# Generalized Canonical Transform method for radio occultation sounding with improved retrieval in the presence of horizontal gradients 

Michael Gorbunov ${ }^{1,2}$, Gottfried Kirchengast ${ }^{3}$, and Kent B. Lauritsen ${ }^{4}$<br>${ }^{1}$ A.M. Obukhov Institute of Atmospheric Physics, Russian Academy of Sciences, Pyzevsky per. 3, 119017, Moscow, Russia<br>${ }^{2}$ Hydrometcenter of Russia, Bol. Prechistensky per. 11-13, 123242, Moscow, Russia<br>${ }^{3}$ Wegener Center for Climate and Global Change (WEGC) and Institute for Geophysics, Astrophysics, and Meteorology/Institute of Physics, University of Graz, Brandhofgasse 5, 8010, Graz, Austria<br>${ }^{4}$ Danish Meteorological Institute, 2100, Copenhagen, Denmark<br>Correspondence: Michael Gorbunov (gorbunov @ifaran.ru)


#### Abstract

By now, a series of advanced Wave Optical (WO) approaches to the processing of Radio Occultation (RO) observations are widely used. In particular, the Canonical Transform (CT) method and its further developments need to be mentioned. The latter include the Full Spectrum Inversion (FSI) method, the Geometric Optical (GO) Phase Matching (PM) method, and the general approach based on the Fourier Integral Operators (FIOs), also referred to as the CT type 2 (CT2) method. The general idea of these methods is the application of a canonical transform that changes the coordinates in the phase space from time and Doppler frequency to impact parameter and bending angle. For the spherically symmetric atmosphere, the impact parameter, being invariant for each ray, is a unique coordinate of the ray manifold. Therefore, the derivative of the phase of the wave field in the transformed space is directly linked to the bending angle, as a single-valued function of the impact parameter. However, in the presence of horizontal gradients, this approach may not work. Here we introduce a further generalization of the CT methods in order to reduce the errors due to horizontal gradients. We describe, in particular, the modified CT2 method denoted CT2A, which complements the former with one more affine transform: a new coordinate that is a linear combination of the impact parameter and bending angle. The linear combination coefficient is a tunable parameter. We derive the explicit formulas for the CT2A and develop the updated numerical algorithm. For testing the method, we performed statistical analyses based on COSMIC RO retrievals and (collocated) ECMWF analysis profiles. We demonstrate that it is possible to find a reasonably optimal value of the new tunable CT2A parameter that mitigates systematic errors in the lower troposphere and allows the practical realization of the improved capability to cope with horizontal gradients.


Discussions

## 1 Introduction

The first step in the development of wave optical (WO) approach to the processing of radio occultation (RO) observations was made by Melbourne et al. (1994), who used the thin screen approximation for the atmosphere combined with the Back Propagation (BP) technique. This approach was further developed under the name of Fresnel Inversion by Mortensen and Høeg (1998). Although the accuracy of this approximation in lower troposphere was insufficient for the practical application, its basic idea was correct. It consisted in the reduction of the influence of the diffraction by using the BP, which made the inversion results independent from the observation distance and canceled the resolution restriction due to the Fresnel zone size.

Later works (Gorbunov et al., 1996; Karayel and Hinson, 1997; Gorbunov and Gurvich, 1998a, b) developed a different understanding of the BP technique. The BP wave field evaluated in some plane was not considered as the actual wave field, but as a representation of the original field observed at the Low Earth orbit (LEO): in this representation, the effects of diffraction and multipath propagation were significantly reduced. This, in a straightforward way, allowed evaluation of the geometric optical (GO) bending angle profile, which was inverted in the framework of the standard GO scheme (Ware et al., 1996; Kursinski et al., 1997).

The further development of the WO approach based on the representation view relied upon the concept of the Canonical Transform (CT) originating from the classical mechanics (Arnold, 1978; Goldstein et al., 2014), generalized for the quantum mechanics by Fock (1978), mathematically substantiated by Egorov (1985); Egorov and Shubin (1993). Further on this concept obtained an extensive mathematical development (Treves, 1982a, b; Hörmander, 1985a, b). The correspondence between the quantum and classical mechanics is the same as the link between the wave optics and geometrical optics.

In both cases, there is a strict mathematical representation (quantum mechanics or wave optics) and its asymptotic solution (classical mechanics or geometrical optics). While the evolution of de Broglie waves of probability or electromagnetic waves is described by the Hamilton operator, the evolution of rays or classical trajectories of particles is described by Hamilton system, where the Hamilton operator is obtained by the substitute of the momentum operator instead of classical momentum. Accordingly, for the classical problem the phase space is introduced, the dimension of which equals doubled geometric dimension, because to each geometrical coordinate we can conjugate the corresponding momentum. For the wave problems momentum is understood as the ray direction vector.

The canonical transforms arise, when we consider the class of the transforms of the phase space that conserve the canonical form of the Hamilton dynamical system. It was first demonstrated by Fock (1978) that these transforms have a very simple implementation in the quantum mechanics: they correspond to linear transforms of the wave function. The kernel of this transform is derived in classical terms, but, still, it describes a short-wave asymptotical solution of the wave problem. This idea was later mathematically developed first by Egorov (1985); Egorov and Shubin (1993) and then by Treves (1982a, b); Hörmander (1985a, b).

The application of the CT approach for the RO observation processing was pioneered by Gorbunov (2002), where it was combined with the BP. The idea of the CT without BP was first developed by Jensen et al. $(2003,2004)$ and later the general view at these results in the framework of the CT approach was developed by Gorbunov and Lauritsen (2004a); Gorbunov et al. (2004).

Discussions

Finally, it was recognized that the different methods: CT (Gorbunov, 2002), Full-Spectrum Inversion (FSI) (Jensen et al., 2003), Phase Matching (PM) (Jensen et al., 2004), and CT of the 2nd type (CT2) (Gorbunov and Lauritsen, 2004a) were, in fact, different approximations of the same solution, for which Fourier Integral Operators (FIOs) provided the general transform approach (Gorbunov and Lauritsen, 2004a).

The idea of the CT approach is as follows. Given the observations or RO complex signal $u(t)$ as function of time $t$, which can be represented through its amplitude $A(t)$ and phase $\phi(t), u(t)=A(t) \exp (i \phi(t))$. It is convenient to use eikonal, or phase path $\Psi(t)=\phi(t) / k$, where $k=2 \pi / \lambda$ is the wavenumber, and $\lambda$ is the wavelength. Thus, $u(t)=A(t) \exp (i k \Psi(t))$, and $k$ is the large parameter. The signal is composed of multiple sub-signals $u_{i}(t)=A_{i}(t) \exp \left(i k \Psi_{i}(t)\right)$ corresponding to interfering rays. For each sub-signal it is possible to introduce the instantaneous frequency $k \dot{\Psi}_{i}=k \sigma_{i}$. However, instantaneous frequency cannot be introduced for their composition.

The multipath propagation problem consists in the de-composition of the signal equal to the sum or different sub-signals, to retrieve the ray structure of the observed field. The solution of this problem discussed in the aforementioned papers consisted in the transform of the observed wave field $u(t)$ into a different representation. The new coordinates in the transformed space were the ray impact parameter $p$ and bending angle $\epsilon$. The transform $(t, \sigma) \rightarrow(p, \epsilon)$ is canonical (Gorbunov and Lauritsen, 2004a), which allows for writing the corresponding linear transform $\widehat{\Phi}_{2}$, where the subscript 2 indicates that it is a CT of the 2nd type (Arnold, 1978; Goldstein et al., 2014), that maps the original field $u(t)$ to field in the impact parameter representation $\hat{u}(p)=\widehat{\Phi}_{2}[u(t)](p)$. The idea of the choice of the ray impact parameter as the new coordinate is based on the fact that in a spherically-symmetric medium, ray impact parameter is the ray invariant, which is known the Bouger law. The locally spherically-symmetric medium is the basic approximation used in the inversion of RO data. For the real atmosphere with horizontal gradients, the dynamic equation for $p$ was derived by Gorbunov and Kornblueh (2001), who demonstrated that derivative of $p$ with respect to the ray arc length is equal to the horizontal component of the refractivity gradient in the occultation plane. Strong horizontal gradients may result in the situation when dependence $\epsilon(p)$ becomes multi-valued (Healy, 2001; Gorbunov and Lauritsen, 2009), which was referred to as the impact multipath (Zou et al., 2019).

The idea explored in the present manuscript consists in the further development of the CT approach by using a generalized transform with the coordinate $p^{\prime}=p+\beta \epsilon$. Unlike the standard CT approach, where the form of the new coordinates in the phase is known in advance, this transform has the tunable parameter $\beta$ that can take into account the statistical impact multipath effect.

The paper is organized as follows. In Section 2 we discuss the canonical transform in wave optics and quantum mechanics in general terms, including brief review of FIOs. Based on this context we discuss in Section 3 the application of the CT method for RO and introduce the particular phase space and the specific choice of coordinates as well as the new generalization adding an affine transform with a tunable parameter for improved the coping capability with horizontal gradients. In Section 4 we discuss the practical modifications needed to readily advance existing numerical implementations of the CT algorithm and present results of our performance evaluation from processing real-observed COSMIC RO data, including how to find an optimal value of the tunable parameter minimizing the systematic errors in the lower troposphere. Section 5 finally provides the summary and main conclusions of the paper.

Discussions

## 2 General concept of Canonical Transform in Wave Optics

We will start with a brief discussion of the Canonical Transform (CT). This concept originated from the classical mechanics (Arnold, 1978; Goldstein et al., 2014), where it referred to a kind of transform of the coordinates and momenta that conserve the Hamiltonian form of the dynamical equation. Fock (1978) introduced the CT in the quantum mechanics. Note, the first
$u(t)=\sum_{j} A^{(j)}(t) \exp \left(i k \Psi^{(j)}(t)\right)$,
where the upper index $j$ enumerates the components, $A^{(j)}(t)$ are their amplitudes, and $\Psi^{(j)}(t)$ are their eikonals. Each component has its own instant amplitude and frequency.

When discussing the CTs, it is necessary to bear in mind that most of the relations have an asymptotic nature, where $k$ is the large parameter (or $\lambda$ is the small parameter). The reason is as follows. Given measurements of wave field, each monochromatic component can be interpreted in terms of wave fronts and rays. Each point has a single ray, and its direction is linked to the normalized frequency $\sigma(t)=\dot{\Psi}(t)$. To this end, it is also necessary to know the position of the transmitter and receiver, as it takes place in RO observation. However, at this stage of the consideration of the problem, we can simply speak about instant tones of the signal.

Therefore, for a specific class of signals, including quasi-monochromatic ones and their superposition, it is possible to introduce a phase space $(t, \sigma)$. Although the original signal is $1-\mathrm{D}$, this space is $2-\mathrm{D}$, and the structure of the signal can be described in terms of the function $\sigma(t)$ which can be both single-valued for quasi-monochromatic signals, or multi-valued for superpositions of such signals.

Consider RO observations as an example. The original signal corresponds to a range of rays starting at the transmitter and the phase space $\sigma(t)$ is a very smooth continuous line. As the signal propagates through the atmosphere its structure gets more and more complicated. Still, in the phase space its topological structure remains the same: it is always a single continuous line, although it may have multiple projections to the axis of time $t$, which corresponds to multipath propaga-
tion (Gorbunov, 2002; Gorbunov and Lauritsen, 2004a). Such a line representing the signal structure is referred to as the ray manifold (Mishchenko et al., 1990).

The outstanding and, still, simple idea of Fock (1978) was that the classical canonical transforms correspond to linear integral transforms of the wave field with oscillating kernels. This class of transforms was later named Fourier Integral Operators (FIO) (Egorov, 1985; Egorov and Shubin, 1993; Treves, 1982a, b; Hörmander, 1985a, b). The general form of such an operator first discussed by Fock (1978) has the following form:
$\hat{u}(p)=\sqrt{-\frac{i k}{2 \pi}} \int a_{2}(p, t) \exp \left(i k S_{2}(p, t)\right) u(t) d t \equiv \Phi_{2}[u(t)](p)$,
where $p$ is a new coordinate in the mapped space. We use notation $\Phi_{2}$ and, accordingly, $a_{2}$ and $S_{2}$, because this type of operators was referred to as the FIO of the second type (Gorbunov and Lauritsen, 2004a), while the FIO of the first type is the composition of a Fourier transform and a second-type FIO (Egorov, 1985; Egorov and Shubin, 1993). This type of operators is linked to the corresponding type of the generating function (Arnold, 1978; Goldstein et al., 2014). Note, historically FIO of the second type appeared first, but in mathematical works it was FIO of the first type that were discussed first.

Considering now $u(t)$ as a quasi-monochromatic signal, we can derive the asymptotic form of transform (3) using the stationary phase principle:
$\hat{u}(p)=\sqrt{-\frac{i k}{2 \pi}} \int a_{2}(p, t) A(t) \exp \left(i k\left(S_{2}(p, t)+\Psi(t)\right)\right) d t \equiv \Phi_{2}[u(t)](p)$,
The stationary phase point $t_{s}(p)$ of this integral satisfies the equation:
$\frac{\partial}{\partial t} S_{2}(p, t)+\dot{\Psi}(t)=0$.
Accordingly, the transformed field, under the assumption that the Eq. (5) has a single solution $t_{s}(p)$, is also quasi-monochromatic and can be written as follows:
$\hat{u}(p)=A^{\prime}(p) \exp \left(i k \Psi^{\prime}(p)\right)=A^{\prime}(p) \exp \left(i k\left(S_{2}\left(p, t_{s}(p)\right)+\Psi\left(t_{s}(p)\right)\right)\right)$.
Its instant frequency equals:

$$
\begin{align*}
\xi(p) & =\dot{\Psi}^{\prime}(p)=\frac{d}{d p}\left(S_{2}\left(p, t_{s}(p)\right)+\Psi\left(t_{s}(p)\right)\right)= \\
& =\frac{\partial}{\partial p} S_{2}\left(p, t_{s}(p)\right)+\left(\frac{\partial}{\partial t} S_{2}\left(p, t_{s}(p)\right)+\frac{\partial}{\partial t} \Psi\left(t_{s}(p)\right)\right) \frac{d t_{s}}{d p}= \\
& =\frac{\partial}{\partial p} S_{2}\left(p, t_{s}(p)\right) \tag{7}
\end{align*}
$$

by virtue of Eq. (5). Recalling that $\Psi(t)=\sigma$, which is the original momentum, we have the following relation between the canonical coordinates $(t, \sigma)$ and $(p, \xi)$, in the original and mapped spaces:
$\frac{\partial S_{2}}{\partial t}=-\sigma, \frac{\partial S_{2}}{\partial p}=\xi$,


Figure 1. Radio occultation observation geometry with relevant geometrical variables indicated (for description see Sect. 3.1)
$d S_{2}=\xi d p-\sigma d t$.
And, vice versa, the requirement that the right-hand part in Eq. (9) should be equal to a full differential $d S_{2}$ of a function $S_{2}(p, t)$ is a necessary and sufficient condition for the transform $(t, \sigma) \rightarrow(p, \sigma)$ to be canonical (Arnold, 1978; Goldstein et al., 2014). The function $S_{2}(p, t)$ is then termed the generating function of the canonical transform.

In terms of FIO, $S_{2}(p, t)$ is referred to as its phase function, and $a_{2}(p, t)$ is its amplitude function. The phase function, which specifies the canonical transform, is of primary importance, while the amplitude function is derived using the energy conservation (Gorbunov and Lauritsen, 2004a). We see, therefore, that using the classical, or geometric optical concepts, it is possible to write down the asymptotic form of the quantum, or wave optical operator implementing the transformation of the original signal into a different representation. If the structure of the original signal is represented as a ray manifold in the phase plane, such a transform is applied to the coordinates in this space. In particular, it may be possible to find such a coordinate system, where the ray manifold geometry will be exceptionally simple.

## 3 The Canonical Transform method for RO and its generalization

Here we discuss the application of the CT technique for the analysis of RO observations (Fig. 1) by first reviewing the different existing variants (3.1) and then introducing the new generalized CT method (3.2) and an application-relevant formulation for readily updating existing algorithms (3.3).

### 3.1 Canonical Transform method in different existing variants

The RO observation geometry is schematically represented in Figure 1. The wave emitted by a transmitter Tx is received by a receiver Rx on a low-Earth orbit. Transmitter is borne by a satellite belonging to one of the modern Global Navigation

Satellites Systems (GNSS), including GPS, GLONASS etc. Due to the movement of the transmitter and receiver the ray descends or ascends in the atmosphere, which allows the derivation of the atmospheric profiles from the bending angles $\epsilon(p)$ (Ware et al., 1996; Kursinski et al., 1997). The CT technique is used for the retrieval of bending angle profile from the wave field measurements.

The first approach of processing RO data, belonging to the class of CT, was the Back Propagation (BP) (Gorbunov et al., 1996; Karayel and Hinson, 1997; Gorbunov and Gurvich, 1998a, b). In this technique the field was linearly transformed to be re-calculated to the BP plane locate at coordinate $x_{B}$ :
$u_{B}(y)=\sqrt{\frac{i k}{2 \pi}} \int \frac{u(t) \exp \left(-i k\left|\mathbf{r}_{B}(y)-\mathbf{r}_{R}(t)\right|\right)}{\left|\mathbf{r}_{B}(y)-\mathbf{r}_{R}(t)\right|^{1 / 2}}\left|\sin \phi\left(\dot{\mathbf{r}}_{R}(t), \mathbf{r}_{B}(y)-\mathbf{r}_{R}(t)\right) \dot{\mathbf{r}}_{R}(t)\right| d t$,
where 2-D vector $\mathbf{r}_{B}(y)$ equals $\left(x_{B}, y\right), \phi(\mathbf{a}, \mathbf{b})$ is the angle between vectors $\mathbf{a}$ and $\mathbf{b}$. This transform is performed under the application of the procedure of the stationarization of the transmitting satellite and projection of the satellite movement to the vertical plane. Note, the same procedure is commonly applied when using CT-like approaches. It is important that the BP field is not the real field in the BP plane, because the BP procedure assumes the vacuum propagation. This procedure results in some representation of the original wave field with reduced diffraction effects due to the reduction of the propagation distance. The new coordinate $y$ is more favorable for finding a unique projection of the ray manifold that disentangles the multipath propagation. Still, this coordinate is not the best choice.

A much better coordinate for the new representation should be the impact parameter $p$, because in a spherically-symmetric medium it is an invariant for each ray due to the Bouger law, and thus it is unique for each ray. A dynamic equation for the variation of $p$ along the ray as a function of the horizontal gradient of refractivity was obtained by Gorbunov and Kornblueh (2001). The idea of complementing the BP technique with one more transform that maps the field to the impact parameter representation was pioneered by Gorbunov (2002). It was the first application of the FIO of the first type, which is linked to the other type of the generating function (Arnold, 1978; Goldstein et al., 2014) and has the form
$\hat{u}(p)=\sqrt{-\frac{i k}{2 \pi}} \int a_{1}(p, \sigma) \exp \left(i k S_{1}(p, \sigma)\right) \tilde{u}(\sigma) d \sigma \equiv \Phi_{1}[u(t)](p)$,
where the only difference with the second type operator is that it acts upon the Fourier-transformed field $\tilde{u}(\xi)$. It can be looked at as the composition of the Fourier transform, which itself is a second type FIO, and the other second type FIO. Because the Fourier transform is a simple rotation of the phase space by $\pi / 2:(t, \sigma) \rightarrow(\sigma,-t)$, the equation for the phase function takes the form:
$d S_{1}=\xi d p+t d \sigma$.
Gorbunov (2002) applied this operator to the back-propagated field. To this end, using the normal vector $\nu=\left(\eta, \sqrt{1-\eta^{2}}\right)$ to the straight ray, we express the impact parameter:
$p=-x \eta+y \sqrt{1-\eta^{2}}$.

Now it is necessary to find the canonical transform $(y, \eta) \rightarrow(p, \xi)$. We look for the first type operator, apply the property of
$\frac{\partial \xi}{\partial \eta} \frac{\partial p}{\partial y}-\frac{\partial \xi}{\partial y} \frac{\partial p}{\partial \eta}=1$,
and additional assumption $\xi=\xi(\eta)$. Then, from Eq. 14, we readily derive:

$$
\begin{align*}
\frac{\partial \xi}{\partial \eta} & =\left(\frac{\partial p}{\partial y}\right)^{-1}=\frac{1}{\sqrt{1-\eta^{2}}} \\
\xi & =\arcsin \eta \tag{15}
\end{align*}
$$

This results in the solution for the phase and amplitude functions:
$S_{2}(p, \eta)=p \arcsin \eta-x \sqrt{1-\eta^{2}}$,
$a_{2}(p, \eta)=\sqrt{\frac{\partial^{2} S}{\partial p \partial \eta}}=\left(1-\eta^{2}\right)^{-1 / 4}$.
This defines the FIO, which is applied to the backpropagated wave field $u_{B}(y)$ and produces the mapped field
$\hat{u}(p)=A^{\prime}(p) \exp \left(i k \int \xi(p) d p\right)$. allows a fast implementation. A similar idea will be applied below.

The complicated nature of the BP+CT algorithm stimulated further studies (Gorbunov and Lauritsen, 2002, 2004b) where the idea was expressed of applying the FIO directly to the observed wave field $u(t)$, without intermediate and numerically expensive steps like BP. Full-Spectrum Inversion (FSI) developed by Jensen et al. (2003) was the first solution of this type, although with some restrictive assumptions. However, the general solution was just one year away: the Phase Matching (PM) was developed by Jensen et al. (2004) and then put in the context of the CT approach by Gorbunov and Lauritsen (2004a), who also introduced an approach based on the linearized canonical transform that reduced the FIO to the composition of non-linear coordinate changes and Fourier transform. This algorithm was termed the 2nd type CT, or CT2.

In order to arrive at the phase function of the FIO of the 2nd type, consider the expression for the derivative of the phase of the observed wave field:

Using Eq. (8), we derive the phase function:

$$
\begin{align*}
S_{2}(p, t) & =-\int\left(p \dot{\theta}+\frac{\dot{r}_{T}}{r_{T}} \sqrt{r_{T}^{2}-p^{2}}+\frac{\dot{r}_{R}}{r_{R}} \sqrt{r_{R}^{2}-p^{2}}\right) d y= \\
& =-\int\left(p d \theta+\frac{d r_{T}}{r_{T}} \sqrt{r_{T}^{2}-p^{2}}+\frac{d r_{R}}{r_{R}} \sqrt{r_{R}^{2}-p^{2}}\right)= \\
& =-p \theta-\sqrt{r_{T}^{2}-p^{2}}+p \arccos \frac{p}{r_{T}}-\sqrt{r_{R}^{2}-p^{2}}+p \arccos \frac{p}{r_{R}} \tag{20}
\end{align*}
$$

$$
\begin{align*}
d \Upsilon & =\left(\frac{\partial p_{0}}{\partial \sigma}\right)^{-1} d t=\frac{\partial \sigma}{\partial p_{0}} d t \\
\eta & =\frac{\partial p_{0}}{\partial \sigma} \sigma \tag{22}
\end{align*}
$$

Finally, we arrive at the following linear canonical transform $(\Upsilon, \eta) \rightarrow(p, \xi)$ :
$\tilde{p}=f(\Upsilon)+\eta$,
$\xi=-\Upsilon$,
The generating function of this canonical transform is easily computed from the differential equation

$$
d S_{2}=\xi d \tilde{p}-\eta d \Upsilon=-\Upsilon d \tilde{p}-(\tilde{p}-f(\Upsilon)) d \Upsilon
$$

$S_{2}(\tilde{p}, Y)=-\tilde{p} Y+\int_{0}^{Y} f\left(Y^{\prime}\right) d Y^{\prime}$.

For the new coordinate $\Upsilon$ we have the following relation:
$d \Upsilon=d \theta-\frac{d r_{G}}{r_{G}} \frac{p_{0}}{\sqrt{r_{G}^{2}-p_{0}^{2}}}-\frac{d r_{L}}{r_{L}} \frac{p_{0}}{\sqrt{r_{L}^{2}-p_{0}^{2}}}$.
$H(\mathbf{r}, \mathbf{p})=\frac{1}{2}\left(\mathbf{p}^{2}-n^{2}(\mathbf{r})\right)$,
where $\mathbf{p}$ is the momentum, and $n(\mathbf{r})$ is the refractivity field. The Hamilton system has the following form:

$$
\begin{gather*}
\dot{\mathbf{r}}=\frac{\partial H}{\partial \mathbf{p}}, \quad \dot{\mathbf{p}}=-\frac{\partial H}{\partial \mathbf{r}}, \quad \dot{\Psi}=\mathbf{p} \dot{\mathbf{r}} \\
\dot{\mathbf{r}}=\mathbf{p}, \quad \dot{\mathbf{p}}=n \nabla n, \quad \dot{\Psi}=n^{2} \tag{29}
\end{gather*}
$$

where $\mathbf{p}$ is the classical momentum. Because $|\mathbf{p}|=|\nabla \Psi|=n$, we arrive at the following differential relation between the parameter $\tau$ of this system, the ray arc length $s$, and the eikonal:
$d \tau=\frac{d s}{n}, \quad d \Psi=n d s$.
Equation (29) has a form that is specific for the Cartesian coordinates. Consider an arbitrary coordinate system with the metric tensor $g_{i j}: d s^{2}=d x^{i} g_{i j} d x^{j}$, where $x^{i}$ are the components of vector $\mathbf{r}$, and we follow the Einstein tensor notation implying the summation over each pair of upper and lower indexes of the same name. If we define the momentum by the relation $p_{i}=g_{i j} \dot{x}^{j}$, the form $\mathbf{p} d \mathbf{r}$ is invariant, the transform to the new coordinates $\left(p^{i}, x^{i}\right)$ is canonical, and the canonical form of the Hamilton system also remains invariant [Arnold1978], provided that the Hamilton function is defined as follows:
$H(\mathbf{r}, \mathbf{p})=\frac{1}{2}\left(p_{i} g^{i j} p_{j}-n^{2}(\mathbf{r})\right)$,
where $g^{i j}$ is the matrix inverse to $g_{i j}$. This results in the following form of the ray equations:
$\dot{x}^{i}=\frac{\partial H}{\partial p_{i}}=g^{i j} p_{j}, \dot{p}_{i}=-\frac{\partial H}{\partial x^{i}}=n \frac{\partial n}{\partial x^{i}}-\frac{1}{2} p_{k} \frac{\partial g^{k j}}{\partial x^{i}} p_{j}$.
The 2-D approximation (Zou et al., 2002) allows treating rays as plane curves. Consider polar coordinates $(r, \theta)$ with the metric tensor:
$g_{i j}=\left(\begin{array}{cc}1 & 0 \\ 0 & r^{2}\end{array}\right), \quad g^{i j}=\left(\begin{array}{cc}1 & 0 \\ 0 & r^{-2}\end{array}\right)$.
Then we have the following equations:
$p_{\theta}=r^{2} \dot{\theta}=n r \frac{r d \theta}{d s}=n r \sin \psi$,
$\dot{p}_{\theta}=n \frac{\partial n}{\partial \theta}$,
$\dot{p}_{r}=\ddot{r}=n \frac{\partial n}{\partial r}+\frac{p^{2}}{r^{3}}$.
where $\psi$ is the angle between vectors $\dot{\mathbf{r}}$ and $\mathbf{r}$. The angular component of the momentum $p_{\theta}$ coincides with the ray impact parameter $p$, which is invariant in a spherically layered medium, but is perturbed by the horizontal gradients (Gorbunov and Lauritsen, 2009).

The variations of the ray impact parameter seem to undermine the elegant idea of the CT approach. Now there is no such a convenient invariant value ascribed to each ray. Still, the CT method can be applied using the same formulas, but the coordinate $p$ will now acquire a different meaning: it will be understood as the "effective impact parameter", i.e. the impact parameter which would result in the observed Doppler frequency shift, if the atmosphere were spherically layered [Gorbunov2019]. Accordingly, the evaluated bending angle will also be the "effective" bending angle. The reason is that for the evaluation of the real bending angle, understood as the angle between the ray directions at the transmitter and receiver, two corresponding values of the impact parameter are required, which cannot be derived from the single variable, the Doppler frequency. This,


Figure 2. Impact multipath, old coordinate (impact parameter) lines, and modified coordinate lines.
by itself, is not a significant problem, because the assimilation of bending angle profiles can be based on the effective values [Gorbunov2019], provided that the observation operator correctly implements their evaluation.

More importantly, horizontal gradients may result in multi-valued ray manifold projections, when using the effective impact parameter $p$ as the coordinate in the mapped space. This situation is termed "impact multipath" [Zou2019]. Theoretically, for any ray manifold perturbation there always exists an unfolding coordinate transform. This follows from the fact that topologically the ray manifold is always a continuous line without self-crossing. However, this coordinate transform will now depend on the a priori unknown horizontal gradients of refractivity.

Typical multi-valued bending angle profile [Gorbunov2009a, Zou2019] is shown in Figure 2. From numerical simulations, it can be inferred that there is a kind of asymmetry: impact multipath manifests itself mostly in ascending spikes, but hardly in descending spikes. Accodingly, in order to better unfold multipath, it must be possible to use another coordinate in such a way that the modified coordinate lines are sloped. Therefore, we modify the transform (23) in order to use another coordinate:
$\widetilde{p^{\prime}}=\tilde{p}+\beta \Upsilon$,
where $\beta$ is a tunable parameter and has a dimension of $\mathrm{km} / \mathrm{rad}$. Although the optimal value of this parameter should be different for individual events, the aforementioned asymmetry results in the conclusion that the preferred value of $\beta$ is expected to be negative. Therefore, it may be possible to find its optimal value that, in the statistical sense, will minimize errors due to impact multipath.

The modified canonical transform (23) is written as follows:
$\widetilde{p^{\prime}}=f(\Upsilon)+\beta \Upsilon+\eta \equiv f^{\prime}(\Upsilon)+\eta$,
$\xi=-\Upsilon$.

Using the modified function $f^{\prime}(\Upsilon)$ instead of the original one, we will obtain the expression for the modified FIO $\widehat{\Phi^{\prime}}{ }_{2}$. The advantage of this approach is that it can be implemented by a very simple modification of the existing CT2 algorithm. Using the numerical implementation of the modified CT will allow us to study its influence upon the RO inversion statistics in the
$\frac{\partial S^{(\beta)}}{\partial \tilde{p}^{\prime}}=\xi$,
$\frac{\partial S^{(\beta)}}{\partial \xi}=\tilde{p}=\tilde{p}^{\prime}+\xi\left(\xi-\xi_{0}\right)$.
From this, we can conclude that
$S^{(\beta)}\left(\tilde{p}^{\prime}, \beta\right)=\tilde{p}^{\prime}\left(\xi-\xi_{0}\right)-\beta \frac{\left(\xi-\xi_{0}\right)^{2}}{2}$.
This phase function defines the FIO of the first type:
350


Figure 3. Statistics for latitude band $0^{\circ}-10^{\circ}$. Left: number of data; right total relative difference of refractivity COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. Both are functions of the parameter $\beta$.

Finally, we can write the operator relation:
$\widehat{\Phi}_{2}^{(\beta)}=\widehat{\Phi}_{1}^{\prime}{ }^{(\beta)} \widehat{\Phi}_{2}^{(0)}$,
which can be used for the modification of the existing version of operator $\widehat{\Phi}_{2}^{(0)}$.
The above derivation allows for one more generalization. We can consider $\beta=\beta(\xi)$. In this case, the phase function is
$S^{(\beta)}\left(\tilde{p}^{\prime}, \xi\right)=\tilde{p}^{\prime}\left(\xi-\xi_{0}\right)-\int \beta(\xi)\left(\xi-\xi_{0}\right) d \xi$.
Using $\beta(\xi)=\sum \beta_{j} \xi^{j}$ results in a simple analytical expression for $S^{(\beta)}$ with a set of tuning parameters $\beta_{j}$. In this work, we, however, use a constant $\beta$.

## 4 Implementation and numerical performance evaluation

Our implementation of the CT2A algorithm was based on the existing program code with addition of the parameter $\beta$ and using the modified function $f^{\prime}(\Upsilon)$ as defined by Eq. (35). Practically, this only required modification of a few lines in the program code that implements the CT2 method, as well as the implementation of one more command line parameter.

In our numerical validation, we retrieved COSMIC refractivity profiles $N_{C}$, using COSMIC data from the year 2008, 1 st and 15th day of every month, leading to a total of 24 days and altogether around 60000 RO events. We used collocated ECMWF refractivity profiles $N_{E}$, i.e., interpolated to the corresponding COSMIC RO event location, as the reference. We used the total relative difference of COSMIC from ECMWF (the difference metric), defined as $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$, which includes both mean (systematic) and fluctuating (random) deviations.



Figure 4. Statistics for latitude band $10^{\circ}-20^{\circ}$. Left: number of data; right total relative difference of refractivity COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. Both are functions of the parameter $\beta$.


Figure 5. Statistics for latitude band $20^{\circ}-30^{\circ}$. Left: number of data; right total relative difference of refractivity COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. Both are functions of the parameter $\beta$.

Figure 3 through 11 show the statistical values of $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$ as function of latitude and parameter $\beta$. We averaged over $10^{\circ}$ wide latitude bands including both South and North hemispheres. The parameter $\beta$ changed in the interval from -4 to $-12 \mathrm{~km} / \mathrm{rad}$ with the step of 1 .

These results indicate that for latitudes $0^{\circ}-50^{\circ}$, in the altitude range from 0.5 km to $1.9-2.5 \mathrm{~km}$, the application of the CT2A algorithm allows minimizing the total relative difference of refractivity profiles COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. The optimal value of parameter $\beta$ is found to be -6 to $-8 \mathrm{~km} / \mathrm{rad}$. The CT2A algorithm also improves the penetration increasing the number of data in the altitude range below 0.5 km . Here the difference metrics for $\beta=0$ and optimal $\beta$ cannot be directly compared, because they are evaluated over different statistical ensembles.



Figure 6. Statistics for latitude band $30^{\circ}-40^{\circ}$. Left: number of data; right total relative difference of refractivity COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. Both are functions of the parameter $\beta$.


Figure 7. Statistics for latitude band $40^{\circ}-50^{\circ}$. Left: number of data; right total relative difference of refractivity COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. Both are functions of the parameter $\beta$.

## 5 Summary and conclusions

In this study we discussed the general idea of the Canonical Transform (CT) method and provided a new generalization adding more flexibility for application in RO processing. The idea came from quantum mechanics, where it was shown that the canonical transforms as they are understood in classical mechanics (geometrical optics) are implemented in quantum mechanics (wave optics) by linear operators with oscillating kernels. Such operators are referred to as Fourier Integral Operators (FIOs). During the past century, this approach acquired a solid theoretical basis. In numerous mathematical monographs, one finds the advanced theory of FIOs. The central role in this theory is played by the concept of the ray manifold and its projections.



Figure 8. Statistics for latitude band $50^{\circ}-60^{\circ}$. Left: number of data; right total relative difference of refractivity COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. Both are functions of the parameter $\beta$.


Figure 9. Statistics for latitude band $60^{\circ}-70^{\circ}$. Left: number of data; right total relative difference of refractivity COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. Both are functions of the parameter $\beta$.

The CT method has been applied for RO observations for a long time. Although there have been many modifications, like original CT combined with Back Propagation (BP), Full-Spectrum Inversion (FSI), Phase Matching (PM), and CT of type 2 (CT2), there is no essential difference between these FIO-based methods. The difference consists in the approximation of the phase function of the FIO, leading to the corresponding approximate representation of the impact parameter and bending angle, and in the specific implementation (such as cut-off, filtering, and quality control procedures). All these methods map the wave field to the representation of the impact parameter $p$. The reason for this choice of the coordinate in the mapped space is that in the case of a spherically-symmetric medium, the impact parameter is always a unique coordinate of the ray manifold.

Because the real atmosphere is not spherically-symmetric, this results in some aggravation. First, in the strict sense, there is no such a quantity as the impact parameter as a unique variable any more. But it is still possible to operate with the effective



Figure 10. Statistics for latitude band $70^{\circ}-80^{\circ}$. Left: number of data; right total relative difference of refractivity COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. Both are functions of the parameter $\beta$.


Figure 11. Statistics for latitude band $80^{\circ}-90^{\circ}$. Left: number of data; right total relative difference of refractivity COSMIC-ECMWF $\sqrt{\left\langle\left(N_{C}-N_{E}\right)^{2}\right\rangle} / N_{E}$. Both are functions of the parameter $\beta$.
impact parameter, derived from Doppler frequency shift using the same relations as for a spherically-symmetric medium. This quantity can be implemented in the observation operator for the variational assimilation of RO observation, cancelling errors due horizontal gradients. However, the above property of the impact parameter, which is supposed to be the unique coordinate of the ray manifold, does not always hold for the effective value. In some cases, the situation referred to as the impact multipath may occur, resulting in retrieval errors in atmospheric profiles derived from RO data.

In order to partially mitigate this fundamental shortcoming, we introduced a generalization of the CT approach. We used a generalized definition of the coordinate in phase space, defined as a linear combination of impact parameter and bending angle. Because this can be understood as an affine transform of the phase space, we coined the abbreviation CT2A for the new method. This transform has a parameter $\beta$, which can be tuned to minimize the retrieval error.

To find such a value of the parameter by statistical performance evaluation under real RO observation conditions including challenging horizontal gradients in the lower troposphere, we processed a large ensemble of COSMIC RO data for the year 2008,1 st and 15 th day of every month, adding up to a total of about 60000 RO events. We used the total relative difference of COSMIC from collocated ECMWF analysis profiles over the lower troposphere as the metric for this evaluation and the tuning parameter estimation.

For latitudes $0^{\circ}-50^{\circ}$, in the altitude range from 0.5 km to $1.9-2.5 \mathrm{~km}$, the application of the CT2A algorithm was used to statistically minimize the COSMC-ECMWF difference metric and the optimal value of parameter $\beta$ is found to be -6 to -8 . We found that the CT2A algorithm as well improves the penetration statistics of RO profile retrievals, increasing the number of data in the altitude range below 0.5 km .

Overall these results suggest that the CT2A method is not only theoretically an innovative generalization of the CT/FIO class of methods but also practically a valuable advancement for RO processing in that it can improve the capability to cope with challenging horizontal gradient conditions in the lower troposphere.

Author contributions. K.B.Lauritsen: The problem formulation, the initial idea of the study, theoretical discussion, contribution to finalizing the manuscript; M.Gorbunov: Theoretical derivations, numerical implementation, statistical study, initial draft of the paper; G.Kirchengast: theoretical discussions, contribution to finalizing the manuscript

Competing interests. The authors declare that no competing interests are present.

Acknowledgements. M.E. Gorbunov is grateful to Russian Foundation for Basic Research (grant No 20-05-00189 A) for the financial support. G. Kirchengast acknowledges support, including for partial co-funding of the work of M.E. Gorbunov, by the Aeronautics and Space Agency of the Austrian Research Promotion Agency (FFG-ALR) under the Austrian Space Applications Programme (ASAP) project ATROMSAF1 (proj.no. 859771) funded by the Ministry for Transport, Innovation, and Technology (BMVIT). K.B. Lauritsen has been supported by the Radio Occultation Meteorology Satellite Application Facility (ROM SAF) which is a decentralized operational RO processing center under EUMETSAT.

Discussions

## References

Arnold, V. I.: Mathematical Methods of Classical Mechanics, Springer-Verlag, New York, 1978.

Egorov, Y. V.: Lectures on Partial Differential Equations. Additional Chapters, Moscow State University Press, Moscow, (In Russian), 1985.
Egorov, Y. V. and Shubin, M. A., eds.: Partial Differential Equations IV, Springer Berlin Heidelberg, https://doi.org/10.1007/978-3-662-09207-1, https://www.springer.com/gp/book/9783540533634, 1993.
Fock, V. A.: Fundamentals of Quantum Mechanics, Mir Publishers, https://archive.org/details/FockFundamentalsOfQuantumMechanicsMir1986, 380 рр., 1978.

Goldstein, H., Poole, C., and Safko, J.: Classical Mechanics, Pearson Education Limited, https://books.google.ru/books?id=Xr-pBwAAQBAJ, 2014.

Gorbunov, M. E.: Canonical transform method for processing radio occultation data in the lower troposphere, Radio Sci., 37, 9-1-9-10, https://doi.org/10.1029/2000RS002592, 2002.
Gorbunov, M. E. and Gurvich, A. S.: Microlab-1 experiment: Multipath effects in the lower troposphere, J. Geophys. Res., 103, $13819-$ 13 826, https://doi.org/10.1029/98JD00806, 1998a.
Gorbunov, M. E. and Gurvich, A. S.: Algorithms of inversion of Microlab-1 satellite data including effects of multipath propagation, Int. J. Remote Sens., 19, 2283-2300, https://doi.org/10.1080/014311698214721, 1998b.
Gorbunov, M. E. and Kornblueh, L.: Analysis and validation of GPS/MET radio occultation data, J. Geophys. Res., 106, 17,161-17,169, https://doi.org/10.1029/2000JD900816, 2001.
Gorbunov, M. E. and Lauritsen, K. B.: Canonical Transform Methods for Radio Occultation Data, Scientific Report 02-10, Danish Meteorological Institute, Copenhagen, Denmark, http://www.romsaf.org/Publications/reports/sr02-10.pdf, http://www.dmi.dk/dmi/Sr02-10.pdf, 2002.

Gorbunov, M. E. and Lauritsen, K. B.: Analysis of wave fields by Fourier integral operators and its application for radio occultations, Radio Sci., 39, RS4010, https://doi.org/10.1029/2003RS002971, 2004a.

Gorbunov, M. E. and Lauritsen, K. B.: Canonical Transform Methods for Radio Occultation Data, in: Occultations for Probing Atmosphere and Climate, edited by Kirchengast, G., Foelsche, U., and Steiner, A. K., pp. 61-68, Institute for Geophysics, Astrophysics, and Meteorology, University of Graz, Springer, Berlin - Heidelberg - New-York, https://doi.org/10.1007/978-3-662-09041-1_6, 2004b.

Gorbunov, M. E. and Lauritsen, K. B.: Error Estimate of Bending Angles in the Presence of Strong Horizontal Gradients, in: New Horizons in Occultation Research, edited by Steiner, A., Pirscher, B., Foelsche, U., and Kirchengast, G., pp. 17-26, Springer, Berlin, Heidelberg, https://doi.org/10.1007/978-3-642-00321-9_2, 2009.

Gorbunov, M. E., Gurvich, A. S., and Bengtsson, L.: Advanced Algorithms of Inversion of GPS/MET Satellite Data and their Application to Reconstruction of Temperature and Humidity, Report No. 211, Max-Planck Institute for Meteorology, Hamburg, https://www.mpimet.mpg.de/en/science/publications/archive/mpi-reports-1987-2004/, 1996.

Gorbunov, M. E., Benzon, H.-H., Jensen, A. S., Lohmann, M. S., and Nielsen, A. S.: Comparative analysis of radio occultation processing approaches based on Fourier integral operators, Radio Sci., 39, RS6004, https://doi.org/10.1029/2003RS002916, 2004.

Healy, S. B.: Radio occultation bending angle and impact parameter errors caused by horizontal refractive index gradients in the troposphere: A simulation study, J. Geophys. Res., 106, 11,875-11,890, https://doi.org/10.1029/2001JD900050, 2001.
Hörmander, L.: The Analysis of Linear Partial Differential Operators, vol. III. Pseudo-Differential Operators, Springer-Verlag, New York, 1985a.

Hörmander, L.: The Analysis of Linear Partial Differential Operators, vol. IV. Fourier Integral Operators, Springer-Verlag, New York, 1985b. Jensen, A. S., Lohmann, M. S., Benzon, H.-H., and Nielsen, A. S.: Full spectrum inversion of radio occultation signals, Radio Sci., 38, 6-1-6-15, https://doi.org/10.1029/2002RS002763, 2003.

Jensen, A. S., Lohmann, M. S., Nielsen, A. S., and Benzon, H.-H.: Geometrical optics phase matching of radio occultation signals, Radio Sci., 39, RS3009, https://doi.org/10.1029/2003RS002899, 2004.
Karayel, T. E. and Hinson, D. P.: Sub-Fresnel scale resolution in atmospheric profiles from radio occultation, Radio Sci., 32, 411-423, https://doi.org/10.1029/96RS03212, 1997.

Kursinski, E. R., Hajj, G. A., Schofield, J. T., Linfield, R. P., and Hardy, K. R.: Observing Earth's atmosphere with radio occultation measurements using the Global Positioning System, J. Geophys. Res., 102, 23 429-23 465, https://doi.org/10.1029/97JD01569, 1997.

Melbourne, W. G., Davis, E. S., Duncan, C. B., Hajj, G. A., Hardy, K. R., Kursinski, E. R., Meehan, T. K., and Young, L. E.: The Application of Spaceborne GPS to Atmospheric Limb Sounding and Global Change Monitoring, JPL Publ. 94-18, Jet Propul. Lab., Pasadena Calif., 1994.

Mishchenko, A. S., Shatalov, V. E., and Sternin, B. Y.: Lagrangian manifolds and the Maslov operator, Springer-Verlag, Berlin - New York, 1990.

Mortensen, M. D. and Høeg, P.: Inversion of GPS occultation measurements using Fresnel diffraction theory, Geophys. Res. Lett., 25, 24412444, https://doi.org/10.1029/98GL51838, 1998.

Treves, F.: Intoduction to Pseudodifferential Operators and Fourier Integral Operators, vol. 1, Pseudodifferential Operators, Plenium Press, New York and London, 1982a.

Treves, F.: Intoduction to Pseudodifferential Operators and Fourier Integral Operators, vol. 2, Fourier Integral Operators, Plenium Press, New York and London, 1982b.

Ware, R., Exner, M., Feng, D., Gorbunov, M., Hardy, K., Herman, B., Kuo, Y.-H., Meehan, T., Melbourne, W., Rocken, C., Schreiner, W., Sokolovsky, S., Solheim, F., Zou, X., Anthes, R., Businger, S., and Trenberth, K.: GPS Sounding of the Atmosphere from Low Earth Orbit: Preliminary Results, Bull. Amer. Meteor. Soc., 77, 19-40, https://doi.org/10.1175/1520-0477(1996)077<0019:GSOTAF>2.0.CO;2, 1996.

Zou, X., Liu, H., and Anthes, R. A.: A Statistical Estimate of Errors in the Calculation of Radio-Occultation Bending Angles Caused by a 2D Approximation of Ray Tracing and the Assumption of Spherical Symmetry of the Atmosphere, J. Atmos. Oceanic Technol., 19, 51-64, https://doi.org/10.1175/1520-0426(2002)019<0051:ASEOEI>2.0.CO;2, 2002.

Zou, X., Liu, H., and Kuo, Y.-H.: Occurrence and detection of impact multipath simulations of bending angle, Quart. J. Roy. Meteor. Soc., https://doi.org/10.1002/qj.3520, 2019.

