



1    **Development of a Universal Correction Algorithm for Filter-Based Absorption Photometers**

2    Hanyang Li<sup>1</sup>, Gavin R. McMeeking<sup>2</sup>, and Andrew A. May<sup>1</sup>

3    1 The Ohio State University Department of Civil, Environmental, and Geodetic Engineering,  
4    Columbus, Ohio, USA

5    2 Handix Scientific, LLC, Boulder, Colorado, USA

6    *Corresponding author:* Andrew A. May ([may.561@osu.edu](mailto:may.561@osu.edu)).

7    **Abstract**

8    Among the various measurement approaches to quantify light absorption coefficient ( $B_{abs}$ ), filter-  
9    based absorption photometers are dominant in monitoring networks around the globe. Numerous  
10   correction algorithms have been introduced to minimize the artifacts due to the presence of the  
11   filter in these instruments. However, from our recent studies conducted during the Fire Influence  
12   on Regional and Global Environments Experiment (FIREX) laboratory campaign, corrected filter-  
13   based  $B_{abs}$  remains biased high by roughly a factor of 2.5 when compared to a reference value  
14   using a photoacoustic instrument for biomass burning emissions. Similar over-estimations of  $B_{abs}$   
15   from filter-based instruments exist when implementing the algorithms on six months of ambient  
16   data from the Department of Energy (DOE) Atmospheric Radiation Measurement (ARM)  
17   Southern Great Plains (SGP) user facility from 2013 (factor of roughly 3). In both datasets, we  
18   observed an apparent dependency on single scattering albedo (SSA) and absorption Ångström  
19   exponent (AAE) in the agreement between  $B_{abs}$  based on existing correction factors and the  
20   reference  $B_{abs}$ . Consequently, we developed a new correction approach that is applicable to any  
21   filter-based absorption photometer that includes light transmission from the filter-based instrument  
22   as well as the derived AAE and SSA. For the FIREX and SGP datasets, our algorithm results in  
23   good agreement between all corrected filter-based  $B_{abs}$  values from different filter-based  
24   instruments and the reference (slopes  $\approx 1$  and  $R^2 \approx 0.98$  for biomass burning aerosols and slopes  
25    $\approx 1.05$  and  $R^2 \approx 0.65$  for ambient aerosols). Moreover, for both the corrected  $B_{abs}$  and the derived  
26   optical properties (SSA and AAE), our new algorithms work better or at least as well as the two  
27   common PSAP-based correction algorithms. The uncertainty of the new correction algorithm is  
28   estimated to be  $\sim 10\%$ , considering the measurement uncertainties of the operated instruments.  
29   Therefore, our correction algorithm is universally applicable to any filter-based absorption  
30   photometer and has the potential to “standardize” reported results across any filter-based  
31   instrument.

32    1. Introduction

33    Light-absorbing atmospheric aerosols directly affect the Earth’s energy budget by absorbing solar  
34   radiation, leading to a warming effect when they are suspended in the atmosphere and to the  
35   melting of snow and ice following deposition (Bond and Bergstrom, 2006; Boucher, 2015; Horvath,  
36   1993). For decades, scientists have conducted field experiments around the globe to investigate  
37   how absorbing aerosols influence the atmospheric radiative balance and interact with clouds (e.g.,  
38   Andrews et al. (2011); Cappa et al. (2016); Lack et al. (2008b); Rajesh and Ramachandran (2018);  
39   Schwarz et al. (2008)). These experiments may be performed at fixed stations (e.g., observation  
40   sites maintained by the Department of Energy (DOE) Atmospheric Radiation Measurement (ARM)  
41   program or the National Oceanic and Atmospheric Administration (NOAA) Global Monitoring



42 Division (GMD)) or on mobile platforms (e.g., car trailer, aircraft, and ship), typically involving  
43 the measurements of aerosol chemical, physical, and optical properties. Crucial to the  
44 quantification of the radiative forcing of absorbing aerosols are measurements of the absorption  
45 coefficient ( $B_{abs}$ ). For example, long-term monitoring of  $B_{abs}$  provides essential data to evaluate  
46 chemistry-climate model simulations (e.g., Chen et al. (2019); Vignati et al. (2010)), while  
47 intensive measurements of  $B_{abs}$  during short-term field campaigns allow for the investigation of  
48 optical properties that govern features of aerosol forcing (e.g., McMeeking et al. (2014); Olson et  
49 al. (2015)).

50 A variety of instruments have been used to measure  $B_{abs}$ , which generally classified into two large  
51 categories: filter-based techniques and photoacoustic techniques (Lack et al., 2014; Moosmüller  
52 et al., 2009). The major difference between the two categories of technique is that  $B_{abs}$  is measured  
53 after the aerosols are deposited on the filter media in the filter-based instruments, while the aerosols  
54 are characterized within an air stream in the photoacoustic instruments. Compared to the filter-  
55 based instruments, the photoacoustic instruments have the advantage of avoiding potential artifacts  
56 due to the contact of aerosols with filters; therefore, they are often used as the reference instruments  
57 in inter-comparison studies of aerosol absorption (e.g., Arnott et al. (2005); Davies et al. (2019);  
58 Jiang et al. (2018); Li et al. (2019); Schmid et al. (2006); Sheridan et al. (2005)).

59 Filter-based absorption photometers have been widely used at observational sites around the world  
60 due to their ease of operation and relatively low cost. Numerous instruments can be classified as  
61 filter-based absorption photometers including the Radiance Research Particle Soot Absorption  
62 Photometer (PSAP), the NOAA Continuous Light Absorption Photometer (CLAP), the Brechtel  
63 Manufacturing Tricolor Absorption Photometer (TAP), the Magee Scientific Aethalometer  
64 (AETH), and the Thermo Scientific Multi-Angle Absorption Photometer (MAAP). Operationally,  
65 all of these instruments are similar in that aerosols are deposited onto a filter and the reduction in  
66 the transmission ( $Tr$ ) of light by the particles (sometimes called attenuation (ATN)) is used to infer  
67  $B_{abs}$ . Where the instruments may differ is that some are multi-wavelength (multi- $\lambda$ ) instruments  
68 (e.g., 3 $\lambda$ -PSAP, CLAP, TAP, 7 $\lambda$ -AETH models), while others are not (e.g., 1 $\lambda$ -PSAP, other AETH  
69 models, MAAP).

70 One challenge with filter-based absorption photometers is that biases can arise due to the presence  
71 of the filter. For example, light scattering by particles loaded onto the filter or by the filter itself  
72 may affect the transmission of light (e.g., (Arnott et al., 2005; Bond et al., 1999)); non-absorbing  
73 material may result in absorption enhancement (e.g., (Cappa et al., 2008)); or organic vapors  
74 adsorbed to the filter may itself absorb light (e.g., (Subramanian et al., 2007)). Consequently,  
75 various correction algorithms exist to minimize these biases, but they are often specific only to  
76 certain instruments. For example, some are applicable to the PSAP, CLAP, and TAP (e.g., (Bond  
77 et al., 1999; Müller et al., 2014; Ogren, 2010; Virkkula, 2010; Virkkula et al., 2005)), while others  
78 are applicable to the AETH (e.g., (Arnott et al., 2005; Collaud Coen et al., 2010; Drinovec et al.,  
79 2017; Kirchstetter and Novakov, 2007; Schmid et al., 2006; Virkkula et al., 2007, 2015;  
80 Weingartner et al., 2003)).

81 Although the equations associated with these existing correction algorithms are different, they  
82 share some commonalities. For example, the filter-based absorption photometers are assessed  
83 using laboratory (e.g., ammonium sulfate, fullerene soot) or ambient aerosols during experiments  
84 which include reference measurements of  $B_{abs}$ . These reference measurements often include either  
85 direct photoacoustic  $B_{abs}$  or inferred  $B_{abs}$  as the difference between the extinction coefficient ( $B_{ext}$ )



86 and the scattering coefficient ( $B_{\text{scat}}$ ). Correction equations are developed by comparing data  
87 between the filter-based instrument and the reference instrument, where the equations often  
88 contain one term that accounts for filter loading effects and another that accounts for multiple-  
89 scattering effects. Consequently, the correction equations frequently incorporate both  $\text{Tr}$  and either  
90  $B_{\text{scat}}$  or the single-scattering albedo (SSA) to account for these effects. However, even when the  
91 correction algorithms are applied, potential issues can remain such as:

- 92 1. Corrected filter-based  $B_{\text{abs}}$  may remain biased high relative to a reference value of  $B_{\text{abs}}$  (e.g.,  
93 (Arnott et al., 2003; Davies et al., 2019; Lack et al., 2008a; Li et al., 2019; Müller et al., 2011a)).
- 94 2. Comparisons between the reference instrument and  $B_{\text{abs}}$  corrected by different algorithms can  
95 yield variable agreement (e.g., (Collaud Coen et al., 2010; Davies et al., 2019; Saturno et al.,  
96 2017)).
- 97 3. Corrected  $B_{\text{abs}}$  from different filter-based absorption photometers may not agree (e.g., (Davies  
98 et al., 2019; Müller et al., 2011a)).
- 99 4. Derived products (such as absorption Ångström exponents (AAE)) may differ based on the  
100 implemented correction algorithm (e.g., (Backman et al., 2014; Davies et al., 2019)).
- 101 5. The agreement between measurements of  $B_{\text{abs}}$  and estimates of  $B_{\text{abs}}$  by chemistry-climate  
102 models may vary based on the implemented correction algorithm (e.g., (Alvarado et al., 2016)).

103 The first three issues in this list may arise due to differences in aerosol optical properties between  
104 those used in deriving the correction equation and those associated with a given aerosol sample,  
105 and these issues can propagate through to the fourth issue. The final issue is arguably most  
106 important because evaluation of chemistry-climate models may be severely affected by the  
107 differences between different correction algorithms, which may inhibit the modeling community  
108 from providing accurate projections of future temperature and precipitation response.

109 In this work, we seek to address some of these issues. First, we evaluate the CLAP, TAP, and  
110 PSAP using two common PSAP-based correction algorithms, namely Bond et al. (1999) as  
111 updated by Ogren (2010) and Virkkula et al. (2005) as updated by Virkkula (2010). For brevity,  
112 we refer to these corrections as “B1999” and “V2005” for Bond et al. (1999) and Virkkula et al.  
113 (2005), respectively, incorporating their respective updates. In addition, we propose “universal”  
114 correction algorithms that are applicable to any filter-based absorption photometer (e.g., CLAP,  
115 TAP, PSAP, and AETH) across multiple wavelengths by combining observed filter-based  $B_{\text{abs}}$   
116 with  $B_{\text{scat}}$  (e.g., from a co-located nephelometer (NEPH)) and reference  $B_{\text{abs}}$  (e.g., from a co-  
117 located photoacoustic instrument). However, in reality (e.g., at long-term observatories), reference  
118 values of  $B_{\text{abs}}$  are rare, and in some cases, complementary  $B_{\text{scat}}$  measurements may not exist;  
119 consequently, we also provide methods to correct filter-based  $B_{\text{abs}}$  data in these scenarios. To our  
120 knowledge, this is the first study to simultaneously evaluate B1999 and V2005 corrections on  
121 PSAP “successors” (i.e., CLAP and TAP) and to present a correction algorithm that is broadly  
122 applicable to any filter-based absorption photometer. Regarding the latter, even if our correction  
123 algorithm has its own limitations, its use can nevertheless standardize the reporting of  $B_{\text{abs}}$  in long-  
124 term datasets.

## 125 2. Methodology

126 We developed the general form for our correction algorithms using CLAP and TAP measurements  
127 collected from biomass burning (65 fires in total) during the Fire Influence on Regional to Global  
128 Environments Experiment (FIREX) laboratory campaign in 2016. By using biomass burning



129 emissions, we considered a dataset spanning a broader range of aerosol optical properties (SSA at  
130 652 nm: 0.14-0.98; AAE: 1.25-4.73) than has traditionally been used in developing these  
131 correction algorithms. We then conducted further evaluation and validation of the model using  
132 ambient data, specifically using CLAP measurements from the DOE ARM Southern Great Plains  
133 (SGP) user facility in Lamont, OK, USA (02/01/13 to 07/09/13). Our algorithms were then  
134 extended to the AETH data from the FIREX laboratory campaign and the PSAP data collected at  
135 the SGP site to verify the “universal” nature of the algorithms.

136 2.1. The FIREX campaign

137 2.1.1. Experimental setup

138 In October and November of 2016, we participated in the laboratory portion of the FIREX  
139 campaign to investigate the wildfire smoke and their impact on the atmosphere. During the  
140 campaign, over 100 burns took place at the U.S. Forest Service’s combustion facility at the Fire  
141 Sciences Laboratory (FSL). The fuels burned in this study are representative of western US  
142 ecosystems, such as spruce, fir, various pines, and “chaparral” biome (e.g., manzanita, chamise).  
143 (see Koss et al. (2018) and Selimovic et al. (2018) for more details).

144 A typical burn lasted for 1-3 hours depending on the smoke sampling strategies (e.g., stack burns  
145 versus room burns). During each burn, one or multiple “snapshots” of smoke (typical  $B_{abs}$  at 652  
146 nm ranged from 100 to 1200  $Mm^{-1}$ ) were transferred from the combustion room at FSL into a  
147 mixing chamber (210 L) through a long transfer duct (30 m in length, 8" in diameter). The smoke  
148 was then diluted by filter air (~230 LPM) in the chamber. Once the concentration in the chamber  
149 was stable (detected by the Photoacoustic Extinctiometer (PAX) which was operated continuously  
150 through all fires), the smoke was passed to a suite of instruments to obtain aerosol and gas phase  
151 parameters. This chamber also served as an intermediate between the transfer duct and the  
152 instrumentation to minimize potential biases that arose due to different sample flow rates and  
153 sample locations of the instruments. A more detailed description of our experiments can be found  
154 in Li et al. (2019).

155 2.1.2. Measurements of aerosol optical properties

156 During the campaign, five instruments provided measurements of  $B_{abs}$  (CLAP, NOAA GMD; TAP,  
157 Brechtel Manufacturing Inc. (BMI); Aethalometer (Model AETH-31), Magee Scientific; and two  
158 PAXs (Model PAX-870 and PAX-405), Droplet Measurement Technologies) and two instruments  
159 provided measurements of  $B_{scat}$  (PAX-870 and PAX-405). The instruments included in the present  
160 work are summarized in Table 1.

161 Both CLAP and TAP provide  $B_{abs}$  measurements of the particles deposited on a filter, similar to  
162 PSAP. Different from PSAP, there are multiple filter spots (8 sample spots and 2 reference spots)  
163 cycling of one filter in CLAP and TAP, enabling the instruments to run continuously through two  
164 or three burns without changing filter. In the CLAP and TAP, sample illumination is provided by  
165 LEDs operated at three wavelengths (467, 528, and 652 nm). Here, we apply both B1999 and  
166 V2005 to CLAP and TAP data, similar to previous work (e.g., (Backman et al., 2014; Davies et  
167 al., 2019)).

168 The key differences between the CLAP and TAP during the FIREX campaign include:



- 169 1. The spot change of the CLAP was manually performed when Tr reached approximately 0.5
- 170 (or ATN decreased to ~69), while the TAP advanced to a new spot automatically with a Tr
- 171 threshold set to be 0.5.
- 172 2. The spot area, flow rate, and LED-detected wavelengths differed slightly (Table 1).
- 173 3. The CLAP recorded  $B_{abs}$  every one minute, while the TAP recorded  $B_{abs}$  every ten seconds. To
- 174 enable the following analysis, we compute the 1-minute averages of TAP-derived parameters.
- 175 4. For the first portion of the campaign (the first 17 days of the 45-day campaign), Pallflex E70-
- 176 2075S filters were used in the CLAP while Azumi filters (model 371M, Azumi Filter Paper
- 177 Co., Japan) were used in the second portion of the campaign (due to a lack of availability of
- 178 the Pallflex filters). The TAP was equipped exclusively with the Azumi filters throughout the
- 179 campaign. We apply the filter correction recommended in Ogren et al. (2017) to the CLAP and
- 180 convert from Pallflex to Azumi filters.
- 181 5. BMI substantially re-engineered the CLAP in their development of the TAP.

182 These differences resulted in variable agreement between the CLAP and TAP during FIREX;  
183 however, the two instruments did largely agree within experimental uncertainty (e.g., see Fig. S8  
184 and Fig. S13 in Li et al. (2019)).

185 A PAX measures  $B_{abs}$  and  $B_{scat}$  simultaneously for suspended particles using a modulated diode  
186 laser. We use these photoacoustic absorption measurements as the reference to evaluate the filter-  
187 based  $B_{abs}$  and develop our correction algorithms. To enable the evaluation of CLAP and TAP  
188 which operate at different wavelengths than the PAXs, we interpolate the measurements of  $B_{abs}$   
189 and  $B_{scat}$  to the wavelengths of 467, 528, and 652 nm using the values of AAE and scattering  
190 Ångström exponents (SAE), similar to Backman et al. (2014) and Virkkula et al. (2005).  
191 Theoretically, AAE and SAE fit absorption and scattering as power law functions of wavelength  
192 (Bergstrom et al., 2007).

193 Due to the numerous correction algorithms for the Aethalometer (e.g. (Arnott et al., 2005; Collaud  
194 Coen et al., 2010; Kirchstetter and Novakov, 2007; Saturno et al., 2017; Schmid et al., 2006;  
195 Virkkula et al., 2007; Weingartner et al., 2003)), we do not evaluate these in the present work to  
196 limit the scope. In fact, the majority of our focus is the B1999 and V2005 corrections to TAP and  
197 CLAP. However, we still test the performance of the new algorithms on the AETH to explore its  
198 applicability to that instrument.

## 199 2.2. Measurements of aerosol optical properties at the SGP observatory

200 The ambient data used in this manuscript are the ground-based aerosol data measured at the SGP  
201 observatory from 02/01/13 to 07/09/13 (archived at <https://www.archive.arm.gov/discovery/>). For  
202 evaluation purposes, we randomly selected a range of dates during which the observations are  
203 valid (without incorrect, suspect, and missing data). This time period was also subsequent to an  
204 upgrade to the 532 nm laser in the three-wavelength photoacoustic soot spectrometer (PASS-3).

205 At the site, an impactor was used to switch the sampling between two cutoffs (particle diameter  
206  $<10 \mu\text{m}$  (PM10) in the first 30 minutes of each hour and  $<1 \mu\text{m}$  (PM1) in the latter 30 minutes of  
207 each hour). The aerosols exiting from the impactor were dried to RH less than 40% and passed to  
208 a CLAP, a PSAP, and two NEPHs. The PASS-3 operated at the site and also provides  $B_{abs}$  and  
209  $B_{scat}$  for aerosols, but these samples did not pass through the impactor (e.g., characterizing total  
210 suspended particles (TSP)). Typical  $B_{abs}$  and  $B_{scat}$  reported at the site ranged from 0 to  $10 \text{ Mm}^{-1}$   
211 and 0 to  $50 \text{ Mm}^{-1}$  at 550 nm, respectively (e.g., (Sherman et al., 2015)). Although the site is rural



212 (clean background air), long-term transport aerosols (such as mineral dust, absorbing organic  
213 aerosols, and secondary organic aerosols (SOA)) may affect the local aerosol properties (Andrews  
214 et al., 2019).

215 We preprocess the SGP data in three steps. First, due to the systematic difference of aerosol sizes  
216 between PASS-derived and filter-based absorption, we only include the PM10 observations,  
217 inherently assuming that any differences in the optical properties of PM10 and TSP are negligible.  
218 Then, we smooth the 1-second data into 10-minute averages. Thirdly, we estimate the detection  
219 limits at each of the three wavelengths in the PASS-3 using the data measured during the  
220 “background zero” periods (Allan, 1966) and discard the observations which are below the  
221 detection limits. With a 10-min-averaging-time, the detection limits ( $3\sigma$ ) for the PASS-3 are  $0.78$   
222  $Mm^{-1}$  (405 nm),  $2.01 Mm^{-1}$  (532 nm), and  $0.30 Mm^{-1}$  (781 nm). For the filter-based instruments,  
223 the detection limits are based on previous studies (See Table 1). Moreover, we only retain the  
224 observations that satisfy  $B_{abs}$  (405 nm)  $> B_{abs}$  (532 nm)  $> B_{abs}$  (781 nm) (or  $AAE > 0$ ), similar to  
225 Fischer and Smith (2018). As with the PAX data from the laboratory, we adjust the PASS-derived  
226  $B_{abs}$  to 467, 528, and 652 nm using the inferred AAE values for each 10-minute average.



227 **Table 1** Summary of specifications for instruments relevant to this work.

Instrument	Flow rate (LPM)	Spot area (cm <sup>2</sup> )	Type of filter	Measured parameters	Response time	Measurement uncertainty	Detection limit (3 $\sigma$ , Mm <sup>-1</sup> )
PAX-870	1.0	-	-	B <sub>abs</sub> and B <sub>scat</sub> (870 nm)	1s	~11% (B <sub>abs</sub> ) ~17% (B <sub>scat</sub> ) (Nakayama et al., 2015)	0.47 (B <sub>abs</sub> ) 0.66 (B <sub>scat</sub> ) <sup>a</sup>
PAX-405	1.0	-	-	B <sub>abs</sub> and B <sub>scat</sub> (405 nm)	1s	4% (B <sub>abs</sub> ) 7% (B <sub>scat</sub> ) (Nakayama et al., 2015)	0.27 (B <sub>abs</sub> ) 0.60 (B <sub>scat</sub> ) <sup>a</sup>
PASS-3 <sup>b</sup>	1.0	-	-	B <sub>abs</sub> and B <sub>scat</sub> (405, 532, and 781 nm)	1s	4 %, 8 %, and 11 % (B <sub>abs</sub> ) (Nakayama et al., 2015)	0.78 (405 nm) 2.01 (532 nm) 0.30 (781 nm) <sup>a</sup>
NEPH <sup>b</sup>	7.5	-	-	B <sub>scat</sub> (450, 550, and 700 nm)	1s	10% (Anderson et al., 1996)	0.29 (450 nm) 0.11 (550 nm) 0.17 (700 nm) (5-min average) (Müller et al., 2011b)
CLAP	0.83 ± 0.02 (FIREX) 0.945 (SGP)	0.199 (FIREX) 0.195 (SGP)	Pallflex E70-2075S and Azumi filter (model 371M) <sup>c</sup>	B <sub>ATN</sub> and Tr <sup>d</sup> (467, 529, and 653 nm)	60s	30% (Ogren et al., 2017)	0.6 (1-min average), 0.12 (10-min average) (Ogren et al., 2017)
TAP	1.26 ± 0.01	0.253	Azumi filter (model 371M) <sup>c</sup>	B <sub>ATN</sub> and Tr <sup>d</sup> (467, 528, and 652 nm)	10s	30% (Laing et al., 2016)	2.67 (467 nm) 4.11 (528 nm) 2.13 (652 nm) (30-s average) (Davies et al., 2019)
AETH	2.4	0.5	quartz fiber sampling tape	B <sub>ATN</sub> and Tr (370, 470, 520, 590, 660, 880, and 950 nm)	120s	10% (Sedlacek, 2016)	0.1 (Sedlacek, 2016)
PSAP	1.0	0.178	Pallflex E70-2075W	B <sub>ATN</sub> and Tr <sup>d</sup> (470, 522, and 660 nm)	60s	-15% (Bond et al., 1999)	0.3 (Springston, 2016)

228 <sup>a</sup> The detection limits of PAX and PASS-3 are determined by Allan deviation analysis (Allan, 1966) of B<sub>abs</sub> during “background zero”.

229 <sup>b</sup> During the analysis of the data collected at the SGP, we use B<sub>abs</sub> derived by the PASS and B<sub>scat</sub> derived by the NEPH to yield the coefficients in the algorithms.

230 <sup>c</sup> Two types of filters were used during the FIREX campaign (See Sect. 2.1.2).

231 <sup>d</sup> The operating wavelengths of CLAP, TAP, and PSAP are stated slightly different by the instrument manufactures. We simply use 467, 528, and 652 nm throughout  
 232 this manuscript.



233 **Table 2** Overview of the studies of B1999 and V2005 and the description of our experiments.

Study	Aerosol source	SSA subset	Range of $B_{abs}$ (Mm $^{-1}$ )	Filter-based instrument for $B_{abs}$	Reference instrument for $B_{abs}$ <sup>a</sup>	Instrument for $B_{scat}$ <sup>a</sup>	Coefficient values in the correction algorithm <sup>b, c</sup>		
The study in B1999	Lab-generated aerosols, including various mixtures of nigrosin and ammonium sulfate	0.5-1 (550 nm)	0-800 (550 nm)	One- $\lambda$ PSAP (550 nm)	The difference between extinction (OEC) and scattering coefficient (NEPH) <sup>d</sup>	NEPH (450, 550, and 700 nm)	$C_1 = 0.016 \pm 0.023$ (550 nm) $C_2 = 1.55 \pm 0.25$ (550 nm) $C_3 = 1.02 \pm 0.17$ (550 nm)		
The laboratory study in V2005	Lab-generated aerosols, including various mixtures of kerosene soot, ammonium sulfate, and polystyrene latex	0.2-0.9 (530 nm)	0-800 (530 nm)	One- $\lambda$ PSAP (550 nm), three- $\lambda$ PSAP (467, 530, 660 nm)	The average of the PA (532 nm and 1064 nm) and the difference between extinction (OEC) and scattering coefficient (NEPH) <sup>d</sup>	NEPH (450, 550, and 700 nm)	467 nm	530 nm	660 nm
The ambient study in V2005	Ambient aerosols measured during RAOS and NEAQS <sup>e</sup>	0.75-1 (530 nm)	0-15 (530 nm)	One- $\lambda$ PSAP (550 nm), three- $\lambda$ PSAP (467, 530, 660 nm)	PA (532 nm and 1064 nm)	NEPH (450, 550, and 700 nm)	$C_1 = 0.015$	$C_1 = 0.017$	$C_1 = 0.022$
FIREX	Biomass burning aerosols under relatively controlled laboratory conditions	0.2-1 (550 nm)	38-1800 (550 nm)	CLAP (467, 529, 652 nm), TAP (467, 528, 653 nm), AETH (370, 470, 520, 590, 660, 880, 950 nm)	PAX (405 nm and 870 nm)	PAX (405 nm and 870 nm)	$C_4 = 0.377 \pm 0.013$	$C_4 = 0.358 \pm 0.011$	$C_4 = 0.352 \pm 0.013$
SGP (02/01/13 to 07/09/13)	Ambient aerosols collected at the SGP user facility in Lamont, OK	0.75-1 (530 nm)	0-8 (550 nm)	CLAP (461, 522, 653 nm), PSAP (470, 522, 660 nm)	PASS (405, 532, and 781 nm)	NEPH (450, 550, and 700 nm)	$C_5 = -0.640 \pm 0.007$	$C_5 = -0.640 \pm 0.007$	$C_5 = -0.674 \pm 0.006$
							$C_6 = 1.16 \pm 0.05$	$C_6 = 1.17 \pm 0.03$	$C_6 = 1.14 \pm 0.11$
							$C_7 = -0.63 \pm 0.09$	$C_7 = -0.71 \pm 0.05$	$C_7 = -0.72 \pm 0.16$
							See Table 4 and Tables S6-S10		

234 <sup>a</sup> The operating wavelengths are based on the manufacturer specifications.

235 <sup>b</sup> The coefficients provided in Table 2 are the values presented in Ogren (2010) and Virkkula (2010), which are updated from Bond et al. (1999) and Virkkula et  
 236 al. (2005), respectively.

237 <sup>c</sup> We reformulate the correction equations in the original publications to agree with Eq. (4)–(6) in this manuscript.  $C_1$  to  $C_7$  are the coefficients in the present work.

238 <sup>d</sup> OEC is optical extinction cell and PA is the instrument using photoacoustic technique.

239 <sup>e</sup> RAOS and NEAQS are Reno Aerosol Optics Study and New England Air Quality Study, respectively.



240 2.3. The correction algorithms

241 In filter-based instruments, the light intensities transmitted through the sample spot and blank spot  
242 of the filter are recorded as  $I_s$  and  $I_b$ , respectively. The logarithmic ratio of the two intensities at  
243 time  $t$  is defined as ATN using the Beer-Lambert law:

244 
$$ATN(t) = -100 \times \ln \frac{I_s(t)}{I_b(t)} \quad 1$$

245 where  $ATN = 0$  when beginning a new filter spot ( $t = 0$ ).

246 The ATN can be related to  $Tr$  by normalizing  $I_s/I_b$  at time  $t$  relative to  $I_s/I_b$  at the start of a new  
247 filter spot ( $t = 0$ ):

248 
$$Tr(t) = \frac{I_s(t)/I_b(t)}{I_s(0)/I_b(0)} = \exp\left(\frac{-ATN(t)}{100}\right) \quad 2$$

249 The change of ATN over a time interval ( $\Delta t$ ) for the instrument operated at a volume flow rate of  
250  $Q$  and spot area of  $A$  yields the attenuation coefficient ( $B_{ATN}$ ) for that time interval:

251 
$$B_{ATN} = \frac{A}{Q \times \Delta t} \times \Delta ATN \quad 3$$

252  $B_{ATN}$  is finally converted to  $B_{abs}$  by applying correction algorithms. The general form of the  
253 correction algorithms presented for the PSAP in Bond et al. (1999) and Virkkula et al. (2005) can  
254 be summarized as:

255 
$$B_{abs} = B_{ATN} \times f(Tr) - C_1 \times B_{scat} \quad 4$$

256 where  $f(Tr)$  is some function of  $Tr$  (that may vary between approaches), correcting for the filter  
257 loading effect.  $C_1$  is a constant that may vary with wavelength; specifically, it is a penalty for the  
258 light scattering by the particles collected on the filter which may contribute to the quantification  
259 of ATN. In most atmospheric and laboratory studies,  $B_{scat}$  is measured independently, typically  
260 using a co-located NEPH.

261 2.3.1. The B1999 correction

262 Bond et al. (1999) was the first study to present the correction algorithm for filter-based  
263 instruments. This empirical correction was originally developed for the PSAP operated at 550 nm  
264 using various mixtures of laboratory-generated nigrosin ( $SSA \approx 0.5$ ) and ammonium sulfate ( $SSA \approx 1$ ) with  $B_{abs}$  ranged from 0 to 800  $Mm^{-1}$ .

266 After calibrating the flow rate and spot area of the PSAP, the authors derived  $C_1 = 0.016$  and

267 
$$f(Tr)_{B1999} = \frac{1}{C_2 \times Tr + C_3} \quad 5$$

268 where  $C_2 = 1.32$  and  $C_3 = 0.87$  (after combining Eq. (3) and Eq. (12) from Bond (1999)).

269 The equation parameters were further clarified in Ogren (2010) who adjusted the B1999-measured  
270 spot area ( $A = 20.43 \text{ mm}^2$ ) to be consistent with the universal area of the PSAP ( $A = 17.83 \text{ mm}^2$ ).  
271 Ogren (2010) also extended the correction to 574 nm using a wavelength dependence of  $B_{abs}$  ( $B_{abs} \sim \lambda^{-0.5}$ ). Consequently,  $C_2$  and  $C_3$  in  $f(Tr)$  were updated to 1.55 and 1.02, respectively. These are  
272 the values used in the present work (Table 2) for B1999. Moreover, Ogren (2010) stated that the  
273 correction forms of Eq. (4) and Eq. (5) were valid for any wavelength, while additional  
274



275 experiments were needed to establish the equation parameters for the wavelengths other than 574  
276 nm.

277 2.3.2. The V2005 correction

278 Virkkula et al. (2005) developed a correction algorithm for both three-wavelength PSAP (467, 530,  
279 and 660 nm) and one-wavelength PSAP (574 nm) using the same functional form as Eq. (4). Since  
280 the operating wavelengths of the photoacoustic instruments and the NEPH were different from  
281 those of the PSAP, the measured photoacoustic  $B_{abs}$  and  $B_{scat}$  was extrapolated or interpolated to  
282 467, 530, and 660 nm, using inferred AAE and SAE respectively. In this study, the authors used  
283 various mixtures of kerosene soot, ammonium sulfate, and polystyrene latex (SSA ranged from  
284 0.2 to 0.9) with  $B_{abs}$  ranging from 0 to 800  $Mm^{-1}$  at 530 nm.

285 Different from the  $f(Tr)$  in the B1999 correction which was a reciprocal function of  $Tr$ , the  $f(Tr)$   
286 presented in V2005 was a multi-variate linear function of the natural logarithm of  $Tr$  and SSA  
287 (including an interaction term between the two):

288 
$$f(Tr(\lambda), SSA(\lambda))_{V2005} = C_4 + C_5 \times (C_6 + C_7 \times SSA(\lambda)) \times \ln(Tr(\lambda)) \quad 6$$

289 where the parameters in Eq. (6) vary with wavelengths. The parameters in V2005 were updated in  
290 Virkkula (2010) by correcting for flowmeter calibration (Table 2).

291 Due to the unknown values of SSA before deriving  $B_{abs}$ , Virkkula et al. (2005) provided a solution  
292 through an iterative procedure. In the iteration,  $B_{abs}$  is first calculated using the B1999 correction  
293 (e.g., Eq. (4) and Eq. (5)) and is then used to compute the initial guess of SSA for use in Eq. (6).  
294 The  $B_{abs}$  and SSA can be updated using Eq. (4) and Eq. (6) until convergence is reached.

295 2.3.3. The new correction

296 We develop a set of new correction algorithms with the same general form as Eq. (4) using the  
297 biomass burning emissions from 65 different burns during the FIREX laboratory study, providing  
298 a broader range of aerosol optical properties and aerosol concentrations than previous work. This  
299 was motivated by the disagreement that remained between filter-based and photoacoustic  
300 instruments, even after applying B1999 to the data (e.g., see Li et al. (2019) Fig. 4 and our Fig. 2  
301 below). These differences may persist because we were effectively extrapolating the B1999  
302 correction equation to values outside the range for which it was developed.

303 This new correction is developed based on multiple linear regression techniques with three  
304 dependent variables of  $\ln(Tr)$ , SSA, and AAE and one independent variable of  $B_{abs}/B_{ATN}$  (Eq. (7)  
305 – (9)). As with other correction equations, this model takes into account the influence of scattering  
306 and weakly-absorbing materials. However, we target two additional aims: 1) extend the correction  
307 a wider range of  $B_{abs}$ ; and 2) develop a model that is applicable to any filter-based instrument.

308 Similar to the B1999 and V2005 corrections, this new model starts with the general form of Eq.  
309 (4), re-written here to define  $B_{scat}$  in terms of SSA and  $B_{abs}$ .

310 
$$B_{abs}(\lambda) = B_{ATN}(\lambda) \times f(Tr(\lambda)) - C_1 \times \frac{SSA(\lambda)}{1-SSA(\lambda)} \times B_{abs}(\lambda) \quad 7$$

311 Re-arranging this equation to move all  $B_{abs}$  terms to the left-hand side yields:

312 
$$B_{abs}(\lambda) = B_{atn}(\lambda) \times g(Tr(\lambda), SSA(\lambda)) \quad 8$$



313 where  $g(\text{Tr}(\lambda), \text{SSA}(\lambda)) = f(\text{Tr}(\lambda)) \times \frac{1-\text{SSA}(\lambda)}{1-(1-C_1) \times \text{SSA}(\lambda)}.$

314 We define the function “g” as a multivariate linear model, introducing AAE as a dependent  
315 variable and including interaction terms between SSA, AAE, and  $\ln(\text{Tr})$ :

316  $g(\text{Tr}(\lambda), \text{SSA}(\lambda), \text{AAE}) = G_0 + G_1 \times \ln(\text{Tr}(\lambda)) + G_2 \times \text{SSA}(\lambda) + G_3 \times \text{AAE} + G_4 \times \ln(\text{Tr}(\lambda)) \times \text{SSA}(\lambda) +$

317  $G_5 \times \ln(\text{Tr}) \times \text{AAE} + G_6 \times \text{SSA}(\lambda) \times \text{AAE} + G_7 \times \text{SSA}(\lambda) \times \text{AAE} \times \ln(\text{Tr}(\lambda)) \quad 9$

318 Equation (9) suggests that different combinations of SSA, AAE and  $\ln(\text{Tr})$  can result in the same  
319 value of “g” (i.e.,  $B_{\text{abs}}/B_{\text{ATN}}$ ); likewise, a given value of  $B_{\text{abs}}/B_{\text{ATN}}$  may have infinitely many points  
320 with distinct slopes passing through it (Fig. S3). Therefore, in order to properly compensate for  
321 the effects of loading and aerosol optical properties, a multiple linear regression with interaction  
322 terms is required.

323 A detailed description of the procedure for the model development (e.g., variable transformation  
324 (from  $\text{Tr}$  to  $\ln(\text{Tr})$ ), variable selection using best-subsets and stepwise approaches, and model  
325 validation) is provided in the Supplementary Material.

326 As in V2005, iteration is required in our algorithm because  $B_{\text{abs}}$  is dependent on knowledge of  
327 SSA and AAE, which themselves are dependent on  $B_{\text{abs}}$ . We propose the following iterative  
328 process to update SSA and AAE in the model.

- 329 1. Initialize AAE from  $B_{\text{ATN}}$  across the three wavelengths ( $B_{\text{ATN}} \sim \lambda^{-\text{AAE}}$ ) and initialize SSA for  
330 each wavelength using  $B_{\text{ATN}}$  from the filter-based absorption photometer and  $B_{\text{scat}}$  from a co-  
331 located NEPH, i.e.,  $\text{SSA}(\lambda) = \frac{B_{\text{scat}}(\lambda)}{B_{\text{scat}}(\lambda) + B_{\text{ATN}}(\lambda)}.$
- 332 2. Yield an initial set of coefficients  $G_0$  through  $G_7$  for each wavelength to calculate  $g(\text{Tr}, \text{SSA},$   
333  $\text{AAE})$  in Eq. (9), using one of the Algorithms described in Sect. 2.4.
- 334 3. Calculate  $B_{\text{abs}}$  for each wavelength using Eq. (8).
- 335 4. Update AAE and SSA using  $B_{\text{abs}}$  calculated in Step 3.
- 336 5. Derive a new set of coefficient values.
- 337 6. Iterate Steps 3-5 until converged.

338 2.4. Application of correction algorithms

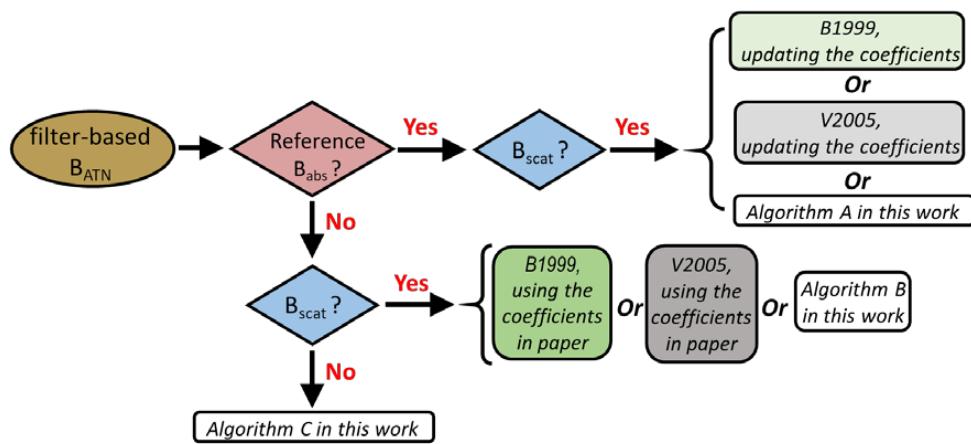
339 In developing a procedure for applying our algorithm, we envision three potential scenarios:

- 340 1. Algorithm A: The filter-based instrument is co-located with a NEPH and reference instrument  
341 providing  $B_{\text{abs}}$ . This scenario facilitates the computation of  $G_0$  through  $G_7$  in Eq. (9) (step 2 in  
342 the iterative process) as well as the derivation of new coefficients for existing correction  
343 algorithms. This scenario can also enable the development of a new set of coefficients that may  
344 be more appropriate for aerosol sources that we do not consider here.
- 345 2. Algorithm B: The filter-based instrument is co-located with a NEPH but not a reference  
346 instrument providing  $B_{\text{abs}}$ , which is perhaps the most likely scenario (at least at many long-  
347 term monitoring sites). This scenario requires an initial guess of the coefficients; we provide  
348 sets of these in Table 4 below for different filter-based instruments and aerosol sources.
- 349 3. Algorithm C: The filter-based instrument is deployed with neither a co-located NEPH nor a  
350 reference instrument providing  $B_{\text{abs}}$ . This scenario is the most challenging, because there are  
351 no measurements of  $B_{\text{scat}}$  to compute SSA; to address this issue, we propose the use of a non-



352 linear relationship between SSA and AAE ( $AAE = a + b \times SSA^c$ ) to provide an initial guess of  
353 SSA in the iterations.

354 To aid in decision-making between algorithms, we developed a flow chart for selecting appropriate  
355 correction algorithm for CLAP, TAP, and PSAP (Fig. 1). Furthermore, an Igor Pro (WaveMetrics,  
356 Inc.) based program for selecting and implementing our correction algorithms can be found in the  
357 Supplemental Material.



358  
359 **Figure 1.** The flow-chart for the application of correction algorithms on PSAP, CLAP, and TAP.  
360 Similar logic is followed for the AETH.

### 361 3. Results and discussion

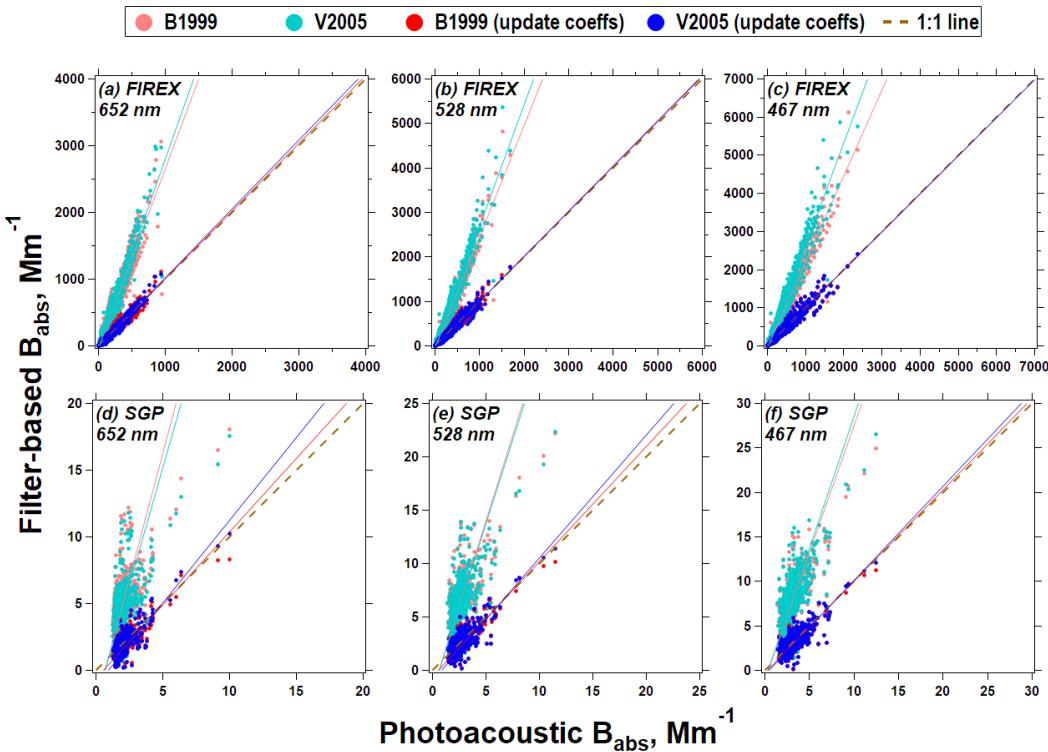
#### 362 3.1. Application of the previous algorithms on different aerosols

363 We first consider the application of the B1999 and V2005 corrections on different combinations  
364 of aerosol type and filter-based absorption photometer. Specifically, we apply the two corrections  
365 to the biomass burning data from the FIREX laboratory campaign (CLAP and TAP) as well as six  
366 months of ambient data from the SGP site (CLAP and PSAP). In doing so, we use the “default”  
367 coefficients recommended in B1999 and V2005 as well as “updated” coefficients that are  
368 estimated via regression techniques. We focus on the results of the CLAP in the main text, because  
369 a CLAP is the only instrument common to deployments for both FIREX and SGP. The results of  
370 the TAP from FIREX and the PSAP from the SGP site can be found in the Supplementary Material  
371 (Table S5 and Fig. S5).

372 Our inter-comparison between the corrected CLAP-derived  $B_{abs}$  and reference  $B_{abs}$  for the FIREX  
373 and SGP data is provided in Fig. 2 and Table 3. For the FIREX measurements, both analyses (using  
374 the “default” coefficients and updating the coefficients) suggest good correlation (coefficient of  
375 determination ( $R^2$ )  $> 0.9$ ) between the CLAP and the reference across all three wavelengths.  
376 Nevertheless, the corrections using the “default” coefficients result in over-prediction of  $B_{abs}$  by  
377 factors of  $\sim 2.5$ . If we update the coefficients in the corrections, there is an obvious improvement  
378 in the agreement (i.e., slope  $\approx 1$ ;  $R^2$  increases). The results are generally similar for SGP, although  
379 the  $R^2$  for ambient data is generally lower for ambient data ( $R^2 < 0.7$ ). Decreased  $R^2$  may be due  
380 to the lower aerosol concentrations measured in ambient air, which could lead to lower signal-to-



381 noise in the instruments. Moreover, it is worth mentioning that for both datasets (FIREX and SGP),  
 382 the corrected  $B_{abs}$  from different filter-based absorption photometers using the “default”  
 383 approaches does not agree with each other (slopes range from 0.69 to 1.40). However, after  
 384 updating the coefficients, the slopes approach unity (Table S6).



385  
 386 **Figure 2.** Inter-comparison between the CLAP-derived  $B_{abs}$  corrected by the B1999 and V2005  
 387 algorithms and the reference  $B_{abs}$  at 652, 528, and 467 nm for both FIREX and SGP data. The solid  
 388 lines represent linear regressions, while the dashed line is a 1:1 line.

389 **Table 3** Relationship between the CLAP-derived  $B_{abs}$  corrected by the B1999 and V2005  
 390 algorithms (including updated coefficients) and the reference  $B_{abs}$  at 652, 528, and 467 nm. The  
 391 relationship is achieved using major axis regression (Ayers, 2001). The value in parentheses  
 392 represents the coefficient of determination ( $R^2$ ) of the linear relationship.

		652 nm	528 nm	467 nm
FIREX	B1999	$y = -39 + 2.69x$ (0.94)	$y = -49 + 2.50x$ (0.96)	$y = -45 + 2.26x$ (0.97)
	V2005	$y = -46 + 2.83x$ (0.96)	$y = -57 + 2.75x$ (0.96)	$y = -56 + 2.68x$ (0.96)
	B1999 (update coeffs)	$y = -8.4 + 1.02x$ (0.96)	$y = -7.7 + 1.01x$ (0.97)	$y = -3.4 + 1.00x$ (0.96)
	V2005 (update coeffs)	$y = -9.4 + 1.03x$ (0.97)	$y = -7.3 + 1.01x$ (0.97)	$y = -3.0 + 1.00x$ (0.96)
SGP	B1999	$y = -2.60 + 3.77x$ (0.41)	$y = -1.90 + 3.20x$ (0.49)	$y = -0.98 + 2.85x$ (0.55)
	V2005	$y = -2.50 + 3.54x$ (0.41)	$y = -2.00 + 3.15x$ (0.48)	$y = -1.10 + 2.96x$ (0.55)
	B1999 (update coeffs)	$y = -0.29 + 1.10x$ (0.60)	$y = -0.29 + 1.08x$ (0.63)	$y = -0.17 + 1.03x$ (0.65)
	V2005 (update coeffs)	$y = -0.57 + 1.24x$ (0.65)	$y = -0.50 + 1.15x$ (0.67)	$y = -0.27 + 1.06x$ (0.67)

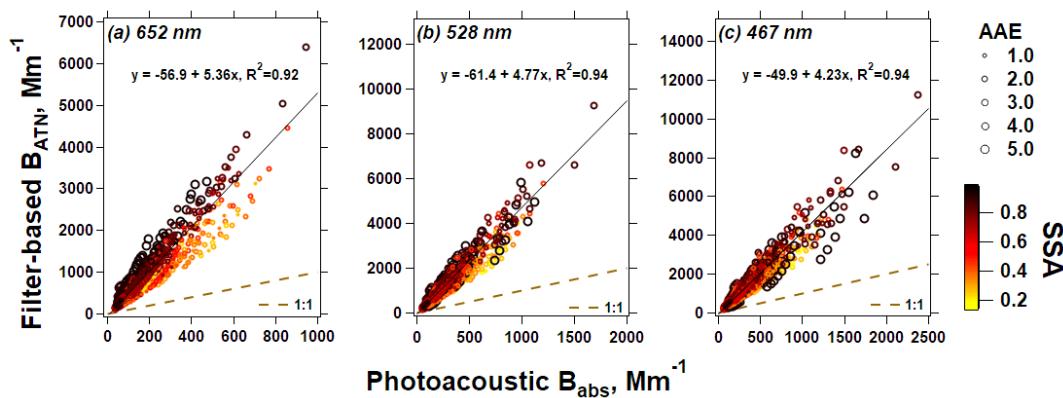


393 In the FIREX data, there is an apparent dependency of the updated coefficients on the wavelength  
 394 of light, but more importantly, on the aerosol optical properties, namely SSA and AAE (Tables  
 395 S7-S9). However, in the ambient data from SGP, the dependency on optical properties is less  
 396 obvious (Tables S10-S11). Nevertheless, all of these coefficients differ from those reported in  
 397 B1999 and V2005 (again, derived for the PSAP rather than the CLAP), which highlights the  
 398 potential need to use coefficient values that are appropriate for the instrument being used, its  
 399 wavelength(s) of light, and optical properties that are representative of the sampled aerosols when  
 400 applying correction factors to  $B_{ATN}$ .

401 3.2. Application of the new algorithms to the FIREX data

402 The co-location of the CLAP, TAP, AETH, and PAX during FIREX allows us to apply each  
 403 algorithm (A, B, C) to these data. Similar to Sect. 3.1, we focus our discussion on the CLAP with  
 404 details on the TAP and AETH presented in the Supplementary Material (Fig. S5-S6). However,  
 405 we provide the recommended initial guesses in the new algorithms and the comparison of  
 406 absorption (corrected filter-based  $B_{abs}$  versus reference  $B_{abs}$ ) for all filter-based absorption  
 407 photometers in Table 4 and Table 5 to help readers quickly retrieve key information of our  
 408 algorithms.

409 Figure 3 provides a comparison between the uncorrected  $B_{ATN}$  from the CLAP at all three  
 410 wavelengths, as well as photoacoustic  $B_{abs}$  interpolated to those wavelengths using AAE. For each  
 411 wavelength, the slopes are significantly greater than one. Moreover, there is an apparent  
 412 dependency on SSA and AAE in the agreement between the instruments. This is most obvious in  
 413 Fig. 3a (652 nm), where data with lower SSA and lower AAE (smaller markers, “brighter” colors)  
 414 fall below the best-fit line, while data with higher SSA and higher AAE (larger markers, “darker”  
 415 colors) fall above the best-fit line. This phenomenon is less clear in Fig. 3b-3c, but an apparent  
 416 dependency on SSA and AAE remains, which highlights the need to include both of these aerosol  
 417 optical properties (and appropriate interaction terms) when correcting  $B_{ATN}$  values.



418

419 **Figure 3.** Comparison of the uncorrected CLAP-derived  $B_{ATN}$  and the reference  $B_{abs}$  at 652, 528,  
 420 and 467 nm for the FIREX data. The data points are colored by the corresponding SSA. The size  
 421 of data points reflects their AAE quantified by the two PAX. The solid line represents the linear  
 422 regression, while the dashed line is a 1:1 line.

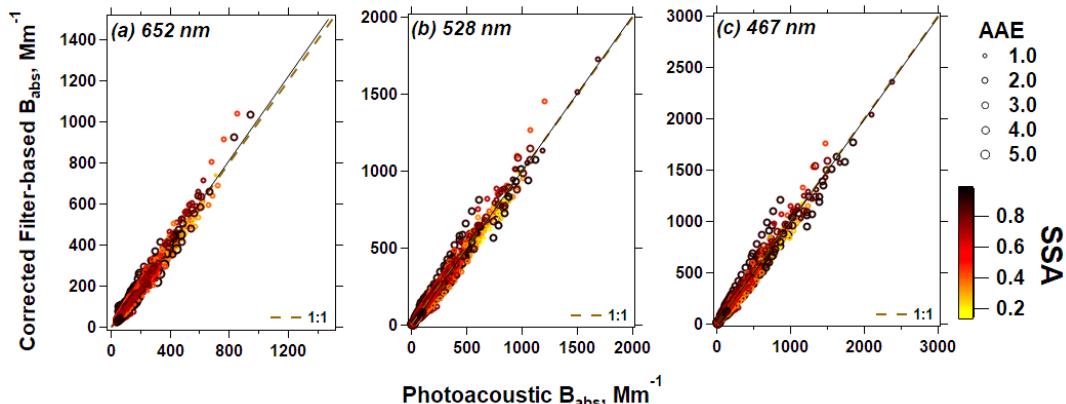


423 We first apply “Algorithm A” to the CLAP BATN data in Fig. 3. Using the reference  $B_{abs}$  values  
 424 from the PAX (in addition to  $B_{scat}$  values), we are able to derive a set of coefficients that enable  
 425 the correction of the data (Table 4). Corrected CLAP values are presented in Fig. 4 with the linear  
 426 relationships presented in Table 5. The slope for each wavelength is very close to the 1:1 line,  
 427 suggesting that our approach works well in correcting these data. Moreover, the heteroscedasticity  
 428 that exists in Fig. 3 has been minimized after correction, and there are no apparent trends in how  
 429 the data are organized in Fig. 4 due to the aerosol optical properties.

430 **Table 4** Coefficient values for Eq. (9) derived using “Algorithm A”. We recommend these as the  
 431 initial guesses when implementing “Algorithm B”.

		$G_0$	$G_1$	$G_2$	$G_3$	$G_4$	$G_5$	$G_6$	$G_7$
CLAP (FIREX)	652 nm	0.27	-0.16	-0.18	-0.05	0.18	0.08	-0.01	0.03
	528 nm	0.30	-0.28	-0.18	-0.07	0.25	0.10	0.13	-0.17
	467 nm	0.32	-0.38	-0.20	-0.08	0.33	0.12	0.24	-0.31
TAP (FIREX)	652 nm	0.45	-0.45	0.07	-0.19	0.94	0.10	0.26	-0.35
	528 nm	0.54	-0.51	0.02	-0.26	0.76	0.20	0.38	-0.44
	467 nm	0.62	-0.59	-0.07	-0.32	0.73	0.29	0.53	-0.60
CLAP (SGP)	652 nm	0.37	-0.18	-0.34	-0.11	0.30	0.18	-0.36	0.41
	528 nm	0.40	-0.15	-0.42	-0.14	0.10	0.24	-0.17	0.25
	467 nm	0.43	-0.16	-0.45	-0.16	0.07	0.27	-0.06	0.12
PSAP (SGP)	652 nm	0.24	0.35	-0.16	-0.04	-0.47	0.07	-0.57	0.73
	528 nm	0.30	0.48	-0.26	-0.10	-0.67	0.17	-0.63	0.77
	467 nm	0.35	0.49	-0.34	-0.15	-0.69	0.23	-0.55	0.79
AETH (FIREX)	950 nm	0.47	0.17	0.01	-0.27	-0.4	0.25	-0.12	0.27
	880 nm	0.34	0.13	0.13	-0.17	0	0.10	-0.13	0.12
	660 nm	0.28	0.09	0.11	-0.12	0.15	0.05	-0.12	0.03
	590 nm	0.16	-0.08	0.26	-0.03	0.59	-0.08	-0.02	-0.19
	520 nm	0.16	-0.05	0.14	-0.01	0.54	-0.07	-0.02	-0.21
	470 nm	0.14	-0.05	0.06	0	0.53	-0.05	-0.02	-0.17
	370 nm	0.13	-0.09	0.11	0	0.59	-0.06	-0.01	0.01

432



433

434 **Figure 4.** As in Fig. 3, but the CLAP-based  $B_{ATN}$  values have been corrected using our “Algorithm  
 435 A”.

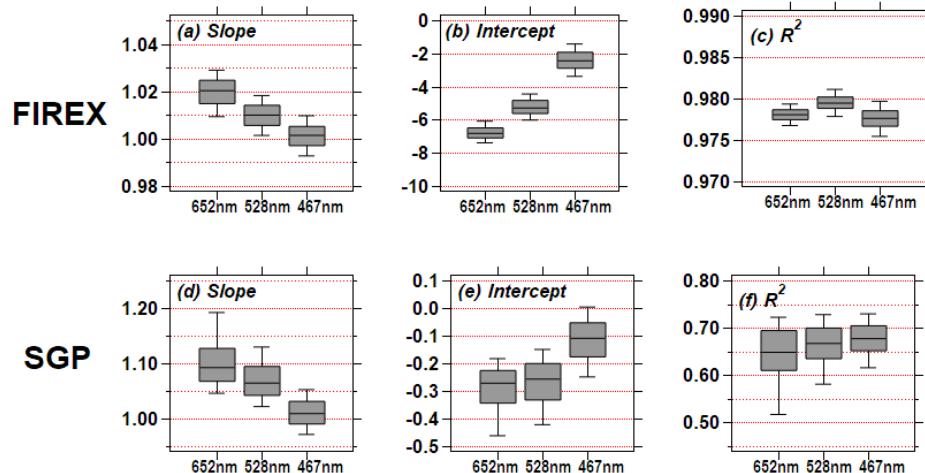


436 **Table 5** Relationship between the filter-based  $B_{abs}$  corrected by “Algorithm A” and the reference  
437  $B_{abs}$  at the operating wavelengths for the filter-based instrument. The relationship is achieved using  
438 major axis regression (Ayers, 2001). The value in the parentheses represents the coefficient of  
439 determination ( $R^2$ ) for the linear relationship.

		652 nm	528 nm	467 nm	
FIREX	CLAP	$y = -7.8 + 1.02x$ (0.98)	$y = -6.2 + 1.01x$ (0.98)	$y = -3.2 + 1.00x$ (0.98)	
	TAP	$y = -10 + 1.00x$ (0.87)	$y = -13 + 0.99x$ (0.87)	$y = -16 + 0.99x$ (0.88)	
SGP	CLAP	$y = -0.25 + 1.08x$ (0.68)	$y = -0.21 + 1.05x$ (0.67)	$y = -0.04 + 0.99x$ (0.68)	
	PSAP	$y = -0.28 + 1.10x$ (0.43)	$y = -0.24 + 1.06x$ (0.55)	$y = -0.07 + 1.00x$ (0.62)	
FIREX	AETH	950 nm	880 nm	660 nm	
		$y = -3.19 + 1.01x$ (0.82)	$y = -3.92 + 1.02x$ (0.85)	$y = -5.97 + 1.03x$ (0.88)	
		520 nm	470 nm	370 nm	
		$y = -2.36 + 0.99x$ (0.90)	$y = 2.93 + 0.95x$ (0.88)	$y = 18.38 + 0.89x$ (0.80)	
				590 nm	
				-	
				-	

440  
441 We next investigate the repeatability of the coefficient values presented in Table 4 by randomly  
442 selecting half of the measurements ( $N = 1338$ ) from the whole FIREX dataset. By implementing  
443 “Algorithm A” to the extracted observations, we obtain new coefficient values for  $G_0$  to  $G_7$ . This  
444 is repeated 1000 times to obtain a distribution of coefficient values (Fig. S7). The extraction  
445 approach mimics the process of obtaining new biomass burning datasets, so that we can estimate  
446 the variability of these derived coefficients. From Fig. S7 and Table S12, the derived coefficients  
447 are mostly insensitive to the different randomly-extracted datasets; most of the quartile deviation  
448 (defined as  $(Q3-Q1)/2$ , where  $Q1$  and  $Q3$  are the first and third quartile respectively) is within 0.05,  
449 except  $G_4$  which has a quartile deviation of  $\sim 0.08$ . Consequently, the coefficient values obtained  
450 in “Algorithm A” appear to be reasonable initial guesses to correct filter-based absorption  
451 measurements during biomass burning events when the reference  $B_{abs}$  is unavailable, such as in  
452 “Algorithm B” and “Algorithm C”.

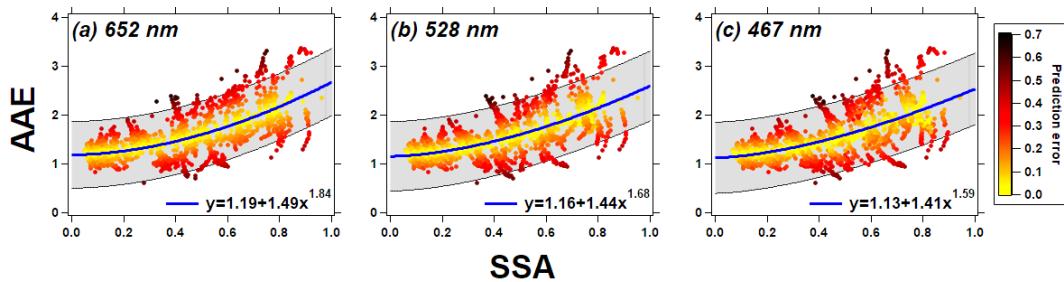
453 We next implement “Algorithm B” to the CLAP  $B_{ATN}$  data from Fig. 3 using the initial guesses of  
454 the coefficients derived from “Algorithm A” (Table 4) along with reference  $B_{scat}$  values. To get a  
455 sense of the variability in the results, we randomly select half of the data and applied the correction;  
456 this process is repeated 1000 times. For each iteration, we compare the corrected  $B_{abs}$  from the  
457 CLAP to the reference  $B_{abs}$  from the PAX; the resulting slope, intercept, and  $R^2$  values are  
458 summarized as box-and-whisker plots in Fig. 5. For all three wavelengths, the slopes are close to  
459 unity, and there is good correlation between the two absorption measurements ( $R^2 \approx 0.98$ ), which  
460 indicates that the good performance seen in Fig. 4 is independent of the reference  $B_{abs}$   
461 measurements and our algorithm is able to correct “new”  $B_{ATN}$ . Consequently, when scattering  
462 measurements are co-located with filter-based absorption measurements, our new correction  
463 algorithm performs well.



464

465 **Figure 5.** The box-and-whisker plots for the slope, intercept, and  $R^2$  of the relationship between  
 466 the CLAP-derived  $B_{abs}$  (corrected by “Algorithm B” in the present work) and PAX-derived  $B_{abs}$  for  
 467 all three wavelengths. For details on how these values were generated, please refer to the text.

468 Lastly, we apply “Algorithm C” to the data in Fig. 3. However, we first require a functional  
 469 relationship between AAE and SSA, because in this scenario, the CLAP  $B_{ATN}$  values are the only  
 470 data input to the algorithm (and therefore, SSA is unknown). Liu et al. (2014) proposed that a  
 471 power function can describe this relationship ( $AAE = a + b \times SSA^c$ ); we present these data from  
 472 FIREX along with power function fits (and associated prediction intervals) in Fig. 6. To define  
 473 AAE in this figure, we fit a power-law relationship to the three  $B_{ATN}$  values from the CLAP;  
 474 similarly, we define SSA using interpolated  $B_{scat}$  from the PAX and  $B_{ATN}$  from the CLAP (The  
 475 rationale for using  $B_{ATN}$  is that if “Algorithm C” were to be implemented in practice, only  $B_{ATN}$   
 476 would be available). In Fig. 6, the data points are colored by “prediction error”, effectively a metric  
 477 to quantify how well the power function reproduces the individual data points. Although there is  
 478 a fair amount of error in some of these points, we still obtain an SSA-AAE relationship required  
 479 to initialize “Algorithm C”.

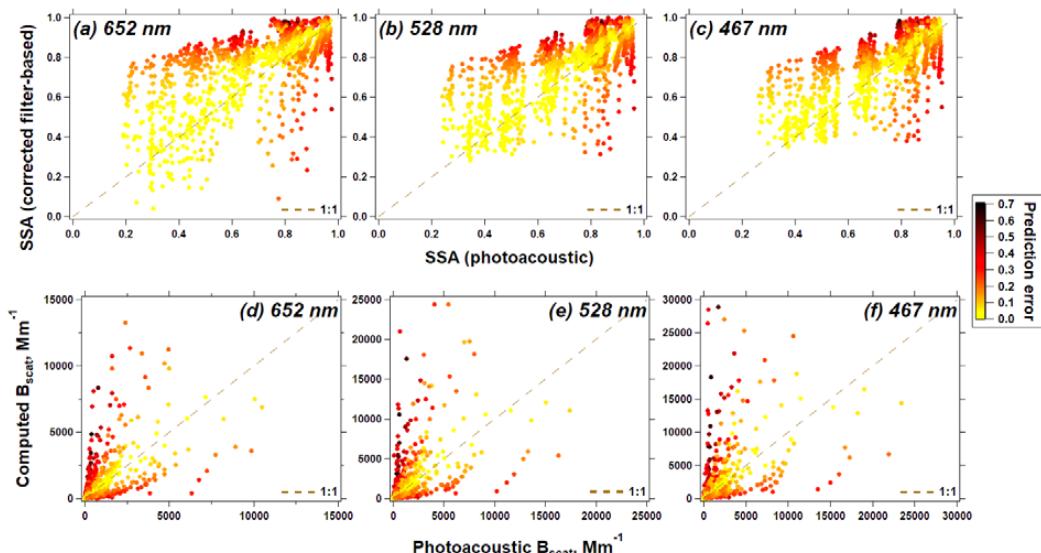


480

481 **Figure 6.** AAE plotted against SSA for the FIREX data. In the figures, AAE was computed using a  
 482 power-law fit across all three wavelengths, and SSA was computed using the interpolated  $B_{scat}$   
 483 from the two PAX and the reported  $B_{ATN}$  from the CLAP. The data points are colored by their  
 484 prediction error ((“true” AAE - “calculated” AAE) / “calculated” AAE).



485 Even though there is uncertainty in the SSA vs. AAE relationship used in “Algorithm C”, after  
486 corrections have been applied, the filter-based  $B_{abs}$  for the CLAP agrees well with the independent  
487 reference  $B_{abs}$ ; the slopes for all wavelengths are slightly greater than 1 (1.03-1.05) and the  $R^2$   
488 values are all high (0.97-0.98). However, even though the absorption measurements are corrected  
489 well, there still remains large uncertainties in values of inferred scattering. Examples of this are  
490 provided in Fig. 7, where we compare the SSA inferred from the PAX to the SSA inferred from  
491 “Algorithm C” as well as  $B_{scat}$  for each wavelength. Generally speaking, data that are better  
492 represented by the SSA vs. AAE relationship (i.e., smaller prediction error) results in better  
493 agreement with the reference for both SSA and  $B_{scat}$ , but there is also a clear divergence from the  
494 1:1 line in Fig. 7a-c as SSA decreases. Therefore, even though “Algorithm C” performs well at  
495 correcting filter-based  $B_{ATN}$  to agree with the reference  $B_{abs}$ , estimates of final SSA values should  
496 be considered to be uncertain.



497

498 **Figure 7.** Comparison of SSA (a-c) and  $B_{scat}$  (d-f) at the three wavelengths for the FIREX data.  
499 Vertical axis: values output from “Algorithm C”; horizontal axis: values calculated using the  
500 photoacoustic  $B_{abs}$  and  $B_{scat}$ .

501 In addition to the CLAP, we apply the new algorithms to the other filter-based absorption  
502 photometers operated during the FIREX study (TAP and AETH). Consistent with what we  
503 observed for the CLAP results, the corrected TAP- and AETH-derived  $B_{abs}$  is in good agreement  
504 with the photoacoustic  $B_{abs}$  (as demonstrated in Table 4 and Table 5, as well as Fig. S5-S6).  
505 Moreover, the corrected  $B_{abs}$  from the three filter-based instruments agrees with each other for all  
506 three wavelengths (Table 6), confirming the universal nature of our algorithm.

507



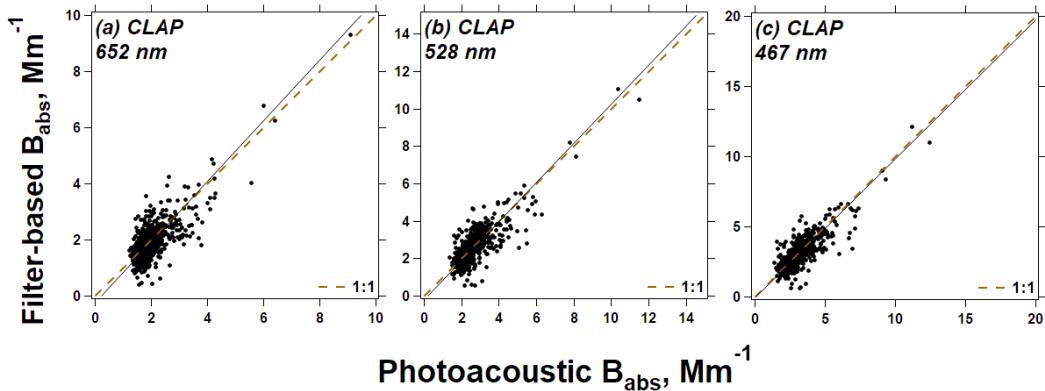
508  
 509 **Table 6** Inter-comparison between different filter-based  $B_{abs}$  corrected by “Algorithm A” in the  
 510 present work. The value in the parentheses represents the coefficient of determination ( $R^2$ ) of the  
 linear relationship.

	FIREX: CLAP vs. TAP	FIREX: CLAP vs. AETH	FIREX: TAP vs. AETH	SGP: CLAP vs. PSAP
652 nm	$y = 1.84 + 1.02x$ (0.89)	$y = 4.17 + 0.94x$ (0.87)	$y = -0.31 + 0.99x$ (0.82)	$y = -0.04 + 0.99x$ (0.70)
528 nm	$y = 5.75 + 1.02x$ (0.88)	$y = 3.70 + 0.91x$ (0.85)	$y = -6.38 + 0.98x$ (0.82)	$y = -0.11 + 1.02x$ (0.73)
467 nm	$y = 10.57 + 1.01x$ (0.88)	$y = 0.45 + 0.98x$ (0.83)	$y = -13.62 + 1.04x$ (0.79)	$y = -0.11 + 1.02x$ (0.76)

511

### 512 3.3. Application of the new algorithms to ambient data

513 To test our algorithms further, we extended our work to ambient data collected the DOE SGP site  
 514 during the time period which the PASS-3 was operational. From the SGP data, we derived a  
 515 different set of coefficients for ambient data using “Algorithm A”, which differ from those derived  
 516 for FIREX (Table 4). The results presented in Fig. 8 and Table 5 suggest that our new algorithm  
 517 works at least as well as B1999 and V2005 on this dataset (both with updated coefficients). The  
 518 repeatability of the coefficient values in “Algorithm A” is confirmed for the SGP measurements  
 519 using the same procedure as described in Sect. 3.2 (see results in Fig. S7 and Table S12).



520

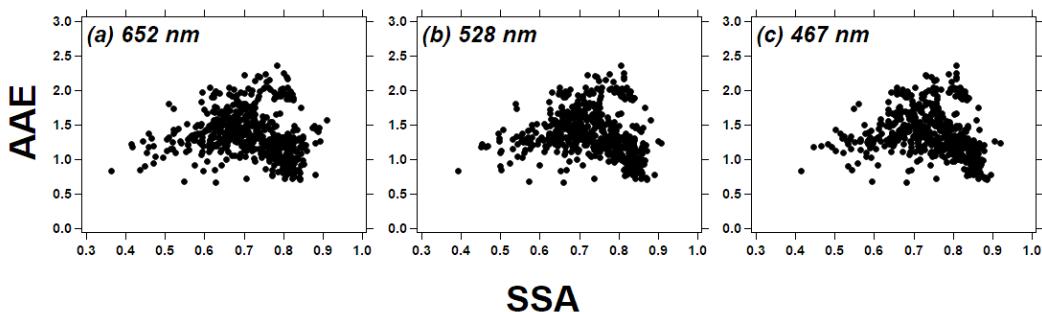
521 **Figure 8.** Inter-comparison between the CLAP-derived  $B_{abs}$  corrected by “Algorithm A” in the  
 522 present work and reference  $B_{abs}$  at 652, 528, and 467 nm for the ambient data at the SGP study  
 523 area. The solid line represents the linear regression, while the dashed line is a 1:1 line.

524 On the SGP data, we see similar performance to the FIREX data when we apply “Algorithm B”,  
 525 where we again sampled half of the CLAP data, used the initial guesses derived in “Algorithm A”,  
 526 and repeated this process 1000 times. Although the slopes tend to be larger than 1 (i.e., the  
 527 corrected CLAP  $B_{abs}$  remains high relative to the PASS  $B_{abs}$ ), the results still represent an  
 528 improvement over B1999 and V2005 using their recommended coefficients for their correction  
 529 equations.

530 Implementing “Algorithm C” is challenging for ambient data, because there is no distinct power  
 531 function relationship in AAE vs. SSA (Fig. 9); this is consistent with other field studies reporting  
 532 both SSA and AAE (e.g., Backman et al. (2014) and Lim et al. (2018)). Our approach described  
 533 here is only appropriate for ambient aerosols that follow a power function, such as sites impacted



534 by biomass burning. Nevertheless, we did apply this to a subset of the SGP data where the AAE-  
535 SSA prediction error is within 30% ( $N = 86$ ), and for this subset of data, “Algorithm C” works  
536 fairly well (slopes  $\approx 0.95$ ; see Fig. S8). Therefore, while “Algorithm C” may have utility for  
537 ambient data, we advise caution when using this algorithm since the aerosols influencing the site  
538 may not be represented by a clear AAE-SSA power function (e.g., when biomass burning and  
539 coarse aerosols are equally prevalent at a long-term monitoring site).



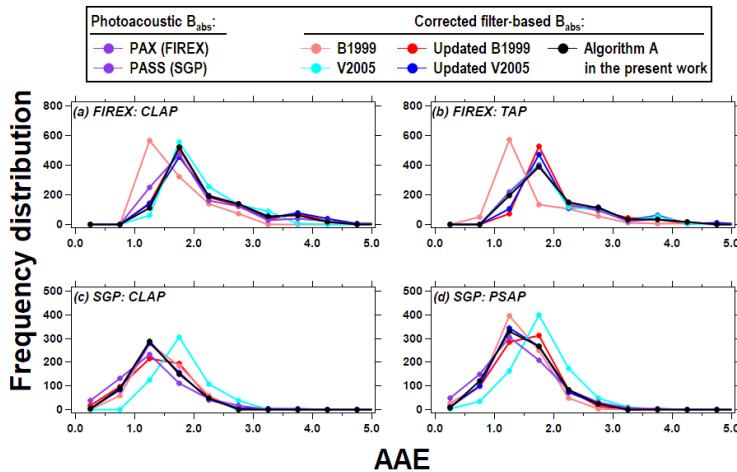
540  
541 **Figure 9.** AAE plotted against SSA for the SGP ambient data. The power law fit ( $AAE = a + b \times SSA^c$ )  
542 is performed on SSA ( $SSA = B_{scat}/(B_{scat} + B_{ATN})$ ) and AAE computed by three-wavelength  $B_{ATN}$ .

543 These new algorithms are also applicable to the PSAP deployed at the SGP site. The results of the  
544 correction for the PSAP are presented in Table 5 and Fig. S5, and the recommended initial guesses  
545 when implementing “Algorithm B” to PSAP- $B_{ATN}$  at ambient environments are given in Table 4.  
546 As expected, there is good agreement between corrected PSAP- and CLAP- $B_{abs}$  (Table 6).

#### 547 3.4. Impact of the implemented correction algorithm on aerosol optical properties

548 In addition to the direct comparisons of  $B_{abs}$  between the filter-based and photoacoustic  
549 measurements, we compare derived optical properties (AAE and SSA) from different instruments  
550 to assess the algorithms’ performance on derived aerosol optical properties. For example, we have  
551 discussed the discrepancy of SSA between the filter-based and photoacoustic measurements when  
552 implementing “Algorithm C” in Sect. 3.2. In this section we will more broadly discuss the impact  
553 of different correction algorithms on AAE and SSA.

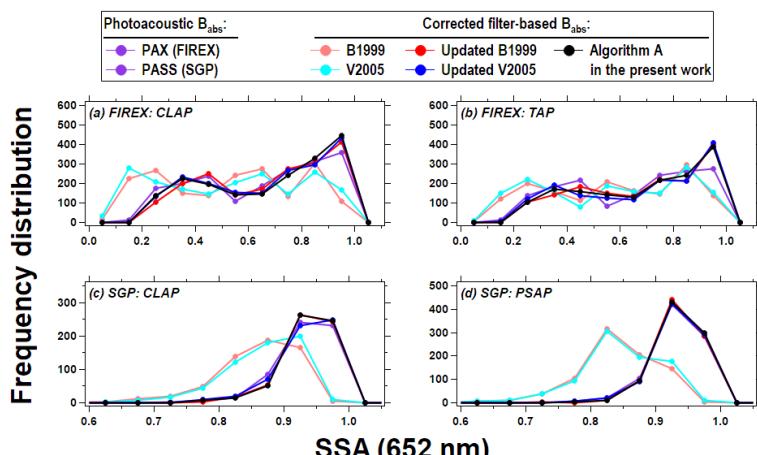
554 In Fig. 10, we present the frequency distribution of AAE for both FIREX and SGP data generated  
555 from different campaign/instrument pairs using different correction approaches. For the FIREX  
556 data (Fig. 10a-b), most corrections (with the exception of the “default” B1999) are consistent with  
557 the photoacoustic data, while for the SGP data (Fig. 10c-d), most corrections (with the exception  
558 of “default” V2005) are consistent with the photoacoustic data. However, updating the coefficients  
559 for B1999 and V2005 improves the agreement with the photoacoustic data. The 50% difference  
560 that exists between the B1999 and V2005 algorithms in all panels in Fig. 10 are consistent with  
561 previous studies. For example, both Backman et al. (2014) and Davies et al. (2019) found that the  
562 V2005-derived AAE is greater than B1999-derived AAE by 33% to 50% for ambient aerosols.  
563 Therefore, we highlight that the default coefficients in B1999 and V2005 may have some  
564 limitations when deriving AAE using the corrected  $B_{abs}$ ; instead, updating the coefficients or using  
565 the new algorithm proposed in this work may yield more robust AAE results.



566

567 **Figure 10.** The frequency distribution of AAE calculated for different instrument/correction  
 568 combinations of multi-wavelength  $B_{abs}$ .

569 Similar to Fig. 10, we also investigate the distribution of SSA computed by using corrected  $B_{abs}$   
 570 along with  $B_{scat}$ . We provide the results at 652 nm as an example in the main text (Fig. 11); figures  
 571 for 528 nm and 467 nm can be found in the Supplementary Material (Fig.S9 and S10). For both  
 572 FIREX and SGP data, the SSA obtained using the new algorithm agree very well with the B1999  
 573 and V2005 but only when their coefficients have been updated. Calculations of SSA using B1999  
 574 and V2005 with their recommended coefficients suggest that these values may be biased low,  
 575 which follows the over-estimation of corrected  $B_{abs}$  demonstrated in Fig. 2.



576

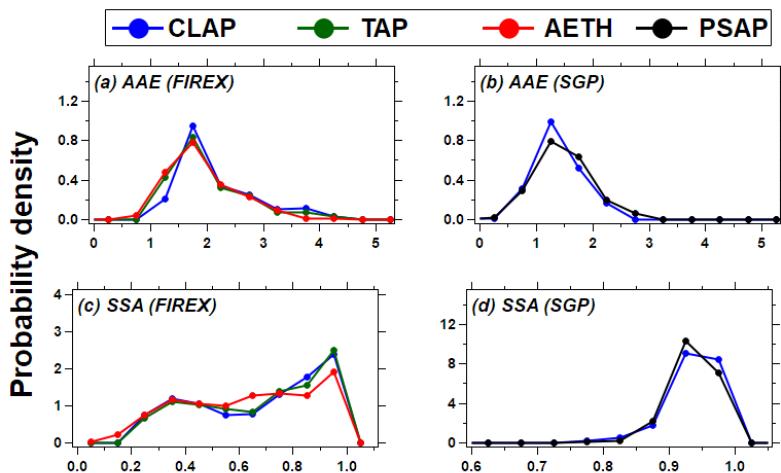
577 **Figure 11.** The frequency distribution of SSA (652 nm) calculated for different  
 578 instrument/correction combinations of  $B_{abs}$  and  $B_{scat}$ .

579 Moreover, we plot similar figures as Fig. 10-11 using all algorithms (A, B, and C). As shown in  
 580 Fig. S11, the results using “Algorithm B” agrees very well with those using “Algorithm A”, but



581 the use of “Algorithm C” results in some obvious discrepancies compared to the photoacoustic  
582 reference, again highlighting the potential for large uncertainty using this algorithm.

583 In Fig. 12, we directly compare the distributions of both AAE and SSA at 652 nm for all of the  
584 filter-based absorption photometers considered here, using our “Algorithm A” to correct the  $B_{ATN}$   
585 data. For both datasets, after the corrections have been applied, there are only marginal differences  
586 of the AAE (Fig. 10a and 10b) derived by different instruments. Similarly, there is good agreement  
587 among the SSA values when using corrected- $B_{abs}$  from different instruments (Fig. 10c and 10d).  
588 Overall, the derived properties using the new correction are consistent across all instruments,  
589 suggesting its universality.



590

591 **Figure 12** The probability density of AAE and SSA (652 nm) derived by different filter-based  
592 photometers  $B_{abs}$  (corrected by “Algorithm A” in the present work). Note that the number of total  
593 observations vary across instruments.

### 594 3.5. Uncertainty of the new algorithms

595 In this section, we estimate the uncertainty of the new algorithms due to both measurement  
596 uncertainties of the instruments and the uncertainties of parameter computation. We then simulate  
597 the propagated uncertainty in the corrected filter-based  $B_{abs}$  reported in this paper.

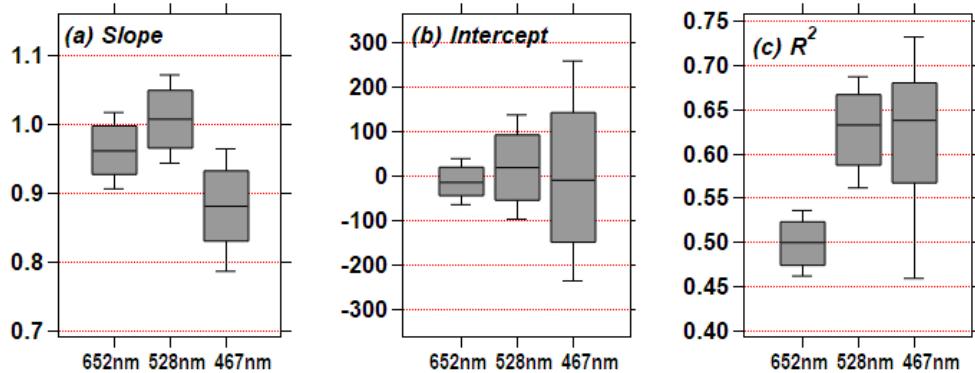
598 Measurement uncertainties of the instruments considered here have been reported in previous work  
599 (e.g., (Anderson et al., 1996; Nakayama et al., 2015; Ogren et al., 2017; Sherman et al., 2015)) and  
600 are summarized in Table 1. The typical sources of measurement uncertainty of the aerosol  
601 instruments include: 1) instrument noise (often associated with the averaging time); 2) calibration  
602 uncertainties (such as the accuracy of the operating wavelengths and the properties of the  
603 calibration materials); 3) standard temperature and pressure (STP) correction uncertainties  
604 (Sherman et al., 2015); and 4) flow rate uncertainties. Additional uncertainties that are specific to  
605 filter-based absorption photometers include spot size and filter medium corrections (Bond et al.,  
606 1999; Ogren et al., 2017). Regardless, these values all tend to be  $\leq 30\%$ , which is consistent with  
607 other commonly-used aerosol instrumentation.



608 Because correction algorithms for filter-based absorption instruments also require aerosol optical  
609 properties, the algorithms' performance will be affected by these values as well. For example,  
610 uncertainties in SSA are directly related to uncertainties associated with  $B_{abs}$  and  $B_{scat}$ , which are  
611 both included in our simulations. However, capturing uncertainties in AAE is more complex, as  
612 AAE can be computed by either "2 $\lambda$  fit" (a linear fit using  $B_{abs}$  at two wavelengths) or "3 $\lambda$  fit"  
613 (same as the power fit used in the present work). Davies et al. (2019) used the 3 $\lambda$  fit to calculate  
614 AAE and compared this to calculations using 662 nm and 785 nm (i.e., AAE<sub>662/785</sub>), finding that  
615 the 3 $\lambda$  results was about 50% greater. Moreover, similar differences (-35% to 85%) can exist  
616 comparing two different 2 $\lambda$  combinations (AAE<sub>440/870</sub> and AAE<sub>675/870</sub>), depending on the  
617 contribution of brown carbon to absorption at 440 nm (Wang et al., 2016). However, based on Fig.  
618 S12 and S13, we demonstrate small (< ~10%) differences in the calculated values of AAE using  
619 our Algorithm A using different 2 $\lambda$  combinations for linear fits and the 3 $\lambda$  power-law fit, when  
620 considering both FIREX and SGP data. Consequently, we do not include AAE calculation  
621 uncertainty in our simulation.

622 In our simulations, the propagated uncertainty of corrected  $B_{abs}$  is estimated by implementing the  
623 new algorithm to datasets in which filter-based  $B_{ATN}$ , reference  $B_{abs}$ , and  $B_{scat}$  are subject to  
624 measurement uncertainties. The full procedure is outlined in the Supplementary Material, but we  
625 provide a brief overview of our Monte Carlo approach here. First, we create a synthetic dataset ( $n$   
626 = 500 records) that defines  $B_{abs}$  at 652 nm and AAE that is intended to represent biomass burning.  
627 Values of  $B_{ATN}$  and SSA are then computed using the relationships presented in Fig. 3 and Fig. 6,  
628 respectively. Respective uncertainties associated with each of these values are applied following  
629 Table 1, assuming that these follow a normal distribution. We then applied "Algorithm B" to the  
630  $B_{ATN}$  dataset, repeated 1000 times, to quantify overall uncertainty associated with our correction  
631 algorithm.

632 Figure 13 provides a graphical summary of our uncertainty simulation results, which was derived  
633 by fitting linear equations to the "true"  $B_{abs}$  value (that we defined) and the "corrected"  $B_{abs}$  values  
634 (outputs of each iteration). Considering the slopes (Fig. 12a), our algorithm can generally  
635 reproduce the "true" value within 10% at 652 nm and 528 nm, but the performance is slightly  
636 degraded at 467 nm. The median intercept for our simulations is close to zero, but the interquartile  
637 range increases with decreasing wavelength (Fig. 12b), suggesting that the uncertainty may  
638 increase at shorter wavelengths. The coefficients of determination (Fig. 12c) range from 0.47 (652  
639 nm) to 0.68 (467 nm), showing that the algorithm may be less precise if large measurement  
640 uncertainties exist. Even though these sources of uncertainty exist when implementing our  
641 correction algorithms and propagate through to the corrected values, we argue that our new  
642 algorithm will "standardize" uncertainties across corrected  $B_{abs}$  values from filter-based absorption  
643 photometers. Moreover, the new algorithms perform, at least, better than the previous algorithms  
644 with "default" coefficients, or as well as the previous algorithms with updated coefficients.



645

646 **Figure 13.** The box-and-whisker plots (slope, intercept, and  $R^2$ ) for the Monte Carlo simulation of  
647 the relationship between the CLAP-derived  $B_{abs}$  (corrected by “Algorithm B” in the present work)  
648 and “true”  $B_{abs}$  for all three wavelengths.

649 4. Conclusions

650 Filter-based absorption instruments are widely used at global observational sites due to their  
651 relatively low cost, fast response, and easy operation. Despite the existence of different correction  
652 algorithms to correct the filter-based  $B_{abs}$  measurements, these are not “standardized” as  
653 differences in corrected  $B_{abs}$  values exist across different instrument/correction combinations, even  
654 when the instruments are co-located. This study provides a systematic evaluation of the previous  
655 correction algorithms (B1999 and V2005 corrections) on the CLAP and similar instruments (TAP  
656 and PSAP) using both laboratory-generated biomass burning emissions and ambient aerosols. We  
657 also developed “universal” correction algorithms that are applicable to any filter-based absorption  
658 photometer (e.g., PSAP, CLAP, TAP, AETH), which will have utility for any historic or future  
659 filter-based absorption measurements and which have the potential to standardize absorption  
660 coefficients across all filter-based instruments. This latter point is demonstrated in Table 6 and Fig.  
661 12 in that there is good agreement across all filter-based absorption photometers when applying  
662 our corrections to both biomass burning and ambient data. In practice, we anticipate that our  
663 Algorithm B will be most common, because at long-term monitoring sites, filter-based absorption  
664 photometers are typically co-located with a nephelometer.

665 Using the existing corrections on our CLAP measurements, we find that the corrected  $B_{abs}$   
666 overestimate photoacoustic  $B_{abs}$  by factors of ~2.6 (biomass burning aerosols) and ~3.2 (ambient  
667 aerosols). Similar overestimations of absorption by filter-based instruments are seen in the results  
668 of TAP from the FIREX study and PSAP deployed at the SGP. Comparing between B1999 and  
669 V2005,  $B_{abs}$  corrected by the two corrections differ by -6% to 18%. These discrepancies in our  
670 results are consistent with those reported for the inter-comparisons between filter-based and  
671 photoacoustic absorption instruments (e.g., (Arnott et al., 2003; Davies et al., 2019; Li et al., 2019;  
672 Müller et al., 2011a)).

673 Overall, our new developed algorithms (A, B, and C) perform well on correcting  $B_{abs}$  for different  
674 filter-based absorption photometers (CLAP, TAP, PSAP, and AETH) from both biomass burning



675 and ambient measurements. Our work suggests that if the filter-based instrument is co-operated  
676 with a reference absorption instrument and a NEPH at field for a period, researchers can compute  
677 site-specific initial guesses (same as “Algorithm A” in the present work). Otherwise, either  
678 “Algorithm B” or “Algorithm C” proposed in this paper can be used to correct the filter-based  
679 measurements. In “Algorithm B” when a filter-based absorption photometer is co-located with a  
680 NEPH but without a reference instrument, the set of coefficients yield in this work (Table 4) can  
681 be used as initial guesses to implement the algorithm. In “Algorithm C” when a filter-based  
682 absorption photometer is operated by itself, a “representative” relationship between AAE and SSA  
683 can be used to estimate SSA from AAE at each step in the iterative process, but we advise caution  
684 if this relationship is not monotonic (e.g., as in the ambient data from SGP and from Backman et  
685 al. (2014) and Lim et al. (2018)). The only scenario not included in the present work is that the  
686 filter-based absorption photometer is co-located with a reference absorption instrument, but no  
687 instrument for scattering. However, under this scenario, one could simply use the photoacoustic  
688  $B_{abs}$  data because no filter-induced biases exist for those instruments.

689 In terms of the aerosol optical properties (AAE and SSA) computed by different corrections, the  
690 new algorithm suggests no bias of AAE and SSA when compared to that derived by updated-  
691 B1999and updated-V2005 for both aerosol datasets.

692 However, the new algorithm is not without limitations. First, we used the photoacoustic  $B_{abs}$  as the  
693 reference to develop the algorithm and the initial guess of the coefficients; meanwhile, some  
694 studies argue that photoacoustic absorption is not a “ground truth” (e.g., (Lack et al., 2006; Lewis  
695 et al., 2008)). Thus, we simulate the propagated uncertainty of our algorithms considering the  
696 measurement uncertainties due to the photoacoustic  $B_{abs}$  (as well as  $B_{ATN}$  and  $B_{scat}$ ) and find that  
697 the corrected  $B_{abs}$  can be biased by -17% to 5%, depending on the operated wavelength. Although  
698 potential bias due to the precision of photoacoustic  $B_{abs}$  cannot be excluded, using the universal  
699 algorithm to correct the filter-based  $B_{abs}$  will at least eliminate correction-related biases among  
700 different filter-based instruments. Second, we only tested the algorithms with data from biomass  
701 burning and ambient measurements. It is unclear how the algorithms will work for other absorbing  
702 aerosols (e.g., dominated by fossil fuel emissions or mineral dust). Further evaluation of the  
703 performance of the new algorithm on other aerosol sources may help to address this issue.  
704 Regardless, we argue that our approach can standardize reported absorption coefficients at long-  
705 term monitoring sites, which has the potential to yield a better data set with which to evaluate  
706 chemistry-climate models.

707 **Code and data availability.** The code for the algorithm has been developed in Igor Pro  
708 (WaveMetrics Inc.). The package is available and fully described in the Supplementary Material.  
709 The FIREX aerosol products are available at <https://esrl.noaa.gov/csd/project/firex>. The SGP  
710 aerosol products are available at <https://www.archive.arm.gov/discovery/> (February 2013–July  
711 2013, 36° 36' 18.0" N, 97° 29' 6.0" W: Southern Great Plains Central Facility, data set accessed  
712 01/16/19).

713 **Author contribution.** AAM and GRM designed experiments, and HL and GRM conducted  
714 experiments. HL conducted data analysis and developed the correction algorithm presented herein.  
715 AAM and HL discussed the results. HL led the efforts to write the manuscript. All authors  
716 contributed to the final manuscript.

717 **Competing interests.** The authors declare that they have no conflict of interest.



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726 and Anthony Prenni from the National Park Service for lending us the AETH during the FIREX  
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729



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