



1 Evaluation of microwave radiances of GPM/GMI for the all-sky 2 assimilation in RTTOV framework

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8 **Abstract.** This study evaluates the all-sky GPM/GMI radiances towards assimilation in regional
9 mesoscale model at 183 ± 7 GHz. The radiative transfer model (RTM) namely RTTOV-SCATT is
10 used for the simulation of three tropical cyclones (hudhud, vardah and kyant respectively). Within
11 the RTM, the performance of non-spherical Discrete Dipole Approximation (DDA) shapes (sector
12 snowflake, 6-bullet rosette, block-column and thinplate) are evaluated. The input data used in
13 RTTOV-SCATT includes vertical hydrometeor profiles, humidity and surface fluxes. In addition,
14 the first guess simulations from Weather Research Forecast (WRF) model were executed at 15 km
15 resolution using ERA-Interim reanalysis datasets. Results indicate that observed minus first guess
16 (FG departures) are symmetric with DDA shapes. The normalized probability density function of
17 FG departures shows large number of spatially correlated samples between clear-sky and poorly
18 forecasted region. Quality control (QC) method was performed to eliminate large FG departures
19 due to instrumental anomalies or poor forecast of clouds and precipitation. The goodness of fit
20 test, h-statistics and skewness of observed and simulated distribution show optimum results for
21 thinplate shape in all the convective events. We also tested the high resolution ERA-5 reanalysis
22 datasets for the simulation of all-sky radiances using thinplate shape. Results illustrate a potential
23 to integrate the GMI sensor data within a WRF data assimilation system.

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29 **1 Introduction**

30 The Numerical weather prediction (NWP) model is widely used for forecasting the evolution of
31 the surface and atmospheric conditions. To predict the state of the atmosphere, the NWP model
32 relies on mathematical models and best initial conditions of the state of the atmosphere. Even
33 though NWP models provide meaningful forecasts, they are biased owing to model structure and
34 approximation of subgrid-scale processes (Shastri et al., 2017). In addition, it is also challenging
35 to define the best initial conditions of the atmosphere state. Recent developments by operational
36 NWP centers state that assimilating satellite radiances improves the forecast skills (Islam et al.,
37 2016; Routray et al., 2016; Saunders et al., 2013; Singh et al., 2016). Studies show that assimilation
38 of all-sky (clear and cloudy) microwave radiances in global NWP models have a large positive
39 impact on temperature and humidity (Geer, 2013; Kazumori et al., 2014; Lean et al., 2017). The
40 satellite radiances from microwave imagers [Tropical Rainfall Measuring Mission (TRMM)
41 microwave imager (TMI) (Kummerow et al., 1998), Aqua Advanced microwave scanning
42 radiometer for earth observing system (Kawanishi et al., 2003), Advanced Microwave Scanning
43 Radiometer-2 (AMSR-2) (JAXA, 2013), Special Sensor Microwave Imager Sounder (SSMIS)
44 (Kunkee et al., 2008) and Global Precipitation Measurement (GPM) Microwave Imager (GMI)
45 (Hou et al., 2014)] contain crucial information on deep and intense convection (Otkin, 2012).

46 Study of rainfall/convective systems involve examining the naturally emitted electromagnetic
47 radiation from the earth which interacts with atmospheric gases like water vapour, hydrometeors
48 (precipitation-sized particles of rainfall, snowfall, ice crystals etc). A radiative transfer model
49 (RTM) uses profiles of all the observed variables to provide the satellite observations which can
50 be either brightness temperature (T_b) (from radiometer) or reflectivity (from radar). An accurate
51 comparison between model and observed forecast relies strongly on the assumptions used for
52 radiation-hydrometeor interaction. To generate improved initial conditions of the model, a crucial
53 role is played by the assumptions made when simulating observations from different instruments
54 in space. To date, the scientific community has not really examined this important aspect.

55 Existing studies state that within the microwave frequency ranges (10-183 GHz), assuming
56 spherical shapes for snow/ice particles in RTM models produce un-realistic scattering in deep
57 convective clouds (Hong et al., 2005). Geer and Baordo, (2014) introduced the realistic 3D discrete
58 dipole approximation (DDA) non-spherical shapes to represent frozen hydrometeors within an



59 RTM model. The DDA shapes examined are usually long column, short column, block column,
60 thick plate, thin plate, 3,4,5 and 6 bullet rosette, sector and dendrite snowflake shapes (Liu, 2008).
61 Geer and Baordo, (2014) claimed that DDA sector snowflake is approximately fit for all
62 frequencies at a global scale but it doesn't perform well at the regional level. Regional case studies
63 show that block column over Indian ocean (Guerbette et al., 2016) and thinplate over
64 Mediterranean region (Rysman et al., 2016) have improved the simulation of low Tb at convective
65 scale for 183 GHz. Hence, careful investigation of DDA shapes is required at the regional level,
66 prior the simulation of cloudy radiances at a higher frequency. Generally, the best choice of DDA
67 shapes at each frequency is based on the statistical analysis of the FG departures which is important
68 in variational data assimilation techniques (Fowler and Van Leeuwen, 2013). Assimilation of
69 satellite radiances offer difficulty due to cloud processes that non-linearly affect the upwelling
70 radiations, non-gaussian FG departures statistics and systematic biases from NWP and radiative
71 transfer models (RTM) (Errico et al., 2007; Okamoto, 2017).

72 The issue of non-gaussian characteristics of FG departure could be resolved using cloud dependent
73 standard deviation (SD_{cloud}) (Geer and Bauer, 2011; Okamoto, 2017; Okamoto et al., 2014). The
74 study conducted by Geer and Bauer, (2011) proposed a symmetric error model between SD_{cloud}
75 and cloud amount predictor for all-sky microwave observations. The authors used the error model
76 for AMSR-E observations at 19 GHz and observed that the probability distribution function (pdf)
77 of normalized FG departures follow a gaussian distribution if normalization is done by SD_{cloud} .
78 Geer, (2013) used the same model at multiple frequencies of TMI and SSMIS channel. In
79 microwave spectrum, the symmetric error model is known to perform well for low frequencies
80 (<50 GHz) as low frequencies are sensitive to cloud liquid droplets and rain-drops (Skofronick-
81 Jackson and Wang, 2000) for which the particle shape and density are pre-defined. At higher
82 microwave frequencies, the backscatter/brightness temperature registered by the sensor is mainly
83 due to scattering from frozen hydrometeors, assuming a spherical shape (Geer and Baordo, 2014).
84 Long-term monitoring of FG departure was found useful for identifying the instrumental error
85 from ground based microwave observation (De Angelis et al., 2017).

86 This study investigates the simulation of all-sky GMI radiances of tropical cyclones over Indian
87 region at 183 ± 7 GHz. In the analysis, we examined the cloud effect to evaluate the normalized
88 FG departures. For the appropriate selection of DDA shapes, we inspect the statistical measure of



89 FG departures. In addition, we also include the analysis of ERA-5 reanalysis datasets (Malardel et
90 al., 2015) to extend the sensitivity to cloud physical processes at a higher resolution. Section 2
91 briefly summarizes the GMI radiance datasets, NWP and RTM experimental setup. The simulation
92 results and error analysis are demonstrated in section 3. Summary and conclusions are provided in
93 section 4.

94

95 **2 Data and Methods**

96 **2.1 GPM/GMI observations**

97 The GMI sensor is a conically scanning passive radiometer on board the GPM satellite (Hou et al.,
98 2014) developed by National Aeronautics and Space Administration (NASA) in collaboration with
99 Japan Aerospace Exploration Agency (JAXA) and successfully launched on 28th February 2014.
100 Being a successor of TRMM (Kummerow et al., 1998) the GPM mission has several advantages.
101 GMI data is acquired in 13 channels in low (10-89 GHz) and high frequency (166-183 GHz) bands
102 [Table 1] while TMI was limited to just 9 low-frequency bands. GMI has additional capabilities
103 of detecting light precipitation and extending the global coverage to the mid-latitude region
104 ($60^{\circ}S - 60^{\circ}N$). The horizontal resolution has been improved in GMI datasets because of the
105 increase in reflector size of GMI (1.2 m) from TMI (61 cm). Figure 1 shows hudhud cyclone event
106 on 9th October 2014, 06 UTC at 10, 89 and 183 ± 7 GHz frequency. The low frequency channel (10
107 GHz) is sensitive to only liquid precipitation and greatly affected by surface emissivity (Hou et
108 al., 2014). The discrimination between land and ocean (Figure 1) is clear in 10 GHz, moderate in
109 89 GHz and insensitive at 183 GHz. Furthermore, 183 ± 7 GHz channel can investigate deeply
110 the atmosphere and it is highly sensitive to frozen hydrometeors. This channel is moderately
111 sensitive to rain and cloud liquid water and also detect the scattering signals from small ice
112 particles (Bennartz and Bauer, 2003; Laviola and Levizzani, 2011). In Figure 1, a strong
113 depression of temperature lower than 100 K can be seen over ocean at 183 ± 7 GHz indicating
114 the presence of frozen hydrometeors in deep convection. For the present study, GMI level 1b
115 radiances for three tropical cyclones at 183 ± 7 GHz-V (hereafter band 13) is used.

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118 2.2 RTTOV-SCATT v12.1 Model

119 The all-sky GMI radiances have been simulated using RTTOV-SCATT (version 12.1) of Radiative
120 Transfer for the television infrared observation vertical sounder (RTTOV) (Hocking et al., 2017;
121 Saunders et al., 2017). The RTTOV is initially developed by ECMWF which was then upgraded
122 within the European Organization for the Exploitation of Meteorological Satellites (EUMETSAT)
123 NWP satellite application facility. This Model is suitable for rapid transformation of a huge
124 number of NWP model outputs into the radiance space. The RTTOV-SCATT is a separate
125 interface for the simulation of cloud and precipitation affected microwave radiances. As an input
126 to RTTOV-SCATT model, the atmospheric profiles (i.e. temperature, water vapour, cloud liquid
127 water, ice, snow and rain) were derived from WRF NWP model output. The delta-eddington
128 approximation is used for solving the radiative transfer equations to simulate the scattering effects
129 of clouds and precipitations (Joseph et al., 1976). The surface emissivity over oceans are calculated
130 by the surface parameters (eg. Temperature, surface wind) using the microwave surface emissivity
131 model (FASTEM-version 6) (Kazumori and English, 2015). The all-sky T_b computed represent
132 the weighted summation of the clear and cloudy independent columns (eq. 1). The weighing
133 criteria is decided by the cloud fraction (Geer et al., 2009) which is based upon the variation in
134 cloud and precipitation at subgrid scale.

$$135 \quad T_b^{all-sky} = C_f * T_b^{cloudy} + (1 - C_f) * T_b^{clear-sky} \quad (1)$$

136 Here, C_f represents the vertical profile of cloud fraction.

137

138 2.3 WRF NWP Model

139 The WRF is specifically designed for regional forecast in operational and research NWP centers
140 (Skamarock et al., 2008). The present study used the version 3.8 of WRF model for the forecasting
141 of tropical cyclones over Indian region. We designed the experimental setup in a single domain
142 from $3^{\circ}N$ to $26^{\circ}N$ and from $73^{\circ}E$ to $103^{\circ}E$ with 213×165 horizontal grids of 15 km resolution
143 [Figure 2 (a)]. This experiment is configured with 51 number of vertical layers and model top is at
144 125 hPa. The initial and boundary conditions are taken from ERA-Interim reanalysis datasets
145 (product of ECMWF) with specification of 71 km spatial resolution at 6 h interval (Simmons et
146 al., 2007). Geographical parameters including land use land cover (LULC), topography, soil type,



147 lake and vegetation parameters are provided by the United States Geological Survey (USGS)
148 global datasets at 30 sec resolution. Three tropical cyclones named “Hudhud” (October7-14,
149 2014), “Vardah” (December 6-12, 2016) and “Kyant” (October 21-27, 2016) over Bay of Bengal
150 (BOB) regions are considered in this study. Their tracks are shown in Figure 2 (b).

151 The physical parameterization schemes used are as suggested by (Routray et al., 2016) over BOB
152 region are; WRF single moment 6-class microphysics scheme (Hong and Lim, 2006), Kain-Fritsch
153 convection scheme (Kain, 2004), Yonsei scheme for planetary boundary layer (Hong et al., 2006),
154 Dudhia shortwave radiation scheme (Dudhia, 1989), rapid radiative transfer model scheme for
155 long-wave radiation (Mlawer et al., 1997) and Noah land surface model scheme (Tewari et al.,
156 2004). This configuration is highly versatile for the prediction of short range forecast over the
157 Indian region (Kumar et al., 2014; Singh et al., 2016).

158 **3 Results and Discussion**

159 In the present study, DDA shapes of sector snowflake is used as first step for initial error analysis
160 in section 3.1-3.4 (Geer and Baordo, 2014). In section 3.5, a statistical investigation is conducted
161 to identify the best shape among the recognized DDA shapes (i.e., Sector snowflake, thinplate, 6-
162 bullet rosette and block-column). In this study, the density of hydrometers (rain and cloud liquid
163 water=1000 kg/m³; ice=917 kg/m³; snow= 50 kg/m³) and particle size distributions by Field et al.,
164 (2007) for snow, marshall-palmer distribution for rain, modified-gamma distribution for cloud
165 liquid water and cloud ice have been used.

166 **3.1 Spatial Distribution of observed and simulated Tb**

167 The Figure 3 shows comparison between the all-sky simulated radiances at band 13 with respect
168 to the observed GMI radiances for three cyclonic events over the BOB region. The microwave
169 observations were averaged to 15 km horizontal resolution to match closely with the effective
170 resolution of NWP model. The increased scattering from frozen hydrometeors at band 13 in deep
171 convective zones results in low temperatures of observed radiances inside the core of cyclone (upto
172 70-80 K). Underestimation was observed using the mie-sphere, sector snowflake and six-bullet
173 rosette shapes. Though the overall pattern and location of convective clouds near the eye of cyclone
174 matched closely with the observations, Tb inside the core can be found to vary with hydrometeor
175 shapes and estimates. This may be attributed to deficiency of frozen hydrometeors at sub-grid scale



176 in Kain-Fritsch convection scheme (Rysman et al., 2016). A study by Wu et al., (2015) found that
177 frozen hydrometeors are underestimated in WRF simulations by all convective parametrization
178 schemes over central and eastern pacific region. Rysman et al., (2016) estimate the
179 underestimation in WRF simulations by a factor of 5 using airborne radar in Mediterranean region.

180 The Figure 4 shows distribution of FG departures in mie-spheres and DDA shapes for all the case
181 study events. A negative departure occurs when the RTToV model is unable to produce realistic
182 representations owing to cloud and precipitation. Within DDA shapes, the pdf curve is found to
183 follow a symmetric distribution. The error is spread equally in both directions due to random
184 forecast errors from first-guess and observations. In case of mie-spheres, the shift towards large
185 negative departures indicates the presence of bias in cloudy region. This is because of insufficient
186 scattering by mie-spheres at band 13 (Geer, 2013). Results show that DDA simulations provide a
187 better realistic scattering in all-sky conditions when FG departures are symmetrical in nature.

188 **3.2 Determination of observation errors with cloud amount**

189 The standard deviations of FG departures in clear-sky assimilation are referred to as observation
190 errors. The observation errors in all-sky radiance assimilation for microwave observations are
191 generally computed from symmetric error models (Geer and Bauer, 2011). Error models are a
192 function of cloud amount predictor at 37 GHz. In the present study, the observed/simulated cloud
193 amounts have been computed from observed/simulated radiances in clear and all-sky conditions.

$$194 \quad C_{37} = 1 - PD_{37}; \quad (2)$$

$$195 \quad PD_{37} = \frac{Tb_v - Tb_h}{Tb_v^{clr} - Tb_h^{clr}} \approx \tau_{cloudy}^2 \quad (3)$$

196 Where, Tb_v & Tb_h are the vertically and horizontal polarised radiances in cloudy condition; Tb_v^{clr}
197 and Tb_h^{clr} are the vertically and horizontal polarized radiances in clear sky condition. PD_{37} is the
198 normalized polarization differences approximately equal to square of transmittance in cloudy
199 region (Petty, 1994). A clear and cloud sky is represented using a PD_{37} of 1 and 0 respectively.
200 For easy interpretation, we preferred cloud amount (C_{37}) which varies from 0 to 1 for the same
201 representation. As the quantities of C_{37obs} and C_{37sim} are affected with sampling error (Geer and
202 Bauer, 2011), their average is considered as the average cloud amount C_{37avg} .



203 Figure 5 shows the SD_{cloud} curve at band 13 using the C_{37avg} on x-axis in a bin range of 0.05.
 204 At $C_{37avg} \sim 0$, both observations and first-guess are free from clouds (i.e. clear-sky condition). As
 205 the C_{37avg} increases, the error is found to initially increase linearly and attain the maxima at
 206 $C_{37avg} \sim 0.48$ in all the meteorological events after which the error starts declining to the
 207 maximum cloud amount ($C_{37avg} = 1$). The sudden peak at $C_{37avg} \sim 0.8$ observed in hudhud and
 208 vardah cyclones was due to poor representation of Tb at higher frequency using DDA sector
 209 snowflake shape in heavy clouds that causes large error.

210 In symmetric error model, the SD_{cloud} curve was piecewise linearly transformed as a function of
 211 C_{37avg} (Geer and Bauer, 2011) (eq. 4).

$$212 \quad SD_{cloud}(C_{37avg}) = \begin{cases} S_{clr} & \text{if } C_{37avg} \leq C_{clr} \\ S_{clr} + \left(\frac{S_{cld} - S_{clr}}{C_{cld} - C_{clr}} \right) (C_{37avg} - C_{clr}) & \text{if } C_{clr} < C_{37avg} < C_{cld} \\ S_{cld} & \text{if } C_{37avg} \geq C_{cld} \end{cases} \quad (4)$$

213 Here, S_{clr} is the minimum SD_{cloud} defined by the threshold C_{clr} in clear-sky region, whereas S_{cld}
 214 is maximum SD_{cloud} in strongly dominating clouds and precipitation region as defined by C_{cld}
 215 threshold. These parameters for each cyclone event at band 13 were summarized in Table 2.

216 3.3 Evaluation of normalized FG departures

217 The bandwidth of FG departures are very high (Figure 4) and finding a symmetric bias in absolute
 218 FG departure is not feasible. Hence, FG departures are normalized with SD_{cloud} (eq. 4) at band 13.
 219 The pdf of normalized FG departures were compared with Gaussian for all the deep convective
 220 events (Figure 6). From Figure 6, it can be seen that, the normalized FG departure curves follow
 221 symmetric distribution but its peak was too high with smaller errors. The main advantage of
 222 symmetric error model is to assign large errors in cloudy conditions without causing difficulty in
 223 all-sky assimilation.

224 3.4 Quality Control (QC)

225 Figure 7 (a), (b) and (c) shows the distribution between observed and simulated Tb using binned
 226 scatter plots in 1.0 K by 1.0 K bin for hudhud, vardah and kyant cyclone respectively. Samples
 227 found to be outside the range of 100-300 K and bins containing less than or equal to 1 sample were
 228 removed from the analysis. The simulated warmer Tb (>240 K) was in good agreement with the



229 observations but samples containing low values of T_b (<240 K) were either from first-guess or
230 from observations having large FG departures. Because of partially random distribution of deep
231 convective clouds, there is a large uncertainty in the prediction of exact location of convective
232 clouds in the model causes the large disagreement (Harnisch et al., 2016). Geer and Bauer, (2011)
233 proposed quality control (QC) method in operational all-sky microwave radiance assimilation to
234 eliminate the large FG departures due to cloud mis-location and instrumental errors, however, their
235 study has not considered the observations wherein normalized FG departures are greater than \pm
236 2.5 K.

237 For the present study at band 13, threshold limits cannot be decided using the normalized FG
238 departures. Hence, we performed QC by removing 2.5% samples from both sides of the tail of
239 normalized FG departures. Samples after QC are shown in Figure 7 (d), (e) and (f) and dashed line
240 represent the window of FG departure at 0, ± 10 and ± 30 K. The low T_b samples removed after
241 QC reduces the variability of FG departure and hence improves the symmetry. Mostly cloudy
242 samples were lie in the error range of ± 30 K. Results has shown the improvement in correlation
243 coefficient after QC. This method also eliminates the negative departures linked with deep
244 convective events. Figure 8 shows the convective clouds on 10th December 2016 at 03 UTC
245 wherein plots (a), (b) and (c) represent the observed and simulated T_b before and after QC
246 respectively. The cloud information remains preserved after the QC.

247 **3.5 Measure of goodness of fit**

248 Accurate simulation of deep convective events at 183 GHz are challenging due to difficulty in
249 modelling of scattering effects from frozen hydrometeors (Geer and Baordo, 2014; Guerbette et
250 al., 2016). This section measures the goodness of fit between observed and simulated radiances
251 using four widely used DDA shapes. It is common practice to use chi-squared or K-S test to
252 statistically measure the discrepancy between two distributions. Geer and Baordo,(2014) proposed
253 an ‘h-statistics’ (eq. 5) for smaller samples arranged into number of bins. The value of h could
254 reach infinite if no samples be present in the bin. This study assigns such bins to 0.1 value.

$$255 \quad h = \frac{(\sum_{bins} |\log_{\# observation} \frac{\# simulated}{\# observation}|)}{total \ no \ of \ bins} \quad (5)$$

256 Here, the bin size is 2.5 K and # denote the numbers.



257 The Figure 9 (a), (b) and (c) shows the $\log \frac{\# \text{ simulated}}{\# \text{ observation}}$ in y-axis for the case study events that lie
258 within the range of 100-280 K Tb in x-axis for all DDA shapes. In the 100-200 K bins, block
259 column and thin-plate have a positive log ratio which means low Tb occurs extensively due to
260 excessive scattering from clouds while six-bullet rosette and sector snowflake with negative log
261 ratio shows very less or no existence of Tb in this range because of insufficient scattering in heavy
262 cloud and precipitation regions. However, in 200-280 K bins, the log ratio closely lies near to zero
263 and overall good agreement is observed between the observed and simulated Tb. The analysis of
264 DDA shapes shows that thinplate has less peak from 0 to either positive or negative side in all the
265 cases. The h-statistics value for each shape is given in Table 3. Less the number, more will be the
266 similarity. Thin-plate have lowest h-value.

267 Figure 9 (d) shows the skewness of FG departure for all convective events for each DDA shapes.
268 Large negative or positive value indicate skew towards left or right from normal distribution curve.
269 Combining h-statistics and skewness, thin plate show optimum results among all DDA shapes over
270 Bay of Bengal at band 13. This result is in accordance with the study by Rysman et al., (2016)
271 which shows that thin plate perform best in simulation of all-sky radiances of Advanced
272 Microwave Sounding Unit (AMSU)-B sounding channels in Mediterranean region. Guerbette et
273 al., (2016) observed the best simulations of all-sky SAPHIR radiances with block column shape
274 over Indian Ocean at 183 ± 7 GHz.

275 **3.6 Sensitive to ERA-5 reanalysis datasets**

276 Simulated Tb using ERA-I datasets were compared with ERA-5 (31 km; 3 hr) simulation for all
277 cases. Figure 10 shows the spatial distribution of observed and simulated Tb using DDA thinplate
278 shape for convective clouds at (a, b, c) 10th October 2014-18 UTC (hudhud), (d, e, f) 8th December
279 2016-15 UTC (vardah) and (g, h, i) 22nd October 2016-18 UTC (kyant). It was observed that the
280 location of clouds and their intensity with ERA-5 datasets was much similar to observations and
281 clear mis-match in distribution of clouds was shown with ERA-I datasets.

282 The observed and simulated Tb is represented using box-plots and histogram in Figure 11 (a) and
283 (b). The total number of samples are 2445 and correlation coefficient has improved drastically
284 from 0.04 (ERA-I) to 0.52 (ERA-5). The ERA-I simulations have large variability of low Tb due



285 to excess scattering from clouds and also decreases the median value. Overall results state that
286 ERA-5 improves the displacement of cloud location, pattern and intensity.

287 **4 Summary and Conclusions**

288 The present study evaluates the simulation of all-sky microwave radiances of GPM/GMI using
289 Weather Research Forecast (WRF) and RTTOV-SCATT (v12.1) radiative transfer model. GMI
290 observations at water vapour sounding channel (183 ± 7 GHz-V) has been considered and
291 spatially averaged over model resolution. This study has been conducted for three tropical cyclones
292 (Hudhud : 07th -14th Oct. 2014; Vardah: 21st -29th Oct. 2016 and Kyant: 06th -12th Dec. 2016) at 3-
293 h interval and 15 horizontal resolution over Bay of Bengal (BOB) region. In the present study, four
294 recognized DDA shapes (sector snowflake, block column, thin plate and six-bullet rosette) were
295 considered for simulation of brightness temperature (Tb). Results show that simulations using mie-
296 spheres produces bias in cloudy region due to inadequate scattering at 183 ± 7 GHz while all DDA
297 shapes have significant scattering at higher frequency (Geer and Baordo, 2014) .

298 We evaluate the cloud effect on FG departures from DDA sector snowflake simulations using
299 symmetric error model. The probability distribution function of normalized FG departures are
300 found to be symmetric. The results show that cloudy samples can offer potential to be assimilated
301 in all-sky radiance assimilation experiments.

302 The present study also conducted the statistical measures to evaluate the performance of DDA
303 shapes for radiance simulation. The h-statistics is performed to measure the consistency between
304 observed and simulated distributions. We also used skewness, the most suitable parameter in large
305 errors situations (Wilks, 2006). This study observed that thinplate simulates all-sky microwave
306 radiances consistently with observations over BOB region. Our finding resonate with Rysman et
307 al., (2016) over mediterranean region at 183 ± 7 GHz. In our simulations, we consider only DDA
308 shapes for snow, however in reality there are also high density particles such as hail, aggregate
309 and graupel hydrometeors which produces very low brightness temperature (Figure 1). Further
310 efforts should be made to include varieties of frozen hydrometeros in RTTOV model. Another
311 improvement can be done in RTTOV-SCATT model to allow multiple DDA shapes of frozen
312 hydrometerors at a time.



313 In the present study, simulation of all-sky GMI radiances is carried out with ERA-1 and ERA-5
314 reanalysis datasets using thinplate shape. Results show improvement in cloud location and
315 intensity near the core when using ERA-5 compared to ERA-I datasets. This can be attributed to
316 the higher spatial and temporal resolution of ERA-5 datasets which when used in WRF model
317 improved the forecast of cloud and precipitation. The initial results using ERA-5 datasets are
318 encouraging and will be a part of ongoing work on radiance simulations. Future work will be
319 focussed on all-sky GMI radiance assimilation in WRF model at higher frequencies for short range
320 forecast over indian region.

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326 DAAC) for providing the GMI radiances and are free available at <https://mirador.gsfc.nasa.gov/>.
327 The ERA-I and ERA-5 datasets are obtained from <http://rda.ucar.edu/datasets/ds627.0> and
328 <https://rda.ucar.edu/datasets/ds630.0/> respectively.

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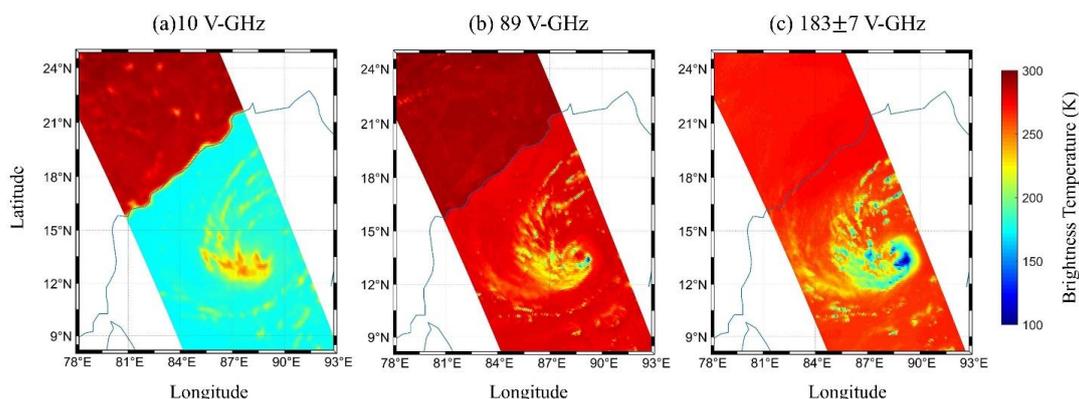
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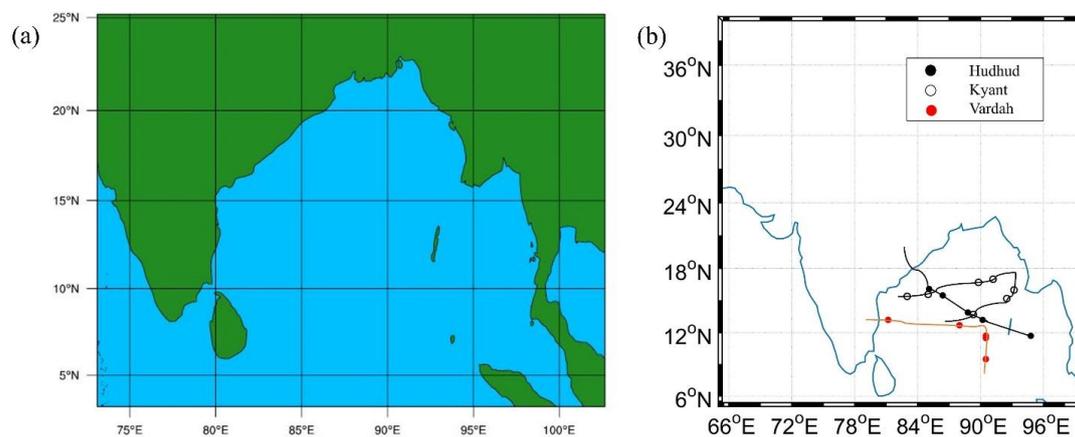
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477 **Figure 1: Brightness Temperature from GPM GMI for (a) 10 V-GHz, (b) 89 V-GHz and (c) 183 ± 7**
 478 **GHz for hudhud cyclone event on 9th October 2018 at 06 UTC. The frozen hydrometeors information**
 479 **are more enhanced at 183 ± 7 GHz frequency.**

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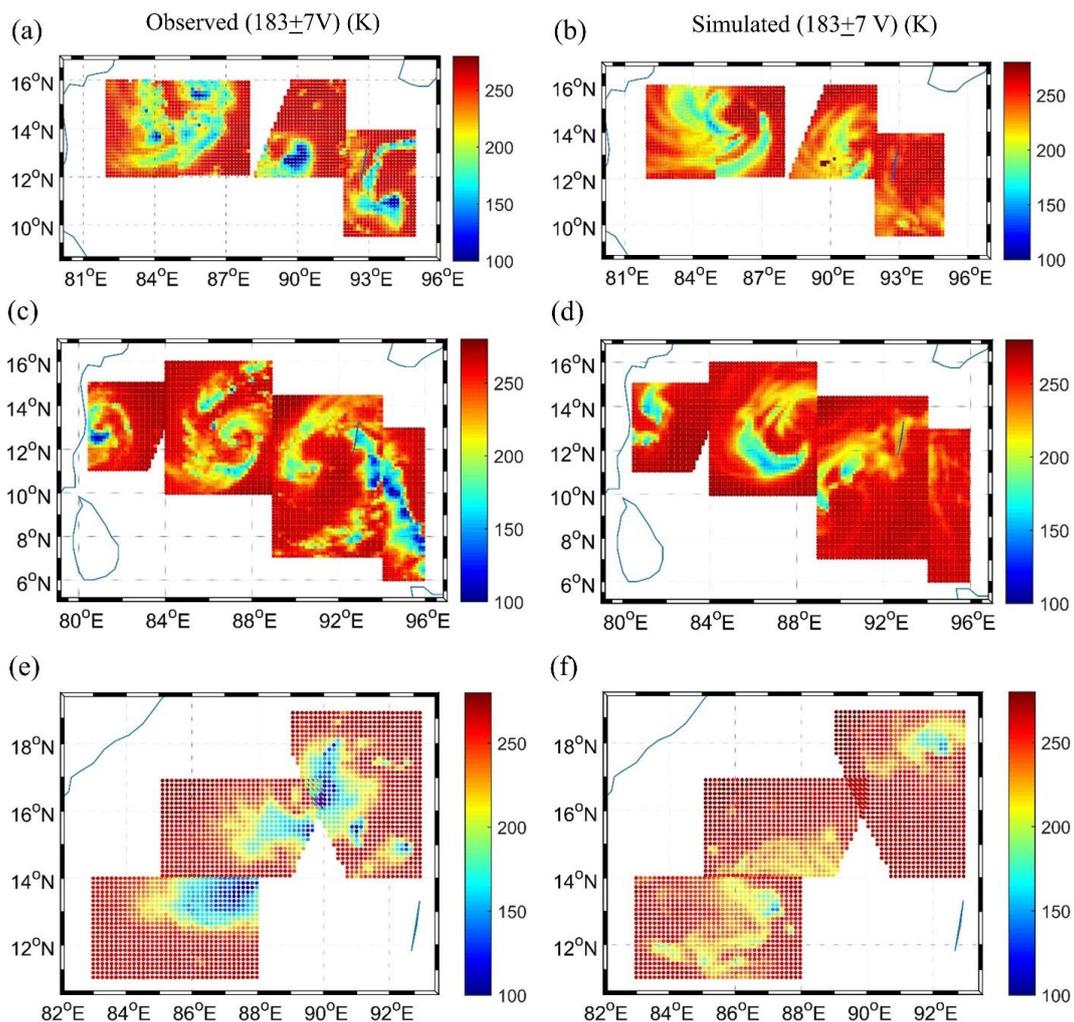
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482 **Figure 2: (a) Single WRF domain used for the simulation of three Tropical cyclones (Hudhud,**
 483 **Vardah and Kyant) over Bay of Bengal and (b) shows the track of cyclones and dot point represent**
 484 **the availability of GMI observations near the eye of cyclone**

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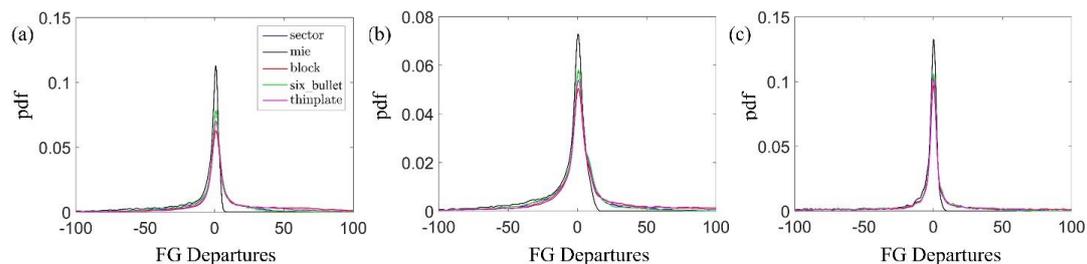
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489 **Figure 3: Spatial distribution of (a, c and e) GMI observed brightness temperature (T_b) and (b, d**
490 **and f) simulated T_b with default DDA sector shape at band 13 for hudhud, vardah and kyant cyclone**
491 **respectively.**

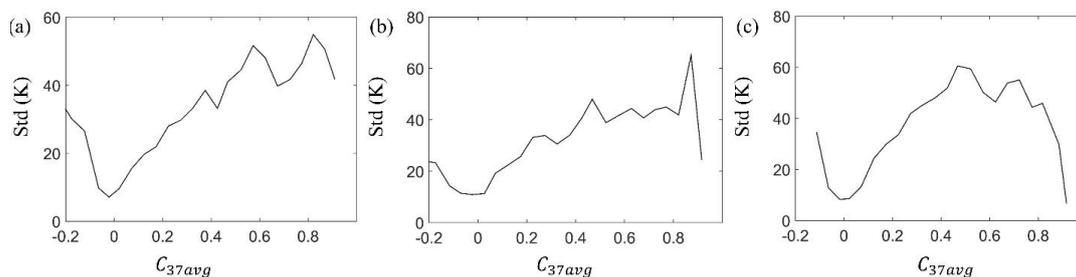
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494 **Figure 4: Probability distribution function (PDF) of observed-background (FG departures) with mie-**
495 **spheres and DDA shapes for (a) hudhud, (b) vardah and (c) kyant cyclone.**

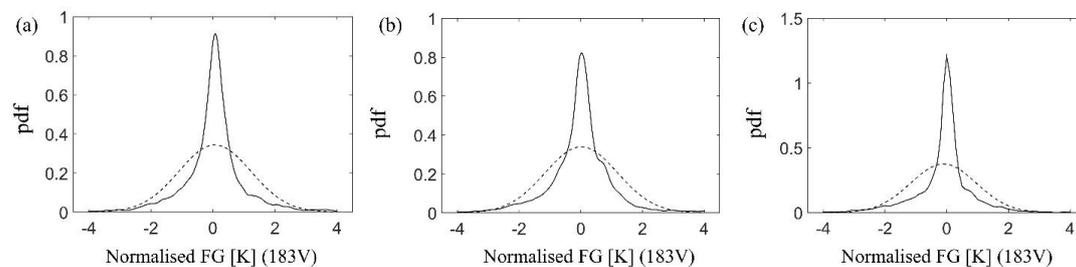
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498 **Figure 5: Standard deviation (SD_{cloud}) curve with respect to average cloud amount at band 13 for**
499 **(a) hudhud, (b) vardah and (c) kyant cyclone. The cloud amount bin is 0.05 at x-axis.**

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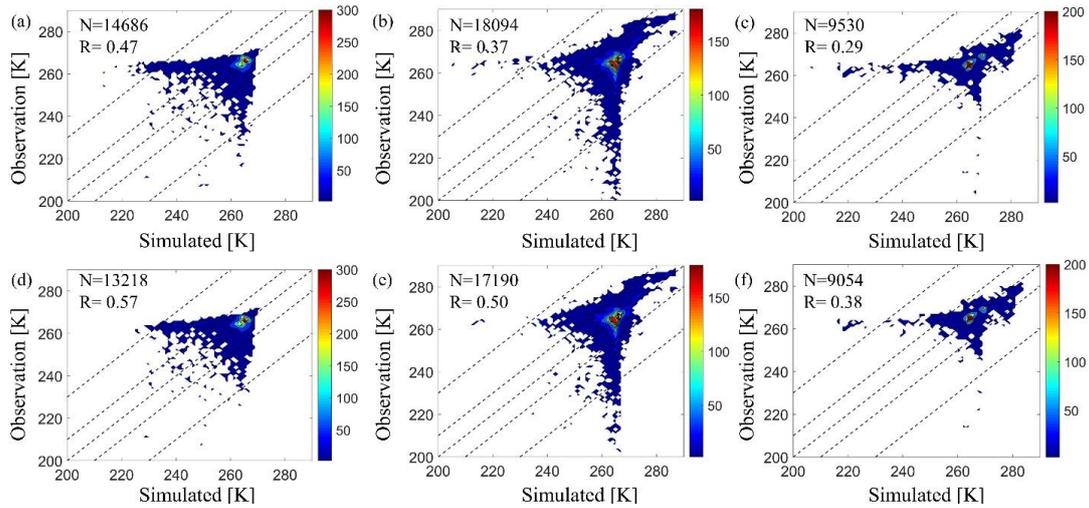


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502 **Figure 6: Probability distribution function (PDF) of FG departure normalized by standard deviation**
503 **as function of average cloud amount for (a) hudhud (b) vardah and (c) kyant cyclone at band 13. The**
504 **dotted curve represent the Gaussian curve.**

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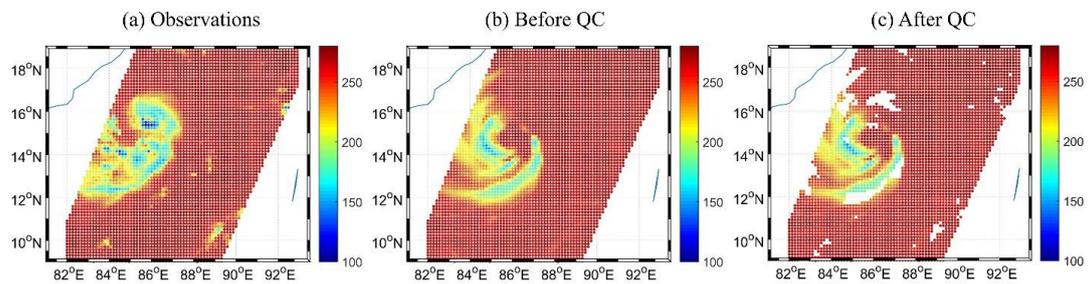


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508 **Figure 7: Binned scatter plots of background and simulated brightness temperature at band 13 for**
 509 **(a, d) hudhud, (b, e) vardah and (c, f) kyant cyclone before and after Quality Control (QC). The**
 510 **samples are counted in 1.0 K by 1.0 K. The colorbar refers to the density in each bin. Dashed line**
 511 **represent the FG departure at 0, ± 10 and ± 30 K.**

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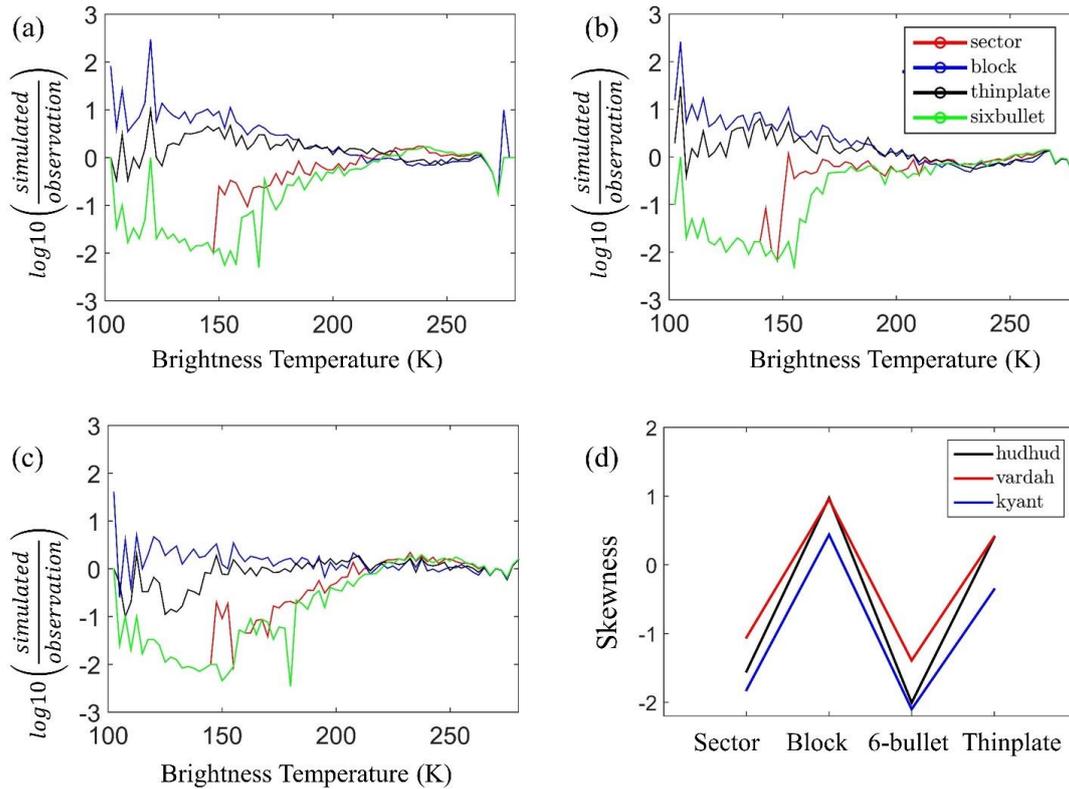
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515 **Figure 8: shows the convective events at 10th October 2014:18 UTC represents (a) GMI observations,**
 516 **(b) Simulated Tb before QC, (c) after QC. The pixels of high FG departures due to mis-match of**
 517 **location were removed.**

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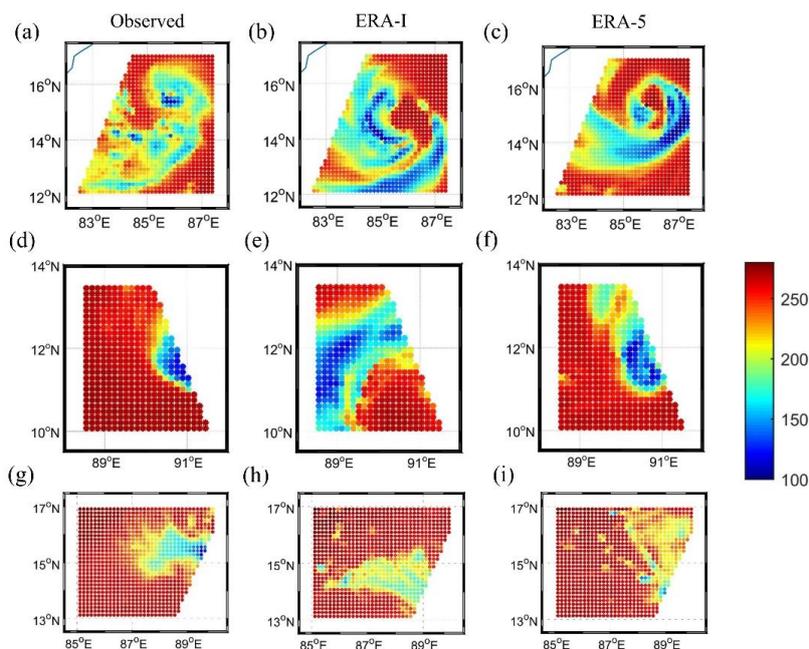
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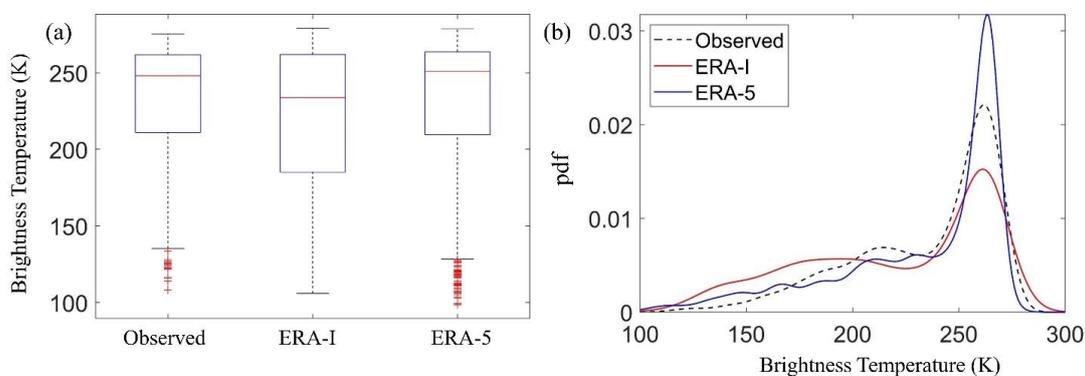
522 **Figure 9: Measure of goodness of fit. The log of the ratio of histograms (simulation divided by the**
 523 **observation) for four different DDA shapes over (a) hudhud, (b) vardah and (c) kyant cyclone. The**
 524 **bin size is 2.5 K. (d) represents the skewness of FG departures. Thinplate performs the best results**
 525 **over Bay of Bengal.**



526

527 **Figure 10: Observed and Simulated Tb with ERA-I and ERA-5 reanalysis datasets for a day event**
 528 **(a, b, c) 10th October 2014, 18 UTC (hudhud cyclone); (d, e, f) 8th December 2016, 15 UTC (vardah**
 529 **cyclone) and (g, h, i) 22nd October 2016, 18 UTC (kyant cyclone).**

530



531

532 **Figure 11: (a) Boxplot of observed and Simulated Tb with ERA-I and ERA-5 reanalysis datasets. The**
 533 **50 percentile of ERA-I simulations is larger than observed data due to excess scattering from the**
 534 **clouds. (b) Histogram of observed and simulated brightness temperature. ERA-5 simulations have**
 535 **similarity with observed data.**

536 **Table 1. GMI sensor characteristics (Hou et al., 2014)**

Channels	Frequency (GHz)	Polarization	Resolution (Km)
1, 2	10.65	V, H	19.4 x 32.2
3,4	18.7	V, H	11.2 x 18.3
5	23.8	V	9.2 x 15.0
6,7	36.5	V, H	8.6 x 15.0
8,9	89.0	V, H	4.4 x 7.3
10,11	166	V, H	4.4 x 7.3
12	183 \pm 3	V	4.4 x 7.3
13	183 \pm 7	V	4.4 x 7.3

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538

539 **Table 2. Standard deviation and threshold for identifying clear-sky and cloudy samples at 183 \pm 7**
540 **V GHz for all tropical cyclones.**

	Hudhud	Vardah	Kyant
S_{ctr}	9.5977	11.2895	8.9561
S_{cld}	51.6524	47.9787	61.7815
C_{cld}	0.5733	0.4677	0.5218
C_{ctr}	0.0209	0.0266	0.0213

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548 **Table 3. shows the h-value and skewness corresponding to the all meteorological events for different**
 549 **DDA shapes. Thinplate has the least h-value and low skewness in all events.**

DDA shapes	Hudhud		Vardah		Kyant	
	h-value	skewness	h-value	skewness	h-value	skewness
Sector	0.6297	-1.5516	0.5477	-1.0571	0.7565	-1.8218
Block	0.4893	0.9733	0.4064	0.9533	0.2522	0.4399
6- bullet	0.8171	-2.0069	0.7028	-1.3925	0.8838	-2.1021
Thinplate	0.2743	0.3982	0.2528	0.4089	0.2397	-0.3575

550