

# A Comparative Evaluation of Aura-OMI and SKYNET Near-UV Single-scattering Albedo Products

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## 24 **ABSTRACT**

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25 The aerosol single-scattering albedo (SSA) retrieved by the near-UV algorithm applied to the  
26 Aura/Ozone Monitoring Instrument (OMI) measurements (OMAERUV) is compared with an  
27 independent inversion product derived from the sky radiometer network SKYNET-a ground-  
28 based radiation observation network with sites in Asia and Europe. The present work continues  
29 previous efforts to evaluate the consistency between the retrieved SSA from satellite and  
30 ground sensors. The automated spectral measurements of direct downwelling solar flux and sky  
31 radiances made by SKYNET Sun-sky radiometer are used as input to an inversion algorithm that  
32 derives spectral aerosol optical depth (AOD) and single-scattering albedo (SSA) in the near-UV  
33 to near-IR spectral range. The availability of SKYNET SSA measurements in the ultraviolet region  
34 of the spectrum allows, for the first time, a direct comparison with OMI SSA retrievals  
35 eliminating the need of extrapolating the satellite retrievals to the visible wavelengths as the  
36 case in the evaluation against the Aerosol Robotic Network (AERONET). An analysis of the  
37 collocated retrievals from over 25 SKYNET sites reveals that about 61% (84%) of OMI-SKYNET  
38 matchups agree within the absolute difference of  $\pm 0.03$  ( $\pm 0.05$ ) for carbonaceous aerosols, 50%  
39 (72%) for dust aerosols, and 45% (75%) for urban-industrial aerosol types. Regionally, the  
40 agreement between the two inversion products is robust over several sites in Japan influenced  
41 by carbonaceous and urban-industrial aerosols; at the biomass burning site *Phimai* in Thailand,  
42 and polluted urban site in *New Delhi*, India. The collocated dataset yields fewer matchups  
43 identified as dust aerosols mostly over the site *Dunhuang* with more than half of the matchup  
44 points confined to within  $\pm 0.03$  limits. Altogether, the OMI-SKYNET retrievals agree within  
45  $\pm 0.03$  when SKYNET AOD (388 or 400 nm) is larger than 0.5 and OMI UV Aerosol Index larger  
46 than 0.2. The remaining uncertainties in both inversion products can be attributed to specific  
47 assumptions made in the retrieval algorithms, i.e., the uncertain calibration constant,  
48 assumption of spectral surface albedo and particle shape, and sub-pixel cloud contamination.  
49 The assumption of fixed and spectrally neutral surface albedo (0.1) in the SKYNET inversion  
50 appears to be unrealistic, leading to underestimated SSA, especially under lower aerosol load  
51 conditions. At higher AOD values for carbonaceous and dust aerosols, however, retrieved SSA

52 values by the two independent inversion methods are generally consistent in spite of the  
53 differences in retrieval approaches.

## 54 **1 INTRODUCTION**

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55 Satellite-based remote sensing of aerosols has become an essential tool to detect, quantify, and  
56 routinely monitor the aerosol optical and size properties over the globe. An accurate  
57 representation of aerosols in the climate models is an essential requirement for reducing the  
58 uncertainty in aerosol-related impact on the Earth's radiation balance (direct and semi-direct  
59 effects) and cloud microphysics (indirect effect) (*IPCC, 2013*). The fundamental aerosol  
60 parameters determining the strength and sign of the radiative forcing are the aerosol optical  
61 depth (AOD) and single-scattering albedo (SSA) in addition to the reflective properties of the  
62 underlying surface. While the columnar AOD represents the total extinction (scattering and  
63 absorption) resulting from the interactions with solar radiation, SSA describes the relative  
64 strength of scattering to the total extinction. Together, both AOD and SSA determine the  
65 magnitude and sign of the aerosol radiative forcing at the top-of-atmosphere. For example, a  
66 decrease in SSA from 0.9 to 0.8 can often change the sign of radiative forcing from negative  
67 (cooling) to positive (warming) that also depends on the albedo of the underlying surface and  
68 the altitude of the aerosols (*Hansen et al., 1997*). Thus, an accurate estimate of both quantities  
69 is a prime requirement for reliable estimates of the net effect of atmospheric aerosols  
70 produced with the anthropogenic as well as natural activities.

71  
72 Launched in July 2004, the Ozone Monitoring Instrument (OMI) onboard NASA's Aura satellite  
73 has produced more than a decade long global record of observations of reflected radiation  
74 from Earth in the 270–500 nm wavelength range of the spectrum on a daily basis. OMI scans  
75 the entire Earth in 14 to 15 orbits with its cross-track swath of ~2600 km at ground level at a  
76 nadir ground pixel spatial resolution of  $13 \times 24 \text{ km}^2$ . OMI observations of the top-of-  
77 atmosphere reflected light at 354 and 388 nm wavelengths are used to derive the UV aerosol  
78 index (UVAI) as well as the AOD and SSA using the OMAERUV algorithm that takes advantage of  
79 the well-known sensitivity to the aerosol absorption in the UV spectral region (*Torres et al.,*  
80 *1998*). While a general description of the OMAERUV algorithm is presented in *Torres et al.*  
81 (*2007*), the recent algorithmic upgrades are documented in *Torres et al. (2013, 2018)*. The most

82 important changes applied in the latest OMAERUV algorithm upgrade includes: 1) use of new  
83 carbonaceous aerosol models that account for the presence of organics in the carbonaceous  
84 aerosols by assuming wavelength-dependent imaginary part of the refractive index (*Jethva and*  
85 *Torres, 2011*), 2) an implementation of robust scheme to identify aerosol type (smoke, dust,  
86 urban/industrial) that combinedly uses the information on carbon monoxide (CO) observations  
87 from the Atmospheric Infrared Sounder (AIRS) and UVAI from OMI (*Torres et al., 2013*), 3) use  
88 of the aerosol height climatology dataset derived from the Cloud-Aerosol Lidar with Orthogonal  
89 Polarization (CALIOP) lidar-based measurements of the vertical profiles of aerosol for the  
90 carbonaceous and dust aerosols (*Torres et al., 2013*), and 4) better treatment of dust particles  
91 assuming realistic spheroidal shape distribution (*Torres et al., 2018*). Additionally, the upgraded  
92 OMAERUV algorithm has adopted a new method to calculate UVAI, which now accounts for the  
93 angular scattering effects of clouds and significantly reduces a scan angle related asymmetry in  
94 UVAI in cloudy scenes (*Torres et al., 2018*).

95

96 The present work continues previous efforts to evaluate the consistency between ground-  
97 based SSA measurements and satellite retrievals from near UV observations. On the first  
98 attempt to intercompare space-based and surface near UV SSA measurements, Earth Probe  
99 TOMS retrievals were compared to AERONET observations acquired during the SAFARI 2000  
100 field campaign (*Torres et al., 2005*). The OMAERUV near-UV aerosol product of AOD and SSA  
101 has been continually assessed and validated against the ground-based measurements acquired  
102 from the globally distributed Aerosol Robotic Network-AERONET (*Torres et al., 2007; Ahn et al.,*  
103 *2008; Jethva and Torres, 2011; Ahn et al., 2014; Jethva et al., 2014*). While the OMAERUV AOD  
104 product was directly validated against the AERONET measurements made in the near-UV (340-  
105 380 nm), as carried out in *Ahn et al. (2014)*, the SSA retrievals have been evaluated by  
106 comparison with the AERONET ground inversion product (*Jethva et al., 2014*). The latter  
107 analysis required OMI retrievals of SSA to be extrapolated to the shortest visible wavelength of  
108 440 nm of AERONET inversion product to make the comparison possible. Such adjustment in  
109 the wavelength of retrievals can introduce uncertainty in the comparison arising from the  
110 inaccuracy of the spectral dependence of absorption assumed in the wavelength conversion.

111  
112 A direct comparison of the column-integrated SSA at 388 nm retrieved from OMI requires  
113 equivalent ground-based columnar retrievals in the near-UV region. The international network  
114 of scanning sun-sky radiometers (SKYNET) fulfills this requirement as it performs the direct Sun  
115 and sky measurements in the near-UV (340-380 nm) as well as visible/near-IR (400-1020 nm)  
116 regions of the spectrum and derives spectral AOD and SSA. Taking advantage of the availability  
117 of ground-based SSA inversions in the near-UV from SKYNET, we inter-compare the OMI and  
118 SKYNET SSA products at several SKYNET sites in Asia and Europe. Since both retrieval  
119 approaches are based on inversion algorithms that rely on assumptions, the resulting level of  
120 agreement can only be interpreted as a measure of consistency (or lack thereof) in the  
121 measurement of the same physical parameter by fundamentally different remote sensing  
122 approaches.

123  
124 The paper is organized as follows: Section 2 describes the satellite and ground-based data sets  
125 assessed in this analysis along with the collocation methodology; the results of OMI-SKYNET  
126 SSA comparison over individual sites, combinedly for each aerosol type, and diagnosis of  
127 differences between them are presented in section 3; the possible sources of uncertainty in  
128 both inversion products are discussed in section 4; the paper is summarized and concluded in  
129 section 5.

130

## 131 **2 DATASETS**

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### 132 **2.1 THE OMI-OMAERUV AEROSOL PRODUCT**

133 The entire record of OMI observations (October 2004 to present) has been reprocessed  
134 recently with the refined OMAERUV algorithm (PGEVersion V1.8.9.1) to derive a comprehensive  
135 aerosol product that includes retrievals of the UV Aerosol Index (UVAI), AOD, SSA, and AAOD  
136 (388 nm) at a pixel resolution of 13 x 24 km<sup>2</sup> at nadir viewing geometry. The retrieved

137 parameters are also reported at 354 nm and 500 nm wavelengths following the spectral  
138 dependence of aerosols assumed in the chosen model. The data set is available in the HDF-  
139 EOS5 format and can be obtained at no cost from NASA Goddard Earth Sciences (GES)-Data and  
140 Information Services Center (DISC) server at <http://daac.gsfc.nasa.gov/>. The recent upgrade has  
141 been documented in detail in the work of *Jethva and Torres (2011)*, *Torres et al. (2013, 2018)*  
142 and *Ahn et al. (2014)*. Here, we use the OMAERUV Level 2 Collection 003 (V1.8.9.1) aerosol  
143 product processed in July 2017. The expected uncertainty limits in the OMAERUV SSA retrievals  
144 are determined to be  $\pm 0.03$  and  $\pm 0.05$ , based on its comparison with AERONET SSA inversion  
145 and sensitivity analysis carried out during the development of the OMAERUV algorithm (*Torres*  
146 *et al., 2007*). Following an early evaluation of OMI aerosol product for a handful of sites, *Jethva*  
147 *et al. (2014)* conducted a global evaluation of SSA product and also carried out a detailed  
148 uncertainty test considering different sources of errors, such as aerosol model, surface albedo,  
149 and aerosol layer height. The results of the sensitivity analysis further confirmed the  
150 uncertainty budget estimated earlier during the early development of the OMAERUV algorithm.  
151 However, note that the errors could attain larger magnitudes when algorithmic assumptions  
152 are far off from the real atmospheric conditions.

153

154 Post-2007, the OMI observations have been affected by a possible external obstruction that  
155 perturbs both the measured solar flux and Earth radiance. This obstruction affecting the quality  
156 of radiance at all wavelengths for a particular viewing direction is referred to as “row anomaly”  
157 (RA) since the viewing geometry is associated with the row numbers on the charge-coupled  
158 device detectors. The RA issue was detected first time in mid-2007 with a couple of rows, which  
159 during the later period of operation expanded to other rows in 2008 and later. At present,  
160 about half of the total 60 rows across the track are identified and flagged as row anomaly  
161 affected positions for which no physical retrievals are performed (*Schenkeveld et al., 2017*). The  
162 details about this issue can be found at  
163 <http://www.knmi.nl/omi/research/product/rowanomaly-background.php>. The RA has  
164 significantly affected the sampling during post-2008 OMI measurements, where row anomaly  
165 flags blanket about half of the OMI swath. As a result, the availability of the number of

166 retrievals since 2009 over a particular site is reduced. Therefore, the OMI-SKYNET matchups are  
167 also expected to be lower during the row anomaly affected period. The OMAERUV algorithm  
168 assigns quality flags to each pixel which carries information on the quality of the retrieval  
169 depending upon the observed condition. We used aerosol retrievals free of RA and flagged as  
170 quality flag '0', which are considered reliable due to higher confidence in detecting aerosols in a  
171 scene with minimal cloud contamination.

## 172 **2.2 THE SKYNET AEROSOL INVERSION PRODUCT**

173 The SKYNET is an international network of scanning sun-sky radiometers (manufactured by  
174 *Prede Co. Ltd.*, Japan) performing routine and long-term measurements of direct and diffuse  
175 solar radiations at several wavelengths spanning UV (340 and 380 nm), visible (400, 500, 675  
176 nm), near-IR region (875, 1020 nm), and in shortwave-IR (1627 nm and 2200 nm) of the  
177 spectrum. The automated measurements of direct and diffuse solar radiations are used to  
178 measure spectral AOD and retrieve SSA and other aerosol optical-microphysical properties  
179 (volume size distribution, refractive index, phase function, and asymmetry parameter) at the  
180 same standard wavelengths of AOD following an inversion algorithm packaged in the  
181 *SKYRAD.pack* software (*Nakajima et al.*, 1996; *Hashimoto et al.*, 2012). Cloudy observations are  
182 screened using the Cloud Screening Sky Radiometer code (*Khatri and Takamura*, 2009).

183

184 The SKYNET radiometers come in two flavors, model POM-01 and model POM-02. The POM-01  
185 instrument carries a total of five wavelength filters covering visible to near-IR (400-1020 nm),  
186 whereas POM-02 instrument has two additional filters in the UV region (340 and 380 nm) along  
187 with the other filters in the visible to shortwave-IR (including 1627 nm and 2200 nm) part of the  
188 spectrum. The calibration of each SKYNET radiometer is performed on-site on a monthly basis  
189 using the improved Langley method (*Nakajima et al.*, 1996; *Campanelli et al.*, 2004, 2007).  
190 Occasionally, the inter-calibration of radiometers is carried out against the master instrument  
191 well-calibrated using the Langley method on a high mountain site, e.g., Mauna Loa. The SKYNET  
192 radiometers are also inter-compared with AERONET Cimel Sunphotometers and precision filter



193 radiometers at three observation sites, i.e., *Chiba University, Valencia (Estelles et al., 2016), and*  
194 *Rome (Campanelli et al., 2018).*

195

196 Studies in the past have compared AODs (*Estellés et al., 2012a*) and SSAs (*Estellés et al., 2012b*)  
197 measured/retrieved from SKYNET and AERONET and shown that AODs are well-correlated and  
198 in good agreement, but the SKYNET SSAs are found to be higher than those of AERONET (*Che et*  
199 *al., 2008; Hashimoto et al., 2012*). *Khatri et al. (2016)* further pinpoints the factors, such as  
200 quality of input data attributed to different calibration and observation protocols, different  
201 quality assurance criteria, the calibration constant for sky radiances, differences in measured  
202 AOD, and surface albedo, responsible for the inconsistent aerosol SSA between AERONET and  
203 SKYNET using observations from the four representative sites, i.e., Chiba (Japan), Pune (India),  
204 Valencia (Spain), and Seoul (South Korea). More discussion on the sources of uncertainties is  
205 presented in section 4.

206

207 In this study, we include the SKYNET data acquired over a total of 25 sites distributed mostly  
208 across Asia and a few in Europe. The dataset is freely accessible from the data portal of the  
209 Center for Environmental Remote Sensing (CERes), Chiba University, Japan  
210 (<http://atmos3.cr.chiba-u.jp/skyenet/data.html>). Figure 1 shows the geographic distribution of  
211 selected sites, whereas Table 1 lists the geo-coordinates of these sites with the associated  
212 sensor type (POM-01 or POM-02) and data periods. The SKYNET aerosol product is derived  
213 using two different Skyrad packs: version 4.2 and version 5, the differences of which are  
214 explained in *Hashimoto et al. (2012)*. In this study, we use the SKYNET Level 2 product retrieved  
215 using version 5 of Skyrad pack. SKYNET retrievals assigned with cloud flag '0' are included in the  
216 analysis since these measurements are believed to be free of cloud contamination considered  
217 as higher quality retrievals. A careful examination of the SKYNET inversion dataset revealed  
218 some irregularities in the measurements for many sites, such as irregular patterns in the shape  
219 of spectral SSAs, identical values of SSA at near-UV and visible wavelengths, and much larger  
220 standard deviation ( $>0.1$ ) in SSA within a few hours. These spurious measurements were  
221 excluded from the present analysis.

### 2.3 THE COLLOCATION OF OMI AND SKYNET MEASUREMENTS

OMI retrievals correspond to a spatial scale of  $13 \times 24 \text{ km}^2$  at nadir representing the atmospheric conditions over an area. Unlike the direct measurements of the spectral AOD, which correspond to columnar point measurements, the retrievals made by SKYNET use the sky radiances measured at several discrete angles azimuthally, therefore representing the sky condition observed over a station which is associated with approximately 5 km radius surrounding the Sun photometer site. SKYNET retrieves aerosol optical-microphysical properties, including spectral SSA, under all cloud-free conditions and at all aerosol loadings. It is expected that the inversion of retrieved parameters from sky radiances offers better accuracy at larger solar zenith angles owing to the longer optical path and better aerosol absorption signal (*Dubovik et al., 2000*). These conditions are best satisfied with the measurements made during the early morning and late afternoon hours. On the other hand, Aura/OMI overpasses a station during the afternoon hours with the local equator-crossing time 1:30 P.M. Therefore, the collocation of the measurements was carried out within a time window of  $\pm 3 \text{ h}$  around OMI overpass time to get sufficient high-quality SKYNET retrievals particularly from early morning/late afternoon measurements. The OMI retrievals of SSA were spatially averaged in a grid area of  $0.5^\circ$  by  $0.5^\circ$  centered at the SKYNET site. Though the spatial averaging area for the OMI retrieval is about  $50 \text{ km}^2$ , due to its larger footprint, the actual area intercepted by OMI pixels around the SKYNET site is likely to be larger.

OMI performs retrieval at 354 nm and 388 nm wavelengths, whereas the SKYNET POM-02 instrument reports SSA at nearby wavelengths of 340, 380, and 400 nm. To compare both SSA products at the same wavelength, SKYNET SSA was linearly interpolated at 388 nm, to match with the wavelength of OMI retrieval, using the measurements at the two nearest wavelengths, i.e., 380 nm and 400 nm. The SKYNET POM-01 instruments don't carry UV wavelength filters, but report the retrievals at the shortest wavelength 400 nm and other visible/near-IR wavelengths. In this case, the OMI retrievals are extrapolated from 388 nm to 400 nm, to match with the wavelength of SKYNET inversion, following the spectral dependence of SSA associated

250 with the chosen aerosol model in the OMI algorithm. It is reasonably fair to assume that the  
251 extrapolation of OMI SSA in a narrow window of 12-nm, i.e., from 388 to 400 nm, should not be  
252 a major source of uncertainty in comparing SSA from OMI and SKYNET.

### 253 **3 RESULTS**

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#### 254 **3.1 OMI-SKYNET COMPARISON OVER INDIVIDUAL STATIONS**

255 Figure 2 displays the OMAERUV versus SKYNET SSA scatterplots for selected sites in Japan. The  
256 comparison was made at 388 nm or 400 nm depending upon the availability of the SKYNET  
257 inversion at those wavelengths, i.e., POM-01 or POM-02 sensors. Legends with different colors  
258 represent the aerosol type selected by the OMAERUV algorithm for the co-located matchups  
259 (N). RMSD is the root-mean-square difference between the two retrievals; Q\_0.03 and Q\_0.05  
260 are the percent of total matchups (N) that fall within the absolute difference of 0.03 and 0.05,  
261 respectively; horizontal and vertical lines for each matchup are the standard deviation of  
262 temporally and spatially averaged SKYNET and OMI SSAs. The comparison includes OMI-SKYNET  
263 matchups with AOD>0.3 (388 or 400 nm) in both measurements simultaneously. The  
264 scatterplots reveal a good level of agreement for matchups identified as carbonaceous/smoke  
265 aerosols over *Chiba University, Cape Hedo, Fukue, Saga, and Etchujima* with the majority of  
266 points confined within the absolute difference of 0.03. The OMI-SKYNET combined dataset is  
267 dominated with matchup points identified as the urban/industrial aerosols in the OMAERUV  
268 algorithm for which the measured UVAI falls below 0.5 representing lower aerosol loading in  
269 the boundary layer with weakly absorbing properties. Under such observed conditions, the  
270 uncertainties in both kinds of measurements are prone to be larger due to lower absorption  
271 signal relative to the instrumental noise and errors in algorithmic assumptions, such as surface  
272 albedo, that could further amplify the overall uncertainty in the retrievals. Despite these  
273 inherent uncertainties, an agreement within the difference of  $\pm 0.03$  for more than half of the  
274 collocated retrievals is encouraging.

275

276 Figure 3 shows the scatterplots of OMI-SKYNET SSA for remaining sites located in South Korea,  
277 China, Thailand, India, and Italy. For the site *Seoul* in South Korea, OMI tends to overestimate  
278 SSA for a number of matchups assigned as the urban/industrial aerosols and for a few with the  
279 carbonaceous/smoke aerosol type such that about 42% of total matchups are falling within the  
280 difference of 0.03. For the *Dunhuang* site located in the desert area of China, a majority of  
281 collocated data points were identified as dust aerosol type providing an overall better  
282 agreement with 50% and 68% matchups bounded within  $\pm 0.03$  and  $\pm 0.05$  differences,  
283 respectively. The *Phimai* site in Thailand is known to be influenced by the springtime biomass  
284 burning activities, where OMI and SKYNET SSAs are found to agree relatively best among all 25  
285 sites providing 71% and 91% of the matchups restricted within  $\pm 0.03$  and  $\pm 0.05$  limits,  
286 respectively. The agreement between the two sensors was robust for the carbonaceous/smoke  
287 aerosol type followed by the urban/industrial aerosols. Over the megacity of New Delhi in the  
288 Indo-Gangetic Plain in India, which is seasonally influenced by the smoke and desert dust  
289 aerosols in addition to the local source of urban pollution, the OMI-SKYNET matchups are found  
290 to agree within  $\pm 0.03$  and  $\pm 0.05$  for 52% and 83% of the evaluated data points respectively.  
291 Over the *Pune* station located near the western boundary of India and the *Bologna* site in Italy,  
292 OMI retrieves higher SSA compared to that of SKYNET yielding 39% and 64%, and 25% and 50%  
293 matchups, respectively, within the two uncertainty limits. Table 1 lists the statistical measures  
294 of the OMI-SKYNET SSA comparison for all 25 sites. A more detailed description of the different  
295 sources of uncertainty is presented in section 4.

### 296 **3.2 COMPOSITES FOR EACH AEROSOL TYPE**

297 Figure 4 displays the composite scatterplots of OMI versus SKYNET SSA derived by segregating  
298 the matchup points for each aerosol type from all 25 sites. The intention here is to evaluate the  
299 consistency between the two retrieval methods for each aerosol type separately and  
300 understand their relative differences. When identified as the carbonaceous/smoke aerosol type,  
301 the OMI-SKYNET matchups reveal relatively best comparison among the three major aerosol  
302 types with 61% and 84% data points falling within the absolute difference of 0.03 and 0.05,  
303 respectively, and providing the lowest (0.035) root-mean-square-difference between the two

304 retrievals. The collocation procedure yields the lowest number of matchups (N=32) for desert  
305 dust aerosol type obtained mostly over the site of *Dunhuang* in China, resulting in 50% and 72%  
306 of data points within the stated uncertainty limits. Among the three aerosol types, the  
307 collocated points assigned with the urban/industrial aerosol type (Figure 4 bottom-left) yield  
308 the maximum number of matchups (N=739) with the relatively weakest agreement  
309 (RMSD=0.052), where OMI tends to overestimate SSA for a significant number of instances  
310 resulting about 45% and 67% data points falling within the two limits of expected uncertainties.  
311 When more than one prescribed aerosol types are selected for OMI pixels around the SKYNET  
312 stations, the matchups between the two sensors resulted in 59% and 77% retrievals within the  
313 uncertainty limits with an RMSD of 0.041—a comparison slightly poorer than ‘smoke-only’ case,  
314 but better than ‘dust-only’ and ‘urban/industrial-only’ retrieval cases. Combined, all three  
315 distinct aerosol types simultaneously yield the total number of matchups (N=1223) with an  
316 RMSD of 0.047 between OMI and SKYNET, resulting 51% and 72% collocated data points falling  
317 within the absolute difference of 0.03 and 0.05 difference, respectively. When the restriction of  
318  $AOD > 0.3$  is removed from the collocation procedure, allowing all matchups regardless of their  
319 respective AOD values, the total number of collocated data points was increased to more than  
320 twice (N=2691) albeit with a relatively weaker agreement yielding an RMSD of 0.06 and percent  
321 data points within the uncertainty limits reducing to 38% and 59%, respectively.

322

### 323 **3.3 COMPOSITES FOR VARYING AEROSOL LOADING AND POM-01 VERSUS POM-02**

324 Figure 5 shows the number density plots comparing OMI-SKYNET SSA matchups obtained from  
325 all sites combined and for varying aerosol loading conditions. The best set of comparison is  
326 achieved under the most restrictive scenario when corresponding OMI-retrieved AOD and UVAI  
327 are constrained to  $>0.3$  and  $>0.5$ , respectively, albeit with a significantly reduced number of  
328 matchups compared to the other two cases with lesser (middle) or no (left) restrictions. The  
329 improved comparison reflected in statistical parameters is a result of avoiding retrievals with  
330 lower aerosol loading when both kinds of measurements might be subjected to larger  
331 uncertainties due to algorithmic assumptions.

332

333 Figure 6 shows the number density plot comparing SSA between SKYNET and OMI for POM-01  
334 and POM-02 sensors separately. Overall, no major difference is noticed in the derived statistics  
335 between the two sets of comparison, except that the number of matchups obtained with  
336 POM02 sensor is 39% more than those with POM01 sensors, and POM02 dataset offers  
337 marginally better comparison (except bias, which is higher with POM02) with OMI SSA  
338 retrievals. This analysis indicates that the interpolation of OMI SSA from 388 nm to 400 nm for  
339 its comparison with POM01 data isn't a significant source of discrepancy between the two SSA  
340 datasets.

341

### 342 **3.4 DIAGNOSIS OF OMAERUV VERSUS SKYNET SSA**

343 The SKYNET algorithm inverts the spectral sky radiances in conjunction with the direct AOD  
344 measurements to retrieve the real and imaginary parts of the refractive index and particle size  
345 distribution of cloud-free observations under all aerosol loading conditions. These inversion  
346 products are believed to be more stable and accurate at higher aerosol loadings and solar  
347 zenith angles due to stronger aerosol absorption signal and longer optical path (*Dubovik et al.*,  
348 2000). Similarly, a sensitivity analysis of the two-channel OMAERUV retrievals suggests that the  
349 retrieved AOD and SSA are susceptible to the small change in surface albedo at lower aerosol  
350 loading (*Jethva et al.*, 2014). For instance, an absolute difference of 0.01 in the surface albedo  
351 leads to a change in AOD approximately by 0.1 and SSA by  $\sim 0.02$ .

352

353 Figure 7 (top) shows the absolute difference in collocated SSA between OMI and SKYNET as a  
354 function of concurrent SKYNET direct AOD (388 or 400 nm) measurements for all aerosol types.  
355 All OMI-SKYNET matchup data obtained from a total of 25 sites under all AOD conditions are  
356 included here. The data are shown in the box and whisker format, where the horizontal lines  
357 represent the median value of each bin of sample size 150, filled circle the mean value, and  
358 shaded vertical bars cover the 25 and 75 percentiles of the population in each data bin. While  
359 for most bins the mean and median values of SSA difference were restricted to within  $\pm 0.03$ ,

360 OMI tends to overestimate SSA relative to that of SKYNET at lower AODs giving larger  
361 differences and spread in the data population. Similar patterns were observed when the  
362 difference in SSA was related to the OMI-retrieved AOD (Figure 7 middle). In both cases, the  
363 differences in SSA minimize at larger AOD values ( $>0.5$ ) suggesting a convergence in both  
364 retrievals. Figure 7 (bottom) shows a similar plot of SSA difference against the concurrent OMI  
365 UVAI. Notably, the differences in SSA exhibit even a stronger relationship to UVAI than that in  
366 the AOD case (top and middle). For UVAI lesser than zero, the differences in the retrieval are  
367 found to be beyond the expected uncertainty in both inversions, at least in the mean sense. For  
368 the lower range of UVAI, OMI algorithm mostly employs the urban/industrial model for the  
369 retrieval where all aerosols are assumed to be confined within the boundary layer ( $<2$  km) with  
370 a vertical profile that follows an exponential distribution. On the other hand, the mean and  
371 median values of the SSA difference for UVAI larger than 0.2 for all bins fall within the 0.03  
372 uncertainty range. The SSA differences approach to near-zero with a reduced spread at larger  
373 magnitudes. Notably, both inversions are found to be in closer agreement for UVAI  
374 measurements  $>0.3$ .

375

376

## 377 **4 SOURCES OF UNCERTAINTY**

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### 378 **4.1 UNCERTAINTIES IN THE GROUND-BASED SKYNET INVERSION PRODUCT**

379 The standard SKYNET inversion algorithm assumes a wavelength-independent surface albedo of  
380 0.1 at all wavelengths across the UV to visible part of the spectrum. However, the algorithm  
381 code allows flexibility to alter the value surface albedo in time and wavelength (Campanelli et  
382 al., 2015). The diffuse light reflected from the ground plays a second-order role in the measured  
383 sky radiances in most situations, however, has a potential to affect the SSA inversion, e.g.,  
384 overestimated (underestimated) surface albedo can underestimate (overestimate) SSA  
385 (*Dubovik et al.*, 2000; *Khatri et al.*, 2012). Using simultaneous inversion data from SKYNET and  
386 AERONET for four representative sites, *Khatri et al.* (2016) have shown that the difference in

387 the prescribed surface albedo between SKYNET and AERONET results in a difference of  $\sim 0.04$  in  
388 SSA at red (675 nm) and near-IR wavelengths retrieved from the two collocated ground sensors.  
389 The difference in SSA can also reach as large as  $\sim 0.08$  when surface albedo differed by 0.3. The  
390 assumed surface albedo value of 0.1 at near-UV (340 and 380 nm) and shorter visible  
391 wavelength (400 nm) seems to be unrealistic for the vegetated and urban surfaces. The surface  
392 albedo database at 354 nm and 388 nm derived from multiyear observations from OMI  
393 suggests that the vegetated surfaces and urban centers are characterized with the lower values  
394 of surface albedo, i.e.,  $\sim 0.02$ - $0.03$  and  $\sim 0.05$ , respectively; for desert surfaces, the albedo could  
395 be as high as 0.08-0.10. Significant differences in the assumed surface albedo values between  
396 OMI and SKYNET at shorter wavelengths could be one of the responsible factors for  
397 discrepancies in SSA noted over several sites, particularly at lower aerosol loading when the  
398 uncertainty in surface characterization can amplify error in the SSA inversion.

399

400 To further investigate this effect, the difference in SSA between OMI and SKYNET as a function  
401 of the simultaneous difference in surface albedo is analyzed and shown in Figure 8 (top). The  
402 data are presented in a standard box and whisker plot format. The analysis reveals a link  
403 between differences in SSA and surface albedo, where increasing differences in SSA  
404 (OMI>SKYNET) are associated with significant negative biases in surface albedo between OMI  
405 and SKYNET. In other words, large overestimation in SKYNET surface albedo causes  
406 underestimation of retrieved SSA, which is consistent with the findings of *Dubovik et al.* (2000)  
407 and *Khatri et al.* (2012, 2016), thereby resulting in a substantial positive difference in SSA  
408 between OMI and SKYNET. Recently, *Mok et al.* (2018) have shown that the use of AERONET  
409 surface albedo dataset at 440 nm in the SKYNET algorithm for the S. Korea region produces SSA  
410 values larger by  $\sim 0.01$  at near-UV wavelengths. Notably, differences in SSA tend to be lower  
411 when the differences in surface albedo are also minimal, such that the mean and median values  
412 of those bins remain within the expected uncertainties of  $\pm 0.03$  in both retrievals. This result,  
413 along with the previous findings cited above, convincingly points out that the SSA inversion  
414 from ground-based sensors, especially at lower aerosol loadings, is likely susceptible to the  
415 prescribed surface albedo. The assumption of a fixed value of spectral surface albedo of 0.1 in



416 the SKYNET algorithm appears to be inappropriate requiring for a revision using more accurate  
417 datasets of spectral reflectance or albedo such as from MODIS and OMI.

418

419 The dependence of SSA difference on the local hour of SKYNET measurements is quantified in  
420 Figure 8 (bottom). The SKYNET dataset accessed from the data server at Chiba University  
421 doesn't contain information on the solar zenith angle. However, the local time of  
422 measurements reported in the data file for each station can serve a proxy for the solar zenith  
423 angle. The OMI-SKYNET matchups exhibit a systematic dependency, where the differences  
424 between the two datasets become relatively minimal when early morning and late afternoon  
425 inversions of SKYNET associated with higher solar zenith angle are collocated with OMI  
426 overpass time around 1:30 PM equator-crossing time. Owing to a longer atmospheric optical  
427 path at higher solar zenith angles, thereby better aerosol absorption signal, the ground-based  
428 aerosol inversions, such as from AERONET and SKYNET, are expected to be more reliable for sky  
429 measurements carried out during early morning/late afternoon.

430

431 SKYNET inversion algorithm (Skyrad.pack Version 4.2 and version 5) assumes aerosols of  
432 spherical shape regardless of the actual aerosol type observed in the scene. Following a  
433 detailed analysis of the effect of non-sphericity of the particles on the difference between the  
434 retrievals carried out assuming spherical and spheroidal size distribution, *Khatri et al.*, (2016)  
435 concluded that the assumed shape of particles has a non-significant impact on the retrieved  
436 SSA. Their study revealed SSA difference of  $\pm 0.01$  for measurements having a maximum  
437 scattering angle  $< 120^\circ$  and difference of up to  $\pm 0.02$  at scattering angle  $> 120^\circ$ , where the  
438 difference in the phase function is significant between spherical and spheroidal size  
439 distributions (*Torres et al.*, 2018). The OMI-SKYNET collocation procedure, as shown in Figure 4,  
440 yields relatively fewer matchups that are identified as dust aerosol type according to the  
441 OMAERUV aerosol type identification scheme. A majority of the collocated data points were  
442 derived over the desert site of *Dunhuang* in China showing a reasonable agreement in SSA  
443 between OMI and SKYNET for dust aerosols further supporting the findings of *Khatri et al.*

444 (2016) that the SSA retrievals are not significantly impacted by the assumption of the shape of  
445 particles, i.e., spherical or spheroidal.

446

447 Apart from the algorithmic assumptions, the calibration constant used for sky radiances  
448 measured by SKYNET instruments can be a potential source of errors in the inversion. *Khatri et*  
449 *al.* (2016) study suggests that the calibration constant for sky radiances determined from the  
450 disk scan method using a solar disk scan area of  $1^\circ \times 1^\circ$  (*Boi et al.*, 1999) may be  
451 underestimated resulting in overestimated sky radiance and thus relatively higher SSA. Some of  
452 the larger differences between in SSA between OMI and SKYNET, where OMI underestimates  
453 SSA relative to the SKYNET, can be attributed to the imperfect calibration applied to the SKYNET  
454 sensors.

#### 455 **4.2 POSSIBLE SOURCES OF UNCERTAINTIES IN OMAERUV RETRIEVALS**

456 Like other satellite-based remote sensing algorithms, OMAERUV also relies on assumptions  
457 about the atmospheric and surface properties for the retrieval of aerosol properties. The single  
458 largest known source of error in the OMI retrievals is the subpixel cloud contamination within  
459 the OMI footprint. Given the footprint of size  $13 \times 24 \text{ km}^2$  for near-nadir pixels, which intercept  
460 an area of about  $338 \text{ km}^2$  on the ground, the presence of subpixel clouds may not be avoided  
461 entirely. Currently, the algorithm assigns quality flags to each pixel, which carries information  
462 on the quality of the retrieval depending upon the observed conditions (*Torres et al.*, 2013).  
463 Aerosol retrievals with the quality flag '0' are considered reliable as this category of flag scheme  
464 largely avoids cloud-contaminated pixels by choosing the appropriate thresholds in reflectivity  
465 and UVAI measurements.

466

467 Over the desert regions, e.g., the *Dunhuang* SKYNET site in China, the frequency of occurrence  
468 of clouds is expected to be minimal. Therefore, it is less likely that the SSA retrievals over these  
469 sites are affected by cloud contamination. A reasonable agreement between the two retrieval  
470 datasets (Figure 3) supports this assumption. The quality flag scheme, however, cannot entirely

471 rule out the presence of small levels of subpixel cloud contamination or the presence of thin  
472 cirrus in the OMI footprint, which can cause overestimation in the retrieval of SSA, such as  
473 noted over the SKYNET sites in *Kasuga, Etchujima, Seoul, Bologna, and Pune*. Larger  
474 uncertainties observed over these sites are associated with the urban-industrial aerosol type,  
475 possibly because the AODs for this aerosol type are the lowest in the analysis, and therefore,  
476 subject to the less sensitivity to absorption and possibly more affected by sub-pixel cloud  
477 contamination.

478

479 Another possible source of uncertainty can be the assumption of the aerosol layer height. The  
480 climatology of aerosol layer height derived from CALIOP measurements adequately describes  
481 the observed mean layer of carbonaceous and desert dust aerosols (*Torres et al., 2013*). It is  
482 particularly robust over the arid and semiarid areas where large numbers of cloud-free  
483 observations were used in the calculation. However, note that the temporal and spatial  
484 coverage of CALIOP is limited to 16-day repeat cycle over the same location. Variations in the  
485 aerosol layer height not observed by CALIOP, therefore, will be missed out in the derived  
486 climatology and thus can be a source of uncertainty. Sensitivity analysis of the OMAERUV  
487 retrievals suggests that an overestimation (underestimation) in the aerosol layer height results  
488 in an overestimated (underestimated) SSA. This is because an increase (decrease) in the  
489 assumed aerosol layer height from the actual one enhances (reduces) absorption in the  
490 radiance look-up table (not in the actual TOA measurements), which the OMAERUV algorithm  
491 compensates by retrieving lower (higher) AOD and higher (lower) SSA to match with the  
492 observations.

493

494 The third source of uncertainty that can affect SSA retrieval is the accuracy of the prescribed  
495 surface albedo. For the surface characterization, the OMAERUV algorithm use a near-UV  
496 surface albedo database derived using the multiyear OMI reflectivity observations. The method  
497 adopts a minimum reflectivity approach, ensuring minimal or no contamination from the  
498 atmosphere, i.e., aerosols and clouds, in the measurements. Afterward, the minimum

499 reflectivity dataset derived from the OMI observations was adjusted in the temporal domain to  
500 the seasonality of surface albedo retrieved in the visible wavelengths from MODIS. The dataset  
501 contains surface albedo values at 354 and 388 nm at a grid resolution of  $0.25^\circ \times 0.25^\circ$ .  
502 Compared to the previous OMAERUV dataset using TOMS-based surface albedo product at  $1^\circ$   
503 grid resolution, the new OMI-based dataset is expected to be more accurate to within 0.005 to  
504 0.01 owing to its higher spatial resolution and the fact that it is contemporary to the OMI  
505 operation. A sensitivity study of the OMAERUV retrievals to the change in surface albedo  
506 described in *Jethva et al. (2014)* suggests that an increase in surface albedo by 0.01 in the near-  
507 UV region over desert areas results in a decrease in the magnitude of retrieved SSA by  $\sim -0.02$ .  
508 The effect of uncertain surface albedo can be more pronounced at lower aerosol loading,  
509 where the reduced signal from the atmosphere makes OMAERUV retrieval more susceptible to  
510 the uncertainty in surface albedo.

511

512 The assumed aerosol microphysical and optical properties could be additional sources of  
513 uncertainty. The particle size distributions assumed in the OMAERUV models are adopted from  
514 long-term AERONET inversion statistics (*Dubovik et al., 2002*), representing areas influenced by  
515 smoke, dust, and urban/industrial aerosols, and therefore are considered realistic  
516 representations of the total atmospheric column. The carbonaceous smoke aerosols are  
517 assumed to be spherical in shape with a bimodal log-normal size distribution and characterized  
518 with a steep absorption gradient, such that the Absorption Angstrom Exponent (AAE) in the  
519 near-UV lies in the range 2.5-3.0, to adequately represent the organics in the biomass burning  
520 smoke particles (*Kirchstetter et al., 2004; Jethva and Torres, 2011*). The desert dust aerosol  
521 model follows bimodal log-normal size distribution with particles comprised of randomly  
522 oriented spheroids with an axis ratio (shape factor) distribution adopted from *Dubovik et al.*  
523 (*2006*). The sensitivity study followed by an actual inversion of OMI data presented in *Torres et*  
524 *al. (2018)* demonstrates that the change in dust particle shape from spherical to spheroidal  
525 distribution improved the AOD retrievals significantly and brought the equivalence between the  
526 retrievals over left and right sides of the OMI swath for the oceanic dust belt region of the  
527 tropical Atlantic. The associated changes in SSA retrievals were noted within  $\pm 0.01$  and  $-0.02$  for

528 the scattering angle up to  $100^{\circ}$ - $150^{\circ}$  and  $>160^{\circ}$ , respectively. The OMAERUV version 1.8.9.1  
529 data product used in the present study adopts a new dust model representing spheroidal  
530 particles based on the work of Dubovik et al. (2006) and Torres et al. (2018). The spectral  
531 dependence of the refractive index in the near-UV assumed in the dust aerosol model is  
532 generally consistent with the in-situ laboratory measurements (*Wagner et al.*, 2012). For  
533 instance, retrieval of AOD and SSA for carbonaceous aerosols using the smoke model with AAE  
534 of 1.90 (10% relative spectral dependence in the imaginary index between 354 and 388 nm)  
535 and 1.0 (no spectral dependence in the imaginary index), instead of the standard AAE  
536 assumption of 2.7, results in a decrease in SSA up to -0.07, respectively, suggesting a marked  
537 sensitivity of the SSA retrieval to the significant changes in the spectral aerosol absorption. Due  
538 to the shortage of ground-based characterization of absorption in the near-UV part of the  
539 spectrum, the regional representation of the spectral absorption properties in the OMAERUV  
540 models is limited. Therefore, spatial and temporal variations in the spectral properties of  
541 aerosols can be a potential source of error in the SSA retrieval.

542

## 543 **5 SUMMARY AND CONCLUSION**

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544 We presented a comparative analysis of the aerosol SSA retrieved from the OMI's two-channel  
545 aerosol algorithm (OMAERUV) against an independent ground-based inversion made by the  
546 SKYNET Sun photometers over selected 25 sites located mainly in Asia and Europe. This study  
547 follows our previous efforts of evaluating the OMI near-UV SSA product carried out using  
548 ground-based AERONET dataset (*Jethva et al.*, 2014). The capability of SKYNET sensors to  
549 measure the Sun and sky radiance at near-UV wavelengths (340-380-400 nm), and  
550 subsequently retrieve the aerosol optical properties, including SSA, at these wavelengths  
551 provide a unique opportunity to directly compare the two near-UV SSA products from ground  
552 and satellite. Ground-based inversion of SSA at the near-UV wavelengths eliminates the need to  
553 adjust and extrapolate satellite retrieval to the visible wavelengths such as the case with  
554 comparison against AERONET. Since the SSA inferred from two different platforms are  
555 essentially retrieved from two fundamentally different inversion algorithms, the present study

556 does not stand as a “validation” exercise for either retrieval data sets. Instead, the purpose of  
557 this analysis was to check the consistency (or lack thereof) between the two retrieved  
558 quantities of the same physical parameter regarding standard statistical comparison, i.e., RMSD  
559 and % of matchups within the expected uncertainties.

560

561 Unlike the AERONET Level 2 inversion product that reports spectral SSA when AOD (440 nm)  
562 exceeds a value of 0.4, SKYNET Level 2 dataset delivers spectral SSA in the near-UV and visible  
563 parts of the spectrum under all cloud-free observations for all AOD conditions. The collocation  
564 procedure that matched temporal inversion data from SKYNET with spatial retrievals from OMI  
565 gave resulted in a total of 2691 collocated data points for AOD>0.0 and 1223 when AOD>0.3  
566 collected from 25 sites representing biomass burning region of Southeast Asia, desert in China,  
567 and urban/industrial areas in Japan, India, and Europe. Combinedly for all 25 sites and under all  
568 AOD conditions, we find 38% and 59% of the total SKYNET-OMI SSA agree within their  
569 estimated uncertainty range of  $\pm 0.03$  and  $\pm 0.05$ , respectively, with an overall root-mean-  
570 square-difference of 0.06. When restricted with condition AOD>0.3 in both measurements, the  
571 agreement of comparison improved to 51% and 72% with root-mean-square-difference of  
572 0.047. When segregated by aerosol type, the agreement between the two sensors is found to  
573 be robust for matchups identified as the carbonaceous aerosols over several sites in Japan,  
574 *Seoul* in South Korea, *Phimai* in Thailand, and *New Delhi* in India, yielding 61% and 84% of data  
575 points falling within the limits of  $\pm 0.03$  and  $\pm 0.05$  with an overall RMSD of 0.035. The  
576 collocation procedure found few matchups for desert dust aerosol, mostly over *Dunhuang* site  
577 in China, showing a reasonable comparison with 50% and 68% data points within expected  
578 uncertainty limits. Among the three major aerosol types, the urban/industrial type aerosols  
579 provide the maximum number of matchup data points with a relatively poorer comparison,  
580 where 45% and 67% data are found to be within the uncertainty limits.

581

582 The differences in SSA between OMI and SKYNET are found to be larger at lower aerosol  
583 loading, where OMI retrieves significantly higher SSA compared to that of SKYNET. However,

584 the differences are minimized at larger AOD values ( $>0.5$ ), suggesting a convergence in both  
585 retrievals at moderate to larger aerosol loading. Similarly, the differences in SSA exhibit a  
586 stronger relationship to UVAI showing larger discrepancies beyond expected uncertainty limits  
587 at lower UVAIs ( $<0$ ), but nearing to zero with a reduced spread in matchups at larger  
588 magnitudes of UVAI ( $>0.2-0.3$ ).

589

590 Much of the inconsistency observed between OMI and SKYNET at lower aerosol loading  
591 indicates retrieval issues due to reduced signal-to-noise ratio and uncertain algorithmic  
592 assumptions. For instance, the OMAERUV retrievals are more susceptible to the changes in  
593 surface albedo at lower AODs, and to the spectral absorption at higher AODs (*Torres and Jethva,*  
594 2011). On the other hand, the SKYNET inversion algorithm assumes a wavelength-independent  
595 surface albedo of 0.1 across the UV to visible-near-IR wavelengths, which appears to be  
596 unrealistic, especially in the UV region where OMI surface albedo dataset reports much lower  
597 values ( $<0.05$ ) over land. Though the reflected light from surface plays a second-order role in  
598 the ground-based retrievals, previous studies and results derived in the present work (Figure 8)  
599 show that the uncertainty in surface albedo can cause non-negligible errors in SSA retrievals  
600 that likely exceed the expected accuracy level of  $\pm 0.03$ .

601

602 Despite the inherent uncertainties associated with both satellite and ground inversion products,  
603 a good level of agreement between the two independent techniques over SKYNET sites under  
604 the favorable conditions, i.e., at higher aerosol loading, higher solar zenith angle, and when the  
605 surface albedo assumption is consistent, is encouraging. We intend to extend the present  
606 analysis to other SKYNET sites whose data are still not directly accessible in the public domain.  
607 Continuing the evaluation of inversion products, both from satellite and ground, is an important  
608 exercise to track the changes and improvements in the algorithms and resulting data products,  
609 and to establish the consistency (or lack thereof) that can help to diagnose further and improve  
610 the accuracy of retrievals.

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**621 AUTHORS' CONTRIBUTIONS**

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622 Dr. Jethva, the leading author, conceptualized the study and wrote the paper. He conducted a  
623 comparative data analysis of OMI- and SKYNET-retrieved single-scattering albedo products  
624 presented in the paper. Dr. Torres (2<sup>nd</sup> author) brought his expertise in interpreting the results  
625 and helped improving the manuscript writeup.

626

**627 Additional Information**

628 The author(s) declare no competing interests, financial or non-financial.

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## 769 TABLES

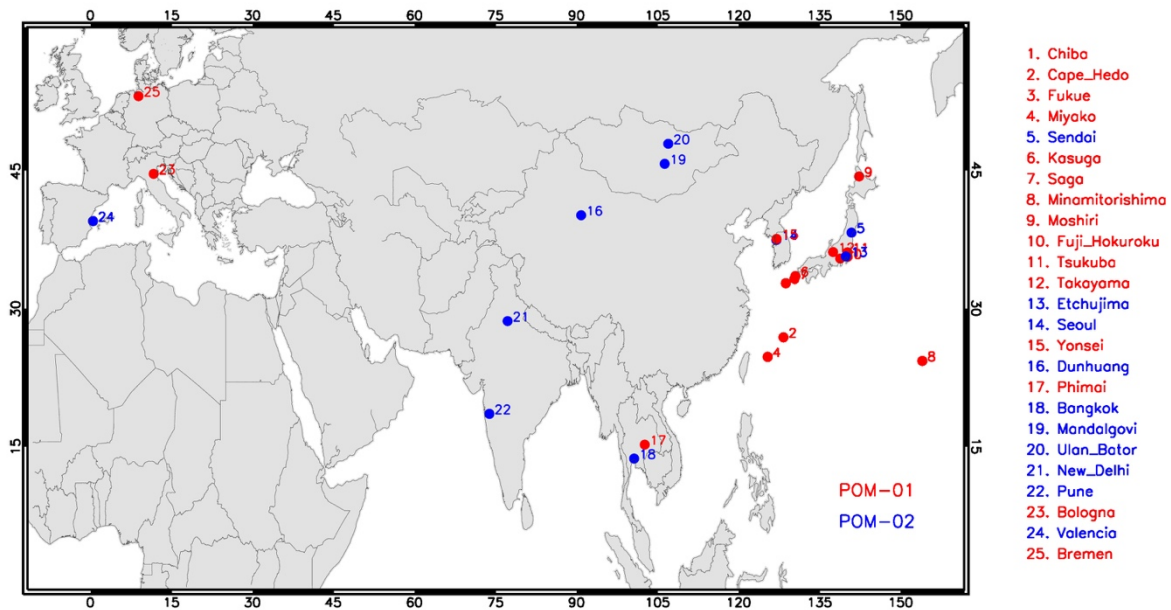
770 *Table 1 A list of SKYNET sites and corresponding dataset used in the present analysis. Sensor*  
 771 *type “POM02” consists of a total of seven wavelength filters, including near-UV bands, i.e., 340,*  
 772 *380, 400, 500, 675, 870, and 1020 nm, whereas “POM01” sensors have a total of five*  
 773 *wavelength filters, i.e., 400, 500, 675, 870, and 1020 nm. The rightmost four columns enlist the*  
 774 *statistical measures of OMI-SKYNET single-scattering albedo matchups.*

775 *Abbreviations: N: number of satellite-ground matchups, RMSD: root-mean-square-difference between OMI and*  
 776 *SKYNET, Q<sub>0.03</sub> and Q<sub>0.05</sub>: percent matchups within an absolute difference of 0.03 and 0.05.*

SKYNET Station Name	Longitude	Latitude	Country	Sensor Type	Data Period	N	RMSD	Q <sub>0.03</sub> (%)	Q <sub>0.05</sub> (%)
<i>Chiba University</i>	140.104°E	35.625°N	Japan	POM02	2005-2017	132	0.039	58	81
<i>Cape Hedo</i>	128.248E	26.867N	Japan	POM02	2005-2017	47	0.044	47	72
<i>Fukue</i>	128.682E	32.752N	Japan	POM02	2008-2017	71	0.041	59	76
<i>Miyako</i>	125.327E	24.737N	Japan	POM02	2004-2017	31	0.059	23	58
<i>Sendai</i>	140.84E	38.26N	Japan	POM01	2009-2017	34	0.052	50	74
<i>Kasuga</i>	130.475E	33.524N	Japan	POM02	2004-2017	159	0.057	40	61
<i>Saga</i>	130.283E	33.233N	Japan	POM02	2011-2017	66	0.044	52	71
<i>Minamitorishima</i>	153.97E	24.3N	Japan	POM02	2006-2009	-	-	-	-
<i>Moshiri</i>	142.260E	44.366N	Japan	POM02	2009-2011	2	0.018	100	100
<i>Fuji Hokuroku</i>	138.750E	35.433N	Japan	POM02	2009-2017	9	0.051	56	67
<i>Tsukuba</i>	140.096E	36.114N	Japan	POM02	2014-2017	5	0.027	80	100
<i>Takayama</i>	137.423E	36.145N	Japan	POM02	2014-2017	3	0.022	67	100
<i>Etchujima</i>	139.796E	35.664N	Japan	POM01	2004-2010	100	0.052	45	66
<i>Seoul</i>	126.95E	37.46N	Republic of South Korea	POM01	2005-2015	182	0.050	42	66
<i>Yonsei</i>	126.980E	37.570N	Republic of South Korea	POM02	2016	5	0.035	40	80
<i>Dunhuang</i>	90.799E	40.146N	China	POM01	1999-2007	40	0.048	50	68
<i>Phimai</i>	102.564E	15.184N	Thailand	POM02	2005-2017	139	0.031	71	91
<i>Bangkok</i>	100.605E	13.667N	Thailand	POM02	2009-2017	15	0.064	47	60
<i>Mandalgovi</i>	106.264E	45.743N	Mongolia	POM01	1998-2009	4	0.087	0	0
<i>Ulan Bator</i>	106.921E	47.923N	Mongolia	POM01	2013-2017	2	0.026	100	100
<i>New Delhi</i>	77.174E	28.629N	India	POM01	2006-2007	63	0.038	52	83
<i>Pune</i>	73.805E	18.537N	India	POM01	2004-2009	94	0.050	39	64
<i>Bologna</i>	11.34E	44.52N	Italy	POM02	2014-2017	114	0.065	25	50
<i>Valencia</i>	0.420E	39.507N	Spain	POM01	2014-2017	4	0.052	25	25
<i>Bremen</i>	8.854E	3.108N	Germany	POM02	2009	-	-	-	-

777

## 778 FIGURES



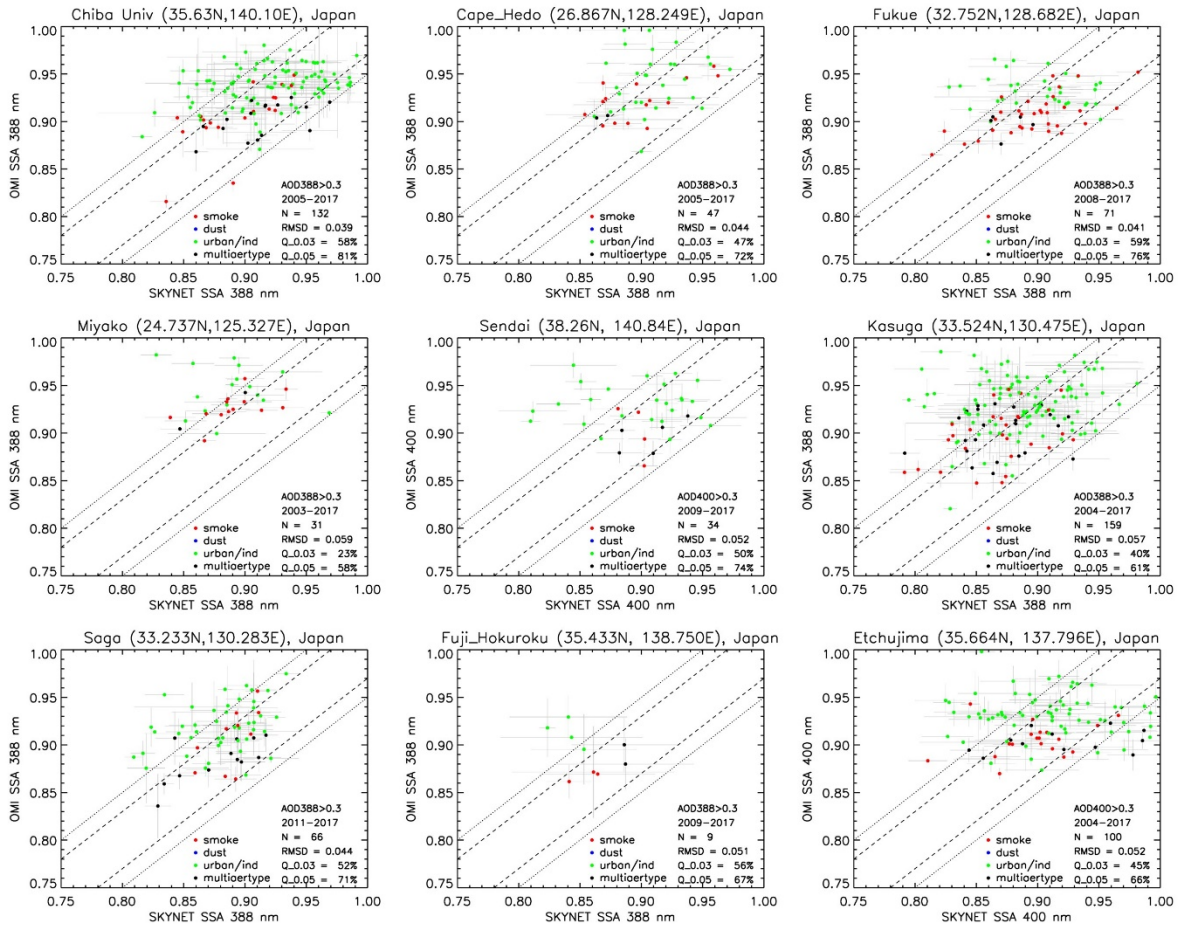
779

780 **Figure 1** Geographical placement of ground-based SKYNET sensors (POM-01 in blue, POM-02 in

781 red) over sites in Asia and Europe The SKYNET dataset for these sites are freely accessible from

782 the Center for Environmental Remote Sensing (CERes), Chiba University, Japan

783 (<http://atmos3.cr.chiba-u.jp/skyenet/data.html>).



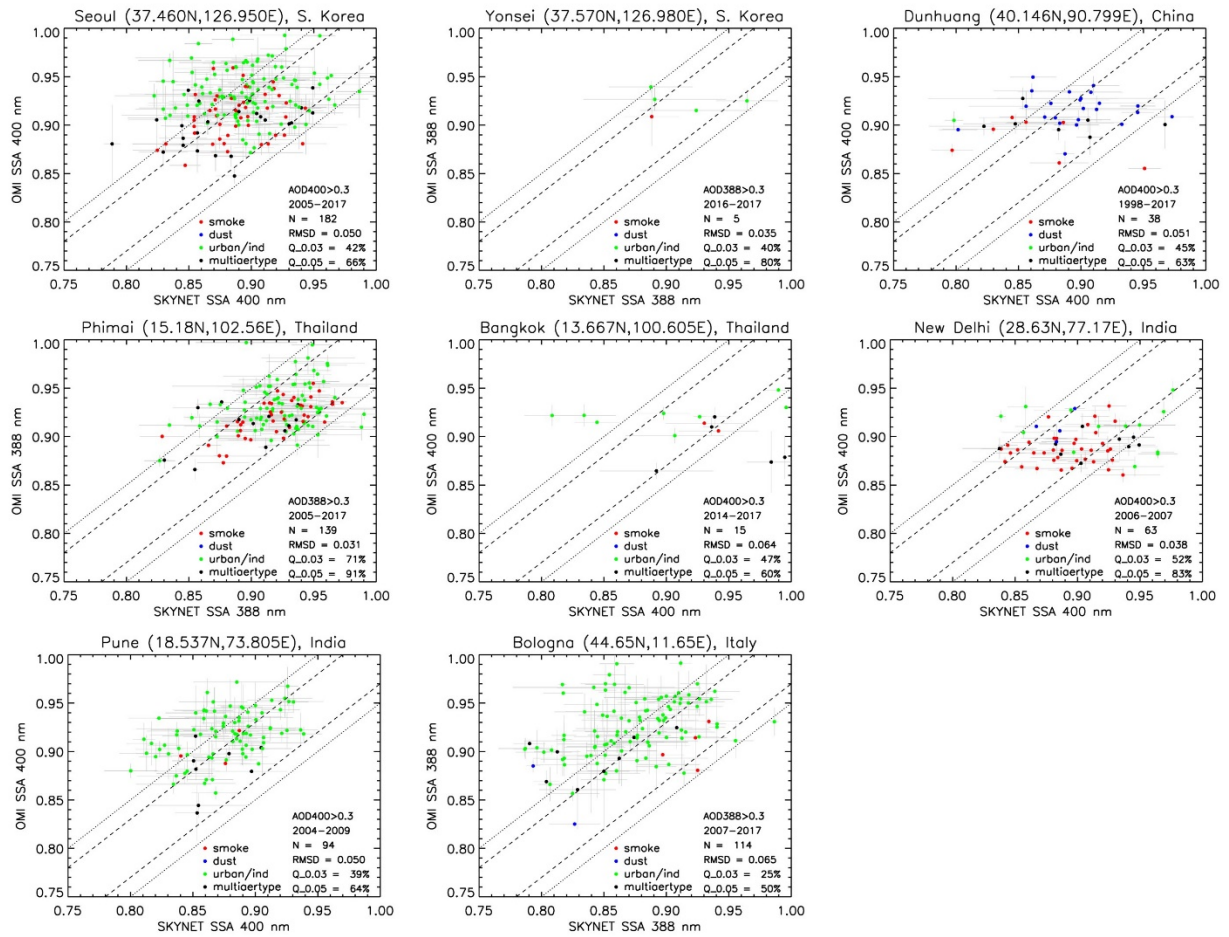
784

785 **Figure 2** OMAERUV versus SKYNET single-scattering albedo comparison for different sites in  
 786 Japan. Legends with different colors represent the aerosol type selected by the OMAERUV  
 787 algorithm for the co-located matchups (N). RMSD is the root-mean-square difference between  
 788 the two retrievals; Q\_0.03 and Q\_0.05 are the percents of total matchups (N) that fall within  
 789 the absolute difference of 0.03 and 0.05, respectively. OMI-SKYNET matchups with AOD>0.3  
 790 (388 or 400 nm) in both measurements are used for comparison.

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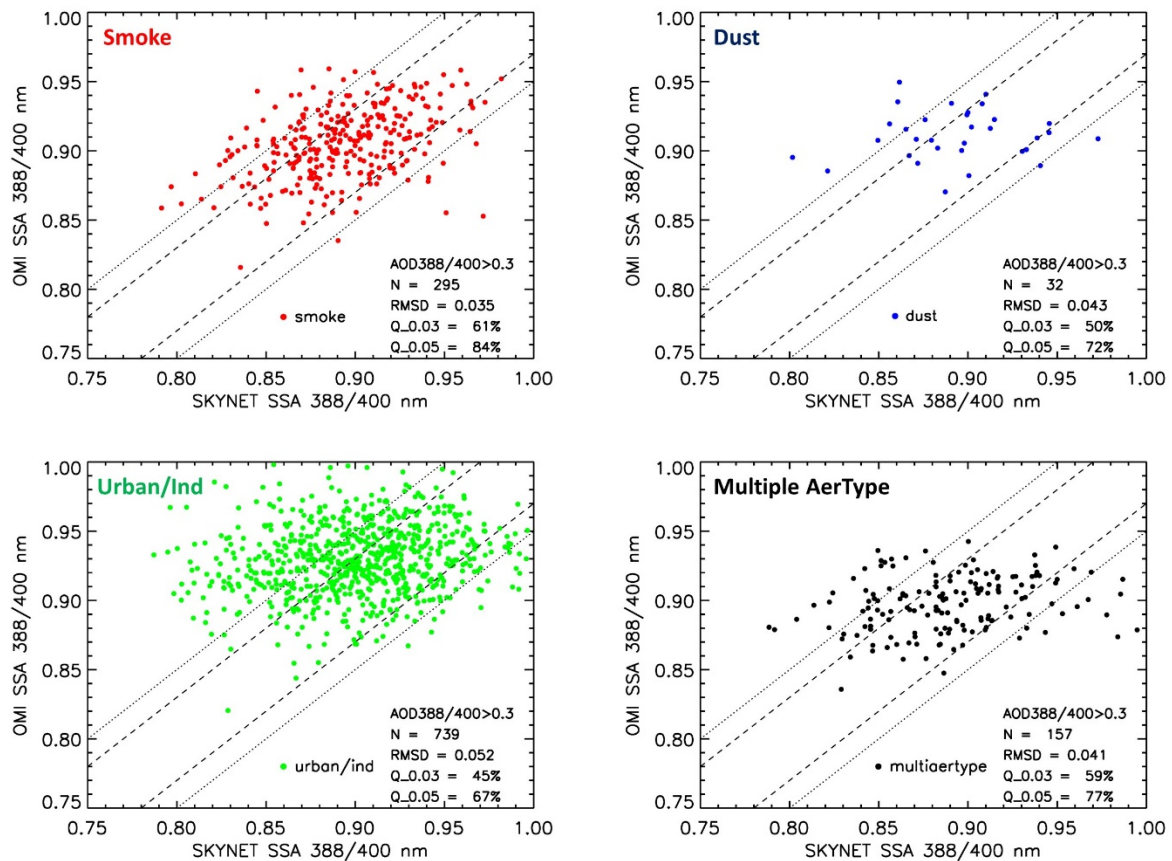




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794 **Figure 3** Same as in Figure 2 but for SKYNET sites in South Korea, China, Thailand, India, and

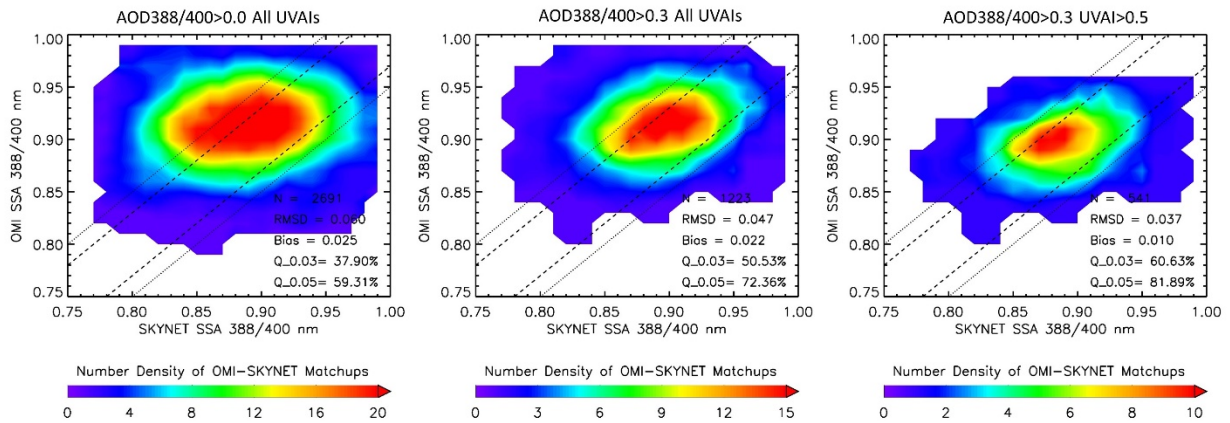
795 Italy.



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797 **Figure 4** Composite scatterplots of OMAERUV versus SKYNET single-scattering albedo (388 or  
 798 400 nm) for the three distinct aerosol types, i.e., smoke, dust, and urban/industrial, as  
 799 identified by the OMAERUV algorithm. OMI-SKYNET matchups with AOD>0.3 (388 or 400 nm) in  
 800 both measurements are used for the comparison.

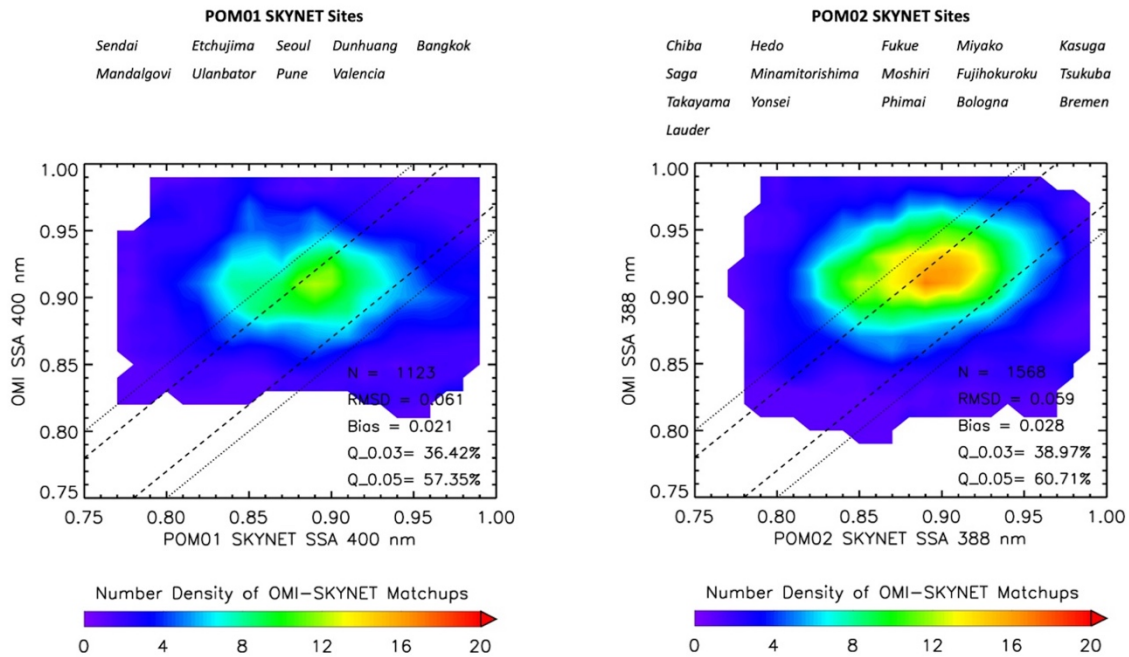
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802

803 **Figure 5** Composite number density plots of SSA comparison between OMI and SKYNET for  
 804 different aerosol loading conditions. The resultant statistics of the comparison are depicted in  
 805 the lower-right in each plot. Note that the scale used for the number density of satellite-ground  
 806 matchups for the three sets of comparisons is different.

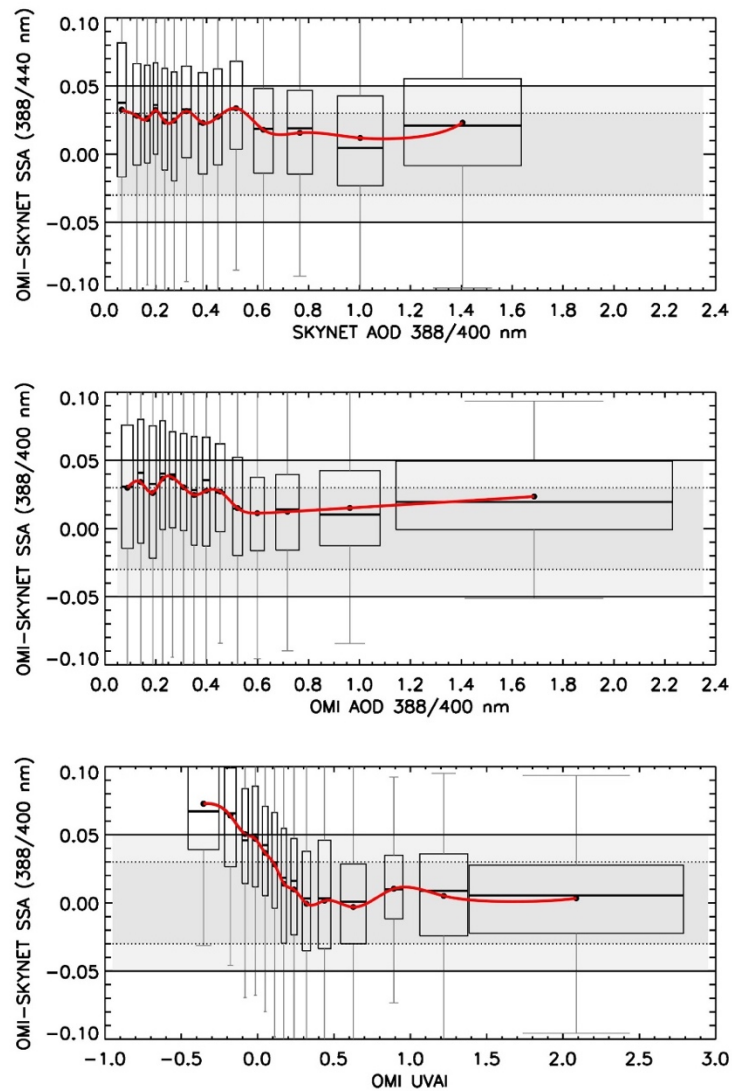
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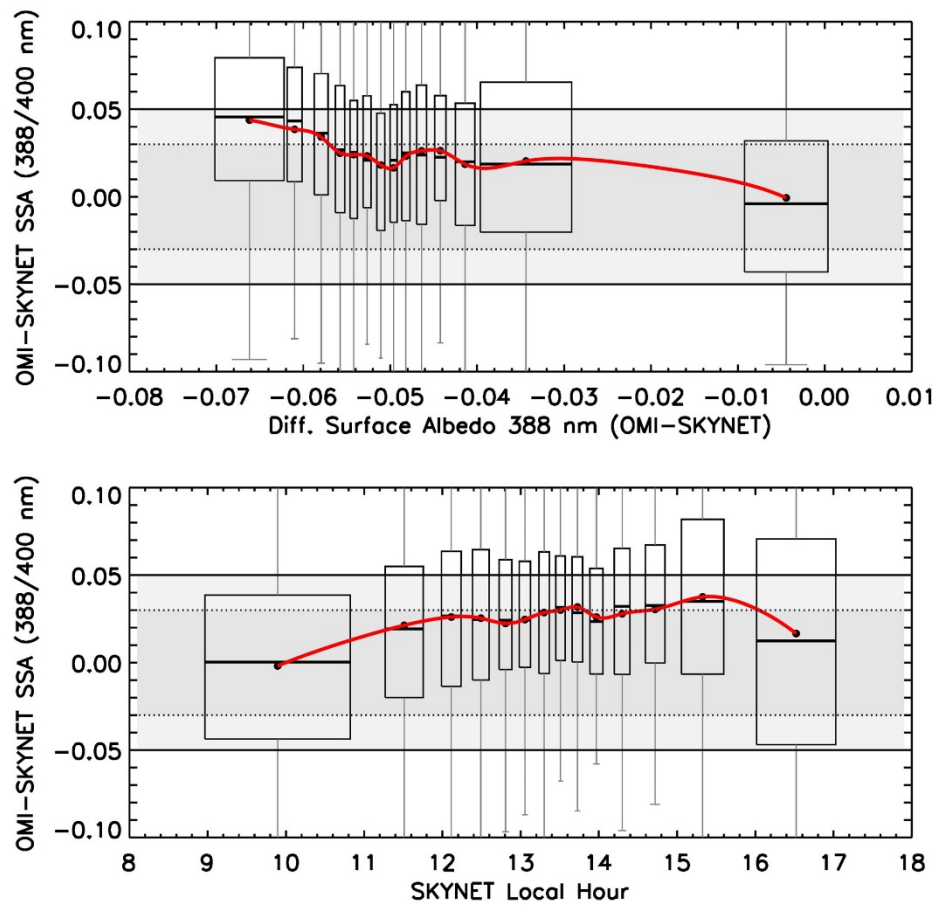
809 **Figure 6** Composite number density contour plots of SSA comparison between OMI and SKYNET  
 810 for different aerosol loading conditions. The resultant statistics of the comparison are depicted  
 811 in the lower-right in both plots.

812



813

814 **Figure 7** Difference in SSA between OMI and SKYNET as a function of the coincident SKYNET-  
 815 measured (top panel) and OMI-retrieved (middle panel) aerosol optical depth and OMI-  
 816 measured UVAI (bottom panel). Filled circles in black are the mean of difference for each AOD  
 817 and UVAI bin with an equal sample size of 200 matchups; horizontal lines represent median of  
 818 the bin samples; shaded area in gray encompasses data within 25 (lower) to 75 (higher)  
 819 percentile range, whereas vertical lines in gray represent 1.5 times interquartile range (25 to 75  
 820 percentile). The dotted and solid horizontal lines are the uncertainty range of  $\pm 0.03$  and  $\pm 0.05$ ,  
 821 respectively. The width of each box represents the 2-standard deviation of the data contained  
 822 in the respective bins.



823

824 **Figure 8** Same as in Figure 7 but the difference in SSA between OMI and SKYNET is related to (a)

825 the difference in surface albedo assumed by the two algorithms and (b) local measurement

826 hour of SKYNET.