

# Supplement: Intercomparison of MAX-DOAS vertical profile retrieval algorithms: studies on field data from the CINDI-2 campaign

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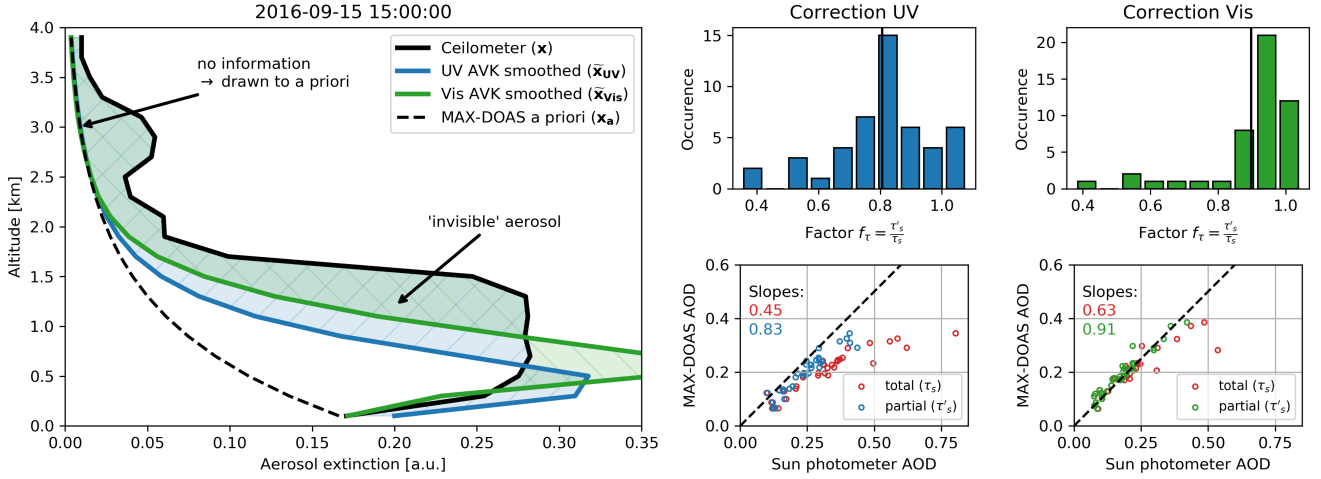
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## S1 M<sup>3</sup> algorithm description

The M<sup>3</sup> algorithm developed at the Ludwig-Maximilians-University in Munich is an OEM algorithm with Newton Gauss optimization. The radiative transfer model LibRadTran (Mayer and Kylling, 2005) serves as forward model for the retrieval. LibRadTran provides several radiative transfer equation solvers which can handle both pseudo spherical and full spherical geometry. The Jacobian is calculated numerically using the finite difference method, while the box air mass factors for trace gas profile retrievals are calculated using the Monte Carlo module of LibRadTran (MYSTIC). The M<sup>3</sup> retrieval used in the CINDI-2 campaign is a modified version of the algorithm described in detail in Chan et al. (2019) with the iterative optimization of the *a priori* profile disabled. Aerosol extinction profiles are retrieved in the linear space.

## S2 Partial AOT correction

As described in the main text Sect. 3.4, ground-based MAX-DOAS observations have limited sensitivity to aerosol at higher altitudes and thus the profile is strongly biased by *a priori* assumptions here. In the case of OEM algorithms, this effect can be accounted for by applying what in the following is referred to by "partial AOT correction" (PAC). This Sect. demonstrates the impact of the PAC on the comparison between MAX-DOAS and sun photometer data. The left panel of Fig. S1 shows an example of an extreme case during the campaign from September 15<sup>th</sup>, 15:00h. Shown are a ceilometer backscatter profile ( $x$ , black) and the same profile smoothed by the MAX-DOAS median OEM averaging kernels for Aerosol UV and Aerosol Vis ( $x_m$ , blue and green), respectively. In this particular case it is expected that a large fraction of the aerosol above 1 km altitude will not be detected by MAX-DOAS instruments, resulting in factors  $f_\tau = \frac{\tau'_s}{\tau_s}$  of 0.67 and 0.78, for the UV and the Vis AOT, respectively. However, a part of the high-altitude aerosol appears to be shifted to lower altitudes here by the retrieval. The right panels in Fig. S1 show information for the UV and the Vis retrieval (2nd and 3rd columns) over the whole campaign. The mean values are  $f_\tau = 0.81 \pm 0.16$  ( $0.9 \pm 0.13$ ) for the UV (Vis) AOTs. It shall be pointed out that for OEM algorithms the necessity for the PAC can be reduced by using improved *a priori* profiles and covariances (e.g. from climatologies and/ or model data).



**Figure S1.** Left panel: example for the smoothing of a ceilometer backscatter profile  $\mathbf{x}$  (according to Eq. (10) in the main text) with particularly heavy aerosol load at high-altitudes retrieved in the UV and Vis, respectively. Right panel: distribution and impact of the correction factor  $f_\tau = \tau'_s / \tau_s$  for the UV and the Vis retrieval. On the top, the distributions of  $f_\tau$  are shown with the solid lines indicating the mean values. At the bottom the correlation plots between sun photometer and MAX-DOAS median AOTs are shown. Red circles represent sun photometer total AOTs A, other dots represent the partial AOT  $\tau_s = f_\tau \cdot \tau'_s$ .

### S3 O<sub>4</sub> scaling factor

By some groups, the O<sub>4</sub> scaling factor  $SF$  is applied to the measured dSCDs before the profile inversion. Initial motivation for its application are previous MAX-DOAS retrieval studies (e.g Wagner et al., 2009; Cl  mer et al., 2010) which report on a significant mismatch between measured and simulated dSCDs and/ or between MAX-DOAS integrated aerosol extinction and simultaneously measured sun photometer AOT that could not yet be explained (Wagner et al., 2019; Ortega et al., 2016). Meanwhile a series of studies use an  $SF$ , empirically determined to values between 0.75 and 0.9.

As described in Sect. 2.1.3, in this study no scaling of O<sub>4</sub> measured dSCDs was applied, except for MPIC-mp0.8 (MAPA algorithm with  $SF = 0.8$ ). For CINDI-2, a  $SF$  of 0.8 was observed to enhance the number of valid profiles retrieved by MAPA and to significantly improve the agreement to the sun photometer total AOT in particularly in the UV (Beirle et al. (2019) and within this study). However, as mentioned before, for MAPA (as a parametrized approach without *a priori* profile and AVKs), a PAC as described in Sect. 3.4, cannot be correctly applied and thus deviations to the sun photometer are expected. To further investigate the impact and necessity of the  $SF$  for CINDI-2 retrievals, also HEIPRO (as an OEM retrieval) was run with different  $SF$ s. In Fig. S2 the impacts of the  $SF$  and the PAC on the agreement between MAX-DOAS profiling results (of HEIPRO and MAPA) and the sun photometer are directly compared. Application of  $SF = 0.8$  or the PAC, respectively, lead to a very similar improvement in the agreement (regarding RMS), while the application of both together results in a clear overcompensation. This suggests that the PAC and  $SF = 0.8$  are equivalent to a large extent and that in the case of MAPA the

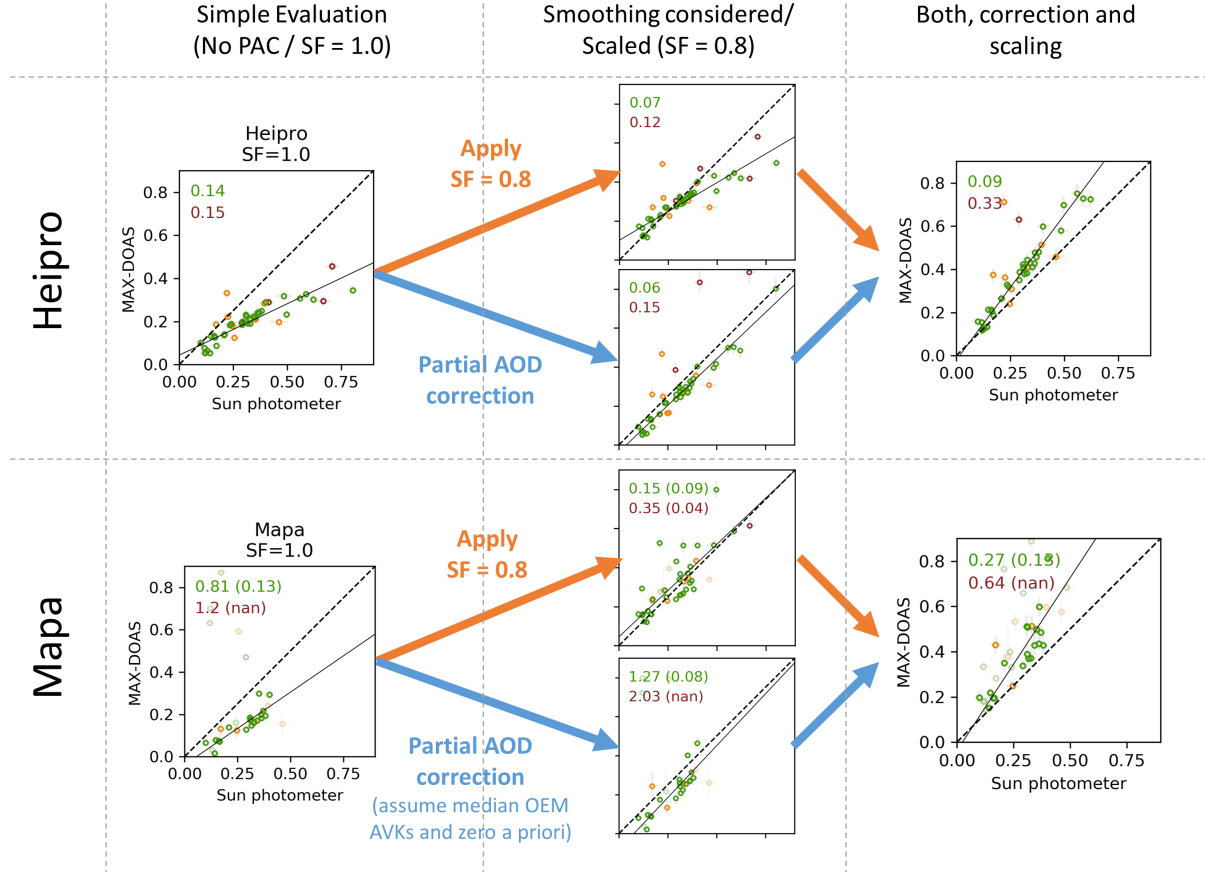
$SF$  is a way to at least partly account for high-altitude aerosol when it comes to retrieving total AOTs. On the other hand a closer look reveals that if only the PAC is applied, a systematic negative offset of  $\approx -0.04$  remains in the correlation (for both algorithms and also other participants, compare to main text Sect. 3.4). Indeed, the top row of Fig. S4 shows that for HEIPRO the best RMSD is observed for  $SF = 0.92 \pm 0.02$  (UV). Regarding Aerosol Vis (Figure S3), the impact of high aerosol is smaller due to the enhanced vertical sensitivity range (see AVKs in main text Sect. 3.1), such that applying the same scaling factor  $SF = 0.8$  as for the UV (without PAC) should already lead to an overcompensation. Indeed, this is observed for HEIPRO as well as MAPA. The bottom row of Fig. S4 shows that, in contrast to the UV spectral range, the best agreement between sun photometer and MAX-DOAS in the Vis is observed for an  $SF > 1$ .

The second indicator for the need of a scaling factor is a significant mismatch between modelled and measured dSCDs. Figure S5 shows, that for HEIPRO the application of an  $SF < 1$  indeed improves the agreement between measured and modelled  $O_4$  UV dSCDs by up to 35% in RMSD under clear-sky conditions. Modelled dSCDs are systematically lower than the measured dSCDs in particular for higher elevation angles. Finally, Wagner et al. (2009) reported, that under low aerosol conditions, measured dSCDs sometimes even exceed dSCDs modelled within an aerosol free atmosphere, where  $O_4$  dSCDs are expected to be the largest possible (regarding clear-sky scenarios only). Therefore, the median measured dSCDs during CINDI-2 at low aerosol load ( $\tau_s < 0.1$ ) were compared to a set of aerosol free modelled dSCDs, showing that this did not happen during CINDI-2 except for the two highest elevation angles  $\alpha = 15^\circ$  and  $30^\circ$ .

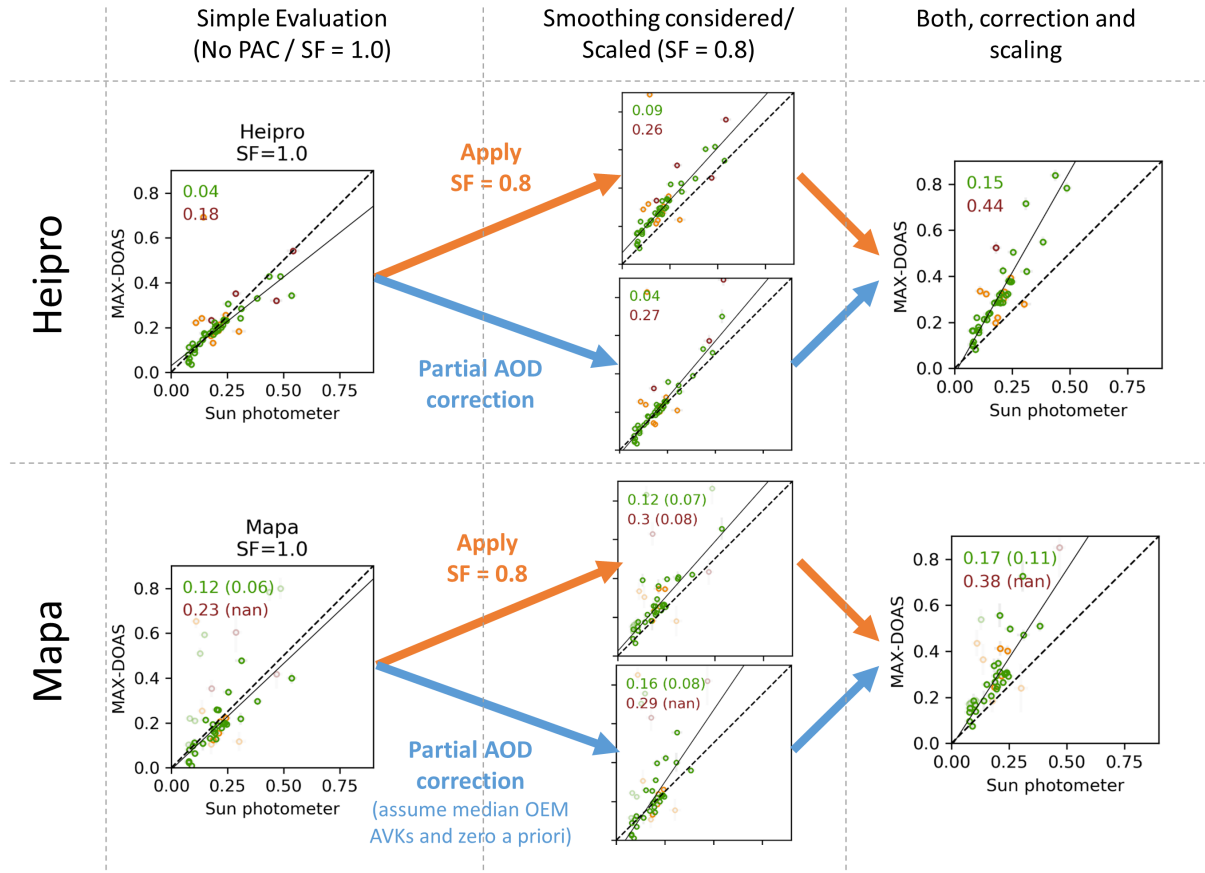
Finally, as already stated by Beirle et al. (2019) applying  $SF = 0.8$  to MAPA leads to an increased number of valid profiles (see Sect. S4), which again indicates that scaling brings the RTM closer to reality.

The conclusions drawn are as follows: Even without  $O_4$  dSCD scaling, reasonable results and agreement with supporting observations are achieved (if a PAC is applied). In general, a scaling factor of 0.8 seems to be too small but might at least partly be used to account for high-altitude aerosol for algorithms, that cannot quantify their sensitivity or the assimilated *a priori* information. However, there are indications that a less extreme scaling ( $0.8 < SF < 1.0$ ) might in general improve the retrieval. Finally, we think that for this study the prescribed  $SF = 1.0$  is justified. Even though it might not be ideal, it is the most straightforward approach and yields reasonable and consistent results within the uncertainties introduced by other factors. To draw more concise conclusions, further studies similar to Wagner et al. (2019) are necessary.

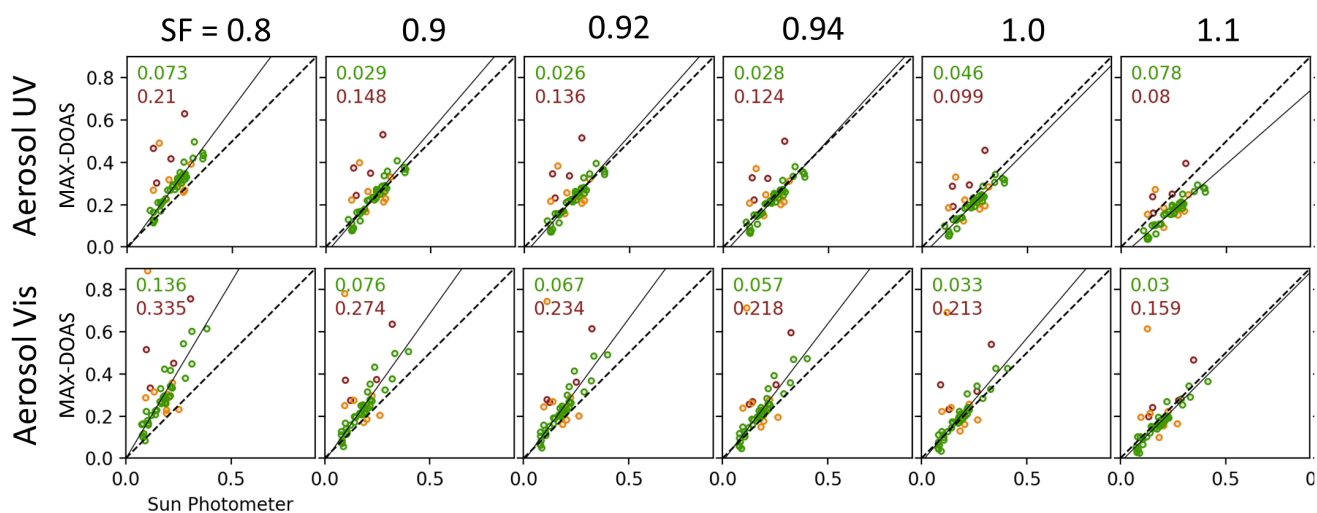




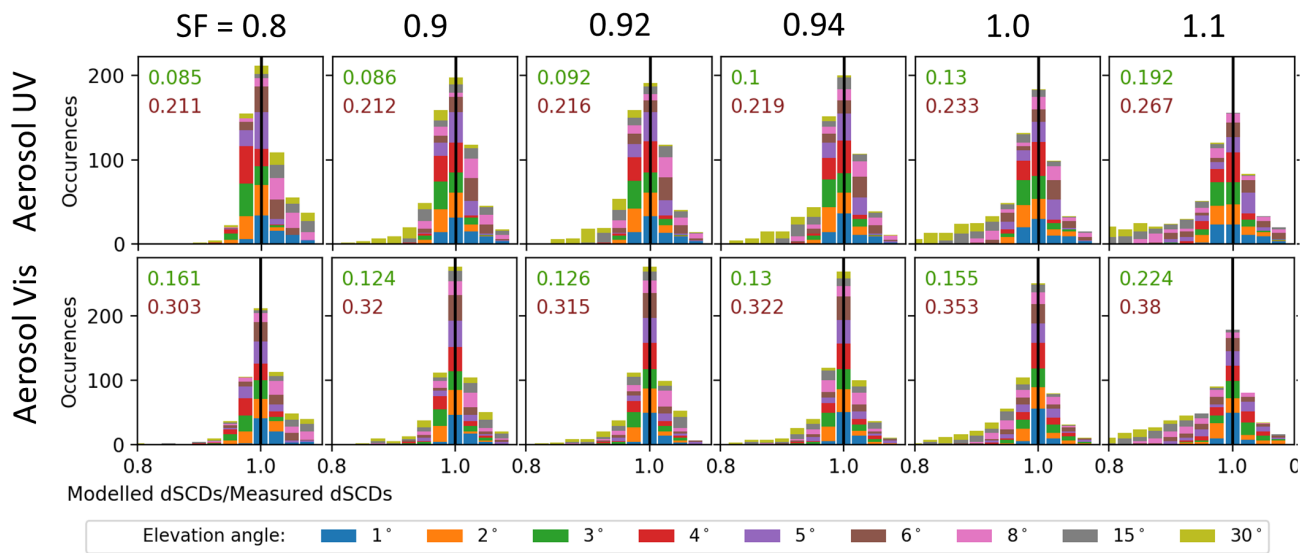
**Figure S2.** The impact of  $SF = 0.8$  and PAC on the agreement between sun photometer and MAX-DOAS AOT (Aerosol UV) in the case of HEIPRO (OEM approach, top row) and MAPA (parametrized approach, bottom row) in a direct comparison. Axes limits and labels of the plots on the left apply for all plots in the figure. Left column: A standard retrieval with  $SF = 1$  yields a clear underestimation of the sun photometer AOT for both algorithms. Middle column: Applying the PAC or  $SF = 0.8$  leads to a significant improvement. Right column: Applying both leads to overcompensation. Note, that the PAC for MAPA incorporates vague assumptions (median AVKs from all OEM algorithms and  $x_a = 0$ ) and is therefore less meaningful.



**Figure S3.** The impact of  $SF$  and PAC on Aerosol Vis results. Description of Fig. S2 applies. Due to the extended vertical sensitivity range of the MAX-DOAS observations (compare to main text Sect. 3.1), the effect of high-altitude aerosol is less significant. While this is accounted for in the PAC, the application of the same  $SF = 0.8$  as for Aerosol UV leads to an overcompensation here already without additionally applying the PAC.



**Figure S4.** Correlation between the sun photometer AOT (PAC applied) and the HEIPRO algorithms for different  $SF$  for Aerosol UV (top row) and Aerosol Vis (bottom row). Symbol colours differentiate between clear-sky (green), and cloudy conditions (orange, red). Numbers in the upper left corners show the corresponding RMSD values. Interestingly, the best agreement for the two spectral ranges is found for different scaling factors, namely for  $SF = 0.92 \pm 0.02$  (UV) and  $SF > 1 \pm 0.02$  (Vis). The thin line depicts the linear fit for clear-sky data only.



**Figure S5.** Agreement between modelled and measured dSCDs (depicted as histograms of the ratio) for HEIPRO with different  $SF$  for Aerosol UV (top row) and Aerosol Vis (bottom row). Here, only clear-sky and valid data are considered. Numbers in the upper left corners show the corresponding RMSD values. There is a clear tendency, that the agreement improves for  $SF < 1$ .

#### S4 Details on profile flagging

Participants were allowed to submit flags, giving them the opportunity to mark profiles as invalid. Four participants submitted flags based on the following criteria:

- 5       – BIRA/ bePRO: Profiles are considered valid if the retrieved degrees of freedom are  $> 1$  and if the difference between measured and modelled dSCDs is smaller than 30 %.
- INTA: Profiles are valid if DOFS  $> 1$  and if the RMSD between measured and simulated dSCDs is smaller than 1.5 times the daily averaged RMS.
- 10     – KNMI: Profiles are invalid if the spread in the ensemble of solutions for AOT or the  $\text{NO}_2$  or HCHO tropospheric column is larger than 15 % of the retrieved value, or if there are less than 5 out of the ensemble of 20 retrievals for which a solution was found.
- MAPA: Flagging is based on a row of criteria (agreement of modelled and measured dSCDs, implausible results, consistency within the ensemble of possible solutions, ...) with carefully chosen thresholds. A detailed description can be found in Beirle et al. (2019).

15     Table S1 shows the statistics regarding the number of submitted profiles and the fraction of invalid profiles for all participants and species.

**Table S1.** Submission and flagging statistics.

		HCHO		NO <sub>2</sub> UV		NO <sub>2</sub> Vis		Aerosol UV		Aerosol Vis	
		Total	Valid [%]	Total	Valid [%]	Total	Valid [%]	Total	Valid [%]	Total	Valid [%]
bePRO	AUTH <sup>o</sup>	170	100	170	100	-	-	170	100	-	-
	BIRA <sup>o</sup>	170	93	170	93	170	87	170	93	170	88
	INTA <sup>o</sup>	170	78	170	75	170	71	170	87	170	80
PRIAM	AIOFM <sub>l</sub> <sup>o</sup>	170	100	170	100	-	-	170	100	-	-
	BSU <sub>l</sub> <sup>o</sup>	170	100	170	100	170	100	170	100	170	100
	CMA <sub>l</sub> <sup>o</sup>	169	100	169	100	169	100	169	100	169	100
	MPIC <sub>l</sub> <sup>o</sup>	170	100	170	100	170	100	170	100	170	100
HEPRO	IUPHD <sub>l</sub> <sup>o</sup>	170	100	170	100	170	100	170	100	170	100
	UTOR <sub>l</sub> <sup>o</sup>	-	-	170	100	170	100	170	100	170	100
BOREAS	IUPB <sub>l</sub> <sup>o</sup>	170	100	170	100	170	100	170	100	170	100
M <sup>3</sup>	LMU <sup>o</sup>	170	100	170	100	170	100	170	100	170	100
MMF	BIRA <sub>l</sub> <sup>o</sup>	170	100	170	100	170	100	170	100	170	100
Realtime	NASA <sup>a</sup>	170	100	170	100	170	100	-	-	170	100
MARK	KNMI <sup>p</sup>	107	38	152	61	168	76	107	43	168	77
MAPA	MPIC-1.0 <sup>p</sup>	170	31	170	32	170	22	170	32	170	33
	MPIC-0.8 <sup>p</sup>	170	52	170	51	170	37	170	52	170	43

## S5 Further details on supporting observations

### S5.1 Aerosol extinction profiles

- The available raw data from the ceilometer are attenuated backscatter coefficient profiles  $\tilde{\beta}(h)$ . For altitudes below 180 m, data are invalid due to insufficient overlap between sending and receiving telescope's FOVs of the instrument, thus  $\tilde{\beta}(h < 180\text{ m})$  was set to  $\tilde{\beta}(h = 180\text{ m})$ . For profiles with simultaneously available sun photometer AOTs  $\tau_s(\lambda)$ , attenuated backscatter profiles  $\tilde{\beta}(h)$  were converted to approximate backscatter profiles  $\beta(h)$ , applying

$$\beta(h) = \tilde{\beta}(h) \cdot \exp \left[ 2 \int_0^h \frac{\tau_s(\lambda = 1064\text{ nm})}{\int \tilde{\beta}(h') dh'} dh \right] \quad (1)$$

Extinction coefficient profiles  $\alpha_\lambda(h)$  at the MAX-DOAS retrieval wavelengths were then obtained by scaling of  $\beta(h)$  with the sun photometer AOT according to:

$$\alpha_\lambda(h) = \frac{\tau_s(\lambda)}{\int \beta(h') dh'} \cdot \beta(h) \quad (2)$$

Integrands with no specified limits in Eq. (1) and (2) indicate integration over the entire available profile. Values for  $\tau_s(\lambda)$  at the desired wavelengths were derived according to Eq. (4) in the main text. In case of missing sun photometer data, for instance

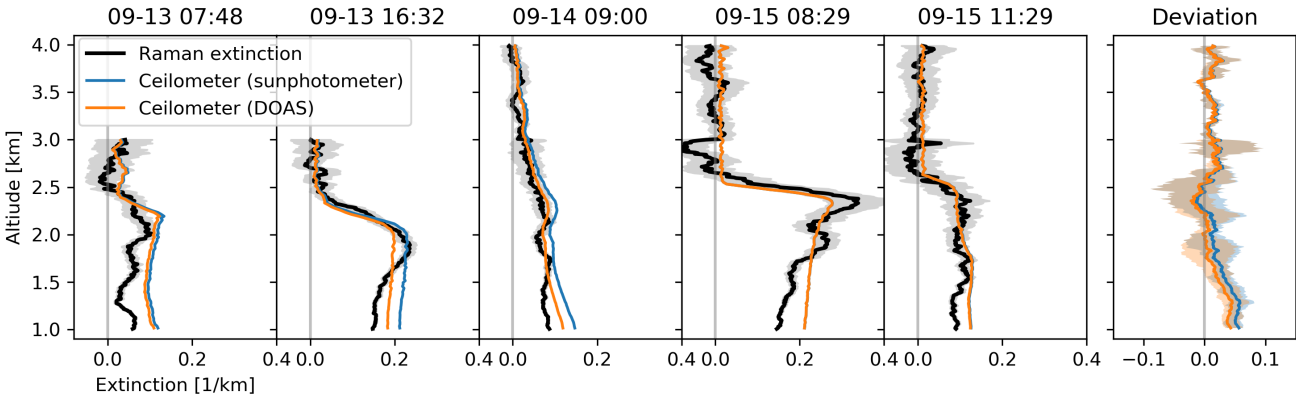
due to clouds, the MAX-DOAS retrieved AOT was used instead of  $\tau_s$ . In this case no attenuation correction (Eq. (1)) could be applied and the integration in Eq. (2) was performed over an averaging kernel smoothed profile (see main text Sect. 2.3.2), to take into account the blindness of MAX-DOAS instruments for higher aerosol layers.

The resulting extinction coefficient profiles at 360 nm could partly be validated with Raman lidar observations at 355 nm (the CESAR Water Vapor, Aerosol and Cloud lidar “CAELI”, operated within the European Aerosol Research lidar Network (EARLINET, Bösenberg et al., 2003; Pappalardo et al., 2014) and described in detail in Apituley et al., 2009). Since for the Raman lidar there is not sufficient telescope FOV overlap for altitudes  $< 1$  km to retrieve reliable extinction profiles this comparison is limited to the altitude range between 1 and 4 km. The average RMSD between scaled ceilometer and Raman lidar profiles is  $\approx 0.03$ . Table S2 summarizes the instrument’s properties. Figure S6 shows the available Raman lidar profiles in comparison to the ceilometer derived profiles scaled with the respective sun photometer and MAX-DOAS AOT respectively.

**Table S2.** Properties of the two lidar instruments.

Instrument	Raman lidar	Ceilometer
Data product	Aerosol extinction profile	Elastic backscatter profile
Operational wavelength	355 nm	1064 nm
Altitude range	1 to 10 km	0.2 to 15 km
Vertical resolution	7.5 m	10 m
Temporal resolution	30 s	12 s
Data coverage	5 profiles between 13.9. and 15.9.	Whole campaign

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**Figure S6.** Comparison of the five available aerosol extinction profiles obtained from Raman lidar measurements (in black, with uncertainties indicated by the grey areas) with AOT scaled ceilometer backscatter profiles. Scaling was performed with AOTs from sun photometer (blue) and DOAS data (orange). In the last plot, lines show the mean deviation, whereas the borders of the coloured areas mark maximum and minimum deviation.

**S5.2 Radiosonde flights**

An overview of the radiosonde flights is given in Fig. S7. Example profiles are shown in the course of a comparison between lidar and radiosonde observations in Supplement S5.5. Another preprocessing step of the sonde profiles shall be mentioned: data quality was affected by calibrational drifting of the sensor, as it was optimized for low weight and cost rather than performance. Even though a calibration against the CE-DOAS was performed directly before each launch, most profiles showed a clear instrumental offset of about  $(1-2) \times 10^{10}$  molec cm<sup>-3</sup> in the free troposphere. The offset was subtracted and the profiles were subsequently rescaled to their initial surface concentration.

**Table S3.** Overview over the radiosonde sampling flights shown in this study.

Launch date	Flight time <sup>a</sup> [min]	Travel distance <sup>a</sup> [km]	Wind direction
9-13 08:42	10	7	SE
9-14 09:03	12	5	SE
9-14 13:06	14	4	SE
9-15 08:04	10	8	E
9-15 10:25	11	8	SE
9-21 07:57	12	10	SE
9-21 10:14	15	5	SE
9-25 06:59	17	7	S
9-25 09:29	12	18	S

<sup>a</sup> Only considering trajectory through the lowest 4 km of the atmosphere.

**S5.3 NO<sub>2</sub> lidar**

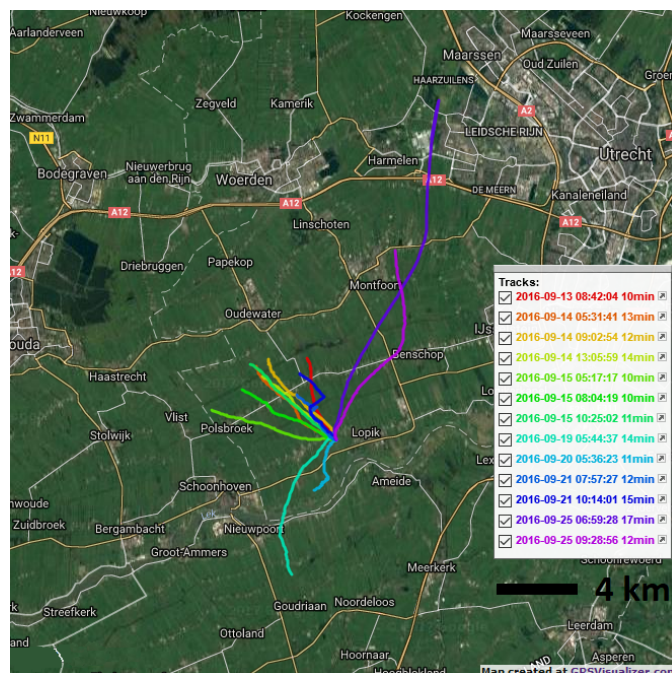
The NO<sub>2</sub> lidar provides profiles consisting of a series of altitude intervals or “boxes” with constant gas concentration between a lower and an upper altitude limit. The conversion to profiles on the MAX-DOAS 200 m grid is demonstrated in Fig. S8. First, the boxes were converted to a continuous profile by linearly interpolating over box overlaps or gaps, which was then averaged down to the 200 m MAX-DOAS retrieval grid resolution.

**S5.4 Long path DOAS**

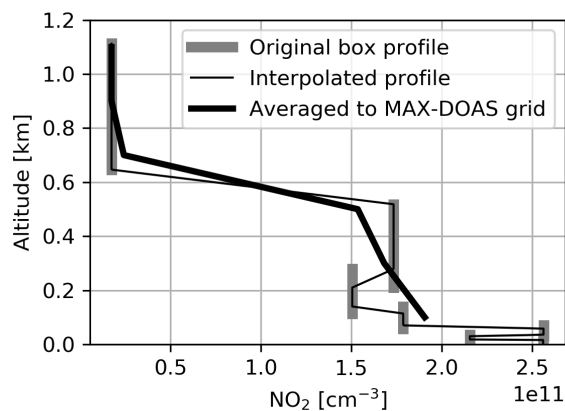
**S5.5 Consistency of supporting observations**

The agreement of redundant supporting observations (processed as described in the main text Sect. 2.2) gives an impression of their reliability and/ or representativeness. In the case of NO<sub>2</sub> several observations of total vertical columns and surface concentration (note again, that throughout this paper “surface concentration” refers to the average concentration in the lowest retrieval layer) are available and compared in the correlation plots in Fig. S10 below. Corresponding time series plots are already

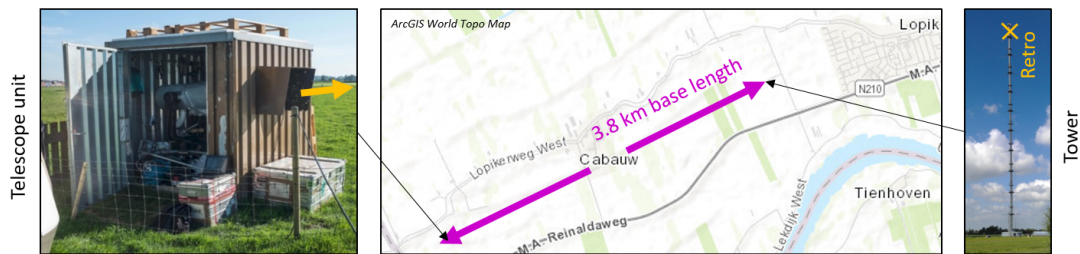




**Figure S7.** Sonde flight paths in the course of the campaign. Only data of the sonde ascent through the lowest 4 km of the atmosphere are shown. Some flights shown here are not included in the comparison as they were launched before 6:30 h. The "minute" values in the legend labels represent the flight time.



**Figure S8.** Regridding of an example  $\text{NO}_2$  lidar box profile (27 September 2016, 11:00) to the MAX-DOAS 200 m vertical resolution. In a first step, gaps and overlaps within the box profile are linearly interpolated. The resulting profile is then averaged within the MAX-DOAS retrieval layers.

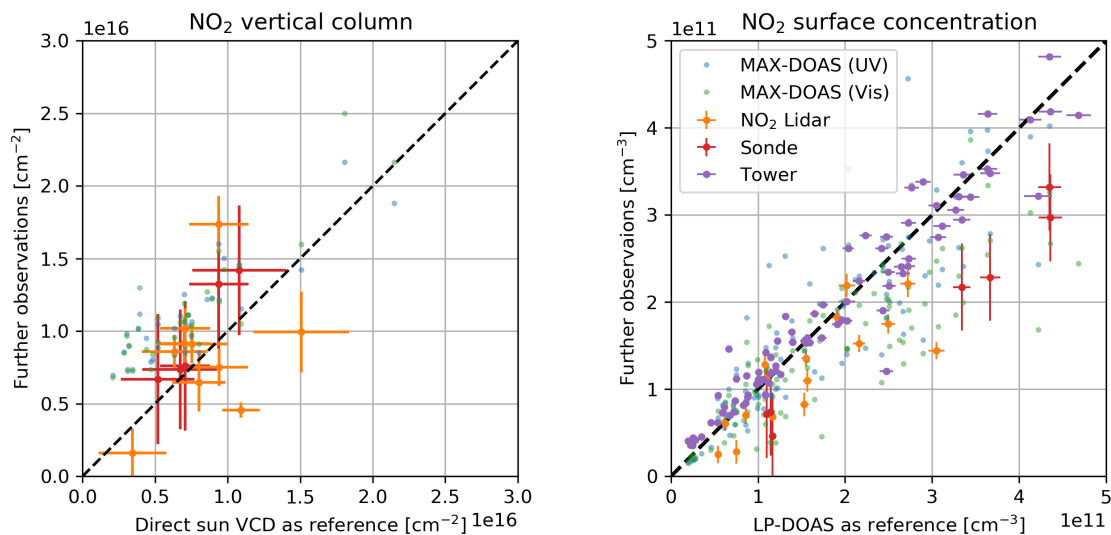


**Figure S9.** Setup of the LP-DOAS system. As shown on the map (Esri et al., 2018), the light sending and receiving telescope unit (left) was located at 3.8 km distance to the meteorological tower (right), resulting in a total light path of 7.6 km. There were several retroreflectors installed on the tower at different altitudes. However for this study, only the one at the very top (207 m altitude) was used to obtain the average gas in the lowest retrieval layer, extending from 0 to 200 m altitude.

shown in the main text in Fig. 16 and 19, respectively. Tables S4 and S5 show the RMSD (as observed) and  $\sigma$  (the expected deviation according to the specified measurement uncertainties) between each possible pair of observations. For the VCDs, the RMSD is close to  $\sigma$  or below. A maximum RMSD of  $1.5 \sigma$  is found between  $\text{NO}_2$  lidar and direct-sun DOAS. For the surface concentrations however, there seem to be systematic deviations which split the observations into two pairs: radiosonde and lidar observations agree well but are both systematically lower than LP-DOAS and tower measurements. MAX-DOAS UV agrees better with LP-DOAS and tower observations, while MAX-DOAS Vis agrees more with sonde and lidar. Between the LP-DOAS and the  $\text{NO}_2$  lidar, an RMSD of more than  $4 \sigma$  is observed. There are several potential explanations:

1. Biases are introduced due to data processing (temporal and spatial regridding, for instance for the lidar profiles described in Sect. S5.3).
2. Spatio-temporal variability of the real gas abundances
3. Imperfect estimates of the measurement uncertainties (in particular systematic deviations)

For the  $\text{NO}_2$  lidar and the radiosondes, there are four simultaneously recorded  $\text{NO}_2$  profiles available over the campaign (simultaneous in the sense that for a single MAX-DOAS profile timestamp, profiles from both systems are available according to the definitions in Sect. 2.2.3 in the main text). They are compared in the top row of Fig. S11. For the first situation, where good spatial and temporal overlap is given, there is mostly an agreement within the specified errors. In the case of bad temporal and/ or spatial overlap, strong deviations occur. For the 2<sup>nd</sup> and the 4<sup>th</sup> plot, there are several lidar profiles available, which are temporally closer to the radiosonde (however further away from the corresponding MAX-DOAS profile), which in contrast show very good agreement again. This shows that the real  $\text{NO}_2$  profile varies strongly even on timescales of  $\approx 30$  minutes (see also (Peters et al., 2019)) and that improved synchronisation between MAX-DOAS and supporting observations should be considered for future campaigns.



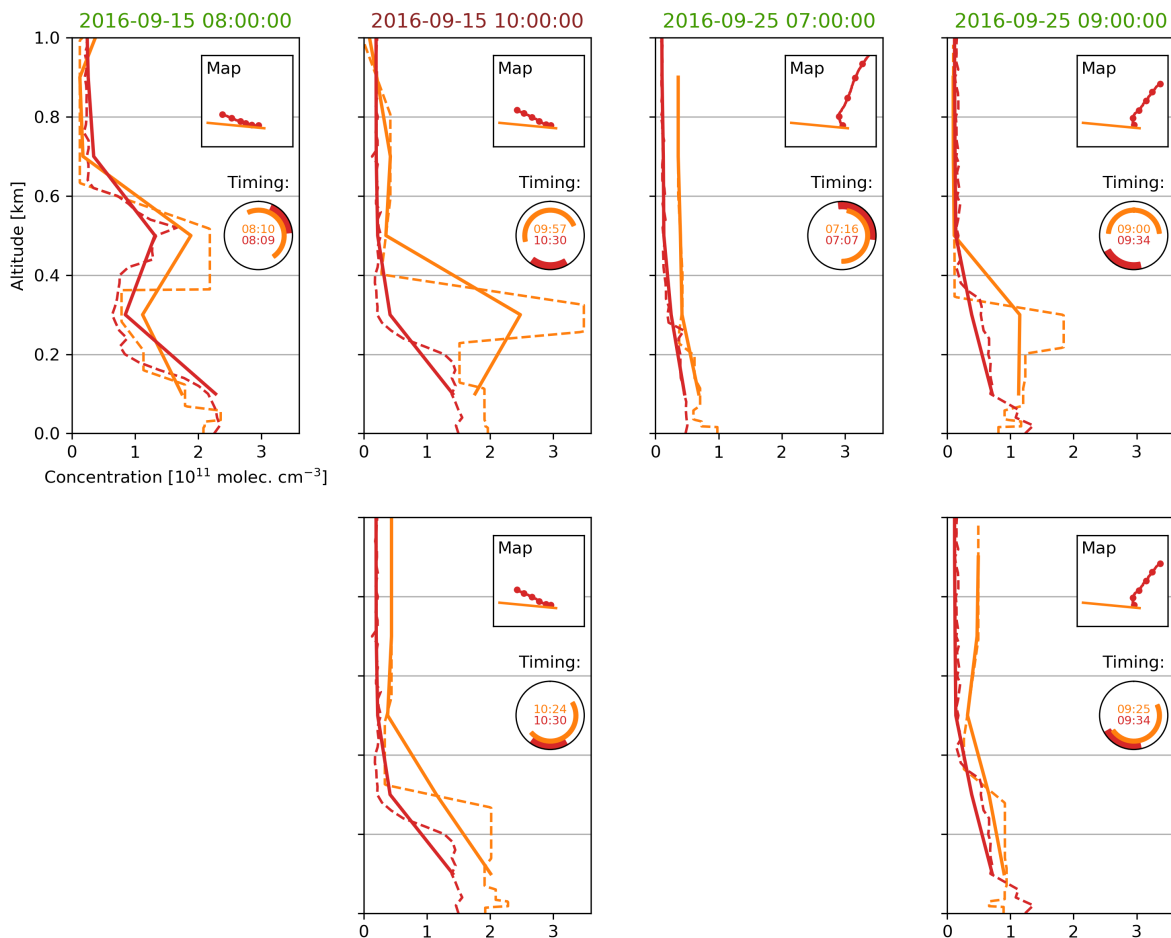
**Figure S10.** Comparison of redundant supporting observations of NO<sub>2</sub> VCDs (left panel) and surface concentration (right panel). MAX-DOAS retrieved values are plotted in the background. To improve visibility, tower measurement uncertainties (vertical error bars of typically  $(6.0 \pm 0.5) \times 10^{10} \text{ molec cm}^{-3}$ ) are not shown.

**Table S4.** Comparison of redundant measurements of the NO<sub>2</sub> surface concentration (in  $10^{11} \text{ molec cm}^{-3}$ ). For each pair of observations, the observed scatter (RMS) is compared to the specified uncertainty ( $\sigma$ ).

	Tower in-situ		Radiosonde		NO <sub>2</sub> -Lidar	
	RMSD	$\sigma$	RMSD	$\sigma$	RMSD	$\sigma$
LP-DOAS	0.32	0.56	1.01	0.51	0.57	0.13
NO <sub>2</sub> -Lidar	0.72	0.57	0.40	0.52	-	-
Radiosonde	0.99	0.78	-	-	-	-

**Table S5.** Comparison of redundant measurements of the NO<sub>2</sub> total columns (in  $10^{16} \text{ molec cm}^{-2}$ ). For each pair of observation, the observed scatter (RMS) is compared to the specified uncertainty ( $\sigma$ ).

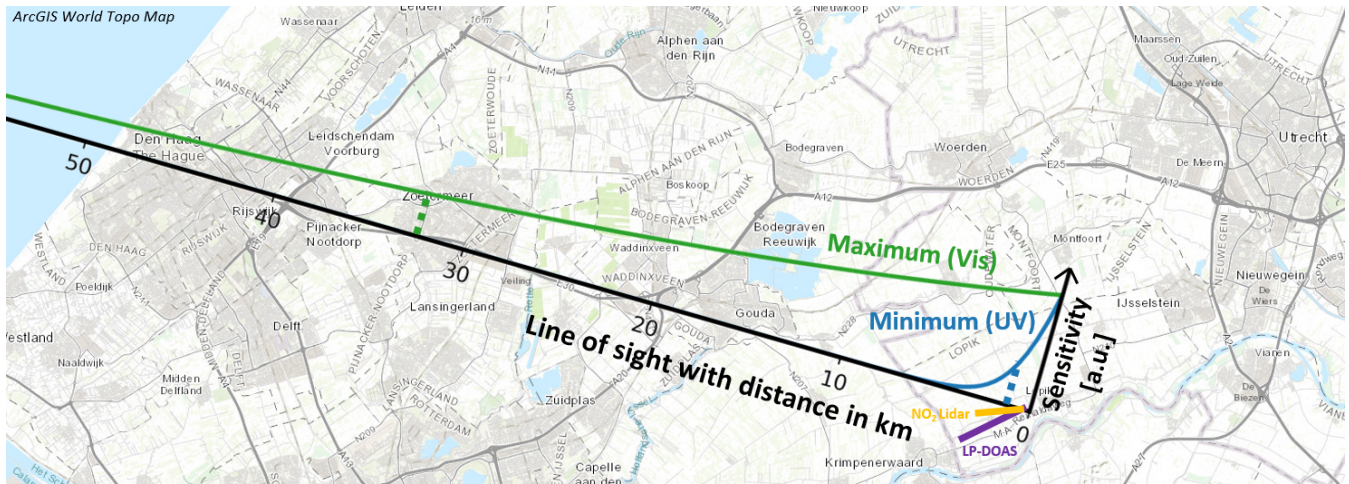
	Radiosonde		NO <sub>2</sub> -Lidar	
	RMSD	$\sigma$	RMSD	$\sigma$
LP-DOAS	0.24	0.51	0.40	0.26
NO <sub>2</sub> -Lidar	0.34	0.48	-	-



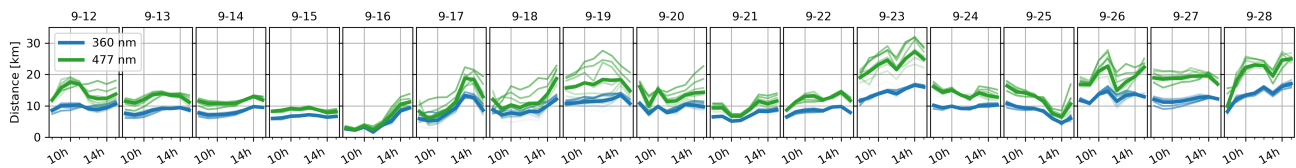
**Figure S11.** Top row: Comparison of NO<sub>2</sub> lidar (orange) and radiosonde (red) profiles, that were assigned to a common MAX-DOAS profile timestamp according to Sect. 2.2.3 in the main text. Dashed lines represent original instrument resolution while thick lines show the concentrations averaged to the MAX-DOAS altitude grid. The colour of the dates indicates the cloud conditions. The rectangular little subplots show a map (4 x 4 km) of the lidar line of sight and the sonde flight path. The dots on the sonde flight path mark the transitions between the retrieval layers. The polar plot (to be read like a clock) shows the temporal overlap between the two observations, together with the middle timestamps of each observation. Lower row: For the 2<sup>nd</sup> and 4<sup>th</sup> timestamp, there were lidar profiles available with improved temporal overlap (however, with a worse overlap with the corresponding MAX-DOAS profile).

## S6 MAX-DOAS viewing distance

Wagner and Beirle (2016) derived polynomial relationships between the “horizontal sensitivity range” (HSR, defined as the distance, at which the box airmass factors dropped to  $1/e$ ) and  $O_4$  differential airmass factors (dAMF). Applying this approach to the CINDI-2  $O_4$  dAMFs yields the HSRs shown in Fig. S13. A constant vertical  $O_4$  column of  $1.19 \cdot 10^{43} \text{ molec}^2 \text{ cm}^{-5}$  was assumed. The HSR for the actually retrieved layers is more complicated and not assessed here, as information aspects (which elevation contributes to information on which layer), geometrical limitations and the atmospheric state (trace gas and aerosol layer height) would have to be taken into account. Depending on the conditions, HSRs vary between a few and tens of kilometres (as shown in Fig. S12) defining whether only air masses over rural areas and/or urban areas (Gouda at 15 km distance, Zoetermeer at 30 km distance and The Hague at 40 km distance. to the measurement site) are sampled. Further, depending on the wind, plumes of Utrecht, Rotterdam or Amsterdam might be sampled.



**Figure S12.** Line of sight of the MAX-DOAS instruments on a map (Esri et al., 2018). The coloured curves indicate the sensitivity for the two extreme (shortest and longest) viewing distances encountered during the campaign. Dashed lines indicate the 1/e-length.



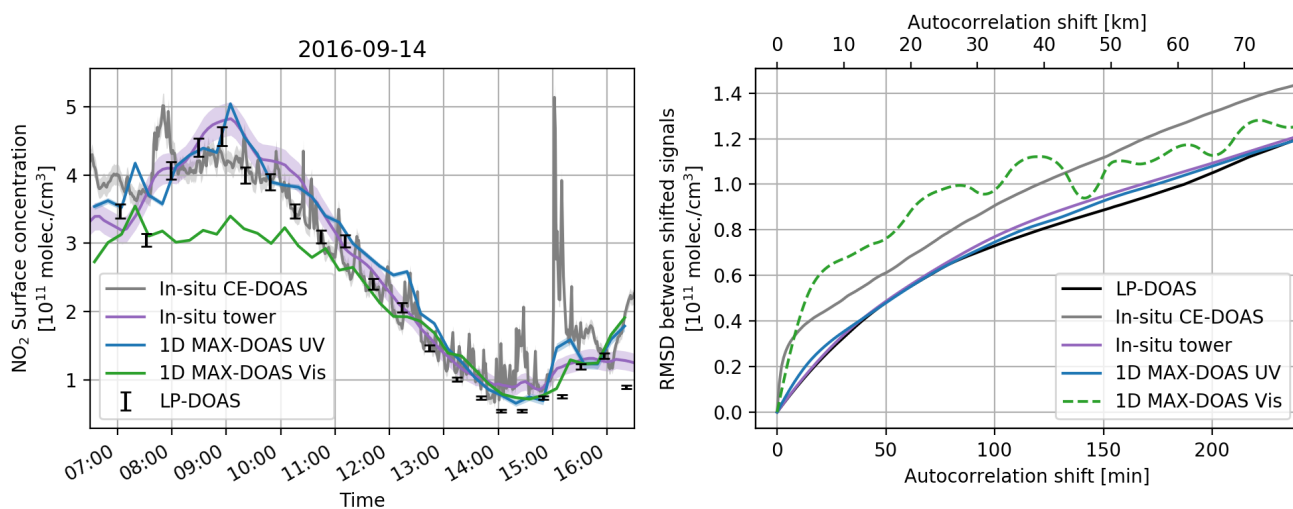
**Figure S13.** Viewing distance (HSR) of MAX-DOAS instruments during CINDI-2. It was calculated for different elevations angles (1, 2, 3, 4, 5, 6 and 8° with increasing transparency of the curves) and the average value for UV and Vis (thick lines).

## S7 Spatio-temporal variability

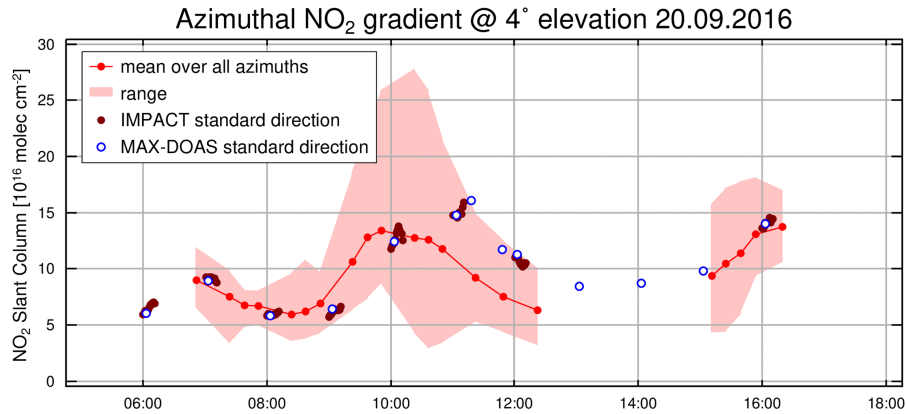
For an impression of the impact of spatio-temporal variability on the comparisons, different observations of the  $\text{NO}_2$  surface concentrations at their original temporal resolution  $\Delta t$  and integration time  $t_{int}$  are shown in the left panel of Fig. S14. The CE-DOAS as a point measurement with  $\Delta t = t_{int} = 1$  min shows very strong variability on short timescales. However, for the tower measurements (all in situ instruments in the tower vertically integrated as described in main text Sect. 2.2.5 at  $\Delta t = 20$  min and  $t_{int} \approx 5$  min), the LP-DOAS ( $\Delta t = 32$  min,  $t_{int} \approx 100$  s) there is already significant smoothing. The 1D-MAX-DOAS data was recorded by DLR (see Supplement S10), who retrieved profiles in the nominal azimuth direction ( $287^\circ$ ) more or less continuously ( $\Delta t = 15$  min,  $t_{int} \approx 10$  min). In all measurements there is significant variation on the sub-hour timescale. Further, spatial variability might be observed in the form of disagreement between UV and Vis observations of the 1D-MAX-DOAS as viewing distance and thus the sampled air volume changes between the two spectral ranges (see Supplement S6). To estimate the order of magnitude, the right panel of Fig. S14 shows a kind of autocorrelation of the total campaign time series of each observation. The RMSD between the original and a temporally shifted signal is calculated. Note, that the 1D MAX-DOAS Vis data is not reliable as it features a lot of gaps. As described in main text Sect. 2.2, typical temporal



shifts/ interpolation intervals between different observations during CINDI-2 are in the order of 30 minutes. RMSD values at 30 minutes shift are around  $3.5 \times 10^{10} \text{ molec cm}^{-3}$  and thus indeed of similar order as typical observed RMSD values in the  $\text{NO}_2$  surface concentration comparisons within this study (approx.  $5 \times 10^{10} \text{ molec cm}^{-3}$ , compare to main text Fig. 23). For an impression on the spatial variability, we refer to data from the IMPACT instrument (Peters et al., 2019), an imaging MAX-DOAS operated by IUP-Bremen (IUPB) which allows to perform elevation "scans" in different azimuth viewing directions in quick succession. During CINDI-2, the IMPACT performed full-azimuthal scans in  $10^\circ$  steps every 15 minutes. Figure S15 exemplarily shows the observed  $\text{NO}_2$  Vis dSCDs at  $4^\circ$  elevation on the 20 September 2016 together with dSCDs measured by the IUPB standard MAX-DOAS instrument in the nominal azimuth direction ( $287^\circ$ , compare main text Sect. 2.1). The red shaded area depicts the variation of the dSCD with azimuth viewing direction. In particular around local noon this variation is tremendous, exceeding a factor of five. Further investigation on this issue can be found in Peters et al. (2019).



**Figure S14.** Left: Different observations of the  $\text{NO}_2$  surface concentrations on 14 September 2016, each at its original temporal resolution to reveal short-term variations. Coloured areas behind the lines indicate the specified uncertainties. Right: RMSD values obtained from a kind of autocorrelation analysis over the whole campaign (night times excluded). For each observation, the RMSD between the original and a temporally shifted signal is calculated. The temporal shift (bottom horizontal axis) was varied between 0 and 4 hours. The temporal shift was roughly converted to its spatial equivalent by multiplication with the average observed wind speed in the surface layer ( $\approx 5 \text{ m/s}$ ), yielding the top horizontal axis.



**Figure S15.** Variation in the  $\text{NO}_2$  Vis dSCDs with different azimuth viewing directions at  $4^\circ$  elevation, as observed by the IMPACT imaging MAX-DOAS (Peters et al., 2019). Around local noon this variation is largest, exceeding a factor of 5. The time is UTC.

## S8 Further details on the comparison results

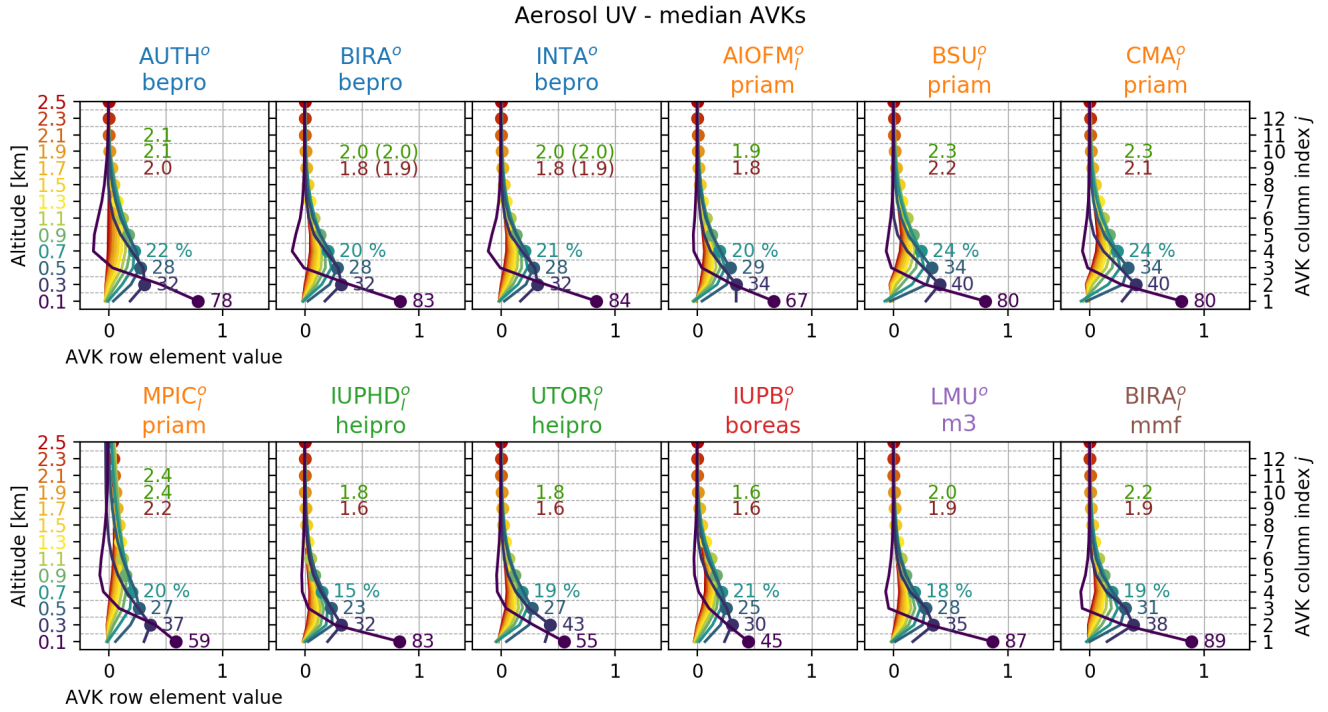
### S8.1 AVKs of individual participants

Figures S36 to S40 show the averaging kernels (AVKs) and retrieved degrees of freedom of signal (DOFS) of each participant for aerosol UV. For explanation of colours and symbols please refer to main text Sect. 3.1. The DOFS values in brackets were calculated considering valid data only.

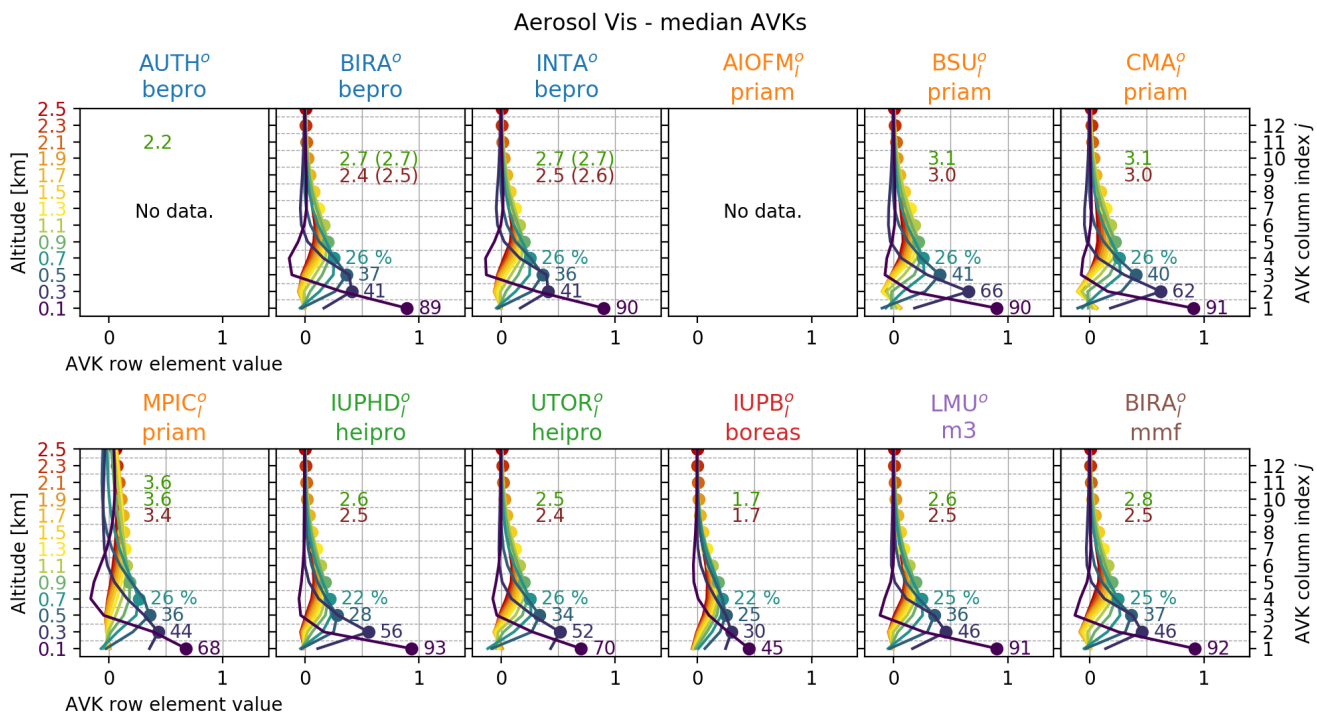
### S8.2 Profile deviation statistics

Figures S21 to S25 show statistics on the observed differences in the retrieved profiles for all five species. The plots on the left compare the retrieved profiles of individual participants  $x$  to the median MAX-DOAS profiles  $\bar{x}$ . While the vertical axes represent altitude, the horizontal axes depicts the difference  $x - \bar{x}$ . The coloured boxes indicate the 25% – 75% percentile, whiskers are 5% – 95%. Black dots indicate the mean value. For each layer there are boxplots for clear-sky (green) and cloudy conditions (red). Note that for aerosol there are two different horizontal axes defined for the two cloud conditions: the green scale at the bottom and the red scale on the top of each plot. Only valid data (flagged) was considered. For aerosol and  $\text{NO}_2$  a plot on the very right shows statistics of the difference of supporting measurements  $x_{anc}$  (lidar/ radiosonde for  $\text{NO}_2$ , sun photometer scaled ceilometer for aerosol) to the median  $\bar{x}$ , hence  $x_{anc} - \bar{x}$ . The numbers in all the plots show RMSD deviation of the three lowest (most sensitive) layers. Dashed lines indicate the median retrieval uncertainty as specified by the participants.

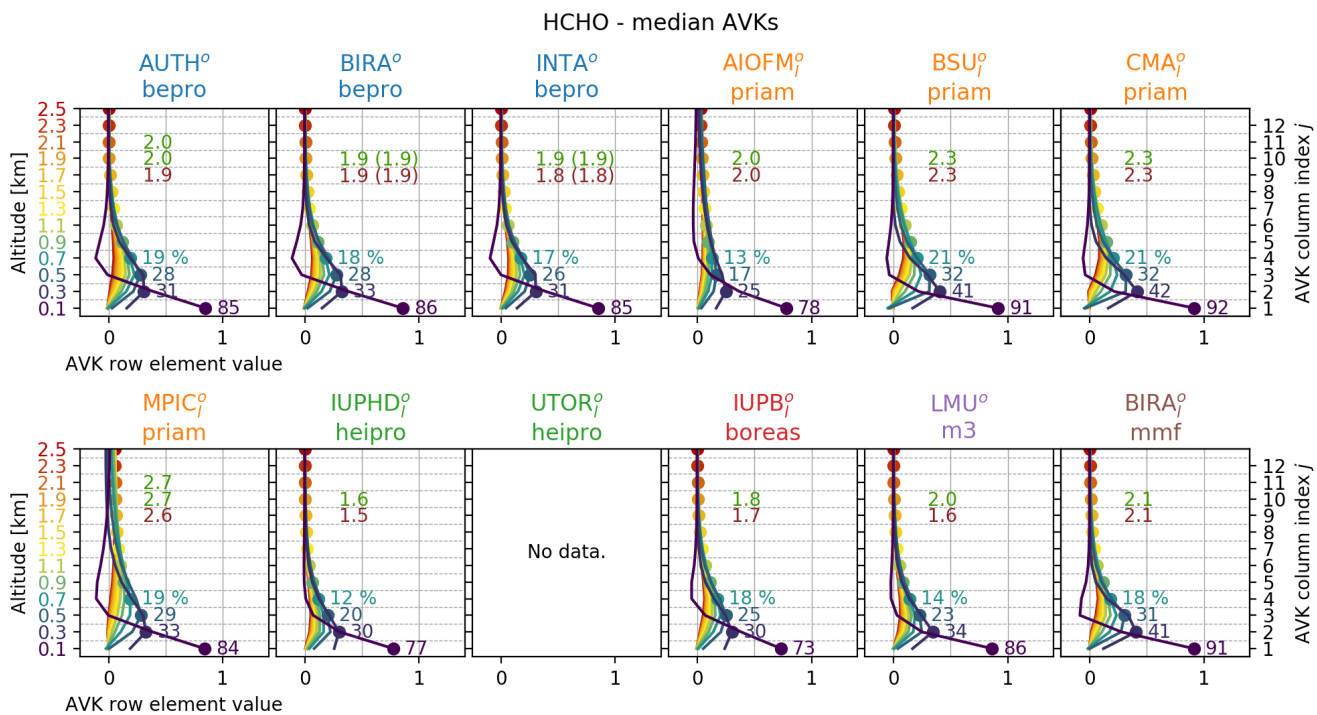




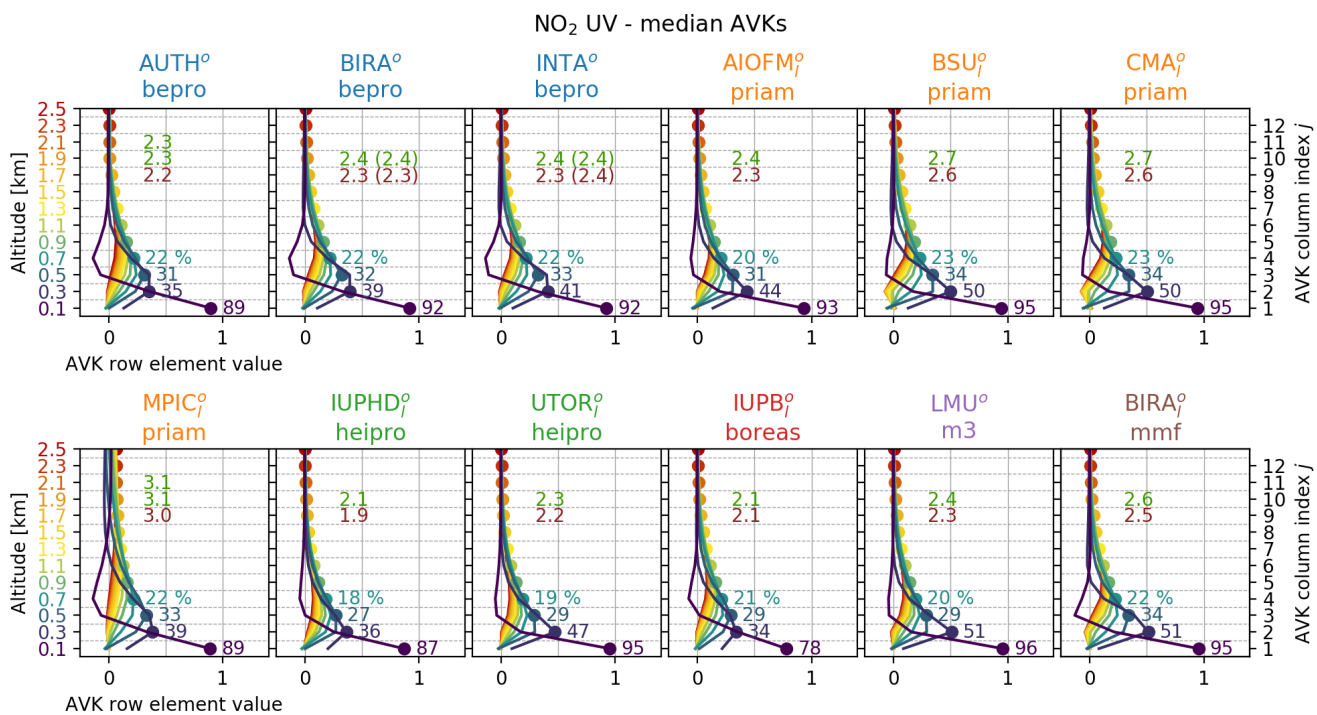
**Figure S16.** Mean averaging kernels for Aerosol UV for each participant. Coloured values at AVK peaks show the amount of retrieved information on the respective layer in percent. "DOFS" numbers are given for clear-sky (green) and cloudy (red) conditions. Values in brackets are DOFS including flagging.



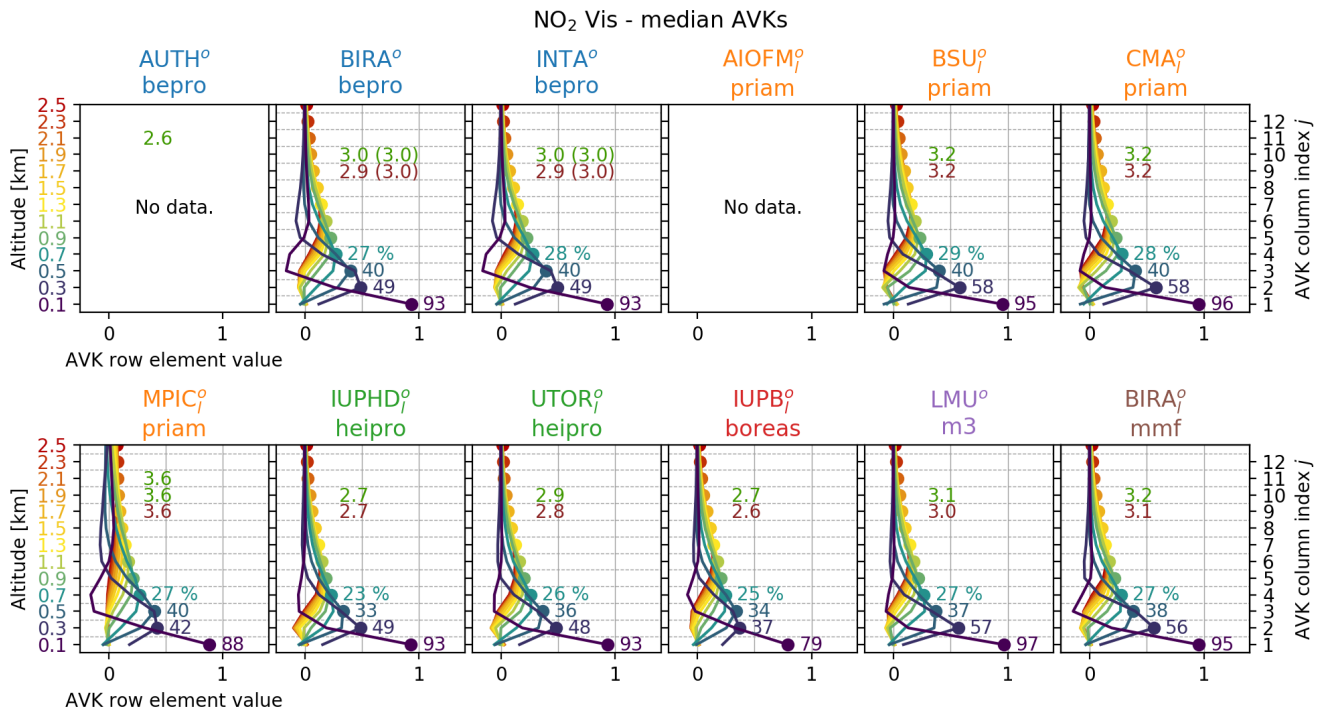
**Figure S17.** Mean averaging kernels for Aerosol Vis for each participant. Description of Fig. S36 applies.



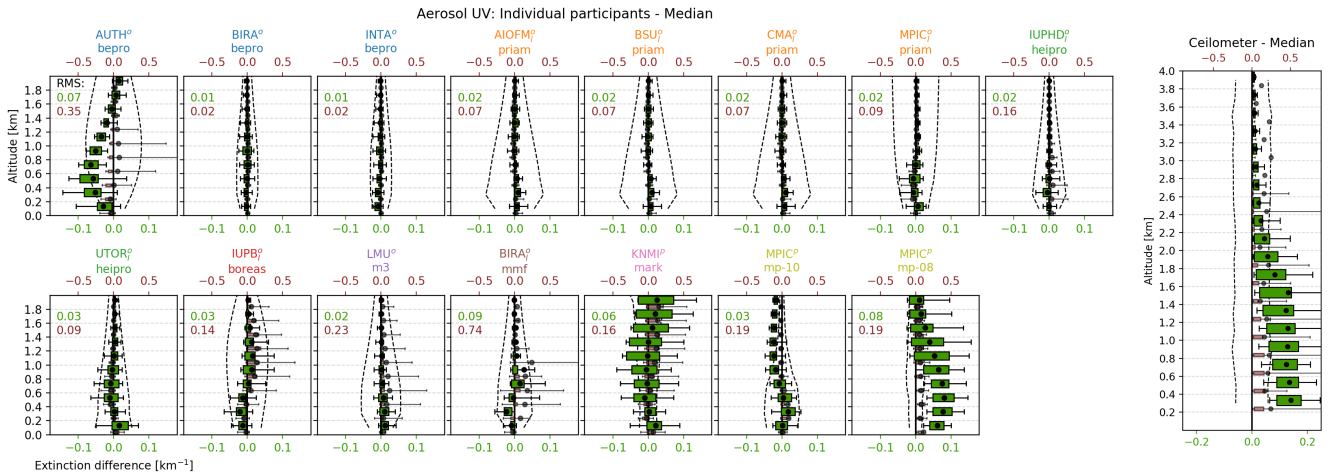
**Figure S18.** Mean averaging kernels for HCHO for each participant. Description of Fig. S36 applies.



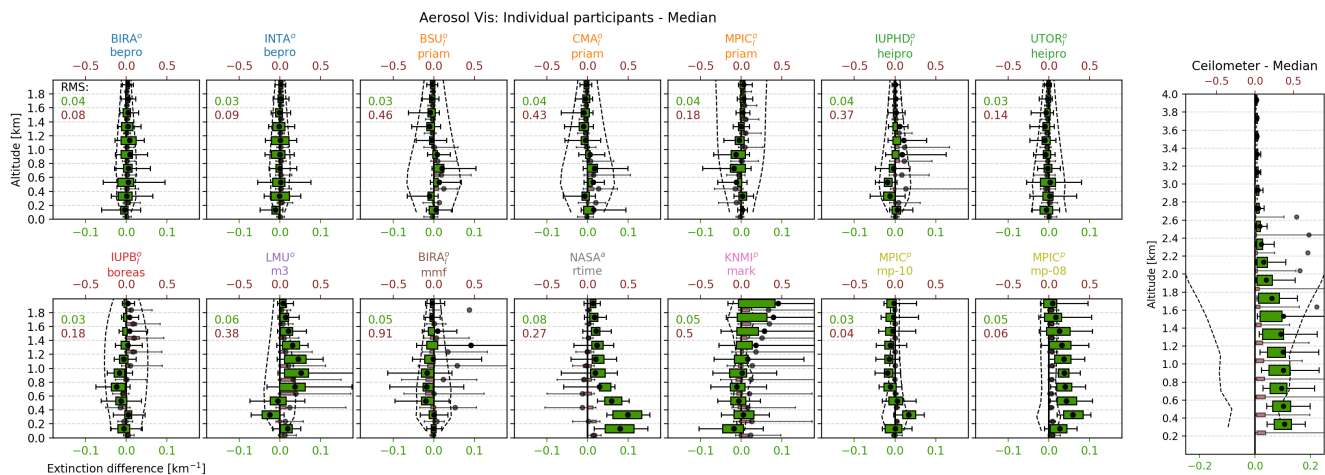
**Figure S19.** Mean averaging kernels for NO<sub>2</sub> UV for each participant. Description of Fig. S36 applies.



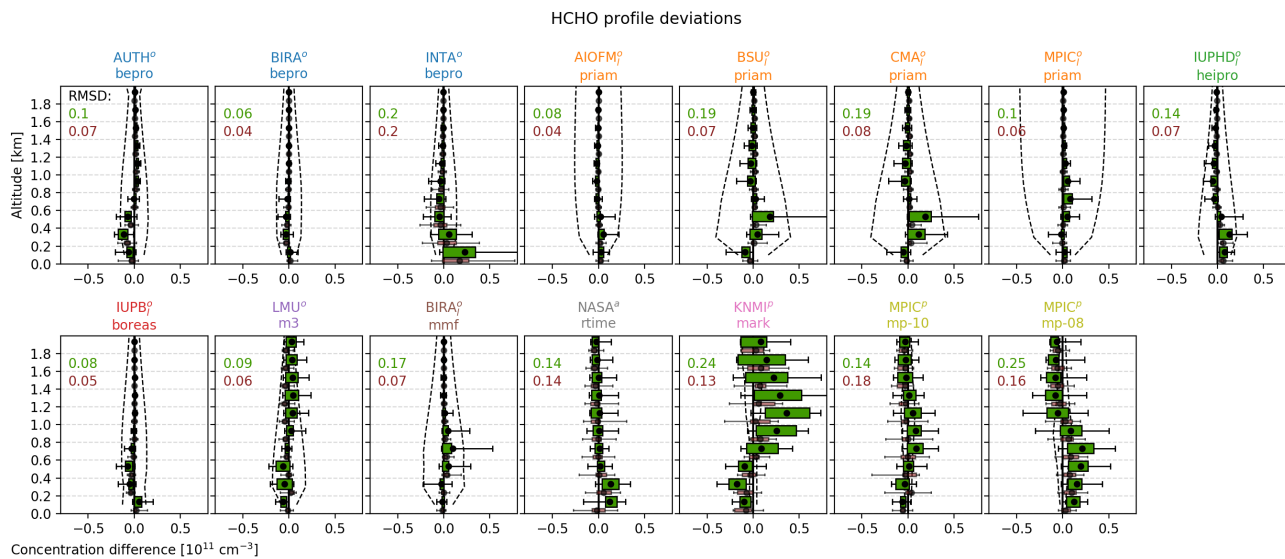
**Figure S20.** Mean averaging kernels for NO<sub>2</sub> Vis for each participant. Description of Fig. S36 applies.



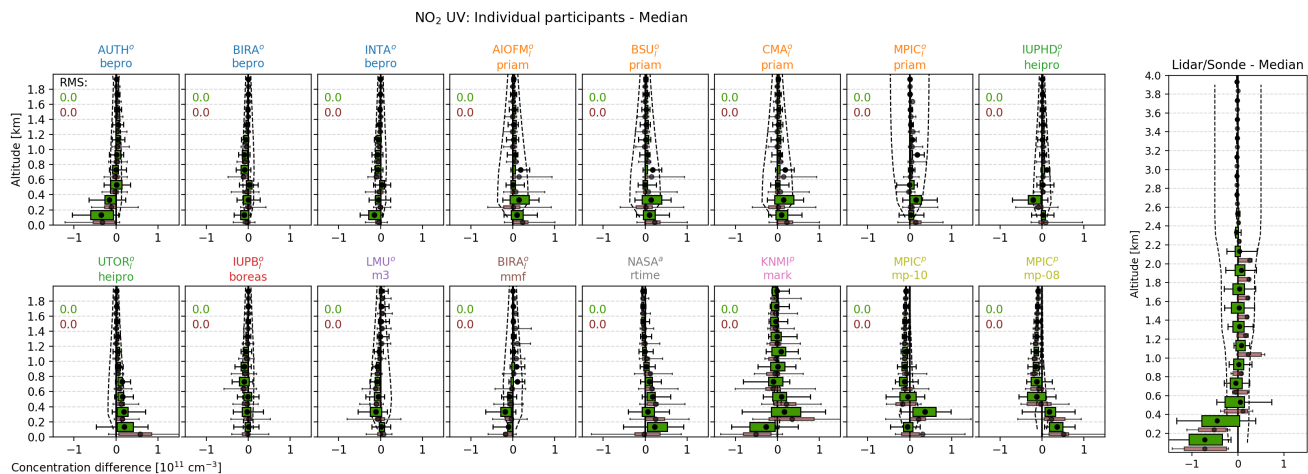
**Figure S21.** Left: Flagged deviations for Aerosol UV of individual participants from the median profiles. Note, that there are different x-scales (on top and bottom of the plot) for different cloud conditions. Right: Deviation of the AOT scaled ceilometer backscatter signal to the median. The numbers in the plots indicate RMSD values for clear sky (green) and cloudy (red) conditions.



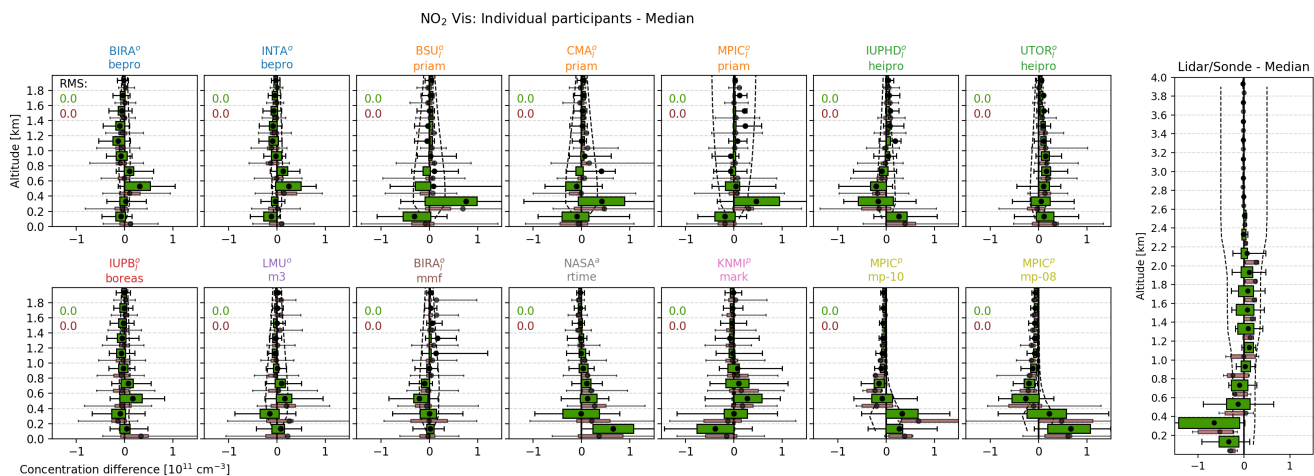
**Figure S22.** Flagged deviations for Aerosol Vis of individual participants from the median profiles. Description of Fig. S21 applies.



**Figure S23.** Flagged deviations for HCHO of individual participants from the median profiles. The description of Fig. S21 applies but for HCHO, there is no independent reference profile available.



**Figure S24.** Flagged deviations for NO<sub>2</sub> UV of individual participants from the median profiles. The description of Fig. S21 applies. On the right, deviation of the median retrieved profiles from the few available NO<sub>2</sub> lidar and sonde profiles are shown.



**Figure S25.** Flagged deviations for NO<sub>2</sub> Vis of individual participants from the median profiles. The description of Fig. S24 applies.



### S8.3 Correlation plots for AOTs, VCDs and surface concentrations

Figures S26 to S33 show the individual correlation plots for the comparisons of AOTs, VCDs and surface concentrations as summarized in Sect. 3.4, 3.5 and 3.6 in the main text, respectively. The colours indicate cloud conditions: clear-sky (green) and cloudy conditions (red). Transparent markers represent data points flagged as invalid. The small grey bars indicate uncertainties in the measurement. For the median values, the bars show the standard deviation among the participants (valid data only). Dashed lines in the correlation plots represent the ideal 1:1 line. Correlations were performed separately against MAX-DOAS median values and supporting observations. For AOTs a third correlation is shown for the partial AOT (with PAC applied).

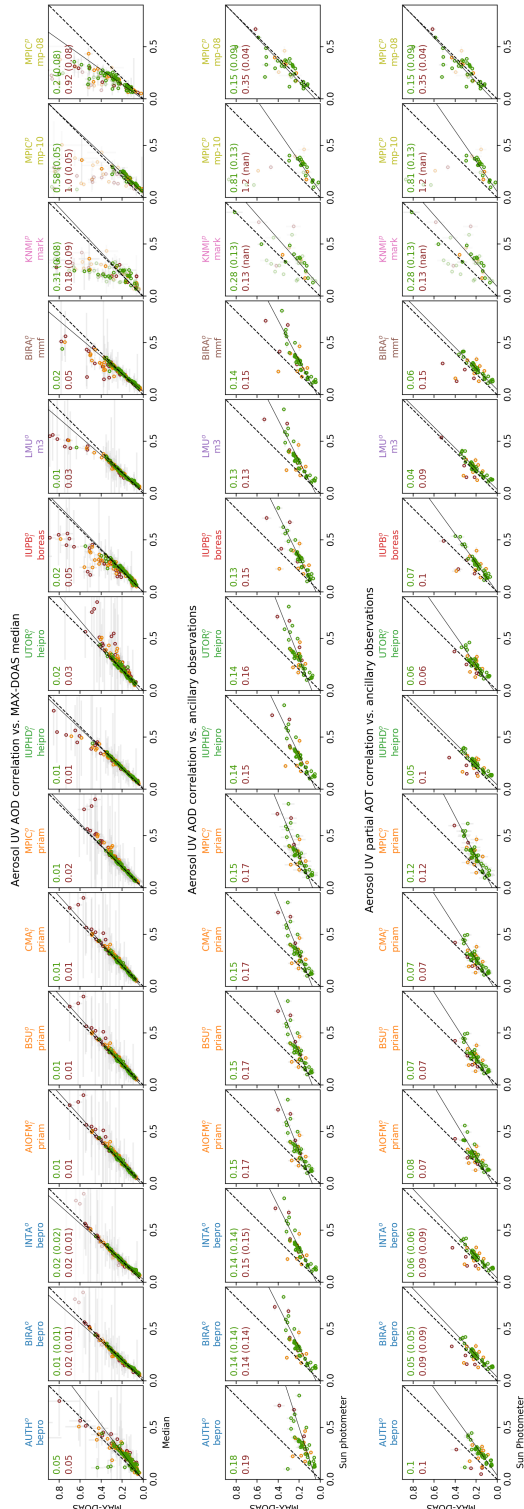


Figure S26. Correlation of Aerosol UV AOTs. Complementary to main text Sect. 3.4



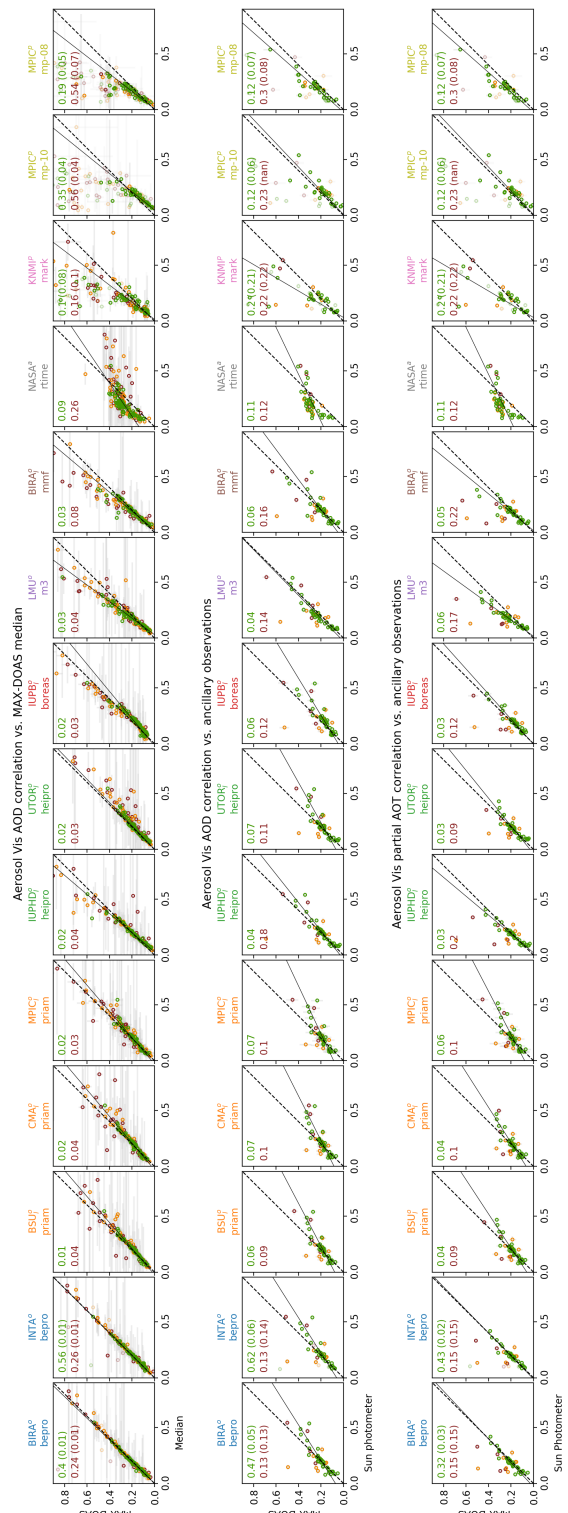


Figure S27. Correlation of Aerosol Vis AOTs. Complementary to main text Sect. 3.4

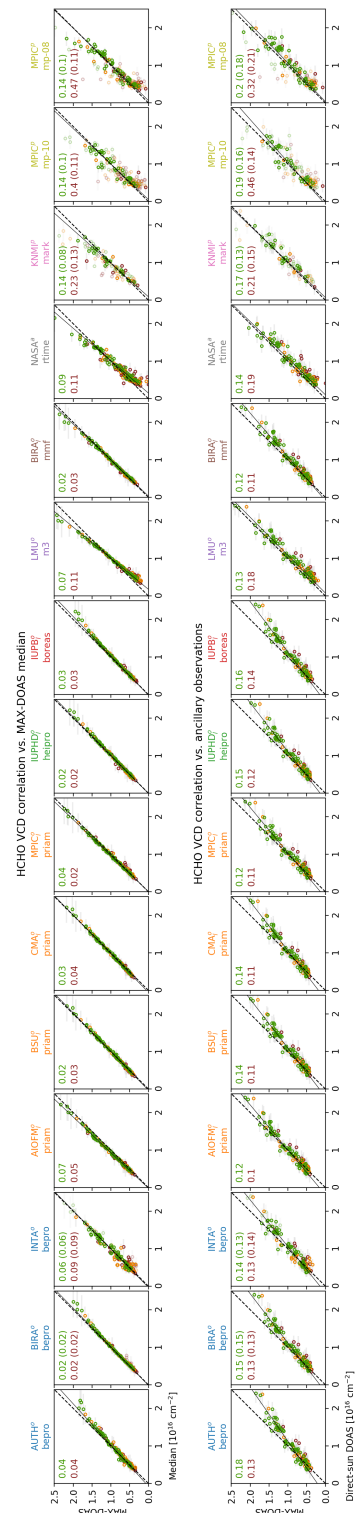


Figure S28. HCHO VCD correlation plots. Complementary to main text Sect. 3.5

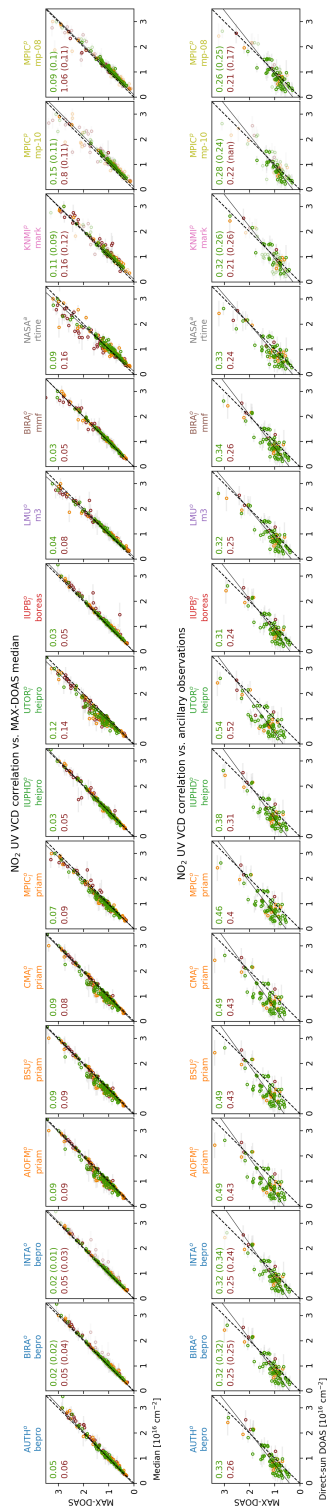


Figure S29. NO<sub>2</sub> UV VCD correlation plots. Complementary to main text Sect. 3.5

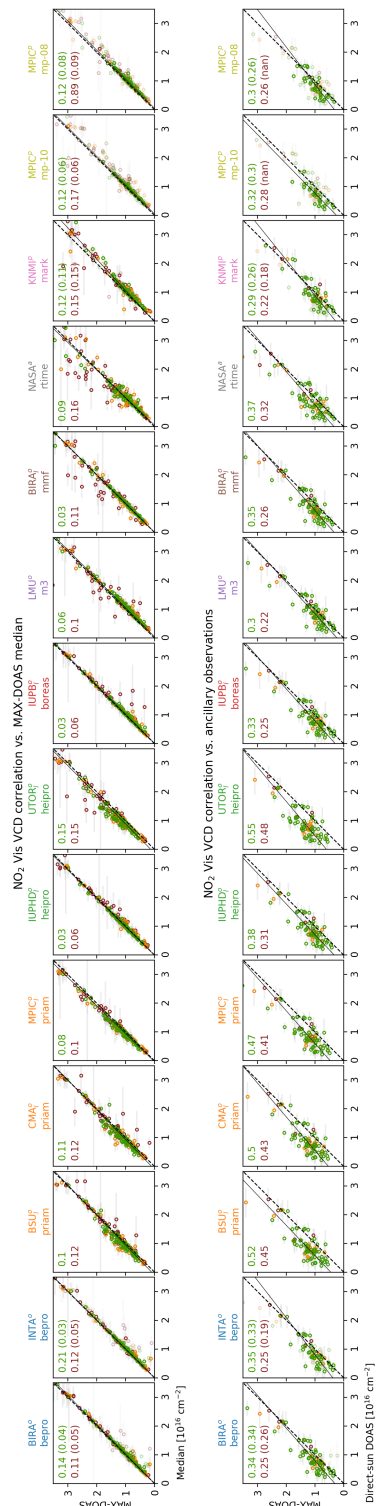
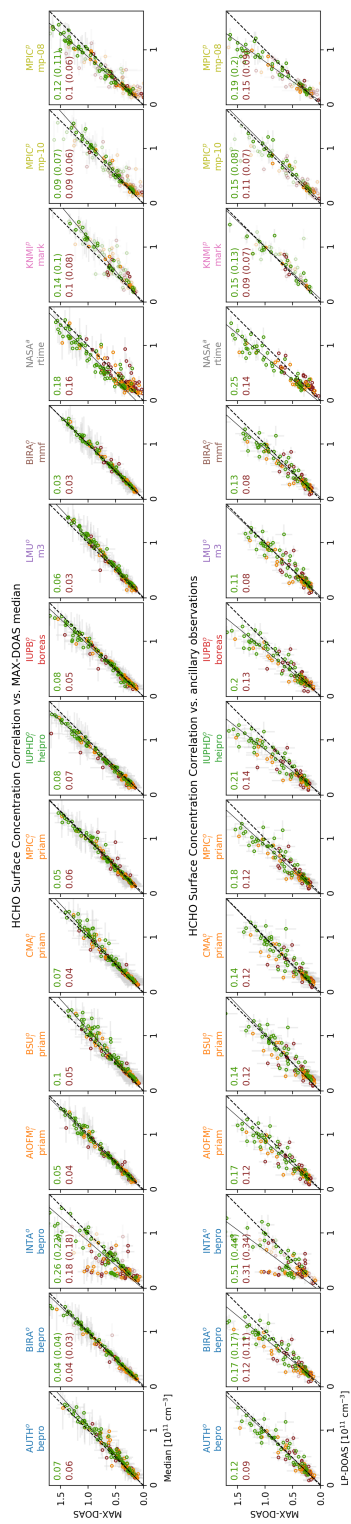
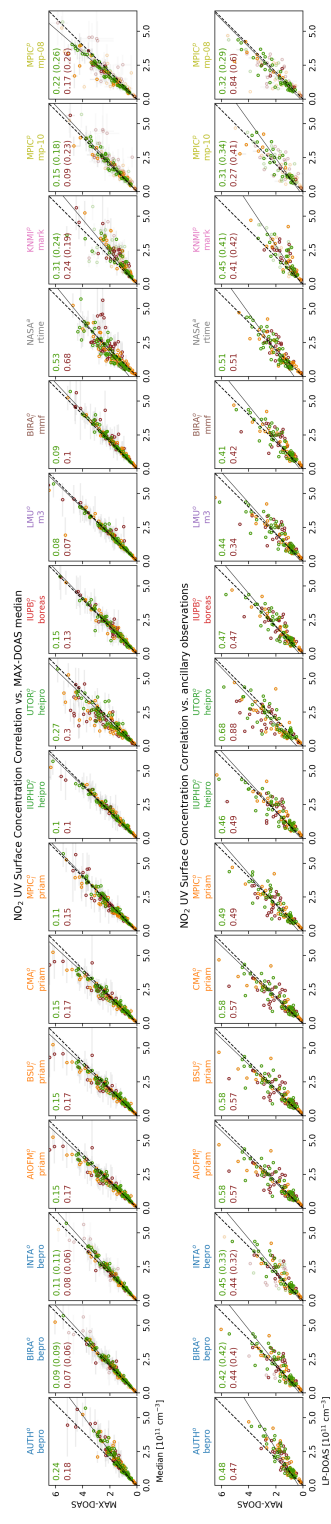


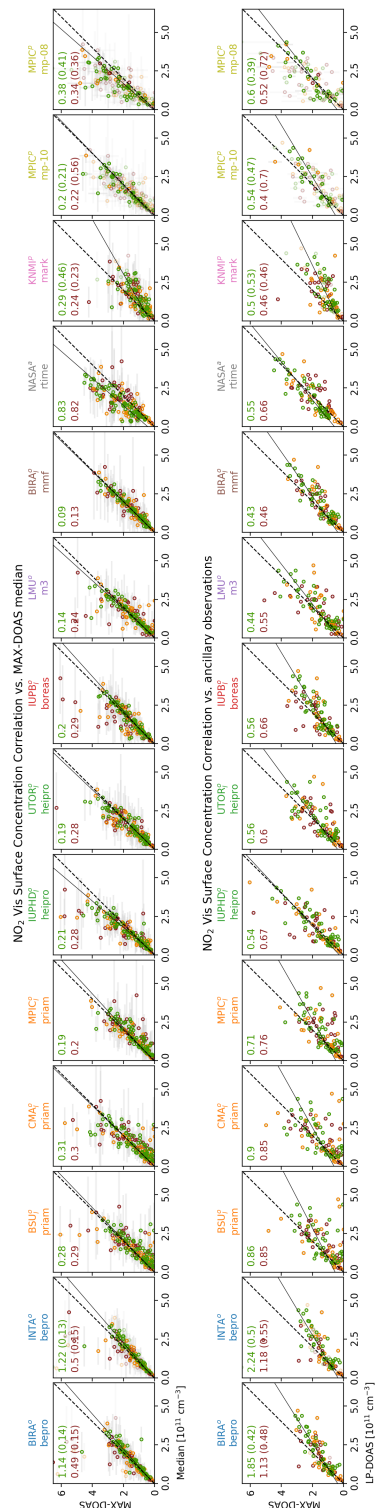
Figure S30. NO<sub>2</sub> Vis VCD correlation plots. Complementary to main text Sect. 3.5



**Figure S31.** HCHO surface concentration correlation plots. Complementary to main text Sect. 3.6



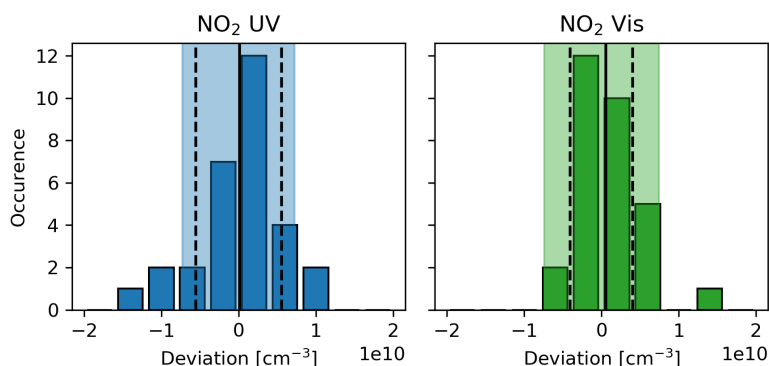
**Figure S32.** NO<sub>2</sub> UV surface concentration correlation plots. Complementary to main text Sect. 3.6



**Figure S33.** NO<sub>2</sub> Vis surface concentration correlation plots. Complementary to main text Sect. 3.6

## S9 Impact of smoothing on surface concentration

For NO<sub>2</sub> the impact of smoothing effects on the surface concentration retrieved by OEM algorithms can be estimated from profiles of the NO<sub>2</sub> lidar and radiosondes. Each profile is smoothed according to Eq. (10) in the main text and the difference in surface concentration between the smoothed and the unsmoothed profile is calculated. Figure S34 shows histograms of the calculated differences. An estimate of the impact of smoothing on the retrieval results is actually provided by the OEM retrievals themselves as the "smoothing error". The specified smoothing errors are also indicated in Fig. S34 and indeed slightly larger than the standard deviation observed in in this test.



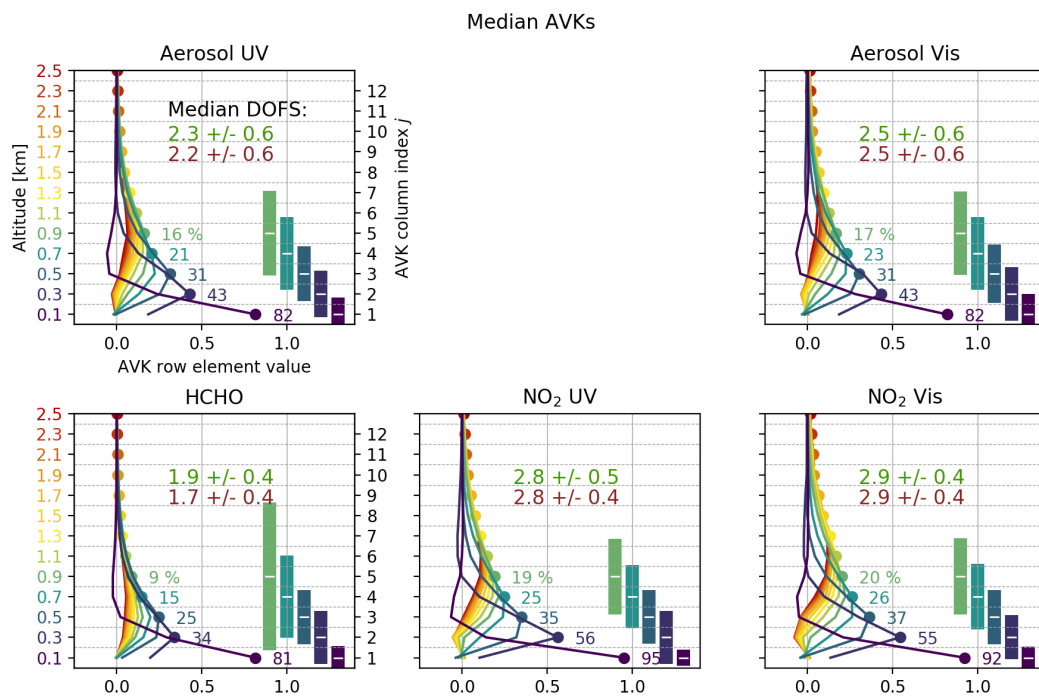
**Figure S34.** Histograms of the observed deviations in surface concentration between raw and smoothed lidar/ radiosonde NO<sub>2</sub> profiles. Solid and dashed lines indicate mean value and standard deviation, respectively. Coloured areas represent the median smoothing errors as specified by the OEM retrievals, which is in good agreement with the deviations obtained from the supporting NO<sub>2</sub> profiles.

## S10 Participant's own dSCD comparison results

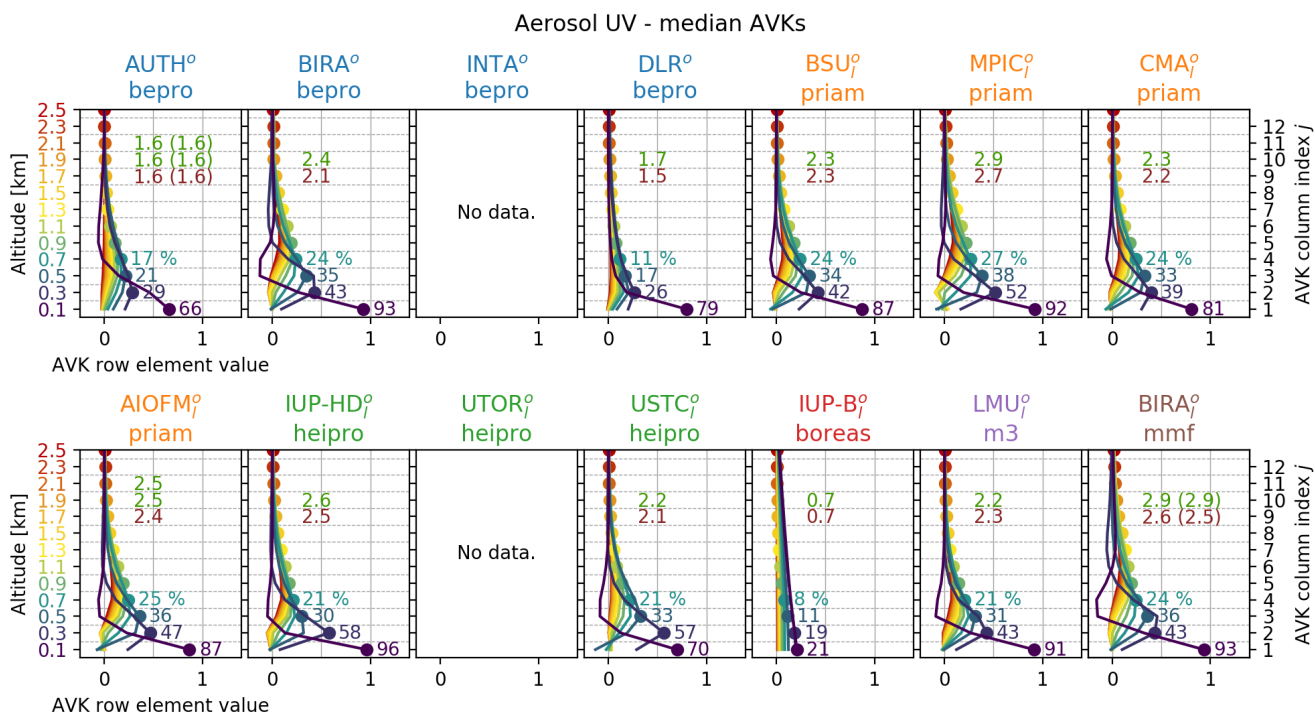
This section shows the comparison results for the case where each participant uses dSCDs measured with his own instrument. Evaluation and plots are fully equivalent to Sect. 3 in the main text. With DLR (German Aerospace Center, Oberpfaffenhofen, Germany, marked by blue squares) and USTC (University of Science and Technology of China, Hefei, China, marked by green squares), two other participants were included here, retrieving profiles with bePRO and HEIPRO, respectively. Gaps in the data are mostly related to instrument malfunction during the campaign. Further, not all instruments covered the spectral range to detect all desired species and the corresponding participants therefore do not appear in the respective plots.

### 15 S10.1 Information content

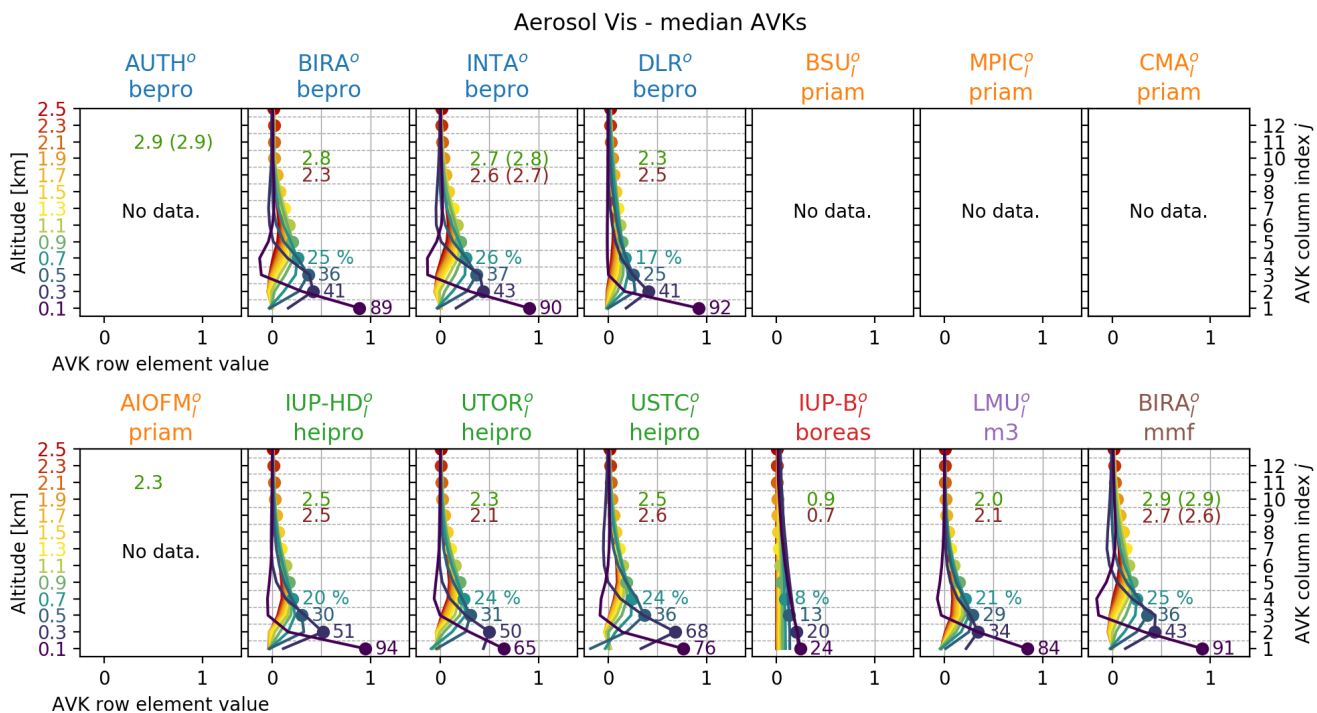
Information on the averaging kernels and DOFS. This section is equivalent to Sect. S10.1 in the main text and Sect. S8.1 in this supplement, respectively.



**Figure S35.** Average AVKs for the retrieved species (median over participants, mean over time). Each altitude and corresponding AVK line are associated with a colour, which is defined by the colour of the corresponding altitude-axis label. The dots mark the AVK diagonal elements. The number next to the dots show the exact value in percent, which corresponds to the amount of retrieved information on the respective layer. In the upper right of each panel, the DOFS (median among institutes, average over time) are given for clear-sky (green) and cloudy conditions (red). The vertical bars indicate the vertical resolution (the "spread" as defined by equation 3.23 in Rodgers (2000)) for the five lowest layers.

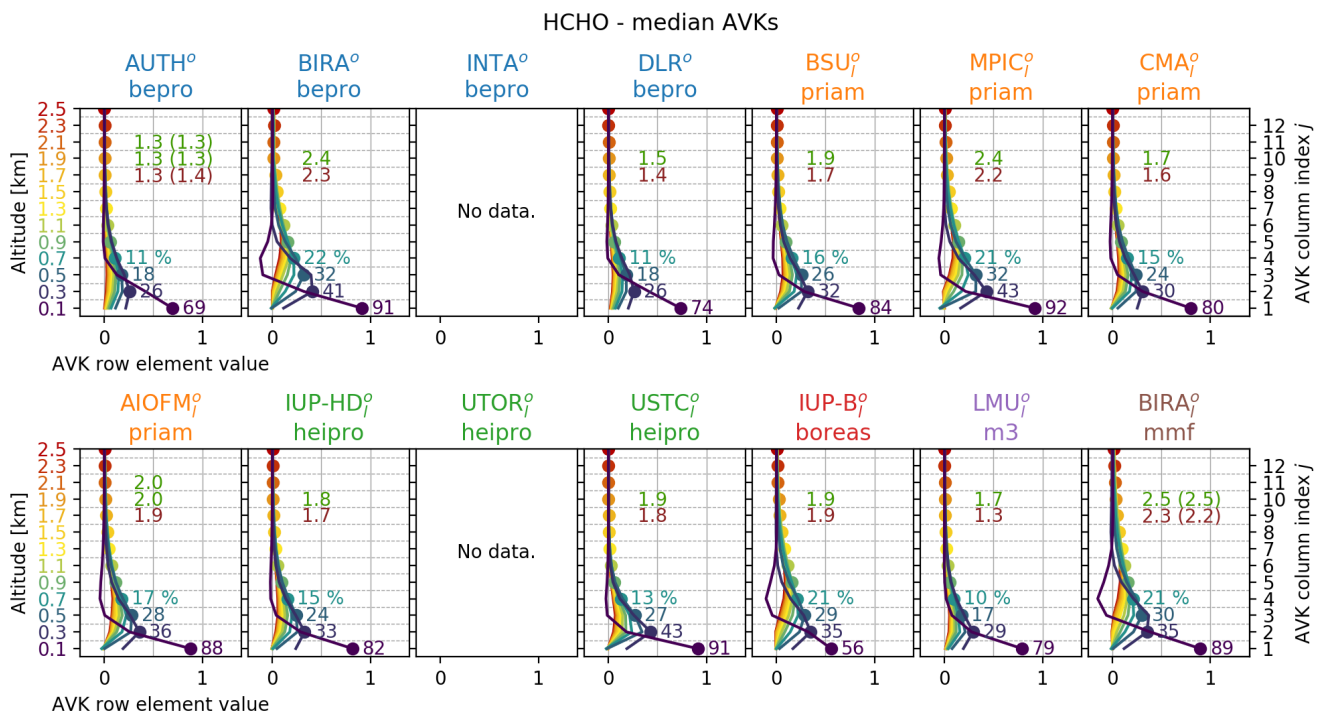


**Figure S36.** Mean averaging kernels for Aerosol UV for each participant retrieving from their own dSCDs. Coloured values at AVK peaks show the amount of retrieved information on the respective layer in percent. "DOFS" numbers are given for clear-sky (green) and cloudy (red) conditions. Values in brackets are DOFS including flagging.

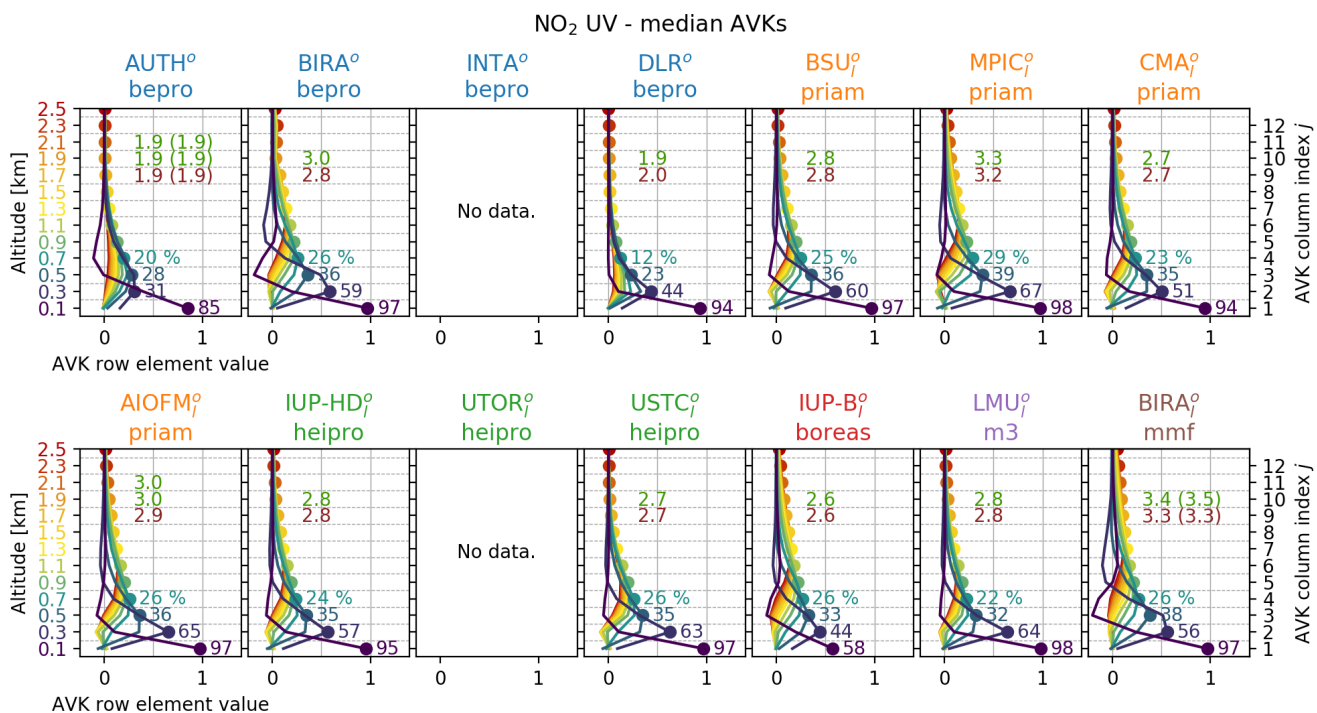


**Figure S37.** Mean averaging kernels for Aerosol Vis for each participant. Description of Fig. S36 applies.

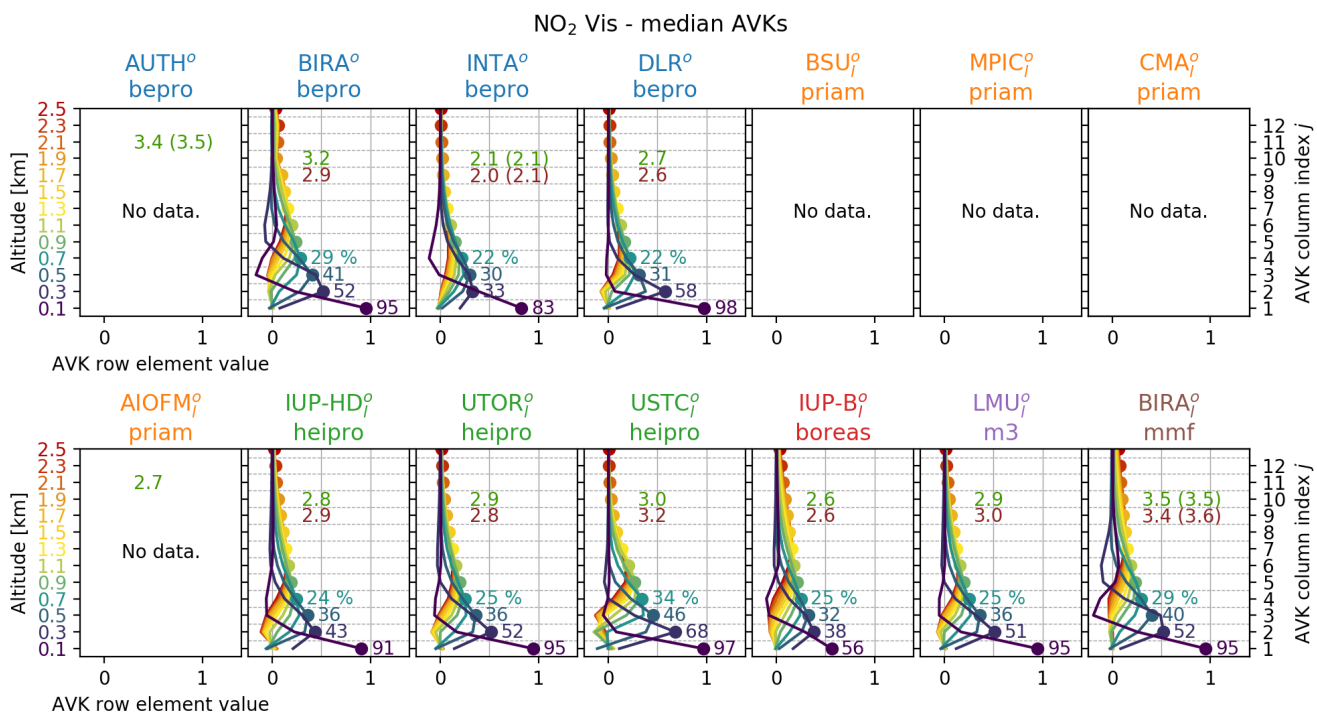




**Figure S38.** Mean averaging kernels for HCHO for each participant. Description of Fig. S36 applies.



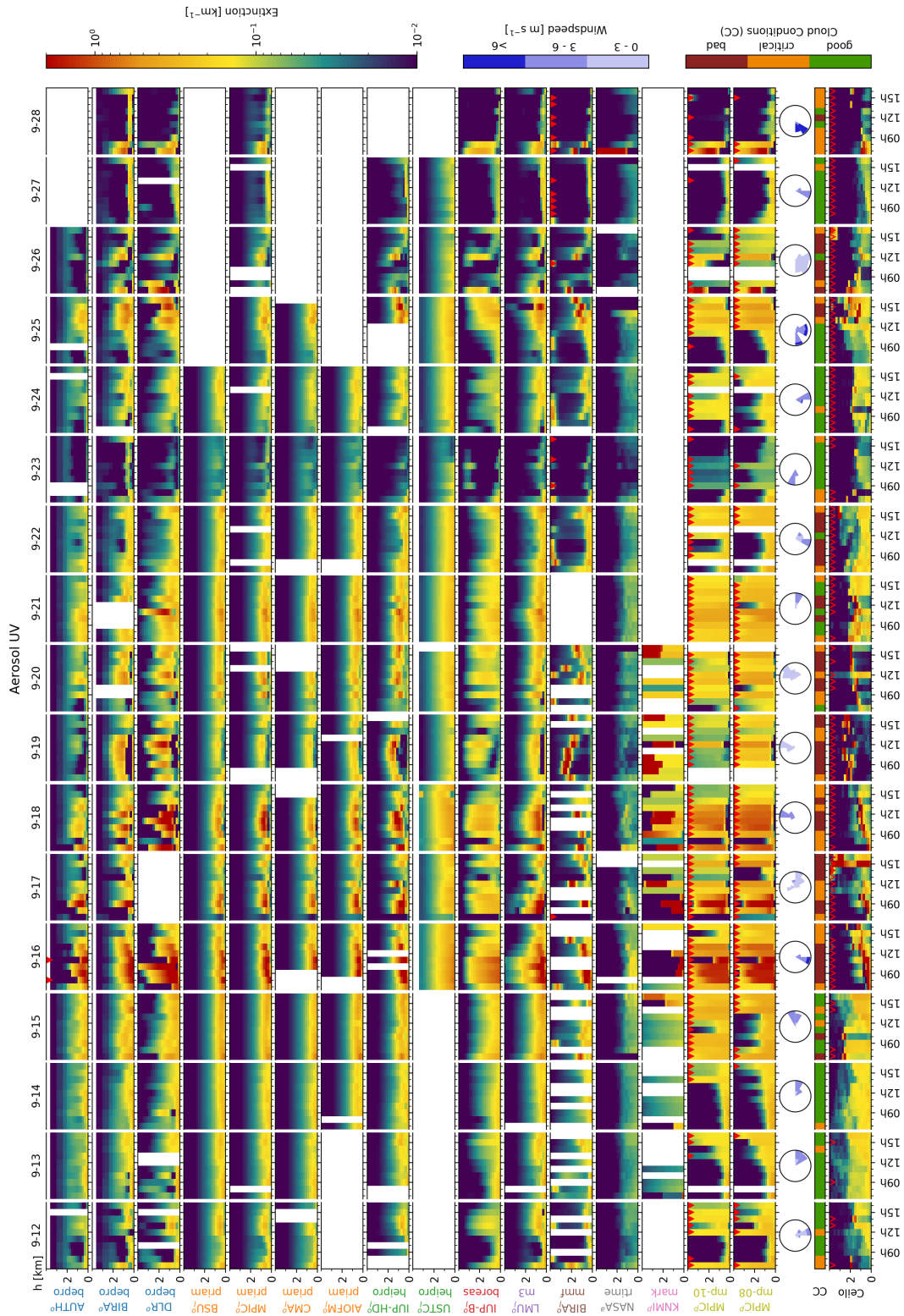
**Figure S39.** Mean averaging kernels for NO<sub>2</sub> UV for each participant. Description of Fig. S36 applies.



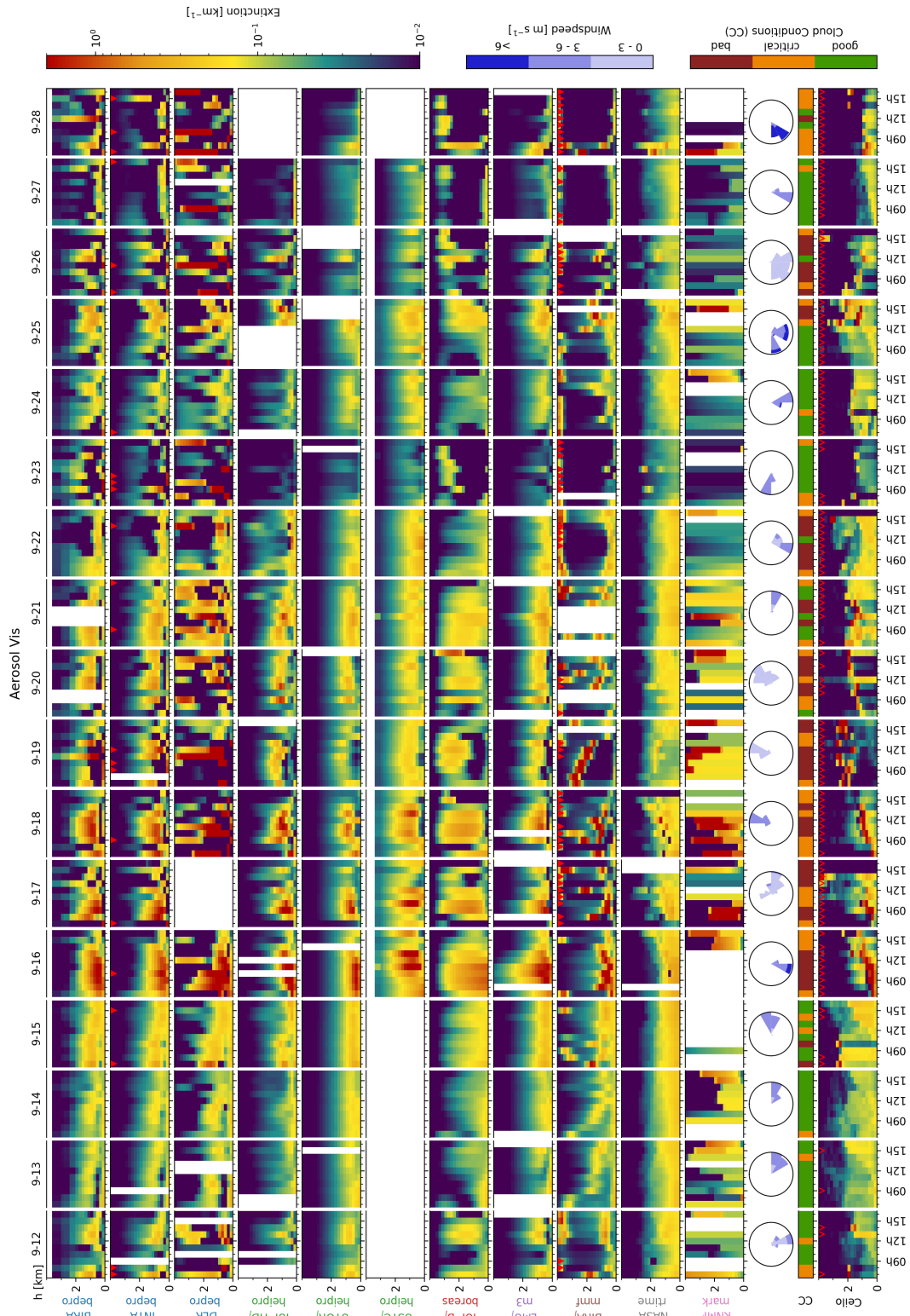
**Figure S40.** Mean averaging kernels for NO<sub>2</sub> Vis for each participant. Description of Fig. S36 applies.

## **S10.2 Overview plots**

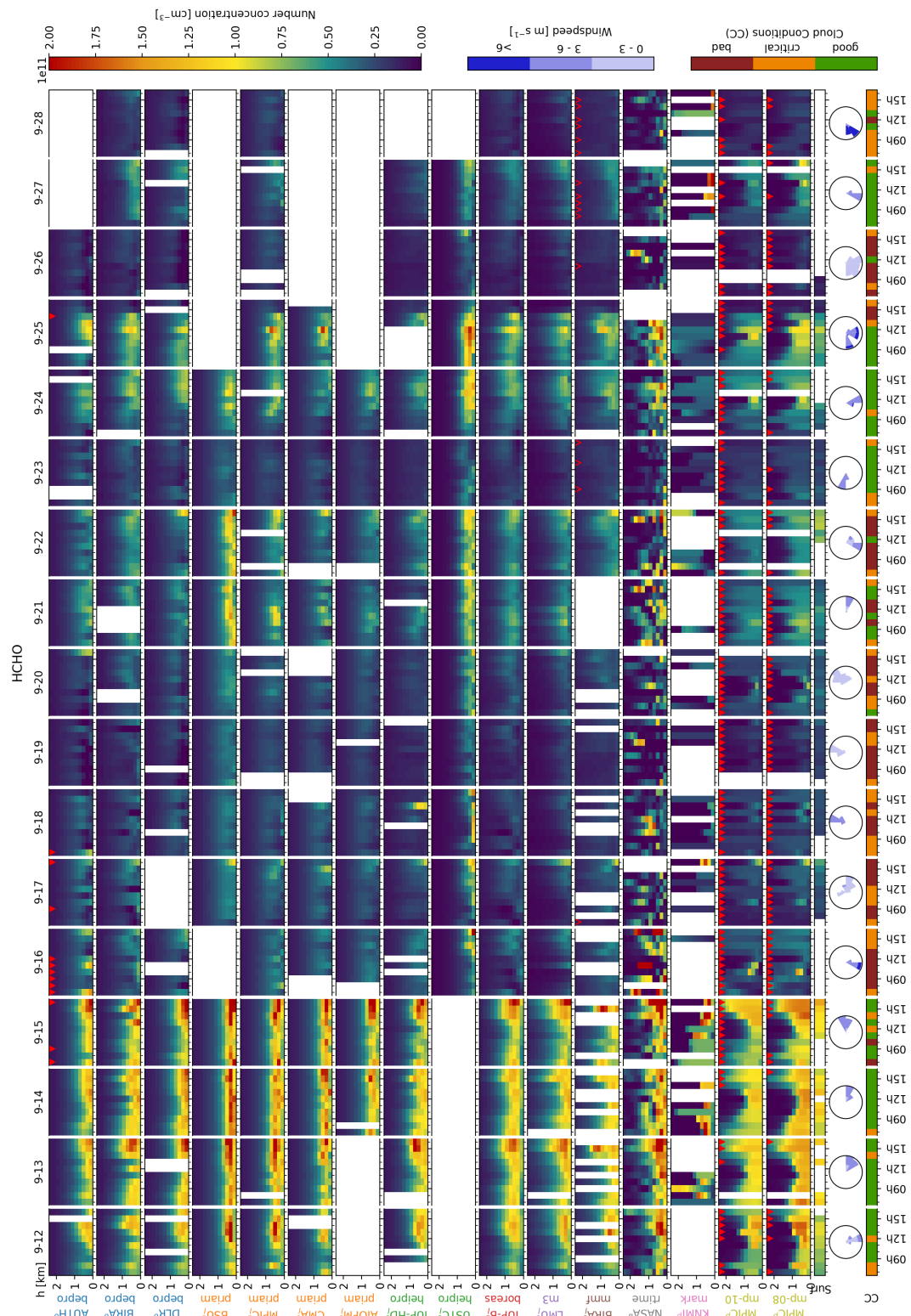
This section is equivalent to Sect. 3.2 in the main text.



**Figure S41.** Aerosol UV extinction profiles retrieved from the participant's own dSCDs. The lowest row shows AOT scaled ceilometer backscatter profiles, calculated as described in Sect. S5.1. Backscatter profiles, which were scaled from MAX-DOAS AOTs (and which are therefore not fully independent) are marked by red triangles. Maximum extinction values reach  $20 \text{ km}^{-1}$ , exceeding the colour scale.



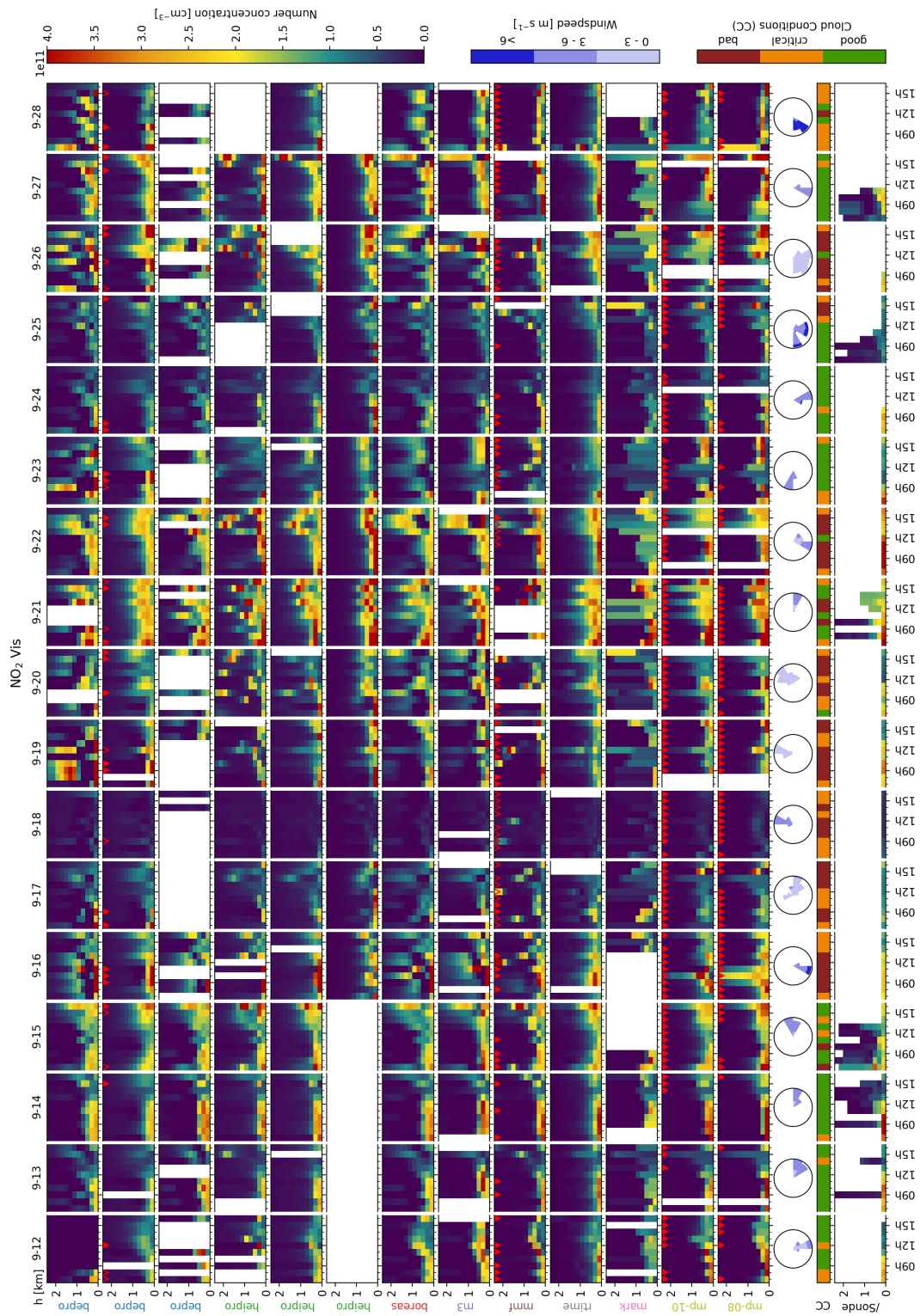
**Figure S42.** Aerosol Vis extinction profiles retrieved from the participant's own dSCDs. The lowest row shows AOT scaled ceilometer backscatter profiles, calculated as described in Sect. S5.1. Backscatter profiles, which were scaled from MAX-DOAS AOTs (and which are therefore not fully independent) are marked by red triangles. Maximum extinction values reach  $20\text{km}^{-1}$ , exceeding the colour scale.



**Figure S43.** HCHO concentration profiles retrieved from the participant's own dSCDs. The "Surf" -row shows LP-DOAS surface concentrations.



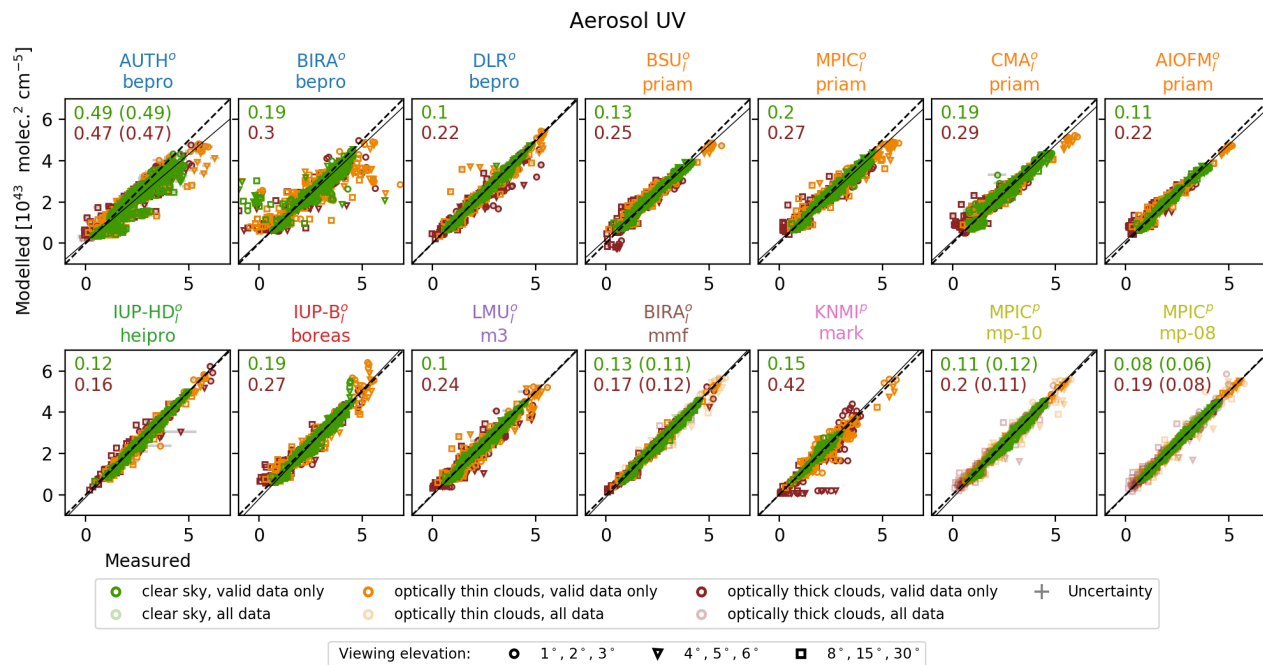




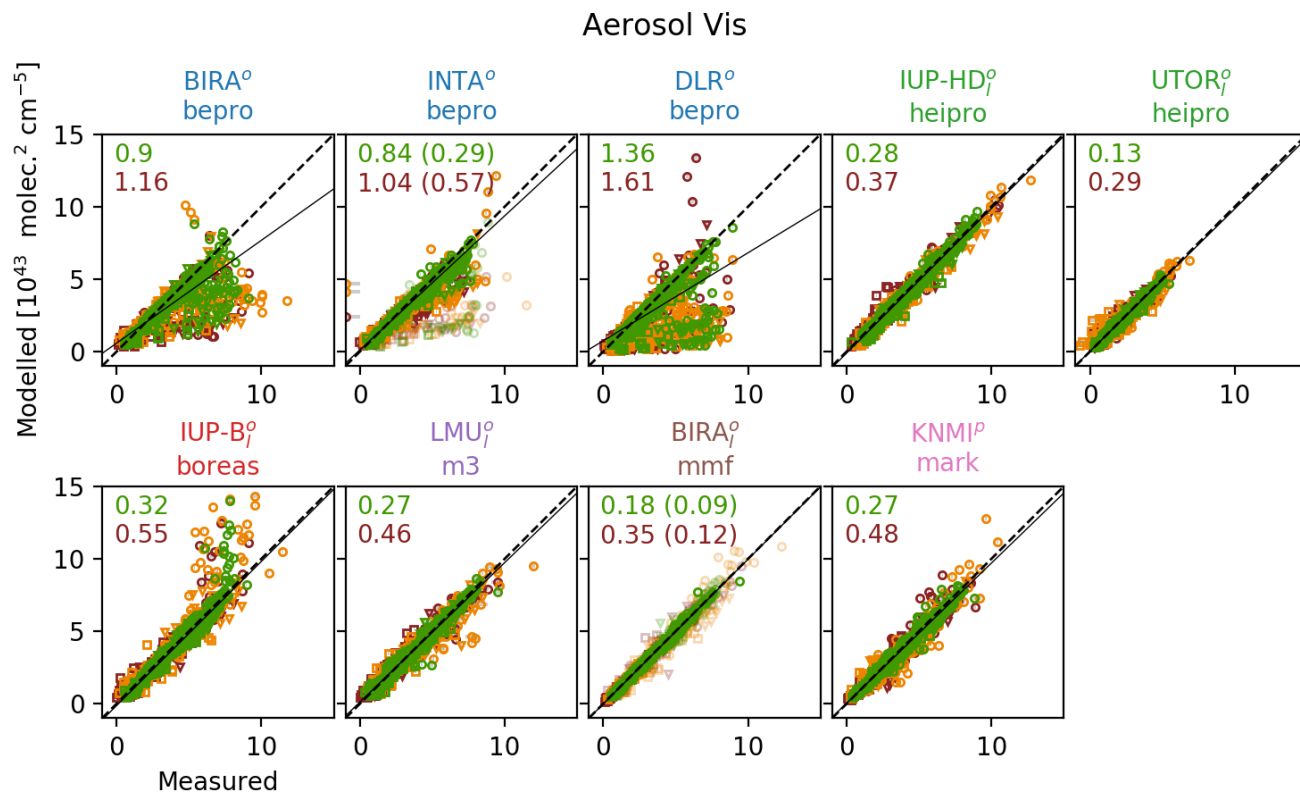
**Figure S45.**  $\text{NO}_2$  Vis concentration profiles retrieved from the participant's own dSCDs. The lowest row shows a combined dataset of  $\text{NO}_2$  lidar, radiosonde, LP-DOAS and tower in-situ data. Redundant surface concentration measurements were averaged.

### S10.3 Modelled and measured dSCDs

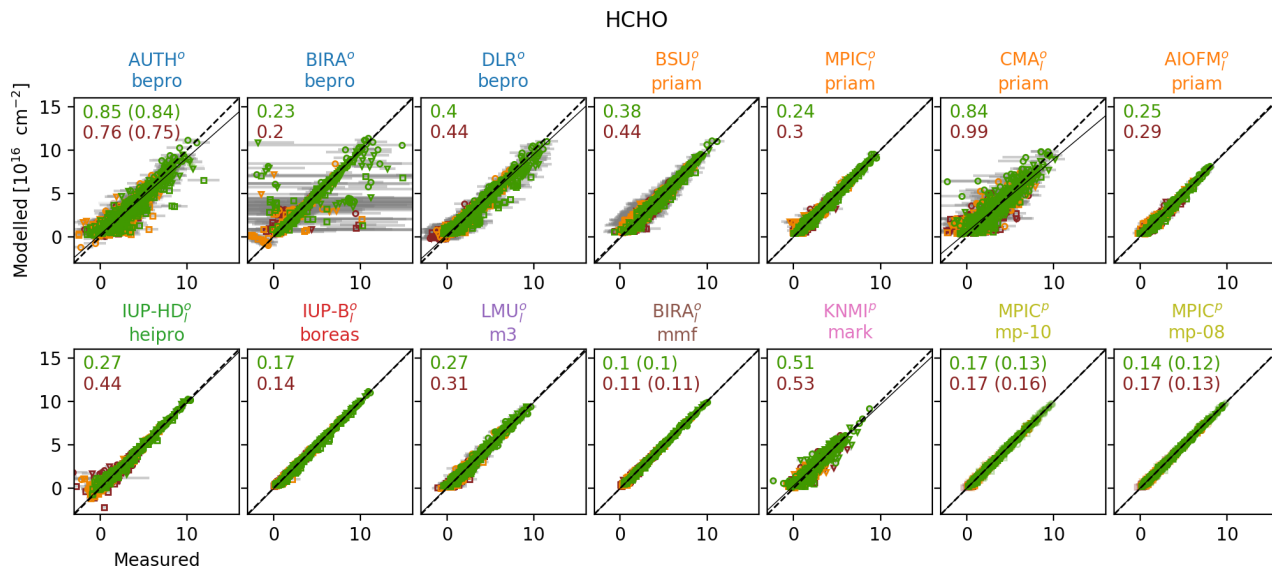
This section is equivalent to Sect. S10.3 in the main text.



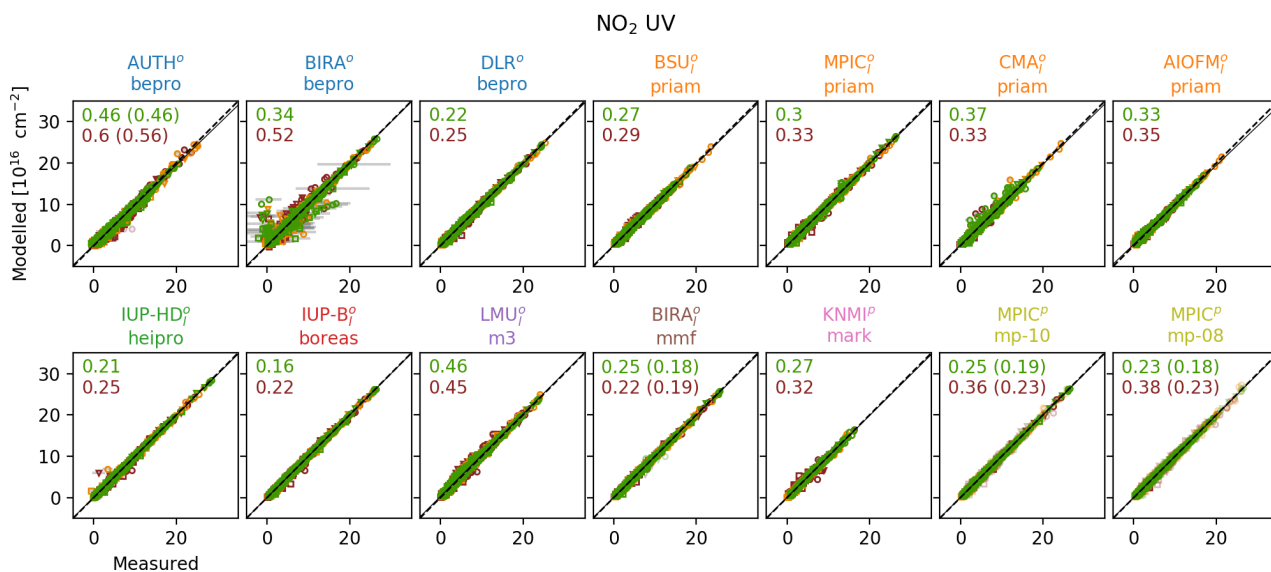
**Figure S46.**  $\text{O}_4$  UV dSCD correlation when profiles are retrieved from the participant's own dSCDs. Marker colours and marker shapes indicate the cloud conditions and viewing elevation angles, respectively. Numbers represent the measurement error weighted RMSD between measured and modelled dSCDs for clear sky (green) and cloudy (red) conditions. Values in brackets were calculated only considering valid data.



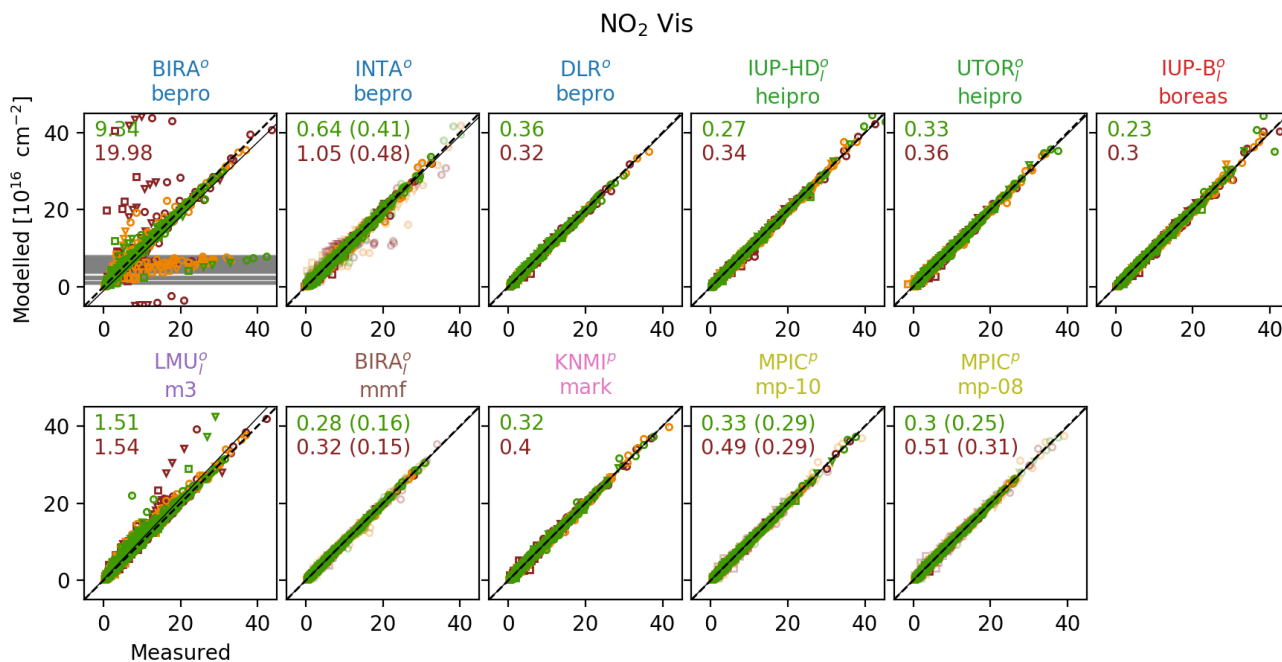
**Figure S47.** O<sub>4</sub> Vis dSCD correlation. Legends of Fig. S46 apply.



**Figure S48.** HCHO dSCD correlation. Legends of Fig. S46 apply.



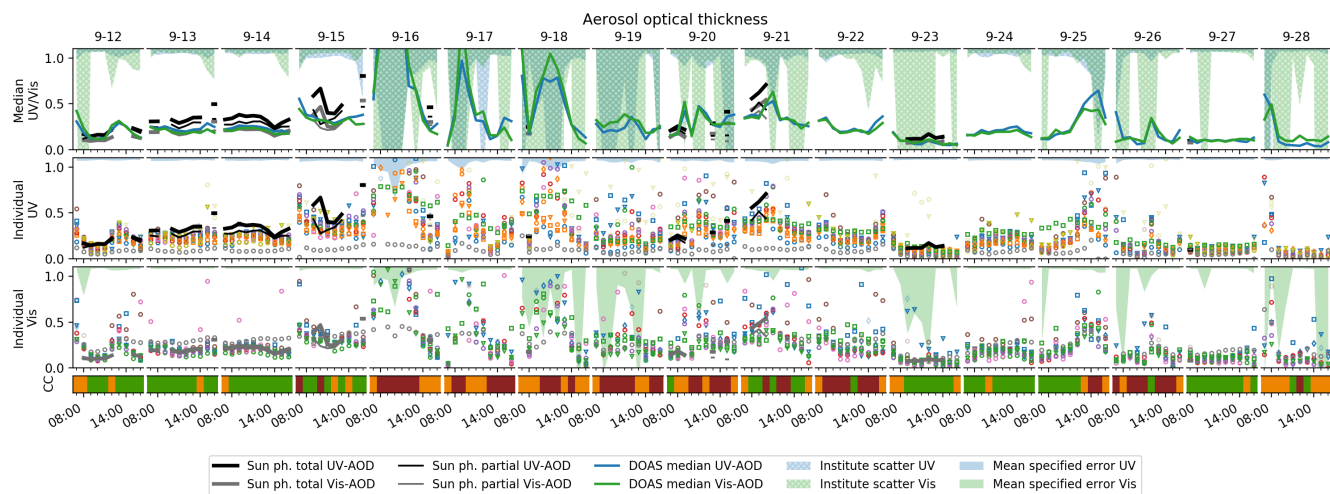
**Figure S49.** NO<sub>2</sub> UV dSCD correlation. Legends of Fig. S46 apply.



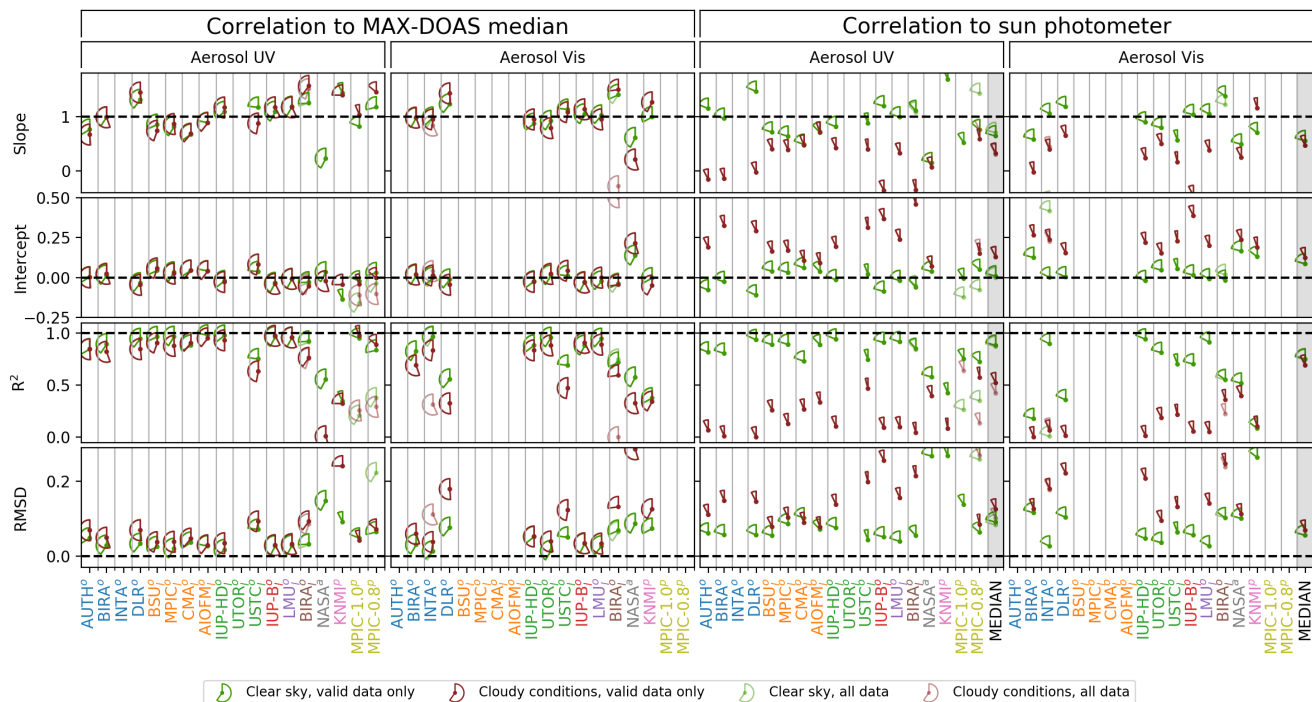
**Figure S50.** NO<sub>2</sub> Vis dSCD correlation. Legends of Fig. S46 apply.

## S10.4 Aerosol optical thickness (AOT)

This section is equivalent to Sect. S10.4 in the main text.



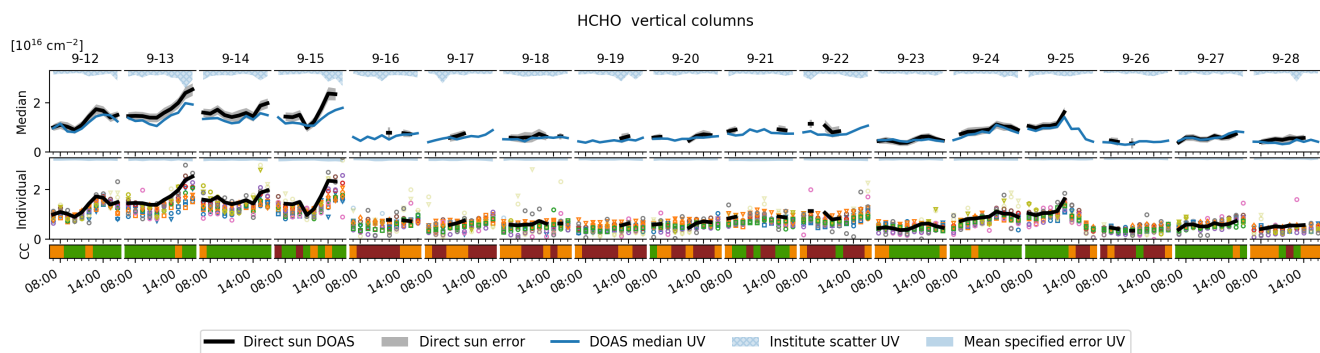
**Figure S51.** MAX-DOAS AOTs retrieved from the participant's own dSCDs in comparison to sun photometer data. Symbol and symbol colours are chosen according to Table 2 in the main text. Transparent symbols indicate data flagged as invalid. Top row: MAX-DOAS median results vs. the available supporting observations, according to the legend below the plot. Hatched areas (starting at the top of the plot) show the scattering (standard deviation) among the participants (only valid data considered). Two lower rows: Comparison of the individual participants for the two spectral retrieval ranges. Here the coloured area is the average retrieval error, as specified by the participants.



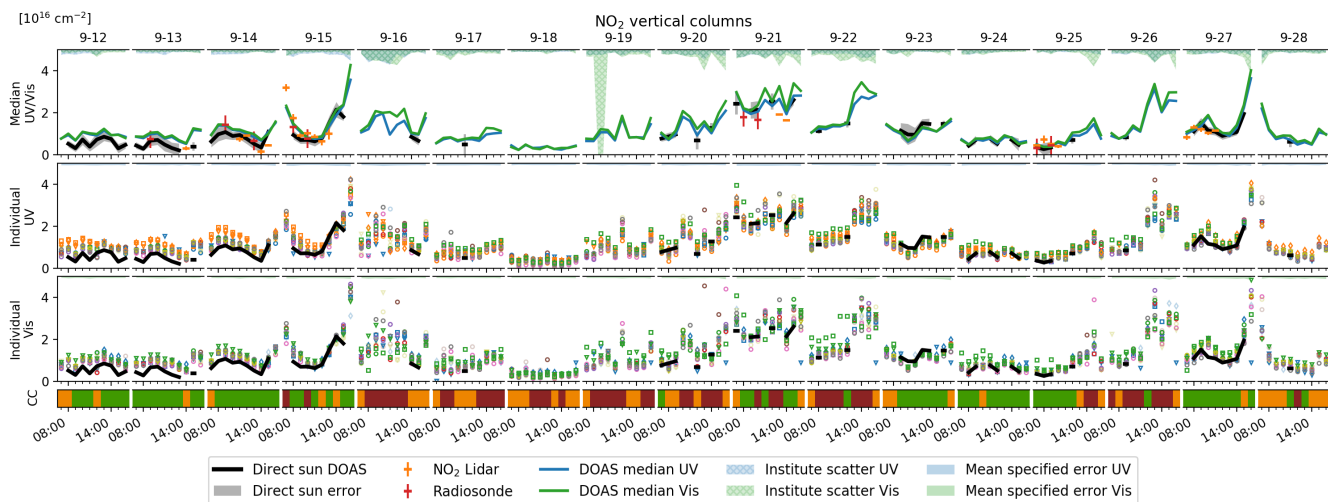
**Figure S52.** Correlation statistics for AOTs retrieved from the participant’s own dSCDs. The two left columns give an impression on the agreement among the institutes, as they show the correlation of the individual participant’s retrieved AOT against the median. The two right columns show the correlation against the sun photometer AOT (partial AOT in the case of OEM retrievals). Green and red symbols represent cloud-free and cloudy conditions, respectively. Transparent symbols represent values for all submitted data, opaque symbols only consider data points flagged as valid. The pies indicate, which fraction of the total number of profiles (170) contributed to the respective. On the right also the correlation between the MAX-DOAS median results and supporting observations are included (grey shaded columns).

### S10.5 Trace gas vertical column densities (VCDs)

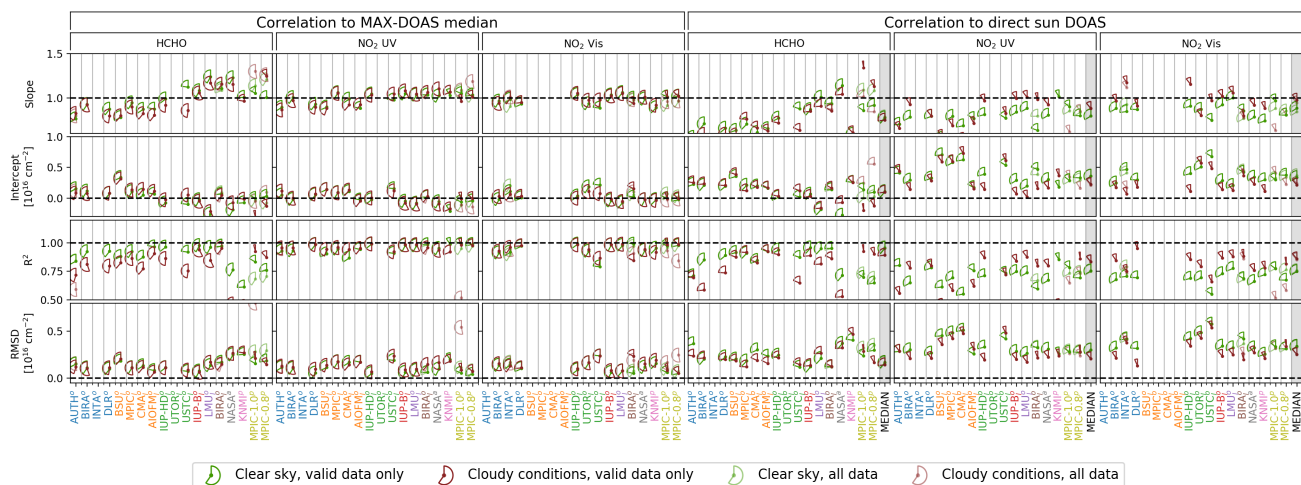
This section is equivalent to Sect. S10.5 in the main text.



**Figure S53.** Comparison of MAX-DOAS HCHO VCDs retrieved from the participant's own dSCDs vs. direct-sun DOAS, NO<sub>2</sub> lidar and radiosonde. Descriptions of Fig. S51 apply.



**Figure S54.** Comparison of MAX-DOAS NO<sub>2</sub> VCDs retrieved from the participant's own dSCDs vs. direct-sun DOAS. Descriptions of Fig. S51 apply.

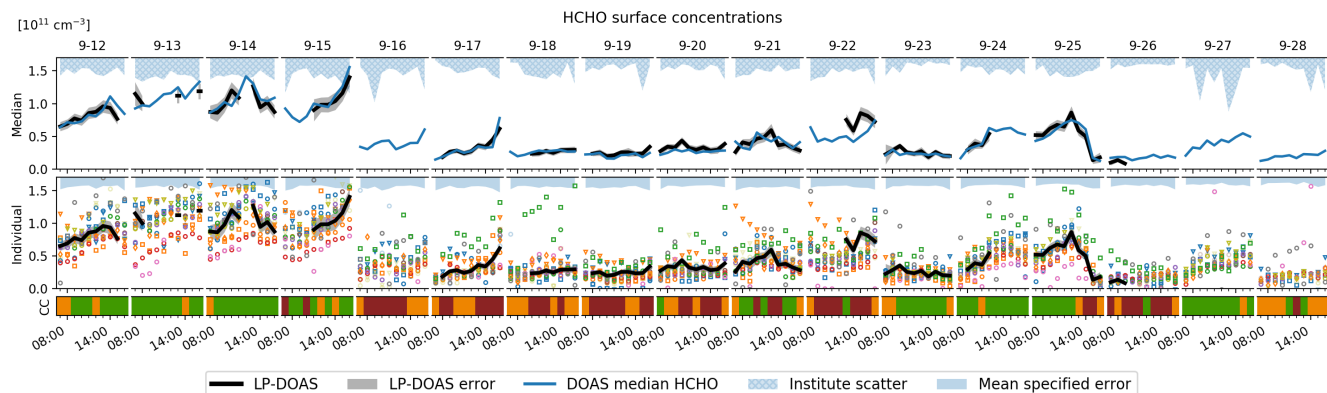


**Figure S55.** Correlation statistics of trace gas VCDs retrieved from the participant's own dSCDs. Legends and description of Fig. S52 apply.

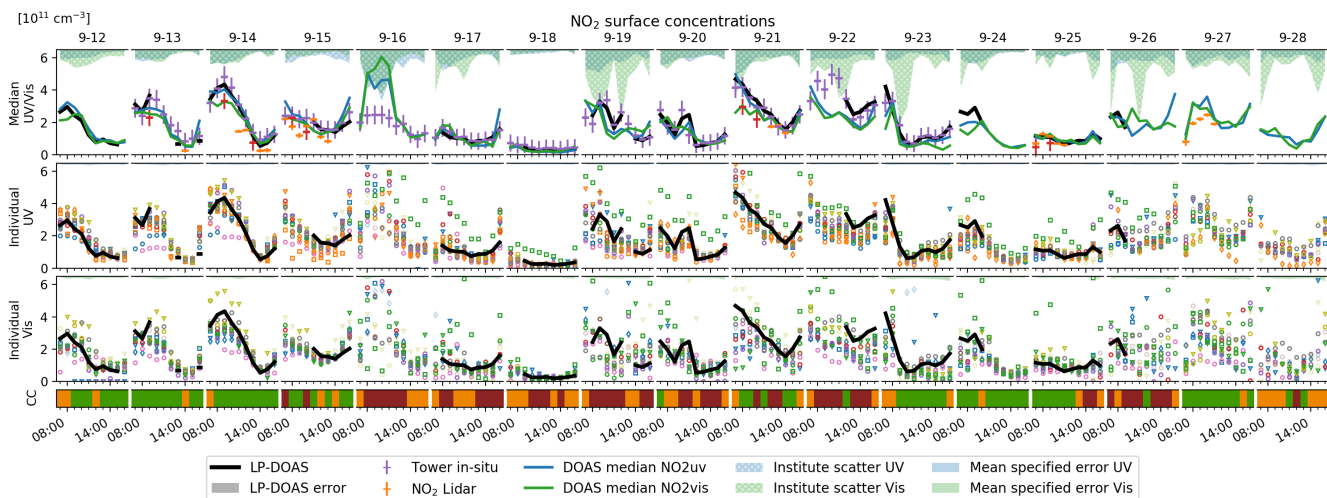


## S10.6 Trace gas surface concentrations

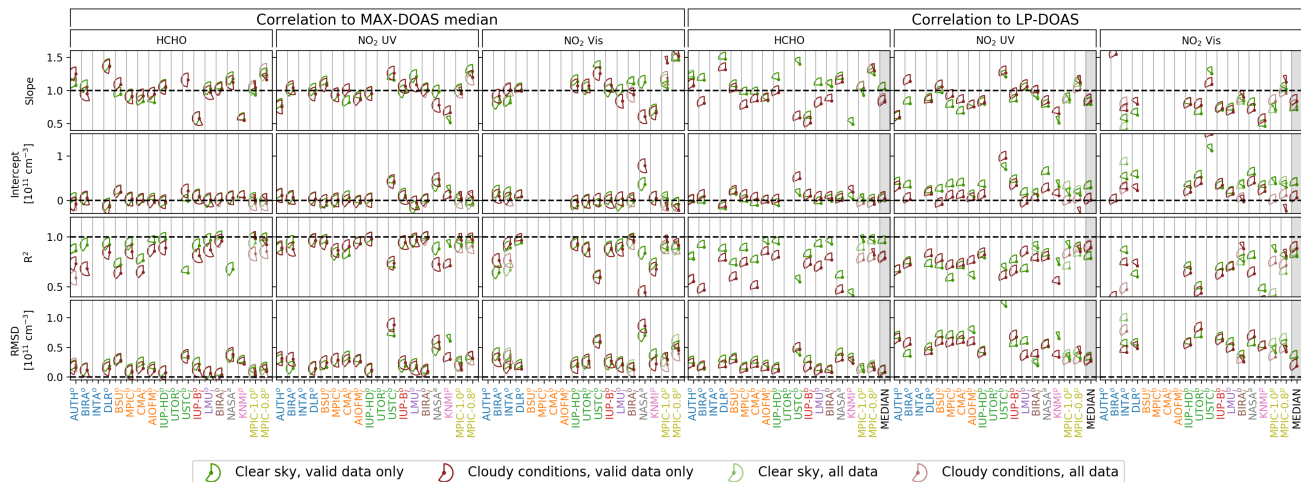
This section is equivalent to Sect. S10.6 in the main text.



**Figure S56.** Comparison of MAX-DOAS HCHO surface concentrations retrieved from the participant's own dSCDs. Description of Fig. S51 applies.



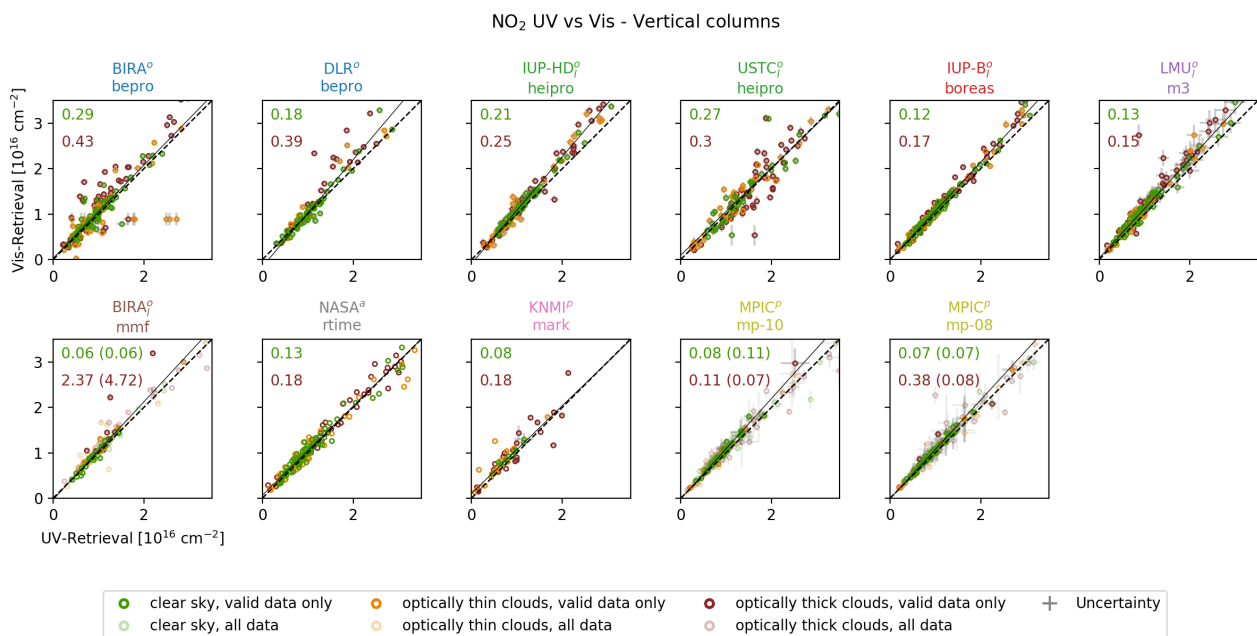
**Figure S57.** Comparison of MAX-DOAS NO<sub>2</sub> surface concentrations retrieved from the participant's own dSCDs. Description of Fig. S51 applies.



**Figure S58.** Correlation statistics of trace gas surface concentrations retrieved from the participant's own dSCDs. Basic description of Fig. S52 applies.

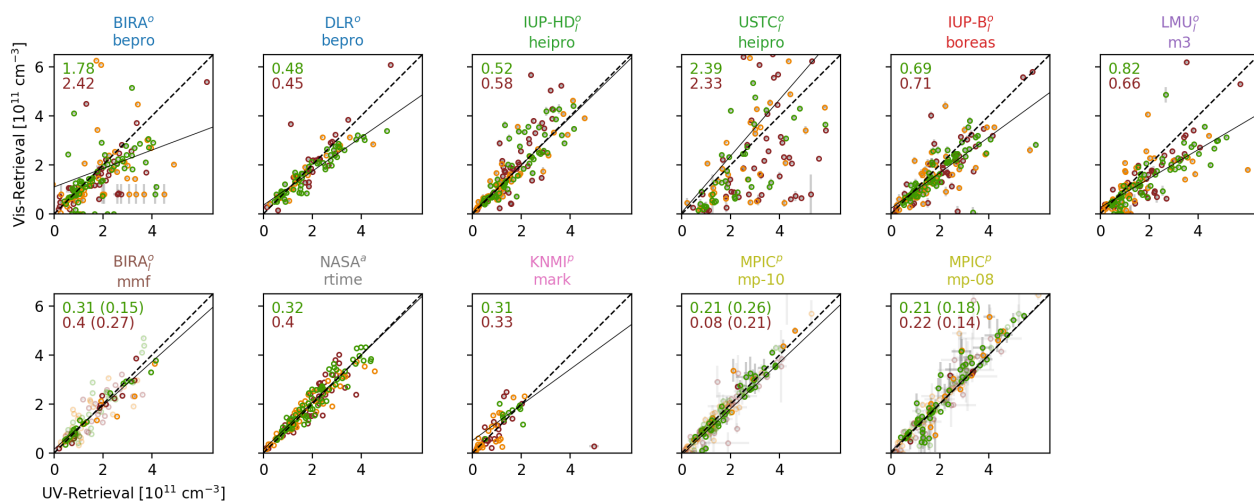
### S10.7 NO<sub>2</sub> UV-Vis comparison

This section is equivalent to Sect. S10.7 in the main text.



**Figure S59.** Correlation of MAX-DOAS VCDs retrieved from the participant's own dSCDs in the UV and the Vis spectral range. Marker colours and marker shapes indicate the cloud conditions.

# NO<sub>2</sub> UV vs Vis - Surface concentrations



**Figure S60.** Correlation of MAX-DOAS surface concentrations retrieved from the participant's own dSCDs in the UV and the Vis spectral range. Legends of Fig. S59 apply.

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