

1 **Identification of Gas-phase Pyrolysis Products in a Prescribed Fire:**  
2 **First Detections Using Infrared Spectroscopy for Naphthalene,**  
3 **Methyl Nitrite, Allene, Acrolein and Acetaldehyde\*\***

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14  
15 **ABSTRACT**

16 Volatile organic compounds (VOCs) are emitted from many sources, including wildland fire.  
17 VOCs have received heightened emphasis due to such gases' influential role in the atmosphere, as  
18 well as possible health effects. We have used extractive infrared (IR) spectroscopy on recent  
19 prescribed burns in longleaf pine stands and herein report **the first** detection of five compounds  
20 using this technique. The newly reported IR detections include naphthalene, methyl nitrite, allene,  
21 acrolein and acetaldehyde. We discuss the approaches used for detection, particularly the software  
22 methods needed to fit the analyte and multiple (interfering) spectral components within the  
23 selected spectral micro-window(s). We also discuss the method's detection limits **and related**  
24 **parameters such as spectral resolution.**

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## 26 1. INTRODUCTION

27 Wildland fire releases significant quantities of trace gases into the environment (Akagi et al., 2011;  
28 Andreae et al., 2001; Crutzen et al., 1979; Yokelson et al., 2013; Andreae, 1991), and such gases  
29 can **profoundly** influence atmospheric chemistry (Crutzen et al., 1990). In some parts of the world,  
30 wildfires are becoming more prevalent as well as increasing in impact (Turetsky et al., 2011; Miller  
31 et al., 2009). **Additionally**, prescribed burning is used as a preventive tool to reduce hazardous  
32 fuel buildups in an effort to reduce or eliminate the risk of such wildfires (Fernandes et al., 2003).  
33 Understanding the products associated with the burning of biomass has received considerable  
34 attention since the emissions can markedly impact the atmosphere. Fourier transform infrared  
35 (FTIR) spectroscopy is one technique that has been extensively used to identify and quantify gases  
36 emitted from burns, generally used in either an open path configuration (Burling et al., 2010; Akagi  
37 et al., 2014; Selimovic et al., 2018; Stockwell et al., 2014) or as an extractive method (Burling et  
38 al., 2011; Akagi et al., 2013; Akagi et al., 2014). Extractive systems typically use a long-path gas  
39 cell coupled to an FTIR instrument so as to increase the sensitivity. Such approaches have been  
40 quite successful; an increasing number of species continue to be identified and quantified due to  
41 the availability of reference gas-phase spectral libraries such as the PNNL **spectral** library (Sharpe  
42 et al., 2004) or the HITRAN database (Gordon et al., 2017). Such libraries contain absorption  
43 cross-sections that make it possible to obtain quantitative results (i.e. mixing ratios) without the  
44 need for calibration gases. To the best of our knowledge, the actual list of biomass burning  
45 chemical species measured by FTIR has remained limited to ca. 36 compounds (Table 1); one goal  
46 of our research was to expand the list of chemical species to which infrared methods could be  
47 applied. All of the compounds detailed in this study have in fact been previously detected using  
48 other analytical methods (Akagi et al., 2013; Gilman et al., 2015; Karl et al., 2007; Koss et al.,

49 2018; Yokelson et al., 2009) such as proton-transfer-reaction time-of-flight mass spectrometry  
50 (PTR-ToF) (Koss et al., 2018) or gas chromatography-mass spectrometry (GC-MS) (Gilman et al.,  
51 2015), but have not as yet been identified using FTIR in burning investigations. We wished to  
52 determine if such species' signatures are also found sequestered in the IR spectra associated with  
53 wildland fire, and are thus amenable to IR detection. A second goal of the present study, whose  
54 biomass burning results are mostly detailed in a separate manuscript, was to better understand  
55 pyrolysis. Pyrolysis is the chemical transformation of material by heat in an oxygen-free or low-  
56 oxygen environment. Wildland fire consists of multiple processes: thermal decomposition  
57 (pyrolysis) of solid wildland fuels into gases, tars, and char is followed by combustion (oxidation)  
58 of pyrolysis products resulting in flame gases and particulate matter in the smoke. The visible  
59 flame is sustained by fuel gases that are produced by pyrolysis (Ward et al., 1991). These two  
60 processes (pyrolysis and combustion) are complementary given that heat released from the  
61 oxidation reactions facilitates further pyrolytic reactions allowing the fire to advance. Description  
62 and measurement (by any means) of the discrete pyrolysis products adjacent to the flames of a  
63 wildland fire has seldom been performed. Non-intrusive measurement of the (pyrolysis) gases in  
64 the near-flame environment is desirable from both a scientific and safety perspective.

65 The major gas-phase compounds emitted from wildland fires are H<sub>2</sub>O, CO<sub>2</sub>, CO and CH<sub>4</sub> (Ward  
66 et al., 1991), all of which are easily identified and quantified via FTIR spectroscopy. Lightweight  
67 hydrocarbons, oxygenated hydrocarbons, nitrogen and sulfur species are all minor products  
68 generated during burns (Yokelson et al., 1996; Lobert et al., 1991; Talbot et al., 1988). A host of  
69 more complex gases which can condense to form tar are also produced by pyrolysis of wildland  
70 fuels (Amini et al., 2019; Safdari et al., 2018). In a gas-phase IR spectrum of such species,  
71 however, peaks associated with the minor products are often obfuscated by more prominent

72 features, such as those from CO<sub>2</sub> or H<sub>2</sub>O and can only be recognized in the residual of a  
73 multicomponent simulated fit once larger features have been removed. Using data from a recent  
74 field campaign to measure pyrolysis products carried out in a pine forest at Fort Jackson, South  
75 Carolina, we have analyzed some of the IR spectra in more detail to search for the signatures of  
76 compounds not found in Table 1. As a partial guide of species for which to investigate, we searched  
77 for those species detected in previous thermogravimetric-FTIR (TG-FTIR) studies (Basilakis et  
78 al., 2001; Taghizadeh et al., 2015). TG-FTIR experiments, however, are typically small-scale and  
79 carried out in controlled environments (in contrast to ambient conditions of prescribed burns or  
80 large-scale laboratory burns) and thus represent burns with different oxidative capacities /  
81 combustion efficiencies (Fang et al., 2006; Yokelson et al., 1996; Akagi et al., 2014). In this study,  
82 we have chosen to examine field fire spectra for species that can be detected and quantified via IR  
83 spectroscopy both to add to the list of compounds, and also to improve the characterization (and  
84 ultimately the detection limits) of the other species listed in Table 1. That is to say, fire IR spectra  
85 are very complex and contain many overlapping peaks; the success of spectral analysis depends  
86 both on the selected spectral region and the proper analysis of all compounds included in the fit to  
87 that domain. The chemometric results become more reliable as signatures of all relevant species  
88 are included in the fit.

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101 **Table 1.** Compounds previously detected in biomass burning studies using FTIR methods (Akagi et al.,  
 102 2013; Alves et al., 2010; Burling et al., 2011; Goode et al., 2000; Karl et al., 2007; Akagi et al., 2014;  
 103 Burling et al., 2010; Christian et al., 2003; Christian et al., 2004; Goode et al., 1999; Selimovic et al.,  
 104 2018; Stockwell et al., 2014; Yokelson et al., 1996; Yokelson et al., 1997; Yokelson et al., 2009; Hatch et  
 105 al., 2017; Gilman et al., 2015).

Compounds		
Carbon monoxide (CO)	Formaldehyde (HCHO)	Nitric oxide (NO)
Carbon dioxide (CO <sub>2</sub> )	Formic acid (HCOOH)	Nitrogen dioxide (NO <sub>2</sub> )
Methane (CH <sub>4</sub> )	Acetic acid (CH <sub>3</sub> COOH)	Nitrous acid (HONO)
Ethane (C <sub>2</sub> H <sub>6</sub> )	Glycolaldehyde (C <sub>2</sub> H <sub>4</sub> O <sub>2</sub> )	Ammonia (NH <sub>3</sub> )
Ethene (C <sub>2</sub> H <sub>4</sub> )	Carbonyls as glyoxal (C <sub>2</sub> H <sub>2</sub> O <sub>2</sub> )	Nitrous oxide (N <sub>2</sub> O)
Acetylene (C <sub>2</sub> H <sub>2</sub> )	Methyl vinyl ether (C <sub>3</sub> H <sub>6</sub> O, MVE)	Hydrogen cyanide (HCN)
Propene (C <sub>3</sub> H <sub>6</sub> )	Acetone (CH <sub>3</sub> COCH <sub>3</sub> )	Peroxyacetyl nitrate (C <sub>2</sub> H <sub>3</sub> NO <sub>5</sub> )**
Isobutene (C <sub>4</sub> H <sub>8</sub> )	Hydroxyacetone (C <sub>3</sub> H <sub>6</sub> O <sub>2</sub> )	Sulfur dioxide (SO <sub>2</sub> )
1,3-Butadiene (C <sub>4</sub> H <sub>6</sub> )	Furan (C <sub>4</sub> H <sub>4</sub> O)	Carbonyl sulfide (OCS)
Isoprene (C <sub>5</sub> H <sub>8</sub> )	2-Methylfuran (C <sub>5</sub> H <sub>6</sub> O)*	Hydrogen chloride (HCl)
Limonene (C <sub>10</sub> H <sub>16</sub> )	Furaldehyde (C <sub>4</sub> H <sub>3</sub> OCHO)	Ozone (O <sub>3</sub> )**
Methanol (CH <sub>3</sub> OH)	Phenol (C <sub>6</sub> H <sub>5</sub> OH)	Water (H <sub>2</sub> O)

106 \* used in the fit, but not analyzed, \*\* secondary components detected downwind  
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## 108 2. EXPERIMENTAL

### 109 2.1 Site description and sampling device

110 In early May 2018, seven prescribed fires were conducted in pine forests at U.S. Army Garrison  
 111 Fort Jackson, adjacent to Columbia, South Carolina, at sites not far from previous smoke emission  
 112 studies (Akagi et al., 2013; Weise et al., 2015). The forest overstory was primarily longleaf pine  
 113 (*Pinus palustris* Mill.) and slash pine (*Pinus elliottii* Engelm.), while sparkleberry (*Vaccinium*  
 114 *arboreum* Marshall) dominated the understory vegetation. During each burn, pyrolyzed gases  
 115 emitted at the base of the flames before ignition were collected using an extractive probe and stored  
 116 in 3-liter Summa canisters. This approach was performed to selectively collect pyrolysis gases  
 117 prior to the onset of combustion. Details regarding the site description and sampling apparatus will  
 118 be provided in a separate paper.

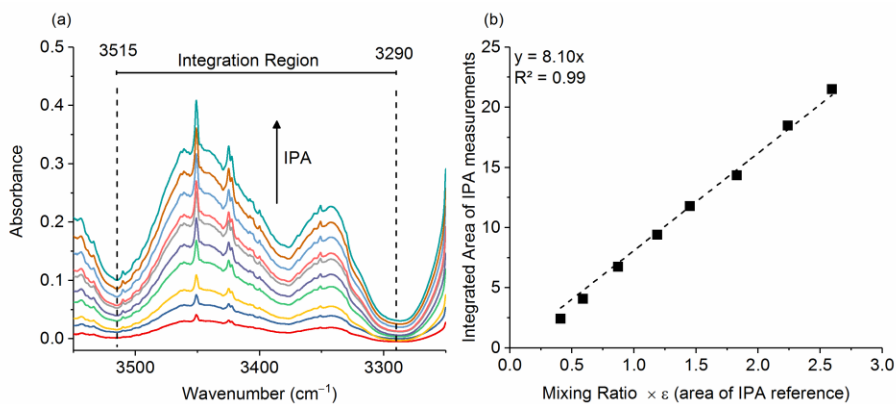
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## 120 2.2 FTIR Spectrometer

121 Gases were analyzed in the laboratory (on the same day or the day following the fire) using an 8-  
122 meter multipass (White) cell (Bruker Optics, A136/2-L) mounted in the sample compartment of a  
123 Bruker Tensor 37 FTIR. Ten canisters were returned from the field to the laboratory and in turn  
124 connected to the gas cell via 3/8" stainless steel tubing. The tubing and gas cell were both heated  
125 to 70 °C to prevent analyte adhesion to the inner surfaces. The White cell (White, 1942) was  
126 equipped with a pressure gauge and temperature probe, both of which were located on the gas  
127 outlet port; the thermocouple wire temperature probe extended into the White cell volume in order  
128 to more accurately measure the gas temperature. Prior to the start of the series of experiments, it  
129 was necessary to calibrate the path length of the variable path gas cell. Measurements **conducted**  
130 **at room temperature** of pure isopropyl alcohol (IPA, Sigma-Aldrich, 99.5%) at ten different  
131 pressures were collected and **a Beer-Lambert Law plot** was created to determine the length. The  
132 IR region from 3515 to 3290  $\text{cm}^{-1}$  was integrated (Figure 1a) **using OPUS v5.5 software**, and the  
133 corresponding areas plotted as a function of the IPA pressure (converted to ppm at 760 Torr)  
134 multiplied by the PNNL reference (Sharpe et al., 2004) integration area for a 1 ppm-meter IPA  
135 burden (Figure 1b). **The y-intercept was set to zero.** The slope is equal to the path length, which  
136 was determined to be  $8.10 \pm 0.1 \text{ m}$  (**standard error of the regression**).

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139 **Figure 1.** a) Multiple burden spectra of dry IPA for 10 measurements at varying pressures. The dashed  
 140 lines represent the integration limits used for spectral integration. b) Calibration plot with regression line  
 141 for IPA measurements. The slope of the regression is the path length in meters.  
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144 The White cell contained analyte smoke for the sample spectrum measurement but was filled with  
 145 ultra-high purity nitrogen gas for the background spectrum measurement (Johnson et al., 2013).  
 146 The FTIR interferometer, detector and sample compartments were purged with dry air from a dry-  
 147 air generator. Inspection of a single beam background spectrum showed no evidence of CO or CH<sub>4</sub>  
 148 contaminants and only negligible amounts of H<sub>2</sub>O and CO<sub>2</sub>. The Tensor 37 was equipped with a  
 149 globar source, a KBr beamsplitter and a broadband liquid nitrogen cooled mercury cadmium  
 150 telluride (MCT) detector, providing spectral coverage from 7,500 to 500 cm<sup>-1</sup>. The spectral  
 151 resolution was 0.6 cm<sup>-1</sup>, and a 2 mm Jacquinot aperture was used. The acquisition mode was set  
 152 to double-sided, forward-backward. For the Fourier transform, data were apodized with a  
 153 Blackman-Harris 3-Term function using a zerofill factor of 4 and phase corrected via the Mertz  
 154 (Mertz, 1967) method.

### 155 2.3 Quantitative Spectral Analysis

156 The program used for quantitative spectral analysis was MALT5 (Griffith, 2016), and it uses both  
 157 broadband reference spectra from PNNL (Sharpe et al., 2004; Johnson et al., 2009; Johnson et al.,  
 158 2006; Lindenmaier et al., 2017; Profeta et al., 2011) and absorption line intensities from HITRAN  
 159 (Gordon et al., 2017) [in units of cm<sup>-1</sup>/(molec × cm<sup>-2</sup>)] to iteratively fit a simulated spectrum to

160 the measured spectrum by optimizing the fit so as to minimize the mean-squared residual, i.e. the  
 161 difference between the measured and simulated spectra. Parameters such as path length, resolution,  
 162 apodization, temperature, pressure, spectral domain, target compounds / overlapping compounds  
 163 are all used as inputs to the spectral fit. **The MALT analysis technique has previously been used**  
 164 **in both open-path and extractive FTIR systems with active sources. (Burling et al., 2010; Burling**  
 165 **et al., 2011; Akagi et al., 2013; Akagi et al., 2014). The program has also been used for ground-**  
 166 **based solar FTIR measurements (Griffith et al., 2003).**

167 During the course of this study, MALT5 was used to identify **multiple** gas-phase species emitted  
 168 during the burns and quantify the gas mixing ratios via IR spectroscopy, **five of them** for the first  
 169 time. Part of the confirmation strategy is to process the experimental spectra both with and without  
 170 the target compound present in the fit and **then to** visually inspect the corresponding residuals.

171 Table 2 summarizes the IR-active vibrational mode used for each species in the spectral fit  
 172 (typically the species' strongest band in the longwave infrared window) along with the spectral  
 173 domain and a list of species with overlapping bands in that domain.

174 **Table 2.** Gas-phase species identified via FTIR, vibrational assignments (Chakraborty et al., 2016; Ghosh  
 175 et al., 1981; Es-Sebbar et al., 2014; Lord et al., 1952; Hamada et al., 1985; Hollenstein et al., 1971), and  
 176 spectral domains used for spectral fit and quantitation.

Target compound	Vibrational bands used for analysis	Spectral region (cm <sup>-1</sup> )	Other species fit in the same region
Naphthalene	$\nu_{46}$ at 782.3 cm <sup>-1</sup>	800–760	C <sub>2</sub> H <sub>2</sub> , CO <sub>2</sub> , HCN and H <sub>2</sub> O
Methyl nitrite	$\nu_8$ at 841.1 ( <i>cis</i> ) and 812.3 ( <i>trans</i> ) cm <sup>-1</sup>	865–775	C <sub>2</sub> H <sub>2</sub> , CO <sub>2</sub> , HCN, naphthalene, C <sub>2</sub> H <sub>4</sub> , allene, and H <sub>2</sub> O
Allene	$\nu_{10}$ at 845.3 cm <sup>-1</sup>	865–775	C <sub>2</sub> H <sub>2</sub> , CO <sub>2</sub> , HCN, naphthalene, C <sub>2</sub> H <sub>4</sub> , methyl nitrite, and H <sub>2</sub> O
Acrolein	$\nu_{10}$ at 1157.7 cm <sup>-1</sup>	1200–1100	Acetic acid (CH <sub>3</sub> COOH), furfural (C <sub>4</sub> H <sub>3</sub> OCHO), acetaldehyde, HCOOH, CH <sub>4</sub> , C <sub>2</sub> H <sub>4</sub> , and H <sub>2</sub> O
Acetaldehyde	$\nu_3$ at 2716.2 cm <sup>-1</sup>	2800–2650	CH <sub>4</sub> , HCHO, C <sub>2</sub> H <sub>2</sub> , acrolein, and H <sub>2</sub> O

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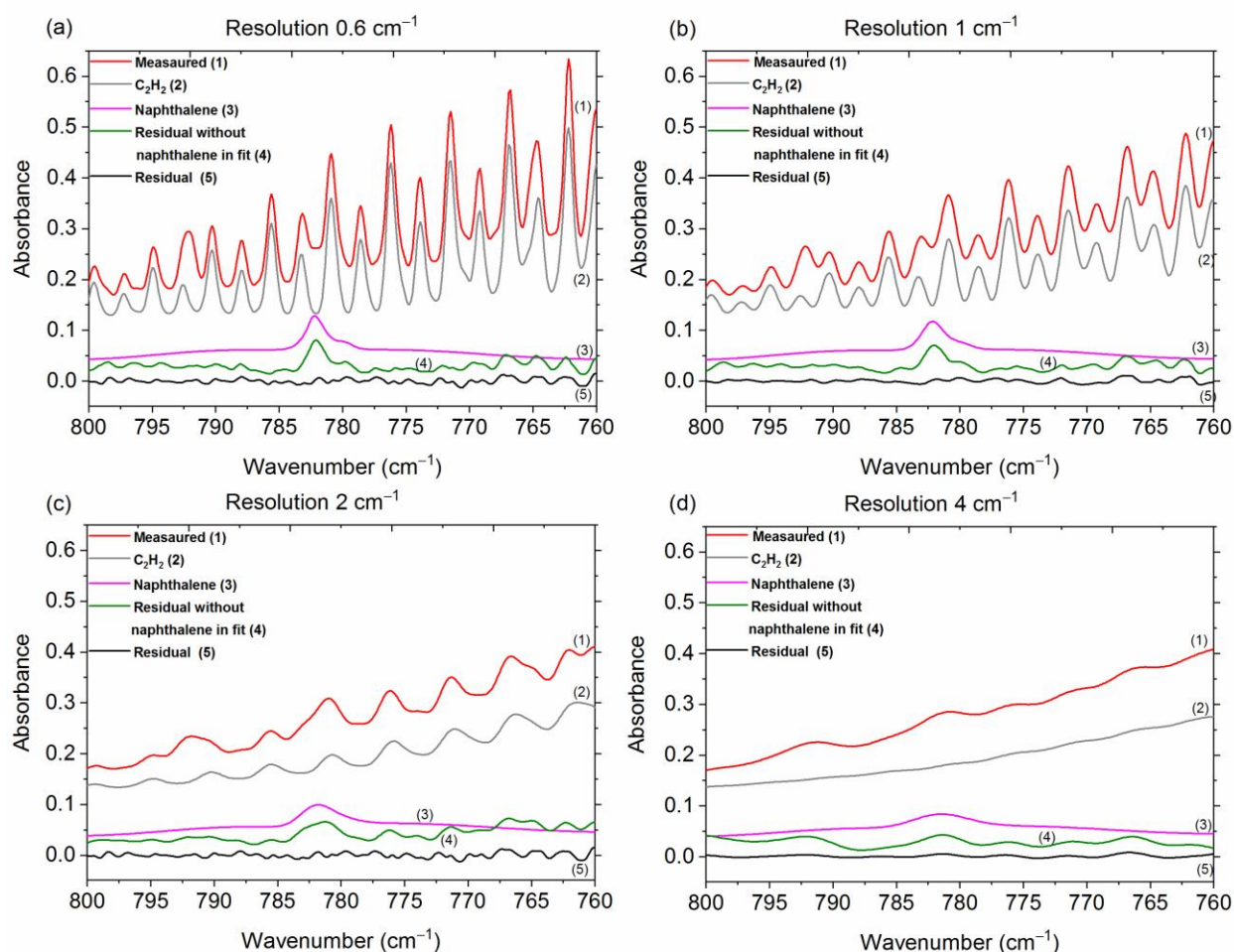


178 The PNNL database provides reference spectra measured at 5, 25 and 50 °C, all of which have  
179 been normalized to a number density of 296 K (~23 °C) and 1 atmosphere. While not perfectly  
180 optimal, the PNNL 50 °C reference spectra were used for evaluation to best match the bandshapes  
181 of the 70 °C experimental data. The fit of the 50 °C PNNL reference data to the 70 °C experimental  
182 spectra is obviously less than ideal. To correctly fit to the experimental spectra, reference data at  
183 70 °C are needed, but short of this knowledge of the temperature, partition function and individual  
184 line assignments are needed, and this changes for each line or set of lines for each molecule used  
185 in the fit. While MALT5 correctly accounts for gas temperature in all cases and for intensities of  
186 the HITRAN line-by-line data, it cannot do so for the PNNL reference data. At higher  
187 temperatures, there can be increases in population/intensity of the high-J lines with decreases for  
188 the lines originating with low J values. The effect is more pronounced for smaller, more rigid  
189 molecules (e.g. allene, acetaldehyde) than for the bands associated with larger, less rigid molecules  
190 of low symmetry. Preliminary estimates for the quality of fit estimate errors in the 2 to 5 percent  
191 range, though the value depends strongly on the species and which waveband is used for the fit.

## 192 **2.4 Spectral Resolution**

193 As mentioned in section 2.2, the spectral resolution was set to 0.6 cm<sup>-1</sup>, which is the highest  
194 resolution obtainable with this instrument. There are many benefits, but also a few disadvantages  
195 to using higher resolution (Herget et al., 1979). Most importantly, higher resolution allows one to  
196 resolve the narrow bands of key analytes **as well as interferences** and discriminate **the analyte signals**  
197 from lines or bands of interferences. For example, the 782 cm<sup>-1</sup> Q-branch of naphthalene was  
198 distinguished from the adjacent absorption lines of C<sub>2</sub>H<sub>2</sub> [Naphthalene's IR bands and results are  
199 discussed in greater depth in Section 3.1]. If a lower resolution were used, the deconvolution of  
200 naphthalene from C<sub>2</sub>H<sub>2</sub> would have been compromised, perhaps unfeasible. To demonstrate, one

201 of the experimental measurements collected at a resolution of  $0.6\text{ cm}^{-1}$  was deresolved to 1, 2, and  
202  $4\text{ cm}^{-1}$  using a Gaussian profile as seen in Figure 2. Those spectra were processed by MALT5 to  
203 check for the presence of naphthalene. Figure 2 displays the measured spectra and the scaled  
204 reference spectra for  $\text{C}_2\text{H}_2$  and naphthalene, and the corresponding residuals with and without  
205 naphthalene included in the fit for the a) original spectrum collected at  $0.6\text{ cm}^{-1}$  and the deresolved  
206 spectra at b)  $1\text{ cm}^{-1}$ , c)  $2\text{ cm}^{-1}$ , and d)  $4\text{ cm}^{-1}$ . With the reference spectra for the original  $0.6\text{ cm}^{-1}$   
207 measurement and the  $1\text{ cm}^{-1}$  deresolved spectrum (Figure 2a and b), the absorption lines for  $\text{C}_2\text{H}_2$   
208 and naphthalene overlap, but the  $782\text{ cm}^{-1}$  feature from naphthalene is still slightly visible in the  
209 original spectra. The naphthalene peak appears clearly in the residuals when it is not included in  
210 the fitting process, but is removed from the residual when naphthalene is included in the fit  
211 (discussed further below). As the resolution is further reduced (Figures 2c and 2d), however, the  
212 features broaden, and the distinction of the naphthalene peak from  $\text{C}_2\text{H}_2$  and other minor  
213 components (i.e.  $\text{CO}_2$ ,  $\text{HCN}$ ,  $\text{H}_2\text{O}$ , spectra not shown) is compromised. The specificity between  
214 compounds is lost and confidence in the identification/quantification of the target species,  
215 particularly for the weaker absorbers, diminishes as the resolution decreases. The well-known  
216 benefits of using a lower resolution are that spectra can be acquired more quickly at an improved  
217 signal-to-noise ratio. For the present measurements,  $0.6\text{ cm}^{-1}$  was deemed an appropriate  
218 resolution.



219  
 220 **Figure 2.** Measured and scaled reference spectra for  $C_2H_2$  and naphthalene, and corresponding residuals  
 221 with and without naphthalene included in the fit for the a) original spectrum collected at  $0.6\text{ cm}^{-1}$  and the  
 222 deresolved spectra at b)  $1\text{ cm}^{-1}$ , c)  $2\text{ cm}^{-1}$ , and d)  $4\text{ cm}^{-1}$ . The reference spectra for  $CO_2$ ,  $HCN$  and  $H_2O$  are  
 223 not shown ( $HCN$  was not included in fit when the resolution was  $4\text{ cm}^{-1}$ ; for resolutions 1, 2 and  $4\text{ cm}^{-1}$ ,  
 224  $H_2O$  was not included in the fit when naphthalene was removed from the fit). Spectra are offset for clarity.  
 225

## 226 2.5 Signal-to-Residual Detection Limits

227 In IR spectroscopy, detection limits often represent the minimum amount of analyte that may be  
 228 detected and are often reported as two to three times the noise level (Griffith et al., 2006). The  
 229 detection limit values presented in this paper are not minimal signal-to-noise limits in the sense  
 230 of a minimal spectral signal against a background of purely stochastic noise sources. In such cases,  
 231 the noise sources are typically of comparable or higher frequencies than the signal (Johnson et al.,

232 1991). Rather, the current limits represent average detection limits for a spectral residual derived  
233 from a convoluted spectrum arising from a gas mixture of differing and fluctuating chemical  
234 composition. The residuals are due to a least-squares fit of (fluctuations in) many complex features  
235 arising from numerous chemicals. That is to say, the residual is not due to just random instrumental  
236 noise but due to spectral features that can arise in the spectra, e.g. imperfectly subtracted features  
237 from strong absorbers or unidentified absorbers. For that reason, we report signal-to-residual  
238 rather than signal-to-noise detection limits. The detection limits for each compound in this study  
239 were thus derived using a value of three times the root-mean-square (RMS) value of the residual  
240 calculated over the corresponding frequency range (e.g. 800–760  $\text{cm}^{-1}$  was used for naphthalene).  
241 The peak-to-peak noise is more sensitive to fluctuations in the fit with levels typically 4 to 5× the  
242 RMS noise (Griffith et al., 2006). For the present data, however, the peak-to-peak values ranged  
243 from 5 to 10× the RMS noise, thus suggesting the peak-peak values tend to overstate the tractable  
244 noise level, i.e. understate the detection limit. The reported detection limits are thus presumably  
245 higher than what would be estimated with an FTIR in clean air conditions (i.e. only the analyte  
246 and dry air). Based on experience, the limits are typically far higher than what can be obtained  
247 with IR laser sensors where the intrinsically narrow laser linewidths allow for the probing of  
248 individual rotational-vibrational lines without drawing in overlapping spectral lines to a congested  
249 spectral fit (Taubman et al., 2004; Phillips et al., 2014; Wagner et al., 2011). While typically far  
250 more sensitive, such laser measurements can only analyze for one or a few species at a time, as  
251 opposed to the 30+ species seen by the broadband FTIR measurements.

### 252 **3. RESULTS AND DISCUSSION**

253 When modeling the burning process (Byram, 1959), complete combustion of 1 kg dry wood  
254 produces 1.82 kg  $\text{CO}_2$  and 0.32 kg  $\text{H}_2\text{O}$  for a total mass of products of 2.14 kg. Incomplete

255 combustion will yield additional products and less CO<sub>2</sub> and H<sub>2</sub>O while combustion of wet fuels  
 256 (Byram, 1959) increases the amount of H<sub>2</sub>O released. For infrared analysis of such smoke, much  
 257 of the challenge arises due not only to the large mole fractions of H<sub>2</sub>O and CO<sub>2</sub>, but to the fact that  
 258 both H<sub>2</sub>O vapor and CO<sub>2</sub> have strong features in the mid-IR that can clutter the spectrum rendering  
 259 certain spectral regions unusable. For burning and other atmospheric studies, ideal compounds for  
 260 detection via IR spectroscopy will thus have strong absorption coefficients that do not overlap with  
 261 the fundamental bands of H<sub>2</sub>O or CO<sub>2</sub>, i.e. are in a spectral window or microwindow (Griffith,  
 262 1996; Esler et al., 2000; Smith et al., 2011) free of strong interferences. Here, we consider five  
 263 such compounds emitted during this prescribed burn, but which had heretofore not been reported  
 264 as being detected by FTIR. **Table 3 presents the range of measured mixing ratios for the target**  
 265 **compounds along with averaged detection limits for 10 measurements collected during the**  
 266 **prescribed burns.** Individual compounds are discussed in turn regarding their formation  
 267 mechanism(s) as well as their detectable IR features and spectral confirmation for this study.  
 268 Lastly, the results are briefly compared with literature values using emission ratios (mixing ratios  
 269 of analyte to excess CO).

270 **Table 3.** Calculated mixing ratios for ten canister FTIR measurements along with average estimated  
 271 residual detection limits for the target compounds derived using 3 times the root-mean-square of the  
 272 residual **in very congested spectra.** Error bars represent standard deviation ( $1\sigma$ ) **from** the mean.

Target compound	Calculated mixing ratios (ppm)			Averaged <b>residual</b> detection limit (ppm) using root-mean-square (RMS) value of the residual
	Min	Max	Average	
Naphthalene*	1.4	19.9	8.5 ± 2.1	1.9 ± 0.5
Methyl nitrite*	2.3	21.0	8.7 ± 2.4	2.2 ± 0.4
Allene	2.2	37.8	13.1 ± 3.6	3.0 ± 0.6
Acrolein	14.7	125.7	43 ± 12	6.1 ± 1.5
Acetaldehyde	34.5	264.8	103 ± 27	11.7 ± 3.2

\*One measurement was below the detection limit.

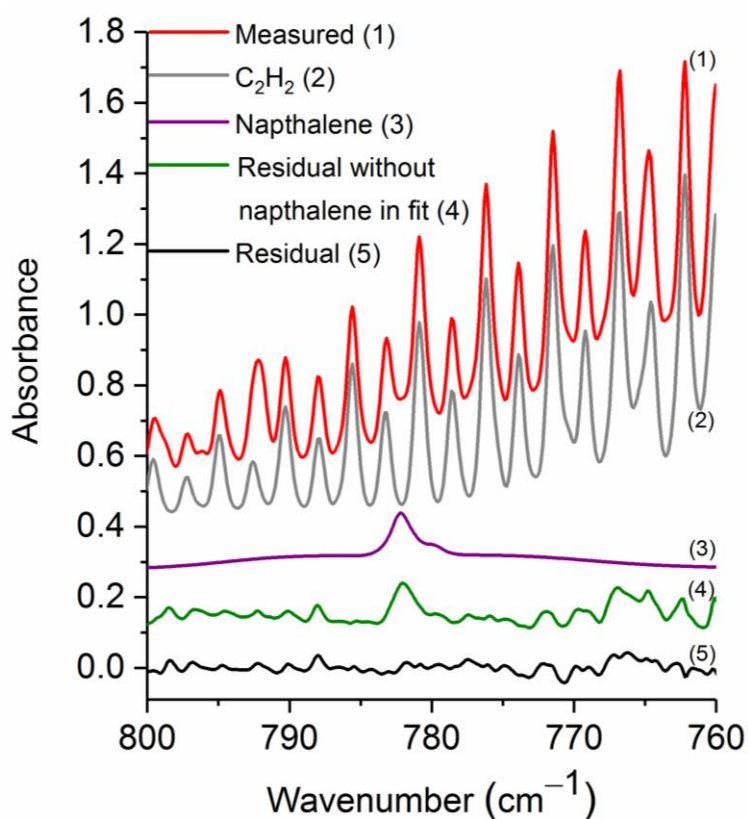
273  
274

### 275 3.1 Naphthalene

276 Naphthalene ( $C_{10}H_8$ ) is a polycyclic aromatic hydrocarbon (PAH) that is emitted from certain  
277 **industrial processes** as well as from the combustion of gasoline and oil (Jia et al., 2010). It is a  
278 condensable hydrocarbon also generated by biomass pyrolysis (Liu et al., 2017). There are a  
279 number of pyrolysis formation routes (Williams et al., 1999; Liu et al., 2017; Lu et al., 2004;  
280 Fairburn et al., 1990; Richter et al., 2000). One proposed mechanism is the generation of single  
281 ring aromatic compounds such as benzene, toluene and styrene via Diels-Alder reaction of alkenes;  
282 the single ring aromatic compound then combines with alkenes to form double-ring PAHs, such  
283 as naphthalene (Fairburn et al., 1990). Naphthalene may even undergo subsequent reactions to  
284 form still larger polyaromatics (Fairburn et al., 1990; Richter et al., 2000). Naphthalene has been  
285 detected (via GC-MS) in tars that were condensed from gas-phase pyrolysis products of both live  
286 and dead southeastern fuels, such as live oak (*Quercus virginiana*) and swamp bay (*Persea*  
287 *palustris*) (Safdari et al., 2018). It has also been detected (Koss et al., 2018; Hosseini et al., 2014;  
288 Aurell et al., 2017) in the gas-phase in laboratory burning experiments. The detection of gas-phase  
289 naphthalene from wildland fire emissions is thus not surprising, but this is the first report of its  
290 identification via IR spectroscopy. The best spectral feature for identification and quantification is  
291 the  $\nu_{46}$  IR mode near  $782.3\text{ cm}^{-1}$ , which corresponds to the H-C-C out-of-plane bend (Chakraborty  
292 et al., 2016). There are other bands at  $3067.7$  and  $3058.0\text{ cm}^{-1}$  previously assigned to  $\nu_{29}$  and  $\nu_{17}$ ,  
293 respectively (Chakraborty et al., 2016). Both of these modes have smaller absorption coefficients  
294 as compared to  $\nu_{46}$ , however, and are located in the C-H stretching region, which is common to  
295 nearly all hydrocarbons and thus provides less specificity.

296 Figure 3 shows a prescribed burn spectrum in the region from  $800$  to  $760\text{ cm}^{-1}$ . The primary  
297 spectral signatures in this plot are those of R-branch rotational-vibrational lines associated with

298 the  $\nu_5$  fundamental of  $C_2H_2$  (Kabbadj et al., 1991), but there are also absorptions due to  $CO_2$ , HCN,  
299  $H_2O$  (individual spectral contributions not shown) and naphthalene. When all of the spectral  
300 components except for naphthalene are included in the fitting process, the residual (green trace,  
301 #4) displays a prominent feature at  $782.3\text{ cm}^{-1}$ , which we ascribe to naphthalene. When  
302 naphthalene is included in the fit, the feature in question is not seen as demonstrated by the black  
303 trace (#5) of Figure 3. Including naphthalene in the analysis clearly improves the fit, which  
304 consequently improves the derived values for the other species. This observation is consistent in  
305 the spectral analyses for all target compounds discussed below.



306  
307 **Figure 3.** Measured spectrum, scaled reference spectra for  $C_2H_2$  and naphthalene, and residuals with and  
308 without naphthalene included in the fit. For clarity, the spectral contributions for  $CO_2$ , HCN, and  $H_2O$  are  
309 not shown. All spectra are at  $0.6\text{ cm}^{-1}$  resolution and have been offset. The calculated mixing ratio of  
310 naphthalene in this measured spectrum is 16.4 ppm.  
311  
312



313 Table 3 displays the range of measured mixing ratios for naphthalene along with averaged  
314 detection limits for the 10 measurements. In the measurements, naphthalene's mixing ratios  
315 ranged from 1.4 to 19.9 ppm, and the averaged RMS-derived detection limit was  $1.6 \pm 0.5$  ppm;  
316 different detection limits were observed for each spectrum. One of the measurements had a mixing  
317 ratio of 2.9 ppm, yet its corresponding RMS-derived detection limit was 3.7 ppm, and is thus below  
318 the estimated detection limit.

### 319 3.2 Methyl Nitrite

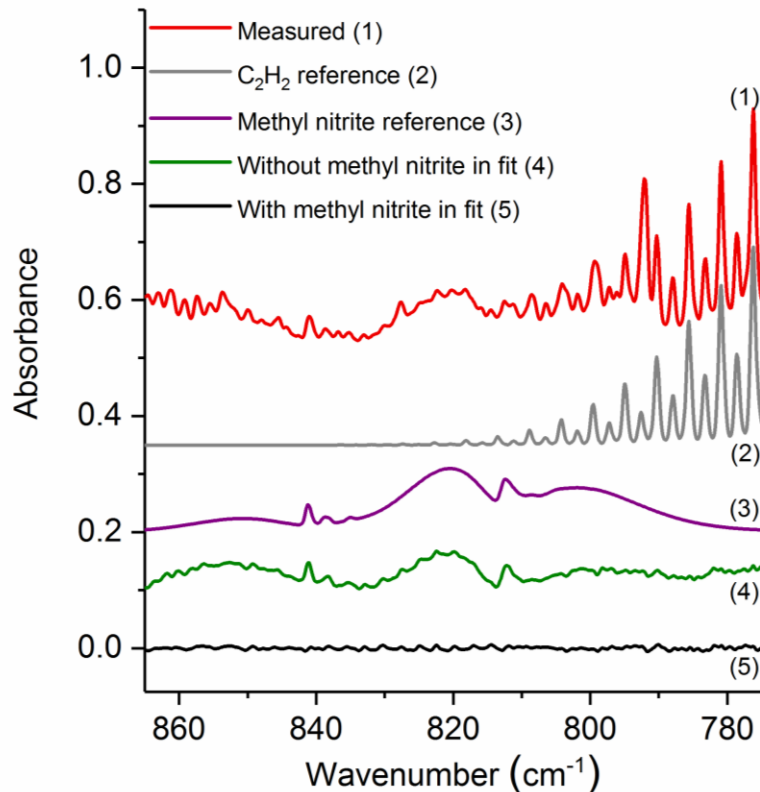
320 A second compound detected for the first time in wildland fire IR spectra was methyl nitrite  
321 ( $\text{CH}_3\text{ON}=\text{O}$ ). Methyl nitrite has previously been observed in aged cigarette smoke (Schmeltz et  
322 al., 1977) and the exhaust of engines fueled by methanol–diesel blends (Jonsson et al., 1982). It  
323 has also been observed as a minor product for the thermal decomposition of both nitrate esters  
324 (Boschan et al., 1955) and isopropyl nitrate at low temperatures and pressures (Griffiths et al.,  
325 1975). Methyl nitrite has moreover been detected in wildland fire emissions by GC-MS (Gilman  
326 et al., 2015). Other nitrogen-containing organic compounds such as acetonitrile ( $\text{CH}_3\text{CN}$ ) emitted  
327 from burns have been previously correlated to the fuel nitrogen content. However, methyl nitrite  
328 [and another oxygenated nitrogen organic compound, isocyanic acid ( $\text{HNCO}$ )] did not show any  
329 significant dependency on fuel N-content (Coggon et al., 2016). It has been suggested that methyl  
330 nitrite is not only a minor direct product of combustion (Finlayson-Pitts et al., 1992), but in  
331 addition is generated *in situ* by the secondary reaction of methanol ( $\text{CH}_3\text{OH}$ ) with nitrogen dioxide  
332 ( $\text{NO}_2$ ).

333 We also note that methyl nitrite is an oxidizing agent and is used as a rocket propellant. It is thus  
334 plausible that the methyl nitrite detected in the present study was not a product of the fire, but  
335 emanated from munitions used in training at Ft. Jackson. However, while the records of the



336 munitions used at all base sites were not complete, a survey of these records did not indicate the  
337 use of methyl nitrite in any munitions at the Ft. Jackson plots where the present burn samples were  
338 collected.

339 With regards to the IR spectra, methyl nitrite exists in equilibrium as a mixture of two conformers-  
340 *cis* and *trans*; at room temperature (25 °C) it is estimated as 58% *cis* and 42% *trans* (Bodenbinder  
341 et al., 1994). The PNNL reference spectrum for methyl nitrite was created using a nascent mixture  
342 of *cis* and *trans*, and the single spectrum contains features from both conformers (Sharpe et al.,  
343 2004). We were able to use the same band associated with both conformers, namely the  $\nu_8$  band,  
344 which is at  $841.1\text{ cm}^{-1}$  for the *cis* conformer and at  $812.4\text{ cm}^{-1}$  for the *trans* conformer (Ghosh et  
345 al., 1981). The  $\nu_8$  mode is associated with the N–O stretch and is very strong for both conformers  
346 (Ghosh et al., 1981). We note that methyl nitrite also has very strong bands at  $627.8\text{ cm}^{-1}$  (*cis*) for  
347  $\nu_9$  ONO bending, as well as at  $1620.1\text{ cm}^{-1}$  (*cis*) and  $1677.4\text{ cm}^{-1}$  (*trans*) due to the  $\nu_3$  N=O stretch  
348 (Ghosh et al., 1981). These bands, however, are of lesser utility for IR detection: The  $\nu_9$  peak is  
349 masked by CO<sub>2</sub> bending mode lines, and the  $\nu_3$  peak is obfuscated by the H<sub>2</sub>O bending mode lines.  
350 The spectral region used for evaluation was  $865\text{--}775\text{ cm}^{-1}$ , which contains the  $\nu_8$  band for both the  
351 *cis* and *trans* conformers (Ghosh et al., 1981). Figure 4 shows the experimental spectrum from  
352 the prescribed burn along with scaled reference spectra for the two major compounds used in the  
353 fit: C<sub>2</sub>H<sub>2</sub> and methyl nitrite. While important, other minor compounds, such as CO<sub>2</sub>, HCN,  
354 naphthalene, C<sub>2</sub>H<sub>4</sub>, allene and H<sub>2</sub>O, were also included in the analysis, but their spectral  
355 contributions are not plotted. Additionally, Figure 4 displays the residuals both when methyl nitrite  
356 was included in the fitting process and when it was excluded. Upon inspection of the residual  
357 spectrum where it was excluded (green trace, #4), it is clear that both the *cis* and *trans* features  
358 from  $\nu_8$  are present, and this confirms methyl nitrite in the pyrolysis smoke.



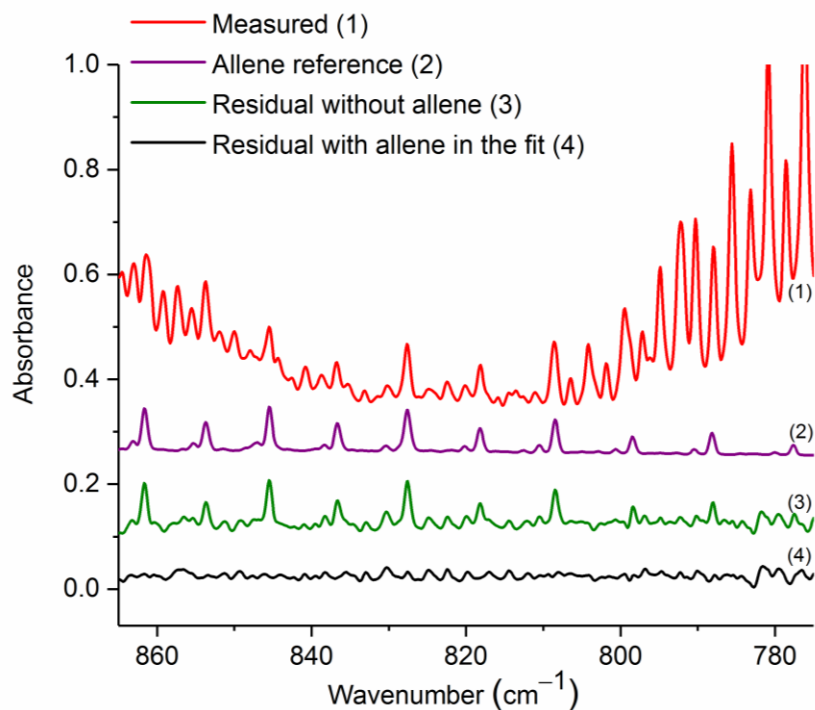
359 **Figure 4.** Measured experimental spectrum and the individual spectral contributions for the major  
 360 components ( $C_2H_2$  and methyl nitrite) and residuals with and without methyl nitrite included in the fit. For  
 361 clarity, the spectral contributions for  $CO_2$ , HCN, naphthalene,  $C_2H_4$ , allene, and  $H_2O$  are not shown. All  
 362 spectra are at  $0.6\text{ cm}^{-1}$  resolution and have been offset for clarity. The calculated mixing ratio of methyl  
 363 nitrite in this measured spectrum is 21.0 ppm.  
 364  
 365

366 The mixing ratio and RMS-derived detection limit for methyl nitrite for the displayed experimental  
 367 spectrum in Figure 4 are 21.0 ppm and 1.4 ppm, respectively. The range for mixing ratios and  
 368 averaged detection limits for methyl nitrite are summarized in Table 3. Methyl nitrite was detected  
 369 with confidence in 9 of the 10 measurements; only one of the measurements was below the RMS-  
 370 derived detection limit.

### 371 3.3 Allene

372 Allene (1,2-propadiene,  $CH_2=C=CH_2$ ) is of high symmetry ( $D_{2d}$ ) and has the two methylene  
 373 groups with their H-C-H planes at right angles to each other (Lord et al., 1952). The compound  
 374 has previously been detected in biomass burning grab samples using GC (Akagi et al., 2013).

375 Allene is a proposed precursor in the burning process that contributes to the formation of both  
376 aromatic compounds and soot (Frenklach et al., 1983; Frenklach et al., 1988). Lifshitz et al. have  
377 observed (at temperatures ranging from 757–847 °C) that the structural isomerization of allene and  
378 propyne ( $\text{CH}_2=\text{C}=\text{CH}_2 \leftrightarrow \text{CH}_3-\text{C}\equiv\text{CH}$ ) will take place via a unimolecular reaction faster than the  
379 decomposition reaction (Lifshitz et al., 1975). Additionally, these same authors investigated the  
380 pyrolysis of allene and propyne and observed that  $\text{C}_2\text{H}_4$  was generated from allene while  $\text{CH}_4$  and  
381  $\text{C}_2\text{H}_2$  were mainly formed from propyne (Lifshitz et al., 1976). Unfortunately, the strongest IR  
382 band for propyne (near  $634\text{ cm}^{-1}$ ) is obscured by  $\text{CO}_2$  bending mode lines. Due to the interferences  
383 we cannot with confidence identify propyne in the measurements; we can, however, detect allene.  
384 In the mid-IR, allene has several strong rotational-vibrational lines near  $845\text{ cm}^{-1}$  associated with  
385 the sub-bands of the perpendicular band  $\nu_{10}$ , which is due to  $\text{CH}_2$  rocking (Lord et al., 1952).  
386 Additionally, allene has a moderately strong band at  $1958.6\text{ cm}^{-1}$  due to the  $\nu_6$  C–C stretching  
387 (Lord et al., 1952). However, the  $\nu_6$  band is not useful for detection due to interference from the  
388  $\text{H}_2\text{O}$  bending mode lines.



389

390 **Figure 5.** Measured absorbance spectrum and residual with and without allene included in the fit, along  
 391 with the scaled reference spectrum for allene. For clarity, the spectral contributions for C<sub>2</sub>H<sub>2</sub>, CO<sub>2</sub>, HCN,  
 392 naphthalene, C<sub>2</sub>H<sub>4</sub>, methyl nitrite, and H<sub>2</sub>O are not shown. All spectra are at 0.6 cm<sup>-1</sup> resolution and have  
 393 been offset for clarity. The calculated mixing ratio of allene in this measured spectrum is 37.8 ppm.  
 394

395 Figure 5 shows the measured absorbance spectrum, scaled allene reference spectrum and the  
 396 associated residual with and without allene included in the fit. The absorption lines associated  
 397 with allene are clearly seen in the resulting spectrum when allene is not included in the fit (green  
 398 trace, #3), thus confirming that allene is one of the primary components contributing to the features  
 399 in this spectral domain. For the experimental spectrum displayed in Figure 5, the calculated mixing  
 400 ratio for allene is 37.8 ppm, and the RMS-derived detection limit is 5.4 ppm.

401

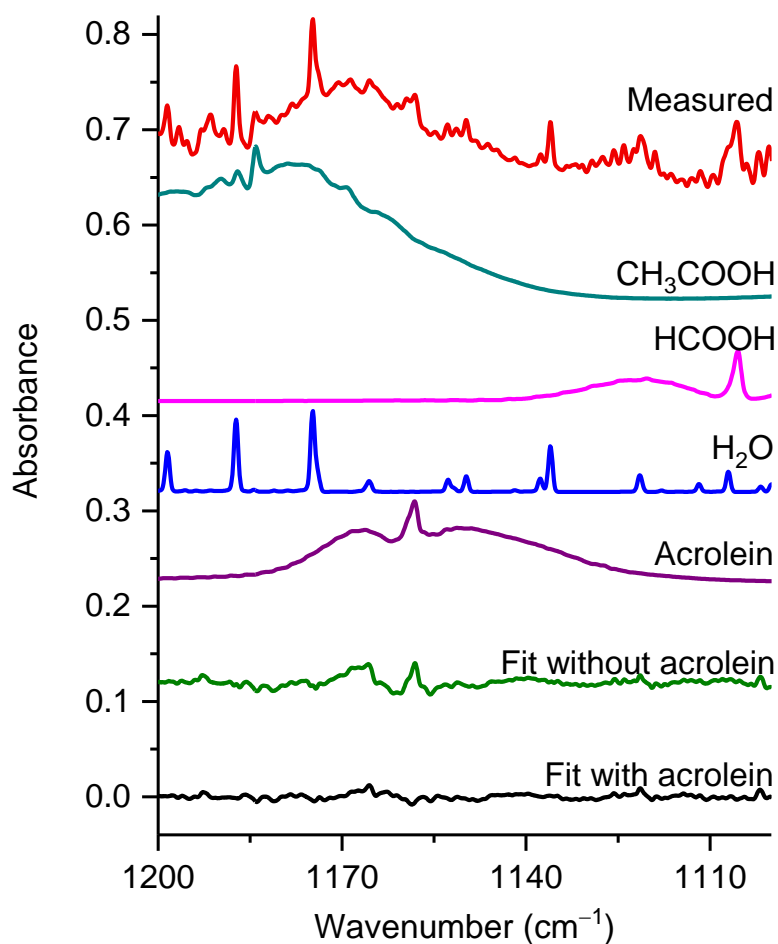
### 402 3.4 Acrolein and Acetaldehyde

403 The two aldehydes, acrolein (CH<sub>2</sub>=CHCHO) and acetaldehyde (CH<sub>3</sub>CHO), have also been  
 404 identified for the first time in burning IR spectra. It has been proposed that both acrolein and

405 acetaldehyde are formed from the pyrolysis of cellulose (a major constituent of biomass) via the  
406 intermediate glycerol, which is a moiety in the structure of levoglucosan, a known pyrolysis  
407 product of cellulose (Stein et al., 1983). Stein et al. observed that acrolein, acetaldehyde and CO  
408 were the initial decomposition products from the pyrolysis of glycerol (Stein et al., 1983). Both  
409 of these compounds have been detected in previous wildland fires studies via methods such as GC  
410 (Akagi et al., 2013) and PTR-ToF (Koss et al., 2018; Brillì et al., 2014), but have not yet been  
411 identified via IR.

412 Acrolein, the simplest unsaturated carbonyl, exists in two forms, *s-cis* and *s-trans*, with *s-trans*  
413 being the more stable, and consequently the more abundant conformer (Wagner et al., 1957). It  
414 has been estimated that the fractions of *s-cis* and *s-trans* are about 4 and 96% at 20 °C, and 7 and  
415 93% at 100 °C, respectively (Alves et al., 1971). The largest IR feature for acrolein is the  $\nu_5$  C=O  
416 stretch (Hamada et al., 1985) at  $1724.1\text{ cm}^{-1}$ , but this band is heavily overlapped by water lines.  
417 There is also the  $\nu_{16}$  band (Hamada et al., 1985) at  $958.8\text{ cm}^{-1}$ , but this feature overlaps with  
418 multiple other strongly absorbing compounds, such as C<sub>2</sub>H<sub>4</sub>. We have therefore focused acrolein's  
419 analysis using the  $\nu_{10}$  band (C–C stretch) (Hamada et al., 1985) at  $1157.7\text{ cm}^{-1}$ .

420 Figure 6 displays a very congested biomass burning spectrum with individual contributions for  
421 several species included in the fit [contributions for furfural (C<sub>4</sub>H<sub>3</sub>OCHO), acetaldehyde, CH<sub>4</sub>, and  
422 C<sub>2</sub>H<sub>4</sub> are included, but not plotted] as well as the residual with and without acrolein included in  
423 the fitting process. When acrolein is not included in the fit, features (both near 1168 and at 1157.7  
424  $\text{cm}^{-1}$ ) that resemble acrolein are observed in the residual spectrum as seen in the green trace in  
425 Figure 6. When acrolein is included in the fit, the features in question are not observed. For  
426 acrolein, no mixing ratios were observed below the RMS-derived detection limits.

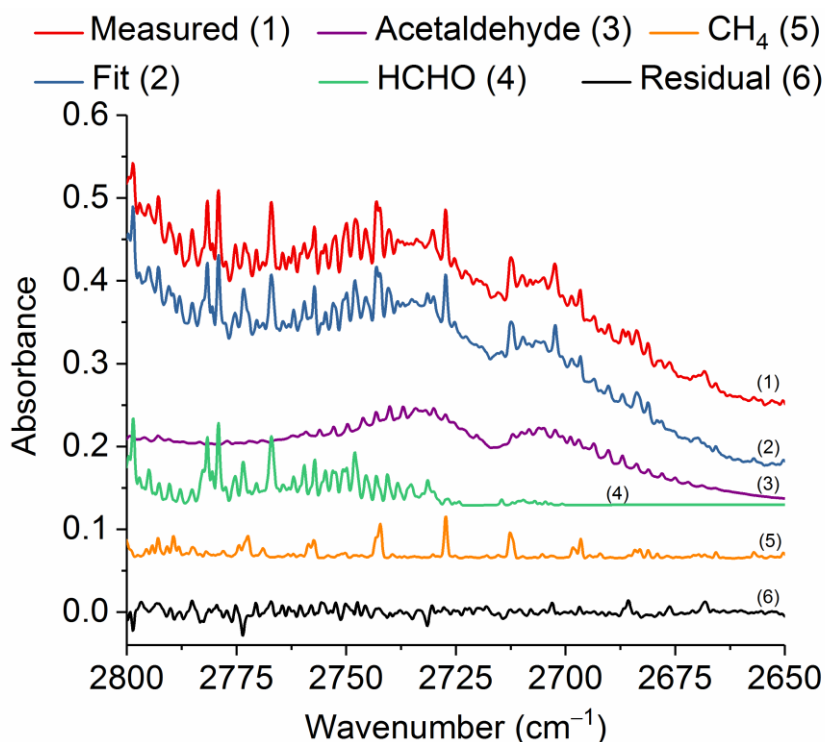


427  
 428 **Figure 6.** Measured spectrum and the individual spectral contributions for the major components and  
 429 associated residual with and without acrolein included in the fit. For clarity, the spectral contributions for  
 430 furfural ( $C_4H_3OCHO$ ), acetaldehyde,  $CH_4$ , and  $C_2H_4$  are not shown. All spectra are at  $0.6\text{ cm}^{-1}$  resolution  
 431 and have been offset for clarity. The calculated mixing ratio of acrolein in this measured spectrum is 99.9  
 432 ppm.  
 433

434 Similar to acrolein, acetaldehyde has its strongest IR feature due to the  $C=O$  stretch (Hollenstein  
 435 et al., 1971), with  $\nu_4$  found at  $1746.1\text{ cm}^{-1}$ . Again, due to the presence of water lines in the  
 436 spectrum, this feature is not practical for detection. The aldehyde  $\nu_3$   $C-H$  stretching band  
 437 (Hollenstein et al., 1971) at  $2716.2\text{ cm}^{-1}$  was instead used for analysis. Figure 7 shows the  
 438 measured and fitted spectra as well as spectral contributions from the major individual components

439 used to calculate the fitted spectrum and the corresponding residual. Other minor components,  
440 such as acrolein, C<sub>2</sub>H<sub>2</sub> and H<sub>2</sub>O, were also included in the fit, but their reference spectra are not  
441 displayed in Figure 7. The spectral profile of acetaldehyde with its P and R branches of  $\nu_3$  is easily  
442 discernable even before deconvolution of the measured spectrum. Similar to acrolein, all of the  
443 mixing ratios for acetaldehyde were above the RMS-derived detection limit.

444



445 **Figure 7.** Measured and fitted spectra as well as the individual components (for clarity, the spectral  
446 contributions for acrolein, C<sub>2</sub>H<sub>2</sub>, H<sub>2</sub>O are not shown) and associated residual in the spectral region 2800–  
447 2650 cm<sup>-1</sup>. All spectra are at 0.6 cm<sup>-1</sup> resolution and have been offset for clarity. The calculated mixing  
448 ratio of acetaldehyde in this measured spectrum is 252.8 ppm.  
449

450

### 451 3.5 Comparison to Other Measurements

452 Preliminary emission ratios (relative to carbon monoxide) for the **five new** reported compounds  
453 are compared to those **presented** in previous wildland burning investigations. **An emission ratio is**  
454 **a standard metric used in fire emission measurements and is defined as the change in the mixing**  
455 **ratio of the target compound relative to the change in mixing ratio of the reference species,**

456 generally either carbon monoxide or carbon dioxide (Urbanski et al., 2008). Here, carbon  
457 monoxide is used as the reference species since the present study focuses on pyrolysis, and prior  
458 fire studies generally provide emission ratios relative to carbon monoxide, which makes it a  
459 convenient quantity for comparison. Table 4 displays the average emission ratios and the standard  
460 deviations ( $1\sigma$ ) for this study as well as emission ratios reported by Koss et al. (2018), Ferek et al.  
461 (1998), Brilli et al. (2014), and Gilman et al. (2015). As shown in the table, there is considerable  
462 variation between the studies due to multiple factors such as different fuel types, analytical  
463 methods, sampling approaches and experimental conditions. For example, the study by Ferek et  
464 al. (1998) focused on collection of airborne samples, while Brilli et al. (2014) measured gases  
465 under nocturnal conditions using a ground-based system. Inspection of the table shows that the  
466 measured emission ratio values are not unprecedented, and in most cases are within range of  
467 previous measurements. Because they have the same molar mass, the mass spectrometric  
468 techniques in some cases cannot distinguish allene from propyne. There are of course advantages  
469 and disadvantages for the various measurement techniques typically used in biomass burning  
470 investigations. For a detailed summary of instrumental methods (including species measured, time  
471 resolution and detection limits), the reader is referred to Table 1 found in Koss et al. (2018). The  
472 FTIR technique presented in that table is open-path (OP-FTIR), which has a lower (better)  
473 detection limit (typically on the order of 10s of ppb) as compared to the extractive method used in  
474 the present study (low ppm, see Table 3). It should be acknowledged that the target compounds  
475 and spectral analysis methods of this study are also fully applicable to both infrared laser systems  
476 and OP-FTIR systems.

477

478



479 **Table 4.** Emission ratios relative to CO and standard deviations ( $1\sigma$ ) for the present study and for three  
 480 other previously published biomass burning studies.

Target compounds	Present average emission ratios to CO (ppb/ppm)	Koss et al. (2018) fire-integrated emission ratio to CO (ppb/ppm)	Ferek et al. (1998) emission ratio to CO (ppb/ppm)	Brilli et al. (2014) emission ratios to CO (ppb/ppm)	Gilman et al. (2015) discrete emission ratios to CO (ppb/ppm)		
					South-western fuels	South-eastern fuels	Northern fuels
Method	FTIR	PTR-ToF-MS	GC-FID*	PTR-ToF-MS	GC-MS	GC-MS	GC-MS
Naphthalene	0.79 ± 0.47	0.20 ± 0.16	n/a	n/a	0.0070 ± 0.0048	0.0040 ± 0.0050	0.022 ± 0.012
Methyl nitrite	0.94 ± 0.85	n/a	n/a	n/a	0.9 ± 1.1	0.52 ± 0.51	0.76 ± 0.90
Acrolein	4.0 ± 1.8	5.4 ± 3.0	n/a	3.14 ± 0.12	0.82 ± 0.68	1.31 ± 0.88	3.5 ± 1.7
Acetaldehyde	9.4 ± 3.6	7.4 ± 5.2	n/a	37.3 ± 1.4	1.6 ± 1.2	2.8 ± 1.8	5.5 ± 3.6
Allene (Propadiene)**	1.05 ± 0.24	n/a	0.1 ± 0.1	8.73 ± 0.28	n/a	n/a	n/a

481 \*GC-FID is gas chromatography with flame ionization detector

482 \*\*Brilli et al. (2014) use both 1-propyne and propadiene to represent C<sub>3</sub>H<sub>4</sub>. Gilman et al. (2015) report emission  
 483 ratios for propyne, but not allene.  
 484

#### 485 4. SUMMARY

486 Gas-phase compounds with appreciable band intensities and appreciable concentrations can be  
 487 both identified and quantified using IR spectroscopy. We have used such spectral information for  
 488 seminal IR detection of five compounds generated during prescribed forest fire burns. Deriving  
 489 the mixing ratios from the congested spectra obtained from wildland smoke samples is more  
 490 challenging due to the multiple overlapping spectral features: Sophisticated software and analysis  
 491 are required in carefully selected spectral windows. We have reported first IR detections of five  
 492 molecules that had previously not been observed by FTIR in ambient measurements of wildland  
 493 emissions. Most of the compounds (excluding acetaldehyde), had their primary features become  
 494 apparent only after the larger spectral features had been fitted and subtracted.

495

496

497

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509

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