are also clearly observable in the data. For instance, in July over Africa south of the equator (ASE), enhanced values are detected in both spectral ranges, consistent with van der A et al. (2008) and Schreier et al. (2014), who found a similar seasonal pattern. Finally, artefacts originating from the simplified stratospheric correction are observed in Fig. 7 (a, b), when enhanced NO<sub>2</sub> values over the Pacific reference area lead to large areas with too low NO<sub>2</sub> values in the subtropics of the northern

15 hemisphere (see Sect. 2.6) in both spectral ranges.

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In order to better compare the two  $NO_2$  retrievals, Fig. 8 shows the relative differences between these results (absolute differences are shown in Fig. S15). Over highly polluted areas,  $NO_2$  values in the UV spectral range are generally lower than in the blue spectral range (see Sect. 2.4). These differences are introduced by the wavelength dependency of the penetration depth (see Sect. 1), which leads to better visibility of lower tropospheric pollution

- 20 to the vis<sub>blue</sub> retrieval. Larger differences between the two NO<sub>2</sub> retrievals can be found in the respective winter hemisphere, e.g., in January over China and the east coast of the USA or in July over HPSA (see also Fig. 7). These large differences are related to both the larger tropospheric NO<sub>2</sub> SCDs and to the larger SZA (see Fig. 1). In addition, large differences are observed in the western South Atlantic, where the retrieval noise in both spectral ranges is strongly enhanced due to the South Atlantic Anomaly (Richter et al., 2011).
- To further analyse the differences between the two spectral windows, we use the ratio of UV to vis<sub>blue</sub> tropospheric NO<sub>2</sub> SCDs which is a first approximation of an altitude dependency; Fig. 9 shows these for (a) January and (b) July 2008. Highly polluted areas are clearly observable, with ratios as low as ~0.2, indicating the NO<sub>2</sub> being located close to the surface. As before, differences between the two retrievals are larger in the respective winter hemisphere, showing as lower ratios. The most obvious example is East China in January, where large areas with very low ratios
- 30 can be detected. However, all strong anthropogenic source regions, e.g., India, northern America, HPSA, and the Middle East, but also individual large cities like Madrid, Moscow, and Mexico City, show low ratios between the two retrievals. Here, it should be noted that due to the rather strict selection criterion on the SZA (see Sect. 2.8), no data are available for the strong source regions in central Europe in January.

In addition to anthropogenic air pollution close to the ground, NO<sub>2</sub> pollution from biomass burning can be detected 35 in the retrieval ratios, e.g., in ANE and ASE (Fig. 9). In July (southern hemispheric, SH, winter), values of  $\sim 0.6$  are



Figure 7. Monthly mean tropospheric NO<sub>2</sub> SCDs for (a, b) January and (c, d) July 2008. (a, c) UV spectral range and (b, d) blue spectral range.

found for the  $UV/vis_{blue}$  ratio over the ASE region. These low values (as well as the absence of a signal in January) correspond to the seasonal pattern of biomass burning in this region.

In the ANE region, the vis<sub>blue</sub> fit results show clearly enhanced NO<sub>2</sub> SCDs in January (NH winter, coinciding with the annual biomass burning peak in this region) but not in July (see Fig. 7). However, no significant differences between the two months are observed in the UV fit results, where slightly enhanced values can be seen throughout the year (not shown). These findings will be further discussed in Sect. 3.2.

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North of the ANE region at approx.  $10^{\circ}-20^{\circ}$  N, SCD ratios of ~0.7 can be found in July (see Fig. 9). This is an indicator for NO<sub>2</sub> enhancement most probably related to soil emissions (Jaeglé et al., 2004; Zörner et al., 2016). While this NO<sub>2</sub> is more clearly observed in the blue spectral region, also the UV results show enhanced SCDs (see

10 Fig. 7). As shown by Delon et al. (2008) and Stewart et al. (2008), NO<sub>x</sub> (NO<sub>x</sub> = NO + NO<sub>2</sub>) from soil emissions is usually well mixed and thus not only located close to the ground, but also in elevated layers, enhancing visibility in the UV.



Figure 8. Relative difference between monthly mean tropospheric  $NO_2$  SCDs in the UV and blue spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no  $NO_2$  values available. Light grey coloured values indicate values where the vis<sub>blue</sub>  $NO_2$  is close to zero, which have been filtered out. This threshold was defined as a smoothed latitude-dependent mean over the reference sector area. The Pacific threshold is one standard deviation of the gridded  $NO_2$  values for both retrievals. To smooth the latitude-dependent threshold, a 5°-running-mean was used. The regions defined in Tab. 3 are depicted in the right figure (b).



Figure 9. Ratio between monthly mean tropospheric SCDs of NO<sub>2</sub> in the UV and blue spectral range. (a) January 2008 and (b) July 2008. Dark grey shaded area: no NO<sub>2</sub> values available. Light grey coloured values indicate values where the vis<sub>blue</sub> NO<sub>2</sub> is close to zero, which have been filtered out (Fig. 8).

Finally, Fig. 9 shows SCD ratios over the well known shipping lane leading from South India to the Strait of Malacca. The shipping signal in this region was first identified in satellite  $NO_2$  data by Beirle et al. (2004) and Richter et al. (2004). Here, while the shipping lanes are clearly observable in the vis<sub>blue</sub> SCDs, they cannot be identified in the UV  $NO_2$  data. This is probably because the  $NO_2$  shipping emissions stay with the shallow marine

5 boundary layer. Consequently, the shipping emissions when averaged over the ground scene are below the detection limit for the UV spectral range (see Fig. 7). Higher horizontal resolution (e.g., in OMI and the upcoming S-5P, S-4, and S-5 missions) might improve detection of shipping lanes in the UV.

From the retrieved ratios of SCDs, a first approximation of the top-altitude of NO<sub>2</sub> can be calculated. For the altitude retrieval, simple box profiles from the surface up to a top altitude are assumed for tropospheric NO<sub>2</sub> and the ratios of these box profile SCDs are compared with the retrieve ratios of the measurements. The altitude of ratio from the box profile SCDs which fits best to the ratio of the measurement is taken as the NO<sub>2</sub> top-altitude. The retrieved top-altitudes for the individual pixels are shown in Fig. 10. A seasonal dependency of the retrieved altitude can be clearly observed in the global maps. Low altitudes are found in regions with low ratios e.g., China in January and HPSA in July. High altitudes are found over the Oceans and in the surroundings of highly polluted areas which is expected from transport events (Richter et al., 2014; Zien et al., 2014).



Figure 10. Monthly mean top-altitudes retrieved from the ratio between the UV and blue spectral range. To retrieve the altitude, box profiles are assumed for the tropospheric  $NO_2$ . Light grey coloured values indicate values where the vis<sub>blue</sub>  $NO_2$  is close to zero, which have been filtered out.

# 3.2 Temporal variability of regional NO<sub>2</sub> slant columns

In this section, we investigate the temporal variability of tropospheric SCDs from the two/three  $NO_2$  retrievals over six regions (see Tab. 3 and Fig. 8 for the regions' definitions) and make inferences about the vertical distribution of NO<sub>2</sub> in the atmosphere. Figure 11 shows monthly mean time series for three natural/biomass burning source regions (a-c) and three anthropogenic source regions (d-f). The seasonal cycle observed in Fig. 11 a – c corresponds to the seasonal pattern of biomass burning activity in these regions (Schreier et al., 2014). Slightly negative SCDs in the UV fit in north Australia (NAUS; Fig. 11 c) and Riyadh (Kingdom of Saudi Arabia; Fig. 11 f) are artefacts caused by the stratospheric correction (see Sect. 2.6). Another artefact in the data is related to degradation of the GOME-2A instrument (see Sect. 2.8; Dikty and Richter, 2011), which is in some regions represented by a slightly decreasing linear trend (for example: in the UV in ANE and HPSA; Fig. 11 b and c).

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Table 3. Selected regions for  $NO_2$  time series with their abbreviation and their location.

| Region                           | Abbreviation | geographical location   |
|----------------------------------|--------------|---|
| Africa south of the equator      | ASE          | $5^{\circ}\mathrm{S} - 20^{\circ}\mathrm{S} \text{ and } 10^{\circ}\mathrm{E} - 40^{\circ}\mathrm{E}$ |
| Africa north of the equator      | ANE          | $0^{\circ}\mathrm{N} - 10^{\circ}\mathrm{N}$ and $15^{\circ}\mathrm{W} - 40^{\circ}\mathrm{E}$        |
| North Australia                  | NAUS         | $15^{\circ}\mathrm{S} - 25^{\circ}\mathrm{S}$ and $120^{\circ}\mathrm{E} - 145^{\circ}\mathrm{E}$     |
| China                            |              | $30^{\circ}\mathrm{N}$ – $40^{\circ}\mathrm{N}$ and $110^{\circ}\mathrm{E}$ – $123^{\circ}\mathrm{E}$ |
| Highveld Plateau in South Africa | HPSA         | $22^{\circ}\mathrm{S}$ – $30^{\circ}\mathrm{S}$ and $27^{\circ}\mathrm{E}$ – $34^{\circ}\mathrm{E}$   |
| Riyadh                           | _            | $23^{\circ}\mathrm{N} - 26^{\circ}\mathrm{N}$ and $45^{\circ}\mathrm{E} - 48^{\circ}\mathrm{E}$       |

In both ASE and NAUS, the seasonal cycles in the UV and blue spectral range are similar (Fig. 11 a and c), but the tropospheric  $NO_2$  SCDs in the UV are smaller than in the blue spectral range. As can be seen in Fig. 12, this results in the strong correlation coefficients (0.87 in both cases, see Tab. 4) and slopes  $\leq 1$  of the regression 10 lines (see Tab. 5). When the seasons are considered individually, the slope differs between 0.65 (biomass burning season) and 1.04 (rainy season). Hence, for both regions in the rainy season the two spectral ranges differ only by an offset whereas in the biomass burning season large differences are observed. Furthermore, in both regions for the individual seasons the correlation coefficient is also high (0.50-0.99; see Tab. 4), especially in the biomass burning season ( $\geq 0.92$ ), which is probably related to the higher signal to noise ratio for this season. In ANE, conversely, 15 significant differences in the seasonal cycle between the two spectral ranges can be observed, as the  $NO_2$  signal from the biomass burning peak in NH winter shows in the blue spectral range only while in the UV, no interannual variability can be detected (Fig. 11 b). The increased  $NO_2$  load observed in the vis<sub>blue</sub> SCDs cannot be detected in the UV SCDs (nevertheless the correlation coefficient is 0.75). In all the other seasons, the correlation coefficient is comparable to that observed in ASE and NAUS (see Fig. 12b and Tab. 4). 20

Several effects could contribute to an explanation for the differences in visibility of  $NO_2$  in ANE in the two spectral ranges:



Figure 11. Time series 2007-2015 of tropospheric NO<sub>2</sub> SCDs for the UV (red line), vis<sub>blue</sub> (blue line), and vis<sub>green</sub> (green line, China only) spectral range for different regions. Note the different y-axes. The vertical lines indicate January of the individual years. (a-c) biomass burning regions and (d-f) regions with high anthropogenic air pollution.

| Region | Annual | DJF  | MAM  | JJA  | SON  |
|--------|--------|------|------|------|------|
| ASE    | 0.87   | 0.50 | 0.69 | 0.92 | 0.80 |
| ANE    | 0.53   | 0.75 | 0.77 | 0.76 | 0.63 |
| NAUS   | 0.87   | 0.76 | 0.75 | 0.87 | 0.90 |
| China  | 0.98   | 0.98 | 0.99 | 0.83 | 0.98 |
| HPSA   | 0.82   | 0.74 | 0.80 | 0.90 | 0.87 |
| Rivadh | 0.95   | 0.73 | 0.86 | 0.70 | 0.95 |

Table 4. Correlation between UV  $NO_2$  SCDs and vis<sub>blue</sub>  $NO_2$  SCDs for the six selected regions.

1. Cloud influence on NO<sub>2</sub> measurements: When considering only those measurements flagged as cloudy (i.e., having a cloud coverage  $\geq 0.3$ ), only a weak seasonal cycle can be found in ANE (Fig. A1), whereas ASE shows a similar seasonal cycle in the cloud covered case as in the cloud free case. This indicates that the biomass



Figure 12. Scatter plot of tropospheric UV NO<sub>2</sub> SCDs against tropospheric visblue NO<sub>2</sub> SCDs for the period 2007–2015 for the six selected regions. Note the different axes. The seasons are colour coded. The related correlation, slope and intercept of the regression line can be found in Tab. 4 and 5. December, January, February: DJF; March, April, May: MAM; June, July, August: JJA; September, October, November: SON.

burning  $NO_2$  in ASE might be located partially in elevated  $NO_2$  layers above the clouds, whereas in ANE  $NO_2$ is located closer to the ground, i.e., usually below and thus partly shielded by the clouds.

- 2. NO<sub>2</sub> layer altitude: The altitude of biomass burning NO<sub>2</sub> emissions influence their visibility (see Fig. 1). According to simulations with the TM5-model (Williams et al., 2017; for details, see App. A2), the layer height of biomass burning NO<sub>2</sub> over ANE is lower than over ASE (Fig. A2). This is consistent with the above result, which shows high and low visibility for the biomass burning NO<sub>2</sub> in the UV for the ASE and ANE regions, respectively. For ASE, several previous studies could show that biomass burning plumes are regularly located at least partly above the boundary layer (Coheur et al., 2007; Rio et al., 2010; Gonzi and Palmer, 2010).
- 10 3. Stratospheric correction: For observations in the UV spectral range, total SCDs (i.e., the SCDs resulting from the DOAS fit, without the stratospheric correction described in Sect. 2.6) over ANE show two distinct peaks over the course of the year: one peak in NH winter and another peak in NH summer (Fig. A3). The NH winter peak in Dec./Jan. falls in the main biomass burning season and is thus expected to be of tropospheric origin. In contrast, the NH summer peak in May – July is of stratospheric origin, as stratospheric  $NO_2$  has its maximum in summer. It is noteworthy that in the vis<sub>blue</sub> spectral range, this NH summer peak is smaller compared to

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the UV, which might be partly related to small differences in the stratospheric sensitivity between the two wavelength ranges. In the UV SCDs, the biomass burning  $NO_2$  might thus be obscured by the stratospheric correction due to the similar magnitude of the stratospheric NH summer peak.

- 4. Cloud influence on data sampling: The data shown in Fig. 11 are filtered to include only measurements not strongly influenced by clouds. In satellite-based cloud retrievals, smoke is often misinterpreted as cloud (Boersma et al., 2004). Assuming that smoke and NO<sub>2</sub> from biomass burning is advected together, this could lead to large parts of the biomass burning NO<sub>2</sub> being filtered out due to apparent cloud contamination of the measurement. However, the vis<sub>blue</sub> NO<sub>2</sub> SCDs are clearly enhanced in the cloud-free measurements, showing that not all biomass burning NO<sub>2</sub> is filtered out. Furthermore, there are only small differences in the cloud coverage between ANE and ASE (not shown); in both regions, the number of pixels being filtered out due to cloud contamination is similar. The minimum fraction of cloud free pixels is about 30 %, observed during the rainy season, while the maximum fraction of about 70–80 % and 80 % in ANE and ASE, respectively, is observed during the dry season, when the biomass burning occurs. It seems thus unlikely that the cloud filtering significantly impacts the visibility of the biomass burning NO<sub>2</sub> over ANE.
- 5. Aerosols: The NO<sub>2</sub> measurements could be influenced by aerosols. According to CALIPSO measurements, the main aerosol-type in ASE is smoke from biomass burning, while other aerosol types contribute only a smaller amount to the total aerosol concentration (Omar et al., 2009; Fig. A4). In contrast, the dominant aerosol-types in ANE are dust and polluted dust (Omar et al., 2009). As in the UV spectral range, the single scattering albedo (SSA) of dust is smaller than that of biomass burning aerosols (Russell et al., 2010; Dubovik et al., 2001; Bergstrom et al., 2007), the biomass burning NO<sub>2</sub> signal in ANE could be shielded by the darker dust aerosol in the UV. However, this hypothesis seems unlikely, as the number of dust aerosols (per area) is even higher in Riyadh compared to ANE, while in Riyadh a clear NO<sub>2</sub> seasonal cycle can be identified in both spectral ranges.

In summary, we conclude that the absence of a biomass burning  $NO_2$  signal in the UV spectral range over ANE 25 is predominantly formed by the lower injection height, causing the  $NO_2$  to be located lower in the atmosphere compared to ASE. Due to the strong differences in sensitivity between UV and blue spectral ranges close to the ground, this could to a large extent explain the invisibility of UV  $NO_2$  in ANE. However, also further effects could influence the visibility. For example, interferences of the seasonal cycles of stratospheric and tropospheric  $NO_2$  over ANE might contribute, while we deem the effect caused by differing prevalent aerosol types in the three regions

30 unlikely to be significant.

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Figures 11 d-f show three regions dominated by anthropogenic air pollution. Here, absolute differences between UV and vis<sub>blue</sub> NO<sub>2</sub> values are often larger than in biomass burning source regions (see also Fig. 9). The seasonal cycle is mostly larger in the blue spectral range compared to the UV spectral range, indicating a seasonal dependency

of the reasons underlying the differences between the spectral regions. In general, the time series of the two spectral ranges show a similar behaviour and are highly correlated (0.82; Tab. 4).

Over East China, the NO<sub>2</sub> SCDs in UV, vis<sub>green</sub>, and vis<sub>blue</sub> have similar shape and show a very high correlation of  $\geq 0.98$  (Fig. 11 d; see also Table S1). Until 2012, increasing NO<sub>2</sub> winter values in the three spectral ranges can be

- 5
- observed; these increases have been analysed in depth by Richter et al. (2005) and Hilboll et al. (2013a). In 2013, there are nearly no changes in NO<sub>2</sub> and afterwards, NO<sub>2</sub> SCDs are slightly decreasing, which was already reported by Richter et al. (2015) and Irie et al. (2016) and is consistent with findings reported by Hilboll et al. (2017) for anthropogenic NO<sub>2</sub> pollution in India. Over the HPSA region, no NO<sub>2</sub> trend is observed (Fig. 11e). A clear seasonal cycle is observed in both spectral ranges, with the highest correlation coefficient in SH winter (0.90; Tab. 4). The 10 detected seasonal cycle in the blue spectral range is similar to the cycle found by Noije et al. (2006) in the year 2000 in data from the Global Ozone Monitoring Experiment (GOME). In Riyadh (Fig. 11 f) a small trend is observed in maximum values. As shown by Hilboll et al. (2013a), during the whole period a slight increase can be detected. The detected month with maximum  $NO_2$  is in agreement with van der A et al. (2008). In this area the  $NO_2$  values derived from the UV and blue spectral range show a similar seasonal cycle and a similar year-to-year variability with

a correlation coefficient of  $\geq 0.95$ . 15

| Region | Annual           | DJF             | MAM               | JJA              | SON              |
|--------|------------------|-----------------|-------------------|------------------|------------------|
|        | 0.66/ -0.19      | 0.98/ -0.45     | 0.96/ - $0.44$    | 0.83/ -0.61      | 0.65/ -0.10      |
| ASE    | $(0.03/ \ 0.05)$ | $(0.23/\ 0.20)$ | $(0.17 / \ 0.16)$ | $(0.07/ \ 0.11)$ | $(0.09/ \ 0.15)$ |
|        | 0.40/ $0.26$     | 1.08/ -0.93     | 0.84/ -0.23       | 0.92/ -0.23      | 0.67/ $0.01$     |
| ANE    | $(0.05/ \ 0.06)$ | $(0.16/\ 0.26)$ | $(0.12/\ 0.15)$   | $(0.14/ \ 0.14)$ | $(0.13/ \ 0.14)$ |
| NAUS   | 0.74/ -0.32      | 0.89/ -0.41     | 1.04/ -0.50       | 0.78/ -0.30      | 0.71/ -0.34      |
|        | $(0.04/ \ 0.03)$ | $(0.13/\ 0.11)$ | $(0.16/\ 0\ 09)$  | $(0.08/ \ 0.04)$ | $(0.06/ \ 0.06)$ |
| China  | $0.37/\ 1.04$    | 0.42/ -0.07     | 0.46/ $0.25$      | 0.49/ $0.52$     | 0.37/ $0.89$     |
|        | $(0.01/ \ 0.08)$ | $(0.02/\ 0.30)$ | $(0.02/ \ 0.14)$  | $(0.06/ \ 0.36)$ | $(0.02/ \ 0.16)$ |
| HPSA   | 0.41/ $0.24$     | 0.74/ -0.37     | 0.46/ $0.21$      | 0.68/ -1.03      | 0.60/ -0.32      |
|        | $(0.03/ \ 0.09)$ | $(0.12/\ 0.26)$ | $(0.06/\ 0.21)$   | $(0.06/ \ 0.27)$ | $(0.06/ \ 0.19)$ |
| Riyadh | 0.74/ -1.26      | 0.63/ -1.15     | 0.64/ -0.97       | 0.46/ $0.06$     | 0.71/ -1.10      |
|        | $(0.02/ \ 0.08)$ | $(0.10/\ 0.26)$ | $(0.07/ \ 0.19)$  | $(0.08/ \ 0.40)$ | $(0.04/ \ 0.17)$ |

Table 5. Slope and intercept between UV NO<sub>2</sub> SCDs and vis<sub>blue</sub> NO<sub>2</sub> SCDs for the six selected regions. Intercept in  $1e15 \operatorname{molec} \operatorname{cm}^{-2}$ . In the brackets the standard errors for both values are shown.

In anthropogenically polluted areas, the differences in the  $NO_2$  signal strength between the UV and blue spectral ranges are a result of the larger sensitivity to the lower troposphere in the blue spectral range. In winter the air pollution is mostly stronger, compared to summer, the SZA is higher, the boundary layer is more shallow (von Engel and Teixeira, 2013), and even within this boundary layer also a strong gradient is expected over high-emission areas

- 5 (Dieudonné et al., 2013). All these effects increase the observed differences between the two spectral ranges. Due to the higher profile sensitivity close to the ground and the resulting larger differences, a lower regression slope (see Fig. 12) is observed for the anthropogenically polluted areas than for the biomass burning regions (Tab. 5) in all seasons. For example, in NAUS and ASE the regression line between all UV and vis<sub>blue</sub> SCDs has a slope of 0.66 and 0.74, respectively, whereas for China and HPSA, a considerably smaller slope ( $\sim 0.4$ ) is found.
- 10 One has to note that the regions ANE and Riyadh do not follow this pattern ANE due to the missing biomass burning NO<sub>2</sub> signal in the UV, and Riyadh probably because of differences in boundary layer height and the maximum in NO<sub>2</sub> seasonal cycle (von Engel and Teixeira, 2013). In Riyadh, the seasonal cycle of NO<sub>2</sub> emission and the cycle of the boundary layer height are in phase, as it is the case for naturally polluted areas, whereas for China and HPSA, the NO<sub>2</sub> maximum can be found in the respective winter season, where a lower boundary layer height 15 is observed (Fig. S16). Consequently, in Riyadh the UV NO<sub>2</sub> signal is stronger and the slope (0.74) is larger than
  - observed in other anthropogenic source regions.

Finally, a seasonal dependency of the regression line slope may result from the combination of the boundary layer height and concentration of NO<sub>2</sub>. For ASE and NAUS, the slope shows a seasonal pattern, which might be influenced by the small signal to noise ratio (see Tab. 5). Slopes of  $\sim 1.0$  can be found in these areas for the rainy

- 20 season, with the largest uncertainties for these areas, possibly related to high cloud cover. The differences are thus most likely introduced by an offset. During biomass burning season, the slope is about 0.83 and 0.71 for ASE and NAUS, respectively. Since the biomass burning  $NO_2$  signal cannot be seen in UV SCDs over ANE, no clear seasonal pattern can be observed there. In East China and the HPSA region, a larger slope can be found in summer, when the boundary layer is higher; spring, autumn and winter show a smaller slope (for China this can also be found for the
- 25 green spectral range, Table S1). This is expected due to the increased sensitivity (see Fig. 1) of the UV measurements to  $NO_2$  in a thick compared to a shallow boundary layer. In Riyadh however, this effect is not observed, possibly due to a decreased signal to noise ratio caused by the smaller area, entailing fewer measurements.

#### 3.3 Spatial distribution of NO<sub>2</sub> vertical columns

The NO<sub>2</sub> vertical column densities (VCDs) are the final result of the DOAS retrieval procedure, as they are a 30 physically meaningful and universally comparable quantity. In this section, we compare NO<sub>2</sub> tropospheric VCDs retrieved from the UV and the blue spectral range. For the calculation of the VCDs, NO<sub>2</sub> vertical profiles from the TM5-model (Williams et al., 2017) are used as a-priori (see Sect. 2.7). Assuming perfect measurements, radiative transfer simulations, and a-priori profiles, the AMF would equalise all differences between the two spectral ranges, leading to identical VCDs for UV and vis<sub>blue</sub> fits. Figure 13 shows maps of monthly mean NO<sub>2</sub> VCDs for January and July 2008. The spatial patterns in both datasets are very similar and agree well with those shown in SCD maps (Fig. 7).



Figure 13. Monthly mean tropospheric NO<sub>2</sub> VCDs for (a, b) January and (c, d) July 2008. (a, c) UV spectral range and (b, d) blue spectral range.

However, also the VCDs of both spectral ranges do not perfectly match, as illustrated in Fig. 14. Compared to the relative differences of the SCDs shown in Fig. 8, the relative differences of the VCDs are smaller (absolute differences are shown in Fig. S17). Especially over areas without significant NO<sub>2</sub> sources, but also over the biomass burning regions, the differences between the two retrievals are reduced compared to the SCDs. Over anthropogenically polluted areas however, e.g., over China in January, the VCDs retrieved from the two spectral regions still show significant differences. The most probable reason for this is that the simulated NO<sub>2</sub> profiles used as a-priori do not represent the actual NO<sub>2</sub> vertical distribution; furthermore, aerosols and the SSR might influence the calculated VCDs, as will be discussed in Sect. 3.4.



Figure 14. Relative difference between monthly mean tropospheric  $NO_2$  VCDs in the UV and blue spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no  $NO_2$  values available. Light grey coloured values indicate values which are filter out. The same filter as for Fig. 8 is used.

# 3.4 Temporal variability of regional NO<sub>2</sub> vertical columns

To better understand the reasons for the differences between the tropospheric VCDs of the two/three spectral regions, we again investigate the temporal variability over the six regions shown before (see Sect. 3.2, Tab. 3, Fig. 8 for the region definitions). The general shape of the seasonal cycles, which was described for SCDs in Sect. 3.2, can also be found in the VCD time series (Fig. 15). Similar to the SCD time series (see Sect. 3.2), the NO<sub>2</sub> VCDs are still

- 5 be found in the VCD time series (Fig. 15). Similar to the SCD time series (see Sect. 3.2), the NO<sub>2</sub> VCDs are still mostly larger in the blue spectral range. Compared to the SCD differences, however, the VCD differences between the UV and the blue spectral range are reduced in all six regions (Fig. 15) which can also be seen in the larger slope of the regression line (Tab. 6). For nearly all regions and seasons, the slope of the regression line (Tab. 7) is still below one, indicating that the calculated AMFs are not representative for the actual state of the atmosphere. The
- 10 SSR and the relative vertical profile as well as aerosols may contribute to the differences of VCDs between the UV and the blue spectral range.

In the ASE region, for example, the seasonal cycle in the VCDs is smoother than in the SCDs, especially for the UV data (Fig. 15 a). The seasonal cycles of the VCDs in the UV and the blue spectral range are more parallel, leading to a high correlation coefficient of 0.97 between the two data sets (see Tab. 6). However, rather than being

15 separated by a simple offset, the VCDs from the two spectral ranges seem to be linearly related as indicated by the regression slopes differing from one (see Tab. 7 and Fig. 16). Compared to SCDs, the slope for the VCDs increases slightly in nearly every season; for the annual values, the slope of VCDs is 0.77 compared to 0.66 for SCDs.



Figure 15. Time series 2007-2015 of tropospheric NO<sub>2</sub> VCDs for the UV (red line), vis<sub>blue</sub> (blue line), and vis<sub>green</sub> (green line, China only) spectral range as well as the TM5-model VCDs (grey line) for different regions. Note the different y-axes, the same as in Fig. 11. The vertical lines indicate January of the individual years. (a-c) biomass burning regions and (d-f) regions with anthropogenic air pollution.

Table 6. Correlation between UV NO<sub>2</sub> VCDs and vis<sub>blue</sub> NO<sub>2</sub> VCDs for the six selected regions.

| Region | Annual | DJF  | MAM  | JJA  | SON  |
|--------|--------|------|------|------|------|
| ASE    | 0.97   | 0.66 | 0.89 | 0.93 | 0.95 |
| ANE    | 0.91   | 0.75 | 0.92 | 0.98 | 0.87 |
| NAUS   | 0.95   | 0.92 | 0.78 | 0.93 | 0.94 |
| China  | 0.98   | 0.96 | 0.98 | 0.96 | 0.97 |
| HPSA   | 0.96   | 0.79 | 0.97 | 0.90 | 0.93 |
| Riyadh | 0.92   | 0.83 | 0.89 | 0.70 | 0.96 |

In the ANE region, the VCDs from the UV and blue spectral range agree considerably better than for the respective SCDs (Fig. 15 b). The UV NO<sub>2</sub> VCDs show a clear seasonal cycle, of similar shape but reduced amplitude compared to the vis<sub>blue</sub> VCDs, leading to reduced differences between the two spectral ranges. Comparing to the SCDs in

the ANE region, this indicates that the biomass burning  $NO_2$  seems to be located too close to the ground to be detected by the UV retrieval; however, as the  $NO_2$  emitted by fires is included in the simulation from which the a-priori profiles are taken, the AMFs introduce the observed seasonal cycle in the UV VCDs. This is reflected by the considerably improved correlation coefficients of 0.91 for all VCDs (see Tab. 6) compared to 0.53 for all SCDs; similar

- 5 improvements can be seen for all seasons individually. Interestingly, when the seasons are considered separately, the regression line has higher slopes between 0.79 (SON) and 1.06 (JJA), showing that in the rainy season, the UV and vis<sub>blue</sub> VCDs only differ by an offset. Nevertheless, discrepancies between the two spectral ranges are still observed, especially during biomass burning season.
- For NAUS, the correlation coefficient between UV and vis<sub>blue</sub> VCDs is higher (0.95) than for the SCDs (0.78).
  10 Also in this area, the slope of the VCD regression lines is constant or increased compared to the SCDs (Tab. 7; Fig. 16). During rainy season, the observed differences between UV and vis<sub>blue</sub> VCDs are mostly related to an offset, whereas during biomass burning season the slope is smaller than one.

For VCD time series over East China (Fig. 15 d), the upward trend which was observed in the winter NO<sub>2</sub> SCDs (see Fig. 11 d) is not present in the VCDs. However, also here the UV NO<sub>2</sub> values are still smaller than the vis<sub>blue</sub>

- 15 NO<sub>2</sub> values. In spite of these discrepancies, the correlation coefficient is highest for this region ( $\geq 0.96$  in all seasons). During all seasons the correlation coefficients are increased or similar compared to those of the SCDs. While the slope of the regression line is still smaller than one (0.59-0.76; Tab. 7), it is still increased compared to SCDs. The green and blue spectral range show a good agreement with a correlation coefficient of 0.98 (see also Table S2). Furthermore, the slope (1.07) agrees well which shows that there are only smaller differences between these spectral
- $20 \quad {\rm ranges.}$

The Highveld (HSPA, Fig. 15 e) region differs from all other regions: a high correlation coefficient (0.90) can be found in SH winter, when also the slope of the regression line (1.00) as well as the offset is high. In SH summer, however, the smallest correlation coefficient is observed for the region (0.79) and also the slope is the smallest with 0.77.

In Riyadh, large relative differences between the UV and blue spectral ranges can be found compared to the other regions (Fig. 15 f). This is reflected by the slope of the regression line which is  $\sim 0.75$ , except for NH winter when the slope is 1.19 with a large offset (Fig. 16 f).

The observed discrepancies between the UV and blue spectral ranges can result from different reasons: they can be related to the a-priori profiles used for VCD calculations, or they can be related to other influencing factors

- 30 like aerosols and surface properties. The importance of the a-priori profile shape for the resulting VCDs was shown by Martin et al. (2002) and Boersma et al. (2004); the resulting uncertainty for this parameter is assumed to be smaller than 15% (Boersma et al., 2004). A potential error in the a-priori profiles from the TM5-model in a given altitude layer will lead to different errors in the VCDs calculated for the UV and the blue spectral range, due to the spectral dependence of the measurement sensitivity (i.e., the a-priori error is weighted with different sensitivities
- 35 in the two spectral ranges; see Fig. S18). Higher  $NO_2$  concentrations close to the surface combined with a lower



Figure 16. Scatter plot of tropospheric UV NO<sub>2</sub> VCDs against tropospheric vis<sub>blue</sub> NO<sub>2</sub> VCDs for the period 2007–2015 for the six selected regions. Note the different axes. The seasons are colour coded. The related correlation, slope and intercept of the regression line can be found in Tab. 6 and 7. December, January, February: DJF; March, April, May: MAM; June, July, August: JJA; September, October, November: SON.

injection height would reduce the differences between the UV and blue spectral range. The a-priori profile is especially relevant in biomass burning season, when the quality of the a-priori profiles is probably lower than usual, since the modelling of biomass burning emissions entails large uncertainties due to spatial resolution, emission uncertainties, and assumptions made on the plume injection height (see, e.g., Archer-Nicholls et al., 2015 and references therein).

- 5 Similarly, since our retrieval does not explicitly account for the effect of aerosols, an aerosol layer at a given altitude will influence the visibility of the NO<sub>2</sub> both above and below the aerosol differently, enhancing the difference in the VCDs retrieved from the two spectral ranges. The same is true for potential errors in the assumed SSR. In order to bring UV and vis retrievals into better agreement, either the UV SSR would have to be lower than assumed or the vis SSR larger than assumed or a combination of both. As shown by Boersma et al. (2004), in March the
- 10 uncertainties of NO<sub>2</sub> concentration are largest in polluted areas with low SSR, but they are generally below 50 % (see also Tab. 2).

Tropospheric model VCDs are plotted in Fig. 15, showing a regional dependency of the discrepancies between the model VCDs and the observed VCDs. For ASE and ANE during biomass burning season, the UV NO<sub>2</sub> VCDs and the model VCDs agree quite well, while the vis<sub>blue</sub> NO<sub>2</sub> values are higher. It should however be kept in mind that in

15 the vis<sub>blue</sub> retrieval, both the uncertainties and the contribution of the a priori are lower, making these results more

| Region | Annual           | DJF              | MAM              | JJA              | SON              |
|--------|------------------|------------------|------------------|------------------|------------------|
| ASE    | $0.77/\ 0.04$    | 1.62/ -0.73      | 0.91/ - $0.15$   | 0.94/ -0.37      | $0.76/\ 0.01$    |
|        | $(0.02/\ 0.03)$  | $(0.30/\ 0.25)$  | $(0.09/\ 0.08)$  | $(0.07/ \ 0.12)$ | $(0.05/ \ 0.08)$ |
| ANE    | 0.69/ $0.15$     | 0.90/ -0.29      | 1.04/ -0.24      | 1.06/ -0.21      | 0.79/ $0.06$     |
|        | $(0.03/ \ 0.04)$ | $(0.14/\ 0.23)$  | $(0.08/\ 0.11)$  | $(0.04/ \ 0.04)$ | $(0.08/ \ 0.09)$ |
| NAUS   | 0.81/ -0.07      | 0.98/ -0.24      | 1.12/ -0.30      | 0.94/ -0.12      | 0.78/ -0.08      |
|        | $(0.03/ \ 0.02)$ | $(0.08/\ 0.09)$  | $(0.16/\ 0.11)$  | $(0.07/ \ 0.05)$ | $(0.06/ \ 0.07)$ |
| China  | 0.69/ -0.15      | 0.75/ -0.85      | 0.72/ -0.11      | 0.76/ -0.12      | 0.59/ $0.79$     |
|        | $(0.01/ \ 0.16)$ | $(0.04/\ 0.78)$  | $(0.03/ \ 0.32)$ | $(0.04/\ 0.29)$  | $(0.03/ \ 0.34)$ |
| HPSA   | 0.94/ -0.54      | $0.77/\ 0.02$    | 0.93/ -0.29      | 1.00/ -0.89      | 0.87/ -0.49      |
|        | $(0.03/ \ 0.11)$ | $(0.11/\ 0.31)$  | $(0.05/ \ 0.19)$ | $(0.09/ \ 0.48)$ | $(0.07/ \ 0.26)$ |
| Riyadh | 0.74/ -0.37      | 1.19/ -1.43      | 0.79/ -0.51      | 0.67/ -0.19      | 0.74/ -0.37      |
|        | $(0.03/ \ 0.09)$ | $(0.15/ \ 0.36)$ | $(0.08/\ 0.20)$  | $(0.12/\ 0.42)$  | $(0.04/ \ 0.12)$ |

**Table 7.** Slope and intercept between UV NO<sub>2</sub> VCDs and vis<sub>blue</sub> NO<sub>2</sub> VCDs for the six selected regions. Intercept in 1e15 molec cm<sup>-2</sup>. In the brackets the standard errors for both values are shown.

reliable. In rainy season, especially in ANE, there are larger discrepancies between the UV and  $vis_{blue}$  NO<sub>2</sub> and the model VCDs. The model values are lower though the UV and blue spectral ranges agree quite well. This confirms the above finding that the differences between the spectral ranges are caused by an offset. The differences during biomass burning season might be related to the aerosol treatment or wrong SSR in AMF calculations.

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Also for NAUS the simulated and observed values agree quite well in the rainy season, whereas in the biomass burning season large differences between the three values are observed. Therefore, in this regions a combination of the different factors, described above, will most likely provide the best explanation.

For China and Riyadh, in the less polluted season the model and UV NO<sub>2</sub> values agree quite well whereas the vis<sub>blue</sub> NO<sub>2</sub> values are larger. In the season with high air pollution, in both regions large discrepancies between 10 the three values are observed. For anthropogenically polluted areas, the discrepancies between the modelled and the measured NO<sub>2</sub> values are well known from previous studies (e.g., Huijnen et al., 2010). For these regions, also aerosols play an important role and can further influence the differences between the two spectral ranges. As shown by Russell et al. (2010), especially dust reduces the visibility in the UV spectral range which could partly explain the differences in Riyadh. Additionally, as shown by Boersma et al. (2004), for these anthropogenically polluted areas

15 also the SSR can introduce large differences. In China only small differences between the blue and green spectral

range were found. The reason for the different behaviour between the two vis spectral ranges might be that the green spectral range is less sensitive to the a-priori profile compared to the UV and blue spectral range (see Fig. S18).

In HPSA, an offset between the model and the observed values can clearly be observed, which could partly explain the differences in SH winter values (Tab. 7). However, also here the SSRs and aerosols might influence the measurements, and therefore, the calculation of AMF.

#### 4 Summary and Conclusion

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In this study, we present a new NO<sub>2</sub> DOAS retrieval for measurements by the GOME-2A instrument, using the ultraviolet (UV) spectral range. Having NO<sub>2</sub> measurements from this wavelength range in addition to the usually exploited blue spectral region gives insights into the vertical distribution of tropospheric NO<sub>2</sub>, as the vertical sensitivity of space-borne NO<sub>2</sub> DOAS retrievals strongly depends on the wavelength.

 $NO_2$  slant column densities (SCDs) can in principle be retrieved from the UV spectral range, as shown by Yang et al. (2014). However, the UV retrievals have larger uncertainty compared to the blue spectral range, caused by the smaller differential absorption and the reduced intensity in the UV spectral range, leading to inferior fit quality, and also due to the larger AMF uncertainties. Nonetheless, the spatial distribution of both retrievals agrees very well.

15 We focused our analysis on six regions: three biomass burning regions (Africa south of the equator / ASE, Africa north of the equator / ANE, and northern Australia / NAUS), and three anthropogenic source regions (Eastern China, the Highveld Plateau in South Africa / HPSA, and the surroundings of Riyadh).

The differences between the UV and  $vis_{blue}$  fits are lower over the biomass burning areas ASE, ANE, and NAUS compared to the regions dominated by anthropogenic pollution. The NO<sub>2</sub> from biomass burning is in some regions

- 20 and periods located in elevated layers, leading to an effectively smaller sensitivity difference between UV and blue spectral ranges. This is reflected by the slope of the regression lines between UV and vis<sub>blue</sub> SCDs, which is 0.66 and 0.74 for the ASE and NAUS regions, respectively. Anthropogenic air pollution on the other hand is mostly located in the boundary layer close to the ground (see Stull, 1988). As the sensitivity of the UV retrieval drops sharply towards the surface, this leads to considerably lower regression line slopes of 0.37 and 0.41 for East China
- and HPSA, respectively. Nonetheless, SCDs in the UV and vis<sub>blue</sub> show high correlation coefficients  $\geq 0.82$  in both cases.

In ANE, the biomass burning  $NO_2$  in NH winter cannot be observed in the UV SCDs. Our analysis of UV  $NO_2$  SCDs over cloudy scenes and of  $NO_2$  profiles simulated by the TM5-MP model both suggest that in ANE the  $NO_2$  is located close to the ground, where sensitivity in the UV is limited. The non-detectability could be increased by

30 the stratospheric  $NO_2$  seasonal cycle, which counteracts the biomass burning  $NO_2$  signal over ANE. Moreover, the ANE biomass burning  $NO_2$  could be shielded by the predominantly dust aerosols, which are dark in the UV and thus lead to decreased sensitivity compared to other biomass burning regions. In ASE, for example, the predominant
aerosol type is smoke, which is brighter than dust in the UV spectral range; combined with the higher altitude of the  $NO_2$  this leads to increased visibility in ASE.

In Riyadh, the seasonal cycles of boundary layer height and tropospheric  $NO_2$  are in phase, contrary to other anthropogenically polluted areas, where  $NO_2$  columns are highest in winter, when the boundary layer is shallow.

5 Therefore, the maximum  $NO_2$  values coincide with high boundary layers and thus better visibility in the UV fit, leading to better agreement between SCDs from the two spectral regions, expressed by a higher slope of the regression line of 0.74.

Even though in theory, given perfect measurements, radiative transfer modelling, and a-priori information, vertical column densities (VCDs) retrieved by the DOAS technique should be independent on the wavelength, VCDs retrieved

- 10 from the UV and vis fits do still show considerable differences. If the a-priori profile does not capture the actual vertical distribution of NO<sub>2</sub> in the troposphere, the vertical measurement sensitivity will be incorrectly aggregated in the air mass factor calculation, which will lead to wavelength-dependent VCDs. Having said this, the observed VCD differences can also be introduced by further input parameters like surface spectral reflectance (SSR; wavelength: 354 nm and 442 nm) and the lack of explicit aerosol treatment. These parameters can lead to large discrepancies
- 15 between the two spectral ranges when assumed wrongly. Even though differences exist between the UV and blue spectral ranges, the correlation coefficient between the VCDs is high everywhere ( $\geq 0.91$ ). Generally, the differences between the two spectral ranges indicate that the NO<sub>2</sub> in the atmosphere is located in lower atmospheric layers than in the model which has stronger inferences in the UV, or the assumed SSR is too high in the blue spectral range.
- While vis<sub>blue</sub> NO<sub>2</sub> VCDs are often larger than UV VCDs, the latter agree with the simulated VCDs in most
  seasons. For China and HPSA, the differences between the UV and blue spectral ranges are pronounced in the respective winter season, whereas for Riyadh large differences can be found throughout the whole year. However in HPSA, in SH winter the correlation coefficient and the slope of the regression line is largest with an high offset. In the biomass burning regions, the agreement is better for the rainy seasons; during dry season, when a larger aerosol load is expected, the differences are larger. After conversion into VCDs, in ANE a clear seasonal cycle can be observed in both spectral ranges, which is mostly introduced by the a-priori assumptions in the AMF calculation.

An additional fitting window in the green spectral range also shows a higher uncertainty compared to the blue spectral range which is related to the smaller absorption lines as well as interferences with liquid water and vegetation, as already reported by Richter and Burrows (2000). Nevertheless, over highly anthropogenically polluted areas, NO<sub>2</sub> retrieved from the green spectral range is even more sensitive to the lower troposphere compared to the blue spectral

30 range. This can be clearly observed in the higher  $vis_{green}$  SCDs over China compared with the UV and blue spectral range. In general, the SCDs of the three spectral ranges agree well with correlation coefficients  $\geq 0.98$ . The VCDs for the green and blue spectral show also a good agreement (correlation coefficient of 0.98 and slope of 1.07) which might be related to the smaller sensitivity to the a-priori profile.

It is currently not possible to retrieve vertical profiles of tropospheric  $NO_2$  from satellite measurements, due to the 35 low information content of the measured spectra (see, e.g., Hilboll et al., 2014). Further developments in space-based observation of the Earth's atmosphere could however increase the capability to retrieve vertical  $NO_2$  distribution. For example, concurrent measurements of the same air mass from different viewing geometries could yield insight on aerosol types and vertical profiles which might be also possible for GOME-2A and GOME-2B observations. This in turn would increase the possibility to retrieve more quantitative information about the vertical distribution of

5  $NO_2$  in the troposphere. Improvements in signal to noise ratio would reduce the uncertainty in the UV retrievals and thus increase the value of the UV slant columns.

However, our study shows that even for the current generation of instrumentation in the global observing system, the difference between  $vis_{blue}$  and UV NO<sub>2</sub> DOAS retrievals can be used to constrain our knowledge of NO<sub>2</sub> vertical distribution in the lower troposphere. For example when coupled with data assimilation techniques and high resolution models the amount of boundary layer NO<sub>2</sub> and its trends can be inferred more accurately.

#### Appendix A: Shielding of northern hemispheric winter values in Africa north of the equator

#### A1 Shielding by clouds

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Figure A1 shows NO<sub>2</sub> time series for cloud free scenes ( $\leq 20 \%$  cloud fraction; a and c) and for cloudy scenes ( $\geq 30 \%$  cloud fraction; c and d). The decrease/increase of NO<sub>2</sub> values in recent years shown in Fig. A1 is probably related to instrumental degradation (Dikty and Richter, 2011); a similar decrease/increase cannot be observed in GOME-2B data (not shown). In ASE, the seasonal cycles for cloud covered and cloud free measurements are similar. In ANE, however, only a weak seasonal cycle can be found for the cloudy scenes in either spectral range, whereas in the cloud free scenes a clear seasonal cycle can be found in the blue but not in the UV spectral range. This implies that the NO<sub>2</sub> layer over ASE could be located also in higher altitudes, while over ANE the NO<sub>2</sub> could be located closer to the ground and thus less sensitive to the UV observations.

# A2 NO<sub>2</sub> profile shape

To compare the vertical distribution of NO<sub>2</sub> over the biomass burning areas ANE, ASE, and NAUS, simulated NO<sub>2</sub> concentration fields from the TM5-MP (Williams et al., 2017) for 2008 at 09:00 LT were used. We assume that the small time difference between the satellite measurements and the model output can be neglected. Figure A2 (a) shows profiles of NO<sub>2</sub> concentration for the regions ANE and ASE for four months. Apparently, NO<sub>2</sub> is located at higher altitudes in ASE, which leads to differences in the visibility of the NO<sub>2</sub> signal between the two spectral ranges. This is partly related to the surface altitude which is on average ~ 830 m above sea level in ASE and only ~ 440 m in ANE. This could explain why the UV fit shows enhanced NO<sub>2</sub> over ASE and not over ANE.



Figure A1. Tropospheric NO<sub>2</sub> SCD time series for ASE (a, c) and ANE (b, d) for cloud free scenes ( $\leq 20\%$  cloud fraction; a, b) and for cloudy scenes ( $\geq 30\%$  cloud fraction; c, d). The vertical lines indicate January of the individual years. (a, b) are identical to Fig. 11 a, b and are replicated here to facilitate direct comparison.



Figure A2. Average profiles of normalised NO<sub>2</sub> concentration for the three regions ANE and ASE for 2008, derived from the TM5-MP for 09:00 LT. The grey dashed lines indicating the surface height of the two regions.

### A3 Seasonality of stratospheric NO<sub>2</sub>

Figure A3 shows time series of detrended total SCDs for the regions ANE and ASE. Detrending is necessary because of the degradation of the instrument (Dikty and Richter, 2011), and is implemented as subtraction of a linear trend. Here, a shortened time series from 2008 to 2012 is used, since changes in the instrument configuration (EUMETSAT, 2015) might introduce an offset after 2012.

In ASE a clear seasonal cycle can be seen for both spectral ranges (Fig. A3 a), which agrees well with the seasonal cycle expected from biomass burning activity. However, comparison with Fig. A3 c shows that a similar seasonal cycle can also be observed in SCD data over the Pacific Ocean (averaged over  $180^{\circ}-210^{\circ}$ E), which is void of any NO<sub>x</sub> emissions and should therefore only have very low tropospheric NO<sub>2</sub> SCDs. This indicates that the seasonality over ASE is at least partly caused by stratospheric fluctuations.

Over ANE, slightly different seasonal cycles can be observed in the UV and blue spectral range (see Fig. A3 b). The peak related to biomass burning in NH winter is clearly observed in the vis<sub>blue</sub> SCD data. In some years a small secondary peak in NH summer can be noticed. Similar seasonal variability can be seen in the UV spectral range, but both peaks have similar magnitude there. This double peak structure can also be found in NO<sub>2</sub> SCDs from the Ozone Monitoring Instrument (OMI, not shown). Over the Pacific Ocean (Fig. A3d), the second peak in NH

15 the Ozone Monitoring Instrument (OMI, not shown). Over the Pacific Ocean (Fig. A3d), the second peak in NH summer can clearly be observed in the UV data, corresponding to the known seasonal variability of stratospheric NO<sub>2</sub>. As the seasonal cycles of biomass burning and stratospheric NO<sub>2</sub> are in-phase and out-of-phase for ASE and ANE, respectively, the visibility of the biomass burning NO<sub>2</sub> differs strongly between the two regions.

#### A4 Aerosol effect

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20 The Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observations (CALIPSO) satellite is in a sun-synchronous polar orbit with an equator crossing time of 13:30 LT (Winker et al., 2007). It was launched in April 2006 and has a repeat cycle of 16 days. On CALIPSO, the Cloud-Aerosol Lidar with Orthogonal Polarisation (CALIOP) instrument is operating since June 2006. CALIOP is a nadir-viewing two wavelength polarization-sensitive lidar, operating at 532 nm and 1064 nm. The horizontal and vertical resolution depends on the altitude. Closer to the ground the resolution increases in both cases (Winker et al., 2007; Winker et al., 2009).

Here, the CALIOP level 3 product "Aerosol Profile All Sky (daytime)" was used. The data are on a 2° latitude times 5° longitude grid (Atmosheric Science Data Center, 2009). For the calculation of regional aerosol contribution for smaller regions, the amount of aerosols in the grid box was weighted by the contribution of the grid box to the region. This dataset provides the six aerosol types "dust", "polluted continental", "smoke", "clean marine", "clean

30 continental", and "polluted dust", where the latter accounts for a mixture of dust and smoke or a mixture of dust and urban pollution (Omar et al., 2009). The aerosol-types were derived using aerosol models based on cluster analysis of an AERONET dataset and the aerosol extinction-to-backscatter ratio (Omar et al., 2009). In the present study, the differences in equator crossing time can be ignored, because especially for ASE and ANE larger areas are averaged.



Figure A3. Time series of detrended total NO<sub>2</sub> SCDs for the UV and blue spectral range for ASE (a) and ANE (b) as well as detrended time series over the Pacific reference sector area for the latitude of ASE (c) and ANE (d). The vertical lines indicate January of the individual years.

Figure A4 shows the sample number for the six aerosol types over the three regions ANE, ASE, and Riyadh. The dominant aerosol-type in ASE is smoke whereas dust dominates in ANE and Riyadh.



Figure A4. Time series for the number of samples of aerosol-types for three different regions (CALIPSO data). To calculate the time series for the individual areas, the amount of aerosol in the grid box was weighted by the contribution of the grid box to the region. Note the different y-axes. The vertical lines indicate January of the individual years.

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## Supplementary material

For the interested reader, this supplementary material contains additional information and figures.



Figure S1.  $NO_2$  cross section and temperature dependency for the UV spectral range. The scaling coefficient of the temperature dependency is calculated as average over all wavelength within the fitting window.



Figure S2.  $NO_2$  cross section and temperature dependency for the blue spectral range. The scaling coefficient of the temperature dependency is calculated as average over all wavelength within the fitting window.



Figure S3.  $NO_2$  cross section and temperature dependency for the green spectral range. The scaling coefficient of the temperature dependency is calculated as average over all wavelength within the fitting window.



Figure S4. Altitude dependency of the  $NO_2$  temperature scaling coefficient for China for the UV and blue spectral range. The profiles are calculated for model data simulated with the TM5-MP model for 2008.



Figure S5. Altitude dependency of the  $NO_2$  temperature scaling coefficient for ASE for the UV and blue spectral range. The profiles are calculated for model data simulated with the TM5-MP model for 2008.



Figure S6. Altitude dependency of the  $NO_2$  temperature scaling coefficient for China for the blue and green spectral range. The profiles are calculated for model data simulated with the TM5-MP model for 2008.



Figure S7. Altitude dependency of the  $NO_2$  temperature scaling coefficient for ASE for the blue and green spectral range. The profiles are calculated for model data simulated with the TM5-MP model for 2008.



Figure S8. Monthly mean total  $NO_2$  SCD for (a, b, c) January and (d, e, f) July 2008 for the (a, d) UV, (b, e) blue, and (c, f) green spectral range.



Figure S9. Relative difference between monthly mean total  $NO_2$  SCDs in the blue and UV spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no  $NO_2$  values available.



Figure S10. Relative difference between monthly mean total  $NO_2$  SCDs for the blue and green spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no  $NO_2$  values available.



Figure S11. Monthly mean tropospheric  $NO_2$  SCD for (a, c) January and (b, d) July 2008 for the (a, b) green and (c, d) blue spectral range.



Figure S12. Absolute difference between monthly mean total  $NO_2$  SCDs in the green and blue spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no  $NO_2$  values available.



Figure S13. Relative difference between monthly mean total NO<sub>2</sub> SCDs in the the green and blue spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no NO<sub>2</sub> values available. Light grey coloured values indicate values where the vis<sub>green</sub> NO<sub>2</sub> is close to zero, which have been filtered out. This threshold was defined as a smoothed latitude-dependent mean over the reference sector area. The Pacific threshold is one standard deviation of the gridded NO<sub>2</sub> values for both retrievals. To smooth the latitude-dependent threshold, a 5°-running-mean was used.



Figure S14. Ratio between monthly mean tropospheric SCDs of  $NO_2$  in the blue and green spectral range. (a) January 2008 and (b) July 2008. Dark grey shaded area: no  $NO_2$  values available. Light grey coloured values indicate values where the vis<sub>green</sub>  $NO_2$  is close to zero, which have been filtered out. This threshold was defined as a smoothed latitude-dependent mean over the reference sector area. The Pacific threshold is one standard deviation of the gridded  $NO_2$  values for both retrievals. To smooth the latitude-dependent threshold, a 5°-running-mean was used.



Figure S15. Absolute difference between monthly mean tropospheric  $NO_2$  SCDs in the UV and blue spectral range. Differences for (a) January and (b) July 2008. Dark gray shaded area: no  $NO_2$  values available.



Figure S16. Planetary boundary layer height from ECMWF data vs monthly mean tropospheric  $NO_2$  SCDs in the UV and blue spectral range.



Figure S17. Absolute difference between monthly mean tropospheric  $NO_2$  VCDs in the UV and blue spectral range. Differences for (a) January and (b) July 2008. Dark grey shaded area: no  $NO_2$  values available.



Figure S18. SCD, VCD, and AMF for different  $NO_2$  profiles. Blue: is assumed as the true profile, the SCDs are calculated for this profile. Other colour: changed input profiles, AMF are calculated for the changed profiles which leads to changes in the retrieved VCDs.

|                 | Annual           | DJF             | MAM          | JJA             | SON          |
|-----------------|------------------|-----------------|--------------|-----------------|--------------|
| UV and green:   |                  |                 |              |                 |              |
| correlation     | 0.98             | 0.96            | 0.98         | 0.85            | 0.97         |
| slope           | $0.25 \ (0.005)$ | $0.27 \ (0.02)$ | 0.32(0.01)   | 0.35(0.04)      | 0.25(0.01)   |
| intercept       | 0.89(0.09)       | 0.19(0.41)      | -0.01 (0.21) | $0.05 \ (0.38)$ | 0.76(0.19)   |
| blue and green: |                  |                 |              |                 |              |
| correlation     | 0.995            | 0.98            | 0.99         | 0.95            | 0.997        |
| slope           | 0.67(0.01)       | 0.64(0.02)      | 0.70(0.02)   | $0.70 \ (0.05)$ | 0.67 (0.01)  |
| intercept       | -0.45 (0.11)     | $0.41 \ (0.62)$ | -0.66 (0.22) | -0.91 (0.44)    | -0.43 (0.17) |

Table S1. Correlation, slope, and intercept (in  $1e15 \text{ molec cm}^{-2}$ ) of SCDs between the UV, blue, and green spectral range for China. In the brackets the standard errors are shown.

**Table S2.** Correlation, slope, and intercept (in  $1e15 \text{ molec cm}^{-2}$ ) of VCDs between the UV, blue, and green spectral range for China. In the brackets the standard errors are shown.

|                 | Annual       | DJF             | MAM            | JJA          | SON             |
|-----------------|--------------|-----------------|----------------|--------------|-----------------|
| UV and green:   |              |                 |                |              |                 |
| correlation     | 0.98         | 0.95            | 0.98           | 0.94         | 0.98            |
| slope           | 0.73(0.01)   | 0.70(0.04)      | $076 \ (0.03)$ | 0.75~(0.05)  | $0.67 \ (0.03)$ |
| intercept       | -0.51 (0.16) | $0.20 \ (0.77)$ | -0.76(0.36)    | -0.69(0.40)  | $0.06\ (0.33)$  |
| blue and green: |              |                 |                |              |                 |
| correlation     | 0.98         | 0.94            | 0.98           | 0.92         | 0.99            |
| slope           | 1.07 (0.02)  | $0.94 \ (0.07)$ | 1.04 (0.04)    | 0.997~(0.08) | 1.13(0.04)      |
| intercept       | -0.95 (0.27) | 1.29(1.16)      | -0.91 (0.49)   | -0.79 (0.64) | -1.19 (0.44)    |