



1                   **Identification of Gas-phase Pyrolysis Products in a Prescribed Fire:**  
2                   **Seminal Detections Using Infrared Spectroscopy for Naphthalene,**  
3                   **Methyl Nitrite, Allene, Acrolein and Acetaldehyde\*\***

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14  
15                   **ABSTRACT**

16                   Volatile organic compounds (VOCs) are emitted from many sources, including wildland fire;  
17                   VOCs have received heightened emphasis due to such gases' influential role in the atmosphere, as  
18                   well as possible health effects. We have used extractive infrared (IR) spectroscopy on recent  
19                   prescribed burns in longleaf pine stands and herein report seminal detection of five compounds  
20                   using this technique. The newly reported IR detections include naphthalene, methyl nitrite, allene,  
21                   acrolein and acetaldehyde. We discuss the approaches used for detection, particularly the software  
22                   methods needed to fit the analyte and multiple (interfering) spectral components within the  
23                   selected spectral micro-window(s). We also discuss the method's detection limits and individual  
24                   species' context in terms of atmospheric chemistry.

25



## 26 1. INTRODUCTION

27 Wildland fire releases significant quantities of trace gases into the environment (Crutzen et al.,  
28 1979; Andreae, 1991; Andreae et al., 2001; Akagi et al., 2011; Yokelson et al., 2013), and such  
29 gases can significantly influence atmospheric chemistry (Crutzen et al., 1990). In some parts of  
30 the world, wildfires are becoming more prevalent as well as increasing in impact (Miller et al.,  
31 2009; Turetsky et al., 2011). In many areas, however, prescribed burning is used as a preventive  
32 tool to reduce hazardous fuel buildups in an effort to reduce or eliminate the risk of such wildfires  
33 (Fernandes et al., 2003). Understanding the products associated with the burning of biomass has  
34 received considerable attention since the emissions can markedly impact the atmosphere. Fourier  
35 transform infrared (FTIR) spectroscopy is one technique that has been extensively used to identify  
36 and quantify gases emitted from burns, generally used in either an open path configuration (Burling  
37 et al., 2010; Akagi et al., 2014; Stockwell et al., 2014; Selimovic et al., 2018) or as an extractive  
38 method (Burling et al., 2011; Akagi et al., 2013; Akagi et al., 2014). Extractive systems typically  
39 use a long-path gas cell coupled to an FTIR instrument so as to increase the sensitivity. Such  
40 approaches have been quite successful; an increasing number of species continue to be identified  
41 and quantified due to the availability of reference gas-phase spectral libraries such as the PNNL  
42 library (Sharpe et al., 2004) or the HITRAN database (Gordon et al., 2017). Such libraries contain  
43 absorption cross-sections that make it possible to obtain quantitative results (i.e. mixing ratios)  
44 without the need for calibration gases. To the best of our knowledge, the actual list of biomass  
45 burning chemical species measured by FTIR has remained limited to ca. 36 compounds (Table 1);  
46 one goal of our research was to expand the list of chemical species to which infrared methods  
47 could be applied. All of the compounds detailed in this study have in fact been previously detected  
48 using other analytical methods (Karl et al., 2007; Yokelson et al., 2009; Akagi et al., 2013; Gilman



49 et al., 2015; Koss et al., 2018) such as proton-transfer-reaction time-of-flight mass spectrometry  
50 (PTR-ToF) (Koss et al., 2018) or gas chromatography-mass spectrometry (GC-MS) (Gilman et al.,  
51 2015), but have not as yet been identified using FTIR in burning investigations. We wished to  
52 determine if such species' signatures are also found sequestered in the IR spectra associated with  
53 wildland fire, and are thus amenable to IR detection. A second goal of the present study, whose  
54 biomass burning results are mostly detailed in a separate manuscript, is to better understand  
55 pyrolysis. Every wildland fire consists of two processes: thermal decomposition (pyrolysis) of  
56 solid wildland fuels into gases, tars, and char followed by combustion (oxidation) of the pyrolysis  
57 products resulting in flame gases and particulate matter in the smoke. Description and  
58 measurement (by any means) of the pyrolysis products adjacent to the flames of a wildland fire  
59 has seldom been performed. Non-intrusive measurement of the (pyrolysis) gases in the near-flame  
60 environment is desirable from both a scientific and safety perspective.

61 The major gas-phase compounds emitted from wildland fires are H<sub>2</sub>O, CO<sub>2</sub>, CO and CH<sub>4</sub> (Ward  
62 et al., 1991), all of which are easily identified and quantified via FTIR spectroscopy. Lightweight  
63 hydrocarbons, oxygenated hydrocarbons, nitrogen and sulfur species are all minor products that  
64 are also generated during burns (Talbot et al., 1988; Lobert et al., 1991; Yokelson et al., 1996). A  
65 host of more complex gases which can condense to form tar are also produced by pyrolysis of  
66 wildland fuels (Safdari et al., 2018; Amini et al., 2019). In a gas-phase IR spectrum of such  
67 species, however, the peaks associated with the minor products are often obfuscated by the more  
68 prominent features, such as those from CO<sub>2</sub>, and can only be recognized in the residual of a  
69 multicomponent simulated fit once the larger features have been removed. Using data from a  
70 recent field campaign to measure pyrolysis products carried out in a pine forest at Fort Jackson,  
71 South Carolina, we have analyzed some of the IR spectra in more detail to search for the signatures



72 of compounds not found in Table 1. As a partial guide of species for which to investigate, we  
 73 searched for those species detected in previous thermogravimetric-FTIR (TG-FTIR) studies  
 74 (Bassilakis et al., 2001; Taghizadeh et al., 2015). TG-FTIR experiments, however, are typically  
 75 small-scale and carried out in controlled environments (in contrast to ambient conditions of  
 76 prescribed burns or large-scale laboratory burns) and thus represent burns with different oxidative  
 77 capacities / combustion efficiencies (Yokelson et al., 1996; Fang et al., 2006; Akagi et al., 2014).  
 78 In this study, we have chosen to examine field fire spectra for species that can be detected and  
 79 quantified via IR spectroscopy both to add to the list of compounds, but also to improve the  
 80 characterization (and ultimately the detection limits) of the other species listed in Table 1. That is  
 81 to say, fire IR spectra are very complex and contain many overlapping peaks; the success of the  
 82 spectral analysis depends both on the selected spectral region and proper analysis of all compounds  
 83 included in the fit to that domain. The chemometric results become more reliable as the signatures  
 84 of all relevant species are included in the fit.

85 **Table 1.** Compounds previously detected in biomass burning studies using FTIR methods (Yokelson et  
 86 al., 1996; Yokelson et al., 1997; Goode et al., 1999; Goode et al., 2000; Christian et al., 2003; Christian et  
 87 al., 2004; Karl et al., 2007; Yokelson et al., 2009; Alves et al., 2010; Burling et al., 2010; Burling et al.,  
 88 2011; Akagi et al., 2013; Akagi et al., 2014; Stockwell et al., 2014; Gilman et al., 2015; Hatch et al.,  
 89 2017; Selimovic et al., 2018).

Compounds				
CO	NO	methanol	phenol	HCOOH
CO <sub>2</sub>	NO <sub>2</sub>	acetic acid	furaldehyde	peroxyacetyl nitrate**
CH <sub>4</sub>	HONO	SO <sub>2</sub>	hydroxyacetone	limonene
C <sub>2</sub> H <sub>2</sub>	NH <sub>3</sub>	furan	1,3-butadiene	carbonyls as glyoxal
C <sub>2</sub> H <sub>4</sub>	HCN	H <sub>2</sub> O	acetone	HCHO
C <sub>2</sub> H <sub>6</sub>	HCl	N <sub>2</sub> O	isoprene	2-methylfuran*
C <sub>3</sub> H <sub>6</sub>	O <sub>3</sub> **	OCS	glycolaldehyde	MVE (methyl vinyl ether)
C <sub>4</sub> H <sub>8</sub>				

90 \* used in the fit, but not analyzed, \*\* secondary components found downwind

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## 92 **2. EXPERIMENTAL**

### 93 **2.1 Site description and sampling device**

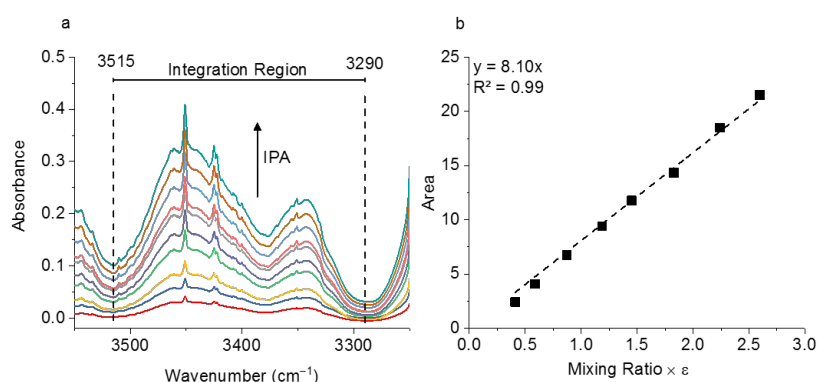
94 In early May 2018 seven prescribed fires were conducted in pine forests at U.S. Army Garrison  
95 Fort Jackson, adjacent to Columbia, South Carolina, at sites not far from previous smoke emission  
96 studies (Akagi et al., 2013; Weise et al., 2015). The forest overstory was primarily longleaf pine  
97 (*Pinus palustris* Mill.) and slash pine (*Pinus elliottii* Engelm.), while sparkleberry (*Vaccinium*  
98 *arboreum* Marshall) dominated the understory vegetation. During each burn, pyrolyzed gases  
99 emitted at the base of the flames before ignition were collected using an extractive probe and stored  
100 in 3-liter Summa canisters. This approach was performed to selectively collect pyrolysis gases  
101 prior to the onset of combustion. Details regarding the site description and sampling apparatus will  
102 be provided in a separate paper.

### 103 **2.2 FTIR Spectrometer**

104 Gases were analyzed in the laboratory (on the same day or the day following the fire) using an 8-  
105 meter multipass (White) cell (Bruker Optics, A136/2-L) mounted in the sample compartment of a  
106 Bruker Tensor 37 FTIR. Ten canisters were returned from the field to the laboratory and in turn  
107 connected to the gas cell via 3/8" stainless steel tubing. The tubing and gas cell were both heated  
108 to 70°C to prevent analyte adhesion to the inner surfaces. The White cell (White, 1942) was  
109 equipped with a pressure gauge and temperature probe, both of which were located on the gas  
110 outlet port; the thermocouple wire temperature probe extended into the White cell volume in order  
111 to more accurately measure the gas temperature. Prior to the start of the series of experiments, it  
112 was necessary to calibrate the path length of the variable path gas cell. Measurements of pure  
113 isopropyl alcohol (IPA, Sigma-Aldrich, 99.5%) at ten different pressures were collected and a  
114 Beer's Law plot was created to determine the length. The IR region from 3515 to 3290 cm<sup>-1</sup> was



115 integrated (Figure 1a), and the corresponding areas plotted as a function of the IPA pressure  
116 (converted to ppm at 760 Torr) multiplied by the PNNL reference library (Sharpe et al., 2004)  
117 integration area for a 1 ppm-meter IPA burden (Figure 1b). The slope is equal to the path length,  
118 which was determined to be 8.10 m.



119

120 **Figure 1.** a) Multiple burden spectra of dry IPA for 10 measurements at varying pressures. The dashed  
121 lines represent the integration limits used for spectral integration. b) Calibration plot with regression line  
122 for IPA measurements. The slope of the line is the path length in meters.

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124 The White cell contained the analyte smoke for the sample spectrum measurement, but was filled  
125 with ultra-high purity nitrogen gas for the reference spectrum measurement (Johnson et al., 2013).

126 The FTIR interferometer, detector and sample compartments were purged with dry air from a dry-  
127 air generator. The Tensor 37 was equipped with a globar source, a KBr beamsplitter and a  
128 broadband liquid nitrogen cooled mercury cadmium telluride (MCT) detector, providing spectral  
129 coverage from 7,500 to 500 cm<sup>-1</sup>. The spectral resolution was 0.6 cm<sup>-1</sup> and a 2 mm Jacquinot  
130 aperture was used. The acquisition mode was set to double-sided, forward-backward. For the  
131 Fourier transform, the data were apodized with a Blackman-Harris 3-Term function using a zerofill  
132 factor of 4, and phase corrected via the Mertz (Mertz, 1967) method.

133



### 134 **2.3 Quantitative Spectral Analysis**

135 The program used for quantitative spectral analysis was MALT5 (Griffith, 2016), which uses both  
136 broadband reference spectra from PNNL (Sharpe et al., 2004; Johnson et al., 2006; Johnson et al.,  
137 2009; Profeta et al., 2011; Lindenmaier et al., 2017) and absorption line intensities from HITRAN  
138 (Gordon et al., 2017) (in units of  $\text{cm}^{-1}/(\text{molec} \times \text{cm}^{-2})$ ) to iteratively fit a simulated spectrum to  
139 the measured spectrum by optimizing the fit so as to minimize the mean-squared residual, i.e. the  
140 difference between the measured and simulated spectra. Parameters such as path length, resolution,  
141 apodization, temperature, pressure, spectral domain, target compounds / overlapping compounds  
142 are all used as inputs to the spectral fit. During the course of this study, MALT5 was used to  
143 identify five gas-phase species emitted during the burns and quantify the gas mixing ratios via IR  
144 spectroscopy for the first time. Part of the confirmation strategy is to process the experimental  
145 spectra both with and without the target compound present in the fit and visually inspect the  
146 corresponding residuals. Table 2 summarizes the IR-active vibrational mode used for each species  
147 in the spectral fit (typically the species' strongest band in the longwave infrared window), along  
148 with the spectral domain and a list of species with overlapping bands in that domain.

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155 **Table 2.** Gas-phase species identified via FTIR, vibrational assignments (Lord et al., 1952; Hollenstein et  
 156 al., 1971; Ghosh et al., 1981; Hamada et al., 1985; Es-Sebbar et al., 2014; Chakraborty et al., 2016), and  
 157 spectral domains used for spectral fit and quantitation.

Target compound	Vibrational bands used for analysis	Spectral region (cm <sup>-1</sup> )	Other species fit in the same region
Naphthalene	$\nu_{46}$ at 782.3 cm <sup>-1</sup>	800–760	C <sub>2</sub> H <sub>2</sub> , CO <sub>2</sub> , HCN and H <sub>2</sub> O
Methyl nitrite	$\nu_8$ at 841.1 ( <i>cis</i> ) and 812.3 ( <i>trans</i> ) cm <sup>-1</sup>	865–775	C <sub>2</sub> H <sub>2</sub> , CO <sub>2</sub> , HCN, naphthalene, C <sub>2</sub> H <sub>4</sub> , allene, and H <sub>2</sub> O
Allene	$\nu_{10}$ at 845.3 cm <sup>-1</sup>	865–775	C <sub>2</sub> H <sub>2</sub> , CO <sub>2</sub> , HCN, naphthalene, C <sub>2</sub> H <sub>4</sub> , methyl nitrite, and H <sub>2</sub> O
Acrolein	$\nu_{10}$ at 1157.7 cm <sup>-1</sup>	1200–1100	Acetic acid (CH <sub>3</sub> COOH), furfural (C <sub>4</sub> H <sub>3</sub> OCHO), acetaldehyde, HCOOH, CH <sub>4</sub> , C <sub>2</sub> H <sub>4</sub> , and H <sub>2</sub> O
Acetaldehyde	$\nu_3$ at 2716.2 cm <sup>-1</sup>	2800–2650	CH <sub>4</sub> , HCHO, C <sub>2</sub> H <sub>2</sub> , acrolein, and H <sub>2</sub> O

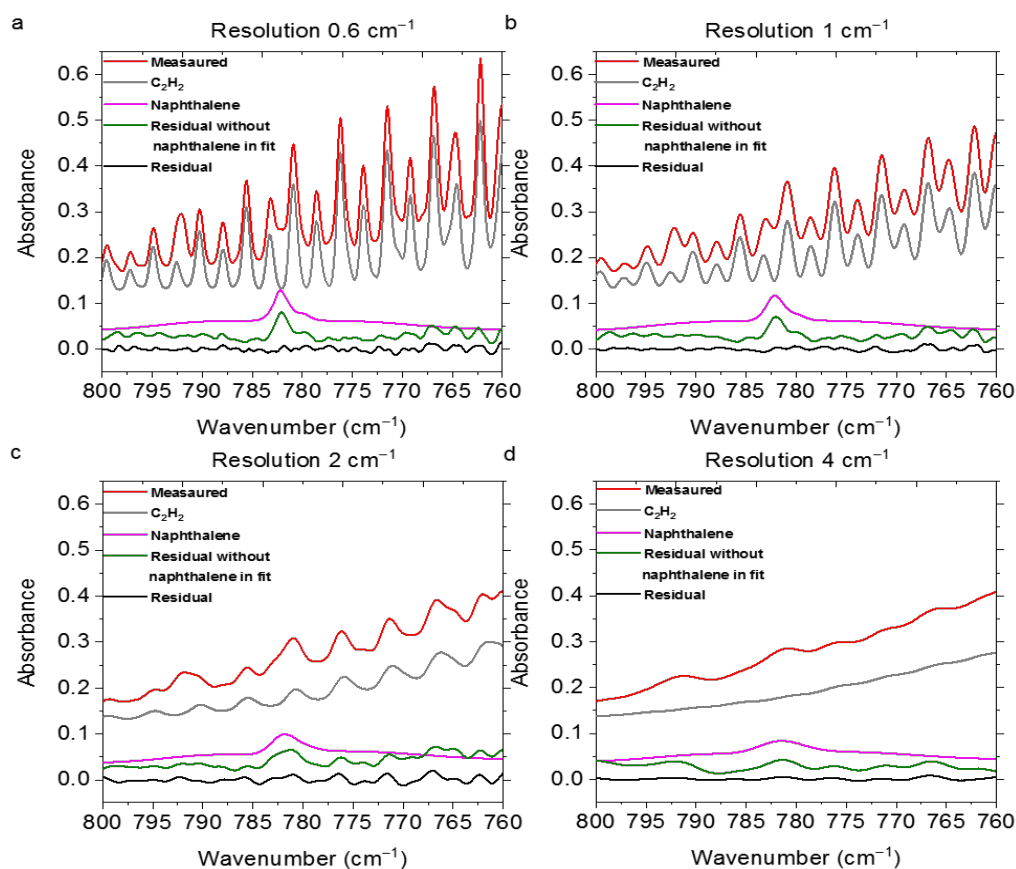
## 158 2.4 Spectral Resolution

159 As mentioned in section 2.2, the spectral resolution was set to 0.6 cm<sup>-1</sup>, which is the highest  
 160 resolution obtainable with this instrument. There are many benefits, but also a few disadvantages  
 161 to using higher resolution (Herget et al., 1979). Most importantly, the higher resolution allows one  
 162 to resolve the narrow bands of key analytes and discriminate them from lines or bands of  
 163 interferents. For example, in the present study the 782 cm<sup>-1</sup> Q-branch of naphthalene was  
 164 distinguished from the adjacent absorption lines of C<sub>2</sub>H<sub>2</sub> [Naphthalene's IR bands and results are  
 165 discussed in greater depth in Section 3.1]. If a lower resolution were used, then the deconvolution  
 166 of naphthalene from C<sub>2</sub>H<sub>2</sub> would have been compromised, perhaps unfeasible. To demonstrate,  
 167 one of the experimental measurements collected at a resolution of 0.6 cm<sup>-1</sup> was deresolved to 1,  
 168 2, and 4 cm<sup>-1</sup> using a Gaussian profile as seen in Figure 2. Those spectra were processed by  
 169 MALT5 to check for the presence of naphthalene. Figure 2 displays the measured spectra and the  
 170 scaled reference spectra for C<sub>2</sub>H<sub>2</sub> and naphthalene, and the corresponding residuals with and  
 171 without naphthalene included in the fit for the a) original spectrum collected at 0.6 cm<sup>-1</sup> and the





172 deresolved spectra at b)  $1\text{ cm}^{-1}$ , c)  $2\text{ cm}^{-1}$ , and d)  $4\text{ cm}^{-1}$ . With the reference spectra for the original  
173  $0.6\text{ cm}^{-1}$  measurement and the  $1\text{ cm}^{-1}$  deresolved spectrum (Figure 2a and b), the absorption lines  
174 for  $\text{C}_2\text{H}_2$  and naphthalene overlap, but the  $782\text{ cm}^{-1}$  feature from naphthalene is still slightly visible  
175 in the original spectra. The naphthalene peak appears clearly in the residuals when it is not included  
176 in the fitting process, but is removed from the residual when naphthalene is included in the fit  
177 (discussed further below). As the resolution is reduced (Figures 2c and 2d), however, the features  
178 broaden and the distinction of the naphthalene peak from  $\text{C}_2\text{H}_2$  and other minor components (i.e.  
179  $\text{CO}_2$ ,  $\text{HCN}$ ,  $\text{H}_2\text{O}$ , spectra not shown) is compromised. The specificity between the compounds is  
180 lost and the confidence in the identification/quantification of the target species, particularly for the  
181 weaker absorbers, diminishes as the resolution decreases. The well-known benefits of using a  
182 lower resolution are that spectra can be acquired more quickly with an improved signal-to-noise  
183 ratio. For the present measurements,  $0.6\text{ cm}^{-1}$  was deemed an appropriate resolution.



184

185 **Figure 2.** Measured and scaled reference spectra for  $C_2H_2$  and naphthalene, and corresponding residuals  
186 with and without naphthalene included in the fit for the a) original spectrum collected at  $0.6\text{ cm}^{-1}$  and the  
187 deresolved spectra at b)  $1\text{ cm}^{-1}$ , c)  $2\text{ cm}^{-1}$ , and d)  $4\text{ cm}^{-1}$ . The reference spectra for  $CO_2$ ,  $HCN$  and  $H_2O$  are  
188 not shown ( $HCN$  was not included in fit when the resolution was  $4\text{ cm}^{-1}$ ; for resolutions 1, 2 and  $4\text{ cm}^{-1}$ ,  
189  $H_2O$  was not included in the fit when naphthalene was removed from the fit). Spectra are offset for clarity.

## 190 2.5 Detection Limits and Signal-to-noise Ratio

191 The detection limit values presented in this paper are not minimal signal-to-noise limits in the  
192 sense of a minimal spectral signal against a background of purely stochastic noise sources. In such  
193 cases, the noise sources are typically of comparable or higher frequencies than the signal (Johnson  
194 et al., 1991). Rather, the current limits represent the average detection limits for a spectral residual  
195 derived from a convoluted spectrum arising from a gas mixture of differing and fluctuating  
196 chemical composition. The residuals are due to the least-squares fit of (fluctuations in) the many



197 complex features arising from numerous chemicals. That is to say, the residual is not due to just  
198 random instrumental noise, but instead, due to spectral features that can arise in the spectra, e.g.  
199 imperfectly subtracted features from strong absorbers or unidentified absorbers. For that reason,  
200 we report signal-to-residual, not signal-to-noise detection limits. The detection limits for each  
201 compound in this study were thus derived using a value of three times the root-mean-square (RMS)  
202 value of the residual calculated over the corresponding frequency range (e.g. 800–760  $\text{cm}^{-1}$  was  
203 used for naphthalene). The peak-to-peak noise is more sensitive to fluctuations in the fit with levels  
204 typically 4 to 5× the RMS noise (Griffith et al., 2006). For the present data, however, the peak-to-  
205 peak values ranged from 5 to 10× the RMS noise, thus suggesting the peak-peak values tend to  
206 overstate the tractable noise level, i.e. understate the detection limit. The reported detection limits  
207 are thus presumably higher than what would be estimated with an FTIR in clean air conditions (i.e.  
208 only the analyte and dry air). Based on experience, the limits are typically far higher than what can  
209 be obtained with IR laser sensors where the intrinsically narrow laser linewidths allow for the  
210 probing of individual rotational-vibrational lines without drawing in overlapping spectral lines to  
211 a congested spectral fit (Taubman et al., 2004; Wagner et al., 2011; Phillips et al., 2014). While  
212 typically far more sensitive, such laser measurements can only analyze for one or a few species at  
213 a time, as opposed to the 30+ species seen by the broadband FTIR measurements.

### 214 **3. RESULTS AND DISCUSSION**

215 When modeling the burning process (Byram, 1959), complete combustion of 1 kg dry wood  
216 produces 1.82 kg  $\text{CO}_2$  and 0.32 kg  $\text{H}_2\text{O}$  for a total mass of products of 2.14 kg. Incomplete  
217 combustion will yield additional products and less  $\text{CO}_2$  and  $\text{H}_2\text{O}$  while combustion of wet fuels  
218 (Byram, 1959) increases the amount of  $\text{H}_2\text{O}$  released. For infrared analysis of such smoke, much  
219 of the challenge arises due not only to the large mole fractions of  $\text{H}_2\text{O}$  and  $\text{CO}_2$ , but the fact that



220 both H<sub>2</sub>O vapor and CO<sub>2</sub> have strong features in the mid-IR that can clutter the spectrum rendering  
221 certain spectral regions unusable. For burning and other atmospheric studies, ideal compounds for  
222 detection via IR spectroscopy will thus have strong absorption coefficients that do not overlap with  
223 the fundamental bands of H<sub>2</sub>O or CO<sub>2</sub>, i.e. are in a spectral window or microwindow (Griffith,  
224 1996; Esler et al., 2000; Smith et al., 2011) free of strong interferences. Here, we consider five  
225 such compounds emitted during this prescribed burn, but which had heretofore not been reported  
226 as being detected by FTIR. Individual compounds are discussed in turn regarding their formation  
227 mechanism(s), detectable IR features and spectral confirmation for this study, along with their  
228 potential fates and atmospheric impacts. Lastly, the results are briefly compared with literature  
229 values using emission ratios (mixing ratios of analyte to excess CO).

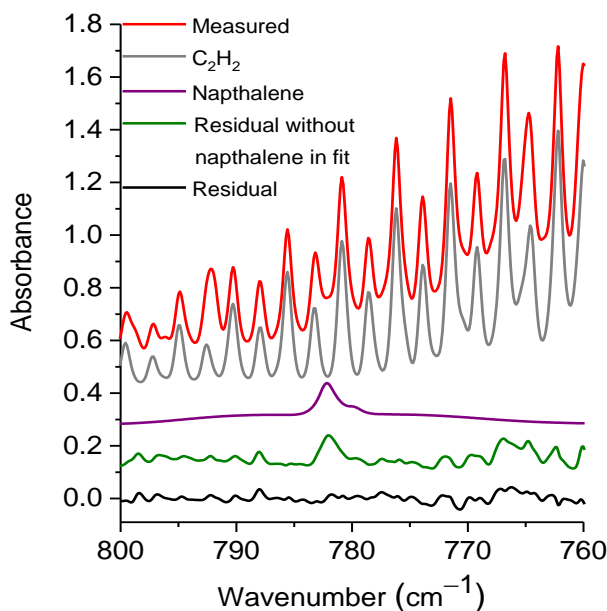
### 230 **3.1 Naphthalene**

231 Naphthalene (C<sub>10</sub>H<sub>8</sub>) is a polycyclic aromatic hydrocarbon (PAH) that is emitted from certain  
232 chemical industries as well as from the combustion of gasoline and oil (Jia et al., 2010). It is a  
233 condensable hydrocarbon also generated by biomass pyrolysis (Liu et al., 2017). There are a  
234 number of pyrolysis formation routes (Fairburn et al., 1990; Williams et al., 1999; Richter et al.,  
235 2000; Lu et al., 2004; Liu et al., 2017). One proposed mechanism is the generation of single ring  
236 aromatic compounds such as benzene, toluene and styrene via Diels-Alder reaction of alkenes; the  
237 single ring aromatic compound then combines with alkenes to form double-ring PAHs, such as  
238 naphthalene (Fairburn et al., 1990). Naphthalene may even undergo subsequent reactions to form  
239 still larger polyaromatics (Fairburn et al., 1990; Richter et al., 2000). Naphthalene has been  
240 detected (via GC-MS) in tars that were condensed from gas-phase pyrolysis products of both live  
241 and dead southeastern fuels, such as live oak (*Quercus virginiana*) and swamp bay (*Persea*  
242 *palustris*) (Safdari et al., 2018). It has been also detected (Hosseini et al., 2014; Aurell et al., 2017;



243 Koss et al., 2018) in the gas-phase in laboratory burning experiments. The detection of gas-phase  
244 naphthalene from wildland fire emissions is thus not surprising, but this is the first report of its  
245 identification via IR spectroscopy. The best spectral feature for identification and quantification is  
246 the  $\nu_{46}$  IR mode near  $782.3\text{ cm}^{-1}$ , which corresponds to the H–C–C out-of-plane bend (Chakraborty  
247 et al., 2016). There are other bands at  $3067.7$  and  $3058.0\text{ cm}^{-1}$  previously assigned to  $\nu_{29}$  and  $\nu_{17}$ ,  
248 respectively (Chakraborty et al., 2016). Both of these modes have smaller absorption coefficients  
249 as compared to  $\nu_{46}$ , however, and are located in the C–H stretching region, which is common to  
250 nearly all hydrocarbons and thus provides less specificity.

251 Figure 3 shows a prescribed burn spectrum in the region from  $800$  to  $760\text{ cm}^{-1}$ . The primary  
252 spectral signatures in this plot are those of the R-branch rotational-vibrational lines associated with  
253 the  $\nu_5$  fundamental (Kabbadj et al., 1991) of  $\text{C}_2\text{H}_2$ , but there are also absorptions due to  $\text{CO}_2$ , HCN,  
254  $\text{H}_2\text{O}$  (individual spectral contributions not shown) and naphthalene. When all of the spectral  
255 components except for naphthalene are included in the fitting process, the residual (green trace)  
256 displays a prominent feature at  $782.3\text{ cm}^{-1}$ , which we ascribe to naphthalene. When naphthalene  
257 is included in the fit, the feature in question is removed as seen in the black trace of Figure 3.



258

259 **Figure 3.** Measured spectrum, scaled reference spectra for C<sub>2</sub>H<sub>2</sub> and naphthalene, and residuals  
260 with and without naphthalene included in the fit. For clarity, the spectral contributions for CO<sub>2</sub>,  
261 HCN, and H<sub>2</sub>O are not shown. All spectra are at 0.6 cm<sup>-1</sup> resolution and have been offset. The  
262 calculated mixing ratio of naphthalene in this measured spectrum is 16.4 ± 0.6 ppm (values obtained  
263 from MALT5 software, and error represents standard error).

264

265

266 Table 3 presents the range of measured mixing ratios for naphthalene along with averaged  
267 detection limits for the 10 measurements collected during the prescribed burns as well as for the  
268 other four reported compounds. In the measurements, naphthalene's mixing ratios ranged from 1.4  
269 to 19.9 ppm, and the averaged RMS-derived detection limit was 1.6 ± 0.5 ppm; different detection  
270 limits were observed for each spectrum. One of the measurements had a mixing ratio of 2.9 ppm,  
271 yet its RMS-derived detection limit was 3.7 ppm, and is thus below the estimated detection limit  
272 (bdl).

273

274



275 **Table 3.** Calculated mixing ratios for ten canister FTIR measurements along with average estimated  
 276 residual detection limits for the target compounds derived using 3 times the root-mean-square of the  
 277 residual. Error bars represent the standard deviation ( $1\sigma$ ) of the mean.

Target compound	Calculated mixing ratios (ppm)			Averaged detection limit (ppm) using root-mean-square (RMS) value of the residual
	Min	Max	Average	
Naphthalene*	1.4	19.9	$8.5 \pm 2.1$	$1.9 \pm 0.5$
Methyl nitrite*	2.3	21.0	$8.7 \pm 2.4$	$2.2 \pm 0.4$
Allene	2.2	37.8	$13.1 \pm 3.6$	$3.0 \pm 0.6$
Acrolein	14.7	125.7	$43 \pm 12$	$6.1 \pm 1.5$
Acetaldehyde	34.5	264.8	$103 \pm 27$	$11.7 \pm 3.2$

278 \*One measurement was below the detection limit.

279

280 Naphthalene emitted from prescribed burns is thus clearly detectable using IR spectroscopy. The  
 281 U.S. Environmental Protection Agency considers naphthalene a potential human carcinogen and a  
 282 hazardous air pollutant (U.S. EPA). Once released, naphthalene may cycle in the atmosphere or  
 283 accumulate in aquatic and terrestrial systems via wet/dry deposition or air-water gas exchange  
 284 (Park et al., 2001). Gas-phase naphthalene's primary atmospheric loss mechanism is its reaction  
 285 with the hydroxyl radical (OH) to form hydroxy-PAHs or nitro-PAHs in the presences of nitrogen  
 286 oxides (Vione et al., 2004). The estimated atmospheric lifetime of naphthalene for reaction with  
 287 OH is 6.8 hours (based on a 12-hour daytime OH level of  $1.9 \times 10^6$  molecules  $\text{cm}^{-3}$ ) (Arey, 1998).

### 288 3.2 Methyl Nitrite

289 A second compound detected for the first time in the wildland fire IR spectra was methyl nitrite  
 290 ( $\text{CH}_3\text{ON}=\text{O}$ ). Methyl nitrite has previously been observed in aged cigarette smoke (Schmeltz et  
 291 al., 1977), and also the exhaust of engines fueled by methanol–diesel blends (Jonsson et al., 1982).  
 292 It has also been observed as a minor product for the thermal decomposition of both nitrate esters  
 293 (Boschan et al., 1955) and isopropyl nitrate at low temperatures and pressures (Griffiths et al.,  
 294 1975). Methyl nitrite has also been detected in wildland fire emissions by GC-MS (Gilman et al.,



295 2015). Moreover, it has been observed that some nitrogen-containing organic compounds such as  
296 acetonitrile ( $\text{CH}_3\text{CN}$ ) and acrylonitrile ( $\text{CH}_2=\text{CHCN}$ ) emitted from burns were directly correlated  
297 to the fuel nitrogen content. However, methyl nitrite [and another oxygenated nitrogen organic  
298 compound, isocyanic acid ( $\text{HNCO}$ )] did not show any significant dependency on fuel N-content  
299 (Coggon et al., 2016). It has been suggested that methyl nitrite is only a minor direct product of  
300 combustion (Finlayson-Pitts et al., 1992), but instead is generated *in situ* by the secondary reaction  
301 of methanol ( $\text{CH}_3\text{OH}$ ) with nitrogen dioxide ( $\text{NO}_2$ ).

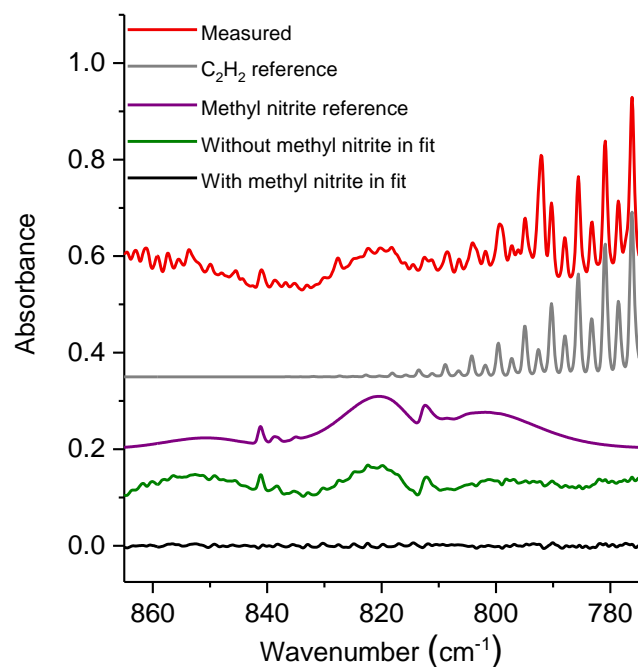
302 We also note that methyl nitrite is an oxidizing agent and is used as a rocket propellant. It is thus  
303 plausible that the methyl nitrite detected in the present study was not a product of the fire, but  
304 emanated from munitions used in training at Ft. Jackson. However, while the records of the  
305 munitions used were not complete, a survey of these records did not indicate the use of methyl  
306 nitrite in any munitions at the Ft. Jackson plots where the present burn samples were collected.

307 With regards to the IR spectra, methyl nitrite exists in equilibrium as a mixture of two conformers-  
308 *cis* and *trans*; at room temperature ( $25^\circ\text{C}$ ) it is estimated as 58% *cis* and 42% *trans* (Bodenbinder  
309 et al., 1994). We were able to use the same band associated with both conformers, namely the  $\nu_8$   
310 band, which is at  $841.1\text{ cm}^{-1}$  for the *cis* conformer and at  $812.4\text{ cm}^{-1}$  for the *trans* conformer  
311 (Ghosh et al., 1981). The  $\nu_8$  mode is associated with the N–O stretch and is very strong for both  
312 conformers (Ghosh et al., 1981). We note that methyl nitrite also has very strong bands at  $627.8$   
313  $\text{cm}^{-1}$  (*cis*) for  $\nu_9$  ONO bending, as well as at  $1620.1\text{ cm}^{-1}$  (*cis*) and  $1677.4\text{ cm}^{-1}$  (*trans*) due to the  
314  $\nu_3$  N=O stretch (Ghosh et al., 1981). These bands, however, are of lesser utility for IR detection:  
315 The  $\nu_9$  peak is masked by  $\text{CO}_2$  bending mode lines, and the  $\nu_3$  peak is obfuscated by the  $\text{H}_2\text{O}$   
316 bending mode lines.





317 The spectral region used for evaluation was  $865\text{--}775\text{ cm}^{-1}$ , which contains the  $\nu_8$  band for both the  
318 *cis* and *trans* conformers (Ghosh et al., 1981). Figure 4 shows the experimental spectrum from  
319 the prescribed burn, along with the scaled reference spectra for the two major compounds used in  
320 the fit:  $\text{C}_2\text{H}_2$  and methyl nitrite. While important, other minor compounds, such as  $\text{CO}_2$ , HCN,  
321 naphthalene,  $\text{C}_2\text{H}_4$ , allene, and  $\text{H}_2\text{O}$ , were also included in the analysis, but their spectral  
322 contributions are not plotted. Additionally, Figure 4 displays the residuals both when methyl nitrite  
323 was included in the fitting process and when it was excluded. Upon inspection of the residual  
324 spectrum where it was excluded (green trace), it is clear that both the *cis* and *trans* features from  
325  $\nu_8$  are present and this confirms methyl nitrite in the pyrolysis smoke



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327

328 **Figure 4.** Measured experimental spectrum and the individual spectral contributions for the major  
329 components ( $\text{C}_2\text{H}_2$  and methyl nitrite) and residuals with and without methyl nitrite included in the fit. For  
330 clarity, the spectral contributions for  $\text{CO}_2$ , HCN, naphthalene,  $\text{C}_2\text{H}_4$ , allene, and  $\text{H}_2\text{O}$  are not shown. All  
331 spectra are at  $0.6\text{ cm}^{-1}$  resolution and have been offset for clarity. The calculated mixing ratio of methyl  
332 nitrite in this measured spectrum is  $21.0 \pm 0.1\text{ ppm}$  (values obtained from MALT5 software, and error  
333 represents standard error).



334 The mixing ratio and RMS-derived detection limit for methyl nitrite for the displayed experimental  
335 spectrum in Figure 4 are 21.0 ppm and 1.4 ppm, respectively. The range for the mixing ratios and  
336 the averaged detection limits for methyl nitrite are summarized in Table 3. Methyl nitrite was  
337 detected with confidence in 9 of the 10 measurements; only one of the measurements was below  
338 the RMS-derived detection limit.

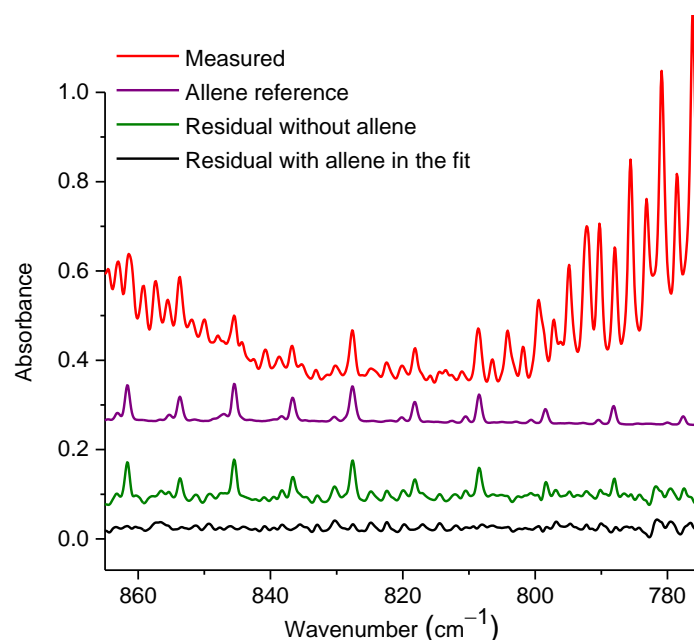
339 We report the detection via IR spectroscopy of methyl nitrite in wildland fire emissions not only  
340 because it is novel, but also because of its influential role in atmospheric chemistry: Methyl nitrite  
341 is a photochemical source of OH. In the atmosphere it undergoes photolysis to form the methoxy  
342 radical ( $\text{CH}_3\text{O}$ ) and nitric oxide (NO) with a quantum yield near unity (Cox et al., 1980). At solar  
343 noon, the photolytic lifetime is only 10–15 min (Seinfeld et al., 2012). The photogenerated  
344 methoxy radical then undergoes subsequent reactions leading to the formation of OH. In turn, both  
345 OH and NO contribute to the production of ozone (Finlayson-Pitts et al., 1999).

### 346 **3.3 Allene**

347 Allene (1,2-propadiene,  $\text{CH}_2=\text{C}=\text{CH}_2$ ) is of high symmetry ( $\text{D}_{2d}$ ) and has the two methylene  
348 groups with their H–C–H planes at right angles to each other (Lord et al., 1952). The compound  
349 has previously been detected in biomass burning grab samples using GC (Akagi et al., 2013).  
350 Allene is a proposed precursor in the burning process that contributes to the formation of both  
351 aromatic compounds and soot (Frenklach et al., 1983; Frenklach et al., 1988). Lifshitz et al. have  
352 observed (at temperatures ranging from 757–847°C) that the structural isomerization of allene and  
353 propyne ( $\text{CH}_2=\text{C}=\text{CH}_2 \leftrightarrow \text{CH}_3-\text{C}\equiv\text{CH}$ ) will take place via a unimolecular reaction faster than the  
354 decomposition reaction (Lifshitz et al., 1975). Additionally, these same authors investigated the  
355 pyrolysis of allene and propyne and observed that  $\text{C}_2\text{H}_4$  was generated from allene while  $\text{CH}_4$  and  
356  $\text{C}_2\text{H}_2$  were mainly formed from propyne (Lifshitz et al., 1976). Unfortunately, the strongest IR



357 band for propyne (near  $634\text{ cm}^{-1}$ ) is obscured by  $\text{CO}_2$  bending mode lines. Due to the interferences  
358 we cannot with confidence identify propyne in the measurements; we can, however, detect allene.  
359 In the mid-IR, allene has several strong rotational-vibrational lines near  $845\text{ cm}^{-1}$  associated with  
360 the sub-bands of the perpendicular band  $\nu_{10}$ , which is due to  $\text{CH}_2$  rocking (Lord et al., 1952).  
361 Additionally, allene has a moderately strong band at  $1958.6\text{ cm}^{-1}$  due to the  $\nu_6$  C–C stretching  
362 (Lord et al., 1952). However, the  $\nu_6$  band is not useful for detection due to interference from the  
363  $\text{H}_2\text{O}$  bending mode lines.



364 **Figure 5.** Measured absorbance spectrum and residual with and without allene included in the fit, along  
365 with the scaled reference spectrum for allene. For clarity, the spectral contributions for  $\text{C}_2\text{H}_2$ ,  $\text{CO}_2$ , HCN,  
366 naphthalene,  $\text{C}_2\text{H}_4$ , methyl nitrite, and  $\text{H}_2\text{O}$  are not shown. All spectra are at  $0.6\text{ cm}^{-1}$  resolution and have  
367 been offset for clarity. The calculated mixing ratio of allene in this measured spectrum is  $37.8 \pm 0.6\text{ ppm}$   
368 (values obtained from MALT5 software, and error represents standard error).  
369  
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371 Figure 5 shows the measured absorbance spectrum, scaled allene reference spectrum and the  
372 associated residual with and without allene included in the fit. The absorption lines associated  
373 with allene are clearly seen in the resulting spectrum when allene is not included in the fit (green



374 trace), thus confirming that allene is one of the primary components contributing to the features in  
375 this spectral domain. For the experimental spectrum displayed in Figure 5, the calculated mixing  
376 ratio for allene is 37.8 ppm and the RMS-derived detection limit is 5.4 ppm.

377

378 Unlike naphthalene and methyl nitrite, allene is not considered a hazardous air pollutant nor is it a  
379 photochemical source of OH. Major loss processes for alkenes include reactions with OH, NO<sub>3</sub>  
380 radical and O<sub>3</sub> (Atkinson et al., 2003). Specifically for allene, the lifetime (calculated from rate  
381 constants from Atkinson et al., 2003, and based on a 12-hour daytime OH level of  $1.9 \times 10^6$   
382 molecules cm<sup>-3</sup> and 24-hour O<sub>3</sub> average of  $7 \times 10^{11}$  molecules cm<sup>-3</sup>) with respect to OH and O<sub>3</sub>  
383 reactions are 1.2 and 89.4 days, respectively. The reaction between OH and allene involves the  
384 initial addition of OH to one of the C=C bonds generating a hydroxyalkyl radical, which then may  
385 undergo subsequent reactions (i.e. reaction with O<sub>2</sub> forming hydroxyalkyl peroxy radical)  
386 contributing to the propagation of radicals in the atmosphere (Atkinson et al., 2003; Daranlot et  
387 al., 2012).

388

### 389 **3.4 Acrolein and Acetaldehyde**

390 The two aldehydes, acrolein (CH<sub>2</sub>=CHCHO) and acetaldehyde (CH<sub>3</sub>CHO), have also been  
391 identified in the burning IR spectra. It has been proposed that both acrolein and acetaldehyde are  
392 formed from the pyrolysis of cellulose (a major constituent of biomass) via the intermediate  
393 glycerol, which is a moiety in the structure of levoglucosan, a known pyrolysis product of cellulose  
394 (Stein et al., 1983). Stein et al. observed that acrolein, acetaldehyde and CO were the initial  
395 decomposition products for the pyrolysis of glycerol (Stein et al., 1983). Both of these compounds

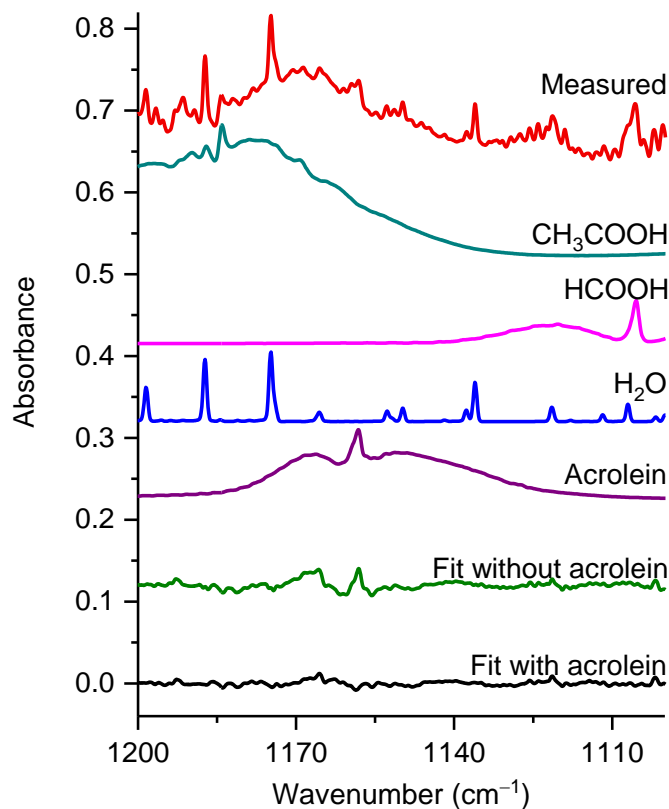


396 have been detected in previous wildland fires studies via methods such as GC (Akagi et al., 2013)  
397 or PTR-ToF (Brilli et al., 2014; Koss et al., 2018), but have not yet been identified via IR.

398 Acrolein, the simplest unsaturated carbonyl, exists in two forms, *s-cis* and *s-trans*, with *s-trans*  
399 being the more stable, and consequently the more abundant conformer (Wagner et al., 1957). It  
400 has been estimated that the fractions of *s-cis* and *s-trans* are about 4 and 96% at 20°C, and 7 and  
401 93% at 100°C, respectively (Alves et al., 1971). The largest IR feature for acrolein is the  $\nu_5$  C=O  
402 stretch (Hamada et al., 1985) at 1724.1  $\text{cm}^{-1}$ , but this band is heavily overlapped by water lines.  
403 There is also the  $\nu_{16}$  band (Hamada et al., 1985) at 958.8  $\text{cm}^{-1}$ , but this feature overlaps with  
404 multiple other strongly absorbing compounds, such as  $\text{C}_2\text{H}_4$ . We have therefore focused acrolein's  
405 analysis using the  $\nu_{10}$  band (C–C stretch) (Hamada et al., 1985) at 1157.7  $\text{cm}^{-1}$ .

406

407 Figure 6 displays the very congested biomass burning spectrum with individual contributions for  
408 several species included in the fit [contributions for furfural ( $\text{C}_4\text{H}_3\text{OCHO}$ ), acetaldehyde,  $\text{CH}_4$ , and  
409  $\text{C}_2\text{H}_4$  are included, but not plotted] as well as the residual with and without acrolein included in  
410 the fitting process. When acrolein is not included in the fit, features (both near 1168 and at 1157.7  
411  $\text{cm}^{-1}$ ) that resemble acrolein are observed in the residual spectrum as seen in the green trace in  
412 Figure 6. When acrolein is included in the fit, the features in question are removed. For acrolein,  
413 no mixing ratios were observed below the RMS-derived detection limits.



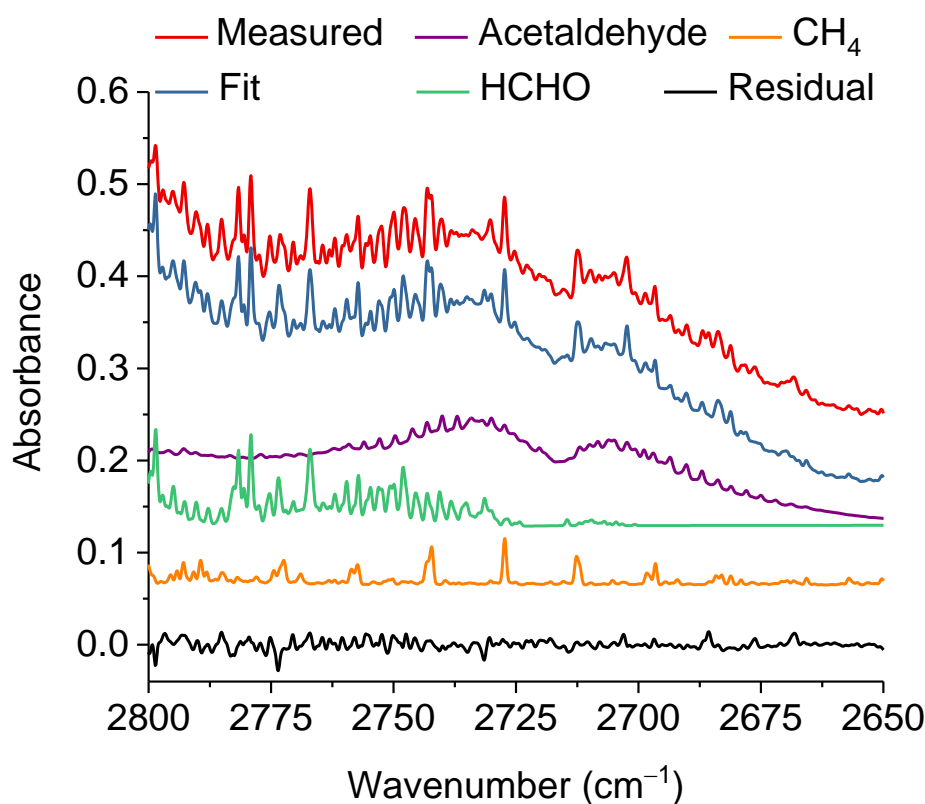
414 **Figure 6.** Measured spectrum and the individual spectral contributions for the major components and  
415 associated residual with and without acrolein included in the fit. For clarity, the spectral contributions for  
416 furfural ( $C_4H_3OCHO$ ), acetaldehyde,  $CH_4$ , and  $C_2H_4$  are not shown. All spectra are at  $0.6\text{ cm}^{-1}$   
417 resolution and have been offset for clarity. The calculated mixing ratio of acrolein in this measured spectrum  
418 is  $99.9 \pm 3.0\text{ ppm}$  (values obtained from MALT5 software, and error represents standard error).  
419

420  
421 Similar to acrolein, acetaldehyde has its strongest IR feature due to the C=O stretch (Hollenstein  
422 et al., 1971), with  $\nu_4$  found at  $1746.1\text{ cm}^{-1}$ . Again, due to the presence of water lines in the  
423 spectrum, this feature is not practical for detection. The aldehyde  $\nu_3$  C–H stretching band  
424 (Hollenstein et al., 1971) at  $2716.2\text{ cm}^{-1}$  was instead used for analysis. Figure 7 shows the  
425 measured and fitted spectra as well as the spectral contributions of the major individual  
426 components used to calculate the fitted spectrum and the corresponding residual. Other minor



427 components, such as acrolein,  $C_2H_2$  and  $H_2O$ , were also included in the fit, but their reference  
428 spectra are not displayed in Figure 7. The spectral profile of acetaldehyde with its P and R branches  
429 of  $\nu_3$  is easily discernable even before deconvolution of the measured spectrum. Similar to  
430 acrolein, all of the mixing ratios for acetaldehyde were above the RMS-derived detection limit.

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**Figure 7.** Measured and fitted spectra as well as the individual components (for clarity, the spectral contributions for acrolein,  $C_2H_2$ ,  $H_2O$  are not shown) and associated residual in the spectral region 2800-2650  $cm^{-1}$ . All spectra are at 0.6  $cm^{-1}$  resolution and have been offset for clarity. The calculated mixing ratio of acetaldehyde in this measured spectrum is  $252.8 \pm 5.5$  ppm (values obtained from MALT5 software, and error represents standard error).

Similar to naphthalene, the U.S. EPA considers both acrolein and acetaldehyde to be hazardous air pollutants (U.S. EPA). Acrolein is toxic to humans, and when inhaled may cause upper respiratory irritation. Acetaldehyde will irritate the eyes, skin and the respiratory tract and is



442 considered a potential human carcinogen (U.S. EPA). Once released into atmosphere, such  
443 aldehydes can either react with O<sub>3</sub> or OH, or undergo photolysis (Seinfeld et al., 2012). For  
444 acrolein, OH reaction is the major loss process with a lifetime of 2.4 hours based on a 12-hour  
445 daytime OH level of  $1.9 \times 10^6$  molecules cm<sup>-3</sup>, (Gierczak et al., 1997) forming products such as  
446 CO, CO<sub>2</sub>, HCHO, glycolaldehyde (Johnson et al., 2013) and acryloylperoxynitrate (APAN)  
447 (Orlando et al., 2002). Similarly, acetaldehyde's lifetime is dominated by OH loss, and that  
448 reaction generates HCHO and CO as well as peroxyacetyl nitrate (PAN) (D'Anna et al., 2003).  
449 Acetaldehyde's estimated tropospheric lifetimes with respect to OH reaction and photolysis are 10  
450 hours (Atkinson et al., 2003) and 5 days (Seinfeld et al., 2012), respectively.

### 451 **3.5 Comparison to Other Measurements**

452 Preliminary emission ratios (relative to CO) for the reported compounds are compared to those  
453 reported in previous wildland burning investigations: Table 4 displays the average emission ratios  
454 and the standard deviations ( $1\sigma$ ) for this study as well as emission ratios reported by Koss et al.  
455 (2018), Ferek et al. (1998), Brillì et al. (2014), and Gilman et al. (2015). As shown in the table,  
456 there is significant variation between the studies due to multiple factors such as different fuel types,  
457 analytical methods, sampling approaches and experimental conditions. For example, the study by  
458 Ferek et al. (1998) focused on the collection of airborne samples, while Brillì et al. (2014)  
459 measured gases under nocturnal conditions using a ground-based system. Inspection of the table  
460 shows that the measured emission ratio values are not unprecedented, but are within range of  
461 previous measurements. Because they have the same molar mass, the mass spectrometric  
462 techniques in some cases cannot distinguish allene from propyne.

463





464 **Table 4.** Emission ratios relative to CO and standard deviations ( $1\sigma$ ) for the present study and for  
 465 three other previously published biomass burning studies.

Target compounds	Present average emission ratios to CO (ppb/ppm)	Koss et al. (2018) fire-integrated emission ratio to CO (ppb/ppm)	Ferek et al. (1998) emission ratio to CO (ppb/ppm)	Brilli et al. (2014) emission ratios to CO (ppb/ppm)	Gilman et al. (2015) discrete emission ratios to CO (ppb/ppm)		
					South-western fuels	South-eastern fuels	Northern fuels
Method	FTIR	PTR-ToF-MS	GC-FID*	PTR-ToF-MS	GC-MS	GC-MS	GC-MS
Naphthalene	0.79 (0.47)	0.20 (0.16)	n/a	n/a	0.0070 (0.0048)	0.0040 (0.0050)	0.022 (0.012)
Methyl nitrite	0.94 (0.85)	n/a	n/a	n/a	0.9 (1.1)	0.52 (0.51)	0.76 (0.90)
Acrolein	4.0 (1.8)	5.4 (3.0)	n/a	3.14 (0.12)	0.82 (0.68)	1.31 (0.88)	3.5 (1.7)
Acetaldehyde	9.4 (3.6)	7.4 (5.2)	n/a	37.3 (1.4)	1.6 (1.2)	2.8 (1.8)	5.5 (3.6)
Allene (Propadiene)**	1.05 (0.24)	n/a	0.1 (0.1)	8.73 (0.28)	n/a	n/a	n/a

466 \*GC-FID is gas chromatography with flame ionization detector

467 \*\*Brilli et al. (2014) use both 1-propyne and propadiene to represent  $C_3H_4$ . Gilman et al. (2015) report emission  
 468 ratios for propyne, but not allene.

#### 470 4. SUMMARY

471 Gas-phase compounds with appreciable band intensities and appreciable concentrations can be  
 472 both identified and quantified using IR spectroscopy. We have used such spectral information for  
 473 seminal IR detection of five compounds generated during prescribed forest fire burns. Deriving  
 474 the mixing ratios from the congested spectra obtained from wildland smoke samples is more  
 475 challenging due to the multiple overlapping spectral features: Sophisticated software and analysis  
 476 are required in carefully selected spectral windows. We have reported seminal IR detection of five  
 477 molecules that had previously not been observed by FTIR in ambient measurements of wildland  
 478 emissions. Most of the compounds (excluding acetaldehyde), had their primary features become  
 479 apparent only after the larger spectral features had been fitted and subtracted.

480

481

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493

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