

REVIEWER 1

Major comments:

1. I know that the terminology used in our community is a little confusing. A LIS “event” is not equivalent to a EUCLID “event“. I would suggest adding a sentence somewhere to emphasize that, so the readers wouldn’t get confused. Something like use the capital Event as a LIS “event.”

=> **Correct. Care must be taken to avoid confusion. Hence, at various places in the text “EUCLID events” is changed into “EUCLID discharges” and ISS-LIS “events” remain “events” where appropriate.**

2. The ISS-LIS data used in this study are non-quality controlled. That could, to some degree, affect the results. I would suggest adding the reference of Blakeslee’s talk at the 2019 GLM meeting (<https://goes-r.nsstc.nasa.gov/home/meetingagenda-2019>), which showed that ISS-LIS is slightly less sensitive than TRMM, and have a few sentences discussing the possible impact.

=> **Although ‘officially’ named as “non-quality controlled” data, it is somewhat misleading. The data does include some quality control in the form of alert/warning flags, however some improvements to the algorithm creating the ISS-LIS data files are still ongoing. The main issue being the sensitivity of ISS-LIS compared to TRMM-LIS, or as formulated by Blakeslee in his 2019 GLM presentation: “The DE for LIS on ISS is on the order of 5% less than the DE for TRMM LIS, reflecting its slightly lower sensitivity. We expect that difference to decrease some when we apply our refined view time corrections to the data associated with the solar panel obscuration presently not being accounted for in view time”. Note that the final quality controlled data are currently unavailable, but will be made available at a later undefined date. The following is added to the text at L127: “The non-quality controlled label is somewhat misleading since the data does include quality information in the form of alert/warning flags, however some improvements to the algorithm creating the ISS-LIS data files is still ongoing. The main issue being the sensitivity of ISS-LIS compared to TRMM-LIS, with the DE for LIS on ISS on the order of 5% less than the DE for TRMM LIS (R. Blakeslee, 2019 GLM Annual Science Team Meeting, Huntsville), impacting to some degree the DE values presented in this paper.”**

3. In Line 206, it says that “on average an ISS-LIS group occurs first.” This looks a little odd to me. LIS group is an accumulation of a 2 ms period and the LIS timestamp is adjusted to the middle of that frame, which is 1 ms earlier than the end of the frame (Bitzer and Christian, 2015). So my impression is that the radio signals should be detected earlier. Zhang et al. (2016) also shows similar results. Do you have an explanation why ISS-LIS group occurs first in here? Possible scenarios could be that the EUCLID didn’t report the initial breakdown processes in the flashes, while LIS reported them. If that’s the case, then what is the detection efficiency of EUCLID reporting initial breakdown processes? Is it known that EUCLID has a poorer DE for initial breakdown than NLDN? I would suggest the authors double-check the results. If the results are correct, then this could lead to some further discussions.

=> **Previously, the best match between a EUCLID stroke and an ISS-LIS group was taken to be the closest in time. However, we believe it should be the closest in space. Changing the ‘best’ choice of match, brings the outcome in line with the expectations whereas now a matched ISS-LIS group is detected slightly later than the corresponding EUCLID stroke. Fig. 5 has been changed accordingly, and the text has been adapted at L170 as follows: “Note that for the stroke/group DE comparisons one ISS-LIS group can be matched to several EUCLID strokes/pulses and vice versa but for time/Location accuracy comparisons only the closest discharge in space is used.” & at L206: “time offset of 0.23 ms (0.11 ms). Thus, on average a EUCLID stroke occurs first.”**

Minor comments:

1. Line 41: Brightness is fine, but the LIS radiance is actually energy density (Koshak, 2010).

=> **The sentence now reads as “The spectral energy density for each event is also available (Koshak et al., 2010), ...”.**

2. Line 216: artifact

=> **Artefact is British, whereas artifact is American spelling. AMT is a European journal, so we remain using artefact in the text.**

3. Line 228: Should it be 12.9% and 20.5%, respectively? The P(EUCLID) in the full domain should be less than in the centre of the network, correct?

=> **Correct, it should have been 12.9% and 20.5%, respectively. However, after some additional consideration and private communication with other scientists after the initial manuscript submission, our current understanding is that the absolute DE values at the level of groups/strokes are not relevant and inappropriate to use. This is due to the fact that the number of groups per LLS report is very large; artificially increasing the ISS-LIS group abs DE value. Therefore we have opted to remove the abs DE values for strokes/groups in Section 4.1 in the reviewed manuscript. The results for the relative DE remain in the text.**

4. The very first paragraph on Page 9 (line number around 255) that discussed the absolute DE of ISS-LIS. The fact that ISS-LIS is less sensitive than TRMM-LIS might also contribute to the findings here. I would suggest adding that in the paragraph.

=> **The text has been changed slightly into: “Additionally, note that in Zhang et al. (2016), the NLDN observations used were restricted to the areas where the NLDN detection efficiency is highest. Hence, P(ISS-LIS) and P(EUCLID) of 63.4% and 69.4 %, respectively, in this work should be compared to P(TRMM-LIS) and P(NLDN) of 81.5% and 58.2 %, respectively, in Zhang et al. (2016). Taking into account the fact that ISS-LIS is somewhat less sensitive compared to TRMM-LIS, would bring the latter values a bit more in line with each other.”**