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Interactive comment on "Statistical analyzing the effect of ionospheric irregularity on GNSS radio occultation atmospheric measurement" *by* Mingzhe Li and Xinan Yue

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Response to Reviewer 1 Thanks for the invaluable comments and suggestions that helped to improve the quality of this manuscript. Considering all the comments, point-to-point responses are addressed in this rebuttal as follows.

General comment: The submitted manuscript quantifies the impact of ionospheric irregularities on failed radio occultation (RO) inverted events and bending angle oscillations, due to the contribution of sporadic E occurrences and the F1-layer. The manuscript provides a very interesting contribution on assessing the impact of smallscall residual ionospheric errors (RIEs) on RO data profiles. Except of a few minor

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revisions, I recommend the manuscript for publication.

Reply: We have checked and revised the manuscript. Thank you very much for the helpful comments.

Specific comments: Literature research (Introduction): Some of the older and newer published literature on assessing RIEs is missing. Regarding older literature the study of Danzer et al. 2013 is very interesting, since it applies a very similar approach to assess the impact of the RIE. The study analyzed the bending angle bias as a difference of the observed bending angle to the MSIS reference climatology, to obtain an estimate of the RIE. This similarity in the approaches should be mentioned. More recent literature is the work by Danzer et al. 2020 and Liu et al. 2020, where the former assesses the climatological impact of the RIE and the latter applies a profile-to-profile study which calculates the RIE through applying four terms (electron density, geomagnetic field, raypath inbound, and raypath outbound effects). Please have a look into it and discuss it in your manuscript. Maybe it could be also relevant to the calculation of small-scale scintillations (kappa and bi-local correction).

Reply: Thanks for the comment. We have discussed the kappa and bi-local correction methods as "Danzer et al. (2013) have analyzed the bending angle bias of CHAMP and COSMIC RO data from 2001-2011 and tried to parameterize bending angle bias versus the solar cycle to make statistical corrections. Healy and Culverwell (2015) found a good correlation between the RIE and the difference of GPS L1 and L2 bending angles at the same impact parameter. They then proposed a correction method using the 'kappa' parameter under the ionospheric spherical symmetry assumption. This method was further tested by Danzer et al. (2015), Angling et al. (2018) and Danzer et al. (2020). It can reduce the systematic errors vary with solar cycle from 0.2 K to 2.0 K at altitudes between 40 km to 45 km. Liu et al. (2018) have analyzed the ionospheric structure influences on RIE in bending angles based on ray tracing simulations and further developed a "Bi-local correction approach" to calculate the RIE through an equation. This method considers both the ionospheric asymmetry effects

as well as the geomagnetic effects on bending angles (Liu et al., 2020)". The related references have also been added in the manuscript.

Line 114: Please provide information about the settings for the NCAR climatology model. How was the setting for the solar activity? Danzer et al. 2013 uses for example a constant solar flux value, in order to assess the solar variations of the RO bending angle to the climatology.

Reply: Thanks for the comment. During the atmospheric RO inversion, the NCAR climatology model is used to provide the neutral refractivity for calculating the model bending angle. Then the climatology model bending angle is used for statistical optimization of the observed bending angle. The bias between the observed bending angle and the climatology model bending angle can be used to evaluate the ionospheric effects on the observed bending angle, on condition that only the neutral atmospheric information is provided by the climatology model bending angle. It means no solar activity setting is needed for the NCAR climatology model. Danzer et al. (2013) use a constant solar flux value for the setting of the NeUoG ionosphere model. They use the operational analysis fields from ECMWF to provide the atmospheric information.

Line 118: Please provide a more detailed introduction of the S4 index, to aid readability and understanding.

Reply: Thanks for the comment. We have added the introduction of the S4 index as "Besides, the amplitude and phase measurements can be used to calculate the ionospheric scintillation index such as the S4 index. The S4 index is defined as the standard deviation of the received signal power normalized to the average signal power, it can represent the occurrence of the ionospheric irregularity (Yue et al., 2016)".

Lines 193-196, Figure 2: I am curious, what is the statistics within a bin? I know it varies with latitude, however, just on average, from low to high latitudes, so that I get an idea of the numbers behind the percentage shown in Figure 2. Please also replot Figure 2, fix the range of the colorbar, e.g. column 2, the range moves from 0 to 20,

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and from 0 to something like 25. Furthermore, a colorbar needs a unit and a label of what is shown, the manuscript always includes this information in the title. Reply: Thanks for the comment. We use the geographical coordinate and the geomagnetic coordinate of the tangent point to decide whether a RO event belongs to a bin. For the ionospheric irregularity in a bin, the occurrence rate is calculated by

"Occurrence rate =" "number of RO events (S4>0.3)" /"number of RO events" ×100% (1)

It means the occurrence rate of a bin is the ratio of the number of RO events with S4 index > 0.3 and the number of all RO events in this bin. Similarly, for the failed inverted RO event, the occurrence rate is calculated by

Occurrence rate =(number of failed inverted RO ("bad RO"))/(number of RO events) \times 100% (2)

The median value of the smean parameter of all RO events in a bin is calculated to represent the bending angle oscillation in this bin. The range of colorbars have been revised, the label and unit have also been added to the colorbars. Please note that we have added a figure after Figure 1 as the new Figure 2. It shows the occurrence rates of Es and FI separately. As a result, the serial numbers of figures after Figure 1 have been changed in orders.

Lines 322-329: I really appreciated the summary, but I think the implication and discussion part could be extended. What are the next best steps to include the ruled-out events in the processing? Bending angle oscillations at high altitudes are usually handled via the high-altitude initialization, to remove such effects. Is the bi-local correction (Liu et al. 2020) an option for correcting such small-scall ionospheric scintillations?

Reply: Thanks for the comment. We think a suitable filter could be helpful in reducing the ionospheric irregularity effects on bending angles and including ruled-out events in the processing. The potential method needs to be more investigated and we hope to

study it in our further work. Besides, the causing of failed inverted RO events in high latitudes is also needed to be studied. As depicted in the new Figure 5, the correlations between the failed inverted RO events and the ionospheric irregularity are not as good as those in low and middle latitudes. We have added a small discussion in the conclusion part as "Overall, the ionospheric irregularity effects on GNSS atmospheric RO measurement exist in terms of failed RO event inversion and bending angle oscillation statistically. It makes the calibration of ionospheric irregularity effects on RO more challenging and urgent. A suitable filter may be helpful in reducing the ionospheric irregularity effects on bending angles and including ruled-out events in the processing. The potential method needs to be more investigated and we hope to study it in our further work. Besides, the causing of failed inverted RO events in high latitudes is also needed to be studied". According to Mannucci et al. (2013) and Liu et al. (2020) have discussed, the "bi-local" correction is mainly used to correct the large-scale RIE caused by the separation of GNSS signals with different frequencies and the RIE caused by the geomagnetic effects. The correction of small-scale ionospheric scintillations still needs to be further studied.

Technical comments: Figures: Please improve readability of the Figures 1,2,5,6 for the reader. For example, Figure 1 misses the information of label and unit on the y-axis (altitude, km). In the third column, the y-axis extent is not constant (it jumps a bit). Provide ticks in-between (third and fourth column, x- and y-axis). And so on ... Figures 5 and 6: Include a unit and label information for the colorbar (r.h.s.), include a larger space between the colorbar and the plots.

Reply: Thanks for the comment. The missed information in Figure 1 has been added. The unit and label information for the colorbar have been added to the new Figure 3, 4, 6, 7. The space between the colorbar and plot has also been adjusted.

References Danzer, J., Scherllin-Pirscher, B., and Foelsche, U.: Systematic residual ionospheric errors in radio occultation data and a potential way to minimize them, Atmospheric Measurement Techniques, 6, 2169-2179, doi.org/10.5194/amt-6-2169-

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2013, 2013. Liu, C., Kirchengast, G., Syndergaard, S., Schwaerz, M., Danzer, J., and Sun, Y.: New Higher-Order Correction of GNSS RO Bending Angles Accounting for lonospheric Asymmetry: Evaluation of Performance and Added Value, Remote Sensing, 12, doi.org/10.3390/rs12213637, 2020. Mannucci, A. J., Ao, C. O., Pi, X., and lijima, B. A.: The impact of large scale ionospheric structure on radio occultation retrievals, Atmospheric Measurement Techniques, 4, 2837-2850, doi.org/10.5194/amt-4-2837-2011, 2011.

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Fig. 1.





Fig. 2.



Fig. 3.





Fig. 4.



Fig. 5.





Fig. 6.



Fig. 7.

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