



- 1 Simultaneous observations by sky radiometer and MAX-
- 2 DOAS for characterization of biomass burning plumes in
- 3 central Thailand in January-April 2016
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# 16 Abstract

17 The first intensive multi-component ground-based remote sensing observations by sky 18 radiometer and Multi-Axis Differential Optical Absorption Spectroscopy (MAX-DOAS) were 19 performed simultaneously at the SKYNET/Phimai site located in central Thailand (15.18°N, 102.56°E) from January to April 2016. The period corresponds to the dry season associated 20 21 with the intense biomass burning (BB) activity around the site. The near-surface concentration of formaldehyde (HCHO) retrieved from MAX-DOAS was found to be a useful tracer for BB 22 23 plumes. As the HCHO concentration tripled from 3 to 9 ppbv, the ratio of gaseous glyoxal to 24 HCHO concentrations in daytime decreased from ~0.04 to ~0.03, responding presumably to the increased contribution of volatile organic carbon emissions from BB. In addition, clear 25 26 increases in aerosol absorption optical depths (AAODs) retrieved from sky radiometer observations were seen with the HCHO enhancement. At a HCHO of 9 ppbv, AAOD at a 27 28 wavelength of 340 nm reached as high as  $\sim 0.15 \pm 0.03$ . The wavelength dependence of AAODs 29 at 340-870 nm was quantified by the absorption Ångström exponent (AAE), providing evidence





for the presence of brown carbon aerosols at an AAE of 1.5±0.2. Thus, our multi-component
 observations around central Thailand are expected to provide unique constraints for
 understanding physical/chemical/optical properties of BB plumes.

4

## 5 1 Introduction

6 It is well recognized that aerosols contribute the largest uncertainty to the estimate of radiative 7 forcing (e.g., IPCC, 2013). Biomass burning (BB) is a substantial source of aerosols to the 8 atmosphere. Black carbon (BC) is a strongly-light-absorbing component of aerosols and can be 9 emitted in large quantities from BB. In addition, about two-thirds of the global primary organic aerosol (OA) that should comprise a large amount of ultraviolet (UV)-light-absorbing OA, 10 11 known as brown carbon (BrC), originates from BB plumes (Bond et al., 2013). Currently, most 12 climate models treat OA as purely scattering. Recent laboratory studies suggested that BrC can 13 enhance net absorption by OA, potentially altering the BB direct radiative forcing from negative 14 to positive (Kirchstetter et al., 2004; Saleh et al., 2014). Moreover, underestimation in aerosol 15 absorption over most BB regions was reported by Hammer et al. (2016), who used a global 3-16 D chemistry transport model (GEOS-Chem), in which OA is regarded as purely scattering. Thus, 17 the potential climate effects of BrC aerosols have received considerable attention recently (e.g., 18 Myhre et al., 2013). In addition, as a result of UV absorption, tropospheric photochemistry can 19 be significantly affected; GEOS-Chem simulation incorporating UV absorption by BrC showed 20 a decrease in tropospheric hydroxyl radical (OH) concentration by up to 15% for BB regions, 21 compared to the simulation without UV absorption by BrC (Hammer et al., 2016). BrC 22 comprises a wide range of poorly characterized compounds that exhibit highly variable 23 absorptivity. Assessing the role of BrC in light absorption is further difficult, because BrC is 24 not only emitted as a primary organic aerosol (POA) but also produced as a secondary organic 25 aerosol (SOA) through complex formation processes from volatile organic compounds (VOCs) 26 originating from BB.

This study focuses on the intense BB activity that occurred around central Thailand from January to April 2016. Characterization for the BB plumes is attempted using our unique remote sensing observations by the sky radiometer (*e.g.*, Nakajima et al., 1996) and the Multi-Axis Differential Optical Absorption Spectroscopy (MAX-DOAS) (*e.g.*, Irie et al., 2011) for both viewpoints of the optical properties of aerosols (aerosol absorption optical depth, AAOD and





- 1 absorption Ångström exponent, AAE) and the organic gas concentrations (formaldehyde,
- 2 HCHO and glyoxal, CHOCHO) in BB plumes.
- 3

#### Observation 4 2

### 2.1 Sky radiometer observation of aerosol optical properties 5

6 We conducted ground-based remote sensing observations using the sky radiometer and the 7 MAX-DOAS at SKYNET/Phimai site (15.18°N, 102.56°E) located in central Thailand from 8 January to April 2016. The period corresponds to the dry season with the occurrence of intense 9 BB around the site. Indeed, satellite data revealed evident enhancements in the carbon 10 monoxide column concentration and the fire radiative power (FRP) around the 11 SKYNET/Phimai site in the dry season (Hoque et al., 2018). The sky radiometer (POM-02; 12 Prede Co., Ltd, Tokyo, Japan), a sun-sky photometer measuring direct and diffuse solar 13 irradiances, is the main instrument of the international ground-based remote sensing network 14 SKYNET (e.g., Takamura and Nakajima, 2004; Nakajima et al., 2007). Measurements of the 15 direct solar and diffuse irradiances within 160° of the center of the Sun were conducted every 16 10 min. Values of aerosol optical depth (AOD), single scattering albedo (SSA), refractive index 17 at 340, 380, 400, 500, 675, and 870 nm were retrieved using the Sky Radiometer analysis 18 package from the Center for Environmental Remote Sensing (SR-CEReS) version 1 (Mok et 19 al., 2018), in which SKYRAD.pack version 5 (Hashimoto et al., 2012) is implemented to 20 retrieve aerosol properties, along with all pre- and post-processing programs for the purpose of 21 the near-real time data delivery. Data at 1020 nm were not used in this study to avoid possible 22 impact by low AAOD and interferace by water vapor (H2O) on the estimate of the AAE. Cloud 23 screening was made by the method of Khatri and Takamura (2009) but without using global 24 irradiance data from a pyranometer.

25 The SKYNET/sky-radiometer has on-site calibration methods, namely the Improved Langley 26 (IL) method determing the calibration constant ( $F_0$ ) (e.g., Campanelli et al., 2007) and the Solar 27 Disk Scan (SDS) method determing the solid view angle (SVA) (e.g., Nakajima et at., 1996; 28 Uchiyama et al., 2018). Recently, Mok et al. (2018) used retrievals with SR-CEReS to compare 29 SKYNET/sky-radiometer AOD and SSA data with those derived from a combination of 30 Aerosol Robotic Netwowk (AERONET), Multifilter Rotating Shadowband Radiometer 31 (MFRSR), and Pandora observations in Seoul, Korea during and after NASA KORUS-AQ 32 (Korea U.S.-Air Quality) campaign in 2016 (Mok et al., 2018 and references therein). For most





cases, their agreements were found to be within ±0.01 and ±0.05 for AOD and SSA data,
 respectively, at all wavelengths from 340 and 870 nm, supporting the ability of the on-site
 calibration methods using IL and SDS.

4 Since the importance of accurate SVA determination was particularly pointed out to better 5 interpret the difference seen in previous SSA comparisons between SKYNET and AERONET 6 (Khatri et al., 2016), sensitivity analysis was made in the present study by conducting additional 7 retrievals using SVAs offset by  $\pm 0.01$  msr ( $\sim \pm 4\%$ ), which is likely to correspond to the 8 uncertainty in SVA determined by a single SDS. Both positive and negative offsets were tested 9 but only the positive offset is discussed here, because the negative offset tented to show only 10 little or no impact on SSA, when SSA was close to unity. This is because a smaller SVA leads 11 to a larger SSA (Hashimoto et al., 2012). The impacts by the SVA offset of +0.01 msr on SSAs 12 13 0.012±0.007, and -0.011±0.008 at 340, 380, 400, 500, 675, and 870 nm, respectively. Thus, 14 overestimation (underestimation) in SVA leads to underestimation (overestimation) in SSA, 15 but the magnitude was found to be very small at about  $\pm 0.01$ , when the uncertainty in SVA was 16  $\sim \pm 0.01$  msr. The small impact on SSAs should be a result of compensation by the associated 17 change in  $F_0$  values; using SVA values offset by +0.01 msr as an input, the IL method employed 18 in SR-CEReS yields smaller  $F_0$  values by about 2.1±0.1%, 1.8±0.2%, 1.7±0.2%, 1.2±0.2%, 19 0.7±0.2%, and 0.5±0.1% for 340, 380, 400, 500, 675, and 870 nm, respectively. The resulting 20 smaller  $F_0$  leads to a larger SSA (Hashimoto et al., 2012), which is an opposite trend of the 21 direct impact that a larger SVA leads to a smaller SSA (Hashimoto et al., 2012). Results from 22 these sensitivity analyses support the agreement of SSAs within  $\pm 0.05$  seen in recent 23 comparisons by Mok et al. (2018) during and after NASA KORUS-AQ campaign.

Using the AOD and SSA data retrieved, AAOD and AAE values were derived as follows. First, for each measurement and for respective wavelengths from 340 to 870 nm, the AAOD value and its uncertainty ( $\varepsilon_{AAOD}$ ) were calculated with the following equations:

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28 
$$AAOD(\lambda) = AOD(\lambda) \cdot [1 - SSA(\lambda)]$$
 (1)

29 
$$\varepsilon_{AAOD(\lambda)} = \sqrt{\left[\left(1 - SSA(\lambda)\right) \cdot \varepsilon_{AOD}\right]^2 + (AOD(\lambda) \cdot \varepsilon_{SSA})^2}$$
(2)







1 For the estimate of  $\varepsilon_{AAOD(\lambda)}$ , uncertainties for AOD( $\lambda$ ) and SSA( $\lambda$ ) ( $\varepsilon_{AOD}$  and  $\varepsilon_{SSA}$ ) were assumed 2 to be 0.01 and 0.05, respectively, based on comparisons by Mok et al. (2018). Since the 3 comparisons by Mok et al. (2018) were made using independent measurements having uncertainties of the same order of magnitude, the actual uncertainties in sky radiometer AOD 4 5 and SSA data would be smaller. AAE is calculated as the slope of the linear fit of  $\ln[(AAOD(\lambda)]]$ 6 versus  $ln(\lambda)$ : 7  $\ln[(AAOD(\lambda)] = a - AAE \cdot \ln(\lambda),$ 8 (3) 9 10 where *a* is an intercept. This equation is equivalent to expression using the power law: 11  $AAOD(\lambda) = K\lambda^{-AAE}$ , 12 (4) 13 14 where K is a constant. To exclude AAE data associated with large uncertainty, only the data,

15 which satisfy the criteria that 1) AAOD( $\lambda$ ) exceeds  $\varepsilon_{AAOD(\lambda)}$  for all wavelengths and 2) the 16 correlation coefficient of the linear fit (R) is high (-1.0  $\leq R \leq -0.9$ ) are used in the analysis below. 17 To refine the data set of AOD, SSA, AAOD, and AAE with reduced uncertainty, the daily mean 18 and its standard deviation with the number of data more than 4 were calculated for 9:00-15:00 19 LT.

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#### 21 2.2 MAX-DOAS observation of trace gases

The MAX-DOAS is an instrument measuring UV-visible spectra of scattered sunlight at several 22 elevation angles between the horizon and zenith (e.g., Hönninger and Platt, 2002; Hönninger et 23 24 al., 2004; Irie et al., 2015). Its measurement is based on the well-established DOAS technique 25 that quantitatively detects narrow band absorption by trace gases by applying Lambert-Beer 26 law (e.g., Platt and Stutz, 2008). Since the pioneering study by Hönninger and Platt (2002) and 27 Hönninger et al. (2004), various types of instruments and algorithms for MAX-DOAS have 28 been developed worldwide. One of the reasons for this is because ground-based MAX-DOAS 29 observations at a low elevation angle provide enhanced signals of concentrations of important





trace gases in the boundary layer (*i.e.*, around the instrument altitude) and the concentrations can be interpreted as being the average over a distance, which is on the same order of or finer than the horizontal resolution of models and satellite observations but coarser that that of in situ observations (Irie et al., 2011).

5 From January to April 2016, our MAX-DOAS system (PREDE, Co., Ltd) (Irie et al., 2011; 6 Hoque et al., 2018) was operated continuously at the SKYNET/Phimai site together with the 7 sky radiometer. The spectrometer Maya2000Pro (Ocean Optics, Inc.) was used to record high-8 resolution spectra (with the full width at half maximum of around 0.3-0.4 nm) from 310 to 515 nm. Measurements were made at six elevation angles of 2°, 3°, 4°, 6°, 8°, and 70° every 30 min. 9 10 MAX-DOAS off-axis elevation angle measurements were limited to below 10° for minimizing 11 the possible systematic error in oxygen collision complex fitting results but keeping the 12 measurement sensitivity in the lowest layer of vertical profiles retrieved high (Irie et al., 2015). 13 Spectral analysis by the so-called DOAS method (Platt and Stutz, 2008) for spectral fitting 14 using the nonlinear least-squares method and the subsequent vertical profile retrievals using the 15 optimal estimation method were performed by our retrieval algorithm, JM2 (Japanese MAX-16 DOAS profile retrieval algorithm, version 2) (e.g., Irie et al., 2008; Irie et al., 2011; Irie et al., 2015). Using the recorded high-resolution UV-visible spectra from 310 to 515 nm, the JM2 17 18 allows us to retrieve lower-tropospheric vertical profile information for 8 quantities, including 19 HCHO, CHOCHO, nitrogen dioxide (NO2), and H2O concentrations, which are analyzed below. 20 Of vertical profiles retrieved, the present study analyzed data for a layer of 0-1 km, which 21 corresponds to the lowest layer with the highest sensitivity owing to the longest light path in 22 profiles retrieved by JM2. The total uncertainties, including random and systematic errors, were 23 estimated to be 24% (HCHO), 19% (CHOCHO), 15% (NO<sub>2</sub>), and 18% (H<sub>2</sub>O) (Irie et al., 2011). 24 Using the retrieved H<sub>2</sub>O concentration, the relative humidity over water (RH<sub>w</sub>) for the layer 0-25 1 km was estimated using NCEP (National Centers for Environmental Prediction) pressure and 26 temperature data. To be consistent with sky radiometer data, the daily mean values for 9:00-27 15:00 LT are analyzed in this study. More detailed descriptions about MAX-DOAS are given 28 by Irie et al. (2015), Hoque et al. (2018), and references therein.

29

# **30 3 Results and Discussion**

31 Time series of multi-components retrieved from sky radiometer and MAX-DOAS observations

32 at the SKYNET/Phimai site for the intense biomass burning period from January to April 2016





1 is shown in Fig. 1. As seen in Fig. 1, RH<sub>w</sub> for a layer of 0-1 km derived from MAX-DOAS 2 observations and the surface RHw from NCEP data confirm that the period was dry around 3 Phimai, particularly from the beginning of February through the middle of April (from days 32 to 110). For the period of January-April 2016, mean AOD values at 340, 500, and 865 nm were 4 5 high at 0.98±0.41, 0.64±0.27, and 0.27±0.11, respectively. The AOD values reached the peak 6 in late March (around days 80-85), when AAOD values and HCHO, CHOCHO, and NO2 7 concentrations were synchronously high. From Fig. 1, positive correlations among them are 8 suggested.

9 In Fig. 2, CHOCHO concentrations, ratios of CHOCHO to HCHO concentrations (R<sub>GF</sub>), NO<sub>2</sub> 10 concentrations, and AAOD values are plotted against the HCHO concentration. As a trace gas 11 having the longest lifetime in the three potential BB-originating trace gases investigated here 12 (i.e., HCHO, CHOCHO, and NO<sub>2</sub>), HCHO was chosen as the standard. We found tight 13 correlations between CHOCHO and HCHO concentrations. The  $R_{GF}$  was suggested to vary 14 responding to different VOC emissions such as BB and biogenic activities (e.g., Hoque et al., 15 2018). As the HCHO concentration tripled from 3 to 9 ppbv, the  $R_{\rm GF}$  decreased from ~0.04 to 16  $\sim 0.03$  and the NO<sub>2</sub> concentration doubled from  $\sim 0.6$  to  $\sim 1.2$  ppby, responding presumably to 17 the increased contribution of VOC emissions from BB. The R<sub>GF</sub> values are slightly greater than 18 those reported by Hoque et al. (2018), whose statistics included data taken in early morning and 19 late evening, when R<sub>GF</sub> values were tended to be low compared to mid-day values analyzed in 20 the present study. At a HCHO concentration of 9 ppbv, AAOD at 340 nm reached as high as 21 ~0.15±0.03. Much larger AAODs were seen at a HCHO concentration higher than 9 ppbv (Fig. 22 2). These results provide strong observational evidence that aerosols in BB plumes (*i.e.*, POA 23 and SOA) are absorptive. In addition, Fig. 2 reveals that HCHO is a good tracer for absorption 24 aerosols from BB.

25 While BC has been shown to have an AAE of about unity in literature, AAE values greater than 26 unity are interpreted as BrC (e.g., Kirchstetter et al., 2004). For the whole period from January 27 to April 2016, the mean AAE was estimated to be 1.57±0.28 for the entire wavelength region 28 from 340 to 870 nm (Fig. 3). Only for a shorter-wavelength range from 340 to 500 nm, the 29 mean AAE was estimated to be 1.93±0.59 (Fig. 3). A larger AAE for a shorter-wavelength 30 range was also reported by Chakrabarty et al. (2010) for BrC in tar balls from smoldering 31 biomass combustion. Also shown in Fig. 3 are data of the imaginary part of refractive index (k)32 retrieved from sky radiometer observations, indicating a strong wavelength-dependence. The





1 wavelength-dependence was quantified similarly to Eq. (3) as the slope (w) of the linear fit of 2  $\ln(k)$  versus  $\ln(\lambda)$ . The k values in the UV region were as high as 0.01-0.03 but one order of 3 magnitude smaller than that of BC (~0.71) (Bond and Bergstrom, 2005). Using the 4 parameterization derived by Saleh et al. (2014) and the k value at 550 nm ( $k_{550}$ ) derived by 5 interpolation in the present study (~0.012), the BC-to-OA ratio of the emissions from BB 6 (RBC/OA) around Phimai was estimated to be 1.9%. A RBC/OA ratio less than 1.9% is suggested 7 for smoldering combustion of duffs investigated by Chakrabarty et al. (2010) as their  $k_{550}$  value 8 is smaller than that estimated for Phimai in the present study.

9 Since HCHO is a good tracer for absorption aerosols from BB as mentioned above, it is 10 interesting to investigate the dependence of AAE on the HCHO concentration. We found, 11 however, that their correlations were weak and the AAE at a HCHO of 3 ppbv (~1.7) tended to 12 be higher than the AAE values at higher HCHO concentrations (~1.5) (Fig. 4). Although 13 uncertainty in the estimate for the single data of daily mean AAE was as large as 0.3-0.5, it can 14 be seen that AAE tended to converge to 1.5±0.2 at a higher HCHO concentration in BB plumes.

15 According to smog chamber experiments by Saleh et al. (2014), aerosol absorptivity depends 16 largely on burn conditions, not fuel type. In addition, the size distribution and the mixing state 17 of BC and OA can be important factors for AAE (e.g., Russel et al., 2010; Kirchstetter et al., 18 2004). It was also reported that non-absorbing shells over BC cores can lead to AAE greater or 19 less than unity (Gyawali et al., 2009). Despite such a complexity in interpretation of the 20 variation in AAE and the uncertainty in sky-radiometer-retrieved AAE as large as 0.3-0.5, we 21 attempted to interpret possible enhancement in AAE at a HCHO of 3 ppbv. For this, using the 22 parameterization of Salah et al. (2014) and the k550 values, the RBC/OC ratio for a low HCHO 23 case at 3 ppbv was calculated to be 0.013, which was smaller than the  $R_{BC/OC}$  ratios at higher 24 HCHO concentrations (e.g., 0.023 at a HCHO of 9 ppbv) (Fig. 4). A smaller R<sub>BC/OC</sub> ratio can 25 be attributed to the lower-temperature BB. In this case, the lower-temperature BB could yield 26 only small values of AOD, AAOD, HCHO, and CHOCHO (i.e., the magnitude of BB emissions 27 was weak) but a high value of AAE about 1.7 (i.e., as a results of a smaller RBC/OA ratio for 28 emissions) (Fig. 4). The other interpretation for the enhancement in AAE at a HCHO of 3 ppbv 29 is that we observed more photochemically-aged BB plumes at smaller HCHO concentrations. 30 As the photochemical aging occurred, more SOA should have been produced, leading to 31 stronger wavelength-dependence of absorption. Considering a large uncertainty in AAE data 32 used here, further investigation using more data from our multi-component observations by sky





radiometer and MAX-DOAS is encouraged to better interpret the characteristics of BB plumes
 observed in this study. In addition, the results presented here are expected to be unique
 constraints for understanding physical/chemical/optical properties of BB plumes.

4

# 5 4 Conclusions

6 We conducted ground-based remote sensing observations using the sky radiometer and the 7 MAX-DOAS at SKYNET/Phimai site in central Thailand from January to April 2016 to 8 characterize optical properties of aerosols and organic gas concentrations in BB plumes. We 9 found that the HCHO concentration for a layer of 0-1 km retrieved from MAX-DOAS was a 10 useful tracer for BB plumes. As the HCHO concentration tripled from 3 to 9 ppbv, the  $R_{\rm GF}$ 11 decreased from ~0.04 to ~0.03, in respond presumably to the increased contribution of VOC 12 emissions from BB. In addition, AAODs increased with HCHO. At a HCHO of 9 ppbv, AAOD 13 at 340 nm reached as high as  $\sim 0.15 \pm 0.03$ . The AAE at 340-870 nm was about  $1.5 \pm 0.2$ , 14 indicating the presence of BrC aerosols. The results from our multi-component observations 15 around central Thailand are expected to be unique constraints for understanding 16 physical/chemical/optical properties of BB plumes.

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Fig. 1. Time series of multi-components retrieved from sky radiometer and MAX-DOAS
observations at Phimai, Thailand for the intense BB period from January to April 2016. Daily
means for 9:00-15:00 LT are plotted. Their 1σ standard deviations are shown by error bars.
AOD, SSA, and AAOD values for different wavelengths are shown in different colors. For RHw,
red symbols indicate MAX-DOAS-derived RHw for a layer of 0-1 km and black symbols
indicate the surface RHw from NCEP data.







1HCHO (ppbv)HCHO (ppbv)2Fig. 2. (a) CHOCHO concentration, (b) RGF, (c) NO2 concentration, and (d) AAOD values3plotted as a function of HCHO concentration. AAOD values at 340, 400, and 870 nm are shown4in blue, green, and black, respectively. The medians of respective quantities for each 2-ppbv5bin of HCHO concentration are shown by open symbols. Error bars represent 67%-ranges.







Fig. 3. Spectra of (a) AAOD and (b) imaginary part of refractive index for the period from
January to April 2016. The power law fitting results for 340-870 nm and 340-500 nm are
shown by solid and dashed lines, respectively. Error bars represent 1σ standard deviations for
each wavelength.

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Fig. 4. AAE values plotted as a function of HCHO concentration.