

# Response to the referee comments

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3 We thank the referees for their valuable and helpful comments. We have  
4 addressed all of them one-by-one in details as listed below. The comments are  
5 in bold and our replies are in regular font. The page/line numbers indicated  
6 in our replies are given with respect to the old manuscript, and may differ  
7 from the revised manuscript.

## 8 **Referee #1**

### 9 **General comments:**

10 **The manuscript is generally well written and well structured. All**  
11 **arguments are clearly described and reasonable, and the conclu-**  
12 **sions are justified. The authors present a very comprehensive**  
13 **study, including spectroscopy, instrument design, modeling, and**  
14 **geophysics. My main recommendation, therefore, is that the au-**  
15 **thors should keep the focus of the manuscript clearer and reduce**  
16 **side-topics or well-known aspects. That would help the reader to**  
17 **keep oversight over the 19 figures and related descriptions. I rec-**  
18 **ommend a minor revision of the manuscript and provide in the**  
19 **following more detailed comments.**

20 The referee's comments helped us to improve our manuscript. We went  
21 through the manuscript, reduced the side-topics and also highlighted the  
22 main questions of this study at the beginning of each of the sub-sections in  
23 the Assessment section, in order to keep the readers focused on the key topics.  
24 We excluded Appendix A since the discussed scale-separation method was  
25 actually not used in our study. We also modified/shortened the Introduction  
26 part.

27 **Minor comments:**

28 **1. The Introduction includes an extensive description of the rele-**  
29 **ance of gravity waves for the understanding of the middle atmo-**  
30 **sphere. I think this is without doubt, and the description can be**  
31 **shortened. The description of the MATS mission is from my point**  
32 **of view not relevant for a feasibility study of another instrument**

33 Following the referee's suggestion, we removed the description about the  
34 MATS mission and rewrote part of the Introduction.

35 **2. l. 184/185, Appendix A: As far as I understand, the described**  
36 **method of Ern et al. is not used in this manuscript and the compar-**  
37 **ison of methods is beyond its scope. Therefore, I suggest removing**  
38 **Appendix A and rephrasing this sentence.**

39 Following the referee's suggestion, we excluded Appendix A describing the  
40 scale separation method, and provided a reference to the paper of Ern et al.  
41 (2018) in former l.170.

42 **3. Fig. 2, Section 2.3: Fig. 2 is very deductive and important for**  
43 **the understanding of the method. I suggest referring the "first**  
44 **question/second question/third question" to the respective up-**  
45 **per/middle/lower yellow diamonds and adding references to Sec-**  
46 **tions 4.2 and 4.3.**

47 We thank the referee for this comment, which makes the topics of this  
48 manuscript more focused. We added "Question 1/Question 2/Question 3"  
49 to the three yellow diamond boxes in Fig. 2, and cross-referred to them back  
50 at the beginning of Sections 4.1, 4.2 and 4.3.

51 **4. l. 505: I do not understand why another cutoff is applied to the**  
52 **simulations compared to the reference. Please describe.**

53 The cut-off eliminates non-reliable fits where wavelengths are much longer  
54 than the analysis volume. In principle, it therefore depends on the number  
55 of tracks specifying one of the horizontal cube dimensions. In order not to  
56 confuse the reader we have now cut all distributions at 2100 km following  
57 the reviewers suggestion. Fig. 10 a-c were updated accordingly.

58 We modified the text in former l.486-487 as:

59 "All spectra are cut off at longer wavelength of around 2100 km horizontally  
60 and 45 km vertically, as the detection upper limit. It results from the limits  
61 when filtering reliable fits, which are up to  $\sim 3$  times the cube size, both for  
62 horizontal and vertical wavelength."

63 **5. l. 507: It does not become clear to me whether in some cases**  
64 **(Fig. 10 c and f) there are no waves below 150 km wavelength,**  
65 **or whether this part of the spectrum is not shown for technical**  
66 **reasons.**

67 For the 15- and 5-track cases (on sampled data), the wavelength spectra  
68 have a clear cutoff of horizontal wavelength close to 150 km at the short  
69 end, while the 4- and 2- cases (on retrieved data) have a wider spread of  
70 wave distribution towards 100 km, particularly at 75 km. This difference  
71 comes from: 1. our simulation is based on the HIAMCM model data, in  
72 which the shortest horizontal wavelength that could be resolved is around  
73 156 km according to Becker et al. (2022). Therefore, a shortest wavelength  
74 of around 150-160 km would be expected in our wave fitting results, and  
75 the 15- and 5-track simulation results do conform to this limit by a sharp  
76 decrease of spectral power. 2. the general cut-off at the short wavelength  
77 side is due to the implementation of the nested interval method in S3D.  
78 After a first guess which is in the low frequency region of the spectrum the  
79 search region for the minimum depends on the spectral resolution provided  
80 by the "natural" spectral grid which a Fourier transform would use. The  
81 more points are used (i.e. 2, 5, 15 pts for as many tracks) the narrower is  
82 this search region around the initial guess. Accordingly, wave solutions are  
83 confined closer to the initial guess and hence in the long wavelength range.  
84 Vice versa, for smaller cubes the search region is wider and gets closer to the  
85 Nyquist limit - never reaching this, though. This is consistent with the fact  
86 that the short-wavelength "cuts" are at wavelength notably longer than the  
87 Nyquist wavelength of 60 km.

88 **6. Section 4.3: I recommend referring at the beginning of this sec-**  
89 **tion back to Fig. 2 and Section 2.1.**

90 We added reference to Fig. 2 and Section 2. at the beginning of Section 4.3.

91 **7. Section 5 discusses in general the relevance of the examination**  
92 **of mesoscale gravity waves independent from the proposed instru-**  
93 **ment or the availability of additional information (wave sources,**  
94 **winds, ...). I suggest either shortening this section or pointing**  
95 **out why these studies cannot be done with other (existing) instru-**  
96 **ments.**

97 In Sect. 5 after former l.609 we now have added a short summary of existing  
98 satellite observations of GWs in the MLT region. None of these data sets  
99 provides 3D information of the observed GWs, which strongly limits the  
100 interpretation of these data.

101 **8. l. 610 – l. 613: I agree that wind information is crucial for the**  
102 **understanding of wave dissipation and other processes. However,**  
103 **it is hardly available on a global scale in the MLT. Wind data from**  
104 **assimilated temperature information may lack precision, especially**  
105 **for non-linear processes. Effects like GW bending cannot be ac-**  
106 **knowledged at all. Please comment on the consequences of limited**  
107 **data availability for the science questions.**

108 Unlike in the stratosphere, mesospheric assimilation systems have far too low  
109 resolution to resolve GWs in any realistic fashion. This is, however, not re-  
110 quired for the interpretation of the data. We need only a reasonably realistic  
111 representation of the global wind fields. This naturally contains gradients,  
112 both in the vertical and in the horizontal and hence will cause ray-tracing  
113 modelling to produce refraction of the wave vector (both horizontal and ver-  
114 tical components) and to generate critical levels when  $\lambda_z \rightarrow 0$ . The question  
115 then is not whether the fields are sufficiently accurate, but sufficiently re-  
116 alistic. Assimilation systems still struggle, in particular in times of special  
117 interest (Harvey et al., 2022), as apparently the information content of the  
118 observations has insufficient weight to correct the model. Still, one can use  
119 geostrophic winds as an approximation of the large scale flow. In addition,  
120 methods were developed to determine the tides (Nguyen and Palo, 2013; Pe-  
121 datella et al., 2016), and based on such results one could adapt tidal models  
122 such as the GSWM in order to gain a complete view of all variables. The  
123 current focus on the MLT rises therefore hope that methods will be found  
124 to determine sufficiently realistic winds to produce diagrams of the ground-  
125 based phase speed, assess critical level filtering and perform ray-tracing to  
126 investigate the fate and characterize the origin of GWs in a general fashion.  
127 Whether we will gain sufficiently accurate winds for backward ray-tracing of  
128 such mesoscale GWs by 60 to 80 km altitude towards individual sources is  
129 a different matter and there I am less optimistic. The need for only global  
130 fields is now emphasized and a shorter version of the state of assimilation  
131 and other methods to determine is included into the text.

132 Of course, it is not sure whether global wind observations, or reliable winds  
133 from data assimilation, will be available in the upper mesosphere/lower ther-  
134 mosphere at the time the instrument will be in operation. This means, as  
135 already stated in former lines 636–638, it is not sure whether we will be  
136 able to perform reliable ray tracing of gravity waves, and identification of  
137 the gravity wave sources. Also gravity dissipation studies in relation to the  
138 background wind would not be possible.

139 However, the gravity wave data set that we expect to obtain from this novel

140 observation method will be quite unique, and of great value in itself. Even  
141 without wind observations, studies based on the observed directional grav-  
142 ity wave momentum flux can be performed in a climatological sense, for  
143 example by comparison with zonal wind climatologies, or climatologies of  
144 atmospheric tides. Particularly, the interaction between gravity waves and  
145 tides is not well understood and offers a wide field of applications. Further,  
146 the novel gravity wave data set can be used to identify cases of excitation of  
147 secondary gravity waves. This can be performed by identifying fishbone-like  
148 structures in along-track/altitude cross sections without the need of having  
149 background wind information. For these kind of studies the relatively short  
150 along-track sampling of 30km, combined with a tomographic retrieval, will  
151 be very beneficial.

152 This reasoning will be added in the revised manuscript after former l.618 &  
153 l.640.

154 **9. l. 643/644: I suggest comparing to other global observations**  
155 **instead of comparing to models. Even GW-resolving GCMs may**  
156 **not display true atmospheric states despite they are good tools for**  
157 **the understanding of atmospheric processes.**

158 At the end, our point is that we need good quantitative global measures  
159 of GWMF. We don't have these, yet. Our current error bar on the global  
160 climatologies which we can deduce e.g. from SABER is about a factor of  
161 3 and with that you can argue for everything between these scales being a  
162 minor contribution and the only thing worthwhile looking at. My best guess  
163 would be about half from scales longer than 100km and half from scales  
164 shorter, but nobody knows. We have included the reference to the shorter  
165 scales here to indicate the contradicting evidence. We have tried to sharpen  
166 the point in the text by including a cross-reference to the introduction, the  
167 short-scale observations you mentioned and making evident that all this is  
168 partly contradictory evidence.

169 A summary of existing satellite observations in the MLT region, and of their  
170 limitations has been added in Sect. 5.

171 **10. l. 653: I suggest describing the effects of the observational filter**  
172 **much earlier.**

173 We added the sentence and reference about the observational filter in the  
174 Introduction after former l.110.

175 **11. l. 676/677: I agree with the statement about zonal mean cli-**  
176 **matologies but suggest removing the two lines including wind data.**

177 **The authors describe their own concerns in L. 636-638.**  
178 We removed the sentence in former l.676/677 in the conclusions.

## 179 **Technical comments and typos:**

180 **1. l. 63: “residual”**

181 We changed the text accordingly.

182 **2. l. 429: “shown by Lehmann et al. (2012)”**

183 We changed the text accordingly.

184 **3. l. 439: “(cf. Section 2.1 and Figure 2)”**

185 We added in the text accordingly.

186 **4. l. 450: “by Ern et al. (2004)”**

187 We changed the text accordingly.

188 **5. l. 470: I suggest adding “(see data flow to the middle yellow**  
189 **diamond in Fig. 2)”**

190 We added in the text accordingly.

191 **6. l. 498: “5-track”**

192 We changed the text accordingly.

193 **7. l. 583: “can be generated” should read “can be calculated”**

194 We changed the text accordingly.

## 195 **References**

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