

# 1 Retrieval of UV-Visible aerosol absorption using AERONET and 2 OMI-MODIS synergy: Spatial and temporal variability across major 3 aerosol environments

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10 **Abstract.** Measuring spectral aerosol absorption remains a challenging task in aerosol studies, especially in the  
11 UV region, where ~~the~~ ground and airborne measurements are sparse. In this research, we introduce an algorithm  
12 that synergizes ground measurements with satellite observations for the derivation of spectral single scattering  
13 albedo (SSA,  $\omega_o$ ) of aerosols in the UV to visible range (340-670 nm). The ~~approach~~ physical basis consists in  
14 explaining satellite measured near-UV radiances (340, 354 and 388 nm) by the Ozone Monitoring Instrument  
15 (OMI), and visible radiances (466 and 646 nm) by MODerate Imaging Spectrometer (MODIS), in terms of  
16 ground-based Aerosol Robotic Network (AERONET) measurements of total column extinction aerosol optical  
17 depth (AOD,  $\tau$ ), and retrieved total column wavelength dependent SSA using radiative transfer calculations.  
18 Required information on aerosol particle size distribution is taken from an AERONET-based seasonal  
19 climatology specifically developed for this project. ~~This~~ The inversion procedure is applied over 110 AERONET  
20 sites distributed worldwide, for which continuous, long-term AERONET measurements are available. Using the  
21 derived data set we present seasonal and regional climatology of  $\omega_o(\lambda)$  for carbonaceous, dust and  
22 urban/industrial aerosol types. The UV-Visible spectral dependence of  $\omega_o$  obtained for the three major aerosol  
23 types from the synergy algorithm is found to be consistent with the insitu measurements reported in the literature.  
24 A comparison to standard AERONET SSA product at ~~646-440~~ nm shows absolute differences within 0.03 (0.05)  
25 for 40% (~~59~~65%) of the compared observations. The derived aerosol  $\omega_o(\lambda)$  data set provides a valuable addition  
26 to the existing aerosol absorption record from AERONET by extending the absorption retrieval capability to the  
27 near-UV region. The combined UV-Visible data set, in addition to improving our understanding of spectral  
28 aerosol absorption properties, also offers wavelength-dependent dynamic aerosol absorption models for use in the  
29 satellite-based aerosol retrieval algorithms.

## 30 1 Introduction

31 Atmospheric aerosols play a significant role in the Earth's climate system through scattering and absorption of  
32 solar radiation, thus capable of perturbing radiation budget. The ratio of the amount of the light scattering to the  
33 total extinction referred to as single scattering albedo (SSA,  $\omega_0$ ) is a fundamental variable used to gauge the  
34 absorbing nature of aerosols. Mie-theory indicates  $\omega_0$  equals to one for purely scattering aerosols and less than  
35 one towards zero for increasingly absorbing nature of aerosols. Studies show that the estimates of net aerosol  
36 radiative forcing is sensitive to the aerosol  $\omega_0$ , and small changes to it could potentially alter the forcing on  
37 atmosphere (Chyacutelek and Coakley, 1974; Hansen et al., 1997). Models-General circulation models are often  
38 fed with essential aerosol properties to estimate the forcing on the atmosphere. These properties include aerosol  
39 optical depth (AOD,  $\tau$ ), complex refractive index, and phase function. Here, the knowledge on spectral  
40 dependence of such properties is crucial in quantifying the overall effects of aerosols. For example, absorbing  
41 aerosols can lead up to a 50% increase-decrease in the near-UV irradiance compared to the similar load of only  
42 scattering aerosols in the atmosphere (Bais et al., 2005). A report by Intergovernmental Panel on Climate Change  
43 suggests that the lack of spectral aerosol absorption is one of the major contributors leading to significant  
44 uncertainties in quantifying the net aerosol radiative effects on the Earth's climate (IPCC, 2013).

45  
46 Developments in ground-based and satellite aerosol retrieval techniques have greatly improved our understanding  
47 of atmospheric aerosols over the last two decades. However, knowledge on spectral aerosol absorption properties  
48 is limited due to difficulties in measurements (e.g., Heintzenberg et al., 1997) and larger uncertainties in remote  
49 sensing retrievals (e.g., Dubovik et al., 2000). Direct measurements of aerosol absorption can be obtained by  
50 using instruments that measure scattering and extinction coefficients. Such measurements are limited to discrete  
51 wavelengths and associated with few ground stations, laboratory measurements, or airborne field campaigns. In  
52 addition insitu techniques often require making corrections of measurements to overcome instrumental challenges  
53 (e.g., Weingartner et al., 2003; Virkkula et al., 2005; Collaud Coen et al., 2010). Aerosol absorption can also be  
54 inferred from the sky radiance and extinction measurements that rely on fitting ground observations to radiative  
55 transfer calculations (Nakajima et al., 1996; Dubovik et al., 1998; Cattrall et al., 2003). The accuracy (or  
56 uncertainty) of the aerosol absorption retrieval through ground-based techniques primarily relies on the  
57 calibration of measurements, while surface characterization has secondary effect on the overall accuracy of the  
58 retrievals. Detailed reviews of measurements and techniques to retrieve aerosol absorption are available in  
59 several papers (e.g., Clarke et al., 1967; Bond and Bergstrom, 2006; Moosmüller et al., 2009). Among the

60 ground-based sensors, currently, AERONET provides long-term aerosol absorption record at four discrete  
61 wavelengths from the visible (Vis) to near Infrared (NIR) spectrum over many sites distributed worldwide.  
62 Known limitation of the currently available AERONET inversion product (Version 3) is the lack of single  
63 scattering albedo at near-UV wavelengths, and the aerosol load threshold ( $\tau_{440} > 0.4$ ) required to obtain reliable  
64 absorption in Vis-NIR spectrum. Like AERONET, a network of ground-based radiation measurement network  
65 with sites in Asia and Europe referred as SKYNET (sky radiometer network) provides aerosol optical depth and  
66 single scattering albedo in the near UV-NIR spectrum (Nakajima et al., 2007). However, the accuracy of the  
67 SKYNET aerosol absorption product is affected by temporally constant and spectrally invariant surface  
68 reflectance used in the inversion procedure (Jethva and Torres, 2019). These limitations incur knowledge gaps in  
69 the complete characterization of aerosol absorption as a function of both wavelength and aerosol load.~~The known~~  
70 ~~limitation of AERONET absorption product is the lack of near-UV wavelengths, besides higher aerosol load ( $\tau_{440}$~~   
71  ~~$> 0.4$ ) and high solar elevation angle required to obtain reliable absorption in Vis-NIR spectrum. These~~  
72 ~~limitations make it imperative to look for alternate data retrieval methods or sources to fill the knowledge gap.~~

73

74 For a few decades now, satellite remote sensing is used as an essential tool to gain a global perspective of  
75 aerosols distribution in the atmosphere. The physical basis of satellite aerosol retrievals is that under cloud-free  
76 conditions after accounting for Rayleigh (molecular) scattering and gaseous absorption effects, the upwelling top-  
77 of-the-atmosphere (TOA) reflectance is a function of aerosol optical depth, particle sizes and composition (i.e.,  
78 complex refractive index). Mathematically for a cloud-free atmosphere overlying an Lambertian surface the  
79 upwelling TOA reflectance ( $L_{TOA}$ ) received by a nadir viewing satellite can be expressed in normalized units as  
80 (Chandrasekhar, 1960):

$$81 \quad L_{TOA}(\theta, \varphi, \lambda) = L_0(\theta, \varphi, \lambda) + \frac{\rho(\lambda) \cdot T(\tau, \theta, \lambda)}{(1 - s(\lambda) \cdot \rho(\lambda))} \quad (1)$$

82 where  $\theta$ ,  $\varphi$ , and  $\lambda$  are the zenith, azimuthal angles of the direction of propagation and wavelength of light,  $\tau$  is the  
83 optical thickness of the atmosphere,  $\rho$  is the surface reflectivity,  $T$  is the total direct and diffuse transmittance of  
84 the light in the atmosphere, and  $s(\lambda)$  is the spherical albedo of the atmosphere when it is illuminated from below.  
85 The first and second term on the right side of the equation represents the atmospheric path radiance and the  
86 amount of light that is reflected to the sensor after encountering the surface, respectively. The satellite measured  
87 TOA reflectances are sensitive to both  $\tau$  and  $\omega_0$ , in addition to the surface reflectance. Therefore, separating the  
88 contributions of atmosphere and surface is of utmost importance to retrieve aerosols from satellite measurements.  
89 For satellites with single-view measurements, aerosol retrieval algorithms rely on prior assumptions on particle

90 ~~sizes (phase function) and  $\omega_0$  to retrieve  $\tau$ .~~ Most satellite aerosol retrieval techniques require making assumptions  
91 ~~on particle sizes (phase function) and  $\omega_0$  to retrieve  $\tau$ .~~ On the other hand, several efforts have been made to  
92 estimate aerosol  $\omega_0$  from direct satellite measurements at visible wavelengths (e.g., Kaufman, 1987; Kaufman et  
93 al., 2002; Satheesh and Srinivasan, 2005; Zhu et al., 2011) and near-UV wavelengths ([Torres et al., 1998, 2007,](#)  
94 [2013](#)~~Torres et al., 2007, 2013~~). However, the variety of natural surface types, choice of wavelengths, and aerosol  
95 models pose limitations on such techniques. In terms of wavelength, enhanced molecular scattering in the near-  
96 UV region acts as a strong attenuating background and helps identify absorbing aerosols. However, to retrieve  
97 aerosol absorption using near-UV measurements, quantitative information on aerosol layer height (ALH) is  
98 required. ~~The e~~Existing satellite aerosol retrieval techniques that rely on observations in the visible spectrum  
99 assume a temporally constant value of  $\omega_0$  that varies regionally (Remer et al., 2005), and for a few algorithms, it  
100 is still assumed wavelength-independent (Hsu et al., 2013)~~value, and for a few sensors, it is still assumed~~  
101 ~~wavelength independent.~~—A review on the commonly used satellite aerosol products singled out aerosol  
102 absorption as an inherent problem common to all sensors (Li et al., 2009). Studies using the evolving ground-  
103 based aerosol record provides evidence that satellite retrieved  $\tau$  can lead to large biases if the assumed aerosol  
104 imaginary index, which derives  $\omega_0$  is wavelength-independent (Jethva and Torres, 2011), and ~~constant seasonally~~  
105 invariant (Lyapustin et al., 2011; Eck et al., 2013). These studies highlight the importance of using wavelength-  
106 dependent aerosol  $\omega_0$  and account for its spatial and temporal variability in the assumptions made for satellite  
107 retrieval of aerosol products.

108

109 In the past, few studies used both ground and satellite measurements to retrieve aerosol absorption properties. Li  
110 et al., (1999) used visible band radiances from AVHRR and in situ measured  $\tau$  during SCAR-B experiment to  
111 derive absorption from biomass burning aerosols. Sinyuk et al., (2003) used UV-radiances from TOMS and  
112 aerosol extinction from AERONET to derive the imaginary refractive index of dust particles over a few stations  
113 in the Sahara belt region. Lee et al., (2007) estimated the aerosol SSA across a few stations over China using  
114 combined ground and satellite (MODIS) measurements at visible wavelengths. Nonetheless, these studies are  
115 limited and do not provide a comprehensive characterization of absorbing aerosols from the UV-Visible  
116 spectrum.

117

118 The objectives of the present work are to derive columnar aerosol  $\omega_0(\lambda)$ , and its spectral dependence in the UV-  
119 Visible part of the spectrum. ~~This approach~~ The proposed inversion procedure makes use of ~~the~~ AERONET  
120 measured wavelength dependent  $\tau$ , and ~~derived-retrieved~~ particle size distribution in conjunction with satellite

121 ~~measured radiances at UV and visible wavelengths by A-train constellation sensors Aqua-MODIS and Aura-OMI~~  
122 ~~(Ozone Monitoring Instrument). from AERONET in an inversion procedure. The A-train constellation of~~  
123 ~~satellites that includes OMI and MODIS makes routine TOA measurements from the UV-Visible spectrum. The~~  
124 ~~n~~Near-simultaneous measurements from A-train satellites~~these sensors~~ provide an excellent opportunity to  
125 combine satellite and ground measurements during the overpass times (local ~~noon~~time, ~13:30 hrs) over the  
126 AERONET sites.

127  
128 The organization of the remaining manuscript is as follows: section 2 describes the ground-based and satellite  
129 measurements used in this work. Section 3 describes the methodology adopted to derive aerosol  $\omega_o(\lambda)$ . Section 4  
130 provides the sensitivity ~~analysis to estimate the expected accuracy of tests conducted to estimate the error~~  
131 ~~incurred in~~ the proposed aerosol absorption retrievals. ~~Section 5 presents a comparison of retrieved SSA to the~~  
132 ~~AERONET aerosol absorption product. Section 6 presents the seasonal variability in regional aerosol  $\omega_o(\lambda)$~~   
133 ~~derived for sites distributed across major aerosol environments worldwide. Section 7 provides a discussion of the~~  
134 ~~regional variability of the UV-Vis aerosol absorption product derived in this work. Section 5 presents the seasonal~~  
135 ~~variability in aerosol  $\omega_o(\lambda)$  derived for sites distributed across major aerosol environments worldwide. Section 6~~  
136 ~~provides a discussion of the regional aerosol absorption models derived in this work. Section 7 presents a~~  
137 ~~comparison of retrieved SSA with the AERONET absorption product.~~ Finally, section 8 provides a summary of  
138 the work, along with the key findings and outlook for further studies.

## 139 **2 Data sets**

140 The details of ground-based and satellite data sets used in this work are provided in Table 1. Our usage of  
141 satellite data is strictly limited to the TOA reflectances, ~~the associate viewing satellite-sun~~ geometry, and other  
142 ancillary information ~~such as~~ (quality flags, ~~and aerosol index. type~~) ~~but not the aerosol retrievals ( $\tau$ ) themselves~~  
143 ~~from OMI and MODIS.~~

### 144 **2.1 AERONET**

145 ~~Aerosol Robotie Network (AERONET)~~ employs an automatic sun-tracking photometer (~~CIMEL Electronique~~  
146 ~~CE-318~~) to measure sun and sky radiances (Holben et al., 1998). The direct sun measurements are made ~~with a~~  
147 ~~1.2° full field of view~~ at nine nominal wavelengths of 340, 380, 440, 500, 675, 870, 940, ~~1020~~ and ~~1020-1640~~ nm  
148 typically for every ~~~5 to~~ 15 minutes interval. ~~Columnar extinction  $\tau$  is computed from these measurements for all~~

wavelengths except for the 940 nm, which is used to retrieve water vapour amounts. The extinction  $\tau$  obtained from these measurements has an estimated uncertainty  $\pm 0.01$  ( $\pm 0.02$ ) at the visible (near-UV) wavelengths, primarily due to calibration uncertainty (Eck et al., 1999). The currently available AERONET Version 3 Level 2 AOD product uses improved cloud screening and quality checks to provide reliable data to the user community (Giles et al., 2019). The extinction  $\tau$  obtained from these measurements are accurate within  $\pm 0.01$  ( $\pm 0.02$ ) at the visible (near UV) wavelengths (Dubovik and King, 2000). In addition to the direct sun measurements, the photometer also measures multi-angular diffuse sky radiances along the almucantar plane at four distinct wavelengths from visible to the near-Infrared spectrum (440, 675, 870 and 1020 nm) with near-hourly frequency. In recent years the newer model of instruments also include sky radiance measurements at 380, 500 and 1640 nm (Sinyuk et al., 2020). An inversion procedure that uses both direct sun and angular sky radiances together is implemented to derive aerosol particle size distribution and complex refractive indices (Dubovik and King, 2000; Dubovik et al., 2006; Dubovik et al., 1998; Dubovik and King, 2000). The uncertainty in the derived spectral aerosol SSA provided by the AERONET inversion Level 2 product is estimated to be  $\pm 0.03$  for  $\tau_{440} > 0.4$  (Dubovik et al., 2000). Since 2018, the release of Version 3 inversion product implements several changes to the traditional AERONET aerosol absorption retrievals. A complete description of the changes implemented in Version 3 inversion products along with the updated uncertainty estimates are available in Sinyuk et al., (2020). It should be noted here that in the currently available AERONET inversion products the shortest wavelength of aerosol SSA is 440 nm. In this work, we use AERONET Version 2 inversion product for constructing a representative aerosol model for the associated sites. For the AOD inputs to our retrieval algorithm and for the comparison of SSA we use the latest Version 3 products. For the locations, where aerosol loading is usually low, the derived aerosol absorption properties have much higher retrieval error (Level 1.5 inversion product). In this work, we use AERONET measured columnar  $\tau$  as a constraint to derive aerosol absorption from satellite TOA measurements. The particle size distributions provided by AERONET products are used for constructing a representative aerosol model for the associated sites. Aerosol absorption properties from AERONET Level 2 Version 3 product are used for comparison with our retrievals at visible wavelengths. Figure 1 shows the location of the total 110 AERONET sites used in this work for which long-term ( $> 7$  years) quality assured Level 2 (version 3) measurements are available.

## 2.2 OMI

Launched in July 2004, the Ozone Monitoring Instrument (OMI) on board NASA's EOS Aura satellite is a nadir-viewing hyper-spectral imaging radiometer (Levelt et al., 2006). OMI measures the TOA radiances in the

179 wavelength range 270-500 nm with a ground pixel spatial resolution of 13 km x 24 km at nadir. OMI achieves  
180 daily global coverage in 14-15 orbits with a swath of 2600 km scanning the entire earth's surface. In this work,  
181 we use OMI radiances (340, 354, and 388 nm) provided in the in-house product OMLERWAVE and publicly  
182 accessible [OMI near UV OMAERUV Level 2 aerosol product \(Version 1.8.9.1, Torres et al., 2018\)](#).~~OMAERUV~~  
183 ~~Level 2 aerosol product (V1.8.9.1)~~. The OMLERWAVE product reports radiances and Lambertian equivalent  
184 reflectivity (LER) at several discrete wavelengths in the near-UV and visible parts of the spectrum. Additionally,  
185 we also use ancillary information on the quality of pixel (cloud contamination, land/sea mask, etc.), aerosol  
186 ~~typeindex~~, LER, [surface pressure](#) and ALH data set used in the operational OMAERUV product. ~~The~~  
187 ~~OMAERUV product primarily relies on the measured near UV aerosol index and AIRS retrieved carbon~~  
188 ~~monoxide information to categorize aerosols into Carbonaceous, Dust, and Urban/Industrial types.~~ Since mid-  
189 2007, OMI suffers from an external obstruction that affects the quality of radiance measurements in a few rows  
190 (cross-track pixels). This is referred to as 'row anomaly' that restricts the current usage of OMI observations for  
191 the scientific purpose to about half in a total of 60 cross-track rows (Schenkeveld et al., 2017). Extensive  
192 documentation about how the row anomaly affects the OMAERUV aerosol product is available at Torres et al.,  
193 (2018).

### 194 2.3 MODIS

195 The ~~MODerate-resolution Imaging Spectro-radiometer (MODIS)~~ on board NASA's EOS Aqua and Terra  
196 satellites are nadir-viewing, multi-spectral radiometer. MODIS measures the TOA radiances in 36 wavelength  
197 bands ranging from 0.41-14.23  $\mu\text{m}$  with a ground pixel spatial resolution between 250-1000 m ([King et al.,](#)  
198 [1992](#)). MODIS scans the earth's surface with a 2300 km wide swath to provide near-global coverage on a daily  
199 ~~basis~~. In this work, we use Aqua-MODIS radiances (at 466 and 646 nm) provided in the 10-km aerosol product  
200 (MYD04\_L2) from the Deep-Blue (DB) aerosol algorithm. This aerosol product provides cloud-free radiances  
201 and ancillary information on the [terrain height/pressure, quality of pixel and estimated cloud fraction \(Hsu et al.,](#)  
202 [2013\)](#). For surface characterization in visible wavelengths we use MODIS MAIAC (Multi-Angle Implementation  
203 [of Atmospheric Correction\) MCD19A1 daily 1-km sinusoidal gridded spectral BRDF \(Bidirectional reflectance](#)  
204 [function\) or surface reflectance product \(Lyapustin and Wang, 2018\)](#).~~quality of pixel and estimated cloud~~  
205 ~~fraction.~~

207 A schematic flowchart shown in Figure 2 illustrates the method adopted in this work to derive wavelength-  
208 dependent aerosol absorption.

209 ~~A schematic flowchart of the method adopted in this work to derive wavelength-dependent aerosol absorption is~~  
210 ~~shown in Figure 2.~~

### 211 3.1 Computation of site-specific Seasonal Look-up tables of TOA reflectances

212 To start, we compile a seasonal climatology of aerosol particle size distributions and real part of the refractive  
213 index (440 nm) for the entire  $\tau$  range from the AERONET Level-2 Version-2 inversion product for each site  
214 considered in the study. Here, we assume that the spectral variability of the real part of the aerosol refractive  
215 index through UV-Visible is minimal and, therefore, values derived at 440 nm were assumed to be wavelength-  
216 independent across the UV-Visible spectrum range considered in this study. The resulting site-specific  
217 climatology-climatologies of aerosol size distribution are fed to a radiative transfer model (RTM) to generate look  
218 up table (LUT's) of outgoing top of the atmosphere (TOA) reflectances at 340, 354, 388, 466, and 646 nm with  
219 varying nodal points of satellite-sun geometry (i.e., SZA–solar zenith angle at 0°, 20°, 40°, 60°, 66°, 72° and 80°;  
220 VZA–viewing zenith angle at 2° interval from 0–88°; RAA–relative azimuth angle at 15° interval from 0–180°).  
221 Reflectance LUTs are created for two values of surface pressure (1013.25 and 600 mb), seven values of  $\tau$  (0.0,  
222 0.1, 0.5, 1.0, 2.5, 4.0 and 6.0), five nodal points on ALH (0, 1.5, 3.0, 6.0 and 10.0 km) for the referenced surface  
223 pressure nodes, and eight values of imaginary component of the refractive index (0.000, 0.008, 0.016, 0.024,  
224 0.032, 0.040, 0.048 and 0.056). The aerosol profiles used in the RTM follow a quasi-Gaussian distribution  
225 centred around the respective modes of ALH. We use a total column ozone of 275 Dobson unit in the RTM to  
226 account for ozone absorption. ~~, surface pressure,  $\tau$ , ALH, and imaginary component of the refractive index.~~ The  
227 Gauss-Seidel radiative transfer code (Mie theory) used for this purpose accounts for gaseous absorption,  
228 molecular and aerosol multiple scattering (Herman and Browning, 1965). Thus, a database of AERONET site-  
229 specific seasonal LUT of reflectances for the aerosols observed over each site in the study is created. Figure 3  
230 shows an example of the calculated net aerosol reflectance at the TOA for selected sun-satellite geometry  
231 (SZA=20°, VZA=40°, RAA=130°) and varying values of  $\tau$  and  $\omega_0$  from our LUT developed for the GSFC site  
232 (38.92° N, 76.84° W). These results illustrate that for a given satellite-sun geometry, observed radiance, and  
233 assumed LUT, multiple combinations of  $\tau$  and  $\omega_0$  can explain the satellite measurements. In addition, it is noted  
234 that the net reflectances are mostly invariant at low optical depths ( $\sim 0.1$ ) regardless of variations in SSA for all



235 wavelengths. This is a typical scenario for the LUT approach to derive aerosol properties, suggesting retrieval of  
236 absorption is likely not reliable at low optical depths. These results also illustrate the *critical reflectance concept*  
237 (Kaufman, 1987), as a particular upwelling reflectance value (also associated with a particular value of surface  
238 reflectance) at which there is no sensitivity to aerosol optical depth and, therefore, theoretically suitable for the  
239 retrieval of aerosol absorption from satellite observations. Nonetheless, to derive the best-fit or unique solution of  
240  $\omega_0$  from satellite measurements, an accurate characterization of  $\tau$ , cloud-free radiances, and surface reflectances  
241 are required. Figure 3 shows the calculated net aerosol reflectance at the TOA over the GSFC site (38.92° N,  
242 76.84° W) using particle sizes derived from the AERONET product and varying values of  $\tau$  and  $\omega_0$ . These results  
243 illustrate that for a given satellite sun geometry and observed radiance, multiple combinations of  $\tau$  and  $\omega_0$  can  
244 explain the satellite measurements. This simulation demonstrates that in order to derive  $\omega_0$  from satellite  
245 measurements, an accurate characterization of  $\tau$ , cloud-free radiances, and surface reflectances are required.  
246

247 The site-specific LUTs developed here assume spherical particle shapes (Mie theory) for carbonaceous and urban  
248 aerosols. However, mineral dust particles are assumed non-spherical and modeled as randomly oriented spheroid  
249 (Dubovik et al., 2006; Torres et al., 2018). To account for the non-spherical-sphericity behavior of dust particles,  
250 a unified dust model LUT is created using particle sizes from selected AERONET sites over Sahara and Arabian  
251 Middle East region that include: Saada, SEDE\_BOKER, Solar\_Village, and Tamanrasset\_INM. These sites were  
252 selected based on the observed prevailing dust aerosol type. The particle sizes and real refractive index obtained  
253 at these sites are used with a pre-computed set of kernels that assume a spheroidal shape with a fixed distribution  
254 of axis ratio to produce phase function (Dubovik et al., 2006). The obtained phase matrix elements are input to  
255 the RTM to create reflectance LUT's. The process of acquiring a non-spherical unified dust model LUT is  
256 necessary to account for the non-spherical shape of particles and save a considerable computational time, which  
257 otherwise would require to create another set of site-specific LUTs.

### 258 **3.2 Collocation of satellite and ground measurements**

259 We use satellite measurements (ground pixels) located within the 50 km radius of each AERONET site. In  
260 essence, we treat the overlying atmospheric aerosols within a 50 km radius of the site as a representative of the  
261 AERONET measured  $\tau$ . We look for valid AERONET  $\tau$  measurements within  $\pm 2$  hours of satellite overpass and  
262 assign the  $\tau$  closest in time to all the ground pixels. It should be noted here that for OMI sensor, the OMAERUV  
263 product provides cloud-free radiances (340, 354 and 388 nm) in the native pixel resolution of 13 km x 24 km,  
264 while the MODIS sensor DB-product provides cloud-free radiances (466 and 646 nm) at 10 km x 10 km

265 ~~resolution. To allow for more sampling, we associate the AERONET observations within  $\pm 2$  hours of satellite~~  
266 ~~overpass to the measured TOA radiances. Here, we do not employ any averaging scheme for the AERONET data~~  
267 ~~and keep it intact. While with satellite measurements, we use native pixel resolution of 13 km x 24 km for the~~  
268 ~~OMI wavelengths (340, 354 and 388 nm) and 10 km x 10 km resolution radiances for the MODIS wavelengths~~  
269 ~~(466 and 646 nm).~~

### 270 **3.3 Retrieval of aerosol $\omega_0(\lambda)$**

271 The proposed technique to derive aerosol absorption follows the procedure of obtaining (a) the AERONET AOD  
272 at desired wavelengths for OMI and MODIS, (b) aerosol type, (c) optimal layer height, (d) surface pressure, (e)  
273 surface reflectance, and (f) best quality assured cloud free-TOA reflectances.

274 ~~the best quality assured cloud free TOA reflectances, identifying the aerosol type, optimal layer height, and~~  
275 ~~characterize surface reflectance. We select over land pixels from both sensors with the best quality flags ('0'~~  
276 ~~OMI, '3' MODIS DB) and cloud fraction  $< 0.2$  in the retrieval procedure. Aerosol type information for the OMI~~  
277 ~~wavelengths is directly adopted from the OMAERUV product. While for the MODIS wavelengths, our algorithm~~  
278 ~~looks for the nearest OMI footprint to obtain and assign the corresponding aerosol type. Once an absorbing~~  
279 ~~aerosol type i.e., carbonaceous smoke or mineral dust is identified, we choose the best estimate of ALH from the~~  
280 ~~joint OMI CALIOP climatology derived from a 30-month long record of collocated observations (Torres et al.,~~  
281 ~~2013). While for a weakly absorbing aerosol (Urban), ALH is characterized with a Gaussian distribution of~~  
282 ~~aerosols with a peak at the surface. This is similar to the procedure adopted in the OMAERUV aerosol retrieval~~  
283 ~~(Torres et al., 2013). For the surface characterization at OMI wavelengths, we use a near UV surface albedo~~  
284 ~~database used in the OMAERUV algorithm. At MODIS wavelengths, surface reflectance provided by MAIAC~~  
285 ~~products (Lyapustin and Wang, 2018) is used. Our retrieval technique gathers all above mentioned information~~  
286 ~~for each pixel along with the associated AERONET  $\tau$  to perform an inversion for each wavelength independently.~~  
287 ~~The inversion procedure solves for the best fit of radiances and  $\tau$  with the prior computed site specific seasonal~~  
288 ~~LUT radiances to derive aerosol  $\omega_0(\lambda)$ .~~

#### 290 **3.3.1 AERONET AOD at sensor wavelengths**

291 The AODs at the nominal wavelengths measured by AERONET along with the Extinction Ångström Exponent  
292 (EAE,  $\alpha$ ) for several wavelength pairs (340-440, 380-500, 440-675, 440-875, etc.) are provided in the AOD  
293 product. We derive the  $\tau$  at our interest of satellite wavelengths using the closest available measurement and  $\alpha$   
294 through the power-law approximation (Ångström, 1929) as shown in equation 2. For OMI wavelengths, the

295 AERONET 340 and 380 nm measurements are readily available while  $\tau_{354}$  is obtained with  $\lambda_{Ref} = 380$  nm and  
296  $\alpha_{\lambda_{Ref}} = \alpha_{340-440}$ . For the few sites with older models of AERONET sunphotometer that does not have direct sun  
297 measurements at 340, and 380 nm (for example Banizoumbou, Avignon, etc.), we use  $\lambda_{Ref} = 440$  nm and  $\alpha_{\lambda_{Ref}} =$   
298  $\alpha_{440-675}$ . Similarly, for MODIS wavelengths the  $\tau$  at 466 and 646 nm are obtained using  $\lambda_{Ref} = 440$  nm and  $\alpha_{\lambda_{Ref}} =$   
299  $\alpha_{440-675}$ .

$$\tau_{\lambda} = \tau_{\lambda_{Ref}} \left( \frac{\lambda}{\lambda_{Ref}} \right)^{-\alpha_{\lambda_{Ref}}} \quad (2)$$

### 302 **3.3.2 Aerosol type and ALH**

303 Aerosol type information is essential to derive absorption properties. We use a combination of Extinction  
304 Ångström Exponent ( $\alpha_{440-870}$ ) derived from AERONET and near UV Aerosol Index (UVAI) from OMAERUV  
305 product to categorize the observed aerosols into three basic types – dust, carbonaceous, and urban/industrial.  
306 Initially, our algorithm uses  $\alpha_{440-870}$  to identify the aerosols as coarse ( $\alpha_{440-870} \leq 0.2$ ) and fine ( $\alpha_{440-870} \geq 1.2$ ) mode  
307 dominated particles. Threshold  $\alpha_{440-870}$  of 0.2 chosen for coarse mode particles unambiguously identifies dust  
308 aerosols. However, the sample of fine mode particles consists of both absorbing carbonaceous and weakly  
309 absorbing urban type aerosols. The near-UV aerosol index is an excellent indicator to identify the presence of  
310 absorbing aerosols. Threshold UVAI value adopted from OMAERUV algorithm is used to separate carbonaceous  
311 (UVAI  $\geq 0.8$ ) and urban (UVAI  $< 0.8$ ) aerosols, respectively. Based on extensive tests on the OMI signal strength  
312 on all surface types it is determined that a minimum UVAI of 0.8 is required to identify absorbing aerosols  
313 (Torres et al., 2007, 2013). Although UVAI is an excellent indicator to identify the presence of absorbing  
314 aerosols, the large OMI footprint (13 km x 24 km) and sub-pixel contamination of signal strength might at times  
315 underestimate UVAI categorizing the observed aerosols as urban type. To derive aerosol absorption for dust,  
316 non-spherical LUT is selected while for carbonaceous and urban aerosols site-specific LUTs are used. In  
317 addition, for the absorbing types of aerosols i.e., carbonaceous and dust, we choose the best estimate of ALH  
318 from the joint OMI-CALIOP (Cloud-Aerosol Lidar with Orthogonal Polarization) product. The joint OMI-  
319 CALIOP product was developed using 30-month long record of collocated observations to gather absorbing  
320 aerosol backscattering profiles at 1064 nm that were gridded on  $1^{\circ} \times 1^{\circ}$  resolution to produce global monthly  
321 average ALH. Prescribed uncertainty in the ALH derived from joint OMI-CALIOP product primarily due to  
322 limited sampling of the CALIOP lidar overpasses is expected to be within  $\pm 1$  km (Torres et al., 2013). For urban  
323 aerosols, an exponential aerosol profile peaking at the surface is employed to perform the inversion procedure.

### 3.3.3 Surface reflectance and pressure

For the surface characterization at OMI wavelengths, we use a near-UV surface albedo database at quarter degree grid resolution provided in the OMAERUV product (Torres et al., 2007). The near-UV surface albedo employed by OMAERUV is derived from minimum Lambertian equivalent reflectance obtained from available long-term measurements. The uncertainty in the near-UV surface albedo from these measurements is expected to be within  $\pm 0.01$  (Torres et al., 2018). At MODIS wavelengths, surface reflectance or BRF provided by MAIAC MCD19A1 product is used. The MAIAC MCD19A1 provides spectral surface BRF over cloud-free and clear-to-moderately turbid atmospheric conditions ( $\tau_{466} < 1.5$ ) for solar zenith angles below  $80^\circ$ . The measurement-based uncertainty in MCD19A1 BRF at visible wavelengths is reported to be in the range of 0.002–0.003 for the combined sources of errors including uncertainties from gridding, cloud detection, and aerosol model properties (Lyapustin et al., 2018). Additionally, the surface or terrain pressure reported in the OMAERUV, and terrain height (converted to pressure) reported in the MODIS aerosol products are used in our SSA retrievals.

Our retrieval technique gathers all the above-mentioned required inputs including the best quality assured cloud-free radiances reported in OMAERUV (Quality flag = 0) and MODIS-DB (Quality flag = 3) products to perform an inversion for each wavelength independently. The inversion procedure interpolates the LUT radiances linearly for the prescribed satellite-sun geometry, ALH,  $\tau$ , and logarithmically over the surface pressure nodes. The obtained LUT radiances as a function imaginary refractive index are then fitted with the satellite measured radiances to derive aerosol  $\omega_o(\lambda)$ .

### 3.3.4 Illustration of retrieved SSA

Figure 4 shows the retrieved aerosol SSA over the GSFC site as a function of AERONET measured  $\tau$ . Located in the vicinity of a metropolitan city area, the prevailing aerosols over the GSFC site are the urban or industrial types that are relatively more scattering in nature. In general, retrieved SSA increases with aerosol loading, except for a small decrease at large AOD's at 466 and 646 nm. The mean aerosol SSA retrieved at the GSFC site for all  $\tau$  observations at 340, 354, 388, 466 and 646 nm are 0.91, 0.93, 0.93, 0.90 and 0.85, respectively. The variability of the retrieved SSA is high at lower aerosol loading for all wavelengths. Particularly notable is the high variability of retrieved SSA in most  $\tau$  bins for the visible wavelengths (i.e., MODIS bands). This is due to the weaker diminishing aerosol signal strength for weakly absorbing urban type aerosols at lower aerosol loading in the visible spectrum, where the measured TOA radiances are dominantly contributed by the underlying surface, notably at 646 nm. The mean aerosol SSA retrieved at the GSFC site for observations with  $\tau_{440} > 0.4$  at 340, 354,

356 388, 466 and 646 nm are 0.94, 0.95, 0.95, 0.94 and 0.93, respectively. These results agrees well with the values  
357 reported for GSFC site using AERONET products at 440 and 675 nm as 0.96 and 0.95, respectively (Giles et al.,  
358 2012). Also shown in figure 4 is the number of collocated observations that were used in the inversion and the  
359 percent of observations for which SSA is retrieved. For about 12 years of the satellite and ground collocated  
360 observations used here, ~~it is clearly evident that~~ the number of observations from OMI is less than MODIS  
361 observations. The difference in the number of collocated observations stems partly from the OMI row-anomaly,  
362 cloud contamination, and the coarser pixel resolution. The percent of SSA retrieved observations varies widely  
363 even within the corresponding sensor wavelengths (OMI: 340, 354, 388 nm and MODIS: 466, 646 nm). At times  
364 depending on the surface ~~albedo used~~reflectance, the computed net aerosol reflectance might exceed the LUT  
365 limits and produce SSA values above one or less than the maximum absorption in the LUT, typically referred to  
366 as out-of-bounds retrieval. We avoid this by constraining our inversion procedure within the LUT limits and do  
367 not allow for any extrapolation of the ~~inputs~~radiances. However, this leads to the unequal number of retrieved  
368 observations within the sensor wavelengths. In other words, for a given observation within the OMI or MODIS  
369 sensor, it is possible to have aerosol SSA retrieved at one wavelength and no retrieval (out-of-bounds) at other  
370 wavelengths. ~~Also, it is worth mentioning that for few sites located along the coasts or in the islands (e.g.,~~  
371 ~~Mauna\_Lao, Ascension\_Island, Nauru), we were either unable to retrieve aerosol SSA or the number of days~~  
372 ~~with retrieval is quite low. This is a consequence of OMI's large pixel size, where the satellite measured~~  
373 ~~radiancees are often contaminated by clouds and mixed signal from the surface that are challenging to resolve and~~  
374 ~~lead to out-of bounds in the inversion.~~

375  
376 To examine the ~~seasonal variation of aerosol SSA and its~~ spectral dependence of aerosol absorption, we created a  
377 subset of the data that includes observations for which aerosol SSA is retrieved for all the corresponding sensor  
378 wavelengths simultaneously (OMI: 340, 354, 388 nm and MODIS: 466, 646 nm) on any given day. This step  
379 reduces the sample size drastically but eliminates the need for making prior assumptions on the wavelength  
380 dependence of aerosol absorption ~~angstrom exponent~~ to fill those gaps. ~~Instead,~~ The obtained subset of aerosol  
381 SSA in the UV-Visible range is used to compute the resulting spectral dependence of aerosol absorption of the  
382 prevailing aerosols over the corresponding AERONET sites in terms of the ~~Aerosol~~ Absorption Ångström  
383 Exponent (AAE), a measure of the spectral dependence of aerosol absorption optical depth (Bond,  
384 2001) using a power-law approximation, analogous to the ~~Angstrom~~ Extinction Ångström Exponent (van de  
385 Hulst, 1957). The spectral dependence of aerosol absorption AAE, is ~~reported as Absorption Angstrom Exponent~~

386 ~~(AAE)~~ defined as the slope of aerosol absorption optical depth with wavelengths on a log-log scale. The aerosol  
387 absorption optical depth  $\tau_{abs}(\lambda)$  is derived as shown in equation ~~(43)~~:

$$388 \quad \tau_{abs}(\lambda) = (1 - \omega_0(\lambda)) \cdot \tau_{ext}(\lambda) \quad (43)$$

389 from which the AAE for wavelength range  $\lambda_1, \lambda_2$  is calculated as shown in equation ~~(42)~~.

$$390 \quad AAE(\lambda_1, \lambda_2) = - \frac{\ln(\tau_{abs}(\lambda_1)/\tau_{abs}(\lambda_2))}{\ln(\lambda_1/\lambda_2)}$$

391 ~~(24)~~

392 The results presented hereafter include only a data subset that meets the following three conditions: (a) SSA  
393 retrievals are available for all five wavelengths on a given day, (b)  $\tau_{440} > 0.4$ , to ensure reliable accuracy spanning  
394 through UV-Visible wavelengths, and (c) there are at least 5 days of observations available per season per aerosol  
395 type.~~In addition, we make use of the AERONET extinction angstrom exponent ( $\alpha$ ) at 440-870 nm to distinguish~~  
396 ~~the particles as coarse ( $\alpha < 0.2$ ), fine ( $\alpha > 1.2$ ), and in between as intermediate or mixed mode. Since our aerosol~~  
397 ~~identification strictly uses three primary types, the use of qualitative indicator  $\alpha$  helps delineate the mixture of~~  
398 ~~aerosols when applicable.~~

#### 399 4 SSA retrieval sensitivity analysis

400 The ~~inversion procedure employed here~~proposed inversion procedure to derive spectral aerosol absorption from  
401 the combined ground and satellite measurements is susceptible to several systematic and random errors. These  
402 error sources include uncertainties in the following input parameters: (a) aerosol extinction measurements, (b)  
403 estimation of particle size distribution, (c) real part of refractive index, (d) calibration of satellite measured TOA  
404 radiances, (e) sub-pixel cloud contamination, (f) any unaccounted trace gaseous absorption in the RTM, (g)  
405 surface reflectance, (h) ALH, and (i) surface pressure.~~(e) surface reflectance, (f) aerosol layer height, and (g) sub-~~  
406 ~~pixel cloud contamination.~~ The retrieved aerosol absorption from our inversion procedure could be affected by all  
407 these sources of uncertainties. ~~Among these, error sources from (a-d) are inevitable for which we do not have any~~  
408 ~~direct control over them. However, we do have control only for the sources from (e-g) in our retrieval.~~ Errors  
409 associated with surface reflectance, aerosol layer height, and cloud contaminations on the satellite retrieved  
410 optical depths are well documented in the literature (e.g., Fraser and Kaufman, 1985; Torres et al., 1998; Jethva et  
411 al., 2014). In summary it is known that an: (i) overestimation (underestimation) of surface reflectance leads to  
412 lower (higher) aerosol SSA, (ii) overestimation (underestimation) of  $\tau$  leads to lower (higher) aerosol SSA, (iii)  
413 overestimation (underestimation) of ALH produces higher (lower) aerosol SSA ~~—(significantly more pronounced~~  
414 ~~in the UV than in visible wavelengths), and (iv) an increase in TOA reflectance due to sub-pixel cloud~~

415 contamination produces higher aerosol SSA. ~~We use sensitivity tests for these key input variables to derive a~~  
416 ~~quantitative estimate of the error percolated in the aerosol SSA retrieval due to changes in these variables.~~

#### 418 4.1 Estimation of theoretical errors in the retrieved aerosol $\omega_o(\lambda)$

419 Here, we conduct sensitivity tests for all sources of errors in the input variables to derive a theoretical estimate of  
420 the error percolated in the aerosol SSA retrieval due to uncertainties in the assumed values of these variables. To  
421 have a controlled setup, we performed tests for a representative fixed satellite-sun geometry (SZA=20°,  
422 VZA=20°, RAA=130°) over the GSFC, Mongu, and Tamanrasset INM. These sites were selected to represent  
423 three distinct aerosol types as well as surface conditions. We assume a fixed value of  $\omega_o = 0.9$  at 388 nm and  
424 aerosol load  $\tau_{440} = 0.2, 0.3,$  and  $0.4$  as our references to estimate errors in the retrieved SSA. To derive  
425 corresponding spectral AODs at remaining wavelengths, we assume an  $EAE_{340-646}$  of 1.9, 0.2, and 1.9 for  
426 carbonaceous, dust and urban aerosols, respectively. Similarly, spectral SSA at other wavelengths is estimated  
427 assuming an  $AAE_{340-646}$  of 1.7, 2.5 and 0.9 for carbonaceous, dust and urban aerosols, respectively. We calculate  
428 the uncertainty of the derived spectral SSA for each aerosol type by perturbing, one at a time, the nominal values  
429 of the nine inputs parameters by an assumed or observationally known uncertainty. The absolute error is  
430 computed as the SSA obtained with altered input minus the assumed SSA. The combined uncertainty of the  
431 derived spectral SSA is given by square root of the summation of the squares of the errors associated with each  
432 parameter.

433  
434 Figure 5 shows the error analysis of the retrieved SSA as a function of wavelength and optical depth given a  
435 change in the input (a)  $\tau_{ext}(\lambda)$ , (b) particle sizes, and (c) real part of refractive index. We perturb the input  $\tau$  with  
436 an absolute value of  $\pm 0.02$  for  $\lambda < 400$  nm and  $\pm 0.01$  for  $\lambda > 400$  nm as prescribed by the AERONET AOD  
437 product. As shown in Fig 5a, AOD over-estimations result in SSA under-estimations whereas AOD  
438 underestimations yields SSA overestimations for all aerosol types over the considered spectral range. The  
439 magnitude of the SSA error ( $\Delta\omega_o$ ) decreases with increasing AOD. It is noted that for all aerosol types an  
440 underestimation (overestimation) of AOD, the magnitude of  $\Delta\omega_o$  is positive (negative) and increases with  
441 wavelength (340 nm to 388 nm, and 466 nm to 646 nm). The higher magnitudes of  $\Delta\omega_o$  noted for visible  
442 wavelengths are attributed to lower spectral AODs where the aerosol absorption signal is diminished for a stable  
443 retrieval, particularly notable for weakly absorbing urban aerosols. For the reference  $\tau_{440} = 0.4$ , perturbation of

444  $\pm 0.02 + \tau$  at 340 nm yields an error  $\Delta\omega_0$  within  $\pm 0.002$ , while a perturbation of  $\pm 0.01 + \tau$  at 646 nm yields an error  
445 within  $\pm 0.011$ .

446  
447 Uncertainties in assumed particle sizes are also expected to affect the retrieval of aerosol absorption. To estimate  
448 the error incurred in our SSA retrieval, we perturb the particles size (volume mean radius, VMR) derived from all  
449 AOD observations by 20%. We chose  $\Delta\text{VMR} = 20\%$  based on examination of seasonal climatology of particle  
450 sizes as a function of  $\tau_{440}$  and the most frequently occurring  $\tau_{440}$  bin. It is noted that an overestimation of particle  
451 radii produces higher aerosol SSA leading to positive  $\Delta\omega_0$  and vice-versa. However, the  $\Delta\omega_0$  noted for dust  
452 aerosols are quite small/negligible in UV wavelengths while the  $\Delta\omega_0$  remains invariant at visible spectral range.  
453 This owes to the size of dust particles that are much higher than the considered spectral range where extinction of  
454 radiation reaches maximum efficiency and remains less variant with additional increase in particle sizes. It should  
455 be noted that at the AODs considered, particle size assumptions here have only a small effect on the retrieved  
456 SSA. Spectrally the magnitude of  $\Delta\omega_0$  is minimum in the UV and increases towards the visible wavelengths. The  
457 magnitude and spectral behavior of  $\Delta\omega_0$  noted here suggests aerosol scattering primarily drives the particle size  
458 effect. For the reference  $\tau_{440} = 0.4$ , perturbation of  $\pm 20\%$  VMR to all aerosol types yields an error  $\Delta\omega_0$  within  
459  $\pm 0.018$ , and  $\pm 0.044$  at 340 nm and 646 nm, respectively.

460  
461 Another aerosol intrinsic property input for our SSA retrieval algorithm obtained from AERONET inversion  
462 product is the real part of refractive index (RRI) – which primarily contributes to the magnitude of scattering. The  
463 prescribed uncertainty in aerosol RRI is estimated to be  $\pm 0.04$  (Dubovik et al., 2000). Our results indicate that an  
464 overestimation of RRI produces lower aerosol SSA and vice-versa. The effect of aerosol RRI perturbation on  
465 retrieved SSA is noted to be higher in the UV-spectrum than in the visible. This is likely a result of strong  
466 competing effects from molecular scattering and aerosol absorption, while aerosol load adds an additional weak  
467 dependence. For the reference  $\tau_{440} = 0.4$ , perturbation of  $\pm 0.04 + \text{RRI}$  yields an error  $\Delta\omega_0$  within  $\pm 0.009$  and  
468  $\pm 0.002$  at 340 nm and 646 nm, respectively.

469  
470 Figure 6 shows the error analysis of the retrieved SSA as a function of wavelength and optical depth given a  
471 change in the (a) TOA radiances due to sensor calibration, (b) sub-pixel cloud contamination, and (c) trace  
472 gaseous absorption. The prescribed uncertainty in the TOA radiance measurements for OMI and MODIS sensors  
473 are expected to be  $\pm 1.8\%$  (Schenkeveld et al., 2017) and  $\pm 1.9\%$  (Guenther et al., 2002), respectively. As  
474 expected, an overestimation of TOA radiances due to sensor calibration produces lower aerosol SSA and vice-



475 versa. Errors in the retrieved SSA due to uncertainties in the sensor calibration increase with decreasing aerosol  
476 optical depth. For the reference  $\tau_{440} = 0.4$ , perturbation of  $\pm 1.8\%$  in TOA radiances at 340 nm yields an error  $\Delta\omega_0$   
477 within  $\pm 0.027$ , while a perturbation of  $\pm 1.9\%$  in TOA radiances at 646 nm yields an error within  $\pm 0.037$ .  
478 However, an overestimation in TOA radiances due to assumed optically thin cloud ( $\tau_{\text{cloud}} = 0.5$ ) produces higher  
479 aerosol SSA. The effect of cloud contamination in TOA radiances is more pronounced in the visible than in UV  
480 spectrum. For observations with  $\tau_{440} = 0.4$ , the sub-pixel cloud contamination yields an error  $\Delta\omega_0$  within  $\pm 0.020$   
481 and  $\pm 0.056$  at 340 nm and 646 nm, respectively.

482  
483 As described in section 1, for the retrieval of aerosol properties from satellite measured radiances it is important  
484 to separate the TOA radiance signal from the underlying surface and atmospheric constituents including trace  
485 gases. It should be noted that the RTM used in this work accounts only for the  $\text{H}_2\text{O}$  and  $\text{O}_3$  gaseous absorption.  
486 For the UV channels from OMI, specific wavelengths at 340, 354 and 388 nm are chosen for aerosol retrieval  
487 that has minimal absorption due to trace gases (Torres et al., 2007). However, for the visible channels at 466 and  
488 646 nm from MODIS the absorption trace gases such as  $\text{NO}_2$  could be significant. To account for the remaining  
489 trace gases (for example  $\text{O}_2$ ,  $\text{NO}_2$ , etc.), we use an estimate of the optical depth due to all trace gases in the  
490 atmosphere to derive the error incurred in our SSA retrievals. The combined optical depth of all traces gases that  
491 has absorption lines in visible wavelengths are estimated to be 0.0042 and 0.0344 for the MODIS wavelengths at  
492 466 nm and 646 nm, respectively (Patadia et al., 2018). To estimate an error in our aerosol SSA retrieval due to  
493 unaccounted trace gases we initially derive the LUT radiances for the prescribed  $\tau_{\text{gases}}$  and retrieve SSA by  
494 perturbing the TOA radiance by that amount. Since our LUT radiances are developed by not accounting for all  
495 trace gases, the retrieval of SSA matches higher radiances than the actual, producing lower aerosol SSA than the  
496 actual. Our analysis shows that for the prescribed  $\tau_{\text{gases}}$  the error in the retrieved aerosol SSA for observations  
497 with  $\tau_{440} = 0.4$  are negligible at 466 nm and within 0.024 at 646 nm.

498  
499 Figure 7 shows the error analysis of the retrieved SSA as a function of wavelength and optical depth given a  
500 change in the (a) surface reflectance, (b) ALH, and (c) surface pressure. We perturb the surface reflectance by an  
501 absolute  $\pm 0.01$  for all wavelengths to derive an estimate of error incurred in the SSA retrieval. Our results  
502 indicate that  $\Delta\omega_0$  increases with increasing wavelength and decreasing  $\tau$  due to changes in surface reflectance for  
503 all aerosol types. For less-absorbing (urban) aerosols, the surface reflectance becomes increasingly important at  
504 the visible wavelengths than compared to absorbing (carbonaceous or dust) aerosols. For the observations with  
505  $\tau_{440} = 0.4$ , it is noted that  $\Delta\omega_0$  are within  $\pm 0.011$  and  $\pm 0.050$  at 340 nm and 646 nm, respectively. In contrast to the

506 surface reflectance, the effect of ALH becomes prominent at near-UV wavelengths under the prescribed  
507 uncertainty of  $\pm 1$  km. The  $\Delta\omega_0$  due to changes in ALH decreases with wavelength because of the gradually  
508 diminishing the intensity of Rayleigh scattering (proportional to  $\lambda^{-4}$ ) and its radiative interactions with aerosols.  
509 For the observations with  $\tau_{440} = 0.4$ , it is noted that  $\Delta\omega_0$  due to  $\pm 1$  km ALH are within  $\pm 0.028$  and  $\pm 0.001$  at 340  
510 nm and 646 nm respectively for both absorbing and non-absorbing aerosols.

511  
512 Another essential input in our SSA retrieval algorithm employed here is the terrain or surface pressure that  
513 determines the contribution of molecular scattering in the simulated TOA radiances through pre-computed LUT.  
514 We assume an uncertainty of  $\pm 100$  m in the terrain height equivalent to  $\pm 12$  mb or hPa pressure to derive an  
515 estimate of error incurred in aerosol SSA. Our analysis indicates that an overestimation of surface pressure  
516 produces lower aerosol SSA to compensate the higher radiances reaching the TOA. For absorbing (carbonaceous  
517 or dust) aerosols, the effect of surface pressure on the retrieved SSA is high in the UV spectrum than in visible,  
518 while for less-absorbing (urban) aerosols the error in retrieved SSA is spectrally invariant. Colarco et al., (2017)  
519 provides detailed examination of the effect of terrain pressure on OMI measurements and reported these effects  
520 are prominent at sites over mountainous regions owing to the coarser OMI footprint. In addition to the surface  
521 elevation uncertainty, the atmospheric effects also alter surface pressure. For the observations with  $\tau_{440} = 0.4$ , it is  
522 noted that  $\Delta\omega_0$  due to  $\pm 12$  hPa surface pressure is within  $\pm 0.011$  and  $\pm 0.006$  at 340 nm and 646 nm, respectively.  
523 ~~Sensitivity tests are performed on observations that were reported with best accuracy (minimal cloud~~  
524 ~~contamination in both OMI and MODIS data sets) for a few selected sites that include GSFC, Avignon,~~  
525 ~~Tamanrasset\_INM, Saada, Alta\_Floresta, and Mongu. These sites were selected to include different types of~~  
526 ~~aerosols observed over these sites. Figure 5 shows the error analysis of the retrieved SSA as a function of~~  
527 ~~wavelength and optical depth given a change in the surface reflectance and ALH by an absolute change of  $\pm 0.01$~~   
528 ~~and  $\pm 1$  km ALH. The absolute error is computed as the SSA obtained with altered input minus the actual SSA.~~  
529 ~~As expected, the error in retrieved SSA increases with increasing wavelength and decreasing  $\tau$  due to changes in~~  
530 ~~surface reflectance for all aerosol types. For less absorbing (Urban) aerosols, the surface reflectance becomes~~  
531 ~~increasingly important at the visible wavelengths than compared to absorbing aerosols. Our analysis shows that~~  
532 ~~for small  $\tau_{440}$  ( $\sim 0.2$ ), the error in retrieved SSA is much higher ( $> \pm 0.05$ ) for visible wavelengths, while that in the~~  
533 ~~near-UV region reaches up to  $\pm 0.03$ . However, for observations with  $\tau_{440} \sim 0.4$ , a reliable accuracy of within~~  
534  ~~$\pm 0.05$  is achieved even for the non absorbing aerosols at 646 nm. In contrast to the surface reflectivity, the effect~~  
535 ~~of ALH becomes prominent at near UV wavelengths. The error in retrieved SSA due to changes in ALH~~  
536 ~~decreases with wavelength because of the gradually diminishing the intensity of Rayleigh scattering and its~~

537 radiative interactions with aerosols. The errors in the retrieved SSA are estimated to be better than  $\pm 0.03$  in the  
538 near-UV and negligible at visible wavelengths for both absorbing and non-absorbing aerosols.

539  
540 Table 2 summarizes the SSA error analysis 340 nm and 646 nm due to uncertainties in most relevant input  
541 variables. Among the nine input variables used in our algorithm, the  $\Delta\omega_o$  at 340 nm arises mostly from (in  
542 descending order) the uncertainties in calibration of TOA radiances, sub-pixel cloud contamination, ALH,  
543 particle sizes and so on. While for the visible wavelength at 646 nm the  $\Delta\omega_o$  arises mostly from (in descending  
544 order) cloud contamination, surface reflectance, calibration of TOA radiances, particle sizes, trace gases, and so  
545 on. These sensitivity tests clearly indicate that  $\Delta\omega_o$  is (a) spectrally dependent due to multiple variables, (b)  
546 decreases with increasing  $\tau$ , and (c) varies with absorbing nature of aerosols. The combined error in retrieved  
547 SSA can be now estimated as the square root of the sum of individual error squares (RMSE). Overall, for the  
548 observations with  $\tau_{440} = 0.4$ , the combined error in the retrieved SSA for absorbing (less-absorbing) aerosols are  
549 within  $\pm 0.043$  ( $\pm 0.038$ ) and  $\pm 0.073$  ( $\pm 0.086$ ) at 340 nm and 646 nm, respectively. However, it should be noted  
550 that depending on the reliability of the input variables, the errors stemming from individual sources could be in  
551 opposite direction resulting in cancellation of errors. Under such scenario, the combined error in the retrieved  
552 spectral SSA is expected to be much lesser than the combined value reported here with our sensitivity tests.  
553 presents the achievable accuracy in our SSA retrievals at 340 nm and 646 nm. Overall, for the observations with  
554  $\tau_{440}$  equals to 0.4, the estimated accuracy in our retrieved SSA is within  $\pm 0.03$  ( $\pm 0.05$ ) for absorbing (non-  
555 absorbing) aerosols through 340-646 nm. While for the observations with  $\tau_{440}$  up to 0.2, the achievable accuracy  
556 in SSA reaches up to  $\pm 0.05$  ( $> \pm 0.05$ ) for absorbing (non-absorbing) aerosols.

#### 558 **4.2 Estimation of theoretical errors in the derived AAE**

559 Similar to the estimation of errors in the retrieval of SSA, we conduct sensitivity tests to determine the errors in  
560 the computed AAE due to certainties in the retrieved SSA. We assume a fixed  $\omega_o(388) = 0.9$ , aerosol load  $\tau_{440} =$   
561 0.4, EAE<sub>340-646</sub> of 1.9, 0.2, and 1.9 for carbonaceous, dust and urban aerosols, respectively to derive the nominal  
562 AAE values at three wavelength pairs 354-388, 466-646, and 340-646. By using fixed  $\tau(\lambda)$ , we perturb the  $\omega_o(\lambda)$   
563 by  $\pm 0.01$  intervals to compute the AAE. The errors are reported as difference in AAE from perturbed SSA minus  
564 the AAE derived from nominal SSA.

566 Table 3 presents the theoretical uncertainty in computed AAE due to uncertainties in the SSA. As expected, the  
567 ΔAAE noted for all wavelength pairs increases with increasing Δω<sub>0</sub> for all aerosol types. It is noted that for fine  
568 mode particles (carbonaceous and urban), an overestimation of ω<sub>0</sub> produces lower AAE (negative errors) and  
569 underestimation of ω<sub>0</sub> produces higher AAE (positive errors). In contrary for coarse mode particles, an  
570 overestimation of ω<sub>0</sub> produces higher AAE (positive errors) and underestimation of ω<sub>0</sub> produces lower AAE  
571 (negative errors). This is due to the fact that large particle size drives the scattering effect producing low aerosol  
572 absorption optical depths and, therefore, further overestimation of ω<sub>0</sub> yields lower single scattering co-albedo (1 –  
573 ω<sub>0</sub>). The magnitude of ΔAAE is higher for overestimation than those noted for underestimation of SSA. In  
574 addition, it is noted that for fine mode particles the errors in ΔAAE<sub>354-388</sub> (UV spectral range) are higher, while for  
575 coarse mode particles the errors in ΔAAE<sub>466-646</sub> (visible spectral range) are higher than the other two wavelength  
576 pairs. For the assumed carbonaceous aerosols, perturbation of Δω<sub>0</sub> = ±0.04 yields a ΔAAE within ±0.13 for all  
577 wavelength pairs. For the urban aerosols, perturbation of Δω<sub>0</sub> = ±0.04 yields a ΔAAE within ±0.70 for all  
578 wavelength pairs. However, for dust a perturbation of Δω<sub>0</sub> = ±0.04 yields a ΔAAE up to ±1.3 for 354-388  
579 wavelength pair and much higher in the 466-646, and 340-646 wavelength pairs. Additional tests were performed  
580 by perturbing only one of the SSA while deriving AAE for any wavelength pair. The resulting ΔAAE is much  
581 higher than for the tests where SSA is perturbed for all wavelengths. It is noted that for even a small perturbation  
582 of Δω<sub>0</sub> = ±0.01 at one of the wavelengths in a pair, the ΔAAE is ±1.2, ±0.2, and ±0.4 for the wavelength pairs at  
583 354-388, 466-646, and 340-646 respectively for all aerosol types. Overall, the errors noted for AAE are  
584 consistent with the wavelength dependence of ω<sub>0</sub> that is function of both size and absorbing nature of the  
585 particles.

## 586

### 587 5-Results

588 The results presented hereafter include only a data subset that meets the following three conditions: (a) SSA  
589 retrievals are available for all five wavelengths on a given day, (b) τ<sub>440</sub> > 0.4, to ensure reliable accuracy spanning  
590 through UV-Visible wavelengths, and (c) there are at least 5 days of observations available per season per aerosol  
591 type. We include results of retrieved SSA for cases when τ<sub>440</sub> ≤ 0.2 as supplementary materials. Table 3 presents  
592 seasonal averages of aerosol SSA and AAE for all sites considered in this work. The spectral dependence of  
593 absorption quantified as AAE is reported for three different wavelength pairs covering the near-UV (354, 388),  
594 Vis (466, 646) and UV-Vis (340, 646) spectra.

## 595 **5.1 Biomass Burning**

### 596 **5.1.1 South America**

597 The AERONET sites located in South America include Alta\_Floresta, Arica, CEILAP\_BA, CUIABA\_Miranda,  
598 Rio\_Branco, Ji\_Parana\_SE, SANTA\_CRUZ\_UTEPSA, Sao\_Paulo and Campo\_Grande\_SONDA. In general,  
599 carbonaceous aerosols over South America are dominantly emitted from biomass burning during southern  
600 hemisphere spring (JJA) and summer (SON), with distinct peaks in August and September. Most aerosol  
601 emissions are associated with biomass burning for land and agricultural management practices. Further, densely  
602 populated places like Sao Paulo and CEILAP\_BA are also affected by vehicular and industrial emissions.  
603 Aerosols from northern parts of the Amazon Basin advecting south or southeast, and long-range transport of aged  
604 smoke from Southern Africa are not uncommon over few locations in South America. Figure 6a shows regional  
605 average SSA for carbonaceous particles at 466 nm is noted to be 0.92 (0.93) during JJA (SON) months. Average  
606  $\tau_{440}$  and  $\alpha_{440-870}$  are about 1 and 1.8 respectively for both seasons, indicating dominantly fine mode nature of these  
607 particles. Spectral SSA increases from 340 nm (0.90) to 388 nm (0.93) followed by a decrease toward the visible  
608 wavelengths. For urban aerosols, the average  $\tau_{440}$  (0.57) and  $\alpha_{440-870}$  (1.70) are lower than those of carbonaceous  
609 particles. The average SSA at 466 nm for urban aerosols is 0.88 (0.92) for JJA (SON) months. We also find that  
610 the urban aerosols in South America are more absorbing during JJA than in SON. Examining individual sites  
611 (Table 3) reveals high absorption during JJA at CUIABA and Sao Paulo. It is likely that the urban aerosol  
612 samples shown here attribute to a mixture of aerosols.

### 613 **5.1.2 Southern Africa**

614 The AERONET sites located in the Southern Africa are Mongu, Pretoria, and Skukuza. In addition to the natural  
615 forest fires and emissions from crop residue burning, heavy industrial facilities and episodic dust commonly  
616 dictate the aerosol amounts over Southern Africa. Figure 6b shows the regional average SSA derived for aerosols  
617 over the Southern Africa. For the carbonaceous and urban aerosols observed over the region, maximum  
618 absorption is found in JJA period. The average SSA for carbonaceous particles increases from 340 nm to 466 nm  
619 and then decreases at longer wavelengths. Distinct seasonality in absorption for carbonaceous particles is  
620 observed with maximum (minimum) value of 0.87 (0.90) at 466 nm for JJA (SON) months. The range of regional  
621 average values of  $\tau_{440}$  and  $\alpha_{440-870}$  for carbonaceous particles are about 0.76 to 1.02 and 1.76 to 1.84 ranges  
622 respectively, while for urban aerosols the variability ranges are 0.55 to 0.73 and 1.68 to 1.74, respectively. As  
623 noted,  $\tau_{440}$  and  $\alpha_{440-870}$  are higher for carbonaceous particulate than for urban aerosols. For the latter, the average

624 SSA shows similar spectral behavior as carbonaceous particles. However, the AAE of carbonaceous aerosols in  
625 the UV-Vis range is noted to be ~1.73, while for urban aerosols it is ~2.2. Notable seasonality for urban aerosols  
626 is observed for DJF where the average SSA is almost flat from 340 nm to 466 nm with a slight increase  
627 thereafter. Aerosols organic components are likely the cause of such absorption spectral feature.

### 628 5.1.3 Australia

629 In general, inland Australia is categorized as arid region vastly covered with deserts. However, Northern and  
630 Western parts of the continent are covered with savanna grasslands, where biomass burning due to natural forest  
631 fires and land management practices are known to produce high aerosol emissions during the dry season (May-  
632 October). The regional average SSA for the aerosols observed over Northern Australia at the sites Jabiru and  
633 Lake Argyle is shown in figure 6c. For the sample obtained in this work, carbonaceous and urban aerosols are  
634 found during spring (SON). The average  $\tau_{440}$  and  $\alpha_{440-870}$  for carbonaceous particles is 0.65 and 1.62, respectively.  
635 Average SSA increases with wavelength from 0.87 (340 nm) to 0.89 (388 nm) and then decreases to 0.87 (646  
636 nm). Such behavior of fine mode particles is likely a result of a mixture of black carbon and organic carbon  
637 amounts in the atmosphere, producing stronger (weaker) absorption in the UV (Vis) wavelengths. Although the  
638 mean  $\alpha_{440-870}$  for both carbonaceous and urban aerosols is similar, the mean  $\tau_{440}$  for urban particles is relatively  
639 low. The UV-Vis spectral dependence of urban aerosols (1.57) follows similar behavior as carbonaceous aerosols  
640 (1.42). Unlike the typical urban aerosol absorption, the spectral SSA observed here indicates a mixture of urban  
641 and carbonaceous particles.

## 642 5.2 Dust

### 643 5.2.1 Sahara

644 The seasonal average SSA of aerosols over the sites Saada and Tamanrasset is shown in figure 7a. As one would  
645 expect from the region, dominantly dust aerosols are observed with average  $\tau_{440}$  and  $\alpha_{440-870}$  ranging from 0.63 to  
646 0.85 and 0.11 to 0.16, respectively. Regional average aerosol SSA obtained at 466 nm is ~0.94. The spectral  
647 dependence of dust aerosols shows increase in SSA with wavelength, i.e., 0.86 at 340 nm to 0.97 at 646 nm. The  
648 UV-Vis AAE obtained for dust aerosols range from 2.8 to 3.3 with no distinct seasonality in the average spectral  
649 SSA. As noted, in addition to the coarse dust both sites constitute intermediate range of aerosols with average  
650  $\alpha_{440-870}$  up to 0.3. Compared to coarse particles the intermediate sizes exhibit typical 'dust' absorption curve but  
651 has relatively low SSA (0.95) at 646 nm and therefore low UV-Vis AAE (2.1 to 2.5). This is consistent with the  
652 known dependence of scattering effects at longer wavelength for the dust particles.

### 653 **5.2.2 Sahel**

654 The AERONET sites located in the Sahel region are Agoufou, Banizoumbou, Dakar, IER\_Cinzana, Ilorin,  
655 Ouagadougou, and Zinder Airport. Regional average SSA for the aerosols over Sahel region is shown in figure 7b.  
656 For dust aerosols, the average spectral SSA resembles typical dust absorption curve (increase in SSA with  
657 wavelength). No distinct seasonality in regional average absorption is observed. However, during SON months  
658 relatively low  $\tau_{440}$  (0.64) and UV-Vis AAE (1.57) is noted compared to other seasons. Aerosols with intermediate  
659 size category ( $0.2 < \alpha_{440-870} < 1.2$ ) are observed for all sites in the Sahel region. The spectral SSA obtained for  
660 these aerosols clearly indicate mixture of dust and carbonaceous particles. Eck et al. (2003) reported smoke  
661 particles found toward southern parts of the West Africa are relatively coarse ( $\alpha$  ranges 0.3 to 0.5) than those  
662 found elsewhere—likely a result of mixing with other aerosol types, coagulation, humidification or combination  
663 of such processes. Typical carbonaceous absorption curve noted here for the wet season (JJA and SON) is  
664 consistent with those reports. The average SSA for these (coarse-) carbonaceous particles is ~0.93 at 466 nm and  
665 exhibit a UV-Vis AAE range 0.5 to 1.2. The average SSA observed for DJF and MAM (dry season) indicates a  
666 mixture of prevailing dust and carbonaceous particles. For the sample obtained over Sahel region, fine-mode  
667 carbonaceous and urban aerosols are observed at Ilorin during DJF and SON. Both these aerosol types show  
668 significant absorption with the average SSA at 340 nm and 646 nm is ~0.86 and 0.87, respectively, with nearly  
669 no spectral dependence. This likely indicates the presence of black carbon amounts over the Ilorin site during  
670 DJF and SON. In addition to the biomass burning, fossil fuel combustion, and vehicular emissions, the vast  
671 number of gas flaring stations (> 300) around the Niger Delta produces high emissions (Onyeuwaoma et al.,  
672 2015). Highly absorbing black carbon amounts observed at Ilorin is possibly a result of such emissions.

### 673 **5.2.3 Arabian Peninsula**

674 The AERONET sites located in the Middle East/Arabian Peninsula region include Solar\_Village,  
675 SEDE\_BOKER, Nes\_Ziona, and Cairo\_EMA\_2. The regional average SSA of aerosols observed for these sites is  
676 shown in figure 7c. For dust aerosols, as expected, the average SSA increases with increasing wavelength. The  
677 regional average SSA ranges from 0.89 to 0.98 at 340 nm to 646 nm, with UV-Vis AAE in the range of 2.7 to  
678 3.8. No distinct seasonality in SSA is found from our sample of observations. However, a slight increase in SSA  
679 at UV wavelengths is noted during winter (DJF). Examining individual sites reveal this feature corresponds to the  
680 aerosols over Solar\_Village. The increase in SSA and high AAE<sub>354-388</sub> noted for Solar\_Village during DJF likely  
681 indicates transport of aerosols from neighboring regions. Intermediate range of particles with  $0.2 < \alpha_{440-870} < 1.2$

682 are observed over all sites. The regional average SSA observed for these particles clearly indicate mixture of  
683 aerosols. Carbonaceous aerosols found over Cairo from our sample have average SSA ranging from 0.89 to 0.91  
684 at 340 nm to 646 nm. The UV-Vis AAE (1.91) obtained for these aerosols likely indicate mixture of black and  
685 organic carbon amounts. For the urban aerosols observed over Cairo and Nes-Ziona the regional average SSA is  
686 found to be 0.89 at both 340 nm and 646 nm indicative of mixture of aerosols. While urban aerosols and pollution  
687 prevail over Nes-Ziona, emissions from crop residue burning (rice straws) over the Nile delta region during  
688 winter (DJF) and heavy pollution dictate the aerosol absorption noted over Cairo.

### 689 **5.3 Urban/Industrial**

#### 690 **5.3.1 Western North America**

691 Dominantly urban type aerosols are observed in Western North America primarily produced from industrial  
692 activities and vehicle emissions. In addition, owing to the general meteorological and geographical setting of the  
693 western North America, the region experiences drier months in the summer and fall that initiates natural forest  
694 fires. The regional average aerosol SSA derived from our sample over the western North America is shown in  
695 figure 8a. The average aerosol absorption for urban aerosols decreases in the wavelength range 340–388 nm  
696 followed by an increase at longer wavelengths. Carbonaceous aerosols are observed in our sample over the  
697 Missoula site located in the State of Montana. The average SSA retrieved for carbonaceous particles increases  
698 with wavelengths from 340 nm to 466 nm (0.89 at 0.94) and then decreases towards 646 nm (0.88). The average  
699  $\tau_{440}$ ,  $\sigma_{440-870}$  and UV-Vis AAE obtained for the carbonaceous particles are 1.74, 1.8 and 1.72 respectively.

#### 700 **5.3.2 Eastern North America**

701 Similar to the western part, atmospheric aerosols found over the eastern North America primarily originated from  
702 the industrial activities and secondary aerosol processes (Malm, 1992). Biomass burning generated carbonaceous  
703 particles and dust or mixture of aerosols over the eastern parts of the continent is a rare occurrence, except in the  
704 events of long range transport of smoke from the west. Thus, the average aerosol SSA retrieved over the  
705 AERONET sites in this region follows a typical ‘Urban’ spectral absorption curve, as shown in figure 8b. It is  
706 observed that aerosol SSA increases from the 340 nm to 388 nm or 466 nm and then decreases attaining a  
707 maximum absorption (~0.90) at 646 nm. The regional average SSA for the MAM and JJA months at 466 nm is  
708 0.89 and 0.87 respectively. It should be noted that for urban aerosols the retrieval error at visible wavelength is  
709 high (up to 0.05). Since there are no notable changes in the aerosol sources, the observed decrease absorption at  
710 646 nm for JJA period is likely attributed from the retrieval uncertainty.



### 5.3.3 Europe

For the AERONET sites located in the Europe, dominantly urban aerosols are observed (figure 8c). The regional average SSA increases from 340 nm to 388 nm and then decreases reaching a minimum value (0.83) at 646 nm for most seasons. The observed aerosol absorption is similar for spring (MAM) and summer (JJA). However, there is an increase in absorption at wavelengths other than 646 nm for fall (SON) that reaches a maximum absorption 0.86 and 0.89 at 340 nm and 466 nm respectively. This likely suggests the presence of organic carbon amounts emitted from local fuel combustion sources. This feature is consistent with Ilias et al., (2019), that reports annual cycle of aerosol absorption over Thessaloniki site. Individual site observations reveal this increase in absorption is prominent for most locations in Europe during fall (Table 3) and winter (Table S1). While the aerosol loading is similar throughout the seasons, the average UV-Vis AAE for urban aerosols range from 1.0 to 1.26. In addition to the urban/industrial aerosols, long-range transport of dust from Sahara is not uncommon over central Europe and Mediterranean Basin. Although our sample over European sites do not constitute any coarse mode particles, intermediate range ( $0.2 < \alpha_{440-870} < 1.2$ ) of aerosols are observed for most sites. The average SSA curve noted here indicates mixture of aerosols.

## 5.4 Mixed aerosol types

### 5.4.1 Mid-Atlantic North America

The seasonal average aerosol SSA obtained over the Mexico City is shown in figure 9a. Unlike typical urban aerosol absorption, the SSA curve obtained here shows steep decrease from 388 nm to 466 nm and remains flat or slightly decreases till the 646 nm. Seasonality in aerosol absorption is observed at UV wavelengths ( $< 400$  nm) with maximum (DJF, SON) and minimum (MAM) absorption of 0.86 and 0.91 at 354 nm. It is clearly evident that such absorption curve and seasonal variation is a result of prevailing mixture of aerosols. Seasonal average  $\alpha_{440-870}$  indicates the prevalence of fine mode particles over the Mexico City. The average SSA at 466 nm for DJF, MAM and SON months are 0.85, 0.86 and 0.84 respectively. Although the seasonal average UV-Vis AAE of aerosols over Mexico City ranges from 0.94-1.27, the AAE<sub>354-388</sub> is higher for (5.13) MAM than compared to (2.2) DJF and (0.6) SON months. In general, Mexico City is a densely populated urban location that is well known for its high pollution levels among the other megacities worldwide. In addition to the high concentration of aerosols from fossil fuel combustion throughout the year, Mexico City also experiences biomass-burning aerosols during the relatively dry months of March-May from local sources (Eck et al., 1998).

### 5.4.2 North-Eastern Asia

The AERONET sites located in the Northeastern Asia include: Beijing, Osaka, Shirahama and XiangHe. Figure 9b shows the regional average SSA derived over Northeastern China. For the samples obtained, dust aerosols possibly mixed with regional pollution are observed over Beijing and XiangHe during spring (MAM). The spectral curve of regional average SSA shows an increase from 0.87 at 340 nm to 0.95 at 646 nm. The UV-Vis AAE obtained for the dust aerosols is 1.56. Among sites considered here, carbonaceous aerosols are observed throughout the year at Beijing and XiangHe. The spectral behavior of carbonaceous aerosols shows increase in SSA from 340 nm to 466 nm and thereafter remains near constant or slightly decreases with an UV-Vis dependence ranging from 1.84 to 2.14. However, significant seasonality is noted with minimum (0.95 at 466 nm) and maximum (0.90 at 466 nm) absorption during JJA and DJF respectively. The increase in SSA likely caused the humidification and secondary aerosol processes during JJA. Carbonaceous and urban aerosols over the region show high absorption in winter (DJF), likely due to high amounts of local fossil fuel combustion and agricultural waste burning. The spectral behavior of urban aerosols is similar to carbonaceous aerosols with decrease in magnitude of average SSA, AOD, and UV-Vis AAE. As expected from the local sources, mixture of aerosols categorized by particle sizes ( $0.2 < \alpha_{440-870} < 1.2$ ) is observed at all sites in the region with widely varying spectral dependence.

### 5.4.3 Northern India

The AERONET sites located in Northern India include Jaipur, Gandhi College and Kanpur. Major source of aerosols over the region includes industrial and vehicular emissions, combustion of biomass and fossil fuels. In addition, desert dust passage from arid and semiarid regions of northwestern India, Pakistan, and Arabian Peninsula is commonly observed during spring and summer months. Figure 9c shows the regional average aerosol absorption observed at Kanpur. As expected from the source regions, dust aerosols are observed during spring (MAM) and summer (JJA) months over Jaipur and Kanpur. The average SSA shows a steep increase from 340 nm (0.88) to 466 nm (0.95), and a relatively smaller increase from 466 nm to 646 nm (0.97). The dust aerosols noted here has average  $\tau_{440}$  0.73 to 0.77 and exhibit UV-Vis AAE between from 2.9 to 3.4. For our sample, carbonaceous aerosols are observed over Kanpur and Gandhi College with similar regional average absorption during SON and DJF. The spectral behavior of carbonaceous aerosols shows increase in average SSA from 340 nm (0.91) to 466 nm (0.93) and slight decrease till 646 nm (0.91). The regional average UV-Vis AAE for carbonaceous aerosols range 1.2 to 1.5. Emissions from crop residue burning during SON and biomass

768 ~~burning for residential heating in DJF prevail over the entire Indo-Gangetic plain and likely result in such~~  
769 ~~absorption. In addition the vehicular and industrial emissions add an extra burden of aerosols to the atmosphere.~~  
770 ~~In other words, urban aerosols over the Northern India can be categorized as carbonaceous aerosols resulting~~  
771 ~~from various carbon emitting, both black and organic carbon, sources such as crop residue burning, local biomass~~  
772 ~~burning for house hold heating purposes in DJF months, vehicular and industrial emissions. These urban aerosols~~  
773 ~~are relatively more absorbing ( $\omega_a(466) \sim 0.89$ ) than the carbonaceous aerosols ( $\omega_a(466) \sim 0.93$ ) from~~  
774 ~~industrial/vehicular activities observed over Northern India. Throughout the seasons, influence of pollution~~  
775 ~~aerosols is clearly evident in the observed urban aerosol absorption. Although the UV-Vis AAE is found to be in~~  
776 ~~the similar range, the  $AAE_{354-388}$  values are low for urban than carbonaceous aerosols indicative of high organic~~  
777 ~~amounts in crop residue/biomass burning emissions. For the group of aerosols in between the fine and coarse~~  
778 ~~mode the spectral variation of SSA varies widely.~~

## 780 **6 Spectral Aerosol Absorption in Major Aerosol Environments**

781 In this section, we describe regional average aerosol absorption and AAE derived from our subset of results over  
782 worldwide regions dominated by carbonaceous, dust, and urban aerosols. The AERONET sites selected for this  
783 analysis are based on the dominant samples observed here and well-known aerosol sources from the literature.

784  
785 The average of aerosol SSA and AAE derived for all sites considered in this work are presented in Table S1 as  
786 supplementary materials. In addition, the corresponding particle size distributions used for developing the LUT  
787 radiances for these sites are presented in supplementary materials as well.

### 788 **6.1 Biomass Burning**

789 Emissions from biomass burning are one of the major contributors of carbonaceous aerosols found in the  
790 atmosphere. These carbonaceous aerosols are primarily composed of black carbon and organic carbon  
791 components in addition to minor fractions of inorganic components (Andreae and Merlet, 2001). Studies show  
792 that black carbon amounts in the atmosphere are high absorbers of solar radiation and have near unity AAE due  
793 to invariant imaginary part of refractive index in the UV-Visible spectrum (Bergstrom, 1973; Bohren and  
794 Huffman, 1983; Bergstrom et al., 2002). The typical spectral behavior for carbonaceous aerosols has decreasing  
795 SSA with increasing wavelength in the visible spectrum (Eck et al., 1998; Reid and Hobbs, 1998). Additionally,

796 the presence of organic carbon amounts shows enhanced absorption in the UV region (Kirchstetter et al., 2004).  
797 Our observations for the carbonaceous aerosols from worldwide biomass burning regions depict these  
798 characteristics very well. The seasonal average of spectral SSA for carbonaceous aerosols found over major  
799 aerosol environments are shown in Figure 11.

800  
801 For the Missoula site located in Northwestern United States (US), carbonaceous aerosols are observed during  
802 JJA. Aerosols observed over Missoula are primarily emitted from natural forest fires of the northwestern US in  
803 the dry season (JJA). The spectral SSA of these aerosols is noted to increase from 340 nm ( $0.89\pm 0.02$ ) to 466 nm  
804 ( $0.94\pm 0.03$ ) followed by a decrease toward the 646 nm ( $0.90\pm 0.06$ ). Average  $\tau_{440}$  and  $\alpha_{440-870}$  are about 1.2 and  
805 1.8 respectively, while the average  $AAE_{340-646}$  is noted as 1.8. Our results are consistent with the insitu  
806 measurements of wildfire smoke at the Missoula ground station for 2017 and 2018 summer, that reports an  
807 average SSA of 0.93-0.94 at 401 nm and AAE 1.7-1.9 over the spectral range 401-870 nm (Selimovic et al.,  
808 2020).

809  
810 Over South America, our subset of carbonaceous aerosols is observed for sites at the Alta Floresta, Cuiaba, Ji  
811 Parana, and Santa Cruz. In general, carbonaceous aerosols over South America are dominantly emitted from  
812 biomass burning during southern hemisphere spring (JJA) and summer (SON), with distinct peaks in August and  
813 September. Most aerosol emissions are associated with biomass burning for land and agricultural management  
814 practices. The regional average SSA of these aerosols at 466 nm is noted to be  $0.92\pm 0.03$  ( $0.93\pm 0.02$ ) during JJA  
815 (SON) months. Spectral SSA increases from 340 nm ( $0.90\pm 0.02$ ) to 388 nm ( $0.93\pm 0.03$ ) followed by a decrease  
816 toward the visible wavelengths. Average  $\tau_{440}$  and  $\alpha_{440-870}$  are about 1 and 1.8 respectively with mean  $AAE_{340-646}$   
817 ranging between 1.5-1.8 for both JJA and SON months. Among the sites considered here, Cuiaba located in the  
818 cerrado ecosystem exhibit highest aerosol absorption ( $\omega_o \sim 0.89\pm 0.03$  at 466 nm), while the remaining sites are  
819 surrounded by tropical rainforest exhibit relatively less absorption ( $\omega_o \sim 0.93\pm 0.03$  at 466 nm). Burning of cerrado  
820 (wooded savanna) and rainforest dominantly happens through flaming and smoldering phase combustion  
821 respectively, resulting in the noted variation of aerosol absorption over these sites (Schafer et al., 2008).

822  
823 Over Southern Africa, our subset of carbonaceous aerosols is observed for sites at the Mongu, and Skukuza.  
824 Emissions from biomass burning primarily for agricultural and land management practices are major source of  
825 aerosols over Southern Africa (Eck et al., 2001, 2003). In addition, crop residue burning, heavy industrial  
826 facilities and episodic dust commonly dictate the aerosol amounts over Southern Africa. Fine mode carbonaceous

827 aerosols noted over these sites shows high average absorption during JJA period. The average SSA for these  
828 aerosols increases from 340 nm to 466 nm and then decreases at longer wavelengths. Distinct seasonality in  
829 absorption for carbonaceous particles is observed with maximum (minimum) value of  $0.87\pm 0.02$  ( $0.90\pm 0.03$ ) at  
830 466 nm for JJA (SON) months. The range of regional average values of  $\tau_{440}$  and  $\alpha_{440-870}$  for these aerosols are  
831 about 0.76 to 1.02 and 1.76 to 1.84 ranges, respectively. Average  $AAE_{340-646}$  of these carbonaceous aerosols is  
832 noted to be  $\sim 1.73$  and 1.58 for JJA and SON months.

833

834 For the sample obtained over Sahel region, carbonaceous aerosols are observed at Ilorin during DJF. Fine mode  
835 aerosols observed over Ilorin are primarily emitted from the biomass burning of the grasslands and savanna in  
836 Sahelian and Sudanian zones during the dry season (November through March). Our results show these aerosols  
837 exhibit significant absorption with the average SSA  $\sim 0.86\pm 0.02$  and  $0.87\pm 0.03$  at 340 nm and 646 nm,  
838 respectively. Average  $\tau_{440}$  and  $\alpha_{440-870}$  are about 1 and 1.3 respectively with mean  $AAE_{340-646}$  1.38 for DJF period.  
839 The high aerosol absorption noted here is consistent with the AERONET data analysis that reports high aerosol  
840 absorption with increasing fine-mode fraction (FMF) of particles (SSA  $\sim 0.80-0.87$  and  $0.81-0.85$  at 440 nm and  
841 675 nm for observations with FMF of 0.75-0.54 at 675 nm) over Ilorin during the dry season (Eck et al., 2010).  
842 Spectrally the SSA is noted to be nearly invariant, suggesting high fractions of black carbon amounts in the  
843 aerosols over Ilorin during DJF. Emissions from burning of grasslands and savanna in the Sahelian and Sudanian  
844 zones dominantly happens through flaming phase combustion producing high amounts of soot (Eck et al., 2010).  
845 In addition to the biomass burning, fossil fuel combustion, and vehicular emissions, the vast number of gas  
846 flaring stations ( $> 300$ ) around the Niger Delta produces high emissions (Onyeuwaoma et al., 2015). Highly  
847 absorbing black carbon amounts observed at Ilorin is likely a result of such emissions.

848

849 For the Cairo site in the Middle East, carbonaceous aerosols are noted in our sample. Cairo in the Middle East is  
850 one of the highly polluted places among the megacities worldwide. Carbonaceous aerosols over Cairo are  
851 primarily emitted from burning of the agricultural waste in the Nile delta during the burning season from  
852 September through December (El-Metwally et al., 2008). The average SSA for these aerosols ranges from  
853  $0.89\pm 0.03$  to  $0.91\pm 0.05$  exhibiting weak spectral dependence from 340 nm to 646 nm. Average  $\tau_{440}$  and  $\alpha_{440-870}$   
854 are about 0.6 and 1.4 respectively with  $AAE_{340-646}$  about 1.9 during DJF. The spectral dependence noted for these  
855 aerosols likely indicate mixture of black and organic carbon amounts.

856

857 Over Northeastern China, carbonaceous aerosols are observed throughout the year at the sites Beijing and  
858 XiangHe. The spectral behavior of carbonaceous aerosols at these sites shows increase in SSA from 340 nm to  
859 466 nm and thereafter remains near constant or slightly decreases with an UV-Vis dependence ranging from 1.84  
860 to 2.14. However, significant seasonality is noted with minimum ( $0.95\pm 0.03$  at 466 nm) and maximum  
861 ( $0.90\pm 0.04$  at 466 nm) absorption during JJA and DJF, respectively. The increase in SSA is likely result of  
862 humidification and secondary aerosol processes during JJA. The high aerosol absorption noted during winter  
863 (DJF) is likely contributed from high amounts of local fossil fuel combustion and agricultural waste burning.

864  
865 For the sample obtained over Northern India, carbonaceous aerosols are observed over Kanpur and Gandhi  
866 College during SON and DJF. Emissions from crop residue burning during SON and biomass burning for  
867 residential heating in DJF prevail over the entire Indo-Gangetic plain and likely result in such absorption. In  
868 addition, industrial activities and vehicular emissions are observed throughout the seasons. The spectral behavior  
869 of these aerosols shows increase in average SSA from 340 nm ( $0.91\pm 0.02$ ) to 466 nm ( $0.93\pm 0.03$ ) and slight  
870 decrease till 646 nm ( $0.91\pm 0.04$ ). The regional average  $AAE_{340-646}$  for carbonaceous aerosols ranges 1.2 to 1.5.  
871 The weak spectral dependence of SSA noted here is consistent with AERONET SSA analysis. For the aerosols  
872 observed over the Kanpur site, spectral dependence of aerosols becomes nearly invariant (SSA  $\sim 0.89$  at 440 and  
873 675 nm) for high fine mode fraction (FMF  $\sim 0.85$ ) of aerosols (Eck et al., 2010).

874  
875 Over Northern Australia, carbonaceous aerosols are observed at the sites Jabiru and Lake Argyle during SON. In  
876 general, Northern and Western parts of the Australia are covered with savanna grasslands, woodlands, and  
877 forests, where biomass burning due to natural fires and land management practices are known to produce high  
878 aerosol emissions during the dry season (Scott et al., 1992; Mitchell et al., 2013). Our results indicate the average  
879 SSA for carbonaceous aerosols over Northern Australia increases with wavelength from  $0.87\pm 0.02$  (340 nm) to  
880  $0.89\pm 0.03$  (388 nm) and then decreases to  $0.87\pm 0.06$  (646 nm). Average  $\tau_{440}$  and  $\alpha_{440-870}$  for these aerosols is 0.65  
881 and 1.62, respectively, while the UV-Vis spectral dependence ( $AAE_{340-646}$ ) is noted to be  $\sim 1.42$ . Such behavior of  
882 fine-mode particles is likely a result of a mixture of black carbon and organic carbon amounts in the atmosphere,  
883 producing stronger (weaker) absorption in the UV (Vis) wavelengths.

## 884 **6.2 Dust**

885 For dust aerosols, minerals such as hematite and other form of oxides play role in scattering/absorption of  
886 particles. The absorbing nature of pure dust aerosols close to the source is sensitive to the presence of hematite

887 than other minerals at shorter wavelengths (Sokolik and Toon, 1999). In addition to the sedimentation of coarse  
888 aggregates, the dust aerosols observed away from the source are sometimes found to have mixed (internally or  
889 externally) with anthropogenic aerosols altering its absorbing nature. Studies show that the typical spectral  
890 behavior of dust absorption decreases with increasing wavelength primarily due to attributed to the larger size of  
891 the particles (Sokolik and Toon, 1996, 1999). However, the absorption of dust from different sources is known to  
892 vary depending on the mineral composition of the soil origin (Di Biagio et al., 2019). The seasonal average of  
893 spectral SSA for coarse-mode dust aerosols found over major aerosol environments are shown in Figure 12.

894  
895 For the dust aerosols sample obtained at Saharan region at the sites Tamanrasset and Saada, the average SSA is  
896 ~0.94 at 466 nm. The average SSA for these dust aerosols increase with wavelength from about  $0.86\pm 0.03$  at 340  
897 nm to  $0.97\pm 0.02$  at 646 nm. The seasonal average  $AAE_{340-646}$  for dust aerosols derived at these sites range from  
898 2.8 to 3.3 with no distinct seasonality in the average spectral SSA. Dust aerosols over the Middle East are  
899 observed for the sites Cairo, Solar Village and SEDE BOKER with an average SSA  $\sim 0.95\pm 0.02$  at 466 nm. From  
900 UV (340 nm) to visible (646 nm), the regional average SSA for dust over the Middle East sites ranges from  
901  $0.89\pm 0.03$  to  $0.98\pm 0.02$ , while the  $AAE_{340-646}$  ranges from 2.7 to 3.8. No distinct seasonality in SSA is found from  
902 our sample for the dust aerosols over Middle East. However, a slight increase in SSA at UV wavelengths is noted  
903 during winter (DJF). Examining individual sites reveal this feature corresponds to the aerosols over Solar\_Village  
904 (Table S1). The increase in SSA and high  $AAE_{354-388}$  noted for Solar\_Village during DJF likely indicates transport  
905 of aerosols from neighbouring regions.

906  
907 Over Sahel, dust aerosols are observed for several sites that include Agoufou, Banizoumbou, Dakar,  
908 IER\_Cinzana, Ilorin, Ouagauodu, and Zinder\_Airport. Located in the middle of Sahelian region through the west  
909 Africa, these sites are influenced by both dust and biomass burning emissions (Basart et al., 2009). It should be  
910 noted that for the identifying dust in this work, we use  $\alpha_{440-870} \leq 0.2$  derived from AERONET. The regional  
911 average spectral SSA for dust aerosols derived here resembles typical dust absorption curve (increase in SSA  
912 with wavelength,  $\sim 0.87-0.91\pm 0.03$  at 340 nm to  $0.95-0.97\pm 0.02$  at 646 nm) with absorption ranging  $\sim 0.93-$   
913  $0.94\pm 0.02$  at 466 nm. No distinct seasonality in absorption of dust aerosols is noted over Sahel. Average  $AAE_{340-}$   
914  $646$  for dust over Sahel ranges 2.0-2.3 for all seasons with an exception during SON where  $AAE_{340-646}$  is noted to  
915 be relatively less 1.57.

916

917 Over Northern India, dust aerosols are observed during spring (MAM) and summer (JJA) months for the sites at  
918 Jaipur and Kanpur, where the former site is in proximity to the Thar desert and the later site is influenced by the  
919 dust transport. The average SSA shows a steep increase from 340 nm ( $0.88\pm 0.02$ ) to 466 nm ( $0.95\pm 0.03$ ), and a  
920 relatively smaller increase from 466 nm to 646 nm ( $0.97\pm 0.02$ ). The dust aerosols noted here has average  $\tau_{440}$   
921  $0.73$  to  $0.77$  and exhibit  $AAE_{340-646}$  between 2.9 to 3.4. Compared to AERONET the absorption for dust aerosols  
922 derived here agrees well. Eck et al (2010) reports the coarse mode particles noted over Kanpur during pre-  
923 monsoon (MAM) months exhibit climatological average SSA  $\sim 0.89$  and  $0.95$  at 440 and 675 nm, respectively.

924  
925 For the samples obtained over northeastern China, dust aerosols are observed over Beijing and XiangHe during  
926 spring (MAM). The spectral curve of regional average SSA shows an increase from  $0.87\pm 0.03$  at 340 nm to  
927  $0.95\pm 0.03$  at 646 nm. The average  $AAE_{340-646}$  obtained for the dust aerosols at these sites is 1.56. Among the  
928 regional dust observations presented here (Figure 12), northeastern China exhibit high absorption in visible  
929 wavelengths ( $\omega_o \sim 0.93\pm 0.03$  and  $0.95\pm 0.03$  at 466 and 646 nm) and low  $AAE_{340-646}$ . It is likely that these coarse  
930 particles are influenced by black carbon components over such highly polluted environments and exhibit  
931 anomalously low AAE than dust particles noted for other regions. However, due to large particle size ( $\alpha_{440-870}$   
932  $\sim 0.09$ ) the spectral SSA noted still shows increasing SSA with wavelength. Chaudhry et al., (2007) reported  
933 insitu measurements of coarse mode particles over XiangHe during March-2005 that exhibits high absorption in  
934 visible wavelengths ( $\omega_o \sim 0.70-0.94$  at 450, 550 and 700 nm). Li et al., (2007) explained the variation in SSA for  
935 coarse particles during March-2005 over XiangHe is a result of synoptic fluctuation – passage of cold fronts that  
936 uplifted ground-level pollution to higher altitudes influencing the aerosol absorption. Similar low  $AAE_{340-646}$   
937 values for coarse mode ( $\alpha_{440-870} < 0.2$ , dust) are noted for few sites over the Sahelian region during DJF (burning  
938 season). It is likely that these coarse particles are influenced by black carbon amounts emitted from biomass  
939 burning.

### 940 **6.3 Urban aerosols**

941 Urban aerosols dominantly constitute sulfates and other forms of nitrate particles. Additionally, industrial  
942 emissions and fossil fuel combustion produces various forms of carbon that contribute to the overall optical  
943 properties of urban aerosols. Further the aerosol size growth due to increase in relative humidity in the  
944 atmosphere and coagulation processes are known to alter the absorbing nature of aerosols. The typical spectral  
945 SSA of urban aerosols decreases with increase in wavelength from UV-Vis spectrum (Bergstrom, 1972). The



946 seasonal average of spectral SSA for carbonaceous aerosols found over major aerosol environments are shown in  
947 Figure 13.

948  
949 Aerosols observed over the Central United States (US) at the sites Sioux Falls and Bondville are primarily  
950 produced from industrial activities and vehicle emissions. The average aerosol SSA for these urban aerosols  
951 shows increase (0.93-0.95±0.03) in the wavelength range 340-388 nm followed by a decrease (0.89±0.06 at 646  
952 nm) towards the visible wavelengths. Average  $\tau_{440}$  and  $\alpha_{440-870}$  are about 0.48 and 1.7 respectively, while the  
953 average  $AAE_{340-646}$  noted as 1.1. Urban aerosols over the Mid-East US are noted for the sites at GSFC,  
954 MD Science Center, and SERC. The average aerosol SSA noted for these sites increases from (0.93±0.02) 340  
955 nm to (0.95±0.03) 388 nm and then decreases attaining a maximum absorption ( $\sim$ 0.87±0.06) at 646 nm. The  
956 regional average SSA for the MAM and JJA months at 466 nm is 0.94±0.03 and 0.92±0.03, respectively.  
957 Seasonally, the average SSA noted for JJA follows typical ‘urban’ absorption curve, while the spectral SSA  
958 shows decrease in aerosol absorption in the visible wavelengths for MAM. Since there are no significant changes  
959 in aerosol sources during MAM and JJA over Mid-East US, the seasonal variation in SSA noted for visible  
960 wavelengths are likely stemming from weak absorption signal insufficient for a robust retrieval. Recall, for  
961 weakly absorbing aerosols the error in SSA retrievals at visible wavelengths are high due to identified factors  
962 from our analysis such as cloud contamination, surface reflectance, particle size, etc.

963  
964 For the site at the Mexico City, Central America urban aerosols are noted in our sample. In general, Mexico City  
965 is a densely populated urban location that is well known for its high pollution levels among the other megacities  
966 worldwide. In addition to the high concentration of aerosols from fossil fuel combustion throughout the year,  
967 Mexico City also experiences biomass-burning aerosols during the relatively dry months of March-May from  
968 local sources. The seasonal average aerosol SSA over Mexico City shows decrease in absorption (0.90±0.03 and  
969 0.91±0.05 at 340 and 388 nm) and increases towards the visible wavelengths during MAM. However, the spectral  
970 SSA noted for DJF (0.88±0.04 and 0.87±0.05 at 340 and 388 nm) and SON (0.88±0.04 and 0.85±0.05 at 340 and  
971 388 nm) deviates from known pattern and exhibit high absorption in the UV spectral range with nearly constant  
972 or slight increase in the visible spectral range. The average SSA at 466 nm for DJF, MAM and SON months are  
973 0.85±0.06, 0.86±0.06 and 0.84±0.06, respectively with average  $\alpha_{440-870}$  about 1.7 throughout the seasons.  
974 Although the seasonal average  $AAE_{340-646}$  of aerosols over Mexico City ranges from 0.94-1.27, the  $AAE_{354-388}$  is  
975 higher for (5.13) MAM than compared to (2.2) DJF and (0.6) SON months. It is likely that our aerosol samples  
976 obtained during MAM are influenced by biomass burning emissions from local neighbour sources exhibiting high

977 AAE<sub>354-388</sub> values suggesting the absorption is driven by organic components (Barnard et al., 2008). While for the  
978 samples noted during SON and DJF the aerosol absorption is primarily driven by the black carbon components  
979 emitted from the heavy industrial and vehicular fleet over the region.

980  
981 Our subset of samples noted urban aerosols at the site Sao Paulo, South America. Sao Paulo is the largest South  
982 American megacity with population exceeding 21 million inhabitants. Heavy industrial and vehicular emissions  
983 are the dominant source of aerosols observed over Sao Paulo. In addition, aerosols from northern parts of the  
984 Amazon Basin advecting south or southeast over Sao Paulo is not uncommon during the peak burning season  
985 (August through September). Average  $\tau_{440}$  and  $\alpha_{440-870}$  noted for these aerosols are  $\sim 0.57$  and  $1.5$ , respectively  
986 during JJA and SON months. Spectral SSA noted here resembles typical urban absorption curve during SON.  
987 However, during JJA spectral SSA shows steep decrease from 340 nm to 388 nm ( $0.87\pm 0.03$  to  $0.85\pm 0.03$ ) and  
988 remains nearly invariant towards the visible spectral range from 466 to 646 nm ( $0.85\pm 0.05$  to  $0.84\pm 0.07$ ). The  
989 average SSA noted at 466 nm is  $0.88\pm 0.06$  ( $0.92\pm 0.06$ ) for JJA (SON) months. It is noted that the urban aerosols  
990 at Sao Paulo are more absorbing during JJA than in SON.

991  
992 For the urban aerosols observed over Cairo and Nes Ziona in the Middle East, the average SSA is noted to be  
993  $0.89\pm 0.03$  at both 340 nm and 646 nm. While urban aerosols and pollution prevail over Nes Ziona, emissions  
994 from crop residue burning (rice straws) over the Nile delta region during winter (DJF) and heavy pollution dictate  
995 the aerosol absorption noted over Cairo. Average  $\tau_{440}$  and  $\alpha_{440-870}$  noted for these aerosols are  $\sim 0.60$  and  $1.3$ ,  
996 respectively throughout the seasons, with mean  $AAE_{340-646}$  ranging between  $1.5-2.0$ . Unlike typical urban aerosol  
997 absorption curve, the spectral SSA noted for these aerosols do not exhibit steep decrease from 388 to 646 nm  
998 indicating mixture of black carbon and organic carbon amounts.

999  
1000 Over Europe, dominantly urban aerosols are observed at the sites Avignon, Carpentras, Ispra, IMS-METU-ERI,  
1001 Lille, Lecce University, Minsk, Modena, Moldova, Moscow MSU, Palaiseau, Rome, and Thessaloniki.  
1002 Primarily industrial activities, vehicular emissions are dominant sources of aerosols over Europe. In addition, fuel  
1003 combustion for residential heating during winter and outbreak of episodic dust aerosols during spring-summer  
1004 over Iberian Peninsula and Mediterranean basin are known to influence the aerosol loading (Basart et al., 2009;  
1005 Mallet et al., 2013). The regional average SSA increases from 340 nm ( $0.91\pm 0.02$ ) to 388 nm ( $0.93\pm 0.03$ ) and  
1006 then decreases reaching a minimum value ( $0.87\pm 0.07$ ) at 646 nm for most seasons. The observed aerosol  
1007 absorption is similar for spring (MAM) and summer (JJA). However, there is an increase in absorption at

1008 wavelengths other than 646 nm for fall (SON) that reaches a maximum absorption  $0.86\pm 0.03$  and  $0.90\pm 0.04$  at  
1009 340 nm and 466 nm, respectively. These highly absorbing aerosols in our sample are noted over polluted urban  
1010 sites at Ispra, Modena, and Rome indicating the mixture of organic and black carbon amounts. Average  $\tau_{440}$  and  
1011  $\alpha_{440-870}$  noted for these aerosols are  $\sim 0.50$  and  $1.5$ , respectively throughout the seasons, with mean  $AAE_{340-646}$   
1012 ranging between 1.0-1.3.

1013  
1014 Over Northeastern China urban aerosols are noted for the sites Beijing, and XiangHe. Significant seasonality in  
1015 aerosol SSA is noted with minimum ( $0.93\pm 0.04$  at 466 nm) and maximum ( $0.88\pm 0.05$  at 466 nm) absorption  
1016 during JJA and DJF, respectively. Seasonal variability in aerosol absorption noted here is likely caused by the  
1017 humidification and secondary aerosol processes. In addition to the high industrial and vehicular emissions  
1018 throughout the year, fuel combustion for residential heating purposes and agricultural waste burning during DJF  
1019 adds additional aerosol burden over the Northeastern China. High aerosol absorption noted during DJF is likely a  
1020 result of such emissions. Average  $\tau_{440}$  and  $\alpha_{440-870}$  noted for these aerosols are  $\sim 0.65$  and  $1.4$ , respectively  
1021 throughout the seasons, with mean  $AAE_{340-646}$  ranging between 1.2-1.7. For the urban aerosols noted at the sites  
1022 Shirahama, and Osaka in Japan during MAM, the spectral aerosol SSA noted follows typical urban absorption  
1023 curve with slight increase in SSA from ( $0.90\pm 0.02$ ) 340 nm to ( $0.92\pm 0.02$ ) 388 nm and then a steep decrease  
1024 towards ( $0.87\pm 0.07$ ) 646 nm. Average  $\tau_{440}$  and  $\alpha_{440-870}$  noted for these aerosols are  $\sim 0.5$  and  $1.5$ , respectively with  
1025 mean  $AAE_{340-646}$  about 1.2.

1026  
1027 Over Northern India urban aerosols are noted for the sites at Jaipur, Kanpur, and Gandhi College. Major source  
1028 of aerosols over the region includes industrial and vehicular emissions, combustion of biomass and fossil fuels.  
1029 The average aerosol SSA increases from ( $0.88\pm 0.02$ ) 340 nm to ( $0.90\pm 0.04$ ) 466 nm and decreases towards 646  
1030 nm ( $0.87\pm 0.07$ ) 646 nm. These aerosols are noted to exhibit relatively high absorption  $\omega_o \sim 0.89\pm 0.04$  at 466 nm.  
1031 Average  $\tau_{440}$  and  $\alpha_{440-870}$  noted for these aerosols are  $\sim 0.60$  and  $1.4$ , respectively throughout the seasons, with  
1032 mean  $AAE_{340-646}$  about 1.3. Throughout the seasons, influence of pollution aerosols is clearly evident in the  
1033 observed aerosol absorption.

## 1034 **6-7 Discussion**

1035 Through extensive studies in the literature, it is known that optical properties of biomass burning aerosols depend  
1036 on fuel/vegetation type, combustion processes, available moisture content (e.g., Ward, 1992; Reid and Hobbs,

1037 1998; Reid et al., 1998; Eck et al., 2001). Such studies reported varying properties of aerosols emitted from two  
1038 phases of vegetation burning: flaming and smoldering. While flaming phase rapidly oxidizes the available  
1039 volatile hydrocarbons in the biomass, smoldering phase mostly requires a surface where slow diffuse oxygen  
1040 converts the biomass through exothermic reaction. In general, burning of grasslands happens dominantly through  
1041 flaming phase combustion process that emits high amounts of ~~soot~~black carbon, while smoldering combustion  
1042 prevail the burning of woodlands/deciduous forest that emits less ~~soot~~black carbon and more organic carbon. The  
1043 observed aerosol absorption at the biomass burning sites (~~figure~~Figure 611) clearly makes this distinction. Over  
1044 South America, in addition to the emissions from burning rainforest (near-by Alta Floresta and Ji Parana),  
1045 Cerrado (wooded grasslands) type vegetation dominates at most sites considered here (Schafer et al., 2008).  
1046 Biomass burning of tropical forests occur through smoldering combustion exhibiting aerosol  $\omega_o$  ( $0.93 \pm 0.03$  at  
1047 466 nm). Such biomass burning occurs through smoldering combustion exhibiting relatively high aerosol  $\omega_o$   
1048 ( $0.93$  at 466 nm). Compared to South America, the aerosols over Southern Africa have distinct seasonality and  
1049 high absorption ( $\omega_o \sim 0.88 \pm 0.02$  at 466 nm~~figure 6b~~). Eck et al (2013) demonstrated that this seasonality in  
1050 aerosol absorption is likely a result of shift in fuel type and combustion process. At the beginning of dry season  
1051 (starting June), the central region is prone to undergo a rapid burning through flaming process, while in the late  
1052 dry season (ends November) the wooded lands located southeastern parts begins to burn through dominantly  
1053 smoldering ~~process~~phase. For the savanna with open grasslands in the northern Australia, biomass burning  
1054 happens dominantly through flaming phase combustion producing high amounts of soot, as also noted in our  
1055 retrievals. Figure 140 shows the ~~regional average~~range of AAE obtained for carbonaceous aerosols at three  
1056 wavelength pairs. Overall, the average slope of absorption in visible ( $AAE_{466\_646}$ ) and UV-Vis ( $AAE_{340\_646}$ ) for  
1057 carbonaceous aerosols is found to be within 2. This is consistent with the studies that report AAE of biomass  
1058 burning aerosols from several field campaigns in the range 1 to 3 (Kirchstetter et al., 2004; Schnaiter et al., 2005;  
1059 Bergstrom et al., 2007; Clarke et al., 2007). However, the average  $AAE_{354\_388}$  obtained is high up to 4 for most  
1060 regions. This is likely a result of higher organic matter in the regional biomass types and highlights the  
1061 importance of UV spectral region in delineating such group of aerosols. Studies show that spectral dependence of  
1062 aerosol absorption in the UV-visible range can be high up to 6 for aerosols with organic compounds (Kirchstetter  
1063 et al., 2004; Bergstrom et al., 2007 and references therein). Among the biomass burning regions, for the  
1064 emissions where contribution of flaming phase combustion is high, the mean  $AAE_{354\_388}$  noted is relatively low–  
1065 Northern Australia (2.1) and Sahel (1.6). For the savanna grasslands in Australia where flaming combustion  
1066 prevails the mean  $AAE_{354\_388}$  is relatively low (2.2 or less) compared to other biomass burning regions. Further, it  
1067 is noted that carbonaceous aerosols observed over Northern India, Northeastern China, Sahel, and the Middle

1068 ~~East For the sample obtained at Ilorin in Sahel and Cairo in the Arabian Peninsula, the mean AAE was found up~~  
1069 ~~to 2 for all wavelength pairs. Further, it is noted that carbonaceous aerosols observed over Northern India,~~  
1070 ~~Northeastern China, Sahel, and Arabian Peninsula~~ has an average  $\alpha_{440-870} \sim 1.4$  (i.e, at the lower end of the fine-  
1071 mode range), while in South America, South Africa and Australia has average  $\sim 1.8$ . This indicates the role of  
1072 aerosol mixing and secondary processes in emanating the observed variability in aerosol absorption other than its  
1073 composition alone.

1074  
1075 Figure 15 shows the range of AAE obtained for dust aerosols at three wavelength pairs. Regional average of the  
1076 UV-Vis spectral dependence ( $AAE_{340-646}$ ) is found to be close to or greater than 3 for all regions, except for the  
1077 Sahel and Northeastern China, where average value ranges 1.5 to 2.5. Although no distinct seasonal variation in  
1078 spectral absorption of dust is noted, the variability in spectral dependence over the regions is quite evident. The  
1079 regional average of SSA for dust aerosols from 340 nm to 646 nm are  $0.87 \pm 0.02$  to  $0.98 \pm 0.02$ ,  $0.89 \pm 0.03$  to  
1080  $0.96 \pm 0.02$ ,  $0.89 \pm 0.02$  to  $0.98 \pm 0.01$ ,  $0.87 \pm 0.02$  to  $0.97 \pm 0.02$ , and  $0.87 \pm 0.03$  to  $0.95 \pm 0.03$  over Sahara, Sahel,  
1081 Middle East, Northern India, Northeastern China, respectively. While the average SSA ranges  $0.87-0.89 \pm 0.03$  at  
1082 340 nm, differences in retrieved SSA are evident with increasing wavelength. The regional average spectral SSA  
1083 noted at 466 nm are  $0.94 \pm 0.02$ ,  $0.93 \pm 0.02$ ,  $0.96 \pm 0.02$ ,  $0.95 \pm 0.02$  and  $0.93 \pm 0.02$  for Sahara, Sahel, MiddleEast,  
1084 Northern India, and Northeastern China, respectively. A recent study uses soil samples collected from over  
1085 different arid regions worldwide to characterize the mineral composition of dust and estimate the spectral SSA  
1086 using the measured scattering absorption coefficients through aethalometer operating at seven discrete  
1087 wavelengths from 370–950 nm (Di Biagio et al., 2019). Their study reports high absorption ( $\sim 0.70-0.75$  at 370  
1088 nm) for dust samples obtained over Niger, Mali, Southern Namibia, and Australia due to the presence of higher  
1089 amounts of iron oxides. While for the samples collected over Bodélé, Northern Namibia, Arizona the estimated  
1090 absorption ( $\sim 0.91-0.96$  at 370 nm) and amounts of iron oxides are relatively low. In comparison, our retrievals  
1091 indicate the dust aerosols noted at Sahel and northeastern China are highly absorbing, while those noted over  
1092 MiddleEast, and Northern India are less absorbing and dust over Sahara shows intermediate absorption. These  
1093 results can be explained by combination of varying mineral composition (iron oxide amounts) and mixing of dust  
1094 with other sources along the transport pathway. However, the magnitude of SSA and AAE reported by Di Biagio  
1095 et al., (2019) are lower than those retrieved here for all regions. The reason for the differences noted in SSA is  
1096 unknown and needs further investigation. In contrary, our retrieved SSA for coarse mode particles agrees well  
1097 with AERONET SSA (59% and 87% observations within  $\pm 0.05$  envelope at 466 nm and 646 nm respectively).  
1098 ~~For dust aerosols, minerals such as hematite and other form of oxides play role in scattering/absorption of~~

1099 particles. The absorbing nature of pure dust aerosols close to the source is sensitive to the presence of hematite  
1100 than other minerals at shorter wavelengths (Sokolik and Toon, 1999). In addition to the sedimentation of coarse  
1101 aggregates, the dust aerosols observed away from the source are often found to have mixed (internally or  
1102 externally) with anthropogenic aerosols altering its absorbing nature. For example,  $\omega_o$  of dust aerosols were  
1103 reported as 0.83 to 0.87 near the source and ~0.9 far away from the source over the tropical North Atlantic Ocean  
1104 using satellite measurements at 331 nm (Torres et al., 2002). For the dust aerosol samples obtained in this work,  
1105 high absorption at 340 nm with a regional average of ~0.86 is noted for Sahara, while relatively lower absorption  
1106 ~0.90 is noted for the Sahel region. The mixing of dust with biomass burning emissions and local pollutants over  
1107 the Sahel likely attributed to the observed low absorption. Similar low absorption (0.90) of dust at 340 nm is  
1108 observed for the Arabian Peninsula region. However, the UV-Vis spectral dependence noted at the Sahel, and  
1109 Arabian Peninsula varies significantly. Figure 11 shows the regional average AAE obtained for dust aerosols at  
1110 three wavelength pairs. Among the dust prone regions considered here, the regional average of the UV-Vis  
1111 spectral dependence (AAE<sub>340-646</sub>) is found to be close to or greater than 3 for all, except the Sahel and Eastern  
1112 China, where average value ranges 1.5 to 2.5. Although no distinct seasonal variation in spectral absorption of  
1113 dust is noted, the variability in spectral dependence over the regions is quite evident. Overall the regional average  
1114 of AAE for dust aerosols observed here is consistent with insitu measurements that report values ranging 1.5-2.0  
1115 to 3.5 (Bergstrom et al., 2004, 2007; Müller et al., 2009; Petzold et al., 2009). However, observations at  
1116 individual sites (Table 3) show that the spectral dependence of the observed dust for few sites is relatively high  
1117 than those reported by insitu measurements. Considering our retrieval method where aerosol absorption is  
1118 derived independently for each wavelength and have computed the dependence, our results agree reasonably well  
1119 with the insitu measurements reported in the literature.

1120  
1121 Figure 16 shows the regional average AAE obtained for urban aerosols at three-wavelength pairs. Urban aerosols  
1122 in highly polluted environments such as over the Mexico City have near unity spectral dependence. While the  
1123 passage of biomass burning emissions over such environment show unusual decrease in absorption at the UV  
1124 region attributing to high AAE<sub>354-388</sub>. While urban aerosols constitute dominantly sulfates and other forms of  
1125 nitrate particles, industrial emissions and fossil fuel combustion produces various forms of carbon that contribute  
1126 to the overall optical properties. Further the aerosol size growth due to increase in relative humidity in the  
1127 atmosphere and coagulation processes are known to alter the absorbing nature of aerosols. Figure 12 shows the  
1128 regional average AAE obtained for urban aerosols at three wavelength pairs. Urban aerosols in highly polluted  
1129 environments such as over the Mexico City have near unity spectral dependence. While the passage of biomass

1130 burning emissions over such environment show unusual decrease in absorption at the UV region attributing to  
1131 high  $AAE_{354\_388}$ . This is consistent with studies that report relatively high AAE in UV region and near unity in  
1132 visible region for the aerosol mixture consisting of organic matter and black carbon amounts (Barnard et al.,  
1133 2008; Vanderlei Martins et al., 2009; Bergstrom et al., 2010; Jethva and Torres, 2011). The urban aerosols found  
1134 in our sample over Northern Indian and Eastern-Northeastern China are highly absorbing exhibiting  $AAE_{340\_646}$   
1135  $\sim 1.5$  than the carbonaceous aerosols with  $AAE_{340\_646} \sim 2$ . These results suggest the combination of magnitude of  
1136 aerosol absorption and its spectral dependence in UV, visible and UV-Visible spectrum could be used to partition  
1137 mixture of aerosol types found in such environments. OverallOverall, the regional average UV-Visible AAE for  
1138 the urban aerosols is found to be near 2.

### 1139 **57 Comparison with AERONET SSA product**

1140 We compare of our aerosol SSA retrievals at the visible wavelengths with ~~that available~~those from the  
1141 AERONET data set. The comparison ~~of SSA retrievals~~is limited AERONET Level 2 reliable retrievals as  
1142 determined the aerosol load at 440 nm ( $\tau_{440} > 0.4$ ) to  $\tau_{440} > 0.4$  and where AERONET Level 2 inversion is  
1143 available. It should be notedWe emphasize here that since AERONET SSA is a derived quantity and cannot be  
1144 considered as ‘ground truth’, this comparison serves as a consistency check rather than a strict validation  
1145 exercise. The nearest AERONET wavelength available for comparison with OMI wavelength (388 nm) retrievals  
1146 is at 440 nm. To facilitate the comparison, we use the AAE computed from our retrievals at 388-466 wavelength  
1147 pair to transform the retrieved SSA at 388 nm to 440 nm ( $440_{OMI}$ ). While at the MODIS wavelengths 466 and 646  
1148 nm, the AERONET SSA were computed by linear interpolation of the values reported at 440 and 675 nm. For  
1149 ease in comparison, we transform the AERONET SSA at 440 and 675 nm to MODIS wavelengths of 466 and  
1150 646 nm respectively following interpolation.This conversion will unlikely introduce any notable bias in our  
1151 comparison as the difference in the nearest wavelengths of both data sets is very low (< 30 nm).

1152  
1153 First, we investigate the consistency of retrieved SSA for selected sites for which the prevailing aerosol types and  
1154 local source environment are well known as documented in several studies in the literature. We use the  
1155 AERONET SSA from time period 2005-2016 for the comparison. Figure 13-8 shows the comparison of average  
1156 spectral SSA (box plots showing lower and upper quartile of observations with white line representing its mean  
1157 value) for three distinct aerosol types sampled from the selected sites. The mean AAE derived for the visible  
1158 wavelength pair at 466-646 nm agrees within 0.5 or less with the AERONET values for all aerosol types and sites

1159 ~~considered. However, it should be noted that AAE is highly susceptible to small changes in the retrieved SSA for~~  
1160 ~~both data sets. The vertical bars shown in the figure corresponds to the uncertainty estimate provided by~~  
1161 ~~AERONET ( $\pm 0.03$ ) and the estimate due to change in  $\pm 0.01$  surface reflectance for our retrieved SSA ( $\pm 0.05$ ).~~  
1162 For dust aerosols, the retrieved average SSA show good agreement with AERONET SSA obtained at Dakar and  
1163 Ouagadougou (within  $\pm 0.008$ ), while the differences in SSA (retrieved minus AERONET) at Tamanrasset and  
1164 Solar Village are  $+0.02$  and  $+0.05$ , respectively at 466 nm. Although our sample is limited for few days that met  
1165 our criteria for subset, the observed differences in SSA are within the overall uncertainty estimates ~~as also noted~~  
1166 ~~from the overlapping error bars presented in the previous section.~~ For carbonaceous particles, the retrieved and  
1167 AERONET SSA agree well with differences of less than 0.015 for Alta\_Floresta, CUIABA and Mongu.  
1168 However, notable difference in SSA (0.045) is observed at the Lake\_Argyle with high absorption for AERONET  
1169 than our retrieved value at both 466 nm and 646 nm. ~~It should be noted that, while AERONET version 3 employs~~  
1170 ~~surface reflectance following BRDF parameters from MODIS BRDF/Albedo CMG Gap Filled Snow Free~~  
1171 ~~Product MCD43GF (Sinyuk et al., 2020), we use MODIS MCD19A1 surface reflectance product. For absorbing~~  
1172 ~~aerosols, the SSA differences observed here is likely a result of different surface reflectances data employed by~~  
1173 ~~the two data sets.~~ For urban/industrial aerosols at GSFC, Avignon, Moldova and Cairo the retrieved SSA agrees  
1174 within the uncertainty estimates. Particularly notable difference ( $-0.05$ ) is found for Avignon at 646 nm. It should  
1175 be noted that, while AERONET Version 3 employs surface reflectance using BRDF parameters from MODIS  
1176 BRDF/Albedo CMG Gap-Filled Snow-Free Product MCD43GF (Sinyuk et al., 2020), we use MODIS  
1177 MCD19A1 BRF/surface reflectance product. The use of different surface reflectances data in the two retrieval  
1178 algorithms likely contributes to some of the observed SSA differences. As discussed in the sensitivity analysis,  
1179 surface reflectance is the second higher source of errors in the retrieved SSA at visible wavelengths after cloud  
1180 contamination. Additionally, for our retrievals as well as with AERONET, uncertainties in SSA increases with  
1181 wavelength for fine mode particles (carbonaceous and urban) with high EAE, notably for weakly absorbing  
1182 aerosols. For example, the  $\pm 0.03$  uncertainty in AERONET SSA at 440 nm for  $\tau_{440} \sim 0.4$  is achieved for the NIR  
1183 wavelength (1020 nm) at  $\tau_{440} \sim 0.6$  for the fine mode particles observed over the GSFC site (Sinyuk et al., 2020).  
1184

1185 Figure 14-9 shows the absolute difference in retrieved SSA comparison of retrieved SSA with AERONET as a  
1186 function of  $\tau_{440}$  for all collocated observations. For the SSA at 440 nm<sub>OMI</sub>, the observations within  $\pm 0.03$  ( $\pm 0.05$ )  
1187 envelopes are 38% (62%), 42% (66%) and 39% (66%) for dust, carbonaceous, and urban aerosol types,  
1188 respectively. As expected, the SSA difference is highest for lower AOD's and decreases with increasing aerosol  
1189 load for all aerosol types. For the SSA at 466 nm, the observations within  $\pm 0.03$  ( $\pm 0.05$ ) envelopes are 25%



1190 (59%), 38% (63%) and 34% (56%) for dust, carbonaceous, and urban aerosol types, respectively. In terms of  
1191 particle sizes, the difference in SSA for fine-mode ( $\alpha_{440-870} > 1.2$ ) are noted to be more scattered than the  
1192 difference in SSA noted for coarse mode particles. For the SSA at 646 nm, the observations within  $\pm 0.03$  ( $\pm 0.05$ )  
1193 envelopes are 73% (87%), 39% (60%) and 28% (45%) for dust, carbonaceous, and urban aerosol types,  
1194 respectively. Spectrally, there is significantly more scatter at 646 nm with SSA differences in the range -0.3 to 0.2  
1195 for urban aerosols. The RMSE of the SSA ranges from 0.04 to 0.09 with lowest error for dust particles at 646 nm  
1196 and highest error for urban aerosols at the same wavelength. This owes to the high spectral AODs for coarse  
1197 mode particles through UV-Visible where sufficient absorption signal is available for the retrieval of SSA. While  
1198 for fine mode particles with the decreasing spectral AODs from UV-Visible the absorption signal becomes weak,  
1199 particularly notable for the less absorbing (urban) aerosols. For the SSA at 466 nm, the observations within  $\pm 0.03$   
1200 ( $\pm 0.05$ ) envelopes are 26% (59%), 38% (63%) and 34% (56%) for dust, carbonaceous, and urban aerosol types,  
1201 respectively. It is observed that for dust aerosols, the observations are well concentrated in a narrow range of SSA  
1202 (retrieved values range from 0.92 to 0.99 at 466 nm), while for carbonaceous (0.86 to 0.95 at 466nm) and urban  
1203 types (0.84 to 0.98 at 466 nm) the observations become increasingly scattered through wide range of SSA. In  
1204 terms of wavelength, the observations at 646 nm are more widely scattered than for the SSA at 466 nm for all  
1205 aerosol types. For the SSA at 646 nm, the observations within  $\pm 0.03$  ( $\pm 0.05$ ) envelopes are 73% (87%), 39%  
1206 (60%) and 28% (45%) for dust, carbonaceous, and urban aerosol types, respectively. The root mean square error  
1207 (RMSE) of the SSA ranges from 0.04 to 0.09 with lowest error for dust particles at 646 nm and highest error for  
1208 urban aerosols at 646 nm. Among the aerosol types, it is noted that our retrieved SSA for dust at 466 nm has  
1209 relatively high scattering than AERONET, while for other types the results are found to be consistent.

1210  
1211 Figure 10a shows the absolute difference in retrieved and AERONET SSA as a function of optical depth by  
1212 combining all aerosol types together. The differences in SSA for all wavelengths at 440<sub>OMI</sub>, 466 and 646 nm are  
1213 higher for lower  $\tau$  and become negligible for higher  $\tau$ . It is observed that at 440<sub>OMI</sub> and 466 nm the observations  
1214 with positive differences are relatively more than that at 646 nm. Comparison of the retrieved SSA with  
1215 AERONET SSA for all aerosol types is shown in Figure 10b. Our retrieved SSA at 440<sub>OMI</sub>, 466 and 646 nm  
1216 agree within  $\pm 0.03$  of AERONET SSA for 39% (0.05), 34% (0.06) and 34% (0.08) of observations (RMSE)  
1217 respectively.

1218 Figure 15a shows the comparison of retrieved SSA with AERONET by combining all aerosol types together. Our  
1219 retrieved SSA at 466 and 646 nm agree with AERONET SSA for 34% (0.06) and 40% (0.07) of observations  
1220 (RMSE) respectively. The absolute difference in retrieved and AERONET SSA as a function of optical depth is

1221 shown in figure 15b. The differences in SSA for both wavelengths at 466 nm and 646 nm are higher for lower  $\tau$   
1222 and become negligible for higher  $\tau$ . It is observed that at 466 nm, the observations with positive differences are  
1223 relatively more than that at 646 nm.

1224 It is important to note here that our retrieval method and that used in the AERONET inversion differ  
1225 fundamentally in several aspects. (i) different source of surface reflectance data used, (ii) instantaneous particle  
1226 sizes derived from sky radiances measurements by AERONET and climatological average particle sizes used in  
1227 our retrievals, (iii) the use multi-spectral sky radiance measurements by AERONET along the almucantar plane  
1228 (multi-angular) and the single-view TOA radiance measurements by the satellites in our retrievals, and (iv) the  
1229 use of weak constraint on spectral variation of imaginary refractive index for the fine mode particles in  
1230 AERONET SSA retrievals, while our retrievals of SSA are carried for each wavelength independently. The use  
1231 of weak constraint on spectral variation of imaginary refractive index for the fine mode particles with high  $\alpha_{440-870}$   
1232 for AERONET inversion owes to the lower spectral AODs and diminished absorption signal strength at higher  
1233 wavelengths insufficient for a robust absorption retrieval (Dubovik et al., 2006). The differences noted between  
1234 our SSA retrievals and that from AERONET at different wavelengths could stem from one or more sources of the  
1235 differences listed above.~~It is important to note here that our retrieval method and that used in the AERONET~~  
1236 ~~inversion differ fundamentally in several aspects. Other than the different source of surface reflectance data used,~~  
1237 ~~one major distinction between the two methods is that the present inversion algorithm retrieves SSA at different~~  
1238 ~~UV and Vis wavelengths independently, whereas the AERONET algorithm internally applies a condition~~  
1239 ~~ensuring the spectral shape of SSA follows an expected pattern for the observed aerosol type (private~~  
1240 ~~communication with Thomas Eck, NASA GSFC). The differences noted between our SSA retrievals and that~~  
1241 ~~from AERONET at different wavelengths could be attributed, at least partially, to this treatment of reported~~  
1242 ~~values of SSA.~~

## 1243 8 Summary

1244 Ground-based measurements of direct and diffuse solar radiation under cloud-free conditions over worldwide  
1245 sites are providing valuable insights into regional aerosol characteristics. Long-term measurements obtained from  
1246 such network, such as from AERONET, are widely used to develop regional aerosol climatology and investigate  
1247 seasonal/annual variability, in addition to providing validation data set for the satellite-based AOD retrievals.  
1248 Satellite measurements of TOA radiances are able to provide global distribution of columnar aerosol amounts.  
1249 However, deriving aerosol optical properties from satellite measurements require constraints on particle sizes and

1250 optical properties. Reliable aerosol measurements from ground-networks and airborne/field campaigns are  
1251 traditionally used to validate and improve the constraints in satellite aerosol retrievals.

1252  
1253 In this work, we use AERONET measured extinction  $\tau$  as constraint in a robust inversion technique that uses  
1254 satellite measured TOA radiances from OMI and MODIS to derive spectral aerosol absorption in the UV-Vis part  
1255 of the spectrum. Other than cloud contamination of the TOA radiances, major sources of error in our retrieved  
1256 SSA come from surface reflectance, and aerosol layer height. We use TOA radiance observations with minimal  
1257 or no cloud-contamination reported by both OMI and MODIS products. Sensitivity tests show that our retrieved  
1258 aerosol SSA has reliable accuracy up to  $\pm 0.043$  from UV-Visible wavelengths for absorbing aerosols  
1259 (carbonaceous and dust) with  $\tau_{440} > 0.4$ . ~~However~~ ~~However~~, for less-absorbing aerosols the error in SSA retrieval  
1260 reaches up to  $\pm 0.075$ . ~~For the sites with low aerosol loading (i.e.,  $\tau_{440}$  up to 0.2) the accuracy of retrieved SSA~~  
1261 ~~reaches up to  $\pm 0.05$  at the UV wavelengths, while in the visible it exceeds  $\pm 0.05$ .~~ Using a subset of results where  
1262 SSA is retrieved independently for 340, 354, 388, 466 and 646 nm wavelengths for the same day with  
1263 observations  $\tau_{440} > 0.4$ , we examine the seasonal variability in aerosol SSA and derive spectral dependence  
1264 (AAE) at three wavelength pairs in the UV-Vis spectrum.

1265

1266 Key observations noted from the spectral aerosol absorption data set derived here are highlighted below:

#### 1267 **Biomass burning aerosols**

- 1268 a) Among sites dominated by biomass burning aerosols, Mongu in Southern Africa has high absorption  $\omega_o$   
1269  $\sim 0.85 \pm 0.02$  and  $0.84 \pm 0.05$  at 340 nm and 646 nm, respectively.  
1270 ~~a)  $\sim 0.85$  to  $0.84$  from 340 nm to 646 nm.~~
- 1271 b) Strong seasonality in absorption of carbonaceous aerosols is evident in Southern Africa indicating the  
1272 role of biomass types and combustion process. The average  $\omega_o$  from 340 nm to 646 nm during JJA and  
1273 SON are  $\sim 0.85 \pm 0.02$  ( $0.88 \pm 0.02$ ) and  $0.83 \pm 0.05$  ( $0.87 \pm 0.05$ ), respectively.  ~~$\sim 0.85$  to  $0.84$ , and  $\sim 0.88$  to~~  
1274  ~~$0.89$ , respectively.~~
- 1275 c) Carbonaceous aerosols found over Northern Australia are as strongly absorbing ( $\omega_o$   $\sim 0.87 \pm 0.03$  and  
1276  $0.86 \pm 0.06$  at 340 nm and  $\sim 0.87$  to  $0.86$  from 340 nm to 646 nm) as smoke over Southern Africa but has  
1277 relatively constant absorption from UV-Vis spectra.
- 1278 d) Carbonaceous aerosols found over Alta\_Floresta in the Amazon Basin have similar absorption ( $\omega_o$   
1279  $\sim 0.89 \pm 0.02$  and  $0.91 \pm 0.06$  at 340 nm and 646 nm  ~~$\sim 0.89$  to  $0.91$  from 340 nm to 646 nm~~) and AAE<sub>340-</sub>  
1280 <sub>646</sub>UV-Vis AAE (1.8) to those found over Missoula in Northestern US-America.

- 1281 e) Highly absorbing carbonaceous aerosols ( $\omega_0 \sim 0.86$  to  $0.87$  from  $340$  nm to  $646$  nm) with weak spectral  
 1282 dependence are found in Cairo ( $\omega_0 \sim 0.89 \pm 0.02$  and  $0.91 \pm 0.06$  at  $340$  nm and  $646$  nm) and Ilorin ( $\omega_0$   
 1283  $\sim 0.86 \pm 0.02$  and  $0.87 \pm 0.03$  at  $340$  nm and  $646$  nm) in the Middle East and Sahel region during winter  
 1284 (DJF). ~~These carbonaceous particles exhibit mean  $\alpha_{440-870} < 1.4$ , indicating possible mixture of fine and~~  
 1285 ~~coarse modes.~~
- 1286 f) Carbonaceous aerosols found over Northern India ( $\omega_0 \sim 0.91 \pm 0.02$  and  $0.92 \pm 0.04$  at  $340$  nm and  $646$  nm),  
 1287 Eastern-Northeastern China ( $\omega_0 \sim 0.87$ - $0.90 \pm 0.03$  and  $0.90$ - $0.94 \pm 0.05$  at  $340$  nm and  $646$  nm), Sahel ( $\omega_0$   
 1288  $\sim 0.86 \pm 0.02$  and  $0.87 \pm 0.03$  at  $340$  nm and  $646$  nm), and Arabian Peninsula exhibiting Middle East ( $\omega_0$   
 1289  $\sim 0.89 \pm 0.03$  and  $0.92 \pm 0.05$  at  $340$  nm and  $646$  nm)  ~~$\sim 0.86$  to  $0.89$  at  $340$  nm) has low average  $\alpha_{440-870}$  ( $< 1.4$ )  
 1290 than over other prominent biomass burning regions, suggesting mixture of fine and coarse modes.~~
- 1291 g) Distinct seasonality in spectral absorption of carbonaceous and urban aerosols is noted for Eastern China.  
 1292 ~~For carbonaceous aerosols, the~~ The maximum and (minimum) absorption is noted during DJF (JJA)  
 1293 exhibiting  $\omega_0 \sim 0.87 \pm 0.02$  and  $0.90 \pm 0.05$  at  $340$  nm and  $646$  nm ( $\omega_0 \sim 0.91 \pm 0.02$  and  $0.94 \pm 0.04$  at  $340$  nm  
 1294 and  $646$  nm), respectively. at  $466$  nm are found during DJF  $\sim 0.90$ , and JJA  $\sim 0.95$ , respectively.

## 1297 Dust aerosols

- 1298 a) For desert dust aerosols, the SSA is known to increase with wavelength from UV to Visible spectrum. No  
 1299 distinct seasonality in SSA is noted. The regional average of SSA for dust aerosols from  $340$  nm to  $646$   
 1300 nm are  $0.87 \pm 0.02$  to  $0.98 \pm 0.02$ ,  $0.89 \pm 0.03$  to  $0.96 \pm 0.02$ ,  $0.89 \pm 0.02$  to  $0.98 \pm 0.01$ ,  $0.87 \pm 0.02$  to  $0.97 \pm 0.02$ ,  
 1301 and  $0.87 \pm 0.03$  to  $0.95 \pm 0.03$  over Sahara, Sahel, Middle East, Northern India, Northeastern China,  
 1302 respectively. The regional averages of SSA for dust aerosols from  $340$  nm to  $646$  nm are  $0.86$  to  $0.98$ ,  
 1303  $0.88$  to  $0.96$ , and  $0.90$  to  $0.99$  over Sahara, Sahel and Arabian Peninsula respectively.
- 1304 b) Among the dust dominated regions considered here, our retrievals indicate relatively high absorption ( $\omega_0$   
 1305  $\sim 0.93 \pm 0.03$  at  $466$  nm) for aerosols noted at Sahel and northeastern China, while those noted over  
 1306 Middle East, and Northern India ( $\omega_0 \sim 0.97 \pm 0.03$  at  $466$  nm) are less absorbing and dust over Sahara ( $\omega_0$   
 1307  $\sim 0.95 \pm 0.02$  at  $466$  nm) shows intermediate absorption. These results can be explained by combination of  
 1308 varying mineral composition (iron oxide amounts) and mixing of dust with other sources along the  
 1309 transport pathway.
- 1310 ~~b) For regions where sources are nearby the sites (Sahara, Middle East/Arabian Peninsula) the AAE for all~~  
 1311 ~~three wavelength pairs considered exhibit high average values ( $>> 3$ ), while for regions where lofted dust~~

1312 through long range transport is observed (Northern India, Eastern China, parts of Sahel) the average  
1313 AAE has a range between 2 to 3.

### 1315 Urban aerosols

- 1316 a) Urban aerosols ( $\omega_o \sim 0.87 \pm 0.04$  and  $0.84 \pm 0.08$  at 340 nm and 646 nm) are highly absorbing and exhibit  
1317 distinct seasonality (higher absorption during JJA than SON) at Sao Paulo, South America. The urban  
1318 aerosols noted here are likely mixtures of carbonaceous particles transported over the region and  
1319 prevailing pollution from local sources.
- 1320 b) Polluted aerosols observed over Mexico City show high absorption in UV extending to visible spectrum  
1321 during DJF ( $\omega_o \sim 0.88 \pm 0.04$  and  $0.84 \pm 0.09$  at 340 nm and 646 nm) and SON ( $\omega_o \sim 0.88 \pm 0.04$  and  
1322  $0.80 \pm 0.08$  at 340 nm and 646 nm) months.
- 1323 c) Polluted aerosols noted over the Middle East are highly absorbing and exhibit weak spectral dependence  
1324 ( $\omega_o \sim 0.87-0.89 \pm 0.03$  and  $0.87-0.90 \pm 0.07$  at 340 nm and 646 nm).
- 1325 d) Polluted aerosols noted over the Northern India are highly absorbing and exhibit weak spectral  
1326 dependence ( $\omega_o \sim 0.88 \pm 0.03$  and  $0.87 \pm 0.07$  at 340 nm and 646 nm).
- 1327 a) ~~Urban aerosols ( $\omega_o(340) \sim 0.87$ ) are more absorbing than the carbonaceous aerosols ( $\omega_o(340) \sim 0.90$ ) and~~  
1328 ~~exhibit distinct seasonality (higher absorption during JJA than SON) in South America at CUIABA and~~  
1329 ~~Sao Paulo. The urban aerosols noted here are likely mixtures of carbonaceous particles transported over~~  
1330 ~~the region and prevailing pollution from local sources.~~
- 1331 b) ~~High absorption from UV ( $\omega_o(340) \sim 0.87$ ) extending to the blue spectral region ( $\omega_o(466) \sim 0.90$ ) is noted~~  
1332 ~~for most sites in Europe during SON and/or DJF months indicating the presence of organic matter due to~~  
1333 ~~local fuel combustion sources.~~
- 1334 e) ~~Polluted aerosols observed over Mexico City shows high absorption in UV ( $\omega_o(340) \sim 0.88$ ) extending to~~  
1335 ~~visible ( $\omega_o(466) \sim 0.85$ ) spectrum during DJF and SON months.~~

1336

1337 As mentioned, the results presented here are limited to our subset of retrievals, where SSA is retrieved for all five  
1338 wavelengths from UV-Visible range with  $\tau_{440} > 0.4$ . ~~However, relaxing the  $\tau_{440}$  up to 0.2 (see supplementary~~  
1339 ~~materials) does not yield significant changes in the derived variability of aerosol absorption and its spectral~~  
1340 ~~dependence.~~ Since one of our objectives is to derive UV-Visible spectral dependence (AAE) of aerosols, ~~without~~  
1341 ~~prior assumptions,~~ and given the inherent sampling bias of the OMI, MODIS and AERONET collocations – the  
1342 analysis method employed here is well justified. In other words, we made use of the best available data synergy

1343 and derive unique aerosol absorption data set with no prior assumptions on wavelength dependence, which  
1344 otherwise is assumed in the standard satellite-based aerosol retrieval algorithm. ~~Although It should be noted that~~  
1345 ~~our results may be biased toward dense pollution/industrial, smoke, and dust events, less absorbing low aerosol~~  
1346 ~~amounts seldom have dependency on the  $\omega_0$ .~~ In addition, the regional aerosol absorption derived here may not be  
1347 representative of the entire region due to limited sampling and fewer sites used. However, these absorption  
1348 models offer essential guidance for selecting spectral absorption in satellite aerosol retrievals using UV (OMI),  
1349 Vis (MODIS), and even spanning the UV-Vis spectrum, such as planned under the upcoming PACE  
1350 mission. ~~Therefore, the regional aerosol absorption models derived here offer essential guidance for selecting~~  
1351 ~~spectral absorption in satellite aerosol retrievals spanning the UV-Vis spectrum.~~ From our analysis of worldwide  
1352 inland sites: (a) it is suggested that satellite aerosol retrieval techniques could employ regional dynamic  
1353 absorption models to avoid potential bias in  $\tau$  retrievals noted in earlier studies, and (b) the spectral dependence  
1354 of aerosol absorption noted here for the UV (354-388 nm), visible (466-646 nm) and UV-Visible (340-646 nm)  
1355 range for all aerosol types other than black carbon varies considerably. Overall, the UV absorption data set well  
1356 compliments and provides more information on the regional aerosol absorption than with the visible data set  
1357 alone.

1358

1359 Given the lack of aerosol absorption information at near-UV wavelengths in the currently existing AERONET  
1360 ~~record-inversion products~~ and limited availability of insitu measurements, the UV-Vis aerosol absorption data set  
1361 developed here, perhaps for the first time, offers a valuable source of information useful for a variety of aerosol  
1362 and trace gas studies. As mentioned earlier the newer models of AERONET sunphotometers include sky radiance  
1363 measurements at 380 nm, and the derived SSA at this near UV wavelength is expected in the future upgrade of  
1364 AERONET inversion product. ~~The analysis presented here focuses on regional aerosol absorption using a subset~~  
1365 ~~of results.~~ The derived spectral dependency can be used with either subset of the results or all SSA retrievals to  
1366 construct and investigate long-term trends in UV-Visible aerosol absorption. Further, the spectral aerosol SSA  
1367 derived here could be used to parameterize absorption in models and better understand the radiative effect of  
1368 aerosols. Our ongoing investigation utilizing the complete data set developed here will explore some of these  
1369 applications in the future.

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1374 Co-PIs of the individual AERONET sites that were used in this work.

1375 **Competing Interests**

1376 The authors declare no conflict of interests.

1377 **Author Contributions**

1378 Omar Torres (OT) and Hiren Jethva (HJ) had conceptualized the research. Vinay Kayetha (VK) developed the  
1379 data set, performed formal analysis, and wrote the manuscript with inputs from OT and HJ. All authors reviewed  
1380 results, helped with the data interpretation and edited the manuscript to make a final version.

1381 **Data Availability**

1382 The spectral aerosol absorption data set developed here will be made available upon request to the authors.  
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1645 Table 1: Description of the ground and satellite data products used in this work.

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<i>Purpose</i>	<i>Instrument</i>	<i>Product</i>	<i>Level &amp; Version</i>	<i>Parameter(s)</i>
For SSA retrieval in this work	AERONET	AOD	L2, V3	AOD and extinction Ångström exponent.
		Inversion	L2, V2	Particle size distributions, and real part of refractive index at 440 nm.
	OMI	OMLERWAVE, OMAERUV	L2, V1.8.9.1	TOA reflectances (with QFs), Aerosol type, LER, Aerosol layer height obtained from CALIPSO.
	MODIS	MYD04	L2, C006	TOA reflectances (with QFs) provided by Deep-Blue algorithm.
		MAIAC - MCD19A1	L2, C006	Surface reflectance at 466 and 646 nm.
Comparison	AERONET	Inversion	L2, V3	SSA at 440 and 675 nm.

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1649 Table 2: Theoretical estimated uncertainties in the retrieval of aerosol SSA due to error in the input  
1650 variables. Configuration of sensitivity tests are: SZA = 20°, VZA = 40°, RAA = 130°, and  $\omega_o(388) = 0.9$ .

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	Input Uncertainty	Theoretical SSA Uncertainty ( $\Delta\omega_0$ ) for $\tau_{440} = 0.4$					
		$\lambda = 340$ nm			$\lambda = 646$ nm		
		Carb.	Dust	Urban	Carb.	Dust	Urban
<i>Extinction AOD</i>	$\lambda < 400$ nm, $\Delta\tau = \pm 0.02$ $\lambda > 400$ nm, $\Delta\tau = \pm 0.01$	0.002	0.001	0.002	0.009	0.007	0.011
<i>Particle sizes</i>	$\Delta\text{VMR} = \pm 20\%$	0.018	0.003	0.014	0.044	0.0006	0.040
<i>Real part of RI.</i>	$\Delta\text{RRI} = \pm 0.04$	0.007	0.007	0.009	0.001	0.002	0.002
<i>Calibration of TOA measurements</i>	OMI = $\pm 1.8\%$ MODIS = $\pm 1.9\%$	0.026	0.021	0.027	0.020	0.027	0.037
<i>Surface reflectance</i>	$\Delta\rho_{\text{surf}} = \pm 0.01$	0.006	0.011	0.006	0.032	0.022	0.050
<i>Aerosol layer hgt.</i>	$\Delta\text{ALH} = \pm 1$ km	0.021	0.028	0.006	0.001	0.001	0.0006
<i>Presence of cloud</i>	$\tau_{\text{cloud}} = 0.5$	0.016	0.020	0.017	0.042	0.041	0.056
<i>Trace gaseous absorption</i>	$\lambda = 466$ nm, $\tau_{\text{gas}} = 0.004$ $\lambda = 646$ nm, $\tau_{\text{gas}} = 0.034$	-	-	-	0.011	0.009	0.024
<i>Surface pressure</i>	$\pm 12$ mb/hPa	0.011	0.011	0.011	0.004	0.0004	0.006
<b>Combined theoretical uncertainty</b>		<b><math>\pm 0.043</math></b>	<b><math>\pm 0.043</math></b>	<b><math>\pm 0.038</math></b>	<b><math>\pm 0.073</math></b>	<b><math>\pm 0.055</math></b>	<b><math>\pm 0.088</math></b>

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**Table 3: Theoretical uncertainties in the computation of Absorption Ångström Exponent.**

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		$\omega_0$ overestimation				$\omega_0$ underestimation		
		$\Delta\omega_0$	$\Delta AAE_{354-388}$	$\Delta AAE_{466-646}$	$\Delta AAE_{340-646}$	$\Delta AAE_{354-388}$	$\Delta AAE_{466-646}$	$\Delta AAE_{340-646}$
<b>Carbon.</b>	$\tau_{440} = 0.4$	0.01	-0.022	-0.021	-0.021	0.018	0.017	0.018
	$EAE_{340-646} = 1.9$	0.02	-0.051	-0.046	-0.048	0.034	0.031	0.032
	$\omega_0(388) = 0.90$	0.03	-0.087	-0.078	-0.081	0.046	0.044	0.045
	$AAE_{340-646} = 1.7$	0.04	-0.135	-0.119	-0.126	0.058	0.054	0.056
<b>Dust</b>	$\tau_{440} = 0.4$	0.01	0.228	0.688	0.488	-0.190	-0.423	-0.325
	$EAE_{340-646} = 0.2$	0.02	0.506	2.067	1.369	-0.351	-0.711	-0.562
	$\omega_0(388) = 0.90$	0.03	0.854	8.769	5.024	-0.489	-0.922	-0.744
	$AAE_{340-646} = 2.5$	0.04	1.302	-	-	-0.609	-1.082	-0.889
<b>Urban</b>	$\tau_{440} = 0.4$	0.01	-0.117	-0.077	-0.092	0.095	0.066	0.077
	$EAE_{340-646} = 1.9$	0.02	-0.265	-0.166	-0.204	0.173	0.124	0.144
	$\omega_0(388) = 0.90$	0.03	-0.458	-0.272	-0.343	0.239	0.175	0.201
	$AAE_{340-646} = 0.9$	0.04	-0.722	-0.399	-0.522	0.295	0.221	0.250

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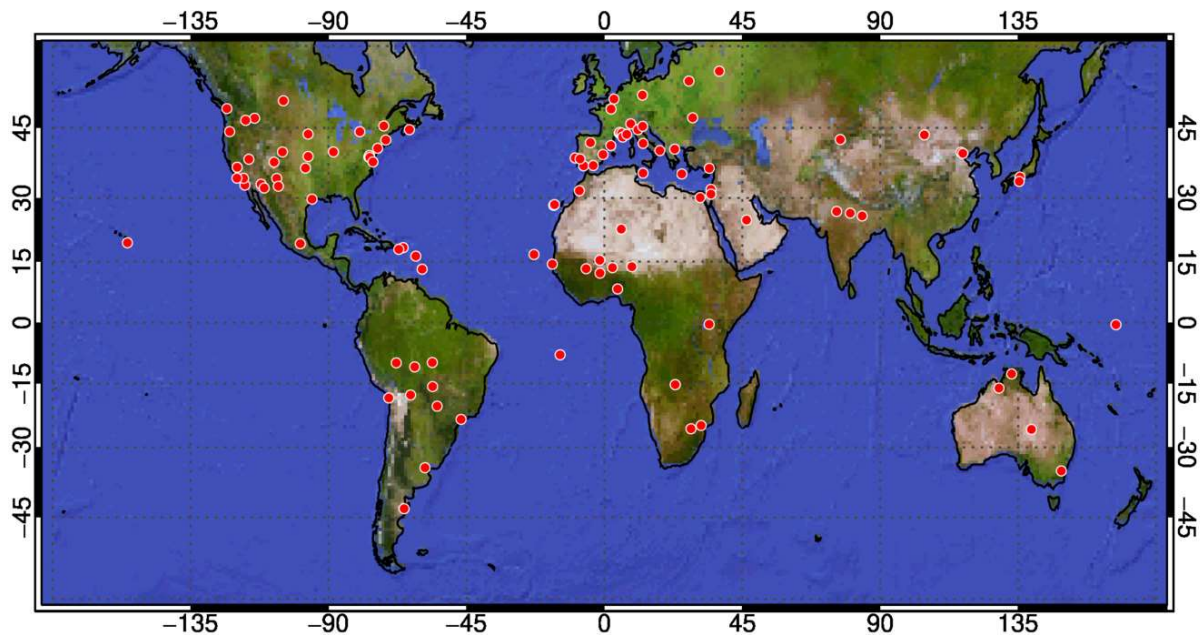
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1663 **Figures**



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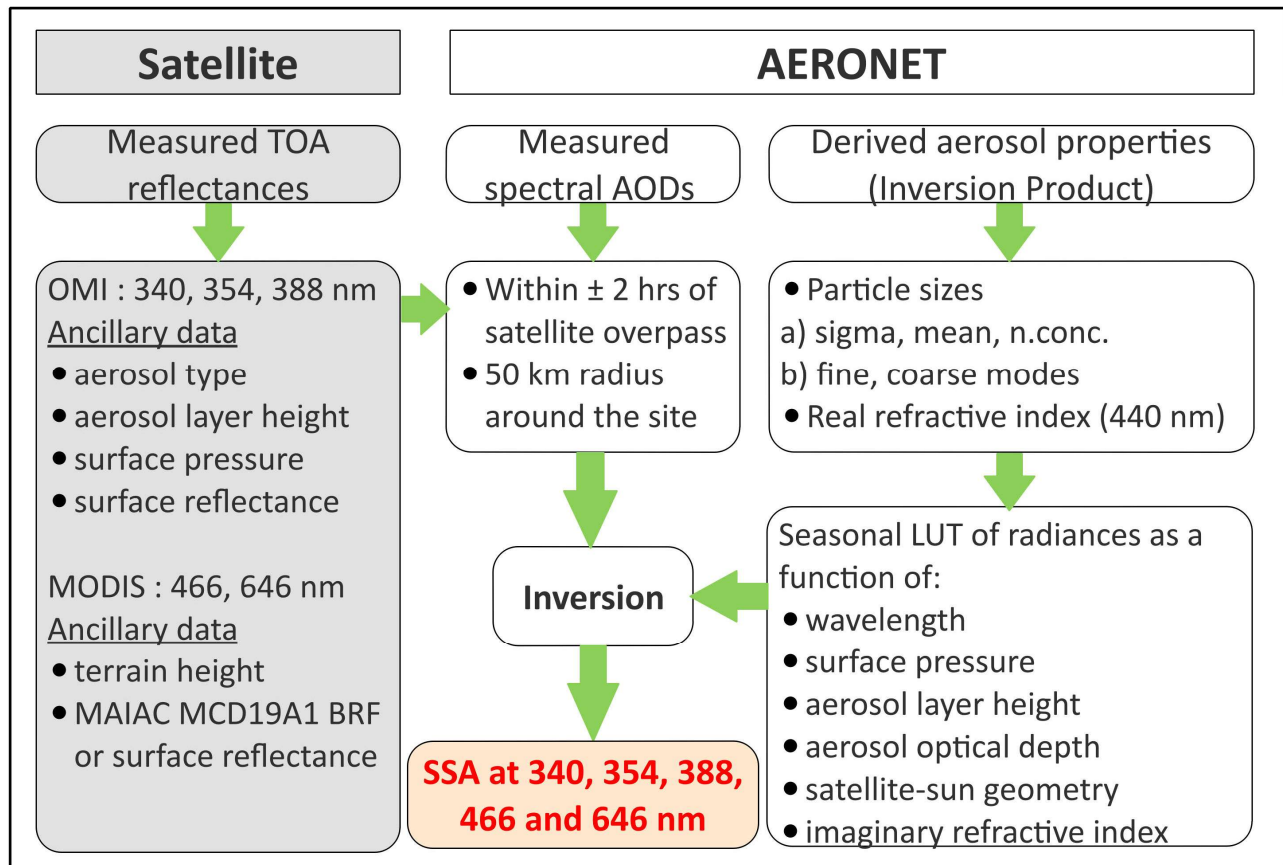
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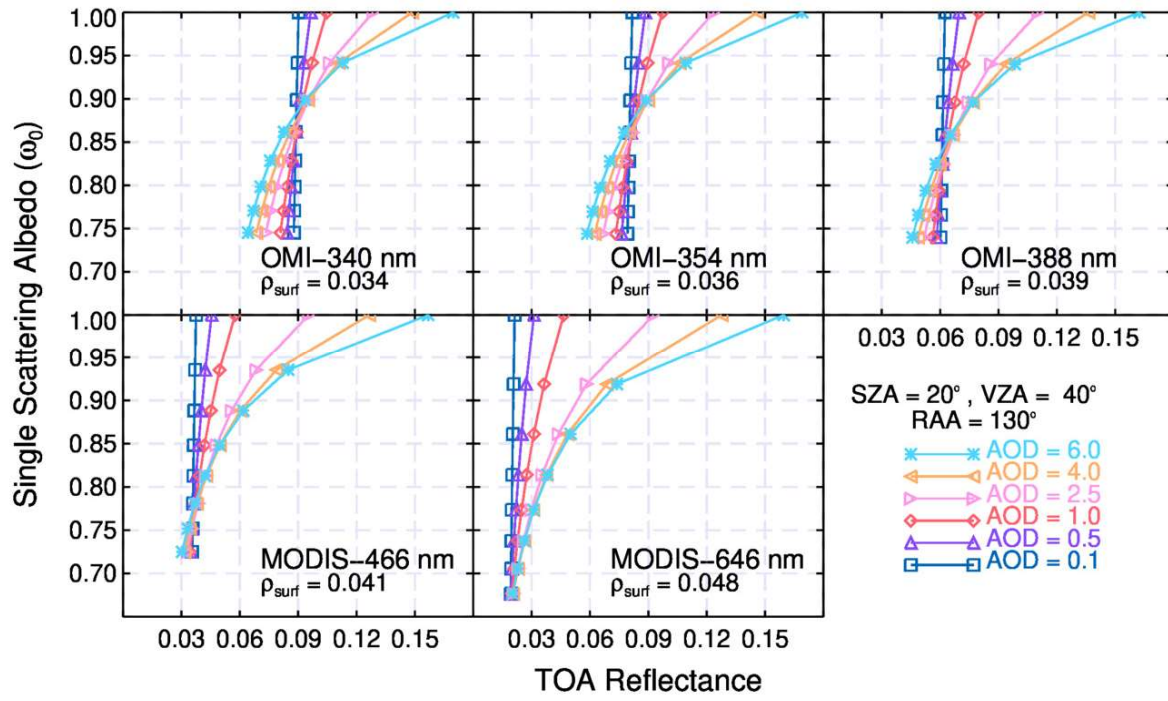
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**Figure 1: The geo-distribution of AERONET sites whose AOD data are used for the retrieval of spectral aerosol single scattering albedo in this work.**



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**Figure 2: Schematic flow chart of the methodology used to retrieve aerosol spectral single scattering albedo.**



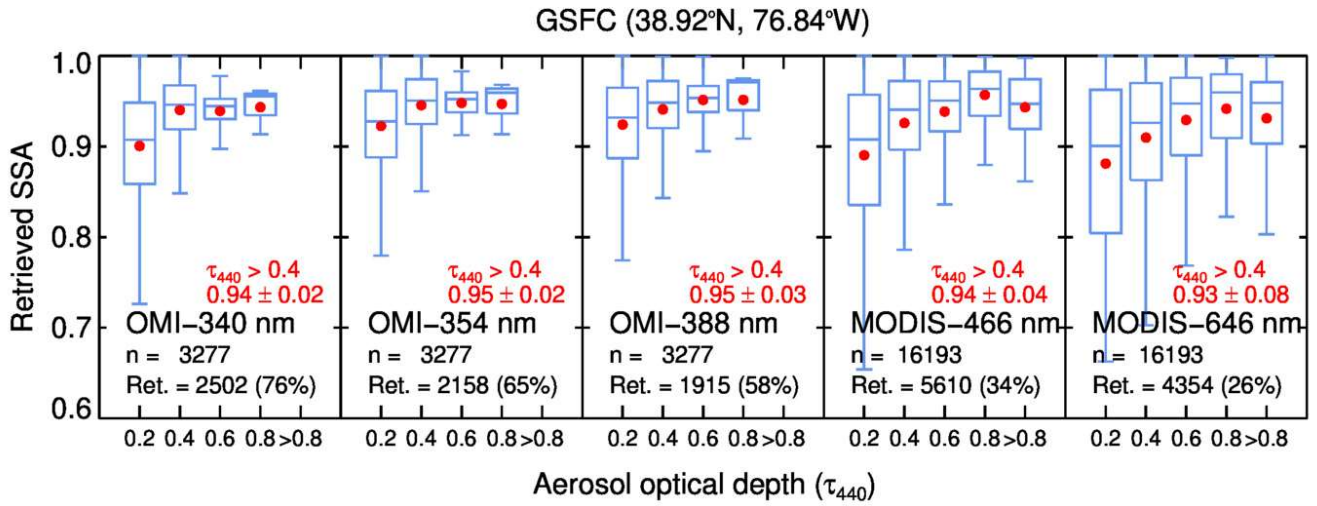
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**Figure 3: Simulated TOA radiances for the aerosols over the GSFC site as a function of  $\tau$  and SSA in the UV-Visible range.**



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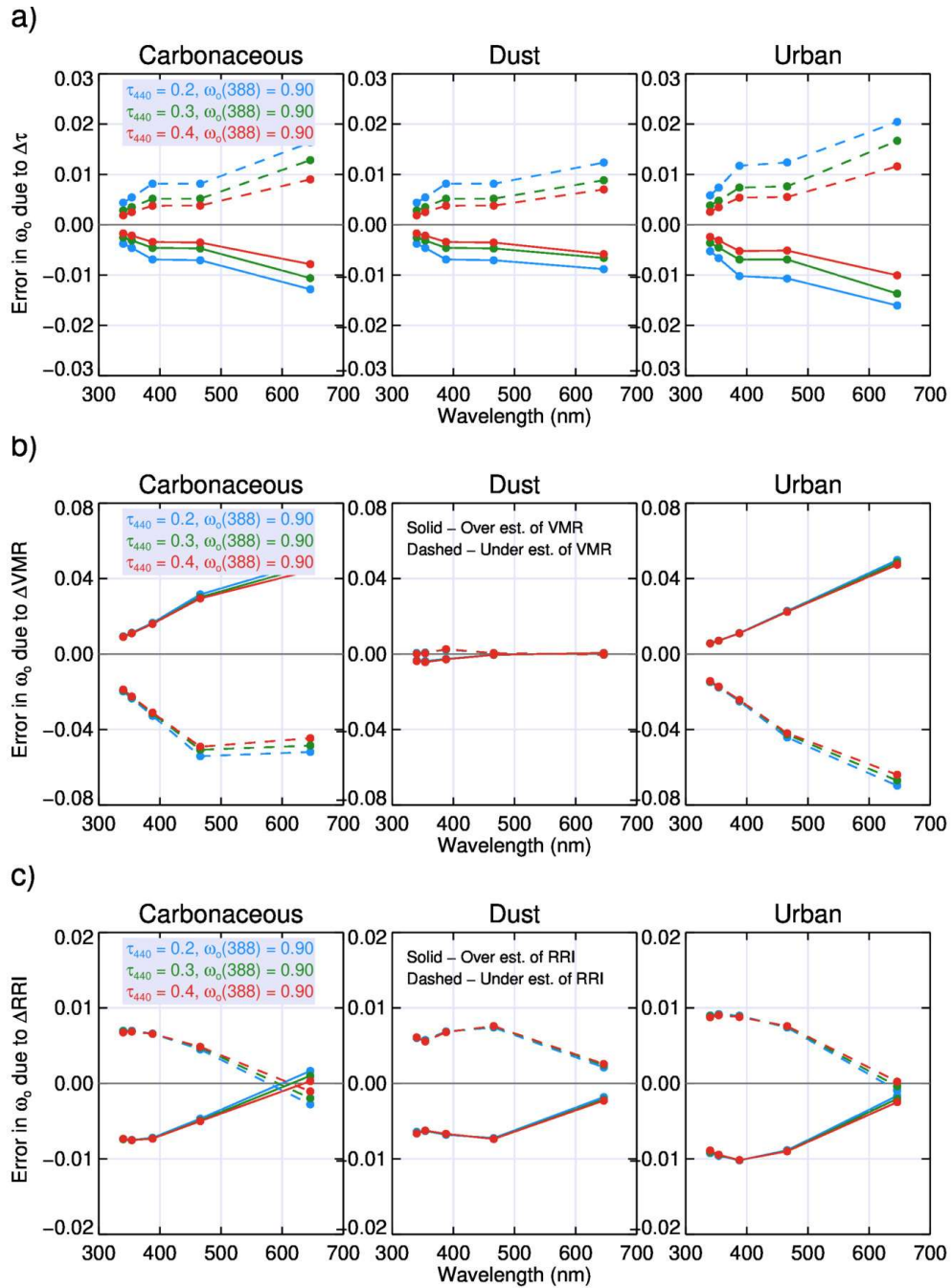
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**Figure 4: Retrieved aerosol SSA over the GSFC site for the satellite observations period of 2005 – 2016. Average and standard deviation of retrieved SSA for the observations with  $\tau_{440} > 0.4$  is shown in red. Abbreviations used: n is the number of satellite-AERONET collocated observations, and ret. is the number of observations for which SSA is retrieved.**



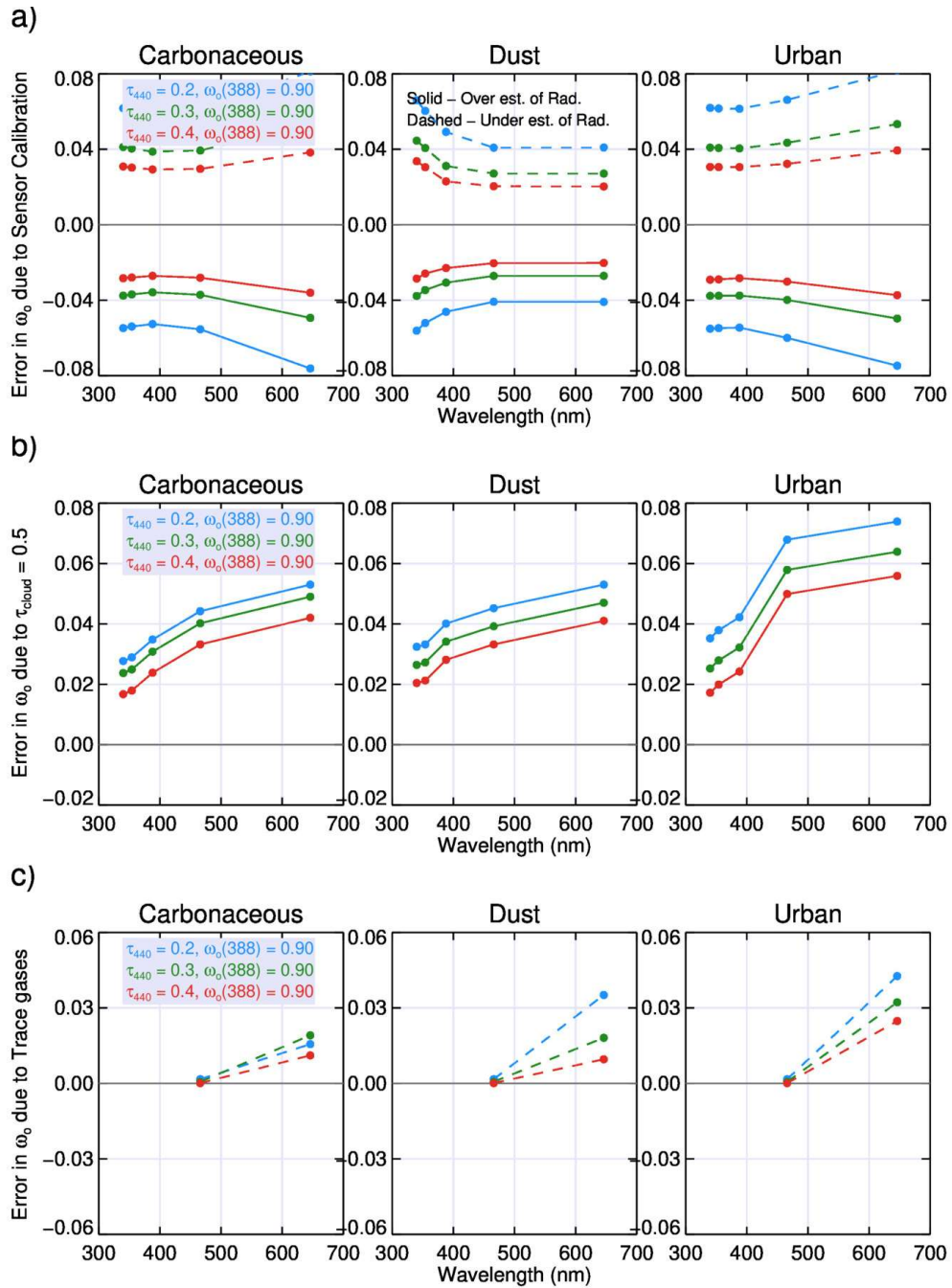
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**Figure 5: Theoretical uncertainty in SSA retrievals due to changes in (a)  $\pm 0.02$  ( $\lambda < 400$  nm) and  $\pm 0.01$  ( $\lambda > 400$  nm) AOD, (b)  $\pm 20\%$  volume mean radius, and (c)  $\pm 0.04$  real part of refractive index.**



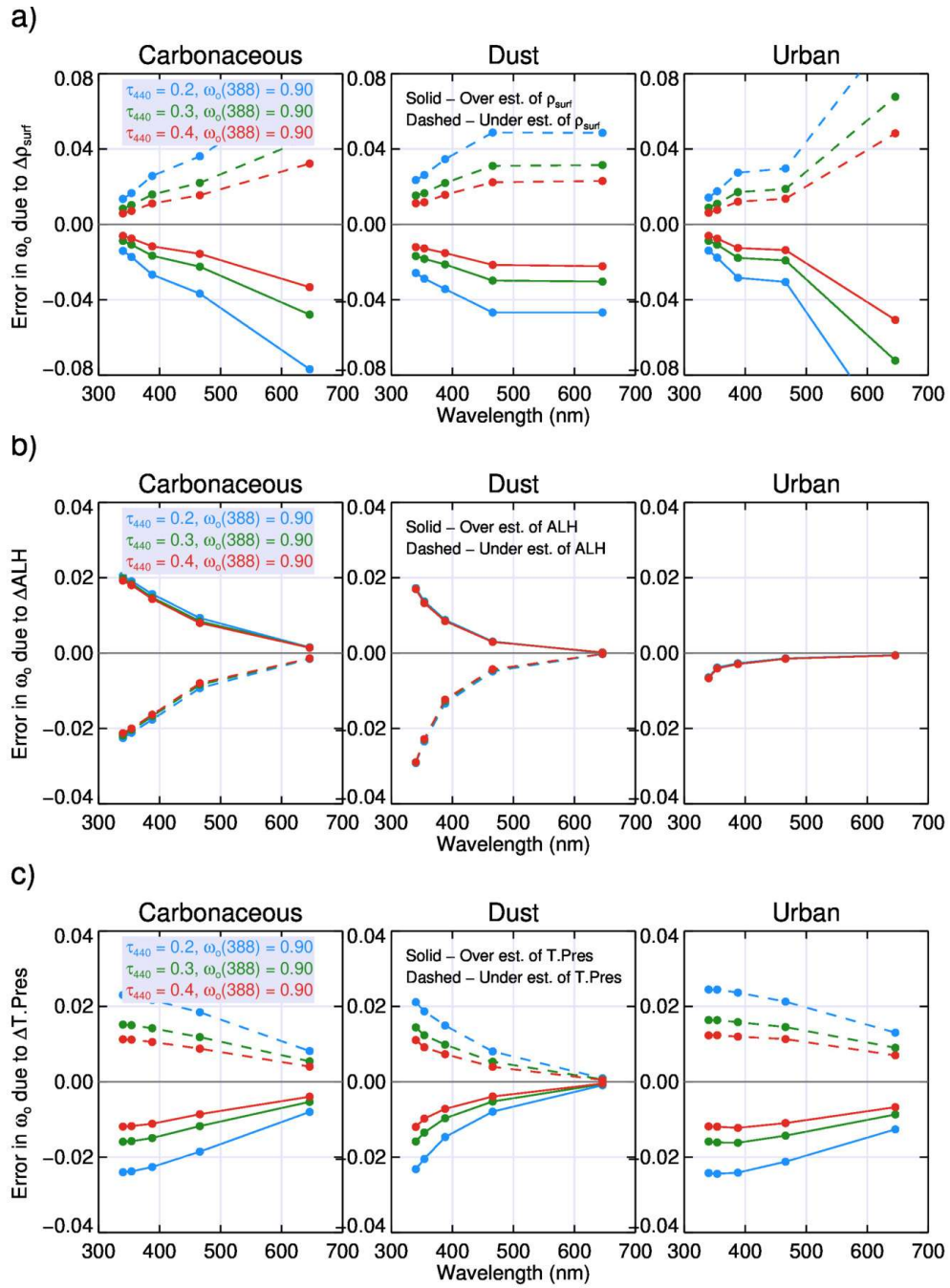
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**Figure 6: Theoretical uncertainty in SSA retrievals due to (a)  $\pm 1.8\%$  (OMI) and  $\pm 1.9\%$  (MODIS) sensor calibration, (b) cloud contamination ( $\tau_{cloud} = 0.5$ ), and (c) trace gases absorption.**



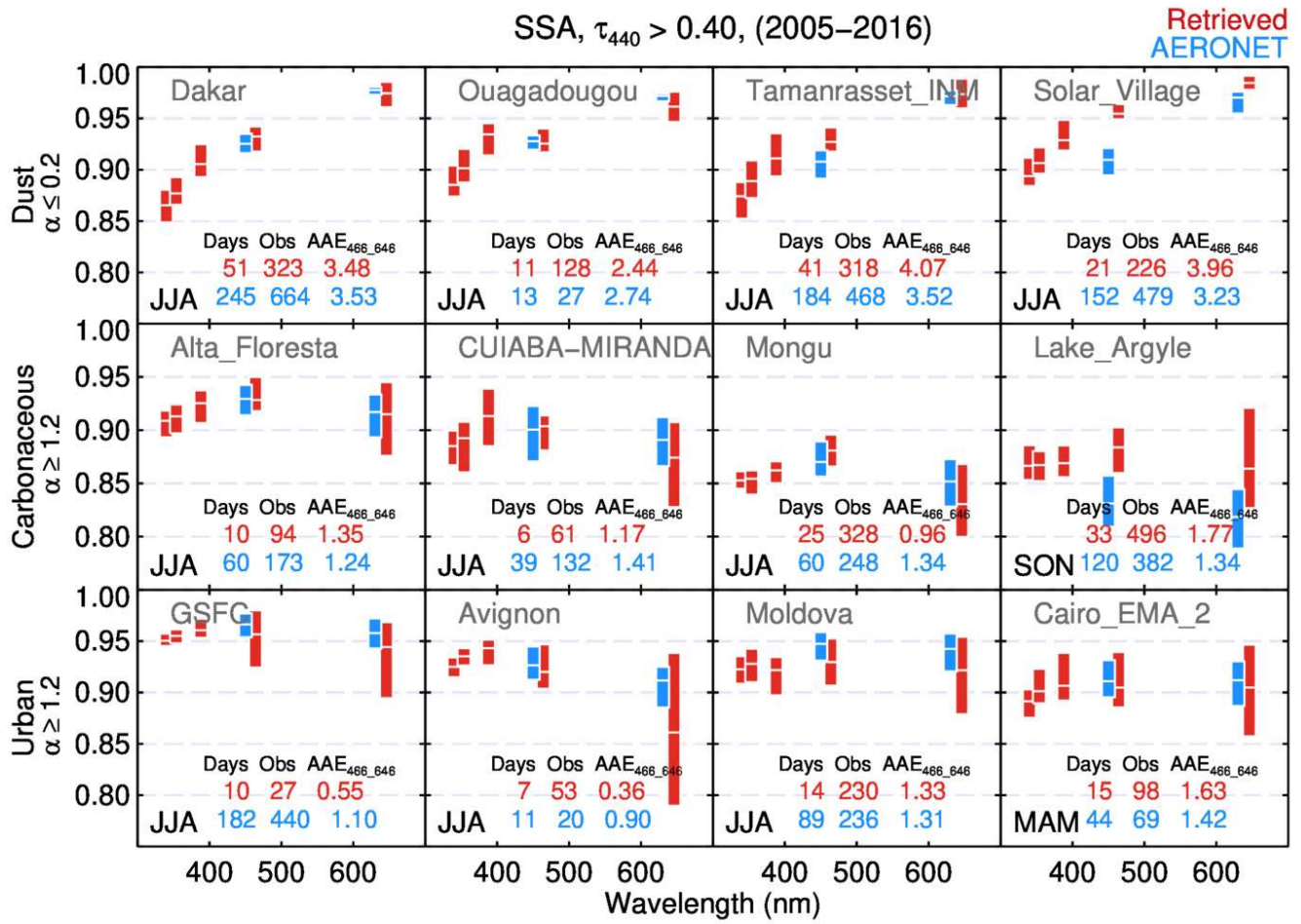
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**Figure 7: Theoretical uncertainty in SSA retrievals due to changes in (a)  $\pm 0.01$  surface reflectance, (b)  $\pm 1$  km ALH, and (c)  $\pm 12$  hPa surface pressure.**



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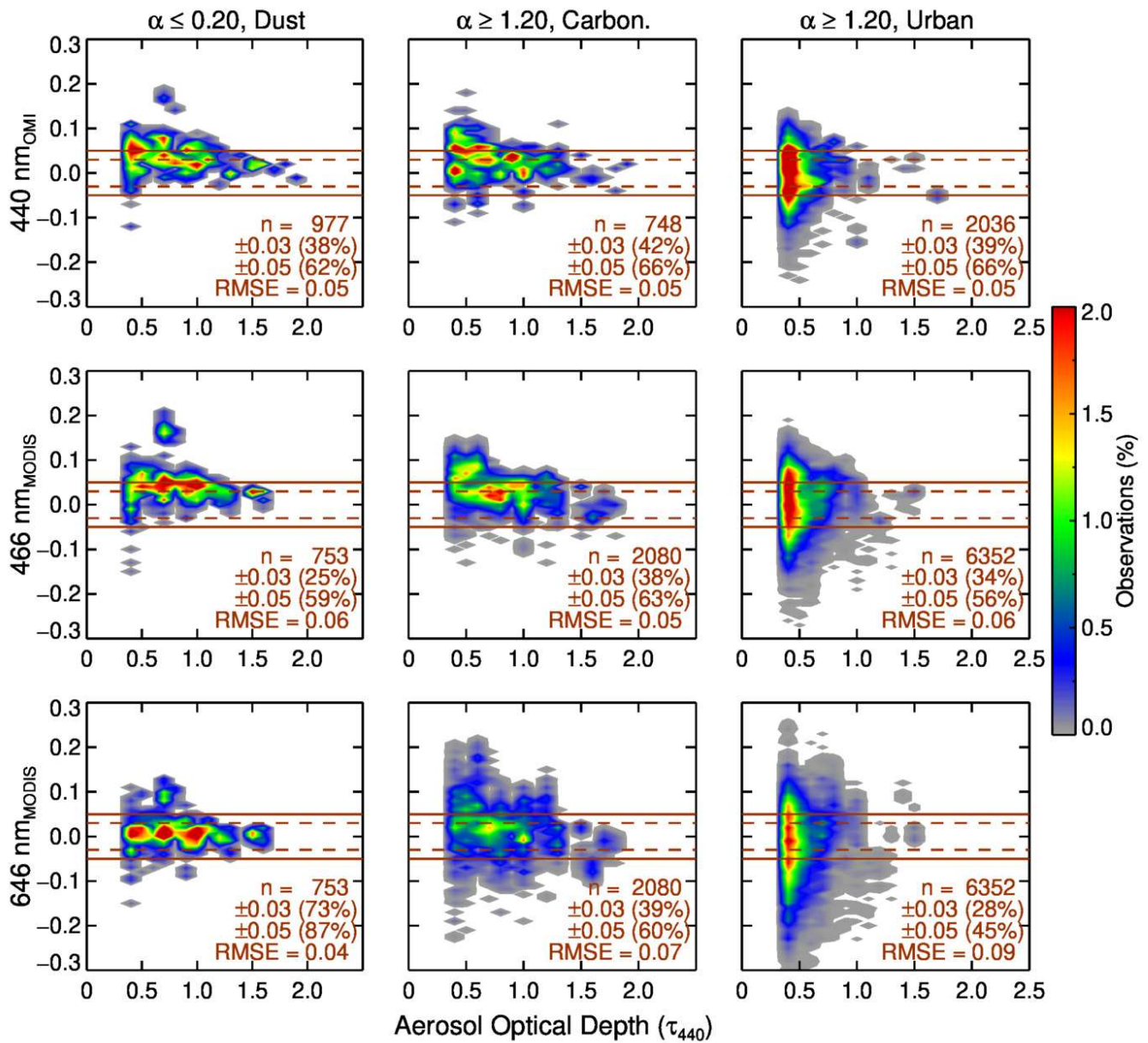
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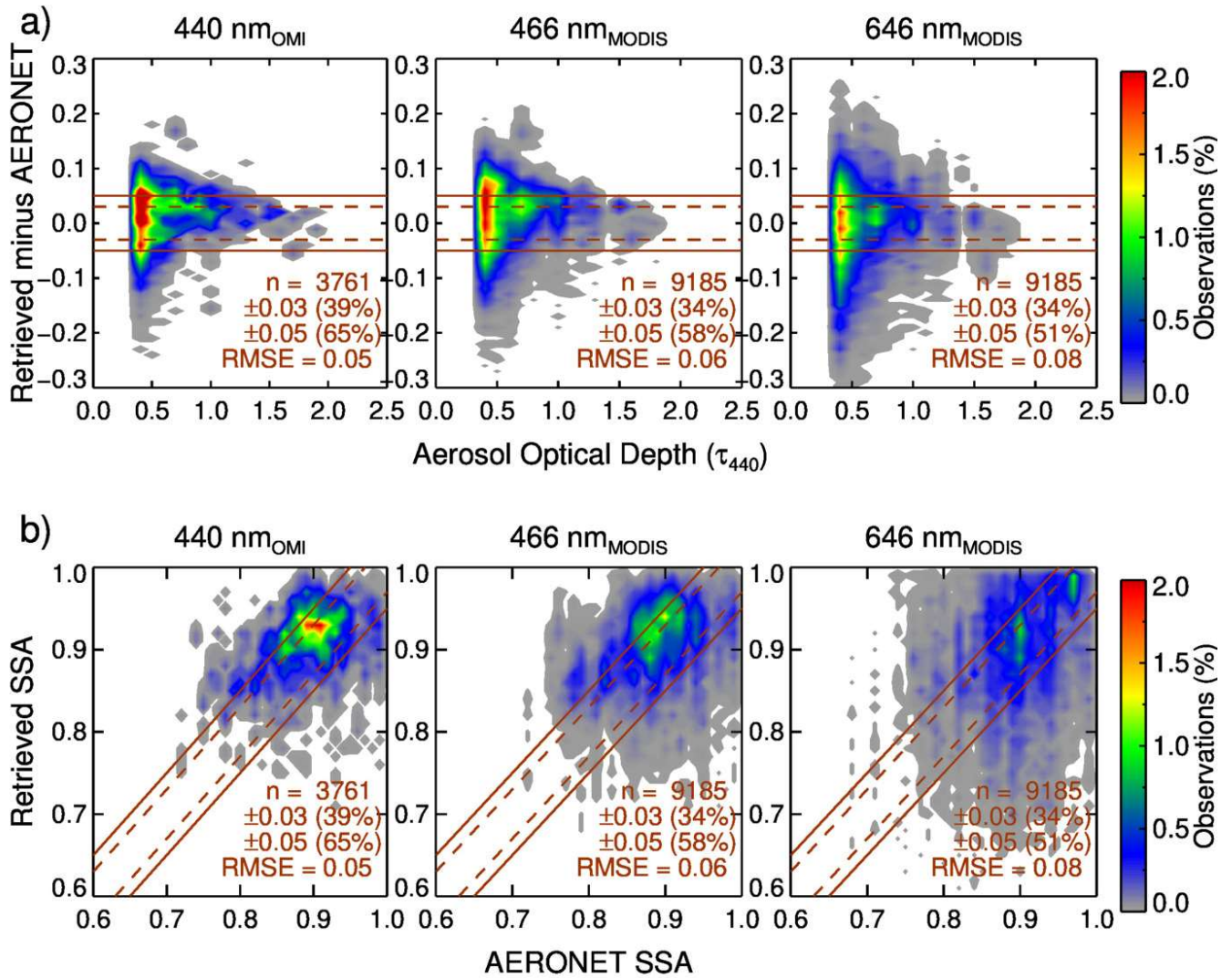
**Figure 8: Comparison of retrieved aerosol SSA with that of AERONET for selected sites. Box plots here represent lower and upper quartile of observations with mean values shown as white line.**





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1708 **Figure 9: Absolute difference in retrieved minus AERONET SSA versus AOD with observations  $\tau_{440} > 0.4$**   
 1709 **for coarse-mode dust, intermediate-mode mixtures, fine-mode carbonaceous, fine-mode urban aerosols at**  
 1710 **440 nm, 466 nm, and 646 nm.**

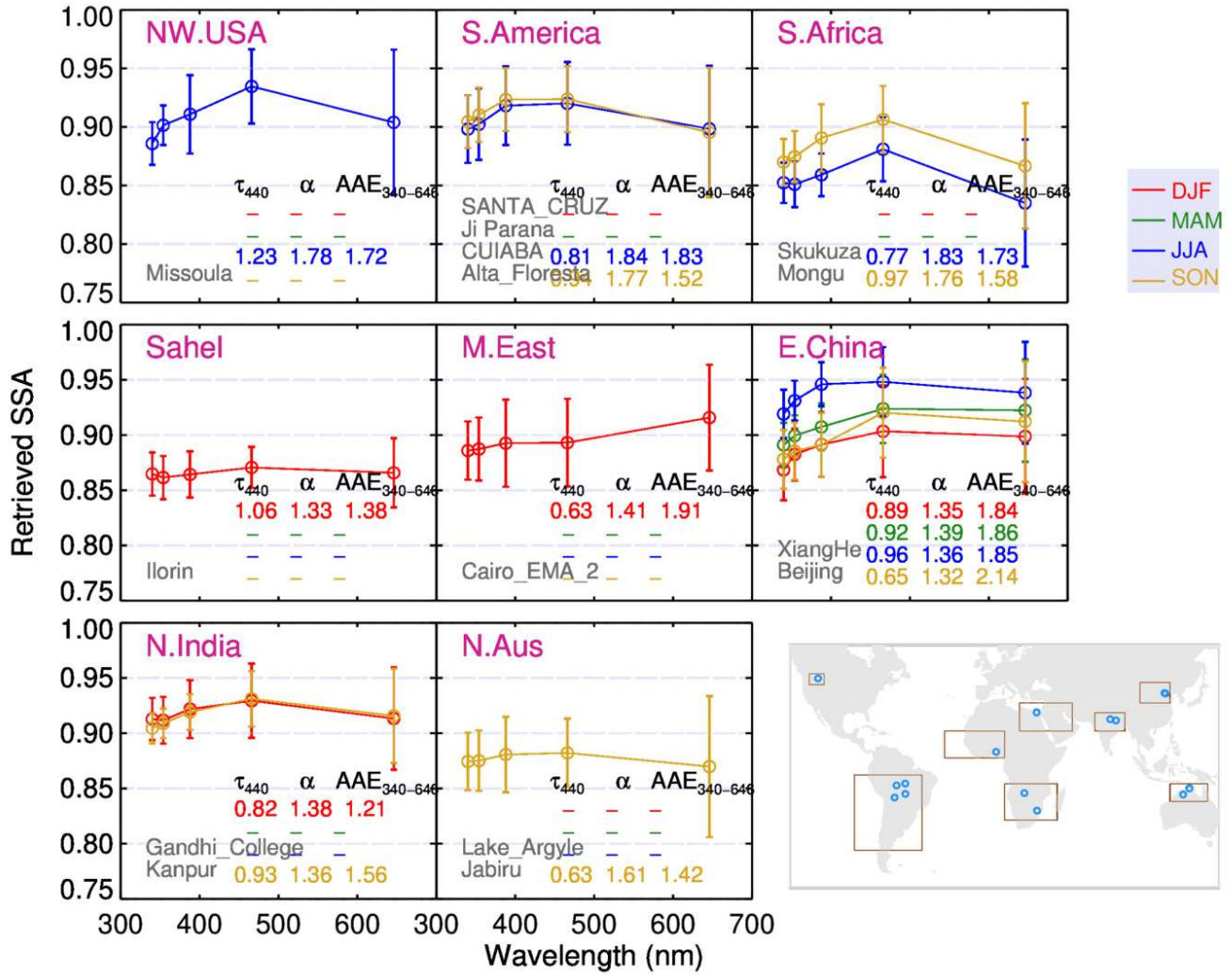


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1713 **Figure 10: (a) Absolute difference in retrieved minus AERONET SSA versus  $\tau_{440}$ , and (b) Retrieved SSA**  
 1714 **versus AERONET for observations with  $\tau_{440} > 0.4$  for the combined aerosol types at 440 nm (OMI), 466**  
 1715 **nm and 646 nm.**

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$\alpha \geq 1.20$ , Carbonaceous ( $\tau_{440} > 0.4$ )



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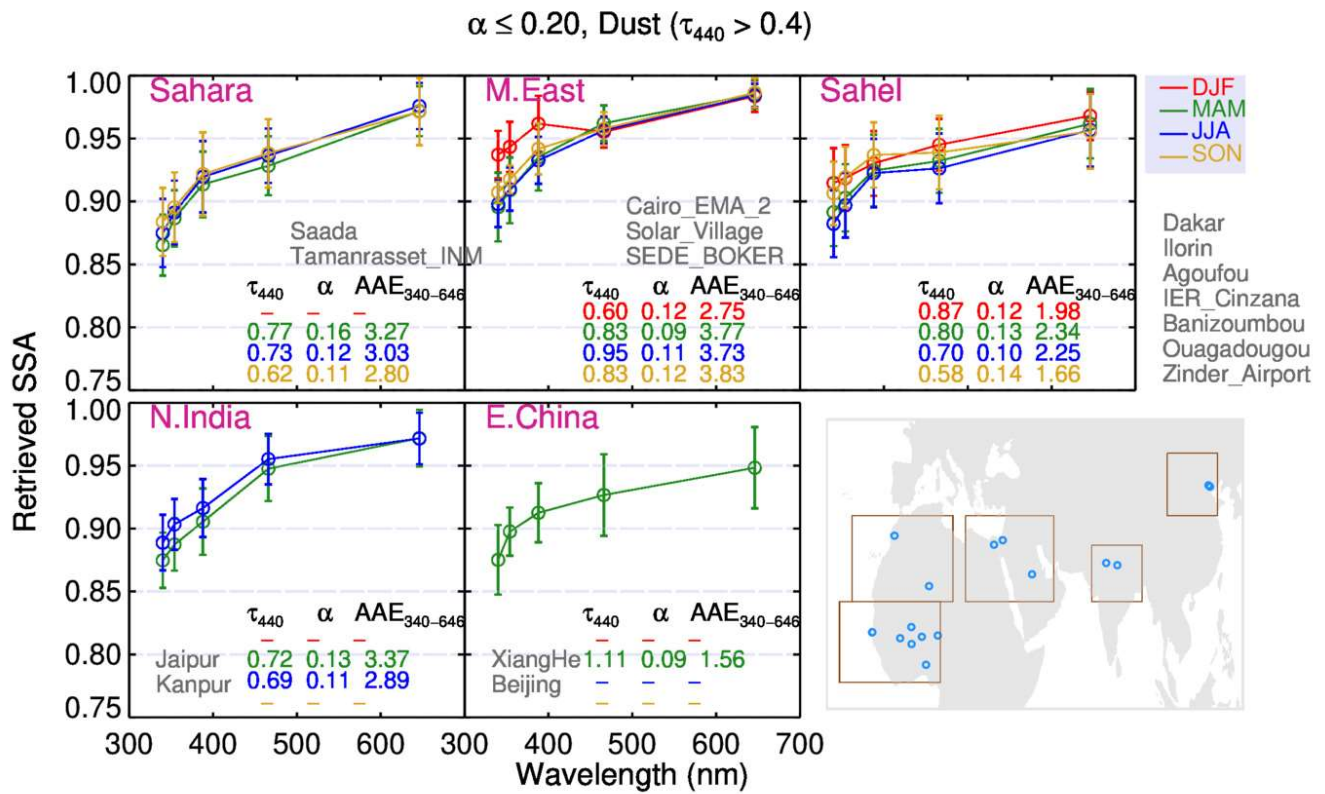
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**Figure 11: Seasonal average of spectral aerosol SSA derived for observations with  $\tau_{440} > 0.4$  over regions dominated by fine-mode carbonaceous aerosols. The error bars represent the standard deviation of the observations.**

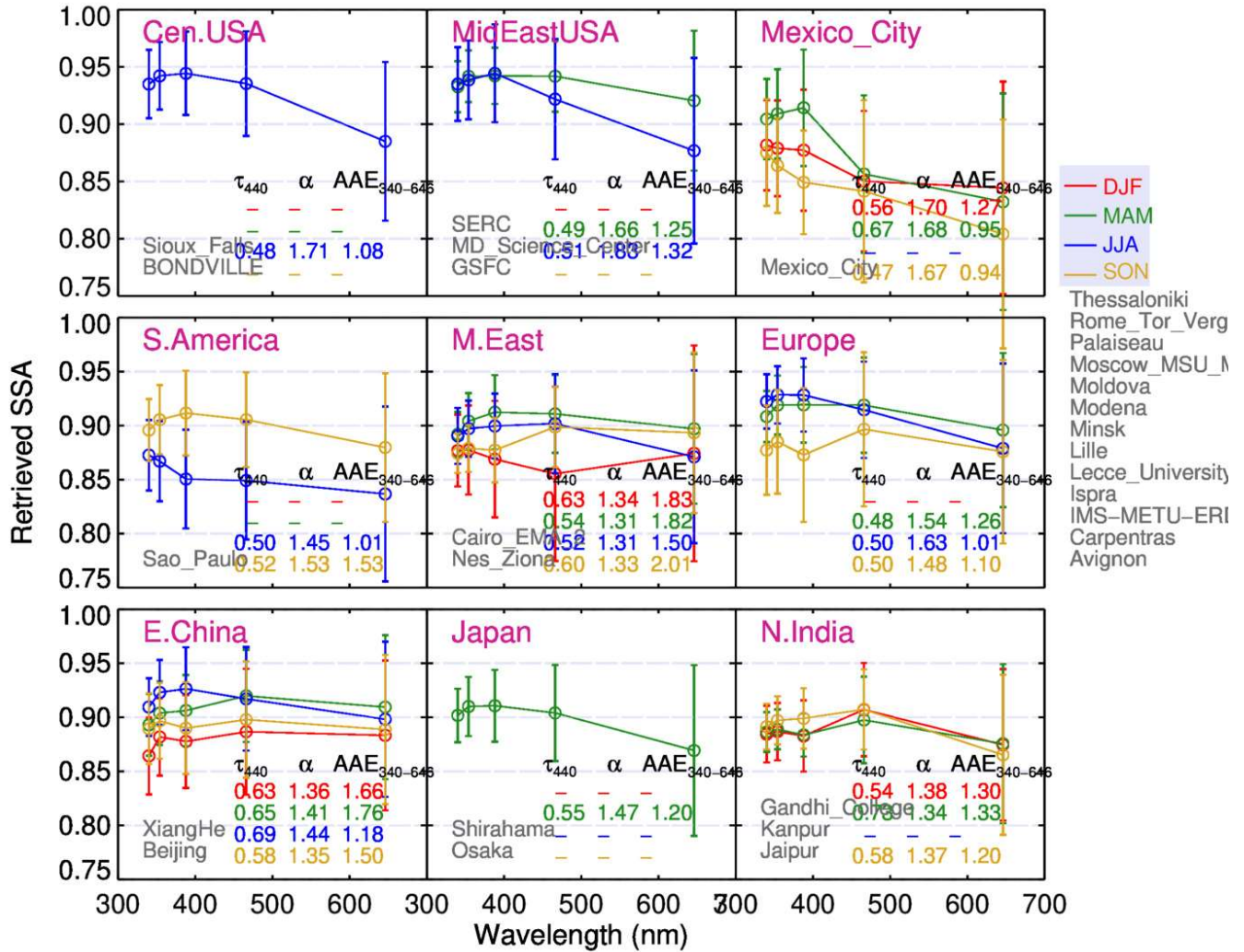


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1725 **Figure 12: Seasonal average of spectral aerosol SSA derived for observations with  $\tau_{440} > 0.4$  over regions**  
 1726 **dominated by coarse-mode dust aerosols. The error bars represent the standard deviation of the**  
 1727 **observations.**

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$\alpha \geq 1.20$ , Urban ( $\tau_{440} > 0.4$ )



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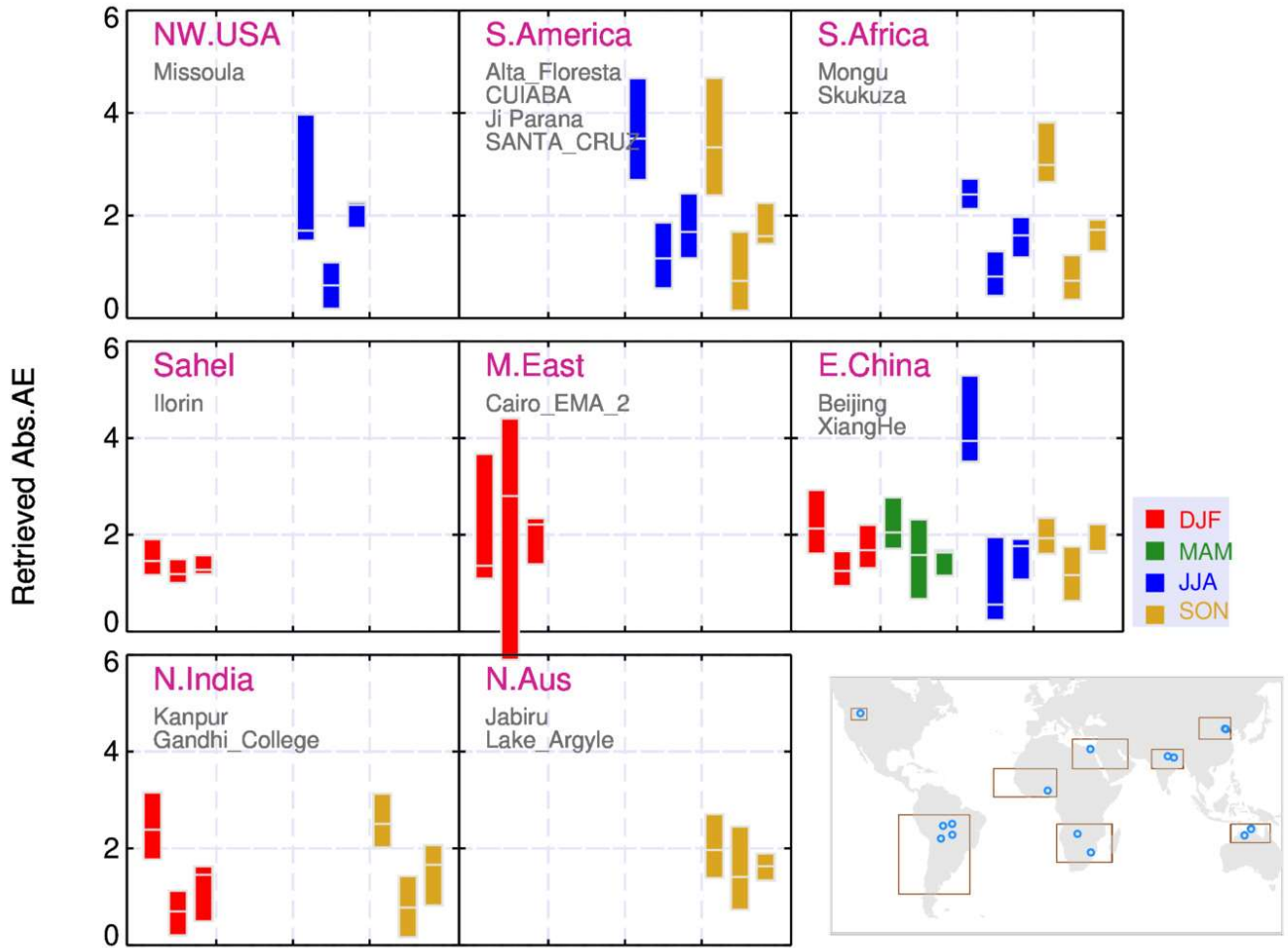
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**Figure 13: Seasonal average of spectral aerosol SSA derived for observations with  $\tau_{440} > 0.4$  over regions dominated by fine-mode urban or mixture of aerosols. The error bars represent the standard deviation of the observations.**

$\alpha \geq 1.20$ , Carbonaceous ( $\tau_{440} > 0.4$ )



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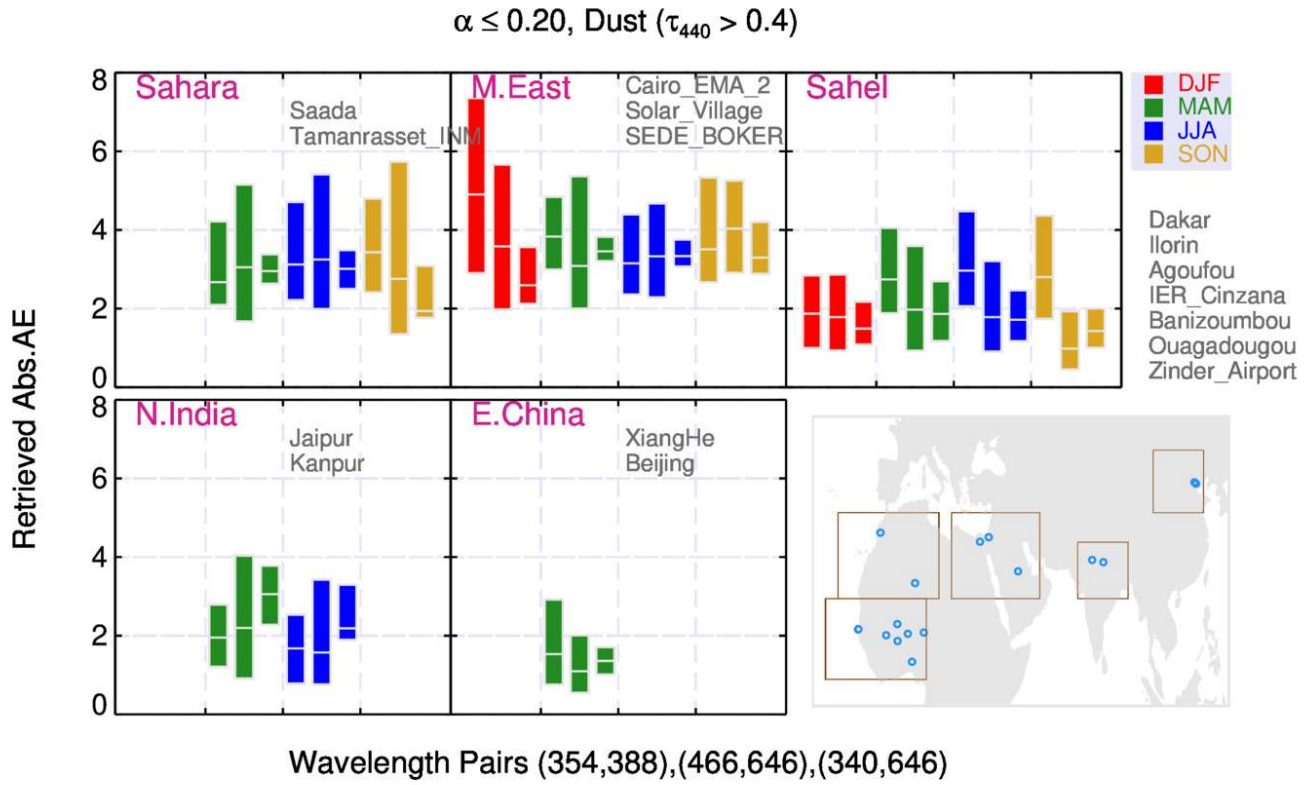
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**Figure 14: Absorption Angstrom exponent derived at three wavelength pairs for observations with  $\tau_{440} > 0.4$  for the fine-mode carbonaceous aerosols. Boxplot represents lower and upper quartile of observations with mean as white line.**

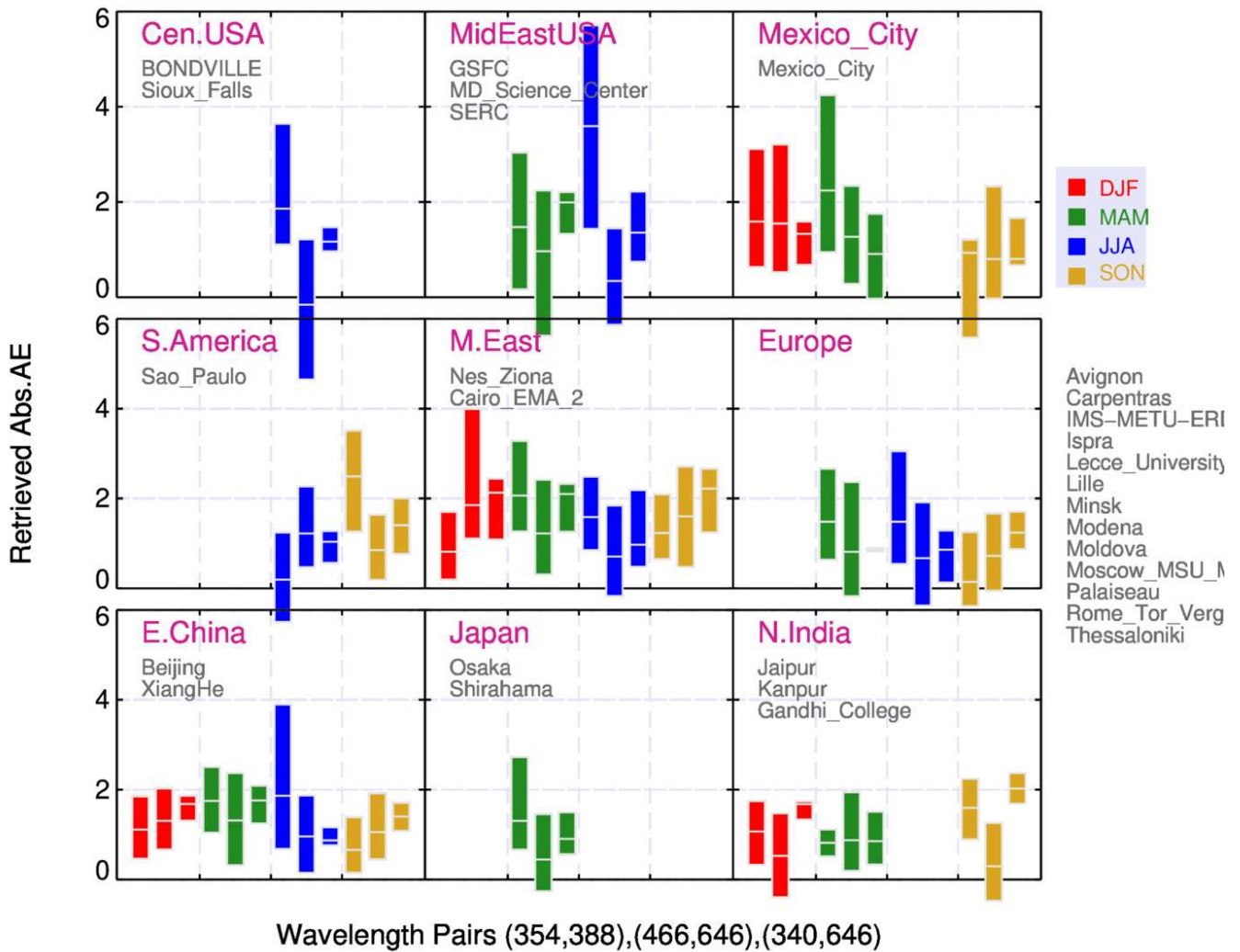


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1743 **Figure 15: Absorption Angstrom exponent derived at three wavelength pairs for observations with  $\tau_{440} >$**   
 1744 **0.4 for the coarse-mode dust aerosols. Boxplot represents lower and upper quartile of observations with**  
 1745 **mean as white line.**

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$\alpha \geq 1.20$ , Urban ( $\tau_{440} > 0.4$ )



1750 **Figure 16: Absorption Angstrom exponent derived at three wavelength pairs for observations with  $\tau_{440} > 0.4$  for the fine-mode urban or mixture of aerosols. Boxplot represents lower and upper quartile of observations with mean as white line.**