

# 1 Cloud top pressure retrieval with DSCOVER-EPIC oxygen A and B bands observation

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## 12 13 **Abstract**

14 An analytic transfer inverse model for Earth Polychromatic Imaging Camera (EPIC)  
15 observation is proposed to retrieve the cloud top pressure (CTP) with considering in-cloud  
16 photon penetration. In this model, an analytic equation was developed to represent the reflection  
17 at top of atmosphere from above cloud, in-cloud, and below-cloud. The coefficients of this  
18 analytic equation can be derived from a series of EPIC simulations under different atmospheric  
19 conditions using a non-linear regression algorithm. With estimated cloud pressure thickness, the  
20 CTP can be retrieved from EPIC observation data by solving the analytic equation. To simulate  
21 the EPIC measurements, a program package using the double-*k* approach was developed.  
22 Compared to line-by-line calculation, this approach can calculate high-accuracy results with a  
23 one-hundred-fold computation time reduction. During the retrieval processes, two kinds of  
24 retrieval results, i.e., baseline CTP and retrieved CTP, are provided. The baseline CTP is derived  
25 without considering in-cloud photon penetration, and the retrieved CTP is derived by solving the  
26 analytic equation, taking into consideration the in-cloud and below-cloud interactions. The  
27 retrieved CTP for the oxygen A and B bands are smaller than their related baseline CTP. At the  
28 same time, both baseline CTP and retrieved CTP at the oxygen B-band are larger than those at  
29 the oxygen A-band. Compared to the difference of baseline CTP between the B-band and A-  
30 band, the difference of retrieved CTP between these two bands is generally reduced. Out of  
31 around 10000 cases, in retrieved CTP between A- and B-bands we found an average bias of 93  
32 mb with standard deviation of 81 mb. The cloud layer top pressure from Cloud-Aerosol Lidar  
33 and Infrared Pathfinder Satellite Observations (CALIPSO) measurements is used to do  
34 validation. Under single-layer cloud situations, the retrieved CTPs for the oxygen A-band agree  
35 well with the CTPs from CALIPSO, which mean difference is within 5 mb in the case study.  
36 Under multiple-layer cloud situations, the CTPs derived from EPIC measurements may be larger  
37 than the CTPs of high level thin-clouds due to the effect of photon penetration.

## 39 1. Introduction

40 The Deep-Space Climate Observatory (DSCOVR) satellite is an observation platform  
41 orbiting within the first Sun-Earth Lagrange point (L1), 1.5 million km from the Earth, carrying a  
42 suite of instruments oriented both Earthward and sunward. One of the Earthward instruments is  
43 the Earth Polychromatic Imaging Camera (EPIC) sensor, which can take images of the Earth  
44 with spatial resolution of 10 km at nadir. The EPIC continuously monitors the entire sunlit Earth  
45 for backscatter, with a nearly constant scattering angle between  $168.5^\circ$  and  $175.5^\circ$ , from sunrise  
46 to sunset with 10 narrowband filters: 317, 325, 340, 388, 443, 552, 680, 688, 764 and 779 nm  
47 (Marshak et al., 2018). Of the 10 narrow-band channels, there are two oxygen absorption and  
48 reference pairs, 764nm versus 779.5nm and 680nm versus 687.75nm, for oxygen A and B bands.  
49 The cloud top pressure (CTP) or cloud top height (CTH) is an important cloud property for  
50 climate and weather studies. Based on differential oxygen absorption, both EPIC oxygen A-band  
51 and B-band pairs can be used to retrieve CTP. It is worth noting that although CTP and CTH  
52 reference the same characteristic of clouds, the conversion between the two depends on the  
53 related atmospheric profile.

54 Although the theory of using oxygen absorption bands to retrieve CTP was proposed  
55 decades ago (Yamamoto and Wark, 1961), it is still very challenging to do the retrieval  
56 accurately due to the complicated in-cloud penetration effect (Yang et al., 2019, 2013; Davis et  
57 al., 2018a, 2018b; Richardson and Stephens, 2018; Loyola et al., 2018; Lelli et al., 2014, 2012;  
58 Schuessler et al., 2013; Rozanov and Kokhanovsky, 2004; Kokhanovsky and Rozanov, 2004;  
59 Kuze and Chance, 1994; O'brien and Mitchell, 1992; Fischer and Grassl, 1991; and etc.). To  
60 estimate the CTP from satellite measurements, many approaches have been designed to retrieve  
61 clouds' effective top pressures without considering in-cloud photon penetration. These  
62 approaches did not consider light penetrating cloud, therefore the derived CTH is lower than the  
63 cloud top, and the effective top pressures is higher than CTP. In the meantime, to improve the  
64 retrieval accuracy of CTP, various techniques have been applied to the retrieval methods with in-  
65 cloud photon penetration. For example, Kokhanovsky and Rozanov (2004) proposed a simple  
66 semi-analytical model for calculation of the top-of-atmosphere (TOA) reflectance of an  
67 underlying surface-atmosphere system, accounting both for aerosol and cloud scattering. Based  
68 on the work of Kokhanovsky and Rozanov (2004), Rozanov and Kokhanovsky (2004) developed  
69 an asymptotic algorithm for the CTH and the geometrical thickness determination using  
70 measurements of the cloud reflection function. This retrieval method was applied by Lelli et al.  
71 (2012, 2014) to derive CTH using measurements from GOME instrument on board the ESA  
72 ERS-2 space platform.

73 Currently, based on the measurements of DSCOVR EPIC sensor, the Atmospheric Science  
74 Data Center (ASDC) at National Aeronautics and Space Administration (NASA) Langley  
75 Research Center archives both calibrated EPIC reflectance ratio data and processed Level 2  
76 cloud retrieval products, including cloud cover, cloud optical depth (COD), cloud effective top  
77 pressure at oxygen A and B bands (Yang et al., 2019). By using EPIC reflectance ratio data at  
78 oxygen A-band and B-band absorption to reference channels, Yang et al (2013) developed a  
79 method to retrieve CTH and cloud geometrical thickness simultaneously for fully cloudy scene  
80 over ocean surface. First their method calculates cloud centroid heights for both A- and B-band  
81 channels using the ratios between the reflectance of the absorption and reference channels, then  
82 derives the CTH and the cloud geometrical thickness from the two dimensional look up tables  
83 that relate the sum and the difference between the retrieved centroid heights for A- and B-bands

84 to the CTH and the cloud geometrical thickness. The difference in the O<sub>2</sub> A- and B-band cloud  
85 centroid heights is resulted from the different penetration depths of the two bands. Compared to  
86 the cloud height variability, the penetration depth differences are much smaller and the retrieval  
87 accuracy from this method can be affected by the instrument noise (Davis et al. 2018a, b).

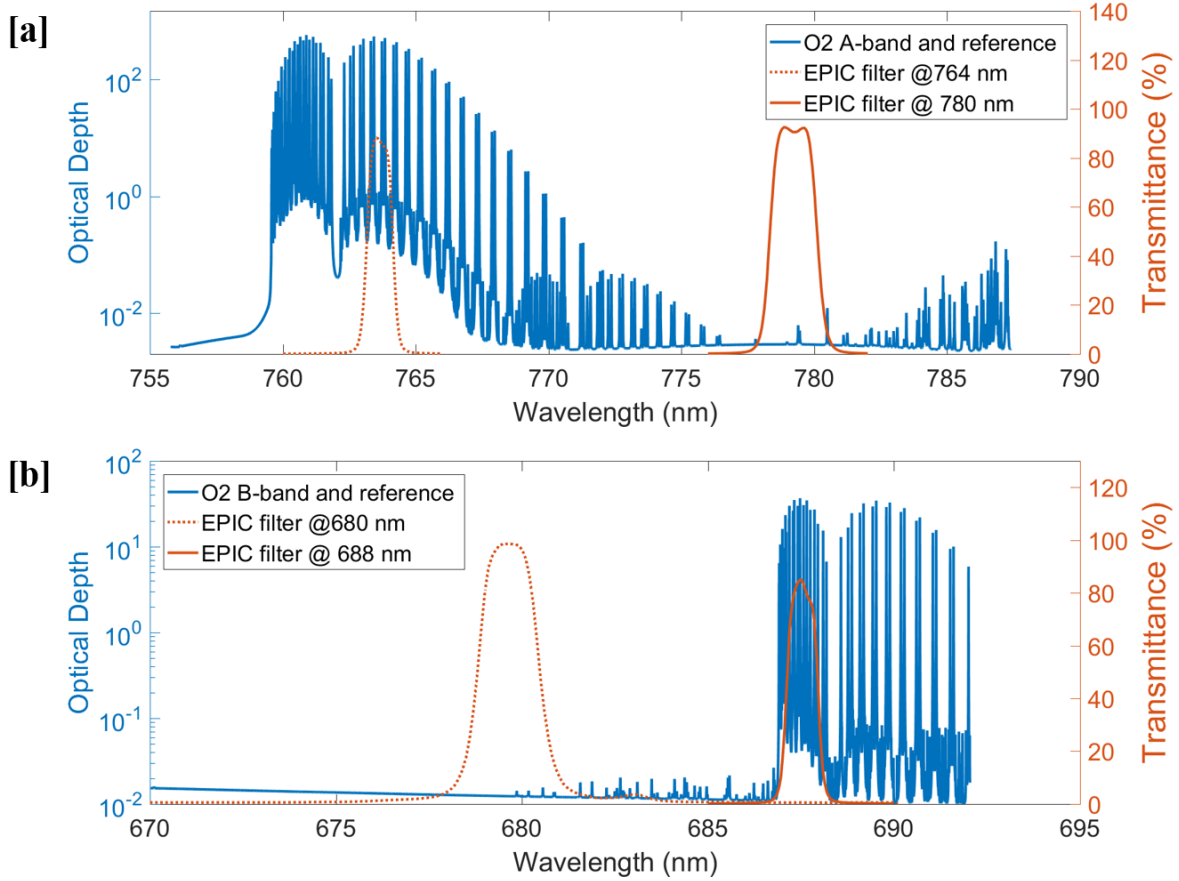
88 In this paper, to address the issue of in-cloud penetration, we proposed an analytic method  
89 to retrieve the CTP by using DSCOVER EPIC oxygen A- and B-band observation. This analytical  
90 method adopted ideas of the semi-analytical model (Kokhanovsky and Rozanov, 2004; Rozanov  
91 and Kokhanovsky, 2004), and developed a quadratic EPIC analytic radiative transfer equation to  
92 analyze the radiative transfer in oxygen A- and B-band channels. The structure of this paper is as  
93 follows: section 2 describes the theory and methods, which includes several subsections, i.e., the  
94 introduction of DSCOVER EPIC oxygen A and B bands filters, the theory of CTP retrieval based  
95 on EPIC oxygen A- and B- band observation, and the detailed retrieval algorithm; section 3  
96 describes the application and validation of the CTP retrieval method, which also includes several  
97 subsections, i.e., case studies of CTP retrieval, validation of the retrieval method, and retrieval of  
98 global observation; and section 4 states the conclusions of this study.

99

## 100 **2. Theory and methods**

### 101 **2.1 DSCOVER EPIC oxygen A and B bands filters**

102 EPIC filters at 764 nm and 779 nm cover the oxygen A-band absorption and reference  
103 bands, respectively (Fig. 1a). The high resolution absorption optical depth spectrum at oxygen A-  
104 band and B-band is calculated by Line-By-Line Radiative Transfer Model (LBLRTM, Clough et  
105 al., 2005) with HITRAN 2016 database (Gordon et al., 2017) for the U.S. standard atmosphere.  
106 In this wavelength range, the O<sub>3</sub> absorption is very weak (O<sub>3</sub> optical depth < 0.003) and there  
107 are no other gas absorptions. The background aerosol and Rayleigh scattering optical depth vary  
108 smoothly within the A-band range; the differences between in-band and reference band are  
109 negligible at nominal EPIC response functions. EPIC filters at 688 nm and 680 nm cover the  
110 oxygen B-band absorption and reference band, respectively (Fig. 1b). Compared to the oxygen  
111 A-band, O<sub>3</sub> absorption is slightly stronger in the oxygen B-band range, with an O<sub>3</sub> optical depth  
112 around 0.01. Any water vapor absorption in the B-band range is negligible. In the standard  
113 atmospheric model, from the oxygen B-band reference band to the absorption band, the O<sub>3</sub>  
114 absorption and Rayleigh scattering optical depth decreased by approximately 0.002 and 0.002,  
115 respectively. This may have some impacts on the CTP retrieval from the oxygen B-band (more  
116 discussion in the later sections). It is worth noting that for EPIC measurements at both oxygen A-  
117 and B-bands, the surface influence cannot be ignored. For examples, in the snow or ice covered  
118 area the surface albedo is high; in the plants covered area, the surface albedo changes  
119 substantially between oxygen A-band and B-band due to the impact of spectral red-edge (Seager  
120 et al., 2005).



121  
122

123 **Figure 1:** High resolution calculated absorption optical depth spectrum at oxygen A-band (a)  
124 and B-band (b) with DSCOVR EPIC oxygen A and B bands in-band and reference filters. Here  
125 the absorption optical depth spectrum is calculated by LBLRTM model with HITRAN 2016  
126 database for the U.S. standard atmosphere.

127 In general, if we use the pair of oxygen A and B absorption and reference bands together,  
128 the impact of other absorption lines, background Rayleigh scattering, and aerosol optical depth  
129 are very limited. At the same time, as a well-mixed major atmospheric component, the vertical  
130 distribution of oxygen in the atmosphere is very stable under varying atmospheric conditions.  
131 Thus, we can use the ratio of reflected radiance (or reflectance) at the TOA of oxygen absorption  
132 and reference bands (i.e.,  $R_{764}$  and  $R_{779}$ ,  $R_{688}$  and  $R_{680}$ ) to study the photon path length  
133 distribution and derive the cloud information. Also, compared to any specific EPIC oxygen  
134 absorption bands (i.e.,  $R_{764}$  and  $R_{688}$ ), the ratios of absorption to reference channels (i.e.,  
135  $R_{764}/R_{779}$  and  $R_{688}/R_{680}$ ) are less impacted by the instrument calibration and other  
136 measurement error. This can be explained by the following reasons: First, the EPIC  
137 measurements at oxygen A and B absorption and reference bands share same sensor and optical  
138 system, when calculating the ratios of them, some preprocessing calibration errors can be  
139 reduced. Second, to calculate  $R_{764}$  and  $R_{688}$ , the ratio of lunar reflectance at neighboring  
140 channels (i.e.,  $F(764,779)$  and  $F(688,680)$ ) and the calibration factors of oxygen A and B  
141 reference bands (i.e.,  $K_{779}$  and  $K_{680}$ ) are used (Geogdzhayev and Marshak, 2018; Marshak et al.,  
142 2018). Therefore, the accuracy of  $R_{764}$  and  $R_{688}$  is determined by the stability of  $F(764,779)$

143 and  $F(688,680)$  and the accuracy of  $K_{779}$  and  $K_{680}$  together. But the accuracy of absorption to  
144 reference ratios is only determined by the stability of  $F(764,779)$  and  $F(688,680)$ .

145

## 146 **2.2 Theory of CTP retrieval based on EPIC oxygen A- and B- band observation**

147 In our study, we tried two methods to retrieve the CTP based on EPIC oxygen A-band and  
148 B-band measurements: (1) Build a lookup table (LUT) for various atmospheric conditions and do  
149 the retrieval by searching the LUT; (2) Develop an analytic transfer inverse model for EPIC  
150 observations and calculate the related coefficients based on a series of simulated values, then use  
151 this analytic transfer inverse model to retrieve the CTP. In this paper, we mainly focus on the  
152 second method.

### 153 **2.2.1 Method 1: LUT based approach**

154 One commonly used method of retrieval for satellite observation is through the building  
155 and usage of LUTs (Loyola et al., 2018, Gastellu-Etchegorry and Esteve, 2003). LUT based  
156 approach can be fast because the most computationally expensive part of the inversion procedure  
157 is completed before the retrieval itself. For DSCOVER EPIC observations, we can build LUTs by  
158 simulating EPIC measurements under various atmospheric conditions, such as different surface  
159 albedo, solar zenith and viewing angles, COD, CTP, and cloud pressure thickness. Comparing  
160 the related simulated reflectance at the oxygen absorption and reference bands, we can obtain  
161 two LUTs for reflectance ratios of absorption/reference at EPIC oxygen A-band and B-band  
162 respectively, which can be used for the CTP retrieval. The detailed information of simulated  
163 reflectance ratio of absorption/reference is stated in Sect. 2.3.3.

164 During the retrieval process, the EPIC measurements (e.g., reflectance at oxygen A and B  
165 bands) with related solar zenith and viewing angles can be obtained from the EPIC level 1B data;  
166 COD information (retrieved from other EPIC channels) can be obtained from EPIC level 2 data.  
167 At the same time, we can get surface albedo from Global Ozone Monitoring Experiment 2  
168 (GOME-2) Surface Lambertian-equivalent reflectivity (LER) data (Tilstra et al., 2017). At this  
169 point the CTP and cloud pressure thickness are the only unknown variables. The cloud pressure  
170 thickness or the cloud vertical distribution has substantial impact on the accuracy of the CTP  
171 retrievals (Carbajal Henken et al., 2015; Fischer and Grassl, 1991; Rozanov and Kokhanovsky,  
172 2004; Preusker and Lindstrot, 2009). In this study, the cloud pressure thickness is used as an  
173 input parameter to retrieve the CTP. However, no related accurate cloud pressure thickness is  
174 provided by other satellite sensors now. To constrain the error from the estimation of cloud  
175 pressure thickness, we related it to the cloud optical thickness. It is reasonable because clouds  
176 with higher optical thickness normally have higher values of pressure thickness. To explore the  
177 correlation between cloud pressure thickness and cloud optical thickness, we use the related  
178 cloud data from Modern-Era Retrospective analysis for Research and Applications Version 2  
179 (MERRA-2, Gelaro et al., 2017), which is a NASA atmospheric reanalysis for the satellite era  
180 using the Goddard Earth Observing System Model Version 5 (GEOS-5) with Atmospheric Data  
181 Assimilation System (ADAS). Based on statistical analysis of one year's single-layer liquid  
182 water clouds over an oceanic region (23.20° S, 170.86° W, 2.11° S, 144.14° W) in 2017, we can  
183 get an equation for cloud pressure thickness approximation, i.e., cloud pressure thickness (mb) =

184 2.5\* COD + 23. The derived correlation coefficients are dependent on the case region and time  
 185 selections. Due to the complexity of cloud vertical distribution, whatever the accuracy of the  
 186 correlation coefficients is, the estimation will certainly bring in error.

187 With an estimated cloud pressure thickness, a multi-variable LUT searching method can  
 188 then be used to interpolate and obtain the CTP. It is worth noting that the reflectance ratio of  
 189 absorption/reference can be seen as a function of surface albedo, solar zenith and viewing angles,  
 190 COD, CTP, and cloud pressure thickness. Some atmospheric variables will have a non-linear  
 191 effect on the reflectance ratio. For example, the reflectance ratio is more sensitive to the variation  
 192 of COD when COD is small. Overall, the reflectance ratio varies monotonically and smoothly  
 193 with these variables (shown in Fig. 3). With a relatively high-resolution simulated table, we can  
 194 use a localized linear interpolation method to estimate the proper values. Multiple interpolations  
 195 are needed for this method to decrease the number of LUT dimensions, which will cost more  
 196 time than the analytic transfer inverse model method. The retrieval error of this method is  
 197 determined by the resolution of the LUT, i.e., the higher the resolution, the higher retrieval  
 198 accuracy. However, for multiple dimensional LUTs, the increase of resolution will increase the  
 199 table size exponentially, which will increase computational cost substantially for the table  
 200 building and inverse searching. Another possible method to increase the retrieval accuracy is  
 201 using different interpolation methods. For example, if the value of LUT varies non-linearly with  
 202 a variable, using high order interpolation method maybe better than using linear interpolation  
 203 method (Dannenber, 1998).

### 204 2.2.2 Method 2: Analytic transfer inverse model

205 For a long time, various efforts have been devoted to the study of radiative transfer in the  
 206 atmosphere, including scattering, absorption, emission, and etc. (Chandrasekhar, 1960; Irvine  
 207 1964; Ivanov and Gutshabash 1974; van de Hulst, 1980, 2012; Ishimaru, 1999; Thomas and  
 208 Stamnes, 2002; Davis and Marshak, 2002; Kokhanovsky et al., 2003; Marshak and Davis, 2005;  
 209 Pandey et al., 2012; and etc.). In this study, we develop an analytic radiative transfer equation to  
 210 analyze the radiative transfer at oxygen A and B bands. Through solving the analytic equation,  
 211 we can retrieve the CTP information directly. The theory of CTP retrieval is similar for EPIC  
 212 oxygen A-band and B-band observation. Here we use oxygen A-band as an example to study the  
 213 radiative transfer model. For oxygen A-band, photon path length distribution is capable of  
 214 describing vital information related to a variety of cloud and atmospheric characteristics.

$$215 I_v(\mu, \varphi; \mu_0, \varphi_0) = I_0(\mu, \varphi; \mu_0, \varphi_0) \int_0^\infty p(l, \mu, \varphi; \mu_0, \varphi_0) e^{-\kappa_v l} dl \quad (1)$$

216 Where,  $p(l)$  is photon path length distribution,  $\kappa_v$  is the gaseous absorption coefficient at wave  
 217 number  $v$ ,  $\mu = \cos(\theta)$ ,  $\mu_0 = \cos(\theta_0)$ ,  $(\theta, \varphi; \theta_0, \varphi_0)$  are zenith and azimuth angles for solar and  
 218 sensor view respectively,  $I_0$  and  $I_v$  are incident solar radiation and sensor measured solar radiation,  
 219 respectively.

220 When clouds exist, the incident solar radiation is reflected to TOA in three primary ways.  
 221 First, incident solar radiation is reflected by cloud top layer directly as a result of single  
 222 scattering. Second, the incident solar radiation will penetrate into the cloud and be reflected back  
 223 to TOA through cloud top via multiple scattering. Third, the incident solar radiation will pass  
 224 through the cloud and arrive at the surface, after that it is reflected back into the cloud and finally

225 scattered back to TOA through the cloud top. Due to the position of the EPIC instrument and the  
 226 long distance between EPIC and Earth, we can consider that solar zenith angle and sensor view  
 227 angle are nearly reverse. At oxygen A-band, the reflected solar radiation will be reduced due to  
 228 oxygen absorption depending on photon path length distributions. Absorption is negligible in  
 229 oxygen A-band's reference band. Oxygen A-band and its reference band are also attenuated by  
 230 air mass and aerosol through Rayleigh scattering and aerosol extinction. In the standard  
 231 atmospheric model, the optical depth of Rayleigh scattering ( $\tau_{Ray}$ ) at oxygen A-band (B-band)  
 232 and its reference band is 0.026 (0.040) and 0.024 (0.042), respectively (Bodhaine et al., 1999).  
 233 The absolute difference of Rayleigh scattering optical depth ( $\Delta\tau_{Ray} = \tau_{Ray}^{In-band} - \tau_{Ray}^{Ref}$ ) between  
 234 them is within 0.002. Compared to Rayleigh scattering, the difference of background aerosol  
 235 optical depth ( $\Delta\tau_{Aer}$ ) between absorbing and reference bands is smaller, within 0.0005.  
 236 Therefore, the attenuations from Rayleigh scattering and aerosol extinction at EPIC oxygen  
 237 absorption and its reference band are close to each other. Thus, when we use the ratio of EPIC  
 238 measured reflectance at oxygen A-band and its reference band to derive the photon path length  
 239 distribution and retrieve cloud information such as CTP, the impact of Rayleigh scattering and  
 240 aerosol extinction can be simplified in the analytic transfer inverse model.

241 To simplify the analytic transfer inverse model for EPIC observations, we made a series  
 242 of assumptions, e.g., isotropic component, a plane-parallel homogenous cloud assumption with  
 243 quasi-Lambertian reflecting surfaces. These assumptions have been widely used in radiative  
 244 transfer calculation for cloud studies. In this model,  $\mu$  and  $\mu_0$  are the same as in Eq. (1),  $\varphi$  is the  
 245 relative azimuth angle between Sun and satellite sensors;  $A_{surf}$  is the surface albedo;  $\tau_{O_2}^{Top}$ ,  $\tau_{O_2}^{Base}$ ,  
 246 and  $\tau_{O_2}^{Surface}$  are oxygen A-band absorption optical depth from TOA to cloud top layer, cloud  
 247 bottom layer, and surface, respectively;  $\Delta\tau_{O_2}^{Above-Cloud}$ ,  $\Delta\tau_{O_2}^{In-Cloud}$  and  $\Delta\tau_{O_2}^{Below-Cloud}$  are layered  
 248 oxygen A-band absorption optical depth above cloud, in cloud, and below-cloud, respectively;  
 249 functions  $f$  mean their contribution to the ratio of measured reflectance at oxygen A-band ( $R_A$ )  
 250 and reference band ( $R_f$ ). The detailed analysis of EPIC analytic transfer inverse model is shown  
 251 as follows:

252 (1) **Above Cloud:** the reflected solar radiation is determined by the oxygen absorption optical  
 253 depth above the cloud and air mass directly.

$$\begin{aligned}
 254 \quad f(\Delta\tau_{O_2}^{Above-Cloud}, \mu_0, \mu, \varphi) &= f(\Delta\tau_{O_2}^{Above-Cloud})f(\mu_0, \mu, \varphi) \\
 255 \quad &= a_0 \tau_{O_2}^{Top} \left( \frac{1}{\mu} + \frac{1}{\mu_0} \right) \quad (2)
 \end{aligned}$$

256 Here,  $a_0$  is a weight coefficient.

257 (2) **Within Cloud:** the reflected solar radiation is not only determined by oxygen absorption  
 258 optical depth above cloud and in-cloud, but also by penetration related factors, e.g., COD. Due to  
 259 photon penetration, oxygen parameter  $\tau_{O_2}^{Top}$  influences the enhanced path length absorption:

$$260 \quad \Delta\tau_{O_2}^{In-Cloud} = \tau_{O_2}^{Base} - \tau_{O_2}^{Top} \quad (3)$$

261 Equivalence theorem (Irvine, 1964; Ivanov and Gutshabash, 1974; van de Hulst 1980) is used to  
 262 separate absorption from scattering:



263  $f(\tau_{O_2}^{Top}, \Delta\tau_{O_2}^{In-cld}, \mu_0, \mu, \varphi) = f(\tau_{O_2}^{Top}, \Delta\tau_{O_2}^{In-cloud})f(\mu_0, \mu, \varphi)$   
 264  $= f(\tau_{O_2}^{Top})f_1(\mu_0, \mu, \varphi) + f(\Delta\tau_{O_2}^{In-cloud})f_2(\mu_0, \mu, \varphi)$  (4)

265  $f(\tau_{O_2}^{Top})$  is determined by two absorption dependences: strong ( $\sim \sqrt{\tau_{O_2}^{Top}}$ ) and weak ( $\sim \tau_{O_2}^{Top}$ ).

266  $f(\tau_{O_2}^{Top}) = a_1 \sqrt{\tau_{O_2}^{Top}} + b_1(\tau_{O_2}^{Top})$  (5)

267 Based on asymptotic approximation (Kokhanovsky et al., 2003; Pandey et al., 2012), the  
 268 reflection of a cloud without considering below cloud interaction is given by Eq. (6):

269  $R(\tau, \mu, \mu_0, T) = R_0^\infty(\tau, \mu, \mu_0) - TK(\mu)K(\mu_0)$   
 270  $= R_0^\infty(\tau, f_1(\mu, \mu_0)) - Tf_2(\mu, \mu_0)$  (6)

271 Here,  $R_0^\infty$  is the reflectance of a semi-infinite cloud,  $K(\mu)$  is the escape function of  $\mu$ ,  $T$  is global  
 272 transmittance of a cloud.  $T$  can be estimated by Eq. (7), with the cloud optical thickness  $\tau_{cld}$ , the  
 273 asymmetry parameter  $g$ , and a numerical constant  $\alpha = 1.07$ .

274  $T = \frac{1}{0.75\tau_{cld}(1-g)+\alpha}$  (7)

275  $f_1$  and  $f_2$  functions have a quadratic form as follows:

276  $f_{i-1} = a_i T + b_i(\mu + \mu_0) + c_i T(\mu + \mu_0) + d_i \mu \mu_0, i = 2, 3$  (8)

277 Combining Eqs. (4), (5) and (8), we can get the Eq. (9):

278  $f(\tau_{O_2}^{Top}, \Delta\tau_{O_2}^{Cld}, \mu_0, \mu, \varphi) = \left( a_1 \sqrt{\tau_{O_2}^{Top}} + b_1(\tau_{O_2}^{Top}) \right) (a_2 T + b_2(\mu + \mu_0) + c_2 T(\mu + \mu_0) + d_2 \mu \mu_0)$   
 279  $+ \Delta\tau_{O_2}^{In-cloud} (a_3 T + b_3(\mu + \mu_0) + c_3 T(\mu + \mu_0) + d_3 \mu \mu_0)$  (9)

280

281 (3) **Below Cloud:** The equivalence theorem used for below cloud is similar to within cloud  
 282 (Kokhanovsky et al., 2003; Pandey et al., 2012).

283  $f(\Delta\tau_{O_2}^{Below-cld}, \mu_0, \mu, \varphi) = T \tau_{O_2}^{Surface} \frac{A_{Surf}}{1+(e_4 * T + f_4) * A_{Surf}}$   
 284  $* (a_4 T + b_4(\mu + \mu_0) + c_4 T(\mu + \mu_0) + d_4 \mu \mu_0)$  (10)

285

286 Combining Eqs. (2), (9) and (10), we can get the total EPIC analytic transfer equation as  
 287 follows

288  $-\log\left(\frac{R_A}{R_f}\right) = f(\Delta\tau_{O_2}^{Above-cld}, \mu_0, \mu, \varphi) + f(\tau_{O_2}^{Top}, \Delta\tau_{O_2}^{Cld}, \mu_0, \mu, \varphi) +$   
 289  $f(\Delta\tau_{O_2}^{Below-cld}, \mu_0, \mu, \varphi) + \Delta\tau_{BG} \left( \frac{1}{\mu} + \frac{1}{\mu_0} \right)$  (11)

290 In Eq. (11),  $\Delta\tau_{BG}$  represents the sum of optical depth difference of background extinction (i.e.,  
 291 Rayleigh scattering  $\Delta\tau_{Ray}$ , aerosol extinction  $\Delta\tau_{Aer}$ , and O3  $\Delta\tau_{O3}$ ) between oxygen in-band and  
 292 reference band, as shown in Eq. (12).



293 
$$\Delta\tau_{BG} = \Delta\tau_{Ray} + \Delta\tau_{Aer} + \Delta\tau_{O3} \quad (12)$$

294 As stated in the previous subsection, in the standard atmospheric model with background aerosol  
 295 loading,  $(\Delta\tau_{Ray}, \Delta\tau_{Aer}, \Delta\tau_{O3})$  is approximately (0.002, 0.0005, -0.0005) and (-0.002, -0.0005, -  
 296 0.002) respectively at oxygen A and B bands, thus  $\Delta\tau_{BG}$  is approximately 0.002 and -0.0045  
 297 respectively at these two bands.

298 In this total analytic equation, there are 17 coefficients  $(a_0, a_1, b_1, a_2, \dots, d_4, e_4, f_4)$ , which  
 299 can be calculated through nonlinear regression algorithm according to a series of simulated  
 300 values for different atmospheric conditions. Based on Eq. (11), we can finally obtain a quadratic

301 equation,  $\mathbf{A}\sqrt{\tau_{O2}^{Top}} + \mathbf{B}\sqrt{\tau_{O2}^{Top}} + \mathbf{C} = \mathbf{0}$ , where the parameters A, B and C can be derived from  
 302 Eq. (11) directly, as shown in Eq. (13).

303 
$$A = a_0 \left( \frac{1}{\mu} + \frac{1}{\mu_0} \right) + b_1(a_2 T + b_2(\mu + \mu_0) + c_2 T(\mu + \mu_0) + d_2 \mu \mu_0) \quad (13.1)$$

304 
$$B = a_1(a_2 T + b_2(\mu + \mu_0) + c_2 T(\mu + \mu_0) + d_2 \mu \mu_0) \quad (13.2)$$

305 
$$C = -\log\left(\frac{R_A}{R_f}\right) - \Delta\tau_{BG} \left( \frac{1}{\mu} + \frac{1}{\mu_0} \right) - \Delta\tau_{O2}^{In-Cloud}(a_3 T + b_3(\mu + \mu_0) + c_3 T(\mu + \mu_0) + d_3 \mu \mu_0)$$
  
 306 
$$-T \tau_{O2}^{Surface} \frac{A_{surf}}{1+(e_4 * T + f_4) * A_{surf}} (a_4 T + b_4(\mu + \mu_0) + c_4 T(\mu + \mu_0) + d_4 \mu \mu_0) \quad (13.3)$$

307 When these parameters (i.e., A, B and C) are obtained from EPIC observation data and  
 308 other data source, we can easily solve the quadratic equation to retrieve cloud top O2 absorption  
 309 depth, and then CTP.

### 310 **2.3 Detailed retrieval algorithm**

311 As previously stated, in method 2, the analytic EPIC equation (i.e., Eq. (11)) is key for the  
 312 CTP retrieval. To derive the coefficients of Eq. (11), a series of model simulations for various  
 313 atmospheric conditions are needed. Thus, developing a radiative transfer model to simulate the  
 314 EPIC measurements at A- and B-bands and their reference bands is the first thing we need to  
 315 complete.

#### 316 **2.3.1 Oxygen A- and B-band absorption coefficients calculation**

317 To simulate the EPIC measurements, one of the most important steps is calculating  
 318 oxygen absorption coefficients at oxygen A-band and B-band. In this step, the HITRAN 2016  
 319 database is used to provide the absorption parameters, and the LBLRTM package is used to  
 320 calculate oxygen absorption coefficients layer by layer. In our algorithm, the whole Earth  
 321 atmosphere is divided by 63 layers.

322 Since oxygen absorption coefficients are pressure (or pressure-squared) and temperature  
 323 dependent, and the line shapes  $(k_i)$  of oxygen A- and B-bands are well fitted as Lorentzian in the  
 324 lower atmosphere, the relationship can be written as follows:

325 
$$k_i = \frac{S_i}{\pi} \frac{\alpha_i}{(v-v_i)^2 + \alpha_i^2} \quad (14)$$

326 
$$\alpha_i = \alpha_i^0 \frac{P}{P_0} \left(\frac{T_0}{T}\right)^{\frac{1}{2}}, S_i = S(T_0) \frac{T_0}{T} \exp \left[1.439E \left(\frac{1}{T_0} - \frac{1}{T}\right)\right] \quad (15)$$

327 Where  $S_i$  is the line intensity,  $\nu_i$  and  $\alpha_i$  are the line center wave number and half width,  
 328 respectively;  $P_0$  and  $T_0$  are standard atmospheric pressure and temperature, respectively.

329 In the simulation of EPIC measurements, the atmospheric layer at a given layer-average  
 330 pressure can have drastically different temperature depending on the atmospheric profile in use.  
 331 To ensure the accuracy of simulation, we need to use the LBLRTM package to calculate oxygen  
 332 absorption coefficients for each pressure/temperature profile, which is a time-consuming process.  
 333 Our goal has been to find a simple and fast method to calculate oxygen absorption coefficients  
 334 for different atmospheric profiles. Based on the study of Chou and Kouvaris (1986), Min et al.  
 335 (2014) proposed a fast method to calculate oxygen absorption optical depth for any given  
 336 atmosphere by using a polynomial fitting function, as shown in Eq. (16).

337 
$$\ln(A_{vLM}) = [a_0(\nu, P) + a_1(\nu, P) \times (T_{LM} - T_{mL}) + a_2(\nu, P) \times (T_{LM} - T_{mL})^2] \times \rho_{O_2} \quad (16)$$

338 Where  $A_{vLM}$  is optical depths for layer L, spectral point  $\nu$ , and atmosphere model M;  $\rho_{O_2}$  is  
 339 molecular column density ( $\frac{\text{molecules}}{\text{cm}^2} \times 10^{-23}$ );  $T_{LM}$  is the average temperature for layer L for a  
 340 given atmosphere; and  $T_{mL}$  is average temperature over all six typical geographic-seasonal model  
 341 atmospheres (M1 to M6, i.e., tropical model, mid-latitude summer model, mid-latitude winter  
 342 model, subarctic summer model, subarctic winter model, and the U.S. Standard (1976) model)  
 343 for layer L. To derive the coefficients  $a_0$ ,  $a_1$ , and  $a_2$ , we first calculated oxygen optical depth  
 344 coefficients for all typical atmospheres (M1 to M6) by using LBLRTM package, and then  
 345 selected three of them (e.g., M1, M5, and M6) to calculate the polynomial fitting coefficients.  
 346 This method has been successfully used by Min et al. (2014) to simulate the high resolution  
 347 oxygen A-band measurements.

### 348 **2.3.2 Fast radiative transfer model for simulating high-resolution oxygen A- and B-bands**

349 At oxygen A and B absorption bands, there are lots of absorption lines, therefore we cannot  
 350 simply calculate narrowband mean optical depth and then calculate the radiation for various  
 351 atmospheric conditions when simulating EPIC narrowband measurements. The correct way is  
 352 described as follows: firstly, simulate the solar radiation spectrum  $S(k(\lambda))$  under specific  
 353 atmospheric conditions, then integrate the spectrum with EPIC narrowband filter  $R(k(\lambda))$  to  
 354 obtain simulated narrowband measurements (Eq. (17)).

355 
$$R(\lambda) = \int S(k(\lambda))R(k(\lambda))d\lambda \neq R(\overline{k(\lambda)}) \quad (17)$$

356 With the high spectrum resolution oxygen absorption coefficient data, we can simulate the  
 357 high resolution upward diffuse oxygen A-band or B-band spectrum through DISORT code  
 358 (Stamnes et al., 1988) for any given atmospheric condition, which has various surface albedo,  
 359 SZA, COD, CTH (CTP), and cloud geometric (pressure) thickness. However, due to the high  
 360 spectrum resolution, it is very time-consuming when performing line by line (LBL) calculations.  
 361 Thus, developing a fast radiative transfer model for simulating high resolution oxygen A-band  
 362 and B-band spectrum is necessary.

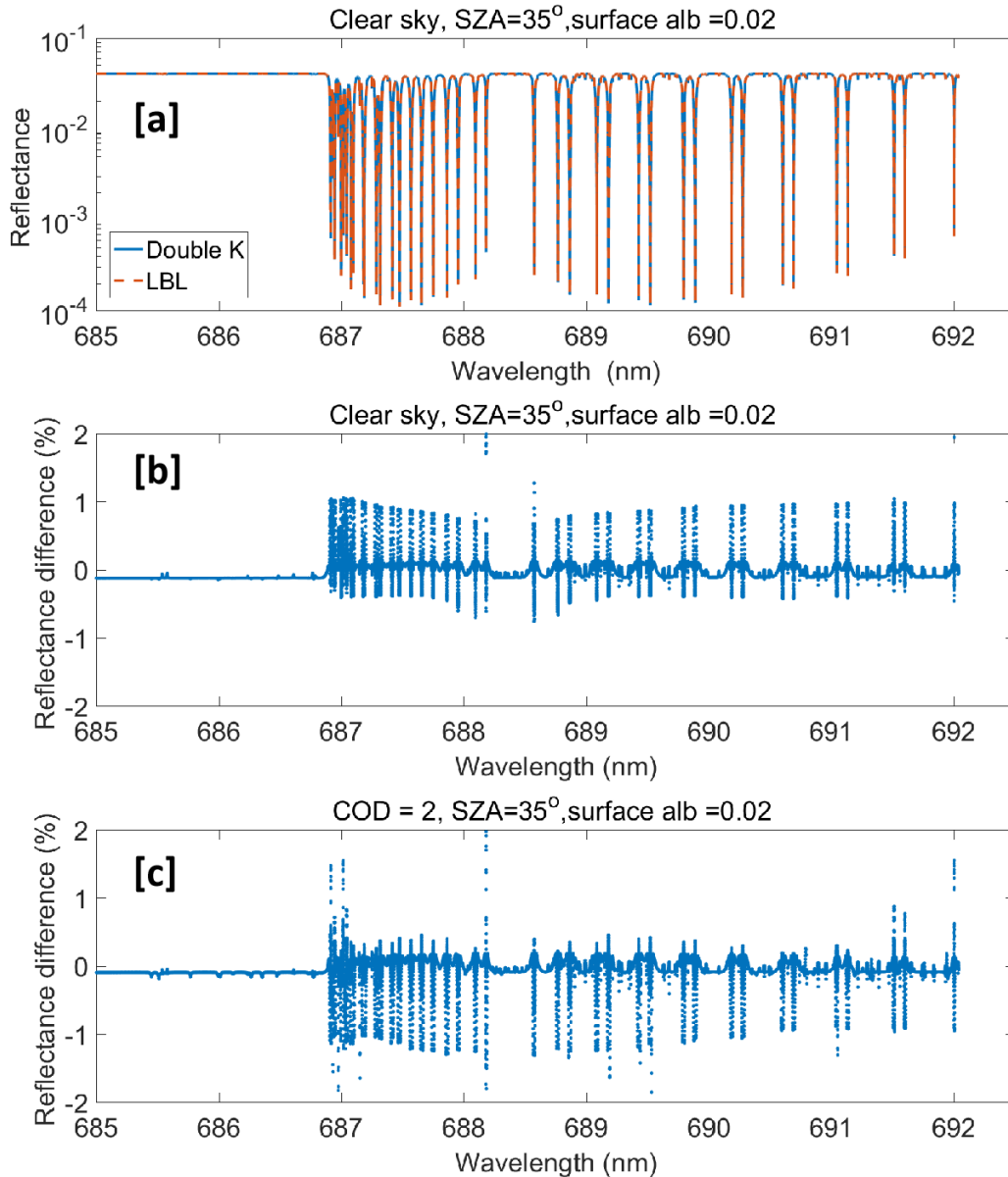
363 In this project, the double- $k$  approach is used to develop a fast radiative transfer model for  
 364 oxygen A-band and B-band respectively. [Min and Harrison 2004; Duan et al, 2005] proposed a

365 fast radiative transfer model. In their approach, the radiation from absorption and scattering  
 366 processes of cloud and aerosol are split into the single- and multiple-scattering components: The  
 367 single scattering component is computed line-by-line (LBL), while multiple scattering (second  
 368 order and higher) radiance is approximated.

$$\begin{aligned}
 369 \quad I &= I^{ss}(\lambda) + I^{ms}(\lambda) \\
 370 \quad &\approx I^{ss}[Z^h(p, T), P^h, \lambda] + I^{ms}[Z^h(p, T), P^h, \lambda] \\
 371 \quad &\approx I^{ss}[Z^h(p, T), P^h, \lambda] + I^{ms}[Z^l(p, T), P^l, \lambda] \\
 372 \quad &\approx I^{ss}[Z^h(p, T), P^h, \lambda] + I^{ms}\{F[Z^l(p, T), P^l, k(\lambda_i)]\} \quad (18)
 \end{aligned}$$

373 Eq. (18) is from Eq. (1) in Duan et al. (2005): *ss* and *ms* mean single and multiple scattering,  
 374 respectively. *Z* is the optical properties of the atmosphere as a function of pressure *p* and  
 375 temperature *T*, with *P* being the phase function of that layer. *h* and *l* represent higher and lower  
 376 number of layers and streams, respectively. *F* is the transform function between wave number  
 377 space and *k* space, defined from a finite set of  $k(\lambda_i)$ .

378 The application of Double-*k* approach in oxygen A-band has been presented in detail in  
 379 Duan et al. 2005. Here we take oxygen B-band as an example. The detailed fast radiative transfer  
 380 model for simulating high-resolution oxygen B-band is as follows: The first order scattering  
 381 radiance is calculated accurately by using a higher number of layers and streams for all required  
 382 wavenumber grid points. The multiple-scattering component is extrapolated and/or interpolated  
 383 from a finite set of calculations in the space of two integrated gaseous absorption optical depths  
 384 to the wavenumber grids: a double-*k* approach. The double-*k* approach substantially reduces the  
 385 error due to the uncorrelated nature of overlapping absorption lines. More importantly, these  
 386 finite multiple-scattering radiances at specific *k* values are computed with a reduced number of  
 387 layers and/or streams in the forward radiative transfer model. To simulate an oxygen B-band  
 388 spectrum with high accuracy, 33 *k* values and 99 calculations of radiative transfer are chosen in  
 389 our program. This results in around a hundred-fold time reduction with respect to the standard  
 390 forward radiative transfer calculation.



391

392 **Figure 2.** (a) High resolution reflectance at EPIC O2 B-Band simulated by fast radiative model  
 393 (double-k) and benchmark (LBL); Difference between simulated reflectance by (b) double-k and  
 394 LBL for a clear sky case and (c) a thin liquid water cloud case with COD=2. Here SZA and view  
 395 angle =35°, surface albedo = 0.02, aerosol optical depth = 0.08, and reflectance difference (%) =  
 396  $100 \times ((\text{double-k}) - \text{LBL}) / \text{LBL}$ .

397 As shown in Fig. 2, under clear sky and thin liquid water cloud situations, the simulated  
 398 high resolution upward diffuse oxygen B-band spectra from LBL calculation and double-k  
 399 approach are compared. The spectrum difference between LBL calculation and double-k  
 400 approach is very small (Fig. 2a). Under both situations, most of the relative difference between  
 401 these two methods are under 0.5%. The obvious relative difference (>1%) occurs only in the  
 402 wavelength range with high absorption optical depth, which has little contribution to the  
 403 integrated solar radiation. Therefore, for the simulated narrowband measurements at EPIC

404 oxygen B-band, the relative difference between LBL and double-k approach is much smaller  
 405 than that of the high resolution spectrum, which is less than 0.1% for clear day. Compared to  
 406 clear sky situation, the relative difference for cloud situations can be bigger. As shown in Table  
 407 1, the relative difference is -0.06% and -0.32% for typical high level optical thin cloud and low-  
 408 level thick cloud situations, respectively. The comparison of simulated narrowband measurement  
 409 at EPIC oxygen A-band channel (764 nm) is also shown in Table 1, the relative differences  
 410 between LBL and double-k approach are -0.06%, 0.21% and 0.23% for clear day, high level thin  
 411 cloud and low level thick cloud cases, respectively. In general, the accuracy of double-k  
 412 approach for both oxygen A and B absorption bands is high.

413 **Table 1.** Comparison of simulated narrowband measurement at EPIC A- and B-Band channels

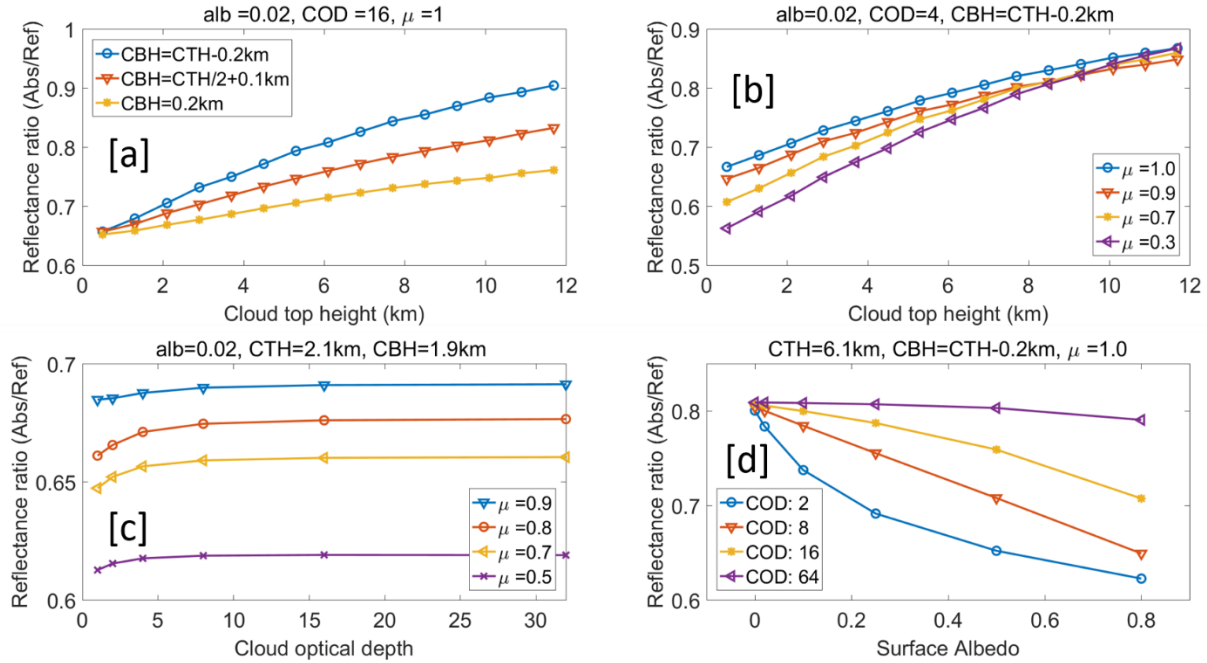
| Case (SZA=35, surface albedo =0.02)      |        | Line by Line | Double k | Relative Difference |
|--|--------|--------------|----------|---------------------|
| Clear Day                                | 688 nm | 0.026963     | 0.026985 | +0.08%              |
|  | 764 nm | 0.013979     | 0.013970 | -0.06%              |
| Thin cloud (COD=2, 8.3-8.5 km, liquid)   | 688 nm | 0.098444     | 0.098131 | -0.32%              |
|  | 764 nm | 0.071359     | 0.071507 | +0.21%              |
| Thick cloud (COD=16, 1.5-2.9 km, liquid) | 688 nm | 0.396354     | 0.396117 | -0.06%              |
|  | 764 nm | 0.233937     | 0.234485 | +0.23%              |

414

### 415 2.3.3 Simulation of oxygen A- and B-bands for different atmospheric conditions

416 Using the EPIC measurement simulation package, we made a series of simulations with  
 417 different settings for surface albedo, solar zenith angle, COD, CTH (CTP), and cloud geometric  
 418 (pressure) thickness (or cloud bottom height). The results of these simulations consist of a data  
 419 table, which can be used not only to calculate the coefficients for the analytic equation, but also  
 420 to study the sensitivity of every variable.

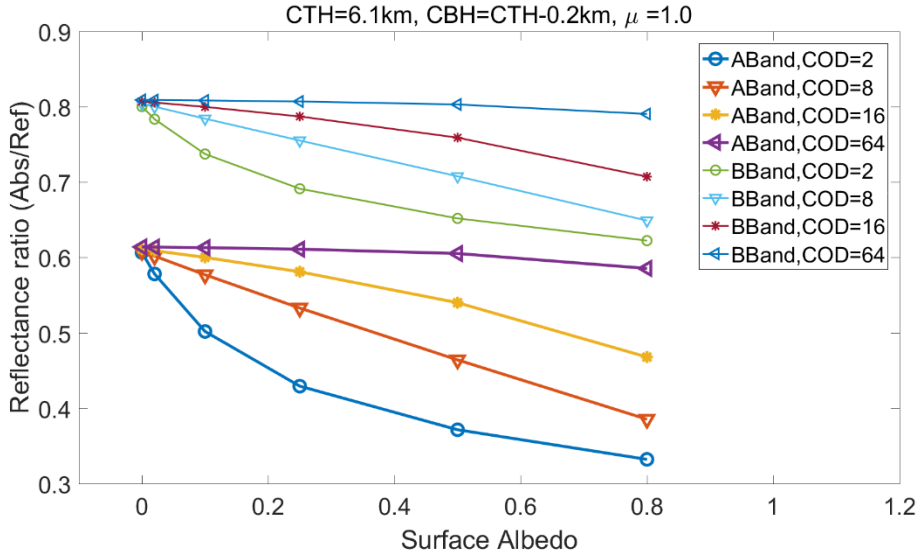
421



422

423 **Figure 3.** Ratio of simulated reflectance measurements for EPIC B-band to B-band reference  
 424 with different surface albedo (alb), COD,  $\mu$  (cosine of solar zenith angle), cloud top height  
 425 (CTH) and cloud bottom height (CBH).

426 According to the previous theory study, the ratio of reflectance radiance (i.e., absorption  
 427 to the reference) at TOA is determined by the photon path length distribution at oxygen A/B  
 428 bands: the larger the mean photon path length, the stronger the absorption, and the smaller the  
 429 reflectance ratio. To make the figures easy to view and understand, we use cloud top and bottom  
 430 geometric height to represent CTP and thickness information in Fig. 3. As shown in Fig. 3a, the  
 431 ratio of upward diffuse radiance at oxygen B-band and its reference band is sensitive to the cloud  
 432 top height (pressure). The higher the CTH, the larger the ratio. At the same time, this ratio is  
 433 affected by the cloud bottom height (or cloud geometric thickness) when the other cloud  
 434 parameters are fixed, the lower the cloud bottom (or the larger the cloud geometric thickness),  
 435 the smaller the ratio. It is consistent with the theory analysis: (1) the higher the CTH, the shorter  
 436 the mean photon path length, and the weaker the absorption; (2) when the COD is given, larger  
 437 cloud geometric thickness means smaller cloud density, then the sunlight can penetrate deeper  
 438 into the cloud, which results in a longer mean photon path length. In Fig. 3b, for clouds with  
 439 given CTH, COD and geometric thickness, the ratio decreases with the solar and view angles.  
 440 This can be understood as: the larger the solar and view angles, the longer the mean photon  
 441 pathlength, and the stronger the absorption. In Fig. 3c, for clouds with given CTH and geometric  
 442 thickness, when the COD is small (e.g., COD < 5), the reflectance ratio increases with COD.  
 443 However, when COD is larger than 16, the effect of COD is small. This is because the larger the  
 444 COD, the shallower the sunlight penetration, and the shorter the mean photon pathlength. In Fig.  
 445 3d, for clouds with given COD, CTP, and geometric thickness, the ratio decreases with surface  
 446 albedo. The smaller the COD, the stronger the impact of the surface albedo. This is because the  
 447 thick cloud prevents the incident sunlight from passing through it to reach the surface, and also  
 448 prevents the reflected light from going back to the TOA.



449

450 Figure 4. Ratio of simulated reflectance measurements for EPIC A and B absorption band to  
 451 reference band with different surface albedo.

452 For oxygen A-band, the ratio of upward diffuse at absorption and reference bands shows  
 453 similar characteristics as for oxygen B-band. Compared to oxygen B-band, under the same  
 454 atmospheric conditions, the oxygen absorption at A-band is stronger, and the ratio of A-band to  
 455 its reference band has smaller values (shown in Fig. 4). As stated previously, for land area that  
 456 covered with plants, the surface albedo may change substantially from oxygen B-band to A-band  
 457 due to the presence of the red edge. Therefore, accurate spectral data of surface albedo for CTP  
 458 retrieval is vitally important, especially for optically thin clouds.

459

### 460 3. Application and validation of the CTP retrieval method

#### 461 3.1 Case studies of CTP retrieval

462 The dataset of DSCOVER EPIC measurements at GMT 00:17:51 on July 25, 2016 is used for  
 463 the case studies. The reflectance at oxygen A and B bands with related solar zenith and viewing  
 464 angles are obtained from the EPIC level 1B data; COD information (retrieved from other EPIC  
 465 channels) is obtained from EPIC level 2 data. The surface albedo data is obtained from Global  
 466 Ozone Monitoring Experiment 2 (GOME-2) Surface Lambertian-equivalent reflectivity (LER)  
 467 data. The detailed information of dataset is shown in the acknowledgements and dataset. To  
 468 reduce the impact of the Earth surface, we selected the region located in spatial range of (75° S  
 469 to 85° N, 177° W to 175° W) for case studies, which is mainly covered by ocean. To constrain  
 470 the influence of surface albedo and broken clouds, only pixels with total cloud cover (i.e., EPIC  
 471 Cloud mask = 4), surface albedo less than 0.05, and liquid assumed COD larger than 3 are  
 472 considered. In the selected region, around 10000 pixels are finally chosen for case studies.

473 In our retrieval algorithm, we have two kinds of retrieval results: baseline CTP and retrieved  
 474 CTP. The baseline CTP is used as a reference for the retrieved CTP. It is similar to the effective  
 475 CTP in Yang et al., (2019), which does not consider cloud penetration. The retrieved CTP is  
 476 calculated by the analytic equation, which considers the in-cloud and below-cloud interaction.

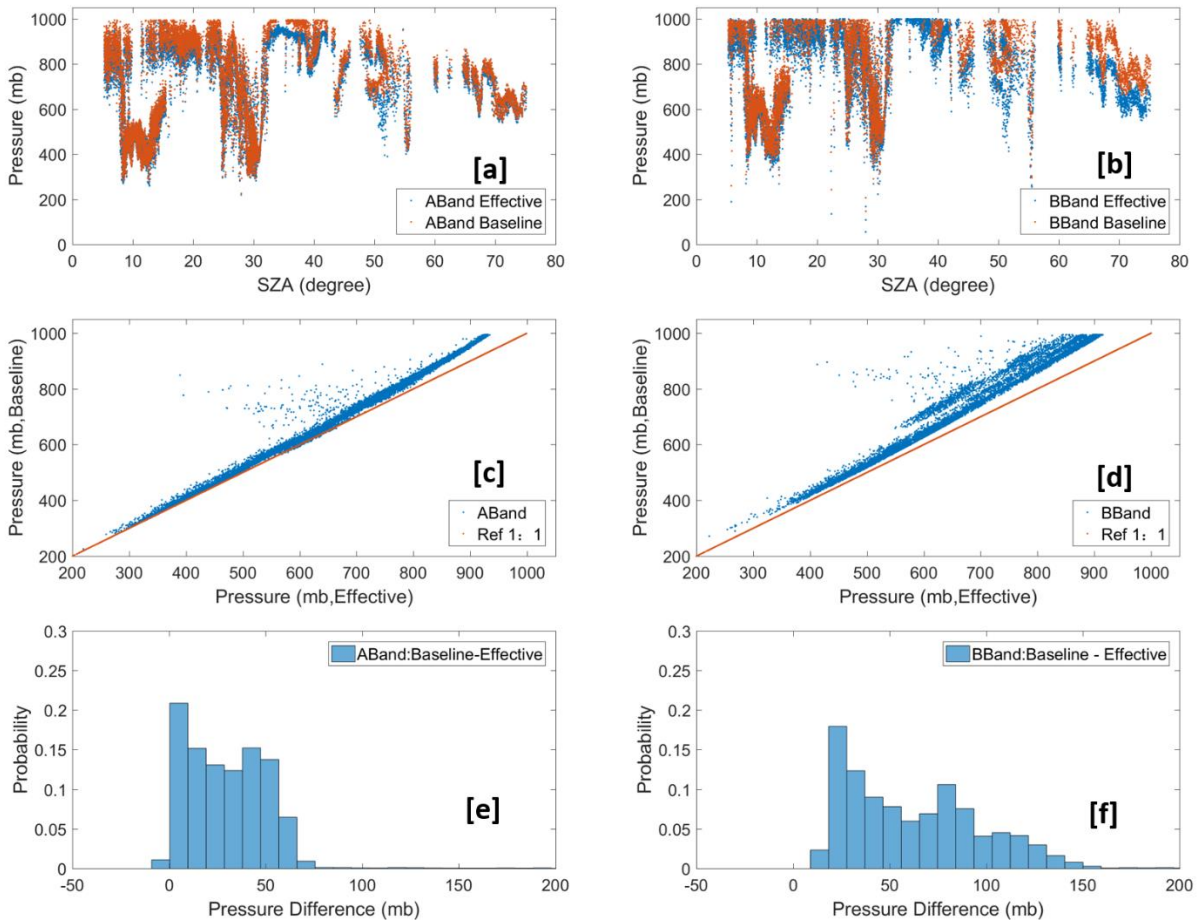


477 During the baseline CTP calculation, the impact of penetration in-cloud is ignored, and the  
 478 incident light that reached cloud top is assumed reflected back directly. As shown in Eq. (19), the  
 479 baseline absorption optical depth  $\tau_{base}$  is derived from the ratio of upward diffuse at absorption  
 480 bands and their reference bands directly. According to the model calculated oxygen A and B  
 481 bands absorption optical depth profile at the specific solar zenith angle, the baseline CTP can be  
 482 derived directly.

$$483 \quad \tau_{base} = \log\left(-\frac{R_{abs}}{R_{ref}}\right) / \left(\frac{1}{\cos(\theta_{sza})} + \frac{1}{\cos(\theta_{view})}\right) \quad (19)$$

484 As shown in Fig. 5, the baseline CTP value at A-band is slightly higher than the effective  
 485 CTP from NASA ASDC L2 data. But the baseline CTP value at B-band is substantially higher  
 486 than the effective CTP from NASA ASDC L2 data. For both A-band and B-band, the difference  
 487 between baseline CTP and effective CTP increases with the CTP. For low-level clouds, the mean  
 488 differences of them are up to 60 mb and 100 mb at A-band and B-band, respectively. The  
 489 difference may be mainly from the calculation of oxygen A and B bands absorption coefficients  
 490 or the absorption optical depth profile.

491



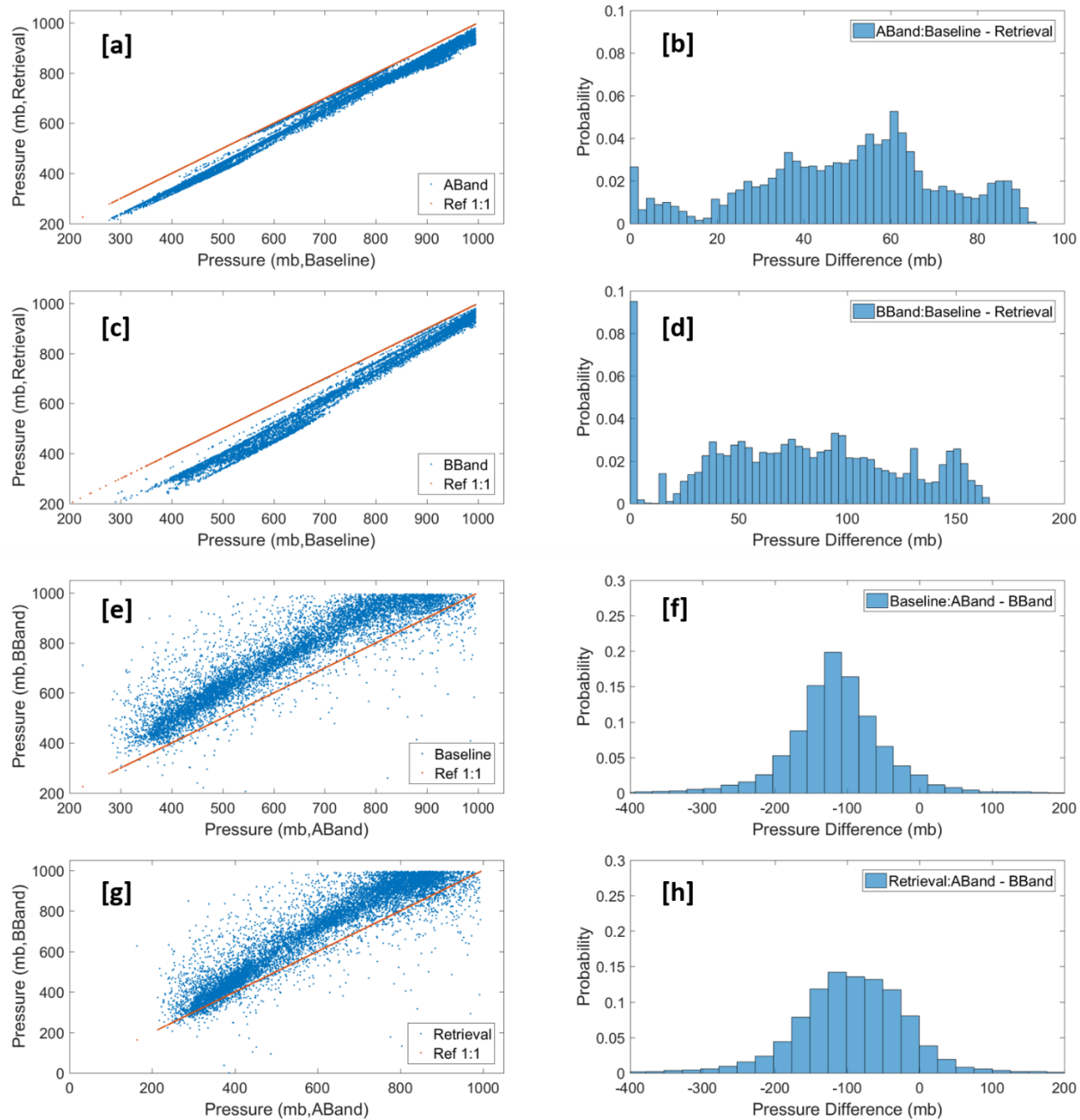
492

493 **Figure 5.** The comparison of effective CTPs (reference from NASA ASDC data) and baseline  
 494 CTPs from our retrieval algorithm for EPIC A and B bands.

495 Based on the simulated reflectance ratio under different atmospheric conditions, we can  
496 calculate the coefficients for the analytic radiative transfer equations by using a nonlinear fitting  
497 algorithm. The coefficients for different SZA's are calculated individually to reduce the fitting  
498 error. Based on the calculated coefficients, we can retrieve the CTP with DISCOVER EPIC  
499 observation data at oxygen A and B bands.

500 During the CTP retrieval, with the exception of the previously mentioned analytic  
501 equation coefficients, we can get the surface albedo data from GOME, obtain reflectance data,  
502 solar zenith and view angles, COD, etc. from the NASA ASDC data file. Another very important  
503 step in the retrieval processing is the acquisition of cloud pressure thickness data, which has a  
504 substantial impact on the retrieval results. We currently use a statistical approach (i.e., cloud  
505 pressure thickness (mb) = 2.5\* COD +23) to estimate the cloud pressure thickness based on  
506 COD. As shown in Figs. 6a-6d, the retrieved CTP when considering cloud penetration is smaller  
507 than baseline CTP. For this case, the mean difference between baseline CTP and retrieved CTP  
508 for oxygen A-band and B-bands are around 57 mb and 85 mb, respectively, which is consistent  
509 with theoretical expectations. For clouds with a given CTP, the mean photon path length will  
510 increase substantially when considering cloud penetration. A decrease in retrieved CTP will  
511 result in order to match the measurement ratio of absorption to reference. Compared to the O2 A-  
512 band, both baseline CTP and retrieved CTP for the O2 B-band are larger (Figs. 6e-6h). This is  
513 because the absorption of solar radiation in the O2 B-band is weaker than that of the O2 A-band,  
514 and the incident light at oxygen B-band can penetrate deeper into the cloud, allowing more light  
515 to pass through. The difference in retrieved CTP between B band and A band (approx. 93 mb  
516 with standard deviation of 83 mb) is generally reduced in comparison to baseline B band and A  
517 band (approx. 114 mb with standard deviation of 73 mb). This indicates, as expected, more  
518 photon penetration correction for B-band than A-band.

519



520

521 **Figure 6.** (a-d) The comparison of retrieved CTPs and baseline CTPs for EPIC A and B bands;  
 522 (e-f) the comparison of retrieved CTPs and baseline CTPs between EPIC A- and B- bands.

523 We also used the LUT based method to do the retrieval for the same observation data,  
 524 because both methods share the same EPIC simulation package and the same simulated data  
 525 table, the results of which are similar.

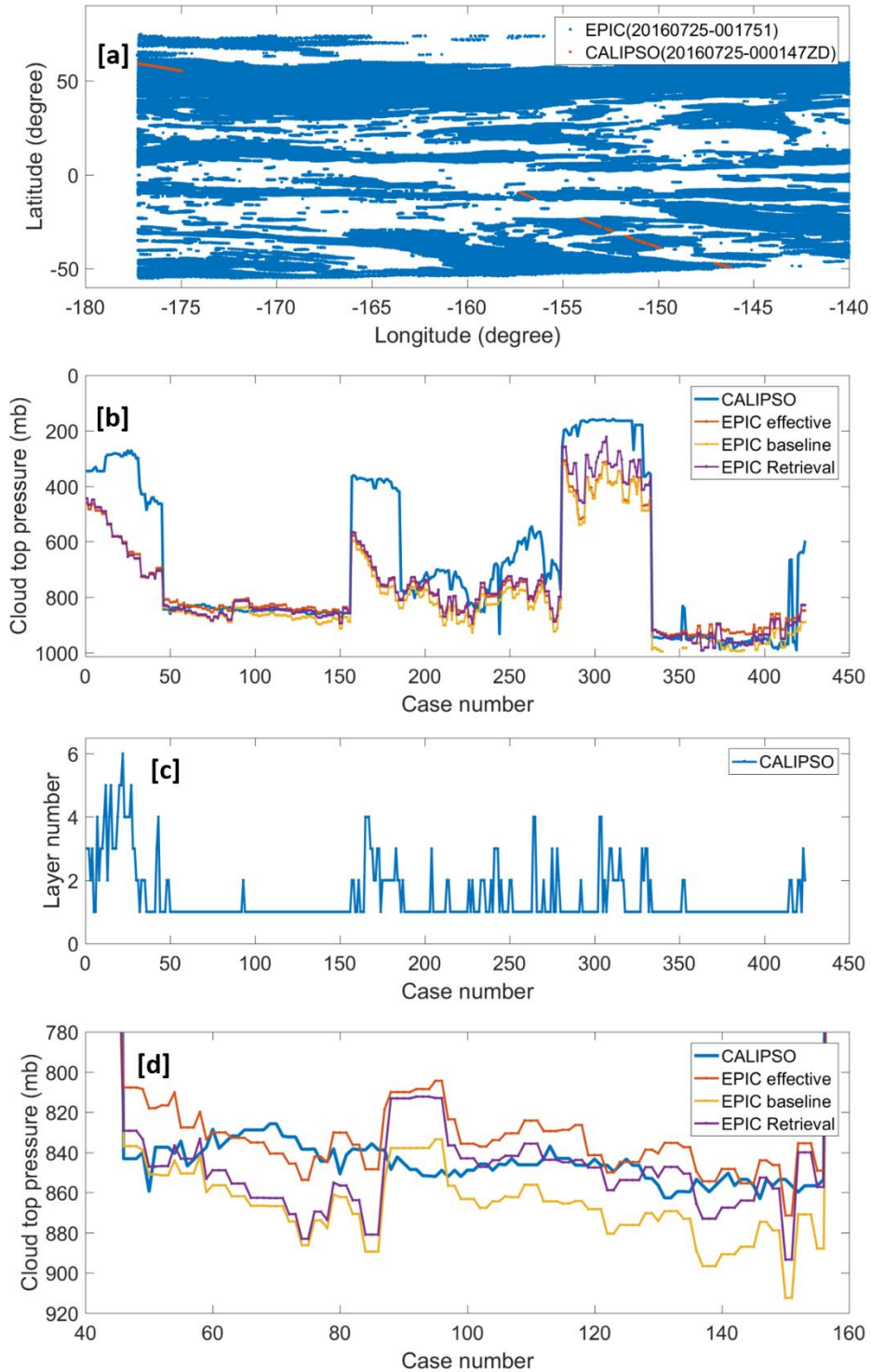
526

### 527 3.2 Validation of the retrieval method

528 To validate the analytic transfer inverse model method for CTP retrieval, we used another  
 529 independent measurement of CTP, i.e., cloud layer top pressure from Cloud-Aerosol Lidar and  
 530 Infrared Pathfinder Satellite Observations (CALIPSO, Vaughan et al., 2014) as a reference. For

531 the previously stated case, i.e., DSCOVER EPIC measurements at GMT 00:17:51 on July 25,  
532 2016, we used the cloud layer data from CALIPSO IIR Version 4.2 Level 2 product with 5 km  
533 resolution at GMT 00:01:47 on July 25, 2016 as its reference to do validation. To constrain the  
534 error from spatial differences between different satellite measurements, we only chose the pixels  
535 of EPIC and CALIPSO measurements with a spatial distance of within  $0.1^\circ$  (degree of latitude or  
536 longitude) to make comparisons. For the EPIC measurements, the same as previously stated,  
537 only pixels with total cloud cover (i.e., EPIC Cloud mask = 4), surface albedo less than 0.05, and  
538 liquid assumed COD larger than 3 are considered. As shown in Fig. 7a, there are a series of  
539 pixels (around 400 cases) from EPIC and CALIPSO measurements can be used for the validation  
540 analysis. For the convenience of reading, we perform the analyses by using the case number as x  
541 axis. Fig. 7b shows the comparisons of cloud layer top pressure from CALIPSO and different  
542 CTPs (i.e., effective CTP, baseline CTP, and retrieved CTP) from EPIC measurements. Fig. 7c  
543 shows the cloud layer number measured by CALIPSO. According to Figs. 7b and 7c, we can get  
544 some results: under single layer cloud situations, the CTPs derived from EPIC measurements are  
545 close to the CTP from CALIPSO; under multi-layer cloud situations, the CTP derived from EPIC  
546 measurements are larger than the CTP from CALIPSO. Fig. 7d shows the expanded view of the  
547 Fig. 7b for some cases under single layer cloud situations. For these single layer cloud cases  
548 (with case number 46 ~ 156), the mean values of CTP of CALIPSO, EPIC effective, EPIC  
549 baseline and EPIC retrieval are 846, 834, 866 and 850 mb, respectively. Compared to the CTP  
550 from CALIPSO measurements, the EPIC effective and baseline CTPs are 12 mb smaller or 20  
551 mb larger, respectively; the EPIC retrieval with consideration of photon penetration is only 4 mb  
552 larger. This shows that our method for the CTP retrieval is valid and accurate under single layer  
553 cloud situations with  $COD > 3$  and low surface albedo. Under multi-level cloud situations, the  
554 high-level clouds are often thin clouds, which can be detected by CALIPSO but hard to derive by  
555 our retrieval method. It is because the EPIC retrieved CTP mainly shows the pressure of cloud  
556 layer that reflects the major part of incident sun light.

557



558

559 Figure 7. (a) The geolocation match of EPIC measurement at GMT 00:17:51 and CALIPSO  
 560 measurement at GMT 00:01:47 on July 25, 2016; (b) the comparisons of cloud layer top pressure  
 561 from CALIPSO measurements and the CTPs derived from EPIC measurements; (c) the cloud  
 562 layer number from CALIPSO measurements; and (d) the expanded view of (b) for some cases  
 563 under single layer cloud situations.

564

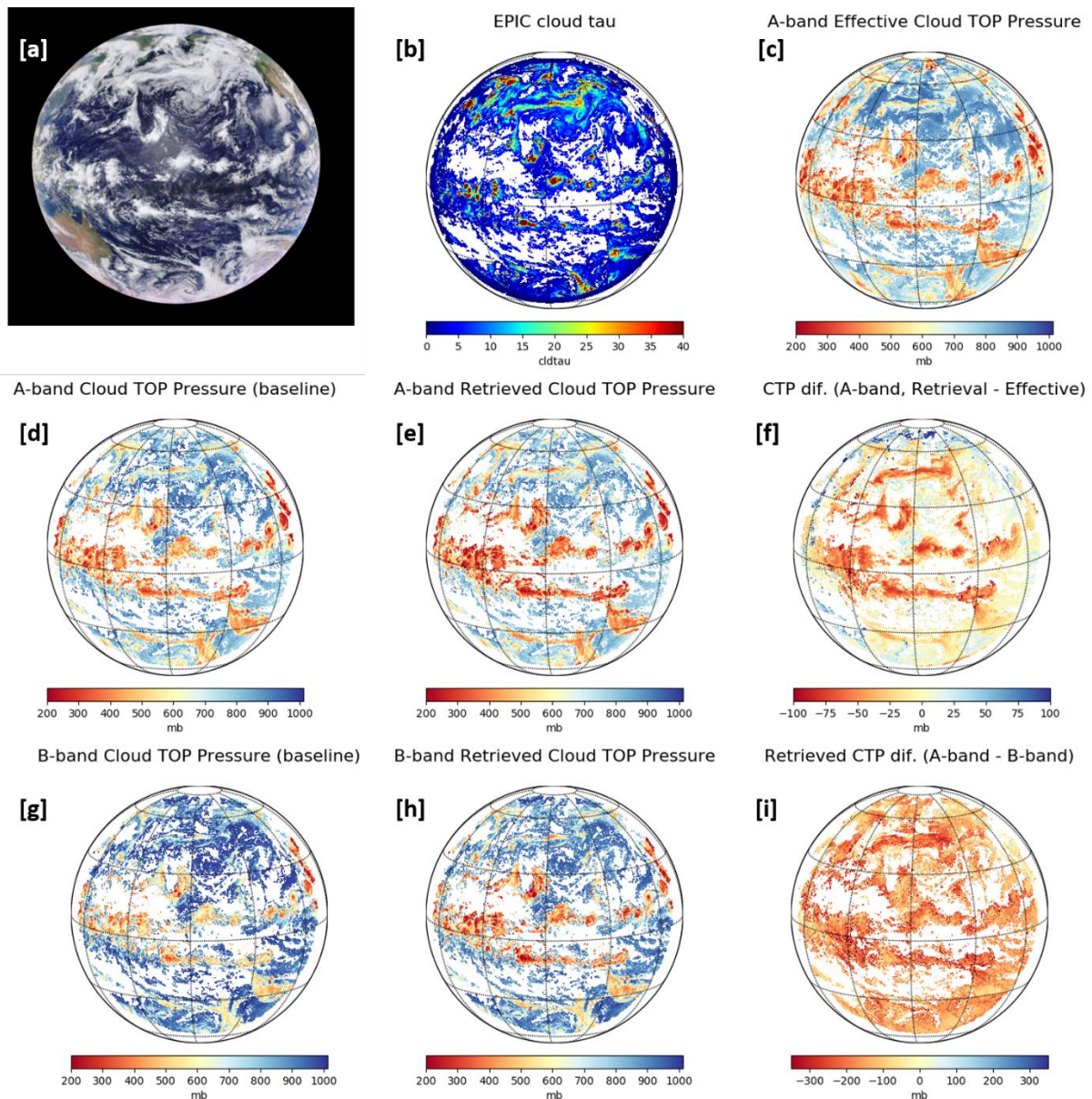
### 565 **3.3 Retrieval of global observation**

566 We applied our retrieval algorithm on the global DISCOVER EPIC measurement data at  
567 oxygen A and B bands. During the retrieval, only pixels with total cloud cover (i.e., cloud mask  
568 index of 4), surface albedo  $< 0.25$ , and COD  $\geq 3$  are considered. To make the pictures easy to  
569 visualize and analyze all invalid values are plot as white (or blank) pixels.

570 Figure 8a shows the synthesized RGB picture of EPIC measurements at GMT time 00:17:51  
571 on July 25, 2016. At this point in time the sun light covers most of the Pacific Ocean. In this  
572 figure, the white pixels represent cloud cover. Figure 8b shows the global COD (NASA ASDC  
573 L2 data), in which the white areas and colorful areas indicate the clear sky areas and cloudy  
574 areas, respectively. On the whole, the cloudy areas are consistent with the RGB image. The  
575 highlight (red) areas indicate that the cloud systems there contain optically heavy clouds. Figure  
576 8c shows the A-band effective CTP (NASA ASDC L2 data), where the white areas indicate clear  
577 sky or no valid values, warm (brown) and cold (blue) color areas indicate high-level and low-  
578 level clouds, respectively. According to the A-band effective CTP, the high-level clouds are  
579 dominant in the equatorial area, and the low-level clouds play a major role in the cloud systems  
580 in the Northern Pacific area. Figures 8d and 8e show the baseline and retrieved CTP at A-band,  
581 respectively, which cloudy areas are consistent with the A-band effective CTP image on the  
582 whole. Due to the filtering setting in the CTP retrieval algorithm, there are more white pixels  
583 (invalid values) in these two figures. The difference of A-band retrieved CTP and A-band  
584 effective CTP is shown in Fig. 8d. The A-band retrieved CTP is overall smaller than A-band  
585 effective CTP, which difference is within 100 mb. The highlighted (brown or red) areas are  
586 located in the high level clouds areas or large COD areas. This indicates that the complexity of  
587 cloud system has significant impact on the CTP retrieval. Figures 8g and 8h show the baseline  
588 and retrieved CTP in B-band respectively, which are similar to, but greater than the A-band. As  
589 shown in Fig. 8i, the retrieved CTP at EPIC B-band is overall significantly larger than the  
590 retrieved CTP at EPIC A-band, which mean difference is up to 200 mb.

591





592

593 **Figure 8.** (a) RGB image from DSCOVER EPIC measurement at GMT time 00:17:51 on July 25,  
 594 2016; (b) and (c) COD (liquid assumption) and A-band effective CTP from NASA ASDC EPIC  
 595 L2 products; (d) and (e) Baseline and retrieved CTP derived from EPIC A-band measurement;  
 596 (f) the difference of A-band retrieved CTP and A-band effective CTP; (g) and (h) Baseline and  
 597 retrieved CTP derived from EPIC B-band measurement; and (i) the difference of retrieved CTP  
 598 between EPIC A-band and B-band.

599 As previously stated in Sect. 3.2: under single-layer cloud situations, the CTPs derived from  
 600 EPIC A-band measurements have good agreement with the CTP from CALIPSO measurements;  
 601 under multiple-layer cloud situations, the CTPs derived from EPIC measurements may be larger  
 602 than the CTPs of high level thin-clouds due to the effect of photon penetration. Therefore, in the  
 603 global range, for the large scale low-level stratus clouds, the retrieved CTPs from EPIC A-band  
 604 measurements should agree well with the actual value of CTPs, but for the complex cloud system



605 with multiple-layer clouds, the CTPs derived from EPIC A-band measurements may be larger  
606 than that of high level thin-clouds.

#### 607 **4. Conclusion**

608 The in-cloud photon penetration has significant impacts on the CTP retrieval when using  
609 DSCOVER EPIC oxygen A- and B- band measurements. To address this issue, we proposed two  
610 methods, (1) the LUT based method and (2) the analytic transfer inverse model method for CTP  
611 retrieval with consideration of in-cloud photon penetration. In the analytic transfer inverse model  
612 method, we build an analytic equation that represents the reflection at TOA from above cloud,  
613 in-cloud, and below-cloud, respectively. The coefficients of this analytic equation can be  
614 derived from a series of EPIC simulations under different atmospheric conditions using a non-  
615 linear regression algorithm. With EPIC observation data, the related solar zenith and sensor view  
616 angle, surface albedo data, COD, and estimated cloud pressure thickness, we can retrieve the  
617 CTP by solving the analytic equation.

618 We developed a package for the DSCOVER EPIC measurement simulation. The high  
619 resolution radiation spectrum must be simulated first and then integrated with the EPIC filter  
620 function in order to accurately simulate EPIC measurements. Because this process is highly time-  
621 consuming, a polynomial fitting function is used when calculating the oxygen absorption  
622 coefficients under different atmospheric conditions. At the same time, the double-k approach is  
623 applied to do the high-resolution spectrum simulation to further reduce time-costs, which can  
624 obtain high accuracy results with hundred-fold time reduction. The results of the EPIC  
625 simulation measurements are consistent with theoretical analysis.

626 Based on the EPIC simulation measurements, we derived a series of coefficients from  
627 various solar zenith angles for the analytic EPIC equations. Using these coefficients, we  
628 performed CTP retrieval for real EPIC observation data. We have two kinds of retrieval results:  
629 baseline CTP and retrieved CTP. The baseline CTP is similar to the effective CTP in Yang et al.,  
630 (2019), which does not consider cloud penetration. The retrieved CTP is derived by solving the  
631 analytic equation, with consideration of the in-cloud and below-cloud interactions. Compared to  
632 the effective CTP provided by NASA ASDC L2 data, the baseline CTP value at A-band is  
633 slightly higher, but the baseline CTP value at B-band is substantially higher. The retrieved CTP  
634 for both oxygen A- and B- bands is smaller than the related baseline CTP. At the same time,  
635 compared to the oxygen A-band, both baseline CTP and retrieved CTP at oxygen B-band is  
636 larger. The cloud layer top pressure from CALIPSO measurements is used to validate the CTP  
637 derived from EPIC measurement. Under single-layer cloud situations, the retrieved CTPs for  
638 oxygen A-band agree well with the CTPs from CALIPSO, which mean difference is within 5 mb  
639 in the case study. Under multiple-layer cloud situations, the CTPs derived from EPIC  
640 measurements may be larger than the CTPs of high level thin-clouds due to the effect of photon  
641 penetration.

642 Currently, this analytical transfer model method can only retrieve CTP, and it still need  
643 cloud pressure thickness as an input parameter. However, in the satellite observations, both CTP  
644 and cloud pressure thickness are unknown. The estimation or assumption of cloud pressure  
645 thickness will bring in extra error in CTP retrieval. In the near future, we plan to address this  
646 issue.

647 **Data availability**

648 Dataset of DSCOVER EPIC Level 1B can be found in  
649 [https://eosweb.larc.nasa.gov/project/dscovr/dscovr\\_epic\\_11b\\_2](https://eosweb.larc.nasa.gov/project/dscovr/dscovr_epic_11b_2); dataset of EPIC Level 2 can be  
650 found in [https://eosweb.larc.nasa.gov/project/dscovr/dscovr\\_epic\\_12\\_cloud\\_01](https://eosweb.larc.nasa.gov/project/dscovr/dscovr_epic_12_cloud_01); dataset of  
651 surface albedo from GOME can be found in [http://temis.nl/surface/gome2\\_ler/databases/](http://temis.nl/surface/gome2_ler/databases/);  
652 dataset of cloud layer data from CALIPSO can be found in  
653 [https://eosweb.larc.nasa.gov/project/calipso/cal\\_lid\\_12\\_05kmclay\\_standard\\_v4\\_20](https://eosweb.larc.nasa.gov/project/calipso/cal_lid_12_05kmclay_standard_v4_20).

654  
655 **Author contributions.**

656 All authors contributed to planning and writing of the paper. QM initiated and led the EPIC  
657 CTP retrieval project, designed the analytic transfer inverse model for EPIC observation. BY  
658 developed a fast radiative transfer model for simulating high resolution oxygen B-band,  
659 implemented the CTP retrieval algorithms by using the analytic transfer inverse model and look-  
660 up table method, and drafted the paper. EM calculated the high resolution Oxygen absorption  
661 coefficients by using the LBLRTM model and HITRAN database. YY provided access to the  
662 EPIC Level 1B and Level 2 products and provided guidance for the evaluation of EPIC CTP  
663 retrieval. AM and ABD conducted the EPIC related studies and provided guidance for the design  
664 of EPIC CTP retrieval algorithm with the consideration of in-cloud photon penetration.

665  
666 **Competing interests**

667 The authors declare that they have no conflict of interest.

668  
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675  
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