Response to Referee # 1

RC: Referee's comment
AR: Author's response

The referee is encouraged to refer our response to comments and concerns raised by the anonymous referee # 2, where we have incorporated additional analysis to the manuscript.

RC: The paper is very interesting because it is the first time a comparison study is performed using a large number of SKYNET sites and products from this network. In addition, the results are very important for the developers of Skyrad pack improvements, particularly for what concern the assumption of fixed value of Surface Albedo.
AR: We appreciate the reviewer for the constructive comments. In the OMI-SKYNET comparison, we have considered all SKYNET stations whose data are freely accessible online from Chiba University SKYNET server. As the referee as stated here, the purpose of such comparison, particularly when both quantities are not directly measured but retrieved using respective algorithms, is to understand and diagnose the (dis)agreement between the two datasets to improve the accuracy of both retrievals.

RC: Lines: 65-66: specify if the change of estimated radiative forcing refers to the top, bottom or middle atmosphere.
AR: The sentence is modified as "Together, both AOD and SSA determine the magnitude and sign of the aerosol radiative forcing at the top-of-atmosphere."

RC: Line 170: both the POMs models take also measurements at 315 and 940 nm for Ozone and water vapour retrieval. Add this information here and in line 178.
AR: The header information given in both POM-01 and POM-02 datasets states that the former sensor carries a total of five wavelength filters covering visible to near-IR (400-1020 nm), whereas the latter has two additional filters in the UV region (340 and 380 nm) along with the other filters in the visible to shortwave-IR (including 1627 nm and 2200 nm) part of the spectrum. These data files do not mention the use of 315 and 940 nm for Ozone and water vapor retrievals.

RC: Line 186: remove and between University and Valencia
AR: Corrected.

RC: Line 255: add "carbonaceous/smoke"
AR: Corrected.

RC: Lines 255-257: it not clear to me according to which parameter has been considered the 5 listed sites better than the others. Moreover looking at Figure 2 the largest percentage of agreement is for Q_0.05 and not 0.03.
AR: The sentence referred here is for the carbonaceous/smoke aerosol type (red dots in the scatterplots), for which the majority of matchups are confined within the difference of 0.03. For

the overall compassion between the sites, we considered RMSD and % matchups falling within 0.03/0.05 as criteria. Lower RSMD and higher % matchups (Q_0.03 & Q_0.05) suggest a better comparison between OMI and SKYNET SSAs.

RC: Line 348, it is better specify that the assumption of fixed ground albedo in Skyrad pack can be changed in time and wavelengths, if necessary. For example in ESR/SKYNET, Antarctica sites are processed with different values.

AR: The first two sentences are revised as “The standard SKYNET inversion algorithm assumes a wavelength-independent surface albedo of 0.1 at all wavelengths across the UV to visible part of the spectrum. However, the algorithm code allows flexibility to alter the value surface albedo in time and wavelength (Campanelli et al., 2015).”

RC: I still suggest using greater characters for Figure 2, 3, 4. It has been difficult reading the statistics values.

AR: Figure 2, 3, and 4 are reproduced with bigger size characters and numbers.

Response to Referee # 2

RC: Referee's comment
AR: Author's response

RC: The present study deals with the assessment of SSA derived by the OMI sensor against the corresponding retrievals obtained from ground SKYNET stations. Most of the issues addressed here have been already discussed in Jethva et al. (2014, JGR) and therefore the contribution of the current work is relatively small. An "innovative" aspect is the availability of coincident OMI-SKYNET UV wavelengths (in contrast to AERONET) thus allowing the direct intercomparison between spaceborne and ground-based measurements. Nevertheless, a significant portion of the ground retrievals is coming from POM01 instruments (longer wavelengths of 400nm) requiring the extrapolation of OMI SSA based on the assumed aerosol model in the retrieval algorithm. Below are listed my comments which should be clarified by the authors prior the publication of the submitted manuscript.

AR: While we agree with the reviewer that the present study follows our earlier paper of OMI-AERONET SSA comparison (Jethva et al., 2014), the present study is the first-ever attempt to evaluate the OMI SSA retrievals against those of SKYNET. The POM-02 sensors owing to its retrievals in the near-UV region facilitates direct comparison with those of OMI at 388 nm. The POM-01 sensors, on the other hand, do not carry UV wavelength filters, but offer retrievals at 400 nm requiring an extrapolation of OMI SSA to only 12-nm window against 52-nm when compared with AERONET at 440 nm. It is reasonably fair to assume that the extrapolation of OMI SSA in a narrow window of only 12-nm should not be a major source of uncertainty in comparing SSA from OMI and SKYNET. An additional analysis comparing OMI SSA with POM01 and POM02 dataset separately, which is discussed later in this response, suggests that the resultant OMI-SKYNET statistics do not change significantly regardless of the use of POM01 or POM02 in the evaluation.

In the revised manuscript, it is stated that,

"It is reasonably fair to assume that the extrapolation of OMI SSA in a narrow window of 12-nm, i.e., from 388 to 400 nm, shouldn't be a major source of uncertainty in comparing SSA from OMI and SKYNET."

RC: 1. Lines 225-227: The selection of degrees (variable with latitude) instead of distance has impact on the area which is averaged for the OMI retrievals. How much important is this for the overall results?
AR: We attempted several spatial windows for the OMI averaging, i.e., 0.1°, 0.25°, 0.5°, or even 1.0°. However, the statistical results in terms of RSMD and % matchups within 0.03/0.05 and derived conclusion didn't alter much. We choose a spatial window of 0.5° centered at SKYNET station to be consistent with our earlier work of comparing OMI and AERONET SSA.

RC: 2. Lines 235-237: How much reliable are the aerosol models considered in the OMI retrieval algorithm? This is critical since an inappropriate selection of an aerosol model can lead to unrealistic SSA retrievals and subsequently will have impact on the intercomparison outputs.

AR: The aerosol size distribution and spectral dependence of AOD assumed for all three standard aerosol types considered in the OMAERUV algorithm are based on the AERONET ground-based inversions reported in Dubovik et al. (2002). The spectral dependence of absorption, often quantified as Absorption Angstrom Exponent, for the carbonaceous smoke aerosols accounts for the presence of organics in biomass burning aerosols and referenced to the in-situ absorption measurements taken during SAFARI-2000 experiment (Kirchstetter et al., 2004; Jethva and Torres, 2011). In the latest OMAERUV product used in the present study, the urban-industrial aerosol model is also revised to adopt the same spectral dependence of absorption that of the smoke model. The expected uncertainty of ±0.03 in the retrieved-SSA accounts for an error in the selection of aerosol model.

RC: Line 244: Please provide an explanation about the uncertainty limits (Q_0.03 and Q_0.05). Have they been arbitrarily selected?

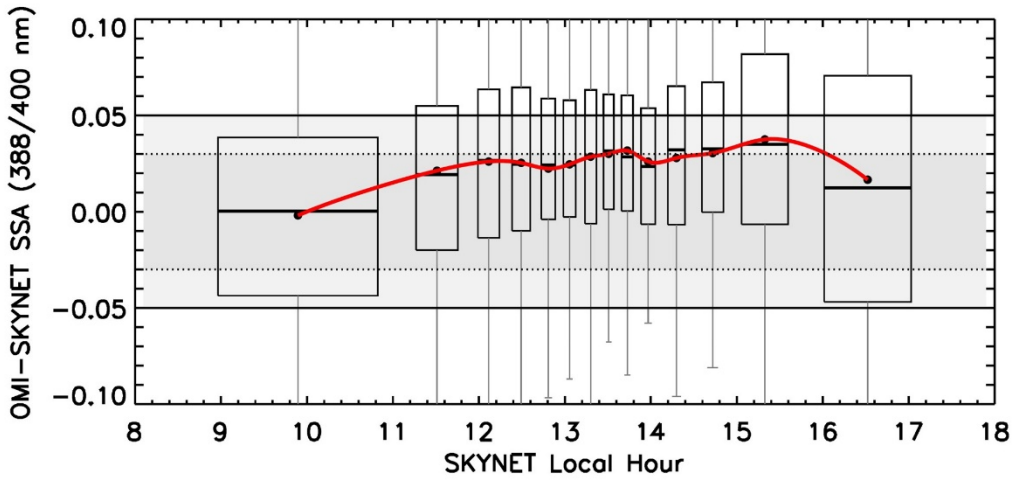
AR: The uncertainty limits of OMAERUV SSA retrievals, i.e., ±0.03 and ±0.05 were determined based on its comparison with AERONET SSA inversion reported in *Torres et al.* (2007). Following the early evaluation of OMI aerosol product for a handful of sites, *Jethva et al.* (2014) conducted a global evaluation of SSA product and also conducted a detailed uncertainty test considering different sources of errors, such as aerosol model, surface albedo, and aerosol layer height. The results of the sensitivity analysis suggested that overall, OMI SSA retrievals are uncertain to within 0.03 to 0.05, however errors could be even larger when algorithmic assumptions are far off from the real atmospheric conditions. This has been clarified in the section 2.1 of the revised paper.

RC: Section 3.3: Are there any quantified dependencies on the solar zenith angle, station altitude or aerosol layer height?

AR: The OMAERUV algorithm does consider the variation in altitude over land in the inversion by interpolating the radiances in terrain pressure in logarithmic space. Therefore, we don't expect significant error in the OMI SSA retrievals due to varying terrain pressure from one station to another. The aerosol layer height (ALH) assumed in each valid retrieval is referenced to the 30-month long CALIOP-OMI 1° x 1° global database [*Torres et al*., 2013]. Results obtained from a sensitivity study reported in *Jethva et al*. [2014] suggested that a decrease in the assumed ALH from 3 km to 1.5 km for the dust aerosol type results in a decrease in SSA by 0.02. Although, the climatology dataset of CALIOP-OMI adequately describes the observed mean layer of carbonaceous and desert dust aerosols, variations in the ALH not observed by CALIOP, therefore, will be missed out in the derived climatology and thus can be a source of uncertainty. Assuming that the CALIOP-OMI ALH dataset is accurate to within ±1 km, the resultant error in SSA retrieval would be ~ ±0.01.

The dependence of SSA difference on the local hour of SKYNET measurements are quantified in the Figure shown below. The SKYNET dataset accessed from the data server at Chiba University

doesn't contain information on the solar zenith angle. However, the local time of measurements reported in data file for each station can serve a proxy for the solar zenith angle. The OMI-SKYNET matchups exhibit a systematic dependency, where the differences between the two datasets become relatively minimal when early morning and late afternoon inversions of SKYNET associated with higher solar zenith angle are collocated with OMI retrievals around 1:30 PM local time overpass. Owing to a longer atmospheric optical path at higher solar zenith angle, thereby better aerosol signal, the ground-based aerosol inversions, such as from AERONET and SKY NET, are expected to be more reliable for sky measurements carried out during early morning/late afternoon.



The difference in SSA between OMI and SKYNET is related to the local measurement hour of SKYNET.

The results obtained in this analysis are added to the discussion in the revised manuscript.

RC: Lines 389-395: Unfortunately, the number of OMI-SKYNET pairs is limited in order to extract robust results. Moreover, how much can affect the irregularity of dust particles' shape? (see Gasteiger et al. (2011); https://www.tandfonline.com/doi/pdf/10.1111/j.1600-0889.2011.00559.x?needAccess=true

AR: The collocation procedure yields fewer OMI-SKYNET matchups for the dust aerosol type identified with the OMAERUV algorithm. The sensitivity study followed by an actual inversion of OMI data presented in *Torres et al.* (2018) demonstrates that the change in dust particle shape from spherical to spheroidal distribution improved the AOD retrievals significantly and brought the equivalency between the retrievals over left and right sides of the OMI swath for the dust belt region of tropical Atlantic. The associated changes in SSA retrievals were noted within ±0.01 and -0.02 for the scattering angle up to 100°-150° and >160°, respectively. The OMAERUV version 1.8.9.1 data product used in the present study adopts spheroidal dust model

based on the work of *Dubovik et al*. (2006) and *Torres et al*. (2018). We have elaborated the discussion on this aspect of the OMI aerosol retrievals in the uncertainty section 4.2.

RC: Lines 431-436: These two sentences are confusing. Since the absorption signal increases with height why the SSA values are also increase (i.e., less absorbing particles)?
AR: We realized that the information presented here is miscommunicated. The corrected sentence should be,

"This is because an increase (decrease) in the assumed aerosol layer height from the actual one enhances (reduces) absorption in the radiance look-up table (not in the actual TOA measurements), which the OMAERUV algorithm compensates by retrieving lower (higher) AOD and higher (lower) SSA to match with the observations."

RC: Lines 442-444: How exactly MODIS is used in the OMI near-UV surface albedo database?
AR: The new OMI-based near-UV surface albedo database was essentially derived based on the minimum reflectivity approach. The 7-year record of OMI observation of Lambertian Equivalent Reflectivity or LER at 388 was used to find its minimum value corresponding to UVAI<0.5 for each grid (0.25°x0.25°) and for each month. The 354-nm LER values are anchored to the corresponding minimum LER at 388 nm. During the persistent cloudy period for grids identified with higher LER values and abrupt temporal variations, the information on monthly variations of high-res MODIS surface albedo dataset aligned with the corresponding temporal record of OMI minimum LER was used to estimate LER in the near-UV.

RC: Section 4.2: In the discussion of this section there are not plots quantifying the OMI-SKYNET declinations with respect to cloud contamination, aerosol layer height and the prescribed surface albedo.
AR: The dependence of the SSA difference between OMI and SKYNET as a function of difference in the surface albedo assumption are already shown in Figure 6 and discussed in the section 4.2 in the manuscript. Regarding the cloud contamination, we use the 'best' quality retrievals from both sensors OMI (final algorithm quality flag=0) and SKYNET (cloud flag=0) in this study largely eliminating the chances of significant cloud interference in both kinds of measurements. Given the larger footprint size of OMI (13x24 km^2), however, the possibility of sub-pixel cloud contamination cannot be ruled out completely. On the other hand, a good level of agreement between the two retrievals at higher aerosol loading, and when the differences in the surface albedo reduces to minimum, are indicative of a consistency under favorable conditions. The explanation on the aerosol layer height assumption is explained in response to one the previous comments.

RC: Figures 4-5-6: Which is the difference on the obtained overall results when POM01 and POM02 SKYNET retrievals are grouped separately?
AR: Figure displayed below shows the number density plot comparing SSA between SKYNET and OMI for POM01 and POM02 sensors separately. Overall, we don't see a major difference in the derived statistics between the two sets of comparison, except that the number of matchups

obtained with POM02 sensor is 39% more than those with POM01 sensors and POM02 dataset offers marginally better comparison (except bias, which is higher with POM02) with OMI SSA retrievals. This analysis indicates that the interpolation of OMI SSA from 388 nm to 400 nm for its comparison with POM01 data isn't a significant source of discrepancy between the two SSA datasets. The revised manuscript includes the new analysis and related discussion.

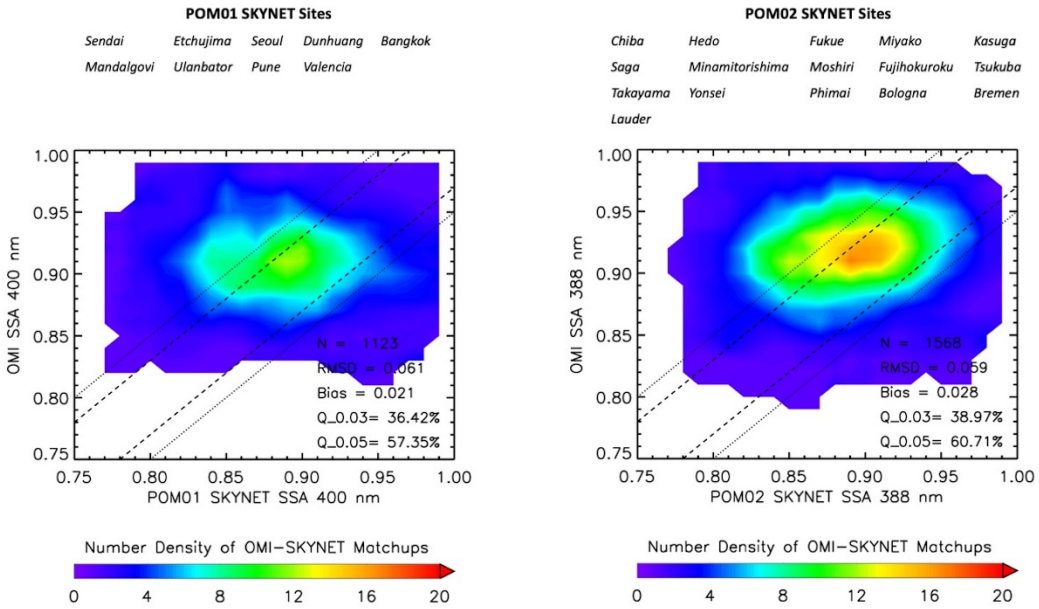


Figure Number density plots of SSA comparison between OMI and SKYNET matchups derived separately using POM-01 (left) and POM-02 (right) sensors. The resultant statistics of the comparison are depicted in the lower-right in each plot.

A Comparative Evaluation of Aura-OMI and SKYNET Near-UV Single-scattering Albedo Products

Hiren Jethva[1,2*], Omar Torres[2]

[1]Universities Space Research Association, Columbia, MD 21044 USA

[2]NASA Goddard Space Flight Center, Greenbelt, MD 20771 USA

Mailing Address:

Room#A422, Building#33,

Laboratory of Atmospheric Chemistry & Dynamics

Earth Science Division

NASA Goddard Space Flight Center,

Greenbelt, MD 20771, USA

[*]**Corresponding Author: Dr. Hiren Jethva**

E-mail: hiren.t.jethva@nasa.gov

ABSTRACT

The aerosol single-scattering albedo (SSA) retrieved by the near-UV algorithm applied to the Aura/Ozone Monitoring Instrument (OMI) measurements (OMAERUV) is compared with an independent inversion product derived from the sky radiometer network SKYNET-a ground-based radiation observation network span over Asia and Europe. The present work continues our efforts to evaluate the consistency between the retrieved SSA from satellite and ground sensors. The automated spectral measurements of direct downwelling solar flux and sky radiances made by SKYNET Sun-sky radiometer are used as input to an inversion algorithm that derives spectral aerosol optical depth (AOD) and single-scattering albedo (SSA) in the near-UV to near-IR spectral range. The availability of SKYNET SSA measurements in the ultraviolet region of the spectrum allows, for the first time, a direct comparison with OMI SSA retrievals eliminating the need of extrapolating the satellite retrievals to the visible wavelengths as the case in the evaluation against the Aerosol Robotic Network (AERONET). An analysis of the collocated retrievals from over 25 SKYNET sites reveals that about 61% (84%) of OMI-SKYNET matchups agree within the absolute difference of ±0.03 (±0.05) for carbonaceous aerosols, 50% (72%) for dust aerosols, 45% (75%) for urban-industrial aerosol types. Regionally, the agreement between the two inversion products was robust over several sites in Japan influenced by carbonaceous and urban-industrial aerosols, at the biomass burning site *Phimai* in Thailand, and polluted urban site in *New Delhi*, India. The collocated dataset yields fewer matchups identified as dust aerosols mostly over the site *Dunhuang* with more than half of the matchup points confined to within ±0.03 limits. Combinedly, the OMI-SKYNET retrievals agree mostly within ±0.03 for the AOD (388 or 400 nm) larger than 0.5 and UV Aerosol Index larger than 0.2. The remaining uncertainties in both inversion products can be attributed to specific assumptions made in the retrieval algorithms, i.e., the uncertain calibration constant, assumption of spectral surface albedo and particle shape, and sub-pixel cloud contamination. The assumption of fixed and spectrally neutral surface albedo (0.1) in the SKYNET inversion appears to be unrealistic, leading to ~~a large~~ underestimated~~ion of retrieved~~ SSA, especially under~~for~~ lower aerosol load conditions. At higher~~large~~ AOD values for carbonaceous and dust

287 aerosols, however, retrieved SSA values by the two independent inversion methods are
288 generally consistent in spite of the differences in retrieval approaches.

1 INTRODUCTION

Satellite-based remote sensing of aerosols has become an essential tool to detect, quantify, and routinely monitor the aerosol optical and size properties over the globe. An accurate representation of aerosols in the climate models is an essential requirement for reducing the uncertainty in aerosol-related impact on the Earth's radiation balance (direct and semi-direct effects) and cloud microphysics (indirect effect) (*IPCC*, 2013). The fundamental aerosol parameters determining the strength and sign of the radiative forcing are the aerosol optical depth (AOD) and single-scattering albedo (SSA) in addition to the reflective properties of the underlying surface. While the columnar AOD represents the total extinction (scattering and absorption) resulting from the interactions with solar radiation, SSA describes the relative strength of scattering to the total extinction. Together, both AOD and SSA determine the magnitude and sign of the aerosol radiative forcing at the top-of-atmosphere. For example, a decrease in SSA from 0.9 to 0.8 can often change the sign of radiative forcing from negative (cooling) to positive (warming) that also depends on the albedo of the underlying surface and the altitude of the aerosols (*Hansen et al*., 1997). Thus, an accurate estimate of both quantities is a prime requirement for reliable estimates of the net effect of atmospheric aerosols produced with the anthropogenic as well as natural activities.

Launched in July 2004, the Ozone Monitoring Instrument (OMI) onboard NASA's Aura satellite has now produced more than a decade long global record of observations of reflected radiation from Earth in the 270–500 nm wavelength range of the spectrum on a daily basis. OMI scans the entire Earth in 14 to 15 orbits with its cross-track swath of ~2600 km at ground level at a nadir ground pixel spatial resolution of 13 × 24 km^2. Satellite observations of the top-of-atmosphere reflected light at 354 and 388 nm wavelengths made by OMI are used to derive the UV aerosol index (UVAI) as well as the AOD and SSA using a near-UV algorithm (OMAERUV) that takes advantage of the well-known sensitivity to the aerosol absorption in the UV spectral region (*Torres et al*., 1998). While a general description of the OMI/OMAERUV algorithm is presented in *Torres et al.* (2007), the recent algorithmic upgrades are documented in *Torres et*

317 *al*. (2013, 2018). The most important changes applied in the latest OMAERUV algorithm
318 upgrade includes: 1) use of new carbonaceous aerosol models that account for the presence of
319 organics in the carbonaceous aerosols by assuming wavelength-dependent imaginary part of
320 the refractive index (*Jethva and Torres*, 2011), 2) an implementation of robust scheme to
321 identify aerosol type (smoke, dust, urban/industrial) that combinedly uses the information on
322 carbon monoxide (CO) observations from the Atmospheric Infrared Sounder (AIRS) and UVAI
323 from OMI (*Torres et al*., 2013), 3) use of the aerosol height climatology dataset derived from
324 the Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) lidar-based measurements of
325 the vertical profiles of aerosol for the carbonaceous and dust aerosols (*Torres et al*., 2013), and
326 4) better treatment of dust particles assuming realistic spheroidal shape distribution (*Torres et*
327 *al*., 2018). Additionally, the upgraded OMAERUV algorithm has adopted a new method to
328 calculate UVAI, which now accounts for the angular scattering effects of clouds and significantly
329 reduces a scan angle related asymmetry in UVAI in~~over~~ cloudy scenes (*Torres et al*., 2018).
330
331 The present work continues our efforts to evaluate the consistency between ground-based SSA
332 measurements and satellite retrievals from near UV observations. On the first attempt to
333 intercompare space-based and surface near UV SSA measurements, Earth Probe TOMS
334 retrievals were compared to AERONET observations acquired during the SAFARI 2000 field
335 campaign (*Torres et al.*, 2005). The OMAERUV near-UV aerosol product of AOD and SSA has
336 been continually assessed and validated against the ground-based measurements acquired
337 from the globally distributed Aerosol Robotic Network-AERONET (*Torres et al*., 2007; *Ahn et al*.,
338 2008; *Jethva and Torres*, 2011; *Ahn et al*., 2014; *Jethva et al*., 2014). While the OMAERUV AOD
339 product was directly validated against the AERONET measurements made in the near-UV (340-
340 380 nm), as carried out in *Ahn et al*. (2014), the SSA retrievals have been evaluated by
341 comparing with an independent ground inversion product of AERONET by *Jethva et al*. (2014).
342 The latter analysis required OMI retrievals of SSA to be extrapolated to the shortest visible
343 wavelength of 440 nm of AERONET inversion product to make the comparison possible. Such
344 adjustment in the wavelength of retrievals can introduce uncertainty in the comparison arising

345 from inaccuracy of the spectral dependence of absorption assumed in the wavelength
346 conversion.

347

348 A direct comparison of the column-integrated SSA at 388 nm retrieved from OMI requires
349 equivalent ground-based columnar retrievals in the near-UV region. The international network
350 of scanning sun-sky radiometers (SKYNET) fulfills this requirement as it performs the direct Sun
351 and sky measurements in the near-UV (340-380 nm) as well as visible/near-IR (400-1020 nm)
352 regions of the spectrum, and derives spectral AOD and SSA at these wavelengths. Taking
353 advantage of the availability of ground-based SSA inversions in the near-UV from SKYNET, we
354 inter-compare the OMI and SKYNET SSA products at several SKYNET sites in Asia and Europe.
355 Since both retrieval approaches are based on inversion algorithms that rely on assumptions, the
356 resulting level of agreement can only be interpreted as a measure of consistency (or lack
357 thereof) in the measurement of the same physical parameter by fundamentally different
358 remote sensing approaches.

359

360 The paper is organized as follows: Ss~~s~~ection 2 describes the satellite and ground-based data sets
361 assessed in this analysis along with the collocation methodology; the results of OMI-SKYNET
362 SSA comparison over individual sites, combinedly for each aerosol type, and diagnosis of
363 differences between them are presented in section 3; the possible sources of uncertainty in
364 both inversion products are discussed in section 4; the paper is summarized and concluded in
365 section 5.

366

367 # 2 DATASETS

368 ## 2.1 THE OMI-OMAERUV AEROSOL PRODUCT

369 The entire record of OMI observations (October 2004 to present) has been reprocessed
370 recently with the refined OMAERUV algorithm (PGEVersion V1.8.9.1) to derive a comprehensive

aerosol product that includes the retrievals of UV Aerosol Index (UVAI), AOD, SSA, and AAOD (388 nm) at a pixel resolution of 13 x 24 km^2 at nadir viewing geometry. The retrieved parameters are also reported at 354 nm and 500 nm wavelengths following the spectral dependence of aerosols assumed in the chosen model. The data set is available in the HDF-EOS5 format and can be obtained at no cost from NASA Goddard Earth Sciences (GES)-Data and Information Services Center (DISC) server at http://daac.gsfc.nasa.gov/. The recent upgrade has been documented in detail in the work of *Jethva and Torres* (2011), *Torres et al.* (2013, 2018) and *Ahn et al.* (2014). Here, we use the OMAERUV Level 2 Collection 003 (V1.8.9.1) aerosol product processed in July 2017. The expected uncertainty limits in the OMAERUV SSA retrievals are determined to be ±0.03 and ±0.05, based on its comparison with AERONET SSA inversion and sensitivity analysis carried out during the development of the OMAERUV algorithm (*Torres et al*., 2007). Following an early evaluation of OMI aerosol product for a handful of sites, *Jethva et al*. (2014) conducted a global evaluation of SSA product and also carried out a detailed uncertainty test considering different sources of errors, such as aerosol model, surface albedo, and aerosol layer height. The results of the sensitivity analysis further confirmed the uncertainty budget estimated earlier during the early development of the OMAERUV algorithm. However, note that the errors could attain larger magnitudes when algorithmic assumptions are far off from the real atmospheric conditions.

Post-2007, the OMI observations have been affected by a possible external obstruction that perturbs both the measured solar flux and Earth radiance. This obstruction affecting the quality of radiance at all wavelengths for a particular viewing direction is referred to as “row anomaly” (RA) since the viewing geometry is associated with the row numbers on the charge-coupled device detectors. The RA issue was detected first time in mid-2007 with a couple of rows which during the later period of operation expanded to other rows in 2008 and later. At present, about half of the total 60 rows across the track are identified and flagged as row anomaly affected positions for which no physical retrievals are performed (*Schenkeveld et al*., 2017). The details about this issue can be found at http://www.knmi.nl/omi/research/product/rowanomaly-background.php. The RA has

significantly affected the sampling during post-2008 OMI measurements, where row anomaly flags blanket about half of the OMI swath. ~~about half of the OMI swath is blanketed by row anomaly flagsr~~. As a result, the availability of the number of retrievals over a particular station is reduced starting in 2009 compared to earlier OMI measurements. Therefore, ~~Consequently,~~ the OMI-SKYNET matchups are also expected to be lower during the row anomaly affected period. The OMAERUV algorithm assigns quality flags to each pixel which carries information on the quality of the retrieval depending upon the observed condition. We used aerosol retrievals free of RA and flagged as quality flag '0', which are considered best in accuracy due to higher confidence in detecting aerosols in a scene with minimal or no cloud contamination.

2.2 The SKYNET Aerosol Inversion Product

The SKYNET is an international network of scanning sun-sky radiometers (manufactured by *Prede Co. Ltd.,* Japan) performing routine and long-term measurements of direct and diffuse solar radiations at several wavelengths spanning UV (340 and 380 nm), visible (400, 500, 675 nm), near-IR region (875, 1020 nm), and in shortware-IR (1627 nm and 2200 nm) of the spectrum. The automated measurements of direct and diffuse solar radiations are used to measure spectral AODs and retrieve SSAs and other aerosol optical-microphysical properties (volume size distribution, refractive index, phase function, and asymmetry parameter) at the same standard wavelengths of AOD following an inversion algorithm packaged in the *SKYRAD.pack* software (*Nakajima et al*., 1996; *Hashimoto et al*., 2012). Cloudy observations are screened using the Cloud Screening Sky Radiometer code (*Khatri and Takamura*, 2009).

The SKYNET radiometers come in two flavors, model POM-01 and model POM-02. The POM-01 instrument carries a total of five wavelength filters covering visible to near-IR (400-1020 nm), whereas POM-02 instrument has two additional filters in the UV region (340 and 380 nm) along with the other filters in the visible to shortwave-IR (including 1627 nm and 2200 nm) part of the spectrum. The calibration of each SKYNET radiometer is performed on-site on a monthly basis using the improved Langley method (*Nakajima et al*., 1996; ~~,~~ *Campanelli et al*., 2004, 2007). Occasionally, the inter-calibration of radiometers is carried out against the master instrument

well-calibrated using the Langley method at a high mountain site, e.g., Mauna Loa. The SKYNET radiometers are also inter-compared with AERONET Cimel Sunphotometers and precision filter radiometers at three observation sites, i.e., *Chiba University*, ~~and~~ *Valencia* (*Estelles et al*., 2016), and *Rome* (*Campanelli et al*., 2018)~~.~~.

Studies in the past have compared AODs (*Estellés et al*., 2012a) and SSAs (*Estellés et al*., 2012b) measured/retrieved from SKYNET and AERONET and shown that AODs are well-correlated and in good agreement, but the SKYNET SSAs are found to be higher than those of AERONET (*Che et al*., 2008; *Hashimoto et al*., 2012). *Khatri et al*. (2016) further pinpoints the factors, such as quality of input data attributed to different calibration and observation protocol, different quality assurance criteria, the calibration constant for sky radiances, differences in measured AOD, and surface albedo, responsible for the inconsistent aerosol SSA between AERONET and SKYNET using observations from the four representative sites, i.e., Chiba (Japan), Pune (India), Valencia (Spain), and Seoul (South Korea). More discussion on the sources of uncertainties is presented in section 4.

In this study, we include the SKYNET data acquired over a total of 25 sites distributed mostly across Asia and a few in Europe. The dataset is freely ~~accesible~~accessible from the data ~~holding~~ portal of the Center for Environmental Remote Sensing (CERes), Chiba University, Japan (http://atmos3.cr.chiba-u.jp/skynet/data.html). Figure 1 shows the geographic distribution of selected sites, whereas Table 1 lists the geo-coordinates of these sites with the associated sensor type (POM-01 or POM-02) and data periods. The SKYNET aerosol product is derived using two different Skyrad packs: version 4.2 and version 5, the differences of which are explained in *Hashimoto et al*. (2012). In this study, we use the SKYNET Level 2 product retrieved using version 5 of Skyrad pack. SKYNET retrievals assigned with cloud flag '0' are included in the analysis~~,~~ since these measurements are ~~-~~believed to be free of cloud contamination considered as higher quality retrievals. A careful examination of the SKYNET inversion dataset revealed some irregularities in the measurements for many sites, such as irregular patterns in the shape of spectral SSAs, identical values of SSA at near-UV and visible wavelengths, and much larger

standard deviation (>0.1) in SSA within a few hours. These spurious measurements were excluded from the present analysis.

2.3 THE COLLOCATION OF OMI AND SKYNET MEASUREMENTS

OMI retrievals correspond to a spatial scale of 13 × 24 km^2 at nadir representing the atmospheric conditions over an area. Unlike the direct measurements of the spectral AOD, which correspond to columnar point measurements, the retrievals made by SKYNET use the sky radiances measured at several discrete angles azimuthally, therefore representing the sky condition observed over a station which is associated with approximately 5 km radius surrounding the Sun photometer site. SKYNET retrieves aerosol optical-microphysical properties, including spectral SSA, under all cloud-free conditions and at all aerosol loadings. It is expected that the inversion of retrieved parameters from sky radiances offers better accuracy at larger solar zenith angles owing to the longer optical path and better aerosol absorption signal (*Dubovik et al.*, 2000). These conditions are best satisfied with the measurements made during the early morning and late afternoon hours. On the other hand, Aura/OMI overpasses a station during the afternoon hours with the local equator-crossing time 1:30 P.M. In order to collocate both types of measurements, therefore we select a time window of ±3 h around OMI overpass time in order to get sufficient high-quality SKYNET retrievals particularly from early morning/late afternoon measurements. The OMI retrievals of SSA were spatially averaged in a grid area of 0.5° by 0.5° centered at the SKYNET site. Though the spatial averaging area for the OMI retrieval is about 50 km^2, due to its larger footprint, the actual area intercepted by OMI pixels around SKYNET site is likely to be larger.

OMI performs retrieval at 354 nm and 388 nm wavelengths, whereas the SKYNET POM-02 instrument reports SSA at nearby wavelengths of 340, 380, and 400 nm. To compare both SSA products at the same wavelength, SKYNET SSA was linearly interpolated at 388 nm, to match with the wavelength of OMI retrieval, using the measurements at the two nearest wavelengths, i.e., 380 nm and 400 nm. The SKYNET POM-01 instruments don't carry UV wavelength filters, but report the retrievals at the shortest wavelength 400 nm and other visible/near-IR

wavelengths. In this case, the OMI retrievals are extrapolated from 388 nm to 400 nm, to match with the wavelength of SKYNET inversion, following the spectral dependence of SSA associated with the chosen aerosol model in the OMI algorithm. It is reasonably fair to assume that the extrapolation of OMI SSA in a narrow window of 12-nm, i.e., from 388 to 400 nm, should not be a major source of uncertainty in comparing SSA from OMI and SKYNET.

3 RESULTS

3.1 OMI-SKYNET COMPARISON OVER INDIVIDUAL STATIONS

Figure 2 displays the OMAERUV versus SKYNET SSA scatterplots for selected sites in Japan. The comparison was made at 388 nm or 400 nm depending upon the availability of the SKYNET inversion at those wavelengths, i.e., POM-01 or POM-02 sensors. Legends with different colors represent the aerosol type selected by the OMAERUV algorithm for the co-located matchups (N). RMSD is the root-mean-square difference between the two retrievals; Q_0.03 and Q_0.05 are the percent of total matchups (N) that fall within the absolute difference of 0.03 and 0.05, respectively; horizontal and vertical lines for each matchup are the standard deviation of temporally and spatially averaged SKYNET and OMI SSAs. The comparison includes OMI-SKYNET matchups with AOD>0.3 (388 or 400 nm) in both measurements simultaneously. The scatterplots reveal a good level of agreement for matchups identified with carbonaceous/smoke aerosols over *Chiba University*, *Cape Hedo*, *Fukue*, *Saga*, and *Etchujima* with the majority of points confined within the absolute difference of 0.03. The OMI-SKYNET combined dataset is~~are~~ dominated with matchup points identified as the urban/industrial aerosols by the OMAERUV algorithm for which the measured UVAI falls below 0.5 representing lower aerosol loading in the boundary layer with weakly absorbing properties. Under such observed conditions, the uncertainties in both kinds of measurements are prone to be larger due to lower absorption signal relative to the instrumental noise and errors in algorithmic assumptions, such as surface albedo, that could further amplify the overall uncertainty in the retrievals. Despite these inherent uncertainties, an agreement within the difference of ±0.03

511 for about or more than half of the collocated retrievals is encouraging. A more detailed
512 description of the different sources of uncertainty is presented in the next section.

513

514 Figure 3 shows the scatterplots of OMI-SKYNET SSA for remaining sites located in South Korea,
515 China, Thailand, India, and Italy. For the site *Seoul* in South Korea, OMI tends to overestimate
516 SSA for a number of matchups assigned with the urban/industrial aerosol type and for a few
517 with the carbonaceous/smoke aerosol type such that about 42% of total matchups are falling
518 within the difference of 0.03. For the *Dunhuang* site located in the desert area of China, a
519 majority of collocated data points were identified as dust aerosol type providing an overall
520 better agreement with 50% and 68% matchups bounded within ±0.03 and ±0.05 differences,
521 respectively. The *Phimai* site in Thailand is known to be influenced by the springtime biomass
522 burning activities, where OMI and SKYNET SSAs are found to agree relatively best among all 25
523 sites providing 71% and 91% of the matchups restricted within ±0.03 and ±0.05 limits,
524 respectively. The agreement between the two sensors was robust for the carbonaceous/smoke
525 aerosol type followed by the urban/industrial aerosols. Over the megacity of New Delhi in the
526 Indo-Gangetic Plain in India, which is seasonally influenced by the smoke and desert dust
527 aerosols in addition to the local source of urban pollution, the OMI-SKYNET matchups are found
528 to agree within ±0.03 and ±0.05 for 52% and 83% of the evaluated data points respectively.
529 Over the *Pune* station located near the western boundary of India and the *Bologna* site in Italy,
530 OMI retrieves higher SSA compared to that of SKYNET yielding 39% and 64%, and 25% and 50%
531 matchups, respectively, within the two uncertainty limits. Table 1 lists the statistical measures
532 of the OMI-SKYNET SSA comparison for all 25 sites.

533 ## 3.2 COMPOSITES FOR EACH AEROSOL TYPE

534 Figure 4 displays the composite scatterplots of OMI versus SKYNET SSA derived by segregating
535 the matchup points for each aerosol type from all 25 sites. The intention here is to evaluate the
536 consistency between the two retrieval methods for each aerosol type separately and
537 understand their relative differences. When identified as the carbonaceous/smoke aerosol type,
538 the OMI-SKYNET matchups reveal relatively best comparison among the three major aerosol

types with 61% and 84% data points falling within the absolute difference of 0.03 and 0.05, respectively, and providing the lowest (0.035) root-mean-square-difference between the two retrievals. The collocation procedure yields the lowest number of matchups (N=32) for desert dust aerosol type obtained mostly over the site of *Dunhuang* in China, resulting 50% and 72% of data points within the stated uncertainty limits. Among the three aerosol types, the collocated points assigned with the urban/industrial aerosol type (Figure 4 bottom-left) yield the maximum number of matchups (N=739) with the relatively weakest agreement (RMSD=0.052), where OMI tends to overestimate SSA for a significant number of instances resulting about 45% and 67% data points falling within the two limits of expected uncertainties. When more than one prescribed aerosol types are selected for OMI pixels around the SKYNET stations, the matchups between the two sensors resulted in 59% and 77% retrievals within the uncertainty limits with an RMSD of 0.041—a comparison slightly poorer than 'smoke-only' case, but better than 'dust-only' and 'urban/industrial-only' retrieval cases. Combined, all three distinct aerosol types simultaneously yield the total number of matchups (N=1223) with an RMSD of 0.047 between OMI and SKYNET resulting 51% and 72% collocated data points falling within the absolute difference of 0.03 and 0.05 difference, respectively. When the restriction of AOD>0.3 is removed from the collocation procedure, allowing all matchups regardless of their respective AOD values, the total number of collocated data points was increased to more than twice (N=2691) albeit with a relatively weaker agreement yielding an RMSD of 0.06 and percent data points within the uncertainty limits reducing to 38% and 59%, respectively.

3.3 Composites for Varying Aerosol Loading and POM-01 versus POM-02

Figure 5 shows the number density plots comparing OMI-SKYNET SSA matchups obtained from all sites combined and for the three aerosol loading conditions. The best set of comparison is achieved under the most restrictive scenario when corresponding OMI-rerieved AOD and UVAI are constrained to >0.3 and >0.5, respectively, albeit with a significantly reduced number of matchups compared to the other two cases with lesser (middle) or no (left) restrictions. The improved comparison reflected in statistical parameters is a result of avoiding retrievals with

lower aerosol loading when both kinds of measurements might be subjected to larger uncertainties due to algorithmic assumptions.

Figure 6 shows the number density plot comparing SSA between SKYNET and OMI for POM-01 and POM-02 sensors separately. Overall, no major difference noticed in the derived statistics between the two sets of comparison except that the number of matchups obtained with POM02 sensor is 39% more than those with POM01 sensors, and POM02 dataset offers marginally better comparison (except bias, which is higher with POM02) with OMI SSA retrievals. This analysis indicates that the interpolation of OMI SSA from 388 nm to 400 nm for its comparison with POM01 data isn't a significant source of discrepancy between the two SSA datasets.

~~3.3~~3.4 DIAGNOSIS OF OMAERUV VERSUS SKYNET SSA

The SKYNET algorithm inverts the spectral sky radiances in conjunction with the direct AOD measurements to retrieve the real and imaginary parts of the refractive index and particle size distribution of cloud-free observations under all aerosol loading conditions. These inversion products are believed to be more stable and accurate at higher~~larger~~ aerosol loadings and solar zenith angles due to stronger aerosol absorption signal and longer optical path (*Dubovik et al.*, 2000). Similarly, a sensitivity analysis of the two-channel OMAERUV retrievals suggests that the retrieved AOD and SSA are susceptible to the small change in surface albedo at lower aerosol loading (*Jethva et al.*, 2014). For instance, an absolute difference of 0.01 in the surface albedo leads to a change in AOD approximately by 0.1 and SSA by ~0.02.

Figure ~~5~~7 (top) shows the absolute difference in collocated SSA between OMI and SKYNET as a function of concurrent SKYNET direct AOD (388 or 400 nm) measurements for all aerosol types. All OMI-SKYNET matchup data obtained from a total of 25 sites under all AOD conditions are included here. The data are shown in the box and whisker format, where the horizontal lines

represent the median value of each data bin of sample size 150, filled circle the mean value, and shaded vertical bars cover the 25 and 75 percentiles of the population in each data bin. While for most bins the mean and median values of SSA difference were restricted to within ±0.03, OMI tends to overestimate SSA relative to that of SKYNET at lower AODs giving larger differences and spread in the data population. Similar patterns were observed when the difference in SSA was related to the OMI-retrieved AOD (Figure 7~~5~~ middle). In both cases, the differences in SSA minimize at larger AOD values (>0.5) suggesting a convergence in both retrievals. Figure 7~~5~~ (bottom) shows a similar plot of SSA difference against the concurrent OMI UVAI. Notably, the differences in SSA exhibit even a stronger relationship to UVAI than that in the AOD case (top and middle). For UVAI lesser than zero, the differences in the retrieval are found to be beyond the expected uncertainty in both inversions, at least in the mean sense. For the lower range of UVAI, OMI algorithm mostly employs the urban/industrial model for the retrieval where all aerosols are assumed to be confined within the boundary layer (<2 km) with a vertical profile that follows an exponential distribution. On the other hand, the mean and median values of the SSA difference for UVAI larger than 0.2 for all bins fall within the 0.03 uncertainty range. The SSA differences approach to near-zero with a reduced spread at larger magnitudes. Notably, both inversions are found to be in closer agreement for UVAI measurements>0.3.

4 SOURCES OF UNCERTAINTY

4.1 UNCERTAINTIES IN THE GROUND-BASED SKYNET INVERSION PRODUCT

The standard SKYNET inversion algorithm assumes a wavelength-independent surface albedo of 0.1 at all wavelengths across the UV to visible part of the spectrum. However, the algorithm code allows flexibility to alter the value surface albedo in time and wavelength (Campanelli et al., 2015). The diffuse light reflected from the ground plays a second-order role in the measured sky radiances in most situations, however, has a potential to affect the SSA inversion, e.g.,

overestimated (underestimated) surface albedo can underestimate (overestimate) SSA (*Dubovik et al*., 2000; *Khatri et al*., 2012). Using simultaneous inversion data from SKYNET and AERONET for four representative sites, *Khatri et al*. (2016) have shown that the difference in the prescribed surface albedo between SKYNET and AERONET results in a difference of ~0.04 in SSA at red (675 nm) and near-IR wavelengths retrieved from the two collocated ground sensors. The difference in SSA can also reach as large as ~0.08 when surface albedo differed by 0.3. The assumed surface albedo value of 0.1 at near-UV (340 and 380 nm) and shorter visible wavelength (400 nm) seems to be unrealistic for the vegetated and urban surfaces. The surface albedo database at 354 nm and 388 nm derived from multiyear observations from OMI suggests that the vegetated surfaces and urban centers are characterized with the lower values of surface albedo, i.e., ~0.02-0.03 and ~0.05, respectively; for desert surfaces, the albedo could be as high as 0.08-0.10. Significant differences in the assumed surface albedo values between OMI and SKYNET at shorter wavelengths could be one of the responsible factors for discrepancies in SSA noted over several sites, particularly at lower aerosol loading when the uncertainty in surface characterization can amplify error in the SSA inversion.

To further investigate this effect, the difference in SSA between OMI and SKYNET as a function of the simultaneous difference in surface albedo is analyzed and shown in Figure ~~6~~8 (top). The data are presented in a standard box and whisker plot format. Figure 8~~6~~ reveals a link between differences in SSA and surface albedo, where increasing differences in SSA (OMI>SKYNET) are associated with significant negative biases in surface albedo between OMI and SKYNET. In other words, large overestimation in SKYNET surface albedo causes underestimation of retrieved SSA, which is consistent with the findings of *Dubovik et al*. (2000) and *Khatri et al*. (2012, 2016), thereby resulting in a substantial positive difference in SSA between OMI and SKYNET. Recently, *Mok et al.* (2018) have shown that the use of AERONET surface albedo dataset at 440 nm in the SKYNET algorithm for the S. Korea region produces SSA values larger by ~0.01 at near-UV wavelengths. Notably, differences in SSA tend to be lower when the differences in surface albedo are also minimal, such that the mean and median values of those bins remain within the expected uncertainties of ±0. 03 in both retrievals. This result, along with the previous findings

cited above, convincingly points out that the SSA inversion from ground-based sensors, especially at lower aerosol loadings, is likely susceptible to the prescribed surface albedo. The assumption of a fixed value of spectral surface albedo of 0.1 in the SKYNET algorithm appears to be inappropriate calling for a revision using more accurate datasets of spectral reflectance or albedo such as from MODIS and OMI.

The dependence of SSA difference on the local hour of SKYNET measurements is quantified in the Figure 8 (bottom). The SKYNET dataset accessed from the data server at Chiba University doesn't contain information on the solar zenith angle. However, the local time of measurements reported in the data file for each station can serve a proxy for the solar zenith angle. The OMI-SKYNET matchups exhibit a systematic dependency, where the differences between the two datasets become relatively minimal when early morning and late afternoon inversions of SKYNET associated with higher solar zenith angle are collocated with OMI overpass time around 1:30 PM equator-crossing time. Owing to a longer atmospheric optical path at higher solar zenith angle, thereby better aerosol absorption signal, the ground-based aerosol inversions, such as from AERONET and SKYNET, are expected to be more reliable for sky measurements carried out during early morning/late afternoon.

SKYNET inversion algorithm (Skyrad.pack Version 4.2 and version 5) assumes aerosols of spherical shape regardless of the actual aerosol type observed in the scene. Following a detailed analysis of the effect of non-sphericity of the particles on the difference between the retrievals carried out assuming spherical and spheroidal size distribution, *Khatri et al.*, (2016) concluded that the assumed shape of particles has a non-significant impact on the retrieved SSA. Their study revealed SSA difference of ±0.01 for measurements having a maximum scattering angle <120° and difference of up to ±0.02 at scattering angle >120°, where the difference in the phase function is significant between spherical and spheroidal size distributions (*Torres et al.*, 2018). The OMI-SKYNET collocation procedure, as shown in Figure 4, yields relatively fewer matchups that are identified as dust aerosol type according to the

678 OMAERUV aerosol type identification scheme. A majority of the collocated data points were
679 derived over the desert site of *Dunhuang* in China showing a reasonable agreement in SSA
680 between OMI and SKYNET for dust aerosols further supporting the findings of *Khatri et al*.
681 (2016) that the SSA retrievals are not significantly impacted by the assumption of the shape of
682 particles, i.e., spherical or spheroidal.

683

684 Apart from the algorithmic assumptions, the calibration constant used for sky radiances
685 measured by SKYNET instruments can be a potential source of errors in the inversion. *Khatri et*
686 *al*. (2016) suggests that the calibration constant for sky radiances determined from the disk
687 scan method using solar disk scan area of 1° × 1° (*Boi et al*., (1999) may be underestimated
688 resulting in overestimated sky radiance and thus relatively higher SSA. Some of the larger
689 differences between in SSA between OMI and SKYNET, where OMI underestimates SSA relative
690 to the SKYNET, can be attributed to the imperfect calibration applied to the SKYNET sensors.

691 ## 4.2 Possible Sources of Uncertainties in OMAERUV Retrievals

692 Like other satellite-based remote sensing algorithms, OMAERUV also relies on assumptions
693 about the atmospheric and surface properties for the retrieval of aerosol properties. The single
694 largest known source of error in the OMI retrievals is the subpixel cloud contamination within
695 the OMI footprint. Given the footprint of size 13 × 24 km^2 for near-nadir pixels which intercept
696 an area of about 338 km^2 on the ground, the presence of subpixel clouds may not be avoided
697 entirely. Currently, the algorithm assigns quality flags to each pixel which carries information on
698 the quality of the retrieval depending upon the observed conditions (*Torres et al*., 2013).
699 Aerosol retrieval with the quality flag '0' are considered to be the best in accuracy as this
700 category of flag scheme largely avoids cloud-contaminated pixels by choosing the appropriate
701 thresholds in reflectivity and UVAI measurements.

702

703 Over the desert regions, e.g., the *Dunhuang* SKYNET site in China, the frequency of occurrence
704 of clouds is expected to be minimal. Therefore, it is less likely that the SSA retrievals over these

sites are affected by cloud contamination. A reasonable agreement between the two retrieval datasets~~s~~ (Figure 3) supports this assumption. The quality flag scheme, however, cannot entirely rule out the presence of small levels of subpixel cloud contamination or the presence of thin cirrus in the OMI footprint, which can cause overestimation in the retrieval of SSA, such as noted over the SKYNET sites in *Kasuga*, *Etchujima*, *Seoul*, *Bologna*, and *Pune*. Largest uncertainties observed over these sites are associated with the urban-industrial aerosol type, possibly due to the fact that the AOD's for this aerosol type are the lowest in the analysis, and therefore, subject to the less sensitivity to absorption and possibly more affected by sub-pixel cloud contamination.

Another possible source of uncertainty can be the assumption of the aerosol layer height. The climatology of aerosol layer height derived from CALIOP measurements adequately describes the observed mean layer of carbonaceous and desert dust aerosols (*Torres et al*., 2013). It is particularly robust over the arid and semiarid areas where large numbers of cloud-free observations were used in the calculation. However, note that the temporal and spatial coverage of CALIOP is limited to 16-day repeat cycle over the same location. Variations in the aerosol layer height not observed by CALIOP, therefore, will be missed out in the derived climatology and thus can be a source of uncertainty. Sensitivity analysis of the OMAERUV retrievals suggests that an overestimation (underestimation) in the aerosol layer height results in an overestimated (underestimated) SSA. ~~This is because an increase (decrease) in the aerosol layer height from the actual one enhances (reduces) absorption signal in the radiance measurements in near-UV, which the OMAERUV algorithm compensates by retrieving higher AOD and SSA to match with the observations.~~ This is because an increase (decrease) in the assumed aerosol layer height from the actual one enhances (reduces) absorption in the radiance look-up table (not in the actual TOA measurements), which the OMAERUV algorithm compensates by retrieving lower (higher) AOD and higher (lower) SSA to match with the observations.

733 The third source of uncertainty that can affect SSA retrieval is the accuracy of the prescribed
734 surface albedo. For the surface characterization, the OMAERUV algorithm use a near-UV
735 surface albedo database derived using the multiyear OMI reflectivity observations. The method
736 adopts a minimum reflectivity approach, ensuring minimal or no contamination from the
737 atmosphere, i.e., aerosols and clouds, in the measurements. Afterward, the minimum
738 reflectivity dataset derived from the OMI observations was adjusted in the temporal domain to
739 the seasonality of surface albedo retrieved in the visible wavelengths from MODIS. The dataset
740 contains surface albedo values at 354 and 388 nm at a grid resolution of 0.25° × 0.25°.
741 Compared to the previous OMAERUV dataset using TOMS-based surface albedo product at 1°
742 grid resolution, the new OMI-based dataset is expected to be more accurate to within 0.005 to
743 0.01 owing to its higher spatial resolution and the fact that it is contemporary to the OMI
744 operation. A sensitivity study of the OMAERUV retrievals to the change in surface albedo
745 described in *Jethva et al*. (2014) suggests that an increase in surface albedo by 0.01 in the near-
746 UV region over desert areas results in a decrease in the magnitude of retrieved SSA by ~ -0.02.
747 The effect of uncertain surface albedo can be more pronounced at lower aerosol loading,
748 where the reduced signal from the atmosphere makes OMAERUV retrieval more susceptible to
749 the uncertainty in surface albedo.

750

751 The assumed aerosol microphysical and optical properties could be additional sources of
752 uncertainty. The particle size distributions assumed in the OMAERUV models are adopted from
753 long-term AERONET inversion statistics (*Dubovik et al*., 2002), representing areas influenced by
754 smoke, dust, and urban/industrial aerosols, and therefore are considered realistic
755 representations of the total atmospheric column. The carbonaceous smoke aerosols are
756 assumed to be spherical in shape with a bimodal log-normal size distribution and characterized
757 with a steep absorption gradient, such that the Absorption Angstrom Exponent (AAE) in the
758 near-UV lies in the range 2.5-3.0, to adequately represent the organics in the biomass burning
759 smoke particles (*Kirchstetter et al*., 2004; *Jethva and Torres*, 2011). The desert dust aerosol
760 model follows bimodal log-normal size distribution with particles comprised of randomly
761 oriented spheroids with an axis ratio (shape factor) distribution adopted from *Dubovik et al*.

(2006). The sensitivity study followed by an actual inversion of OMI data presented in Torres et al. (2018) demonstrates that the change in dust particle shape from spherical to spheroidal distribution improved the AOD retrievals significantly and brought the equivalency between the retrievals over left and right sides of the OMI swath for the dust belt region of tropical Atlantic. The associated changes in SSA retrievals were noted within ±0.01 and -0.02 for the scattering angle up to 100°-150° and >160°, respectively. The OMAERUV version 1.8.9.1 data product used in the present study adopts spheroidal dust model based on the work of Dubovik et al. (2006) and Torres et al. (2018). The spectral dependence of the refractive index in the near-UV assumed in the dust aerosol model is generally consistent with the in-situ laboratory measurements (*Wagner et al*., 2012). For instance, retrieval of AOD and SSA for carbonaceous aerosols using the smoke model with AAE of 1.90 (10% relative spectral dependence in the imaginary index between 354 and 388 nm) and 1.0 (no spectral dependence in the imaginary index), instead of the standard AAE assumption of 2.7, results in a decrease in SSA up to -0.07, respectively, suggesting a marked sensitivity of the SSA retrieval to the significant changes in the spectral aerosol absorption. Due to the shortage of ground-based characterization of absorption in the near-UV part of the spectrum, the regional representation of the spectral absorption properties in the OMAERUV models is limited. Therefore, spatial and temporal variations in the spectral properties of ~~actual~~ aerosol~~s spectral properties~~ can be a potential source of error in the SSA retrieval.

5 SUMMARY AND CONCLUSION

We presented a comparative analysis of the aerosol SSA retrieved from the OMI's two-channel aerosol algorithm (OMAERUV) against an independent ground-based inversion made by the SKYNET Sun photometers over selected 25 sites located mainly in Asia and Europe. This study follows our previous efforts of evaluating the OMI near-UV SSA product carried out using ground-based AERONET dataset (*Jethva et al*., 2014). The capability of SKYNET sensors to measure the Sun and sky radiance at near-UV wavelengths (340-380-400 nm), and subsequently retrieve the aerosol optical properties, including SSA, at these wavelengths

provide a unique opportunity to directly compare the two near-UV SSA products from ground and satellite. Ground-based inversion of SSA at the near-UV wavelengths eliminates the need to adjust and extrapolate satellite retrieval to the visible wavelengths such as the case with comparison against AERONET. Since the SSA inferred from two different platforms are essentially retrieved from two fundamentally different inversion algorithms, the present study does not stand as a "validation" exercise for either retrieval data sets. Instead, the purpose of this analysis was to check the consistency (or lack thereof) between the two retrieved quantities of the same physical parameter regarding standard statistical comparison, i.e., RMSD and % of matchups within the expected uncertainties.

Unlike AERONET Level 2 inversion product that reports spectral SSA when AOD (440 nm) exceeds a value of 0.4, SKYNET Level 2 dataset delivers spectral SSA in the near-UV and visible parts of the spectrum under all cloud-free observations for all AOD conditions. The collocation procedure that matched temporal inversion data from SKYNET with spatial retrievals from OMI gave resulted in a total of 2691 collocated data points for AOD>0.0 and 1223 when AOD>0.3 collected from 25 sites representing biomass burning region of Southeast Asia, desert in China, and urban/industrial areas in Japan, India, and Europe. Combinedly for all 25 sites and under all AOD conditions, we find 38% and 59% of the total SKYNET-OMI SSA agree within their estimated uncertainty range of ±0.03 and ±0.05, respectively, with an overall root-mean-square-difference of 0.06. When restricted with condition AOD>0.3 in both measurements, the agreement of comparison improved to 51% and 72% with root-mean-square-difference of 0.047. When segregated by aerosol type, the agreement between the two sensors is found to be robust for matchups identified as the carbonaceous aerosols over several sites in Japan, *Seoul* in South Korea, *Phimai* in Thailand, and *New Delhi* in India, yielding 61% and 84% of data points falling within the limits of ±0.03 and ±0.05 with an overall RMSD of 0.035. The collocation procedure found few matchups for desert dust aerosol, mostly over *Dunhuang* site in China, showing a reasonable comparison with 50% and 68% data points within expected uncertainty limits. Among the three major aerosol types, the urban/industrial type aerosols

provide the maximum number of matchup data points with a relatively poorer comparison, where 45% and 67% data are found to be within the uncertainty limits.

The differences in SSA between OMI and SKYNET are found to be large~~rst~~ at lower aerosol loading, where OMI retrieves significantly higher SSA compared to that of SKYNET. However, the differences are minimized at larger AOD values (>0.5) suggesting a convergence in both retrievals at moderate to larger aerosol loading. Similarly, the differences in SSA exhibit a stronger relationship to UVAI showing larger discrepancies beyond expected uncertainty limits at lower UVAIs (<0), but nearing to zero with a reduced spread in matchups at larger magnitudes of UVAI (>0.2-0.3).

Much of the inconsistency observed between OMI and SKYNET at lower aerosol loadings indicate retrieval issues due to reduced signal-to-noise ratio and uncertain algorithmic assumptions. For instance, the OMAERUV retrievals are more susceptible to the changes in surface albedo at lower AODs, and to the spectral absorption at higher AODs (*Torres and Jethva*, 2011). On the other hand, the SKYNET inversion algorithm assumes a wavelength-independent surface albedo of 0.1 across the UV to visible-near-IR wavelengths, which appears to be unrealistic especially in the UV region where OMI surface albedo dataset shows much lower values (<0.05) over land. Though the reflected light from surface plays a second-order role in the ground-based retrievals, previous studies as well as shown in the present work (Figure ~~6~~8), uncertainty in surface albedo can cause non-negligible errors in SSA retrievals that likely exceed the expected accuracy level of ±0.03.

Despite the inherent uncertainties associated with both satellite and ground inversion products, a good level of agreement between the two independent techniques over SKYNET sites at increasing aerosol loading is encouraging. We intend to extend the present analysis to other SKYNET sites whose data are still not directly accessible in the public domain. Continuing the

evaluation of inversion products, both from satellite and ground, is an important exercise to track the changes and improvements in the algorithms and resulting data products, and to establish the consistency (or lack thereof) that can help to diagnose further and improve the accuracy of retrievals.

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AUTHORS' CONTRIBUTIONS

Dr. Jethva, the leading author, conceptualized the study and wrote the paper. He conducted comparative data analysis of OMI- and SKYNET-retrieved single-scattering albedo products presented in the paper. Dr. Torres (2nd author) brought his expertise in interpreting the results and helped improving the manuscript writeup.

Additional Information

The author(s) declare no competing interests, financial or non-financial.

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TABLES

Table 1 A list of SKYNET sites and corresponding dataset used in the present analysis. Sensor type "POM02" consists of a total of seven wavelength filters, including near-UV bands, i.e., 340, 380, 400, 500, 675, 870, and 1020 nm, whereas "POM01" sensors have a total of five wavelength filters, i.e., 400, 500, 675, 870, and 1020 nm. The rightmost four columns enlist the statistical measures of OMI-SKYNET single-scattering albedo matchups.

Abbreviations: N: number of satellite-ground matchups, RMSD: root-mean-square-difference between OMI and SKYNET, Q_0.03 and Q_0.05: percent matchups within an absolute difference of 0.03 and 0.05.

SKYNET Station Name	Longitude	Latitude	Country	Sensor Type	Data Period	N	RMSD	Q_0.03 (%)	Q_0.05 (%)
Chiba University	140.104°E	35.625°N	Japan	POM02	2005-2017	132	0.039	58	81
Cape Hedo	128.248E	26.867N	Japan	POM02	2005-2017	47	0.044	47	72
Fukue	128.682E	32.752N	Japan	POM02	2008-2017	71	0.041	59	76
Miyako	125.327E	24.737N	Japan	POM02	2004-2017	31	0.059	23	58
Sendai	140.84E	38.26N	Japan	POM01	2009-2017	34	0.052	50	74
Kasuga	130.475E	33.524N	Japan	POM02	2004-2017	159	0.057	40	61
Saga	130.283E	33.233N	Japan	POM02	2011-2017	66	0.044	52	71
Minamitorishima	153.97E	24.3N	Japan	POM02	2006-2009	-	-	-	-
Moshiri	142.260E	44.366N	Japan	POM02	2009-2011	2	0.018	100	100
Fuji Hokuroku	138.750E	35.433N	Japan	POM02	2009-2017	9	0.051	56	67
Tsukuba	140.096E	36.114N	Japan	POM02	2014-2017	5	0.027	80	100
Takayama	137.423E	36.145N	Japan	POM02	2014-2017	3	0.022	67	100
Etchujima	139.796E	35.664N	Japan	POM01	2004-2010	100	0.052	45	66
Seoul	126.95E	37.46N	Republic of South Korea	POM01	2005-2015	182	0.050	42	66
Yonsei	126.980E	37.570N	Republic of South Korea	POM02	2016	5	0.035	40	80
Dunhuang	90.799E	40.146N	China	POM01	1999-2007	40	0.048	50	68
Phimai	102.564E	15.184N	Thailand	POM02	2005-2017	139	0.031	71	91
Bangkok	100.605E	13.667N	Thailand	POM02	2009-2017	15	0.064	47	60
Mandalgovi	106.264E	45.743N	Mongolia	POM01	1998-2009	4	0.087	0	0
Ulan Bator	106.921E	47.923N	Mongolia	POM01	2013-2017	2	0.026	100	100
New Delhi	77.174E	28.629N	India	POM01	2006-2007	63	0.038	52	83
Pune	73.805E	18.537N	India	POM01	2004-2009	94	0.050	39	64
Bologna	11.34E	44.52N	Italy	POM02	2014-2017	114	0.065	25	50
Valencia	0.420E	39.507N	Spain	POM01	2014-2017	4	0.052	25	25
Bremen	8.854E	3.108N	Germany	POM02	2009	-	-	-	-

FIGURES

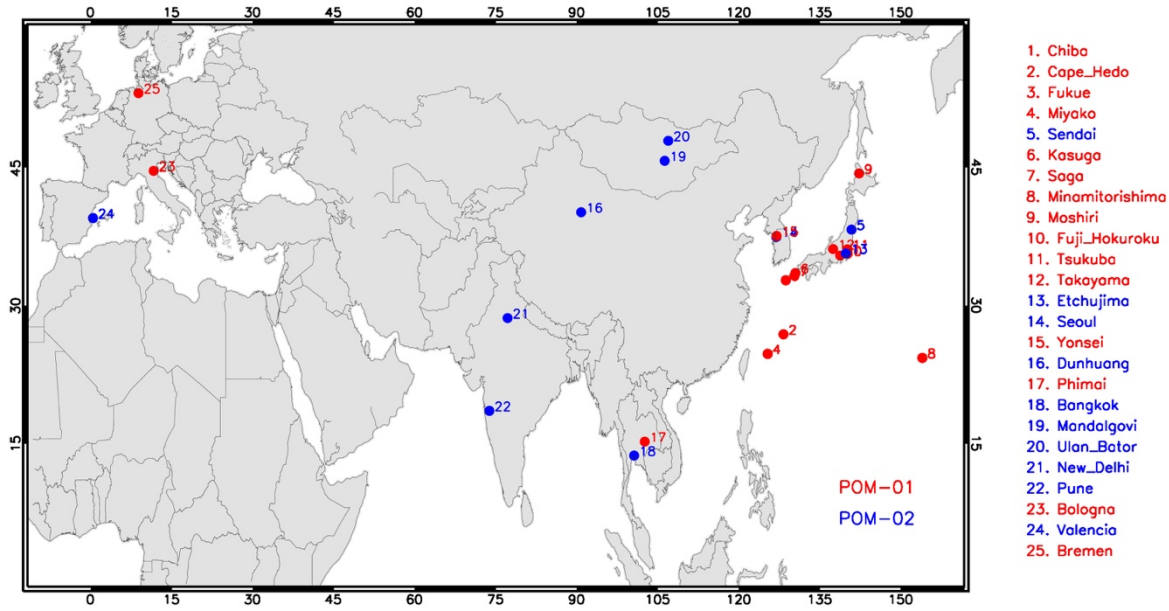

Figure 1 Geographical placement of ground-based SKYNET sensors over sites in Asia and Europe The SKYNET dataset for these sites are freely accessible from the Center for Environmental Remote Sensing (CERes), Chiba University, Japan (http://atmos3.cr.chiba-u.jp/skynet/data.html).

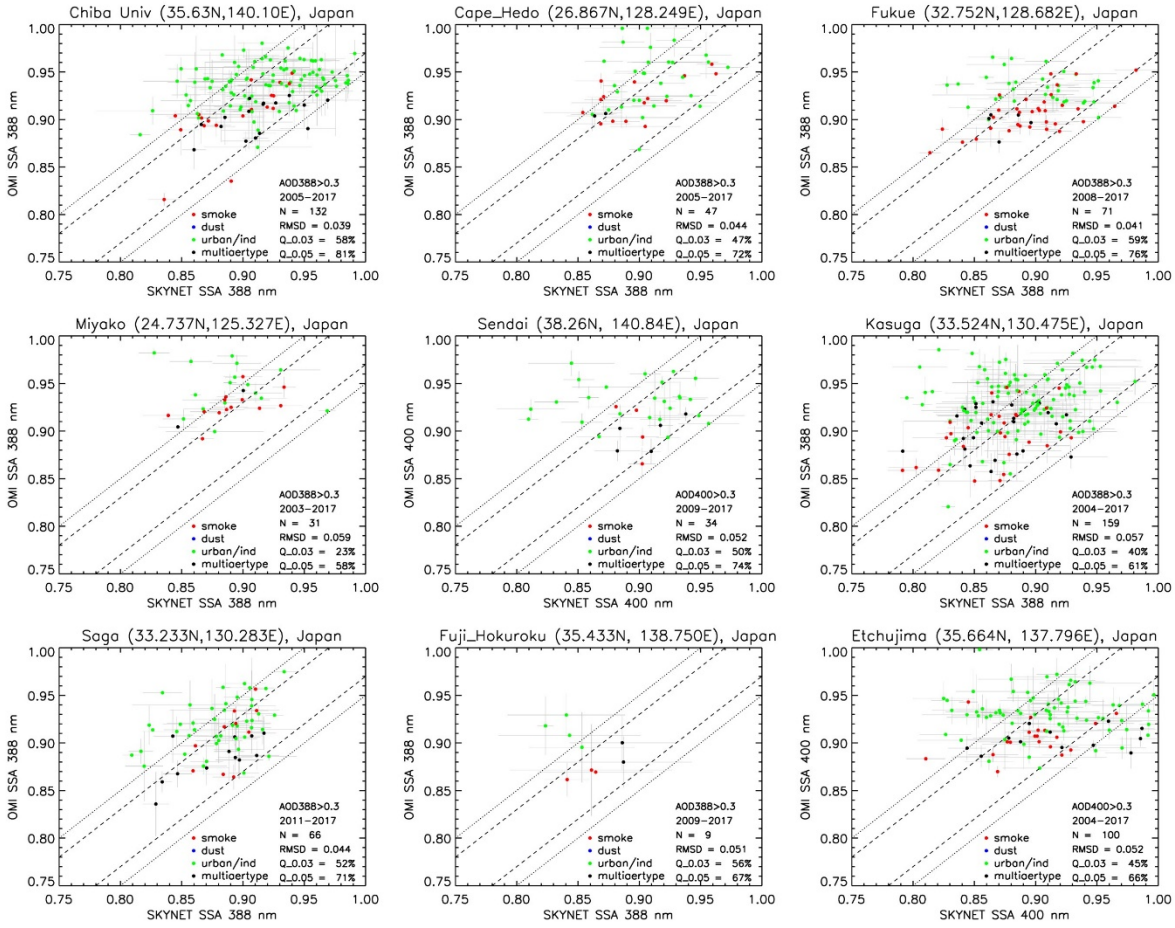

Figure 2 OMAERUV versus SKYNET single-scattering albedo comparison for different sites in Japan. Legends with different colors represent the aerosol type selected by the OMAERUV algorithm for the co-located matchups (N). RMSD is the root-mean-square difference between the two retrievals; Q_0.03 and Q_0.05 are the percents of total matchups (N) that fall within the absolute difference of 0.03 and 0.05, respectively. OMI-SKYNET matchups with AOD>0.3 (388 or 400 nm) in both measurements are used for comparison.

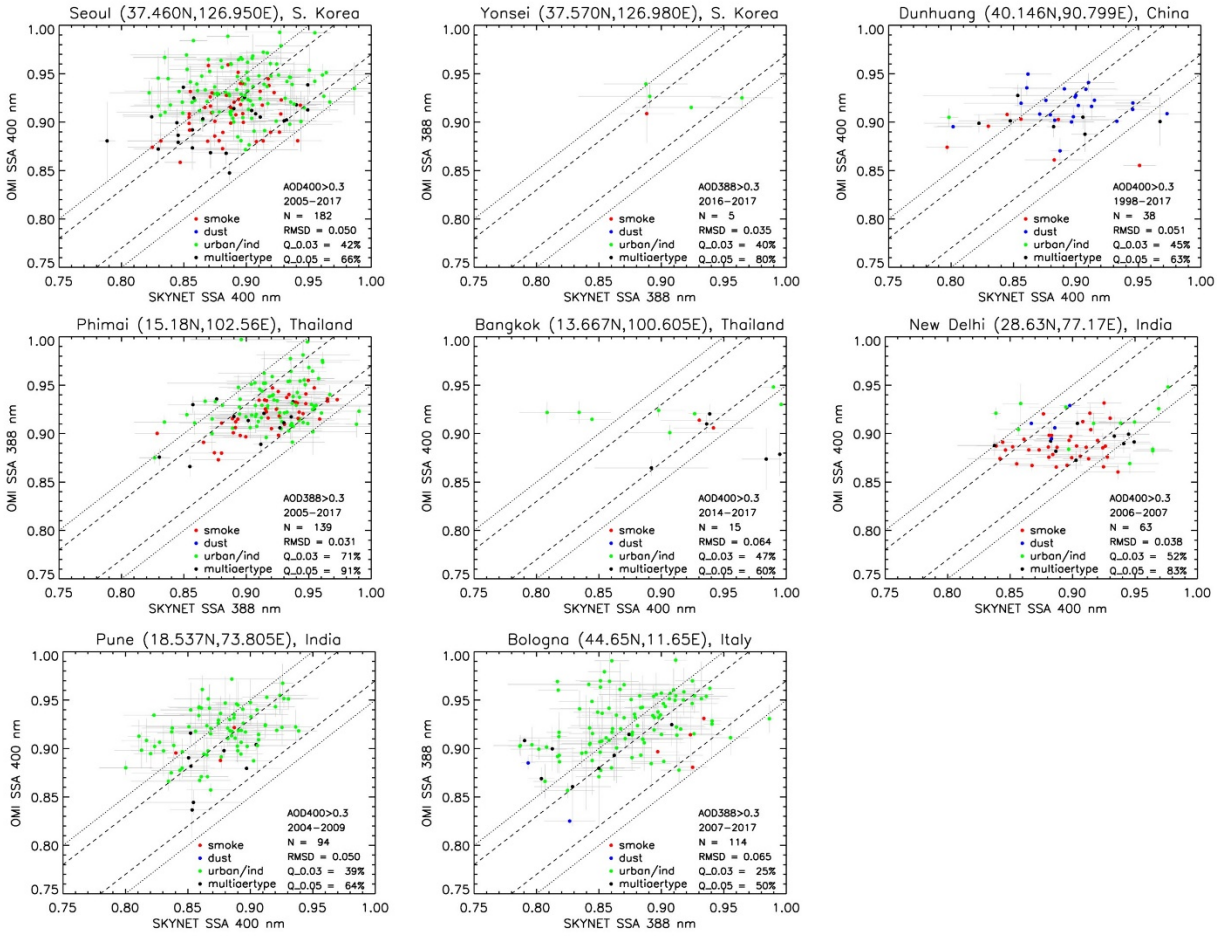

Figure 3 Same as in Figure 2 but for SKYNET sites in South Korea, China, Thailand, India, and Italy.

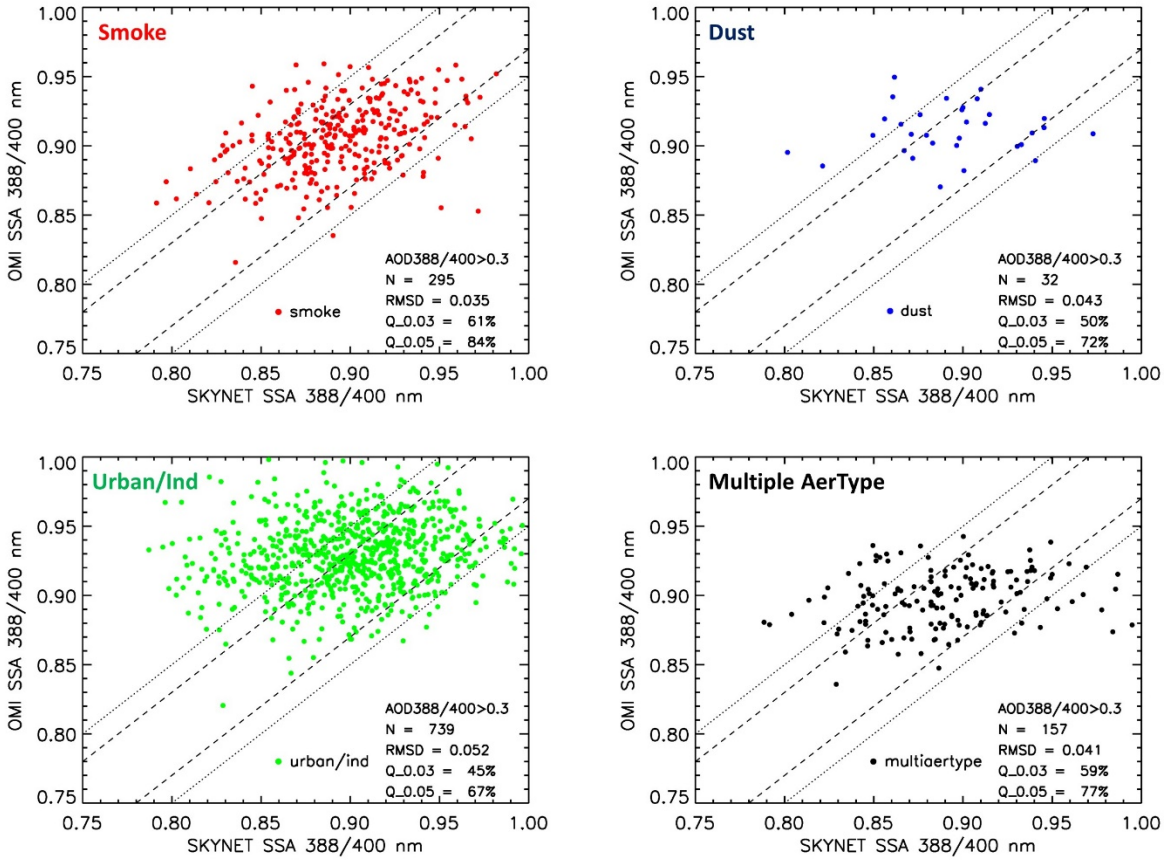

Figure 4 Composite scatterplots of OMAERUV versus SKYNET single-scattering albedo (388 or 400 nm) for the three distinct aerosol types, i.e., smoke, dust, and urban/industrial, as identified by the OMAERUV algorithm. OMI-SKYNET matchups with AOD>0.3 (388 or 400 nm) in both measurements are used for the comparison.

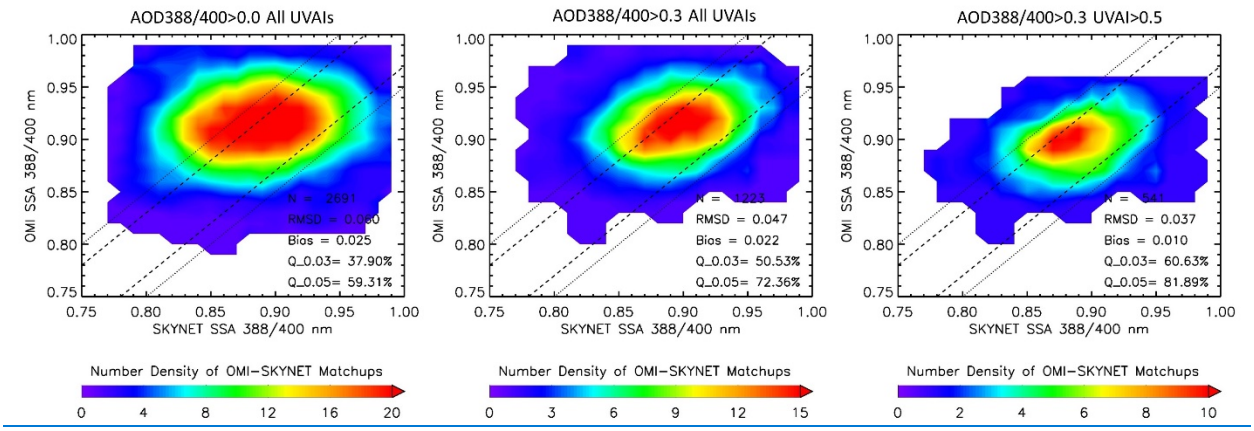

Figure 5 Composite number density ~~contour~~ plots of SSA comparison between OMI and SKYNET for different aerosol loading conditions. The resultant statistics of the comparison are depicted in the lower-right in each plot. Note that the scale used for number density of satellite-ground matchups for the three sets of comparisons are different.

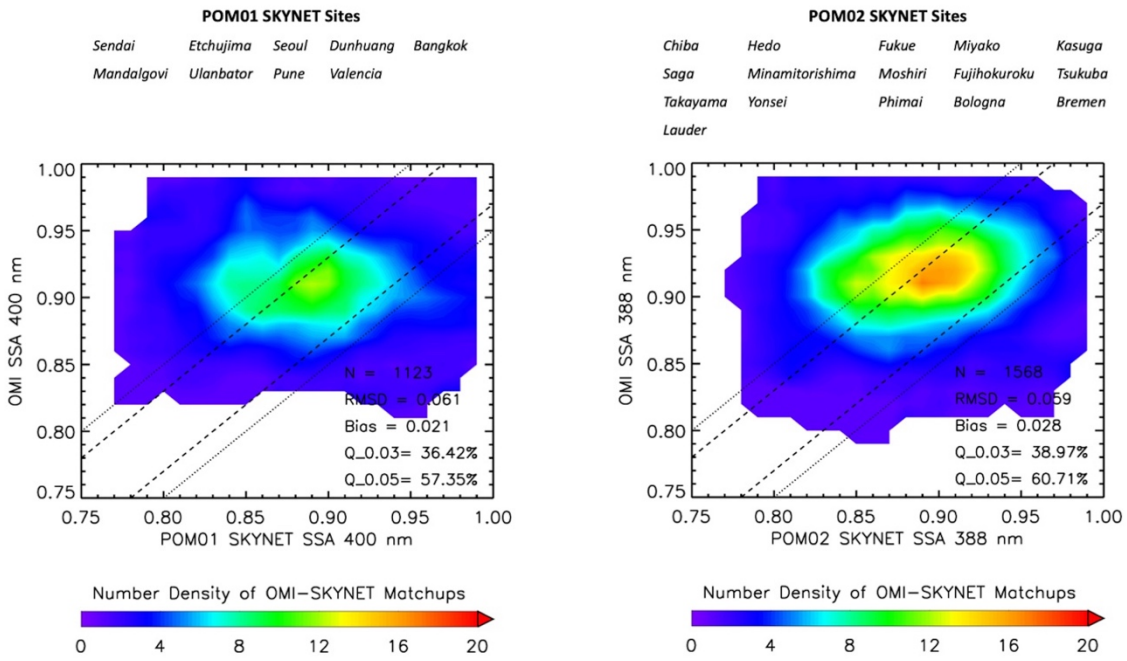

Figure 6 Composite number density contour plots of SSA comparison between OMI and SKYNET for different aerosol loading conditions. The resultant statistics of the comparison are depicted in the lower-right in both plots.

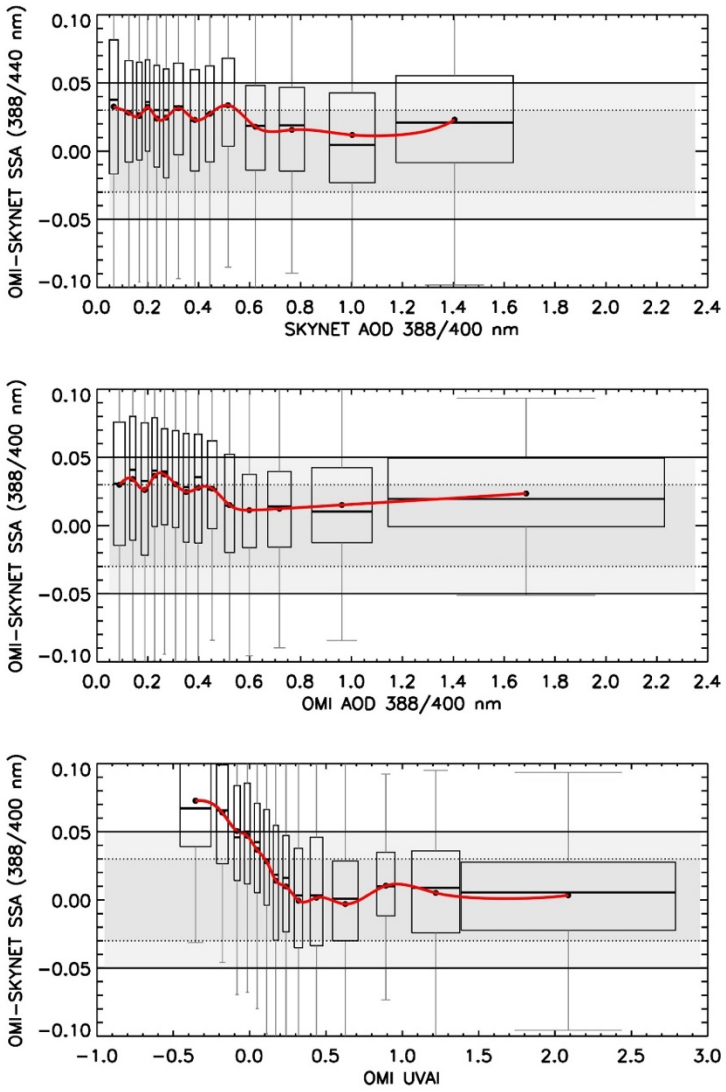

Figure ~~6~~5 7 Difference in ~~single-scattering albedo~~ SSA between OMI and SKYNET as a function of the coincident SKYNET-measured (top panel) and OMI-retrieved (middle panel) aerosol optical depth and OMI-measured UVAI (bottom panel). Filled circles in black are the mean of difference for each AOD and UVAI bin with an equal sample size of ~~150~~200 matchups; horizontal lines represent median of the bin samples; shaded area in gray encompasses data within 25 (lower) to 75 (higher) percentile range, whereas vertical lines in gray represent 1.5 times inter-quartile range (25 to 75 percentile). The dotted and solid horizontal lines are the uncertainty range of ±0.03 and ±0.05 respectively. The width of each box represents 2-standard deviation of the data contained in the respective bins.

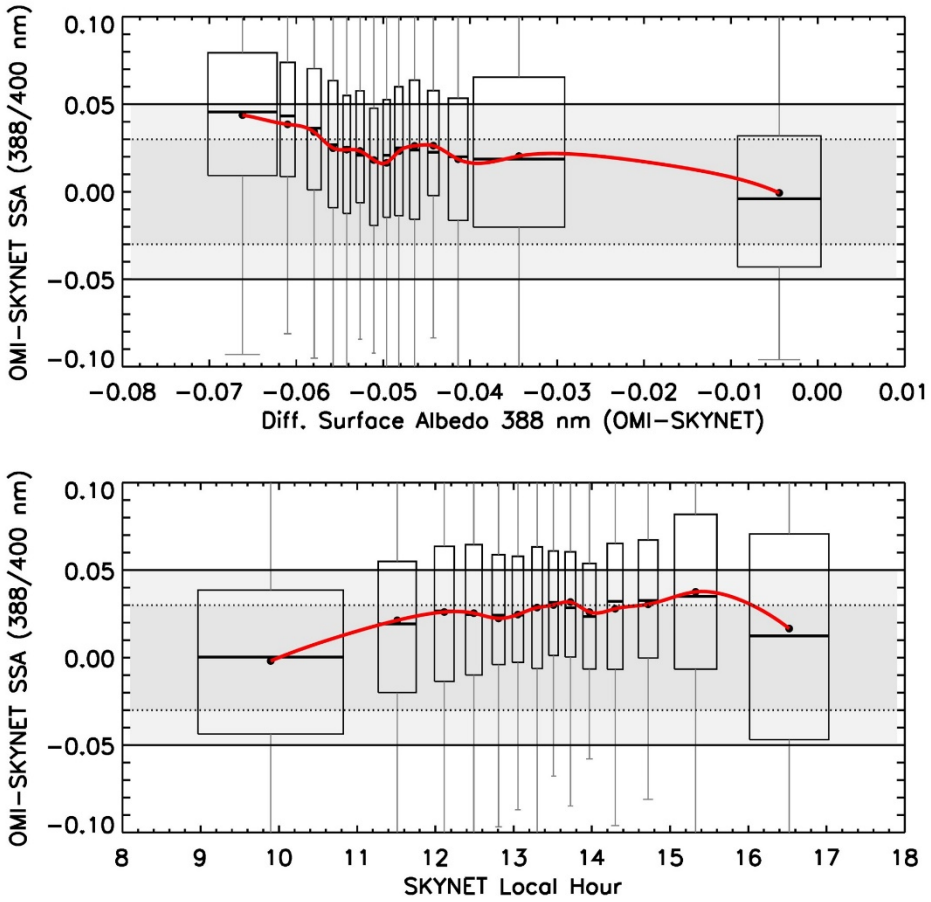

Figure ~~7~~6~~8~~ Same as in Figure ~~7~~5 but the difference in SSA between OMI and SKYNET is related to (a) the difference in surface albedo assumed by the two algorithms and (b) local measurement hour of SKYNET.