

1 Global Evaluation of the Precipitable Water Vapor Product from 2 MERSI-II onboard the Fengyun-3D Satellite

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10 **Abstract.** Atmospheric water vapor plays a key role in the Earth's radiation balance and hydrological cycle, and the precipitable
11 water vapor (PWV) product under clear sky condition has been routinely provided by the advanced Medium Resolution
12 Spectral Imager (MERSI-II) onboard FengYun-3D since 2018. The global evaluation of the PWV product precipitable water
13 vapor (PWV) derived from ~~MERSI-II the advanced Medium Resolution Spectral Imager (MERSI-II) onboard FengYun-3D~~ is
14 performed herein by comparing with the PWV from the Integrated Global Radiosonde Archive (IGRA) based on a total of 462
15 sites (57,219 match-ups) during 2018–2021. The monthly averaged PWV from MERSI-II presents a decreasing distribution
16 of PWV from the tropics to the polar regions. In general, a sound consistency exists between the PWVs of MERSI-II and
17 IGRA, and their correlation coefficient is 0.951 and root mean squared error (RMSE) is 0.36 cm. The histogram of mean bias
18 (MB) shows that the MB is concentrated around zero and mostly located within the range from -1.00 cm to 0.50 cm. For most
19 sites, the PWV is underestimated with the MB between -0.41 cm and 0.05 cm. However, there is also overestimated PWV,
20 which is mostly distributed in the surrounding area of the Black Sea and the middle of South America. There is a slight
21 underestimation of MERSI-II PWV for all seasons with the MB value below -0.18 cm, with the bias being the largest magnitude
22 in summer. This is probably due to the presence of thin clouds, which weaken the radiation signal observed by the satellite.
23 We also find that there is a larger bias in the Southern Hemisphere, with a large value and significant variation of PWV. The
24 binned error analysis revealed that the MB and RMSE increased with the increasing value of PWV, but there is an
25 overestimation for PWV smaller than 1.0 cm. In addition, there is a higher MB and RMSE with a larger spatial distance
26 between the footprint of the satellite and the IGRA station, and the RMSE ranged from 0.33 cm to 0.47 cm. There is a notable
27 dependency on solar zenith angle of the deviations between MERSI-II and IGRA PWV products.

28 1 Introduction

29 Water vapor is an important component of the atmosphere and widely known as an important greenhouse gas since it can

30 significantly affect climate change, the radiation balance, and the hydrological cycle (Kiehl & Trenberth, 1997; Held & Soden,
31 2000; Dessler & Wong, 2009; Zhao et al., 2012). The spatiotemporal variations of water vapor are essential for understanding
32 the formation of clouds and mesoscale meteorological systems in that cloud, and precipitation always rely on changes in water
33 vapor (Trenberth et al., 2003). Furthermore, water vapor can also influence the atmospheric transmittance, and the upward
34 radiance measured by the satellite sensor. Therefore, the information of water vapor is highly required to correct atmospheric
35 effects in the satellite-based retrieval algorithm for land surface temperature (Meng et al., 2017).

36 Considering the critical role of water vapor, technologies aiming at the measurement of atmospheric water vapor have
37 been developed. The precipitable water vapor (PWV), which means the integrated water vapor contained in a vertical column
38 of a cross-sectional area, is an important indicator for the total atmospheric water vapor condition. The two major methods
39 used for measuring PWV are satellite-based and ground-based technologies. Several ground-based measurements, such as
40 radiosonde (Durre et al., 2009), global position system (GPS) receivers (Bevis et al., 1992), microwave radiometer (MWR)
41 (Westwater, 1978) and sun photometer (Alexandrovet al., 2009), have been deployed to monitor the variability of water vapor.
42 However, the spatial coverage of ground-based measurements is limited and inhomogeneous, and it is difficult to obtain a wide
43 range of observations from multiple sources to support studies for the distribution of PWV on both a regional and global scale.
44 This is because the uncertainties of different measurements are not completely consistent, and they have distinct discrepancies
45 and magnitudes (Chen & Liu, 2016; Wang et al., 2016). Different from ground-based measurements, the satellite-based
46 measurement is more useful for the temporal analysis of PWV over a wide area. In particular, the polar-orbiting satellite-based
47 measurements of water vapor have a considerable advantage due to their global coverage with satisfactory temporal and spatial
48 resolutions. Therefore, the polar-orbiting satellite-based PWV product is widely used for understanding the global distribution
49 of water vapor. As is commonly known, the well knowledge of global water vapor distribution is especially important for
50 global atmospheric models aiming to predict weather or climate. Thus, the water vapor products retrieved via polar-orbiting
51 satellites have become essential input parameters to sustain numerical models of the atmosphere, especially where global water
52 vapor information is required within a short time span, and the assimilation of PWV has been proven to help improve
53 precipitation forecasts (Rakesh et al., 2009).

54 Over the past few decades, satellite-based PWV retrieval algorithms have been developed with observations from
55 different sensors, which can be divided into four main types according to the spectral region: (1) visible (VIS), (2) near-infrared
56 (NIR), (3) thermal infrared (TIR), and (4) microwave (MW). There are three major satellite-borne sensors that can provide
57 the global NIR PWV product. The Moderate Resolution Imaging Spectroradiometer (MODIS) onboard the Terra and Aqua
58 polar-orbiting satellite platforms is one of the most important instruments for obtaining global PWV, and has been widely used
59 for a few decades since the launching of Terra spacecraft in 1999. The Medium Resolution Imaging Spectrometer (MERIS) is
60 one of ten instruments built in Envisat, which was launched on 1 March 2002, but the mission was terminated on 8 April 2012
61 due to the loss of contact with the satellite. For the Chinese FengYun 3 (FY-3) meteorological series satellite, one of the major

62 payloads onboard is the Medium Resolution Spectral Imager (MERSI), which primarily monitors the ocean, land, atmosphere,
63 etc. FY-3D is the second-generation Chinese polar-orbiting meteorological satellite, equipped with the advanced MERSI
64 (MERSI-II), and was launched on 15 November 2017. For MERIS, the PWV retrieval algorithm employs the ratio of top of
65 atmosphere (TOA) radiance at one water vapor absorption channel (around 900 nm) to the TOA radiance at the atmospheric
66 window channel such as 885 nm (Bennartz and Fischer, 2001). However, both the algorithms for NIR PWV derivation of
67 MODIS and MERSI-II adopt the reflected solar radiance ratios between three NIR water vapor absorption channels and two
68 non-absorption channels (Gao and Kaufman, 2003; Wang et al., 2021). The setup of non-absorption channels of MERSI-II is
69 the same as that of MERSI but the absorption channels of MERSI-II are similar to those of MODIS. Furthermore, the prelaunch
70 and orbital calibration and characterization of MERSI-II were conducted to ensure the quality of its products (Xu et al., 2018).

71 It is necessary to evaluate the satellite-based PWV product ahead of its application in atmospheric science research. The
72 PWV from MODIS has been extensively evaluated by comparing it with the PWV derived from other measurements. The
73 GPS PWV is widely used for the evaluation of PWV derived from MODIS (Liu et al., 2006; Prasad and Singh, 2009; Lu et
74 al., 2011). Ground-based MWR, which can measure integrated water vapor with high temporal resolution and has a reliable
75 measurement under clear sky condition, is also used for the evaluation of MERIS PWV (Li et al., 2003). In addition, the
76 radiosonde PWV, calculated from the integration of specific humidity, has been recognized to be a useful benchmark, being
77 used for the evaluation of the MODIS PWV in China (Liu et al., 2015), the Iberian Peninsula (Sobrino et al., 2014), and Hong
78 Kong (Liu et al., 2013). However, up to now, few studies have focused on the evaluation of the MERSI-II PWV, and the lack
79 of effective assessments greatly limits the application of the MERSI-II PWV product, since the accuracy of the product has
80 not been fully acknowledged.

81 [The](#) Integrated Global Radiosonde Archive (IGRA) is the greatest and most comprehensive collection dataset of historical
82 and near real-time global quality-assured radiosonde observations. It has been used extensively in a variety of studies, including
83 model verification, atmospheric processes, and climate research. Moreover, the radiosonde PWV is also widely applied in the
84 assessments of measurements from other platforms, especially satellite-derived PWV around the world (Adeyemi and Schulz,
85 2012; Antón et al., 2015; Niilo et al., 2016). Consequently, the IGRA data are selected for the evaluation of the PWV derived
86 from MERSI-II in this study.

87 The purpose of this paper is to evaluate the MERSI-II PWV globally by comparing it with the global IGRA observations.
88 We are seeking to explore the global performance of FY-3D MERSI-II PWV and analyze the influence factors in the evaluation.
89 The structure of this paper is arranged as follows: Data sources and details are discussed in Section 2. Section 3 presents the
90 methodology of the merging procedures applied in the sample selection. The evaluation results of MERSI-II PWV against the
91 PWV from IGRA are presented in Section 4. In the final section, a discussion and conclusion of the aforementioned results
92 are given.

93 **2 Data description**

94 The satellite-based PWV product used in this paper is derived from FY-3D MERSI-II, and the ground-based
 95 measurements are the PWV data derived from AERONET and IGRA.

96 **2.1 MERSI-II PWV**

97 FY-3D, which was successfully launched on 15 November 2017, is the fourth and latest satellite of the second-generation
 98 Chinese polar-orbiting meteorological satellite. It is operated in a sun-synchronous orbit at an average altitude of 830.73 km,
 99 passing over the equator at 13:40 local time (Yang et al., 2019). ~~The MERSI is one of the major instruments carried by FY-3~~
 100 ~~series satellites, and the MERSI II onboard FY-3D is an upgraded version of the first generation instrument.~~ A series of
 101 comprehensive prelaunch calibrations have been operated to ensure the high quality of the products from MERSI-II (Xu et al.,
 102 2018), which is an advanced version of MERSI and has been significantly improved with high-precision on-board calibration
 103 and lunar calibration capabilities (Wu et al., 2020). Besides, MERSI-II has 25 channels with a spectral coverage from 0.412
 104 μm to 12.0 μm , and the NIR PWV products of FY-3D are retrieved with three water absorption channels (bands 16, 17, and
 105 18) and two non-absorption channels (bands 15 and 19) in the 0.8–1.3 μm range with a spatial resolution of 1 km \times 1 km at
 106 nadir (Wang et al., 2021). The positions and widths of NIR channels used by MERSI-II and MODIS are given in Table 1. The
 107 water vapor absorption channels of MERSI-II, which is now similar to those of MODIS, are reselected because the three
 108 absorption bands have different sensitivities to various water vapor conditions. Therefore, MERSI-II is more useful in the
 109 retrieval of water vapor under different conditions (dry, medium, and humid).

110 Table 1 Characteristics of NIR channels used in PWV retrievals of MERSI-II and MODIS

MERSI-II				MODIS			
Band No.	Position (nm)	Width (nm)	Window channel	Band No.	Position (nm)	Width (nm)	Window channel
15	865	20	yes	2	865	40	yes
16	905	20	no	17	905	30	no
17	936	20	no	18	936	10	no
18	940	50	no	19	940	50	no
19	1030	20	yes	5	1240	20	yes

111 For the NIR channels, ~~typically with a small the~~ aerosol optical thickness ~~is typically small and the path scattered~~
 112 ~~radiance that~~ can be ignored (Gao and Kaufman, 2003; Wang et al., 2021,-). Hence, the TOA radiance observed by a downward-
 113 looking satellite sensor can be calculated as the following:

114 $TOA_{\lambda} = T_{\lambda} \times \rho_{\lambda}$, (1)

115 where TOA_{λ} , T_{λ} and ρ_{λ} are the apparent reflectance, total atmospheric transmittance and surface bidirectional reflectance

116 at the channel with a wavelength of λ , respectively. The term $T(\lambda)$ contains information of the total amount of water vapor in
117 the Sun–surface–sensor path.

118 For ~~five major soil types~~~~most types of land surfaces~~, a reflectance between 850 and 1250 nm changes approximately
119 linearly with the wavelength, therefore, the transmittance of the absorption channel will be calculated by a three-channel ratio
120 of the absorption channel with a combination of two window channels. For the iron-rich soil, the vegetation and snow, although
121 the reflectance does not linearly correlate with the wavelength, reasonable estimates of water vapor transmittances over these
122 surface types can also be given with the three-channel ratio techniques (Gao and Kaufman, 2003).

123 By using the MODerate resolution atmospheric TRANsmission (MODTRAN), the transmittances of the five MERSI-II
124 channels as a function of the total water vapor amount under six different atmospheric conditions were calculated, according
125 to the six standard atmospheric models defined in MODTRAN4.3. Furthermore, the results are defined as the transmittance–
126 water vapor lookup table. Based on the solar zenith angle and surface temperature, the atmosphere model can be selected from
127 the six standard atmospheric models, and then the combined two-way water vapor content is derived using a table-searching
128 procedure. Note that there are no PWV retrievals in the region with a solar zenith angle above 72° , which means the observation
129 time is close to night, due to the weak energy at the satellite’s entrance. Subsequently, the derived total water vapor amount
130 will be converted to the vertical column water vapor amount based on the solar and observational geometries. The absorption
131 coefficients of atmospheric water vapor are very different over the three absorption channels. As a result, the derived water
132 vapor values from the three channels are different even under the same atmospheric condition. In order to solve this problem,
133 a mean water vapor value is obtained from the water vapor values derived from three absorption channels, by multiplying with
134 the corresponding weight in each channel. A more detailed description of the algorithm of MERSI-II PWV can be found in
135 Wang et al. (2021).

136 The NIR PWV products derived from MERSI-II have been routinely produced at the National Satellite Meteorological
137 Center, China and can be accessed on the website of <http://satellite.nsmc.org.cn/PortalSite/Data/Satellite.aspx>. The operational
138 NIR PWV products include the Level-2 5-min granule product and Level-3 global daily, 10-day, and monthly mean products.
139 The Level-2 products are generated on a pixel-by-pixel basis (i.e., $1 \text{ km} \times 1 \text{ km}$) from standard MERSI-II L1B radiance datasets
140 as well as ancillary data from the L1B geolocation and the cloud mask (CLM) product of MERSI-II. The outputs from the
141 Level 2 product include the total weighted column water vapor amount on a pixel-by-pixel basis, independently derived PWV
142 from one of the water vapor absorption channels, and an associated quality assurance parameter that indicates whether the
143 inversion algorithm has a two-channel or three-channel ratio and whether a pixel is clear or cloudy. The Level-2 5-min granule
144 PWV product is evaluated in this study, and the data span is from September 2018 to June 2021 with a spatial resolution of 1
145 $\text{km} \times 1 \text{ km}$.

146 **2.2 Radiosonde**

147 The Integrated Global Radiosonde Archive (IGRA), which is a collection of historical and near real-time global
148 radiosonde observations, is archived and distributed by the National Centers for Environmental Information (NCEI), formerly
149 known as the National Climatic Data Center (NCDC), and can be accessed at <ftp://ftp.ncdc.noaa.gov/pub/data/igra>. Version 2
150 of IGRA (IGRA 2) is used in this study. A total of 33 data sources, including 10 out of 11 source datasets used in IGRA 1,
151 have been integrated into IGRA 2, which was fully operational on August 15, 2016, and has a higher spatial and temporal
152 coverage. Therefore, compared to IGRA 1, the IGRA 2 contains nearly twice as many sounding stations and 30% more
153 soundings. Sounding-derived parameters are recorded according to separated station files and continue to be updated daily,
154 and PWV is one of the derived parameters. PWV will be calculated if the pressure, temperature, and dew point depression are
155 available from the surface to a level of 500 hPa (Durre et al., 2009). The calculation involves the acquirement of specific
156 humidity at each observation level and then the integration of specific humidity between the surface and the level of 500 hPa,
157 so that IGRA-derived PWV is recognized as surface-to-500-hPa PWV. As discussed by Turner et al. (2003), the PWV obtained
158 from radiosonde has an approximate 5% dry bias compared to that derived from the MWR. Therefore, there is an
159 underestimation of PWV evaluation for taking the IGRA-derived PWV as a reference, and the bias found in the tropical area
160 is ~9% (Zhang et al. 2018). Due to the time range of IGRA data, there are only 462 out of 1535 global IGRA stations that can
161 be matched with the FY-3D MERSI-II PWV products.

162 **2.3 AERONET**

163 The federated Aerosol Robotic Network (AERONET) is a network of ground-based Cimel Electronique Sun photometry,
164 which can measure beam irradiance and directional sky radiance routinely during the daytime in clear conditions (Holben et
165 al., 1998). AERONET was established by NASA and PHOTONS (PHOTométrie pour le Traitement Opérationnel de
166 Normalisation Satellitaire), primarily aiming to provide public domain dataset of global aerosol optical and microphysical
167 properties. In addition, based on the measurements at the 940 nm water-vapor channel and the atmospheric window bands
168 centered at 870 nm and 1020 nm, PWV was also calculated (Che et al., 2016). The AERONET version 3 database provides
169 three levels of data: Level 1.0 (unscreened), Level 1.5 (cloud-screened), and Level 2.0 (cloud-screened and quality-assured),
170 and can be accessed at <https://aeronet.gsfc.nasa.gov>. Level 2.0 dataset, which is used in this study, signifies an automatically
171 cloud-cleared, manually quality-controlled dataset with pre- and post-field calibrations applied. All the instruments in the
172 AERONET are annually calibrated with reference to the world standard: the Mauna Loa Observatory (Malderen et al., 2014).
173 Thus, the measuring accuracies of different AERONET stations are accurate and consistent (Liu et al., 2013). As discussed by
174 Pérez-Ramírez et al. (2014), PWV obtained from AERONET has a dry bias of approximately 0.16 cm against the radiosonde
175 PWV and it is reasonable for meteorological studies.

176 3 Methodology

177 3.1 Statistical indicators

178 The common statistical indicators, such as the mean bias (MB, perfect value = 0), the mean relative bias (MRB, perfect
179 value = 0), correlation coefficient (CC, perfect value = 1) and the root mean squared error (RMSE, perfect value = 0), are used
180 to evaluate the precision of the retrieved PWV from MERSI-II. All calculations of indicators are presented as follows:

$$181 \quad MB = \frac{1}{N} \sum_{i=1}^N (PWV_{si} - PWV_{gi}) , \quad (2)$$

$$182 \quad MRB = \frac{1}{N} \sum_{i=1}^N \left(\frac{PWV_{si} - PWV_{gi}}{PWV_{gi}} \right) \times 100\% , \quad (3)$$

$$183 \quad CC = \frac{\sum_{i=1}^N (PWV_{si} - \overline{PWV_{si}})(PWV_{gi} - \overline{PWV_{gi}})}{\sqrt{\sum_{i=1}^N (PWV_{si} - \overline{PWV_{si}})^2 \sum_{i=1}^N (PWV_{gi} - \overline{PWV_{gi}})^2}} , \quad (4)$$

$$184 \quad RMSE = \sqrt{\frac{1}{N} \sum_{i=1}^N (PWV_{si} - PWV_{gi})^2} \quad (5)$$

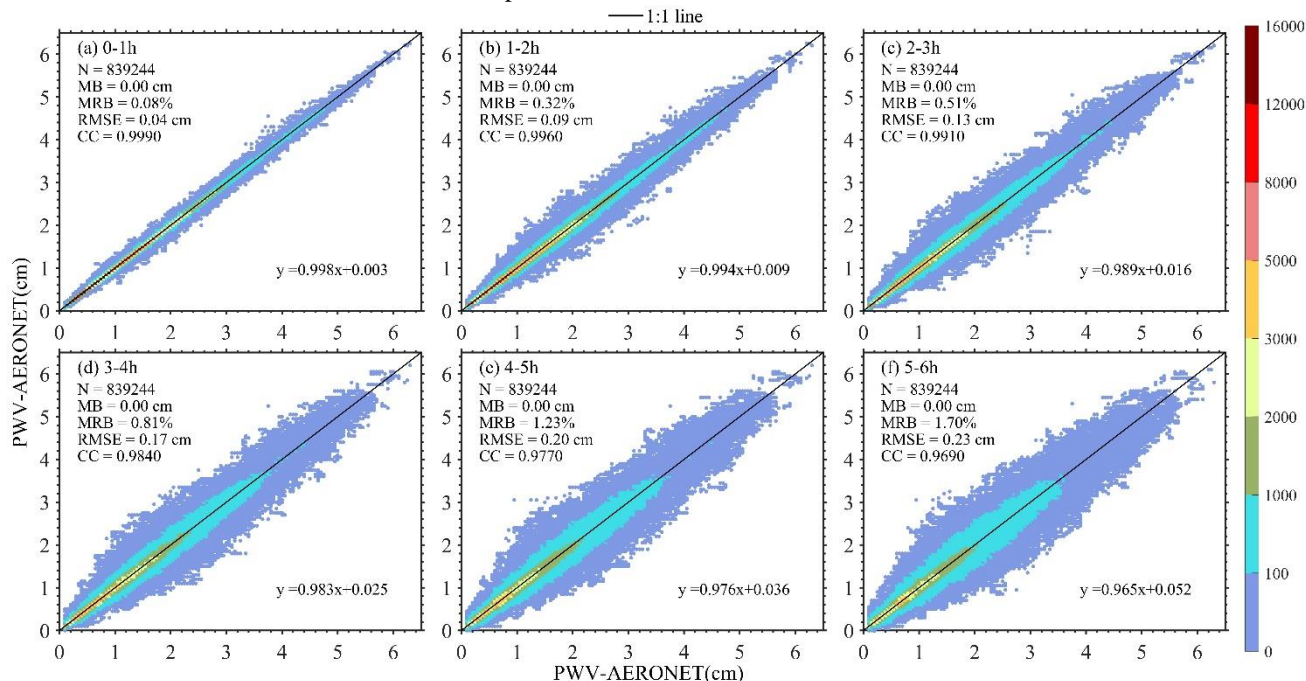
185 where PWV_s is the MERSI-II PWV product, PWV_g is the IGRA PWV product, and N is the total number of match-up.

186 3.2 Collocation strategy

187 As we have mentioned above, FY-3D is operated in a Sun-synchronous orbit with an equator crossing time at 13:40 local
188 time. However, the radiosonde is released at 00:00 UTC and 12:00 UTC and there is a significant temporal discrepancy
189 between satellite and radiosonde at most sites. Furthermore, the distribution of radiosonde sites is sparse over the globe. For
190 the evaluation of PWV from global reanalysis models with a temporal resolution of 6 h, temporal window of ± 3 h and distance
191 of ± 50 km is employed in the comparison with PWV from Maritime Aerosol Network (Pérez-Ramírez et al., 2019).

192 In order to determine the temporal collocation window that can adequately match the satellite observations with the
193 ground-based measurements, the consistencies between the existing AERONET PWV and the temporal averaged AERONET
194 PWV in various temporal discrepancy intervals from 1 h to 6 h with a step of 1 h, that is, 0–1 h, 1–2 h, etc., are analyzed
195 respectively. In processing, only the existing AERONET PWV, which has the matching averaged AERONET PWV in each
196 temporal discrepancy interval, is selected for the determination of the temporal collocation window. Therefore, there is the
197 same number of collocations for all the temporal discrepancy intervals. The results are presented in Figure 1, and as evidenced,
198 there is a good consistency in all situations with the CC larger than 0.969 and the slope larger than 0.965. Although MRB and
199 RMSE become larger with the increasing temporal interval, their values are less than 1.7% and 0.23 cm, respectively. Moreover,

200 it can be observed that the MB values of all comparisons are 0.00 cm, which suggest that the biases are distributed equally
 201 around zero. Therefore, we conclude that the temporal collocation window for PWV evaluation can be set to 6 h.



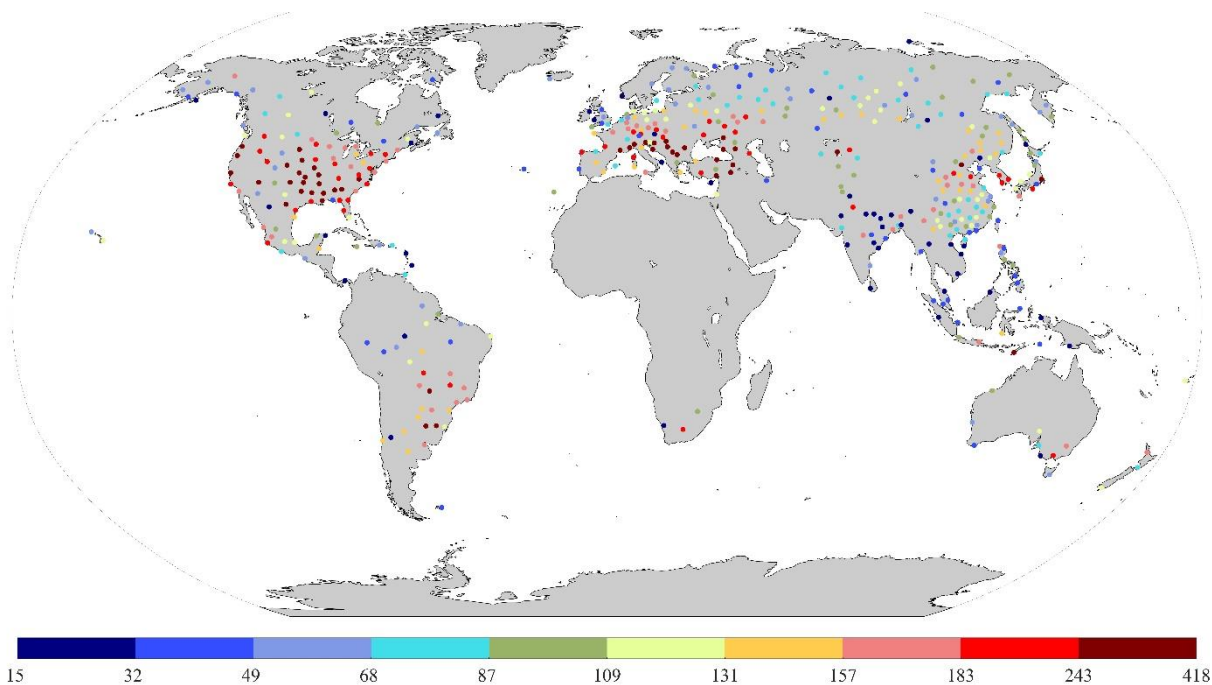
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203 Figure 1 Scatter plots of PWV derived from AERONET in different temporal discrepancy intervals and (a)–(f) present the
 204 temporal discrepancy of 0–1 h, 1–2 h, 2–3 h, 3–4 h, 4–5 h and 5–6 h, respectively. The solid line represents the 1:1 line. The
 205 colorbar depicts the number density of match-ups for each bin of PWV in a 0.01 cm x 0.01 cm grid.

206 For the MERSI-II, the spatial resolution at nadir is 1 km x 1 km for NIR bands, which are used for the retrieval of PWV.
 207 It is not perfectly justifiable that the PWV in an image pixel represents the surrounding averaged PWV (Ichoku et al., 2002),
 208 and during the ascending of balloon, its lateral drift should also be considered. Therefore, the spatial averaged PWV within a
 209 box of 9x9 pixels was calculated. Furthermore, only when all pixels within the box are confidently proved to be clear by the
 210 MERSI-II CLM product, the pixel is marked as reliable and the PWV of the central pixel is replaced by the spatially averaged
 211 PWV value. Otherwise, the pixel is marked as unreliable and will not be selected for the comparison. Unfortunately, there is
 212 no cloud measurement in the radiosonde observation, so the cloud detection method with the relative humidity threshold of
 213 sounding is employed here (Zhang, 2010), and then the cloudless radiosonde PWV dataset is established.

214 In this study, the threshold of the horizontal distance between an IGRA station and the footprint of MERSI-II is set to be
 215 50 km (Qin et al., 2012; Pérez-Ramírez et al., 2019). In processing, all the reliable PWV retrievals derived from MERSI-II
 216 within ±6 h of the radiosonde release time are collected and only the spatial closest PWV retrieval within a 50 km distance
 217 from the IGRA site is selected and matched with the IGRA-derived PWV. Figure 2 illustrates the available sample numbers

218 of radiosonde sites over the globe from 2018–2021, with a total of 462 sites. The sample numbers of all sites vary from 15 to
219 418, and observations are concentrated in the Northern Hemisphere. Around the equator, few samples are obtained due to the
220 high occurrence frequency of clouds and precipitation. Most frequently sampled places are China, Europe, and North America,
221 where IGRA sites are densely distributed, while there are few match-ups in Africa since radiosonde stations associated with
222 IGRA in this region are sparse (Durre et al., 2018).
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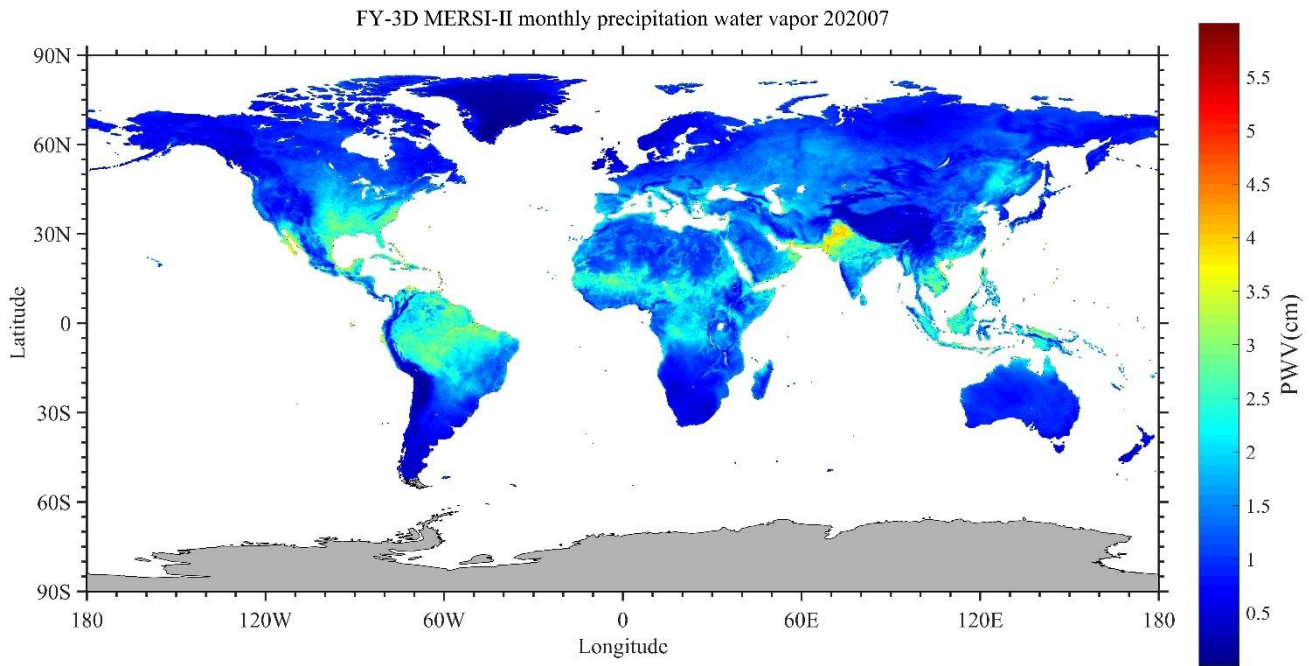
225 Figure 2 Number of matchups between MERSI-II and IGRA PWV observations for each site from 2018–2021.

226 4 Results and Discussion

227 4.1 Global evaluation of PWV derived from MERSI-II

228 Figure 3 illustrates the global monthly averaged PWV obtained from the MERSI-II for the month of July in 2020 under
229 clear sky condition. In general, the averaged PWV derived from MERSI-II shows a decreasing distribution of PWV with
230 increasing latitude, and large PWV values are mostly found in the tropics but rare in high latitudes. High PWV values (> 3.0
231 cm) are mostly detected in the Amazon rainforest of South America, West and South Asia, southeastern China, Southeast
232 Asian islands, and central Africa. The significantly high PWV center in West Asia is mainly contributed by the Indian moon

233 during the summer season. The PWV contents over the Qinghai–Tibet Plateau and Greenland are small because of the high
234 altitude and high latitude, respectively. Furthermore, due to the winter season of July in the Southern Hemisphere, small PWV
235 contents are also presented over the southern parts of South America and Africa, as well as Australia.

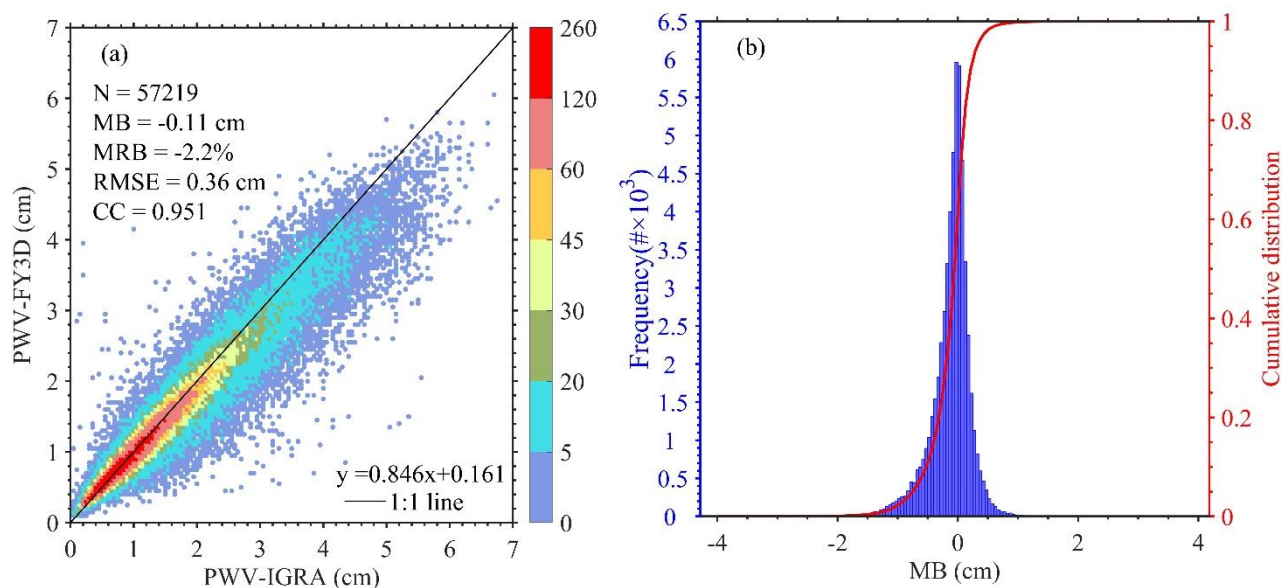


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237 Figure 3 Global averaged Level-3 NIR precipitable water vapor (PWV) image derived from FY3D MERSI-II in July 2020
238 [under clear sky condition](#).

239 Figure 4a shows the scatter plot of PWV derived from MERSI-II against IGRA observations. There are 57,219 match-
240 ups in total and the MERSI-II (IGRA) PWV ranges from 0.11(0.05) cm to 6.07 (6.78) cm, with a high number density between
241 0.20 cm and 2.00 cm. Moreover, it is found that the MERSI-II and IGRA PWV measurements are well correlated with CC of
242 0.951, while the retrieved PWV from MERSI-II is slightly underestimated, with an MB of -0.11 cm and an MRB of -2.2%.
243 Furthermore, the RMSE is 0.36 cm and the statistical biases are slightly larger than those from the evaluation of MODIS over
244 globe by comparing with the observations of AERONET (Martins et al., 2019). It is considered that the satellite has a larger
245 temporal discrepancy with radiosonde than AERONET, which has a high temporal resolution of about 1 min, and from the
246 discussion in section 3.2, a large temporal discrepancy will cause an increasing bias in the evaluation of the PWV product.
247 Figure 4b reveals the distribution of MB between FY-3D MERSI-II and IGRA with a bin width of 0.05 cm, and notably, the
248 MB is concentrated around zero and there is a small flattening towards negative values. Moreover, there are 20.8% of all points

249 within the interval from -0.05 cm to 0.05 cm, and the standard deviation (STD) of MB is 0.34 cm. It can be concluded that
 250 there is high accuracy for the MERSI-II PWV product, as evidenced by the low MB and STD which are similar to those in the
 251 evaluation of ground-based GPS PWV against radiosonde PWV (Wang et al., 2007). Although the MB is mainly distributed
 252 between -1.00 cm and 0.50 cm, it is observed that there are also some points with a large MB value. As presented in figure 4b,
 253 the large MB is mostly with a negative value, and this is mainly due to the different situations observed by the radiosonde and
 254 satellite because of the radiosonde drift and the large temporal discrepancy between MERSI-II and IGRA observations. Due
 255 to the lack of cloud measurement in the radiosonde observation, the PWV from IGRA possibly contains the point in cloudy
 256 conditions, which is proved to have a larger water content than clear conditions (Zhang et al., 2015). For most large positive
 257 MRB, there is a large temporal discrepancy of more than 4 h, and this is recognized as the primary reason for the high positive
 258 MRB.



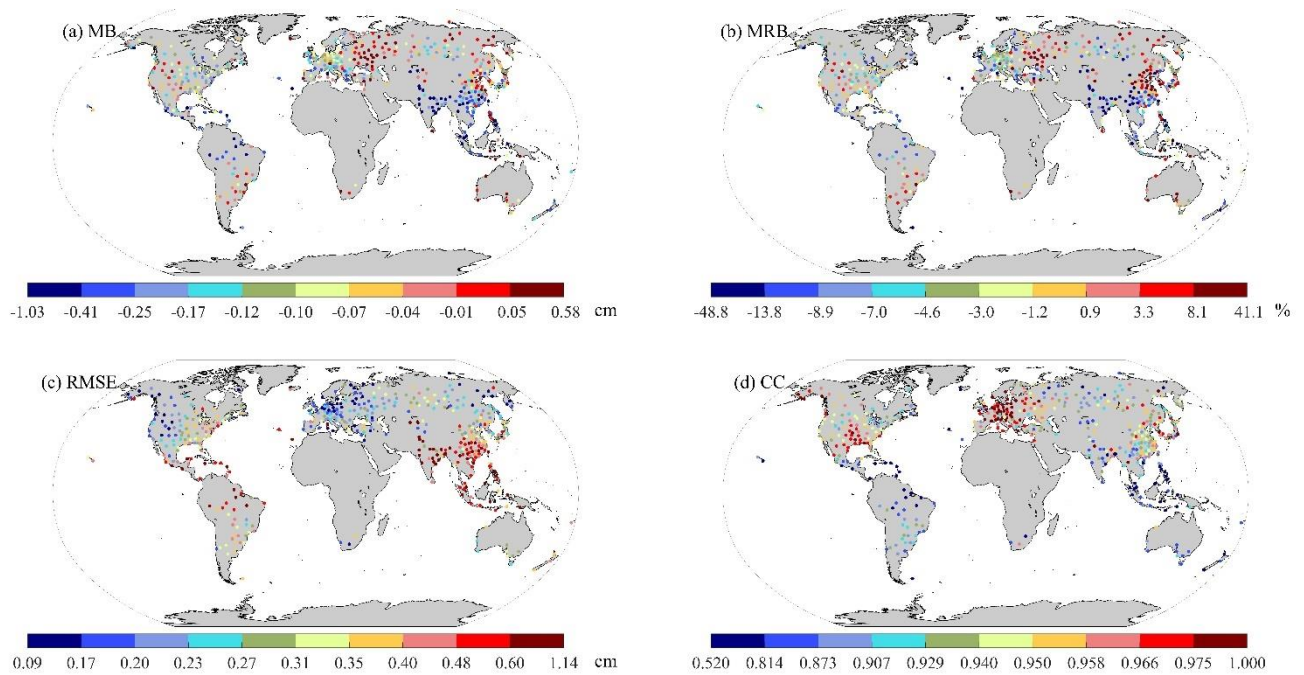
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260 Figure 4 (a) Total density scatterplot of the PWV derived from MERSI-II against that of IGRA for each bin of PWV in a 0.05
 261 cm×0.05 cm grid. The solid line represents the 1:1line. (b) Frequency histogram of -MB between MERSI-II and IGRA PWV
 262 superimposed on a cumulative distribution curve.

263 4.2 Evaluation of MERSI-II PWV product in different locations

264 Figure 5 shows the geographical distributions of PWV comparison statistics of 462 sites between MERSI-II and IGRA
 265 over the globe. In order to equally present the statistical indicators, all sites are separated into approximately ten equal parts,
 266 that is, the site number of each part is ~43. Consequently, the steps of comparison statistics are not equidistant in the

267 presentation. As we can see from the MB distribution in Figure 5a, the MB mostly presents a low value between -0.41 cm and
268 0.05 cm. About 80% of all sites have negative MB values, and this indicates that PWVs derived from MERSI-II are primarily
269 underestimated compared with IGRA PWV values. Most sites with an MB value larger than -0.41 cm are distributed in the
270 west and south of Asia, with a large mean PWV content but a small number of match-ups. Furthermore, attributed to the
271 monsoon climate, there is a large seasonal variation over this region, particularly in the south foothill of the Himalayas (Chen
272 and Liu, 2016). Those sites with overestimated PWV values of MERSI-II are mostly distributed in the surrounding areas of
273 the Black Sea and central South America, and most of them have MB values larger than 0.05 cm. It is also found in the
274 evaluation of the PWV product derived from MODIS onboard Terra that the MB of MERSI-II is slightly smaller (Martins et
275 al., 2019). In general, the distribution of the MRB (Figure 5b) is similar to that of the MB at most sites. However, there are
276 two areas that have slight discrepancies between them. One area is in eastern Russia and northeastern China, where there are
277 some sites with a large MRB value above 3.3%, although with a small MB value ranging from -0.10 cm to 0.05 cm. As we
278 can see from figure 3, there is a low averaged PWV value in this region, and this is the dominant reason for the great-large
279 MRB values but with small MB values over this area. Another area is central South America, where the sites have large MB
280 values and comparatively low MRB values, and this is because of the large mean PWV values in this region. The larger
281 evaluation bias of PWV derived from MODIS and reanalysis products has also been found in the middle of South America,
282 with most sites having the MB and RMSE both larger than 0.40 cm (Lu, 2019; Wang et al., 2020). Figure 4e-5c depicts the
283 distributions of RMSE for all sites and most sites have a small RMSE below 0.48 cm. The large RMSE values are primarily
284 found at low latitudes, mostly in South and Southeast Asia. However, in the east of Europe, there are small RMSE with values
285 below 0.21 cm at most sites. In general, there is a good agreement between MERSI-II and IGRA PWV at most sites with the
286 CC value above 0.873. The highly correlated sites are mainly distributed around the east of Europe and have CC values larger
287 than 0.958, while low CC values smaller than 0.814 are predominantly concentrated around the equator. There are large biases
288 but small CC values over the equator, and that is possibly due to the following: 1) large residual IGRA PWV above 500 hPa
289 (Boukabara et al., 2010); 2) high content and variation of PWV (Chen and Liu, 2016); 3) the covered surface with the
290 reflectance does not linearly correlate with the wavelength (Gao and Kaufman, 2003); 4) a small number of samples. In addition,
291 the temporal discrepancy can also lead to bias because the discrepancy in the equatorial region is slightly larger than in other
292 regions overall. As discussed by Alraddawi et al (2018), for MODIS PWV, there are also noteworthy latitudinal decreases for
293 MB, MRB and RMSE.



294

295 Figure 5 The geographical distributions of PWV comparison statistics between MERSI-II and IGRA. (a) MB; (b) MRB; (c)
 296 RMSE; (d) CC.

297 4.3 The annual performance of MERSI-II PWV product

298 As we mentioned above, PWV presents a notable temporal variation. Therefore, the annual performance of the MERSI-
 299 II PWV product by season was evaluated and the results are given in Table 2. It is noted that the number of samples in each
 300 season is significantly different.

301 Table 2 Seasonal statistics of comparison between PWVs derived from MERSI-II and IGRA in the Northern (Southern)
 302 Hemisphere

Season	N	Slope	MB (cm)	MRB (%)	RMSE (cm)	CC
MAM	11866(1232)	0.833(0.859)	-0.09(-0.10)	-1.8(-1.4)	0.31(0.39)	0.954(0.953)
JJA	15935(2187)	0.807(0.873)	-0.18(-0.05)	-5.3(1.7)	0.41(0.34)	0.931(0.953)
SON	16196(2176)	0.836(0.858)	-0.11(-0.06)	-2.5(0.3)	0.34(0.41)	0.945(0.933)
DJF	6558(1069)	0.852(0.799)	-0.05(-0.17)	3.8(-3.9)	0.31(0.49)	0.944(0.921)

303 The number of match-ups ranges from 6558 (1069) to 16196 (2187) in the Northern (Southern) Hemisphere for all
 304 seasons. In all seasons, the slope values are all less than 0.873, which is the fit-slope in winter in the Southern Hemisphere.
 305 The MERSI-II PWV is underestimated for all seasons, and the MB is less than -0.18 cm. The MB is obviously large in
 306 the warm season, and it is more significant during the summer. With abundant water vapor in summer, clouds are easily
 307 to form, however, thin clouds are difficult to be measured by satellite due to their low optical depth (Solbrig, 2009;

308 Naumann and Kiemle, 2020). Therefore, the higher underestimation of PWV in summer is probably triggered by the
 309 weakened or covered radiation signal under the thin cloud. For MRB, the variation is within a large range, and the largest
 310 MRB is in summer over both the Northern and Southern Hemisphere, with values of -5.3% and -3.9%, respectively. For
 311 the largest MRB during winter in the Northern Hemisphere, this might be related to the points with a small PWV value
 312 but a high positive MB, because we firstly calculate the MRB for each match-up and then average all MRB values in
 313 seasons. In addition, MRB is mostly negative during the warm season, but positive in the cold season. The RMSE in the
 314 Northern Hemisphere is slightly smaller than that in the Southern Hemisphere, where the greatest RMSE value is 0.49
 315 cm in summer. There is a large oceanic coverage in the Southern Hemisphere, with a larger mean PWV than that in the
 316 Northern Hemisphere (Chen and Liu, 2016). Thus, this is a possible reason accounting for large RMSE in the Southern
 317 Hemisphere, considering the increasing bias of the remote sensing PWV with the larger PWV value. Moreover, there is
 318 an improved correlation between PWV derived from MERSI-II and IGRA, and all CC values are larger than 0.921.

319 4.4 Influencing factors on evaluation

320 As mentioned above, there is a higher bias with the larger PWV value, and this is found in the evaluation of MODIS
 321 PWV product (Martins et al., 2019). Furthermore, the impact of the spatial distance between the footprint of the satellite and
 322 the IGRA station on the evaluation of PWV is also considered in the validation of HY-2A CMR PWV (Wu et al., 2020b). [A](#)
 323 [strong solar zenith angle \(SZA\) dependence was also found in the comparisons of MODIS and GOME-2 integrated water](#)
 324 [vapor products with GPS measurements at the Iberian Peninsula \(Roman et al., 2015; Vaquero-Martinez et al., 2017\)](#). In this
 325 section, influencing factors such as the value of IGRA PWV, ~~and~~ the spatial distance between the footprint of the satellite and
 326 the IGRA station ~~and~~ [SZA](#), are all explored in order to quantify their effects on the evaluation of MERSI-II PWV. Table 3
 327 illustrates the evaluation results of the MERSI-II PWV in different intervals of IGRA PWV, ~~and~~ spatial distance ~~and~~ [SZA](#). All
 328 metrics are calculated using all match-ups.

329 Table 3 Statistics of the global evaluation of MERSI-II PWV in different PWV and distance ranges.

	Intervals	N	Slope	MB (cm)	MRB (%)	RMSE (cm)	CC
IGRA PWV (cm)	(0, 1]	15528	0.869	0.04	9.0	0.17	0.754
	(1, 2]	21698	0.878	-0.06	-3.7	0.26	0.709
	(2, 3]	11831	0.842	-0.19	-7.8	0.41	0.556
	(3, 4]	5493	0.867	-0.35	-10.1	0.57	0.483
	(4, 5]	2122	0.664	-0.55	-12.3	0.72	0.358
	>5	547	0.636	-0.94	-17.4	1.10	0.347
Spatial distance (km)	(0, 5]	31216	0.860	-0.08	-0.6	0.33	0.952
	(5, 10]	19972	0.844	-0.14	-3.4	0.38	0.953
	(10, 20]	4800	0.813	-0.19	-6.1	0.42	0.951

	>20	1231	0.792	-0.20	-5.9	0.47	0.933
Solar zenith angle	(0°, 20°]	4197	0.831	-0.23	-6.0	0.50	0.933
	(20°, 30°]	12230	0.823	-0.19	-5.5	0.45	0.935
	(30°, 40°]	13440	0.835	-0.13	-3.2	0.37	0.942
	(40°, 50°]	13140	0.852	-0.08	-1.5	0.31	0.938
	(50°, 60°]	9764	0.829	-0.05	0.3	0.25	0.919
	>60°	4448	0.829	0.01	6.5	0.23	0.869

330 Firstly, the IGRA PWV is binned and compared with the MERSI-II PWV. Most match-ups are located at the IGRA PWV
331 interval ranging from 1.0 to 2.0 cm, and there are not many samples above 5.0 cm. The MB and RMSE gradually increase
332 with the increasing content of PWV, from 0.04 cm and 0.17 cm to -0.94 cm and 1.10 cm, respectively. Moreover, the fit-slope
333 value is generally decreasing with the increasing content of PWV, and it can be concluded that there is an obvious
334 underestimation when the PWV is larger than 5.0 cm. This result is similar to the conclusion drawn by Martins et al. (2019) in
335 the evaluation of the PWV from MODIS, however, the slope is smaller. There is a good agreement between the IGRA PWV
336 and MERSI-II PWV for dry conditions (<1.0 cm), with the highest CC value of 0.754. However, a slightly large discrepancy
337 is observed in wet conditions (> 5.0cm). For MRB, there is a positive value of 9.0% in the 0.0–1.0 cm interval, and this is also
338 caused by the small PWV value. In contrast, all MRB is negative and the value of MRB increases with the content of PWV
339 above 1.0 cm.

340 The results of the MERSI-II PWV and IGRA PWV comparison in different distance intervals are also presented in Table
341 3. Most points are located within the 0–5 km distance interval, and the number of points is 31,216 out of all 57,219 points. The
342 MB increases with the extension of the distance between the IGRA station and the footprint of MERSI-II, and the largest MB
343 is -0.20 cm when the distance is larger than 20 km. For the MRB, a more obvious difference is present within the 0–20 km
344 distance range, as the value increases from -0.6% to -6.1% with the increasing distance. However, there is a slightly smaller
345 MRB when the distance is larger than 20 km, and this probably has a relationship with the small number of samples. The
346 RMSE has a value ranging from 0.33 cm to 0.47 cm and becomes larger with the increasing distance. The large RMSE in the
347 distance above 20 km is mainly caused by the obvious underestimation of MERSI-II PWV at some points. Overall, a good
348 correlation exists between MERSI-II PWV and IGRA PWV with the CC value larger than 0.933.

349 The solar radiation changes with SZA and the performance of MERSI-II PWV product during different SZA intervals is
350 studied in this section. It can be seen from the Table 3 that there are the greatest number of match-ups with SZA ranging from
351 20° to 50°, and the deviations between MERSI- II and IGRA PWV highly depend on SZA. The MB presents the decreasing
352 tendency with the increasing SZA, from -0.23 cm to 0.01 cm. For the MRB, there is large value both in small (<20°) and large
353 (>60°) SZA conditions, with the values of -6.0% and 6.5%, respectively. In small (<20°) SZA condition, a large RMSE is
354 found with the value of 0.50 cm, and it is probable due to the large STD of MB, which was also presented by Roman et al.
355 (2015) in the comparison between GOME-2 PWV and radiosonde/radiosounding measurements. However, the slope value and

356 CC value are similar at each SZA interval.

357 **5 Summary and Conclusions**

358 In this paper, we evaluated the global clear sky PWV product derived from FY-3D/MERSI-II by comparing with the
359 PWV from 462 IGRA stations, with 57,219 match-ups during the period from September 2018 to June 2021. The monthly
360 averaged PWV derived from MERSI-II shows a distribution of decreasing values with an increasing latitude.

361 Overall, PWVs derived from MERSI-II and IGRA are in good agreement with the CC value of 0.951. However, there is
362 a slight underestimation for the FY-3D/MERSI-II PWV, and the values of MB and MRB are -0.11 cm and -2.2%, respectively,
363 while the RMSE is 0.36 cm. The histogram of MB indicates that MB value approaches zero and mostly distributes between -
364 1.00 cm and 0.50 cm with a left-skewed distribution pattern.

365 For all sites, the MB value is low and most sites have a value between -0.41 cm and 0.05 cm. In the west and south of
366 Asia, the MERSI-II PWV is obviously underestimated with an MB value larger than -0.41 cm. However, the overestimated
367 PWV are mostly distributed in the surrounding areas of the Black Sea and central South America. Large MRB values are
368 mostly located in eastern Russia, northeastern China, and central South America. Most sites have a small RMSE below 0.48
369 cm, and CC values above 0.873. Lastly, it is observed that there are large MB and RMSE values while there are small CC
370 values around the equator, especially in South and Southeast Asia.

371 The MERSI-II PWV is in good agreement with the PWV obtained from IGRA with all CC values larger than 0.921. There
372 is a slight underestimation of MERSI-II PWV for all seasons with an MB value below -0.18 cm, and it is significant in the
373 summer. In addition, the MRB and RMSE also have the largest magnitude in summer. The underestimation of PWV in summer
374 is probably due to the presence of thin clouds, which weaken the radiation signal observed by the satellite. We found the MRB
375 with a positive value in the winter, and this is mainly due to the low PWV then. For RMSE, there is a larger value in the
376 Southern Hemisphere and the greatest RMSE value is 0.49 cm in summer.

377 In addition, the influencing factors on the evaluation are also discussed. First of all, there is an obvious effect of binned
378 IGRA PWV on the evaluation, and in general, the MB and RMSE are both increasing with the IGRA PWV. In the dry condition
379 (<1.0 cm), there is a positive MRB value of 9.0%, and this is also mainly due to the low PWV value. Nevertheless, the MRB
380 is all negative and increases with an IGRA PWV above 2.0 cm. Subsequently, the evaluations within different distance intervals
381 are presented in order to reveal the effect of distance between the footprint of the satellite and the IGRA site location. The MB
382 varies positively with the increasing distance, and the largest MB is -0.20 cm within the distance above 20 km. The MRB
383 increases from -0.6% to -6.1% with the distance increases from 0 to 20 km. The RMSE also increases with the distance
384 increased and the large RMSE is mainly caused by the obvious underestimation of MERSI-II PWV at some points with the
385 spatial distance larger than 20 km. The MB decreases with the increasing SZA, from -0.23 cm to 0.01 cm. There are large

386 MRB value both in small (<20°) and large (>60°) SZA conditions, and the largest RMSE appears in small (<20°) SZA
387 condition.

388 The global evaluation of the MERSI-II PWV product can explore a wide variety of applications of this product, and the
389 analysis of the influencing factors on the evaluation can be helpful for improving the PWV retrieval algorithm. Although we
390 have partially explained the underestimation of the PWV from MERSI-II, other influencing factors, such as the solar zenith
391 angle, the precision of the transmittance calculation and the uncertainty of the radiation signals should be studied in the future.
392 Furthermore, how to quantitate the influence of aerosols (e.g., dust, haze) and thin cirrus clouds on the PWV retrieval is also
393 a key problem that should be explored in the application of the PWV product.

394 **Data availability**

395 The MERSI-II PWV product is available from <http://satellite.nsmc.org.cn/PortalSite/Data/Satellite.aspx>, the IGRA data is
396 available from <ftp://ftp.ncdc.noaa.gov/pub/data/igra>, and the global AERONET data are provided at
397 <https://aeronet.gsfc.nasa.gov>. The altitude data set is provided by Geospatial Data Cloud site, Computer Network Information
398 Center, Chinese Academy of Sciences at <http://www.gscloud.cn>. The processed data are available from Zenodo
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400 **Author contributions**

401 Conceptualization, ZWG and WL; data curation, WL, YY and HQX; formal analysis, ZWG, YY and XGR; writing-original
402 draft preparation, ZWG; writing-review and editing, ZWG and WL; supervision, XGR and HXQ; funding acquisition, XGR
403 and CCG. All authors have reviewed and agreed on the final version of the manuscript.

404 **Competing interests**

405 The authors declare that they have no conflict of interest.

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