

Response to Anonymous Referee #1.

We thank the Anonymous reviewer #1 for a thorough review and useful suggestions, which have been carefully implemented in the revised manuscript. The detailed answers to all comments are provided below.

5

General comments

1) The manuscript lacks some details on the upgrade and development of the lidar and retrieval algorithms (including calibration) during the recent years since 2012 (after the last publications from this group by Soupravaren et al. 1999). E.g. please state clearly, which part of the instrument design in ch. 2.2. is new and also provide more details of the upgrade.

10 10 Please be more specific on instrument details (see also my specific comments) to this part, e.g. in ch. 3.1 it is stated that FPI plates were reconditioned, but not further explained. I would also recommend providing more details on the calibration (L79-L83), as this is essential for wind retrieval and wind bias. E.g. is the spectral tuning of the FPI only used for monitoring, or is it used during the wind retrieval (as mentioned in L125). If yes how are these functions used (measured, fitted), and used for wind retrieval from the actual measurements of the same day. Also a short description on how

15 calibration constant C in equation (1) is obtained is missing.

The Section 2 regarding the instrument design, measurement principle and instrument calibration has been entirely reworked, please see the revised text and the answers to specific comments below.

20 20 2) I have two comments to the statistical comparison approach. I am wondering about a justification of using weighted distances for deriving bias and standard deviation in Ch. 3. I would like to see a clear justification of this approach, because I consider this as unusual for instrument intercomparisons, and provide a short description (e.g. equation), how this was implemented. But overall I would recommend deriving these statistical numbers on bias/std. with and without this weighted approach.

25

As a matter of fact, the comparison statistics figures in Table 1. are obtained without the horizontal offset weighting. Besides, as mentioned in Sect. 3, the weighting only affects the average standard deviation, whereas the average bias and correlation coefficient are not affected. Please see the answer to specific comment below.

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My second comment here is related to a missing statistical comparison of the horizontal wind speed (from u and v components, and possibly wind direction). I would propose to add this quantity to chapter 3, and specifically provide a scatterplot (as Fig. 4d) and statistical numbers (as part of Table 1). I would also propose to add the statistics of all radiosonde comparisons to Table 1 as an additional row, and discuss these numbers in the text.

35 The results and discussion of statistical comparison for the total wind and wind direction have been added into Sect. 3 and Table 1.

Specific comments

40 *Lines 12, 523 Provide numbers for vertical and temporal resolution; “high resolution” is different for several application areas*

The respective fragment in the abstract and the summary have been modified: **“After a recent upgrade, the instrument gained the capacity of wind profiling between 5 and 75 km altitude with 45 vertical resolution up to 115 m and temporal resolution up to 5 minutes.”**

Line 29 Provide a reference for deriving wind speed on regular bases from space-borne temperature measurements using geostrophic assumptions.

Reference provided: **Oberheide J, Lehmacher G A, Offermann D, Grossmann K U, Manson 50 A H, Meek C E, Schmidlin F J, Singer W, Hoffmann P, Vincent R A. Geostrophic wind fields in the stratosphere and mesosphere from satellite data. J. Geophys. Res. 107(D23):8175. doi: 10.1029/2001JD000655, 2002.**

Line 38 I would propose to add some more references in the introduction of wind lidars using molecular backscattering, 55 especially here also mention ALADIN and its airborne demonstrator

The following text has been added: **“The direct-detection technique for wind profiling has been successfully realized in an airborne Doppler lidar – A2D, Aeolus Airborne Demonstrator (Reitebuch et al., 2009). A2D instrument served a prototype for the most ambitions endeavor in the 60 context of lidar wind profiling – the first ever satellite-borne Doppler lidar ALADIN (Atmospheric Laser Doppler INstrument) (ESA, 2008; Stoffelen et al., 2005), that has been successfully launched by European Space Agency (ESA) in August 2018 (Kanitz et al., 2019).”**

Line 58 Add 1-2 references for Aeolus here

65 Done

Line 70 Parameters of the FPI are introduced here, while the operating wavelength is not stated (at this place of the manuscript).

70 *Line 91, eq 2 The introduction of parameter $P(z, 40^\circ)$ is missing in the text. 95-99 The vertical pointing beam is used to compensate for laser frequency drifts, with a value, which is constant for each altitude (average over 15-25 km). Please discuss, if there are or not altitude dependent effects in the calibration, which need to be compensated.*

75 *Line 100ff Please provide more instrumental details, as laser frequency stability (shot-to-shot), laser divergence (at output of beam expander) and laser linewidth. Also FOV of telescope should be provided, as well as diameter of multimode fiber. The method of mode scrambling should be shortly introduced. Also the “reconditioning” of the FPI plates (as mentioned in ch. 3.1) should be explained here (new coating? New polishing?)*

All of the above comments have been carefully implemented in the Sect. 2, please see the revised text. The reconditioning of the OHP wind lidar FPI has not actually been carried out. It was not necessary since the auxiliary experiments have shown that the spectral characteristics of the FPI have remained unchanged. The reconditioning of FPI has only been done for the La Reunion wind lidar.

80

Line 150 Is this equation of the error in units of m/s? Is C the same constant as introduced in eq. (1)?

This equation describes the error in the response profile R , which is unitless. C is the same constant as in eq. (1). This has been clarified in the text.

85 *Line 200 Figure 3: black circles are hardly visible, e.g. use different colour*

The colour of the circles in Fig. 3 (now Fig. 4) has been changed.

Line 220 Do you provide numbers for correlation coefficient as r or r^2 . Please state explicitly in the text and in Table I.

90 We provide the Pearson’s correlation coefficient r . This has been specified in the text and in the Table 1 caption.

95 *212 Please explain the rationale to compute the comparison statistics, by “weighting” the difference with the horizontal offset between the measurements. I think this is very unusual. I would propose to provide statistics without weighting, or at least show both the non-weighted or weighted results. The weighting should be shortly explained (e.g. via an equation).*

We have used the horizontal offset weighting for computing the averages of comparison statistics in order to evaluate the effect of the spatial variability of the horizontal wind components. The results of intercomparison in Sect. 3 and Table 1 are computed as ordinary arithmetic averages. The following text and equation have been added:

100 **“For evaluating the effect of the horizontal offset between the lidar and RS measurements we computed the offset-weighted averages of the intercomparison statistics and compared them with the ordinary averages. The weight for each individual value is defined as $w=1-D/\bar{D}_{\text{max}}$, where D is the mean distance between the lidar and RS sampling locations and D_{max} is the maximum distance amounting to 69 km (Table 1). We note that the horizontal-offset weighting of the 105 differences neither affects the mean difference nor the mean correlation but reduces the standard deviation for the wind components and total wind by about 0.2 m/s.”**

Line 240 Figure 4: y-intercept also in units of m/s

The figures' legend has been modified accordingly

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314-316 Please explain, how a possible Mie-induced bias would be recognized in the profiles, e.g. too high or too low values? Do you correct for the Mie-induced bias in the wind retrieval (or any QC), or is it only compensated by the FPI spectral configuration (spacing, FWHM)?

115 The Mie-induced bias would appear as sharp enhancement in the wind profile towards higher absolute values. Such a bias may appear in case of the spectral detuning of the FPI bandpasses with respect to the laser backscattered line. The Mie bias can be corrected for, however in reality this is required only in the case of cirrus clouds with scattering ratio above 20 or so. Otherwise, the correction is unnecessary as we demonstrate in the article.

The following paragraph has been added in the beginning of Sect. 3.1:

120 **“Although the Mie-backscattered line is narrow (0.08 pm) compared to the thermally-broadened Rayleigh line (2 – 2.4 pm) the intensity of the former may be substantially higher and thereby alter the spectral shape of the return signal. In this case, a disproportionately larger flux would be transmitted through one of the FPI bandpasses, affecting its calibration function and introducing a bias into the wind retrieval within the particle layer. The sensitivity to Mie scattering**

125 can be reduced by increasing the FPI spectral spacing, however this also reduces the sensitivity to the Doppler shift. The optimal spectral configuration of the FPI has been established on the base of a theoretical model carried out by Souprayen et al. (1999b). They found that for observable stratospheric wind velocities, the residual Mie-induced error is less than 1 m/s for the scattering ratio $R=10$, which is characteristic of a cirrus cloud readily visible to an unaided eye.”

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Line 505 Fig caption 10: provide date of comparison and mean distance of Aeolus observations to OHP; it would be also good to include Aeolus track in Fig. 3

Date and mean distance have been added to the figure caption. The Aeolus track has been added into Fig. 3 (now Fig. 4).

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Line 515 Please provide distance for altitudes below 5 km of OHP and Aeolus track for spatial variability. Causes could be also related to preliminary nature of Aeolus observations. Have you checked error estimates within Aeolus data products, and potentially exclude data with too high errors (e.g. 8-10 m/s)? Have you checked presence of aerosol or cloud layers, which might influence Aeolus Rayleigh wind retrieval?

140 The distance between RS and Aeolus (91 km) has been provided in the text. The preliminary nature of the Aeolus data is clearly articulated in the text and we refrained from a further discussion on the ALADIN data quality. Indeed, the variability of lower-tropospheric winds on a scale of 100 km in the vicinity of Alps can be much larger than the measurement errors. The Rayleigh-clear profiles that were used for this particular validation case did not feature any error anomalies. From the ground-based lidar 145 and AERONET measurements, we noted an absence of clouds and tropospheric aerosol layers during the period of measurements.

526 Please state that this number of 6 m/s refers to random error.

Done.

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553 Could you be more specific, how this finding should be considered for spatial and temporal collocation requirement for performing comparisons for space-borne wind lidars as Aeolus

155 The respective sentence has been modified: “**This finding is to be considered for Aeolus wind validation activities in a sense that a precise temporal collocation may be more important than the spatial collocation of the measurements.**”

527 *Please specify which optics could be replaced to improve performance*

The respective sentence has been modified: “**We note that the vertical range can potentially be extended to 3 – 80 km through replacement of the beam-commuting and beam-splitting mirrors, for which the resources are available.**”

160 533 *The std of 2.2. m/s refers only to 1 component and not the horizontal wind speed. This should be clarified. I would also propose to add statistics for the horizontal wind speed in the conclusion (see my general comment 2).*

We have provided the statistical figures for the total wind speed and direction in the summary section.

605, Table 1 *Please add in Table caption if you use R or R=r2 as correlation coefficient; I would also propose to add at least columns for mean difference and standard deviation for horizontal wind speed (squared sum of u,v; and possibly wind direction) and also another row with mean quantities over all days of comparison.*

All done.

References: I would propose to add a few more references related to Aeolus (ESA 2008, Stoffelen et al. 2005) and its actual performance (Kanitz et al. 2019, Reitebuch et al. 2019).

All the suggested references have been added into the introduction and Sect. 5.

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Editorials

All done.

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190 **Response to Anonymous Referee #2.**

We thank the Anonymous reviewer #2 for a positive review and useful comments. We provide below the detailed answers to them.

Line 12: It would be good to get some numbers for the improvement, e.g. from typically xx m/s uncertainty to yy m/s uncertainty.

195 *Or were the improvements just technical - then also indicate what has improved.*

The abstract has been modified to clarify the improvement: **“A direct-detection Rayleigh-Mie Doppler lidar for measuring horizontal wind speed in the middle atmosphere (10 to 50 km altitude has been deployed at Observatoire de Haute Provence (OHP) in southern France since 1993. After a recent upgrade, the instrument gained the capacity of wind profiling between 5 and 75 km altitude 200 with vertical resolution up to 115 m and temporal resolution up to 5 minutes.”**

Line 15: Instead of "The evaluation" I suggest to write "An initial evaluation". The present paper, in my opinion, does not provide a full and comprehensive evaluation. With only a few radiosondes and ECWMF profiles, the statistics are not very comprehensive yet.

We believe that 12 spatiotemporally-collocated radiosoundings conducted in various atmospheric 205 conditions over a period of 4 years is sufficient for the instrument performance evaluation. Note that all the intercomparison measurements have been conducted in a “campaign” regime, which is resourceful and costly.

Line 39: drop "there exists"; "with" → "have"; "which" → "and";

Line 56: drop "preparation of"

210 *Line 218: "The both" → "both"*

All done.

Section 2, 2.1: I think it is necessary to give a bit more background on the wind-lidar measurement principle. I strongly suggest to add a schematic Figure showing the two (A and B) Fabry-Perot band-passes spectral shapes, as well as the spectral shape of the backscattered Rayleigh and Mie radiation. Also explain that a Doppler shift of the return signal will enhance one

215 channel (A or B) while reducing the other (B or A). How is spectral calibration obtained? I assume by de-tuning the laser with a wavemeter, and observing the zenith pointing return channel. Please also explain.

The Section 2 regarding the instrument design, measurement principle and instrument calibration has been entirely reworked, please see the revised text. A figure showing the spectral shapes of backscattered line and FPI bandpasses has been added.

220 Around line 70: Please give the manufacturer of the Fabry-Perot interferometer.

It is StigmaOptique, a small French company that does not exist anymore, we mentioned the name in the text.

Around line 151: You might want to say here that the uncertainty scales with $1/\sqrt{t_{\text{acquisition}}}$ and / or with $1/\sqrt{\Delta z}$, where Δz is the vertical resolution chosen for data processing.

225 Thank you for the suggestion. Done.

Fig. 2a: It would be good to show both the NA and NB profiles (or their difference if they are very similar). Maybe also show a raw $R(\theta, z)$ profile?

The NA and NB profiles look identical on such a plot, whereas the difference between them in MHz doesn't have much physical meaning. Meanwhile, the response profile R is not much different from the 230 horizontal wind profile as the latter is obtained by multiplying R by the instrumental constant, which only slightly varies with temperature.

Line 172: By "noise level" you probably mean the "background noise level"? If yes, change text. I would assume that the total noise level would increase a bit at lower altitudes, e.g. at the altitude where low and high gain channels are spliced together.

Yes, we mean the background noise level. The splicing of low and high gain signals is done where the 235 former is orders of magnitude above the background level.

Around line 177: 12 Comparisons over a 4-year period are not a lot. Please add some statement why only so few RS comparisons are made, especially since nearby Nimes launches one or two radiosondes every day.

The wind lidar validation experiment was conceived to rely exclusively on the reference measurements by GPS radiosondes collocated with the lidar acquisition in time and, as close as possible, in space. The 240 Nimes radiosoundings are too far away (>100 km) and not always collocated in time, which would make the attribution of the discrepancies in the wind profiles ambiguous.

Around line 254: By eye, Fig. 4a and 4b seem to indicate increasing standard deviation from about 10 to 30 km. How do standard deviation profiles compare to the estimated uncertainty profile from shot noise? Always a good idea to check such estimates. Maybe this warrants an additional Figure?

245 This aspect is discussed in the original version of the article around line 255. We have included the statistical error profile in Fig. 4 (now fig. 5).

Section 3.1: Please add some explanation, that the very narrow Mie line alters the spectral shape of the return signal, and that this might affect/ alter the calibration function in Section 2.1.

The following paragraph has been added in the beginning of Sect. 3.1:

250 **“Although the Mie-backscattered line is narrow (0.08 pm) compared to the thermally-broadened Rayleigh line (2 – 2.4 pm) the intensity of the former may be substantially higher and thereby alter the spectral shape of the return signal. In this case, a disproportionately larger flux would be transmitted through one of the FPI bandpasses, affecting its calibration function and introducing a bias into the wind retrieval within the particle layer. The sensitivity to Mie scattering 255 can be reduced by increasing the FPI spectral spacing, however this also reduces the sensitivity to the Doppler shift. The optimal spectral configuration of the FPI has been established on the base of a theoretical model carried out by Souprayen et al. (1999b). They found that for observable stratospheric wind velocities, the residual Mie-induced error is less than 1 m/s for the scattering ratio $R=10$, which is characteristic of a cirrus cloud readily visible to an unaided eye.”**

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Figure 6: I suggest that the authors be more critical here. The largest differences between RS wind and Doppler lidar wind do occur near 12 and 17 km, very close to the aerosol / cirrus layers. I don't think the authors should ignore that and simply claim no effect. Could the Mie effect be reduced / quantified by wavelength scanning the zenith return signal in the presence of aerosol layers, and assume negligible vertical wind?

265 The Mie-induced bias would appear as sharp enhancement in the wind profile towards higher absolute values. It would be closely correlated with the scattering ratio, which is obviously not the case here. Such a bias may appear in case of the spectral detuning of the FPI bandpasses with respect to the laser backscattered line. The Mie bias can be corrected for, however in reality this is required only in the case of cirrus clouds with scattering ratio above 20 or so. Otherwise, the correction is unnecessary as we 270 demonstrate in the article.

Line 366: But ECWMF also assimilates stratospheric and mesospheric radiance measurements from satellites, providing a large amount of information on the temperature fields. Since the atmosphere is close to a geostrophic state in the stratosphere and mesosphere, it is not surprising to me that ECWMF winds are quite realistic up to 60 or 70 km.

To the best of our knowledge, there are no operational radiance measurements in the mesosphere that are
275 assimilated into ECMWF model. The highest channel of AMSU (ch14) peaks at around 43 km.

*lines 386/386: Is it the mirrors, or is it the darker sky in the North? Should "due to a better condition . . . mirrors of" be
replaced by "due to the darker sky seen by"?*

It is the mirrors and the alignment issues that lower the signal strength for the East line-of-sight. The sky
background is the same for both directions

280 *Line 534: I am not sure if you have really demonstrated that results are "insensitive" to aerosol. I think
"not very sensitive" would be a better statement.*

A statement "not very sensitive" would have to be quantified, whereas this is not possible as we did not
see any measureable effect.

Around line 535: Can you not measure the temperature profile as well (using the Chanin Hauchecorne method)?

285 To measure the temperature profile using the Chanin Hauchecorne method it is necessary to acquire the
full spectrum of the backscattered signal proportional to the atmospheric density. In the case of our
Doppler lidar the signal is convoluted with the spectral transmission of the Fabry-Pérot interferometer.
The temperature retrieval would thus be prone to a much larger error than using the dedicated LTA lidar
instrument at OHP.

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Line 538: I don't think the authors have provided "insight". They only showed "examples" . Replace the word?

Done.

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300 **Doppler lidar at Observatoire de Haute Provence for wind profiling
up to 75 km altitude: performance evaluation and observations**

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Abstract. A direct-detection Rayleigh-Mie Doppler lidar for measuring horizontal wind speed in the middle atmosphere (10

310 to 50 km altitude has been deployed at Observatoire de Haute Provence (OHP) in southern France since 1993. After a recent upgrade, the instrument gained the capacity of wind profiling between 5 and 75 km altitude with vertical resolution up to 115
m and temporal resolution up to 5 minutes. The lidar comprises a monomode Nd:Yag laser emitting at 532 nm, three telescope assemblies, and a double-edge Fabry-Perot interferometer for detection of the Doppler shift in the backscattered light. In this article, we describe the instrument design, recap retrieval methodology and provide an updated error estimate for horizontal wind. The evaluation of the wind lidar performance is done using a series of twelve time-coordinated radiosoundings conducted at OHP. A point-by-point intercomparison shows a remarkably small average bias of 0.1 m/s between the lidar and the radiosonde wind profiles with a standard deviation of 2.2 m/s. We report examples of a weekly and an hourly observation series, reflecting various dynamical events in the middle atmosphere, such as a Sudden Stratospheric Warming event in January 2019 and an occurrence of a stationary gravity wave, generated by the flow over the Alps. A qualitative comparison between 315 the wind profiles from the lidar and the ECMWF Integrated Forecast System is also discussed. Finally, we present an example of early validation of the ESA Aeolus space-borne wind lidar using its ground-based predecessor.

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1 Introduction

Vertically-resolved measurements of the wind velocity in the middle atmosphere are essential for understanding the global circulation driven by dynamical processes such as gravity and planetary waves interacting with the atmospheric flow (Holton 1983). While weather balloon soundings provide regular observations of horizontal wind profiles up to about 30 km altitude, the region of upper stratosphere and lower mesosphere (USLM, ~30 - 75 km) is poorly covered by observations. The only information on the wind field in this layer available on the regular basis is inferred from horizontal pressure gradients derived from space-borne temperature measurements using geostrophic balance assumption (e.g. Oberheide et al., 2002), however this does not allow characterizing regional-scale dynamical processes.

Of particular challenge are the wind measurements in the so-called radar gap between 20 – 60 km (Baumgarten, 2010). Until the early 1990s, the only source of wind measurements in USLM region were the rocket soundings, on which the middle atmosphere wind field climatology was based (Schmidlin, 1986). The high cost of rocket operations has fostered development of remote sensing techniques for wind profiling of the middle atmosphere. Pioneering work in the remote sensing of wind profiles up to the stratopause was conducted by Chanin et al. (1989) at Observatoire de Haute Provence (OHP, 43.9°N, 5.7°E) using incoherent Doppler Rayleigh lidar. Since then, several methods for ground-based lidar measurements of wind using molecular backscattering have been proposed and demonstrated (Bills et al., 1991; Abreu et al., 1992; Tepley et al., 1994; Rees et al., 1996; Friedman et al., 1997; Gentry et al., 2000; Yan et al., 2017). The direct-detection technique for wind profiling has been successfully realized in an airborne Doppler lidar – A2D, Aeolus Airborne Demonstrator (Reitebuch et al., 2009). A2D instrument served a prototype for the most ambitious endeavour in the context of lidar wind profiling – the first ever satellite-borne Doppler lidar instrument ALADIN (Atmospheric Laser Doppler INstrument) (ESA, 2008; Stoffelen et al., 2005), that has been successfully launched by European Space Agency (ESA) in August 2018 (Kanitz et al., 2019).

While the necessity of high resolution (<1 km) wind profiling of USLM region is well recognized (e.g. Meriwether and Gerrard, 2004; Dörnbrack et al., 2017), presently very few instruments have such capacity which are operated on a regular or quasi-regular basis. These are Doppler lidar at ALOMAR observatory in northern Norway (Baumgarten, 2010; Hildebrand et al., 2017), LiWind lidar at high-altitude Maito observatory at La Reunion island (Baray et al., 2014) and LIOvent lidar at Observatoire de Haute Provence (OHP), where the pioneering lidar measurements of wind up to 50 km altitude were conducted by Chanin et al. (1989).

The OHP wind lidar was originally designed to cover the height range of 25 – 50 km (Garnier and Chanin, 1992), i.e. where the contribution of Mie scattering by aerosol particles can be neglected in most cases. After the eruption of Pinatubo volcano in 1991 polluting the stratosphere with aerosol up to 35 km, the OHP wind lidar was redesigned to minimize the effect of Mie scattering (Souprain et al., 1999a,b). The new Rayleigh-Mie Doppler lidar named LIOvent was deployed at OHP in late 1993 and was operated on a regular basis during 1995 – 1999. The observations were used for retrieval of gravity wave parameters and stratospheric wind climatology at OHP (Souprain et al., 1999a; Hertzog et al., 2001) as well as for a study of the effect of gravity waves on ozone fluctuation in the lower stratosphere (Gibson-Wilde et al., 1997).

After a long period of sporadic operation and limited maintenance, the upgrade of OHP wind lidar was started in 2012. At the same time, a similar wind lidar instrument was deployed at Maito observatory at La Reunion island and passed a thorough performance evaluation, which will be presented in a companion paper. The upgrade of both wind lidars included replacement of the laser, optical filtering elements, and detectors. These efforts were largely motivated by the ESA Aeolus satellite mission carrying ALADIN instrument, which exploits the same principle of Doppler shift detection, i.e. double-edge Fabry-Perot interferometer (Stoffelen et al., 2005; Reitebuch, 2012).

This study aims at characterizing the performance and capacities of OHP LIOvent Doppler lidar after its upgrade. Chapter 2 describes the instrument design; Chapter 3 reports the results of LIOvent validation using 12 collocated radiosoundings;

Supprimé: A relatively recent development is the Doppler microwave radiometry capable of USLM wind profiling, although with a coarse vertical resolution of ~10 km (Rufenacht et al., 2012).

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Chapter 4 provides examples of weekly and hourly observation series; Chapter 5 presents an early case of Aeolus validation using OHP lidar; Chapter 6 concludes the study and sketches the outlook.

2 Instrument design and measurement principle

375 A detailed description of the OHP Doppler lidar (LIOvent) and the methodology for retrieving wind profiles was provided by Souprayen et al. (1999a). Here, we recap the general design of the instrument and its sub-systems after its upgrade, the measurement principle, the error budget as well as the principal stages of data retrieval.

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Supprimé: Measurement principle and i

2.1 Instrument design

380 The lidar instrument senses the horizontal wind velocity by measuring the Doppler shift between the emitted and backscattered light of the laser. The Doppler shift corresponds to the projection of the horizontal wind components onto the line-of-sight of the laser inclined 40° off-zenith. The detection of the Doppler shift is performed by means of a double-edge Fabry-Perot interferometer (FPI), as detailed in the following section.

Supprimé: the general design of the instrument and its sub-systems,

385 The general design of transmitter-receiver system is shown in Fig. 1. The transmitter of the lidar is based on a Quantaray Pro290 Q-switched, injection-seeded Nd:YAG laser emitting at 532 nm with a repetition rate of 30 Hz and 800 mJ per pulse energy. In seeded mode, the linewidth of the laser beam is less than 0.003 cm⁻¹, whereas the shot-to-shot frequency stability is better than 10 MHz. During the measurement, the laser beam is commuted successively to each of the three fixed mosaic telescope assemblies, respectively zenith (1), North (2) and East (3), using a galvanometric scanner mirror (4). Each telescope assembly has a field of view of 0.1 mrad and is comprised of a central transmitter shaft with a beam expander (5), ensuring the beam divergence of 35 µrad (full angle at half maximum) and four collecting parabolic mirrors of 500 mm diameter (6), which translates to the total collective area for each telescope of 0.78 m².

390 The backscattered light is collected by means of 200 µm multimode optical fibers located at the focal point of each mirror (1500 mm focal distance) and linked to an optical commutation chamber (7), which transfers the collected light from a given telescope to the entrance of the spectral analysis sub-system through a 600 µm fiber. The latter (not shown) comprises the FPI etalon in a thermally-stabilized pressure-controlled chamber, a 0.3 nm interference filter for reducing the sky background and 395 a mode scrambler, which serves for homogenizing the incidence angles of light projected onto the FPI. The homogeneity of the flux angular distribution is important because the transmission function of the FPI depends on the angular incidence. The scrambler module comprises two lenses, the first collecting the light from the input fiber and the second projecting its image onto the output fiber.

400 The detection of the spectrally-processed light is done with two pairs of cooled super-bialkali Hamamatsu R9880-110 photomultipliers (PMTs), receiving respectively 95% and 5% of the flux (high- and low-gain channels). The high-gain PMTs are electronically gated at 100 µs, i.e. 15 km radial distance. The acquisition is done using a four-channel Licel transient recorder featuring 32760 gates of 50 ns width (i.e. 7.5 m radial resolution).

The upgrade of the lidar (with respect to the design described by Souprayen et al. (1999a)) was carried out during 2012-2018 period and included replacement of various parts. The essential improvements that allowed extending the vertical range for wind profiling are due to the following upgrades: a higher-power laser (24W versus 10W), a new interference filter (0.3 nm vs 1 nm), and the new PMTs with faster response and lower dark current. Additionally, a new Licel transient recorder (50 ns versus 1 μ s gate bins) and a new cooling system have been introduced.

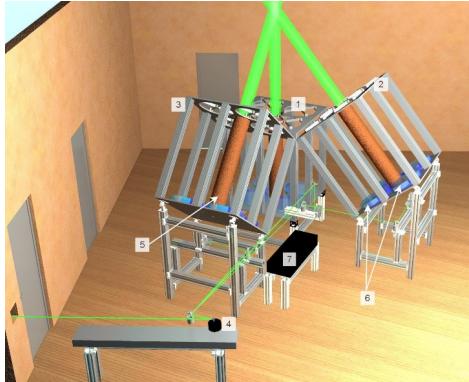


Figure 1. Three-dimensional representation of the transmitter-receiver system of LIOvent lidar comprising three telescope assemblies for zenith (1), north (2) and east (3) lines-of-sight with four collecting mirrors (6) and a beam expander (5) each. The emitted laser beam is commuted between the telescopes using rotating scanner mirror (4). The backscattered light collected by each mirror is transferred via fibers into the optical commutation chamber (7) for aggregating the fluxes from the four mirrors of each telescope assembly.

2.2 Doppler shift detection

The detection of Doppler shift is done using a double-edge Fabry-Perot interferometer (FPI). The FPI etalon (manufactured by StigmaOptique) is assembled by molecular contact and features two half-disc areas with slightly different air gaps, which results in two distinct bandpasses on the both sides of the Rayleigh-Mie backscattered line, as shown in Fig. 2. The Doppler shift of the backscattered line (shown as dashed black in Fig. 2) enhances the signal transmitted through the channel A whilst reducing that of the channel B. The interferometer is set in a sealed pressure-controlled chamber, allowing a spectral tuning of the FPI A and B bandpasses relative to the backscattered line, through variation of the air-gap refractive index. Depending on the atmospheric temperature, the backscattered Rayleigh line obtains a FWHM between about 2 and 2.4 pm, whereas the Mie line is assumed to have the same spectral width as that of the laser (<0.08 pm). The spectral spacing of

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Supprimé: corresponding to relative difference between the air-gap thicknesses of 10^{-5}
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Supprimé: on both sides of
Supprimé: Rayleigh-Mie
Supprimé: ,
Supprimé: whose
Supprimé: spectral width
Supprimé: varies
Supprimé: depending on atmospheric temperature
Déplacé (insertion) [1]
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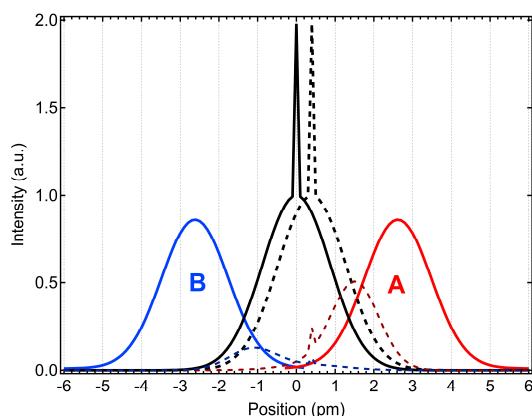
the FPI A and B bandpasses of 5.24 pm is determined by the difference in optical thickness of the respective half-disc areas of
465 the interferometer (34.5 ± 0.1 nm), whereas the FWHM of the FPI bandpasses depends on its finesse and amounts to 2.03 ± 0.01 pm according to a series of experiments by Souprayen et al., (1999a,b).

The Doppler shift response profile $R(z, \theta)$ for a given line of sight is calculated as:

$$R(z, \theta) = \frac{CN_A(z, \theta) - N_B(z, \theta)}{CN_A(z, \theta) + N_B(z, \theta)}, \quad (1)$$

where $N_A(z, \theta)$ and $N_B(z, \theta)$ are the number of photons received from altitude z and transmitted through the bandpasses

470 A and B respectively; and C is the corrective factor accounting for a possible imbalance between the signals in channels A and B due to a difference in detectors' sensitivity. The corrective factor corresponds to the ratio between the channels A and B and is obtained by comparing N_A and N_B signals from a continuous white source. The Doppler shift (in units of pm) is deduced from the response profile through the instrumental calibration function, which accounts for the temperature broadening of the Rayleigh backscattered line.▼



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495 **Figure 2. Spectral shapes of the thermally-broadened Rayleigh backscatter line with the thin Mie line on top of it (solid black) and the FPI bandpasses A and B (solid red and blue). The Doppler-shifted backscatter line (corresponding to an imaginary wind speed of 175 m/s) is shown as black dashed curve. The signals transmitted through the A and B bandpasses for the Doppler-shifted backscattering are illustrated by the dashed red and blue curves.**

Supprimé: The uncertainty in the FPI characteristics induces an uncertainty of $\pm 0.3\%$ on the horizontal wind velocity, whereas the effect of temperature uncertainties does not exceed 0.07% per Kelvin (Souprayen et al., 1999a).

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The procedure of FPI calibration is thoroughly described by Souprayen et al. (1999b). Briefly, it makes use of the pressure scanning system, which allows for spectral sampling across the two FPI bandpasses by sequentially shifting their spectral position with respect to the spectrally-stable laser line. With the constant and known spectral spacing between FPI bandpasses, one can relate the scanner motor steps to the unit of pm. This relation is then used to retrieve the FWHM of each bandpass.

which together with the FPI spectral spacing and temperature-dependent spectral width of the backscattered line yields the instrumental calibration function. The calibration function is linear in the central zone (between -0.2 and 0.2 pm) and obtains the slope of 0.755 pm^{-1} at 210 K. The uncertainty in the FPI characteristics induces an uncertainty of $\pm 0.3\%$ on the horizontal wind velocity, whereas the effect of temperature uncertainties does not exceed 0.07% per Kelvin (Souprayen et al., 1999a).

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An absolute measurement of the wind velocity requires a careful determination of the null Doppler shift reference, which is done through 1-minute zenith-pointing acquisition within each 5-minute cycle. This enables accounting for the possible drift in the emitted laser wavelength, typically of $0.03\text{--}0.08 \text{ pmh}^{-1}$. The horizontal wind components are then obtained by subtracting the line-position profile for the vertical pointing $\overline{P(z, 0^\circ)}$, from those of the tilted pointings $P(z, 40^\circ)$:

$$v_h(z) = \frac{c}{2\lambda_0 \sin 40^\circ} [P(z, 40^\circ) - \overline{P(z, 0^\circ)}], \quad (2)$$

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where λ_0 is the emitted wavelength; and $\overline{P(z, 0^\circ)}$ is the reference (null Doppler-shift) position given by an average in the altitude range of 15–25 km, a region where the vertical wind velocity is negligibly small. Expectedly, the line-position profile for the vertical pointing does not vary with altitude therefore the resulting wind profiles are insensitive to the choice of the vertical range for the null Doppler shift reference.

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The spectral tuning of the FPI bandpasses with respect to the laser wavelength is verified at the beginning of each measurement session through adjustment of the air pressure inside the FPI chamber using a stepper motor. The temperature inside the FPI housing module is maintained at 30° C at all times and we note that over many months of lidar operation, the spectral tuning remains stable, except after an occasional laser maintenance.

2.3 Signal processing and statistical uncertainty

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The measurement cadence is such that the zenith, north and east lines of sight are alternated in a cycle of 1-2-2 minutes respectively. A typical acquisition lasts 5 hours during nighttime, that is 2 h integration for each tilted pointing, which ensures signal-to-noise ratio better than 2 all the way up to about 80 km altitude a.s.l. Figure 3a shows an example of raw lidar return profile from the North pointing obtained by stitching the low- and high-gain signals. The vertical range of the useful signal spans between about 5 and 80 km. The lower boundary is due to strong returns from the lower troposphere saturating the detectors in addition to an incomplete geometrical overlap below 2 km.

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The photons received from the transient recorder are aggregated over 1-minute intervals and downsampled to 1 μs bins (150 m radial resolution). The off-line signal pre-processing includes subtraction of background due to sky light and PMT thermal noise as well as dead-time correction, after which the response profiles are calculated for each line-of-sight according to Eq. (1). Then, the Doppler shift (line-position profile) is computed using the instrument calibration function with account for atmospheric temperature profile, provided by operational analysis by European Center for Medium-range Weather Forecast (ECMWF). Finally, the zonal and meridional wind components are obtained by comparing the tilted East and North pointings to the corresponding zenith pointing using Eq. (2).

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Déplacé vers le haut [1]: The spectral configuration of FPI is conceived to minimize the effect of Mie backscattering (Souprayen et al., 1999b, see Sect. 3.1 for further detail).

Déplacé vers le bas [2]: The spectral tuning of the FPI bandpasses is verified at the beginning of each measurement session through adjustment of the air pressure (i.e. adaptation of the air-gap refractive index) inside the FPI chamber using a stepper motor. The temperature inside the FPI housing module is maintained at 30 C at all times and we note that over many months of lidar operation, the spectral tuning remains stable, except after an occasional laser maintenance.

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Supprimé: During a measurement session, the backscattered light is acquired successively from the zenith, north and east lines of sight. The Doppler shift response profile $R(z, \theta)$ for a given line of sight is calculated as ¶

$$R(z, \theta) = \frac{CN_A(z, \theta) - N_B(z, \theta)}{CN_A(z, \theta) + N_B(z, \theta)},$$

(1) ¶

where $N_A(z, \theta)$ and $N_B(z, \theta)$ are the number of photons received from altitude z and transmitted through the bandpasses A and B respectively; and C is the corrective factor accounting for a possible imbalance between the signals in channels A and B due to a difference in detectors' sensitivity. One then deduces the line-position profile (in units of pm) from the response profile through the instrumental calibration function, which accounts for atmospheric temperature profile (Souprayen et al., 1999a). T

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An absolute measurement of wind velocity requires a careful determination of the null Doppler-shift reference.

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Supprimé: 2.2 Instrument design and data processing ¶ The transmitter of the lidar is based on a Quanta-Ray Pro290 Q-switched, injection-seeded Nd:YAG laser emitting at 532 nm with a repetition rate of 30 Hz and 800 mJ per pulse energy. The general design of transmitter-receiver system is shown in Fig. 1. The laser beam is commuted successively to the three fixed telescope assemblies, respectively zenith (1), North (2) and East (3) telescopes using a galvanometric scanner mirror (4). Each telescope assembly is comprised of a central transmitter shaft with a beam expander (5) and four collecting parabolic mirrors of 500 mm diameter (6), which translates to the total collective area for each telescope of 0.78 m^2 . ¶ The backscattered light is collected by means of multimode optical fibers located at the focal point of each mirror and linked to an optical commutation chamber (7), which transfers the ... [1]

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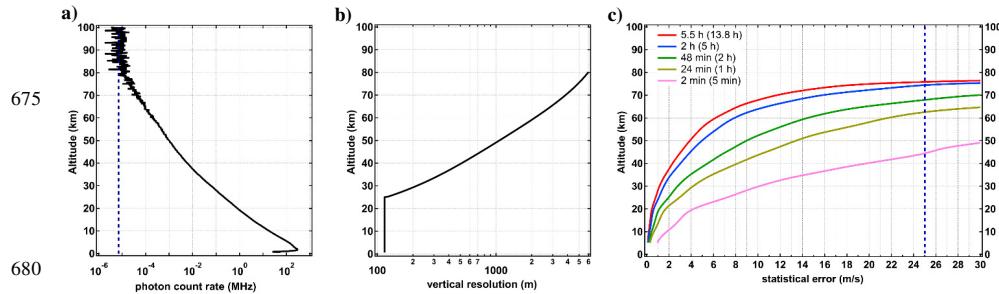
Statistical error due to Poisson noise (shot error) increases with altitude proportionally to the exponential decay of molecular backscatter. As the error scales with $1/\sqrt{\Delta z}$ (where Δz is the vertical resolution), we use a height-dependent Δz , which is set to 115 m (150 m radial) below 25 km and then increased quasi-exponentially with altitude, from 500 m at 40 km

660 to 4000 m at 70 km (Fig. 3b). For a given vertical resolution, we compute the statistical error of an individual response profile:

$$\sigma_R = 2C \frac{\sqrt{N_A(N_B-Fc_B)^2+N_B(N_A-Fc_A)^2}}{[C(N_A-Fc_A)+(N_B-Fc_B)]^2}, \quad (3)$$

where Fc_A and Fc_B is the background signals in channels A and B and C is the corrective factor introduced in the previous section.

665 Figure 3c shows the altitude profiles of statistical error for different integration times. For a typical lidar acquisition lasting 5 hours (i.e. 2 hours of a given tilted pointing acquisition, blue curve), the statistical error is less than 2 m/s below 33 km and does not exceed 6 m/s throughout the stratosphere. In the mesosphere, the error increases from 6 m/s at 55 km to 16 m/s at 70 km. A longer acquisition (13.8 h, red curve) reduces the error yet does not extend the vertical coverage: at 75 km altitude, the statistical error for 5 h and 13.8 h acquisition are nearly the same. We use the statistical error value of 25 m/s as a threshold for cut-off altitude of retrieved wind profiles. Given such threshold, the top of vertical range for a standard (5 h) lidar 670 acquisition is \sim 75 km, whereas a 5-minute acquisition (pink curve), corresponding to a single north-east-zenith measurement cycle, enables coverage up to about 44 km.



675 Figure 3 (a): nightly-average raw signal vertical profile obtained in a 7-hour lidar acquisition on 28.01.2019. Dashed blue curve indicates the background noise level. (b) altitude-variable vertical resolution used in the retrieval. (c) statistical error profiles 680 computed for different acquisition times of a tilted (north) pointing. The values in brackets correspond to the total duration of lidar acquisition, including zenith, north and east pointings in cycle of 1-2-2 minutes. The dashed blue line marks the cut-off error threshold.

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3 Comparison with collocated radiosoundings

Over the 4-yr period, spanning June 2015 to June 2019, the validation of the LIOvent wind lidar was conducted using 12 radiosonde (RS) ascents performed at OHP during the time of lidar acquisition. We used Meteomodem M10 radiosondes 695 equipped with GNSS receiver, launched under TOTEX 1200 gr weather balloons. The balloons were reaching on average 29.9 km altitude, whereas the horizontal drift during ascent did not exceed 90 km from the launch point.

Figure 4 displays the altitude-coded trajectories of the 12 radiosonde ascents as well as the ground projections of LIOvent tilted pointings. The horizontal displacement of the radiosondes with respect to the lidar sampling location at every 700 altitude level was calculated separately for the East and North pointings and amounted respectively to 27 ± 19.5 km and 39 ± 25 km (1σ), the largest being 117 km for the North pointing. Generally, the displacement increased with altitude as the balloons were drifting away from the lidar sampling locations, as Fig. 4 suggests.

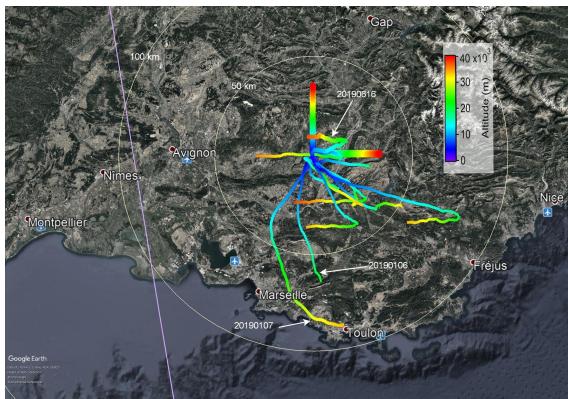


Figure 4. OHP wind lidar sampling location along north and east lines-of-sight (thick lines) and trajectories of 12 radiosondes 715 launched from OHP for wind lidar validation (thinner curves) color-coded by the altitude a.s.l. Grey circles indicate the distance from OHP. Particular radiosonde flights are tagged by white arrows with indication of the flight date. The magenta line shows the ground track of Aeolus lidar (see Sect. 5).

For setting up the intercomparison, lidar measurements were averaged over the time period of radiosonde ascent (110 720 minutes from the ground to 33 km altitude at 5 m/s ascent rate), whereas the radiosonde measurements, reported at 1 Hz frequency, were downsampled to match the vertical resolution of lidar profiles (115 m – 320 m depending on the altitude). The results of intercomparison are reported in Table 1 as absolute difference between RS and LIOvent wind profiles, standard deviation of the differences and correlation for each particular sounding. The intercomparison exercise is done separately for the zonal and meridional wind components as well as for the total wind speed and wind direction.

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The mean differences obtained from the individual comparison cases varies between -1.3 and 0.9 m/s for the zonal wind, and between -2 and 0.9 m/s for the meridional wind. For the total wind and direction, the differences vary between -1.1 and 0.7 m/s and between -4.9 and 9.6 degrees respectively. The averages of all intercomparison cases amount to +0.1 m/s and -0.1 m/s respectively for the zonal and meridional components, 0.0 m/s for the total wind and 0.3 degrees for the wind direction.

Figure 5 (top panels) shows altitude profiles of the absolute difference between LIOvent and RS for each sounding as well as the median profile. The point-by-point differences rarely exceed 5 m/s, whereas the median values never exceed 2 m/s. While the median difference profiles do not indicate any altitude-dependent bias, the scatter of differences appears to increase with altitude. This is due, on one hand, to larger horizontal offset between measurements at higher altitudes and, on the other hand, due to increase of statistical error with altitude shown as dashed curves in the a) and b) panels of Fig. 5.

The bottom panels of Fig. 5 display the scatter plots of the wind velocities measured by the lidar and the radiosondes with associated regressions and correlation coefficients. For both wind components, the slope of the regression line is close to 1, which affirms the credibility of the FPI calibration function relating the Doppler shift response to wind velocity. The Pearson correlation coefficient r deduced from the ensemble of collocated measurements amounted to 0.97 and 0.96 respectively for zonal and meridional wind velocities, 0.97 for the total wind and 0.89 for the wind direction.

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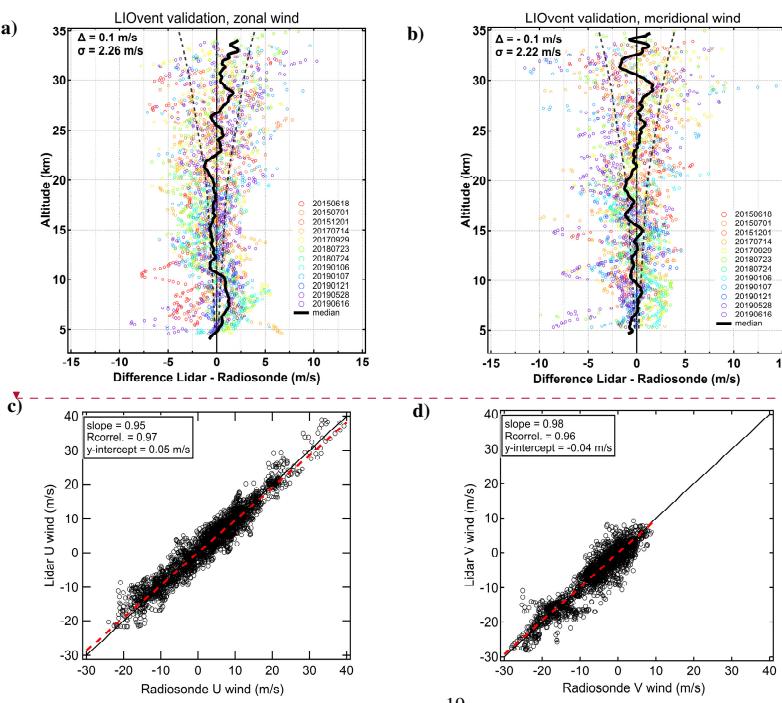


Figure 5. Summary of LIOvent lidar validation using radiosonde ascents from OHP. Top panels: absolute difference between

LIOvent and radiosonde zonal (a) and meridional (b) wind velocity for each sounding (date-colored circles, dates provided in the

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785 legend) and median profile (black curve). The dashed line indicate the statistical uncertainty estimated for a 2-hour lidar acquisition.

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The mean difference Δ and the mean standard deviation of the difference (σ) are indicated in the top-left corner of a) and b) panels.

Bottom panels: scatter plots of the zonal (c) and meridional (d) wind velocities measured by the lidar and the radiosondes. The 1:1 line is shown in solid black, the linear regression line is shown in dashed red.

790 The mean standard deviation of the differences for the 12 collocated soundings amounts to 2.26 m/s for the zonal and 2.22 m/s for the meridional wind profiles. These values are consistent with the estimated shot error for a 2 hours lidar acquisition (i.e. duration of a radiosounding), which increases from 0.2 m/s to 3.4 m/s in the altitude range of lidar-radiosonde

795 intercomparison, as can be inferred from Fig. 5a,b. For evaluating the effect of the horizontal offset between the lidar and RS measurements we computed the offset-weighted averages of the intercomparison statistics and compared them with the ordinary averages. The weight for each individual value is defined as $w = 1 - \bar{D}/D_{max}$, where \bar{D} is the mean distance between the lidar and RS sampling locations and D_{max} is the maximum distance amounting to 69 km (Table 1). We note that the horizontal-offset weighting of the differences neither affects the mean difference nor the mean correlation but reduces the standard deviation for the wind components and total wind by about 0.2 m/s. This suggests that horizontal variability of the wind field on a scale of few tens of kilometres is small.

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800 Figure 6 presents examples of summer and winter zonal wind intercomparison cases, both showing directional wind shear in the lower stratosphere but of opposite sign. Remarkably, the small-scale fluctuations, presumably caused by gravity waves propagation and/or breaking, are reproduced by the lidar just as accurately as measured by the balloon sonde, carried by those winds.

805 At higher altitudes, the fine-scale fluctuations resolved by the lidar appear at times out of phase with those seen by the radiosonde. This is more prominent in the Summer case (Fig. 6a) despite the closer collocation of the measurements. In this case, the RS ascent closely follows the lidar's line-of-sight up until 19 km (cf. Fig. 4) while the LIOvent zonal profile precisely tracks the one of RS up to about the same level. At 19 km, the zonal wind reverses, the balloon makes a U-turn and progressively drifts westward and away from the lidar. Above 30 km, the RS and LIOvent profiles start to get out of phase whilst both showing an increasing easterly wind between 30 – 35 km. The lidar profile, extending above the top of 810 radiosounding, reveals a typical signature of a gravity wave, supposedly propagating in the zonal direction (considering a relatively unperturbed meridional wind profile in this layer, not shown).

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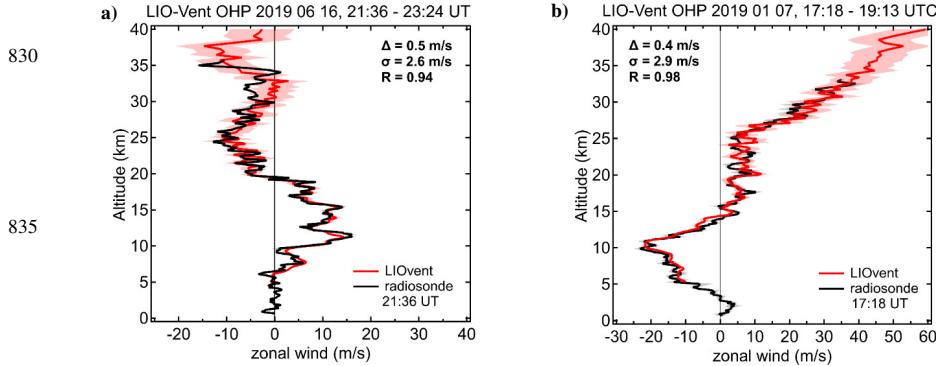
815 While the statistical error of the lidar measurement becomes comparable to the observed variations at these levels, the dephasing of LIOvent and RS profiles in Fig. 6a is likely due to spatially-offset sampling of the gravity wave front. Interestingly, the dephasing above 30 km is less obvious in the winter case (Fig. 6b) despite a considerably larger spatial offset, compared to the summer case (88 km vs 36 km). This may be explained in consideration of the much stronger zonal wind in the winter case (38 m/s versus -2 m/s), damping the amplitude of the wave-induced perturbations.

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840 **Figure 6.** Selected cases of the lidar-radiosonde intercomparison of the zonal wind profiles in June (a) and January (b) 2019. The lidar measurement dates and times are given above the panels, the time of radiosonde launch on the same date is provided in the legend.

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3.1 Sensitivity to Mie scattering

845 Although the Mie-backscattered line is narrow (0.08 pm) compared to the thermally-broadened Rayleigh line (2 – 2.4 pm) the intensity of the former may be substantially higher and thereby alter the spectral shape of the return signal. In this case, a disproportionately larger flux would be transmitted through one of the FPI bandpasses, affecting its calibration function and introducing a bias into the wind retrieval within the particle layer. The sensitivity to Mie scattering can be reduced by increasing the FPI spectral spacing, however this also reduces the sensitivity to the Doppler shift. The optimal spectral 850 configuration of the FPI has been established on the base of a theoretical model carried out by Soupravayen et al. (1999b). They found that for observable stratospheric wind velocities, the residual Mie-induced error is less than 1 m/s for the scattering ratio $R=10$, which is characteristic of a cirrus cloud readily visible to an unaided eye.

855 In this study, we have experimentally revisited the aspect of FPI sensitivity to particulate scattering. The eruption of the Raikoke stratovolcano (22 June 2019, Kuril Islands, 48° N, 153° E) has polluted the lower stratosphere with a large amount of sulfuric aerosol (NASA Earth Observatory, 2019). The aerosol plumes were observed by OHP lidars every night since 10 July 2019 (and at the time of writing) between 12 and 20 km altitude, which provided an opportunity for testing the sensitivity of wind lidar to Mie-scattering in the stratosphere.

860 Figure 7 displays lidar measurements of aerosol scattering ratio (SR) and zonal wind velocity on 20 July. The SR profiles were obtained from an aerosol channel (532 nm) of LTA (Lidar Temperature Aerosol) instrument (Keckhut et al., 1993) and from the zenith acquisition of LIOvent lidar using aerosol retrieval method described by Khaykin et al. (2017 and

Supprimé: A thorough study on the effect of Mie scattering on the wind measurement using the double-edge Fabry-Perot interferometer of the OHP wind lidar was carried out by Soupravayen et al. (1999b).

Supprimé: Using a theoretical model for the FPI, they recovered the true spectral properties of the interferometer and estimated the particle-induced error for varying values of scattering ratio.

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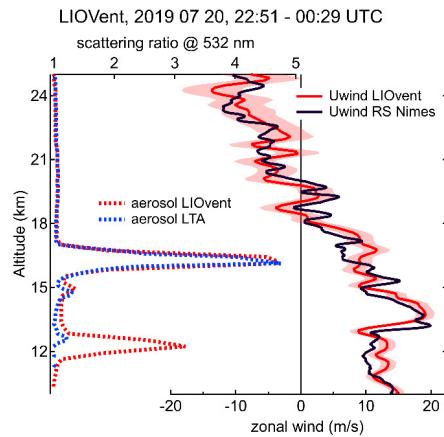
references therein). The LIOvent operation was started after the end of LTA operation since the lidars share the same laser and cannot be operated simultaneously. Both lidars consistently show an aerosol layer at 16.2 km altitude with SR reaching 4.7 and an estimated optical depth of 0.03 which is comparable to a thin cirrus cloud (Hoareau et al., 2013). In addition, the LIOvent measurement reveals a cirrus cloud at 12.2 km, which occurred only towards the end of LTA acquisition and thereby left a weaker imprint in the average SR profile of LTA.

The LIOvent wind measurement in the presence of ice crystals and volcanic aerosol is compared in Fig. 7 to a time-
875 collocated radiosoundings conducted from Nimes airport, situated ~100 km west from OHP (cf. Fig. 4). While the vertical
880 structures in the LIOvent and RS wind profiles are at times out of phase (which may be explained by spatial variability), the
lidar profile does not show any indications of Mie-induced bias, neither due to a thin cirrus cloud nor due to a volcanic aerosol
layer. Such a bias would appear as a sharp feature in the wind profile coinciding the with the SR enhancement, which is
obviously not the case here. This result confirms that the spectral configuration of the FPI allows accurate wind measurements
in the presence of particles in the middle atmosphere.

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900 **Figure 7.** Zonal wind profiles (bottom axis) measured by the LIOvent wind lidar (red solid) and by MeteoFrance routine
radiosounding launched from Nimes airport (100 km away from OHP, see Fig. 4) at 00 UT on 21.07.2019 (black solid). The date and
time of LIOvent measurement is shown at the panel top. The aerosol scattering ratio profiles obtained using LIOvent zenith pointing
(red dashed) and LTA lidar (blue dashed), showing volcanic aerosol layer at 16.2 km are plotted versus top axis. See text for details.

915 **4 Observations**

During the 2015-2019 period, the LIOvent instrument was operated on 52 nights, mostly during early summer and winter seasons. This section reports examples of successive nightly-mean profiles reflecting the wind variability in the middle atmosphere during opposite seasons as well as a particular case of temporally-resolved wind profiling.

4.1 January 2019 series

920 An interesting dynamical event in the USLM was observed in January 2019 during an intensive measurement campaign dedicated to Aeolus validation (AboVE-OHP, Aeolus Validation Experiment at OHP). A strong perturbation of the Arctic circumpolar vortex has occurred as a result of a major Sudden Stratospheric Warming event during the first week of January 2019. According to potential vorticity maps (not shown) based on ECMWF Integrated Forecast System (IFS), the vortex started to split around 1st January and evolved into two separate vortices above Europe and Canada by 4th January. The European 925 counterpart was displaced southward and its edge - at 850 K potential temperature level (~50 km) - reached OHP around 6th January, that is when the AboVE-OHP measurement campaign was started.

Figure 8 shows ensembles of the zonal and meridional wind profiles obtained during 6 - 9 January period. The plots include indications of the stratopause level which was progressively descending from 47 to 43 km during that period, as inferred from simultaneous temperature profiling using LiO3S ([Lidar Q3 stratosphérique](#)) differential absorption lidar (Godin-930 Beekmann et al., 2003; Wing et al., 2018). The 6th January wind profiles (red curves) reflect the perturbed conditions when the vortex edge was located above OHP and both zonal and meridional components were maximizing at 80 m/s around the stratopause. As the edge of vortex was moving back north of OHP during the following days, the measurements show weakening winds throughout the USLM and reversal of both wind components in the lower mesosphere by 9th January. The rapidly weakening winds form an envelope of profiles with a bottom at ~27 km for the zonal wind and ~38 km for the 935 meridional component. Below this envelope, neither of the wind components show significant change over the 4-day period.

The observed evolution of wind profiles is reproduced by the ECMWF T1279/L137 operational analysis represented by cross-circles in Fig 8. The wind change envelope and the vertical structure of the wind profiles are both well resolved by the model. The ECMWF profiles reproduce the observed vertical fluctuations on a scale of a few kilometres up to the stratopause, which is remarkable since the regular radiosoundings assimilated into the model hardly reach 30 km altitude. We 940 note that the consistency between ECMWF and LIOvent is better for the zonal wind, whereas the vertical structures in meridional wind are somewhat less consistent with the observations in the USLM. Analogous results, inferred from intercomparison between ALOMAR wind lidar and ECMWF forecast winds, were reported by Rüfenacht et al. (2018). The damping and dephasing of the vertical structures by ECMWF becomes more prominent above about 50 km, which might owe to both the temporal averaging over 5-13 hours by the lidar and the coarse model resolution in the mesosphere. A detailed 945 comparison between wind lidar observations, ECMWF IFS and reanalysis data will be the subject of a separate study.

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4.2 May 2019 series

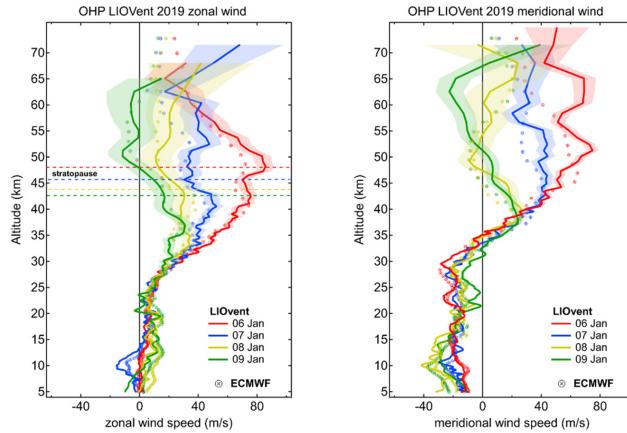
950 Figure 9 shows a sequence of wind profiles acquired during late May 2019. While the winds appear highly variable in
 the upper troposphere due to the dynamics of the jet stream, the middle stratosphere remains relatively calm and stable over
 the considered 4-day period. The zonal wind reverses at around 36 km and the easterlies pick up until \sim 65 km, that is when
 the wind shear reverses. The meridional wind is very weak throughout the stratosphere except for a small envelope at 30 km,
 which accompanies the zonal wind perturbations in this layer. The ECMWF IFS accurately resolves this envelope, however
 955 the observed vertical structures above, in the USLM, are not reproduced by the model. In particular, the reversal of wind shear
 at around 65 km in both wind components is missed by the model, whereas the lidar profiles consistently report this feature,
 significant at the permitted statistical error of 25 m/s.

The imposed error threshold of 25 m/s determines the cut-off altitude of the wind profiles reported in Figures 8 and 9.
 The top altitude varies between 65 and 75 km depending on the presence of cirrus clouds inhibiting the return signal from
 960 above. We note that the meridional profiles normally reach higher altitudes, which is due to a better condition of the collecting
 mirrors of the north-pointing telescope assembly.

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 Figure 8. Ensembles of nightly-mean vertical profiles of zonal (left) and meridional (right) wind profiles obtained by LIOvent
 in January 2019 (solid curves) with statistical uncertainty shown as shading and the corresponding ECMWF Integrated Forecast
 System profiles (cross-circles). Horizontal dashed lines in the left panel indicate the stratopause altitude obtained from simultaneous
 temperature lidar measurements.

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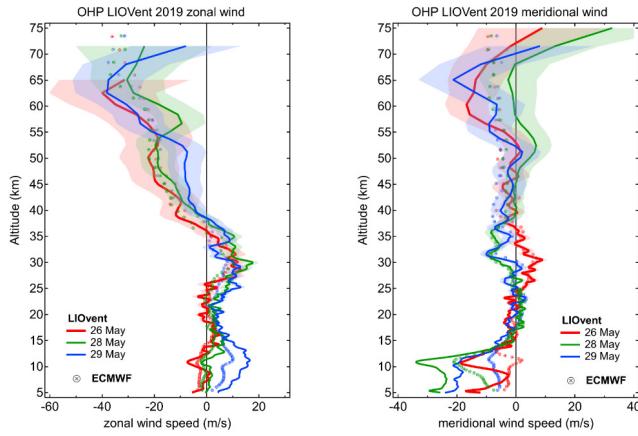


Figure 9. Same as Fig. 8 but for May 2019

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4.3 Time-resolved wind profiling

An important advantage of the Doppler lidar technique is the capacity to provide temporally-resolved vertical profiling of the atmosphere, which enables characterization of high-frequency fluctuations in the wind profile, inaccessible with snapshot-like radiosonde measurements. Figure 10a provides an example of meridional wind profile variation over the course of a continuous whole-night lidar acquisition lasting nearly 14 hours. Superimposed onto the lidar time-altitude section is a radiosonde ascent, plotted using the same color map as the lidar wind curtain.

The LIOvent and RS profiles are remarkably consistent as can be seen from the color map (correlation coefficient amounts to 0.99 in this case). With that, the lidar wind curtain shows important variation of the wind velocity over the course of 14-hour acquisition. The peak-to-peak variation at any level below 30 km altitude is between 10 and 20 m/s, increasing to ~30 m/s towards 40 km. The wind change rate in any 3 km thick layer is reaching 10 m/s per hour, which points to the predominance of temporal variability of the wind field over its spatial variability. Indeed, the maximum deviation of the lidar profile from the RS one in this case did not exceed 4 m/s at any given level, all the while that the RS measurements were taken as far as 71 km away from the lidar sampling location (cf. Fig. 4).

In the upper-middle stratosphere (i.e. around 35 km), where the meridional wind reverses, one can discern wind patterns slowly propagating downward. This pattern is a typical signature of upward-propagating gravity waves with a non-zero ground-based phase speed. A somewhat different pattern is observed in the lower-middle stratosphere (15 – 30 km), where the vertical structures appear to remain constant with altitude. Figure 10b provides a deeper insight into the variable structure of

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1025 this layer by showing a sequence of 6 wind profiles, obtained by integrating over successive 135-minute temporal intervals as well as the corresponding RS profile. One can see three layers of stronger southward wind at around 17 km, 23 km and 30 km altitude interleaved by two layers of weaker wind at around 20 and 26 km.

1030 The persistence of the observed structures in both temporal and vertical dimensions suggests the occurrence of stationary gravity waves, most likely generated by the flow over the Alpine mountains. Indeed, the circulation in the lower troposphere at that time (not shown) was such that the OHP site was downwind of the Alps. The stationary gravity waves, generated by the flow over the mountain range, could propagate freely into the stratosphere because of the absence of directional wind shear all the way up to 35 km. The amplitude of these wave-induced perturbations appears to increase with altitude from ~ 5 m/s to ~ 10 m/s, which is expected from the linear theory of atmospheric waves.

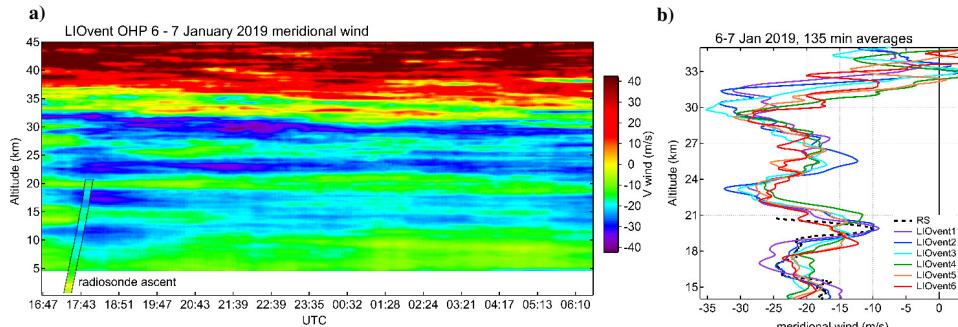
1035 The orographic nature of the gravity wave, identified using time-resolved lidar measurements, can be verified in consideration of the vertical wavelength. For a stationary wave, the vertical wavelength λ_z can be deduced from the horizontal wind speed v_h and the buoyancy frequency N :

$$\lambda_z = 2\pi \frac{v_h}{N} \quad (4)$$

Given the observed $v_h \sim 20$ m/s and $N \sim 0.02$ s $^{-1}$, we obtain the vertical wavelength of ~ 6.5 km, which corresponds well to what can be deduced directly from the wind profiles in Fig. 10.

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1045 Figure 10. (a) Temporal variation of the meridional wind profile over the course of a continuous whole-night LIOvent acquisition started on 6 January 2019. Superimposed onto the lidar time-altitude section is the corresponding radiosonde ascent from OHP, plotted using the same color map as the lidar wind curtain. (b) successive 135-minute averages of meridional wind measured by the lidar (solid curves) and the radiosonde profile (black dashed).

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5 First results of Aeolus validation

Aeolus is the ESA's satellite mission designed to measure wind and aerosol profiles in the troposphere and lower stratosphere on a global scale (Stoffelen et al., 2005; [ESA, 2008](#)). Launched on 22 August 2018, the Aeolus satellite carries the Atmospheric LAser Doppler INstrument (ALADIN) which features a telescope of 1.5 m diameter and a laser emitting at 355 nm with a repetition rate of 50 Hz and ~65 mJ per pulse energy. ALADIN instrument has two detection channels for measuring Doppler shift using the molecular (Rayleigh) and particulate (Mie) backscattering. The Rayleigh channel makes use of a double edge Fabry-Perot interferometer, that is the measurement principle exploited by the OHP wind lidar.

The ALADIN telescope is pointed 35° away from the orbital plane in order to sense the backscattered light perpendicular to the trajectory of the satellite. This enables measuring the so-called horizontal line-of-sight (HLOS) wind velocity, which is close to the zonal wind component except at high latitudes. The Aeolus satellite has a Sun-synchronous dusk/dawn orbit with a 7-day repeat cycle, passing near the OHP station (within 100 km) twice per week along the successive ascending and descending orbits, at around 17:50 and 5:50 UT respectively.

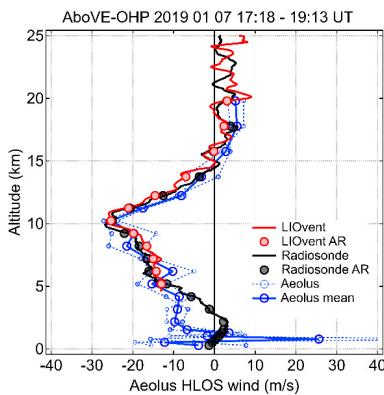
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As Aeolus and the OHP wind lidar exploit the same measurement technique, the LIOvent instrument is an important contributor to Aeolus Cal/Val activity. Since January 2019 and by the time of writing, LIOvent has been operated on 27 nights, providing 8 measurements collocated with Aeolus overpasses. Some of the Aeolus-collocated LIOvent acquisitions were accompanied by simultaneous radiosonde ascents. While a comprehensive validation exercise will be the subject of a separate study, here we provide an example of comparison between collocated Aeolus [level 2B](#) 'Rayleigh-clear' and LIOvent wind profiles. One should bear in mind that Aeolus wind data processing is still being improved by optimizing the in-orbit instrument calibration, therefore the presented validation case is to be considered as preliminary.

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Figure 1 displays a collocation case from 7 January 2019 when the satellite was sampling the atmosphere around 100 km west of OHP along the ascending orbit. The plot includes two successive Aeolus wind profiles (blue dashed) obtained by 12 seconds integration (i.e. 90 km along-track distance) as well as their mean (blue solid). The LIOvent and RS wind components are converted to Aeolus HLOS wind and reported at their native vertical resolution (solid red and black) as well as after downsampling to Aeolus vertical resolution (circles).

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1095 **Figure 11. Example of the validation of preliminary L2B HLOS wind of the ESA Aeolus ALADIN lidar using the LIOvent lidar and a radiosounding within AboVE-OHP campaign. The LIOvent and RS wind components are converted to Aeolus HLOS wind and reported at their native vertical resolution (solid red and black) as well as after downsampling to Aeolus vertical resolution (circles, marked AR in the legend). The Aeolus overpass near OHP took place on 7 January 2019 at 17:50 UT (ascending orbit). The mean distance between LIOvent and ALADIN sampling locations is 106 km. The lidar acquisition time (corresponding to the radiosonde ascent time) is provided in the panel top. See text for further details.**

1100 The profiles are found to be in good agreement, consistently reproducing the peak in the eastward wind of -25 m/s at 10 km, which corresponds to an anticyclonic feature of the jet stream (not shown). In the middle troposphere, successive Aeolus profiles appear somewhat scattered around their mean, with the latter being in better agreement with the ground-based measurements. Below 5 km, the Aeolus profile deviates from the RS, which may be caused by a stronger spatial variability of 1105 the wind field in the lower troposphere (note that the minimum horizontal distance between the RS and Aeolus measurements is 91 km). In the lower stratosphere, that is above about 11 km, Aeolus follows well the downsampled measurements by OHP lidar and radiosonde. The average difference between the mean Aeolus and downsampled LIOvent HLOS wind profiles in the overlapping range of 5 - 20 km amounts in this case to +1.5 m/s with a standard deviation of 3.2 m/s and a correlation coefficient of 0.96. Similar values were obtained from other collocations during AboVE-OHP campaign in January 2019.

1110 6 Summary and outlook

The OHP wind lidar presented here was a unique instrument at the time of its creation and remains one of the very few instruments capable of wind profiling in the middle atmosphere with vertical resolution up to 115 m and temporal resolution up to 5 minutes. In this paper, we have described the design of the instrument after its upgrade and evaluated its capacities using a dozen time-collocated radiosoundings launched from OHP. We have shown that the lidar is capable of measuring 1115 horizontal wind velocity between 5 and 75 km altitude with a random error of less than 6 m/s up to the top of the stratosphere. We note that the vertical range can potentially be extended to 3 – 80 km through replacement of the beam-commuting and beam-splitting mirrors, for which the resources are available.

A noticeable result of the lidar-radiosonde intercomparison is a remarkably small average bias of ± 0.1 m/s for the both wind components and 0.3 degrees for the wind direction. This finding affirms the reliability of the on-the-run calibration 1120 (through periodical zenith pointing) as well as the stability of the FPI calibration function. Also remarkable is that the small-scale wind fluctuations are reproduced by the lidar just as accurately as measured by the balloon sondes, carried by those winds. The average standard deviation of the differences for the total horizontal wind was found to be only ~ 2.3 m/s, which is consistent with the error estimates for the considered altitude range. The correlation coefficient obtained from the ensemble of collocated measurements amounted to 0.97 for the total wind speed and to 0.89 for the wind direction.

1125 We have shown that wind profiling with the LIOvent lidar has little or no sensitivity to the presence of particles (thin cirrus clouds or stratospheric aerosols) and we have demonstrated the capacity of the wind lidar to measure vertical profiles of

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aerosol backscattering. In addition, using the 3 different lines-of-sight, one can obtain information on the fine-scale horizontal
1140 variability of stratospheric aerosol.

The examples of successive nightly-mean wind profiles given in Sect. 4.1 and 4.2 provide interesting example of the
wind variability in the upper stratosphere and lower mesosphere, the atmospheric layer of exceptionally poor observational
coverage. The observed vertical structures and the rapidly changing wind shear reflect the complex dynamics of the USLM
layer and its two-way interactions with the upward-propagating gravity waves, whose manifestation may be not well
1145 reproduced by atmospheric models. We note though that ECMWF operational model tends to reproduce, in most cases, the
observed vertical structures at least up to 50 km altitude.

The example of time-resolved wind profiling presented in Sect. 4.3 highlights the capacity of LIOvent instrument to
detect high-frequency fluctuations in the wind profile, indicative of various types of gravity waves. This rare capacity enables
a comprehensive characterization of the high-frequency part of the wave spectrum, inaccessible with any other measurement
1150 technique. At OHP, the wind lidar acquisitions are typically accompanied by temperature profiling using Rayleigh lidar, which
altogether provides a complete suite of thermodynamical parameters in the middle atmosphere on a regular basis and on a long
term.

Using the time-resolved wind profiling and simultaneous radiosoundings, we have found that the temporal variability
of wind profile in the free atmosphere at a scale of 1 hour may be at least twice as large as the spatial variability on a scale of
1155 50-100 km, as deduced from the lidar-radiosonde intercomparison. This finding is to be considered for Aeolus wind validation
activities in a sense that a precise temporal collocation may be more important than the spatial collocation of the wind
measurements.

We have presented the first preliminary case of Aeolus validation using the LIOvent lidar. We note that while the
Aeolus data processing and calibration may be subject to further improvement, the first results of intercomparison between the
1160 ground-based and space-borne Doppler lidars are encouraging. The validation of Aeolus is to be continued at OHP on a regular
basis for monitoring of the long-term stability of the satellite lidar, whereas a dedicated Aeolus validation study will be
conducted in a separate article.

Further studies exploiting LIOvent observations will address the characteristics of gravity waves retrieved from
simultaneous wind and temperature profiling at OHP as well as intercomparison with operational analysis and new reanalysis
1165 data sets such as ECMWF ERA5. The lidar wind profiling is also to be used in conjunction with infrasound measurements
carried out at OHP (Le Pichon et al., 2015) for studies of middle atmosphere dynamics.

Data availability. The wind lidar data will be soon made available through the AERIS data portal at <https://cds-espri.ipsl.upmc.fr/NDACC/station?methodName=viewDataOhp>

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Authors contributions. SK and AH conceived the study and conducted the LIOvent measurements, RW conducted the
LIOvent measurements and extracted Aeolus data, PK offered scientific insight, SGB offered scientific insight and provided

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request.

1180 access to LiO3S lidar data, JP conceived the optical design of LIOvent instrument, JFM and JS upgraded and optimized the
wind lidar. The paper is written by SK with contributions from all co-authors.

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(DLR) for providing high-resolution PV polar projections based on ECMWF IFS data. The work related to Aeolus validation
has been performed in the frame of Aeolus Scientific Calibration & Validation Team (ACVT) activities. Results are based on
1190 preliminary (not fully calibrated/validated) Aeolus data, that have not yet been publically released. Further data quality
improvements, including in particular a significant product bias reduction, will be achieved before the public data release.
Aeolus is a European Space Agency (ESA) mission from the Earth Explorer Program of research satellites. We thank Anne
Grete Straume (Aeolus mission scientist) and Jonas von Bismark (Aeolus data quality manager) for the fruitful exchange on
this study,

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Date	Mean difference, m/s or degrees (φ)				Stand. deviation m/s or degrees (φ)				Correlation coefficient r_s				Distance, km		Top of RS, km ₂₀₀
	U	V	V _h	φ	U	V	V _h	φ	U	V	V _h	φ	U	V	
2015 06 18	-1.3	0.1	-0.9	2.9	2.9	2.3	2.6	16.8	0.95	0.96	0.97	0.93	31	47	32.5
2015 07 01	0.9	-0.8	-0.1	-4.1	2.1	1.2	2.4	13.5	0.92	0.89	0.96	0.95	16	21	28.0
2015 12 01	-0.4	0.6	-0.2	-1.4	1.8	2.0	1.9	18.9	0.96	0.96	0.98	0.92	15	27	33.0
2017 07 14	0.1	-0.3	0.5	1.9	2.6	3.1	2.5	20.3	0.98	0.79	0.99	0.96	46	69	33.0205
2017 09 29	-0.0	0.65	0.3	-4.6	2.7	2.7	2.5	27.8	0.93	0.85	0.96	0.75	18	38	33.0
2018 07 23	0.29	0.1	-0.3	0.9	3.2	2.9	2.5	19.4	0.97	0.88	0.98	0.94	21	38	37.3
2018 07 24	0.2	0.0	0.5	-2.0	1.6	1.3	1.5	14.6	0.99	0.89	0.99	0.97	8	22	20.1
2019 01 06	-0.1	0.9	-1.1	0.5	2.5	2.1	3.4	10.7	0.90	0.99	0.98	0.91	37	45	20.9
2019 01 07	0.4	-0.7	0.7	-0.0	2.9	3.8	3.0	8.4	0.98	0.98	0.99	0.98	57	67	33.5
2019 01 21	0.1	0.0	0.2	-1.7	0.9	1.5	1.2	6.7	0.99	0.98	0.99	0.99	13	20	18.0
2019 05 28	-0.2	-2.0	0.4	9.6	2.7	3.6	2.6	24.8	0.94	0.89	0.97	0.71	27	49	32.5
2019 06 16	0.6	-0.5	0.3	1.3	2.5	2.5	2.0	20.9	0.94	0.72	0.96	0.73	12	15	37
Average	0.1	-0.1	0.0	0.3	2.3	2.2	2.3	16.9	0.97	0.96	0.97	0.89	27	35	29.9 ₁₂₁₅

Table 1. Summary of intercomparison between LIOevent lidar and time-collocated radiosoundings launched at OHP. The results are shown separately for zonal (U) and meridional (V) wind measurements as well as for the total wind speed (V_h) and wind direction (φ). Provided for each case of intercomparison (from left to right) are: measurement date, mean absolute difference, standard deviation of the differences, Pearson's correlation coefficient r_s , mean horizontal distance between the lidar and radiosonde sampling locations and top altitude of radiosounding. The average values are provided in the bottom row.

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2.2 Instrument design and data processing

The transmitter of the lidar is based on a Quanta-Ray Pro290 Q-switched, injection-seeded Nd:YAG laser emitting at 532 nm with a repetition rate of 30 Hz and 800 mJ per pulse energy. The general design of transmitter-receiver system is shown in Fig. 1. The laser beam is commuted successively to the three fixed telescope assemblies, respectively zenith (1), North (2) and East (3) telescopes using a galvanometric scanner mirror (4). Each telescope assembly is comprised of a central transmitter shaft with a beam expander (5) and four collecting parabolic mirrors of 500 mm diameter (6), which translates to the total collective area for each telescope of 0.78 m².

The backscattered light is collected by means of multimode optical fibers located at the focal point of each mirror and linked to an optical commutation chamber (7), which transfers the collected light from a given telescope to the entrance of the spectral analysis sub-system. The latter (not shown) comprises a mode scrambler for homogenizing the incidence angles of light projected onto FPI, a 0.3 nm interference filter for reducing the sky background, and the FPI in a thermally-stabilized pressure-controlled chamber.

The detection of the spectrally-processed light is done with two pairs of cooled Hamamatsu R9880-110 photomultipliers (PMTs), receiving respectively 95% and 5% of the flux (high- and low-gain channels). The high-gain PMTs are electronically gated at 100 μ s, i.e. 15 km radial distance. The acquisition is done using a four-channel Licel transient recorder featuring 32760 gates of 50 ns width (i.e. 7.5 m radial resolution).

The measurement cadence is such that the zenith, north and east lines of sight are alternated in a cycle of 1-2-2 minutes respectively. A typical acquisition lasts 5 hours during nighttime, that is 2 h integration for each tilted pointing, which ensures signal-to-noise ratio better than 2 all the way up to about 80 km altitude a.s.l. Figure 2a shows an example of raw lidar return profile from the North pointing obtained by stitching the low- and high-gain signals. The vertical range of the useful signal spans between about 5 and 80 km. The lower boundary is due to strong returns from the lower troposphere saturating the detectors in addition to an incomplete geometrical overlap below 2 km.

The photons received from the transient recorder are aggregated over 1-minute intervals and downsampled to 1 μ s bins (150 m radial resolution). The off-line signal pre-processing includes subtraction of background due to sky light and PMT thermal noise as well as dead-time correction, after which the response profiles are calculated for each line-of-sight according to Eq. (1). Then, the Doppler shift (line-position profile) is computed using the instrument calibration function with account for atmospheric temperature profile, provided by operational analysis by European Center for Medium-range Weather Forecast (ECMWF). Finally, the zonal

and meridional wind components are obtained by comparing the tilted East and North pointings to the corresponding zenith pointing using Eq. (2).

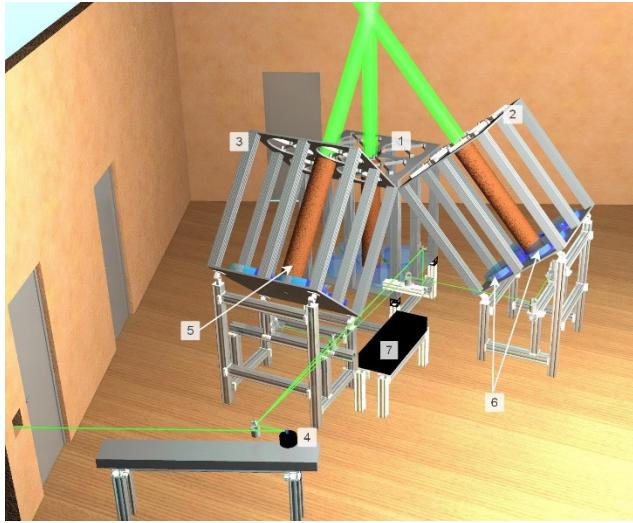


Figure 1. Three-dimensional representation of the transmitter-receiver system of LIOvent lidar comprising three telescope assemblies for zenith (1), north (2) and east (3) lines-of-sight with four collecting mirrors (6) and a beam expander (5) each. The emitted laser beam is commuted between the telescopes using rotating scanner mirror (4). The backscattered light collected by each mirror is transferred via fibers into the optical commutation chamber (7) for aggregating the fluxes from the four mirrors of each telescope assembly.