

1 Investigating the dependence of mineral dust depolarization on 2 complex refractive index and size with a laboratory polarimeter at 3 180.0° lidar backscattering angle

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7
8 **Abstract.** In this paper, the dependence of the particles depolarization ratio (*PDR*) of mineral dust on the complex refractive
9 index and size is for the first time investigated through a laboratory π -polarimeter operating at 180.0° backscattering angle and
10 at (355, 532) nm wavelengths for lidar purposes. The dust *PDR* is indeed an important input parameter in polarization lidar
11 experiments involving mineral dust. Our π -polarimeter provides sixteen accurate ($< 1\%$) values of the dust lidar *PDR* at
12 180.0° corresponding to four different complex refractive indices, studied at two size distributions (fine, coarse) ranging from
13 10 nm to more than 10 μm , and at (355, 532) nm wavelengths, while accounting for the highly irregular shape of mineral dust,
14 which is difficult to model numerically. At 355 nm, the lidar *PDR* of coarser silica, the main oxide in mineral dust, is equal to
15 $(33 \pm 1)\%$ while that of coarser hematite, the main light absorbent in mineral dust, is $(10 \pm 1)\%$. This huge difference is
16 here explained by accounting for the high imaginary part of the hematite complex refractive index. In turn, Arizona dust
17 exhibits higher depolarization than Asian dust, due to the higher proportion in hematite in the latter. As a result, when the
18 strong light absorbent hematite is involved, the dust lidar *PDR* primarily depends on the particles complex refractive index
19 and its variations with size and shape are less pronounced. When hematite is less or not involved, the dust lidar *PDR* increases
20 with increasing sizes, though the shape dependence may then also play a role. The (355, 532) nm wavelength dependence of
21 the dust lidar *PDR* then allows discussing on the involved particle sizes, thus highlighting the importance of dual-wavelength
22 (or more) polarization lidar instruments. We believe these laboratory findings will help improving our understanding of the
23 challenging dependence of the dust lidar *PDR* with complex refractive index and size to help interpret the complexity and the
24 wealth of polarization lidar signals.

25 1 Introduction

26 With worldwide annual emissions between 1000 to 3000 Tg (Monge et al., 2012), mineral dust is a highly important constituent
27 of the atmosphere, which contributes to ice cloud formation by acting as a freezing nucleus and to the carbon cycle by fertilizing
28 nutrient poor ecosystems such as the Amazon rainforest after long-range transport (Bristow et al., 2010). As underscored in
29 the latest IPCC report (2021), mineral dust also contributes to the Earth's radiative budget through light scattering and

30 absorption, by reducing the amount of energy reaching the Earth's surface (Kosmopoulos et al., 2017). The radiative impact
31 associated with a Saharan dust storm has been recently quantified by Francis et al. (2022). This climatic impact is however
32 subject to large uncertainties, mainly due to the great complexity in size, shape and mineralogy of mineral dust. In the
33 atmosphere, the size distribution of mineral dust is mainly determined by the distance from the dust source region. Two freshly
34 uplifted dust aerosols may indeed exhibit different size distributions at far-range remote sites (Ryder et al., 2013), due to the
35 rapid removal of the largest particles by gravitational settling. Mineral dust particles also exhibit a high degree of complexity
36 in shape. Electron microscopic images (Kandler et al., 2011) indeed highlight the nonspherical and highly irregular shape of
37 mineral dust particles, with sharp edges, sometimes even surface roughness (Nousiainen, 2009). The mineral dust surface is
38 itself subject to photo-catalytic reactions leading even to new particle formation events (Dupart et al., 2012). The third degree
39 of complexity of mineral dust related to this study lies in its mineralogy. Mineral dust indeed consists in a heterogeneous
40 mixture of various chemical oxides among which the most predominant is silica oxide. Aluminum and iron oxides are also
41 present in proportions depending on the dust source region. As an example, the desert in Central Australia is iron oxides rich
42 (Bullard and White, 2002). This diverse mineralogy results in a diversity of complex refractive indices for mineral dust.

43

44 In the atmosphere, mineral dust is additionally often mixed with other aerosols. To face such a complexity, ground and satellite-
45 based polarization lidar instruments, based on light backscattering by nonspherical particles, have been developed
46 (Freudenthaler et al., 2009; Tesche et al., 2009; Sugimoto and Lee, 2006; Winker et al., 2009; Miffre et al., 2019; Hofer et al.,
47 2020; Hu et al., 2020) to discern the mineral dust contribution to two-component particles external mixtures, by applying lidar
48 partitioning algorithms such as the $1\beta + 1\delta$ algorithm (Tesche et al., 2009; Mehri et al., 2018). Such lidar-based retrievals are
49 however under-constrained and depend on prior knowledge regarding input parameters such as the lidar particles'
50 depolarization ratio (*PDR*). The lidar *PDR* quantifies the mineral dust particles deviation from isotropy and is key for aerosol
51 typing (Hofer et al., 2020; Burton et al., 2012). As explained in light scattering textbooks (Bohren and Huffman, 1983;
52 Mishchenko et al., 2002), it depends on the particles size, shape and complex refractive index. The size dependence of the
53 lidar *PDR* was studied in field by Hofer et al. (2020). The downside of such field measurements is that the observed aerosol is
54 nevertheless that of a particles mixture, which may induce some discrepancies in the retrieved dust lidar *PDR* (Miffre et al.,
55 2011). As an alternate, for accurate retrievals of the mineral dust lidar *PDR*, light backscattering numerical simulations have
56 been developed, by assuming a particles shape model such as the spheroidal shape model, computed with the T-matrix
57 numerical code (Mishchenko and Travis, 1998), as successfully applied for mineral dust during the SAMUM field campaign
58 (Müller et al., 2013) or, by considering more realistic shapes, based on stereograms, computed with the discrete-dipole-
59 approximation (Lindqvist et al., 2014; Gasteiger et al., 2011). Depending on the assumed shape model, the lidar *PDR* can be
60 very different with induced variations in the lidar-retrieved dust mass concentrations (Mehri et al., 2018). Recently, Luo et al.
61 (2022) and Huang et al. (2022) discussed on the ability of the spheroidal model to mimic the complex shape of mineral dust.
62 Likewise, Zubko et al. (2013) found spheroids inadequate for describing the dust particles' spectral dependence of the lidar
63 *PDR*. Such light scattering numerical simulations nonetheless rely on simplifying assumptions that should be carefully

64 checked. Laboratory experiments on natural dust samples at 180.0° lidar exact backscattering angle are then looked-for as they
65 provide quantitative evaluations of the mineral dust lidar *PDR* within experimental error bars. Indeed, in laboratory, the
66 retrieved lidar *PDR* is, by construction, that of pure mineral dust and the dependence of the dust lidar *PDR* with size and
67 mineralogy can be evaluated. Moreover, the complex shape of mineral dust is then accounted for. However, existing laboratory
68 light scattering experimental set-ups (Glen and Brooks, 2013; Järvinen et al., 2016; Gautam et al., 2020; Liu et al., 2020;
69 Kahnert et al., 2020; Gómez Martín et al., 2021) can only provide approximate values of the dust lidar *PDR* for the following
70 reasons:

71

72 – Such apparatuses operate at near backscattering angles only ($< 180.0^\circ$), without covering the exact lidar
73 backscattering angle of 180.0° . The retrieved lidar *PDR* is then extrapolated to 180.0° following simplifying
74 numerical assumptions, ignoring the complexity in shape of mineral dust (Liu et al., 2020; Gómez Martín et al., 2021).
75 To provide accurate values of the dust lidar *PDR*, such assumptions must be carefully discussed as the lidar *PDR*
76 actually depends on the scattering angle in an unpredictable way, as underscored in light scattering textbooks (Bohren
77 and Huffman, 1983; Mishchenko et al., 2002), due to the complex shape of mineral dust. For that, a laboratory
78 measurement of the dust lidar *PDR* at 180.0° is mandatory.

79 – Also, most of the above apparatuses operate at a single wavelength, either 442, 488, 552, 632, 647 or 680 nm, which
80 differs from the (355, 532, 1064 nm) wavelengths which are applied in polarization lidar field experiments. As for
81 Raman lidars, such wavelength extrapolations up to the (355, 532, 1064 nm) lidar wavelengths are a source of
82 discrepancy as the dust lidar *PDR* actually depends on the complex refractive index, which is wavelength dependent
83 (Bohren and Huffman, 1983; Mishchenko et al., 2002). For that, a laboratory measurement at the lidar wavelengths
84 is mandatory.

85

86 In this paper, accurate values ($< 1\%$) of the dust lidar *PDR* are provided from a laboratory π -polarimeter operating at 180.0°
87 lidar exact backscattering angle and at 355, 532 nm wavelength, to account for the importance of the spectral dependence of
88 the lidar *PDR* to better constrain lidar inversions and aerosol typing (Burton et al., 2016; Haarig et al., 2022). Since the
89 scattering angle and the wavelengths are determined for lidar purposes, we here investigate the dependence of the mineral dust
90 lidar *PDR* on the dust particles size and complex refractive index (*CRI*), the latter being particularly important as related to
91 light absorption. Light absorption by mineral dust preferentially occurs in the UV and VIS spectral domains, being nearly null
92 in the near-infrared spectral range (Di Biagio et al., 2019), noticeably in the presence of iron oxides (Formenti et al., 2014;
93 Caponi et al., 2017). By absorbing short-wave radiations, such oxides hence play a critical role in determining the overall
94 impact of dust aerosol on climate forcing (Go et al., 2022). We hence focused on 355 and 532 nm lidar wavelengths and
95 considered four dust samples differing in their *CRI*, thus in mineralogy: i) silica oxide (SiO_2), as the most abundant mineral
96 oxide present in mineral dust, ii) iron oxide (hematite, Fe_2O_3), as the main light absorbent present in mineral dust (Gautam et

97 al., 2020; Zong et al., 2021; Go et al., 2022), iii) and iv) two heterogeneous mixtures of the above two oxides in various
98 proportions, as detailed in Section 2. The dependence of the lidar *PDR* with size is then likewise investigating by accounting
99 for the fine and coarse modes of the particles size distribution (*SD*), to which lidar instruments are sensitive (Mamouri and
100 Ansmann, 2017), thus extending the size range of our previous laboratory findings (Miffre et al., 2016) to particles sizes larger
101 than 800 nm and to other mineralogy, as asked for in Tesche et al. (2019). According to the manufacturer, the size distribution
102 of our dust samples ranged from 10 nm to more than 10 μm in diameter. Our work provides sixteen laboratory-derived accurate
103 dust lidar *PDR* values, corresponding to four mineral dust samples differing in mineralogy, given at two *SD* (fine, coarse) and
104 at two wavelengths (355, 532 nm). Moreover, the role of the imaginary part of the hematite *CRI*, which may lead to
105 modifications in the lidar *PDR*, is here for the first time quantified and discussed.

106

107 The paper is structured as follows. In Section 2, the complex refractive indices and size distributions of our four dust samples
108 are presented. The laboratory π -polarimeter at 180.0° lidar backscattering angle is then presented in Section 3, together with
109 the dust lidar *PDR* retrieval methodology, derived from the scattering matrix formalism (Mishchenko et al., 2002). The main
110 findings are outlined in Section 4 where the sixteen values of dust lidar *PDR* are given and a discussion is proposed to
111 investigate the dependence of the dust lidar *PDR* on the imaginary part of the dust *CRI*. As in elastic lidar applications, we
112 here consider the elastic backscattering of electromagnetic radiation of wavelength λ by an ensemble of mineral dust particles
113 of complex refractive index $m = n + i\kappa$ embedded in ambient air.

114 **2. Mineral dust samples**

115 **2.1 Refractive indices**

116 Mineral dust is a complex mixture of several chemical oxides presenting various complex refractive indices. To investigate
117 the dependence of the dust lidar *PDR* on the complex refractive index (*CRI*), we consider the four following case studies :

118

- 119 – Silica, or silicon oxide (SiO_2) is here considered as being the main pure chemical component present in mineral dust.
120 The silica *CRI* as given by Longtin et al. (1988) is equal to 1.546, hence exhibiting no absorptive component.
- 121 – Iron oxide, or hematite (Fe_2O_3), is in contrast here selected as being a climatically significant light absorbent in the
122 shortwave spectral region, that can be transported far from source regions with similar efficiency as black carbon
123 particles (Lamb et al., 2021). It recently regained in interest with papers specifically dedicated to this constituent
124 (Gautam et al., 2020; Zong et al., 2021). Hematite is unique among all chemical oxides present in mineral dust due
125 its strong *CRI*. Both n and κ are large for hematite, with κ -values more than 100 times those of other soil mineral
126 components at lidar wavelengths. Hence, hematite dominates absorption while other minerals can be considered as

127 non-absorbing (Go et al., 2022). The real and imaginary part of the hematite CRI is provided by Scanza et al. (2015):
128 $m = 2.13 + 0.94i$ at 355 nm wavelength ($3.07 + 0.55i$ at 532 nm wavelength)..

- 129 – Arizona Test Dust (hereafter called Arizona dust) is likewise considered as an example of natural mineral dust sample
130 that is a mixture of the above two oxides. According to the manufacturer (Power Technology Inc.), Arizona Test Dust
131 is composed of silica (68-76 %), while hematite is only weakly present in Arizona dust (2-5 %). In short, Arizona
132 dust is hence rather silica-rich. As given by the manufacturer, the Arizona dust CRI is $m = 1.51 + 10^{-3}i$, without
133 however any given on its spectral dependency. Effective medium theories can alternately be applied to account for
134 the sample inhomogeneity as calculated in Miffre et al. (2016), who arrived to $m = 1.57 + 10^{-2}i$ at 355 nm
135 wavelength and $1.55 + 5.10^{-3}i$ at 532 nm wavelength. As a result, the Arizona dust sample CRI is characterized by
136 $n \sim 1.5$ and a low absorbing component $\kappa \sim 5.10^{-3}$.
- 137 – Asian dust is finally also considered as an important case study of natural mineral dust sample, presenting however a
138 lower proportion of silica (34-40 %) and a higher proportion in hematite (17-23 %). For Asian dust, we use a
139 commercial sample provided by Powder Technology (commercial name: Kanto Loam), commonly used as a dust
140 interferon in pollen light scattering measurements in Japan (Iwai, 2013), hence representative of observed atmospheric
141 Asian dust. In this way, we symmetrized our approach by dealing with both Arizona Test Dust and Asian Test Dust.
142 The CRI of Asian dust, evaluated from effective medium approximation, is $m = 1.70 + 0.09i$ at 355 nm wavelength
143 and $1.72 + 0.03i$ at 532 nm wavelength. Hence, compared with Arizona dust, Asian dust is more hematite-rich and
144 hence exhibits a larger imaginary part for its CRI.

145
146 Other chemical oxides are also present in our dust samples in various percentages, but with negligible imaginary parts of CRI
147 compared with that of hematite. Investigating the PDR of these oxides is then beyond the scope of this paper. Their percentage
148 in (Arizona Test Dust, Asian Dust) is given for clarity: Al_2O_3 (11 %, 29 %), CaO (4 %, 1.5 %), K_2O (3.5 %, 0 %), Na_2O (2 %, 0 %),
149 MgO (1.5 %, 5 %), TiO_2 (0.5 %, 2 %). The solid dust samples, provided by Sigma Aldrich and Powder Technology
150 manufacturers, were embedded in laboratory ambient air by using a solid dust generator supplied with dried compressed air
151 ($RH < 10\%$) to get dry solid dust particles embedded in laboratory ambient air at a constant number concentration, before
152 injecting the dust samples into the light scattering volume, as presented in Section 3.

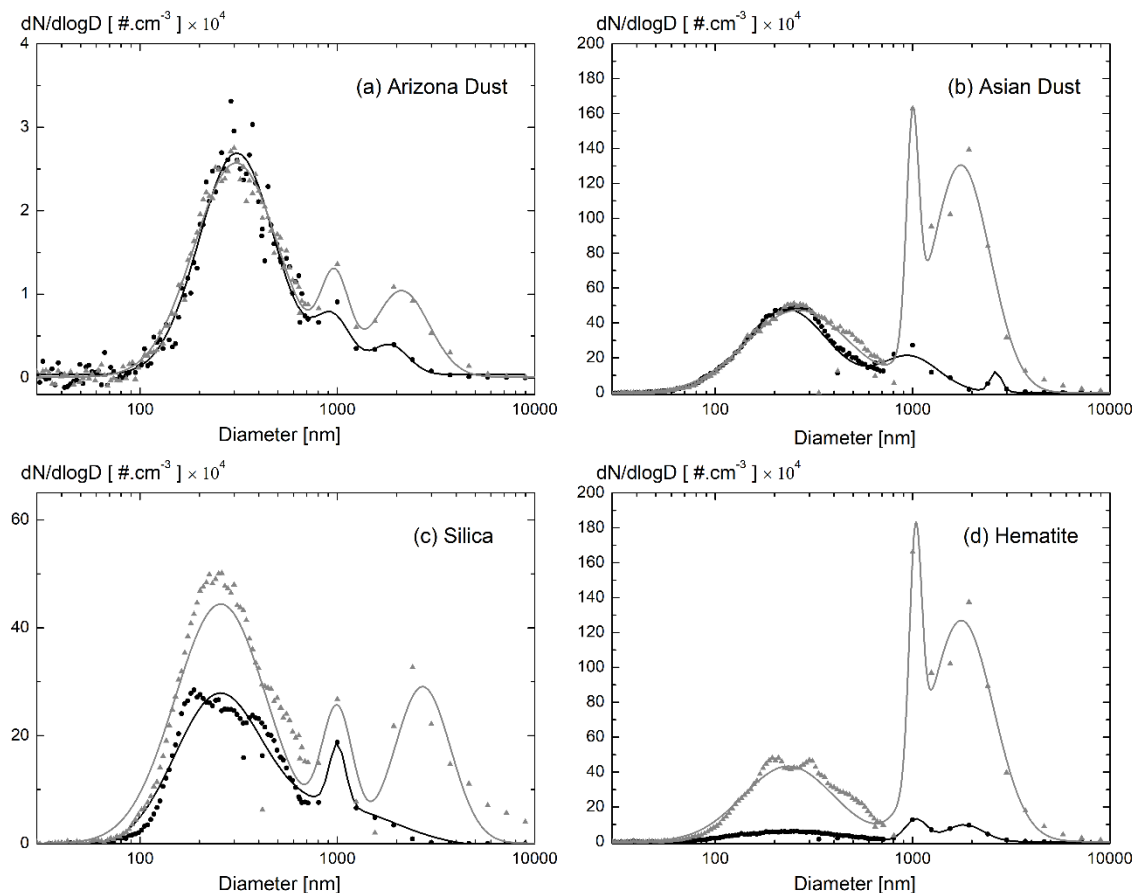
153 **2.2 Size distribution (SD)**

154 For each above dust sample, we consider two size distributions (SD) to likewise investigate the dependence of the dust lidar
155 PDR on the particles size:

- 156 – The coarser SD, represented in grey in Fig. 1. This SD is more representative of mineral dust particles close to dust
157 regions, although it does not cover the full range of large dust particles measured close to dust sources, showing
158 particles with diameters $> 50 \mu\text{m}$ (Ryder et al., 2019),

159 – A finer *SD*, plotted with a black line in Fig. 1, aimed at being more representative of mineral dust particles after long-
160 range transport, i.e. farther from the dust source regions.

161 The *SD* were obtained by adding / removing a cyclone to our experimental set-up allowing to add / remove particles with
162 diameter above 800 nm, thus exploring particles size ranges below and above 800 nm, as asked for in Tesche et al. (2019).
163 More precisely, the two considered *SD* correspond to a size distribution with and without coarse mode. The *SD* were measured
164 with an optical particles sizer (OPS 3330) coupled with a scanning mobility particles sizer (SMPS 3081), which selects the
165 dust particles as a function of their electric mobility, this latter quantity being diameter-dependent. As in Järvinen et al. (2016),
166 our size instruments could not measure dust particles with diameter above 10 μm . According to the manufacturer, such giant
167 particles (Ryder et al., 2019) are however present in our dust samples, at a low number concentration. The measured *SD* are
168 representative of what is observed in atmosphere, with a low number concentration of more than 10 μm particles, as observed
169 by Weinzierl et al. (2017). The particles *SD* displayed in Fig. 1 are in agreement with the specifications provided by the
170 manufacturers.



171

172 Figure 1: Dust particles size distributions (*SD*) for: (a) Arizona dust, (b) Asian dust, (c) Silica (SiO_2), (d) Hematite (Fe_2O_3) in
173 the presence / absence of the added cyclone (finer *SD*, in solid black) / (coarser *SD*, in dotted grey). The retrieved *SD*, obtained
174 by log-normal adjustments, agree with the specifications provided by the manufacturers.

175 3 Methodology

176 In this section, we detail our methodology for accurate laboratory evaluations of the dust lidar *PDR* at lidar exact
 177 backscattering angle of 180.0° for accurate lidar *PDR*-retrievals.

178 3.1 Scattering matrix formalism

179 The dust lidar *PDR* can be evaluated in the framework of the scattering matrix formalism (Mishchenko et al., 2002; Bohren
 180 and Huffman, 1983). In this formalism, the polarization state of the incident and scattered radiations are described by their
 181 respective Stokes vectors $\mathbf{St}_i = [I_i, Q_i, U_i, V_i]^T$ and $\mathbf{St} = [I, Q, U, V]^T$, defined with respect to the scattering plane, used as a
 182 reference plane (Mishchenko et al., 2002). The first Stokes component I corresponds to the light intensity, Q and U describe
 183 linear polarization, while V accounts for circular polarization. At a distance d from the mineral dust samples, if single-
 184 scattering and particles random orientation are assumed, for macroscopically isotropic and mirror-symmetric mediums, the
 185 incident and scattered Stokes vectors relate with a bloc-diagonal scattering matrix (Mishchenko et al., 2002; Bohren and
 186 Huffman, 1983):

$$187 \begin{pmatrix} I \\ Q \\ U \\ V \end{pmatrix} = \frac{1}{k^2 d^2} \begin{bmatrix} F_{11,\lambda}(\theta) & F_{12,\lambda}(\theta) & 0 & 0 \\ F_{12,\lambda}(\theta) & F_{22,\lambda}(\theta) & 0 & 0 \\ 0 & 0 & F_{33,\lambda}(\theta) & F_{34,\lambda}(\theta) \\ 0 & 0 & -F_{34,\lambda}(\theta) & F_{44,\lambda}(\theta) \end{bmatrix} \begin{pmatrix} I_i \\ Q_i \\ U_i \\ V_i \end{pmatrix} \quad (1)$$

189
 190 Where the matrix elements $F_{ij,\lambda}(\theta)$ ($i, j = 1 - 4$) depend on the wavelength λ of the radiation (hereafter noted as a subscript)
 191 and carry information on the mineral dust particles size, shape and *CRI*. The scattering angle is $\theta = (\mathbf{k}_i, \mathbf{k})$, where $k = k_i =$
 192 $2\pi/\lambda$ is the wave vector of the electromagnetic wave. In lidar applications, the scattering angle is equal to π (i.e. exact
 193 backscattering angle). To highlight the need for laboratory measurements at the specific 180.0° lidar backscattering angle, near
 194 backscattering angles (i.e. $\theta < \pi$) are also considered in this section. Indeed, at $\theta = \pi$, $F_{33,\lambda} = -F_{22,\lambda}$ and $F_{12,\lambda} = F_{34,\lambda} = 0$
 195 (Zubko et al., 2013; David et al., 2013) while $F_{44,\lambda} = F_{11,\lambda} - 2F_{22,\lambda}$ due to the backscattering theorem (van de Hulst, 1957),
 196 so that Eq. (1) simplifies as follows for lidar applications:

$$197 \begin{pmatrix} I \\ Q \\ U \\ V \end{pmatrix} = \frac{1}{k^2 d^2} \begin{bmatrix} F_{11,\lambda}(\pi) & 0 & 0 & 0 \\ 0 & F_{22,\lambda}(\pi) & 0 & 0 \\ 0 & 0 & -F_{22,\lambda}(\pi) & 0 \\ 0 & 0 & 0 & F_{11,\lambda}(\pi) - 2F_{22,\lambda}(\pi) \end{bmatrix} \begin{pmatrix} I_i \\ Q_i \\ U_i \\ V_i \end{pmatrix} \quad (2)$$

199
 200 As a result, it is only at elastic lidar exact backscattering angle ($\theta = \pi$) that $F_{12,\lambda} = 0$ so that the scattering matrix reduces to
 201 only two non-vanishing elements $F_{11,\lambda}(\pi)$ and $F_{22,\lambda}(\pi)$.

202 3.2 Lidar particles depolarization ratio *PDR*

203 The expression of the so-called particles linear depolarization ratio (*PDR*) at wavelength λ and scattering angle θ can be
204 found in light scattering textbooks (Mishchenko et al., 2002; Schnaiter et al., 2012):

$$206 \quad PDR_{\lambda}(\theta) = \frac{1 - F_{22,\lambda}(\theta)/F_{11,\lambda}(\theta)}{1 \pm 2F_{12,\lambda}(\theta)/F_{11,\lambda}(\theta) + F_{22,\lambda}(\theta)/F_{11,\lambda}(\theta)} \quad (3)$$

207
208 where the positive (resp. negative) sign corresponds to *p*-polarized (resp. *s*-polarized) incident electromagnetic radiation. The
209 *PDR* stated in Eq. (3) is the linear *PDR*, which can be related to the circular *PDR* if need be (Mishchenko et al., 2002). Since
210 $F_{11,\lambda}$, $F_{12,\lambda}$ and $F_{22,\lambda}$ may vary with the scattering angle, depending on the dust sample, the dust *PDR* at near backscattering
211 angles ($\theta < \pi$) differs from that obtained at specific lidar backscattering angle ($\theta = \pi$). The deviation of $F_{11,\lambda}$, $F_{12,\lambda}$ and $F_{22,\lambda}$
212 from their value at exact backscattering angle cannot be quantified since no analytical light scattering theory exists for such
213 complex-shaped particles as mineral dust. Therefore, a laboratory experiment at specific lidar exact backscattering angle ($\theta =$
214 π) is required for precise evaluations of the dust lidar *PDR*. At specific lidar backscattering angle of π , Eq. (3) becomes:

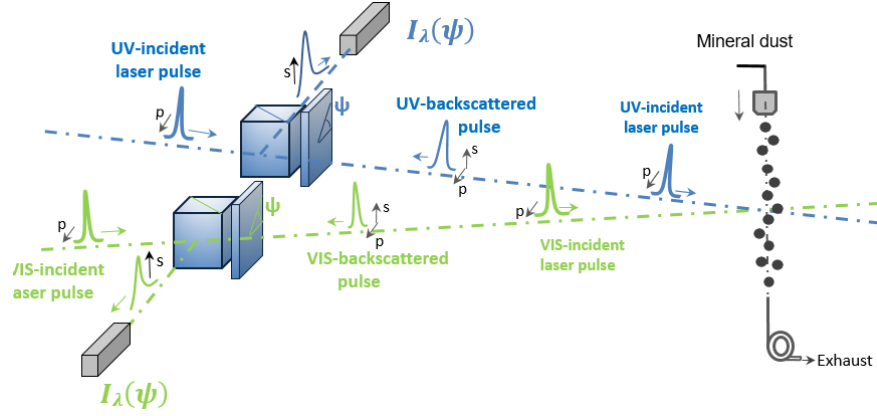
$$216 \quad PDR_{\lambda}(\pi) = \frac{1 - F_{22,\lambda}(\pi)/F_{11,\lambda}(\pi)}{1 + F_{22,\lambda}(\pi)/F_{11,\lambda}(\pi)} \quad (4)$$

217
218 Hence, accurate evaluations of the dust lidar *PDR* rely on accurate determinations of the ratio $F_{22,\lambda}/F_{11,\lambda}$ at specific lidar $\pi -$
219 angle. As for the ratio $F_{22,\lambda}/F_{11,\lambda}$, the dust lidar *PDR* is size, shape and refractive index dependent and this dependency is
220 discussed in Section 4. Spherical particles, for which $F_{22,\lambda}/F_{11,\lambda} = 1$, lead to $PDR_{\lambda}(\pi) = 0$. In what follows, to ease the
221 reading, the dust lidar *PDR* will be noted PDR_{λ} without reference to scattering angle ($\theta = \pi$).

222 3.3 Laboratory π -polarimeter for retrieving the lidar *PDR* of mineral dust

223 In (Miffre et al., 2016), for the first time to our knowledge, a laboratory π -polarimeter was built to address light
224 backscattering by aerosol particles. We here recall its main characteristics for clarity. The aerosols π -polarimeter is schemed
225 in Fig. 2. As in lidar applications, pulsed laser light is used to measure the time-of-flight taken by a laser pulse to reach the
226 dust sample and be detected after light backscattering. The backscattering geometry is set by inserting a well-characterized
227 polarizing beam splitter cube (*PBC*) between the emission and the dust samples, with a precision of 1 mm out of 10 meters to
228 ensure the π -polarimeter covers the lidar exact backscattering direction with accuracy: $\theta = (180.0 \pm 0.2)^{\circ}$. The laboratory
229 aerosol π -polarimeter is actually composed of two identical polarimeters, one per wavelength, to evaluate the lidar *PDR* of a
230 given dust sample at 355 and 532 nm wavelength simultaneously. Moreover, to decrease the retrieval uncertainty on the dust
231 *PDR*, the polarization state of the backscattered radiation is analysed for a set of incident polarization states of the incident

232 light using a quarter-wave plate (*QWP*). To validate the laboratory π -polarimeter, we carefully checked that homogeneous
 233 spherical particles, such as ammonium sulfate particles, which scatter light as described by the Mie theory (Bohren and
 234 Huffmann, 1983), were indeed providing zero lidar *PDR* when following the methodology described in the below section.
 235



236
 237 Figure 2: Scheme of the laboratory π -polarimeter operating at lidar exact backscattering angle of $(180.0 \pm 0.2)^\circ$ allowing accurate retrievals
 238 of the lidar *PDR* at 355 and 532 nm wavelength simultaneously for an aerosol sample (Miffre et al., 2016). The (p, s) polarization
 239 components are defined with respect to the laser scattering plane and ψ is the angle between the fast axis of the *QWP* and the laser scattering
 240 plane, counted counter-clockwise for an observer looking from the *PBC* to the particles. The dust lidar *PDR* is then evaluated from the ratio
 241 $F_{22,\lambda}/F_{11,\lambda}$ at specific π -angle, following the methodology described in Section 3.4.

242 3.4 Laboratory retrievals of mineral dust lidar *PDR*

243 We can formulate the *PDR* measurements of dust particles, using successive Mueller matrices denoting to the optical
 244 elements of the π -polarimeter and the scattering medium, encountered by the laser pulse from the laser source to the dust
 245 particles sample then back to the light detector. The measured backscattered intensity is :

$$247 I_\lambda(\psi) = \frac{\eta_\lambda P_\lambda}{d^2} [1, 0, 0, 0]^T [\mathbf{PBC}][\mathbf{QWP}(-\psi)][\mathbf{F}_\lambda][\mathbf{QWP}(\psi)][\mathbf{PBC}](St_i) \quad (5)$$

248
 249 Where η_λ is the optoelectronics efficiency of the light detector and P_λ is the laser power density, while $(St_i) = [1, 1, 0, 0]^T$ is
 250 the Stokes vector of the incident laser light. The expression of the dust backscattering matrix $[\mathbf{F}_\lambda]$ at wavelength λ is is given
 251 in Eq. (2), while $[\mathbf{PBC}]$ and $[\mathbf{QWP}(\pm\psi)]$ are the Mueller matrices of the *PBC* and the *QWP* respectively (Shurcliff, 1962).
 252 To develop Eq. (5), it is then advised to first calculate the raw vector $[1, 0, 0, 0]^T [\mathbf{PBC}][\mathbf{QWP}(-\psi)][\mathbf{F}_\lambda]$ then multiply it with
 253 the Stokes vector of the incident laser light $[\mathbf{QWP}(\psi)][\mathbf{PBC}](St_i)$ equal to $[1, \cos^2(2\psi), -\sin(4\psi)/2, -\sin(2\psi)]^T$, with ψ
 254 the modulation angle of the *QWP*. After a few calculations, the dust backscattered light intensity I_λ at wavelength λ is
 255 calculated as shown in Eq. 6:

$$257 I_\lambda(\psi) = I_{\lambda,0} \times [a_\lambda - b_\lambda \cos(4\psi)] \quad (6)$$

258

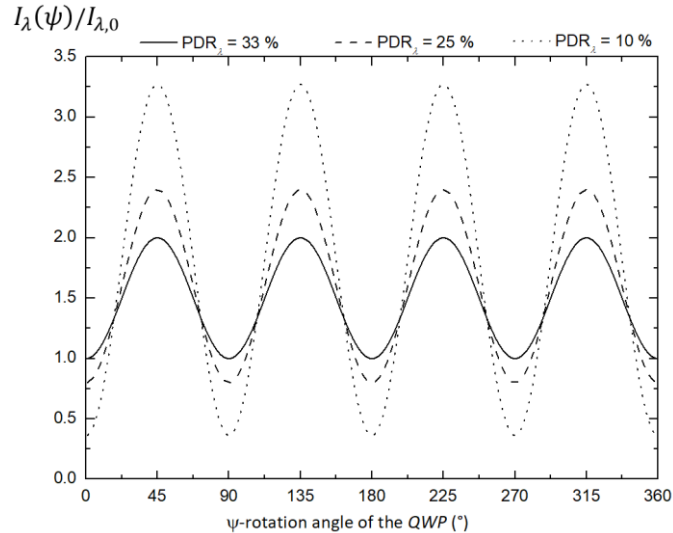
259 where the intensity $I_{\lambda,0} = \eta_{\lambda} P_{\lambda} / (4d^2)$, while coefficients a_{λ} and b_{λ} are equal to $a_{\lambda} = F_{11,\lambda} + F_{22,\lambda}$ and $b_{\lambda} = 3F_{22,\lambda} - F_{11,\lambda}$.
 260 Hence, $F_{22,\lambda}/F_{11,\lambda} = (1 + b_{\lambda}/a_{\lambda}) / (3 - b_{\lambda}/a_{\lambda})$ so that the ratio $F_{22,\lambda}/F_{11,\lambda}$ at π -angle can be determined from the ratio
 261 b_{λ}/a_{λ} . This ratio can be obtained from measurements of $I_{\lambda}(\psi)$, for different ψ -angles of the QWP, then adjusting these
 262 variations with Eq. (6) to get accurate determinations of $I_{\lambda,0} a_{\lambda}$ and $I_{\lambda,0} b_{\lambda}$, then b_{λ}/a_{λ} . Evaluations of the dust lidar PDR are
 263 then finally retrieved from Eq. (4):

264

$$265 \quad PDR_{\lambda} = (1 - b_{\lambda}/a_{\lambda})/2 \quad (7)$$

266

267 Within our methodology, the dust lidar PDR is independent of $I_{\lambda,0}$. For that reason, in Section 4, the applied voltage to the UV
 268 and VIS-photodetectors is adjusted to each dust SD and mineralogy to gain in accuracy in the retrieved dust lidar PDR by
 269 improving the signal-to-noise ratio on I_{λ} . For example, Fig. 3 provides simulations of $I_{\lambda}(\psi)/I_{\lambda,0}$ for the three following dust
 270 lidar PDR case studies : 33 % dust lidar PDR (in full lines, i.e. $F_{22,\lambda}/F_{11,\lambda} = 0.5$), 25 % dust lidar PDR (in dashed-lines, i.e.
 271 $F_{22,\lambda}/F_{11,\lambda} = 0.6$), 10 % dust lidar PDR (in dotted lines, i.e. $F_{22,\lambda}/F_{11,\lambda} = 0.82$). The curve minima, which are equal to
 272 $I_{\lambda,m}/I_{\lambda,0} = a_{\lambda} - b_{\lambda} = F_{11,\lambda} - F_{22,\lambda}$, are shape-dependent : each curve hence exhibits non-vanishing minima since mineral dust
 273 particles are nonspherical. Likewise, the curve maxima are equal to $I_{\lambda,M}/I_{\lambda,0} = a_{\lambda} + b_{\lambda} = 2F_{22,\lambda}$ and are size-dependent,
 274 though it is also shape dependent. The dust lidar PDR is determined from $I_{\lambda,m}$ and $I_{\lambda,M}$ since, following Eq. (7),
 275 $PDR_{\lambda} = I_{\lambda,m} / (I_{\lambda,m} + I_{\lambda,M})$, independently of $I_{\lambda,0}$.



276

277 Figure 3: Numerical simulation of the dust backscattered light intensity $I_{\lambda}(\psi)/I_{\lambda,0}$ as a function of the orientation ψ of the QWP at a given
 278 wavelength at the three following case studies: $PDR_{\lambda} = 33\%$ (in full lines, corresponding to $F_{22,\lambda}/F_{11,\lambda} = 0.50$), $PDR_{\lambda} = 25\%$ (in
 279 dashed-lines, $F_{22,\lambda}/F_{11,\lambda} = 0.60$), $PDR_{\lambda} = 10\%$ (in dotted lines, $F_{22,\lambda}/F_{11,\lambda} = 0.82$).

280 3.5 Accuracy on the retrieved laboratory mineral dust lidar PDR

281 Special care has been taken to quantify the uncertainties on the retrieved dust lidar *PDR*. The systematic errors in the π -
282 polarimeter are that encountered in 2λ -polarization lidar experiments, which we extensively studied (David et al., 2012) and
283 can also be found in polarization lidar reference papers (Freudenthaler, 2016). To summarize, systematic errors arise from:

- 284 • *Imperfect definition of the polarization state of the incident radiation.* In the π -polarimeter, the polarization state of
285 the electromagnetic radiation emitted from the laser is precisely set to $[1, 1, 0, 0]^T$ (i.e. with no remaining ellipticity)
286 by using two successive *PBC*.
- 287 • *Polarization cross-talk between the emitter and the detector polarization axes.* Likewise, on the detector side, to
288 account for the imperfections of the retro-reflecting *PBC* ($R_s > 99.5\%$, $T_p > 90\%$), a secondary *PBC* is inserted
289 between the retro-reflecting *PBC* and the light detector to ensure polarization cross-talk or undesired fraction $R_p T_s$
290 originating from the *p*-component of the backscattered radiation are fully negligible. Hence, the π -polarimeter is
291 sensitive to the *s*-component of the backscattered radiation only. Also, the emitting *PBC* being used as retro-reflecting
292 *PBC*, any possible mismatch between the *s*-polarization axis of the emitted and detected backscattered radiations
293 cannot occur.
- 294 • *Spectral cross-talk between the UV and the VIS-backscattered radiations.* Likewise, wavelength cross-talk is
295 minimized by using selective interference filters exhibiting a higher than 5 optical density, at 355 nm wavelength in
296 the VIS π -polarimeter and at 532 nm wavelength in the UV π -polarimeter.
- 297 • *Multiple scattering can induce further light depolarization.* However, the single-scattering approximation is ensured
298 in our laboratory backscattering experiment (Mishchenko et al., 2007) where the particles are moving in a thin (2.5
299 mm) wide beam so that the volume element is optically thin in contrary to atmospheric chambers (1100 cm^{-3} for the
300 coarser SD).

301

302 Finally, to account for potential fluctuations in the dust particle number concentration that may cause variations in the dust
303 backscattered light intensity I_λ , a normalization channel has been added to the π -polarimeter by including a polarization-
304 insensitive light detector operating at scattering angle $\theta_0 = 165^\circ$. The corresponding scattered light intensity $I_\lambda(\theta_0)$ is
305 quantified similarly to Eq. 5 considering a scattering angle of θ_0 : $I_\lambda(\theta_0) = [1, 0, 0, 0]^T [\mathbf{F}_\lambda(\theta_0)] [\mathbf{QWP}(\psi)] [\mathbf{PBC}] [1, 1, 0, 0]^T$,
306 where $[\mathbf{F}_\lambda(\theta_0)]$ is the scattering matrix at angle θ_0 . There, the *QWP* and the *PBC* only act on the detector side while
307 (St_i) equals $[1, 1, 0, 0]^T$. Hence, $I_\lambda(\theta_0) = I_{\lambda,0} \times [2F_{11,\lambda}(\theta_0) + F_{12,\lambda}(\theta_0) + F_{12,\lambda}(\theta_0)\cos(4\psi)]$. Once the variations of $I_\lambda(\theta_0)$
308 with ψ -angle are recorded, the $\cos(4\psi)$ -dependency of $I_\lambda(\theta_0)$ can be removed by applying a numerical low-pass filter on
309 $I_\lambda(\theta_0)$, to get a light intensity proportional to the dust particles number concentration. As a result, in the light backscattering
310 curves presented in Section 4, the plotted quantity is the normalized backscattered light intensity $I_{\lambda,N} = I_\lambda(\pi)/I_\lambda(\theta_0)$, which
311 is insensitive to potential fluctuations in the dust particles number concentration. The scattered light intensities $I_\lambda(\pi)$ and

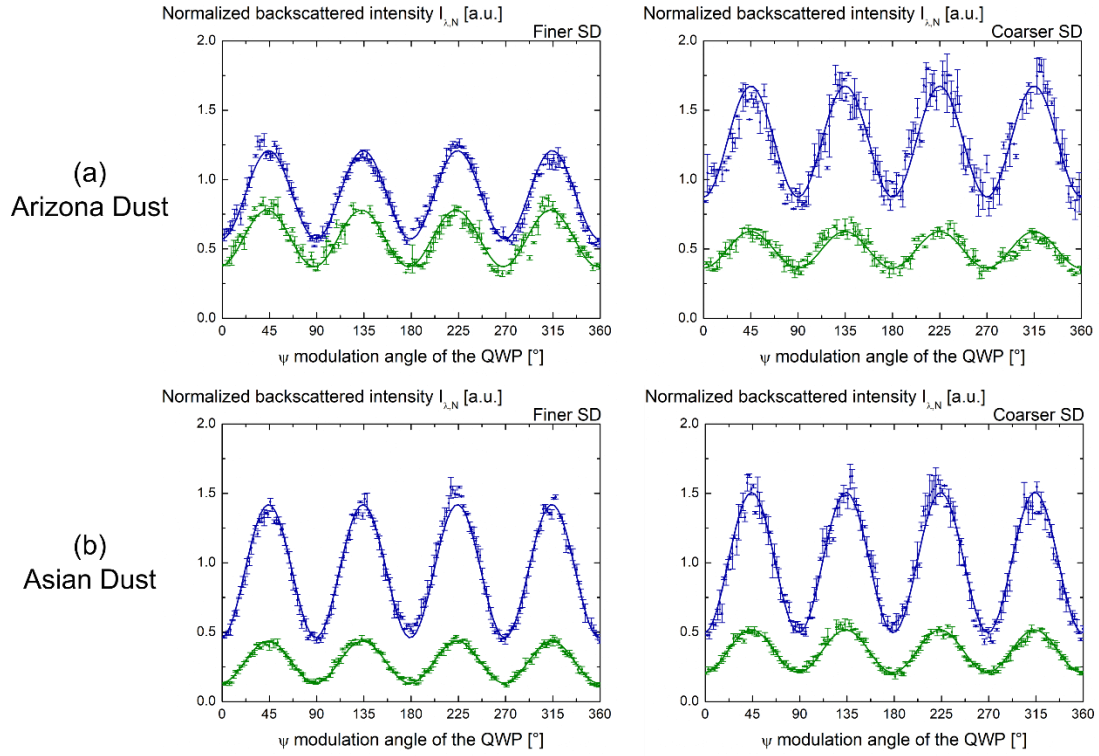
312 $I_\lambda(\theta_0)$ being correlated, the standard deviation σ_N on $I_{\lambda,N}$ was calculated by considering the covariance $\sigma_{I_\lambda I_\lambda(\theta_0)}$ of I_λ and
313 $I_\lambda(\theta_0)$. This covariance contributes to the uncertainty on $I_{\lambda,N}$ at a rate $-2I_\lambda \sigma_{I_\lambda I_\lambda(\theta_0)} / I_\lambda^3(\theta_0)$. Moreover, to gain in accuracy in
314 the dust lidar *PDR* retrievals, $I_{\lambda,N}$ was measured for a complete ψ -angle rotation, while averaging the acquired backscattered
315 light intensity over several thousand laser shots per ψ -angle, with resulting mean and standard deviations on $I_{\lambda,N}$ as plotted in
316 Fig. 4 and 5.

317 **4. Results and discussion**

318 In this section, using the methodology presented in Section 3, the lidar *PDR* of Arizona dust, Asian dust, silica and hematite
319 is evaluated and discussed at 355 and 532 nm wavelength for the finer and the coarser *SD*.

320 **4.1 Laboratory evaluation of the lidar *PDR* of Arizona and Asian dust**

321 Figure 4 displays the variations of $I_{\lambda,N}$ for Arizona (Fig. 4a) and Asian dust (Fig. 4b) as a function of the ψ -rotation angle
322 of the *QWP* for the finer (left panels) and the coarser *SD* (right panels) at 355 and 532 nm wavelength. The observed variations
323 are related to a determined size and shape distribution of the dust sample: indeed, as explained in Section 3.4, if the size (resp.
324 the shape) of the dust sample was varying during our acquisitions, the maxima (resp. the minima) of the curves would not
325 remain constant. As a result, the observed variations of $I_{\lambda,N}$ reveal the spectral and polarimetric backscattering characteristics
326 of each considered dust sample. Therefore, the experimental data points could be fitted with Eq. (6) to evaluate $F_{22,\lambda}/F_{11,\lambda}$
327 then the dust lidar *PDR* by applying Eq. (7). Table 1 presents the retrieved values of $F_{22,\lambda}/F_{11,\lambda}$ and of dust lidar *PDR*. The
328 uncertainty on $F_{22,\lambda}/F_{11,\lambda}$ results from the measurement errors of the laboratory π -polarimeter and leads to accurate evaluations
329 of the dust lidar *PDR*. Within experimental error bars, the lidar *PDR* of Arizona and Asian dust clearly differ, whatever the
330 chosen wavelength. The generally admitted value of around 33 % for the dust lidar *PDR* (Tesche et al., 2009) is only obtained
331 for Arizona dust: Asian dust exhibits a lower *PDR* in the range from 24 to 28 % depending on the considered *SD* and
332 wavelength. This suggests that the dust lidar *PDR* is primarily governed by the dust mineralogy and hence particles refractive
333 index. The sensitivity of the dust lidar *PDR* with the considered *SD* is indeed less pronounced: from the coarser to the finer
334 *SD*, a reduction in the dust lidar *PDR* of below 5 % is observed at 532 nm wavelength. At 355 nm wavelength however, the
335 Arizona and Asian dust lidar *PDR* seem practically insensitive to variations in the considered *SD*.



336

337 Figure 4: Normalized backscattered light intensity $I_{\lambda,N} = I_{\lambda}(\boldsymbol{\pi})/I_{\lambda}(\boldsymbol{\theta}_0)$ of Arizona (a) and Asian dust (b) for finer *SD* (left panels) and
 338 coarser *SD* (right panels), using the laboratory π -polarimeter at lidar exact backscattering angle ($\boldsymbol{\theta} = \boldsymbol{\pi}$) at 355 (blue) and 532 (green) nm.
 339 The experimental data points are fitted with Eq. 6 to derive $F_{22,\lambda}/F_{11,\lambda}$ and then the dust lidar *PDR* is derived using Eq. 7. Care should be
 340 taken when comparing $I_{\lambda,N}$ for Arizona and Asian dust since the applied voltage to the UV and VIS-photodetectors was adjusted to increase
 341 the signal-to-noise ratio, as explained in Section 3.4. The Arizona dust lidar *PDR*, retrieved from $I_{\lambda,m}/(I_{\lambda,m} + I_{\lambda,M})$, is higher than that of
 342 Asian dust.

343

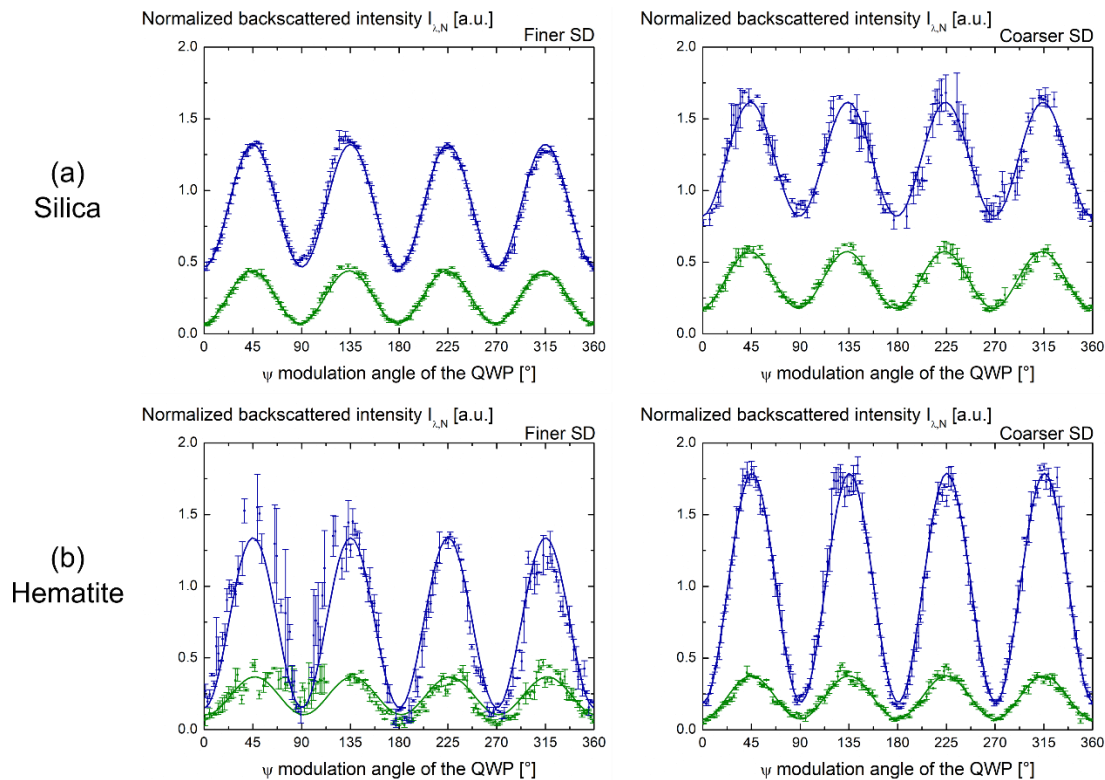
344 Tab. 1: Laboratory measurement of the *PDR* of Arizona and Asian dust at 355 (blue) and 532 nm (green), for the finer and the coarser *SD*.
 345 The *PDR* is calculated with Eq. 7 after the derivation of $F_{22,\lambda}/F_{11,\lambda}$ using the laboratory π -polarimeter presented in Section 3.2. The
 346 uncertainty on $F_{22,\lambda}/F_{11,\lambda}$ is deduced from the evaluation of b_{λ}/a_{λ} , itself deduced from the least-square fit adjustment of I_{λ} . The uncertainty
 347 on $F_{22,\lambda}/F_{11,\lambda}$ is mostly dominated by statistical uncertainties since our biases are minimized, as explained in Section 3.5.
 348

Mineralogy	λ (nm)	Finer <i>SD</i>		Coarser <i>SD</i>	
		$F_{22,\lambda}/F_{11,\lambda}$	PDR_{λ} (%)	$F_{22,\lambda}/F_{11,\lambda}$	PDR_{λ} (%)
Arizona dust	355	0.514 ± 0.007	32.1 ± 0.6	0.489 ± 0.012	34.3 ± 1.0
	532	0.512 ± 0.012	32.3 ± 1.0	0.464 ± 0.012	36.6 ± 1.1
Asian dust	355	0.603 ± 0.009	24.7 ± 0.6	0.603 ± 0.011	24.8 ± 0.8
	532	0.622 ± 0.009	23.3 ± 0.7	0.558 ± 0.011	28.4 ± 0.8

349

350 4.2 Laboratory evaluation of the lidar *PDR* of silica and hematite

351 By applying the same methodology, we obtain the *PDR* of silica and hematite, as presented in Fig. 5 and Table 2.
352 Accordingly, Fig. 5 is the analog of Fig. 4 for silica (Fig. 5a) and hematite (Fig. 5b). As for Arizona and Asian dust samples,
353 the lidar *PDR* of silica and hematite primarily depends on the particles *CRI*, at least at 355 nm wavelength where the silica
354 lidar *PDR* ranges from 23 to 33 % depending on the considered *SD* while the hematite lidar *PDR* reaches 10 % only. The silica
355 lidar *PDR* also strongly depends on the particles diameter: from the coarser to the finer *SD*, the silica dust lidar *PDR* reduces
356 by 10 % at both wavelengths. The dependence of the hematite dust lidar *PDR* with the *SD* is less pronounced, especially at
357 355 nm wavelength. The silica and hematite lidar *PDR* also strongly depend on the chosen lidar wavelength, with higher
358 depolarization observed at 355 nm wavelength for silica and at 532 nm wavelength for hematite.
359



360
361 Figure 5: Same as Fig. 4 for silica (a-plots) and for hematite samples (b-plots).

362

363

364
365
366

Tab. 2: Same as Table 1 for silica and hematite.

Mineralogy	λ (nm)	Finer SD		Coarser SD	
		$F_{22,\lambda}/F_{11,\lambda}$	PDR_{λ} (%)	$F_{22,\lambda}/F_{11,\lambda}$	PDR_{λ} (%)
Silica	355	0.622 ± 0.014	23.3 ± 0.9	0.506 ± 0.011	32.8 ± 1.0
	532	0.751 ± 0.016	14.2 ± 0.9	0.618 ± 0.016	23.6 ± 1.1
Hematite	355	0.805 ± 0.050	10.8 ± 2.5	0.823 ± 0.015	9.7 ± 0.7
	532	0.652 ± 0.055	21.1 ± 3.5	0.715 ± 0.019	16.6 ± 1.1

367

368 4.3 Discussion

369 Comparing our laboratory findings with other laboratory experiments is not straightforward, since as explained in the
370 introduction, none operates at 180.0° lidar exact backscattering angle, while the dust lidar PDR differs from near to exact
371 backscattering angles, especially when light absorbents are present (Cholleton et al., 2022). Moreover, the PDR is wavelength-
372 dependent and the size distributions (SD) used are different from other studies. Lidar field experiments provide accurate values
373 of the lidar PDR after accurate calibration procedure based on the scattering matrix (Freudenthaler, 2016; Belegante et al.,
374 2018; Miffre et al., 2019). Although in such lidar field experiments, the measured PDR is usually that of dust mixtures (Miffre
375 et al., 2011), the comparison with our laboratory findings remains interesting. In lidar retrievals (see for example (Tesche et
376 al., 2009)), a dust lidar PDR of 30 % is often used. The laboratory π -polarimeter verifies this statement by providing the silica
377 PDR , which is the main oxide present in mineral dust, equal to (33 ± 1) % for the coarser SD at 355 nm. In comparison,
378 within our experimental error bars, the hematite lidar PDR , equal to (10 ± 1) %, is clearly lower. The real part n and the
379 imaginary part κ of the hematite CRI , which are large compared with that of other chemical oxides present in mineral dust (see
380 Section 2.1), can be responsible for the observed difference in the silica and hematite lidar PDR . Indeed, n and part κ modify
381 the backscattering matrix elements, so does the corresponding dust lidar PDR . To highlight the role of κ on the hematite lidar
382 PDR , the lidar PDR of rutile was measured with our π -polarimeter. Indeed, the real part of the rutile CRI is as large as that of
383 hematite but its imaginary part is negligible compared with that of hematite. As a result, the rutile lidar PDR substantially
384 differed from that of hematite, showing the key role played by light absorption in the measured hematite lidar PDR . In turn,
385 Arizona dust exhibits a higher PDR than Asian dust, due to the higher proportion in hematite in the latter. Hence and as a
386 conclusion, our laboratory findings show that, when the light absorbent hematite is present, it mainly governs the dust lidar
387 PDR , which hence primarily depends on the particles mineralogy, with less pronounced variations with the particles size and
388 wavelength. This finding is in line with (Kahnert, 2015; Kahnert et al., 2020) numerical findings, who highlighted that the dust
389 PDR is strongly modulated by the particles inhomogeneity, especially in the lidar backward scattering direction and in the
390 presence of hematite. We here quantify this effect with a laboratory experiment that accounts for the real shape of mineral

391 dust. The shape dependence of the hematite PDR is weak due to its large imaginary part of complex refractive index: following
 392 Wiscombe and Mugnai (1986) or Mishchenko et al. (1997), the effect of particle shape becomes weaker with increasing
 393 imaginary part of the refractive index, a conclusion also drawn by Meland et al. (2011). In contrast, when the proportion of
 394 hematite becomes negligible, as is the case for silica and Arizona dust, our laboratory findings show that the dust lidar *PDR*
 395 then increases with increasing the particles size, though the shape dependence may then also play a role. Also, it would be
 396 interesting to investigate giant dust particles (Ryder et al., 2019). Likewise, in the literature (Sakai et al., 2010; Hofer et al.,
 397 2020; Järvinen et al., 2016; Mamouri and Ansmann, 2017), the dust lidar *PDR* is usually found to increase with the particles
 398 size from the fine to the coarse mode of the *SD*. The (355, 532) nm wavelength dependence of the dust lidar *PDR* then becomes
 399 key for discussing on the involved particle sizes, thus underlying the importance of dual-wavelength (or more) polarization
 400 lidar instruments. We here establish this result in laboratory at 180.0° and (355, 532) nm wavelength, and moreover, show that
 401 this consideration holds only when hematite, which is a strong light absorbent, is not involved : the hematite lidar *PDR* is
 402 indeed higher in the finer mode of the *SD*.

403

404 To go further and discuss on the role of light absorption in the retrieved dust lidar *PDR*, we here propose a basic partitioning
 405 model in which the dust particles mixture (d) = $\{Abs, \overline{Abs}\}$ is comprised of two components: an absorbing component (*Abs*),
 406 mainly corresponding to hematite particles, and a non-absorbing component (\overline{Abs}), mainly corresponding to silica-particles.
 407 For simplicity, we here resume the absorbing (resp. non-absorbing) component to hematite (resp. silica)-particles with
 408 respective abbreviations (*Hmt*) and (*Sil*). We focus on the 355 nm wavelength at which hematite is an efficient light absorber
 409 and on the coarser *SD* as the dependence of the dust lidar *PDR* with size is less pronounced than with the particles mineralogy.
 410 In Appendix A is detailed the derivation of the lidar *PDR* of such a dust-particles mixture (d) = $\{Hmt, Sil\}$ (hereafter noted
 411 δ_d , as in lidar applications). This Appendix is an extension of our previous works (Miffre et al., 2011; David et al., 2013, 2014;
 412 Mehri et al., 2018) for the case study where both components $\{Hmt, Sil\}$ are nonspherical. The lidar *PDR* of such a dust-
 413 particles mixture relates to that of its pure components (hereafter noted δ_{Sil} and δ_{Hmt}) as follows:

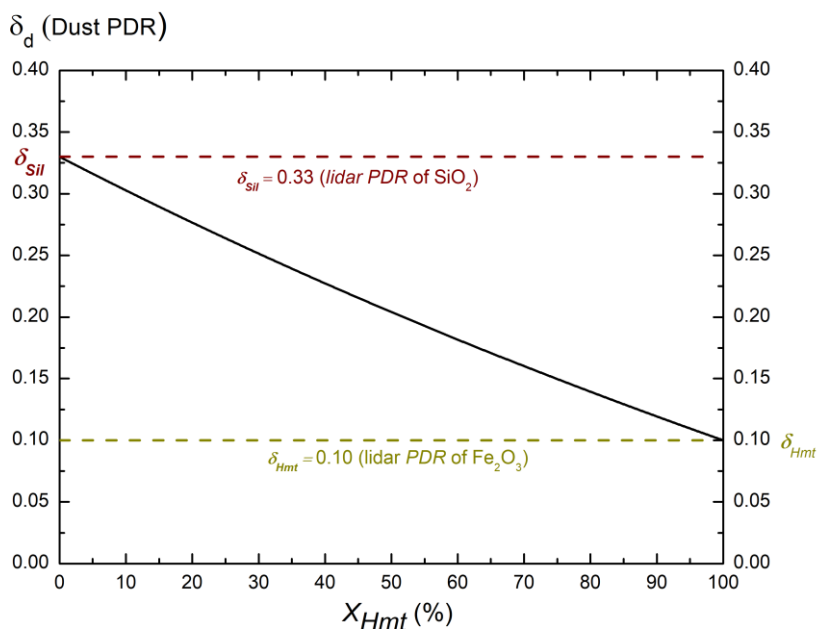
414

$$415 \quad \delta_d = \frac{-e + (c + e)X_{Hmt}}{f - (d + f)X_{Hmt}} \quad (8)$$

416

417 where the expressions of the c, d, e and f -coefficients are provided in Appendix A and independently on the depolarization
 418 ratios δ_{Sil} and δ_{Hmt} of silica and iron oxides. X_{Hmt} is the fraction of *Hmt* to dust particles backscattering. Following Eq. (8)
 419 and Appendix A, Fig. 6 displays the variation of δ_d as a function of X_{Hmt} when considering $\delta_{Sil} = 33\%$ and $\delta_{Hmt} = 10\%$, as
 420 obtained in our laboratory findings at 355 nm wavelength with the coarser *SD*. As shown in Fig. 6, the dust lidar *PDR* lies in
 421 between δ_{Sil} and δ_{Hmt} and equals δ_{Sil} (resp. δ_{Hmt}) only when $X_{Hmt} = 0$ (resp. 1), depending on the fraction X_{Hmt} of light
 422 corresponding to the absorbent of the dust particle mixture. Hence, Arizona dust, which contains a lower fraction of hematite,
 423 exhibits a higher lidar *PDR* compared with Asian dust, at least at 355 nm wavelength where hematite is strongly absorbing.

424 Though rather simple, our model interestingly highlights the key role played by light absorption in the retrieved Asian dust
 425 lidar *PDR*. To go further and provide a quantitative analysis, this simple model should be refined, by considering also the other
 426 chemical oxides present in mineral dust, other lidar wavelengths, as well as other *SD* and the effect of shape. To handle such
 427 a complex issue, more laboratory experiments are required on other chemical oxides, ideally also at 1064 nm wavelength. This
 428 work is however beyond the scope of this paper. Still as is, our model provides an interpretation of the laboratory-observed
 429 differences in the dust lidar *PDR* when the light absorbent hematite is involved. In the most general case, the dust lidar *PDR*
 430 hence appears as a complex function of the particles mineralogy, *SD*, wavelength and shape. Comparison with lidar field
 431 experiments, involving particle mixtures, with a more complex distribution of sizes and refractive indices, is then not
 432 straightforward, as underscored by comparison with Hu et al. (2020) who reported $0.28 - 0.32 \pm 0.07$ at 355 nm wavelength.
 433 Though this complex dependence is difficult to disentangle, our laboratory findings show that the dust lidar *PDR* is primarily
 434 affected by the particles mineralogy, at least when hematite is involved.
 435



436

437 Figure 6: Numerical simulation of the 355 nm Lidar *PDR* of a two-component particles mixture (d) = {*Hmt*, *Sil*}, composed of hematite
 438 (*Hmt*) and silica (*Sil*) oxides as a function of the $X_{Hmt} = \beta_{Hmt}/\beta_d$ fraction of *Hmt* to *d*-particles backscattering, following Eq. (8) and
 439 Appendix A, by accounting for our laboratory experimental findings for $\delta_{Sil} = 33\%$ and $\delta_{Hmt} = 10\%$ (see Table 2 at 355 nm wavelength
 440 with coarser *SD*).

441 5 Summary and conclusion

442 In this paper, the dependence of the lidar particles depolarization ratio (*PDR*) of pure mineral dust with complex refractive
443 index (*CRI*) and size is for the first time investigated through a laboratory π -polarimeter operating at 180.0° lidar
444 backscattering angle and (355, 532) nm wavelengths for lidar purposes. The goal of this work is to improve the knowledge on
445 the dust lidar *PDR*, which is an important input parameter involved in lidar partitioning algorithms, which are widely applied
446 to reveal the contribution of mineral dust in particles external mixtures (Tesche et al., 2009; Mehri et al., 2018). While mineral
447 dust exhibits a complex and highly irregular shape, which is difficult to model mathematically and numerically, our laboratory
448 approach allows accounting for the real shape of mineral dust. Our laboratory π -polarimeter is likewise a good complement to
449 lidar field experiments, which provide accurate retrievals of the lidar *PDR* of particles mixtures involving mineral dust.
450 Another advantage of our laboratory π -polarimeter lies in its ability to provide accurate retrievals of the lidar *PDR* of pure
451 mineral dust samples, differing in *CRI* and size. The π -polarimeter indeed operates at 180.0° lidar backscattering angle and at
452 (355, 532) nm lidar wavelengths: no assumption is made to retrieve the dust lidar *PDR*. This is a key novelty of our study.
453 Indeed, the variation of the dust lidar *PDR* with scattering angle and wavelength cannot be analytically calculated (Bohren and
454 Huffman, 1983; Mishchenko et al., 2002) for complex-shaped particles such as mineral dust. Hence, our π -polarimeter
455 improves the knowledge on the dust *PDR*, provided in the literature at non 180.0° backscattering angle and / or at wavelengths
456 differing from (355, 532 nm). Our work provides sixteen accurate dust lidar *PDR*-values, corresponding to four different
457 complex refractive indices, studied at two size distributions (fine, coarse) and at (355, 532) nm wavelengths (see Section 4).
458 The precision on the retrieved dust *PDR* from the laboratory π -polarimeter is detailed in Section 3. To investigate the
459 dependence of the dust lidar *PDR* with *CRI*, hematite, the main light absorbent present in mineral dust, was considered in
460 addition to silica oxide, the main chemical oxide present in mineral dust, which is practically nonabsorbent. At 355 nm, our
461 laboratory π -polarimeter provides values of the *PDR* of coarser silica of $(33 \pm 1) \%$ while that of coarser hematite is only
462 $(10 \pm 1) \%$. In Section 4, this large difference is explained by accounting for the high imaginary part of the hematite *CRI*. In
463 turn, Arizona dust exhibits a higher depolarization ratio than Asian dust, due to the higher proportion in hematite in the latter.
464 As a result, when the strong light absorbent hematite is involved, the dust lidar *PDR* is primarily governed by the particles
465 mineralogy and the variations of the dust lidar *PDR* with size are less pronounced. The dependence of the PLDR on the
466 particles shape is not pronounced in our experiment where hematite, which exhibits a large imaginary part of complex
467 refractive index, plays a key role (Wiscombe and Mugnai, 1986, Mishchenko et al., 1997, Meland et al., 2011). When hematite
468 is less or not involved, the dust lidar *PDR* increases with increasing sizes and the (355, 532) nm wavelength dependence of
469 the dust lidar *PDR* then becomes key for discussing on the involved particle sizes, thus underscoring the importance of dual
470 wavelengths (or more) polarization lidar instruments. To further disentangle the complex dependence of the dust lidar *PDR*
471 with complex refractive index and size, our methodology should be extended to other chemical oxides, other natural mineral
472 dust samples, other *SD* and other wavelengths, as well as other shape distributions. Giant dust particles, whose importance has
473 been highlighted by Ryder et al. (2019), would likewise be interesting to study specifically. This is however far beyond the

474 scope of this paper : we here focused on (355, 532) nm wavelengths, since mineral dust slightly absorb light in the near infra-
 475 red (Di Biagio et al., 2019). Still, the above laboratory findings underscore the importance of accounting for the wavelength
 476 dependence of the dust lidar *PDR*, whatever the hematite proportion. The spectral dependence of the dust lidar *PDR* is indeed
 477 instructive (Burton et al., 2016; Haairig et al., 2022; Miffre et al., 2020). Outlooks of this work are obviously also interesting,
 478 as underscored by recent papers (Kahnert et al., 2020; Luo et al., 2022), discussing on the ability of the spheroidal model to
 479 mimic light scattering by complex-shaped mineral dust.

480 Appendix A

481 The goal of this Appendix is to establish the expression of the lidar *PDR* of a two-component particle mixture (p) =
 482 $\{ns_1, ns_2\}$ composed of two non-spherical components ns_1 and ns_2 . As in lidar applications, the lidar *PDR* of p , ns_1 and ns_2 -
 483 particles are respectively noted δ_p , δ_{ns_1} and δ_{ns_2} . The starting point is given by the set of four equations:

484

$$485 \beta_{p, //} = \beta_{ns_1, //} + \beta_{ns_2, //} \quad (\text{A-1-a})$$

$$486 \beta_{p, \perp} = \beta_{ns_1, \perp} + \beta_{ns_2, \perp} \quad (\text{A-1-b})$$

$$487 \delta_{ns_1} = \beta_{ns_1, \perp} / \beta_{ns_1, //} \quad (\text{A-1-c})$$

$$488 \delta_{ns_2} = \beta_{ns_2, \perp} / \beta_{ns_2, //} \quad (\text{A-1-d})$$

489

490 where $\beta_{p, //}$ and $\beta_{p, \perp}$ are the lidar particles backscattering coefficients, evaluated from a polarization lidar experiment carried
 491 out at wavelength λ (here omitted to ease the reading). The backscattering coefficient β_{ns_1} of ns_1 -particles is then retrieved by
 492 noting that $\beta_{ns_1} = \beta_{ns_1, //} + \beta_{ns_1, \perp} = \beta_{ns_1, \perp}(1 + 1/\delta_{ns_1})$ (Miffre et al., 2011; David et al., 2013). Moreover, $\beta_{ns_1, \perp}$ can be
 493 expressed as a fonction of $\beta_{p, //}$ and $\beta_{p, \perp}$ since $\beta_{ns_1, \perp} = \beta_{p, \perp} - \beta_{ns_2, \perp} = \beta_{p, \perp} - \delta_{ns_2}\beta_{ns_2, //} = \beta_{p, \perp} - \delta_{ns_2}(\beta_{p, //} - \beta_{ns_1, \perp}/\delta_{ns_1})$
 494 using Eqs. (A-1). Hence, $\beta_{ns_1, \perp} = (\beta_{p, \perp} - \delta_{ns_2}\beta_{p, //}) / (1 - \delta_{ns_2}/\delta_{ns_1})$. By applying the same methodology to ns_2 - particles,
 495 we finally get:

496

$$497 \begin{pmatrix} \beta_{ns_1} \\ \beta_{ns_2} \end{pmatrix} = \begin{bmatrix} c & d \\ e & f \end{bmatrix} \begin{pmatrix} \beta_{p, //} \\ \beta_{p, \perp} \end{pmatrix} \quad (\text{A-2})$$

498

499 where the c, d, e and f -coefficients only depend on the depolarization ratios δ_{ns_1} and δ_{ns_2} :

500

$$501 c = -\delta_{ns_2}(1 + 1/\delta_{ns_1}) / (1 - \delta_{ns_2}/\delta_{ns_1}) \quad (\text{A-3-a})$$

$$502 d = (1 + 1/\delta_{ns_1}) / (1 - \delta_{ns_2}/\delta_{ns_1}) \quad (\text{A-3-b})$$

$$503 e = -\delta_{ns_1}(1 + 1/\delta_{ns_2}) / (1 - \delta_{ns_1}/\delta_{ns_2}) \quad (\text{A-3-c})$$

$$504 \quad f = (1 + 1/\delta_{ns_2})/(1 - \delta_{ns_1}/\delta_{ns_2}) \quad (\text{A-3-d})$$

505

506 The 2 x 2 matrix introduced in Eq. (A-2) can be inverted to get the expression of $\beta_{p, //}$ and $\beta_{p, \perp}$ and hence $\delta_p = \beta_{p, \perp}/\beta_{p, //}$. By
 507 introducing $X_{ns_2} = \beta_{ns_2}/(\beta_{ns_1} + \beta_{ns_2})$ the fraction of ns_2 to p -particles backscattering, we finally get the relationship
 508 between δ_p and δ_{ns_1} and δ_{ns_2} :

509

$$510 \quad \delta_p = \frac{-e + (c + e)X_{ns_2}}{f - (d + f)X_{ns_2}} \quad (\text{A - 4})$$

511

512 In the specific case where ns_2 -particles are spherical (i.e. $\delta_{ns_2} = 0$), the expressions of the c, d, e and f -coefficients simplify
 513 and the relationship between δ_p and $X_{ns_2} = X_{ns}$ becomes identical to that we already published in (Miffre et al., 2011; David
 514 et al., 2013). This new material is hence as an extension of our previous works (Miffre et al., 2011; David et al., 2013, 2014;
 515 Mehri et al., 2018) to the case study where both components of the particles mixture (p) = $\{ns_1, ns_2\}$ are nonspherical.

516

517 **Author contribution**

518 **Alain Miffre**: Conceptualization, Formal analysis, Investigation, Methodology, Supervision, Writing - original draft, Writing
 519 - review & editing **Danaël Cholleton**: Formal analysis, Investigation, Software, Visualization, Writing - review & editing.
 520 **Clément Noël**: Software, Writing - review & editing **Patrick Rairoux**: Project administration, Supervision, Writing - review
 521 & editing.

522 **Competing interests**

523 The authors declare that they have no conflict of interest.

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