Impact of 3D Cloud Structures on the Atmospheric Trace Gas Products from UV-VIS Sounders - Part III: bias estimate using synthetic and observational data

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Abstract. Three-dimensional (3D) cloud structures may impact atmospheric trace gas products from ultraviolet-visible (UV-VIS) sounders. We used synthetic and observational data to identify and quantify possible cloud-related bias in NO_2 tropospheric vertical column densities (TVCD). The synthetic data were based on high-resolution large eddy simulations which were input to a 3D radiative transfer model. The simulated visible spectra for low-earth orbiting and geostationary geometries

5 were analysed with standard retrieval methods and cloud correction schemes that are employed in operational NO_2 satellite products. For the observational data the NO_2 products from the TROPOspheric Monitoring Instrument (TROPOMI) were used while the Visible Infrared Imaging Radiometer Suite (VIIRS) provided high spatial resolution cloud and radiance data. Cloud shadow fraction, cloud top height, cloud optical depth, solar zenith and viewing angles, were identified as the metrics being the most important in identifying 3D cloud impacts on NO_2 TVCD retrievals. For a solar zenith angle less than about $40°$

10 the synthetic data show that the NO_2 TVCD bias is typically below 10%. For larger solar zenith angles both synthetic and observational data often show NO_2 TVCD bias on the order of tens of %. Specifically, for clearly identified cloud shadow bands in the observational data, the NO_2 TVCD appears low-biased when the cloud shadow fraction > 0.0 compared to when the cloud shadow fraction is zero. For solar zenith angles between 50-60°, about 16% of TROPOMI pixels with high quality value NO_2 TVCD retrievals, were found to be impacted by cloud effects larger than 20%.

15 ## 1 Introduction

Operational retrievals of tropospheric trace gases from space–borne spectrometers are based on the use of 1D radiative transfer models. To minimize cloud effects, generally only partially cloudy pixels are analysed using simplified cloud contamination treatments based on radiometric cloud fraction estimates (e.g. Grzegorski et al., 2006; Stammes et al., 2008) and photon path length corrections based on oxygen collision pair (O_2–O_2) (Acarreta et al., 2004; Veefkind et al., 2016) or O_2–A absorption

20 band measurements (for example the FRESCO and OCRA/ROCINN algorithms, see Koelemeijer et al., 2001; Wang et al., 2008; Loyola et al., 2018; Liu et al., 2021). In reality however the impact of clouds can be much more complex, involving

unresolved sub-pixel clouds, scattering of clouds in neighbouring pixels and cloud shadow effects. In a model study Merrelli et al. (2015) showed that 3D radiation scattering from unresolved boundary layer clouds may give significant biases in Orbiting Carbon Observatory-2 (OCO-2) retrievals of CO_2 concentration. Massie et al. (2017, 2020) provided observational evidence of 3D cloud effects in OCO-2 CO_2 retrievals and found them consistent with 3D radiative transfer simulations.

5   In general, space-borne measurements of trace gases may be cloud contaminated and the presence of clouds may result in both positive and negative biases. It is thus vital to quantify the impact of clouds on trace gas retrievals and, if possible, envisage correction methods. To exclude cloud contaminated pixels from analysis is not a viable option as this, for example, may give bias in long-term averages, as shown by Geddes et al. (2012) for NO_2. Furthermore, the loss in data coverage when excluding partially cloudy scenes may become critical, especially for low/medium resolution sensors and for regions where the

10 probability of cloud occurence is high (e.g. Germany and other western European countries).

     This paper is one of a series of three papers discussing the impact of 3D cloud structures on the atmospheric trace gas products from ultraviolet-visible (UV-VIS) sounders. In this paper, the bias due to 3D clouds is quantified using both synthetic and observational data. The paper by (Emde et al., 2021) describes the synthetic data which is based on 3D radiative transfer model simulations utilizing realistic 3D clouds as input and is designed for validation of remote sensing trace gas retrievals.

15 Finally, Yu et al. (2021) discuss trace gas retrieval and mitigation strategies in the presence of 3D cloud structures.

     Specifically, we studied NO_2, as it is an important measure of air quality and a key tropospheric trace gas measured by the atmospheric Sentinels (sentinel.esa.int). The synthetic data, which ignore stratospheric NO_2, were used to identify the cloud situations that give bias in tropospheric NO_2 retrievals. The observational satellite data from TROPOMI were used to investigate the presence of 3D cloud radiative transfer effects in real data. While focus is on tropospheric NO_2, the results are

20 expected to be valid for other UV-VIS derived trace gas products.

     From the HD(CP)2 (hdcp2.eu) project large eddy simulations (LES) based on the ICOsahedral Non–hydrostatic atmosphere model (ICON Dipankar et al., 2015; Zängl et al., 2015) are available for a region including Germany, the Netherlands and parts of other surrounding countries. This unique synthetic data set provide realistic input for 3D radiative transfer modelling. Hence, we adopted the study region to be the area covered by the LES. This region also covers part of the footprints of

25 the upcoming Sentinel-4 and Sentinel-5 missions. Using the LES output, satellite radiances were simulated. To include 3D radiative effects we used the 3D MYSTIC Monte Carlo radiative transfer model (Mayer, 2009; Emde et al., 2011) to generate synthetic observations (Emde et al., 2021). From the simulated spectra the slant NO_2 column amounts were obtained using the QDOAS retrieval algorithm (Blond et al., 2007; De Smedt et al., 2008; Yu et al., 2021). The slant column densities were converted to vertical column densities using layer air mass factors (AMF). The AMFs included state-of-the-art independent

30 pixel approximation cloud correction schemes (using the O_2-O_2 or O_2-A band) and were calculated with the VLIDORT 1D radiative transport model (Spurr, 2006). The retrieved NO_2 amount using standard 1D algorithms, was compared to the amount input to the 3D radiative transfer simulations and possible 3D radiative effects were identified and quantified.

     The observational satellite data were used to investigate the presence of 3D cloud radiative transfer effects in real data. NO_2 tropospheric vertical column density data were taken from the TROPOMI/S5P operational L2 tropospheric column product.

35 S5P does not include an imager for cloud information. However, it flies in tandem with the Suomi National Polar-orbiting

Partnership (S-NPP) satellite. The S-NPP payload includes the Visible Infrared Imaging Radiometer Suite (VIIRS) instrument which may be used as an imager for TROPOMI. We utilized data from both sensors.

The paper first discusses the synthetic and observational data sets, section 2. This is followed by a description of the various metrics used to identify cloud impacts, section 3. In section 4 the results are presented and discussed before the paper ends by

5 some concluding remarks and an outlook, section 5.

2 Data

The study area covers approximately Germany, the Netherlands and parts of other surrounding countries. For this region we have generated synthetic satellite data using high resolution LES data as input for 3D radiative transfer simulations.

2.1 Synthetic satellite data

10 The synthetic satellite data comprise results from three sources: 1) high-resolution LES cloud data; 2) 3D radiative transfer modelling of satellite radiances with the LES cloud data as input; and 3) NO_2 retrieval using the synthetic satellite radiances.

Within the HD(CP)2 project LES were made with the ICOsahedral Non–hydrostatic atmosphere model (ICON; Dipankar et al., 2015; Zängl et al., 2015) which simulated realistic liquid and ice clouds. The results have a spatial resolution of approximately 1.2×1.2 km^2 for a region including Germany, the Netherlands and parts of other surrounding countries and was

15 validated against ground and satellite–based observational data by Heinze et al. (2017). Several weeks of simulations are available including all kinds of weather situations in Europe. We, however, utilize only one from 29 July 2014, 12:00 UTC, due to the computational burden of the radiative transfer simulations. The simulated scene includes all cloud types that are typical for Europe, such as shallow cumulus, cirrus, stratus, and also convective clouds. While focus is on Europe, the results/methods are expected to be general and thus applicable elsewhere.

20 The LES results were input to the 3D MYSTIC Monte Carlo radiative transfer model (Mayer, 2009; Emde et al., 2011) run within the libRadtran package (Mayer and Kylling, 2005; Emde et al., 2016) to generate synthetic observation spectra in the visible spectral range from 400–500 nm and in the O_2-A band region from 755–775 nm (for further details see: Emde et al., 2021; Yu et al., 2021). The spatial resolution of the simulated sensor was set to approximately 7×7 km, corresponding to 98×104 pixels for the full LES domain. Note, that each simulated sensor pixel includes 36 cloud pixels, hence the simulations include

25 sub-pixel cloud inhomogeneity. The synthetic spectra were input to a NO_2 retrieval algorithm which included two steps: first a differential optical absorption spectroscopy (DOAS) fit was performed using the QDOAS retrieval algorithm (Blond et al., 2007; De Smedt et al., 2008) to get the NO_2 slant column densities; second the slant column densities were converted to vertical column densities using layer air mass factors based on the VLIDORT 1D radiative transport model (Spurr, 2006). The fitting window was between 425 and 495 nm, similar to the one used by Richter et al. (2011). The air mass factor was calculated

30 at the middle of the fitting window, that is 460 nm. Cloud corrections were made using both O_2-O_2 and O_2-A band based methods. An example of retrieved NO_2 TVCD for a synthetic case is shown in Fig. 1a. Note that the "true" NO_2 is constant over the scene with column density of 1.6×10^{16} molec/cm^2 corresponding to an European tropospheric polluted NO_2 profile
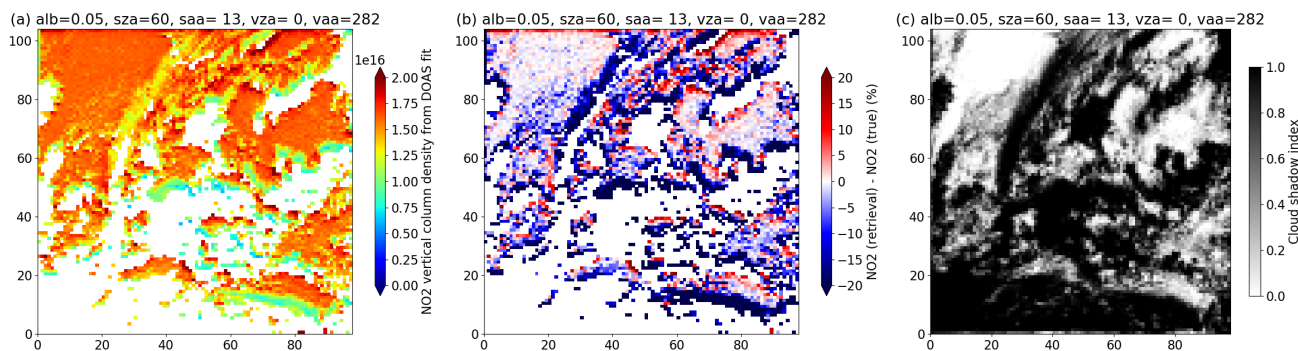
Figure 1. (a) The retrieved NO$_2$ TVCD for the low-earth orbit geometry case with albedo=0.05, solar zenith angle=60°, solar azimuth angle=13°, satellite viewing angle=0°, and satellite azimuth angle=282° (identical for all pixels). White pixels are cloudy regions for which the retrieval was not performed. (b) The difference between the retrieved and the true NO$_2$ TVCD. (c) The cloud shadow index, see section 3.3 for details. The units on the axes are in pixel numbers.

from Levelt et al. (2009). Thus any differences between the retrieved and "true" NO$_2$ TVCDs are due to the presence of clouds, Fig. 1b. The NO$_2$ retrieval is further discussed by Yu et al. (2021).

While we discuss our synthetic results in connection with observational results from satellites in low-earth-orbit (LEO), simulations were also made for a geostationary orbit (GEO). Azimuth and zenith viewing and solar angles were chosen to resemble geometries for the study region when viewed by the TROPOspheric Monitoring Instrument (TROPOMI, Veefkind et al., 2012) and the future Ultra-violet Visible Near-infrared (UVN, https://sentinel.esa.int/web/sentinel/missions/sentinel-4) instrument to be in geostationary orbit. In total 15 and 36 combinations of viewing and solar angles were simulated for the GEO-case and LEO-case, respectively (Emde et al., 2021). The surface was assumed to be snow free and with constant albedo to simplify the interpretation of the results. In the visible (400-500 nm) simulations were made with albedos of 0,0.05 and 0.2, while in the O$_2$-A band region an additional albedo of 0.5 was included to account for the potentially larger albedo in this part of the spectrum. Combing the sun-sensor geometries and visible albedo values our simulated data set thus include a total of 45 GEO-cases and 108 LEO-cases.

2.2 Observational satellite data

Both observational satellite spectrometer and imager data were utilized. NO$_2$ TVCD data were taken from the TROPOMI/S5P operational L2 NO$_2$ tropospheric column product. TROPOMI/S5P was launched 2017-10-13 and L2 NO$_2$ data are available from 2018-06-28. It passes the equator at 1330 local time. Imager data were taken from VIIRS and we used both L1b and L2 data. The L1b reflectances were used for RGB plots and to produce various metrics, see section 3. The L2 data include various cloud products of which we used the cloud mask, cloud shadow, cloud optical thickness and cloud top height products.
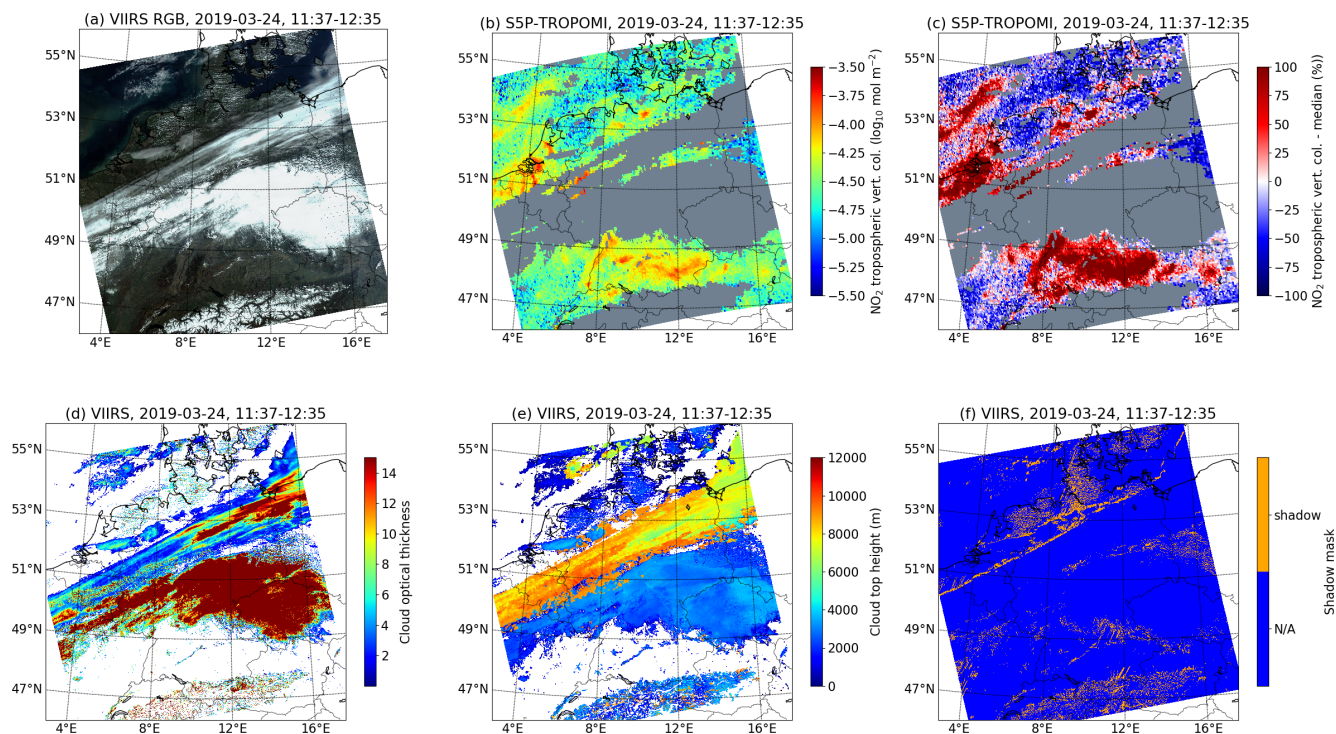
Figure 2. (a) RGB composite of VIIRS bands M3, M4 and M5 (centred at 0.488, 0.555 and 0.672 μm). (b) The tropospheric NO_2 vertical column density from TROPOMI. Only pixels with data quality value > 0.95 are displayed. (c) The percentage difference of the NO_2 column from the median over the study area. (d) The VIIRS cloud optical thickness. (e) The VIIRS cloud top height. (f) The VIIRS cloud shadow mask.

The TROPOMI data includes the latitude and longitude of the corners of each TROPOMI pixel. This information was used to identify VIIRS pixels within each TROPOMI pixel.

Data within the study region roughly covering Germany and parts of neighbouring countries were collected for the periods 28 June 2018 - 15 October 2018 and 1 March 2019 - 30 June 2019 when the sun was sufficiently high on the horizon to avoid 5 problems with NO_2 retrievals for low sun (solar zenith angle $< 60°$). An example of data from near simultaneous TROPOMI and VIIRS overpasses of Germany is shown in Fig. 2.

3 Methods

Several metrics were calculated to quantify cloud features and their possible connection with NO_2-biases due to 3D cloud effects.

3.1 Cloud geometric and radiance fractions

Cloud fractions may be defined in several ways. We calculate the geometric cloud fraction, CF_g, the radiometric cloud fraction, CF_r, and the weighted radiometric cloud fraction, CF_w.

The geometric cloud fraction for a TROPOMI pixel is defined as

$$CF_g = \sum CM_i / N, \tag{1}$$

where the sum is over all N VIIRS pixels within the TROPOMI pixel and CM_i is the VIIRS cloud mask where $CM_i = 1$ for pixels identified as cloudy and $CM_i = 0$ otherwise.

The radiometric cloud fraction is the fraction of measured radiance reflected from clouds in a pixel (see also Grzegorski et al., 2006)

$$CF_r = \begin{cases} 0 & R \leq R_s \\ \frac{R - R_s}{R_c - R_s} & R_s < R < R_c \\ 1 & R \geq R_c \end{cases} \tag{2}$$

Here R is the observed reflectance, R_s is the reflectance for a cloudless sky and R_c the reflectance for an opaque cloud. For the O_2-O_2 cloud correction, CF_r is calculated based on the reflectance at 460nm which is in the middle of the DOAS fitting window for NO_2. For the FRESCO algorithm, CF_r is determined by the reflectance in the 758-759nm window band. Further details are described by Yu et al. (2021). We also define an average radiometric cloud fraction CF_r^{VIIRS} using the average of CF_r calculated for each VIIRS M3 band pixel, centred at 0.488 μ, within a TROPOMI pixel:

$$CF_r^{VIIRS} = \sum CF_{r,i} / N, \tag{3}$$

where the sum is over all N VIIRS pixels within the TROPOMI pixel.

Finally, the weighted radiometric cloud fraction is defined as (Yu et al., 2021)

$$CF_w = \frac{CF_r R_c}{CF_r R_c + (1 - CF_r) R_s} \tag{4}$$

3.2 Cloud mask

No cloud mask was available for the LES based synthetic data. In principle this may be calculated from the cloud optical thickness assuming some treshold. However, we chose to calculate a cloud mask similar to how it is done for several operational satellite products. We thus assume that pixels with reflectance larger than some threshold, $R > R_t$, are cloudy. The cloud mask was calculated from high spatial resolution radiance simulations at 0.55 μm of the synthetic cases described by Emde et al. (2021). A threshold of $R_t = 0.25$, similar to Heinze et al. (2017), was adopted. It is noted that Barker et al. (2017) used $R_t = 0.15$ for the GOES-13 0.65 μm band. Yang and Di Girolamo (2008) has shown that there may be overlap between clear and cloudy pixels and hence some mis-classification is unavoidable. Our slightly higher value potentially includes more cloud contaminated pixels, but also reflects that we are mainly working over land where surface albedo is higher than over ocean.

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3.3 Cloud shadow fraction and cloud shadow index

The VIIRS cloud shadow mask algorithm is geometry-based and described by Hutchison et al. (2009). They compared the spectral procedure used in the MODIS MOD35 product with the geometry-based approach and found the latter to be far superior. The cloud shadow fraction, CSF, for a TROPOMI pixel is defined as

5
$$CSF = \sum CSM_i/N, \tag{5}$$

where the sum is over all N VIIRS pixels within the TROPOMI pixel and CSM_i is the VIIRS cloud shadow mask where $CSM_i = 1$ for pixels identified as cloud shadow and $CSM_i = 0$ otherwise.

For the LES we define the cloud shadow index (CSI) as a surrogate for a cloud shadow product as follows:

$$CSI = E_0 - E_{\theta,\phi}/\cos(\theta) \tag{6}$$

10   Here $E_0 = 1$ is the direct transmittance at the surface for an atmosphere with no molecular absorption nor clouds or aerosol. The direct transmittance at the surface for a solar zenith angle of θ and solar azimuth angle ϕ, $E_{\theta,\phi}$, was calculcated with the MYSTIC model. It also does not include molecular absorption nor aerosol, however, it includes clouds, which are fully accounted for in 3D geometry and with the 3D cloud fields in higher spatial resolution than the TROPOMI pixel size. The CSI is zero for pixels with no cloud shadow and 1 if a pixel is fully in the cloud shadow. The CSI clearly and unambigously 15   identify cloud shadow regions as it is based solely on sun, cloud and surface geometry. An example of the CSI is given in Fig. 1c.

3.4 H-metric

For a TROPOMI pixel the H-metric is the standard deviation of VIIRS band reflectances divided by the mean of the VIIRS reflectances within the TROPOMI pixel. It is an estimate of the variation of radiance within the TROPOMI pixel and has been 20   used earlier by for example Massie et al. (2017) as metric for the impact of 3D clouds on CO_2 retrievals. In the presence of sub-pixel clouds the H-metric is expected to increase as the horizontal cloud inhomogeneity increase. In the absence of clouds the H-metric is expected to be small. However, for cloud free pixels with large variations in surface albedo the H-metric may still be large, and conversely, for a completely cloudy pixel the H-metric may be small. Thus, it may not be used without care to unambiguously identify cloud inhomogeneity. From the VIIRS data we calculate the H-metric for the M3, M4 and M5 bands 25   centred at 0.488, 0.555 and 0.672 μm respectively.

3.5 Absorbing aerosol index

Clouds have an effect on the UV absorbing aerosol index (AAI). It is thus of interest to investigate possible relationships between the AAI and the cloud shadow fraction and the NO_2 TVCD. We use the TROPOMI AAI product where the AAI has "positive values of the reflectance residue between an absorbing aerosol loaded atmosphere and a clear atmosphere. Negative 30   values are associated with an atmosphere that contains more scattering particles than a clear atmosphere" (Kooreman et al., 2020; Penning de Vries et al., 2009).
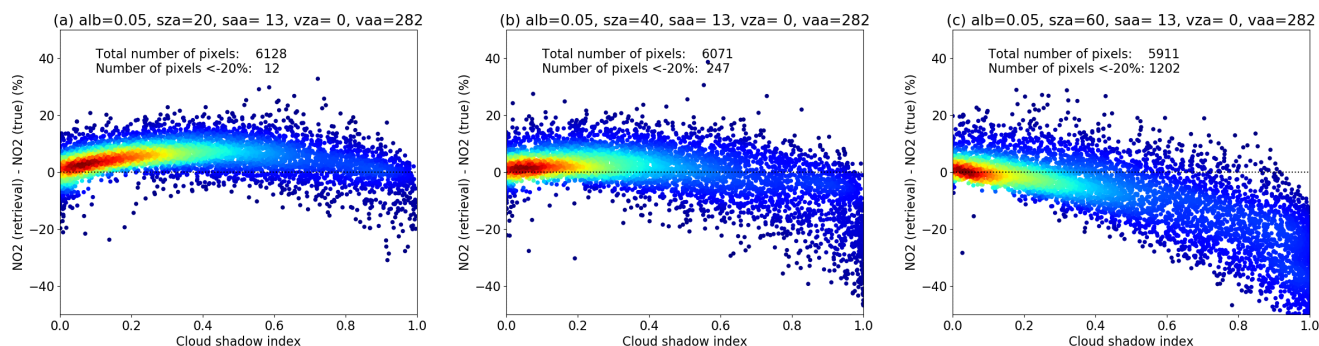
Figure 3. The difference between the retrieved and "true" NO_2 TVCDs versus the cloud shadow index for three low earth-orbiting-geometry cases. Results are shown for solar zenith angles of (a) 20, (b) 40 and (c) $60°$, albedo=0.05, solar azimuth angle=$13°$, satellite azimuth angle=$0°$, and satellite viewing angle=$282°$. The data in (c) are for the case presented in Fig. 1.

4 Results

For both the synthetic and observational data the aim is to compare the NO_2 TVCD for cloud affected cases with the "true" NO_2 TVCD and identify potential biases. This is straightforward for the synthetic data as the "true" NO_2 TVCD is known. For the observational data the "true" NO_2 TVCD is in general not known. One approach to estimate the "true" NO_2 TVCD from

5 the observational data is discussed in section 4.2, which also include the analysis of the observational data.

4.1 Synthetic satellite data

For the synthetic data the fully cloudy pixels were not simulated due to the computational burden, see Emde et al. (2021) for details. To further filter for the presence of clouds the NO_2 TVCDs were retrieved for pixels where the radiometric cloud fraction $CF_r < 0.3$ for all LEO and GEO cases. The retrieval tries to correct for the presence of remaining clouds by including

10 a standard photon path length correction based on absorption by the oxygen collision pair O_2-O_2 or the O_2-A band as described by Yu et al. (2021). An example of the retrieved NO_2 TVCD and the bias is shown in Figs. 1a and b, while the cloud shadow index is shown in Fig. 1c.

For three LEO cases the difference between the retrieved and "true" NO_2 TVCDs versus the cloud shadow index is shown in Fig. 3. The cloud shadow impact is seen to increase as the solar zenith angle increases with the number of pixels with NO_2

15 TVCD differences $< -20\%$ being $0.1\%, 4.\%$ and 20.3% for solar zenith angles of $20°$ (Fig. 3a), $40°$ (Fig. 3b) and $60°$ (Fig. 3c), respectively. As the solar zenith angle increases a linear relationship appears between the NO_2 TVCD difference and the CSI.

For the cloud correction of AMF the weighted radiometric cloud fraction, CF_w, is used as described by Yu et al. (2021). The effect of increasing CF_w is shown in Fig. 4 where the NO_2 AMF bias (1D AMF - 3D AMF) is plotted as a function of CF_w for all GEO and LEO geometries. The retrieval bias is binned in CF_w bins with 2% steps which gives more than 5000 pixels

20 in each bin when the $CF_w < 75\%$. For both geometries the NO_2 AMF bias is high (median of 20%) and with a large range
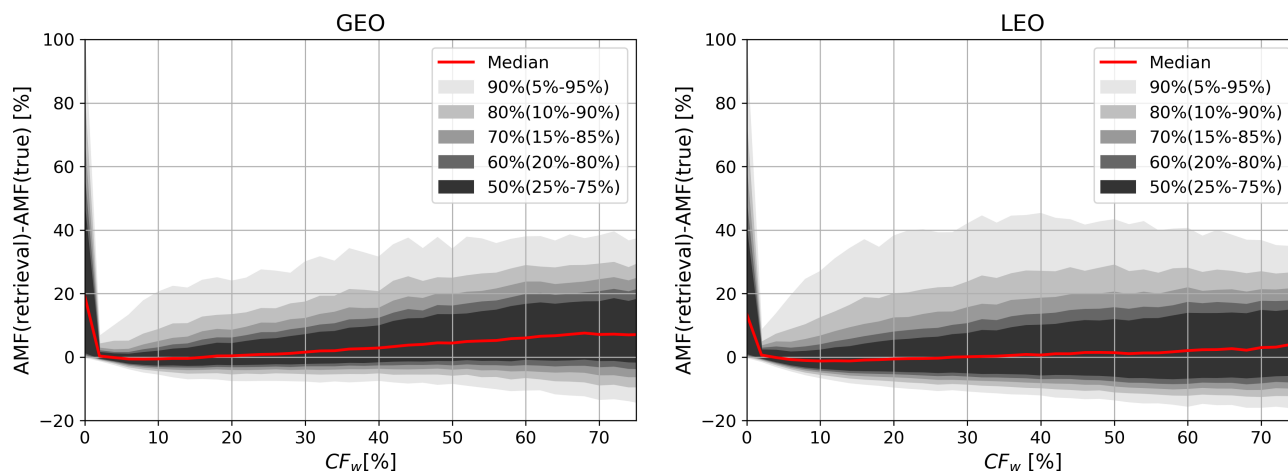
Figure 4. Distribution of the NO$_2$ AMF bias as a function of the weighted radiometric cloud fraction, CF_w, for geostationary and low earth orbiting geometries. Results are shown for the retrieval using the O$_2$-O$_2$ cloud correction. Results are similar for the O$_2$-A band based cloud correction.

(0-65%) for $CF_w < 1\%$. The bias decrease to 0% when the CF_w is between 1-3%. The bias range increases slightly with inreasing CF_w: about 75% of the pixels have a positive bias for large CF_w for GEO geometry, while there are comparatively more pixels with a negative bias for LEO geometry. The solar zenith angle is the same for the LEO and GEO cases. However, the solar azimuth angle and the sensor viewing zenith and azimuth angles are different, thus giving different sensitivity to cloud

5   shadows for LEO and GEO geometries. This is discussed in more detail below. More than 15% of the pixels have a bias larger than 20% for $CF_w > 50\%$ for both GEO and LEO geometries. As discussed by Yu et al. (2021), the NO$_2$ retrieval error due to the use of a simple cloud correction scheme is generally less than 20% for nearly cloud-free cases ($CF_w < 50\%$). Therefore, pixels with a NO$_2$ retrieval bias >20% are likely to be affected by cloud shadow effects. Figure 5 shows the number of pixels that satisfy these conditions for all the LEO and GEO geometries. For the geostationary geometry, the number of pixels with

10   retrieval bias > 20% is up to about 1000 or 10.6% (out of 9400 pixels for a single case). This number increases with surface albedo and SZA, the difference between surface albedos of 0.05 and 0.2 is about 150 to 400 pixels. This number also strongly depends on the solar azimuth angle (SAA), and the difference between SAAs of 315° and 45° is more than 25%. For low earth orbiting geometry, the number of pixels with bias > 20% increases with surface albedo and SZA as well, and it is up to 1600 (17%) for high surface albedo of 0.2 and SZA of 60°.

15   To identify the localisation of the pixels with NO$_2$ AMF bias > 20%, maps of the number of cases with NO$_2$ AMF bias > 20% were made as shown in Fig. 6. Generally, the majority of pixels with large NO$_2$ AMF bias are found in cloud free regions close to cloud edges. Figure 7 shows maps of the maximum retrieval bias for each pixel. As above, cloudy data ($CF_w > 50\%$) are excluded from the analysis. The maximum bias is usually less than 60% over the northwest region (x=30-50, y=70-100) where there is an ice cloud with small optical thickness (< 10). In the center of the map, the bias is often higher than 100%,
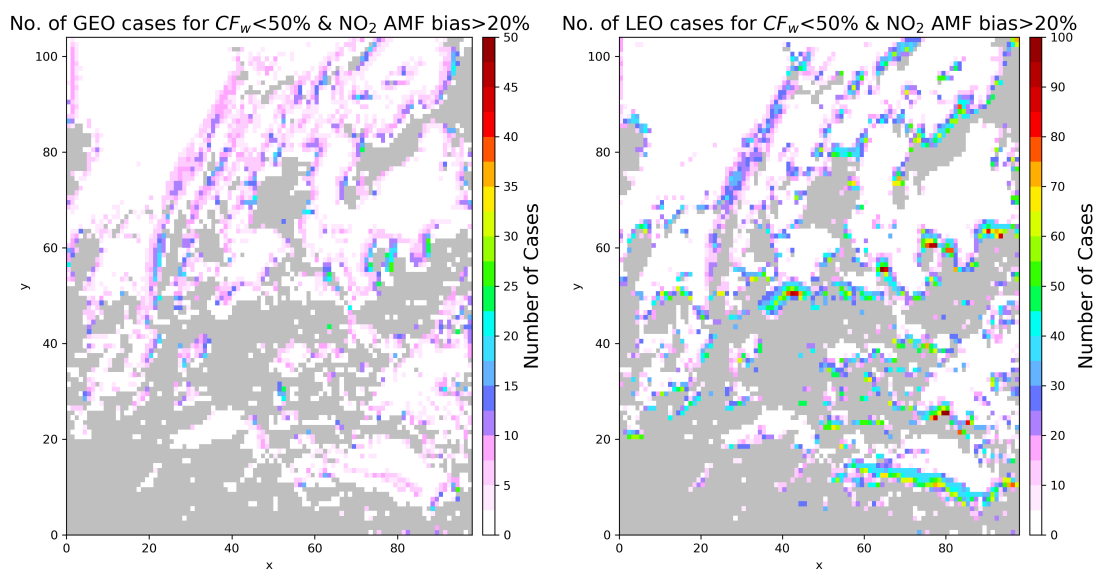
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Figure 5. Number of pixels with weighted radiometric cloud fraction $CF_w < 50\%$ and NO_2 AMF bias $> 20\%$ for geostationary (top) and low earth orbiting (bottom) geometries. The solar zenith angle (SZA), solar azimuth angle (SAA), viewing zenith angle (VZA) and viewing azimuth angle (VAA) for each case is given on the x-axes labels. The total number of pixels for a single case is 9400.

Figure 6. The number of cases with $CF_w < 50\%$ and NO_2 AMF bias $> 20\%$ for each pixel for GEO (left) and LEO geometries (right). The x- and y-axes represent pixel number in the longitude/latitude direction. Grey shaded pixels indicate cloudy pixels.
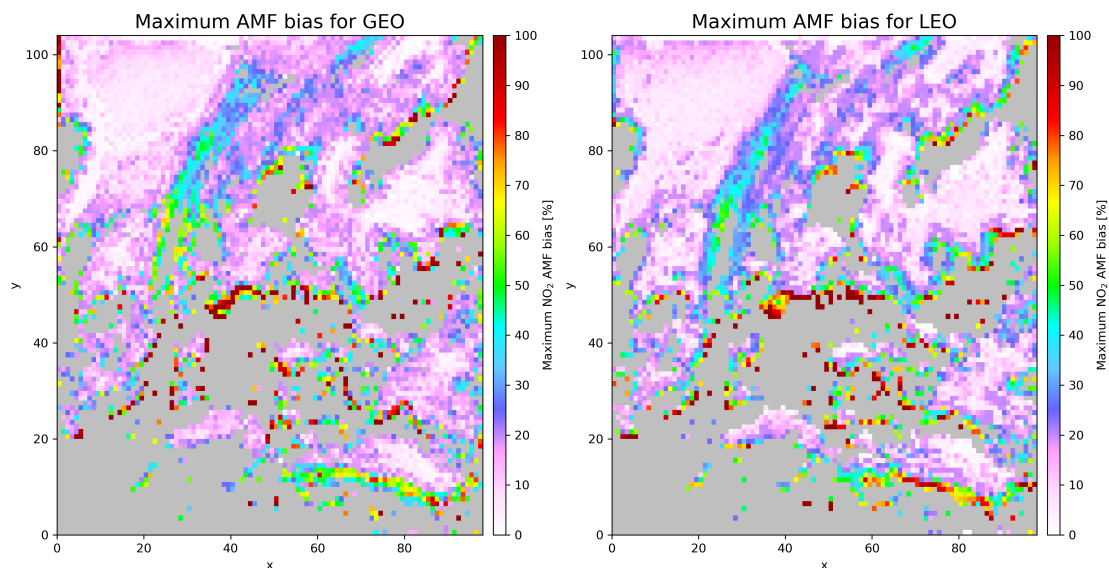

Figure 7. Maximum NO_2 AMF bias for each pixel (left: geostationary geometry; right: low earth orbiting geometry).
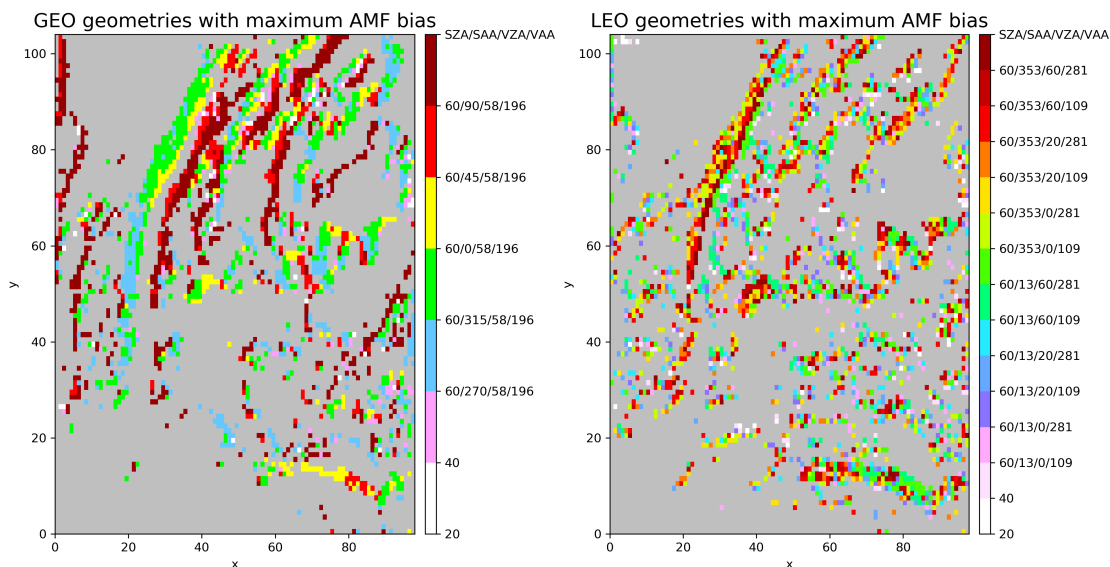
Figure 8. Geometry for the largest AMF bias for each location (left: geostationary geometry; right: low earth orbiting geometry). Only the pixels with maximum bias > 20% are shown.

the pixels around are covered by a convective cloud with large vertical extent and optical thickness > 100. Fig. 8 shows under which geometry the maximum bias is observed. Pixels with maximum biases < 20% are not shown. In general, the maximum bias is obtained at high SZA ($60°$), and for GEO cases, the bias also depends on the solar azimuth angle. Maximum biases are found on the east/west side of the cloud when the sun is in the west/east. Note that this dependency for GEO geometry is
5 particular for the region, and thus solar and viewing conditions, studied. This dependency is not seen for the LEO geometry as for LEO geometry the local daily revisit time is the same and thus also the solar azimuth.

Generally, as shown above, the NO_2 AMF retrieval bias is mostly within 20%, and the bias increases slightly with the weighted radiometric cloud fraction. The NO_2 profile and surface albedo input gives differences between the LIDORT and MYSTIC RTM on the order of 1% (Emde et al., 2021; Yu et al., 2021). Therefore, the retrieval bias is mainly due to the
10 simple cloud correction and 3D cloud effects. For the clear pixels ($CF_w < 1\%$), there is a significant number of pixels with the retrieval bias > 20% and this is due to cloud shadow effects.

To summarize the overall differences between the retrieved and "true" NO_2 TVCDs for all LEO and GEO cases, various statistics were calculated. These include the median, mean, standard deviation and skewness of the difference distribution. Furthermore, the number n of pixels with differences d: $d < -20; -20 < d < -10; -10 < d < -5; -5 < d < 5; 5 < d < 10; 10 <$
15 $d < 20; d > 20\%$ were calculated. These results are summarized in Fig. 9 for the low-earth-orbiting geometries and in Fig. 10 for the geostationary geometries. For the low-earth-orbiting geometry the bias varies with solar and viewing zenith angles (red solid line, Fig. 9b). The average median bias is -0.5%. Between 53% (large solar zenith angle) and 89% (small solar zenith angle) of the retrieved NO_2 TVCDs are within ±10% of the true column. The number of pixels with differences $< -20\%$
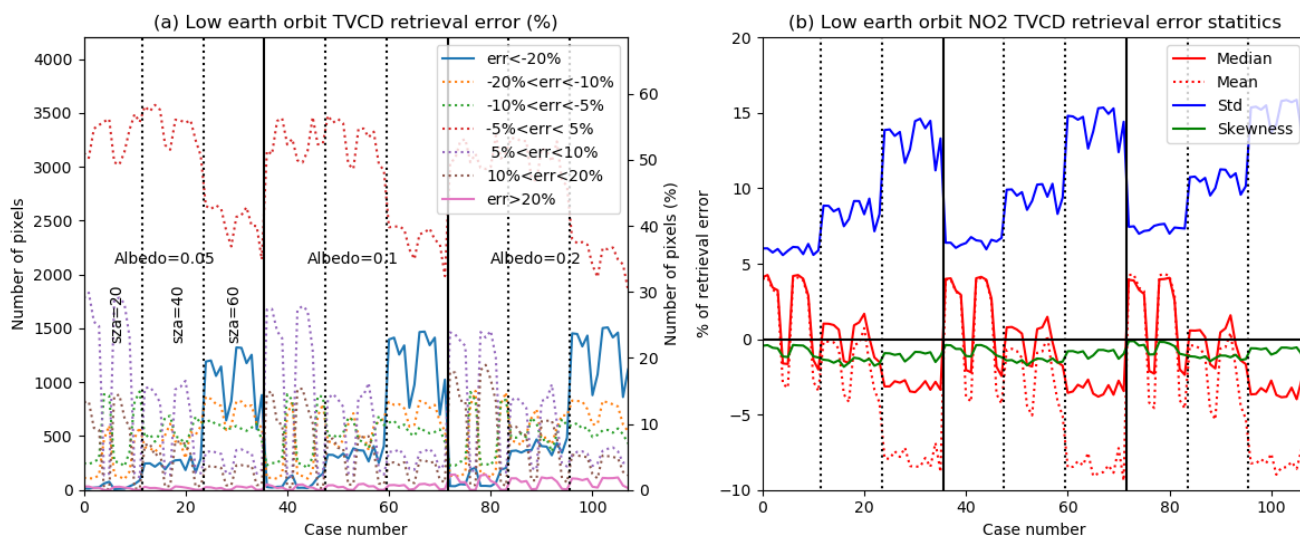
Figure 9. Summary of NO$_2$-retrieval differences for the low-earth-orbiting geometry. (a) The number of pixels with NO$_2$ TVCD differences larger than 20% (magenta solid line) and less then -20% (blue solid line). Also shown are pixels in other difference intervals as indicated in the legend. Note that the % is with respect to the total number of pixels for which a NO$_2$ retrieval has been done. This number is smaller than the number of pixels in the scene as cloud contaminated pixels are excluded from the retrieval. (b) The median, mean, standard deviation and skewness of the difference between the retrieved and input NO$_2$ TVCDs for each case. The x-axes give the simulation case number. See text for further details.

increase from about 0.1% for low albedo and high sun to up to 26% for high albedo and low sun (solid blue line, Fig. 9a). Within each solar zenith angle interval the solar azimuth angle takes two values and the results are similar for these two angles. The largest differences within each solar zenith angle interval are found when the viewing zenith angle is large (60°). The viewing azimuth angle has less influence.

5   For the geostationary geometry, there is an overall negative bias of about -0.9% (red line, Fig. 10b). Between 61% (large solar zenith angle) and 93% (small solar zenith angle) of the retrieved NO$_2$ TVCD are within 10% of the "true" column. The number of pixels with differences $< -20\%$ increase from about 0.2% for low albedo and high sun to up to 22% for high albedo and low sun (Fig. 10b). The variations within each solar zenith angle interval is due to different solar azimuth angles. The skewness is always negative, indicating a skewed distribution.

10   These results show that the solar zenith angle is of prime importance for 3D cloud impacts and that the impact increases with increasing solar zenith angle. For similar reasons, the viewing zenith angle is also of importance. Both under- and over-estimates occur in pixels close to clouds. Cloud shadows are a cloud feature metric that may be used to identify affected pixels, Fig. 3. However, while for large solar zenith angles pixels affected by cloud shadows are mostly underestimated, overestimates
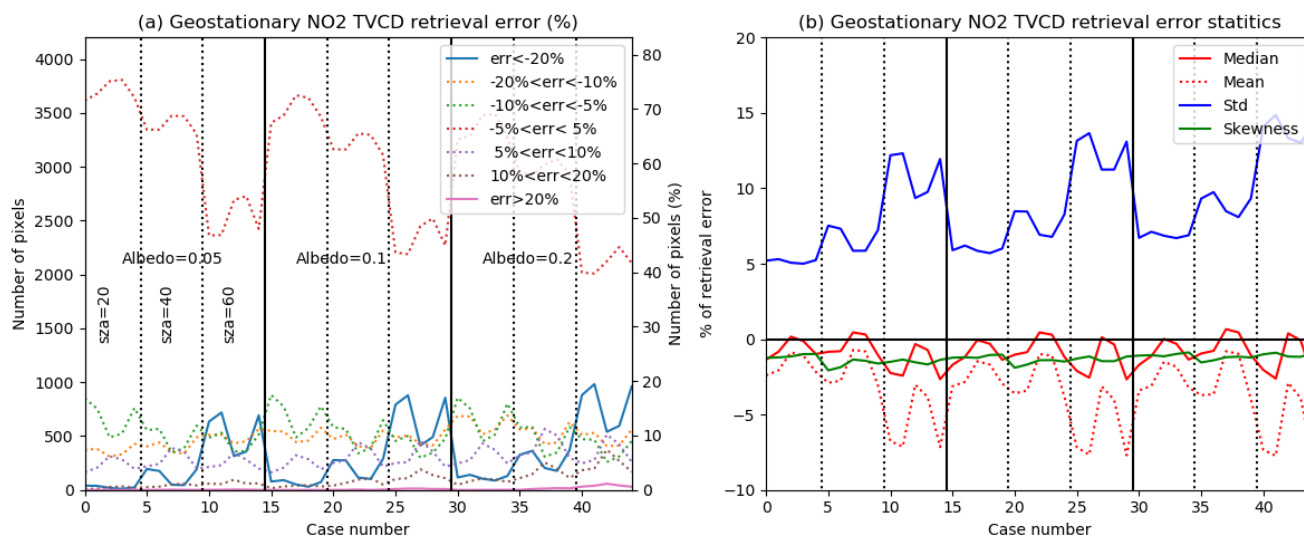
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Figure 10. As in Fig. 9, but for geostationary orbit geometry.

also occurr. Thus, cloud features in neighbouring pixels, such as cloud top altitude and cloud optical thickness, are also of importance.

According to box cloud simulations presented by Emde et al. (2021) and Yu et al. (2021), the cloud enhancement effect is of similar magnitude as the cloud shadow effect. For the synthetic data the enhancement effect is present, but is smaller in

5 magnitude than the cloud shadow effect, see Figs. 9 and 10. This is most likely due to cloud enhancement affected pixels being identified as cloudy and thus not analysed. Also the synthetic data indicate that the smaller the solar zenith angle is, the larger are the chances for cloud enhancements between 5-10% for low earth orbit satellites ($\theta = 20 - 30°$, purple dotted line Fig. 9a). The differences are largest for large satellite viewing angles (60°) and may thus indicate enhancement due to the satellite partly viewing sun-illuminated cloud sides. For geostationary satellite this effect is not present (Fig. 10a). This is due to differences

10 in the solar azimuth angles for the two geometries. However, in magnitude, this effect has a much smaller impact than cloud shadows.

For comparison Lorente et al. (2017) estimated the structural uncertainty (differences in retrieval methodology) in the tropospheric NO_2 AMF to be 42% over polluted regions and 31% over unpolluted regions, and mostly driven by the uncertainty in the a-priori NO_2 profile, cloud properties and surface albedo. Furthermore, Yu et al. (2021) showed that the NO_2 AMF bias

15 using the simple cloud correction is generally within 20% for cases with $CF_w < 50\%$. Yu et al. (2021) found that the retrieval bias significantly increases for high albedo (albedo=0.3) for 1D cloud layer simulations, whereas for box-cloud sensitivity studies, the 3D bias decreases with increasing surface albedo. The bias of the NO_2 retrieval has two origins: the simplistic treatment of clouds in the retrieval algorithm, and 3D effects. Due to the different results for the 1D and box-cloud results, the high bias for high albedo cases is believed to be due to the simplistic treatment of clouds in the retrieval algorithm. The

different cloud correction schemes (O_2-O_2 and O_2-A band) are normally within 10% (not shown) and this number may be interpreted as the level of uncertainty of the uncertainty introduced by these correction schemes. For real measurement, the surface albedo is usually different between the visible (for O_2-O_2) and NIR band (for O_2-A band), and the comparison thus more complicated.

5 4.2 Observational satellite data

4.2.1 Cloud shadow band cases

For the observational data the "true" NO_2 TVCD unaffected by clouds is not known. In order to have a reference or "true" NO_2 TVCD to compare with, we look at neighbour pixels in a 3×3 pixel matrix where the pixel of interest is in the centre. The "true" NO_2 TVCD is then taken to be the average of cloudfree neighbours with NO_2 retrieval quality value > 0.95. Obviously

10 this choice of "true" NO_2 TVCD has its problems including that neighbours may also be affected by clouds and that NO_2 TVCD may have large spatially gradients, see for example Fig. 2b which shows the NO_2 TVCD, and Fig. 2c which shows the percentage difference of the NO_2 TVCD to the area median NO_2 TVCD.

Based on the findings from the analysis of the synthetic data we searched TROPOMI and VIIRS data for cases with cloud shadows and large solar zenith angles. An example of a cloud shadow band is shown in Fig. 11. The NO_2 TVCD, Fig. 11b,

15 is markedly higher in the Amsterdam area, but is otherwise overall slowly varying over the region. The H-metric, Fig. 11c, is sensitive to inhomogeneous clouds, but also to variations in the surface albedo, compare Fig. 11a and 11c. The geometric cloud fraction, based on the VIIRS cloud mask, shows larger variability than the average radiometric cloud fraction CF_r^{VIIRS}, as expected, compare Fig. 11d and 11e and see also the distributions of these quantities in Fig. 11h. The cloud shadow fraction, Fig. 11f, is between 0.2 and 0.5 in the area with scattered clouds. The cloud shadow fraction also clearly delineates the large

20 cloud structures. The NO_2 TVCD versus the cloud shadow fraction is shown in Fig. 11g. For a cloud shadow fraction < 0.5 the NO_2 TVCD decrease with increasing cloud shadow fraction. However, the decrease is well within the variability of the NO_2 TVCD (the red line with error bars indicates the average NO_2 TVCD ±standard deviation). The distribution of the H-metric for various CF_r^{VIIRS} is shown in Fig. 11i. For CF_r^{VIIRS} between 0.4 and 0.6 (green bars) maxima are found both for low (scattered clouds) and high (homogeneous clouds or surface) H-metrics. Thus indicating that the CF_r^{VIIRS} may not

25 unambigiously be used to identify scattered cloud cases. For the absorbing aerosol index no obvious dependence between the AAI and the NO_2 TVCD is present for this region, see Appendix A for further details.

For the cloud shadow band pointed to by the red arrow in Fig. 11a, further analysis is provided in Fig. 12. A RGB zoom in of the cloud shadow band is shown in Fig. 12b where red marks indicate pixels with cloud shadow. The VIIRS cloud top height and the TROPOMI NO_2 TVCD are shown in Figs. 12c and 12d. The cloud top height, cloud optical thickness and

30 the cloud fraction is shown for row 267 in Fig. 12e. The cloud shadow is just north of a cloud band with optical thickness up to 10 (slant optical thickness of about 15) and an altitude between 9-10 km. North of the cloud shadow band there are also some clouds at the same height but with a smaller optical thickness of around 1.5-2. These thinner clouds are not easily seen in the VIIRS RGB, Fig. 11a, but are present in the VIIRS cloud mask and thus the geometric cloud and averaged radiometric
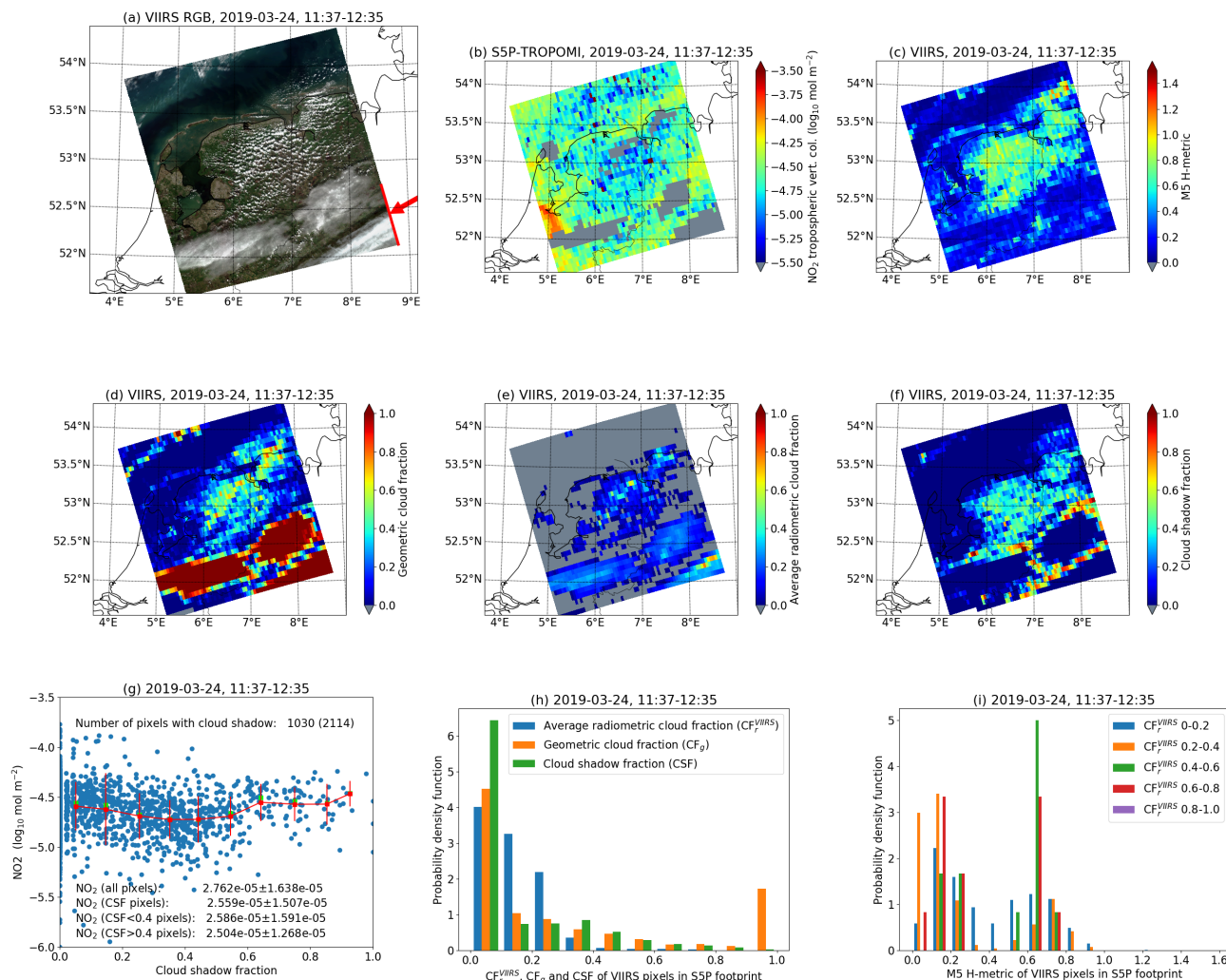
Figure 11. Examples of VIIRS and TROPOMI data and various CFMs for the cloud shadow band in Fig. 2. (a) VIIRS RGB. The red marked pixels correspond to the data shown in Fig. 12. (b) NO$_2$ TVCD from TROPOMI. (c) The H-metric from VIIRS band M5. (d) The cloud geometric fraction from VIIRS. (e) The cloud radiance fraction from VIIRS. (f) The cloud shadow fraction from VIIRS. (g) The NO$_2$ TVCD versus the cloud shadow fraction. The red lines is the average NO$_2$ TVCD with standard deviation (vertical lines). Lime green squares indicate the median NO$_2$ TVCD. (h) The distribution of the cloud radiance, cloud geometric and cloud shadow fractions. (i) The distribution of the H-metric for various cloud radiance fractions.

cloud fractions, Figs. 11d-e. The NO$_2$ TVCD for row 267 using FRESCO and OCRA/ROCINN cloud correction algorithms, is shown in Fig. 12f. The NO$_2$ TVCD inside the cloud shadow is low by about a factor $2-3$ compared with the NO$_2$ TVCD north of the cloud shadow. Fig. 12g displays the NO$_2$ TVCD south of the shadow, in the shadow and north of the shadow for rows

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262-269 and the average of these. A shadow pixel is defined as having a CSF $> 50\%$ and $CF_w < 50\%$. For row 263 no pixels satisfied this criteria and does no data is shown in the shadow region for this row. Pixels to be included in the regions north and south of the shadow band (up to 4 TROPOMI pixels north and south of shadow band), where required to have CSF $< 25\%$ and $CF_w < 50\%$. For the cloud shadow band the NO_2 TVCD is on average reduced by 17%. There is no clear dependence of the

5   NO_2 difference on the cloud shadow fraction. The cloud optical thickness of the cloud causing the cloud shadow varies little, the average cloud optical thickness being 9.0 ± 1.6. Thus it is not possible to say anything about the dependence of the NO_2 difference on the cloud optical thickness for this case.

Another example of a cloud shadow band and 3D cloud effects on NO_2 retrieval is shown in Fig. 13. The VIIRS RGB image (Fig. 13a) shows the cloud coverage over Northern Germany on December 30, 2019, when the solar zenith angle during the

10 VIIRS and TROPOMI overpass time is larger than $70°$. The cloud shadow is just north of a large cloud band, and most of the northern region is completely cloudless, which is similar to the idealized box cloud cases presented by Emde et al. (2021) and Yu et al. (2021). The cloud height from west to east is 1 km to 8 km (Fig. 13c). The cloud optical thickness is not available for this case, but from the RGB the cloud is clearly optically thick enough to make the ground not visible from space. To look for cloud shadow effects, we select the region within the red box in Fig. 13a). In Fig. 13e is shown the VIIRS reflectance and

15 cloud top height for TROPOMI row 394 as a function of latitude. The pixels are identified as cloudy, cloud shadow and clear based on the reflectance; with the reflectance being about 0.25 over the clear region, down to 0.18-0.24 in the cloud shadow, and higher than 0.25 for cloudy pixels. There are four TROPOMI pixels in the cloud shadow where the NO_2 TVCD is low by 20-60% compared with the NO_2 TVCD to the north and south of the cloud shadow (cloud and clear pixels). There is a slight difference in the NO_2 retrieval using different cloud corrections, Fig. 13f. In order to reduce the influence of the signal to noise

20 ratio, the NO_2 TVCD is averaged over cloudless (four pixels near the cloud shadow), cloud shadow and cloudy pixels (four pixel near the cloud shadow) for rows 393 to 398, as shown in Fig. 13g. For the cloudy pixels, the difference is within 10% between NO_2 averaged over all the pixels and NO_2 averaged for the pixels with high quality retrieval ($CF_w < 50\%$). All cases show that the NO_2 TVCD in the cloud shadow is lower by 8-46% (average of 25%) compared with the NO_2 TVCD around the shadow. Here, the NO_2 TVCD around the shadow represents NO_2 retrieval unaffected by 3D cloud, which is an average of

25 NO_2 over cloudy and cloudless pixels, in order to reduce the impact of spatial variation of NO_2.

If it is assumed that the clouds are the main reason for the variations in the NO_2 TVCD over the cloud shadow bands, then these cases are examples of how cloud shadows give underestimates of NO_2 TVCD, in agreement with the theoretical idealized box cloud results presented by Emde et al. (2021) and Yu et al. (2021).

4.2.2 Cloud effects for general cases

30 As shown in Figs. 9 and 10 for the synthetic data, the NO_2 bias may be on the order of tens of percent and are largest for large solar zenith angles. Hence, for the months of October 2018 and March 2019 when the solar zenith angle is between 50-60° for the study region, we looked for NO_2 TVCD biases due to clouds. For October 2018 and March 2019 this amounted to a total of 1,023,081 TROPOMI pixels. To quantify possible cloud shadow effects and factors that impact it, TROPOMI pixels with NO_2 retrieval data quality value > 0.95 and cloud shadows were selected for further analysis. A NO_2 retrieval with the data quality

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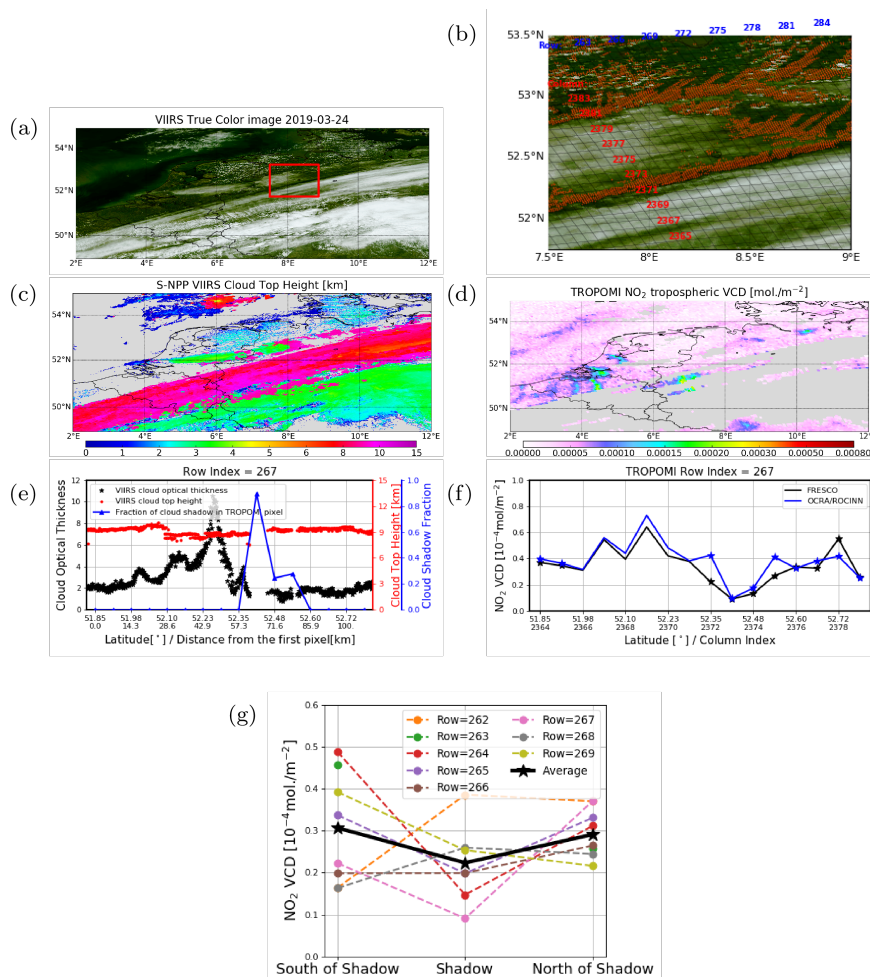Discussions
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Figure 12. Another view of the cloud shadow band example in Fig. 11. (a) VIIRS RGB image; (b) zoom-in of cloud shadow band with TROPOMI footprint, red marks indicate pixels with cloud shadow; (c) the VIIRS cloud top height; (d) the TROPOMI NO_2 TVCD; (e) the VIIRS cloud optical thickness, cloud top height and cloud shadow fraction for TROPOMI row 267 (the row localisation is shown in panel b); (f) the NO_2 TVCD using FRESCO and OCRA/ROCINN cloud correction algorithms, as a function of latitude. The star marks represent pixels with $CF_w < 50\%$; (g) averaged NO_2 TVCD for pixels south of the shadow, in the shadow and north of the shadow for rows 262-269 and the average of these. See text for further details.

value >0.95 was reported for 367,584 (36%) of the pixels. Of these, 129,180 (35%) were affected by cloud shadows according to the VIIRS cloud shadow product. Note that this number pertains to months of the year where we expect cloud shadow effects to be large. For months of the year with smaller solar zenith angles this number may well be smaller. The pixels affected by cloud shadows were further analysed to understand which parameters that have the largest impact on the NO_2 TVCD. In order

5 to have a reference NO_2 TVCD to compare the centre pixel value with, only centre pixels which have one or more cloudfree
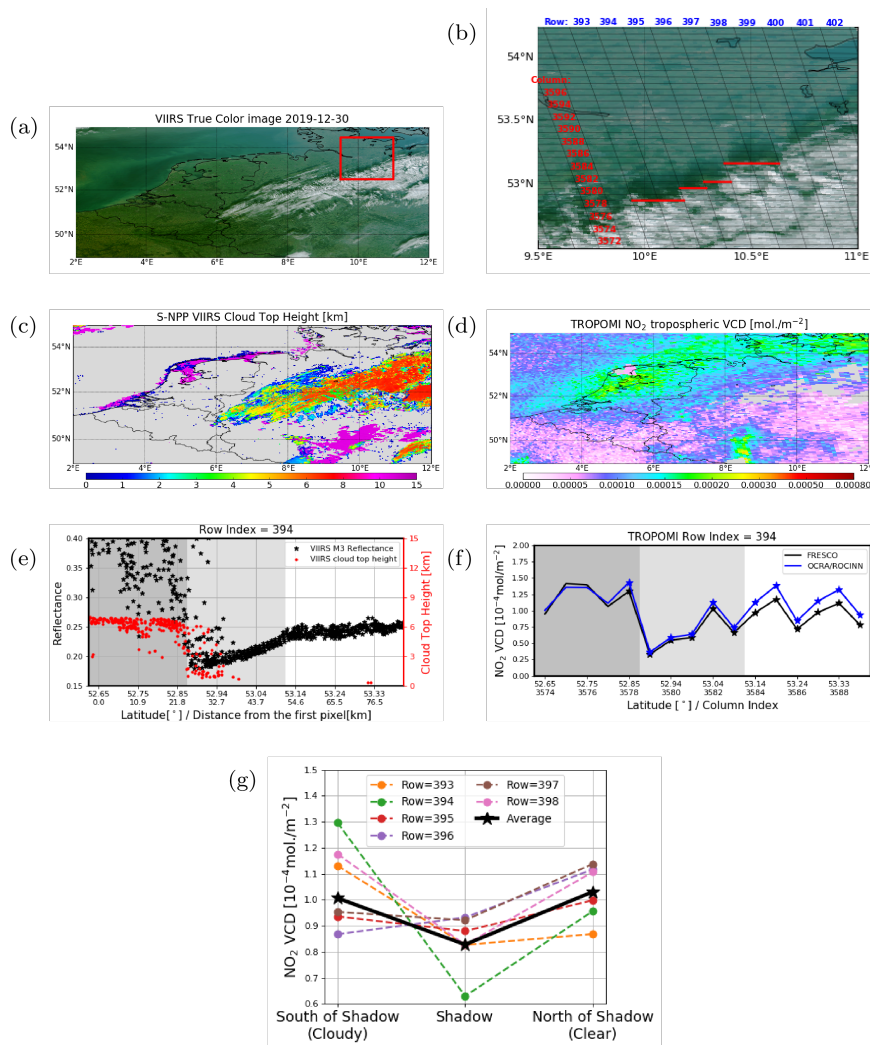
Figure 13. Similar to Fig. 12, but data for 30 December 2019. Furthermore, in (b) the red lines indicate the cloud edge in along-track direction for rows 393 to 398; in (e) no cloud shadow nor cloud opticall thickness information is included, instead the VIIRS M3 reflectance is plotted; in (e)-(f) cloud, shadow and clear regions are identified as dark gray/light gray/white regions.
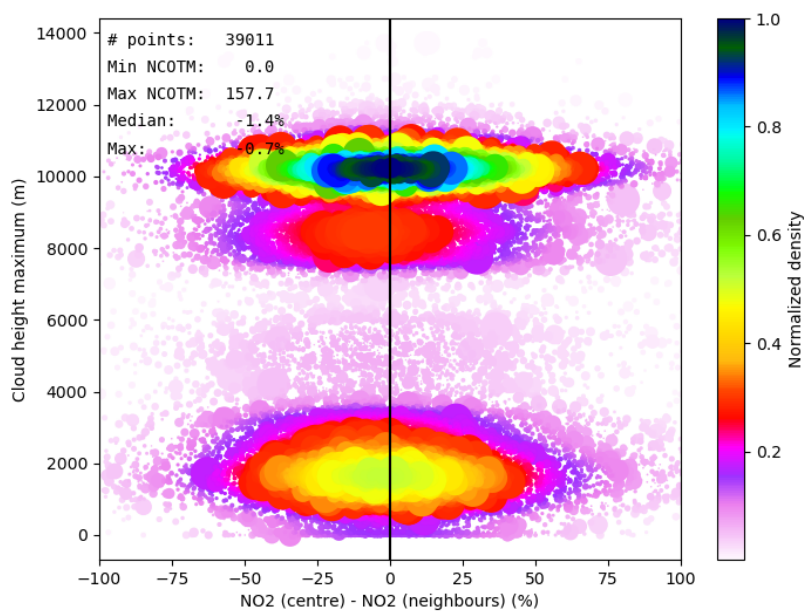
Figure 14. The difference between the centre pixel NO$_2$ TVCD and the average of the NO$_2$ TVCD in the cloudfree neighbours as a function of cloud height maximum of neighbour pixels. The median of the difference is given in the legend together with the maximum of a Gaussian kernel-density probability density function estimate of the data points. The size of the data points illustrates the maximum cloud optical thickness in neighbour pixels (NCOTM). It varies between the minimum and maximum NCOTM values given in the legend.

neighbours with NO$_2$ retrieval quality value > 0.95, were included. This restriction reduced the number of TROPOMI pixels to be analysed from 129,180 to 39,011. For these pixels the difference between the centre pixel NO$_2$ TVCD and the average of the NO$_2$ TVCD in the cloud free neighbours (ΔNO$_2$) is shown as a function of cloud height maximum of neighbour pixels in Fig. 14. For this data subset clouds are mostly found at high (about 10,000 m) and low (about 1,900 m) altitudes, with relatively

5   more clouds at the higher altitudes. This is qualitatively in agreement with Noel et al. (2018). They reported cloud fractions as estimated from a lidar on board the International Space Station. For Europe (their Fig. 5b) they report a similar cloud vertical behaviour, albeit for a different time of year (July, June and August). The subset of TROPOMI pixels presented in Fig. 14 shows both high and low NO$_2$ TVCD biases with a median NO$_2$ TVCD bias of -1.4% and maximum of the Gaussian kernel density probability density function estimate maximum of -0.7%. Thus, for this subset of TROPOMI pixels no signifcant cloud

10   shadow effect was visible in the NO$_2$ TVCD. As stated above, no true NO$_2$ TVCD is available as for the synthetic data and this may be the reason for the lack of a clear cloud shadow effect in this data subset.

This subset of data were further binned according to the maximum cloud height of neighbour pixels and the maximum slant cloud optical thickness (SCOT) of neighbour pixels as given in Table 1. The maximum slant cloud optical thickness is the

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Table 1. Binning of parameters used to quantify the cloud shadow effect.

Parameter	Bin borders
Maximum cloud height of neighbour pixels (NCH)	0, 3000, 9000, 20000
Maximum slant cloud optical thickness of neighbour pixels (SCOT)	0, 3, 5, 7, 10, 15, 300

neighbour pixel maximum cloud optical thickness divided by the cosine of the solar zenith angle for that pixel. Furthermore, only TROPOMI pixels where the maximum slant cloud optical thickness and maximum cloud top height are in the same neighbour pixel were included. This reduced the data set to 18,029 pixels.

The NO_2 TVCD bias density as a function of the cloud shadow fraction for the SCOT bins and maximum cloud heights is shown in Fig. 15. As the cloud height increases (from bottom row to top row in Fig. 15) the cloud shadow fraction increase because generally the cloud shadow within a pixel geometrically increase with cloud height when the cloud is larger than the pixel size. For the low clouds the median NO_2 TVCD bias varies between -1.2 to 0.2% for SCOT<15. For SCOT>15 the median NO_2 TVCD bias is -2.9%. For the medium height clouds the median NO_2 TVCD bias for SCOT>15 increase in magnitude to -5.8%. For smaller SCOT the median NO_2 TVCD bias is negative and varies between -0.8 and -3.4%. For the high clouds the median NO_2 TVCD bias is between -0.3 and 1.6% for SCOT<7. For larger SCOT the median NO_2 TVCD bias is -0.9, -8.7 and -15.1% (7<SCOT<10, 10<SCOT<15, 15<SCOT<300).

For all cloud heights the median NO_2 TVCD bias is negligible for SCOT<7. For low and medium clouds there is a negative bias for the largest SCOT. For the high clouds the bias is pronounced for SCOT>10. Thus, the median NO_2 TVCD bias increases with cloud height and with slant cloud optical thickness. It is noted that for individual pixels the bias may be larger and both positive and negative. We use a neighbour pixel as the reference NO_2-value. This assumes that only the cloud shadow effect is the reason for the NO_2 TVCD bias. In reality there are horizontal gradients in the NO_2 TVCD for numerous other causes as well, including local emissions and wind transport differences.

Finally it is noted that for the 129,180 pixels which were affected by cloud shadows, the number of pixels with cloud free neighbours where $| \Delta NO_2 |> 20\%$ is 45.7%. Under the assumption that this number is applicable to all TROPOMI pixels affected by cloud shadows, about 16% of TROPOMI pixels for which NO_2 TVCD retrievals with a high quality value were made, may be impacted by cloud effects larger than 20% for solar zenith angles between 50-60°. Of these about half of the NO_2 TVCD are overestimated and half underestimated. The underestimate is clearly linked to cloud shadow effects. The overestimate may be due to in-scattering or horizontal gradients in NO_2 concentrations and thus wrong cloud shadow free NO_2 TVCD reference. These two processes may, however, not be distinguished in the observed data set.

5 Conclusions

In this study we have investigated the impact of 3D clouds on NO_2 TVCD retrievals from UV-VIS sounders. Both synthetic and observational data have been used to identify and quantify possible biases in NO_2 TVCD retrievals. The synthetic data were based on high-resolution LES results which are input to the MYSTIC 3D radiative transfer model. The simulated visible spectra
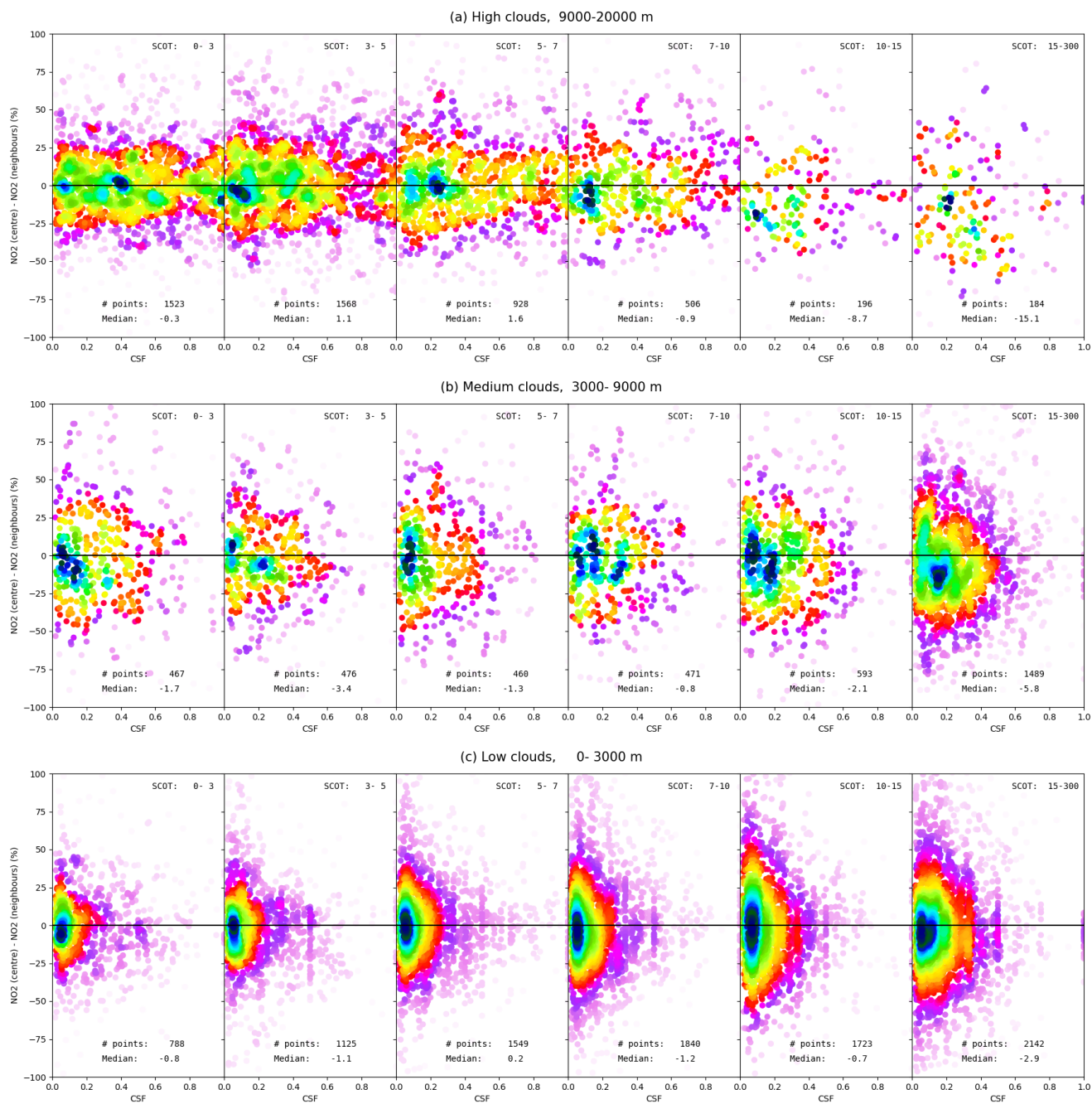
Figure 15. The NO$_2$-bias density as a function of the cloud shadow fraction. The data are binned into slant cloud optical thickness bins for high (a), medium (b) and low (c) clouds. See text for further details.

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for low-earth orbiting and geostationary geometries were analysed with standard NO_2 retrieval methods including operational cloud corrections. Profiles of NO_2 for polluted conditions were considered as cloud shadow effects are not important for background NO_2 conditions. For the observational data the NO_2 products from TROPOMI were used while VIIRS provided high spatial resolution cloud data. Both single cases and overall statistics were calculated. The main findings are:

5   – The following metrics were identified as being the most important in identifying 3D cloud impacts on NO_2 retrievals: cloud shadow fraction, cloud top height, cloud optical thickness, solar zenith and viewing angles.

   – Analysis of the synthetic data show that for low-earth and geostationary orbit geometries, 89 and 93%, respectively, of the retrieved NO_2 TVCD are within 10% of the actual column for small solar zenith angles. For large solar zenith angles the numbers decrease to 53 and 61%.

10   – The synthetic data shows that in general the NO_2 TVCD bias is slightly larger for low-earth orbiting geometries than for geostationary geometries. This is due to differences in viewing geometry where, for the mid-latitude targets studied here, the sun-target-detector geometry overall sees fewer cloud shadows for geostationary geometry.

   – For a solar zenith angle less than about 40° the synthetic data show that the NO_2 TVCD bias is below 10%. For larger solar zenith angles both synthetic and observational data show NO_2 TVCD bias on the order of tens of %.

15   – For clearly identified cloud shadow bands in the observational data, the NO_2 TVCD appears low-biased when the cloud shadow fraction > 0.0 compared to when the cloud shadow fraction is zero. If it is assumed that the clouds are the main reason for the variations in the NO_2 TVCD over the cloud shadow band, i.e. the NO_2 field is assumed to be horizontally homogeneous, then these cloud shadow band cases are examples of how cloud shadows give underestimates of NO_2 TVCD, in agreement with the theoretical findings.

20   – For solar zenith angles between 50-60°, about 16% of TROPOMI pixels (spatial resolution 3.5×7 km^2 at nadir) for which NO_2 retrievals with a high quality value were made, were found to be impacted by cloud effects larger than 20%.

The above conclusions suggest that further work is required on 3D cloud radiative impacts. For this future work the following topics may be considered:

   – As shown in this study, 3D radiative simulation of high-resolution satellite instrument spectra with realistic 3D cloud
25   input is fully achievable. However, to cover all possible atmospheric situations on Earth, possibly more high-spatial resolution LES results may be needed. Furthermore, 3D RT modelling is demanding on computer resources and requires careful interpretation and analysis of the output. It is expected that such modelling efforts may prove useful not only for trace gas retrieval algorithm studies, but also for cloud detection and cloud microphysics retrievals and more.

   – For future missions, including the Meteosat Third Generation (MTG) Flexible Combined Imager (FCI) and the Mete-
30   orological Operational Satellite Second Generation METimage (formerly known as VII-Visible and Infrared Imager), a cloud shadow product is needed for assessing and mitigating 3D cloud impacts. Experience with the VIIRS cloud
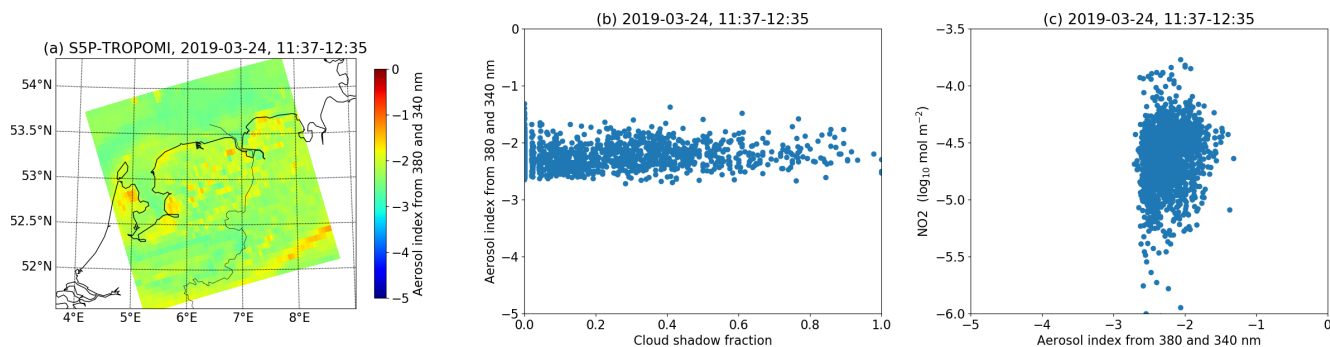
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Figure A1. (a) The TROPOMI aerosol index. (b) The TROPOMI aerosol index versus the cloud shadow fraction. (c) The TROPOMI aerosol index versus the NO_2 TVCD. For (b) and (c) only data points where the NO_2 TVCD data quality flag > 0.95 and the AAI data quality flag > 0.5, are included. All data from 24 March 2019.

shadow product (large changes between versions) suggests that independent verification with ground measurements may be of use. Such validation is non-trivial and possibly require new experimental approaches for measurements of both cloud shape and trace gas spatial variation at sub-pixel resolution.

Code availability. The libRadtran software used for the radiative transfer simulations is available from www.libradtran.org. The QDOAS
5   software for DOAS retrieval of trace gases is available from uv-vis.aeronomie.be/software/QDOAS.

Data availability. VIIRS data were accessed through the NOAA Comprehensive Large Array-Data Stewardship System (CLASS, https://www.class.noaa.gov). TROPOMI data were downloadded from https://s5phub.copernicus.eu/

Appendix A: Absorbing aerosol index

Recently Kooreman et al. (2020) identified and explained AAI large scale effects such as the cloud bow, sunglint and viewing
10   zenith angle dependence. In addition they reported small-scale negative AAI values in partly cloudy areas and attributed this to 3D cloud structures casting shadow, but left the in-depth analysis for a future study,

As cloud shadow impact NO_2 TVCD retrievals it is of interest to investigate possible relationships between the AAI and the cloud shadow fraction and the NO_2 TVCD. For the partly cloudy scene in Fig. 11, the TROPOMI AAI from 380 and 340 nm is shown in Fig. A1a. Overall the AAI is negative indicating the absence of absorbing aerosol, but the presence of scattering
15   particles. Indeed, over land the AAI is more negative over cloudy pixels, compare Fig. 11d and Fig. A1a. For pixels with cloud

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shadows, the AAI does not vary with the cloud shadow fraction as shown in Fig. A1b. It is noted that the NO_2 TVCD varies considerably for these pixels, Fig. A1c. Thus, while the NO_2 TVCD shows some dependcy on cloud shadow fraction, Fig. 11g, there appears to be no obvious dependence between the AAI and the NO_2 TVCD.

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Competing interests. The authors declare that no competing interests are present.

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