

S5P/TROPOMI NO₂ slant column retrieval: method, stability, uncertainties, and comparisons against OMI

by Jos van Geffen, et al.; amt-2019-471

Response to anonymous referee #1

RC Any early assessment of the data quality for a major remote-sensing mission (such as TROPOMI) is always welcome, and the study in review, "S5P/TROPOMI NO₂ slant column retrieval: method, stability, uncertainties, and comparisons against OMI", by van Geffen et al., delivers a timely and conscientious report on the particulars of the NO₂ slant column density retrievals from the TROPOMI data. The article is well-written and reasonably compact. I perceive no major problems that may delay a prompt publication. My comments are mostly of a clarifying or editorial nature.

We thank the referee for the kind words and for reading the manuscript in great details.

Changes to the manuscript are based on the comments and suggestions of three referees. In addition we have extended the data record of the paper by 3 months, which has led to updating some figures and numbers, but has *not* affected the conclusions of the paper.

In the following we answer the specific comments of referee #1.

RC P.8,L.11: The smaller-pixel size doubles the percentage of flagged NO₂ retrievals. This leads to additional $\sim 2 * 10^3$ flags/orbit compared to the larger-pixel case. Please comment, if you may, on the possible cause of this doubling. Are these flagged retrievals randomly distributed across the orbit? Are they mostly related to the saturated pixels?

From the way its written it indeeds seems like the number of pixels with SCD error $> 33 \mu\text{mol}/\text{m}^2$ doubles with the coming of the smaller pixel size, compared to the total number of pixels for which an SCD error is computed. The numbers are based on comparing complete large pixel orbits from 20190403 ($0.05 \pm 0.03\%$; in absolute numbers: 627 ± 357) against the small pixel test orbits from 20180405 ($0.12 \pm 0.05\%$; absolute: 1872 ± 817).

Due to the reduces pixel size, the SCD error increases, as the paper shows, by about 10%. Hence, it would be more fair to compare the number of small pixels with SCD error $> 33 \mu\text{mol}/\text{m}^2$ to large pixels with SCD error $> 30 \mu\text{mol}/\text{m}^2$: the latter is $0.08 \pm 0.03\%$ (absolute: 1037 ± 430). In other words, the relative number of high SCD error pixels increases from 0.08% to 0.12% due to the pixel size reduction.

BUT here we are talking about averages over the full distribution of SCD errors, while the high SCD error values are in the tail of that distribution and the behaviour of tails of distributions may differ quite a bit from the behaviour of the average. Hence it's hard to say whether this increase from 0.08% to 0.12% is unexpected or not. In any case, it would be better to phrase this more carefully in the paper, also keeping in mind that comparing numbers from different days is not trivial due to differences in atmospheric circumstances.

In both types of orbits, the orbits passing over the South Atlantic Anomaly have the largest number of cases of larger SCD errors. Other concentrations of pixels with larger SCD errors are found along the edges of high bright clouds, where they are associated with saturation.

New text [P8,L18-20 in revised version]:

SCD error values this large occur rarely: usually $< 0.1\%$ of the pixels per orbit with original ground pixel sizes; for the smaller size pixel orbits there are about 50% more pixels with high SCD error values (based on one test day of data), taking into account that the SCD error itself increases with reduced pixel size.

RC P.9,1.16. Please comment on the evolution of the radiance-irradiance wavelength difference along the TROPOMI orbit.

The irradiance wavelength shifts in Fig. 2a are derived from the one irradiance measurement, hence that does not change along the orbit. The radiance w_s in Fig. 2a is, as mentioned, an average over the 30-degree (sub-satellite) Tropical Latitude range, so as to filter out large scanline-to-scanline variation due to atmospheric circumstances: the presence of clouds may lead to large along-track variations in w_s that are not related to possible instrumental issues, which is what we're looking for here.

When taking other 30-degree latitude ranges the across-track shape shown in Fig. 2a does not change visibly, while there is a very small change in the average $\overline{w_s}$ (i.e. the dotted line in Fig. 2a) from about 23 to about 24×10^{-3} nm, going south to north. Why $\overline{w_s}$ increases from south to north is not known, but the change is very small (5% at most) and well within requirements, so we can safely ignore it further. Text added to the manuscript [P9,L26-28]:

When taking a different latitude range the across-track shape of the radiance wavelength shift shown in Fig. 2a does not noticeably change, while the absolute value of the average shifts increases by about 5% going south to north – it is not known what causes this small increase, but it is well within instrument specifications.

RC P.9,1.23: "A similar seasonal variation, though larger in magnitude, is seen in the OMI wavelength calibration data...". Fig. 34 from Schenkeveld et al. (2017) shows comparable (early-mission OMI data) magnitudes of the seasonal $< w_s >$ variability, without any detectable long-term trends in the OMI VIS channel. Please re-phrase ll. 23-27 accordingly.

Thank you for pointing this out, you are quite right.

The seasonal amplitude is comparable – the manuscript text has been adapted, with the remark on the temperature of the optical bench removed [P10,L3-5]:

A similar seasonal variation of similar amplitude is seen in the wavelength calibration data of OMI's visible channel (Schenkeveld et al., 2017, Fig. 34). Both for TROPOMI and OMI this amplitude does not exceed scatter levels and is thus well within instrument requirements.

RC P.11, Eq. (9) As written, the intensity offset implies a [potentially] non-linear change of the wavelength sampling for radiances. Please check.

Oops, misplaced parenthesis in Eq. (9) – sorry: has been corrected: $(I(\lambda) + P_{\text{off}}(\lambda)) \cdot S_{\text{off}}$

RC P.12,1.1. "For OMI/QA4ECV both a shift and stretch are fitted;..." Is the stretch term essential for OMI? Please comment.

For by far most OMI ground pixels the stretch is $\times 10^{-4}$ or less; only close to begin and end of orbits, i.e. at high SZA, the stretch may be a little larger, up to a few $\times 10^{-3}$.

RC P.12, Sect.3.4 - Please specify whether OMNO2A v2.0 uses the intensity offset term. It does not – added that to the text [P12,L20-21].

RC P.12, Sect.4.1, 1st par. - Please comment on how representative this particular TROPOMI orbit is. Have similar comparisons been done for different seasons? Can the shown trends be safely extrapolated on a \sim year-round sample? This in

particular applies to the tests shown in Fig.3b,e. To some extent, this question is answered in P.19. It is worthwhile to make a general statement earlier on.

Note first of all that we do *not* extrapolate trends: we show daily Pacific Ocean orbits data.

The orbit used in Fig. 3 is an arbitrary choice. Obviously, the average GCD value itself varies over time due to atmospheric circumstances, but the overall characteristics of the GCD and the SCD error – the across-track shape and stripiness – are representative as is shown by the stability of the wavelength calibration shifts (Fig. 2), of the stripe amplitude (Fig. 4), and of statistical and DOAS uncertainties (Fig. 8).

Some words on the representativity are added at the beginning of Sect. 4.1 [P14,L8-9]:

The TROPOMI orbit used here is representative for all Pacific Ocean orbits in across-track shape and variability, as is shown in subsequent sections by the stability of stripe amplitude (Sect. 4.3) and slant column uncertainties (Sect. 4.6).

RC P.14, last par. Perhaps it would be worthwhile to point to the big difference in the striping patterns delivered by various OMI algorithms, also mentioning that the forthcoming Sect. 4.3 quotes exclusively the OMI/QA4ECV stats.

Good point – text has been adapted [P14,L32ff]:

Note that the across-track striping in the OMI results differs markedly between the different processor results, which is related to a combination of processor differences and the response to instrumental issues (OMI striping data quoted in Sect. 4.3 is taken from OMI/QA4ECV).

RC P.15,1.5. "The increased SCD error visible in the TROPOMI data of Fig. 3d-e around $\theta \approx +20\text{deg}$ is related to the presence of saturation effects above bright clouds along this particular orbit." Nevertheless, most of the corresponding saturation-affected GCD values seem normal. Does this imply that the 405-465nm retrieval range retains enough of the saturation-free data to provide ~normal-looking SCDs? Is there anything else helping to stabilize the saturation-affected retrievals? Please comment on.

Spectral pixels flagged in the L1B spectra as saturated are removed from the fit, and in v1.2-v1.3 of the TROPOMI processor ground pixels with more than 3 such flags are not processed (the 405–465 fit window has about 305 spectral pixels).

Ground pixels with 1 to 3 flags will give more or less normal-looking SCDs but are likely to give markedly higher SCD error levels. In addition, ground pixels next to saturation cases may suffer from blooming, which possibly affects the SCDs and certainly will increase the SCD error.

As mentioned in Sect. 3.2 ground pixels with large SCD error values are flagged in the final data product as unreliable. Filtering on the SCD error was *not* done for the data in Fig. 3.

In the forthcoming v2.1 processing, the L1B flagging will be improved and the NO2 algorithm will use an outlier removal, leading to more spectral pixels being removed from the fit. As a result of this, we can allow for more spectra to enter the SCD retrieval and we can expect cases around saturation effects to give us more reliable SCDs, also for previously discarded ground pixels, albeit with somewhat elevated SCD error levels.

RC P.16,1.18. The 4-times difference in the QDOAS/TROPOMI RMS estimates is hard to overlook, indeed. Considering the typical $S/N \sim 1500$ in the TROPOMI data, one may side with the $RMS \sim 8 * 10^{-4}$ provided by QDOAS. Similar- magnitude

RMS $\sim 0.5 - 1.0 \times 10^{-4}$ is frequently quoted by various groups working with \sim similar S/N data sets from various spacecrafts. The cited (Table 2) RMS $\sim 0.2 \times 10^{-4}$ seems like an overly optimistic assessment. Any comment?

The point behind this is that the RMS error definitions in the IF and ODF modes differ. QDOAS ODF uses Eq. (7) and TROPOMI IF used Eq. (4); the relation between these two is discussed in Appendix B. QDOAS IF gives a 9% higher RMS error than QDOAS ODF (mentioned in the preceding "dash" point), hence that mode seems to be using an RMS error definition different from both Eq. (4) and Eq. (7).

The manuscript is improved by rephrasing these two "dash" points [P17,L1-5]:

- The RMS error calculation of the TROPOMI IF mode and the QDOAS ODF mode, given in Eq. (4) and Eq. (8), respectively, lead to different results; a relation between these two is given in App. B.
- Given that the RMS error in the QDOAS IF mode is $\sim 9\%$ higher than in the QDOAS ODF mode, indicates that the RMS definitions of these two QDOAS modes may be slightly different for the two modes, and that the definition of the QDOAS IF mode is different from the TROPOMI IF mode.

RC P.18,1.19 "Fig. 4c shows this RMS...". Please provide more details on how this value was calculated.

Actually there is not more to it than just that: the RMS of the SCD stripe amplitude, i.e. the blue line in Fig. 4a. The equation has been added [P19,L3-4]:

A measure for the stability of the SCD stripe amplitude is the RMS of the across-track stripe amplitude, i.e. of the blue line in Fig. 4a: $\sqrt{\{\sum_i (N_{s,i}^{\text{str}})^2\}}$, with summation over rows $i = 0, 1, \dots, 449$.

RC P.19,1.9 Is de-stripping applied both to TROPOMI and OMI? Please clarify in the text.

Yes, and added [P19,L17]:

... after applying the respective de-stripping of the SCDs described in the previous subsection on both datasets.

RC Fig.6. Since the intensity offset has the highest impact over the cloud-free, clear-water ocean areas (Sect. 5), one may test this by segregating the data in Fig.6a into two cases (cloudy and cloud-free; probably, selecting even more extreme cases than in Fig.8) and commenting on. I consider this as an important test, in light of the findings from Oldeman (2018).

Fig. 6 only shows cloud-free cases (cloud radiance fraction < 0.5), but is indeed a good idea to add cloudy cases – that has been done in the form of a new figure, Fig. 7, and the text has been extended accordingly [P16,L18ff]:

Fig. 6 shows the scatter plot of the TROPOMI and OMI/Q4ERCV GCDs of (almost) clear-sky ground pixels (i.e. cloud radiance fraction < 0.5) for July 2018 for two regions: the remote Pacific Ocean and the polluted area covering India and China on the Northern Hemisphere; the definition of these two areas is included in the figure panel legends. Both regions show a very good correlation with $R^2 \approx 0.99$. Over the Pacific Ocean area (Fig. 6a) the clear-sky TROPOMI GCD is on average $2.20 \pm 1.65 \mu\text{mol}/\text{m}^2$ ($1.33 \pm 0.99 \times 10^{14} \text{ molec}/\text{cm}^2$) or $5.23 \pm 3.93\%$ larger than the OMI/QA4ECV GCD. For Jan. 2019 the result (not shown) is quite similar: the clear-sky TROPOMI GCD over the Pacific Ocean is on average $2.19 \pm 1.56 \mu\text{mol}/\text{m}^2$ or $5.78 \pm 4.61\%$ larger than OMI/QA4ECV. Over the polluted India-to-China area (Fig. 6b) the clear-sky TROPOMI GCD is on average $2.02 \pm 2.08 \mu\text{mol}/\text{m}^2$ or $3.79 \pm 4.06\%$ larger than OMI/QA4ECV, i.e. the relative difference is a little smaller than from the Pacific Ocean.

For cloudy pixels (i.e. cloud radiance fraction > 0.5) the difference between the

TROPOMI and OMI/QA4ECV GCD is smaller, both in absolute and in relative terms, and the scatter is less, as can be seen from Fig. 7. Over the Pacific Ocean area (Fig. 7a) the cloudy TROPOMI GCD is on average $1.27 \pm 0.93 \mu\text{mol}/\text{m}^2$ ($0.76 \pm 0.56 \times 10^{14} \text{ molec}/\text{cm}^2$) or $3.04 \pm 2.39\%$ larger than the OMI/QA4ECV GCD. Over the polluted India-to-China area (Fig. 7b) the clear-sky TROPOMI GCD is on average $1.38 \pm 1.26 \mu\text{mol}/\text{m}^2$ or $2.74 \pm 2.37\%$ larger than OMI/QA4ECV.

Over land (Fig. 6b and Fig 7b) the differences are less than over ocean (Fig. 6a and Fig 7a), and over clouds (Fig. 7) the differences are less than over clear-sky cases (Fig. 6). Since clear-water cases show larger differences than land and cloudy cases, it thus seems that the intensity offset correction may be correction some absorption effect in open waters but not only such effects. The discussion in Sect. 5.1 has been extended in this sense [P26,L22ff]:

The quantitative comparison discussed in Sect. 4.4 revealed that for clear-sky cases (Fig. 6) the differences are a little larger than for the cloudy cases (Fig. 7), and for clear-sky cases the difference is larger for the remote Pacific Ocean area (almost completely water) than for the polluted India-to-China area (mainly land surface), while for the cloudy cases the differences are comparable for the two areas. These differences thus seems to indicate that the IOC may be correcting for some absorption effects in ocean waters, but not only for such absorption effects given that the reduction in GCD is also seen over land and over clouds.

RC P.22,1.2 "The fact that the GCD value itself (Fig. 7c) is not appreciably affected by the time difference is very reassuring..." Actually, I find this really puzzling: cf. Fig.7b and 7c. The SCD error, as anticipated, linearly increases in time in Fig.7b, but remains essentially flat in Fig.7c. Please help me (and the readers) to interpret this.

Fig. 7c does (in the revised version it's Fig. 8) *not* show the SCD error, I'm afraid. Instead:

- the SCD error and its stripiness (stddev) are given in Fig. 7b,
- the GCD and its stripiness (stddev) are given in Fig. 7c.

Fig. 7c shows that the GCD itself (red) remains the same – which is the reassuring point made – while at the same time the stripiness increases with increasing time difference between radiance and irradiance.

The confusion may have come from the fact that in the Fig. 7b and Fig. 7c *red and blue solid* lines are used for one case of SCD error or GCD and its stddev, while *red and blue dashed* lines are used for the other case. This is a bit counter-intuitive as one would expect the quantity to be given by a solid line and the stddev by a dashed line. But in that case the solid lines for the quantities almost overlap in Fig. 7b and fully overlap in Fig. 7c – that's the background for the original choice.

But your comment and the comment of referee #3 has shown that the choice made is too confusing, hence the more intuitive approach is used now, with an updated figure caption [P23].

RC P.23,1.12 Table 3: Please specify whether the TROPOMI (presumably, yes) and OMI (?) SCDs were de-striped.

No, the SCD data is *not* de-striped for the statistical uncertainty analysis.

De-striping would introduce model information in the SCD data and we want to investigate the uncertainty if the SCD data itself, i.e. coming directly from the instrument. Note further that the DOAS uncertainty is linked to the not de-striped SCDs, i.e. if de-striped SCDs were used for the statistical uncertainty, the link between the quantities would be weaker.

RC P.23,l.28 "From Fig. 8 it is furthermore clear that the statistical and the DOAS uncertainties of TROPOMI appear to be stable over the currently available data period." This is not what Fig.8 shows (the Aug. 6 jump). Please re-phrase.

No – the Fig. 8 does show that the instrument is stable over time.

The jump at 7 Aug. is *not* an instrumental issue/problem, but a configuration choice. This is also indicated by the fact that the standard deviation of the quantities given in Table 3 are almost the same before and after the pixel size change. New text [P26,L1-5]:

From Fig. 9 and Table 3 it is furthermore clear that the statistical and the DOAS uncertainties of TROPOMI appear to be stable over the currently available data period: the standard deviation of the quantities given in Table 3 are small and Fig. 9 shows no systematic change over time. The jumps in the quantities on 6 Aug. 2019 are caused by the along-track pixel size change, not by an instrumental issue, and this change has not affected the stability: the standard deviations of the quantities given in Table 3 are not markedly different between the two measurement modes.

RC P.25,l.14 "...also because instrumental effects such as straylight and dark current are adequately corrected for in the spectral calibration in the level 0-to-1bprocessor."

Either remove the statement or provide the references that address the subject and are based on assessment of the post-launch data (both radiances and irradiances).

Word "adequately" removed (stating the spectra are calibrated suffices) and two references added: (Kleipool et al, 2018; Ludewig et al., 2020); the latter is a paper in review for the same AMT TROPOMI Special Issue since early Feb. 2020.

RC Section 5.2 I have a problem linking the statement "Since NO₂ over the Pacific Ocean is primarily stratospheric NO₂, validation of stratospheric NO₂ essentially is also validation of Pacific Ocean NO₂ SCDs." Besides the point that it should be re-phrased, I do not see any connections between the discussed TROPOMI retrievals over Pacific and the ZSL-DOAS/SAOZ S5PVT network that completely avoids the Pacific basin. If the authors employed somewhat different approach than in Lambert et al. (2019), then they should provide a detailed description of the validation process. If this Section provides a summary of the Lambert et al. report and nothing besides, it should say so.

The section argues that the NO₂ over the Pacific Ocean is primarily located in the stratosphere, in view of the absence of (anthropogenic) sources of NO₂ in the region. Hence, validation of stratospheric NO₂ can also be seen as validation of Pacific Ocean SCDs. The sentence quoted by the referee has been reformulated somewhat [P27,L4-5]:

Since NO₂ over the Pacific Ocean, i.e. away from anthropogenic sources of NO₂, is primarily located in the stratosphere, validation of stratospheric NO₂ can also be seen as validation of Pacific Ocean NO₂ SCDs.

Validation of stratospheric NO₂ is described by Lambert et al. (2019) and we quote from that. We have performed no additional validation.

RC P.26,l.7 "...and the temperature dependence of the NO₂ reference spectrum (usually corrected for a-posteriori in the AMF application) may show spectral structures..."

Do you imply: "...and the temperature correction of the NO₂ VCDs (usually introduced a-posteriori in the AMF application) may result in spectral artifacts in the fitting residuals that are linked to the temperature dependence of the NO₂ reference spectrum." ? The NO₂ reference spectrum always shows spectral structures in the 405-465nm range...

The formulation is indeed not clear. It was formulated by Richter et al. (2014) in the poster abstract a "the spectral signature of the temperature dependence of the NO₂

absorption cross-section could be detected.” The point thus is probably that the temperature dependency depends on the wavelength. Hence ”may show spectral structures” is replaced by ”may be wavelength dependent” [P27,L25].

RC P.27,1.7 ”Since NO₂ over the Pacific Ocean is primarily stratospheric NO₂, validation of stratospheric NO₂ essentially is also validation of Pacific Ocean NO₂SCDs. As reported by Lambert et al. (2019...” Again, I fail to relate the discussed Pacific Ocean retrievals to the Lambert et al. report.

See above.

RC P.28,1.3 There is no QDOAS Case 6 in Table 2. Please correct.

Corrected: it’s case 3 [P29,L20]; the ”case 3” mentioned a few lines down [P29,L25] has to be ”case 2” (leftovers from earlier manuscript version; sorry).

RC P.28,1.23 ”After removal of such outliers...” Please specify how the removal is done: are the pixels corrected and re-used or completely removed from the fit?

Spectral pixels with outliers are completely removed from the fit, just as spectral pixels suffering from (too much) saturation or flagged as ”bad” in the level-1b spectra. The text has been adjusted somewhat [P30,L9-10]:

Spectral pixels with such outliers are removed completely from the measured reflectance and the DOAS fit is redone to provide ...

RC P.28,1.25 If this concerns the total, spatial x spectral number (I presume, though, that the authors speak of the spatial domain only – please clarify!), ~5% is, actually, a sizable population of the pixels. It may grow in time, eventually leading to much noisier and probably biased (depending on the preferential location of the spikes) retrievals. Any comments on this?

We find in ~5% of the ground pixel spectra one or more spectral pixel with an outlier (some 80,000 of the about 1.5 million ground pixels per orbit). This is indeed a sizeable portion, but since most of these outliers are related to either the South Atlantic Anomaly transients or saturation effects over clouds, this does not seem to be much of a problem nor an issue that is likely to grow over time. See also next point.

RC P.28,1.26 ”...most of which have less than 5 outliers per ground pixel...” Please clarify that these 5 outliers happen in the spectral domain.

Indeed, 5 spectral pixels showing outliers; text adapted, along with the changes from the preceding point [P30,L11ff]:

Outliers occur only in a small fraction of the ground pixels: usually ~ 5% of the successfully processed ground pixels show one or more outliers in their spectrum, and most of these ground pixels with outliers have less than 5 spectral pixels showing outliers per ground pixel; the largest effects ...

In the following we answer the technical comments of referee #1.

RC P.4,1.23: ”...away from anthropogenic sources of NO₂...”

Done.

RC P.9,1.28: ”The dominant term in the overall magnitude of the radiance is the inhomogeneous illumination...” This needs to be clarified: ”For a given field-of-view (row), the dominant term in the overall magnitude...” - is this what the authors meant?

Indeed, for a given field-of-view (ground pixel, not just row). The presence of clouds will likely cause differences in w_s along-track, across-track and day-to-day. The text has been adapted as suggested [P10,L6].

RC P.9, l.29 "The magnitude of the day-to-day variation in the average is much smaller than this long-term oscillation..." By 'this long-term oscillation' do you mean ...the seasonal (Fig.2b) oscillation...?

We indeed mean the oscillation in Fig. 2b; that has been added to the text [P10,L9].

RC P.15,l.19 "...the TROPOMI processor reports 10.2% ..."

Done.

RC P.15,l.27 "TROPOMI level-1b version 1.0.0 spectra..."

Done.

RC P.16,l.6 "...with other configuration settings as much as possible matching those of the TROPOMI processor..."

Done.

RC P.16,l.16 "...that the RMS definition may be different..."

Done. Thanks for reading the manuscript so carefully.

RC P.16,l.20 "As a reference..."

Done

RC P.18,l.4 "... is smooth rather than stripe-like over the non-contaminated areas..."

Good idea to add that.

RC P.25,l.4 "...the intensity offset corrections are..."

Done.

RC P.25,l.14 adequately

Done.