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- 2 Raindrop Fall Velocities from an Optical Array Probe and 2D-Video
- 3 Disdrometer
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12 Abstract

We report on fall speed measurements of rain drops in light-to-heavy rain events from 13 two climatically different regimes (Greeley, Colorado, and Huntsville, Alabama) using 14 the high resolution (50 µm) Meteorological Particle Spectrometer (MPS) and a 3rd 15 generation (170 µm resolution) 2D-video disdrometer (2DVD). To mitigate wind-effects, 16 especially for the small drops, both instruments were installed within a 2/3-scale Double 17 Fence Intercomparison Reference (DFIR) enclosure. Two cases involved light-to-18 moderate wind speeds/gusts while the third case was a tornadic supercell and several 19 squall-lines that passed over the site with high wind speeds/gusts. As a proxy for 20 turbulent intensity, maximum wind speeds from 10-m height at the instrumented site 21 recorded every 3 s were differenced with the 5-min average wind speeds and then 22 squared. The fall speeds versus size from 0.1-2 mm and >0.7 mm were derived from 23 the MPS and the 2DVD, respectively. Consistency of fall speeds from the two 24 instruments in the overlap region (0.7-2 mm) gave confidence in the data quality and 25 processing methodologies. Our results indicate that under low turbulence, the mean fall 26 27 speeds agree well with fits to the terminal velocity measured in the laboratory by Gunn and Kinzer from 100 µm up to precipitation sizes. The histograms of fall speeds for 0.5, 28 29 0.7, 1 and 1.5 mm sizes were examined in detail under the same conditions. The histogram shapes for the 1 and 1.5 mm sizes were symmetric and in good agreement 30 31 between the two instruments with no evidence of skewness or of sub- or super-terminal fall speeds. The histograms of the smaller 0.5 and 0.7 mm drops from MPS while 32 generally symmetric showed that occasional occurrences of sub- and super-terminal fall 33 speeds could not be ruled out. In the supercell case, the very strong gusts and inferred 34 high turbulence intensity caused a significant broadening of the fall speed distributions 35 with negative skewness (for drops of 1.3, 2 and 3 mm). The mean fall speeds were also 36 found to decrease nearly linearly with increasing turbulent intensity attaining values 37 about 25-30% less than the terminal velocity of Gunn-Kinzer, i.e. sub-terminal fall 38 39 speeds.

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42 1 Introduction

Knowledge of the terminal fall speed of raindrops as a function of size is important in 43 44 modelling collisional break-up and coalescence processes (e.g., List et al., 1987), in the radar-based estimation of rain rate, in retrieval of drop size distribution using Doppler 45 spectra at vertical incidence (e.g., Sekhon and Srivastava, 1971) and in soil erosion 46 studies (e.g., Rosewell 1986). In these and other applications it is generally accepted 47 that there is a unique fall speed ascribed to drops of a given mass or diameter and that 48 it equals the terminal speed with adjustment for pressure (e.g., Beard 1976). 49 The terminal velocity measurements of Gunn and Kinzer, 1949) under calm laboratory 50 conditions, and fits to their data (e.g., Atlas et al., 1973; Foote and du Toit, 1969; Beard 51 and Pruppacher, 1969) are still considered the standard against which measurements 52 using more modern optical instruments in natural rain are compared (Löffler-Mang and 53 Joss, 2000; Barthazy et al., 2004; Schönhuber et al., 2008; Testik and Rahman, 2016; 54 Yu et al., 2016). More recently, the broadening and skewness of the fall speed 55 distributions of a given size (3 mm) in one intense rain event were attributed to mixed-56 57 mode amplitude oscillations (Thurai et al., 2013). Super- and sub-terminal fall speeds in intense rain shafts have been detected and attributed, respectively, to drop breakup 58 59 fragments (sizes < 0.5 mm), and high wind/gusts (sizes 1-2 mm) (Montero-Martinez et al., 2009; Larsen et al., 2014; Montero-Martinez and Garcia-Garcia, 2016). Thus, there 60 is some evidence that rain drops may not fall at their terminal velocity except under 61 calm conditions and that the concept of a fall speed distribution for a drop of given mass 62 (or, diameter) might need to be considered which is the topic of this paper. The 63 implications are rather profound especially for numerical modeling of collision-64 coalescence and breakup processes which are important for shaping the drop size 65 distribution. 66

The fall speeds and concentration of small drops (< 1 mm) in natural rain are difficult to 67 measure accurately given the poor resolution (>170 µm) of most optical disdrometers 68 and/or sensitivity issues. While cloud imaging probes (with high resolution 25-50 µm) 69 on aircraft have been used for many years they generally cannot measure the fall 70 71 speeds. A relatively new instrument, the Meteorological Particle Spectrometer (MPS) is a droplet imaging probe that was built by Droplet Measurements Technologies (DMT, 72 Inc.) under contract from the US Weather Service specifically designed for drizzle as 73 small as 50 µm and rain drops up to 3 mm. This instrument in conjunction with a lower 74 75 resolution 2D-Video Disdrometer (Schoenhuber et al., 2008) is used in this paper to 76 measure fall speed distributions in natural rain.

This paper briefly describes the instruments used, presents fall speed measurements from two sites under relatively low wind conditions, and one case from an unusual tornadic supercell with high winds and gusts and ends with a brief discussion andsummary of the results.

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82 2 Instrumentation and Measurements

The principal instruments used in this study are the MPS and 3rd generation 2D-video 83 disdrometer (2DVD), both located within a 2/3-scale Double Fence Intercomparison 84 85 Reference (DFIR: Rasmussen et al., 2012) wind shield. As reported in (Notaros et al., 2016), the 2/3-scale DFIR was effective in reducing the ambient wind speeds by nearly 86 a factor of 2-3 based on data from outside and inside the fence. The flow field in and 87 88 around the DFIR has been simulated by (Theriault et al., 2015) assuming steady 89 ambient winds. They found that depending on the wind direction relative to the octagonal fence, weak vertical motions could be generated above the sensor areas. For 90 5 m/s speeds, the motions could range between -0.4 (down draft) to 0.2 m/s (up draft). 91

The instrument set-up was the same for the two sites (Greeley, Colorado and Huntsville, Alabama). Huntsville has a very different climate from Greeley, and its altitude is 212 m MSL as compared with 1.4 km MSL for Greeley. According to the Köppen–Trewartha climate classification system (*Trewartha and Horn,* 1980), this labels Greeley as a semiarid-type climate, whereas Huntsville is a humid subtropical-type climate (*Belda et al.,* 2014).

The MPS is an optical array probe (OAP) that uses the technique introduced by 98 Knollenberg (1970, 1976, 1980) and measures drop diameter in the range from 0.05-3.1 99 mm. A 64 element photo-diode array is illuminated with a 660 nm collimated laser 100 101 beam. Droplets passing through the laser cast a shadow on the array and the decrease 102 in light intensity on the diodes is monitored with the signal processing electronics. A two dimensional image is captured by recording the light level of each diode during the 103 period that the array is shadowed. The fall velocity is derived using two methods. One 104 uses the same approach as described by (Montero-Martinez et al., 2009) where the fall 105 velocity is calculated from the product of the true air speed clock and ratio of the image 106 height -to-width. Note that "width" is the horizontal dimension parallel to the array and 107 "height" is along the vertical. The second method computes the fall velocity from the 108 maximum horizontal dimension (spherical drop shape assumption) divided by the 109 amount of time that the image is on the array, a time measured with a 2 MHz clock. In 110 order to be comparable to the results of (Montero-Martinez et al., 2009), their approach 111 112 is implemented here for sizes > 250 μ m. The fall velocity of smaller, slower moving 113 droplets, is measured using the second technique.

The limitations and uncertainties associated with OAP measurements have been well 114 documented (Korolev et al., 1991; 1998; Baumgardner et al., 2017). There are a 115 number of potential artifacts that arise when making measurements with optical array 116 probes (Baumgardner et al., 2017): droplet breakup on the probe tips that form satellite 117 118 droplets, multiple droplets imaged simultaneously, and out-of-focus drops whose images are usually larger than the actual drop (Korolev, 2007). The measured images 119 have been analyzed to remove satellite droplets whose interarrival times are usually too 120 short to be natural drops, multiple drops are detected by shape analysis and removed, 121 122 and out-of-focus drops are detected and size corrected using the technique described by (Korolev 2007). The sizing and fall speed errors primarily depend on the digitization 123 error (\pm 25 µm). The fall speed accuracy according to the manufacturer (DMT) is <10% 124 for 0.25 mm and <1% for sizes greater than 1 mm, limited primarily by the accuracy in 125 126 droplet sizing.

The 3rd generation 2DVD is described in detail by (Schoenhuber et al., 2007; 2008) and 127 its accuracy of size and fall speed measurement has been well documented (e.g., 128 Thurai et al., 2007; 2009; Huang et al., 2008; Bernauer et al., 2015). Considering the 129 horizontal pixel resolution of 170 µm and other factors (such as "mis-matched" drops), 130 the effective sizing range is D> 0.7 mm. To clarify the "mis-matched" drop problem: it is 131 132 very difficult to match a drop detected in the top light-beam plane of the 2DVD to the corresponding drop in the bottom plane for tiny drops resulting in erroneous fall speeds. 133 The fall velocity accuracy is determined primarily by the accuracy of calibrating the 134 distance between the two orthogonal light "sheets" or planes and is < 5% for fall velocity 135 In our application, we utilize the MPS for measurement of small drops 136 <10 m s⁻¹. with D < 1.2 mm. The measurements from the MPS are compared with those from the 137 2DVD in the overlap region of $D \approx 0.7-2.0$ mm to ensure consistency of 138 observations. The only fall velocity threshold used for the 2DVD is the lower limit set at 139 0.5 m s⁻¹ in accordance with the manufacturer guidelines for rain measurements. 140

141 2.1 Fall Speeds from Greeley, Colorado

We first consider a long duration (around 20 h) rain episode on 17 April 2015 which consisted of a wide variety of rain types/rates (mostly light stratiform < 8 mm h⁻¹) as described in Table 2 of (*Thurai et al.*, 2017). Two wind sensors at a height of 1 m were available to measure the winds outside and inside the DFIR. Average wind speeds were, respectively, < 1.5 m s⁻¹ inside the DFIR and < 4 m s⁻¹ outside with light gusts. These wind sensors were specific to the winter experiment described in (*Notaros et al.*, 2016) and were unavailable for the rain measurement campaign after May 2015.

Figure 1(a) shows the fall speeds versus D from the 2DVD (shown as contoured frequency of occurrence), along with mean and $\pm 1\sigma$ standard deviation from the MPS. Also shown is the (*Foote and du Toit 1969*) (henceforth FT fit) to the terminal fall speed

measurements of (Gunn and Kinzer, 1949) at sea level and after applying altitude 152 corrections (*Beard*, 1976) for the elevation of 1.4 km MSL for Greeley. Panels (b,c) 153 shows the histogram of fall speeds for diameter intervals (0.5±0.1) and (1±0.1 mm), and 154 (0.7±0.1) and (1.5±0.1 mm), respectively. Panel (a) demonstrates the excellent "visual" 155 156 agreement between the two instruments in the overlap size range (0.7-2 mm) which is quantified in Table 1. However, the altitude-adjusted FT fit is slightly higher than the 157 measured values as shown in Table 1. Notable in Fig. 1a is the remarkable agreement 158 in mean fall speeds between the FT fit and the MPS for D< 0.5 mm down to near the 159 lower limit of the instrument (0.1 mm). Few measurements have been reported of fall 160 speeds in this size range. 161

Table 1: Expected fall velocities for various diameter intervals (bin width of 0.2 mm) from (*Foote and du Toit*, 1969) with altitude adjustment, and the measured mean fall velocities with $\pm 1\sigma$ (standard deviation)

D range (mm)	Expected (m s ⁻¹)	MPS (m s ⁻¹⁾	2DVD (m s ⁻¹)
(Greeley)	at 1.4 km	Mean $\pm 1\sigma$	Mean $\pm 1\sigma$
0.6 to 0.8	2.6 to 3.5	$\textbf{2.6} \pm \textbf{0.6}$	2.5 ± 0.8
0.8 to 1.0	3.5 to 4.3	$\textbf{3.4}\pm\textbf{0.6}$	3.3 ± 0.9
1.0 to 1.2	4.3 to 4.9	$\textbf{4.2}\pm\textbf{0.6}$	4.1 ± 0.9
1.2 to 1.4	4.9 to 5.5	4.9 ± 0.5	5.0 ± 0.8
1.4 to 1.6	5.5 to 6.1	5.6 ± 0.5	5.7 ± 0.7
1.6 to 1.8	6.1 to 6.6	$\textbf{6.1}\pm\textbf{0.4}$	6.2 ± 0.7
1.8 to 2.0	6.6 to 7.0	$\textbf{6.7}\pm\textbf{0.4}$	6.6 ± 0.8
D range (mm)	Expected (m/s)	MPS (m/s)	2DVD (m/s)
(Huntsville)	at 0 km	Mean \pm Std_dev	Mean ± Std. dev
0.6 to 0.8	2.5 to 3.3	$\textbf{2.6}\pm\textbf{0.6}$	2.5 ± 0.7
0.8 to 1.0	3.3 to 4.0	$\textbf{3.4}\pm\textbf{0.5}$	$\textbf{3.3}\pm\textbf{0.7}$
1.0 to 1.2	4.0 to 4.6	$\textbf{4.2}\pm\textbf{0.6}$	4.1 ± 0.8
1.2 to 1.4	4.6 to 5.2	$\textbf{4.9}\pm\textbf{0.4}$	4.9 ± 0.7
1.4 to 1.6	5.2 to 5.7	5.4 ± 0.4	5.4 ± 0.6
1.6 to 1.8	5.7 to 6.1	6.0 ± 0.3	5.8 ± 0.6
1.8 to 2.0	6.1 to 6.5	6.5 ± 0.4	6.3 ± 0.5

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The histograms in Fig. 1(b,e) show good agreement between 2DVD and MPS for 1 mm and 1.5 mm drop sizes, respectively, with respect to the mode, symmetry, spectral width and lack of skewness in the distributions. For the 1 mm size histogram, the mean is 3.8 m s⁻¹ while the spectral width or standard deviation from MPS data is 0.6 m s⁻¹. The corresponding coefficient of variation (ratio of standard deviation to mean) is 15.7%. The finite bin width used (0.9-1.1 mm) causes a corresponding fall speed "spread" of around 0.6 m s⁻¹ which is clearly a significant contributor to the measured coefficient of

variation. Similar comments apply to the fall speed histogram for the 1.5 mm size shown 173 in Fig. 1c. The definition of sub- or super-terminal fall speeds by (Montero-Martinez et 174 al., 2009) is based on fall speeds that are, respectively, less than 0.7 times the mean 175 value or greater than 1.3 times the mean value (i.e., exceeding 30% threshold on either 176 side of the mean terminal fall speed). From examining the 1 mm size fall speed 177 histogram there is negligible evidence of occurrences with fall speeds < 2.66 m s⁻¹ 178 (sub) or > 4.94 m s⁻¹ (super). Similar comment also applies for the 1.5 mm size based 179 on the corresponding histogram. 180



182 <u>Figure 1.</u> (a) Fall velocity versus diameter (D). The contoured frequency of occurrence from 183 2DVD data is shown in color (log scale). The mean fall velocity and $\pm 1\sigma$ standard deviation bars 184 are from MPS. The dark dashed line is from the fit to the laboratory data of Gunn and Kinzer 185 (1949) and the purple line is the same except corrected for the altitude of Greeley, CO (1.4 km

186 *MSL*). (b) Relative frequency histograms of fall velocity for the 0.5 ± 0.1 mm and 1 ± 0.1 mm 187 bins.(c) as in (b) except for the 0.7 ± 0.1 mm and 1.5 ± 0.1 mm bins.

The histogram from MPS for the 0.5 mm sizes shows positive skewness with mean of 188 1.8 m s⁻¹, spectral width of 0.65 m s⁻¹ and corresponding coefficient of variation nearly 189 doubling to 35% (relative to the 1 mm size histogram). The finite bin width (0.4-0.6 mm) 190 causes a corresponding fall speed "spread" of 0.4 m s⁻¹ which contributes to the 191 measured coefficient of variation. Nevertheless, it is not possible to rule out the low 192 frequency of occurrence of sub- or super-terminal fall speeds, respectively, less than 193 1.26 m s⁻¹ or exceeding 2.34 m s⁻¹ based on our data. Examination of the MPS-based 194 fall speed histogram for the 0.7 mm size indicates negative skewness. As with the 0.5 195 mm drops it is not possible to rule out the occurrences of fall speeds < 1.8 m s⁻¹ or > 3.4196 m s⁻¹, i.e., sub- or super-terminal fall speeds. 197

198 2.2 Fall Speeds from Huntsville, Alabama

The first Huntsville event occurred on 11 April 2016 and consisted of precipitation 199 associated with the mesoscale vortex of a developing squall line that moved across 200 northern Alabama between 1800 and 2300 UTC and produced over 25 mm of rainfall in 201 the Huntsville area. Figure 2(a) shows the ambient 10-m height wind speeds (3 s and 5-202 min averaged) recorded at the site. Maximum speeds were less than 5 m s⁻¹ and wind 203 gusts were light. As no direct in situ measurement of turbulence was available we use 204 the approach by (Garrett and Yuter, 2014) who estimate the difference between the 205 maximum wind speed, or gust, that was sampled every 3 s, and the average wind 206 speed derived from successive 5 min intervals. The estimated turbulent intensity is 207 proportional to $E = (Gusts - AverageWind)^2/2$. Figure 2(b) shows the E values which 208 were small (maximum $E < 0.4 \text{ m}^2 \text{ s}^2$) and indicative of low turbulence. Also, shown in 209 Fig. 2(b) is the 2DVD-based time series of rainfall rate (R) averaged over 3 mins; the 210 maximum R is around 10 mm h^{-1} . 211



213 <u>Figure 2</u>: (a) 3-s raw and 5-min averaged wind speeds at 10-m height. (b) turbulent 214 intensity estimates E, and 3-min averaged R.

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Figure 3(a) shows the fall velocity versus D comparison between the two instruments while panels (b,c) show the histograms for the 0.5 and 1 mm, and 0.7 and 1.5 mm sizes, respectively. Similar to the Greeley event, the mean fall speed agreement between both instruments in the overlap region is excellent (see Table 1) and consistent with the FT fit to the Gunn-Kinzer laboratory data. As in Fig. 1(a), the MPS data in Fig. 3(a) is in excellent agreement with FT fit for sizes < 0.5 mm.

The 0.5 and 1 mm histogram shapes in Fig. 3(b) are quite similar to the Greeley case 222 shown in Fig. 1(b). The mean and standard deviations from the MPS data for the 0.5 223 and 1 mm bins are, respectively, $[2 \pm 0.62]$ and $[3.88 \pm 0.44]$ m s⁻¹. The values for the 224 0.7 and 1.5 mm bins are, respectively, [2.6 \pm 0.6] and [5.4 \pm 0.4] m s⁻¹. There is 225 negligible evidence of sub- or super-terminal fall speed occurrences based on the 1 and 226 1.5 mm histograms. The comments made earlier with respect to Fig. 1(b,c) of the 227 Greeley event for the 0.5 and 0.7 mm histograms are also applicable here, i.e., we 228 cannot rule out the occasional occurrences of sub- or super-terminal fall speeds based 229 230 on our data.



<u>Figure 3</u>. (a) as in Fig. 1(a) except for 11 April 2016 event. The dashed line is fit to Gunn-Kinzer at sea level. (b,c) as in Fig. 1(b,c) except for 11 April 2016 event.

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The second case considered is from 30 November 2016 wherein a supercell passed 235 over the instrumented site from 0300-0330 UTC producing about 15 mins later a long-236 lived EF-2 tornado. Strong winds were recorded at the site with 5-min averaged speeds 237 reaching 10-12 m s⁻¹ between 0320-0330 and E values in the range to 7-8 m² s⁻² 238 indicating strong turbulence (Fig. 4a,b). The rain rates peaked at 70 mm h⁻¹ during this 239 time (Fig. 4b). About 3 h later several squall-line type storm cells passed over the site 240 from 0700-0900 UTC again with strong winds but considerably lower E values 2-4 m² s⁻² 241 and maximum R of 80 mm h⁻¹. After 1000 UTC the E values were much smaller (< 0.5 242 $m^2 s^{-2}$) indicating calm conditions. The peak R is also smaller at 30 mm h^{-1} at 1000 UTC. 243

Figure 4 panels (c), (d) and (e) show the mean and $\pm 1\sigma$ of the fall speeds from the 245 2DVD for the 1.3, 2 and 3 mm drop sizes, respectively. The MPS data are not shown here since during this event it was located outside the DFIR on its turntable and we did
not want to confuse the wind-effects between the two instruments. It is clear from Fig.
4(c) that during the supercell passage (0300-0330 UTC) the mean fall speed for 1.3 mm
drops decreases (from 5 to 3.5 m s⁻¹) and the standard deviation increases (from 0.5 to
1.5 m s⁻¹). The histogram shapes also show increasing negative skewness (not shown).
The same trend can be seen for the subsequent squall-line rain cell passage from 07000900 UTC. Similar trends are noted in panels (d) and and less so in panel (e).

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255 <u>Figure 4</u>. (a) as in Fig. 2(a) except for 30 Nov 2016 event. (b) as in Fig. 2(b). (c) mean 256 and $\pm 1\sigma$ standard deviation of fall speeds from 2DVD for 1.3 \pm 0.1 mm sizes. (d,e) as in

257 (c) except for 2 ± 0.1 and 3 ± 0.1 mm sizes, respectively.

To expand on this observed correlation, Fig. 5 shows scatterplots of the mean fall speed 259 and standard deviation versus E for the 1.3 mm drops (panels a,b), while panels (c,d) 260 and (e,f) show the same but for the 2 and 3 mm drops, respectively. The mean fall 261 speed decreases with increasing *E* nearly linearly for $E > 1 \text{ m}^2 \text{ s}^{-2}$ but less so for the 3 mm 262 size drops (Stout et al., 1995). This decrease relative to Gunn-Kinzer terminal fall 263 speeds is termed as "sub-terminal" and our data is in general agreement with (Montero-264 Martinez and Garcia-Garcia 2016) who found an increase in the numbers of sub-265 terminal drops with sizes between 1-2 mm under windy conditions using a 2D-266 Precipitation probe with resolution of 200 µm (similar to 2DVD) but without a wind fence. 267 The standard deviation of fall speeds (σ_f) versus E is shown in panels 5 (b,d,f). When 268 E>1 m² s⁻², the σ_f is nearly constant at 1.5 m s⁻¹ for both 1.3 and 2 mm drop sizes and 269 constant at 1 m s⁻¹ for the 3 mm size. For E < 1, the σ_f is more variable and essentially 270 uncorrelated with E. From the discussion related to Fig. 1(b,c) and 3(b,c), σ_f values 271 exceeding approximately 0.5 m s⁻¹ can be attributed to physical, not instrumental or 272 finite bin width effects (see, also, Table 1). Thus, the fall speed distributions are 273 considerably broadened when $E>1 \text{ m}^2 \text{ s}^{-2}$ due to increasing turbulence levels which is 274 again consistent with the findings of (Montero-Martinez and Garcia-Garcia, 2016) as 275 well as those of (Garett and Yuter, 2014). The latter observations, however, were of 276 graupel fall speeds in winter precipitation using a multi-angle snowflake camera (Garrett 277 et al., 2012). 278



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<u>Figure 5.</u> (a,b) mean fall speed and standard deviation, respectively, versus E for 1.3
 mm sizes. (c,d) same but for 2 mm sizes. (e,f) same but for 3 mm.

283

284 3 Discussion and Conclusions

We have reported on raindrop fall speed distributions using a high resolution (50 μ m) droplet spectrometer (MPS) collocated with moderate resolution (170 μ m) 2DVD (with both instruments inside a DFIR wind shield) to cover the entire size range (from 0.1 mm onwards) expected in natural rain. Turbulence intensity (*E*) was derived from wind/gust

data at 10-m height following (Garrett and Yuter, 2014). For low turbulent intensities (E 289 $< 0.4 \text{ m}^2 \text{ s}^{-2}$), in the overlap region of the two instruments (0.7-2 mm), the mean fall 290 speeds were in excellent agreement with each other for both the Greeley, CO and 291 Huntsville, AL sites giving high confidence in the quality of the measurements. For 292 293 D<0.5 mm and down to 0.1 mm, the mean fall speeds from MPS from both sites were in remarkable agreement with the (Foote and du Toit, 1969) fit to the laboratory data of 294 (Gunn and Kinzer, 1949). In the overlap region, the mean fall speeds from the two 295 instruments were in excellent agreement with the FT fit for the Huntsville site (no 296 297 altitude adjustment required) and good agreement for the Greeley site (after adjustment for altitude of 1.4 km). For D>2 mm, the mean fall speeds from 2DVD were in excellent 298 299 agreement with the FT fit at both sites.

Our histograms of fall speeds for 1 and 1.5 mm sizes under low turbulence intensity 300 conditions ($E < 0.4 \text{ m}^2 \text{ s}^2$) from both MPS and 2DVD were in good agreement and did 301 not show any evidence of either sub- or super-terminal speeds, rather the histograms 302 were symmetric with mean close to the Gunn-Kinzer terminal velocity with no significant 303 broadening over that ascribed to instrument and/or finite bin width effects. (Note: sub-304 terminal implies fall speeds < 0.7 times the terminal fall speed whereas super-terminal 305 implies > 1.3 times terminal value; Montero-Martinez et al., 2009). However, for the 0.5 306 and 0.7 mm sizes, from the histogram of fall speeds using the MPS under the same 307 conditions occasional occurrences of both sub- and super-terminal fall speeds, after 308 accounting for instrumental and finite bin width effects, cannot be ruled out. 309

The only comparable earlier study is by (Montero-Martinez et al., 2009) who used 310 collocated 2D-cloud and precipitation probes (2D-C, 2D-P) but restricted their data to 311 calm wind conditions. Their main conclusion was that the distribution of the ratio of the 312 measured fall speed to the terminal fall speed for 0.44 mm size, while having a mode at 313 1 m s⁻¹ was strongly positively skewed with tails extending to 5 m s⁻¹ especially at high 314 rain rates. In our data for the 0.5 and 0.7 mm sizes shown in Fig. 1(b,c) and 3(b,c), no 315 such strong positive skewness was observed in the fall speed histograms, and the 316 317 corresponding ratio of MPS-measured fall speeds to terminal values does not exceed 1.5 to 2. 318

Another study by *Larsen et al.,* (2014) appears to confirm the ubiquitous existence of super-terminal fall speeds for sizes < 1 mm using different instruments one of which was a 2DVD similar to the one used in this study. However, it is well-known that "mismatched" drops cause erroneous fall speed estimates from 2DVD for drops <0.5 mm (*Schoenhuber et al.,* 2008; Appendix in *Huang et al.,* 2010; *Bernauer et al.,* 2015). It is not clear if (*Larsen et al.,* 2014) accounted for this problem in their analysis. In addition, their 2DVD was not located within a DFIR-like wind shield.

In a later study using only the 2D-P probe, (Montero-Martinez and Garcia-Garcia, 2016) 326 found sub-terminal fall speeds and broadened distributions under windy conditions for 327 1-2 mm sizes in general agreement with our results using the 2DVD. Stout et al., (1995) 328 simulated the motion of drops subject to non-linear drag in isotropic turbulence and 329 330 determined that there would be a significant reduction of the average drop settling 331 velocity (relative to terminal velocity) of greater that 35% for drops around 2 mm size when the ratio of rms velocity fluctuations (due to turbulence) relative to drop terminal 332 velocity is around 0.8. Whereas we did not have a direct measure of the rms velocity 333 334 fluctuations, the proxy for turbulence intensity (E) related to wind gusts during supercell passage (very large E around 7 m² s⁻²) and two squall-line passages (moderate E 335 between 2-5 m² s⁻²) clearly showed a significant reduction in mean fall speeds of 25-336 30% relative to terminal speed for 1.3 and 2 mm sizes (and less so for 3 mm drops), 337 with significant broadening of the fall speed distributions relative to calm conditions by 338 nearly a factor of 1.5 to 2. 339

While our dataset is limited to three events they cover a wide range of rain rates, wind conditions and two different climatologies. One caveat is that the response of the DFIR wind shield to ambient winds in terms of producing subtle vertical air motions near the sensor area is yet to be evaluated as future work. Analysis of further events with direct measurement of turbulent intensity, for example using a 3D-sonic anemometer at the height of the sensor, would be needed to generalize our findings.

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- 347 Data Availability
- 348 Data used in this paper can be accessed from:
- 349 ftp://lab.chill.colostate.edu/pub/kennedy/merhala/Bringi_et_al_2017_GRL_datasets/
- 350

351 Competing interests

VNB and MT declare they have no conflict of interest. DB is employed by Droplet Measurements Technologies, Inc. (Longmont, Colorado, USA) who manufacture the Meteorological Particle Spectrometer used in this study.

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