

# Quantifying CH<sub>4</sub> emissions from hard coal mines using mobile sun-viewing Fourier transform spectrometry

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**Abstract.** Methane (CH<sub>4</sub>) emissions from coal production amount to roughly one-third of European anthropogenic CH<sub>4</sub> emissions in the atmosphere. Poland is the largest hard coal producer in the European Union with the Polish side of the Upper Silesian Coal Basin (USCB) as the main part of it. Emission estimates for CH<sub>4</sub> from the USCB for individual coal mine ventilation shafts range between 0.03 kt/a and 20 kt/a, amounting to a basin total of roughly 440 kt/a according to the European Pollutant Release and Transfer Register (E-PRTR, <http://prtr.ec.europa.eu/>, 2014). We mounted a ground-based, portable, sun-viewing FTS (Fourier Transform Spectrometer) on a truck for sampling coal mine ventilation plumes by driving cross-sectional stop-and-go patterns at 1 to 3 km from the exhaust shafts. Several of these transects allowed for estimating CH<sub>4</sub> emissions based on the observed enhancements of the column-averaged dry-air mole fractions of methane (XCH<sub>4</sub>) using a mass balance approach. Our resulting emission estimates range from  $6 \pm 1$  kt/a for a single shaft up to  $109 \pm 33$  kt/a for a subregion of the USCB, which is in broad agreement with the E-PRTR reports. Three wind lidars were deployed in the larger USCB region providing ancillary information about spatial and temporal variability of wind and turbulence in the atmospheric boundary-layer. Sensitivity studies show that, despite drawing from the three wind lidars, the uncertainty of the local wind dominates the uncertainty of the emission estimates, by far exceeding errors related to the XCH<sub>4</sub> measurements themselves. Wind-related relative errors on the emission estimates typically amount to 20%.

## 15 1 Introduction

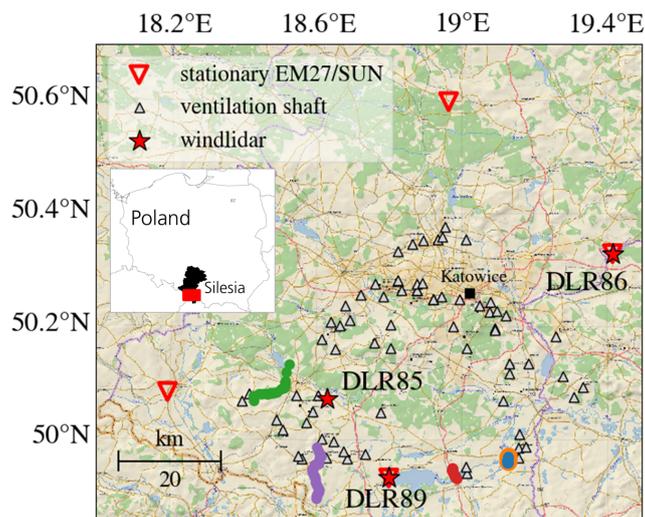
Atmospheric methane (CH<sub>4</sub>) causes the second largest radiative forcing of the long-lived greenhouse gases after carbon dioxide (IPCC, 2013). Except for the period from about 1999 to 2006, the global atmospheric CH<sub>4</sub> burden has been rising since 1750

and is now a factor of roughly 1.5 higher than in the pre-industrial era (Nisbet et al., 2014; Miller et al., 2013). Reasons for the renewed rise since 2007 are debated. Increased anthropogenic emissions are likely to be part of the answer (Bousquet et al., 2006; Kirschke et al., 2013; Schwietzke et al., 2016; Nisbet et al., 2016; Helmig et al., 2016; Worden et al., 2017).

Generally, about 20% of the global methane source is thought to be caused by the fossil fuel industry (Schwietzke et al., 2016; Bousquet et al., 2006). The E-PRTR (European Pollutant Release and Transfer Register, <http://prtr.ec.europa.eu/>, 2014) reports the total European anthropogenic CH<sub>4</sub> emission as 1874kt for the year 2014. Methane emitted by underground mining and related operations amounts to 686 kt/a for all reporting facilities in Europe. With emissions of 466 kt/a the Upper Silesian Coal Basin (USCB) in Poland is a CH<sub>4</sub> emitting hot spot in Europe. The USCB roughly comprises an area of 50 km × 50 km centered at 50.1N, 18.8E with about 50 active CH<sub>4</sub> ventilating hard coal mining shafts emitting between 0.03 kt/a and 20 kt/a (E-PRTR, 2014). The USCB was the main target of the CoMet (The Carbon dioxide and Methane mission 2018) campaign, which covered roughly three weeks from 23 May to 12 June 2018. During CoMet several aircraft and ground-based instruments were co-deployed to evaluate strategies on how to verify the local CH<sub>4</sub> emissions. Here, we focus on the CoMet measurements of a single mobile ground-based, sun-viewing FTS.

The sun-viewing FTS of the type EM27/SUN developed by the Karlsruhe Institute of Technology in collaboration with Bruker Optics (Gisi et al., 2012) generally are used to measure the column-averaged dry-air mole fractions of methane (XCH<sub>4</sub>) and other gases by measuring direct-sun absorption spectra in the shortwave-infrared spectral range (around 6000 cm<sup>-1</sup>). Recently, Hase et al. (2015) and Chen et al. (2016) combined several of these FTS instruments into ad-hoc networks in the vicinity of major cities to estimate urban carbon dioxide (CO<sub>2</sub>) and CH<sub>4</sub> emissions respectively. Viatte et al. (2017) used a similar FTS configuration together with large-eddy simulations to estimate the amount of CH<sub>4</sub> emitted by dairies in southern California. Toja-Silva et al. (2017) verified power plant emissions with computational fluid dynamics simulations and differential column measurements using two EM27/SUN instruments in Munich. Klappenbach et al. (2015) demonstrated mobile deployment on a research vessel requiring a custom-built solar tracker to compensate for the motion of the platform. Butz et al. (2017) mounted the EM27/SUN on a small truck to measure the volcanic CO<sub>2</sub> plume of Mt. Etna by recording plume transects in stop-and-go patterns. Kille et al. (2019) separated natural, from agricultural CH<sub>4</sub> emissions using a network of four EM27/SUN instruments.

Based on a mobile FTS setup similar to Butz et al. (2017), we explore the feasibility of estimating CH<sub>4</sub> emissions for individual coal mine ventilation shaft and groups of shafts. To this end, we mounted the EM27/SUN on a truck and, we drove transects through the CH<sub>4</sub> plumes in the USCB while driving stop-and-go patterns. We estimated emissions from the recorded XCH<sub>4</sub> enhancements using a mass balance method, which was assisted by wind information from three wind lidars deployed in the broader region. Section 2 summarizes the campaign setup, describes the deployed instruments and the data analysis, and explains the emission estimation method. Emission estimates of CH<sub>4</sub> are summarized in section 3 followed by the analysis of the errors in section 4. We evaluate the suitability of the method used, compare the estimated emissions to the E-PRTR database, and conclude this work with section 5.

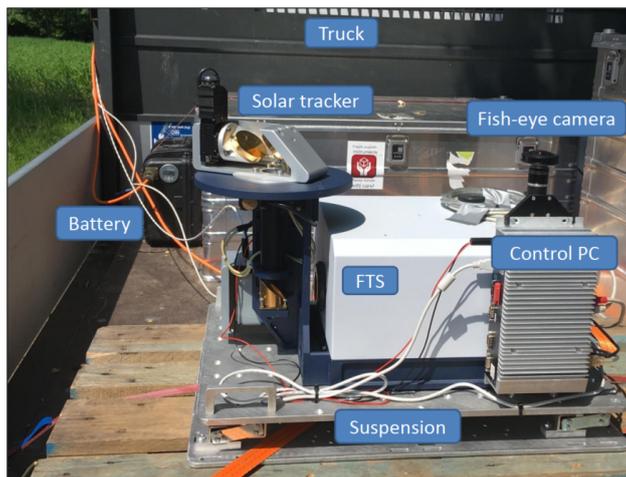


**Figure 1.** Map of the USCIB in Southern Poland (see small inset on the left with black illustrating the region of Silesia and red depicting the map excerpt of the USCIB). Ventilation shafts are gray triangles. Colored dots depict five plume transects measured using the mobile FTS on 24 May at around 7-8 am (orange), on 24 May around noon (blue), on 1 June (green), on 6 June during morning hours (red), and on 6 June around noon (purple). The most active CH<sub>4</sub> emitters are assumed to be located in the southern part of the USCIB, which is why mobile FTS measurements are focused on this area. Stationary EM27/SUN FTS locations are marked as red triangles; the three wind lidars DLR85, DLR86, and DLR89 are marked as red stars. Eastern and southern wind lidars are placed at the same locations as the respective EM27/SUNs. Background map from ESRI (2019).

## 2 Campaign setup, instruments, and methods

Figure 1 shows a map of the USCIB. The coal mine ventilation shafts group into a northern part in the vicinity of the city of Katowice and into a southern part close to the Czech border. The reported operations of the mobile FTS were carried out in the southern part where the largest emitters are expected. In the framework of CoMet, we also operated three wind lidars and four stationary, sun-viewing FTS of the same model as the mobile one at stations surrounding the USCIB. In-situ sensors were deployed in cars sampling throughout the region. Aircraft carrying in-situ (Kostinek et al., 2019) and remote sensing instrumentation (Gerilowski et al., 2011) was operated out of Katowice airport, and a long-distance aircraft (Amediek et al., 2017) visited the USCIB operating out of the airport at Oberpfaffenhofen close to Munich, Germany.

Here, we use the mobile FTS measurements (section 2.1) and the wind lidar data (section 2.2) together with the cross-sectional flux method (section 2.3) to estimate CH<sub>4</sub> emissions. Future studies will target at the other CoMet data and at combining them.



**Figure 2.** Mobile FTS (EM27/SUN) mounted on a truck. Direct sunlight is collected via the solar tracker. The battery supplies the measurement notebook inside the vehicle, the spectrometer including solar tracker, and the control PC. A fish-eye camera automatically detects the solar disc in the sky above the spectrometer. The control PC steers the solar tracker towards the solar disc.

## 2.1 Mobile FTS observatory

The mobile FTS observatory consists of a Bruker EM27/SUN Fourier Transform spectrometer with a custom-built solar tracker deployed on a truck similar to the setup used by Butz et al. (2017). The custom-built solar tracker enables efficient stop-and-go patterns since it supports start-up and repointing within a few seconds once the solar-tracking is interrupted. During driving, however, significant high frequency mechanical disturbances occur, which is not yet compensated for by the tracker. The mobile FTS observatory was operated out of its campaign-base at the site Pustelnik in the south of the USCB between 23 May and 12 June 2018 whenever weather conditions were promising for direct-sun viewing. The mobile instrument followed public roads on approximately perpendicular tracks to the assumed plume 1 to 3 km downwind of the targeted CH<sub>4</sub> sources (Fig. 1). We chose the transects depending on the wind direction and aimed for isolated shafts with preferably no other shafts upwind. However, we also examined subregions within the USCB containing several mines and shafts upwind. For the plume transects, it was necessary to sample the CH<sub>4</sub> background on both sides of the plume in order to reliably remove the XCH<sub>4</sub> background and to derive the above-background enhancements, which are used for the emission estimation. In total, we report on five transects. One morning transect and one noon transect for 24 May, one transect on 1 June, and one morning transect and one noon transect on 6 June.

The EM27/SUN used here has a spectral resolution of 0.5 cm<sup>-1</sup> and operates in the spectral range 4000 to 11000 cm<sup>-1</sup>. Frey et al. (2019) observed an average instrument-to-instrument difference of 0.8 ppb for XCH<sub>4</sub> for an ensemble of instruments. According to the Allan deviations described by Chen et al. (2016) the precision of XCH<sub>4</sub> for 120 s integration time is 0.3 ppb, which is roughly 0.02% of the total column. The accuracy amounts to roughly 4.7 ppb based on a 3 hours side-by-side measurements with the Karlsruhe TCCON station in spring 2017. We recorded double-sided interferograms with a sampling rate

of 10kHz and we typically coadded ten double-sided interferograms resulting in roughly 120s total integration time per stop. Typical dwell times per stop in our stop-and-go patterns were 2min 30s. For the last part of our deployment on June 06 we increased the sampling rate to 40kHz, which resulted in average dwell times of 60s. As depicted in Fig. 2 the FTS is mounted on a suspension plate. The rubber bearings absorb minor shocks and protect the instrument from major bumps. A 12V battery powers the whole system. We remotely operated the instrument from inside the driver's cabin, which enabled us to measure alongside busy roads as we did not have to leave the car to start the measurements. The solar tracker is the standard two mirror system developed by Gisi et al. (2011) supplemented with a two-stage tracking control mechanism. The coarse tracking stage uses a fish-eye-camera that identifies the position of the sun and gears the tracker into the vicinity of the sun. The fine-tracking system is the camera system developed by Gisi et al. (2011), which takes over once the coarse tracking is successful. Generally, the fine-tracking errors amount to 10arc s, i.e., the tracking is precise to within fractions of the apparent diameter of the solar disk.

For the retrieval of  $XCH_4$  from the FTS measurements we use the software package PROFFIT (Hase et al., 2004), which is in routine use for accurate trace gas retrievals from shortwave-infrared direct-sun absorption spectra (e.g., Gisi et al., 2012; Frey et al., 2015; Klappenbach et al., 2015; Chen et al., 2016; Butz et al., 2017; Frey et al., 2019). Quality filtering, based on the unmodulated (DC) part of the recorded interferograms, removes all measurements that are affected by unstable pointing conditions, e.g., due to intermittent thin clouds disturbing solar-tracking (e.g., Klappenbach et al., 2015). We set the DC fluctuation threshold to 1% for discarding corrupt interferograms. The spectral bandwidth of the EM27/SUN allows for measurements of  $H_2O$ ,  $O_2$ ,  $CO_2$ ,  $CH_4$ ,  $CO$ , ( $CO$  in combination with an additional channel described by Hase et al. (2016)) and other gases if present in significant amounts (Butz et al., 2017). Related species for this study are  $O_2$  and  $CH_4$ . The spectral retrieval windows for these target absorbers are  $7765$  to  $8005\text{cm}^{-1}$  and  $5897$  to  $6145\text{cm}^{-1}$  for  $O_2$  and  $CH_4$ , respectively. The absorption line parameters are taken from Toon (2017) and Rothman et al. (2009). Assuming that the targeted  $CH_4$  plumes mainly reside in the planetary boundary layer (PBL), the PROFFIT retrieval was set up to estimate the total column of methane [ $CH_4$ ], in units of molecules  $\text{m}^{-2}$ , by only scaling the lower part ( $< 1700\text{m a.g.l.}$ ) of the a priori vertical profile taken from the EMAC model. According to Butz et al. (2017) who only scaled the relevant plume layers of mount Etna in Sicily, we only scaled the relevant layers of the  $CH_4$  plumes. Here, we use EMAC simulation results from a simulation similar to the simulation described by Jöckel et al. (2016). Figure A1 compares  $CH_4$  profiles of both, a-priori and retrieved  $CH_4$  columns which are congruent above  $1700\text{m a.g.l.}$

For the [ $O_2$ ] retrieval we scaled the full a priori profile. The column-averaged dry-air mole fraction of methane ( $XCH_4$ ) is then calculated through

$$XCH_4 = \frac{[CH_4]}{[O_2]} \times 0.20942 \quad (1)$$

where 0.20942 is the atmospheric  $O_2$  mole fraction.

It is crucial to remove the  $XCH_4$  background reliably to derive the above-background enhancements  $\Delta XCH_4$ , which are then used to estimate the emissions. Similar to Butz et al. (2017), we used  $XCH_4$  measurements before and after crossing a

plume to fit the local background linearly in time. We conducted several sensitivity studies to best estimate the background and to quantify the error associated with background removal (see section 4).

On average, we recorded ten spectra per stop and averaged the retrieved XCH<sub>4</sub> for every stop location. The relative standard deviations of all measurements recorded at every stop range from 0.12% (roughly 2ppb) on 6 June, when most observations were taken far (>40km, stops every 500m) from any source to 0.26% (roughly 4ppb) on 24 May, when we sampled the plume within 2km of the source, stopping approximately every 70m. We take the stop-wise standard deviations as a metric for the measurement error that propagates into the emission estimates, which is conservative compared to the small noise error of individual measurements (0.3ppb).

On 6 June, we additionally operated a Picarro CRDS (Cavity Ring-Down Spectroscopy) analyzer for in-situ CO, CO<sub>2</sub>, CH<sub>4</sub>, H<sub>2</sub>O on board the mobile observatory to discriminate background from plume enhancements of CH<sub>4</sub> in real-time. The CRDS analyzer provides additional support to identify the horizontal extent of the plume without relying on forecast wind fields.

## 2.2 Wind lidars

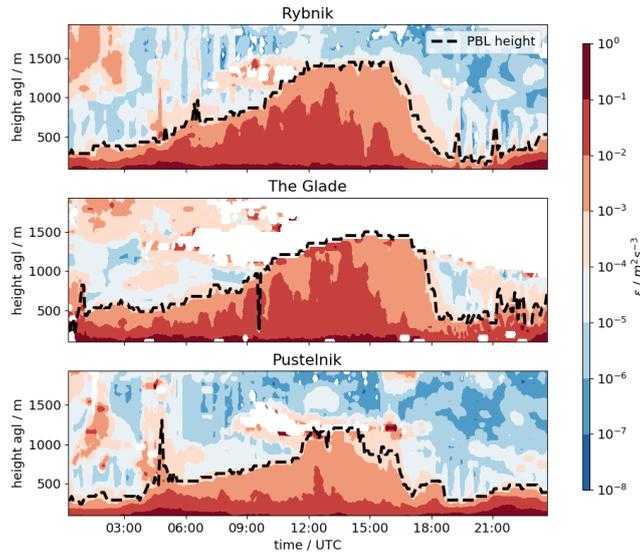
Three Leosphere Windcube 200S Doppler wind lidars (Vasiljević et al., 2016; Wildmann et al., 2018) were deployed during the CoMet measurement campaign. The instruments were distributed to the east, south, and mid-west of the USCB. The exact locations can be found in Tab. 1 and Fig. 1.

Instr.	Site	Lat. (°N)	Lon. (°E)	m a.s.l
DLR85	Rybnik (W)	50.0725	18.6298	253
DLR86	The Glade (E)	50.3292	19.4155	303
DLR89	Pustelnik (S)	49.9326	18.7998	266

**Table 1.** Locations of wind lidars deployed during CoMet. The second column lists the site names with their respective cardinal directions relative to the USCB in brackets.

The measurement sites differ to some extent in the surrounding land surface and vegetation. DLR85 is deployed in the west of the USCB next to a private airfield mostly used for parachuting activities and flight training schools. The land surface towards the south-westerly direction is flat. However, the airfield is close to a forest located to the north-west of the wind-lidar. DLR86 is placed in the east of the USCB, 62km from DLR85. The area is surrounded by forest. The southern instrument DLR89 is deployed in the vicinity of the barrier lake Goczalkowickie. The linear distance between the instrument and the bank is roughly 500m.

The Doppler wind lidars were programmed to perform velocity azimuth display (VAD) scans continuously. To retrieve vertical profiles of not only wind but also turbulence, the VAD scans were performed with two different elevation angles. A series of 24 VAD scans with an elevation angle of 75° are followed by six scans with an elevation angle of 35.3°. The 75° scans were chosen to allow retrievals of mean wind speed and direction profiles with a small cone angle, covering at least the whole boundary layer. Filtered sine-wave fitting (FSWF) according to Smalikho (2003) was used to calculate wind speed and



**Figure 3.** Wind lidar profiles of eddy dissipation rate (EDR) on 6 June at the three locations Rybnik/DLR85 (top), The Glade/DLR86 (middle), Pustelnik/DLR89 (bottom). We performed mobile FTS transects on this day in the morning from 7 to 8am and around noon (9:30 am to 10:30 am). Dashed lines represent the respective PBL heights. EDR greater than  $10^{-4}\text{m}^2\text{s}^{-3}$  corresponds to PBL values. Note, that the PBL height at station Pustelnik/DLR89 is shallower compared to the other two stations. This is generally observed throughout the measurement campaign.

wind direction from line-of-sight velocities. While the  $75^\circ$ -scans is also used to retrieve turbulence kinetic energy dissipation rate, an elevation angle of  $35.3^\circ$  is necessary to derive turbulent kinetic energy, integral length scale and momentum fluxes according to Smalikho and Banakh (2017). As described in Stephan et al. (2018b), boundary-layer height can be estimated through detection of the height at which the eddy dissipation rate (EDR) drops below a value of  $10^{-4}\text{m}^2\text{s}^{-3}$ . As an example,

5 Fig. 3 depicts EDR and PBL height as retrieved from  $75^\circ$ -scans of the three wind lidars on 06 June 2018. The uncertainty of wind speed retrievals with the FSWF-method depends on the signal-to-noise ratio of line-of-sight velocity measurements (Stephan et al., 2018a). We consider an uncertainty of  $0.2\text{ms}^{-1}$ , which is particularly critical under low wind conditions.

### 2.3 Cross-sectional flux method

The cross-sectional flux method as discussed by Varon et al. (2018) is an established tool typically used for airborne in-situ

10 observations (White et al., 1976; Mays et al., 2009; Cambaliza et al., 2014, 2015; Conley et al., 2016). It has also been applied to airborne  $\text{CH}_4$  (partial) column measurements (Krings et al., 2011, 2013; Tratt et al., 2011, 2014; Amediek et al., 2017). Based on the mass balance assumption, the strength  $Q$  of a  $\text{CH}_4$  point source such as a coal mine ventilation shaft can be calculated by integrating the product of  $\text{CH}_4$  concentration enhancements and wind speed along a plume cross-section. Approximating the integral by a sum over all cross-plume measurements  $i$ , the source strength  $Q$  in units  $[\text{kg s}^{-1}]$  is given by

$$Q = \sum_i \Delta\Omega(x_i, y_i) U_{\text{eff}}(x_i, y_i) \Delta y_i \quad (2)$$

where we assume the horizontal coordinates  $x_i$  and  $y_i$  along and across plume direction, respectively.  $\Delta\text{XCH}_4$  [ppm] is translated into the plume enhancement  $\Delta\Omega$  [ $\text{kg m}^{-2}$ ] by

$$\Delta\Omega = \Delta\text{XCH}_4 \frac{[\text{O}_2]}{0.20942 N_a} 10^3 M_{\text{CH}_4} \quad (3)$$

5 with the molar mass  $M_{\text{CH}_4}$  and the Avogadro constant  $N_a$ .  $U_{\text{eff}}(x_i, y_i)$  is an effective wind speed representative of plume dispersion. The distance  $\Delta y_i$  is the cross-plume segment for which the  $\Delta\Omega(x_i, y_i)$  enhancement is assumed to be representative. We choose the measurements  $i$  to be individual stops of the stop-and-go patterns i.e., we average quantities over individual stops. For estimating wind speed and wind direction, we choose the distance-weighted average of all three wind lidar profiles interpolated to the time of observation, as the baseline wind scenario. The vertical profiles of wind speed (wind direction) are  
 10 then averaged over the PBL to represent  $U_{\text{eff}}(y_i)$  in Eq. (2).

There are several caveats and sources of error to this method. (1) The measurement vehicle has to follow public roads, which are not always perpendicular to the plume direction. Therefore, we calculate the cross-plume segment  $\Delta y_i$  by

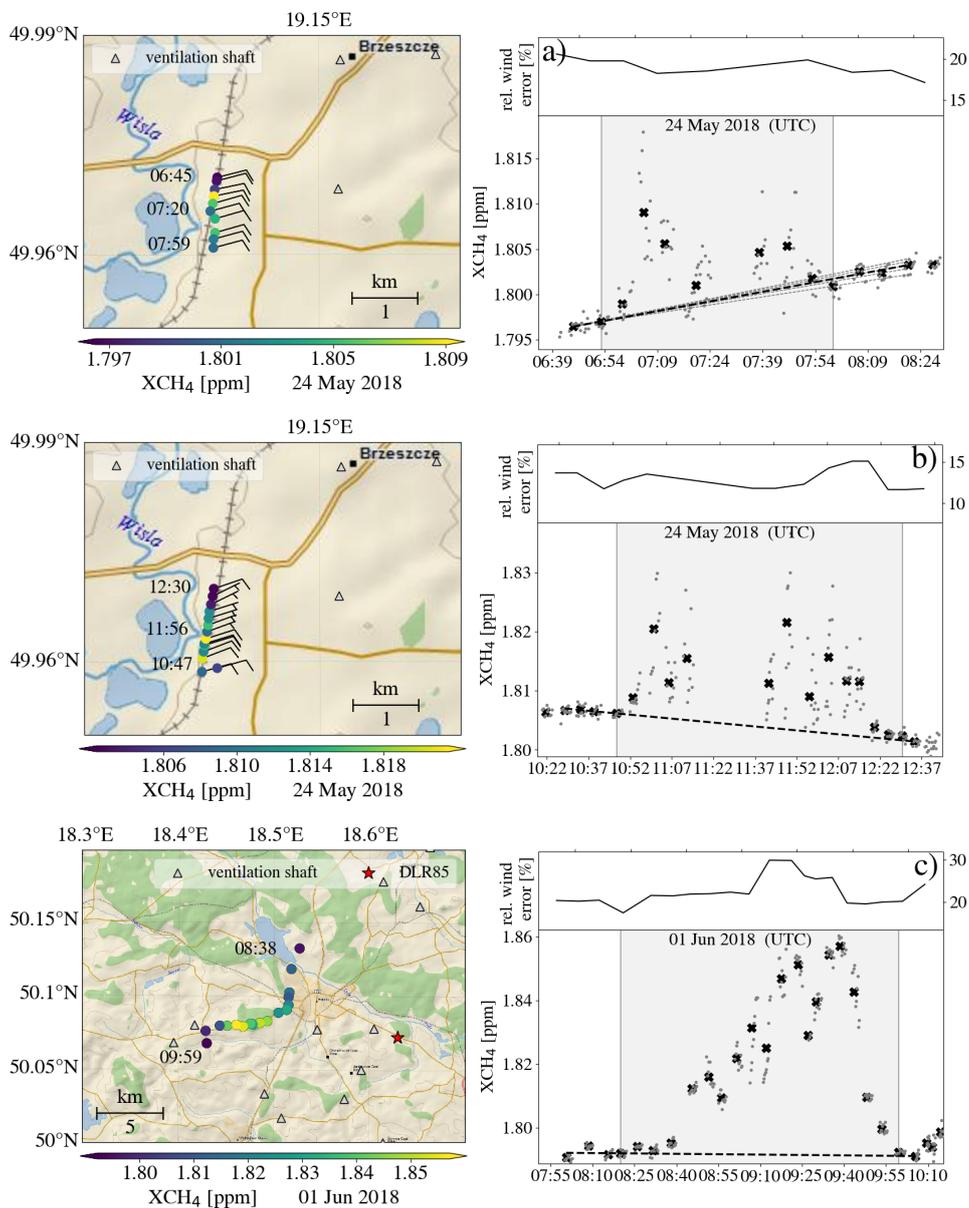
$$\Delta y_i = ds_i \sin(\alpha_i) \quad (4)$$

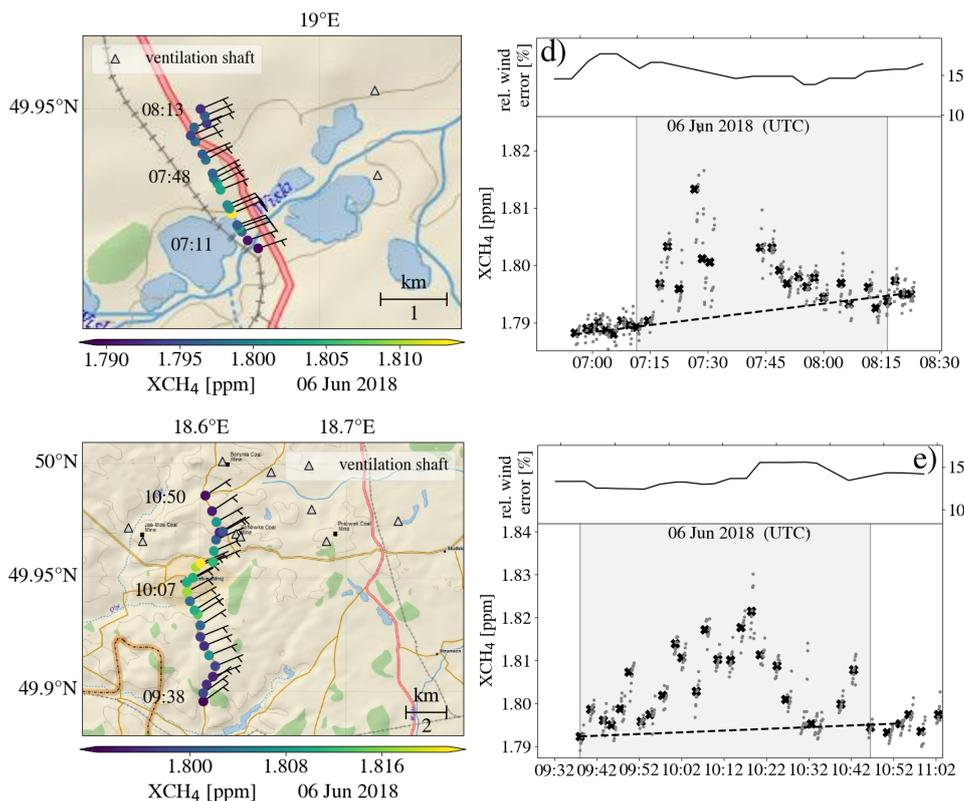
where  $ds_i$  is the driven distance between two stops of the stop-and-go pattern and  $\alpha_i$  is the relative angle between the PBL-  
 15 averaged wind direction and the driving direction. (2) The calculation of  $\Delta\text{XCH}_4$  requires a well-defined  $\text{XCH}_4$  background removal, which translates into the operational requirement to sample background air on both sides of the plume. (3) We assume that the wind direction remains constant over the sampling duration, which typically is 1 to 1.5 hours. (4) The effective wind speed  $U_{\text{eff}}$  must be accurate and representative for the observed plume enhancement  $\Delta\text{XCH}_4$ . Relative errors in effective wind speed or its representativeness equal relative errors in estimated emissions, which is particularly striking under low wind speed  
 20 conditions. We discuss these caveats along with the data and errors in section 3 and 4.

### 3 Estimated $\text{CH}_4$ emissions

Figure 4 displays five transects: one morning and one noon transect on 24 May, one transect on 1 June, and again one morning  
 and one noon transect on 6 June covering different mines and shafts in the USCB. The panels on the left illustrate the locations  
 of observations relative to the nearest mining ventilation shafts. The wind speed and direction is indicated as small wind barbs.  
 25 Note, that for 1 June the wind speed was below 5 kn (i.e., always below  $2.5 \text{ ms}^{-1}$ ) and therefore wind barbs are not displayed. However, wind direction measured by the wind lidars reveals a general south-easterly wind direction. Right-hand side panels show the timeline of all measurements of the EM27/SUN as gray dots together with their respective local stop-wise averages as black crosses. The background  $\text{XCH}_4$  concentration (dashed black lines in Fig. 4) is based on a linear least squares (in time)

fit of the measurements before and after the plume transect. For our analysis, we selected all cases for which (1) we measured a clear background signal before and after crossing the plume, (2) the wind direction was roughly constant as indicated by the wind-lidars, and (3) the overall transect duration does not exceed 1.5 hours. All transects and their best-estimate emissions calculated with the baseline wind scenario are described in detail in this section and summarized in Table 2. Discussion of errors follows in section 4.





**Figure 4.**  $XCH_4$  measured for five transects downwind of the coal mine ventilation in the USCBA (panel rows a) to e)) Left panels display the driven path and the locations of each observation. Wind bars indicate the wind direction. Note, that on 1 June wind speed was very low ( $<2.5\text{ m s}^{-1}$ ) and the wind bars are not resolved. However, the general wind direction was south-east. The targeted mining ventilation shafts are marked as gray triangles. Right hand panels display measured  $XCH_4$  as gray dots and respective local stop-wise averages as black crosses. The best-estimate background  $XCH_4$  is indicated as dashed black line. Gray areas represent the measurements identified as intra-plume. The black line on top of every right hand side panel display the combined, relative wind error for each measurement. Dashed gray lines in panel a) illustrate all background options contributing to the background related error. Geographic maps from ESRI (2019).

On 24 May in the morning (panels a) in Fig. 4), we drove stop-and-go patterns with stops roughly every 100m about 2km west of a ventilation shaft. The target shaft belongs to a mine with four shafts in total. The next closest ventilation shaft was located about 2km to the north, and no other known shafts were located upwind. Observed plume enhancements  $\Delta XCH_4$  were up to 15ppb with a maximum around 7am. The best-estimate emissions  $Q$  are  $6 \pm 1\text{ kt/a}$ . Later the same day (panels b) in Fig. 4), we observed the same ventilation shaft but driving from south to north. The transect was interrupted for nearly half an hour due to technical issues from 11:15 to 11:41 UTC, however, we finished the transect afterwards. Maximum  $\Delta XCH_4$  reached roughly 30ppb around 11:45am. Estimated emissions  $Q$  are  $10 \pm 1\text{ kt/a}$ , nearly 70% higher than for the morning transect, which is particularly examined in section 5. Throughout our observations in the morning and around noon, the  $XCH_4$  background south of the plume is roughly 4 to 5ppb higher compared to the north.

Date and time	estimated emissions [kt/a]	combined $\sigma$ [kt/a]	E-PRTR 2014 [kt/a]	relative standard deviation ( $1\sigma$ ) due to averaging of:								
				FTS obs. in ppb		wind speed ( $U_{\text{eff}}$ ) in %			wind direction ( $dy_i$ ) in %			
				XCH <sub>4</sub>	backgr.	vert.	horz.	time	vert.	horz.	time	
24 May 7 to 8 am	<b>6</b>	<b>1</b>	19	<b>9.63</b>	2	0.2	13	10	8	2	3	2
24 May noon	<b>10</b>	<b>1</b>	15	<b>9.63</b>	4	0.3	8	10	3	5	3	3
01 June 8 to 10 am	<b>109</b>	<b>33</b>	30	-	3	0.6	18	10	9	14	3	12
06 June 7 to 8 am	<b>17</b>	<b>3</b>	16	<b>24.3</b>	2	0.3	10	10	6	2	3	2
06 June noon	<b>81</b>	<b>13</b>	16	<b>~ 80</b>	2	0.4	8	10	4	4	3	2

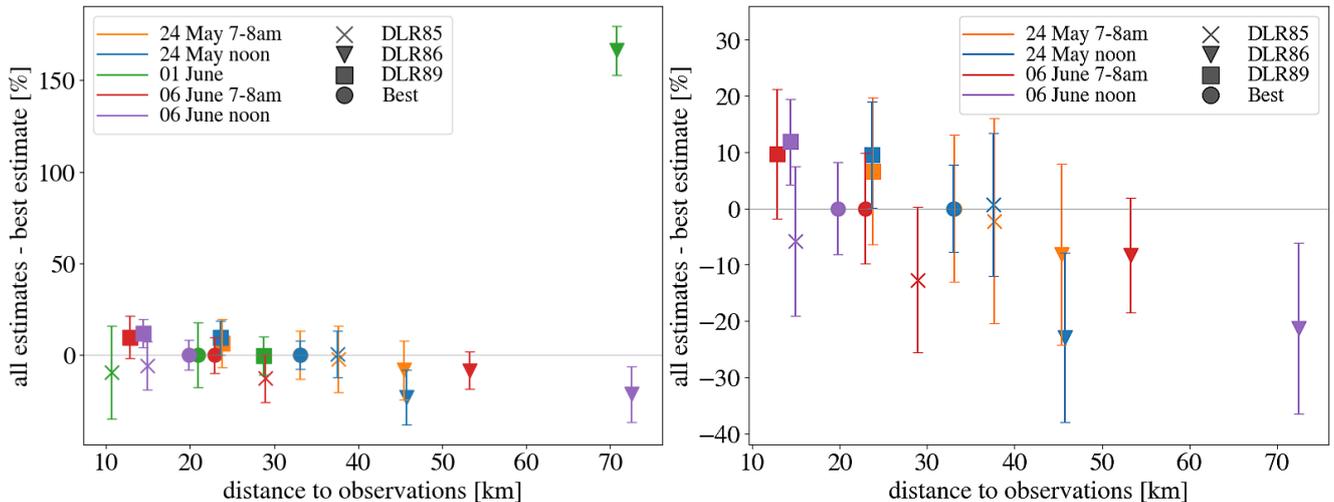
**Table 2.** Overview of successful plume transects along with relative standard deviations of the primary sources of uncertainty. Error estimation procedures are explained in the main text. Bold numbers represent estimated emissions and errors together with the respective E-PRTR 2014 entries in the fifth column, which are the mine-wise reported values distributed evenly to every single listed shaft of each mine. Several upwind sources do not allow accurate source attribution on 1 June, hence no E-PRTR estimate is reported. E-PRTR CH<sub>4</sub> sources contributing to the observations of the noon transect on 6 June amount to roughly 80 kt/a, if only considering mining ventilation shafts in the near surroundings (20 km radius). However, far upwind (60 km) ventilation shafts can influence the measurements although all mines located in direct wind direction are listed with 0 kt/a and reach 13 kt/a somewhat north of the mean wind direction.

On 1 June (panels c) in Fig. 4), we performed a transect at the western border of the USCB (green dots in Fig. 1). The wind lidar data from the nearby DLR85 (11 km distance to the observations) revealed particularly low wind speeds ( $<2.5 \text{ ms}^{-1}$ ). Source attribution to individual ventilation shafts is challenging since a group of upwind mining shafts contributes to the measured enhancements.  $\Delta\text{XCH}_4$  values peaked at about 60 ppb around 9:30 UTC at the end of the 1.5 h north to south transect.

5 The distance of our observatory to the closest mining shaft was always greater than 4 km. The CH<sub>4</sub> emission for the group of shafts is estimated as  $109 \pm 33 \text{ kt/a}$ .

Clear sky conditions on the 6 June 2018 enabled us to observe two targets on one day, one in the morning and one around noon. In the morning around 7 am UTC, we started observing a mine with two shafts at the south-east border of the USCB (red dots in Fig. 1). With the wind blowing from east-northeast, we sampled 1 to 2 km west of the shafts (panels d) in Fig. 4). With the CRDS instrument on board, we were able to assess the rough position of the plume as we could see ground-based in-situ enhancements online while moving the truck. Once we reached background CH<sub>4</sub> levels in the south, we started sampling with the FTS and moved northward with an average step size of 100 m. Maximum enhancements  $\Delta\text{XCH}_4$  reached roughly 35 ppb around 7:30 am. Estimated emissions  $Q$  amount to  $17 \pm 3 \text{ kt/a}$ . We measured a difference of +4 ppb between southern and northern background values.

15 For the second transect on 6 June (purple dots in Fig. 1 and panels e) in Fig. 4), the wind was also from north-easterly directions. We aimed at measuring a cross section through the southern part of the USCB sampling roughly every 500 m moving from south to north. Downwind (by about 1 to 2 km) of a group of four shafts close to our track, we measured XCH<sub>4</sub> enhancements exceeding 30 ppb, from which we calculate total emissions  $Q$  of  $81 \pm 13 \text{ kt/a}$ .



**Figure 5.** Relative differences between the best-estimate (weighted averaged) wind speed and sensitivity calculations using wind speed from individual wind lidars. Different colors represent different transects. Individual symbols refer to the wind information source used. The right panel is a zoom of the left panel, but it omits data from 1 June.

#### 4 Error analysis

The errors reported for the emission estimates in section 3 are composed of several contributions from terms in equation (2): the estimated errors for  $\Delta\Omega$  which partition into the measurement error and the error for background removal; the errors associated with effective wind speed  $U_{\text{eff}}$ , which partition into errors related to vertical, horizontal, and temporal averaging of the wind-lidar measurements; and the errors in  $\Delta y_i$ , which are dominated by the errors in the relative angle between wind directions and driving directions.

As a measure for the measurement error attributed to  $\text{XCH}_4$ , we take the standard deviation among all measurements collected during an individual stop of our stop-and-go pattern. These standard deviations exhibit a maximum of 4 ppb for the noon transect on 24 May. Short distances to the mining shaft and wind speeds around  $6 \text{ ms}^{-1}$  imply large variability in the  $\text{CH}_4$  column above the FTS. For the same day in the morning, the relative  $\text{CH}_4$  standard deviations for a transect at the same shaft were smaller (2 ppb). The actual instrument precision of the FTS  $\text{XCH}_4$  measurements is on the order of roughly 0.3 ppb (0.02%). Thus, the error estimate is driven by atmospheric variability.

The error associated with  $\text{XCH}_4$  background removal is estimated as the standard deviation of all possible combinations of non-plume signals observed before and after crossing the plume during the generally 1 to 1.5 h long transects. We define an observation as intra-plume if the absolute difference between two consecutive measurements is at least 0.5 ppb greater than the average standard deviation of all measurements. We then linearly fit all possible background estimates as illustrated by gray dashed lines in panel a) of Fig. 4. The average of all these background options is defined as the best estimate background. Their standard deviation is considered in the error budget. The best-estimate  $\text{XCH}_4$  background is then used to obtain the plume

enhancements  $\Delta XCH_4$  by subtracting the background from the measured  $XCH_4$ . Overall the background removal standard deviation is smaller than 0.6 ppb which is small compared to other sources of error. For 6 June, we rely on the onboard in-situ measurements to locate the start and end point of the plume. We observed the in-situ concentrations on the fly and started with the FTS measurements once we observed constantly low  $CH_4$  concentrations.

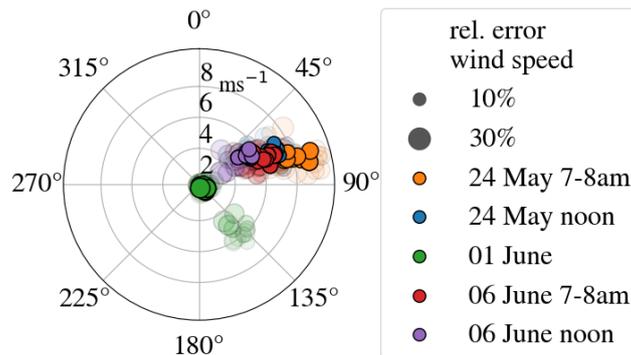
5 Errors related to the estimate of effective wind speed  $U_{\text{eff}}$  typically dominate the error budget. Our baseline wind scenario averages the lidar wind profiles vertically throughout the PBL and considers the errors arising from temporal wind speed variability throughout one plume transect and from distance-weighted averaging between the three wind lidars. We take the standard deviations among all wind speed measurements inside the PBL to quantify the error related to vertical averaging. The respective error estimates range from 8% to 18%, with larger errors occurring under low wind speed conditions as observed on  
10 1 June ( $U_{\text{eff}} < 2.5 \text{ ms}^{-1}$ ). The error due to temporal averaging of wind speed is estimated by the respective standard deviations of wind speed during the averaging period. This error ranges from 3% to 9%.

Having the advantage of three wind lidars co-deployed in the USCB, we can also assess  $U_{\text{eff}}$  errors related to horizontal variability of wind speeds. To this end, we conducted a sensitivity study that calculates the emissions  $Q$  with wind information from each single wind lidar instrument, instead of using a horizontal average that is weighted by the distance between lidar  
15 and FTS as in our baseline scenario. Distances between  $XCH_4$  observations and wind measurements range between 11 km and 72 km. Figure 5 shows the relative differences of calculated wind speeds  $U_{\text{eff}}$  for all sensitivity runs with respect to the best estimate baseline scenario (reported in section 3).

Differences between the wind speeds generally increase the larger the distance between the FTS and the wind lidar locations. Differences range between 1% and 13% except for distances exceeding 40 km and for low wind speed conditions such as  
20 encountered on 1 June, which implies larger errors. For 1 June with  $U_{\text{eff}} < 2.5 \text{ ms}^{-1}$ , relative wind speed differences differ substantially if wind information is taken from measurements more than a few ten kilometers away. Figure 5 also reveals, that using wind information from DLR89 causes emission estimates that are higher than for the other wind information sources. Wind measurements of DLR89 are likely influenced by the nearby lake, especially during the prevailing easterly winds. Low surface roughness and low friction above the lake surface could cause the wind speeds – and therefore the emission estimates  
25 – to be higher in lower levels compared to the other sites located close to forests. Since the distance of the FTS to the closest wind lidar never exceeded 33 km and since we used distance-weighted averages of the wind measurements, we estimate the error due to horizontal wind speed averaging as 10%.

Errors due to wind direction translate into errors of the cross-plume segment  $\Delta y_i$ . Similar to the error estimation procedures for wind speed, we estimated wind direction errors due to vertical, temporal and horizontal averaging. In general, these errors  
30 are smaller compared to wind speed errors. Vertical averaging yields a standard deviation of 2% to 5% except for the low wind speed day 1 June (14%). Temporal averaging induces errors ranging between 2% and 3% again except for 1 June with 12%. We estimate the error due to horizontal wind direction averaging with 3% according to the procedure we used to estimate the horizontal wind speed averaging error.

Generally, we take the above individual error contributions, and we propagate them into the reported emission errors for  
35  $Q$  by Gaussian error propagation. The error contributions from estimating the effective wind-speed  $U_{\text{eff}}$  dominate the error



**Figure 6.** The wind rose indicates wind direction and wind speed as occurred during the transects. Different colors mark different days. Size of the circles corresponds to the relative standard deviation in % resulting from averaging wind speed horizontally, within the PBL, and over time. Solid circles mark the best-estimate wind information whereas transparent circles mark wind information that stems from the individual wind lidars directly. During the measurement periods, easterly winds prevailed. Note that for 1 June wind speeds are exceptionally low.

budget. An overview of all wind situations for all transects is given in Figure 6. Every circle refers to a certain time during the transect and the circle size indicates the total relative  $U_{\text{eff}}$  error (calculated through Gaussian error propagation from the individual contributions given in Table 2). Solid circles mark best-estimate wind information. Transparent circles indicate data from our sensitivity study based on individual wind lidars. Mainly easterly winds were observed for the transects reported in this study. Wind speeds varied between  $4\text{ ms}^{-1}$  and  $8\text{ ms}^{-1}$  for all cases except for 1 June, which was challenging for emission estimation since wind speeds were generally low and variable in the deployment region.

## 5 Discussion & Conclusion

We demonstrate ground-based remote sensing of  $\text{XCH}_4$  enhancements in the plumes of hard coal mining ventilation shafts several kilometers downwind of the sources. Our observatory truck was equipped with an EM27/SUN FTS measuring direct-sunlight near-infrared absorption spectra. With transect periods of 1 to 1.5 h, the plumes were crossed in stop-and-go patterns. In combination with the data of 3 co-deployed wind lidars we were able to estimate the emissions for five transects on three days.

Our error analysis includes errors related to the FTS measurements and errors related to vertical, horizontal, and temporal averaging of wind speed and wind direction. Combined through Gaussian error propagation these errors range between 15% and 30%.

Generally, our best-estimate emissions and the respective errors show broad agreement with the E-PRTR report (European Pollutant Release and Transfer Register, <http://prtr.ec.europa.eu/>, 2014), which contains emission data reported annually by industrial facilities over Europe including the coal mining branch. This data set provides annual mean values of  $\text{CH}_4$  emissions for each mining facility registered. The annual means refer to whole mines, which typically have several shafts to ventilate  $\text{CH}_4$ . We distributed the annual means reported by E-PRTR evenly to all known shafts of every mine. However, our measurements are

restricted to a few hours per shaft, and ventilating CH<sub>4</sub> of coal mines is a variable process depending on the concentrations in the mine. Furthermore, turbulent processes influence the shape and position of the plume and therefore the measurements. Thus, comparing our snapshot-like measurements with the disaggregated E-PRTR reports can only provide a rough comparison.

Tab. 2 lists the estimated and the reported, and disaggregated emissions for all five transects. On 24 May, we observed a single shaft in the south-east of the USCB. The reported mine emissions are 9.63 kt/a. The cross-sectional flux method yields 6 ± 1 kt/a for the morning transect and 10 ± 1 kt/a for the transect around noon. Background CH<sub>4</sub> exhibits a south to north gradient, indicating that the measurements could have suffered from CH<sub>4</sub> inflow from unknown sources. An observed low-level jet is likely influencing the morning transect. However, differences between single transects are likely as the ventilation system is not emitting constantly. Source attribution becomes challenging measuring downwind of various mines and groups of shafts. Hence, no specific reported emission can be assumed for the transect on 1 June. However, reported emissions likely range between 50 kt/a (sum of nearest upwind mines < 10 km) and 150 kt/a (sum of south-eastern upwind part of the USCB). Our best estimate using the mobile FTS dataset is 109 ± 33 kt/a. Reported (24.3 kt/a) and estimated emissions (17 ± 3 kt/a) generally agree for the morning transect of 6 June. The noon transect on 6 June crossed several possible CH<sub>4</sub> sources although the reported emissions are roughly 80 kt/a. Our best estimate using the mobile FTS data is 81 ± 13 kt/a, which includes the reported value in the error range. However, many sources are located upwind in the northern part of the USCB within about 60 km distance of the measurements, which also could have influenced the observations but are not included in the sum of the reported emissions for that day.

Enhancing the sampling frequency of the FTS decreased the dwell times significantly (only affects 6 June). This made faster transects possible and helped to avoid changing wind and emission conditions. Faster tracking of the sun would be necessary to allow measuring while driving. Wind information is crucial when using the cross-sectional flux method as relative changes in wind speed equal relative changes in estimated emissions. Particularly low wind speed situations suffer from large discrepancies in the estimates when using wind information from different instruments. Thus, the error budget is dominated by atmospheric variability. Precise wind measurements on board the observational truck would help to reduce the wind induced errors. CH<sub>4</sub> imaging instruments could guide the mobile FTS providing position and extent of the target plume.

Summarized, our approach enables the emission estimation of CH<sub>4</sub> with good confidence (15 to 30%). However, the method is restricted to direct sunlight and stable wind conditions. Together with detailed wind information and in-situ CH<sub>4</sub> measurements, a modified mobile FTS is a flexible, fast (1 to 1.5 h), and accurate (combined relative error of 15 to 30%) possibility to estimate coal mine CH<sub>4</sub> emissions reliably.

## 6 Data availability

The data are available from the author upon request.

*Author contributions.* Andreas Luther wrote the paper except for section 2.2 which was written by Norman Wildmann who also took an active part during the campaign deploying the wind lidars and retrieving and providing wind lidar data. Andreas Luther,

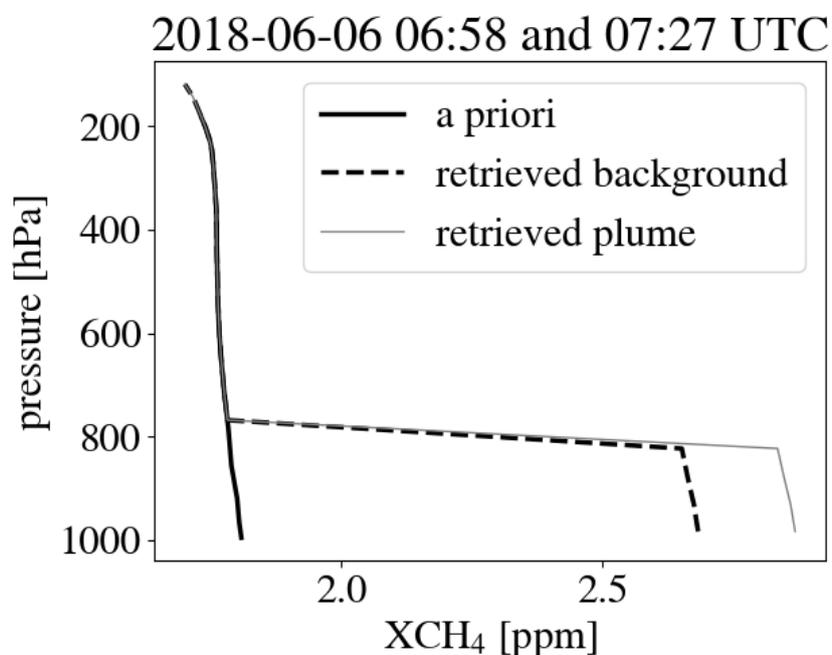
Ralph Kleinschek, Leon Scheidweiler, Sebastian Wolff, Sara Defratyka, Mila Stanisavljevic, Andreas Forstmaier, Alexandru Dandocsi, and Darko Dubravica operated the EM27/SUN spectrometer in the field during the campaign and collected and shared the data. Ralph Kleinschek and Julian Kostinek developed the suntracker software which made mobile measurements possible. Patrick Jöckel and Anna-Leah Nickl supported the selection of the target shafts with CH<sub>4</sub> forecasts for the region.

- 5 Theresa Klausner provided shaft-wise E-PRTR geolocation and emission data retrieved from the E-PRTR dataset (E-PRTR 2014). André Butz, Frank Hase, Matthias Frey, Jia Chen, and Florian Dietrich supported preparations for the measurement campaign, contributed to the spectral retrievals, and assisted with data postprocessing. Jarosław Nęcki and Justyna Swolkień provided detailed information about coal mining and CH<sub>4</sub> ventilation and functioned as local advisers. Andreas Fix coordinated the CoMet campaign operations. André Butz and Anke Roiger developed the research question.

10

*Competing interest.* The authors declare that they have no conflict of interest.

## Appendix A



**Figure A1.** A-priori versus retrieved CH<sub>4</sub> profile. Note, that the retrieval only scales the lower part of the a-priori profile up to the expected maximum PBL-height of roughly 800 hPa ( 1700m above ground). The gray line represents an intra plume measurement.

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