



## 1                    **Evaluating the PurpleAir monitor as an aerosol light scattering** 2                    **instrument**

3    James R. Ouimette<sup>1</sup>, William C. Malm<sup>2</sup>, Bret A. Schichtel<sup>3</sup>, Patrick J. Sheridan<sup>4</sup>, Elisabeth  
4    Andrews<sup>4,5</sup>, John A. Ogren<sup>6</sup>, W. Patrick Arnott<sup>7</sup>

5    <sup>1</sup>Sonoma Ecology Center, Sonoma Ecology Center, Eldridge, CA 95431, USA

6    <sup>2</sup>Cooperative Institute for Research in the Atmosphere, Colorado State University, Fort Collins,  
7    CO 80523, USA

8    <sup>3</sup>National Park Service Air Resource Division, Fort Collins, CO 80523, USA

9    <sup>4</sup>NOAA Global Monitoring Laboratory, Boulder, CO 80305, USA

10   <sup>5</sup>Cooperative Institute for Research in Environmental Sciences, University of Colorado, Boulder,  
11   Colorado 80309, USA

12   <sup>6</sup>NOAA Global Monitoring Laboratory, Boulder, CO 80305, USA (Retired)

13   <sup>7</sup>Department of Physics, University of Nevada, Reno, NV 89557, USA

14   Correspondence to Elisabeth Andrews (betsy.andrews@noaa.gov)

15

### 16   **Abstract**

17   The Plantower PMS5003 sensors (PA-PMS) used in the PurpleAir (PA) monitor PA-II-SD  
18   configuration are equivalent to cell-reciprocal nephelometers using a 657 nm perpendicularly  
19   polarized light source that integrates light scattering from 18 to 166 degrees. Yearlong field data  
20   at the National Oceanic and Atmospheric Administration's (NOAA) Mauna Loa Observatory  
21   (MLO) and Boulder Table Mountain (BOS) sites show that the 1 h average of the PA-PMS first  
22   size channel, labeled ">0.3  $\mu\text{m}$ " ("CH1") is highly correlated with submicrometer aerosol  
23   scattering coefficients at the 550 nm and 700 nm wavelengths measured by the TSI 3563  
24   integrating nephelometer, from 0.4  $\text{Mm}^{-1}$  to 500  $\text{Mm}^{-1}$ . This corresponds to an hourly average  
25   submicrometer aerosol mass concentration of approximately 0.2 to 200  $\mu\text{g m}^{-3}$ . A physical-  
26   optical model of the PA-PMS is developed to estimate light intensity on the photodiode,  
27   accounting for angular truncation as a function of particle size. Predictions are then compared  
28   with yearlong fine aerosol size distribution and scattering coefficient field data at the BOS site. It  
29   is shown that CH1 is linearly proportional to the model-predicted intensity of the light scattered  
30   by particles in the PA-PMS laser to its photodiode over 4 orders of magnitude. This is consistent  
31   with CH1 being a measure of the scattering coefficient and not the particle number concentration  
32   or particulate matter concentration. Field data at BOS confirm the model prediction that the ratio  
33   of CH1 to the scattering coefficient would be highest for aerosols with median scattering  
34   diameters <0.3  $\mu\text{m}$ . The PA-PMS detects aerosols smaller than 0.3  $\mu\text{m}$  diameter in proportion to  
35   their contribution to the scattering coefficient. The model predicts that the PA-PMS response to  
36   particles >0.3  $\mu\text{m}$  decreases relative to an ideal nephelometer by about 75% for particle  
37   diameters  $\geq 1.0 \mu\text{m}$ . This is a result of using a laser that is polarized, the angular truncation of the  
38   scattered light, and particle loss in the instrument before reaching the laser. The results of this  
39   study indicate that the PA-PMS is not an optical particle counter and that its six size fractions are  
40   not an accurate representation of particle size distribution. The relationship between the PA-PMS  
41   1 h average CH1 and  $b_{\text{sp}1}$ , the scattering coefficient in  $\text{Mm}^{-1}$  due to particles below 1  $\mu\text{m}$   
42   aerodynamic diameter, at wavelength 550 nanometers, is found to be  $b_{\text{sp}1} = 0.015 \pm 2.07 \times 10^{-5} \times$   
43   CH1, for relative humidity below 40%. The coefficient of determination  $R^2$  is 0.97. This



44 suggests that the low-cost and widely used PA monitors can be used to measure and predict the  
45 aerosol light scattering coefficient in the mid-visible nearly as well as integrating nephelometers.

46 Keywords: PurpleAir, Plantower PMS5003, nephelometer, low-cost sensor, physical-optical  
47 model, PM2.5, scattering coefficient, visibility, atmospheric aerosol

## 48 1. Introduction

49 Currently there are tens of thousands of low-cost aerosol monitors used by atmospheric research  
50 groups, air quality monitoring and regulatory organizations, and individual citizen scientists  
51 around the world. The recent explosion in the number of these sensors (see, for example, AAQR,  
52 20(2), “Special Issue on Low-cost Sensors for Air Quality Monitoring” and papers therein) is a  
53 result of the increased research, regulatory, and citizen interest over the past few years. For  
54 example, there are over 9,000 active PurpleAir (PA) aerosol monitors (PurpleAir LLC, Draper,  
55 UT), with sampling locations on almost every continent. The large geographic coverage of this  
56 array of low-cost sensors presents enormous potential for obtaining valuable information on  
57 atmospheric aerosol properties and transport processes.

58 The majority of these low-cost aerosol sensors are used to monitor the mass concentration of  
59 particles with aerodynamic diameters  $<2.5 \mu\text{m}$  (PM2.5) (Kelly et al., 2017; Gupta et al., 2018;  
60 Sayahi et al., 2018; Zheng et al., 2018; Malings et al., 2019; Barkjohn et al., 2020; Holder et al.,  
61 2020; Jayaratne et al., 2020; Mehadi et al., 2020). However, these sensors do not actually  
62 measure aerosol mass concentrations but light scattered by the aerosols and thus are dependent  
63 on the aerosol particle size distribution, morphology, and composition. Recently, Hagan and  
64 Kroll (2020) developed a framework and computer model to estimate the effects of relative  
65 humidity and aerosol refractive index on PM2.5 estimated by a number of low-cost sensors.  
66 Their model assumed that the low-cost sensor lasers were not polarized and could be modeled  
67 with Mie theory. The PMS5003 (PMS) was included in their classification scheme as an example  
68 of a sensor that behaved more like a nephelometer than an optical particle counter.

69 Three recent laboratory studies showed that the PMS response decreases with particle size. He et  
70 al. (2020) measured the PMS response to monodisperse ammonium sulfate aerosol particles  
71 having diameters of 0.1, 0.3, 0.5, and 0.7  $\mu\text{m}$ . The PMS was able to detect 0.1  $\mu\text{m}$  particles. They  
72 derived a transfer function that showed that the PMS  $>0.3 \mu\text{m}$  channel (CH1) response was  
73 maximum at particle diameter 0.26  $\mu\text{m}$  but decreased significantly below this size. They  
74 concluded that the PMS behaved more like a nephelometer than an optical particle counter.  
75 Kuula et al. (2020) generated monodisperse dioctyl sebacate oil droplets from 0.5 to 20  $\mu\text{m}$  and  
76 measured the PMS CH1 response versus particle diameter using an aerosol particle sizer (APS).  
77 Their data showed that the PMS relative response decreased for particles  $>0.5 \mu\text{m}$  diameter.  
78 Tryner et al. (2020) evaluated three low-cost particulate matter sensors, including the PMS, by  
79 exposing them to five different types of aerosols in the laboratory. They found that the ratios of  
80 PMS-reported to filter-derived PM2.5 mass concentrations were inversely proportional to mass  
81 median diameter (MMD). Wood smoke had the smallest MMD, 0.42  $\mu\text{m}$ ; its PMS PM2.5  
82 averaged 2.5 times the filter-derived PM2.5. Conversely, oil mist had the largest MMD, 2.9  $\mu\text{m}$ ;  
83 its PMS PM2.5 averaged only 0.23 times the filter-derived PM2.5.

84 Climate modeling requires a robust set of models and atmospheric measurements for predicting  
85 anthropogenic aerosol radiative forcing. Currently, there are uncertainties in the modeling



86 results, due in part to the sparseness of ground-based data used to evaluate and refine the models  
87 (e.g., Gliss et al., 2021). Satellite observations provide global coverage that can be used for  
88 model evaluation, but satellite data require further assessment, particularly when trying to  
89 provide information about surface aerosol properties. The Surface Particulate Matter Network  
90 (SPARTAN) (<https://www.spartan-network.org/>; Snider et al., 2015) was specifically designed  
91 to assess and improve algorithms to relate satellite retrievals to surface aerosols. SPARTAN  
92 operates collocated filter-based PM<sub>2.5</sub>, aerosol scattering coefficient via nephelometer, and  
93 aerosol optical depth (AOD) measurements at approximately 20 sites around the world. Model  
94 and satellite uncertainties can be reduced using a distributed set of low-cost sensors that can  
95 provide aerosol light scattering estimates at a higher spatial and temporal resolution than is  
96 possible using nephelometers alone. Low-cost sensors are increasingly being used along with  
97 satellite data to estimate global aerosol impacts (Gupta et al., 2018).

98 There is ongoing scientific debate about the accuracy and precision of these low-cost sensors and  
99 their limitations (Morawska et al., 2018; Jayaratne et al., 2020). Many of the recent papers  
100 discuss performance evaluations or ‘calibrations’ of these low-cost sensors by comparing their  
101 measurements with traditional, research-grade aerosol measurements (Papapostolou et al., 2017;  
102 Barkjohn et al., 2020). The concerns over data quality, stemming largely from inexpensive  
103 components, lack of transparency of signal processing, and inadequate quality control and testing  
104 at the factory, must be weighed against the advantages of low cost and wide spatial coverage.

105 The actual measurement in the PA monitor with its two PMS5003 sensors (PA-PMS), and in  
106 many other low-cost aerosol monitors, is of light scattered by particles (Kelly et al., 2017), which  
107 traditionally has been done in atmospheric research and aerosol monitoring programs using  
108 integrating nephelometers. Aerosol light scattering extinction measurements are useful in many  
109 applications, including determination of the radiative forcing effects of aerosols on climate  
110 change, atmospheric visibility, wildfire/smoke impacts, and validation of model outputs and  
111 satellite retrievals (e.g., Malm et al., 1994; Sherman et al., 2015; Snider et al., 2015; Gliss et al.,  
112 2021). Even though most low-cost aerosol sensors use light scattering as the basis of their  
113 operation, almost none have been evaluated as a low-cost nephelometer to estimate atmospheric  
114 light scattering. Markowicz and Chilinski (2020) conducted a 3-year evaluation of two low-cost  
115 sensors versus the Aurora 4000 polar integrating nephelometer at a site in southeast Poland.  
116 They found that the mass concentration of particles with aerodynamic diameters <10 μm (PM<sub>10</sub>)  
117 from the DfRobot SEN0177 and the Alphasense OPC-N2 were highly correlated ( $R^2 > 0.89$ )  
118 with the aerosol scattering coefficient measured by the nephelometer. They were able to estimate  
119 the 1 h average aerosol scattering coefficient from the low-cost sensors with a root mean square  
120 error (RMSE) of 20 Mm<sup>-1</sup>, corresponding to 27% of the mean aerosol scattering coefficient.

121 Unfortunately, due to cost, availability, and the expertise required to run them, integrating  
122 nephelometers are not operated in great numbers around the world. A recent analysis by Laj et  
123 al. (2020) showed 56 long-term monitoring stations reporting their nephelometer data to the  
124 World Meteorological Organization (WMO) Global Atmosphere Watch (GAW) World Data  
125 Centre for Aerosols. This count includes nephelometers operated in several monitoring  
126 networks, including the National Oceanic and Atmospheric Administration’s (NOAA) Federated  
127 Aerosol Network (NFAN, Andrews et al., 2019), the Aerosols, Clouds and Trace Gases Research  
128 Infrastructure (ACTRIS) network (e.g., Pandolfi et al., 2018), and the Interagency Monitoring of  
129 Protected Visual Environments (IMPROVE) network (Malm et al., 1994). While there are more



130 nephelometers in use around the world for short-term field and laboratory studies, the number  
131 almost certainly does not exceed a few hundred. This is small compared with the number of low-  
132 cost aerosol monitors in use globally.

133 This paper presents an evaluation of the performance characteristics of the low-cost PA-PMS  
134 monitor to measure the integrated aerosol light scattering coefficient. It is shown that the PA-  
135 PMS sensor configuration is similar to a cell-reciprocal nephelometer. A physical-optical model  
136 based on Mie theory and the PMS geometry is created that predicts scattered light intensity on  
137 the PMS photodiode and aerosol forward and backward light scattering truncation. Model  
138 predictions are then compared with yearlong field data at NOAA's Mauna Loa Observatory  
139 (MLO) in Hawaii and the Boulder Table Mountain (BOS) site in Colorado. These PA-PMS  
140 measurements are also compared to measured light scattering coefficients, and an empirical  
141 relationship is developed to estimate the light scattering and uncertainty from the PA-PMS data.

142 With a better understanding of what the PA measures, how it works, and its uncertainties, the  
143 large network of PA-PMSs could be used to estimate the submicrometer aerosol scattering  
144 coefficient at visible wavelengths throughout the world. These data could then be used to  
145 improve chemical transport and general circulation models, advance climate change predictions,  
146 and provide for better air quality forecasts.

## 147 **2. Instrument description**

148 In this section we first describe the physical and optical characteristics of the PA-PMS to place it  
149 in the context of nephelometry. We then provide a brief overview of integrating nephelometers,  
150 which are instruments designed specifically to measure light scattering.

### 151 **2.1 PA-PMS nomenclature**

152 The PMS sensor outputs 14 fields that are processed and reported by the PA. Each of these  
153 fields will be referred to as a channel. For instance, the PA-reported number concentration of  
154 particles  $>0.3 \mu\text{m}$  is referred to as CH1 in the remainder of this paper, number concentrations  
155  $>0.5 \mu\text{m}$  as channel two (CH2), and so forth. Furthermore, the two PMS sensors embedded in  
156 the PA will be referred to as either sensor A or sensor B. Therefore, the number concentration of  
157 particles  $>0.3 \mu\text{m}$  derived from sensor A will be referred to as CH1A and those from sensor B as  
158 CH1B. The average of CH1A and CH1B will be referred to as CH1avg. The PMS reports the  
159 CH1 units as “#/dl”, which is the number of particles having diameters  $>0.3 \mu\text{m}$  per deciliter. In  
160 this paper the PMS units for CH1 are not used.

### 161 **2.2 Description of the PA and its PMS 5003 sensors**

162 The PA monitor integrates two PMS sensors, a Bosch BME280 pressure, temperature, and  
163 relative humidity sensor and an ESP 8266 chip (<https://www2.purpleair.com/pages/technology>).  
164 The available specifications of the PMS are incomplete, and the processing algorithms are  
165 unknown (He et al., 2020). The following is based on available information and, where needed,  
166 professional judgment. Each PMS includes a small laser, a photodiode, a small fan to draw air  
167 across the laser, a microprocessor control unit (MCU), and probably an operational amplifier.  
168 The MCU processes the signal from the photodiode and outputs the following data fields  
169 approximately once per second:  $>0.3 \mu\text{m}$ ,  $>0.5 \mu\text{m}$ ,  $>1.0 \mu\text{m}$ ,  $>2.5 \mu\text{m}$ ,  $>5 \mu\text{m}$ ,  $>10 \mu\text{m}$ , PM1,  
170 PM2.5, and PM10. The PMS denotes the first six data fields as particle number concentrations



171 above the designated cutpoint and the last three data fields as mass concentrations of particles  
172 below the designated cutpoints; the PM data fields are reported for two different conditions,  
173 “standard particles” and “under atmospheric environment”. The PA ESP8266 chip calculates 2  
174 min averages of the PMS and BME280 signals. It transmits them wirelessly and writes them as a  
175 CSV file on a microSD card.

### 176 **2.2.1 Airflow and particle losses**

177 The recommended orientation of the PA results in aerosol being drawn upward by a small fan  
178 through four 3-mm diameter entrance holes in each PMS. The aerosol then enters a 9.4 cm<sup>3</sup>  
179 chamber (Fig. S1a) and flows upward, parallel to and exposed to the circuit board as shown in  
180 Fig. S1b. Particles then make a 180 degree turn through three exit holes at the top of the chamber  
181 to emerge on the other side of the circuit board and flow downhill through a channel that is  
182 illuminated by the laser. The PMS volumetric flow rate is estimated to be 1.5 cm<sup>3</sup> s<sup>-1</sup> (~0.090  
183 lpm) based on measurements described in Supplemental Materials Sect. S1. The estimated inlet  
184 velocity through the entrance holes is estimated to be 5.3 cm s<sup>-1</sup>.

185 The PMS inlet orientation 90 degrees to the wind, upward flow, and the low inlet velocity  
186 through the sampling holes can result in significant aspiration losses of larger particles (Hangal  
187 and Willeke, 1990). Aspiration losses are greater at higher wind speeds because it is more  
188 difficult for the larger particles to follow the streamlines into the low velocity PMS inlet. This  
189 can result in a lower concentration of larger particles entering the PMS than are in the ambient  
190 air. Particle aspiration losses are proportional to the particle Stokes number and the ratio of the  
191 wind velocity to the inlet face velocity (Hangal and Willeke, 1990). More details are provided in  
192 Supplemental Materials Sect. S1.

193 At typical wind velocities of 1–3 m s<sup>-1</sup>, the ratio of PMS inlet face velocity to wind speed is only  
194 0.02 to 0.05, much lower than typical sampling ratios of 0.5 to 6.0 (Kulkarni et al., 2011). Pawar  
195 and Sinha (2020) addressed this problem for the Laser Egg low-cost sensor by putting it in a box  
196 and adding a 40 lpm fan to increase the inlet-to-wind velocity ratio and to direct the airflow  
197 upward to the Laser Egg inlet. During calm winds, large particle aspiration losses may occur by  
198 particle gravitational settling, acting against the PMS upward flow (Grinshpun et al, 1993). The  
199 actual wind conditions in the ambient air and in the PA near the PMS sample inlet are turbulent.  
200 Hangal and Willeke (1990) found in their wind tunnel experiments that turbulence intensity had  
201 a negligible effect on aspiration efficiency. Calculations using Eq. S1 (see Fig. S2) predict that at  
202 a wind speed of 1 m s<sup>-1</sup>, the PMS aspiration losses for particles >2 μm may be significant.  
203 However, it must be cautioned that the literature does not include data for the very low 5.3 cm s<sup>-1</sup>  
204 PMS face velocity and actual measurements of the PMS aspiration efficiencies were not made.  
205 They may be significantly different from these calculated efficiencies.

206 Inside the PMS 9.4 cm<sup>3</sup> chamber, the air has an average velocity of 0.57 cm s<sup>-1</sup> and Reynolds  
207 number of 6.1, resulting in an average residence time of 6.3 s. The average air velocity in the  
208 chamber is equal to the sedimentation velocity of a spherical 10 μm diameter particle with a  
209 density of 2 g cm<sup>-3</sup> in air at STP (standard temperature and pressure; values used in this analysis  
210 are 273.15 K and 1010.25 hPa, respectively). This suggests that some 2 g cm<sup>-3</sup> density particles  
211 with diameters >10 μm that enter the PMS would settle out in the chamber and not make it to the  
212 three exit holes at the top of the chamber. Ultrafine particles can also be lost to the walls of the  
213 chamber and the printed circuit board due to convective diffusion. Calculations using the



214 equation for diffusional losses (Friedlander, 1977) show that less than 1% of the 0.01  $\mu\text{m}$   
215 diameter aerosols would be lost in the chamber due to convective diffusion, with even smaller  
216 diffusional losses for larger particles.

217 Loss of particles due to inertial impaction on the wall opposite the three holes (Fig. S1b) was  
218 estimated by the local air flow Reynolds number near the three holes and the aerosol Stokes  
219 number. The local Reynolds number is calculated to be 23, and the Stokes number for 10  $\mu\text{m}$   
220 particles is  $8.2 \times 10^{-4}$ . At these low numbers, the calculated loss to impaction is less than 1% for  
221 all particles less than 10  $\mu\text{m}$  diameter (Hering, 1995).

222 The average flow velocity through the laser beam is approximately  $3.0 \text{ cm s}^{-1}$ . By the time the  
223 air flows through the laser beam, it has lost most of the particles over 10  $\mu\text{m}$  diameter. Further  
224 particle losses due to gravitational settling over the photodiode would be very small, since the  
225 gravitational force is parallel to the photodiode.

### 226 **2.2.2 Laser**

227 The wavelength and power of three PMS diode lasers were measured using an Ocean Optics Red  
228 Tide USB650 spectrometer and Melles Griot Universal Optical Power Meter, respectively. The  
229 wavelength averaged  $657 \pm 1 \text{ nm}$ , and the power averaged  $2.36 \pm 0.04 \text{ mW}$ . The laser is  
230 polarized parallel to the plane of the photodiode detector. This results in the aerosol-scattered  
231 light being polarized perpendicular to the plane of incidence. Figure S3 shows that  
232 perpendicular polarization results in significantly greater scattering intensity from 0.3  $\mu\text{m}$   
233 particles compared to natural or parallel polarization.

234 The PMS laser beam profile is not a simple plane wave, but complex in shape. The laser has a 3  
235 mm diameter lens that focuses the laser over the photodiode. The beam profile evolves  
236 significantly as it goes through the focal region (Naqwi and Durst, 1990). The laser beam  
237 diameter in the laser sensing region over the photodiode was not measured. It was estimated by  
238 eye to be 0.5 to 1.0 mm, with significant uncertainty. The PMS MCU turns the laser on and off  
239 every 800 msec or 2.5 s, depending on aerosol concentration. The laser pulses are 600–900  
240 msec, with the laser power on continuously during this time. We hypothesize that the PMS MCU  
241 gathers data during laser on, processes it during laser off, and uses the difference of the  
242 photodiode output during these stages to obtain and subtract any electronic or stray light (other  
243 than the laser) background signal to the photodiode.

### 244 **2.2.3 Photodiode detector**

245 The actual photodiode model in the PMS is unknown. The photodiode appearance is similar to  
246 the BPW34 silicon PIN photodiode. In this paper the specifications of the BPW34 are used to  
247 estimate the likely properties of the detector in the PMS. It has a very large dynamic range when  
248 operated with reverse bias. The dependence of the photodiode current on the light intensity is  
249 very linear over 6 or more orders of magnitude, e.g., in a range from a few nanowatts to tens of  
250 milliwatts. Silicon PIN photodiodes have low dark current, a 20 nanosecond rise time, and good  
251 wavelength sensitivity between roughly 400 and 1000 nm. ([https://www.rp-  
252 photonics.com/photodiodes.html](https://www.rp-photonics.com/photodiodes.html)). At a wavelength of 657 nm, the BPW34 produces  
253 approximately 0.4 microampere current per microwatt of incident radiant power  
254 (<https://www.fiberoptics4sale.com/blogs/archive-posts/95046662-pin-photodetector->

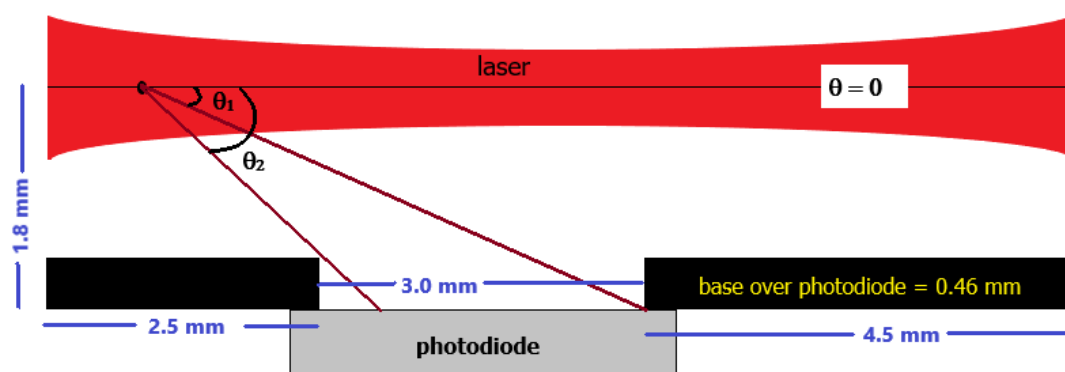


255 [characteristics-for-optical-fiber-communication](#)). The PMS does not have any optical elements  
256 to capture and focus the aerosol-scattered light on its photodiode.

257 The photodiode does not have a cosine corrector in front and is probably not a true cosine  
258 detector. However, the relative spectral sensitivity is advertised to be a cosine response by the  
259 manufacturers  
260 ([https://www.osram.com/ecat/DIL%20BPW%2034%20B/com/en/class\\_pim\\_web\\_catalog\\_1034](https://www.osram.com/ecat/DIL%20BPW%2034%20B/com/en/class_pim_web_catalog_103489/prd_pim_device_2219537/)  
261 [89/prd\\_pim\\_device\\_2219537/](https://www.vishay.com/docs/81521/bpw34.pdf) and <https://www.vishay.com/docs/81521/bpw34.pdf>).

## 262 2.2.4 Laser and photodiode geometry

263 The PMS geometry is very similar to a cell-reciprocal nephelometer. Figure 1 shows the PMS  
264 laser and photodiode geometry. The distance from the laser exit hole to the light trap is 10 mm;  
265 the perpendicular distance from the center of the laser to the photodiode is 1.8 mm; the diameter  
266 of the exposed photodiode area is 3.0 mm; and the thickness of the mask over the photodiode is  
267 0.46 mm.  $\theta_1$  is the lower angular scattering limit, and  $\theta_2$  is the upper angular scattering limit for a  
268 particle in the laser.



269

270 **Figure 1.** Schematic of the PMS sensor geometry highlighting the dimensions of laser beam  
271 (red) and photodiode (gray) and the various relevant distances between the two.

272 Due to the PMS geometry, the upper and lower angular scattering limits for  $\theta$  depend on the  
273 location,  $x$ , of a particle in the laser. This can be seen on Fig. S4. For example, at  $x = 0$  mm, at  
274 the laser exit, the upper and lower scattering limits for  $\theta$  are 18–38 degrees. At  $x = 4.0$  mm, over  
275 the center of the photodiode, the angular integration limits are 50–130 degrees. The PMS  
276 photodiode is not capable of detecting light scattered from particles at less than 18 degrees.

277 Figures S5–S8 provide more detail about the PMS dimensions and geometry.

## 278 2.2.5 PMS5003 sensing volume

279 The sensing volume is the volume in which the aerosol is irradiated by the laser. The sensing  
280 volume extends the length of the laser where the aerosol flows through it, approximately 10 mm.  
281 The sensing volume is shown in Fig. S9. The average residence time of a particle in the laser



282 beam is approximately 0.03 s. Some of the scattered light is detected by the photodiode and  
283 creates a voltage pulse approximately 30 msec wide. It appears that the photodiode is detecting  
284 either a cloud of particles from the sensing volume or individual pulses, depending on the  
285 concentration. At low concentrations, the aerosol concentration within the sensing volume is  
286 unlikely to be uniform, resulting in large relative changes in output per second.

### 287 **2.2.6 Signal processing and electronics**

288 It is not known how the PMS MCU differentiates and processes the photodiode signals. The  
289 PMS MCU sends the PA a signal approximately every second in the form of a digital sequence  
290 of unsigned 16 bit binary data words, and CH1 is thought to be proportional to the photodiode  
291 current. The photodiode current was not measured in this study. The PA creates 80 s (Firmware  
292 Version 3) or 120 s (Firmware Version 4 and higher) averages and writes them to its microSD  
293 card. We measured an average percentage difference of 0.3% between the 2 min averages  
294 reported by the PA and the 2 min averages calculated from the 1 s values from the PMS. The  
295 results are shown in Fig. S10. It is apparent that the processing done by the PA to calculate its  
296 reported 2 min averages does not bias the results.

### 297 **2.2.7 PMS CH1 variability in sampling filtered air**

298 We found significant variability in PMS response to filtered air. We exposed 21 PAs containing  
299 42 PMS sensors to filtered air for 2 to 94 hours. The results are summarized in Table S1. Hourly  
300 average CH1 ranged from 0.10 to 377. Eleven PAs had both PMS CH1A and CH1B averages  
301 below 2, while seven PAs had at least one CH1 average over 26. We recommend that before  
302 deployment the PAs sample filtered air for at least 4 hours to identify and eliminate PAs with  
303 CH1 hourly averages over 2 in filtered air. Removing PAs with high CH1 offsets in filtered air  
304 reduces uncertainty and improves precision, particularly in cleaner ambient air.

### 305 **2.2.8 PMS CH1 unresponsive to CO<sub>2</sub> and Suva<sup>®</sup>**

306 Filtered air, CO<sub>2</sub>, and Suva<sup>®</sup> (DuPont<sup>™</sup> Suva<sup>®</sup> 134a refrigerant) are often used to calibrate  
307 integrating nephelometers (Anderson et al., 1996). The Rayleigh scattering coefficients of  
308 filtered air, CO<sub>2</sub>, and Suva at 657 nm and at STP (°C and 1013.25 hPa) are 5.5, 13.3, and 46.2  
309 Mm<sup>-1</sup>, respectively. We found that the PMS was unresponsive to 100% CO<sub>2</sub> (Fig. S11) and Suva.  
310 The CH1 for each gas was the same as filtered air. These results indicate that the PMS signal  
311 processing zeroes out a constant scattering signal and cannot be used to measure the scattering  
312 coefficient of gases that are commonly used in calibrating nephelometers. Furthermore, the  
313 method used by the PMS to subtract light scattering by air molecules in the sampling volume is  
314 unknown.

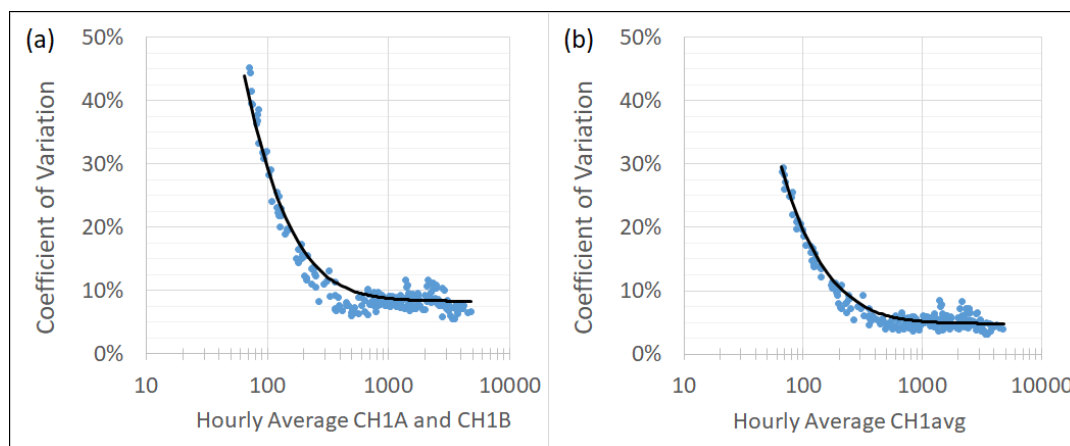
### 315 **2.2.9 PMS CH1 and CH1avg precision**

316 The PMS CH1 precision was measured by collocating ten PA monitors on the roof of the NOAA  
317 building in Boulder, Colorado, between 22 January 2021 and 1 February 2021. These monitors  
318 were not checked with filtered air before deployment. It was found that two of the PMS sensors  
319 had large offsets and two had moderate offsets at low CH1 values. One PMS sensor was found to  
320 produce errant data and was removed from the analysis, resulting in valid data from 19 CH1A  
321 and CH1B sensors in the ten PAs.





322 The hourly CH1A-CH1B and CH1avg precisions were estimated as the coefficient of variation  
 323 for the 19 CH1A-CH1B and 9 CH1avg values for each hour, which are plotted against the  
 324 average CH1 values in Fig. 2. As shown, above CH1 values of 500, the precision is relatively  
 325 constant with an average of 8% and 4.8% for CH1A-CH1B and CH1avg, respectively. Below  
 326 CH1 values of 500, the uncertainties increase rapidly with decreasing CH1 values.



327  
 328 **Figure 2.** Precision estimated as the coefficient of variation of the hourly CH1A-CH1B (a) and  
 329 CH1avg values (b) for the 19 collocated sensors and 9 PAs.

330 The data in Fig. 2 can be modeled by the sum of squares of an additive ( $Un_{add}$ ) and multiplicative  
 331 uncertainty ( $Un_{mult}$ ) (Currie, 1968; JCGM100:GUM, 2008; Hyslop and White, 2008):

$$332 \quad \text{Uncertainty} = \sqrt{Un_{add}^2 + Un_{mult}^2 * CH1} \quad (1)$$

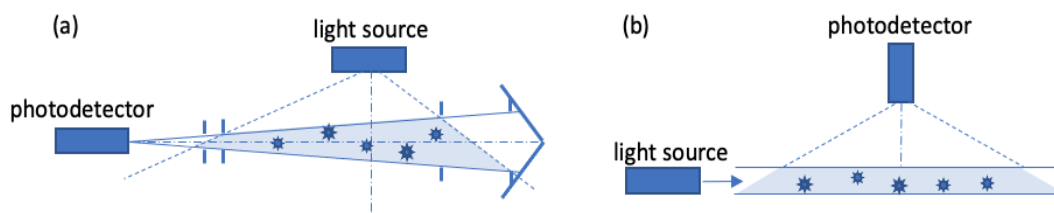
333 Equation 1 was fitted to the precision data in Fig. 2 where the  $Un_{mult}$  was set to the average  
 334 precision at high CH1 values, and  $Un_{add}$  was set to 28 and 19 for the A and B sensors and  
 335 CH1avg, respectively, to fit the highest variances. The  $Un_{add}^2$  is the variance in CH1 as CH1  
 336 approaches zero and is assumed to be equivalent to the uncertainty in values below the  
 337 instrument minimum detection limit (MDL) or that of blanks (Currie, 1968), which were 0.08  
 338 and 0.048 for the A and B sensors and CH1avg, respectively. The coefficient of determination in  
 339 the model fit for both sets of data was  $r^2 = 0.96$ . Defining the MDL as the 99% confidence  
 340 interval of the  $Un_{add}$  (Code of Federal Regulations, 40 CFR 136, <https://ecfr.io/Title-40/Part-136>), MDLs for the individual CH1 sensors and CH1avg were 65 and 44, respectively.

342 As shown in Sect. S3, the  $Un_{mult}$  and  $Un_{add}$  are highly dependent on the systematic biases  
 343 between the individual CH1 sensors and CH1avg and the four CH1 sensors with data offsets as  
 344 the CH1 approaches zero. Removing these four sensors and normalizing the data for each CH1  
 345 sensor by its average reduced the  $Un_{add}$  and  $Un_{mult}$  to 9% and 3%, respectively, for the CH1  
 346 sensors and 6% and 1.9%, respectively, for the CH1avg data. These results correspond to an  
 347 MDL of 21 and 14 for the normalized CH1 sensor and CH1avg data, respectively. Based on  
 348 these results, an “off the shelf” PA will have a CH1avg MDL of about 44 and precision of less  
 349 than 4.3%, but the careful selection of a PA without an offset and that is relatively calibrated will  
 350 have an MDL of 14 and precision of less than 1.9%.



### 351 2.3 Overview of cell-direct and cell-reciprocal nephelometers

352 The integrating nephelometer was invented during World War II (Beuttell and Brewer, 1949). It  
353 provides a direct measure of aerosol light scattering integrated over a large angular range, the  
354 “aerosol light scattering coefficient”. This measure requires no assumptions about aerosol  
355 composition, size distribution, refractive index, or shape. The most common nephelometer  
356 configurations are the “cell-direct” and “cell-reciprocal”. Figure 3 presents schematics of the two  
357 types of nephelometers. The geometrical relationship between the laser and the photodetector in  
358 the PMS resembles a cell-reciprocal nephelometer (Fig. 3b).



359  
360 **Figure 3.** Diagrams of the (a) cell-direct nephelometer and (b) cell-reciprocal nephelometer,  
361 simplified from Peñaloza-Murillo (1999).

362 Middleton (1952) was the first to show that the cell-direct nephelometer with a Lambertian  
363 (cosine-adjusted diffuser) light source directly measures the aerosol light scattering coefficient.  
364 Anderson et al. (1996), following the derivation in Butcher and Charlson (1972), added  
365 geometrical diagrams to make Middleton’s derivation much clearer. Mulholland and Bryner  
366 (1994) proved that the cell-reciprocal nephelometer with a Lambertian diffuser followed by a  
367 photodiode placed at the center of the cell-reciprocal nephelometer also directly measures the  
368 aerosol scattering coefficient. This put both the cell-direct and cell-reciprocal nephelometers on  
369 equal theoretical footing.

370 There are a number of cell-direct nephelometers in use today. They include the TSI 3563 (St.  
371 Paul, MN, USA; Anderson et al., 1996), the Ecotech Aurora Models 3000 and 4000 (Knoxfield,  
372 Australia; Müller et al., 2011), the Radiance Research M903 (Seattle, WA, USA; Heintzenberg  
373 et al., 2006), and the Optec NG-2 (Lowell, MI, USA; Molenaar, 1997). In contrast, cell-  
374 reciprocal nephelometers have more limited commercial availability. The photoacoustic  
375 extinctions (PAX; Droplet Measurement Technologies, Inc., Longmont, CO, USA) and the  
376 three-wavelength photoacoustic soot spectrometer (PASS-3) use a cell-reciprocal nephelometer  
377 to measure aerosol light scattering coefficient (Arnott et al., 2006). A cosine corrector followed  
378 by a photomultiplier tube is placed at the center of the cell-reciprocal nephelometer (Abu-  
379 Rahmah et al., 2006; Nakayama et al., 2015).

380 A “perfect nephelometer” is one in which the nephelometer is able to see the scattered light over  
381 the entire angular range from 0 to 180 degrees. In practice, this cannot be achieved for the cell-  
382 direct and cell-reciprocal nephelometers. Both the forward and backward scattering angles are  
383 truncated. For example, the TSI 3563 nephelometer has measured angular truncation below  
384 about 7 degrees in the forward direction and above 170 degrees in the backward direction  
385 (Anderson et al., 1996; Heintzenberg and Charlson, 1996). For the PASS-3, Nakayama et al.  
386 (2015) found that both the large effective truncation angle (21 degrees) as well as the



387 perpendicular polarization of the 532 nm laser relative to the scattering plane contribute to the  
388 large particle size dependence of measured scattering. Light scattering from ammonium sulfate  
389 particles of 0.71  $\mu\text{m}$  diameter was truncated 50%. Truncation generally results in nephelometers  
390 underestimating the contribution of particles larger than approximately 1  $\mu\text{m}$  diameter to the  
391 scattering coefficient, although corrections have been developed to account for angular  
392 nonidealities (e.g., Anderson and Ogren, 1998; Müller et al., 2011).

### 393 **3. A physical-optical model of the PMS5003**

394 To gain insight into how the PMS responds to ambient aerosol properties, a model was  
395 developed to estimate the intensity of scattered light impinging on the PMS photodiode. The  
396 primary purpose of the model was to predict how the PMS performance compares to other  
397 instruments designed to measure the aerosol scattering coefficient, such as integrating  
398 nephelometers. The model makes simplifying assumptions about the laser that allow the  
399 application of Mie theory to the light scattered from particles in the laser. Details of the model  
400 are presented in the Appendix.

401 The equation describing the intensity of light scattered from a particle in the laser is (Middleton,  
402 1952; Anderson et al., 1996)

$$403 \quad I(\theta) = F_{dv} \beta_p(\theta) dv \quad (2)$$

404 where  $I(\theta)$  is the intensity of light at angle  $\theta$  scattered from a particle in the volume element  $dv$   
405 (with units of  $\text{W sr}^{-1}$ );  $\beta_p(\theta)$  is the volume scattering function ( $\text{m}^{-1} \text{sr}^{-1}$ );  $F_{dv}$  is the incident laser  
406 flux density ( $\text{W m}^{-2}$ ) impinging on the volume element  $dv$ ; and  $dv$  is the volume element within  
407 the laser.

408 The volume scattering function for a single particle in the laser beam is a function of aerosol  
409 diameter  $D_p$ , complex refractive index  $m$ , laser wavelength  $\lambda$ , and scattering angle  $\theta$ :

$$410 \quad \beta_p(\theta) = (\lambda/2\pi)^2 (1/dv) |S_1(m, \lambda, \theta, D_p)|^2 \quad (3)$$

411 where  $|S_1(m, \lambda, \theta, D_p)|^2$  is the Mie scattering intensity function for laser light polarized parallel to  
412 the photodiode surface and perpendicular to the plane of incidence (Bohren and Huffman, 1983).

413 The scattered light intensity from a single particle in the laser beam to a narrow strip across the  
414 middle of the photodiode and from all positions in the scattering volume is integrated to predict  
415 the total power received by the photodiode as a function of particle diameter  $D_p$  and refractive  
416 index  $m$ :

$$417 \quad P(m, D_p) = K \int_{x=0}^{x=10mm} \int_{\theta_1(x)}^{\theta_2(x)} |S_1(m, \theta, D_p)|^2 \sin(\theta) d\theta dx. \quad (4)$$

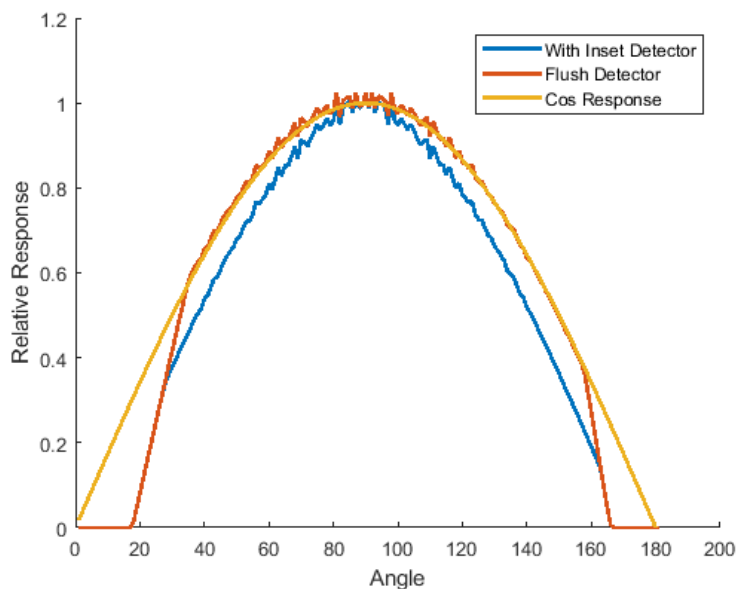
418 Due to the PMS geometry, the upper and lower angular scattering limits for  $\theta$  depend on the  
419 location,  $x$ , of a particle in the laser. Details are provided in the Appendix. This approach can be  
420 used to estimate the amount of scattered energy detected from mixtures of particles of varying  
421 diameters and indices of refraction, as shown in Eq. (5)



$$P = K \int_{D_p} \int_{x=0}^{x=10mm} \int_{\theta_1(x)}^{\theta_2(x)} |S_1(m, \theta, D_p)|^2 \sin(\theta) N(D_p, m) d\theta dx dD_p. \quad (5)$$

### 3.1 Model predictions - Deviation from a perfect cosine response

As discussed above, the PMS has a photodetector that is about 1.8 mm below the laser, resulting in forward scattering and backscattering truncation angles of 18 and 166 degrees, respectively. Furthermore, the photodetector is recessed 0.46 mm below the scattering chamber base. Equation 4 is used to explore the deviation from a perfect cosine response resulting from the truncated scattering volume and recessed detector. It is shown in Fig. 4. For these calculations,  $S_1(m, \theta, D_p)$  is set equal to 1, which corresponds to isotropic scattering or a volume scattering function that is constant over all scattering angles. It is assumed that the detector has a Lambertian response, i.e., the light detected is independent of the direction of the incident energy, which results in a detector cosine response. Figure 4 shows a perfect cosine response in yellow, while the red line shows the deviation from a perfect cosine response due to angular truncation. The blue line shows the effect of both angular truncation and an inset detector that is 0.46 mm below the chamber base. All curves have been normalized to one at 90 degrees.



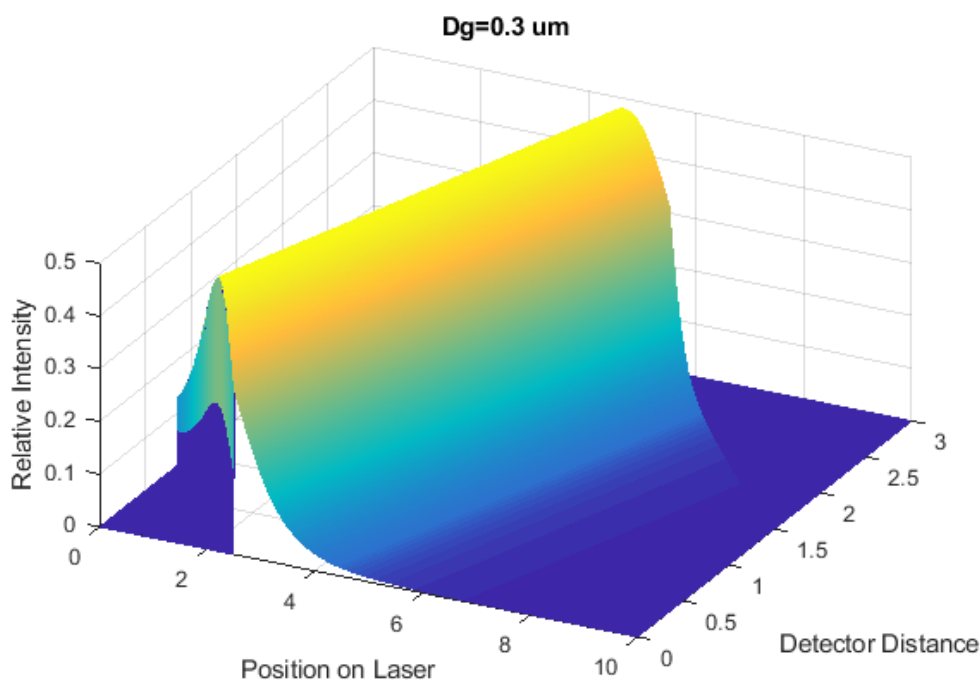
**Figure 4.** Relative response of the photodetector resulting from truncated scattering angles and a recessed photodetector. See explanation of the different curves in text.

### 3.2 Model predictions - Intensity versus position on the detector

Figure 5 provides an example of the energy distribution on the photodiode as a function of position in the laser and on the diode resulting from scattering from particles represented by a lognormally distributed aerosol volume size distribution with a volume mean diameter of 0.33  $\mu\text{m}$  and geometric standard deviation of 1.7. Figure 5 shows model predictions of the relative



443 intensity of scattered light, where the values are proportional to energy flux impinging on the  
444 detector.



445 **Figure 5.** Relative intensity of radiant energy scattered by a lognormally distributed aerosol  
446 volume size distribution with a volume mean diameter of  $0.33 \mu\text{m}$  and geometric standard  
447 deviation of 1.7 as a function of location of scattering event in the laser and as a function of  
448 position on the photodiode. Assumed laser wavelength was 650 nm, and the particle index of  
449 refraction was assumed to be 1.53. Positions in the laser and detector are from left to right as in  
450 Fig. 1.  
451

452 The masking resulting from a recessed detector truncates the scattering both in the most forward  
453 and backward scattering angles. This masking is shown as the triangular area corresponding to  
454 distance down the laser and detector of 0.0–2.5 mm and 0.0–0.78 mm, respectively, for the  
455 forward scattering angles and 5.6–10 mm and 1.44–3.0 mm, respectively, for backscattering.  
456 Because the laser is parallel to the photodetector, which is assumed to have a  $\cos(90-\theta)$  response,  
457 the maximum energy scattered to the detector is approximately at  $\theta = 90$  degrees. However,  
458 more energy is scattered to the detector for scattering angles less than 90 degrees, which  
459 corresponds to forward scattering, and very little energy is detected by the photodiode for  
460 particles in the laser that are greater than about 8 mm down the laser beam, even though the  
461 detector is exposed to particles in the laser that are 10 mm away from the laser exit hole. These  
462 distances down the laser correspond to backscattering. The total energy detected by the  
463 photodiode is the sum or integral across both the detector surface and position in the laser and  
464 corresponds to the volume under the curve depicted in Fig. 5.

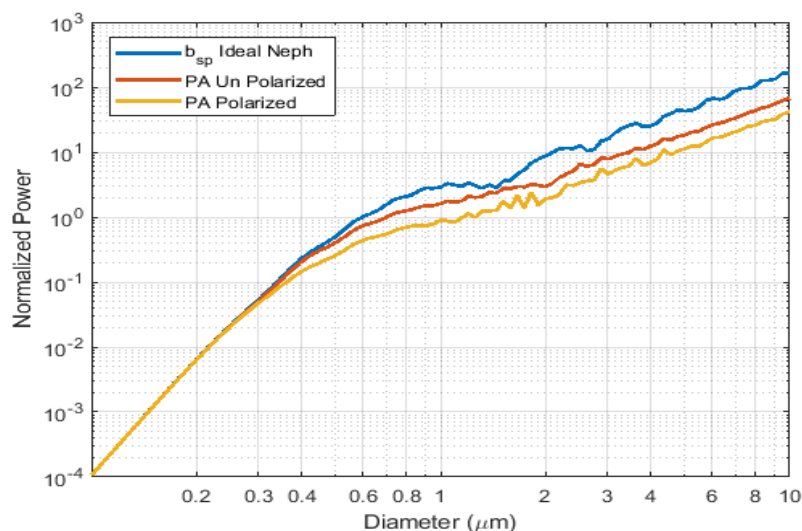


### 465 3.3 Model predictions - Predicted photodiode response as a function of particle diameter

466 The PMS differs from a perfect nephelometer in at least five important ways:

- 467 1. The laser is polarized, whereas the nephelometer light source is unpolarized.
- 468 2. The laser beam profile is not a simple plane wave, but complex in shape. The laser beam  
469 profile evolves significantly as it is focused over the photodiode.
- 470 3. The photodiode likely does not have a perfect cosine response.
- 471 4. The PMS geometry limits the photodiode to receiving scattered light between  
472 approximately 18 and 166 degrees, whereas a perfect nephelometer measures all energy  
473 scattered between 0 and 180 degrees.
- 474 5. The unknown PMS signal processing removes the light scattering signal from CO<sub>2</sub>, Suva,  
475 and filtered air. These gases are used to calibrate nephelometers, but cannot be used to  
476 calibrate the PMS.

477 The effects of these differences can be seen in Fig. 6, which shows predicted photodiode  
478 response as a function of particle diameter. The perfect nephelometer response is in blue, and  
479 the PMS response is in yellow. The red line predicts PMS response if the laser were not  
480 polarized. Relative intensities have been normalized to an ideal nephelometer measurement of a  
481 0.1  $\mu\text{m}$  diameter particle, which is akin to adjusting the laser power such that the scattered power  
482 at a diameter equal to 0.1  $\mu\text{m}$  is the same for all configurations. Scattering as a function of  
483 particle diameter is nearly the same for all three configurations from 0.1  $\mu\text{m}$  to about 0.3  $\mu\text{m}$ . At  
484 about 0.8 to 1.0  $\mu\text{m}$ , the response of a PMS with an unpolarized laser is about half that of an  
485 ideal nephelometer, and the use of a polarized laser reduces its response to about 30% to that of  
486 an ideal nephelometer. For particles above 2  $\mu\text{m}$  in diameter, the PMS response compared to an  
487 ideal nephelometer is decreased by about 75%. Additionally, the PMS manual (Zhou, 2016)  
488 quotes a lower detection limit diameter of 0.3  $\mu\text{m}$ . The model predicts that particles smaller than  
489 0.3  $\mu\text{m}$  in diameter would be detected by the PMS, in direct proportion to their contribution to  
490 the scattering coefficient.

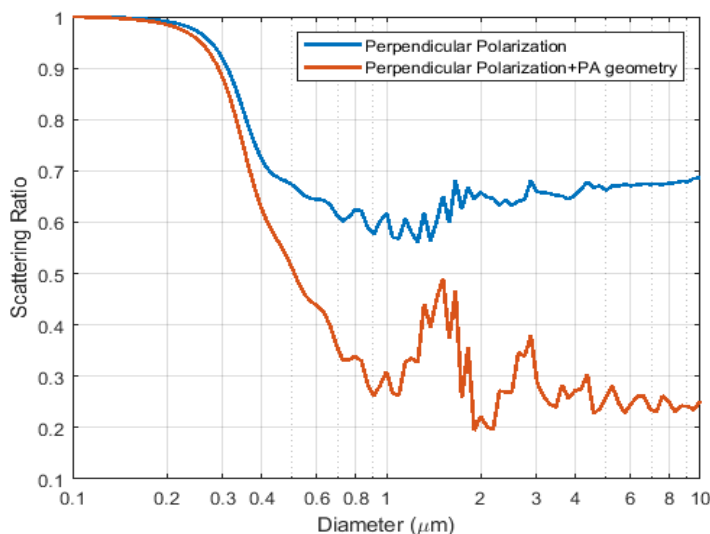


491



492 **Figure 6.** Normalized power detected by an ideal integrating nephelometer, a PMS with an  
493 unpolarized light source, and a PMS with a perpendicularly polarized light source plotted as a  
494 function of particle diameter. Modeled light source wavelength is 657 nm, and the particle index  
495 of refraction is 1.55. See explanation of the different curves in text.

496 These differences in geometry and optics from an ideal nephelometer are further highlighted in  
497 Fig. 7. To highlight the effect of polarization, the blue line shows the ratio of an ideal  
498 nephelometer with a laser light source that is perpendicularly polarized to an ideal nephelometer  
499 with an unpolarized light source. The red line shows the effects that polarization and PMS  
500 geometry have on the measured scattering signal. Again, all hypothetical instrument responses  
501 have been normalized to a particle diameter of 0.1  $\mu\text{m}$ . Relative to scattering for a 0.1  $\mu\text{m}$   
502 particle, the polarization alone reduces the scattering signal of an ideal nephelometer by 40% for  
503 particles with diameters in the 0.8–1.5  $\mu\text{m}$  size range. The additional effect of PMS scattering  
504 geometry reduces the scattering signal at 0.8–1.0  $\mu\text{m}$  by about another 30% relative to an ideal  
505 nephelometer.



506 **Figure 7.** Ratio of scattering of a “perfect” nephelometer to a nephelometer with a light source  
507 that is perpendicularly polarized (blue) and to a perpendicularly polarized nephelometer with  
508 PMS geometry (red) as a function of particle diameter.  
509

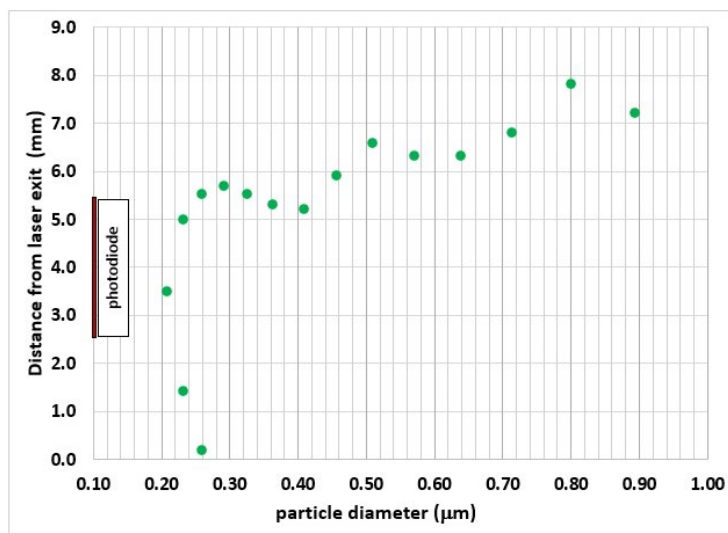
### 510 3.4 Model predictions – Differentiating by particle size

511 The irradiance received by the PMS photodiode from a particle of a given diameter and  
512 refractive index depends on the particle’s location in the laser beam. The model predicts that  
513 particles of different sizes may contribute the same irradiance to the photodiode, depending on  
514 their location in the beam, or conversely, light scattered by a particle of a given size can vary by  
515 more than an order of magnitude.

516 As an example, the model predicts that all of the particles in Fig. 8 contribute the same irradiance  
517 to the PMS photodiode. The smaller particles contribute the same irradiance by scattering in the



518 more effective forward scattering regime. The larger particles contribute the same irradiance by  
519 scattering in the less effective backscattering regime. The photodiode and its associated  
520 electronics would not be able to differentiate between them. As a result, the model predicts that  
521 the values reported in the six PA-PMS particle size channels from  $>0.3 \mu\text{m}$  to  $>10 \mu\text{m}$  cannot  
522 correctly represent the aerosol size distribution.



523  
524 **Figure 8.** The model predicts that different size particles can generate the same irradiance on the  
525 photodiode, depending on their location in the laser beam. In this example, each of the particles  
526 would create  $1.7 \times 10^{-2}$  picowatts of scattered irradiance on the photodiode.

## 527 4. Experimental – Field studies

528 Field experiments were conducted at two of the NFAN aerosol monitoring stations: the Mauna  
529 Loa Baseline Observatory in Hawaii and the Table Mountain Test Facility in Colorado. Both  
530 sites have large suites of aerosol instrumentation and daily access for scientists and technicians to  
531 inspect, calibrate, and maintain the instruments. These sites also have integrating nephelometers  
532 (TSI 3563, St. Paul, MN, USA) against which to evaluate the PA monitors.

### 533 4.1 Description of Mauna Loa site

534 The Mauna Loa Baseline Observatory (MLO) is located on the north flank of the Mauna Loa  
535 volcano, on the Big Island of Hawaii ( $19.536^\circ\text{N}$ ,  $155.576^\circ\text{W}$ , 3397 m asl). The observatory is a  
536 premier atmospheric research facility that has been continuously monitoring and collecting data  
537 on global background conditions and atmospheric change since the 1950s  
538 (<https://www.esrl.noaa.gov/gmd/obop/mlo/>). Continuous aerosol measurements at MLO began in  
539 the mid-1970s with the installation of condensation particle counters and an integrating  
540 nephelometer (Bodhaine and Mendonca, 1974; Bodhaine et al., 1981). MLO lies above the  
541 strong marine temperature inversion layer present in the region, which separates the more-  
542 polluted lower portions of the island atmosphere from the marine boundary layer from the much  
543 cleaner free troposphere. MLO experiences a diurnal wind pattern (Ryan, 1997) that is strongly  
544 influenced by the daily heating and nighttime cooling of the dark volcanic lava rock that makes





545 up the mountain. This “radiation wind” brings air up from lower elevations during the daytime,  
546 when atmospheric measurements reflect the local mountain environment. In contrast, during the  
547 nighttime, downslope winds develop, and the measurements at MLO are typically dominated by  
548 clean, free-tropospheric conditions (Chambers et al., 2013). At these times, the aerosol  
549 measurements at MLO often reflect some of the cleanest conditions at any station in the northern  
550 hemisphere. It has long been known, however, that episodic long-range transport of Asian  
551 pollution and dust aerosols occurs, most frequently in the springtime (Shaw, 1980; Miller, 1981;  
552 Harris and Kahl, 1990), and these aerosol events can influence both the daytime and nighttime  
553 measurements at MLO. Consequently, the aerosol levels at MLO vary over a large range, from  
554 extremely low to at times mildly elevated. Here we use observations from the MLO integrating  
555 nephelometer to evaluate the PMS sensor.

#### 556 **4.2 Description of Boulder Table Mountain site**

557 The Table Mountain Test Facility (BOS) is a large restricted-access federal complex located 14  
558 km north of Boulder, Colorado (40.125 °N, 105.237 °W, 1689 m asl). NOAA conducts  
559 atmospheric research at this site, and in addition to its NFAN station, it is one of the Global  
560 Monitoring Laboratory’s seven U.S. Surface Radiation Network (SURFRAD) sites  
561 (<https://www.esrl.noaa.gov/gmd/grad/surfrad/tablemt.html>). Many instruments for measuring  
562 surface and column aerosol properties are maintained at this location and used for long-term  
563 monitoring of the atmosphere.

564 The BOS site lies just east of the Front Range foothills of the Rocky Mountains and is typical of  
565 a semi-arid, high plains environment. It is a high mesa of predominantly grassland with some  
566 desert scrub vegetation. The location is well suited for sampling of wildfire smoke plumes during  
567 fire season in the western United States (summer and autumn), dust events at any time of the  
568 year, and occasional urban pollution episodes. The NFAN station at BOS  
569 (<https://www.esrl.noaa.gov/gmd/aero/net/bos.html>) was completed in September 2019. BOS  
570 operates an integrating nephelometer and a differential mobility particle spectrometer (DMPS).  
571 Both provided useful data for evaluating some of the predictions from the physical-optical model  
572 we developed for the PMS sensor.

#### 573 **4.3 PA monitors**

574 PA-PMS monitors were installed on the aerosol towers at the MLO and BOS stations, just below  
575 the main aerosol inlets. MLO had two PA-PMS monitors, one gently heated and one unheated,  
576 whereas BOS had one gently heated PA-PMS monitor. Prior to deployment, the monitors were  
577 tested in a filtered air chamber for 4 hours to ensure that the 1-h average CH1 values were less  
578 than 1 when no particles were present. One of the PMS sensors in the unheated MLO PA had 1-h  
579 average CH1 values of 27 when no particles were present. The heated monitors were wrapped  
580 with heating tape and powered by small DC power supplies. All the monitors were covered with  
581 stainless steel flashing 5 cm below the bottom to prevent rain and snow from entering the inlet  
582 (Fig. S13).

583 The PA-PMS monitors were warmed in an effort to reduce the sample relative humidity (RH) to  
584 be closer to that of the nephelometer, which is unavoidably heated by the warmth of the  
585 laboratory and by the nephelometer’s halogen lamp to above ambient temperatures. Because of  
586 this warming, the RH inside the nephelometers rarely exceeded 40%. Both MLO and BOS are



587 low-RH environments under normal conditions, although occasionally moist air masses are  
588 encountered. The heating of the monitors increased the sample temperatures by 5–8 C, which  
589 helped to lower the sample RH. While the PA heating was not controlled to achieve an RH  
590 match with the nephelometer, it brought the sample RH of the two measurements closer together.

591 Due to internet protocols at both sites, PA's wireless data transmission feature was not used, and  
592 the data were stored on the internal micro-SD card. At approximately 1 month intervals, the data  
593 were downloaded from the micro-SD cards, and the PAs were returned to service. Outputs from  
594 the two PMS sensors were then compared at these intervals to determine if the PAs were still  
595 functioning properly. In this study, the 80 s or 2 min averages were used to create 1 h averages to  
596 compare the PA observations to those of the nephelometer and the DMPS.

#### 597 **4.4 Integrating nephelometer**

598 The integrating nephelometer (TSI Inc., model 3563) measures the aerosol light scattering  
599 coefficient at three wavelengths (450, 550, and 700 nm). At both sites, the sample flow path is  
600 switched every 6 minutes between 1 and 10  $\mu\text{m}$  aerodynamic diameter, multijet, Berner-type  
601 impactors. Here, the scattering coefficients at 550 nm for both the PM1 and PM10 size fractions  
602 are used for comparison with the PMS measurements. These are referred to as  $b_{\text{sp}1}$  and  $b_{\text{sp}10}$ ,  
603 respectively.

604 There are two quality checks of the nephelometer operation made in the field. First, the  
605 nephelometer automatically samples filtered air once per hour. This provides a record of the  
606 stability of the instrument background measurement. Second, the nephelometer calibration is  
607 manually checked on a monthly basis using  $\text{CO}_2$  and filtered air (Anderson et al., 1996). The 1 h  
608 average  $b_{\text{sp}1}$  in filtered air is  $0.01 \text{ Mm}^{-1}$  with a standard deviation of  $0.12 \text{ Mm}^{-1}$ , based on 125  
609 hours of sampling filtered air.

610 The nephelometer measurements were corrected for angular truncation (Anderson and Ogren,  
611 1998) and reported at STP. Weekly data review provides quality assurance of the nephelometer  
612 data. Scattering coefficient data were averaged to 1 min resolution for logging and were further  
613 averaged to hourly resolution for comparison with the PA data. The 1 h average  $b_{\text{sp}1}$  uncertainties  
614 of the nephelometer measurements are  $\sim 0.13 \text{ Mm}^{-1}$  for scattering coefficients less than  $1.0 \text{ Mm}^{-1}$   
615 and  $\sim 10\%$  for scattering coefficients greater than  $1 \text{ Mm}^{-1}$  (Sherman et al., 2015).

#### 616 **4.5 Differential mobility particle spectrometer (DMPS)**

617 The DMPS was provided by the Institute for Atmospheric and Earth System Research,  
618 University of Helsinki, Finland. It was checked and calibrated by the World Calibration Centre  
619 for Aerosol Physics (WCCAP) at Leibniz Institute for Tropospheric Research (IfT), Leipzig,  
620 Germany, just prior to deployment at NOAA's Table Mountain site. After shipment from IfT to  
621 NOAA, the DMPS was again checked by aerosolizing polystyrene latex spheres and confirming  
622 that the peaks occurred in the correct size bins. The DMPS was housed inside the same building  
623 as the nephelometer at BOS and sampled aerosols through the same inlet, although the DMPS  
624 flow did not pass through the aerosol impactors.

625 The DMPS provides 40 channels of particle concentration versus size, ranging from mobility  
626 diameters of  $0.01 \mu\text{m}$  to  $0.8 \mu\text{m}$ . The  $0.1 \mu\text{m}$  to  $0.8 \mu\text{m}$  channels of the DMPS were used to



627 calculate hourly-average fine aerosol scattering coefficient distributions and the total fine aerosol  
628 scattering coefficient. The hourly-average, DMPS-calculated fine aerosol scattering coefficients  
629 were compared to the nephelometer-measured fine aerosol scattering coefficients to check  
630 operational consistency (Fig. S14). No operational changes were made to the DMPS during this  
631 field study. This study did not measure coarse aerosol size distributions. The DMPS hourly-  
632 average fine aerosol scattering coefficient distributions were used with the PMS physical-optical  
633 model to predict total 1 h average scattered irradiance on the photodiode.

## 634 5. Results

635 This section describes our evaluation of the PA-PMS using field data from MLO and BOS. First,  
636 we provide an overview of the observational data. We then assess how well the model described  
637 in Sect. 3 is able to represent the observed data and show consistency with results previously  
638 reported in the literature. Next, we present results showing the potential of the PA-PMS to  
639 perform as a nephelometer. Finally, we note how the size information output by the PA is not  
640 correct due to the PA's primary measurement being a scattering measurement. For the results  
641 presented below, data from the PA-PMS, nephelometer and DMPS were averaged to hourly  
642 frequency and merged prior to analysis.

### 643 5.1 Field data overview

644 Heated PA monitors were deployed at the MLO and BOS observatories for 15 and 11 months,  
645 respectively (Table 1). At both sites weather had no impact on the operation of the PA  
646 instrument, and downtime only occurred during data downloading.

647 **Table 1.** Summary of PA, TSI nephelometer, and DMPS data coverage.

site	number of hours				percent coverage			Time period
	PA-PMS	TSI neph	DMPS	overlap	PA-PMS	TSI neph	DMPS	
MLO	9371	9204	na	9204	97.6	95.9	na	6 May 2019 to 5 June 2020
BOS	7716	7479	7045	6901	97.7	94.7	89.2	13 February 2020 to 6 January 2021

648 These two deployments provide an excellent dataset for assessing PA performance in both a  
649 clean location (MLO) and in an environment with more elevated particle concentration (BOS).  
650 As shown in Table 2, during the field study at MLO, the median CH1 was 26.7. The median  $b_{sp1}$   
651 was  $0.76 \text{ Mm}^{-1}$  at 550 nm which is approximately 10% of Rayleigh scattering at the MLO  
652 altitude. The reported PM<sub>2.5</sub> mass concentration from the PA was zero for most of the MLO  
653 deployment. The CH1 and  $b_{sp1}$  are adjusted to STP in Table 2. The air quality at BOS was less  
654 pristine than at MLO and is more representative of nonurban continental air quality. The very  
655 high maximum CH1 and  $b_{sp1}$  at BOS reported in Table 2 occurred during smoke events in the  
656 summer and autumn of 2020. One of the BOS PMS sensors experienced approximately 10%  
657 degradation in sensitivity after one year in the field (Fig. S15).  
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**Table 2.** Summary of PA-PMS and nephelometer hourly observations at MLO and BOS.

Site	1h median (average)				1h Range min-max			
	PA-PMS		TSI nephelometer		PA-PMS		TSI nephelometer	
	PM2.5 μg m <sup>-3</sup>	CH1	b <sub>sp10</sub> Mm <sup>-1</sup>	b <sub>sp1</sub> Mm <sup>-1</sup>	PM2.5 μg m <sup>-3</sup>	CH1	b <sub>sp10</sub> Mm <sup>-1</sup>	b <sub>sp1</sub> Mm <sup>-1</sup>
MLO	0.000 (0.12)	26.7 (75.2)	1.19 (2.82)	0.76 (1.50)	0.0 - 21.6	0.26 - 1649	-0.35 - 35.2	-0.29 - 34.2
BOS	3.37 (8.42)	720 (1422)	14.6 (32.4)	9.9 (20.9)	0.0 - 571	7.38 - 63340	-0.11 - 4097	-0.44 - 2596

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## 5.2 Relationship between model predictions and field data

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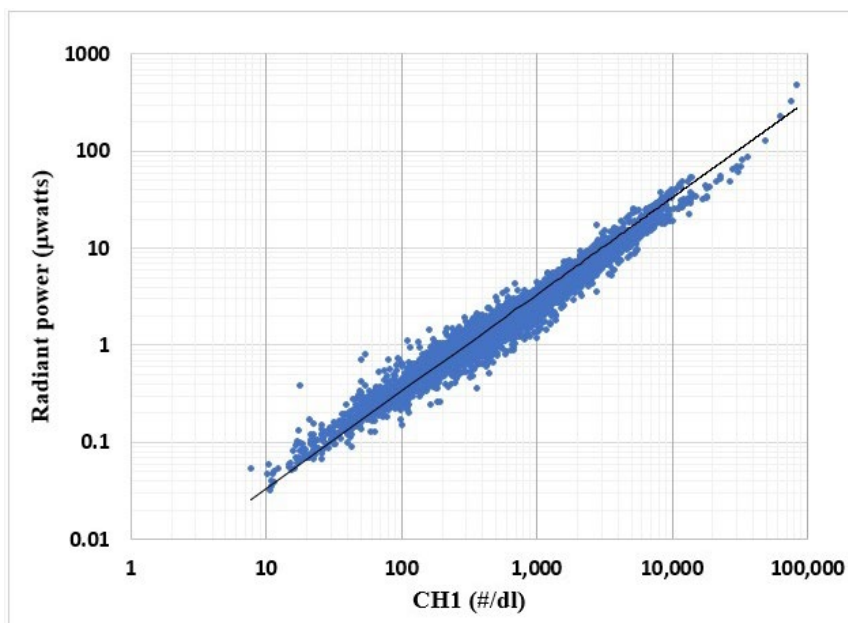
The PMS sensor is described by the manufacturer as a particle counter that measures particles between 0.3 μm and 10 μm in six size bins. Based on the theoretical characterization of the PMS sensor described in Sect. 3, the sensor is more akin to a polarized, reciprocal integrating nephelometer than a particle counter. Below, the field data and theoretical model are used to demonstrate that the raw PMS sensor signal is an integrated scattering measurement that is sensitive to particles smaller than 0.3 μm but relatively insensitive to particles larger than 1.0 μm.

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### 5.2.1 Predicted photodiode irradiance versus CH1 field data at BOS

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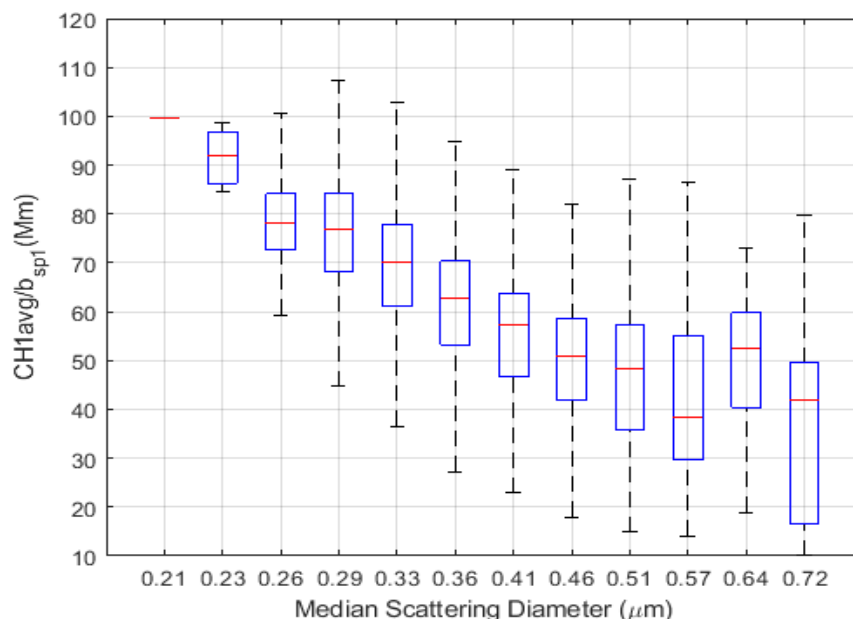
Our model described in Sect. 3 and the Appendix predicts a value proportional to the scattered irradiance impinging on the PMS photodiode as a function of particle diameter and concentration. This was done using the DMPS size distribution data from BOS. The modeled PMS photodiode output is plotted against the PMS CH1 output (Fig. 9). The predicted photodiode output is linearly correlated ( $R^2 = 0.90$ ) with CH1 over 4 orders of magnitude. The linear relationship between CH1 and modeled photodiode response suggests the likelihood that the CH1 output is directly related to what the photodiode is sensing (i.e., scattering from all particles in the scattering volume). The PA-PMS reported values, such as concentrations of particle numbers in various size ranges or PM concentrations, are quantities derived from the scattering signal and the use of an undescribed algorithm.



681  
682 **Figure 9.** One hour average CH1 reported values are plotted against model-predicted radiant  
683 power (or energy) in  $\mu\text{watts}$  on the photodiode. Both the CH1 and DMPS data were adjusted to  
684 STP conditions. The ordinary least squares regression line is also shown.

### 685 **5.2.2 Predicted aerosol size truncation versus field data at BOS**

686 The PMS physical-optical model described in Sect. 3 predicts that if CH1 is proportional to the  
687 photodiode power, then its signal will be truncated relative to a perfect nephelometer. Thus, the  
688 ratio  $\text{CH1}/b_{\text{sp1}}$  should decrease as median scattering diameter increases. To test this prediction,  
689 the DMPS data from BOS were used to calculate hourly-average aerosol scattering coefficient  
690 distributions for diameters between  $0.1 \mu\text{m}$  and  $0.8 \mu\text{m}$ . A wavelength of  $657 \text{ nm}$  and a particle  
691 refractive index of  $1.53 - 0.0i$  were used for the calculations. The median scattering diameter  
692 (MSD) was calculated for each hour. The MSD is the aerosol diameter at which approximately  
693 half of the light scattering is due to particles smaller than the MSD and the other half to particles  
694 larger than the MSD. The MSD was then compared to the ratio of the measured CH1 and  $b_{\text{sp1}}$   
695 values, i.e.,  $\text{CH1}_{\text{avg}}/b_{\text{sp1}}$ , for each of these hours. The results shown in Fig. 10 are consistent  
696 with the PMS physical-optical model. The highest  $\text{CH1}_{\text{avg}}/b_{\text{sp1}}$  ratios tend to occur for aerosols  
697 with the lowest MSD and decrease as MSD increases. The results show, as suggested above,  
698 that the PMS can efficiently detect particles below  $0.3 \mu\text{m}$  in diameter in proportion to their  
699 contribution to the scattering coefficient.



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**Figure 10.** Observed decrease in  $\text{CH1avg}/b_{\text{sp1}}$  ratio as function of median scattering diameter. Red line is the median value, while the upper and lower edges of the blue box represent the 75<sup>th</sup> and 25<sup>th</sup> percentile values, respectively. Whiskers extend to the 9<sup>th</sup> and 91<sup>st</sup> percentiles. Overall average  $\text{CH1avg}/b_{\text{sp1}}$  ratio for all hours is 65. Data are for 6839 hourly averages at BOS. Approximately 67% of the MSDs observed at BOS were between 0.29  $\mu\text{m}$  and 0.36  $\mu\text{m}$ , and 98% of MSDs were between 0.26  $\mu\text{m}$  and 0.46  $\mu\text{m}$ .

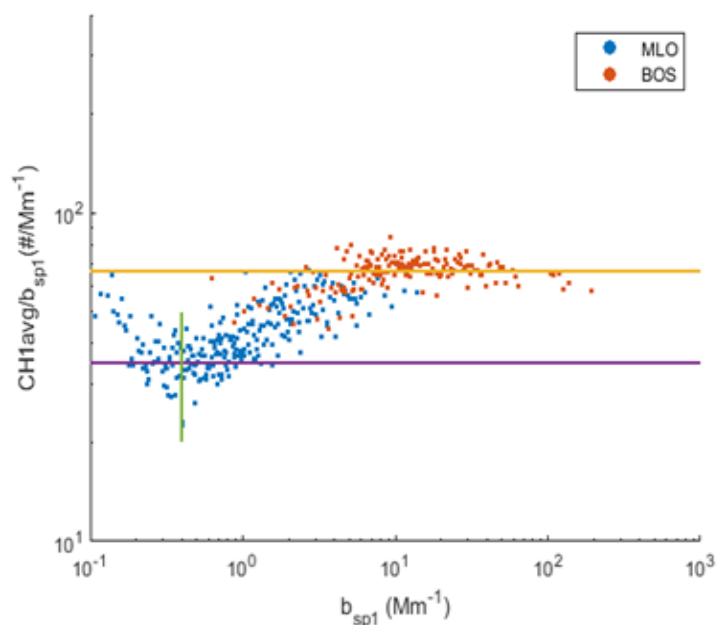
### 707 5.2.3. Estimating the scattering coefficient minimum detection limit of the PA-PMS

708 The precision analysis in Sect. 2 indicates that the PA monitors used in this study estimated 1 h  
709 average CH1 and CH1avg MDLs of approximately 21 and 14, respectively. The estimated 1 h  
710 average MDL  $b_{\text{sp1}}$  of the TSI 3563 nephelometer is approximately 0.20  $\text{Mm}^{-1}$ , based on filtered  
711 air tests. Further analysis of the relationship between CH1 and  $b_{\text{sp1}}$  at low levels was performed  
712 by plotting the ratio,  $\text{CH1avg}/b_{\text{sp1}}$ , for the combined MLO and BOS dataset, as a function of  $b_{\text{sp1}}$ .  
713 This relationship is shown graphically in Fig. 11. The data values were first averaged over 6  
714 hours because hourly  $b_{\text{sp1}}$  values near zero included many small negative  $b_{\text{sp1}}$  values due to the  
715 very clean conditions occasionally observed at MLO. The averaging eliminated all but five  
716 negative  $b_{\text{sp1}}$  values, which were removed from the dataset. The  $\text{CH1avg}/b_{\text{sp1}}$  and  $b_{\text{sp1}}$  values  
717 were further averaged over six data points after sorting the data on  $b_{\text{sp1}}$  levels to more clearly  
718 show the relationship between CH1avg and  $b_{\text{sp1}}$ . At  $b_{\text{sp1}} > 5 \text{ Mm}^{-1}$ , the  $\text{CH1avg}/b_{\text{sp1}}$  ratio is  
719 relatively constant at 67, the yellow line in Fig. 11. The yellow line is the slope of CH1avg  
720 versus  $b_{\text{sp1}}$  at  $b_{\text{sp1}}$  values greater than 5  $\text{Mm}^{-1}$ . The  $\text{CH1avg}/b_{\text{sp1}}$  ratio systematically decreases  
721 from its highest values to about 35, the slope of CH1avg versus  $b_{\text{sp1}}$  at  $b_{\text{sp1}} = 0.4 \text{ Mm}^{-1}$ . For  $b_{\text{sp1}}$   
722  $< 0.4 \text{ Mm}^{-1}$  the  $\text{CH1avg}/b_{\text{sp1}}$  ratio then increases significantly as  $b_{\text{sp1}}$  decreases, consistent with  
723 CH1avg values staying approximately constant below 0.4  $\text{Mm}^{-1}$ . Both the CH1avg and  $b_{\text{sp1}}$  are



724 below MDL for  $b_{sp1} < 0.2 \text{ Mm}^{-1}$ . A  $\text{CH1avg}/b_{sp1}$  ratio of approximately 35 at  $b_{sp1} = 0.4 \text{ Mm}^{-1}$   
725 and a  $\text{CH1avg}$  value of about  $14 \pm 5$  is consistent with the estimated  $\text{CH1avg}$  MDL of 14.

726 Based on these results, the 1 h average  $\text{CH1}$  sensor MDL for hourly data in units of scattering is  
727 approximately  $0.4 \text{ Mm}^{-1}$  at MLO. Laboratory tests challenging the PAs with known low-level,  
728 spiked aerosol concentrations and defined size distributions are needed to further refine the  
729 estimated MDL.



730  
731 **Figure 11.** Ratios of  $\text{CH1avg}$  and measured scattering,  $b_{sp1}$ , as a function of measured  $b_{sp1}$  for  
732 MLO and BOS. Green line corresponds to  $0.4 \text{ Mm}^{-1}$  while the purple line, a ratio of 35,  
733 corresponds to the additive uncertainty of 14. Yellow line corresponds to a  $\text{CH1avg}/b_{sp1}$  ratio of  
734 approximately 67, the slope of  $\text{CH1avg}$  vs.  $b_{sp1}$  above about  $5 \text{ Mm}^{-1}$ .

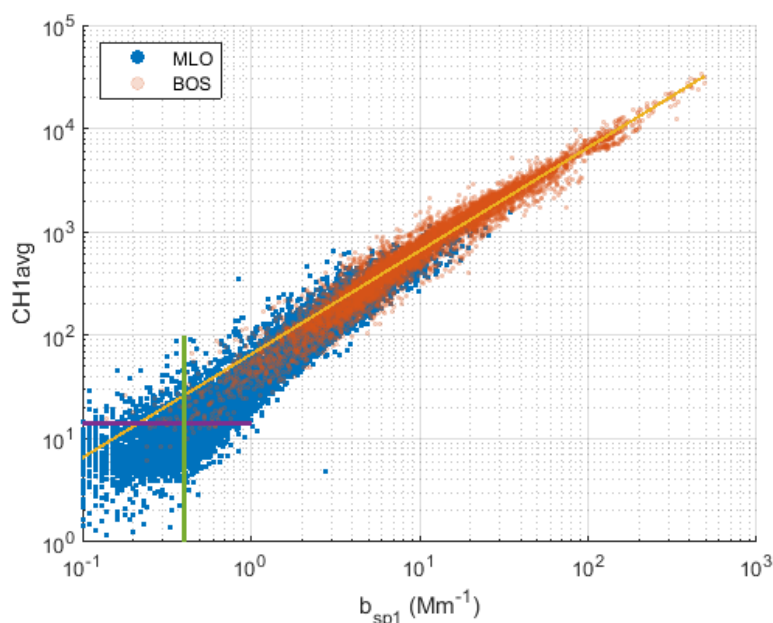
#### 735 5.2.4 Evaluating the use of the PA-PMS as a nephelometer

736 The MLO and BOS hourly-average  $\text{CH1avg}$  are plotted against  $b_{sp1}$ , measured at 550 nm, in Fig.  
737 12. Also shown in Fig. 12 is an ordinary least squares (OLS) regression line with the intercept  
738 set equal to zero using the BOS and MLO combined dataset but with values associated with  $b_{sp1}$   
739 less than  $0.4 \text{ Mm}^{-1}$  and greater than  $500 \text{ Mm}^{-1}$  removed. Results of the regression for the  
740 combined datasets as well as for the individual BOS and MLO datasets are presented in Table 3.  
741 There is good agreement for both datasets (Table 3) with an  $R^2$  of 0.97 and 0.85 for the BOS and  
742 MLO datasets, respectively, and 0.97 for the combined datasets. The relationship deviates  
743 somewhat from linear with increasing slopes and scatter at lower values of atmospheric  
744 scattering coefficient, particularly for the MLO data. The slopes for all data, MLO, and BOS, are  
745  $0.015 \pm 2.07 \times 10^{-5}$ ,  $0.017 \pm 5.72 \times 10^{-5}$ , and  $0.015 \pm 2.68 \times 10^{-5}$ , respectively. In the following analysis,  
746 a PA-derived atmospheric scattering ( $b_{sp1,PA}$ ) for both MLO and BOS is estimated using  $b_{sp1,CH1} =$   
747  $0.015 \times \text{CH1avg}$  at a wavelength of 550 nm. The “calibration” value of 0.015 corresponds to the



748 yellow horizontal line in Fig. 11 of 67.0 (1/0.015) and corresponds to a median scattering  
 749 diameter of about 0.33  $\mu\text{m}$  (Fig. 10).

750 Figure S16 shows that the submicron aerosol scattering coefficients at 550 nm and 700 nm are  
 751 highly correlated, with the 700 nm scattering coefficient averaging 52% of the 550 nm scattering  
 752 coefficient. This results in  $b_{\text{sp1,CH1}} = 0.0078 \times \text{CH1}_{\text{avg}}$  at a wavelength of 700 nm.



753  
 754 **Figure 12.** Fine aerosol scattering coefficient from TSI nephelometer vs. CH1avg value from  
 755 PA. Yellow line represents the fit to all data. Purple line shows the additive uncertainty of 14  
 756 while  $b_{\text{sp1}}$  values less than the green line were removed for the regression analysis.

757 **Table 3.** Ordinary least square regression coefficients with a zero intercept and standard error for  
 758  $b_{\text{sp1}}$  and CH1 as the dependent and independent variables, respectively, for the BOS, MLO, and  
 759 combined datasets. CH1 and  $b_{\text{sp1}}$  reported at STP. Also shown are the respective coefficients of  
 760 determination,  $R^2$ .

Site	slope	standard error	1/slope	$R^2$
BOS	0.015	$2.68 \times 10^{-5}$	67.0	0.97
MLO	0.017	$5.72 \times 10^{-5}$	59.0	0.85
Both BOS&MLO	0.015	$2.07 \times 10^{-5}$	67.0	0.97

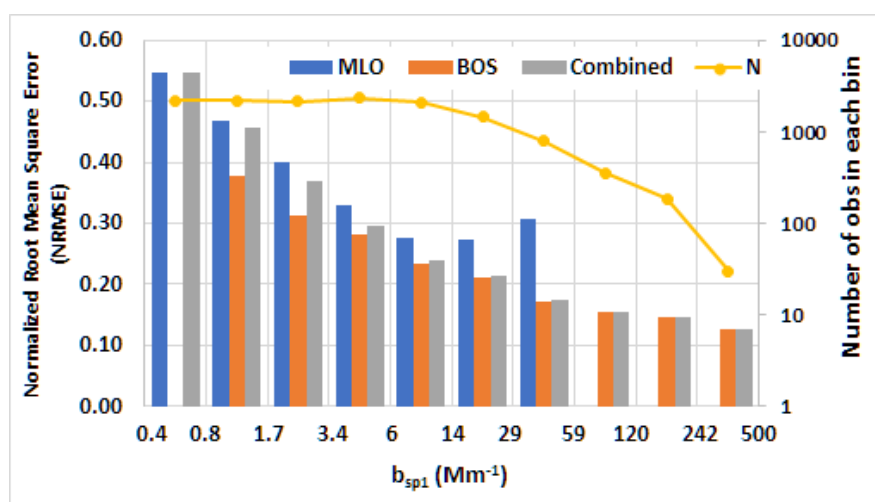
761  
 762 As discussed above, the regression coefficient between  $b_{\text{sp1}}$  and CH1 for the combined dataset of  
 763 0.015 is used to estimate the  $b_{\text{sp1,PA}}$  derived from the CH1 channel. The data for each dataset and  
 764 the combined dataset were binned into ten bins based on measured  $b_{\text{sp1}}$  levels that ranged from  
 765 0.4  $\text{Mm}^{-1}$  to 500  $\text{Mm}^{-1}$ . Values of  $b_{\text{sp1}}$  above 500  $\text{Mm}^{-1}$  were removed from the dataset. For each





766 bin the normalized root mean square error (NRMSE) between  $b_{sp1,PA}$  and measured  $b_{sp1}$   
767 calculated. The NRMSE values as a function of the  $b_{sp1}$  bins are plotted in Fig. 13 for the  
768 combined dataset represented as the gray bars and BOS and MLO represented by blue and  
769 orange bars, respectively.

770 For  $b_{sp1}$  levels less than  $0.8 \text{ Mm}^{-1}$ , the NRMSE is 45–55%, and for  $b_{sp1}$  levels greater than 10  
771  $\text{Mm}^{-1}$ , the NRMSE is about 25% or less. For  $b_{sp1}$  levels greater than  $60 \text{ Mm}^{-1}$ , the NRMSE  
772 approaches 15%.



773  
774 **Figure 13.** Normalized root mean square error between measured and estimated scattering from  
775 CH1 values plotted as a function of binned  $b_{sp1}$  for the BOS, MLO, and combined datasets.  
776 Yellow line is referenced to the right axis to provide the number of observations in each bin.  
777 Numbers on the x-axis represent the lower and upper levels of each scattering bin.

778 As discussed in Sect. 2.2.9, the uncertainty for high CH1avg values is small (1.9% to 4.8%). The  
779 precision of the TSI 3563 nephelometer is also similarly high, and together they account for  
780 about 10% NRMSE at high  $b_{sp1}$  values.

781 The overall normalized error is likely due to a variety of sources, primarily the variability in the  
782 CH1 values due to using a polarized light source and truncation errors due to the geometry of the  
783 PA-PMS sensors. Also, the variability in aerosol characteristics such as size distribution,  
784 refractive index, and shape may be important. At extremely low levels, uncertainty may also be  
785 due to a nonuniform distribution of particles in the PMS laser beam.

### 786 5.2.5 PA-PMS size distributions and PM2.5

787 The aerosol number concentrations from the six PMS size channels are unrealistic. The BOS  
788 field data showed that the concentration of particles larger than  $0.3 \mu\text{m}$  diameter calculated from  
789 the DMPS averaged 10 times higher than CH1 (Fig. S17). The other PMS size channels are so  
790 highly correlated with CH1 that they provide no additional information (Table S3). Furthermore,  
791 it appears that the PMS creates an approximately invariant normalized aerosol number  
792 distribution across a wide range of sites (Table S4, Fig. S18). Although the overall CH1



793 concentration can vary over 6 orders of magnitude (column 3 in Table S4), the shape of the PMS  
794 size distribution remains fairly constant. This suggests that the values in the channels above CH1  
795 are software generated and indicates that the most relevant output from the PMS is from the CH1  
796 channel. The bottom row of Table S4 shows that the PMS bin fractions above 1  $\mu\text{m}$  increased by  
797 only a factor of 2–5 in a high-PM<sub>2.5</sub> windblown dust episode at Keeler, California. This is  
798 consistent with the PMS model prediction that PMS coarse aerosol response is small relative to a  
799 perfect nephelometer.

800 The PM<sub>2.5</sub> was not measured by Federal Reference Method (FRM) or Federal Equivalent  
801 Method (FEM) instruments at MLO and BOS during this study. As a result, this study cannot  
802 compare the PA-PMS PM<sub>2.5</sub> channel results with these methods. Figure S19 shows that the PA-  
803 PMS PM<sub>2.5</sub> channel is reasonably well correlated with  $b_{\text{sp}1}$  for values greater than about 10–20  
804  $\mu\text{g m}^{-3}$ , typical of many moderately polluted locations, with a calculated mass scattering  
805 efficiency of approximately  $2.5 \text{ m}^2 \text{ g}^{-1}$ . However, it is likely that the PA-PMS underestimates  
806 PM<sub>2.5</sub> for very clean areas where  $b_{\text{sp}1}$  is often less than  $10 \text{ Mm}^{-1}$ . For example, the PA-PMS  
807 PM<sub>2.5</sub> was zero for 1099 of the hours in this study when  $b_{\text{sp}1}$  was greater than  $1 \text{ Mm}^{-1}$ .

808 The results above indicate that CH1 is the primary source of aerosol information from the PMS  
809 sensor. However, consistent with the sensor behaving like a cell-reciprocal nephelometer, it was  
810 found that CH1 was not the number concentration of particles having diameters greater than 0.3  
811  $\mu\text{m}$ . CH1 was approximately a factor of 10 lower than the DMPS number concentration for a  
812 similar size range.

## 813 6. Summary, Discussion, and Future Work

814 We have demonstrated that the PMS sensor inside the PA monitor (PA-PMS) appears to behave  
815 as an imperfect, reciprocal integrating nephelometer. As a scattering sensor, the PMS cannot  
816 directly count nor size particles in the air stream. The PMS uses an unknown algorithm to  
817 convert the scattering signal to a near-constant normalized number distribution from which PM  
818 concentrations are derived.

819 The scattering coefficient that is measured by an ideal integrating nephelometer does not need  
820 correction for any aerosol attributes such as shape, chemical composition, refractive index, or  
821 diameter. It is a valuable measure for visibility and global climate monitoring. Simple low-cost  
822 sensors such as the PA-PMS can play a role in estimating aerosol scattering coefficients and  
823 improving global coverage. Yearlong field data at NOAA's Mauna Loa Observatory and  
824 Boulder Table Mountain sites show that the 1 h average of the PA-PMS CH1 is highly correlated  
825 with a nephelometer-measured fine aerosol scattering coefficient at 550 nm,  $b_{\text{sp}1}$ , over a wide  
826 scattering coefficient range of  $0.4 \text{ Mm}^{-1}$  to  $500 \text{ Mm}^{-1}$ . The relationship between CH1 and  $b_{\text{sp}1}$  at  
827 550 nm is found to be  $b_{\text{sp}1} = 0.015 \times \text{CH1}$  when both quantities are adjusted to the same  
828 temperature and pressure.

829 The physical-optical model developed in this paper for the PMS and the agreement with field  
830 data may motivate users of other low-cost sensors to develop similar models. It is possible that  
831 some of the other low-cost sensors also use polarized lasers in a cell-reciprocal configuration like  
832 the PMS. Such models would improve the understanding of sensor operation and help users  
833 better recognize the opportunities and limitations of other low-cost sensors in applications such  
834 as monitoring the scattering coefficient.



835 The strong relationship between  $b_{sp1}$  and CH1 and the agreement between the model and field  
836 data support characterizing the PA-PMS as an imperfect truncated cell-reciprocal nephelometer.  
837 The results demonstrate that it is possible to use the PA-PMS to estimate the 1 h average fine  
838 aerosol scattering coefficient across a wide range of aerosol scattering concentrations, provided  
839 the aerosol median scattering diameter is between 0.26  $\mu\text{m}$  and 0.46  $\mu\text{m}$ . The CH1 and  $b_{sp1}$   
840 relationship is dependent on the size distribution, and it is expected to change for locations and  
841 times where the particle size shifts to larger or smaller sizes than those measured at BOS and  
842 MLO.

843 We found that the PA-PMS has important limitations compared to integrating nephelometers. It  
844 measures the light scattering over a smaller angular range, causing a significant truncation of the  
845 scattering signal in the forward and backward directions. Nephelometers calibrate their scattering  
846 coefficient with  $\text{CO}_2$  or Suva, but the PMS is unresponsive to these gases. As a result, there is  
847 currently no convenient way to calibrate the PMS to ensure its accuracy. Neither PA nor  
848 Plantower provide technical support. Quality assurance and control are not as robust as one  
849 encounters for regulatory and scientific monitoring instruments. For this reason, it is useful to  
850 test the PMS sensors in filtered air before using them and to limit field use to those sensors that  
851 have 1 h average CH1 values less than 2. While sampling, it is necessary to compare 1 h  
852 averages from the two PMS sensors in each PA monitor to become aware of any changes and, if  
853 needed, to replace them in a timely fashion.

854 This study limited its findings to low-RH air, because both the PA monitors and the  
855 nephelometers were heated to reduce RH. Since RH plays such an important role in water uptake  
856 by hygroscopic aerosols and the concomitant increase in the scattering coefficient, future work is  
857 planned to compare unheated PA monitors with an unheated nephelometer that does not reduce  
858 RH before sampling. Our model predicts that the PMS may not be as responsive to hygroscopic  
859 growth as an unheated nephelometer. This would imply that the PMS might underestimate high-  
860 RH, low-visibility aerosol scattering coefficients.  
861



## 862 Appendix

863 The PMS physical-optical model makes some simplifying assumptions. The actual PMS laser  
864 beam profile is not a simple plane wave but complex in shape. The model assumes the laser is a  
865 plane wave with a constant laser beam irradiance profile. This allows the use of Mie theory to  
866 predict the light scattered by particles in the laser. Secondly, the model calculates the light  
867 scattered to a narrow strip across the middle instead of the entire photodiode. It assumes that the  
868 irradiance received by the narrow strip is representative of the entire photodiode.

869 The intensity of light scattered from a particle in the laser is

$$870 \quad I(\theta) = F_{dv} \beta(\theta) dv \quad (A1)$$

871 where  $I(\theta)$  is the intensity of light at angle  $\theta$  scattered from a particle in the volume element  $dv$   
872 (with units of Watt  $sr^{-1}$ );  $\beta(\theta)$  is the volume scattering function ( $m^{-1} sr^{-1}$ );  $F_{dv}$  is the incident laser  
873 flux density (Watt  $m^{-2}$ ) impinging on the volume element  $dv$ ; and  $dv$  is the volume element  
874 within the laser.

875 The volume scattering function for a monodisperse aerosol having a diameter  $D_p$  and number  
876 concentration  $N(D_p)$  in the PMS laser is

$$877 \quad \beta(m, \lambda, \theta, D_p) = (\lambda/2\pi)^2 N(D_p) |S_1(m, \lambda, \theta, D_p)|^2 \quad (A2)$$

878 where  $|S_1(m, \lambda, \theta, D_p)|^2$  is the perpendicular scattering intensity function;  $\lambda$  is the laser  
879 wavelength;  $m$  is the particle complex refractive index;  $\theta$  is the scattering angle; and  $D_p$  is the  
880 aerosol diameter. Note that  $\theta = 0$  in the direction of the laser, and  $\theta = 90$  degrees perpendicular to  
881 the laser and photodiode.

882 For one particle of size  $D_p$  in the volume element  $dv$ ,  $N(D_p) dv = (1/dv) \times (dv) = 1$ .

883 The incremental power  $dP$  (Watt) scattered from a particle in the volume element  $dv$  across a  
884 solid angle  $d\Omega$  subtended on the surface of a sphere at distance  $r$  from the particle, and normal to  
885  $r$ , is

$$886 \quad dP = I(\theta) d\Omega. \quad (A3)$$

887  $d\Omega = dA_0 / r^2$ , where  $dA_0$  is the incremental area on the sphere at distance  $r$  from the particle and  
888 normal to  $r$ .  $dP$  is then

$$889 \quad dP = I(\theta) dA_0 / r^2.$$

890 For the PMS model,  $dA_0$  is a small rectangle with width  $w$  and height  $rd\theta$ , where  $w$  is the width  
891 of the strip on the photodiode, and  $d\theta$  is the differential scattering angle.

892  $dA_0 = r d\theta \times w$ , where  $w$  is the width of the strip on the photodiode. From Fig. A1,  $r = b/\sin(\theta)$ ,  
893 where  $b$  is the distance from the laser to the photodiode.



894  $d\Omega = dA_0 / r^2 = (r d\theta \times w) / r^2 = d\theta \times (w/r) = (w/b) \times \sin(\theta)d\theta.$

895 The incremental power across the solid angle  $d\Omega$  normal to  $r$  is then

896  $dP = I(\theta) \times dA_0 / r^2 = I(\theta) \times (w/b) \times \sin(\theta)d\theta. \quad (A4)$

897 Substituting for  $I(\theta)$ ,

898  $dP(g,x,\theta) = [F_0 (\lambda/2\pi)^2 |S_1(\theta,D_p)|^2] \times (w/b) \times \sin(\theta) d\theta. \quad (A5)$

899 Equation A5 can be further simplified by combining the constants into  $K = (\lambda/2\pi)^2 F_0 w/b$ , where  
 900  $K$  has units of watts:

901  $dP(g,x,\theta) = K |S_1(\theta,D_p)|^2 \sin(\theta) d\theta. \quad (A6)$

902 The power received by the photodiode from a particle of diameter  $D_p$  in the volume element at  $x$   
 903 is obtained by numerically integrating across  $\theta$  on the photodiode:

904  $P(m,D_p) = K \int_{\theta_1(x)}^{\theta_2(x)} |S_1(m,\theta,D_p)|^2 \sin(\theta) d\theta. \quad (A7)$

905 Due to the PMS geometry, the upper and lower angular scattering integration limits for  $\theta$  depend  
 906 on the location  $x$ . This can be seen in Fig. S4. For example, at  $x = 0$  mm, the upper and lower  
 907 integration limits for  $\theta$  are 18 to 38 degrees. At  $x = 4.0$  mm, over the center of the photodiode,  
 908 the angular integration limits are 50 to 130 degrees.

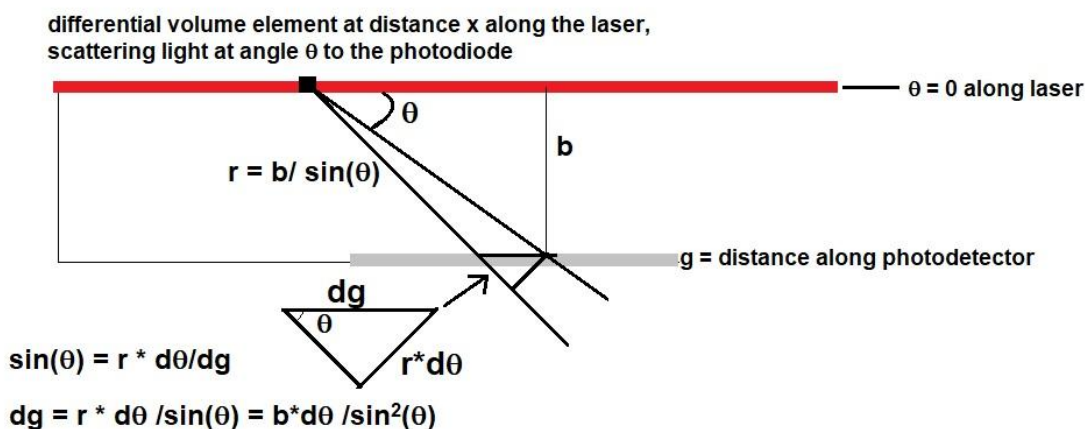
909 The total power  $P$  in Watts received by the photodiode from the light scattered by all the  
 910 particles of diameter  $D_p$  in the laser is obtained by carrying out the numerical integration in Eq.  
 911 A7 for all  $x$  from 0 to 10 mm:

912  $P(m,D_p) = K \int_{x=0}^{x=10mm} \int_{\theta_1(x)}^{\theta_2(x)} |S_1(m,\theta,D_p)|^2 \sin(\theta) d\theta dx. \quad (A8)$

913 The result for carrying out this calculation for the power per particle of size  $D_p$  is in Table S5 for  
 914 wavelength 657 nm and particle refractive index 1.53 - 0.015i. The total power received at the  
 915 photodiode by a distribution of particles is obtained by summing up the power per particle of size  
 916  $D_p$  times the number of particles  $N(D_p, m)$  in the size interval  $D_p$  to  $D_p + dD_p$ .

917  $P = K \int_{D_p} \int_{x=0}^{x=10mm} \int_{\theta_1(x)}^{\theta_2(x)} |S_1(m,\theta,D_p)|^2 \sin(\theta) N(D_p, m) d\theta dx dD_p. \quad (A9)$

918 Figure A1 shows the PMS geometry. The distance along the laser is the variable  $x$ , which ranges  
 919 from 0 to 10 mm. The distance along the photodiode is the variable  $g$ , which ranges from 0 to 3.0  
 920 mm. The distance between the photodiode and the laser is  $b$ , approximately 1.8 mm.



921 **Figure A1.** Sketch of PMS5003 geometry.

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## 929 Disclaimer

930 The assumptions, findings, conclusions, judgments, and views presented herein are those of the  
 931 authors and should not be interpreted as necessarily representing NOAA or National Park  
 932 Service policies. Reference to any companies or specific commercial products does not  
 933 constitute endorsement by NOAA or the National Park Service.

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