



# 1 Investigating the dependence of mineral dust depolarization on 2 complex refractive index and size with a laboratory polarimeter at 3 180.0° lidar backscattering angle

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8 **Abstract.** In this paper, the dependence of the particles depolarization ratio (*PDR*) of mineral dust on the complex refractive  
9 index and size is for the first time investigated through a laboratory  $\pi$ -polarimeter operating at 180.0° backscattering angle and  
10 at (355, 532) nm wavelengths for lidar purposes. The dust *PDR* is indeed an important input parameter in polarization lidar  
11 experiments involving mineral dust. Our  $\pi$ -polarimeter provides sixteen accurate values of the dust lidar *PDR* at 180.0°  
12 corresponding to four different complex refractive indices, studied at two size distributions (fine, coarse) and at (355, 532) nm  
13 wavelengths, while accounting for the highly irregular shape of mineral dust, which is difficult to model numerically. At 355  
14 nm, the lidar *PDR* of coarser silica, the main oxide in mineral dust, is equal to  $(33 \pm 1)$  % while that of coarser hematite, the  
15 main light absorbent in mineral dust, is  $(10 \pm 1)$  %. This huge difference is here explained by accounting for the high  
16 imaginary part of the hematite complex refractive index. In turn, Arizona dust exhibits higher depolarization than Asian dust,  
17 due to the higher proportion in hematite in the latter. As a result, when the strong light absorbent hematite is involved, the dust  
18 lidar *PDR* primarily depends on the particles complex refractive index and its variations with size are less pronounced. When  
19 hematite is less or not involved, the dust lidar *PDR* increases with increasing sizes and the (355, 532) nm wavelength  
20 dependence of the dust lidar *PDR* then allows discussing on the involved particle sizes, thus highlighting the importance of  
21 dual-wavelength (or more) polarization lidar instruments. We believe these laboratory findings will help improving our  
22 understanding of the challenging dependence of the dust lidar *PDR* with complex refractive index and size to help interpret  
23 the complexity and the wealth of polarization lidar signals.

## 24 1 Introduction

25 With worldwide annual emissions between 1000 to 3000 Tg (Monge et al., 2012), mineral dust is a highly important constituent  
26 of the atmosphere, which contributes to ice cloud formation by acting as a freezing nucleus and to the carbon cycle by fertilizing  
27 nutrient poor ecosystems such as the Amazon rainforest after long-range transport (Bristow et al., 2010). As underscored in  
28 the latest IPCC report (2021), mineral dust also contributes to the Earth's radiative budget through light scattering and  
29 absorption, by reducing the amount of energy reaching the Earth's surface (Kosmopoulos et al., 2017). The radiative impact



30 associated with a Saharan dust storm has been recently quantified by (Francis et al., 2022). This climatic impact is however  
31 subject to large uncertainties, mainly due to the great complexity in size, shape and mineralogy of mineral dust. In the  
32 atmosphere, the size distribution of mineral dust is mainly determined by the distance from the dust source region. Two freshly  
33 uplifted dust aerosols may indeed exhibit different size distributions at far-range remote sites (Ryder et al., 2013), due to the  
34 rapid removal of the largest particles by gravitational settling. Mineral dust particles also exhibit a high degree of complexity  
35 in shape. Electron microscopic images (Kandler et al., 2011) indeed highlight the nonspherical and highly irregular shape of  
36 mineral dust particles, with sharp edges, sometimes even surface roughness (Nousiainen, 2009). The mineral dust surface is  
37 itself subject to photo-catalytic reactions leading even to new particle formation events (Dupart et al., 2012). The third degree  
38 of complexity of mineral dust related to this study lies in its mineralogy. Mineral dust indeed consists in a heterogeneous  
39 mixture of various chemical oxides among which the most predominant is silica oxide. Aluminum and iron oxides are also  
40 present in proportions depending on the dust source region. As an example, the desert in Central Australia is iron oxides rich  
41 (Bullard and White, 2002). This diverse mineralogy results in a diversity of complex refractive indices for mineral dust.

42

43 In the atmosphere, mineral dust is additionally often mixed with other aerosols. To face such a complexity, ground and satellite-  
44 based polarization lidar instruments, based on light backscattering by nonspherical particles, have been developed  
45 (Freudenthaler et al., 2009; Tesche et al., 2009; Sugimoto and Lee, 2006; Winker et al., 2009; Miffre et al., 2019) to discern  
46 the mineral dust contribution to two-component particles external mixtures, by applying lidar partitioning algorithms such as  
47 the  $1\beta + 1\delta$  algorithm (Tesche et al., 2009; Mehri et al., 2018). Such lidar-based retrievals are however under-constrained and  
48 depend on prior knowledge regarding input parameters such as the lidar particles' depolarization ratio (*PDR*). The lidar *PDR*  
49 quantifies the mineral dust particles deviation from isotropy and is key for aerosol typing (Hofer et al., 2020). As explained in  
50 light scattering textbooks (Bohren and Huffman, 1983; Mishchenko et al., 2002), it depends on the particles size, shape and  
51 complex refractive index. The size dependence of the lidar *PDR* was studied in field by (Hofer et al., 2020). The downside of  
52 such field measurements is that the observed aerosol is nevertheless that of a particles mixture, which may induce some  
53 discrepancies in the retrieved dust lidar *PDR* (Miffre et al., 2011). As an alternate, for accurate retrievals of the mineral dust  
54 lidar *PDR*, light backscattering numerical simulations have been developed, by assuming a particles shape model such as the  
55 spheroidal shape model, computed with the T-matrix numerical code (Mishchenko and Travis, 1998), as successfully applied  
56 for mineral dust during the SAMUM field campaign (Müller et al., 2013) or, by considering more realistic shapes, based on  
57 stereograms, computed with the discrete-dipole-approximation (Lindqvist et al., 2014). Depending on the assumed shape  
58 model, the lidar *PDR* can be very different with induced variations in the lidar-retrieved dust mass concentrations (Mehri et  
59 al., 2018). Recently, (Luo et al., 2022) discussed on the ability of the spheroidal model to mimic the complex shape of mineral  
60 dust. Likewise, (Zubko et al., 2013) found spheroids inadequate for describing the dust particles' spectral dependence of the  
61 lidar *PDR*. Such light scattering numerical simulations nonetheless rely on simplifying assumptions that should be carefully  
62 checked. Laboratory experiments on natural dust samples at  $180.0^\circ$  lidar exact backscattering angle are then looked-for as they  
63 provide quantitative evaluations of the mineral dust lidar *PDR* within experimental error bars. Indeed, in laboratory, the



64 retrieved lidar *PDR* is, by construction, that of pure mineral dust and the dependence of the dust lidar *PDR* with size and  
65 mineralogy can be evaluated. Moreover, the complex shape of mineral dust is then accounted for. However, existing laboratory  
66 light scattering experimental set-ups (Glen and Brooks, 2013; Järvinen et al., 2016; Gautam et al., 2020; Liu et al., 2020;  
67 Kahnert et al., 2020; Gómez Martín et al., 2021) can only provide approximate values of the dust lidar *PDR* for the following  
68 reasons:

- 69
- 70 – Such apparatuses operate at near backscattering angles only ( $< 180.0^\circ$ ), without covering the exact lidar  
71 backscattering angle of  $180.0^\circ$ . The retrieved lidar *PDR* is then extrapolated to  $180.0^\circ$  following simplifying  
72 numerical assumptions, ignoring the complexity in shape of mineral dust (Liu et al., 2020; Gómez Martín et al., 2021).  
73 To provide accurate values of the dust lidar *PDR*, such assumptions must be carefully discussed as the lidar *PDR*  
74 actually depends on the scattering angle in an unpredictable way, as underscored in light scattering textbooks (Bohren  
75 and Huffman, 1983; Mishchenko et al., 2002), due to the complex shape of mineral dust. For that, a laboratory  
76 measurement of the dust lidar *PDR* at  $180.0^\circ$  is mandatory.
  - 77 – Also, most of the above apparatuses operate at a single wavelength, either 442, 488, 552, 632, 647 or 680 nm, which  
78 differs from the (355, 532, 1064 nm) wavelengths which are applied in polarization lidar field experiments. As for  
79 Raman lidars, such wavelength extrapolations up to the (355, 532, 1064 nm) lidar wavelengths are a source of  
80 discrepancy as the dust lidar *PDR* actually depends on the complex refractive index, which is wavelength dependent  
81 (Bohren and Huffman, 1983; Mishchenko et al., 2002). For that, a laboratory measurement at the lidar wavelengths  
82 is mandatory.

83

84 In this paper, accurate values of the dust lidar *PDR* are provided from a laboratory  $\pi$ -polarimeter operating at  $180.0^\circ$  lidar exact  
85 backscattering angle and at 355, 532 nm wavelength, to account for the importance of the spectral dependence of the lidar  
86 *PDR* to better constrain lidar inversions and aerosol typing (Burton et al., 2016; Haarig et al., 2022). Since the scattering angle  
87 and the wavelengths are determined from lidar purposes, we here investigate the dependence of the mineral dust lidar *PDR* on  
88 the dust particles size and complex refractive index (*CRI*), the latter being particularly important as related to light absorption.  
89 Light absorption by mineral dust preferentially occurs in the UV and VIS spectral domains, being nearly null in the near-  
90 infrared spectral range (Di Biagio et al., 2019), noticeably in the presence of iron oxides (Formenti et al., 2014; Caponi et al.,  
91 2017). By absorbing short-wave radiations, such oxides hence play a critical role in determining the overall impact of dust  
92 aerosol on climate forcing (Go et al., 2022). We hence focused on 355 and 532 nm lidar wavelengths and considered four dust  
93 samples differing in their *CRI*, thus in mineralogy: i) silica oxide ( $\text{SiO}_2$ ), as the most abundant mineral oxide present in mineral  
94 dust, ii) iron oxide (hematite,  $\text{Fe}_2\text{O}_3$ ), as the main light absorbent present in mineral dust (Gautam et al., 2020; Zong et al.,  
95 2021; Go et al., 2022), iii) and iv) two heterogeneous mixtures of the above two oxides in various proportions, as detailed in  
96 Section 2. The dependence of the lidar *PDR* with size is then likewise investigating by accounting for the fine and coarse



97 modes of the particles size distribution ( $SD$ ), to which lidar instruments are sensitive (Mamouri and Ansmann, 2017), thus  
98 extending the size range of our previous laboratory findings (Miffre et al., 2016) to particles sizes larger than 800 nm and to  
99 other mineralogy, as asked for in (Tesche et al., 2019). Our work provides sixteen laboratory-derived accurate dust lidar  $PDR$   
100 values, corresponding to four mineral dust samples differing in mineralogy, given at two  $SD$  (fine, coarse) and at two  
101 wavelengths (355, 532 nm). Moreover, the role of the imaginary part of the hematite  $CRI$ , which may lead to modifications  
102 in the lidar  $PDR$ , is here for the first time quantified and discussed. The paper is structured as follows. In Section 2, the complex  
103 refractive indices and size distributions of our four dust samples are presented. The laboratory  $\pi$ -polarimeter at  $180.0^\circ$  lidar  
104 backscattering angle is then presented in Section 3, together with the dust lidar  $PDR$  retrieval methodology, derived from the  
105 scattering matrix formalism (Mishchenko et al., 2002). The main findings are outlined in Section 4 where the sixteen values  
106 of dust lidar  $PDR$  are given and a discussion is proposed to investigate the dependence of the dust lidar  $PDR$  on the imaginary  
107 part of the dust  $CRI$ . As in elastic lidar applications, we here consider the elastic backscattering of an electromagnetic radiation  
108 of wavelength  $\lambda$  by an ensemble of mineral dust particles of complex refractive index  $m = n + i\kappa$  embedded in ambient air.

## 109 2. Mineral dust samples

### 110 2.1 Refractive indices

111 Mineral dust is a complex mixture of several chemical oxides presenting various complex refractive indices. To investigate  
112 the dependence of the dust lidar  $PDR$  on the complex refractive index ( $CRI$ ), we consider the four following case studies :

113

- 114 – Silica, or silicon oxide ( $\text{SiO}_2$ ) is here considered as being the main pure chemical component present in mineral dust.  
115 The silica  $CRI$  as given by (Longtin et al., 1988) is equal to 1.546, hence exhibiting no absorptive component.
- 116 – Iron oxide, or hematite ( $\text{Fe}_2\text{O}_3$ ), is in contrast here selected as being a climatically significant light absorbent in the  
117 shortwave spectral region, that can be transported far from source regions with similar efficiency as black carbon  
118 particles (Lamb et al., 2021). It recently regained in interest with papers specifically dedicated to this constituent  
119 (Gautam et al., 2020; Zong et al., 2021). Hematite is unique among all chemical oxides present in mineral dust due  
120 its strong  $CRI$ . Both  $n$  and  $\kappa$  are large for hematite, with  $\kappa$ -values more than 100 times those of other soil mineral  
121 components at lidar wavelengths. Hence, hematite dominates absorption while other minerals can be considered as  
122 non-absorbing (Go et al., 2022). Reference literature for the hematite  $CRI$  is  $m = 3.102 + 0.0925i$  by (Longtin et  
123 al., 1988). More recently, the real and imaginary part of the hematite  $CRI$  were reviewed by (Go et al., 2022): from  
124 their Figure 1, we conclude that  $m = 2.25 + 0.9i$  at 355 nm wavelength ( $m = 3.10 + 0.6i$  at 532 nm wavelength).
- 125 – Arizona dust is likewise considered as an example of natural mineral dust sample involving a mixture of the above  
126 two oxides. According to the manufacturer (Power Technology Inc.), Arizona Test Dust is composed of silica (68-76  
127 %), while hematite is only weakly present in Arizona dust (2-5 %). In short, Arizona dust is hence rather silica-rich.



128 As given by the manufacturer, the Arizona dust *CRI* is  $m = 1.51 + 10^{-3}i$ , without however any given on its spectral  
129 dependency. Effective medium theories can alternately be applied to account for the sample inhomogeneity as  
130 calculated in (Miffre et al., 2016), who arrived to  $m = 1.57 + 10^{-2}i$  at 355 nm wavelength and  $1.55 + 5.10^{-3}i$  at  
131 532 nm wavelength. As a result, the Arizona dust sample *CRI* is characterized by  $n \sim 1.5$  and a low absorbing  
132 component  $\kappa \sim 5.10^{-3}$ .

133 – Asian dust is finally also considered as an important case study of natural mineral dust sample, presenting however a  
134 lower proportion of silica (34-40 %) and a higher proportion in hematite (17-23 %). Hence, compared with Arizona  
135 dust, Asian dust is more hematite-rich and hence exhibits a larger imaginary part for its *CRI*.

136

137 The solid dust samples, provided by by Sigma Aldrich and Powder Technology manufacturers, were embedded in laboratory  
138 ambient air by using a solid dust generator supplied with dried compressed air ( $RH < 10\%$ ) to get dry solid dust particles  
139 embedded in laboratory ambient air at a constant number concentration, before injecting the dust samples into the light  
140 scattering volume, as presented in Section 3.

## 141 2.2 Size distribution (*SD*)

142 For each above dust sample, to likewise investigate the dependence of the dust lidar *PDR* on the particles size, we consider  
143 two size distributions (*SD*):

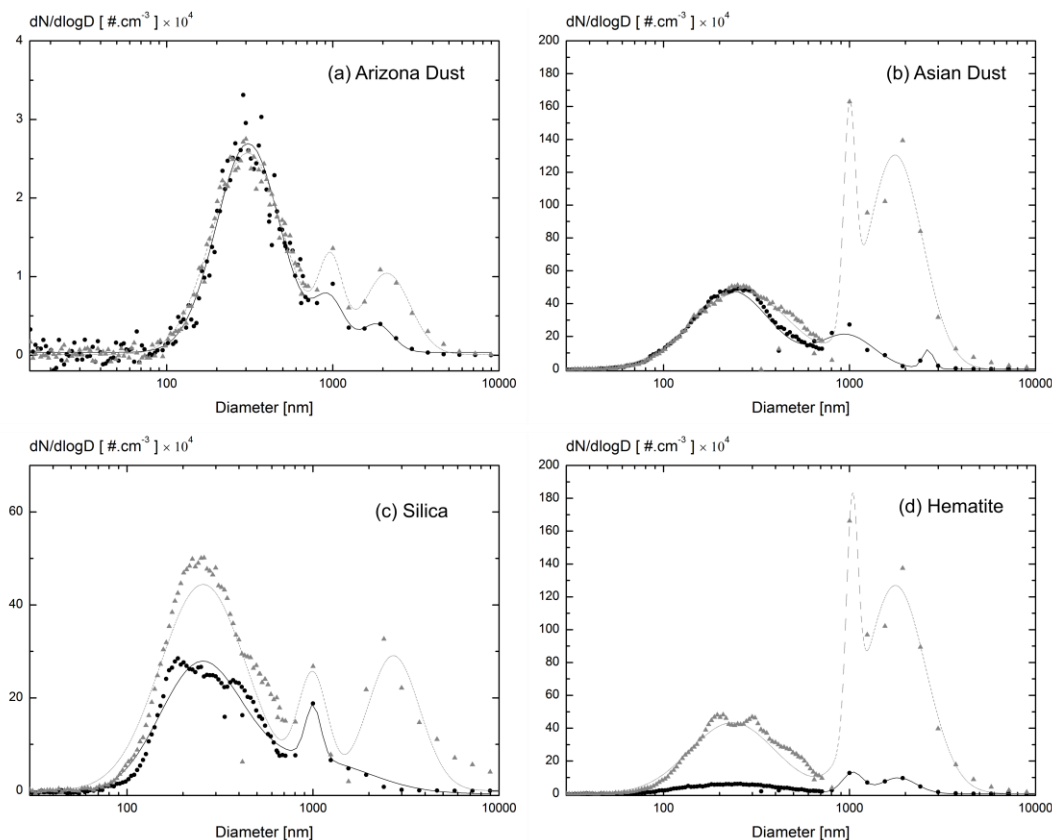
- 144 – The coarser *SD*, represented in grey in Figure 1. This *SD* is aimed at being more representative of mineral dust  
145 particles close to dust source regions,
- 146 – A finer *SD*, represented in black in Figure 1, aimed at being more representative of mineral dust particles after long-  
147 range transport, i.e. farther from the dust source regions.

148 The *SD* were obtained by adding / removing a cyclone to our experimental set-up allowing to add / remove particles with  
149 diameter above 800 nm, thus exploring particles size ranges below and above 800 nm, as asked for in (Tesche et al., 2019).

150 The *SD* were measured with an optical particles sizer (OPS 3330) coupled with a scanning mobility particles sizer (SMPS  
151 3081), which selects the dust particles as a function of their electric mobility, this latter quantity being diameter-dependent.

152 The particles *SD* displayed in Figure 1 are in agreement with the specifications provided by the manufacturers.

153



154

155 Figure 1: Dust particles size distributions ( $SD$ ) for: (a) Arizona dust, (b) Asian dust, (c) Silica ( $\text{SiO}_2$ ), (d) Hematite ( $\text{Fe}_2\text{O}_3$ ) in  
156 the presence / absence of the added cyclone (finer  $SD$ , in solid black) / (coarser  $SD$ , in dotted grey). The retrieved  $SD$ , obtained  
157 by log-normal adjustments, agree with the specifications provided by the manufacturers.

### 158 3 Methodology

159 In this section, we detail our methodology for accurate laboratory evaluations of the dust lidar  $PDR$  at lidar exact  
160 backscattering angle of  $180.0^\circ$  for accurate lidar  $PDR$ -retrievals.

#### 161 3.1 Scattering matrix formalism

162 The dust lidar  $PDR$  can be evaluated in the framework of the scattering matrix formalism, which is the dedicated formalism  
163 for polarization-resolved elastic light scattering measurements, as recommended in light scattering textbooks (Mishchenko et  
164 al., 2002; Bohren and Huffman, 1983). In this formalism, the polarization state of the incident and scattered radiations are  
165 described by their respective Stokes vectors  $\mathbf{St}_i = [I_i, Q_i, U_i, V_i]^T$  and  $\mathbf{St} = [I, Q, U, V]^T$ , defined with respect to the scattering  
166 plane, used as a reference plane (Mishchenko et al., 2002). The first Stokes component  $I$  corresponds to the light intensity,  $Q$   
167 and  $U$  describe linear polarization, while  $V$  accounts for circular polarization. At a distance  $d$  from the mineral dust samples,



168 if single-scattering and particles random orientation are assumed, the incident and scattered Stokes vectors relate with a bloc-  
 169 diagonal scattering matrix (Mishchenko et al., 2002; Bohren and Huffman, 1983):

170

$$171 \begin{pmatrix} I \\ Q \\ U \\ V \end{pmatrix} = \frac{1}{k^2 d^2} \begin{bmatrix} F_{11,\lambda}(\theta) & F_{12,\lambda}(\theta) & 0 & 0 \\ F_{12,\lambda}(\theta) & F_{22,\lambda}(\theta) & 0 & 0 \\ 0 & 0 & F_{33,\lambda}(\theta) & F_{34,\lambda}(\theta) \\ 0 & 0 & -F_{34,\lambda}(\theta) & F_{44,\lambda}(\theta) \end{bmatrix} \begin{pmatrix} I_i \\ Q_i \\ U_i \\ V_i \end{pmatrix} \quad (1)$$

172

173 Where the matrix elements  $F_{ij,\lambda}(\theta)$  ( $i, j = 1 - 4$ ) depend on the wavelength  $\lambda$  of the radiation (hereafter noted as a subscript)  
 174 and comprise the information on the mineral dust particles size, shape and *CRI*. The scattering angle is  $\theta = (\mathbf{k}_i, \mathbf{k})$ , where  
 175  $k = k_i = 2\pi/\lambda$  is the wave vector of the radiation. In lidar applications, the scattering angle is equal to  $\pi$  (i.e. exact  
 176 backscattering angle). To highlight the need for laboratory measurements at the specific  $180.0^\circ$  lidar backscattering angle, near  
 177 backscattering angles (i.e.  $\theta < \pi$ ) are also considered in this section. Indeed, at specific lidar backscattering angle ( $\theta = \pi$ ),  
 178  $F_{33,\lambda} = -F_{22,\lambda}$  and  $F_{12,\lambda} = F_{34,\lambda} = 0$  (Zubko et al., 2013; David et al., 2013) while  $F_{44,\lambda} = F_{11,\lambda} - 2F_{22,\lambda}$  due to the  
 179 backscattering theorem (van de Hulst, 1957), so that Eq. (1) simplifies as follows for lidar applications:

180

$$181 \begin{pmatrix} I \\ Q \\ U \\ V \end{pmatrix} = \frac{1}{k^2 d^2} \begin{bmatrix} F_{11,\lambda}(\pi) & 0 & 0 & 0 \\ 0 & F_{22,\lambda}(\pi) & 0 & 0 \\ 0 & 0 & -F_{22,\lambda}(\pi) & 0 \\ 0 & 0 & 0 & F_{11,\lambda}(\pi) - 2F_{22,\lambda}(\pi) \end{bmatrix} \begin{pmatrix} I_i \\ Q_i \\ U_i \\ V_i \end{pmatrix} \quad (2)$$

182

183 As a result, it is only at elastic lidar exact backscattering angle ( $\theta = \pi$ ) that  $F_{12,\lambda} = 0$  so that the scattering matrix reduces to  
 184 only two non-vanishing elements  $F_{11,\lambda}(\pi)$  and  $F_{22,\lambda}(\pi)$ .

### 185 3.2 Lidar particles depolarization ratio *PDR*

186 The expression of the so-called particles depolarization ratio (*PDR*) at wavelength  $\lambda$  and scattering angle  $\theta$  can be found in  
 187 light scattering textbooks (Mishchenko et al., 2002; Schnaiter et al., 2012):

188

$$189 PDR_\lambda(\theta) = \frac{1 - F_{22,\lambda}(\theta)/F_{11,\lambda}(\theta)}{1 \pm 2F_{12,\lambda}(\theta)/F_{11,\lambda}(\theta) + F_{22,\lambda}(\theta)/F_{11,\lambda}(\theta)} \quad (3)$$

190

191 where the positive (resp. negative) sign corresponds to *p*-polarized (resp. *s*-polarized) incident electromagnetic radiation. For  
 192  $F_{11,\lambda}$ ,  $F_{12,\lambda}$  and  $F_{22,\lambda}$  vary with the scattering angle  $\theta$ , so does the dust *PDR*. Since  $F_{11,\lambda}$ ,  $F_{12,\lambda}$  and  $F_{22,\lambda}$  may vary with the  
 193 scattering angle, depending on the dust sample, the dust *PDR* at near backscattering angles ( $\theta < \pi$ ) differs from that obtained  
 194 at specific lidar backscattering angle ( $\theta = \pi$ ). The deviation of  $F_{11,\lambda}$ ,  $F_{12,\lambda}$  and  $F_{22,\lambda}$  from their value at exact backscattering



195 angle cannot be quantified since no analytical light scattering theory exists for such complex-shaped particles as mineral dust.  
196 Therefore, a laboratory experiment at specific lidar exact backscattering angle ( $\theta = \pi$ ) is required for precise evaluations of  
197 the dust lidar *PDR*. At specific lidar backscattering angle of  $\pi$ , Eq. (3) becomes:

$$199 \quad PDR_{\lambda}(\pi) = \frac{1 - F_{22,\lambda}(\pi)/F_{11,\lambda}(\pi)}{1 + F_{22,\lambda}(\pi)/F_{11,\lambda}(\pi)} \quad (4)$$

200

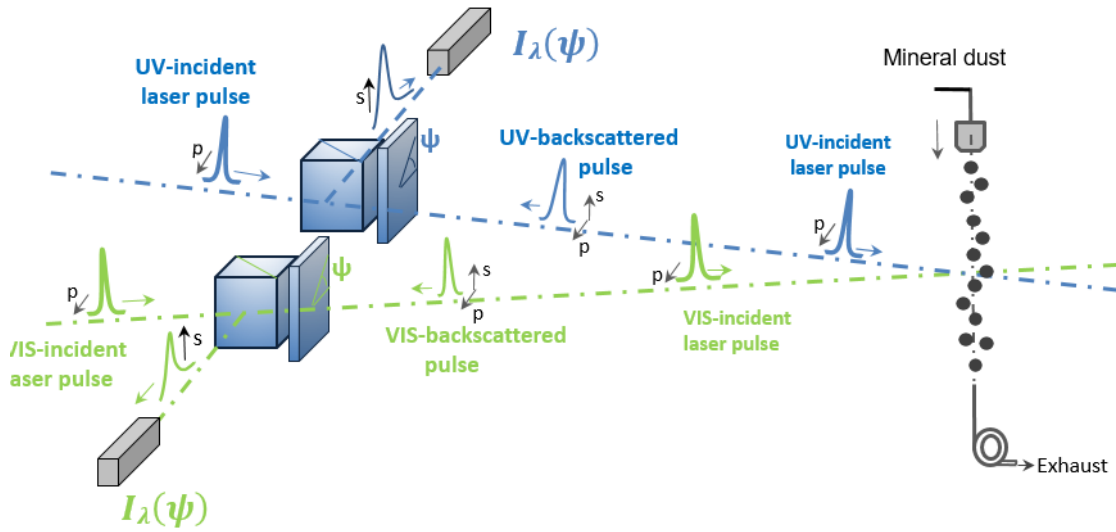
201 Hence and as a result, accurate evaluations of the dust lidar *PDR* rely on accurate determinations of the ratio  $F_{22,\lambda}/F_{11,\lambda}$  at  
202 specific lidar  $\pi$  – angle. As for the ratio  $F_{22,\lambda}/F_{11,\lambda}$ , the dust lidar *PDR* is size, shape and refractive index dependent and this  
203 dependency is discussed in Section 4. Spherical particles, for which  $F_{22,\lambda}/F_{11,\lambda} = 1$ , lead to zero depolarization. In what  
204 follows, to ease the reading, the dust lidar *PDR* will be noted  $PDR_{\lambda}$  without reference to scattering angle ( $\theta = \pi$ ).

### 205 3.3 Laboratory $\pi$ -polarimeter for retrieving the lidar *PDR* of mineral dust

206 In (Miffre et al., 2016), for the first time to our knowledge, a laboratory  $\pi$ -polarimeter was built to address light  
207 backscattering by aerosol particles. We here recall its main characteristics for clarity. The aerosols  $\pi$ -polarimeter is schemed  
208 in Figure 2. As in lidar applications, pulsed laser light is used to measure the time-of-flight taken by a laser pulse to reach the  
209 dust sample and be detected after light backscattering. The backscattering geometry is set by inserting a specified polarization  
210 beam splitter cube (*PBC*) on the way from the laser pulse to the dust samples, with a precision of 1 mm out of 10 meters to  
211 ensure the  $\pi$ -polarimeter to cover the lidar exact backscattering direction with accuracy:  $\theta = (180.0 \pm 0.2)^{\circ}$ . The laboratory  
212 aerosol Pi-polarimeter is actually composed of two identical polarimeters, one per wavelength, to evaluate the lidar *PDR* of a  
213 given dust sample at 355 and 532 nm wavelength simultaneously. Moreover, to gain in accuracy in the lidar *PDR*-retrieval,  
214 the polarization state of the backscattered radiation is analysed for a set of incident polarization states, obtained by modulating  
215 the incident polarization state with a quarter-wave plate (*QWP*). To validate the laboratory  $\pi$ -polarimeter, we carefully checked  
216 that homogeneous spherical particles, such as ammonium sulfate particles, which follow the Mie theory, were indeed providing  
217 zero lidar *PDR* when following the methodology described in the below section.

218





219

220 Figure 2: Scheme of the laboratory  $\pi$ -polarimeter operating at lidar exact backscattering angle of  $(180.0 \pm 0.2)^\circ$  allowing accurate  
 221 retrievals of the lidar  $PDR$  of an aerosol at 355 and 532 nm wavelength simultaneously (Mishchenko et al., 2002). The  $(p, s)$   
 222 polarization components are defined with respect to the laser scattering plane and  $\psi$  is the angle between the fast axis of the  $QWP$  and the  
 223 laser scattering plane, counted counter-clockwise for an observer looking from the  $PBC$  to the particles. The dust lidar  $PDR$  is then evaluated  
 224 from the ratio  $F_{22,\lambda}/F_{11,\lambda}$  at specific  $\pi$ -angle, following the methodology described in Section 3.4.

### 225 3.4 Laboratory retrievals of mineral dust lidar $PDR$

226 Interestingly, the laboratory  $\pi$ -polarimeter can be described in the framework of the scattering matrix formalism  
 227 (Mishchenko et al., 2002). Hence, to retrieve the dust  $PDR$ , we account for the successive Mueller matrices encountered by  
 228 the laser pulse from the laser source to the dust particles sample then back to the light detector to get the expression of the the  
 229 detected backscattered intensity

230

$$231 \quad I_\lambda(\psi) = \frac{\eta_\lambda P_\lambda}{d^2} [1, 0, 0, 0]^T [\mathbf{PBC}] [\mathbf{QWP}(-\psi)] [\mathbf{F}_\lambda] [\mathbf{QWP}(\psi)] [\mathbf{PBC}] (St_i) \quad (5)$$

232

233 Where  $\eta_\lambda$  is the optoelectronics efficiency of our light detector and  $P_\lambda$  is the laser power density, while  $(St_i) =$   
 234  $[1, 1, 0, 0]^T$  sets the laser incident polarization state. The expression of the dust backscattering matrix  $[\mathbf{F}_\lambda]$  at wavelength  $\lambda$  is  
 235 is given in Eq. (2), while  $[\mathbf{PBC}]$  and  $[\mathbf{QWP}(\pm\psi)]$  are the Mueller matrices of the  $PBC$  and the  $QWP$  respectively (Shurcliff,  
 236 1962). To develop Eq. (5), it is then advised to first calculate the raw vector  $[1, 0, 0, 0]^T [\mathbf{PBC}] [\mathbf{QWP}(-\psi)] [\mathbf{F}_\lambda]$  then multiply  
 237 it with the Stokes vector of the radiation incident  $[\mathbf{QWP}(\psi)] [\mathbf{PBC}] (St_i)$  equal to  $[1, \cos^2(2\psi), -\sin(4\psi)/2, -\sin(2\psi)]^T$ , if  
 238  $\psi$  is the modulation angle of the  $QWP$ . After a few calculations, the dust backscattered light intensity  $I_\lambda$  at wavelength  $\lambda$  finally  
 239 expresses as follows:

240

$$241 \quad I_\lambda(\psi) = I_{\lambda,0} \times [a_\lambda - b_\lambda \cos(4\psi)] \quad (6)$$



242

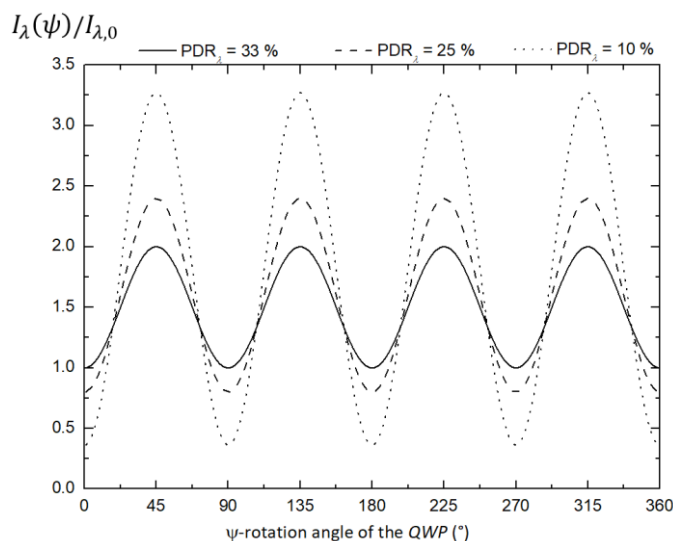
243 where the intensity  $I_{\lambda,0} = \eta_{\lambda} P_{\lambda} / (4d^2)$ , while coefficients  $a_{\lambda}$  and  $b_{\lambda}$  are equal to  $a_{\lambda} = F_{11,\lambda} + F_{22,\lambda}$  and  $b_{\lambda} = 3F_{22,\lambda} - F_{11,\lambda}$ .  
 244 Hence,  $F_{22,\lambda} / F_{11,\lambda} = (1 + b_{\lambda} / a_{\lambda}) / (3 - b_{\lambda} / a_{\lambda})$  so that the ratio  $F_{22,\lambda} / F_{11,\lambda}$  at  $\pi$ -angle can be determined from the ratio  
 245  $b_{\lambda} / a_{\lambda}$ . This ratio can be obtained by recording the variations of  $I_{\lambda}(\psi)$ , then adjusting these variations with Eq. (6) to get  
 246 accurate determinations of  $I_{\lambda,0} a_{\lambda}$  and  $I_{\lambda,0} b_{\lambda}$ , then  $b_{\lambda} / a_{\lambda}$ . Accurate evaluations of the dust lidar  $PDR$  are then finally retrieved  
 247 from Eq. (4):

248

$$249 \quad PDR_{\lambda} = (1 - b_{\lambda} / a_{\lambda}) / 2 \quad (7)$$

250

251 Within our methodology, the dust lidar  $PDR$  is independent of  $I_{\lambda,0}$ . For that reason, in Section 4, the applied voltage to the UV  
 252 and VIS-photodetectors will be adjusted to each dust  $SD$  and mineralogy to gain in accuracy in the retrieved dust lidar  $PDR$   
 253 by improving the signal-to-noise ratio on  $I_{\lambda}$ . To fix ideas, we numerically simulated in Figure 3 the variations of  $I_{\lambda}(\psi) / I_{\lambda,0}$  for  
 254 the three following dust lidar  $PDR$  case studies : 33 % dust lidar  $PDR$  (in full lines, i.e.  $F_{22,\lambda} / F_{11,\lambda} = 0.5$ ), 25 % dust lidar  
 255  $PDR$  (in dashed-lines, i.e.  $F_{22,\lambda} / F_{11,\lambda} = 0.6$ ), 10 % dust lidar  $PDR$  (in dotted lines, i.e.  $F_{22,\lambda} / F_{11,\lambda} = 0.82$ ). The curve  
 256 minima, which are equal to  $I_{\lambda,m} / I_{\lambda,0} = a_{\lambda} - b_{\lambda} = F_{11,\lambda} - F_{22,\lambda}$ , are shape-dependent : each curve hence exhibits non-  
 257 vanishing minima since mineral dust particles are nonspherical. Likewise, the curve maxima are equal to  $I_{\lambda,M} / I_{\lambda,0} = a_{\lambda} +$   
 258  $b_{\lambda} = 2F_{22,\lambda}$  and are size-dependent. The dust lidar  $PDR$  is determined from  $I_{\lambda,m}$  and  $I_{\lambda,M}$  since, following Eq. (7),  
 259  $PDR_{\lambda} = I_{\lambda,m} / (I_{\lambda,m} + I_{\lambda,M})$ , independently of  $I_{\lambda,0}$ .



260

261 Figure 3: Numerical simulation of the dust backscattered light intensity  $I_{\lambda}(\psi) / I_{\lambda,0}$  as a function of the orientation  $\psi$  of the  $QWP$  at a given  
 262 wavelength at the three following case studies:  $PDR_{\lambda} = 33\%$  (in full lines, corresponding to  $F_{22,\lambda} / F_{11,\lambda} = 0.50$ ),  $PDR_{\lambda} = 25\%$  (in  
 263 dashed-lines,  $F_{22,\lambda} / F_{11,\lambda} = 0.60$ ),  $PDR_{\lambda} = 10\%$  (in dotted lines,  $F_{22,\lambda} / F_{11,\lambda} = 0.82$ ).



### 264 3.5 Accuracy on the retrieved laboratory mineral dust lidar PDR

265 Special care has been taken to quantify the uncertainties on the retrieved dust lidar *PDR*. The systematic errors in the  $\pi$ -  
266 polarimeter are that encountered in  $2\lambda$ -polarization lidar experiments, which we extensively studied (David et al., 2012) and  
267 can also be found in polarization lidar reference papers (Freudenthaler et al., 2009). To summarize, systematic errors arise  
268 from:

- 269 • *Imperfect definition of the polarization state of the incident radiation.* In the  $\pi$ -polarimeter, the polarization state of  
270 the electromagnetic radiation emerging from the laser is precisely set to  $[1, 1, 0, 0]^T$  (i.e. with no remaining ellipticity)  
271 by using two successive *PBC*.
- 272 • *Polarization cross-talks between the emitter and the detector polarization axes.* Likewise, on the detector side, to  
273 account for the imperfections of the retro-reflecting *PBC* ( $R_s > 99.5\%$ ,  $T_p > 90\%$ ), a secondary *PBC* is inserted  
274 between the retro-reflecting *PBC* and the light detector to ensure polarization cross-talk or undesired fraction  $R_p T_s$   
275 originating from the *p*-component of the backscattered radiation to be fully negligible. Hence, the  $\pi$ -polarimeter is  
276 sensitive to the *s*-component of the backscattered radiation only. Also, the emitting *PBC* being used as retro-reflecting  
277 *PBC*, any possible mismatch between the *s*-polarization axis of the emitted and detected backscattered radiations  
278 cannot occur.
- 279 • *Spectral cross-talks between the UV and the VIS-backscattered radiations.* Likewise, wavelength cross-talks are  
280 minimized by using selective interference filters exhibiting a higher than 5 optical density, at 355 nm wavelength in  
281 the VIS  $\pi$ -polarimeter and at 532 nm wavelength in the UV  $\pi$ -polarimeter.
- 282 • *Multiple scattering can induce further light depolarization.* However, the single-scattering approximation is rather  
283 safe in our laboratory backscattering experiment (Mishchenko et al., 2007) where the particles are moving in a thin  
284 (2.5 mm) wide beam so that the volume element is optically thin in contrary to atmospheric chambers.

285

286 Finally, to account for potential fluctuations in the dust particle number concentration that may cause variations in the dust  
287 backscattered light intensity  $I_\lambda$ , a normalization channel has been added to our experiment by considering a polarization-  
288 insensitive light detector operating at scattering angle  $\theta_0 = 165^\circ$ . The corresponding scattered light intensity  $I_\lambda(\theta_0)$  can be  
289 retrieved from the Mueller matrices successively encountered by the laser pulse at scattering angle  $\theta_0$  (there, the *QWP* and the  
290 *PBC* only act on the detector side while  $(St_i)$  equals  $[1, 1, 0, 0]^T$ ) to get  $I_\lambda(\theta_0) = I_{\lambda,0} \times [2F_{11,\lambda}(\theta_0) + F_{12,\lambda}(\theta_0) +$   
291  $F_{12,\lambda}(\theta_0)\cos(4\psi)]$ . Once the variations of  $I_\lambda(\theta_0)$  with  $\psi$ -angle are recorded, the  $\cos(4\psi)$ -dependency of  $I_\lambda(\theta_0)$  can be  
292 removed by applying a numerical low-pass filter on  $I_\lambda(\theta_0)$ , to get a light intensity proportional to the dust particles number  
293 concentration. As a result, in the light backscattering curves presented in Section 4, the plotted quantity is the normalized  
294 backscattered light intensity  $I_{\lambda,N} = I_\lambda(\pi)/I_\lambda(\theta_0)$ , which is insensitive to potential fluctuations in the dust particles number  
295 concentration. The scattered light intensities  $I_\lambda(\pi)$  and  $I_\lambda(\theta_0)$  being correlated, the standard deviation  $\sigma_N$  on  $I_{\lambda,N}$  was



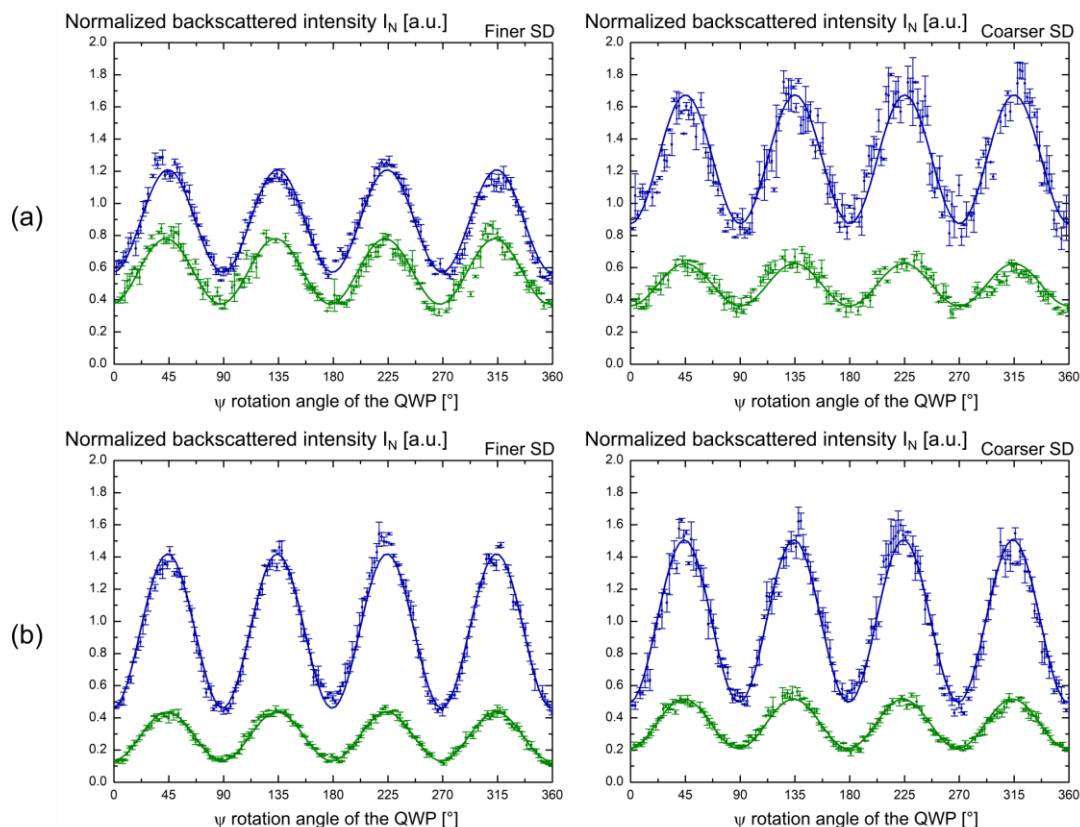
296 calculated by considering the covariance of  $I_\lambda$  and  $I_\lambda(\theta_0)$ . Moreover, to gain in accuracy in the dust lidar *PDR* retrievals,  
297  $I_{\lambda,N}$  was measured for a complete  $\psi$ -angle rotation, while averaging the acquired backscattered light intensity over several  
298 thousand laser shots per  $\psi$ -angle, with resulting mean and standard deviations on  $I_{\lambda,N}$  as plotted in Figures 4 and 5.

#### 299 4. Results and discussion

300 In this section, using the methodology presented in Section 3, the lidar *PDR* of Arizona dust, Asian dust, silica and hematite  
301 is evaluated and discussed at 355 and 532 nm wavelength for the finer and the coarser *SD*.

##### 302 4.1 Laboratory evaluation of the lidar *PDR* of Arizona and Asian dust

303 Figure 4 displays the variations of  $I_{\lambda,N}$  for Arizona (Fig. 4a) and Asian dust (Fig. 4b) as a function of the  $\psi$ -rotation angle  
304 of the *QWP* for the finer (left panels) and the coarser *SD* (right panels) at 355 and 532 nm wavelength. The observed variations  
305 are related to a determined size and shape distribution of the dust sample: indeed, as explained in Section 3.4, if the size (resp.  
306 the shape) of the dust sample was varying during our acquisitions, the maxima (resp. the minima) of the curves would not  
307 remain constant. As a result, the observed variations of  $I_{\lambda,N}$  reveal the spectral and polarimetric backscattering characteristics  
308 of each considered dust sample. Therefore, the experimental data points could be adjusted with Eq. (6) to evaluate  $F_{22,\lambda}/F_{11,\lambda}$   
309 then the dust lidar *PDR* by applying Eq. (7). Table 1 presents the retrieved values of  $F_{22,\lambda}/F_{11,\lambda}$  and of dust lidar *PDR*. The  
310 accuracy on  $F_{22,\lambda}/F_{11,\lambda}$  results from the accuracy of the laboratory  $\pi$ -polarimeter and leads to accurate evaluations of the dust  
311 lidar *PDR*. Within experimental error bars, the lidar *PDR* of Arizona and Asian dust clearly differ, whatever the chosen  
312 wavelength. The generally admitted value of around 33 % for the dust lidar *PDR* (Teschke et al., 2009) is only obtained for  
313 Arizona dust: Asian dust exhibits a lower *PDR* in the range from 24 to 28 % depending on the considered *SD* and wavelength.  
314 This suggests that the dust lidar *PDR* is primarily governed by the dust mineralogy and hence particles refractive index. The  
315 sensitivity of the dust lidar *PDR* with the considered *SD* is indeed less pronounced: from the coarser to the finer *SD*, a reduction  
316 in the dust lidar *PDR* of below 5 % is observed at 532 nm wavelength. At 355 nm wavelength however, the Arizona and Asian  
317 dust lidar *PDR* seem practically insensitive to variations in the considered *SD*.



318

319 Figure 4: Normalized backscattered light intensity  $I_{\lambda,N}$  by Arizona (a) and Asian dust (b) for the finer  $SD$  (left panels) and the coarser  $SD$   
 320 (right panels) using the laboratory  $\pi$ -polarimeter at lidar exact backscattering angle ( $\theta = \pi$ ) at 355 (blue) and 532 (green) nm wavelength.  
 321 The experimental data points are adjusted with Eq. (6) to evaluate  $F_{22,\lambda}/F_{11,\lambda}$  then the dust lidar  $PDR$  by applying Eq. (7). Care should be  
 322 taken when comparing  $I_{\lambda,N}$  for Arizona and Asian dust for the applied voltage to the UV and VIS-photodetectors were adjusted to increase  
 323 the signal-to-noise ratio, as explained in Section 3.4. Hence, the Arizona dust lidar  $PDR$ , retrieved from  $I_{\lambda,m}/(I_{\lambda,m} + I_{\lambda,M})$ , is higher than  
 324 that of Asian dust.

325 Tab. 1: Laboratory measurement of the lidar  $PDR$  of Arizona and Asian dust at 355 (blue) and 532 nm (green) wavelength for the finer and  
 326 the coarser  $SD$ . The lidar  $PDR$  is retrieved by applying Eq. (7) after accurate evaluations of  $F_{22,\lambda}/F_{11,\lambda}$  obtained with the laboratory  $\pi$ -  
 327 polarimeter (Miffre et al., 2016) presented in Section 3.2.

328

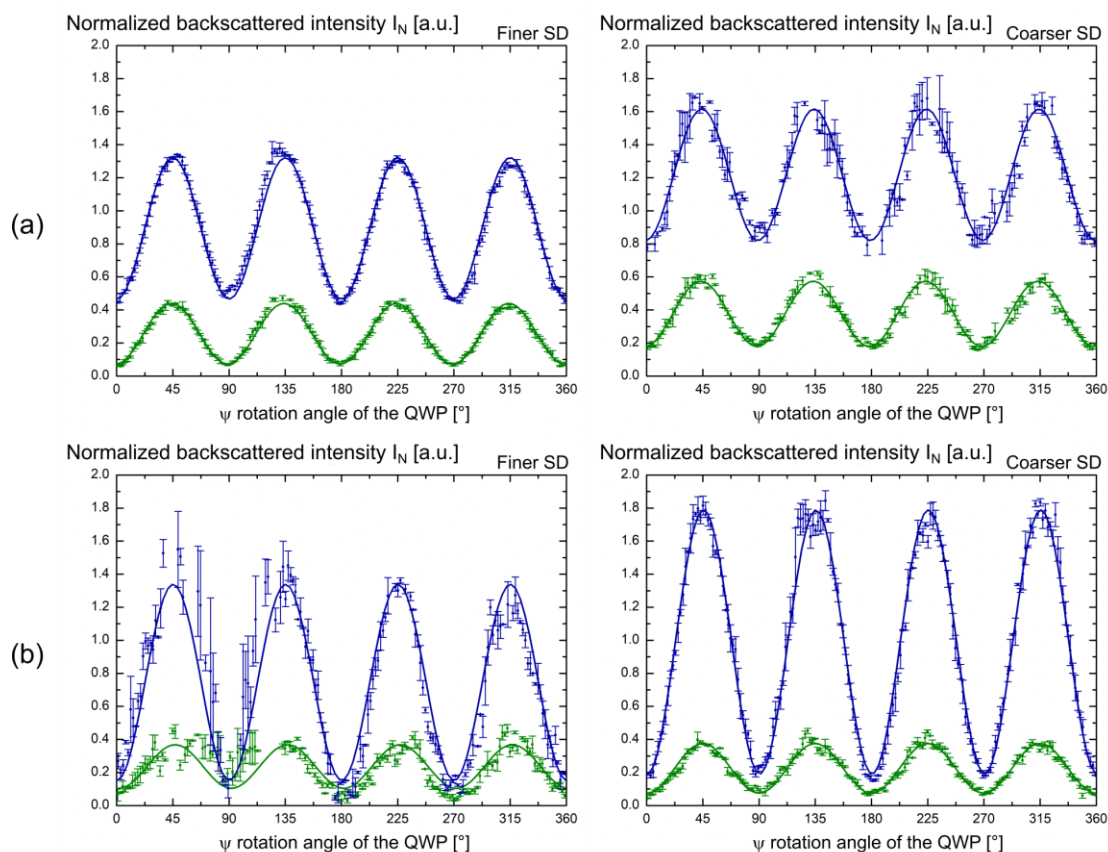
Mineralogy	$\lambda$ (nm)	Finer $SD$		Coarser $SD$	
		$F_{22,\lambda}/F_{11,\lambda}$	$PDR_{\lambda}$ (%)	$F_{22,\lambda}/F_{11,\lambda}$	$PDR_{\lambda}$ (%)
Arizona dust	355	$0.514 \pm 0.007$	<b><math>32.1 \pm 0.6</math></b>	$0.489 \pm 0.012$	<b><math>34.3 \pm 1.0</math></b>
	532	$0.512 \pm 0.012$	<b><math>32.3 \pm 1.0</math></b>	$0.464 \pm 0.012$	<b><math>36.6 \pm 1.1</math></b>
Asian dust	355	$0.603 \pm 0.009$	<b><math>24.7 \pm 0.6</math></b>	$0.603 \pm 0.011$	<b><math>24.8 \pm 0.8</math></b>
	532	$0.622 \pm 0.009$	<b><math>23.3 \pm 0.7</math></b>	$0.558 \pm 0.011$	<b><math>28.4 \pm 0.8</math></b>

329



#### 330 4.2 Laboratory evaluation of the lidar *PDR* of silica and hematite

331 By applying the same methodology, the lidar *PDR* of silica and hematite were obtained, as presented in Table 2 which is  
332 the analogue of Table 1 for silica and hematite. Accordingly, Figure 5 is the analog of Figure 4 for silica (Fig. 5a) and hematite  
333 (Fig. 5b). As for Arizona and Asian dust samples, the lidar *PDR* of silica and hematite primarily depends on the particles *CRI*,  
334 at least at 355 nm wavelength where the silica lidar *PDR* ranges from 23 to 33 % depending on the considered *SD* while the  
335 hematite lidar *PDR* reaches 10 % only. The silica lidar *PDR* also strongly depends on the *D* : from the coarser to the finer *SD*,  
336 the silica dust lidar *PDR* reduces by 10 % at both wavelengths. The dependence of the hematite dust lidar *PDR* with the *SD* is  
337 less pronounced, especially at 355 nm wavelength. The silica and hematite lidar *PDR* also strongly depend on the chosen lidar  
338 wavelength, with higher depolarization observed at 355 nm wavelength for silica and at 532 nm wavelength for hematite.



339  
340 Figure 5: Same as Figure 4 for silica (a-plots) and for hematite (b-plots).

341  
342



343 Tab. 2: Same as Table 1 for silica and hematite.

344

Mineralogy	$\lambda$ (nm)	Finer <i>SD</i>		Coarser <i>SD</i>	
		$F_{22,\lambda}/F_{11,\lambda}$	$PDR_{\lambda}$ (%)	$F_{22,\lambda}/F_{11,\lambda}$	$PDR_{\lambda}$ (%)
Silica	355	$0.622 \pm 0.014$	<b><math>23.3 \pm 0.9</math></b>	$0.506 \pm 0.011$	<b><math>32.8 \pm 1.0</math></b>
	532	$0.751 \pm 0.016$	<b><math>14.2 \pm 0.9</math></b>	$0.618 \pm 0.016$	<b><math>23.6 \pm 1.1</math></b>
Hematite	355	$0.805 \pm 0.050$	<b><math>10.8 \pm 2.5</math></b>	$0.823 \pm 0.015$	<b><math>9.7 \pm 0.7</math></b>
	532	$0.652 \pm 0.055$	<b><math>21.1 \pm 3.5</math></b>	$0.715 \pm 0.019$	<b><math>16.6 \pm 1.1</math></b>

345

### 346 4.3 Discussion

347 Comparing our laboratory findings with other laboratory experiments is not straightforward, since as explained in the  
 348 introduction, none operates at 180.0° lidar exact backscattering angle, while the dust lidar *PDR* differs from near to exact  
 349 backscattering angles. Otherwise, the lidar *PDR* is wavelength-dependent and the size distributions (*SD*) are different. Lidar  
 350 field experiments provide accurate values of the lidar *PDR* after accurate calibration procedure based on the scattering matrix  
 351 (Freudenthaler, 2016; Belegante et al., 2018; Miffre et al., 2019). Though in such lidar field experiments, the measured *PDR*  
 352 is nevertheless that of particle mixtures (Miffre et al., 2011), comparison with our laboratory findings remains interesting as a  
 353 complement. In lidar retrievals (see for example (Tesche et al., 2009)), a dust lidar *PDR* in the range of 30 % is often  
 354 considered. The laboratory  $\pi$ -polarimeter validates that statement by providing the silica lidar *PDR*, which is the main oxide  
 355 present in mineral dust, equal to  $(33 \pm 1)$  % for the coarser *SD* at 355 nm. In comparison, within our experimental error bars,  
 356 the hematite lidar *PDR*, equal to  $(10 \pm 1)$  %, is clearly lower. The real part  $n$  and the imaginary part  $\kappa$  of the hematite *CRI*,  
 357 which are large compared with that of other chemical oxides present in mineral dust (see Section 2.1), can be responsible for  
 358 the observed difference in the silica and hematite lidar *PDR*. Indeed,  $n$  and part  $\kappa$  modify the backscattering matrix elements,  
 359 so does the corresponding dust lidar *PDR*. To highlight the role of  $\kappa$  on the hematite lidar *PDR*, the lidar *PDR* of rutile was  
 360 measured with our  $\pi$ -polarimeter. Indeed, the real part of the rutile *CRI* is as large as that of hematite but its imaginary part is  
 361 negligible compared with that of hematite. As a result, the rutile lidar *PDR* substantially differed from that of hematite,  
 362 showing the key role played by light absorption in the measured hematite lidar *PDR*. In turn, Arizona dust exhibits a higher  
 363 *PDR* than Asian dust, due to the higher proportion in hematite in the latter. Hence and as a conclusion, our laboratory findings  
 364 show that, when the light absorbent hematite is present, it mainly governs the dust lidar *PDR*, which hence primarily depends  
 365 on the particles mineralogy, with less pronounced variations with the particles size and wavelength. This finding is in line with  
 366 (Kahnert, 2015; Kahnert et al., 2020) numerical findings, who highlighted that the dust *PDR* is strongly modulated by the  
 367 particles inhomogeneity, especially in the lidar backward scattering direction and in the presence of hematite. We here quantify  
 368 this effect with a laboratory experiment that accounts for the real shape of mineral dust. When the proportion of hematite  
 369 becomes negligible, as is the case for silica and Arizona dust, our laboratory findings show that the dust lidar *PDR* then



370 increases with increasing the particles size. Likewise, in the literature (Hofer et al., 2020; Järvinen et al., 2016; Mamouri and  
371 Ansmann, 2017), the dust lidar *PDR* is usually found to increase with the particles size from the fine to the coarse mode of the  
372 *SD*. The (355, 532) nm wavelength dependence of the dust lidar *PDR* then becomes key for discussing on the involved particle  
373 sizes, thus underlying the importance of dual-wavelength (or more) polarization lidar instruments. We here establish this result  
374 in laboratory at 180.0° and (355, 532) nm wavelength, and moreover, show that this consideration holds only when hematite,  
375 which is a strong light absorbent, is not involved : the hematite lidar *PDR* is indeed higher in the finer mode of the *SD*.

376

377 To go further and discuss on the role of light absorption in the retrieved dust lidar *PDR*, we here propose a basic partitioning  
378 model in which the dust particles mixture ( $d$ ) = { $Abs$ ,  $\overline{Abs}$ } is resumed to two components: an absorbing component ( $Abs$ ),  
379 mainly corresponding to hematite particles, and a non-absorbing component ( $\overline{Abs}$ ), mainly corresponding to silica-particles.  
380 For simplicity, we here resume the absorbing (resp. non-absorbing) component to hematite (resp. silica)-particles with  
381 respective abbreviations ( $Hmt$ ) and ( $Sil$ ). We focus on the 355 nm wavelength at which hematite is an efficient light absorber  
382 and on the coarser *SD* as the dependence of the dust lidar *PDR* with size is less pronounced than with the particles mineralogy.  
383 In Appendix A is detailed the derivation of the lidar *PDR* of such a dust-particles mixture ( $d$ ) = { $Hmt$ ,  $Sil$ } (hereafter noted  
384  $\delta_d$ , as in lidar applications). This Appendix is an extension of our previous works (Miffre et al., 2011; David et al., 2013, 2014;  
385 Mehri et al., 2018) to the case study where both components { $Hmt$ ,  $Sil$ } are nonspherical. The lidar *PDR* of such a dust-  
386 particles mixture relates to that of its pure components (hereafter noted  $\delta_{Sil}$  and  $\delta_{Hmt}$ ) as follows:

387

$$388 \quad \delta_d = \frac{-e + (c + e)X_{Hmt}}{f - (d + f)X_{Hmt}} \quad (8)$$

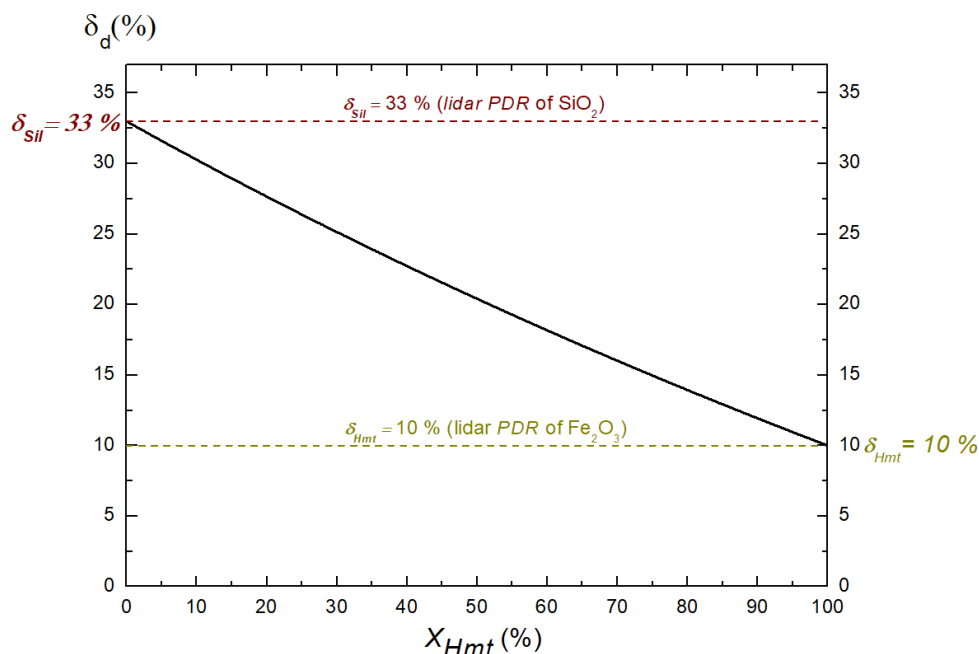
389

390 where the expressions of the  $c$ ,  $d$ ,  $e$  and  $f$ -coefficients are given in Appendix A and only depend on the depolarization ratios  
391  $\delta_{Sil}$  and  $\delta_{Hmt}$  of silica and iron oxides.  $X_{Hmt}$  is the fraction of  $Hmt$  to dust particles backscattering. Following Eq. (8) and  
392 Appendix A, Figure 6 displays the variation of  $\delta_d$  as a function of  $X_{Hmt}$  when considering  $\delta_{Sil} = 33$  % and  $\delta_{Hmt} = 10$  %, as  
393 obtained in our laboratory findings at 355 nm wavelength with the coarser *SD*. As shown in Figure 6, the dust lidar *PDR* lies  
394 in between  $\delta_{Sil}$  and  $\delta_{Hmt}$  and equals  $\delta_{Sil}$  (resp.  $\delta_{Hmt}$ ) only when  $X_{Hmt} = 0$  (resp. 1), depending on the fraction  $X_{Hmt}$  of light  
395 absorbent in the dust particle mixture. Hence, Arizona dust, which contains a lower fraction of hematite, exhibits a higher lidar  
396 *PDR* compared with Asian dust, at least at 355 nm wavelength where hematite is strongly absorbing. Though rather simple,  
397 our model interestingly highlights the key role played by light absorption in the retrieved Asian dust lidar *PDR*. To go further  
398 and become quantitative, this simple model should be refined, by considering also the other chemical oxides present in mineral  
399 dust, as well as other *SD* and other lidar wavelengths. To handle such a complex issue, more laboratory experiments are  
400 required on other chemical oxides, ideally also at 1064 nm wavelength. This work is however beyond the scope of this paper.  
401 Still as is, our model provides an interpretation of the laboratory-observed differences in the dust lidar *PDR* when the light





402 absorbent hematite is involved. In the most general case, the dust lidar *PDR* hence appears as a complex function of the  
 403 particles mineralogy, *SD* and wavelength. Though this triple dependence is difficult to disentangle, our laboratory findings  
 404 show that the dust lidar *PDR* is primarily affected by the particles mineralogy, at least when hematite is involved.



405  
 406 Figure 6: Numerical simulation of the 355 nm Lidar *PDR* of a two-component particles mixture ( $d$ ) = {*Hmt*, *Sil*}, composed of a light  
 407 hematite (*Hmt*) and silica (*Sil*) oxides as a function of the  $X_{Hmt} = \beta_{Hmt}/\beta_d$  fraction of *Hmt* to *d*-particles backscattering, following Eq.  
 408 (8) and Appendix A, by accounting for our laboratory experimental findings for  $\delta_{Sil} = 33\%$  and  $\delta_{Hmt} = 10\%$  (see Table 2 at 355 nm  
 409 wavelength with coarser *S*).

## 410 5 Conclusion

411 In this paper, the dependence of the lidar particles depolarization ratio (*PDR*) of pure mineral dust with complex refractive  
 412 index (*CRI*) and size is for the first time investigated through a laboratory  $\pi$ -polarimeter operating at  $180.0^\circ$  lidar  
 413 backscattering angle and (355, 532) nm wavelengths for lidar purposes. The goal of this work is to improve the knowledge on  
 414 the dust lidar *PDR*, which is an important input parameter involved in lidar partitioning algorithms, which are widely applied  
 415 to reveal the contribution of mineral dust in particles external mixtures (Tesche et al., 2009; Mehri et al., 2018). While mineral  
 416 dust exhibits a complex and highly irregular shape, which is difficult to model mathematically and numerically, our laboratory  
 417 approach allows accounting for the real shape of mineral dust. Our laboratory  $\pi$ -polarimeter is likewise a good complement to  
 418 lidar field experiments, which provide accurate retrievals of the lidar *PDR* of particles mixtures involving mineral dust.  
 419 Another advantage of our laboratory  $\pi$ -polarimeter lies in its ability to provide accurate retrievals of the lidar *PDR* of pure  
 420 mineral dust samples, differing in *CRI* and size. The  $\pi$ -polarimeter indeed operates at  $180.0^\circ$  lidar backscattering angle and at  
 421 (355, 532) nm lidar wavelengths: no assumption is made to retrieve the dust lidar *PDR*. This is a key novelty of our study.



422 Indeed, the variation of the dust lidar *PDR* with scattering angle and wavelength cannot be quantified (Bohren and Huffman,  
423 1983; Mishchenko et al., 2002) for complex-shaped particles such as mineral dust. Hence, our  $\pi$ -polarimeter improves the  
424 knowledge on the dust *PDR*, which are given in the literature at non  $180.0^\circ$  backscattering angle and / or at wavelengths  
425 differing from (355, 532 nm). Our work provides sixteen accurate dust lidar *PDR*-values, corresponding to four different  
426 complex refractive indices, studied at two size distributions (fine, coarse) and at (355, 532) nm wavelengths (see Section 4).  
427 The precision on the retrieved dust lidar *PDR* originates from the scattering matrix formalism, on which the laboratory  $\pi$ -  
428 polarimeter relies, as detailed in Section 3. To investigate the dependence of the dust lidar *PDR* with *CRI*, hematite, the main  
429 light absorbent present in mineral dust, was considered in addition to silica oxide, the main chemical oxide present in mineral  
430 dust, which is practically nonabsorbent. At 355 nm, our laboratory  $\pi$ -polarimeter proves that the lidar *PDR* of coarser silica is  
431 equal to  $(33 \pm 1) \%$  while that of coarser hematite is only  $(10 \pm 1) \%$ . In Section 4, this huge difference is explained by  
432 accounting for the high imaginary part of the hematite *CRI*. In turn, Arizona dust exhibits a higher depolarization ratio than  
433 Asian dust, due to the higher proportion in hematite in the latter. As a result, when the strong light absorbent hematite is  
434 involved, the dust lidar *PDR* is primarily governed by the particles mineralogy and the variations of the dust lidar *PDR* with  
435 size are less pronounced. When hematite is less or not involved, the dust lidar *PDR* increases with increasing sizes and the  
436 (355, 532) nm wavelength dependence of the dust lidar *PDR* then becomes key for discussing on the involved particle sizes,  
437 thus underscoring the importance of dual wavelengths (or more) polarization lidar instruments. To further disentangle the  
438 complex dependence of the dust lidar *PDR* with complex refractive index and size, our methodology should be extended to  
439 other chemical oxides, other natural mineral dust samples, other *SD* and other wavelengths. This is however far beyond the  
440 scope of this paper : we here focused on (355, 532) nm wavelengths, since mineral dust slightly absorb light in the near infra-  
441 red (Di Biagio et al., 2019). Still, the above laboratory findings underscore the importance of accounting for the wavelength  
442 dependence of the dust lidar *PDR*, whatever the hematite proportion. The spectral dependence of the dust lidar *PDR* is indeed  
443 instructive (Burton et al., 2016; Haarig et al., 2022; Miffre et al., 2020). Numerical outlooks of this work are obviously also  
444 interesting, as underscored by recent papers (Kahnert et al., 2020; Luo et al., 2022), discussing on the ability of the spheroidal  
445 model to mimic light scattering by complex-shaped mineral dust.

## 446 Appendix A

447 The goal of this Appendix is to establish the expression of the lidar *PDR* of a two-component particle mixture ( $p$ ) =  
448  $\{ns_1, ns_2\}$  composed of two non-spherical components  $ns_1$  and  $ns_2$ . As in lidar applications, the lidar *PDR* of  $p$ ,  $ns_1$  and  $ns_2$ -  
449 particles are respectively noted  $\delta_p$ ,  $\delta_{ns_1}$  and  $\delta_{ns_2}$ . The starting point is given by the set of four equations:

450

$$451 \beta_{p, //} = \beta_{ns_1, //} + \beta_{ns_2, //} \quad (\text{A-1-a})$$

$$452 \beta_{p, \perp} = \beta_{ns_1, \perp} + \beta_{ns_2, \perp} \quad (\text{A-1-b})$$



$$453 \quad \delta_{ns_1} = \beta_{ns_1,\perp} / \beta_{ns_1, //} \quad (\text{A-1-c})$$

$$454 \quad \delta_{ns_2} = \beta_{ns_2,\perp} / \beta_{ns_2, //} \quad (\text{A-1-d})$$

455

456 where  $\beta_{p, //}$  and  $\beta_{p,\perp}$  are the lidar particles backscattering coefficients, evaluated from a polarization lidar experiment carried  
 457 out at wavelength  $\lambda$  (here omitted to ease the reading). The backscattering coefficient  $\beta_{ns_1}$  of  $ns_1$ -particles is then retrieved by  
 458 noting that  $\beta_{ns_1} = \beta_{ns_1, //} + \beta_{ns_1,\perp} = \beta_{ns_1,\perp}(1 + 1/\delta_{ns_1})$  (Miffre et al., 2011; David et al., 2013). Moreover,  $\beta_{ns_1,\perp}$  can be  
 459 expressed as a fonction of  $\beta_{p, //}$  and  $\beta_{p,\perp}$  since  $\beta_{ns_1,\perp} = \beta_{p,\perp} - \beta_{ns_2,\perp} = \beta_{p,\perp} - \delta_{ns_2}\beta_{ns_2, //} = \beta_{p,\perp} - \delta_{ns_2}(\beta_{p, //} - \beta_{ns_1,\perp}/\delta_{ns_1})$   
 460 using Eqs. (A-1). Hence,  $\beta_{ns_1,\perp} = (\beta_{p,\perp} - \delta_{ns_2}\beta_{p, //}) / (1 - \delta_{ns_2}/\delta_{ns_1})$ . By applying the same methodology to  $ns_2$ - particles,  
 461 we finally get:

462

$$463 \quad \begin{pmatrix} \beta_{ns_1} \\ \beta_{ns_2} \end{pmatrix} = \begin{bmatrix} c & d \\ e & f \end{bmatrix} \begin{pmatrix} \beta_{p, //} \\ \beta_{p,\perp} \end{pmatrix} \quad (\text{A-2})$$

464

465 where the  $c, d, e$  and  $f$ -coefficients only depend on the depolarization ratios  $\delta_{ns_1}$  and  $\delta_{ns_2}$  :

466

$$467 \quad c = -\delta_{ns_2}(1 + 1/\delta_{ns_1}) / (1 - \delta_{ns_2}/\delta_{ns_1}) \quad (\text{A-3-a})$$

$$468 \quad d = (1 + 1/\delta_{ns_1}) / (1 - \delta_{ns_2}/\delta_{ns_1}) \quad (\text{A-3-b})$$

$$469 \quad e = -\delta_{ns_1}(1 + 1/\delta_{ns_2}) / (1 - \delta_{ns_1}/\delta_{ns_2}) \quad (\text{A-3-c})$$

$$470 \quad f = (1 + 1/\delta_{ns_2}) / (1 - \delta_{ns_1}/\delta_{ns_2}) \quad (\text{A-3-d})$$

471

472 The 2 x 2 matrix introduced in Eq. (A-2) can be inverted to get the expression of  $\beta_{p, //}$  and  $\beta_{p,\perp}$  and hence  $\delta_p = \beta_{p,\perp} / \beta_{p, //}$ . By  
 473 introducing  $X_{ns_2} = \beta_{ns_2} / (\beta_{ns_1} + \beta_{ns_2})$  the fraction of  $ns_2$  to  $p$ -particles backscattering, we finally get the relationship  
 474 between  $\delta_p$  and  $\delta_{ns_1}$  and  $\delta_{ns_2}$  :

475

$$476 \quad \delta_p = \frac{-e + (c + e)X_{ns_2}}{f - (d + f)X_{ns_2}} \quad (\text{A-4})$$

477

478 In the specific case where  $ns_2$ -particles are spherical (i.e.  $\delta_{ns_2} = 0$ ), the expressions of the  $c, d, e$  and  $f$ -coefficients simplify  
 479 and the relationship between  $\delta_p$  and  $X_{ns_2} = X_{ns}$  becomes identical to that we already published in (Miffre et al., 2011; David  
 480 et al., 2013). This new material is hence as an extension of our previous works (Miffre et al., 2011; David et al., 2013, 2014;  
 481 Mehri et al., 2018) to the case study where both components of the particles mixture ( $p = \{ns_1, ns_2\}$ ) are nonspherical.

482



483 **Author contribution**

484 **Alain Miffre**: Conceptualization, Formal analysis, Investigation, Methodology, Supervision, Writing - original draft, Writing  
485 - review & editing **Danaël Cholleton**: Formal analysis, Investigation, Software, Visualization, Writing - review & editing.  
486 Clément Noël: Software, Writing - review & editing **Patrick Rairoux**: Project administration, Supervision, Writing - review  
487 & editing.

488 **Competing interests**

489 The authors declare that they have no conflict of interest.

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