

# 1 Volcanic SO<sub>2</sub> Effective Layer Height Retrieval for OMI Using a 2 Machine Learning Approach

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11  
12 **Abstract.** Information about the height and loading of sulfur dioxide (SO<sub>2</sub>) plumes from  
13 volcanic eruptions is crucial for aviation safety and for assessing the effect of sulfate aerosols on  
14 climate. While SO<sub>2</sub> layer height has been successfully retrieved from backscattered Earthshine  
15 ultraviolet (UV) radiances measured by the Ozone Monitoring Instrument (OMI), previously  
16 demonstrated techniques are computationally intensive and not suitable for near real-time  
17 applications. In this study, we introduce a new OMI algorithm for fast retrievals of effective  
18 volcanic SO<sub>2</sub> layer height. We apply the Full Physics Inverse Learning Machine (FP\_ILM )  
19 algorithm to OMI radiances in the spectral range of 310-330 nm. This approach consists of a  
20 training phase that utilizes extensive radiative transfer calculations to generate a large dataset of  
21 synthetic radiance spectra for geophysical parameters representing the OMI measurement  
22 conditions. The principal components of the spectra from this dataset in addition to a few  
23 geophysical parameters are used to train a neural network to solve the inverse problem and  
24 predict the SO<sub>2</sub> layer height. This is followed by applying the trained inverse model to real OMI  
25 measurements to retrieve the effective SO<sub>2</sub> plume heights. The algorithm has been tested on  
26 several major eruptions during the OMI data record. The results for the 2008 Kasatochi, 2014  
27 Kelud, 2015 Calbuco, and 2019 Raikoke eruption cases are presented here and compared with  
28 volcanic plume heights estimated with other satellite sensors. For the most part, OMI-retrieved  
29 effective SO<sub>2</sub> heights agree well with the lidar measurements of aerosol layer height from Cloud-  
30 Aerosol Lidar and Infrared Pathfinder Satellite Observations (CALIPSO) and thermal infrared  
31 retrievals of SO<sub>2</sub> heights from the infrared atmospheric sounding interferometer (IASI). The  
32 errors in OMI retrieved SO<sub>2</sub> heights are estimated to be 1-1.5 km for plumes with relatively large  
33 SO<sub>2</sub> signals (> 40 DU). The algorithm is very fast and retrieves plume height in less than 10 min  
34 for an entire OMI orbit. ~~This approach offers a promising prospect of using physics-based  
35 machine learning applications to other instruments.~~

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## 38 **1 Introduction**

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40 The observation and tracking of emissions from volcanic eruptions are crucial for both air traffic  
41 safety and for assessing climate forcing impacts from volcanic sulfate aerosols. In the last 10  
42 years, volcanoes have emitted roughly 20-25 million metric tons of sulfur dioxide (SO<sub>2</sub>) per year  
43 through passive degassing (Carn et al, 2017). Explosive volcanic eruptions, however, can  
44 additionally release large SO<sub>2</sub> amounts high into the atmosphere. SO<sub>2</sub> can be converted to sulfate  
45 aerosols within 2-3 days in the troposphere (Lee et al., 2011) and within a few weeks in the  
46 lower stratosphere (von Glasow et al., 2009, Krotkov et al., 2010). Sulfate aerosols are known to  
47 have a cooling effect on climate, especially if an SO<sub>2</sub> plume is injected into the lower  
48 stratosphere and remains there for longer periods of time. This is demonstrated by significant  
49 eruptions such as Mt. Pinatubo in 1991 that temporarily reduced global temperatures by up to  
50 0.5°C (McCormick et al, 1995). Aside from releasing SO<sub>2</sub>, volcanoes also emit large amounts of  
51 ash into the atmosphere which can have adverse impacts on air travel. Ash from volcanic plumes  
52 can often interfere with flight paths, greatly reduce visibility near the ground, and cause damage  
53 to the aircraft including engine failure (Carn et al., 2009). In addition, SO<sub>2</sub> causes sulfidation in  
54 the engines, an effect that can reduce their lifetimes in the long term. From 1953 to 2009, over  
55 120 aviation incidents involving volcanic activity were reported, with roughly 80 of them  
56 involving serious damage to the airframe or engine (Guffanti et al., 2010). There is also the  
57 possibility of highly concentrated volcanic SO<sub>2</sub> plumes producing acidic aerosols which can  
58 cause irritation of the eyes, nose and respiratory airways of occupants inside airplanes (Schmidt  
59 et al., 2014). In many cases SO<sub>2</sub> and ash are often collocated, thus making estimates of SO<sub>2</sub> layer  
60 height useful for aviation hazard mitigation and volcanic plume forecasting. Lastly, the accurate  
61 determination of SO<sub>2</sub> height can ideally aid in producing accurate SO<sub>2</sub> vertical column depth  
62 (VCD) estimates given that those retrievals typically use a fixed *a priori* vertical distribution of  
63 SO<sub>2</sub> in the absence of additional information on SO<sub>2</sub> height.

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65 With remote sensing, these volcanic plumes can be regularly observed from space. In  
66 particular, hyperspectral spectrometers such as the Ozone Monitoring Instrument (OMI),  
67 GOME-2, OMPS, TROPOMI and others, have provided frequent and increasingly accurate  
68 observations of global SO<sub>2</sub> amounts, through retrieval algorithms from backscattered radiance

68 measurements. The OMI instrument, a Dutch-Finish contribution to the NASA Aura satellite,  
69 has been operational since 2004. OMI has 60 cross track positions (rows) and has a  $13 \times 24 \text{ km}^2$   
70 spatial resolution at the nadir position (Levelt et al., 2006). The instrument uses two UV channels  
71 and one visible channel to measure backscattered radiances from the Earth's atmosphere. About  
72 half of the OMI rows are affected by the row anomaly which affects the quality of OMI Level 1  
73 and Level 2 data. This anomaly affects individual rows and slowly evolves over time. It is  
74 thought to occur due to a physical obstruction caused by the loosening of material on the interior  
75 of sensor (Torres et al., 2018). In general,  $\text{SO}_2$  slant column amounts are retrieved from these  
76 measurements through the differential optical absorption spectroscopy (DOAS) technique and  
77 then converted to vertical columns using Air Mass Factors (AMFs). The 310.5-340 nm range in  
78 OMI's UV2 channel is used in retrieving  $\text{SO}_2$ , with focus on the 310.8 and 313 nm wavelengths.  
79 The band residual algorithm (Krotkov et al., 2006) and the Linear Fit (LF) algorithm (Yang et  
80 al., 2007) were first used as the OMI operational algorithms for retrieving planetary boundary  
81 layer (PBL)  $\text{SO}_2$  and volcanic  $\text{SO}_2$  vertical column densities (VCDs) respectively. These were  
82 replaced with the principal component analysis (PCA) based algorithm (Li et al., 2013) which  
83 retrieves  $\text{SO}_2$  amounts directly from spectral radiance measurements. The same technique was  
84 also applied to OMI volcanic  $\text{SO}_2$  retrievals (Li et al., 2017). This data-driven approach does not  
85 rely on extensive radiative transfer modeling and has led to reduced biases and significant  
86 improvements (Fioletov et al., 2015). For volcanic retrievals, algorithms still have uncertainties  
87 in  $\text{SO}_2$  mass in volcanic plumes, especially in the presence of relatively larger errors in the  
88 assumed *a priori* profiles.

89 In addition to column amounts, backscattered radiances can provide important  
90 information about the height of an  $\text{SO}_2$  layer. Conceptually, a change in altitude of an  $\text{SO}_2$  plume  
91 alters the number of backscattered photons going through the layer. If a plume is high in the  
92 atmosphere, more photons that are scattered from below the layer pass through the absorbing  
93  $\text{SO}_2$  plume. This results in larger  $\text{SO}_2$  absorption structures in the measured radiance spectra,  
94 especially in the 310-320 nm range where Rayleigh scattering is dominant. Relative to the  $\text{SO}_2$   
95 amount, obtaining a fast retrieval of the height of a volcanic plume presents a greater challenge.  
96 Until recently, retrieval techniques have involved a direct spectral fitting approach that use  
97 backscattered ultraviolet (BUV) measurements in conjunction with extensive forward radiative  
98 transfer modeling. For instance, the Iterative Spectral Fitting (ISF) algorithm (Yang et al., 2009)

99 for OMI was utilized to determine the altitude of SO<sub>2</sub> layer by adjusting the height while  
100 minimizing the differences between measured radiances and forward RT calculations. Another  
101 study has utilized an optimal estimation algorithm along with the VLIDORT radiative transfer  
102 (RT) model to retrieve SO<sub>2</sub> density and plume height from the GOME-2 instrument (Nowlan et  
103 al., 2011). Sulfur dioxide amounts and plume heights have also been estimated with the infrared  
104 atmospheric sounding interferometer (IASI), through brightness temperature changes and  
105 relative intensities of absorption lines (Clarisse et al., 2008). For these techniques, extensive  
106 radiative transfer modeling is needed, in addition to a variety of assumptions including a  
107 reasonable first guess for the plume altitude. Newer schemes were later developed for GOME-2  
108 using the SOPHRI algorithm (Rix et al., 2012), a DOAS based technique that included  
109 minimizing differences between plume height from simulated spectra and the assumed height  
110 from measured spectra. This technique allowed for reasonably fast retrievals that could be used  
111 in near real-time, thanks to the use of pre-calculated GOME spectra stored in a look up table  
112 classified according to SO<sub>2</sub> column, SO<sub>2</sub> heights and other physical parameters. An updated  
113 algorithm was also developed for IASI (Clarisse et al., 2014), this time implementing an optimal  
114 estimation fit approach with pre-calculated Jacobians. Faster and more efficient methods for  
115 GOME-2 (Efremenko et al., 2017) and TROPOMI (Hedelt et al., 2019) have made use of  
116 machine learning algorithms, specifically neural networks (NNs), to develop a trained, full-  
117 physics inverse learning machine (FP\_ILM) for retrieving SO<sub>2</sub> plume height. This approach has  
118 shown good accuracy and speed fast enough for near-real-time operations. The FP\_ILM has also  
119 been used for retrieving ozone profile shapes (Xu et al., 2017) and geometry-dependent  
120 Lambertian equivalent reflectivity (Loyola et al., 2020). The primary advantage of this approach  
121 is the execution speed. By separating the training phase, which involves large amounts of time  
122 consuming radiative transfer computations and machine learning model training, from the  
123 application phase, the desired parameter can be retrieved within milliseconds for a single satellite  
124 ground pixel using the inverse model. However, similar methods of retrieving SO<sub>2</sub> layer height  
125 have not yet been implemented for OMI. Now in this study, the FP\_ILM has been applied to  
126 OMI to estimate SO<sub>2</sub> layer height from backscattered Earthshine radiance measurements. The  
127 retrieval was tested on four past volcanic eruption cases and performance was assessed through  
128 machine learning metrics, as well as comparisons to other datasets such as those from  
129 TROPOMI, IASI and CALIOP lidar instruments.

## 130 **2 Methodology:**

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132 The FP\_ILM approach consists of two parts, the training phase and the application (or  
133 operational) phase. The training phase starts with the generation of a synthetic training dataset of  
134 top of the atmosphere (TOA) reflectance spectra from a radiative transfer model. This spectral  
135 dataset is then used to train a Multi-Layer Perceptron Regression (MLPR) NN model to predict  
136 the SO<sub>2</sub> layer height as an output. In the application phase, the trained inverse model is applied to  
137 real OMI radiance measurements. This inverse model is optimized from the training, and the  
138 predictions of SO<sub>2</sub> layer height based on the model are very fast as compared with the time-  
139 consuming RT calculations during the training phase. The main steps of the algorithm are shown  
140 in a flowchart (Figure 1) and discussed in detail in the next sections.

### 141 142 **2.1 Forward Radiative Transfer Model**

143  
144 The first step in the training phase is to build a large data set of synthetic backscattered  
145 Earthshine reflectance spectra from forward radiative transfer (RT) calculations. These  
146 calculations are performed using the Linearized Discrete Ordinate Radiative Transfer (LIDORT)  
147 model with the rotational Raman scattering (RRS) capability (Spurr et al., 2008). This version of  
148 the model treats first-order inelastic Raman scattering in addition to all orders of elastic  
149 (Rayleigh) scattering processes. Rotational Raman scattering occurs when a photon is scattered  
150 at lower or higher energy levels than the incident radiation. RRS cannot be neglected; it is known  
151 to be responsible for the Ring effect (Grainger and Ring 1962), a spectral interference signature  
152 characterized by the filling-in of Fraunhofer lines and telluric-absorber features. Allowing for  
153 RRS in the RT model leads to differences in calculated radiances compared to those made with  
154 purely elastic scattering, as characterized by the filling-in factor. This quantity is generally of the  
155 order of a few percent, consistent with estimates that 4% of the total scattering in the  
156 atmospheric is inelastic (Young, 1981). Fundamentally the SO<sub>2</sub> layer height information can be  
157 retrieved by backscattered radiance spectra because the amount of scattering occurring in the  
158 overlying atmosphere is determined by the height of the volcanic SO<sub>2</sub> plume. This is  
159 demonstrated by comparing two otherwise identical RT calculations with different SO<sub>2</sub> layer  
160 heights (Figure 2a). At shorter wavelengths where Rayleigh scattering is stronger, there is less  
161 backscattered radiance for the case with higher SO<sub>2</sub> plume height, particularly at shorter  
162 wavelengths < 320 nm (Figure 2b). Likewise, the filling-in factor (Figure 2c) shows the

163 importance of including RRS in the RT calculations as in some cases there can be 2-3%  
164 difference between the Raman and elastic calculations.

165 All LIDORT-RRS calculations in this study were performed for the 310-330 nm spectral  
166 range, which captures strong SO<sub>2</sub> and ozone absorption features. The model is supplied with  
167 ozone (Daumont et al., 1992) and SO<sub>2</sub> absorption (Bogumil et al., 2003) cross sections,  
168 atmospheric profile, ozone profile and a high resolution Fraunhofer solar irradiance spectrum.  
169 The atmospheric profile has 48 layers and contains a temperature/pressure/height grid from the  
170 standard US atmosphere, with an increased vertical resolution of 0.5 km below 12 km. The  
171 ozone profile is determined by the total column amount, latitude zone and month as specified in  
172 the TOMS V7 ozone profile climatology (Bhartia, 2002), while the SO<sub>2</sub> profile is assumed to be  
173 a Gaussian shape with a full width half maximum (FWHM) of 2.5 km. The solar spectrum is a  
174 re-gridded version of the high resolution synthetic solar reference spectrum (Chance and Kurucz,  
175 2010), originally with a spectral resolution of 0.01 nm. The re-gridded version has a resolution of  
176 0.05 nm, finer than that for OMI (0.16 nm sampling for a FWHM spectral resolution of ~0.5  
177 nm). The advantage of using this reference spectrum over the instrument-measured irradiance is  
178 that only one set of calculations is needed; they can be applied to multiple instruments and  
179 instrument cross track positions without utilizing unique measured solar flux spectra for each  
180 situation. Using instrument-measured solar flux data may carry less potential error and be able to  
181 better handle issues with instrument degradation. However, the downside is that the inverse  
182 model would need to be re-trained whenever a new measured solar flux spectrum is used. Since  
183 we expect the retrieval to be primarily sensitive to SO<sub>2</sub> absorption signatures, the radiative  
184 transfer calculation was performed for a molecular atmosphere with no aerosol scattering.

185 In order to obtain a large number of different spectra, eight key physical parameters were  
186 varied for the LRRS calculations. These parameters include solar zenith angle (SZA), relative  
187 azimuth angle (RAA), viewing zenith angle (VZA), surface albedo, surface pressure, O<sub>3</sub> column  
188 amount, SO<sub>2</sub> column amount and SO<sub>2</sub> layer height. The ranges of these parameters are given in  
189 Table 1.

190 The number of calculations and the parameter sets for each simulation were determined through  
191 a smart sampling technique (Loyola et al. 2016). A selective parameter grid with sets of  
192 parameters for each simulation was established through the use of Halton sequences (Halton,  
193 1962) in 8 dimensions. The calculations are continued until the moments of the output data,

194 mean and median converged across all wavelengths. In total around 200,000 calculations were  
195 done to achieve sufficiently comprehensive sample size for the variation in the eight parameters  
196 across all rows of OMI. This sampling was done in order to ensure that 1) each set of parameters  
197 was unique and training data is diverse; and 2) that the sample size of the entire dataset is large  
198 enough for the machine learning application.

## 199 200 **2.2 Data pre-processing**

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202 After the RT calculations are completed, the spectra are convolved with OMI instrument slit  
203 function. Since each cross-track position of OMI contains a unique slit function, the appropriate  
204 function was applied based on the VZA input for that particular calculation. The VZA ranges  
205 from 0-70° across all rows in the OMI swath, with the middle (nadir) rows having a VZA of  
206 close to 0. For each row, only spectra within +/- 3° of the actual VZA were convolved with the  
207 appropriate slit functions. In addition, Gaussian noise with a signal-to-noise ratio (SNR) of 1000  
208 was added to the spectra. While the SNR of OMI tends to be lower (Schenkeveld et al., 2017),  
209 adding too much noise can greatly decrease performance of the machine learning (Table 2). The  
210 root mean squared error (RMSE) and mean absolute difference (MAE) between the SO<sub>2</sub> height  
211 from the RT calculation parameter sets and the height predicted by the neural network were used  
212 as metrics (see Section 3). At SNRs of less than 500, the performance starts to increasingly  
213 degrade. Between 1000 and 500 SNR, there is an increase of around 0.1 km in RMSE. However,  
214 adding some degree of noise is necessary to account for errors in satellite instrument  
215 measurements.

216 Next, principal component analysis (PCA) was applied to the spectral dataset for each  
217 row, in order to extract the most significant features of the spectra, and to reduce dimensionality.  
218 Since each convolved sample consists of 142 wavelength points, the dimensionality of this  
219 problem becomes very large. However, PCA transforms each sample to a set of weights based on  
220 8 principal components (PCs). These principal components explain 99.998% of the variance in  
221 the synthetic dataset (Figure 3). Including additional PCs does not add any significant value to  
222 the retrieval and may even lead to overfitting. Prior to starting the machine learning process, the  
223 dataset is split into a training subset (90%) and a testing subset (10%). The training subset is used  
224 for the neural network learning, while the testing subset only deployed verifying the performance  
225 of the network to predict the output.

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### 2.3 Machine Learning using a Neural Network

The 8 PCs, and selected parameters including the SZA, RAA, VZA, surface pressure and surface albedo were used as input for training a MLPR, which is sometimes referred to as a deep neural network. The output layer of the NN contains the effective SO<sub>2</sub> layer height. Column amounts of SO<sub>2</sub> and O<sub>3</sub> were not included in the training or in the application stage because of the large dependency of column amounts on SO<sub>2</sub> layer height and due to biases in OMI ozone retrieval in the presence of the enhanced SO<sub>2</sub> plume, respectively. To improve stability, the inputs (PC weights, SZA, VZA, etc.) and output (effective SO<sub>2</sub> height) are scaled between -0.9 and 0.9 according to the minimum and maximum of each input variable prior to input into the NN. In a NN, the input and output layers are connected by hidden layers containing neurons (also known as nodes). Each neuron is connected to others by a series of weights, by means of which the input data is passed to the next level as a weighted sum of all inputs. Inside the neural network, the Adam optimizer with a stochastic gradient descent algorithm (Kingman et al., 2014) is used to minimize the loss function, in this case the mean squared error (MSE) between the result of each iteration and the actual SO<sub>2</sub> layer height used to generate the synthetic spectral sample. With each iteration, the partial derivative of the MSE with respect to each node is calculated; this is used to update the weights. The training of a NN progresses by cycling through iterations of the entire training dataset, called epochs, until the training and validation MSE is minimized and there is no improvement to be obtained from further training. Throughout the training, the NN uses 10% of the training subset for validation to assess the performance with each iteration. This validation set is different from the independent test data that was set aside from training. The “tanh” (hyperbolic tangent) activation function is applied at the hidden layers to further increase stability in the NN. Other activation functions (e.g., ReLU and PReLU) were tested, however tanh was found to produce slightly better NN performance. There is also considerable flexibility in the structure of the NN, in particular the number of hidden layers and nodes in each layer. The final configuration of the NN in this study includes 2 hidden layers with 20 and 10 nodes in the first and second layer, respectively. This was determined through testing and analyzing the errors of the NN with respect to the synthetic test data set and the quality of the retrieval results after application to satellite measurements. During testing, more complex configurations of hidden layers and number of neurons were found to have worse performance



258 when after applying to OMI data. Hence the relatively simple configuration was chosen as the  
259 final setup for this study.

260 In neural networks a common problem known as overfitting often occurs when the  
261 machine learning model is tuned so closely to the training inputs that it does not perform well on  
262 new data. During training this can be diagnosed if the validation error (loss) decreases at a much  
263 slower rate than the training error. To reduce overfitting the L2 regularization technique was  
264 implemented in the training. The regularization reduces the effect of small and very large weight  
265 values by penalizing the MSE loss function. For this study, the training was done separately for  
266 each OMI row due to the different VZAs and slit functions between rows; however, the  
267 configuration of the NN was kept constant between rows. The only difference in the training is  
268 the number of training epochs conducted for each row before the solution becomes optimal for  
269 that row. The number of epochs varies slightly but is in the 200-300 range for all rows. The final  
270 trained version of the NN, the inverse operator, contains the optimal weights needed to predict  
271 the SO<sub>2</sub> layer height from an input of separate test data.

272 An important aspect for neural network performance is the number of training samples.  
273 Aside from smart sampling, the appropriate number of samples for training can be determined by  
274 comparing errors from training runs where different percentages of training samples were  
275 removed (e.g. 10%, 20%, 50%) beforehand. The mean absolute error between height predicted  
276 by the NN and the test set height was calculated when using different numbers of input samples.  
277 With a 50% reduction in training samples, the absolute error went up by around 0.3 km. In  
278 contrast, reducing the training set by 10% had little impact on the error (see Table A1). These  
279 results provide confirmation that for this case the training data are adequate, and that there would  
280 likely be diminishing returns in NN performance with a larger training dataset.

## 281 **2.4 Application to satellite measurements**

282 In the application phase of the retrieval, the inverse operator is applied to OMI radiance  
283 spectra, resulting in a predicted SO<sub>2</sub> layer height for each ground pixel in the OMI swath. For  
284 this the OMI L1B Geolocated Earthshine radiance dataset is used. Since OMI only provides  
285 absolute radiances, these data were normalized with respect to the same solar flux spectrum as  
286 used in the generation of the synthetic spectra. In other words, the measured input becomes the  
287 fraction of backscattered radiance to the incoming solar irradiance (i.e., reflectance spectrum).  
288 Prior to normalizing, the irradiance spectrum was convolved with an OMI slit function for the

289 particular OMI row and orbit. The output is a predicted SO<sub>2</sub> layer height based on the input of a  
290 radiance spectra and associated parameters, including VZA, SZA, RAA, surface albedo and  
291 surface pressure, for a single OMI pixel. The irradiance spectrum is convolved with the  
292 appropriate OMI slit function in order to have consistency in wavelength points between the  
293 measured radiances, synthetic radiances and irradiance of each row. To follow the same  
294 procedure as was used in the training step, the PCA operator from the training phase is applied to  
295 the OMI spectra to perform the dimensionality reduction and obtain a set of PC weights for each  
296 sample. The other inputs are VZA, SZA, RAA, albedo and surface pressure parameters from the  
297 OMI data files. As in the training phase, all inputs are scaled to the [-0.9, 0.9] range. After SO<sub>2</sub>  
298 heights are retrieved separately for each row, one height value is given for each pixel (and  
299 spectral sample). The application phase of the retrieval takes only 2-3 seconds for a given row.  
300 This short duration includes the application of the training phase PCA operator to OMI  
301 measurements, the scaling of inputs and the deployment of the inverse operator. The whole  
302 process is repeated for each row in order to get a prediction for an entire OMI swath. For some  
303 rows the retrieval is unreliable due to the row anomaly, which negatively affects the quality of  
304 the OMI L1B radiance data at all wavelengths and consequently L2 retrievals.

### 305 **3 Impacts of various parameters on the performance of the trained inverse model**

306 From the training phase, it becomes clear that the performance of the algorithm will  
307 depend on several factors. As demonstrated in Fig. 3, an important factor is the SO<sub>2</sub> column  
308 amount. Overall, the NN makes better predictions for the test data subset for SO<sub>2</sub> amounts > 40  
309 DU. Below 40 DU, information content on the layer height to be retrieved becomes increasingly  
310 small, as evidenced by large differences between predicted heights and those in the actual test set  
311 (Figure 5a). Additionally, larger SO<sub>2</sub> loadings result in greater sensitivity between two heights,  
312 as seen by comparisons of SO<sub>2</sub> height Jacobians for multiple amounts (Figure 4). Quantitatively,  
313 if samples with SO<sub>2</sub> amounts less than 40 DU are excluded, the RMSE decreases from 1.48 to  
314 1.15 km (Table 3). As with other sensitivity analyses, the RMSE and MAE in Table 3 are  
315 calculated between the predicted output from NN and the height from the independent test  
316 dataset. We can therefore expect the retrieval to produce reasonable results for moderate to large  
317 volcanic eruptions. In widely dispersed plumes where the SO<sub>2</sub> VCD is low or for volcanic  
318 degassing events, the retrieval would be less accurate. The second major dependency is on SZA.  
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321 The problem here stems from the occurrence of relatively large errors in RT modeling due to  
322 shallow light paths and lower OMI SNR at the higher SZAs. Reasonably accurate results are to  
323 be expected only for  $SZA < 75^\circ$ . Figure 2b shows significant differences in predicted and actual  
324 heights in spectra associated with large SZAs, after removal of low VCD samples. For the final  
325 training approach, it was therefore necessary to exclude spectra with large SZAs. Dependencies  
326 on other physical parameters are small when compared with these two issues discussed here,  
327 although there is some evidence that high surface albedo also increases error. If we remove  
328 spectra with albedo  $> 0.6$  there is a minor improvement in RMSE from 0.93 to  $\sim 0.89$  km.  
329 However, even with strong volcanic  $SO_2$  signals, we can realistically expect that on average the  
330 absolute error to be at least 1 km, due to inherent simplifications in the neural network retrieval  
331 approach. The errors in actual retrievals using OMI data are expected to be larger (see Section  
332 4.4).

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#### 334 **4. OMI $SO_2$ Effective Layer Height Results**

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336 For testing the FP\_ILM retrieval on OMI data, four volcanic eruption cases with sufficiently  
337 strong  $SO_2$  signals were selected (i.e. where peak  $SO_2$  VCDs were greater than 40 DU). Each  
338 case is described in detail in the following subsections. For each case, comparisons were made to  
339 other satellite-derived datasets where available, for example the CALIOP lidar onboard  
340 CALIPSO, the IASI  $SO_2$  layer height retrieval (Clarisse et al., 2014), and the GOME-2  
341 (Efremenko et al., 2017) and TROPOMI retrievals (Hedelt et al., 2019). It is important to note  
342 that the CALIOP lidar only indicates the height of the ash plume and not the  $SO_2$  height.  
343 Although ash and  $SO_2$  plumes are often collocated, this is not always the case, making direct  
344 comparisons difficult.

345

##### 346 **4.1 Kasatochi (2008)**

347 Kasatochi is a volcano located on the Aleutian Islands of Alaska ( $52.178^\circ N, 175.508^\circ W$ ). It  
348 underwent a series of eruptions beginning late in the day on August 7<sup>th</sup>, 2008, which injected  
349 great amounts of ash and  $SO_2$  into the stratosphere. Overall the explosion released roughly 2  
350 million tons of  $SO_2$ , at the time the highest  $SO_2$  loading since the Mt Pinatubo eruption (Yang et  
351 al, 2010).  $SO_2$  effective layer heights retrieved using the machine learning model for OMI (orbit  
352 21650) on August 10<sup>th</sup>, 2008, were around 11-12 km with some portions being slightly lower

353 (Figure 6a). This is in reasonable agreement with previous SO<sub>2</sub> height retrievals of 9-11 km  
354 which used the ISF algorithm for OMI (Yang et al., 2010), considering that the uncertainty of  
355 both retrievals are around 2 km. Likewise, Nowlan et al. (2011) showed that the majority of the  
356 plume was around 10 km, and up to 15 km in some parts. There is also agreement with IASI  
357 (Figure 6b) and CALIOP data (Figure 6d) which showed plume heights of 10-12 km and 12.5  
358 km respectively. It is important to note that the IASI overpass occurred later in the day than those  
359 for OMI and CALIPSO. Another verification source we used was the GOME-2 SO<sub>2</sub> layer height  
360 retrieval that uses FP\_ILM (Efremenko et al., 2017). The study found a height of around 10 km  
361 and up to 14 km in areas of high SO<sub>2</sub> loading for August 10<sup>th</sup> (Figure 6c). The GOME-2 overpass  
362 occurred 4 hours earlier than OMI. The mean, median, standard deviation and the inner quartile  
363 range (IQR) of the three retrievals (Table 4) also show good agreement for this case. Although  
364 the OMI results agree well in general with the results of these studies and datasets, the retrieval is  
365 less sensitive with respect to detecting variability in the SO<sub>2</sub> layer height within the plume.

#### 366 367 **4.2 Kelud (2014)**

368 Kelud, a stratovolcano located in East Java, Indonesia (7.935°S, 112.315°E), erupted on  
369 February 13<sup>th</sup>, 2014 at 1550 UTC, in the process depositing ash in a 500 km diameter around the  
370 volcano and leading to mass evacuations from nearby towns. Even though this case has  
371 somewhat lower SO<sub>2</sub> VCDs than those from Raikoke and Kasatochi, the peak SO<sub>2</sub> VCDs of ~60-  
372 70 DU should still allow for retrievals with reasonable accuracy (see section 2). The OMI  
373 retrieval results indicate that the maximum height of the main plume was 18-19 km (Figure 7a),  
374 although other studies suggest that several smaller layers of SO<sub>2</sub> and ash were located as high as  
375 26 km (Vernier et al., 2016) on the previous day. However, the SO<sub>2</sub> loading at that level was  
376 most likely too low for an accurate retrieval using OMI radiances. CALIOP lidar detected ash  
377 plumes at around 19.5 km and the IASI retrievals registered the plume at 17.5 km over the same  
378 area as that for OMI. The height of the ash plume from this eruption was also estimated using  
379 Multifunctional Transport Satellite (MTSAT 2) observations and transport modeling (Kristiansen  
380 et al., 2015). That study found an injected height of around 17 km, which is in agreement with  
381 the OMI result, especially when considering the most probable heights on the PDF (Figure 8b).  
382 We note here that only a small portion of the plume was retrieved with our algorithm, given the  
383 relatively low SO<sub>2</sub> VCDs and interference due to the OMI row anomaly. It is promising to note

384 that the OMI retrieval was able to identify heights at the upper end of the height range used in  
385 the training phase. On the other hand, while the retrieval can extrapolate to heights above 20 km,  
386 the accuracy would likely degrade due to the lack of training data with heights outside of this  
387 limit.

### 388 389 **4.3 Calbuco (2015)**

390  
391 The Calbuco volcano is located in Chile (41.331°S, 72.609°W). The primary eruption  
392 had a volcanic explosivity index (VEI) of 4 and occurred on April 22nd with little warning. The  
393 primary plume ascended higher than 15 km, while plumes from smaller subsequent eruptions  
394 stayed in the troposphere. The volcanic plume spread northeast in the following days, resulting in  
395 flight cancellations at Uruguayan and south Brazilian airports. The OMI-retrieved SO<sub>2</sub> effective  
396 layer heights in the area of greatest VCD was in the 15-17 km range. In the same region, IASI  
397 results (Figure 7c) show similar plume heights, approximately around 15 km, although as with  
398 the previous events, the overpass times of the two instruments are different. CALIOP lidar shows  
399 the ash plume to at roughly 17 km (Figure 7e). Unfortunately, the overpass of CALIPSO occurs  
400 over an area of OMI's swath that is affected by the row anomaly, and this makes a direct  
401 comparison unfeasible. Nevertheless, the CALIPSO aerosol layer height is still comparable to  
402 OMI-retrieved effective SO<sub>2</sub> layer heights for the portion of the plume further to the west. The  
403 retrieval for OMI is consistent with the other instruments for SO<sub>2</sub> plumes, with the exception of  
404 that part of the plume with SO<sub>2</sub> below 30-40 DU (see Figure A1), for which results were not  
405 plotted in Figure 7a due to lower biases.

### 406 407 **4.4 Raikoke (2019)**

408 The eruption of the Raikoke stratovolcano (48.2932°N, 153.254°E), located on the Kuril Islands  
409 of Russia, occurred on June 21<sup>st</sup>, 2019 at 1800 UTC. A series of explosions during the eruption  
410 sent large amounts of ash and SO<sub>2</sub> into the lower stratosphere. Maximal loadings of SO<sub>2</sub>  
411 measured by OMI and other sensors exceeded 500 DU. In the following days the plume  
412 underwent dispersion and spread out over the northern Pacific Ocean and later over eastern  
413 Russia. Early estimates of plume injection height for the eruption were predominantly in the 10-  
414 13 km range with potentially larger heights in some areas of the plume. In Figures 9a and 9b, the  
415 SO<sub>2</sub> effective layer heights retrieved from OMI data are shown for the Raikoke plume on June

416 23<sup>rd</sup> and June 24<sup>th</sup> respectively. The plume heights for both days are predominantly in the range  
417 10-12 km, although some areas of the plume had estimated peak heights of 13-14 km. In  
418 comparison, the TROPOMI results show slightly larger heights (13-14 km) for June 24<sup>th</sup> and  
419 similar heights to OMI for June 23<sup>rd</sup> (Figure 9c and d). The IASI SO<sub>2</sub> height product also shows  
420 fairly good agreement, with heights mainly at the 10-11 km level (Figure 9e and f). It is also  
421 useful to look at a distribution of heights predicted for the domain (Figure 10) in order to get a  
422 more quantitative comparison between the datasets. Based on this distribution, there is clearly at  
423 least 2 km difference between the most probable heights from OMI and those from TROPOMI  
424 for June 24<sup>th</sup> (Figure 10b and d) and slightly lower heights in the distribution for IASI. This is  
425 also displayed in Table 4 which shows a 2-3 km difference in the mean and median of retrieved  
426 heights between OMI and TROPOMI. Additionally the IQR and standard deviation provide a  
427 quantitative measure of the variation in the distribution of the retrieved heights, which can  
428 change from one orbit to another. Note that points with lower than 30 DU are not included in the  
429 PDFs for all sensors. The results are also compared with CALIOP lidar onboard CALIPSO,  
430 which shows ash plume heights of 12-13 km for both days (Figure 11a and 11b). Although there  
431 is overestimation for some OMI pixels, especially for June 24<sup>th</sup>, the section of the plume with the  
432 CALIPSO flyover has similar heights (around 12.5 km) to lidar-determined aerosol layer  
433 altitudes. Lastly, we note that a recent study highlighted probabilistic height retrievals using the  
434 Crosstrack Infrared Sounder (CrIS) for Raikoke. This study found a median height of 10-12 km  
435 across a large part of the plume, however with some areas upwards of 15 km. While there are  
436 some notable differences across all of the datasets, the OMI retrieval for this case falls within the  
437 general consensus of plume height estimates for this volcanic event.

438

#### 439 **4.5 Discussion of errors**

440 It is clear that predicting SO<sub>2</sub> layer height with FP\_ILM is an efficient process, but one that is not  
441 flawless in terms of accuracy. As comparisons between retrievals have shown, on average there  
442 were 1-2 km differences in heights, especially for the Raikoke event, although we consider this  
443 to be good agreement given the estimated MAE and RMSE associated with this retrieval. In this  
444 regard, the retrieval is an approximate estimate of the SO<sub>2</sub> plume height rather than a precise  
445 determination. Differences in the retrieved heights between different studies/algorithms result  
446 from differences in instruments, forward model assumptions and retrieval techniques as well as

447 uncertainties in each retrieval. For instance, IASI is a thermal IR instrument and its retrieval does  
448 not use FP\_ILM. Therefore exact agreement with IASI results is difficult to achieve, especially  
449 since the IASI retrieval itself has a stated error range of  $\pm 2$  km, although its retrievals serve as a  
450 good verification dataset. The stated uncertainty for TROPOMI retrievals (Hedelt et al., 2019) is  
451  $\sim 2$  km for SO<sub>2</sub> amounts of greater than 20 DU, similar to our estimated uncertainties for OMI.  
452 While the general retrieval approach for TROPOMI (Hedelt et al., 2019) is similar to that for  
453 OMI in the present study, there are also important instrument differences that can lead to  
454 differences in the retrieved heights between the two instruments, such as the pixel size, noise,  
455 radiometric accuracy and the level of degradation. TROPOMI has a much finer spatial resolution  
456 compared to OMI, with footprints typically  $5.5 \times 3.5$  km<sup>2</sup> up to maximum size  $7 \times 3.5$  km<sup>2</sup>;  
457 TROPOMI also has larger maximal SO<sub>2</sub> signals. Consequently, TROPOMI is better able to  
458 resolve localized variations in the height throughout the plume, and is likely to be more accurate  
459 overall due to better SNR. However, current TROPOMI L1 data are known to have issues with  
460 instrument degradation and radiometric accuracy in the UV spectral range (Ludewig et al.,  
461 2020); this could be a potential contributing factor the differences between the two instruments.  
462 OMI retrievals show more or less uniform height levels across the entire plume with the peak  
463 heights in areas with the best SO<sub>2</sub> signal. Note, CALIOP lidar profiles sometimes show  
464 disagreements with OMI retrieved heights, because CALIOP only identifies the height of the ash  
465 or aerosol plume. It also offers a comparison for only a single cross section of the entire plume  
466 per orbit. Despite the uncertainties, the consensus provided by different instrumental datasets can  
467 provide a reasonable estimate for the SO<sub>2</sub> layer height, and if done in near real time, can aid in  
468 decision making with regards to aviation safety.

469 Another source of error is present in the training phase. One difficulty here is finding the  
470 ideal choice of neural network setup. With many parameters to consider, such as the number of  
471 input PCs, number of layers, number of nodes, learning rate, regularization, weight initialization,  
472 etc., it is very time consuming to optimize the neural network setup. We have found a relatively  
473 simply configuration that performed reasonably well with both test data and real OMI  
474 measurements for all scenarios and events considered. However, even after optimization of the  
475 parameters, random error inherently exists in the neural network. A measure of random error can  
476 be obtained by altering the random state of the neural network whilst keeping other parameters  
477 constant. For ten trial runs with different random seeds, variations of the MAE error were around

478 0.15-0.2 km (see Table A2). Although the differences in the errors calculated with the synthetic  
479 test data are relatively small, larger changes can be expected during the application phase.  
480 Indeed, when applying the inverse models to OMI, there is noticeable, up to 1 km variation in the  
481 retrieved height for the same pixels. It is thus difficult to improve results further than  $\sim 1$  km  
482 absolute error, even in the training phase. In the application phase, some additional error comes  
483 from the differences between synthetic spectra and real satellite measurements with noise errors.  
484 For example, with an SNR of 500 used in training, which is a typical noise level for OMI, the  
485 RMSE of the neural network prediction is around 1.25 km (Table 3). This can be considered the  
486 lower limit of retrieval error when the inverse operator is used on OMI measurements. Lastly,  
487 some deviations between the measured and synthetic training spectra originate from the RT  
488 modeling. The calculations contain several assumptions including the SO<sub>2</sub> plume shape,  
489 atmospheric profiles, gas profiles, and a molecular scattering atmosphere. Further testing is  
490 required in order to determine if the inclusion of aerosols in RT calculations would improve the  
491 algorithm performance.

## 492 493 **5 Conclusion** 494

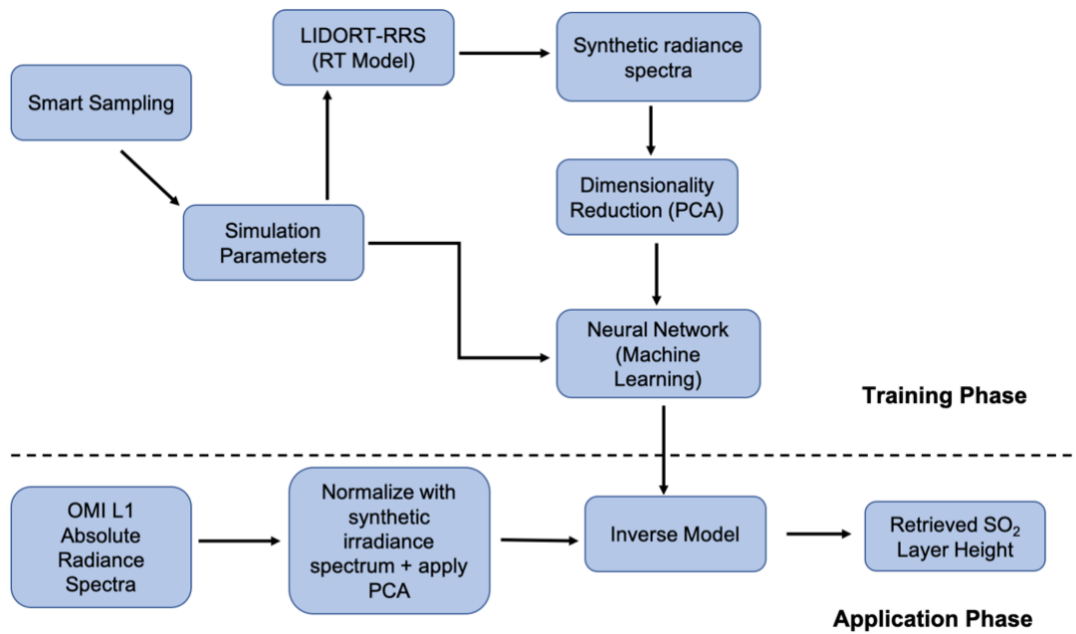
495 In this study we have introduced a new algorithm for OMI retrievals of the volcanic SO<sub>2</sub>  
496 effective layer height from UV earthshine radiances. This algorithm is based on an existing  
497 FP\_ILM method which combines a computationally time-consuming training phase with full  
498 radiative transfer model simulations and a machine learning approach to develop a fast inverse  
499 model for the extraction of plume height information from radiance spectra. Fast performance  
500 means that the algorithm can be considered for operational deployment, given that the retrieval  
501 of a SO<sub>2</sub> layer height prediction from the inverse model takes only a matter of milliseconds for a  
502 single OMI ground pixel. For the training, a synthetic dataset of earthshine radiance spectra were  
503 created with the LIDORT-RRS RT model for a variety of conditions based on choices of 8  
504 physical parameters determined with smart sampling techniques. A dimensionality reduction was  
505 performed through PCA in order to reduce the complexity of the problem and to separate those  
506 features that best capture the great majority of variance of the dataset; 8 principal components  
507 were sufficient for this purpose. Dimensionally-reduced data together with the associated  
508 parameters were used to train a double hidden-layer neural network to predict SO<sub>2</sub> plume height  
509 from any given input data. The PCA from the training phase and the inverse operator resulting



510 from the optimal NN framework were then applied to real satellite radiance spectra and  
511 parameters to get retrieved values of SO<sub>2</sub> plume heights for several volcanic eruption events.

512 Through comparisons with CALIPSO lidar overpasses, as well as TROPOMI and IASI  
513 retrievals, it was shown that the retrieval for OMI can estimate reasonable SO<sub>2</sub> layer height for  
514 all the events considered, with absolute errors in the range of 1-2 km . These results can give an  
515 indication of plume heights achieved during medium- to large-scale eruptions, and guide  
516 important decisions in aviation hazard mitigation. For all events treated in this study, there was  
517 general agreement with CALIOP lidar, although SO<sub>2</sub> could not be retrieved for the locations of  
518 the CALIPSO flight path for the Kelud and Calbuco cases due to OMI row anomaly issues.

519 Uncertainties and sources of error in using this approach open up possibilities for future  
520 work in improving the accuracy of the retrieval. We assumed that ash and sulfur dioxide plumes  
521 are mostly collocated when using CALIPSO as a source to verify the plume height. Although  
522 this is often true, dispersion of the plume in the days following the eruption can separate the two  
523 components. Therefore, tracking these plumes become challenging when using reflectance  
524 spectra alone; further analysis may need to include trajectories or wind data. The model was  
525 trained on synthetic spectra calculated for molecular atmosphere conditions in the absence of any  
526 aerosol loading. The impact of including aerosols in the simulations is another subject for a  
527 follow-up study. We also intend to generate data sets of synthetic spectra by using a vector RRS  
528 model to account for polarization. For improving the performance and efficiency of the machine  
529 learning, the use of neural network ensembles and a further optimized setup of NN parameters  
530 will be explored. Other future work will include extending the application of FP\_ILM to the  
531 Suomi-NPP OMPS instrument as well as exploring the ability to predict multiple outputs at once  
532 from this approach.



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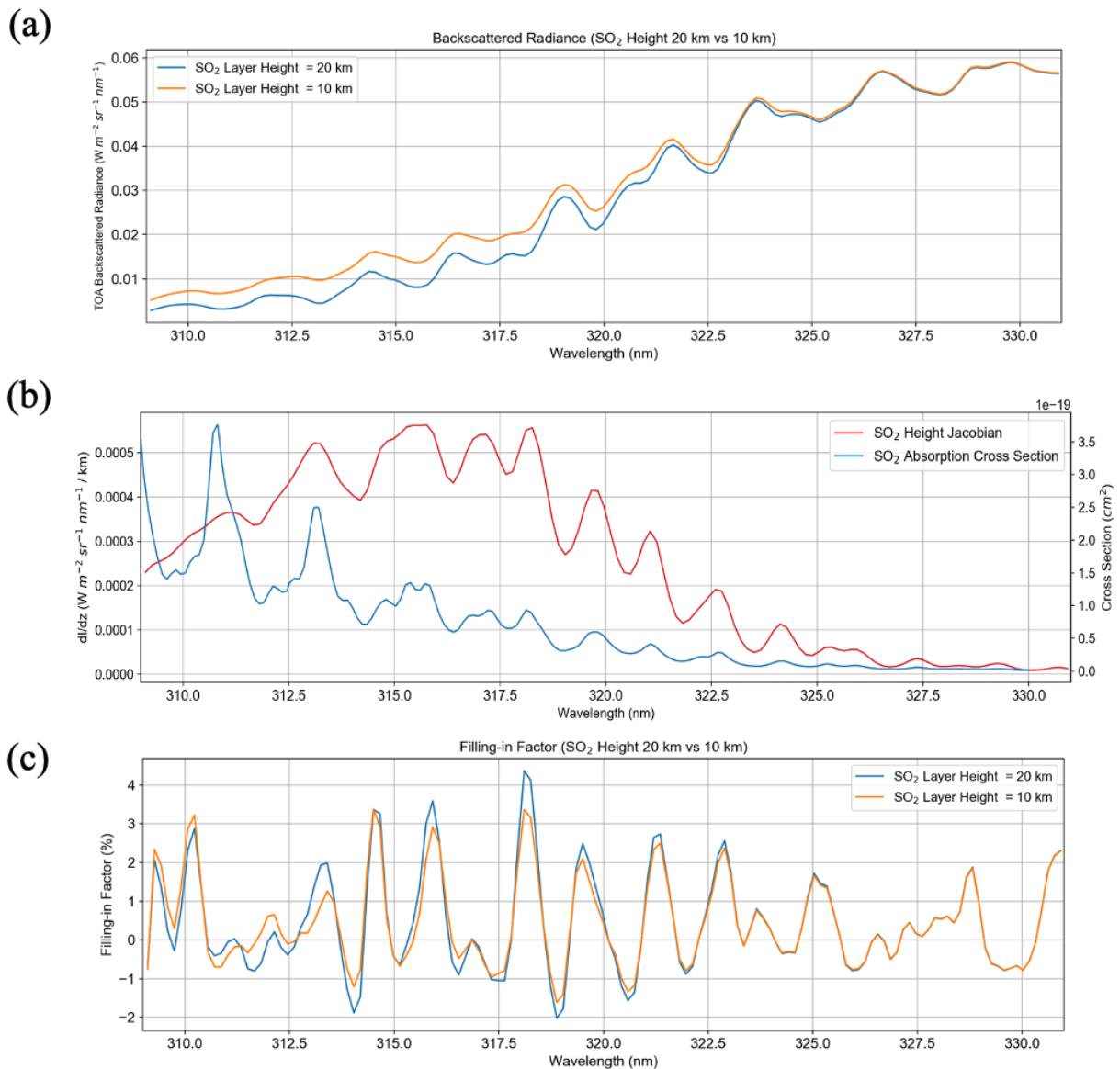
534 **Figure 1:** The flowchart of the FP\_ILM methodology for retrieving OMI SO<sub>2</sub> Effective Layer Height.

535 The steps above the dashed line are part of the training phase done prior to incorporation of OMI

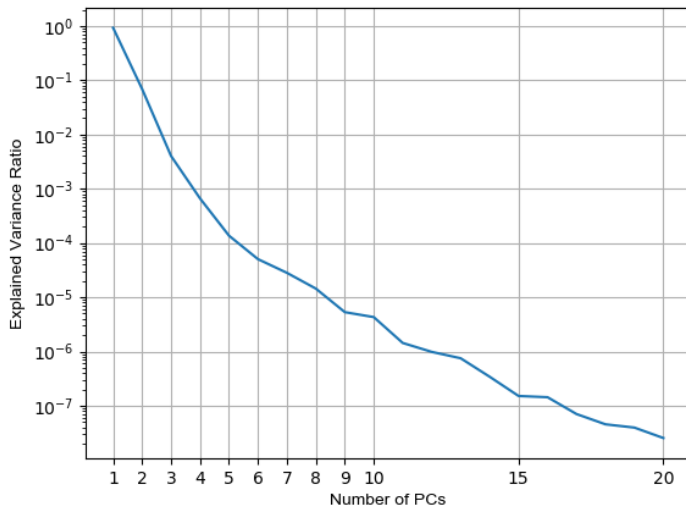
536 measurements. The application phase involves deployment of the trained model to the OMI radiance

537 measurements to obtain estimates of effective volcanic SO<sub>2</sub> layer heights.

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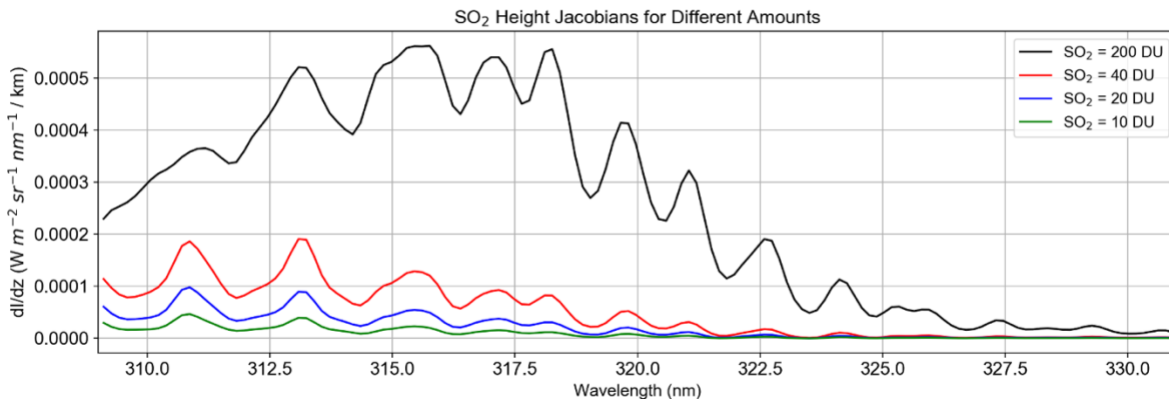


539  
 540 **Figure 2:** (a) Simulated top of the atmosphere (TOA) Earthshine radiances for two different SO<sub>2</sub> layer  
 541 heights (10 km and 20 km) from the LIDORT-RRS model. Also shown: (b) the SO<sub>2</sub> height Jacobian  
 542 (change in radiance per km between the two spectra) along with the absorption cross-sections of SO<sub>2</sub> for  
 543 reference; (c) the filling-in factor. The filling-in factor is defined as the difference between the total and  
 544 elastic-only radiance results, divided by the total radiance, expressed as a percentage. An SO<sub>2</sub> column  
 545 amount of 200 DU was used in the two calculations and all other parameters were kept constant except  
 546 for the SO<sub>2</sub> layer height.  
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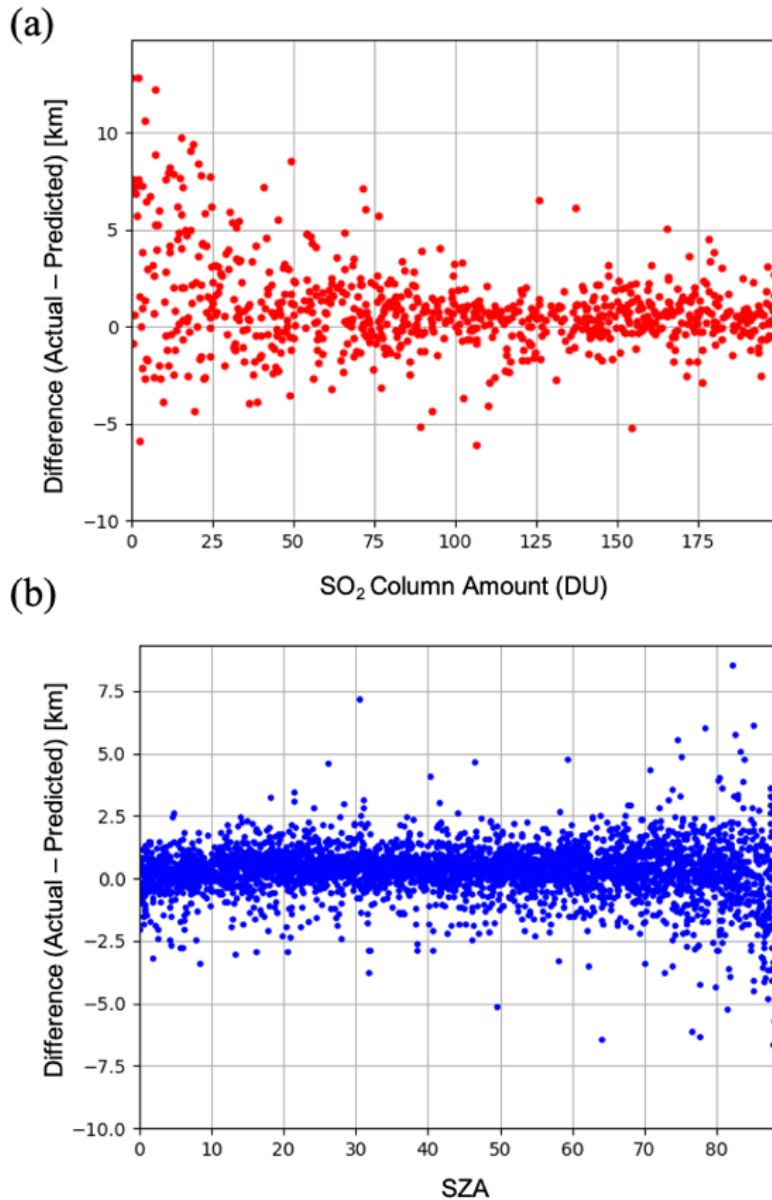
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**Figure 3:** Explained variance ratio as a function of the number of principal components of the spectral dataset.

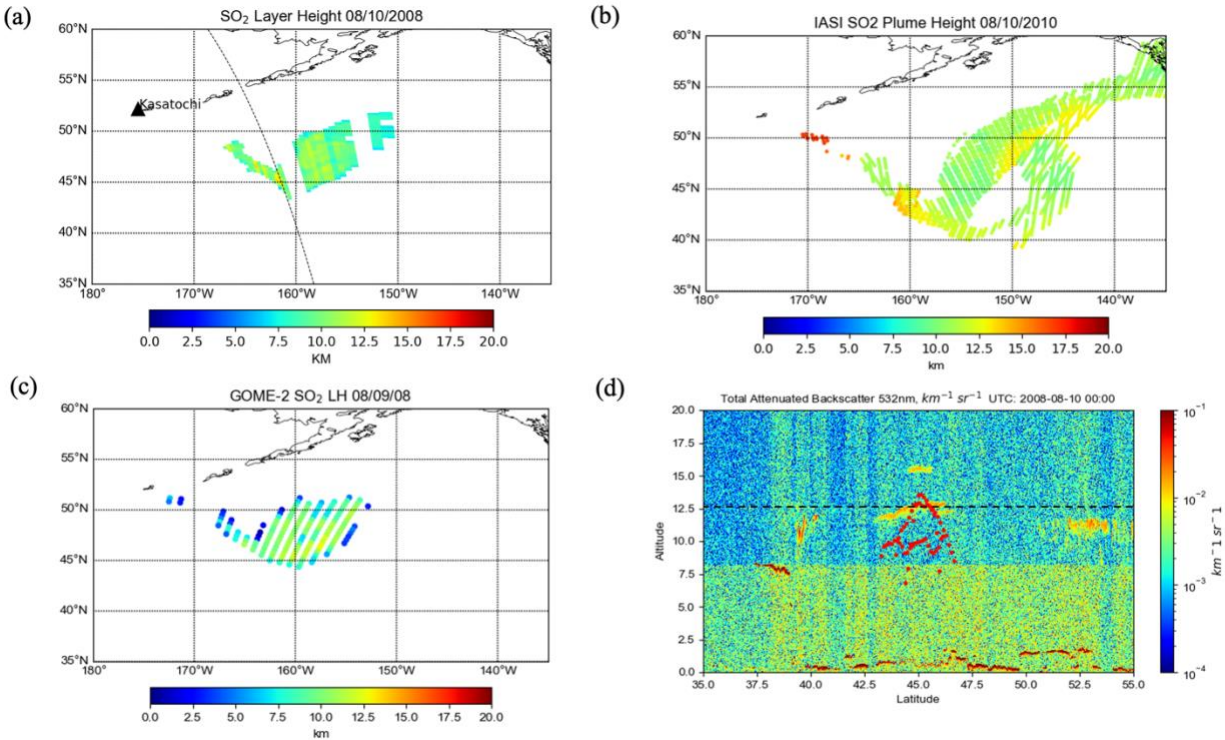


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**Figure 4:** SO<sub>2</sub> Height Jacobians ( $dI/dz$ ) for 4 different assumed SO<sub>2</sub> column amounts. The Jacobians were calculated from the difference between two radiance spectra with 10 km and 20 km SO<sub>2</sub> height. All other physical parameters were identical in the calculation of the spectra.

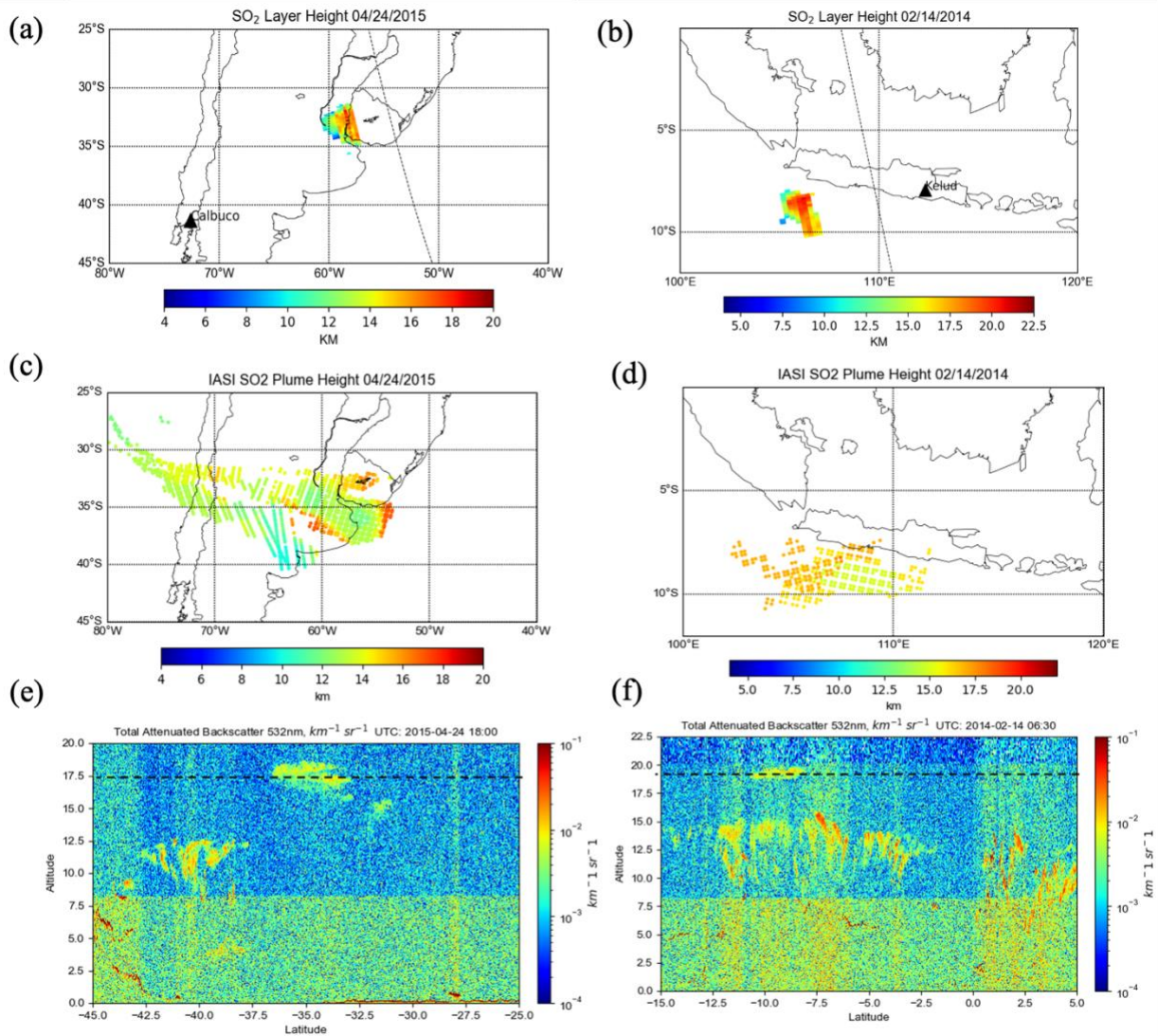


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 560 **Figure 5:** Dependence of retrieval errors on (a) SO<sub>2</sub> amount and (b) SZA for cases with SO<sub>2</sub> VCD > 40  
 561 DU. The error is defined as the difference between the SO<sub>2</sub> layer height predicted by the neural network  
 562 using inputs from the independent test set, and the actual height from the same samples. The test set  
 563 comprises 10% of the original spectral dataset withheld from training the neural network. The plots show  
 564 that the retrieval error is mostly within +/- 2.5 km for SZA < 70, but increases significantly for large  
 565 SZAs.  
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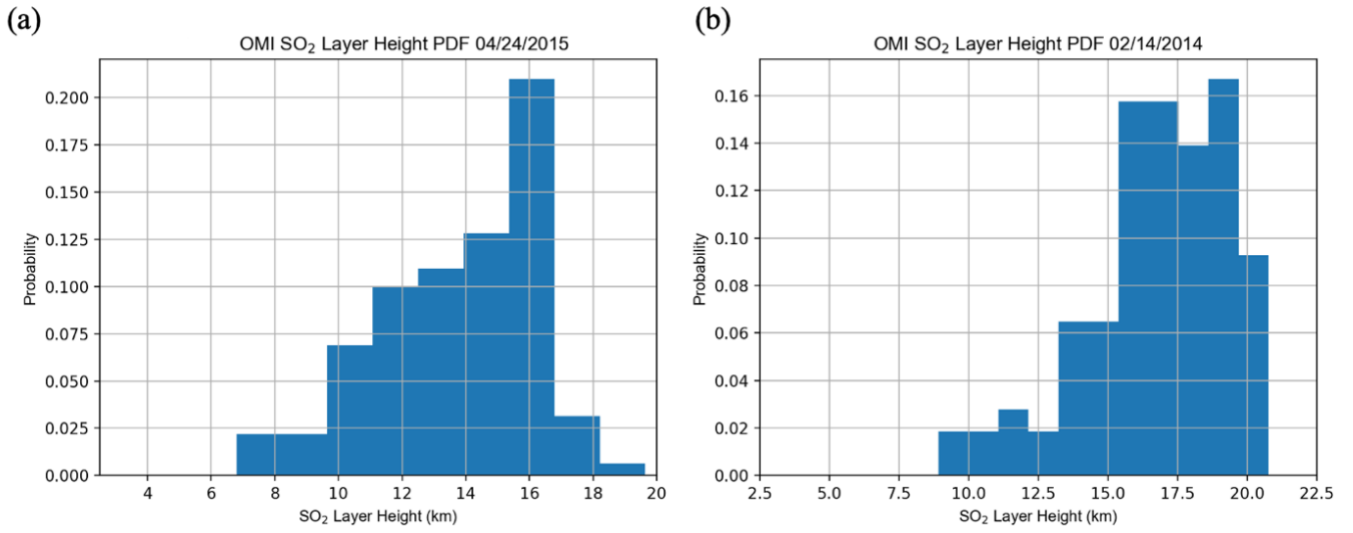


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**Figure 6:** Comparison between the volcanic plume heights from (a) OMI, (b) IASI, (c) GOME-2 and (d) CALIOP lidar 532-nm attenuated backscatter, for the 2008 Kasatochi eruption. The black dotted line in (a) shows the CALIPSO track. Some rows of OMI in this case were affected by the row anomaly, as seen by the gaps in the plume. The red dots in (d) show the OMI retrieval near the CALIPSO path and the black dashed line denotes the height of the ash plume observed by CALIPSO.



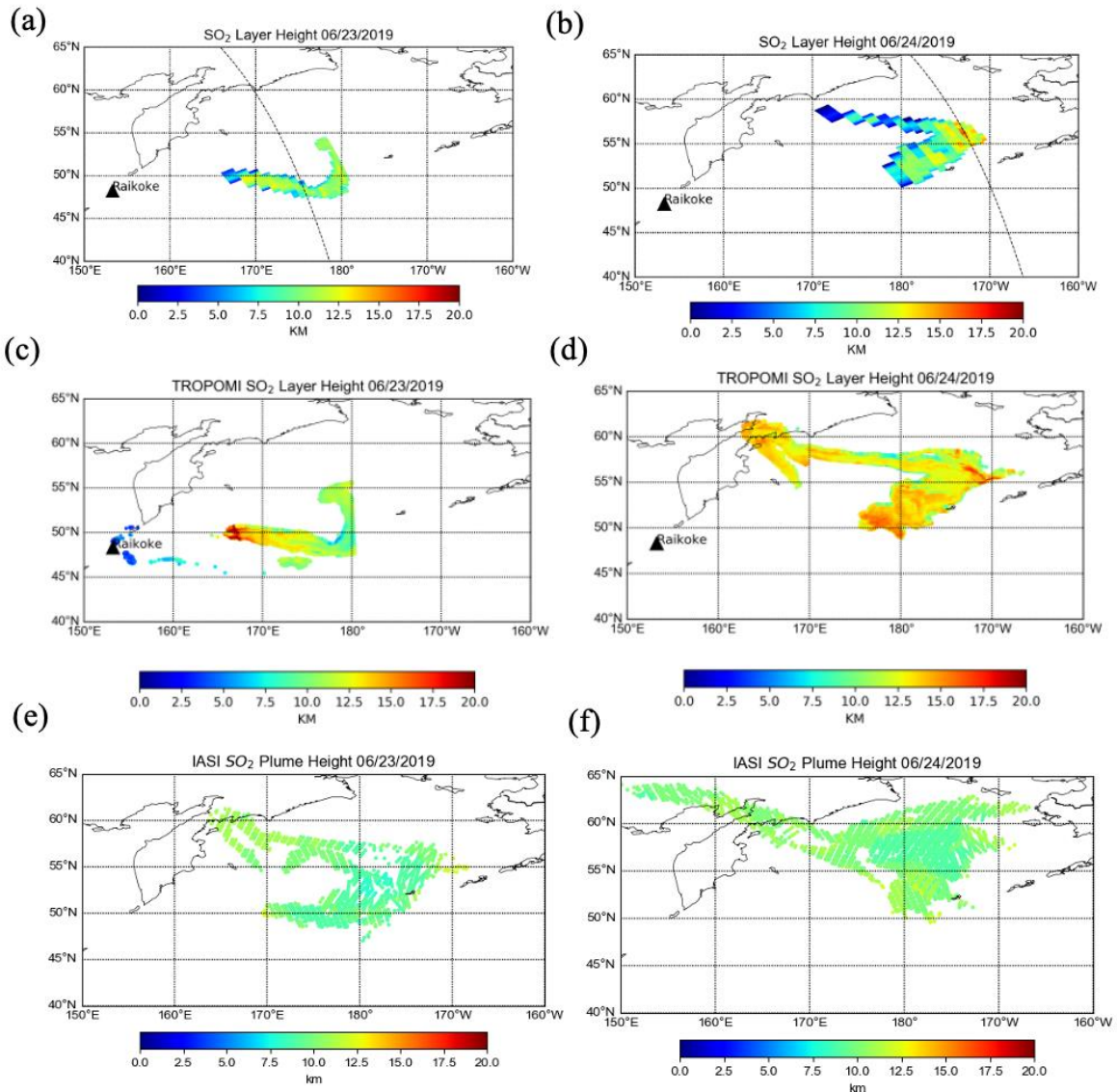
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 576 **Figure 7:** Comparisons of plume heights for the 2015 Calbuco eruption (left) and the Kelud  
 577 eruption (right) for OMI (a,b), IASI (c,d) and 532-nm total attenuated backscatter from the  
 578 CALIOP lidar (e,f). For OMI, only pixels with > 30 DU of SO<sub>2</sub> are shown and retrievals were  
 579 unavailable for some parts of the plume due to the row anomaly. The black dotted line in (a) and  
 580 (b) marks the CALIPSO track. The white rectangles in (e) and (f) show the location of the plume  
 581 in the lidar profile. Direct comparison with CALIPSO was not possible due to obstruction by the  
 582 OMI row anomaly  
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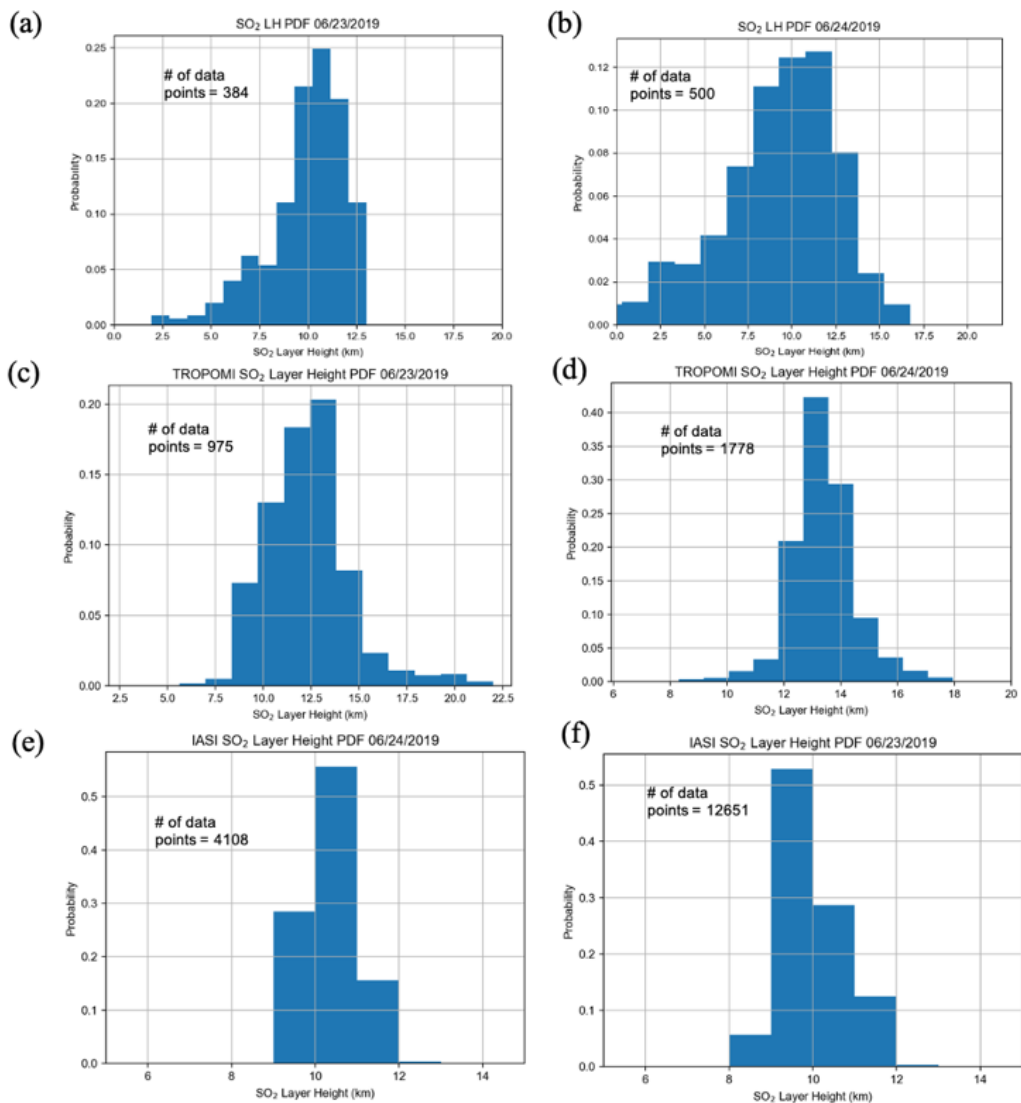
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**Figure 8:** Probability histograms of SO<sub>2</sub> effective layer height retrievals for (a) the Calbuco eruption on April 24, 2015 and (b) the Kelud eruption on February 14, 2014.

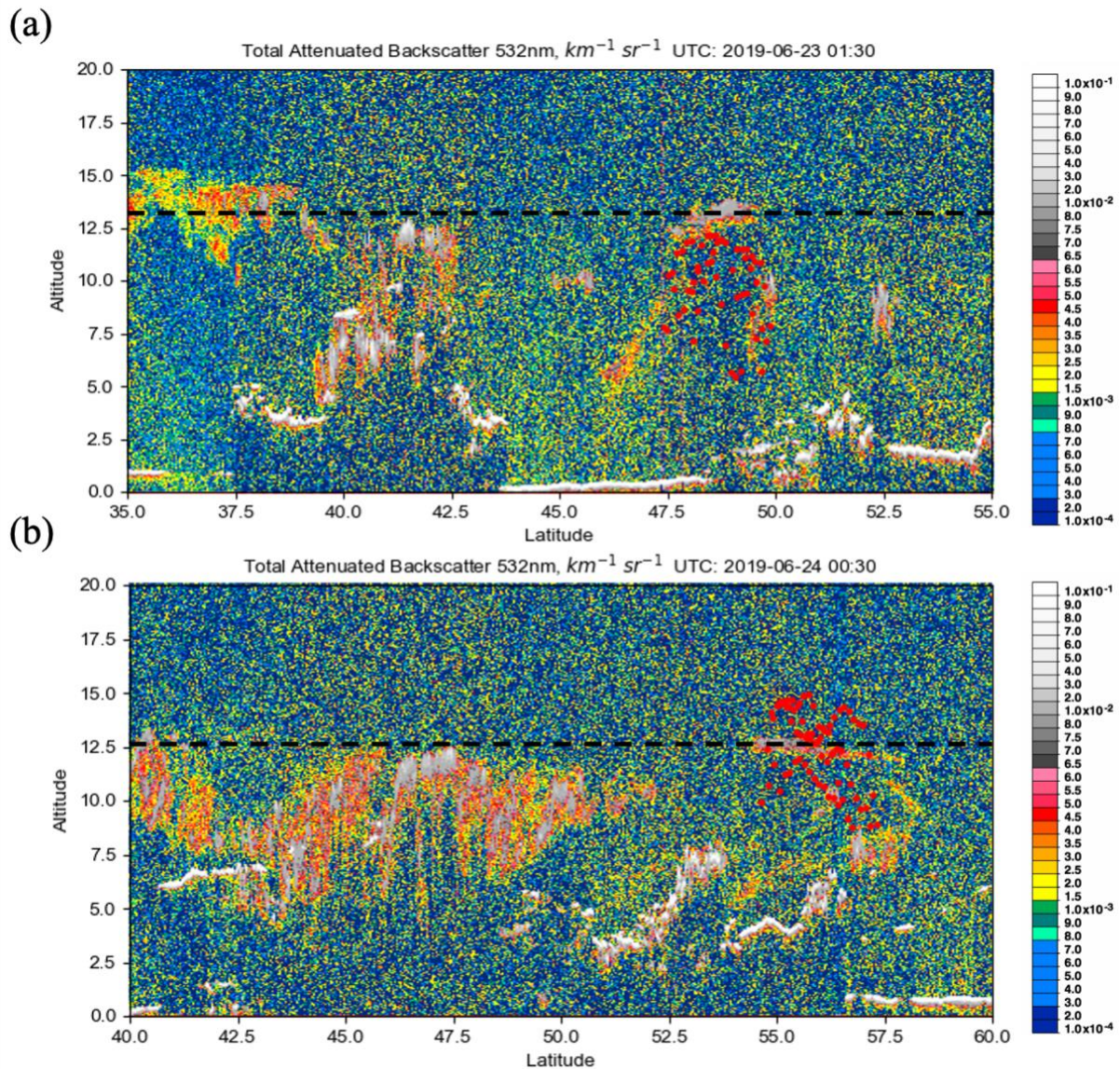




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 589 **Figure 9:** The SO<sub>2</sub> layer height retrieval for the Raikoke eruption plume on June 23<sup>rd</sup>, 2019 (left)  
 590 and June 24<sup>th</sup>, 2019 (right) for the OMI (a, b), TROPOMI (c, d) and IASI (e, f) instruments. For  
 591 all 3 sensors, only pixels where SO<sub>2</sub> VCD > 30 DU are shown.  
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 594 **Figure 10:** Probability histograms of SO<sub>2</sub> layer height retrievals for (a,b) OMI and (c,d),  
 595 TROPOMI on June 23<sup>rd</sup>, 2019 (left) and June 24<sup>th</sup>, 2019 (right) and (e,f) IASI. Only pixels with  
 596 SO<sub>2</sub> column amount greater than 30 DU are included. These plots correspond to the results  
 597 plotted in Figures 4a-f.  
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 600 **Figure 11:** CALIPSO lidar 532-nm attenuated backscatter for the Raikoke eruption on (a) June  
 601 23<sup>rd</sup> and (b) June 24<sup>th</sup>, 2019. The black dashed line symbolizes the height of ash plume seen by  
 602 CALIPSO and red dots show the results from the OMI retrieval along CALIPSO's flight path.  
 603 The flyovers occurred shortly after 01:30 and 00:30 UTC on June 23<sup>rd</sup> and 24<sup>th</sup> respectively,  
 604 around the same time as OMI.

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613 **Table 1:** Ranges of the eight physical parameters varied in LIDORT-RRS for the synthetic  
 614 spectra calculations.

Parameter	Range
Solar Zenith Angle	0-90°
Viewing Zenith Angle	0-70°
Relative Azimuth Angle	0-180°
Surface albedo	0-1
Surface pressure	250-1013.25 hPa
O <sub>3</sub> VCD	225-525 DU
SO <sub>2</sub> VCD	0-1000 DU
SO <sub>2</sub> Layer Height	2.5-20 km

615

616 **Table 2:** The RMSE and the mean absolute difference (km) of all data points in the independent  
 617 test set after adding noise as indicated by different SNR values. All other parameters and input  
 618 data were kept constant. SZA < 75 degrees and SO<sub>2</sub> VCD > 40 DU were excluded from the test  
 619 set for these comparisons.

	No noise	SNR=1000	750	500	200	100
Mean Absolute Difference (y_known - y_pred) (km)	0.894	0.904	0.939	0.996	1.114	1.362
RMSE (km)	1.454	1.498	1.521	1.632	1.807	2.143
R-coefficient	0.988	0.985	0.983	0.980	0.972	0.955

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623 **Table 3:** The RMSE and the mean absolute difference of all data points in the test set under  
 624 different conditions. For each condition, the appropriate points were removed and excluded in  
 625 error calculations. All cases in this table used synthetic training spectra with added SNR 1000.

	All cases	SO <sub>2</sub> > 20 DU	SO <sub>2</sub> > 40 DU	SO <sub>2</sub> > 60 DU	SZA < 75°	SO <sub>2</sub> > 40 DU and SZA < 75°	Albedo < 0.6	SO <sub>2</sub> > 40 DU , SZA < 75° , Albedo < 0.6
RMSE	1.487	1.216	1.150	1.109	1.281	0.931	1.524	0.895
Absolute Mean Difference (km) (Predicted – Actual)	0.910	0.834	0.803	0.782	0.795	0.697	0.895	0.667

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627 **Table 4:** Statistical comparisons of the SO<sub>2</sub> height retrievals for two days of the Raikoke  
 628 eruption and the Kasatochi eruption cases.

Metric (km)	Raikoke (June 23 <sup>rd</sup> , 2019)			Raikoke (June 24 <sup>th</sup> , 2019)			Kasatochi		
	OMI	IASI	TROPOMI	OMI	IASI	TROPOMI	OMI	IASI	GOME-2
Std. Deviation	1.67	0.85	1.96	2.38	0.65	1.04	1.39	0.72	1.29
Median	10.60	9.00	12.10	10.30	10.00	13.24	9.70	10.00	10.21
Mean	10.20	9.63	12.15	10.00	9.83	13.30	9.84	10.40	10.02
IQR	1.79	1.00	2.71	2.68	1.00	1.20	1.36	1.00	1.67

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631 **Appendix A**

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633 **Table A1:** Mean absolute difference and RMSE for different reductions of the original training  
 634 dataset. The test was performed on training sets for five different OMI rows and the errors were  
 635 averaged.

% of samples withheld	0	10	20	30	40	50
Mean Abs Difference	0.95	0.98	1.02	1.08	1.12	1.24
RMSE	1.46	1.45	1.62	1.69	1.79	2.00

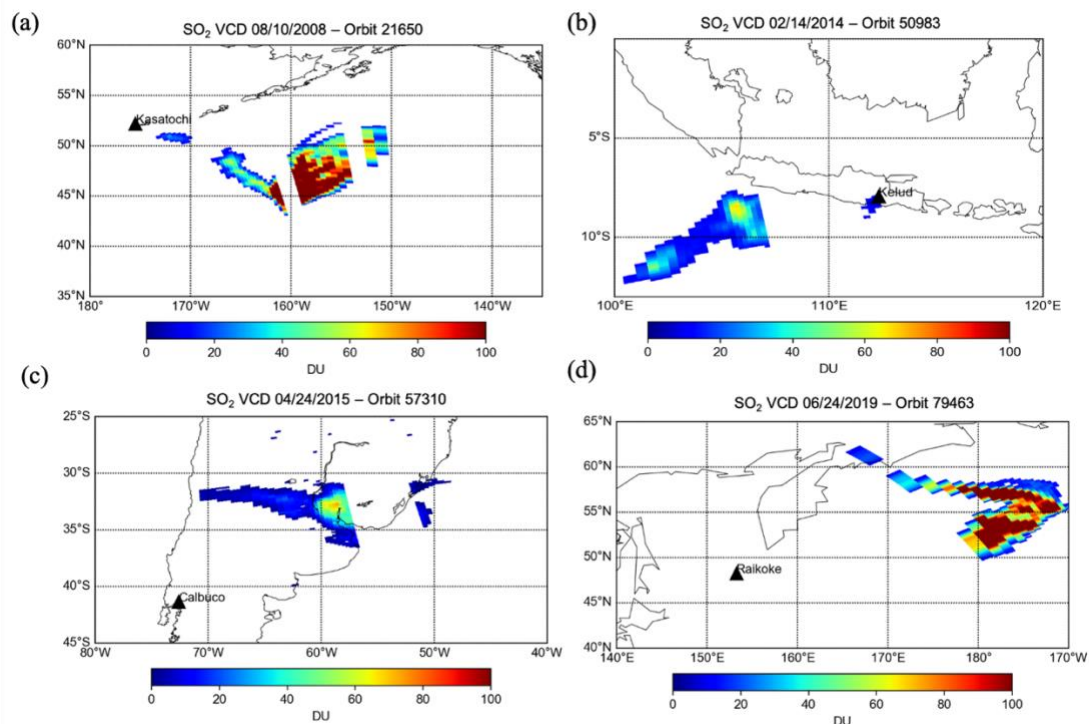
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638 **Table A2:** Effect of altering random seed number on error obtained using the test dataset, and  
 639 the SO<sub>2</sub> height retrieval result after application to OMI. For the results, heights for two different  
 640 pixels within the orbit from the Raikoke event (June 24<sup>th</sup>, 2019) are shown. Heights were  
 641 retrieved using separate inverse models trained using 10 random states.

	Random Seed Number	1	2	3	4	5	6	7	8	9	10
	Absolute										
NN Training error	Mean Error	0.98	1.14	1.03	1.16	1.08	1.18	1.05	1.01	1.12	0.98
	RMSE	1.69	1.85	1.71	1.78	1.79	1.92	1.71	1.67	1.73	1.70
Application (Raikoke - OMI Orbit 79463)	Sample pixel 1	10.52	10.69	10.49	9.72	9.98	10.23	10.53	10.19	10.07	10.48
	Sample pixel 2	12.42	13.15	12.08	11.70	11.88	12.01	12.38	11.22	11.94	12.16

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 647 **Figure A1:** OMI SO<sub>2</sub> VCD for the four volcanic cases: (a) Kasatochi on August 10<sup>th</sup>, 2008, (b)  
 648 Kelud on February 14<sup>th</sup>, 2014, (c) Calbuco on April 24<sup>th</sup>, 2015 and (d) Raikoke on June 24<sup>th</sup>,  
 649 2019. In these maps, only pixels with SO<sub>2</sub> > 10 DU are shown.

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 652 **Data availability.** OMI SO<sub>2</sub> L1 and L2 data can be accessed via the Goddard Earth Sciences Data and  
 653 Information Services Center (GES DISC) at <https://earthdata.nasa.gov/eosdis/daacs/gesdisc>. IASI SO<sub>2</sub>  
 654 LH data is available via the IASI AERIS portal <https://iasi.aeris-data.fr/>. NASA CALIPSO data can be  
 655 downloaded from <https://www-calipso.larc.nasa.gov/> and images can be found at [https://www-](https://www-calipso.larc.nasa.gov/products/lidar/browse_images/production)  
 656 [calipso.larc.nasa.gov/products/lidar/browse\\_images/production](https://www-calipso.larc.nasa.gov/products/lidar/browse_images/production). TROPOMI L2 SO<sub>2</sub> data can be obtained  
 657 at <https://s5phub.copernicus.eu/dhus/#/home> while the LH is experimental and is not yet publicly  
 658 available online. The results of OMI SO<sub>2</sub> layer height retrieval presented in this study can be obtained  
 659 from the author by request.

660  
 661 **Author contributions.** NF wrote the manuscript and performed most computational and model work in this  
 662 study. The project was conceived and overseen by CL and NK. DL and PH provided the TROPOMI SO<sub>2</sub> LH  
 663 retrieval data and input on the comparisons in the paper. PH also offered support relating to the machine  
 664 learning aspect of the study. RS is the original developer of the LIDORT-RRS code and provided related  
 665 support, as well as input to the relevant sections of the manuscript. RD is an advisor of NF and provided  
 666 additional input to the paper and was involved in project planning.

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668 **Competing interests.** The authors declare that they have no conflict of interest.

669

670 **Acknowledgements.** We would like to acknowledge the NASA Earth Science Division (ESD) Aura Science  
671 Team program for funding of the OMI SO<sub>2</sub> product development and analysis (Grant # 80NSSC17K0240).

672 OMI is a Dutch/Finish contribution to the NASA Aura mission. The OMI project is managed by the Royal  
673 Meteorological Institute of the Netherlands (KNMI) and the Netherlands Space Agency (NSO).

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