# Compositional data analysis (CoDA) as a tool to evaluate a new low-cost settling-based $PM_{10}$ sampling head in a desert dust source region

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**Abstract.** This paper presents a new sampling head design and the method used to evaluate it. The elemental composition of aerosols collected by two different sampling devices in a semi-arid region of Tunisia is compared by means of compositional perturbation vectors and biplots. This set of underused mathematical tools belongs to a family of statistics created specifically to deal with compositional data. The two sampling devices operate at a flow rate in the range of one cubic meter per hour, with

- 5 a cut-off diameter of 10 μm. The first device is a low-cost laboratory-made system, where the largest particles are removed by gravitational settling in a vertical tube. This new system will be compared to the second device, a brand-new standard commercial PM<sub>10</sub> sampling head, where size segregation is achieved by particle impaction on a metal surface. A total of 44 elements (including rare earth elements(REE), together with Al, As, Ba, Be, Ca, Cd, Co, Cr, Cu, Fe, K, Li, Mg, Mn, Mo, Na, Ni, P, Pb, Rb, S, Sc, Se, Sr, Ti, Tl, U, V, Zn, and Zr), was analysed in sixteen paired samples, collected during a two-week
- 10 field campaign in Tunisian dry lands, close to source areas, with high levels of large particles. The contrasting meteorological conditions encountered during the field campaign allowed a broad range of aerosol compositions to be collected, with very different aerosol mass concentrations. The CoDA tools show that no compositional differences were observed between samples collected simultaneously by the two devices. The mass concentration of the particles collected was estimated through chemical analysis. Results for the two sampling devices were very similar to those obtained from an on-line aerosol weighing system,
- 15 TEOM (tapered element oscillating micro-balance), installed next to them. These results suggest that the commercial  $PM_{10}$  impactor head can therefore be replaced by the decanter, without any measurable bias, for the determination of chemical composition, and for further assessment of  $PM_{10}$  concentrations in source regions.

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## 1 Introduction

- 20 At a global scale, mineral dust or mineral aerosols could represent about 40% of the total amount of particles injected into the atmosphere each year (Boucher et al., 2013; Huneeus et al., 2011). Studying atmospheric mineral dust, which modifies atmospheric radiation and alters cloud properties, thus impacting climate, is essential to better understand the evolution of Earth's climate system (e.g. Mahowald et al., 2011). Mineral dust is also an important source of nutrients necessary for phytoplankton growth in the open ocean (e.g. Okin et al., 2011) and for terrestrial plant development (e.g. Okin et al., 2004). Most of the
- 25 mineral dust present in the atmosphere comes from West Africa (Prospero and Nees, 1986; N'Tchayi Mbourou et al., 1997), with the Sahara as the main source (e.g. Ginoux et al., 2004). Accurate measurement of the chemical composition of aerosols is necessary for source tracing in aeolian studies (e.g. Scheuvens et al., 2013), which require aerosol data to assess global land degradation and climate change (e.g. Chappell et al., 2018).

In source regions of dry erodible material, high local wind speeds can move the largest and heaviest coarse soil particles
(between 50 and 200 μm in diameter) on the soil surface, while the smaller particles (less than 70 μm in diameter) move by saltation, a jumping movement near the soil surface. Collisions between these particles and aggregates of the finest particles present at the soil surface release a large spectrum of smaller particles into the air (Marticorena and Bergametti, 1995; Marticorena et al., 1997; Alfaro and Gomes, 2001). These fine particles, particularly those smaller than 10 μm in diameter (PM<sub>10</sub>), can be transported by wind at higher altitudes over long distances (Gillette, 1981; Gomes et al., 1990; Shao et al., 1993; Shao, 2008). These particles are also a key parameter in air quality control (Kuklinska et al., 2015).

Efforts are made in atmospheric sciences to develop devices able to prevent unwanted collection of the largest particles with a 10 µm cut-off diameter. Commercially available standard sampling devices are commonly used to collect fine particles. One of the most popular is the PM10 sampling head, where size segregation is obtained by removal of the largest particles through impaction on an aluminium alloy plate. This process may however contaminate aerosol samples with metal particles, because of

- 40 friction between coarse particles and the metallic parts of the system. This is not an issue for simple aerosol mass determination, but could generate problems if the objective is to define the chemical composition of airborne particles. It is well known that  $PM_{10}$  impactor inlet systems must be cleaned regularly: deposited particles that do not stick well on the impaction surface can bounce or can be de-agglomerated and re-entrained downstream, leading to oversampling (Le et al., 2019; Faulkner et al., 2014). Among other aerosol sampling head systems is the cyclone sampling device, where particles are separated by centrifugal
- 45 force. Cyclone walls may be made of glass instead of metal, thus reducing potential secondary emission effects. This system is nevertheless difficult to manage, because of its sensitivity to air pump flow rate (Haig et al., 2016). Impinger systems present a liquid impaction surface (Yu et al., 2016), and are well adapted for bio-aerosols, but not for mineral particles. In this study, the potential of a new PM<sub>10</sub> sampling head is evaluated in terms of mass collection efficiency and chemical composition accuracy. This new inlet uses the decantation principle; it can be build at low cost, using local materials, because of its simple design and
- 50 the broad availability of its components. Particle separation in this 125 mm-diameter Vertical Tube Decanter (VTD) system is based on gravitational settling counteracted by upward airflow. This system prevents collision between airborne particles and aerosol collector surfaces, so that sample contamination by metallic surface abrasion is minimized.

Source regions are good places to test possible biases introduced by the sampling head device in fine aerosol sampling, because coarse aerosols larger than 10  $\mu$ m are often present. Differences in cut-off diameter tuning will lead to differences in

- aerosol sample mass and chemical composition, as different amounts of the coarse particles present in the source zone will be collected. It is for this reason that we decided to compare the performance of two different sampling heads in a dry region of Tunisia. Aerosol chemical composition, including Rare Earth Elements (REE), and mass concentration of aerosols were measured at the same time using two sampling devices: the newly designed stainless steel decanter, VTD, and a brand-new aluminium alloy commercial PM<sub>10</sub> (hereafter PM10), both operating at a flow rate of about 1 m<sup>3</sup> h<sup>-1</sup>. The chosen sampling
- 60 station is part of the International Network to study Deposition and Atmospheric composition in AFrica (INDAAF), and is equipped with a reference instrument for mass concentration measurements, a  $PM_{10}$  automatic weighing device (Tapered Element Oscillating Microbalance, TEOM). Masses deduced from elemental analysis of samples collected by each device were compared with one another, and also with this third system, operating within the same flow rate range. The objective of this paper is to show that a low-cost decanter tube can replace an impaction-based  $PM_{10}$  sampling head for proper aerosol sampling.
- 65 To achieve this objective, we use Compositional Data Analysis (CoDA), an innovative tool for geochemical data analyses.

## 2 Materials and methods

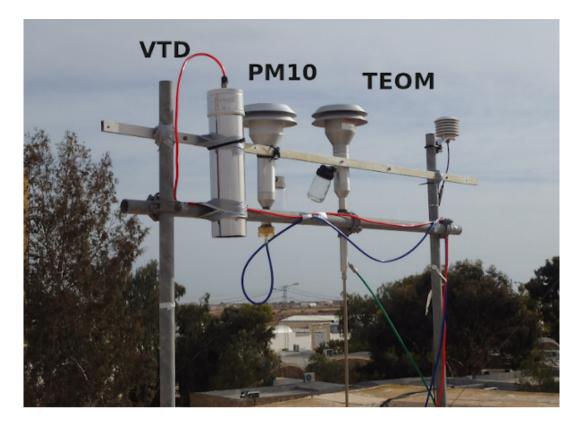
## 2.1 Aerosol sampling and direct measurements

Sixteen paired samples were collected during a two-week field experiment, at the Institut des Régions Arides campus, 20 km north of the city of Medenine, Tunisia. The collection site (33°29'58.62" N - 10°38'35.2" E), surrounded by dry lands, is 5 km
south-west of the Boughara Gulf. The two sampling devices were fixed to the roof of the highest building on campus, about 20 m above ground level. Both VTD and PM10 were attached to a tubular stand, with a distance of about 30 cm between them (Figure 1), to facilitate comparison of results. Aerosol samples were collected continuously from 2016/03/29 to 2016/04/07 using polysulfone open-face 47 mm filter holders (Nalgene®), and mixed cellulose ester filters, with a pore size of 0.45 µm (Whatman®). The filters were changed twice a day for each device at the same time: around 8:30 AM and 7:30 PM, except for the pair YX29/30, which was exposed for 24 hours.

Figure 2 shows the internal structure of the commercial PM10 sampling head (Tecora, Paris, France) installed, for the present study, with an aluminium alloy sampling plate. In the VTD system installed beside it (Figure 3), air is pumped at the top of the tube and enters from the bottom of the tube. Fine particles are dragged upwards by the airflow and collected by the filter, but the largest particles do not reach the filter because of their weight. The terminal settling velocity for a particle of diameter D in a gravitational field is calculated using Stokes' law (e.g. Calvert, 1990):

$$v_g = \frac{D^2(\rho_p - \rho_{air})g}{18\mu_{air}}$$

where  $v_g$  is the velocity of the particle when the steady state is reached;  $\rho_p$  is particle density;  $\rho_{air}$  is air density; g is gravitational acceleration; and  $\mu_{air}$  is the dynamic viscosity of air.



**Figure 1.** From left to right, the VTD system, the PM10 sampling head, and TEOM. Both TEOM and PM10 heads are from the same brand: Tecora<sup>TM</sup>  $PM_{10}$ 

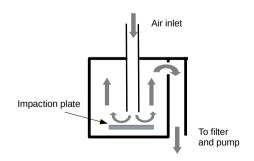


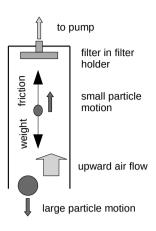
Figure 2. Internal structure of PM10 sampling head



(a) VTD installation on the roof of the building



(b) Two bottom-up views of filter on filter holder inside the decanter tube



(c) Diagram of decantation system

## Figure 3. The VTD system

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When a particle is in the upward airflow, it is pulled up unless its gravitational settling velocity is greater than the airflow velocity, in which case it will settle down. A cut-off point occurs when gravitational velocity is equal to air velocity: only particles smaller than this cut-off size can reach the top of the VTD system and thus be collected on the filter. With a flow rate of  $1 \text{ m}^3 \text{ h}^{-1}$ , the Reynolds number is equal to  $\approx 50$  inside the VTD. A laminar flow can be assumed, and therefore a constant air velocity in the tube. The steady state settling velocity of a particle is then reached when:

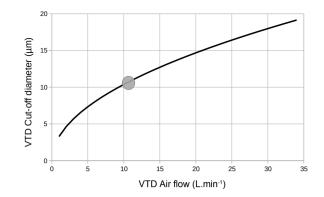
$$v_g = v_{air} = \frac{F_{air}}{\pi r^2}$$

where  $v_{air}$  is the upward air velocity,  $F_{air}$  is the pumped air flux, and r is the radius of the cylindrical VTD system, which is about six times smaller than its height. The cut-off diameter  $(D_{cut-off})$  can thus be rewritten as follows:

$$D_{cut-off} = \sqrt{\frac{18\mu_{air}F_{air}}{(\rho_P - \rho_{air})g\pi r^2}}$$

The  $D_{cut-off}$  value varies as a function of the pumped air flux when all the other parameters are fixed (Figure 4), so that it can easily be tuned to 10 µm. In an ambient air loaded with particles including a significant amount larger than 10 µm, perfect systems should exclude these largest fractions and therefore collect the same aerosol mass concentration with the same composition.

A Tapered Element Oscillating Microbalance (TEOM, Thermo Scientific), equipped with the same commercial PM10 head, was also installed beside the VTD and the PM10 systems (Figure 1). It measures the mass concentration of airborne particles directly, providing values considered as references for further comparison. A Portable Laser Aerosol Spectrometer (OPC,



**Figure 4.** Calculated particle cut-off diameter ( $\mu$ m) for VTD as a function of airflow. Calculations are performed using Stoke's equations for a vertical cylinder with a diameter of 125 mm. The grey dot shows actual operating conditions with measured airflows varying between 10 and 12 L min<sup>-1</sup> leading to a cut-off diameter between 10 and 11  $\mu$ m.

Model 1.108/1.109, Grimm), which measures particle size distribution over a large size range, was also installed ca. 3 m away from these three systems. A 1.111 Radial symmetric sampling head (Grimm) was installed at the air inlet of the instrument to ensure reasonable capture efficiency for large particles. The OPC measures the number of particles within 15 diameter intervals between 16 diameter channels of 0.30, 0.40, 0.50, 0.65, 0.80, 1.0, 1.6, 2.0, 3.0, 4.0, 5.0, 7.5, 10, 15, 20 and 25 µm. Counting by OPC is converted into mass, assuming lognormal distribution of spherical particles. The volume of  $N_i$  particles in the  $[d_i, d_{i+1}]$ diameter interval is equal to  $V_i = N_i \frac{\pi \overline{d}^3}{6}$ , where  $\overline{d}$  is the geometric mean of  $d_i$  and  $d_{i+1}$ . With a particle density  $\rho$ , commonly chosen to equal 2.2 g cm<sup>-3</sup>, the PM<sub>10</sub> mass,  $m_{10}$ , is equal to the sum of all the channels under 10 µm:

$$m_{10} = \sum_{d \le 10 \mu \mathrm{m}} \rho V_i$$

while the mass of coarse particles larger than  $10 \,\mu\text{m}$  is obtained by summing the channels over  $10 \,\mu\text{m}$ . This coarse particle mass fraction should not be sampled by our sampling devices.

## 2.2 Washing procedure for sampling instruments

- Prior to the field experiment, in the laboratory, 50 Petri dishes (PALL, filter storage box) were washed with detergent, and rinsed with tap water, after which they were soaked in osmosed water containing 2% of Decon90<sup>®</sup> for at least 15 hours. They were then thoroughly rinsed with tap water followed by osmosed water, before being soaked in acidified (HCl 1%) osmosed water for 3 days. Finally, the Petri dishes were rinsed with MilliQ<sup>®</sup> water (18 M $\Omega$  cm<sup>-1</sup>) and dried in an ISO-2 laminar flow hood. The filter holders and their PP boxes were cleaned using the same procedure. The PM10 head was disassembled and each
- 90 part was washed with tap water and detergent, and then soaked in osmosed water containing Decon90<sup>®</sup> for several minutes. Finally, each part was washed with osmosed and MilliQ water ( $18 \text{ M}\Omega \text{ cm}^{-1}$ ), and dried in the laminar flow hood. The tube of the VTD was washed with detergent, and rinsed with Milli-Q water, before being dried in the laminar flow hood.

## 2.3 Sample digestion

The filters coated with dust samples were brought back to the laboratory (ISO-7 clean room) and dissolved in sealed Teflon®

- 95 (PTFE) digestion vessels by 3 mL of a mixture of sub-boiled HNO<sub>3</sub>/HF (9:1) for 18 h on a heater plate at 125°C. All the Teflon vessels were previously cleaned with the detergent/acid procedure described above, completed with blank digestion. At the end of digestion, each vessel was opened, and the temperature of the heater plate was raised to 135°C, until complete evaporation of all liquid. The temperature of the heater plate was then lowered to 80°C, and 3 mL of a 30% nitric acid solution was added to each vessel, which was then sealed. Two hours later, the content of each vessel was transferred into a 60 mL
- 100 polypropylene bottle (thoroughly detergent/acid cleaned), by adding Milli-Q water. Laboratory blanks (no filter), 4 field blanks (pristine filter), and 2 finely ground geostandards (SCO-1 and MAG-1 from USGS) were also prepared following the same digestion procedure.

## 2.4 Chemical analyses

An ARCOS (Spectro-Ametek) ICP-AES, equipped with a CETAC ultrasonic nebuliser, was used for elemental determination
of Al, Ba, Ca, Cr, Fe, K, Li, Mg, Mn, Na, P, Sc, S, Sr, Ti, Zn, and Zr. A Field-Sector High-Resolution Inductively Coupled Plasma-Mass-Spectrometer (FS-HR-ICP-MS) Thermo Element 2, equipped with a concentric micro-nebuliser in a cyclonic nebulisation chamber was used for elemental determination of As, Be, Cd, Co, Cu, Mo, Ni, Pb, Rb, Se, Tl, U, V, and REE. External linear calibration was performed for all elements analysed with ICP-AES, by measuring a set of multi-elementary solutions with concentrations up to 250 μgL<sup>-1</sup>. The intercept was computed as the average of eight replicates of a blank
sample (ultra-pure nitric acid diluted in Mili-Q water). High-resolution analysis avoids polyatomic interference for elements

- lighter than arsenic, and also for REE (Heimburger et al 2013). The FS-HR-ICP-MS was externally calibrated for all elements analysed, with fourteen replicates of a blank solution and five replicates of a 1  $\mu$ g L<sup>-1</sup> multi-elementary solution. The first analytical detection limit was obtained with analytical blanks, and digestion with dilution water and acid reagents only, while the second field detection limit was obtained with blank filters transported to the field. For most of the elements, quantities
- 115 found in blank filters were higher than analytical detection limits, so that blank correction used the average quantity found in blank filters. For a few elements (Pr, Eu, Tb, Dy, Ho, Tm, and Lu), blanks were below detection limits so no blank correction was made. Seven elements (As, Cd, Cr, Mo, Ni, Sc, and Se) are not discussed, because they cannot be handled by the statistical tools used here, as at least one measured value was below the field or analytical detection limit. Analytical results are provided in Appendix A: Tables A1, A2, A3, A4, A5, A6 and A7.

## 120 2.5 Validation of analytical methods

There is no commercially available certified reference material comparable to the fine aerosols collected on filters. Two geostandards were therefore used as proxies: SCO–1 (typical of Upper Cretaceous silty marine shale), and MAG–1 (a fine-grained grey-brown clayey mud with low carbonate content, from the Wilkinson Basin of the Gulf of Maine). They were hand-crushed for 30 min in an agate mortar to approximate aerosol grain size. The powders produced were deposited on a filter at the small125 est amount that can be weighed (around 10 mg with an accuracy of 0.2 mg), to obtain a mass as close as possible to field aerosol samples. A table with individual recovery rates, as well as individual measurement results for each certified element and aerosol sample is proposed in Appendix B, Table B2. Recovery rates for most elements ranged from 80% to 120% for SCO–1 and MAG–1, but could not be calculated for S, Se, and Tm, because no value was available for comparison.

## 2.6 Computation of total aerosol mass concentration

130 The PM<sub>10</sub> mass concentration was not directly measured because of the low expected weight and the nature of the celluloseester filters which are sensitive to moisture. That is why a TEOM was installed, as it directly provides aerosol mass concentration in air. In this region, almost all the particle mass can be assumed to be carried by silicate crustal particles, sea-salts, sulphuric acid (H<sub>2</sub>SO<sub>4</sub>), and additional calcium in the form of calcium carbonate (CaCO<sub>3</sub>). A chemical reaction occurs between calcium carbonate and sulphuric acid, producing gypsum (CaSO<sub>4</sub>.2H<sub>2</sub>O), and preventing the simultaneous presence of sulphuric acid and calcium carbonate (Mori et al., 1998). If carbonate predominates over sulphuric acid, the total particle mass

concentration is computed as:

$$[particles]_{air} = [crust \ particles]_{air} + [sea \ salt]_{air} + [CaSO_4 \cdot 2H_2O]_{air} + [CaCO_3]_{air}$$

If sulphuric acid predominates over carbonate, then:

$$[particles]_{air} = [crust \; particles]_{air} + [sea \; salt]_{air} + [CaSO_4 \cdot 2H_2O]_{air} + [H_2SO_4]_{air} + [H_2SO$$

 $[crustparticles]_{air}$  is estimated using aluminium, and a crustal composition model where aluminium accounts for 7.01% of the mass (Bowen, 1966). This value is consistent with that of 7.09  $\pm$  0.79% observed by Guieu et al. (2002) for Saharan 140 dust.

$$[crust \ particles]_{air} = \frac{[Al]_{air}}{(X_{Al})_{crust \ model}} = \frac{[Al]_{air}}{7.01\%}$$

[sea salt] is estimated using sea-salt sodium, and a seawater composition model (Dickson and Goyet, 1994), where sodium accounts for 30.9% of sea-salt mass. Sea-salt sodium is deduced by subtracting crustal sodium from total sodium, crustal sodium being deduced from aluminium (Rahn, 1976), and a crustal composition model where the Na/Al ratio is equal to 0.0887 (Bowen, 1966):

$$[Na_{crustal}]_{air} = [Al]_{air} \left(\frac{[Na]}{[Al]}\right)_{crust\ model} = [Al]_{air} \cdot 0.0887$$
$$[Na_{sea\ salt}]_{air} = [Na_{total}]_{air} - [Na_{crustal}]_{air}$$
$$[sea\ salt]_{air} = \frac{[Na_{sea\ salt}]_{air}}{(X_{Na})_{seawater\ model}} = \frac{[Na_{sea\ salt}]_{air}}{30.9\%}$$

145  $[CaSO_4 \cdot 2H_2O]$ ,  $[CaCO_3]$  and  $[H_2SO_4]$  are calculated using additional calcium and additional sulphur not included in crustal and sea-salt estimation. Ca\* and S\* are defined respectively as calcium and sulphur of neither sea-salt nor crustal origin. Ca\* and S\* are computed using the same crustal and sea-salt composition models previously used:

$$\begin{split} [Ca^*]_{air} &= [Ca]_{air} - [Na_{sea\ salt}]_{air} \cdot \left(\frac{[Ca]}{[Na]}\right)_{sea\ salt\ model} - [Al]_{air} \cdot \left(\frac{[Ca]}{[Al]}\right)_{crustal\ model} \\ [Ca^*]_{air} &= [Ca]_{air} - [Na_{sea\ salt}]_{air} \cdot 0.037 \cdot [Al]_{air} \cdot 0.193 \\ [S^*]_{air} &= [S]_{air} - [Na_{sea\ salt}]_{air} \cdot \left(\frac{[S]}{[Na]}\right)_{sea\ salt\ model} - [Al]_{air} \cdot \left(\frac{[S]}{[Al]}\right)_{crustal\ model} \end{split}$$

$$[S^*]_{air} = [S]_{air} - [Na_{sea\ salt}]_{air} \cdot 0.0843 \cdot [Al]_{air} \cdot 0.0099$$

Depending on the resulting products of calcium carbonate with sulphuric acid reaction, the mass associated to additional calcium and sulphur is computed as follows:

$$[CaSO_4 \cdot 2H_2O]_{air} + [CaCO_3]_{air} = [S^*]_{air} \frac{M_{CaSO_4 \cdot 2H_2O}}{M_S} + \left( [Ca^*]_{air} - [S^*]_{air} \frac{M_{Ca}}{M_S} \right) \frac{M_{CaCO_3}}{M_{Ca}}$$
$$[CaSO_4 \cdot 2H_2O]_{air} + [H_2SO_4]_{air} = [CaS^*]_{air} \frac{M_{CaSO_4 \cdot 2H_2O}}{M_{Ca}} + \left( [S^*]_{air} - [Ca^*]_{air} \frac{M_S}{M_{Ca}} \right) \frac{M_{H_2SO_4}}{M_S}$$

where  $M_X$  is the molar mass of the compound or element X.

# 2.7 Multivariate analysis for compositional data (CoDA)

- 150 Compositional data are, by nature, difficult to handle straightforwardly. Any given component cannot vary independently from the others, because the sum of all components, equal to 100%, includes components that are not measured. If this closure constraint is not taken into account, spurious correlations and biased conclusions are to be expected (Van der Weijden, 2002). Appropriate mathematical tools must therefore be selected to overcome this drawback. These questions are extensively discussed in several papers (Aitchison, 1986, 1992, 2005; Barceló-Vidal et al., 2001; Filzmoser et al., 2009; Egozcue, J.J. et al., 2003).
- Briefly, the suitable sample space of any compositional vector x, representing a D-part subset of a whole  $\mathbf{x} = [x_1, \dots, x_D]$ , is the simplex  $S^D$ , as defined by Aitchison (1986). This technique is particularly well adapted to situations where elemental ratios are more relevant than absolute concentrations.

Let  $\mathbf{x} = [x_1, \dots, x_D]$  and  $\mathbf{y} = [y_1, \dots, y_D]$  denote two compositional vectors in  $S^D$ . Then  $\mathbf{z}$ , corresponding to the perturbation of  $\mathbf{x}$  by  $\mathbf{y}$ , in  $S^D$  is given by:

$$\mathbf{z} = \mathbf{x} \oplus \mathbf{y} = C[x_1y_1, ..., x_Dy_D]$$

with C the closure-to-unity operation defined as:

$$C(\mathbf{x}) = \begin{bmatrix} x_1 \\ \sum_{i=1}^{D} x_i, \dots, \frac{x_D}{\sum_{i=1}^{D} x_i} \end{bmatrix}$$

The neutral element of the perturbation is  $\mathbf{e} = C[1, ..., 1] = \begin{bmatrix} \frac{1}{D}, ..., \frac{1}{D} \end{bmatrix}$ , and  $\mathbf{x} = \mathbf{x} \oplus \mathbf{e}$ , while the perturbation vector expression compositional change from  $\mathbf{y}$  to  $\mathbf{x}$ , noted  $\mathbf{x} \oplus \mathbf{y}$ , is equal to  $\mathbf{x} \oplus \mathbf{y}^{-1}$ , with  $\mathbf{y}^{-1} = C[y_1^{-1}, ..., y_D^{-1}]$  (von Eynatten et al., 2002; Aitchison and Ng, 2005). The centred log-ratio (*clr*) transformation is commonly performed to open the data before applying any multivariate techniques based on correlation:

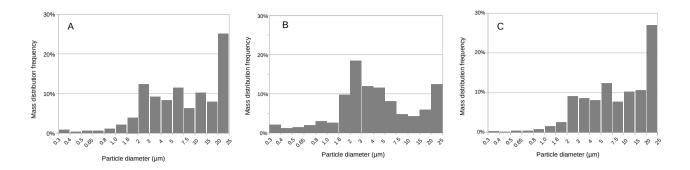
$$clr(\mathbf{x}) = \left[ln\frac{x_1}{gm(\mathbf{x})}, \dots, ln\frac{x_D}{gm(\mathbf{x})}\right]$$

where  $gm(\mathbf{x})$  denotes the geometric mean of the D parts:  $gm(\mathbf{x}) = \left(\prod_{i=1}^{D} x_i\right)^{\frac{1}{D}}$ . A Principal Component Analysis (PCA) can then be computed on transformed data to summarize the structure of the data in a lower dimensional space (ideally two for the sake of simplicity of projection on a plane). A compositional biplot, where both samples and variables are plotted in the same space can be used as a user-friendly graphical representation, but it differs from the original biplot by Gabriel (1971) in the sense that rays formed by the variables are proportional to the standard deviation of their log-ratios, and that the length of a link between arrow heads of two rays represents the standard deviation of the log-ratio between these compositional parts (Suárez et al., 2016). Practically, the "acomp" (closure operation) and "princomp" (PCA projection) functions used here were provided by the "compositions" package for the R software (R Core Team, 2014), which was specifically designed to analyse compositional data (van den Boogaart et al., 2014). This data processing based on log-ratio computing is named "Compositional Data Analysis (CoDA)".

## 3 Results and discussion

### 3.1 Variability of sampling conditions

- 170 The sampling site can be influenced by local and remote soil dust emission, by sea salt, and by anthropogenic emissions. During the sampling campaign, a broad variety of meteorological conditions was observed, allowing different aerosol sources to be sampled. Average local wind speed varied from about 1 to  $7 \text{ m s}^{-1}$  with no preferred direction (Appendix C, Figure C1). Backward air trajectories are presented for each sample pair in Appendix C (Figures C2, C3, C4, C5), indicating their differences in origin, leading to a variety of conditions for aerosol loading. Atmospheric aerosol loading presented a large range of
- values, from 21 to 679  $\mu$ g m<sup>-3</sup> (Table 1, TEOM values), with great variations between marine versus crustal proportions in any given sample pair.



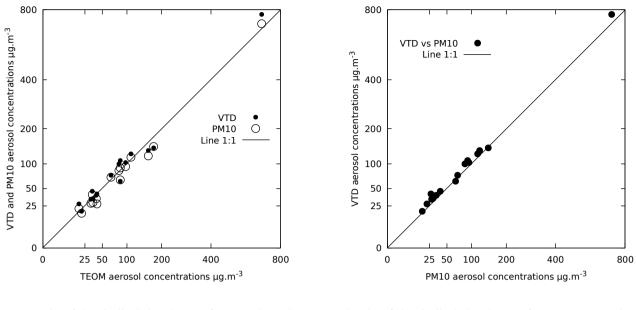
**Figure 5.** Daily average particle mass concentration size distribution in air on March 31 (A, aerosol concentration ca.  $40 \,\mu g \,m^{-3}$ ), April 6 (B, aerosol concentration ca.  $100 \,\mu g \,m^{-3}$ ) and April 2, morning (C, aerosol concentration ca.  $700 \,\mu g \,m^{-3}$ ). Measured using the Grimm OPC

## 3.2 Size distribution of the sampled aerosol

The fraction of particles larger than 10 µm suspended in the air is shown by OPC measurements. For the entire field experiment this coarse fraction represents, on average, 34% of the total mass concentration of aerosols as plotted in Figure 5, for three given periods with various dust concentrations. The presence of a significant amount of large particles in air makes the systems sensitive to possible inaccuracy and variations in their cut-off diameters: if the cut-off diameter was not the same for each sampling head in a given sample pair, the amount of large particles collected would not be the same and would produce differences in sampled aerosol mass concentration. Because chemical composition may be dependent on particle size, differences in cut-off diameters would also produce differences in chemical composition.

## 185 3.3 Total aerosol mass concentration in air

Comparisons of the measured mass concentrations between VTD, PM10, and the reference instrument TEOM are shown in Figure 6 and Table 1. Mass concentrations are averaged during each collection period. Plotted concentrations vary from 21 to 680 μg m<sup>-3</sup> within a range that can be cleverly plotted using a square root scale (Verrall and Bell, 1969). Masses of particles collected by VDT and PM10, deduced from calculations using Al, Na, S, and Ca, fit the TEOM values (Figure 6a). Similar results are observed for each VTD and PM10 sample pair (Figure 6b), suggesting the same collection efficiency for both sampling heads, and hence the same cut-off diameter. The median value of the relative mass differences between VTD and PM10 is +12%, and values range from -3% to +22%. Such variability is of the same magnitude as that observed by Heal et al. (2000) or Hitzenberger et al. (2004) in PM<sub>10</sub> and PM<sub>2.5</sub> inter-comparison exercises or by Motallebi et al. (2003) in a comparison of entire monitoring networks. An orthogonal regression, also known as total least square, was performed on the data presented here by treating the variances of x and y symmetrically. Orthogonal regressions were performed twice, with and without the highest point, which could potentially be considered as an outlier. In each case, slope and intercept are not significantly different from 1 and zero, respectively, suggesting that if any bias really occurs, it is too small to be identified with our data (see Tables D1 and D2 in Appendix D). To summarize, the differences observed between aerosol masses measured



(a) Plot of chemically deduced mass of VTD and PM10 sampling heads versus TEOM measurement.

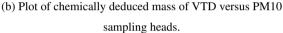


Figure 6. Comparisons of sample masses using square-root scale. The lines y = x are also shown

by the three sampling systems are much lower than the daily variability observed during the field experiment. The coherence
between direct measurement of masses (TEOM) and "chemical" weighing shows that substances not taken into account in our chemical budget (ammonium and organic molecules) do not significantly contribute to the total aerosol mass here.

## 3.4 Compositional data

The aim is now to compare chemical compositions of samples collected simultaneously by both VDT and PM10, as differences may appear due to contamination or size segregation of particles. Note that major and trace elements are treated separately from the REE in the following, because of the particular importance of REE as tracers of mineral particle origin (Wang et al., 2017).

## 3.4.1 Major and trace elements

The first two axes of the compositional biplot built from major and trace elements, without REE, explain 77% of the total variance (61% and 16% respectively), a high value, considering that 23 variables are taken into account for the analysis (Figure 7a). The variability between each pair of samples (i.e. collections by PM10 and VTD on the same day), figured by the

210 segment linking the two samples of the same pair, appears to be much lower than the variability observed within the entire set of samples. In other words, each dust event can be characterised properly with respect to the others, independently of the sampling device used. This finding is in good agreement with a close examination of compositional changes between PM10 and VTD for each pair of samples, expressed as perturbation vectors:  $VTP \ominus PM10$ , with  $VTP \ominus PM10 \in S^{23}$  (Figure 7b).

Sample Name	e Date start	Date stop	TEOM	VTD	PM10	VTD	VTD	PM10	PM10
			$\mu \mathrm{g}\mathrm{m}^{-3}$	$\mu gm^{-3}$	$\mu \mathrm{g}\mathrm{m}^{-3}$	Sea salt %	Crustal %	Sea salt %	Crustal %
YX03/04	29/03/16 18:40	30/03/16 09:12	21	19	17	14%	46%	24%	45%
YX05/06	30/03/16 09:47	30/03/16 18:26	18	27	22	25%	37%	25%	37%
YX07/08	30/03/16 19:08	31/03/16 09:03	33	33	28	16%	61%	13%	60%
YX09/10	31/03/16 09:38	31/03/16 18:29	86	63	66	8%	73%	10%	71%
YX11/12	31/03/16 19:01	01/04/16 09:09	41	40	34	23%	55%	25%	54%
YX13/14	01/04/16 09:39	01/04/16 18:19	175	140	145	14%	75%	15%	75%
YX15/16	01/04/16 18:49	02/04/16 09:38	111	124	116	4%	92%	4%	93%
YX17/18	02/04/16 10:08	02/04/16 19:23	679	769	711	1%	95%	1%	95%
YX19/20	02/04/16 19:49	03/04/16 09:52	82	99	84	18%	69%	15%	71%
YX21/22	03/04/16 10:19	03/04/16 18:29	66	75	70	29%	50%	31%	47%
YX23/24	03/04/16 18:59	04/04/16 10:01	42	41	35	55%	22%	59%	18%
YX25/26	04/04/16 10:31	04/04/16 19:31	97	103	94	17%	69%	19%	69%
YX27/28	04/04/16 20:01	05/04/16 10:00	36	35	29	33%	43%	34%	40%
YX29/30	05/04/16 10:24	06/04/16 09:06	157	133	120	6%	86%	6%	86%
YX31/32	06/04/16 10:33	06/04/16 18:51	85	108	91	24%	62%	23%	64%
YX33/34	06/04/16 19:19	07/04/16 07:16	35	45	40	40%	42%	39%	40%

Table 1. Sampling dates (local time) and aerosol mass concentrations directly measured by TEOM and calculated from chemical analysis of samples collected by VTD and PM10 respectively. Masses derived from chemical analyses are computed using equations presented in section 2.6. The last four columns (right) display mass proportion of sea salt and crustal aerosol for VTD and PM10 samples. Detailed results are shown in Tables E1 and E2 (Appendix E).

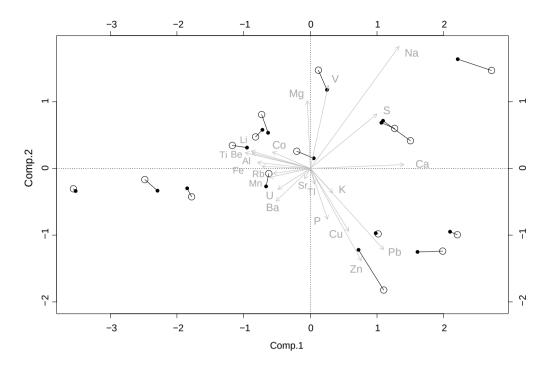
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Interestingly, the neutral element  $\mathbf{e} = \begin{bmatrix} \frac{1}{23}, \dots, \frac{1}{23} \end{bmatrix} = \begin{bmatrix} 0.043, \dots, 0.043 \end{bmatrix}$ , which indicates no perturbation, is included inside all the box plot quartiles. No systematic compositional shift, in terms of elemental ratios, can therefore be observed between the two sampling heads, at least for these elements, and it can be concluded that sample composition is not affected by the type of sampling head. Note, however, that Zn exhibits the greatest variability, suggesting noticeable random contamination. The slight differences observed between the two sampling heads in each paired sample are found to be correlated neither to air aerosol concentrations nor to wind speed. Potential contamination issues due to aluminium impaction plates were among the 220 main reasons why sampling heads were tested in the field with natural aerosols. No systematic compositional differences were observed between the two sampling heads although they are made of different alloys. This observation strongly suggests that neither of the two devices (brand new PM10 and VTD) would contaminate natural samples collected during this campaign.

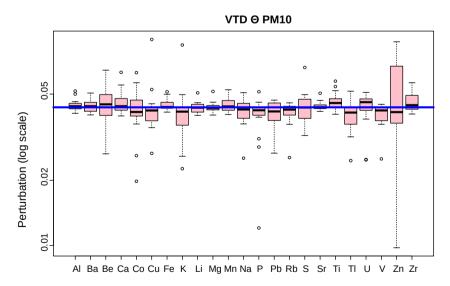
#### 3.4.2 **Rare Earth Elements (REE)**

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In the compositional biplot built from REE, only 51% of the total variance is explained by the first two axes (Figure 8a). This value is much lower than that obtained above for the other chemical elements (77%), but with only half the number of variables. The corresponding perturbation vector diagram again shows no systematic difference between the two sampling



(a) Biplot for the two sampling devices, PM10 samples are figured with circles, solid disks represent VTD samples. Lines between PM10 and VTD symbols link paired samples. Percentages of variability explained by the first two components are 61% and 16%, a total of 77%.



(b) Perturbation diagram as box plots for paired samples, measured by PM10 and VTD, for all elements except REE. The horizontal blue line represents no perturbation.

Figure 7. Biplot and perturbation diagrams for all elements except REE

heads (Figure 8b). Because REE essentially come from a stable crustal source, log-ratios between these elements vary little within the sample set (almost ten times less than the variability observed for the other elements). This stability explains why the percentage of variance expressed by the first two principal components is so low.

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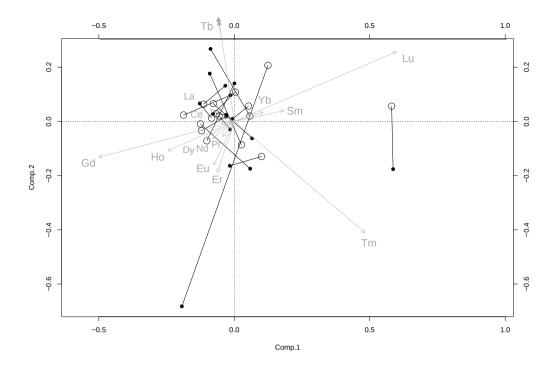
To test if the differences observed between the two systems might be explained solely by analytical error, the behaviour of identical duplicate samples was simulated: 16 new pairs of compositions were generated, by pairing each VTD sample with a modified sample, where each REE measurement was randomly shifted inside the given uncertainty interval of that REE. These new pairs of simulated samples were then represented as a biplot (Figure F1, Appendix F), producing results very similar to those observed for the real (VTD and PM10) paired samples. During this field campaign, the REEs profiles were found to be stable and unaffected by the design of the sampling head.

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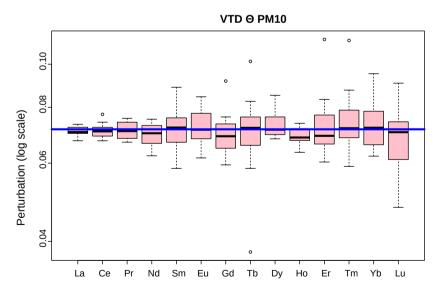
## 4 Conclusions

The main advantage of this new PM<sub>10</sub> inlet is its simple design associated with its low cost and the broad availability of the components making this new inlet easy to build locally by everyone. A second possible reason to use VTD is easier maintenance. Compositional Data Analysis tools have been used to present large sets of measurements at a glance, allowing us to perceive the compositional similarity of paired samples quickly and directly. No significant differences between the laboratory-made decanter sampling head and the commercial PM<sub>10</sub> sampling head (based on impaction) were observed in terms of aerosol composition (including REE) and total mass concentration, for samples collected in a source region of mineral dust, under very different meteorological conditions. In the source region investigated, where particle mass concentrations ranged from 20 to 700 µg m<sup>-3</sup> according to TEOM values, the chemical composition of the PM<sub>10</sub> aerosol fraction was therefore unaffected by the sampling head design. Consequently, both devices can be used for the determination of mass and chemical composition of aerosols in source regions, or even simply to determine mass by gravimetry. An aerosol survey network can therefore be built using a combination of the two sampling devices without any measurable consequences on data reliability or

consistency. This would also be the case for a time series if a PM10 was replaced by a VTD, or vice versa.



(a) REE biplot for the two sampling devices. PM10 samples are figured with circles, solid disks represent VTD samples. Lines between circles and disks link paired samples. Percentages of variability explained by the first two components are 29% and 22%, a total of 51%.



(b) REE perturbation diagram as box plots for paired samples, measured by PM10 and VTD. The horizontal blue line represents no perturbation.

Figure 8. Biplot and perturbation diagrams for REE

## Appendix A: Air concentrations, measured values

250 Raw data of the paper are presented in Tables A1, A2, and A3 for ICP-AES measurements, in Tables A4 and A5 for ICP-MS measurements, and in Tables A6 and A7 for REE measured with ICP-MS. "DL" is "Detection Limit" expressed in mass on the filter. "<" is "less than concentration detection limit"; this concentration detection limit must be calculated by dividing the DL value (expressed in mass) by the air volume. Uncertainties are given for a 95% confidence interval. The air volume uncertainty is constant at 1% and not displayed.

# Table A1. Elemental air concentrations measured with ICP-AES

Element		Al	Ca	Fe	K	Mg	Na
Wavelength (nm)		396.2	396.847	238.2	766.491	279.553	589
Analytical DL (ng	3)	0.02	0.005	0.1	0.2	0.003	1
Field DL (ng)		0.5	13	0.8	3	1	79
Sample name	Air volume (m <sup>3</sup> )	$\mu gm^{-3}$	$\mu gm^{-3}$	$\mu \rm gm^{-3}$	$\mu gm^{-3}$	$\mu gm^{-3}$	$\mu gm^{-3}$
YX03 (VTD)	10.52	$0.63\pm0.02$	$1.7\pm0.1$	$0.4\pm0.01$	$0.86\pm0.04$	$0.32\pm0.01$	$0.9\pm0.3$
YX04 (PM10)	12.59	$0.55\pm0.02$	$1.6\pm0.1$	$0.34\pm0.01$	$1.46\pm0.05$	$0.29\pm0.01$	$1.3\pm0.2$
YX05 (VTD)	5.89	$0.71\pm0.02$	$2.6\pm0.1$	$0.44\pm0.02$	$2.05\pm0.08$	$0.33\pm0.02$	$2.2\pm0.5$
YX06 (PM10)	6.68	$0.58\pm0.02$	$2.2\pm0.1$	$0.34\pm0.02$	$2.74\pm0.1$	$0.27\pm0.01$	$1.8\pm0.4$
YX07 (VTD)	10.02	$1.43\pm0.04$	$2.5\pm0.1$	$0.87\pm0.03$	$1.77\pm0.06$	$0.65\pm0.02$	$1.7\pm0.3$
YX08 (PM10)	10.8	$1.19\pm0.04$	$2.1\pm0.1$	$0.73\pm0.02$	$0.76\pm0.03$	$0.54\pm0.02$	$1.2\pm0.3$
YX09 (VTD)	6.04	$3.2\pm0.1$	$3.7\pm0.2$	$1.78\pm0.06$	$1.58\pm0.07$	$1.41\pm0.05$	$1.8\pm0.5$
YX10 (PM10)	6.77	$3.3\pm0.1$	$3.5\pm0.2$	$1.81\pm0.06$	$1.72\pm0.07$	$1.5\pm0.05$	$2.3\pm0.4$
YX11 (VTD)	9.59	$1.54\pm0.05$	$2.4\pm0.1$	$0.94\pm0.03$	$0.8\pm0.04$	$0.98\pm0.03$	$2.9\pm0.3$
YX12 (PM10)	10.92	$1.31\pm0.04$	$2.1\pm0.1$	$0.81\pm0.03$	$0.76\pm0.03$	$0.85\pm0.03$	$2.7\pm0.3$
YX13 (VTD)	5.71	$7.5\pm0.2$	$4.2\pm0.2$	$4.1\pm0.1$	$9.32\pm0.3$	$3.3\pm0.1$	$6.9\pm0.6$
YX14 (PM10)	6.64	$7.7\pm0.2$	$3.5\pm0.2$	$4.3\pm0.1$	$13.7\pm0.4$	$3.4\pm0.1$	$7.6\pm0.6$
YX15 (VTD)	10.85	$8.1\pm0.2$	$2.2\pm0.1$	$4.4\pm0.1$	$2.27\pm0.08$	$2.08\pm0.07$	$2.3\pm0.3$
YX16 (PM10)	11.75	$7.6\pm0.2$	$2\pm0.1$	$4.1\pm0.1$	$2.05\pm0.07$	$1.9\pm0.06$	$2.1\pm0.3$
YX17 (VTD)	6.53	$52\pm2$	$3.4\pm0.2$	$29.6\pm0.9$	$13.8\pm0.4$	$14.1\pm0.4$	$6.7\pm0.6$
YX18 (PM10)	7.12	$48\pm1$	$3.2\pm0.1$	$27.4\pm0.8$	$12.7\pm0.4$	$12.6\pm0.4$	$6.1\pm0.5$
YX19 (VTD)	10.33	$4.9\pm0.1$	$3\pm0.1$	$2.61\pm0.08$	$2.5\pm0.09$	$1.78\pm0.06$	$5.8\pm0.4$
YX20 (PM10)	10.96	$4.2\pm0.1$	$2.2\pm0.1$	$2.29\pm0.07$	$1.98\pm0.07$	$1.59\pm0.05$	$4.4\pm0.3$
YX21 (VTD)	5.48	$2.63\pm0.08$	$3.7\pm0.2$	$1.48\pm0.05$	$1.03\pm0.06$	$1.58\pm0.05$	$6.9\pm0.6$
YX22 (PM10)	6.08	$2.36\pm0.07$	$3.5\pm0.2$	$1.35\pm0.04$	$1.12\pm0.06$	$1.43\pm0.05$	$6.9\pm0.6$
YX23 (VTD)	10.76	$0.64\pm0.02$	$1.9\pm0.1$	$0.39\pm0.02$	$1.41\pm0.05$	$1.14\pm0.04$	$7.1\pm0.4$
YX24 (PM10)	11.95	$0.44\pm0.01$	$1.17\pm0.06$	$0.27\pm0.01$	$1.44\pm0.05$	$0.98\pm0.03$	$6.4\pm0.4$
YX25 (VTD)	6.79	$5.1\pm0.2$	$3.4\pm0.2$	$2.97\pm0.09$	$1.93\pm0.07$	$2.41\pm0.08$	$5.9\pm0.5$
YX26 (PM10)	7.65	$4.6\pm0.1$	$3.2\pm0.1$	$2.68\pm0.08$	$1.67\pm0.07$	$2.17\pm0.07$	$5.8\pm0.5$
YX27 (VTD)	10.73	$1.06\pm0.03$	$1.98\pm0.09$	$0.66\pm0.02$	$0.53\pm0.03$	$0.85\pm0.03$	$3.6\pm0.3$
YX28 (PM10)	11.8	$0.83\pm0.03$	$1.64\pm0.08$	$0.49\pm0.02$	$0.47\pm0.03$	$0.67\pm0.02$	$3.1\pm0.3$
YX29 (VTD)	16.66	$8.1\pm0.2$	$1.37\pm0.06$	$4.87\pm0.15$	$4.6\pm0.1$	$3.88\pm0.12$	$3.4\pm0.2$
YX30 (PM10)	19.25	$7.3\pm0.2$	$1.2\pm0.05$	$4.43\pm0.13$	$5.2\pm0.2$	$3.42\pm0.1$	$3\pm0.2$
YX31 (VTD)	5.92	$4.8\pm0.1$	$4.2\pm0.2$	$2.61\pm0.08$	$3\pm0.1$	$2.64\pm0.09$	$8.5\pm0.7$
YX32 (PM10)	7.49	$4.1\pm0.1$	$3.2\pm0.1$	$2.26\pm0.07$	$2.17\pm0.08$	$2.27\pm0.07$	$6.7\pm0.5$
YX33 (VTD)	8.31	$1.37\pm0.04$	$2.2\pm0.1$	$0.82\pm0.03$	$0.69\pm0.04$	$1.16\pm0.04$	$5.7\pm0.5$
YX34 (PM10)	9.15	$1.11\pm0.03$	$1.8\pm0.1$	$0.65\pm0.02$	$0.61\pm0.03$	$1.01\pm0.03$	$4.9\pm0.4$

# Table A2. Elemental air concentrations measured with ICP-AES, continued

Element	Ba	Li	Mn	Р	S	Sc	Sr
Wavelength (nm)	233.527	670.78	257.611	177.495	182.034	335.373	460.733
Analytical DL (ng)	) 0.001	0.0002	0.001	0.01	0.9	0.001	0.002
Field DL (ng)	0.02	0.002	0.1	0.2	85	—	0.05
Sample name	$\mathrm{ng}\mathrm{m}^{-3}$						
YX03	$6.2\pm0.3$	$0.53\pm0.03$	$7.3\pm0.4$	$56\pm 2$	$1.6\pm0.3$	$0.2\pm0.1$	$10.8\pm0.5$
YX04	$5.4\pm0.2$	$0.47\pm0.02$	$6.5\pm0.4$	$49\pm2$	$0.9\pm0.2$	<	$9.4\pm0.4$
YX05	$7.7\pm0.4$	$0.54\pm0.04$	$8.2\pm0.6$	$74\pm4$	$2.2\pm0.5$	$0.2\pm0.2$	$12.2\pm0.7$
YX06	$6.6\pm0.3$	$0.43\pm0.04$	$5.4\pm0.5$	$60\pm3$	$1.5\pm0.4$	<	$9.9\pm0.6$
YX07	$12.7\pm0.4$	$1.19\pm0.05$	$16.2\pm0.7$	$39\pm2$	$1.2\pm0.3$	$0.2\pm0.1$	$20.7\pm0.8$
YX08	$10.6\pm0.4$	$0.96\pm0.04$	$14.1\pm0.6$	$34\pm2$	$1.3\pm0.3$	$0.2\pm0.1$	$17.2\pm0.7$
YX09	$22.2\pm0.8$	$2.65\pm0.09$	$31\pm1$	$62\pm3$	$2.2\pm0.5$	$0.9\pm0.2$	$45\pm2$
YX10	$23.0\pm0.8$	$3.0\pm0.1$	$31\pm1$	$62\pm3$	$2.7\pm0.5$	$0.6\pm0.2$	$47\pm2$
YX11	$16.3\pm0.6$	$1.28\pm0.05$	$18.7\pm0.8$	$41\pm2$	$1.9\pm0.3$	$0.4\pm0.1$	$24.7\pm0.9$
YX12	$13\pm0.4$	$1.17\pm0.04$	$16.03\pm0.7$	$36\pm2$	$1.5\pm0.3$	$0.3\pm0.1$	$21.5\pm0.8$
YX13	$48\pm2$	$6.9\pm0.2$	$71\pm3$	$129\pm5$	$3.8\pm0.6$	$1.5\pm0.3$	$86\pm3$
YX14	$51\pm2$	$7.2\pm0.2$	$72\pm3$	$132\pm5$	$4.3\pm0.5$	$1.5\pm0.2$	$90\pm3$
YX15	$47\pm1$	$6.4\pm0.2$	$65\pm2$	$99\pm4$	$1.4\pm0.3$	$1.4\pm0.1$	$38\pm1$
YX16	$45\pm1$	$6.1\pm0.2$	$62\pm2$	$92\pm3$	$1.2\pm0.3$	$1.3\pm0.1$	$34\pm1$
YX17	$348 \pm 11$	$62\pm2$	$446\pm14$	$684\pm21$	$11.4\pm0.7$	$10.2\pm0.4$	$318\pm10$
YX18	$319\pm10$	$58\pm2$	$411\pm13$	$627\pm19$	$10.6\pm0.7$	$9.2\pm0.3$	$295\pm9$
YX19	$30\pm1$	$4.6\pm0.1$	$38\pm1$	$83\pm3$	$3.7\pm0.4$	$0.8\pm0.2$	$39\pm1$
YX20	$25.6\pm0.8$	$4.2\pm0.1$	$34\pm1$	$81\pm3$	$3.4\pm0.3$	$0.9\pm0.1$	$34\pm1$
YX21	$17.8\pm0.7$	$2.4\pm0.09$	$24\pm1$	$55\pm3$	$4.1\pm0.6$	$0.5\pm0.3$	$27\pm1$
YX22	$15.9\pm0.6$	$2.08\pm0.08$	$21\pm1$	$50\pm3$	$3.8\pm0.5$	$0.3\pm0.2$	$25\pm1$
YX23	$4.2\pm0.2$	$0.64\pm0.03$	$6.9\pm0.4$	$17\pm2$	$2.8\pm0.3$	$0.2\pm0.2$	$11.9\pm0.5$
YX24	$3.1\pm0.2$	$0.53\pm0.03$	$4.8\pm0.3$	$21\pm1$	$2.7\pm0.3$	<	$9.1\pm0.4$
YX25	$23.1\pm0.8$	$4.5\pm0.1$	$49\pm2$	$80\pm4$	$3.9\pm0.5$	$1.2\pm0.2$	$43\pm2$
YX26	$21.0\pm0.7$	$3.6\pm0.1$	$45\pm2$	$70\pm3$	$2.9\pm0.4$	$0.8\pm0.2$	$38\pm1$
YX27	$7.5\pm0.3$	$0.87\pm0.04$	$11.3\pm0.6$	$22\pm2$	$2.0\pm0.3$	$0.2\pm0.1$	$11.5\pm0.5$
YX28	$5.4\pm0.2$	$0.71\pm0.03$	$8.7\pm0.5$	$67\pm3$	$1.8\pm0.3$	$0.2\pm0.1$	$8.9\pm0.4$
YX29	$52\pm2$	$8.4\pm0.3$	$79\pm3$	$152\pm5$	$3.5\pm0.3$	$1.53\pm0.09$	$69\pm2$
YX30	$46\pm1$	$7.6\pm0.2$	$71\pm2$	$143\pm5$	$3.2\pm0.2$	$1.43\pm0.08$	$62\pm2$
YX31	$29\pm1$	$4.1\pm0.1$	$41\pm2$	$80\pm4$	$3.8\pm0.6$	$0.7\pm0.2$	$44\pm2$
YX32	$25.8\pm0.9$	$3.7\pm0.1$	$37\pm1$	$69\pm3$	$3.1\pm0.4$	$0.8\pm0.2$	$37\pm1$
YX33	$9.6\pm0.4$	$1.27\pm0.05$	$12.3\pm0.7$	$29\pm2$	$2.1\pm0.4$	$0.3\pm0.2$	$15.6\pm0.7$
YX34	$8.3 \pm 0.3$	$1.05\pm0.04$	$10.3\pm0.6$	$34 \pm 2$	$2.3\pm0.4$	$0.3\pm0.2$	$13.1\pm0.6$

Element	Ti	Zn	Zr
Wavelength (nm)	334.187	213.86	339.2
Analytical DL (ng)	0.01	0.001	0.003
Field DL (ng)	0.2	0.1	0.01
Sample name	$\mathrm{ng}\mathrm{m}^{-3}$	$\mathrm{ng}\mathrm{m}^{-3}$	$\mathrm{ng}\mathrm{m}^{-3}$
YX03	$29\pm2$	$9.1\pm0.4$	$1.3\pm0.3$
YX04	$26\pm1$	$11.3\pm0.5$	$1.1\pm0.2$
YX05	$28\pm2$	$11.9\pm0.6$	<
YX06	$20\pm2$	$5.7\pm0.4$	<
YX07	$76\pm3$	$16.7\pm0.6$	$3.1\pm0.3$
YX08	$64\pm2$	$61.7\pm2$	$2.6\pm0.3$
YX09	$177\pm 6$	$11\pm0.6$	$7.1\pm0.5$
YX10	$189\pm 6$	$9\pm0.5$	$8\pm0.5$
YX11	$85\pm3$	$25.3\pm0.9$	$4.1\pm0.3$
YX12	$71\pm3$	$25.5\pm0.9$	$3.3\pm0.3$
YX13	$430\pm14$	$29\pm1$	$17.3\pm0.7$
YX14	$438\pm14$	$17.7\pm0.7$	$17.5\pm0.7$
YX15	$452\pm14$	$19.4\pm0.7$	$17.8\pm0.6$
YX16	$431\pm13$	$12.1\pm0.5$	$16.7\pm0.6$
YX17	$3145\pm95$	$75\pm2$	$124\pm4$
YX18	$2871\pm87$	$71\pm2$	$111\pm3$
YX19	$275\pm9$	$14.4\pm0.6$	$10.8\pm0.5$
YX20	$228\pm7$	$15.1\pm0.6$	$9.5\pm0.4$
YX21	$157\pm 6$	$10.4\pm0.6$	$5.5\pm0.5$
YX22	$129\pm5$	$5.3\pm0.4$	$5\pm0.5$
YX23	$32\pm2$	$7.5\pm0.4$	$1.5\pm0.3$
YX24	$19\pm1$	$6.8\pm0.3$	$1.1\pm0.2$
YX25	$336\pm11$	$18.4\pm0.8$	$15.9\pm0.6$
YX26	$266\pm9$	$9.5\pm0.5$	$12.2\pm0.5$
YX27	$68\pm3$	$8.3\pm0.4$	$2.9\pm0.3$
YX28	$46\pm2$	$7.4\pm0.3$	$2.1\pm0.2$
YX29	$488\pm15$	$37\pm1$	$21.1\pm0.7$
YX30	$440\pm13$	$34\pm1$	$19.1\pm0.6$
YX31	$265\pm9$	$27\pm1$	$11\pm0.6$
YX32	$236\pm8$	$25.2\pm0.9$	$9.5\pm0.5$
YX33	$73\pm3$	$9.5\pm0.5$	$3.2\pm0.4$
YX34	$58\pm2$	$10\pm0.5$	$2.1\pm0.3$

Element	Be	Cd	Co	Cr	Cu	Мо	Ni
Isotope	9	111	59	52	63	95	60
Analytical DL (n	g) 0.02	0.04	0.04	0.4	0.1	0.3	3
Field DL (ng)	0.1	0.1	2	643	15	3	141
Sample name	$\mathrm{ng}\mathrm{m}^{-3}$	$\mathrm{ng}\mathrm{m}^{-3}$	$\mathrm{ng}\mathrm{m}^{-3}$	${\rm ngm^{-3}}$	$\mathrm{ng}\mathrm{m}^{-3}$	$\mathrm{ng}\mathrm{m}^{-3}$	${\rm ngm^{-3}}$
YX03	$0.018\pm0.003$	$0.06\pm0.01$	$0.19\pm0.02$	<	$1.28\pm0.1$	<	<
YX04	$0.012\pm0.002$	$0.06\pm0.01$	$0.13\pm0.02$	<	$1.31\pm0.1$	<	<
YX05	$0.03\pm0.01$	$0.09\pm0.01$	$0.27\pm0.04$	<	$1.6\pm0.2$	<	<
YX06	$0.015\pm0.004$	$0.08\pm0.01$	$0.15\pm0.03$	<	$1.3\pm0.1$	<	<
YX07	$0.05\pm0.01$	$0.15\pm0.01$	$0.36\pm0.03$	<	$3.4\pm0.2$	<	<
YX08	$0.03\pm0.01$	$0.14\pm0.01$	$0.31\pm0.03$	<	$2.9\pm0.2$	<	$1\pm 2$
YX09	$0.11\pm0.02$	$0.07\pm0.01$	$0.75\pm0.07$	<	$1.9\pm0.2$	<	<
YX10	$0.11\pm0.01$	$0.07\pm0.01$	$0.83\pm0.07$	<	$2.5\pm0.2$	<	$3\pm3$
YX11	$0.04\pm0.01$	$0.22\pm0.02$	$0.44\pm0.04$	<	$4.3\pm0.2$	$0.19\pm0.04$	$3\pm 2$
YX12	$0.04\pm0.01$	$0.2\pm0.02$	$0.35\pm0.03$	<	$3.1\pm0.2$	$0.15\pm0.04$	$2\pm 2$
YX13	$0.14\pm0.02$	$0.05\pm0.01$	$1.01\pm0.08$	<	$2\pm0.1$	$0.15\pm0.04$	$2\pm 2$
YX14	$0.28\pm0.04$	$0.08\pm0.01$	$2\pm0.2$	$10\pm3$	$4.1\pm0.3$	$0.35\pm0.07$	$6\pm4$
YX15	$0.25\pm0.03$	$0.04\pm0.005$	$1.7\pm0.1$	<	$3.1\pm0.2$	$0.19\pm0.04$	$3\pm 2$
YX16	$0.24\pm0.03$	$0.04\pm0.005$	$1.6\pm0.1$	$7\pm2$	$2.9\pm0.2$	$0.16\pm0.04$	$3\pm 2$
YX17	$1.7\pm0.2$	$0.24\pm0.02$	$11.4\pm0.7$	$48\pm4$	$17.6\pm0.8$	$1.5\pm0.2$	$26\pm 6$
YX18	$1.7\pm0.2$	$0.22\pm0.02$	$11.2\pm0.7$	$42\pm4$	$16.4\pm0.7$	$1.2\pm0.1$	$24\pm 6$
YX19	$0.17\pm0.02$	$0.06\pm0.01$	$1.11\pm0.08$	<	$2.3\pm0.2$	$0.17\pm0.05$	$4\pm 2$
YX20	$0.14\pm0.02$	$0.05\pm0.01$	$1.06\pm0.08$	$6\pm 2$	$2\pm0.1$	$0.18\pm0.04$	$4\pm 2$
YX21	$0.08\pm0.01$	$0.09\pm0.01$	$0.61\pm0.06$	<	$1.5\pm0.2$	<	$4\pm4$
YX22	$0.07\pm0.01$	$0.09\pm0.01$	$1.2\pm0.1$	<	$1.4\pm0.1$	$0.35\pm0.07$	$5\pm4$
YX23	$0.026\pm0.004$	$0.04\pm0.01$	$0.2\pm0.02$	<	$1.1\pm0.1$	<	$4\pm 2$
YX24	$0.014\pm0.003$	$0.04\pm0.005$	$0.15\pm0.02$	<	$1.1\pm0.1$	$0.12\pm0.03$	$3\pm 2$
YX25	$0.15\pm0.02$	$0.07\pm0.01$	$1.1\pm0.09$	<	$2.5\pm0.2$	$0.24\pm0.06$	$2\pm3$
YX26	$0.16\pm0.02$	$0.06\pm0.01$	$0.99\pm0.08$	<	$2.2\pm0.2$	$0.22\pm0.05$	$4\pm3$
YX27	$0.03\pm0.01$	$0.05\pm0.01$	$0.28\pm0.03$	<	$1.4\pm0.1$	$0.16\pm0.04$	$2\pm 2$
YX28	$0.028\pm0.004$	$0.05\pm0.01$	$0.22\pm0.02$	<	$1.1\pm0.1$	<	$2\pm 2$
YX29	$0.33\pm0.04$	$0.13\pm0.01$	$2\pm0.1$	$10\pm1$	$4.5\pm0.2$	$0.35\pm0.04$	$6\pm 2$
YX30	$0.29\pm0.04$	$0.12\pm0.01$	$1.8\pm0.1$	$9\pm1$	$4.8\pm0.2$	$0.33\pm0.04$	$6\pm 2$
YX31	$0.17\pm0.03$	$0.14\pm0.02$	$1.1\pm0.1$	<	$6.7\pm0.4$	<	$4\pm4$
YX32	$0.14\pm0.02$	$0.12\pm0.01$	$1.04\pm0.07$	<	$2.7\pm0.2$	<	$3\pm3$
YX33	$0.04\pm0.01$	$0.06\pm0.01$	$0.36\pm0.04$	<	$1.8\pm0.1$	<	$2\pm 2$
YX34	$0.03\pm0.01$	$0.06\pm0.01$	$0.36\pm0.04$	<	$1.6\pm0.1$	<	$2\pm 2$

# Table A4. Elemental air concentrations measured with ICP-MS.

# Table A5. Elemental air concentrations measured with ICP-MS, continued.

Element	Pb	Rb	Sb	Se	Tl	U	V
Isotope	208	85	121	77	205	238	51
Analytical DL (ng)	0.01	0.1	0.02	0.5	0.002	0.01	0.4
Field DL (ng)	2	4	0.6	_	0.1	0.04	2
Sample name	$\mathrm{ng}\mathrm{m}^{-3}$						
YX03	$4\pm1.1$	$1.04\pm0.05$	$0.6\pm0.04$	$0.4\pm0.1$	$0.017\pm0.003$	$0.04\pm0.01$	$2.8\pm0.1$
YX04	$3.3\pm0.9$	$0.98\pm0.05$	$0.55\pm0.04$	$0.29\pm0.09$	$0.014\pm0.003$	$0.03\pm0.01$	$2.9\pm0.1$
YX05	$3.4\pm0.9$	$1.11\pm0.07$	$0.51\pm0.04$	$0.18\pm0.09$	$0.011\pm0.003$	$0.03\pm0.01$	$2.2\pm0.1$
YX06	$3.3\pm0.9$	$1.05\pm0.06$	$0.49\pm0.04$	$0.17\pm0.08$	$0.012\pm0.003$	$0.04\pm0.01$	$2.1\pm0.1$
YX07	$4.9\pm1.3$	$2\pm0.09$	$1.2\pm0.08$	$0.14\pm0.06$	$0.023\pm0.004$	$0.07\pm0.01$	$3.3\pm0.2$
YX08	$4.4\pm1.2$	$1.63\pm0.07$	$1.12\pm0.07$	$0.14\pm0.06$	$0.02\pm0.004$	$0.05\pm0.01$	$2.7\pm0.1$
YX09	$3.4\pm1$	$4.0\pm0.2$	$0.43\pm0.03$	<	$0.03\pm0.01$	$0.14\pm0.03$	$5.6\pm0.3$
YX10	$3.5\pm0.9$	$4.3\pm0.2$	$0.38\pm0.03$	$0.09\pm0.07$	$0.03\pm0.01$	$0.14\pm0.03$	$5.9\pm0.3$
YX11	$5.9\pm1.5$	$2.3\pm0.1$	$1.11\pm0.07$	$0.3\pm0.09$	$0.027\pm0.005$	$0.08\pm0.02$	$3.6\pm0.2$
YX12	$5.2\pm1.4$	$2.01\pm0.09$	$1.03\pm0.07$	$0.22\pm0.07$	$0.024\pm0.004$	$0.07\pm0.01$	$3.1\pm0.1$
YX13	$2\pm0.6$	$5.6\pm0.2$	$0.28\pm0.02$	$0.12\pm0.06$	$0.03\pm0.01$	$0.13\pm0.03$	$6.4\pm0.3$
YX14	$4\pm1.1$	$11.7\pm0.5$	$0.53\pm0.04$	$0.3\pm0.1$	$0.07\pm0.01$	$0.28\pm0.06$	$13.6\pm0.6$
YX15	$3.1\pm0.8$	$8.7\pm0.3$	$0.4\pm0.03$	$0.18\pm0.06$	$0.05\pm0.01$	$0.28\pm0.06$	$11.2\pm0.5$
YX16	$2.6\pm0.7$	$8\pm0.3$	$0.36\pm0.03$	$0.15\pm0.06$	$0.05\pm0.01$	$0.24\pm0.05$	$10.3\pm0.5$
YX17	$15.3\pm4$	$63\pm2$	$0.45\pm0.03$	$0.4\pm0.1$	$0.33\pm0.05$	$1.9\pm0.4$	$80\pm4$
YX18	$13\pm3$	$60\pm2$	$0.4\pm0.03$	$0.4\pm0.1$	$0.31\pm0.05$	$1.6\pm0.3$	$77\pm3$
YX19	$3.2\pm0.9$	$5.7\pm0.2$	$0.22\pm0.02$	$0.5\pm0.2$	$0.05\pm0.01$	$0.16\pm0.04$	$11.9\pm0.5$
YX20	$2.6\pm0.7$	$4.9\pm0.2$	$0.21\pm0.02$	$0.6\pm0.2$	$0.05\pm0.01$	$0.14\pm0.03$	$10.7\pm0.5$
YX21	$2.3\pm0.7$	$3.3\pm0.1$	<	$0.9\pm0.2$	$0.03\pm0.01$	$0.11\pm0.02$	$11.2\pm0.5$
YX22	$2\pm0.6$	$3\pm0.1$	<	$0.5\pm0.1$	$0.03\pm0.01$	$0.1\pm0.02$	$9.9\pm0.5$
YX23	$1.5\pm0.4$	$1.02\pm0.05$	$0.23\pm0.02$	$0.6\pm0.2$	$0.01\pm0.002$	$0.02\pm0.01$	$9.7\pm0.4$
YX24	$1.5\pm0.4$	$0.83\pm0.05$	$0.21\pm0.02$	$0.8\pm0.2$	$0.011\pm0.002$	$0.017\pm0.005$	$9.2\pm0.5$
YX25	$4.2\pm1.1$	$5.9\pm0.2$	<	$0.72\pm0.2$	$0.05\pm0.01$	$0.18\pm0.04$	$11.8\pm0.5$
YX26	$3.7\pm1$	$5.6\pm0.2$	<	$0.5\pm0.2$	$0.05\pm0.01$	$0.14\pm0.03$	$11.6\pm0.5$
YX27	$2.7\pm0.7$	$1.37\pm0.06$	$0.34\pm0.02$	$0.5\pm0.1$	$0.017\pm0.003$	$0.05\pm0.01$	$7.4\pm0.3$
YX28	$2.6\pm0.7$	$1.09\pm0.05$	$0.33\pm0.02$	$0.7\pm0.2$	$0.017\pm0.003$	$0.03\pm0.01$	$6.7\pm0.3$
YX29	$6\pm 1$	$10.7\pm0.4$	$0.56\pm0.04$	$0.6\pm0.2$	$0.07\pm0.01$	$0.25\pm0.06$	$18.1\pm0.8$
YX30	$6\pm 1$	$10.4\pm0.4$	$0.54\pm0.03$	$0.5\pm0.1$	$0.06\pm0.01$	$0.24\pm0.05$	$16.8\pm0.8$
YX31	$5\pm 2$	$5.8\pm0.3$	<	$0.6\pm0.2$	$0.04\pm0.01$	$0.15\pm0.04$	$10.6\pm0.6$
YX32	$5\pm 1$	$5.4\pm0.2$	$0.33\pm0.02$	$0.5\pm0.1$	$0.04\pm0.01$	$0.14\pm0.03$	$9.3\pm0.4$
YX33	$3.3\pm0.9$	$1.80\pm0.08$	<	$0.8\pm0.2$	$0.024\pm0.004$	$0.05\pm0.01$	$6.2\pm0.3$
YX34	$2.8\pm0.8$	$1.46\pm0.07$	$0.26\pm0.02$	$0.8\pm0.2$	$0.017\pm0.003$	$0.04\pm0.01$	$5.3\pm0.3$

# Table A6. REE air concentrations measured with ICP-MS.

Element	La	Ce	Pr	Nd	Sm	Eu	Gd
Isotope	139	140	141	146	147	153	157
Analytical DL (1	ng) 0.01	0.004	0.001	0.01	0.01	0.01	0.004
Field DL (ng)	0.4	0.5	0.1	0.2	0.05	0.03	0.1
Sample name	$\mathrm{ng}\mathrm{m}^{-3}$						
YX03	$0.37\pm0.07$	$0.8\pm0.2$	$0.09\pm0.02$	$0.3\pm0.1$	$0.07\pm0.02$	$0.013\pm0.004$	$0.05\pm0.01$
YX04	$0.32\pm0.06$	$0.6\pm0.1$	$0.07\pm0.01$	$0.3\pm0.1$	$0.06\pm0.02$	$0.011\pm0.003$	$0.05\pm0.01$
YX05	$0.41\pm0.08$	$0.8\pm0.2$	$0.1\pm0.02$	$0.4\pm0.1$	$0.09\pm0.02$	$0.016\pm0.005$	$0.04\pm0.01$
YX06	$0.36\pm0.07$	$0.7\pm0.2$	$0.08\pm0.02$	$0.3\pm0.1$	$0.07\pm0.02$	$0.012\pm0.004$	$0.04\pm0.01$
YX07	$0.9\pm0.2$	$1.7\pm0.4$	$0.19\pm0.03$	$0.8\pm0.3$	$0.13\pm0.03$	$0.028\pm0.009$	$0.11\pm0.03$
YX08	$0.7\pm0.1$	$1.4\pm0.3$	$0.16\pm0.03$	$0.6\pm0.2$	$0.12\pm0.03$	$0.024\pm0.007$	$0.1\pm0.03$
YX09	$2.0\pm0.3$	$3.8\pm0.8$	$0.45\pm0.08$	$1.9\pm0.7$	$0.33\pm0.08$	$0.08\pm0.02$	$0.27\pm0.07$
YX10	$2.0\pm0.3$	$4.0\pm0.9$	$0.47\pm0.08$	$1.8\pm0.7$	$0.36\pm0.09$	$0.07\pm0.02$	$0.28\pm0.07$
YX11	$0.9\pm0.2$	$2\pm0.4$	$0.22\pm0.04$	$0.9\pm0.3$	$0.18\pm0.04$	$0.04\pm0.01$	$0.13\pm0.03$
YX12	$0.8\pm0.1$	$1.6\pm0.3$	$0.19\pm0.03$	$0.7\pm0.3$	$0.15\pm0.04$	$0.031\pm0.009$	$0.11\pm0.03$
YX13	$2.2\pm0.4$	$4.4\pm0.9$	$0.53\pm0.09$	$1.9\pm0.7$	$0.4\pm0.1$	$0.07\pm0.02$	$0.29\pm0.07$
YX14	$4\pm0.7$	$7\pm2$	$0.9\pm0.2$	$4\pm 1$	$0.7\pm0.2$	$0.15\pm0.05$	$0.6\pm0.1$
YX15	$4.7\pm0.8$	$10\pm 2$	$1.1\pm0.2$	$4\pm 2$	$0.8\pm0.2$	$0.16\pm0.05$	$0.6\pm0.1$
YX16	$4.2\pm0.7$	$9\pm 2$	$0.9\pm0.2$	$4\pm 1$	$0.7\pm0.2$	$0.14\pm0.04$	$0.6\pm0.1$
YX17	$31\pm 5$	$60\pm10$	$7\pm1$	$24\pm9$	$5\pm1$	$1\pm0.3$	$4.1 \pm 1$
YX18	$28\pm5$	$60\pm10$	$7\pm1$	$25\pm9$	$5\pm1$	$0.9\pm0.3$	$3.7\pm0.9$
YX19	$2.6\pm0.4$	$5\pm 1$	$0.6\pm0.1$	$2\pm0.7$	$0.4\pm0.1$	$0.08\pm0.02$	$0.33\pm0.08$
YX20	$2.5\pm0.4$	$5\pm 1$	$0.56\pm0.09$	$2.2\pm0.8$	$0.4\pm0.1$	$0.08\pm0.02$	$0.33\pm0.08$
YX21	$1.5\pm0.3$	$3.1\pm0.7$	$0.33\pm0.06$	$1.4\pm0.5$	$0.28\pm0.07$	$0.05\pm0.02$	$0.19\pm0.05$
YX22	$1.5\pm0.3$	$3\pm0.6$	$0.34\pm0.06$	$1.4\pm0.5$	$0.26\pm0.06$	$0.06\pm0.02$	$0.19\pm0.05$
YX23	$0.36\pm0.07$	$0.7\pm0.2$	$0.08\pm0.01$	$0.3\pm0.1$	$0.05\pm0.01$	$0.013\pm0.004$	$0.06\pm0.02$
YX24	$0.27\pm0.05$	$0.5\pm0.1$	$0.06\pm0.01$	$0.23\pm0.09$	$0.04\pm0.01$	$0.008\pm0.003$	$0.034\pm0.009$
YX25	$3.2\pm0.5$	$6\pm 1$	$0.7\pm0.1$	$3\pm 1$	$0.6\pm0.1$	$0.14\pm0.04$	$0.5\pm0.1$
YX26	$2.7\pm0.5$	$6\pm 1$	$0.6\pm0.1$	$2.5\pm1$	$0.5\pm0.1$	$0.1\pm0.03$	$0.4\pm0.1$
YX27	$0.7\pm0.1$	$1.3\pm0.3$	$0.15\pm0.03$	$0.6\pm0.2$	$0.12\pm0.03$	$0.026\pm0.008$	$0.08\pm0.02$
YX28	$0.52\pm0.09$	$1\pm0.2$	$0.12\pm0.02$	$0.5\pm0.2$	$0.09\pm0.02$	$0.02\pm0.006$	$0.07\pm0.02$
YX29	$4.6\pm0.8$	$9\pm 2$	$1.1\pm0.2$	$4\pm 1$	$0.8\pm0.2$	$0.15\pm0.04$	$0.6\pm0.2$
YX30	$4.3\pm0.7$	$9\pm 2$	$1.0\pm0.2$	$4\pm 1$	$0.8\pm0.2$	$0.15\pm0.04$	$0.6\pm0.1$
YX31	$2.7\pm0.5$	$5\pm1$	$0.6\pm0.1$	$2.1\pm0.8$	$0.5\pm0.1$	$0.09\pm0.03$	$0.34\pm0.09$
YX32	$2.2\pm0.4$	$4.5\pm1$	$0.51\pm0.09$	$1.9\pm0.7$	$0.37\pm0.09$	$0.08\pm0.02$	$0.26\pm0.07$
YX33	$0.8\pm0.1$	$1.7\pm0.4$	$0.19\pm0.03$	$0.8\pm0.3$	$0.14\pm0.04$	$0.028\pm0.009$	$0.12\pm0.03$
YX34	$0.7\pm0.1$	$1.3\pm0.3$	$0.14\pm0.02$	$0.6\pm0.2$	$0.1\pm0.03$	$0.021\pm0.006$	$0.08\pm0.02$

# Table A7. REE air concentrations measured with ICP-MS, continued.

Element	Tb	Dy	Но	Er	Tm	Yb	Lu
Isotope	159	163	165	166	169	172	175
Analytical DL (ng	g) 0.003	0.004	0.001	0.001	0.01	0.002	0.001
Field DL (ng)	0.02	0.05	0.01	0.02		0.04	0.01
Sample name	$\mathrm{ng}\mathrm{m}^{-3}$						
YX03	$0.007\pm0.002$	$0.05\pm0.01$	$0.009\pm0.002$	$0.027\pm0.007$	$0.004\pm0.002$	$0.019\pm0.005$	$0.003 \pm 0.0008$
YX04	$0.007\pm0.002$	$0.04\pm0.01$	$0.008 \pm 0.002$	$0.019\pm0.005$	$0.003\pm0.002$	$0.018\pm0.004$	$0.0019 \pm 0.0005$
YX05	$0.007\pm0.002$	$0.05\pm0.02$	$0.008 \pm 0.002$	$0.026 \pm 0.007$	$0.006\pm0.003$	$0.025\pm0.006$	$0.004\pm0.001$
YX06	$0.007\pm0.002$	$0.04\pm0.01$	$0.007\pm0.002$	$0.024\pm0.007$	$0.005\pm0.003$	$0.023\pm0.006$	$0.005\pm0.001$
YX07	$0.019\pm0.004$	$0.11\pm0.03$	$0.019\pm0.004$	$0.05\pm0.01$	$0.007\pm0.002$	$0.05\pm0.01$	$0.007\pm0.002$
YX08	$0.013\pm0.003$	$0.08\pm0.02$	$0.016\pm0.003$	$0.05\pm0.01$	$0.007\pm0.002$	$0.034\pm0.008$	$0.006\pm0.001$
YX09	$0.06\pm0.01$	$0.23\pm0.07$	$0.042\pm0.009$	$0.12\pm0.03$	$0.019\pm0.005$	$0.10\pm0.02$	$0.014\pm0.003$
YX10	$0.042\pm0.009$	$0.24\pm0.07$	$0.045\pm0.009$	$0.13\pm0.04$	$0.02\pm0.006$	$0.12\pm0.03$	$0.017\pm0.004$
YX11	$0.02\pm0.004$	$0.14\pm0.04$	$0.023\pm0.005$	$0.06\pm0.02$	$0.01\pm0.003$	$0.05\pm0.01$	$0.008 \pm 0.002$
YX12	$0.015\pm0.003$	$0.1\pm0.03$	$0.018\pm0.004$	$0.06\pm0.02$	$0.006\pm0.002$	$0.05\pm0.01$	$0.007\pm0.002$
YX13	$0.045\pm0.009$	$0.24\pm0.07$	$0.049\pm0.01$	$0.13\pm0.04$	$0.018\pm0.005$	$0.13\pm0.03$	$0.017\pm0.004$
YX14	$0.08\pm0.02$	$0.5\pm0.1$	$0.09\pm0.02$	$0.26\pm0.07$	$0.032\pm0.009$	$0.2\pm0.05$	$0.028 \pm 0.007$
YX15	$0.09\pm0.02$	$0.5\pm0.2$	$0.09\pm0.02$	$0.28\pm0.07$	$0.038\pm0.01$	$0.25\pm0.06$	$0.034\pm0.008$
YX16	$0.07\pm0.01$	$0.5\pm0.1$	$0.09\pm0.02$	$0.25\pm0.07$	$0.034\pm0.009$	$0.19\pm0.04$	$0.029\pm0.007$
YX17	$0.6\pm0.1$	$3\pm1$	$0.7\pm0.1$	$1.7\pm0.5$	$0.27\pm0.07$	$1.5\pm0.3$	$0.23\pm0.05$
YX18	$0.5\pm0.1$	$3\pm1$	$0.6\pm0.1$	$1.7\pm0.5$	$0.22\pm0.06$	$1.5\pm0.3$	$0.22\pm0.05$
YX19	$0.05\pm0.01$	$0.3\pm0.09$	$0.05\pm0.01$	$0.15\pm0.04$	$0.022\pm0.006$	$0.13\pm0.03$	$0.019\pm0.005$
YX20	$0.045\pm0.009$	$0.29\pm0.09$	$0.05\pm0.01$	$0.14\pm0.04$	$0.018\pm0.005$	$0.13\pm0.03$	$0.019\pm0.004$
YX21	$0.029\pm0.006$	$0.18\pm0.06$	$0.033\pm0.007$	$0.1\pm0.03$	$0.012\pm0.004$	$0.09\pm0.02$	$0.011\pm0.003$
YX22	$0.032\pm0.007$	$0.18\pm0.06$	$0.034\pm0.007$	$0.09\pm0.02$	$0.013\pm0.004$	$0.08\pm0.02$	$0.013\pm0.003$
YX23	$0.004\pm0.001$	$0.05\pm0.01$	$0.009\pm0.002$	$0.024\pm0.007$	$0.004\pm0.002$	$0.021\pm0.005$	$0.0022 \pm 0.0006$
YX24	$0.005\pm0.001$	$0.03\pm0.01$	$0.006\pm0.001$	$0.011\pm0.003$	$0.002\pm0.001$	$0.017\pm0.004$	$0.0024 \pm 0.0007$
YX25	$0.06\pm0.01$	$0.4\pm0.1$	$0.08\pm0.01$	$0.21\pm0.06$	$0.03\pm0.008$	$0.19\pm0.04$	$0.028 \pm 0.007$
YX26	$0.06\pm0.01$	$0.32\pm0.1$	$0.06\pm0.01$	$0.17\pm0.04$	$0.026\pm0.007$	$0.15\pm0.03$	$0.022\pm0.005$
YX27	$0.012\pm0.003$	$0.09\pm0.03$	$0.015\pm0.003$	$0.05\pm0.01$	$0.006\pm0.002$	$0.03\pm0.007$	$0.005\pm0.001$
YX28	$0.009 \pm 0.002$	$0.06\pm0.02$	$0.012\pm0.002$	$0.035\pm0.01$	$0.005\pm0.002$	$0.028 \pm 0.007$	$0.004\pm0.001$
YX29	$0.09\pm0.02$	$0.5\pm0.2$	$0.1\pm0.02$	$0.26\pm0.07$	$0.035\pm0.009$	$0.24\pm0.05$	$0.033\pm0.008$
YX30	$0.08\pm0.02$	$0.5\pm0.2$	$0.09\pm0.02$	$0.27\pm0.07$	$0.035\pm0.009$	$0.22\pm0.05$	$0.032\pm0.008$
YX31	$0.05\pm0.01$	$0.3\pm0.1$	$0.05\pm0.01$	$0.17\pm0.04$	$0.022\pm0.006$	$0.14\pm0.03$	$0.022\pm0.005$
YX32	$0.039 \pm 0.008$	$0.28\pm0.09$	$0.048 \pm 0.01$	$0.12\pm0.03$	$0.019\pm0.005$	$0.11\pm0.02$	$0.014\pm0.003$
YX33	$0.014\pm0.003$	$0.08\pm0.03$	$0.018\pm0.004$	$0.05\pm0.01$	$0.007\pm0.003$	$0.05\pm0.01$	$0.007\pm0.002$
YX34	$0.013\pm0.003$	$0.07\pm0.02$	$0.015\pm0.003$	$0.04\pm0.01$	$0.006\pm0.002$	$0.035\pm0.008$	$0.006\pm0.001$

# 255 Appendix B: Geostandard recovery rates

Recoveries of geostandards MAG-1 in Table B1 and SCO-1 in Table B2.
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Analysed by	Measured	Recovery rate	Analysed by	Measured	Recovery rate
ICP-AES	$\mu {\rm g} {\rm g}^{-1}$		ICP-MS	$\mu {\rm g}/{\rm g}$	
Al	80 000	92%	Be	2.99	93%
Ba	433	90%	Rb	160	107%
Ca	9 300	95%	Мо	1.22	76%
Fe	49 000	102%	Cd	0.265	130%
Κ	34 000	110%	Sb	0.873	91%
Li	100	130%	Pb	25.1	105%
Mg	19 000	106%	U	2.69	100%
Mn	784	130%	V	159	110%
Na	19 000	68%	Cr	103	107%
Р	826	120%	Со	23	120%
Sc	16.7	98%	Ni	80.3	150%
Sr	122	82%	Cu	29	97%
Ti	3.82 000	85%	As	8.31	90%
Zn	187	140%	La	44	110%
Zr	144	110%	Ce	91	109%
			Pr	10.7	120%
			Nd	41	115%
			Sm	7.7	109%
			Eu	1.49	98%
			Gd	6.2	104%
			Tb	0.87	90%
			Dy	4.8	90%
			Но	0.88	82%
			Er	2.5	103%
			Tm	0.34	76%
			Yb	2.3	82%
			Lu	0.33	76%

**Table B1.** MAG–1 recovery rates. Elements have a recovery rate between 68% and 130%, except for Zn and Ni. The very low amount of CRM used (<10 mg) could explain the difference observed in recovery rates, because subsampling heterogeneity is possible with such small amounts. Zn and Ni are overestimated, probably due to contamination.

Analysed by	Measured	Recovery rate	Analysed by	Measured	Recovery rate
ICP-AES	$\mu \mathrm{g}  \mathrm{g}^{-1}$		ICP-MS	$\mu \mathrm{g}  \mathrm{g}^{-1}$	
Al	58 500	81%	Be	1.87	104%
Ba	426	75%	Rb	135	120%
Ca	13 100	70%	Мо	1.46	104%
Fe	30 300	84%	Sb	2.73	109%
Κ	22 100	96%	Pb	32.6	105%
Li	50	110%	V	160	120%
Mg	13 100	80%	Cr	76.3	110%
Mn	340	83%	Со	12.7	120%
Na	10 100	150%	Ni	28.6	106%
Р	827	90%	Cu	31.6	109%
Sc	10.06	92%	As	12.2	101%
Sr	127	75%	La	32	104%
Ti	2.84	75%	Ce	63	98%
Zn	117	120%	Pr	7.7	110%
Zr	129	81%	Nd	29	110%
			Sm	5.6	102%
			Eu	1.20	113%
			Gd	4.6	101%
			Tb	0.66	87%
			Dy	3.8	96%
			Но	0.72	76%
			Er	2.1	82%
			Tm	0.31	72%
			Yb	2.0	84%
			Lu	0.31	79%

**Table B2.** SCO–1 recovery rates. Elements, except Na (recovery rate= 150%), have a recovery rate between 70% and 130%. The very low amount of CRM used (<10 mg) could explain the difference observed in recovery rates, because subsampling heterogeneity is possible with such small amounts.

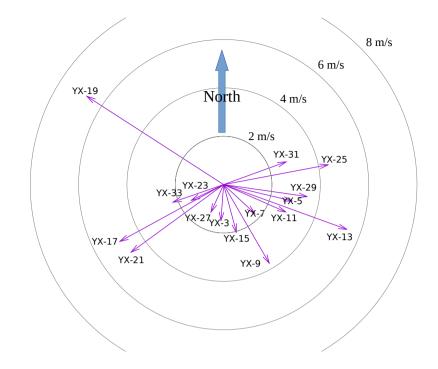
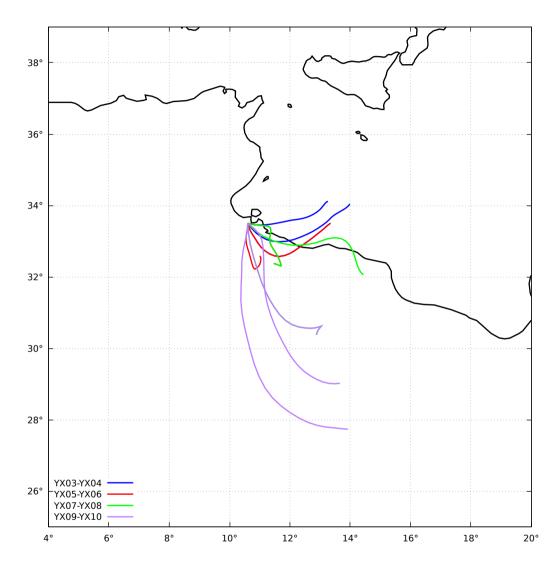


Figure C1. Vector representing local wind conditions at the sampling station. The length of the vector represents wind speed average, and its angle indicates the average direction during each sampling period.

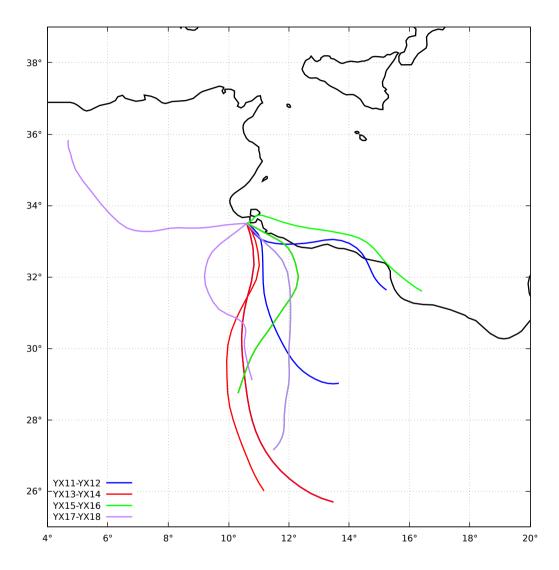
# Appendix C: Local meteorological conditions and air trajectories

Wind speed and direction are measured continuously at the sampling location, and backward air trajectories are calculated using the on-line facility at NOAA HYSPLIT model web pages (Stein et al., 2015; Rolph et al., 2017). Trajectories for a 24-hour period are calculated every 6 hours (at midnight, 6 AM, noon, and 6 PM).

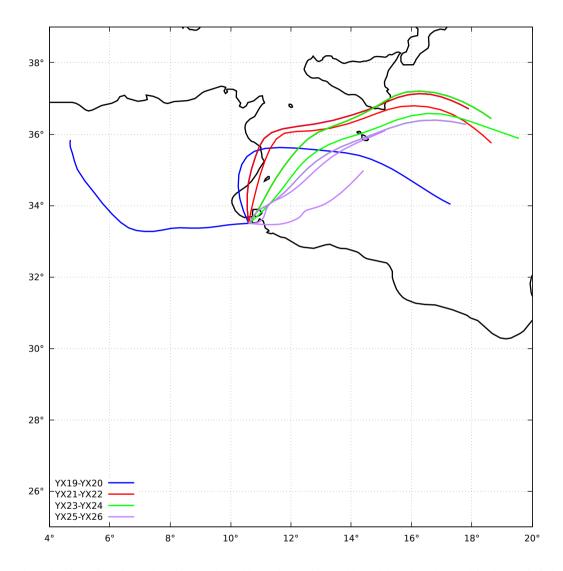
260



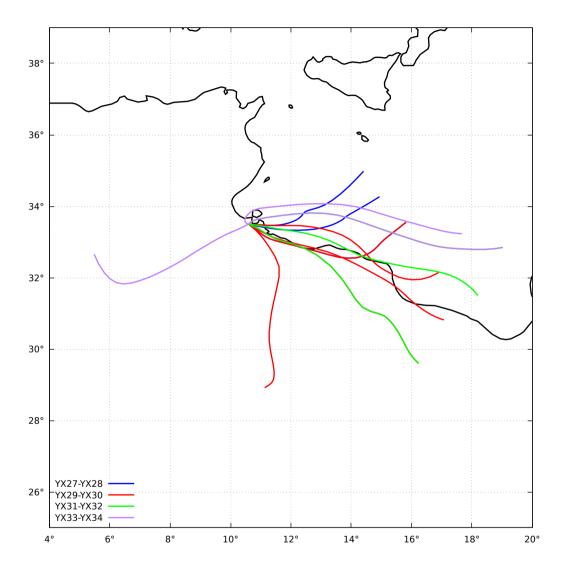
**Figure C2.** Backward-trajectories of sample pairs YX03-YX04, YX05-YX06, YX07-YX08, and YX09-YX10. The x axis is longitude, the y axis is latitude. Two or three trajectories are associated with a given sample pair of  $\approx$ 12 h duration.



**Figure C3.** Backward-trajectories of sample pairs YX11-YX12, YX13-YX14, YX15-YX16, and YX17-YX18. The x axis is longitude, the y axis is latitude. Two or three trajectories are associated with a given sample pair of  $\approx$ 12 h duration.



**Figure C4.** Backward-trajectories of sample pairs YX19-YX20, YX21-YX22, YX23-YX24, and YX25-YX26. The x axis is longitude, the y axis is latitude. Two or three trajectories are associated with a given sample pair of  $\approx$ 12 h duration.



**Figure C5.** Backward-trajectories of sample pairs YX27-YX28, YX29-YX30, YX31-YX32, and YX33-YX34. The x axis is longitude, the y axis is latitude. Two or three trajectories are associated with a given sample pair of  $\approx$ 12 h duration, except for the pair YX29-YX30, for which four trajectories are necessary because the sampling duration was 24 hours.

# **Appendix D:** Mass comparison statistical parameters

	Slope	Intercept ( $\mu g  m^{-3}$ )
	95% confidence interval	95% confidence interval
VTD = f(TEOM)	[0.78, 1.16]	[-18, +15]
PM10 = f(TEOM)	[0.79, 1.07]	[-17, +8]
VTD = f(PM10)	[0.96, 1.10]	[-1, +9]

Table D1. Optimal slope and intercept using orthogonal regressions, including the heavy loaded sample

	Slope	Intercept $(\mu g  m^{-3})$		
	95% confidence interval	95% confidence interval		
VTD = f(TEOM)	[0.76, 1.20]	[-8, +19]		
PM10 = f(TEOM)	[0.77, 1.11]	[-9, +10]		
VTD = f(PM10)	[0.94, 1.12]	[-0.2, +11]		

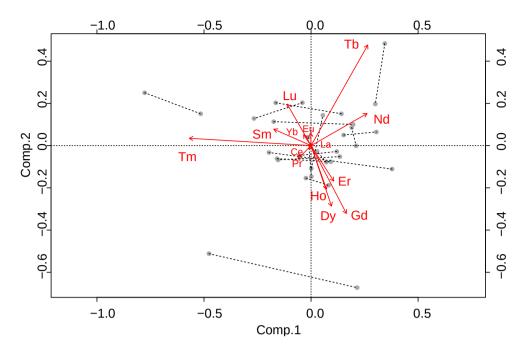
Table D2. Optimal slope and intercept using orthogonal regressions, excluding the heavy loaded sample

Sample name	[Total VTD]	[Sea salt]	[Crustal]	[Calcium species]
	$\mu \rm gm^{-3}$	$\mu \rm gm^{-3}$	$\mu \rm gm^{-3}$	$\mu { m g}{ m m}^{-3}$
YX03	$19\pm1$	$3\pm1$	$9\pm0.3$	$8\pm0.4$
YX05	$27\pm2$	$7\pm2$	$10\pm0.3$	$10 \pm 1$
YX07	$33\pm2$	$5\pm 1$	$20\pm1$	$8\pm 2$
YX09	$63\pm4$	$5\pm 1$	$45\pm1$	$12\pm3$
YX11	$40\pm1$	$9\pm1$	$22\pm1$	$9\pm1$
YX13	$140\pm4$	$20\pm2$	$106\pm3$	$15\pm1$
YX15	$124\pm4$	$5\pm 1$	$115\pm3$	$5\pm1$
YX17	$769\pm22$	$7\pm2$	$729\pm22$	$33\pm2$
YX19	$99\pm3$	$18\pm1$	$68\pm2$	$13 \pm 1$
YX21	$75\pm3$	$22\pm2$	$37\pm1$	$16\pm1$
YX23	$41\pm2$	$23\pm1$	$9\pm0.3$	$10 \pm 0.4$
YX25	$103\pm3$	$18\pm2$	$71\pm2$	$14 \pm 1$
YX27	$35\pm1$	$11 \pm 1$	$15\pm0.5$	$8\pm0.5$
YX29	$133\pm4$	$9\pm1$	$115\pm3$	$10 \pm 0.4$
YX31	$108\pm3$	$26\pm2$	$67\pm2$	$15\pm1$
YX33	$45\pm2$	$18\pm1$	$19\pm1$	$8 \pm 1$

 Table E1. VTD aerosol mass concentrations derived from chemical analyses with associated analytical uncertainties (95% confidence interval).

Sample name	[Total VTD]	[Sea salt]	[Crustal]	[Calcium species]
	$\mu \rm gm^{-3}$	$\mu \rm gm^{-3}$	$\mu \rm gm^{-3}$	$\mu { m gm}^{-3}$
YX04	$17\pm2$	$4\pm1$	$8\pm0.2$	$5\pm1.6$
YX06	$22\pm3$	$6\pm 1$	$8\pm0.3$	$8\pm3$
YX08	$28\pm2$	$4\pm1$	$17\pm1$	$7\pm2$
YX10	$66\pm2$	$6\pm 1$	$46\pm1$	$13 \pm 1$
YX12	$34\pm2$	$8\pm1$	$18\pm1$	$7\pm2$
YX14	$145\pm4$	$22\pm2$	$108\pm3$	$14 \pm 1$
YX16	$116\pm3$	$5\pm 1$	$107\pm3$	$4\pm 0$
YX18	$711\pm20$	$6\pm 2$	$674\pm20$	$31\pm2$
YX20	$84\pm2$	$13\pm1$	$60\pm 2$	$12 \pm 1$
YX22	$70\pm2$	$22\pm2$	$33\pm1$	$15\pm1$
YX24	$35\pm1$	$21\pm1$	$6\pm0.2$	$8\pm0.3$
YX26	$94\pm3$	$17\pm2$	$65\pm2$	$11 \pm 1$
YX28	$29\pm1$	$10\pm1$	$12\pm0.4$	$7\pm0.4$
YX30	$120\pm3$	$8\pm1$	$104\pm3$	$9\pm0.3$
YX32	$91\pm3$	$21\pm2$	$58\pm2$	$12 \pm 1$
YX34	$40 \pm 1$	$16 \pm 1$	$16\pm0.5$	$8\pm0.5$

Table E2. PM10 aerosol mass concentrations derived from chemical analyses with associated analytical uncertainties (95% confidence interval).



**Figure F1.** REE Biplot for two VTD simulation results. Percentages of variability explained by the first two components are 27% and 19%, a total of 46%.

Appendix F: REE biplot simulated with the observed analytical uncertainty.

265 Competing interests. No competing interest are present

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280

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