

1 **ICE-CAMERA: a flatbed scanner to study inland Antarctic polar precipitation.**

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6 **Abstract.**

7 Studying precipitation at very high latitudes is difficult because of the harsh environmental conditions that limit the external  
8 activity of humans and instruments, especially in the polar winter. The direct monitoring of ice crystal habits and size  
9 distribution in antarctic precipitation is important for the validation of the algorithms used for retrieving precipitation from  
10 ground-based and satellite-borne radar instruments, and for the improvement of the climatological modeling of polar areas.  
11 The paper describes an automated device (ICE-CAMERA) specifically developed for the imaging, measurement and  
12 classification of ice precipitation on the Antarctic high plateau. The instrument gives detailed information on precipitation on  
13 an hourly basis. The article provides a description of the device and its image processing software. Starting in 2014, the  
14 instrument operates *almost* unattended all year round at Concordia station, Antarctica (75°S, 123°E, 3220 m altitude).

15

16 **1. Introduction.**

17 In Antarctica, the characteristics of ice precipitation depend greatly on the region. In coastal areas, precipitation is influenced  
18 by synoptic scale features, such as cyclones and fronts (Bromwich, 1988). In the interior (> 2500 m), a significant part of the  
19 precipitation falls in the form of small ice crystals (“diamond dust”, DD) under clear-sky conditions (Fujita and Abe, 2006).  
20 Snow particles over Antarctica are generally smaller compared to other regions of the world. The largest particles are found  
21 close to the coast, where more water vapour is available and diameters up to 10 mm are recorded (Konishi et al., 1992) with  
22 particle shapes similar to mid-latitude ones (Satow, 1983). Most of the bigger particles are aggregates (some can be found in  
23 the dataset of Grazioli et al.,2022). More inland stations record snowflakes of much smaller sizes, ranging from particles  
24 smaller than 100  $\mu\text{m}$  at South Pole (Walden et al., 2003, Lawson et al., 2006) till hundreds of  $\mu\text{m}$  at other inland stations  
25 (Lachlan-Cope et al., 2001).

26 In situ measurements of precipitation are rare in Antarctica and affected by large uncertainties. This is particularly true in the  
27 high plateau, where less than 20 cm of snow accumulates every year (Palerme et al.,2014). As a result, the global precipitation  
28 products that rely on these observations (i.e. the Global Precipitation Climatology Centre (GPCC), (Schneider et al., 2017))  
29 have no coverage over this region. Other observational products, such as the Global Precipitation Climatology Project (GPCP)  
30 (Huffman et al.,2001) , that uses GPCC for bias correction over land, has relied on satellite-only precipitation estimates.  
31 Satellite products also face large uncertainties over cold regions such as Antarctica due to insufficient sensitivity of sensors to  
32 detect and estimate precipitation signals, complex surface emissivities, and poor understanding of precipitation microphysics.  
33 Ground based K-band radars ( $\sim$ 1-cm wavelength) are robust instruments successfully employed for studying precipitation in  
34 coastal Antarctic sites (Souverijns et al., 2017) but are quite blind to the sub-millimetre ice particles encountered on the plateau,  
35 due to the relationship  $D^6$  between the radar scattering cross-section and the particle diameter (D).

36 The satellite-borne radar CloudSat (Liu, 2008) did provide a quantum leap in observing ice in the Antarctic atmosphere (up  
37 to 82 °S), but being a single-frequency radar (like K-band radars), the retrieval of precipitation quantities relies on many  
38 assumptions about the properties of particles, resulting into  $\pm 50\%$  uncertainties for IWC (Heymsfield et al.,2008). The  
39 microphysical assumptions (shapes and size distribution of particles) are the biggest causes for IWC, IWP, and snowfall rate  
40 retrieval uncertainty (Hiley et al.,2011, Wood et al.,2015). Moreover, CloudSat bins close to the ground cannot be used for  
41 precipitation retrieval, resulting into a severe underestimation of the diamond-dust and blowing-snow contribution to Antarctic  
42 snow balance (Palm et al., 2018). Despite these uncertainties, in absence of ground validation CloudSat data are now used as

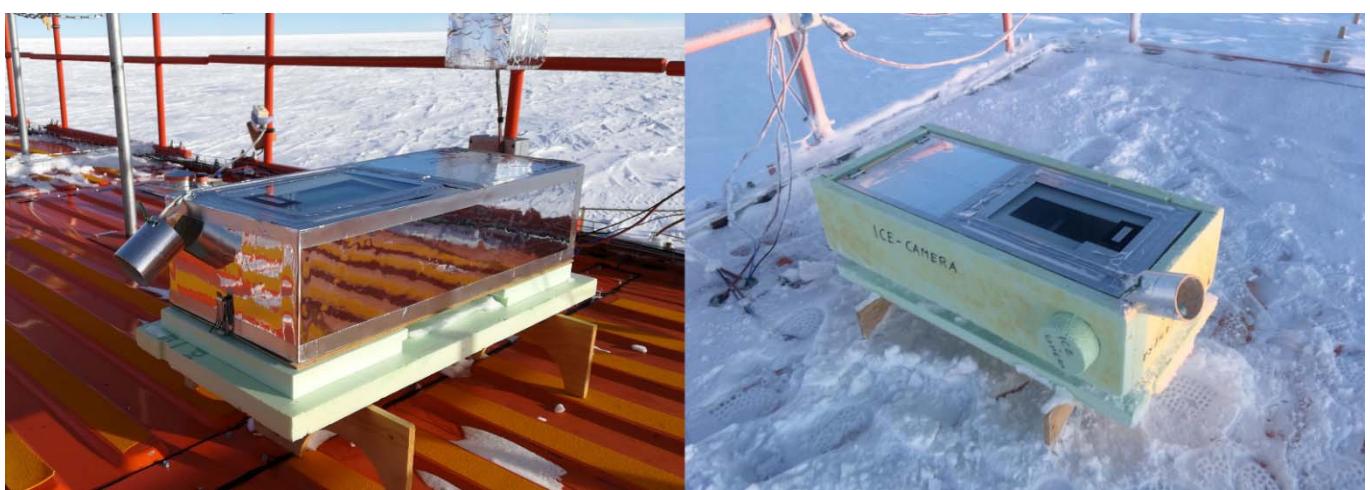
43 independent dataset for the validation of precipitation models in Antarctica (Palerme et al.,2014, Palerme et al.,2017).  
44 The direct observation and the continuous monitoring of habit and size distribution of precipitation is therefore required in  
45 order to validate both precipitation models, CloudSat and radar algorithms on the Antarctic plateau.  
46 Disdrometers are robust *in-situ* devices, increasingly used in Antarctic coastal areas (Souverijns et al., 2017, Bracci et al.,2022).  
47 They provide the size distribution and falling speed of hydrometers, but they give no direct information about the shape. The  
48 evolution of disdrometers into 2D-disdrometers gave access to some shape indications about hydrometeors (Grazioli et  
49 al.,2014). A further evolution of disdrometers into Imaging-disdrometers, such as the Snowflake Video Imager (SVI) (Newman  
50 et al., 2009), provided realistic images of the crystals. Grazioli et al., (2017), as part of a multidisciplinary field campaigns,  
51 deployed a multi-angle snowflake camera (MASC) to take photographs of individual snow particles. This instrument,  
52 representing a further advance in the field of imaging disdrometers, collects high-resolution stereoscopic photographs of  
53 snowflakes in free fall while they cross the sampling area (Garrett et al., 2012), thus providing information about snowfall  
54 microphysics (Praz et al., 2017). The optical structure of the imaging-disdrometer and the MASC makes these instruments  
55 reliable in the presence of millimetre-sized hydrometeor precipitation. In Antarctica, their practical application is mostly  
56 limited to coastal zones where particles are coarse (e.g. MASC resolution is 33  $\mu\text{m}$ ).  
57 The direct observation of inner Antarctic particles requires imaging techniques with resolution of a few microns. Photographic  
58 studies of precipitation in the interior of the Antarctic are quite rare, carried out primarily at the South Pole Station (SPS)  
59 through formvar replicas. In early works with formvar, Hogan (1975) identified at SPS millimetre-sized columnar crystals and  
60 column- and bullet-rosettes in cloud precipitation, and smaller ( $\geq 100 \mu\text{m}$  diameter) platelike particles in clear-sky  
61 precipitation. Satow (1983), working with formvar replicas on Mizuho plateau found prevalently single bullets and  
62 combination of bullets. Long solid column crystals were also found (with an air temperature range from -42°C to -56°C) with  
63 a mean length of 290  $\mu\text{m}$  and a maximum length of 1.2 mm, with a mean aspect ratio of 18. Small (50-400  $\mu\text{m}$ ) hexagonal,  
64 triangular, scalene and square plates were also observed. Kikuchi and Hogan (1979) collected formvar replicas of DD in the  
65 summer at SPS, finding columnar crystals of 90  $\mu\text{m}$  average lengths and plates as small as 50  $\mu\text{m}$  in diameter. Ohtake and  
66 Yogi (1979) classified winter ice crystal precipitation in Antarctica under six categories. These included large rosettes, bullets  
67 and columns (millimetre-sized), thin hexagonal plates and columns (200  $\mu\text{m}$  or less), and smaller crystals of various shapes  
68 including triangular and polyhedral. Shimizu (1963) observed "long column" crystals in the winter at Byrd Station (80S,  
69 120W). Size distributions of Antarctic DD in winter and spring were reported by Smiley et al. (1980) for particles larger than  
70 50  $\mu\text{m}$ : they observed the same ice crystal forms that were reported earlier. Walden et al. (2003) studied DD, blowing snow,  
71 and cloud precipitation in winter, at SPS, by collecting crystals on slides and analyzing them using microphotography. In their  
72 study, columns with an average length of 60  $\mu\text{m}$  and plates with an average diameter of 30  $\mu\text{m}$  were found in DD. The direct  
73 observation of ice precipitation on the plateau was typically carried out by means of formvar replicas and/or microphotography,  
74 but these techniques take time, are difficult to implement throughout the year and are necessarily limited to short field  
75 campaigns and samples of very limited size. Designing automatic instruments for the continuous, photographic study of  
76 precipitation in such a harsh environment necessarily require several compromises between the high resolution of  
77 microphotography and the robustness of outdoor optical instruments such as disdrometers. Lawson et al. (2006) worked at  
78 SPS, in summer, using innovative Cloud Particle Imagers (CPIs), which replaced formvar replicas. This technique allowed the  
79 automatic analysis of around 700,000 DD crystal images in terms of caliper size, aspect ratio and other shape parameters. An  
80 automatic classification software, based on shape parameters, was used to categorize the images into nine simplified classes:  
81 small plates and spheroids, columns, thick plates, plates, budding rosettes, rosettes, complex with side planes, irregulars.  
82 Concordia International Station, located on the Dome-C (DC, 75°S,123°E, 3220 m above sea level) is a special location to test  
83 new instruments for precipitation studies. Surface temperatures seldom exceed -25°C in summer, whereas winter temperatures  
84 can reach -85°C. The 3 m average wind speed is 3  $\text{ms}^{-1}$  for Aristidi (2005) and 4.5  
85  $\text{ms}^{-1}$  (hourly-averaged) for Argentini et al. (2014). The strongest winds (up to 15  $\text{ms}^{-1}$ , hourly-averaged) blow from the

86 continental regions. These winds are due to gravity flows from the inner plateau regions south of Dome C, and are more often  
87 observed during the winter, especially in coincidence with warming events. The circulation at the surface during the summer  
88 is affected, especially in daytime, by the synoptic circulation. In summer the wind speed oscillates during the day, with values  
89 increasing (by a few  $\text{ms}^{-1}$ ) in the afternoon, when a convective layer develops, leading to the increase of the wind speed  
90 (Argentini et al., 2014). Relative humidity relative to ice is typically around 55-85% (Genthon et al., 2022). In these conditions,  
91 precipitation of ice crystals can be studied by simply collecting them on horizontal surfaces. This is done at DC by hand,  
92 starting in 2008, collecting precipitation on flat surfaces ("benches") and visually inspecting it. This analysis is restricted to  
93 one observation per day, a rate that is difficult to increase, especially in winter. The analysis of these samples is also time-  
94 consuming and often subject to biases due to ice re-processing and sublimation, hoar formation, and subjective judgement of  
95 the shape and relative abundance of ice particles. Schlosser et. al (2017) relied on this manual observation and classification  
96 of ice particles in his analysis of precipitation isotope data at DC. They classified the ice grains into diamond dust, drifting  
97 snow, snow and frost (hoar). The prevalence of hoar in the observed daily precipitation record (with temperatures below -  
98 50°C) indicates the limitations of this manual technique if detailed information on DC precipitation particles is desired.  
99 Detailed work was carried out in DC on a few individual DD and cloud precipitation crystal replicas by means of SEM electron  
100 microscopy by Santachiara et al. (2016). They also analyzed very small particles (10-50  $\mu\text{m}$ ), in a size range inaccessible to  
101 ordinary optical methods.  
102 The purpose of developing ICE-CAMERA was to fill a gap in precipitation monitoring at Concordia with a robust instrument  
103 capable of monitoring with continuity, all-year round, habit and size of ice particles in precipitation, while avoiding some of  
104 the problems associated with the visual inspection of precipitation. This was achieved through the combined development of  
105 robust camera equipment and machine learning techniques for sizing and classifying ice crystals.  
106  
107

## 108 **2. The instrument.**

### 109 **2.1 Overview of ICE-CAMERA.**

110 ICE-CAMERA is a flatbed scanner (Zheleznyak et al, 2015), whose operating principle is the same as that of ordinary  
111 flatbed scanners in offices. In the case of ICE-CAMERA, it specially designed for observing polar precipitation in the harsh  
112 environmental conditions of Concordia station (Fig. 1). Within this work, the term "precipitation" will include both  
113 "diamond dust" and cloud precipitation. The 'Deposition Surface' (DS) is defined as the horizontal, glass surface of the  
114 instrument facing the sky and collecting precipitation.  
115



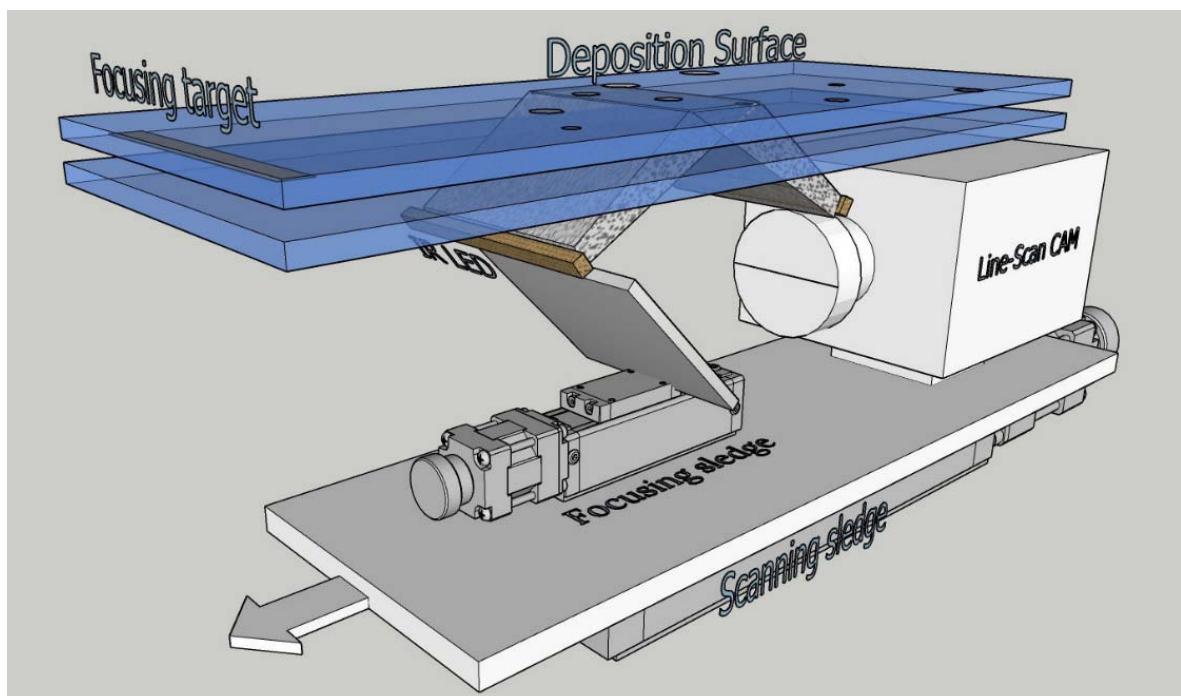
116  
117 **Fig. 1: ICE-CAMERA with its summer sun-shield (left) and with the winter coat (right).**  
118

119

120 The principle is simple: at the low temperatures and low wind speeds typically encountered at DC, precipitation falling on a  
 121 horizontal glass surface accumulates with time until it sublimates, leaving enough time for scanning the DS for counting and  
 122 measuring individual ice particles. The DS is the external surface of a special glass with electric heating (2.7). A second sheet  
 123 of glass together with the DS glass creates a double glass window that isolates the DS from the heated parts of the instrument  
 124 (Figs.2, 3). The scanning, like in ordinary flatbed scanners, is performed by means of a line-scan camera (Sect.2.2). moved by  
 125 a motorized scan sledge, and looking up at the DS through a 45° mirror (Fig. 2). The focus of the camera is adjusted by a  
 126 small motorized focusing sledge moving the 45° mirror (Sect.2.3). During the scan, the image is sent to the PC, located inside  
 127 the shelter.

128 After a complete scan of the DS, the glass is heated and the precipitation sublimated (section 2.7). Once cooled down, the  
 129 clean SD begins to accumulate new particles. This cycle takes place every hour. After each image acquisition, the MATLAB  
 130 image processing code is called to process the DS image, and a summary-image containing only segmented particles (if  
 131 present) is stored for post-processing. (Sect. 4.1.3). Every particle is also automatically measured through image processing  
 132 (Sect. 4.1) and classified through machine learning (Sect.4.2). Individual particle data are stored in rows in a text file, along  
 133 with weather and housekeeping data, for post-processing and statistical analysis.

134



135

136 **Fig. 2: ICE-CAMERA basics: the scan sledge moves the image-acquisition line along the deposition surface. The focusing sledge  
 137 adjusts the focus.**

138

139 All basic operations of ICE-CAMERA, (with the exception of CAM acquisition) are driven by a custom microprocessor  
 140 (Microchip PIC) logic board (Fig. 3). The same PIC board reads the housekeeping temperature sensors (attached to the DS  
 141 and placed inside and outside the instrument), drives the stepper motors of the sledges as well as pumps and fans. The PIC  
 142 Board communicates with the main computer (located inside the shelter) through RS232. NI Labview software controls image  
 143 acquisition, reads maintenance data, and monitors PIC operations along the RS232 line. The line-scan camera communicates  
 144 with the PC via Gigabit Ethernet.

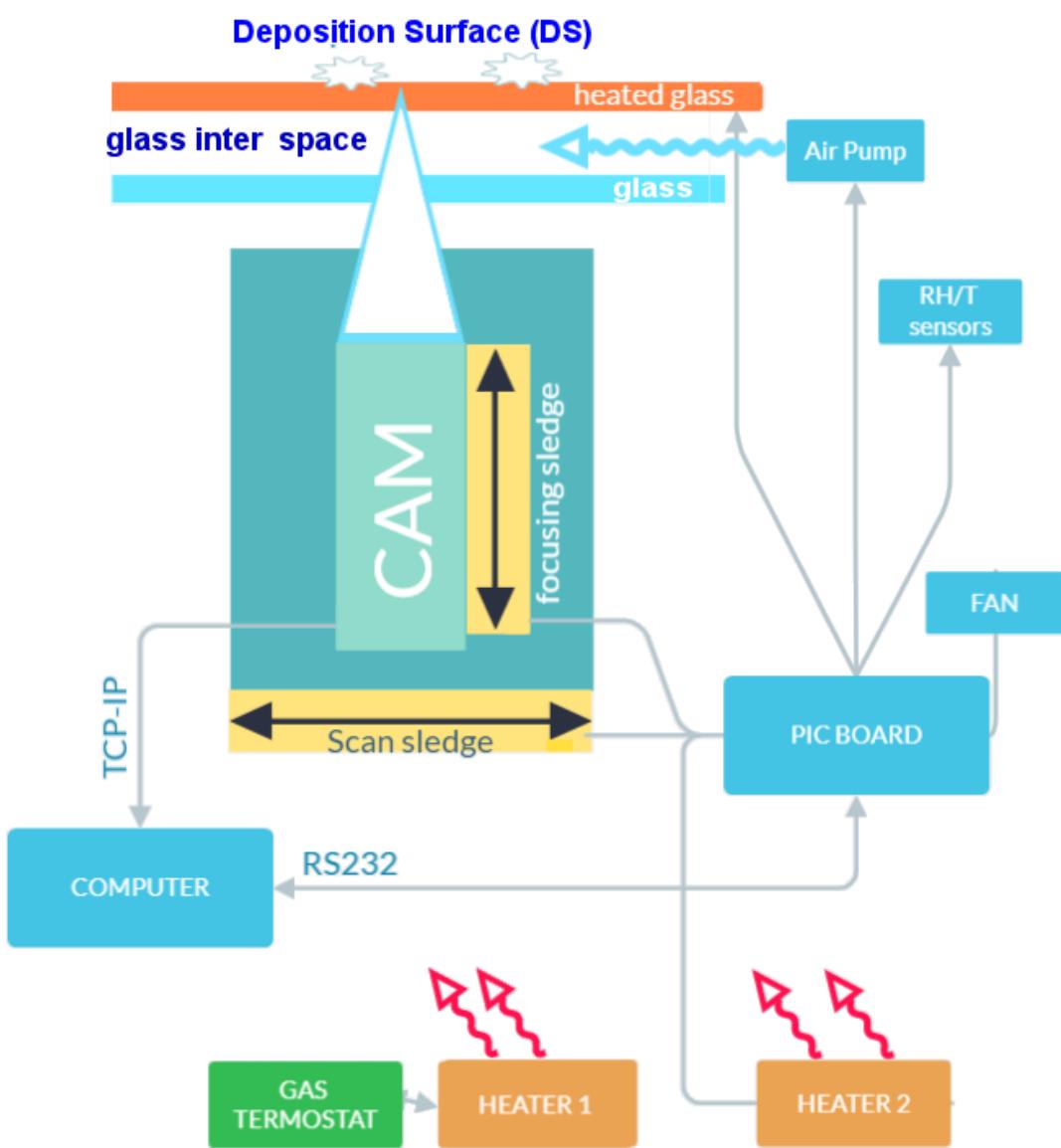
145 The instrument is placed outdoor, on the roof of the "Physique" shelter, approximately 6 m above the ground.

146 ICE-CAMERA was first installed in Concordia in 2012, but replaced in 2014 with its improved version, described here.

147 From then on, the instrument works year-round to produce precipitation data, every hour. Standard meteorological data are  
148 automatically obtained from local weather station AWS MILOS 520.

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## 155 2.2 The line scan Camera.

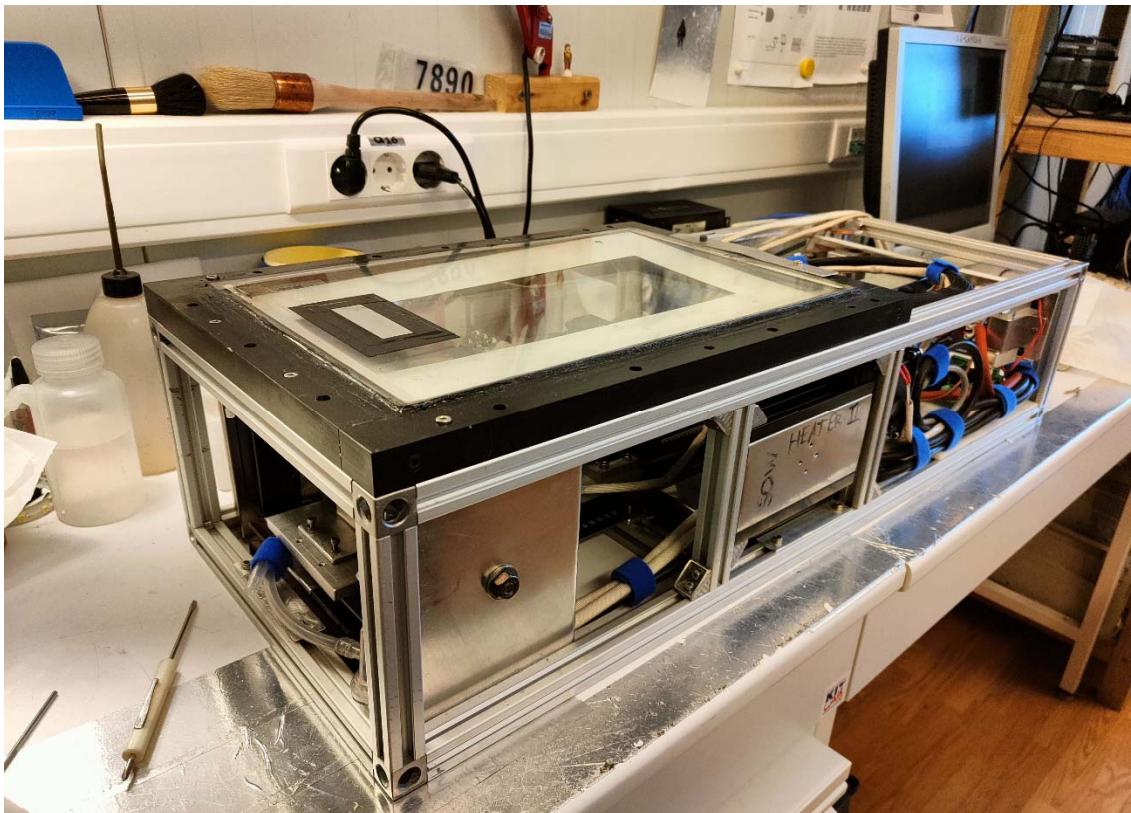
156 A linear scanning GigE Vision monochrome camera (Schafter-Kirchoff SK7500VTF-XB (52.5 mm sensor, 7500 pixels,  
157 7x7 $\mu$ m pixels, 8.2kHz line frequency), equipped with a 1:1 macro lens (APO-Rodagon D1X, f5.6) is used for the acquisition.  
158 The optics were designed by Schafter-Kirchoff in order to have a resolution equivalent to the 7 $\mu$ m pixel size. The 45 deg  
159 mirror is used to look upward. The illumination is ensured by 850 nm LEDs. A colour filter (Schott RG715, 800-1000 nm  
160 band-pass) was used on the CAM lens, in order to have a fully solar blind instrument. The line-scan camera assembly is moved,  
161 hourly, by a motorized sledge at a speed of 8 mm s<sup>-1</sup> in order to scan the rectangular DS ( 55x200 mm), located at the center  
162 of the window. The final image is 7500\*30000 pixel, 12 bits, monochrome. A fine calibration of the actual pixel size of the  
163 DS image was achieved by scanning a calibrated grid (0.1mm spacing) placed on the DS. This is necessary because the  
164 effective resolution of the image produced by the moving linear camera along the sledge direction depends on how fast the  
165 sledge moves. After the correction of this effect, the image pixel size resulted in 6.97 x 6.9  $\mu$ m, which was extremely close to

166 the simulated size of  $7 \times 7 \mu\text{m}$ . For the Nyquist sampling theorem, details less than  $14 \mu\text{m}$  cannot be detected in the image  
167 (under optimal focusing conditions). This resolution is enough, for example, for the observation of the hexagonal edges of the  
168 smallest plates detected by the instrument.

169 **2.3 The Focusing.**

170 In working conditions, the focal depth is  $\pm 0.5 \text{ mm}$ . A preliminary and accurate alignment of the motorized sledge plane to  
171 the DS ensures uniformity of focus across the DS at room temperature. A motorized focusing sledge, moving the bending  
172 mirror, allows to adjust the focus in operating conditions (Fig. 2). As ICE-CAMERA works outdoor at DC, it can  
173 experience a broad internal temperature range, from  $+5^\circ\text{C}$  in summer to  $-45^\circ\text{C}$  in winter, with quite large temperature  
174 gradients across the structure. Thermal expansion and changes in optical refractive indexes result in unpredictable changes in  
175 the focal plane. The correction of the focus is thus automatically performed, every 6 hours, by bringing the measuring sledge  
176 outside the DS, where a focusing spot (a sandpaper strip) is glued to the window (Fig. 4).

177



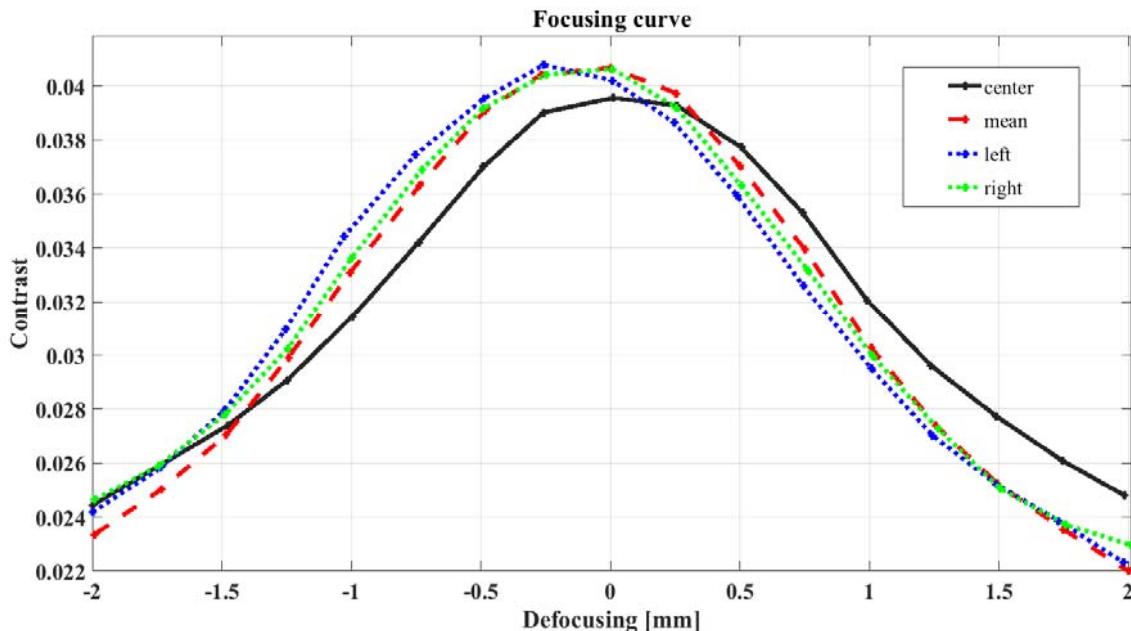
178

179 **Fig. 4: ICE-CAMERA out-of-the-box. The focus target is fixed onto the DS.**

180

181 The porous structure of the sandpaper has a length-scale of the order of  $0.1 \text{ mm}$ , comparable with the size of the measured  
182 ice particles. While calibrating, the focusing sledge is moved by  $\pm 2 \text{ mm}$  around the actual position in  $0.25 \text{ mm}$  steps.  
183 Successive images of the sandpaper are taken and their contrast (defined as the standard deviation of the intensity of the  
184 pixels) is measured. After a Gaussian-fit of the contrast as a function of defocusing (Fig. 5), the position corresponding to  
185 the maximum contrast is obtained, and the mirror sledge is moved into that position. The typical focal spot adjustment  
186 between two consecutive calibrations is  $0-0.25 \text{ mm}$ . The calibration takes approximately 5 minutes. For this reason, it is not  
187 done after each measurement, so as to save PC resources for data processing.

188



189

190

191 **Fig. 5: Typical focus calibration: Contrast is calculated in three sectors of the image: center, left and right. The contrast**  
192 **throughout the image is also displayed (red). The slight difference in focus (0.2mm) between the center of the image and the side**  
193 **wings is a normal lens effect.**

194

195 2.4 Illumination.

196 Lighting is supplied by two 850 nm LED (TSHG6200) strips. Both arrays illuminate the scan line symmetrically and  
197 approximately 45° from the optic axis in order to minimize multiple reflections in the double glass and within the camera lens.  
198 Infrared illumination was chosen in order to work in solar-blind conditions. This is particularly important, as the linear scanning  
199 camera always looks upward, to the sky. The uniformity of lighting along the linear CCD image was tested by taking an image  
200 of the same sandpaper used in the focus. The intensity profile along the CCD image was measured, and the intensity of the  
201 LEDs eventually changed to have a final intensity uniformity across the entire frame of less than 15%.

202

## 203 2.5 The Deposition Surface (DS).

204 The DS is the external surface of a 10 mm thick, electrically-heated glass (E-GLAS, Saint-Gobain). The glass is a sandwich  
205 with an electrically conductive layer pressed between two usual glass sheets. This glass is transparent at 850 nm, and can be  
206 electrically heated with 45 V ac, 95 W. A second, 2 mm thick, optically graded glass sheet (an ordinary flatbed scanner optical  
207 glass), placed 13 mm under the DS, makes up with the DS a double glass. This arrangement is necessary in order to keep the  
208 DS thermally insulated from the heated, interior of the instrument. A thermocouple is attached to the DS, while other  
209 thermocouples monitor double glass inter-space temperatures. A temperature of (at least) 3 °C above air temperature is enough  
210 to prevent, in DC, the formation of frost on surfaces in any season, as suggested by the work of Tremblin et al., (2011). In  
211 ICE-CAMERA, the temperature of the DS is usually 4 to 5°C above room temperature, which keeps the DS free of frost in all  
212 seasons (Fig. 6).



214

215 **Fig. 6: ICECAMERA at  $-70^{\circ}\text{C}$ , Concordia station winter: the DS if free of frost**

216

217

218 During the sublimation period (Sect.2.7), ambient air is pumped for five minutes by means of a  $3.5 \text{ l m}^{-1}$  miniature pump  
 219 through the double glass inter space, in order to keep the internal surfaces of the double glass always free of frost. Using  
 220 inert gases such as argon in the double glass space for the same purpose proved unsuccessful in Concordia at the extremely  
 221 low winter temperatures. In order to avoid the eventual accumulation of wind-drifted snow, the DS has no walls or obstacles  
 222 all around. Furthermore, the instrument is located on the roof of a shelter, almost 6 meters above the ground, an altitude  
 223 where blowing snow is not normally important at Concordia. Libois et al. (2014) identify drifting snow events at Dome C  
 224 when the 10-m wind speed exceeds  $7 \text{ m s}^{-1}$ . Assuming a logarithmic wind speed profile between the surface and 10-m and  
 225 an aerodynamic roughness length value of 1 mm (Vignon et al., 2016), this corresponds to a wind speed threshold value of 5  
 226  $\text{m s}^{-1}$  at 6 m above the ground. Winds below this threshold (near the annual average wind speed in DC) are not expected to  
 227 carry blowing snow to the DS. In addition, blowing snow impacts the flat horizontal and smooth DS at very small angles,  
 228 with a very limited chance of sticking to it. As a consequence, ice particles collected on the DS can be considered  
 229 representative of precipitation. In case of strong winds, not only the attachment of blowing snow to the DS is very low, but  
 230 also the collection of eventual precipitation is reduced. Since DS is warmer than air, there is no secondary growth in  
 231 deposited ice. Instead, the partial sublimation of ice particles before scanning could not be excluded, especially in summer.  
 232 This topic needs additional field work and will be modelled in Sect. 3.2.

233

## 234 **2.6 The thermal control.**

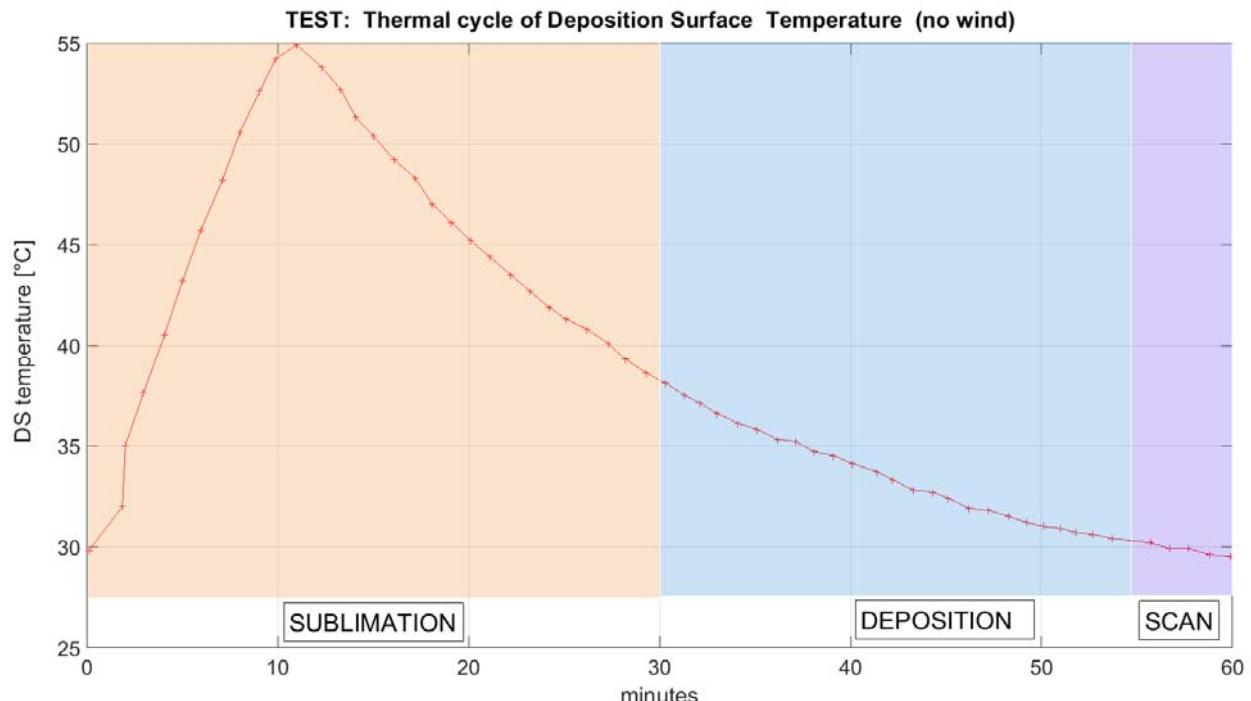
235 The temperatures measured by the ICE-CAMERA sensors are continuously transferred to the computer. The NI-Labview  
 236 software controls the internal temperature of ICE-CAMERA above  $-40^{\circ}\text{C}$  (by driving the 200W, ventilated air heater "Heater  
 237 2" of Fig. 3), and the DS temperature always under  $-5^{\circ}$  (by eventually disabling the "heated glass" of Fig.3). These conditions  
 238 are maintained throughout the year during every phase of the measuring cycle. An independent 200W thermostat ("Heater 1"  
 239 in Fig. 3) provides emergency temperature control in case of computer or PIC board failure. After a black-out, when the power  
 240 is restored, a timer is used to heat the inside of the instrument before turning on the electronics. This is important at Concordia  
 241 to prevent damage to standard electronics with typical operating temperatures of  $-40^{\circ}\text{C}$ .  
 242 In winter, a 40 mm thick Styrofoam coat is added around the instrument for increasing thermal insulation, whereas in summer

243 a Mylar sunscreen prevents overheating of the instrument and allows keeping the DS below -5°C in the warmest days (Fig.  
244 1). Additionally, in warm weather, outdoor air is carried inside the box with a tangential fan, for better cooling of the  
245 instrument. The DS temperature (without heating applied to DS glass) is on the average 4 to 5°C above the air temperature.  
246

247 **2.7 Sublimation-deposition cycle.**

248 After an entire scan of the DS, electricity is applied to the DS glass to sublimate the particles. The heating rate of the DS  
249 depends primarily on the electrical power applied to the glass and its thermal constant (approximately  $0.8 \text{ W m}^{-1} \text{ K}^{-1}$ ), and  
250 secondarily on the wind speed. An indoor test (Fig. 7), showed a heating of rate of  $2.5^\circ\text{C min}^{-1}$ , and a cooling rate of  $1^\circ\text{C}$   
251  $\text{min}^{-1}$ .

252



253 **Fig. 7 : Indoor test of DS heating-cooling within a 60 min cycle. For reasons of simplicity, the periods of sublimation and**  
254 **deposition are well separated..**

255

256 The cooling rate is at most only about 50% of the heating rate. Cooling is passive through heat transfer to ambient air, with a  
257 heat transfer coefficient of approximately  $k=0.024 \text{ W m}^{-1} \text{ K}^{-1}$  in still air.  $k$  increases with radiation cooling, convection and  
258 wind. During glass heating, heat is quickly transferred to the DS from the electrically heated inner layer, while during  
259 cooling the heat transfers from the DS to the air occurs slowly, with a thermal constant  $k$ . This explains the asymmetrical  
260 curve of Fig.7.

261 Outdoor tests carried out in summer at DC (-30°C air temperature) showed a heating rate of  $3^\circ\text{C min}^{-1}$  in still air,  $2.5^\circ\text{C}$   
262  $\text{min}^{-1}$  with  $2.5 \text{ m s}^{-1}$  wind speed, and  $1.8^\circ\text{C min}^{-1}$  with  $5 \text{ m s}^{-1}$  wind speed. In all cases, the cooling rate was approximately  
263  $1.5^\circ\text{C min}^{-1}$ .

264 An outdoor sublimation test (-30°C air temperature, RH=60%, wind speed  $<3 \text{ m s}^{-1}$ ) performed with snow manually spread  
265 on the DS showed that, after applying heating for 10 minutes (up to a DS temperature of -8°C) the sublimation of the  
266 majority of particles (diameter  $<1000 \mu\text{m}$ ) was complete within 20 minutes after turning off the heating, with just a few big  
267 grains (initial diameter  $>1000 \mu\text{m}$ ) still present after 30 minutes.

268 After these tests, the glass heating period was set at 10 minutes (the heating is stopped anyway if the DS temperature exceeds  
269 -5°C to avoid melting of the ice in summer). At the peak of the sublimation period, DS resulted warmer than air of about  $dT$

270 = 20°C. Once the heater is turned off, and after a cooling time of approximately 20 minutes, the DS temperature comes back  
271 to be warmer than the air by only 4-5°C. At this point the "sublimation period" (of approximately 30 minutes) is considered  
272 complete, and ice particles start accumulating again on the DS, with no relevant sublimation, i.e. the "deposition period"  
273 begins (as sketched in fig.7 for the indoor test). At the end of the deposition period, a scan of the DS is carried out, for a  
274 duration of one minute. If no ice particles were detected on the previous scan, the DS heater is not applied, sublimation is not  
275 needed. The effective deposition period depends on the temperature, wind and exposure to the sun in summer. This  
276 uncertainty, combined with occasional wind removal and particulate sublimation (Sect. 3.2) during the deposition period,  
277 prevents the use of ICE-CAMERA for rigorous quantitative precipitation studies.

278 DS surface temperature is actually measured by using a small thermocouple. This measurement implies great uncertainties due  
279 to the radiant warming of the sensor in summer and the difficult thermal coupling with the glass surface. A non-contact  
280 measurement of DS temperature by means of IR sensors would also be uneffective in winter conditions.

281

### 282 **3. Ice particles and the deposition surface.**

#### 283 **3.1 Adhesion of ice particles on the DS.**

284 The adhesion of ice crystals to the smooth DS is caused by two principal reasons: Van der Waals and electrostatic forces.  
285 Eidevåg et al (2020) studied the adhesion of dry snow particles after 90° impact to different wall materials (gravity is a minor  
286 force in this application, and therefore their work applies to any wall orientation). They considered models for normal direction,  
287 tangential sliding, and tangential rolling that account for the adhesive interaction of spherical ice particles (25-275  $\mu\text{m}$   
288 diameter) and their aggregates. The Johnson-Kendall-Roberts (JKR) model for adhesion was used. Their findings showed that  
289 the maximum normal velocity at which spherical ice particles adhere to a glass surface (critical stick velocity) decreases with  
290 decreasing particle diameter. Spherical particles of 100  $\mu\text{m}$  would adhere for speeds less than 0.02  $\text{m s}^{-1}$ . The numerical method  
291 of Eidevåg et al (2020) has been applied in the present work to compute the critical sticking velocity of spherical ice particles  
292 up to 1000  $\mu\text{m}$  diameter. The sedimentation velocity of ice crystals up to 1000  $\mu\text{m}$  was also calculated using the formulation  
293 of Böhm (1989) for DC conditions ( $\text{Tair}=-50^\circ\text{C}$ , air density=1.03  $\text{kg m}^{-3}$ , dynamic viscosity= 1.45E-5  $\text{Pa s}$ ). Particles  
294 impacting on the DS with a sedimentation velocity lower than the critical sticking velocity are immediately captured with  
295 100% efficiency by Van der Waals forces. The simulation shows that only spherical particles smaller than 50  $\mu\text{m}$  in diameter  
296 fall on the DS with a speed smaller than the critical stick velocity, and thus stick immediately. This result does not change with  
297 the different forms of particles, as for these sizes the speed of sedimentation is close to that of Stokes. Particles above 50  $\mu\text{m}$   
298 in diameter have an excess of kinetic energy to bind effectively immediately on the DS. However, the impact surface (DS) is  
299 horizontal, so that the excess kinetic energy is rapidly dissipated in one or more vertical rebounds, until the critical sticking  
300 velocity is achieved (Chokshi et al., 1993). So the adherence of ice crystals to the DS could also be explained by the forces of  
301 Van der Waals alone. Ryzhkin and Petrenko (1997) showed that static charges, naturally transported by ice crystals, increase  
302 adhesion. The electrostatic interaction between the ice and the surface is significantly stronger than the van der Waals forces  
303 at distances greater than the inter-molecular forces. Electrostatic forces are therefore expected to significantly improve the  
304 adhesion of large ice particles to the DS. Once attached to the DS, the weak winds generally observed at DC cannot detach the  
305 particles from the DS. Particulates are protected by the boundary layer (BL) that forms on the DS. The 99% thickness of the  
306 laminar BL (Blasius solution) at the centre of the DS (0.15 m distance from the glass edge) is expected to be 7 mm at -50°C  
307 with a wind speed of 1  $\text{m s}^{-1}$ , decreasing to 2 mm at 10  $\text{m s}^{-1}$ . As a result, the particles deposited on the DS are protected against  
308 the wind.

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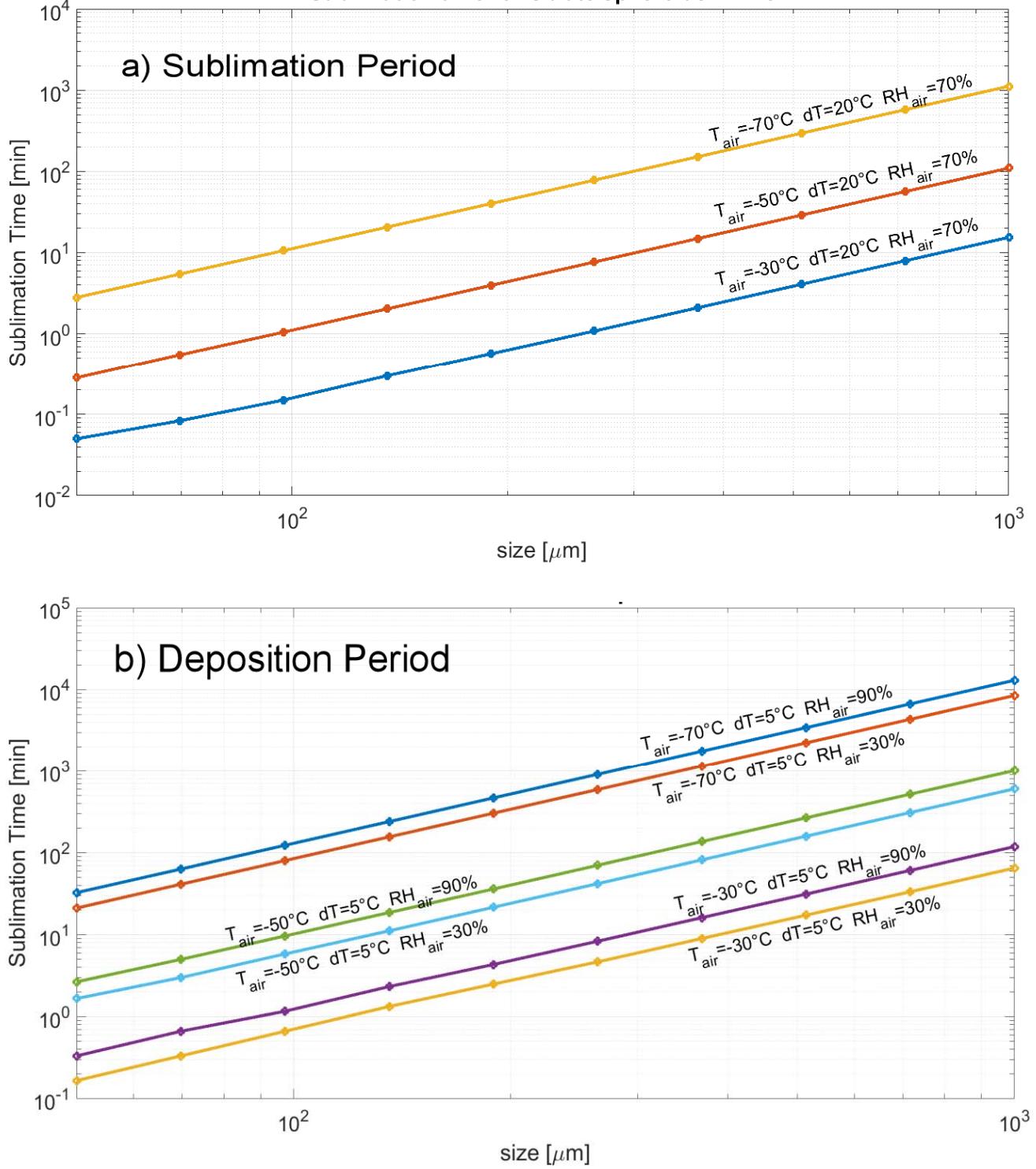
313 **3.2 Sublimation of ice particles.**

314 The DS is always warmer than the surrounding air. This is necessary to eliminate hoar, enabling the device to be used in all  
315 DC conditions. The undesirable effect is the accelerated natural sublimation of deposited particles. A wide range of  
316 experimental and theoretical research efforts has characterized the effects of temperature and super-saturation on ice crystal  
317 growth rates and morphology under conditions relevant to atmospheric processes (for example Lamb and Hobbs, 1971;  
318 Libbrecht, 2005; Libbrecht, 2017). The wide variety of ice crystals found in nature has sparked an interest. Sublimation was  
319 sometimes regarded either as the opposite process, or a less intriguing process, and was less visited in lab studies.  
320 Nelson,(1988) sublimated numerous, 100  $\mu\text{m}$  diameter plate crystals ( $0.1^\circ\text{C} > \text{T} > -18^\circ\text{C}$ , 0.05% to 5% sub-saturation) showing  
321 that the crystals first lost sharp edges, and finally evolved into spheroidal particles, and the aspect ratio remained almost  
322 constant. The sublimation rates were accurately predicted by the diffusion equation with the surface vapour density at the  
323 equilibrium value for a uniform surface temperature. The sublimating crystal reaches a self-preserving shape that is one of the  
324 shape preserving solutions of the diffusion equation. Ham (1959) showed that ellipsoids and thus spheroids preserve shape  
325 during growth and sublimation if the grain surface has a uniform temperature. Jambon-Puillet et al.(2018) also showed  
326 experimentally and theoretically that sublimation first smooths out regions of sharp curvature, leading to an ellipsoid. The  
327 second stage is the sublimation of the self preserved ellipsoid shape. The entire process may be modelled as a vapour diffusion  
328 problem, mathematically equivalent to the resolution of the electrical potential around a charged conductor. Using this analogy,  
329 they provided a mathematical method for simulating the sublimation of the ice particle. The sublimation of the ellipsoid turned  
330 out mathematically simple, and their method was adopted in this work to numerically simulate the second stage of sublimation  
331 of ICE-CAMERA particles.

332 Monodispersed oblate spheroids with an aspect ratio (AR) of 5, in thermal equilibrium with the DS, were assumed in the  
333 simulations as a surrogate for ice plates. The two major spheroid axes coincide with the «diameter» of the oblate spheroid, D.  
334 In the model, D, DS temperature, air temperature and relative humidity with respect to ice (RH<sub>air</sub>) can be changed. The  
335 sublimation time required for full sublimation of a spheroidal ice particle was computed. As sublimation accelerates when the  
336 particle is going to vanish, the time necessary for the complete sublimation is only slightly larger than the time necessary to  
337 reduce the particle to the minimum particle size (D=60  $\mu\text{m}$ ) accepted by ICE-CAMERA image processing. The simulations  
338 assume that the preliminary sublimation of the high-curvature parts of the particle (sharp edges, corners, surface irregularities)  
339 was already completed, so that the calculated time of sublimation must be considered as a lower limit for real-world crystals,  
340 and probably almost one half of the overall duration of sublimation (Jambon-Puillet et al.,2018). Simulations also assume the  
341 thermal equilibrium between the particle and DS, a condition which is not necessarily satisfied on the thermally insulating  
342 glass surface of the DS. Figure 8a shows the total sublimation time with the DS heated  $\text{dT}=+20^\circ\text{C}$  above air temperature  
343 (sublimation period). The humidity resulted irrelevant in this case, and only results for 70% RH<sub>air</sub> are shown. Results show  
344 that at  $-30^\circ\text{C}$  air temperature (summer conditions in DC) complete sublimation can occur within a few minutes after attaining  
345 the DS sublimation temperature, for all particle sizes up to 1 mm. At lower air temperatures, the sublimation time increases:  
346 at  $-70^\circ\text{C}$  (winter temperature in DC), particles smaller than 100  $\mu\text{m}$  in diameter still disappear within 10 minutes, while larger  
347 particles can survive along the sublimation period. Simulations showed that, at  $-70^\circ\text{C}$ ,  $\text{dT}$  should be increased to  $\text{dT}=60^\circ\text{C}$  in  
348 order to ensure the complete sublimation of ice particles up to 1 mm diameter during the sublimation period. This is actually  
349 not possible with the electric heated glass adopted, but could probably be achieved by microwave heating.

350

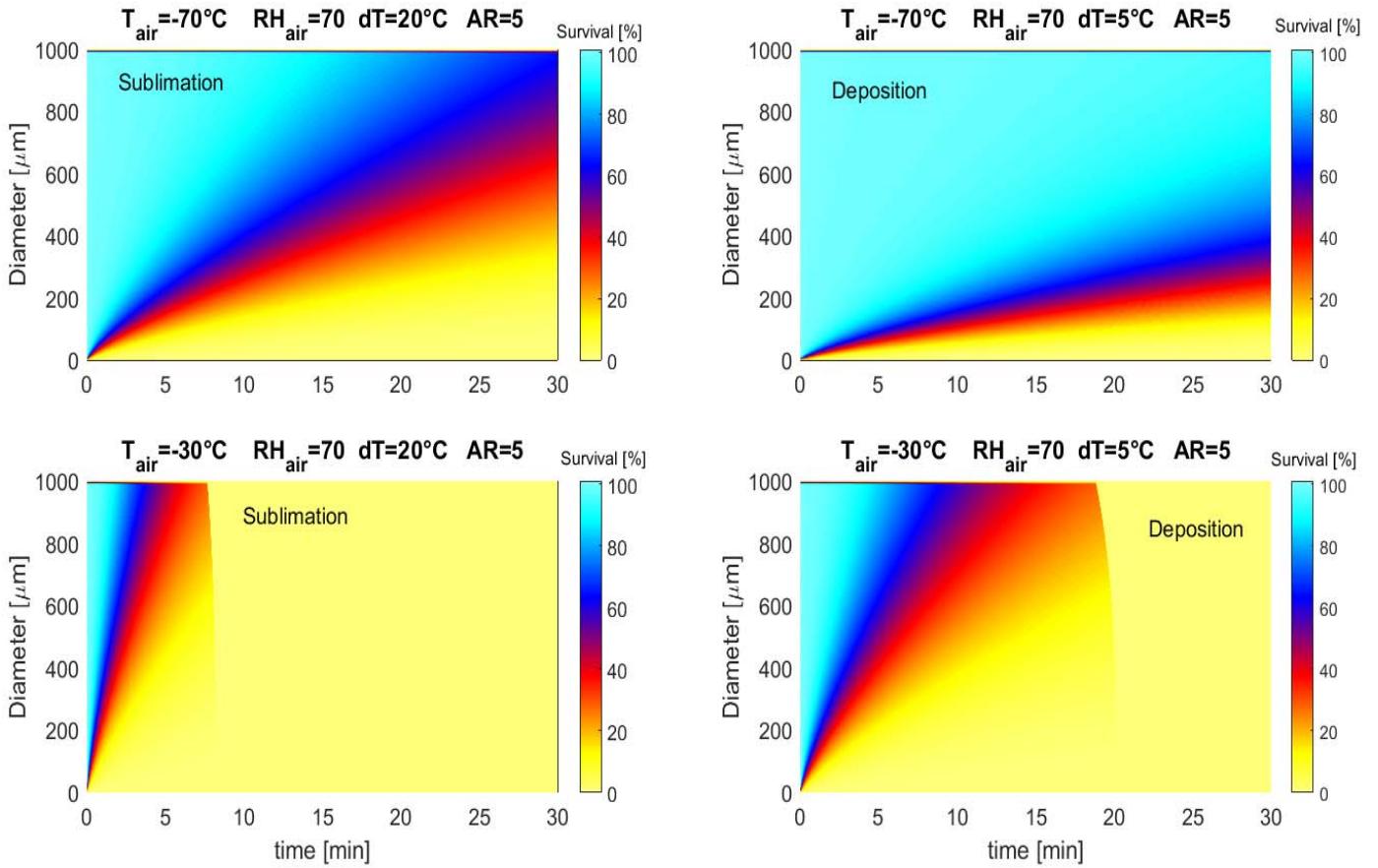
### Sublimation time for Oblate spheroids AR=5



351  
352 **Fig. 8: Sublimation time of monodisperse oblate spheroids at varying air temperatures, with a)  $dT = 20^{\circ}\text{C}$   
353 (sublimation period) and b)  $dT=5^{\circ}\text{C}$  (deposition period).**  
354

355 After the sublimation period, most of the particles previously collected on the DS are sublimated, and a new deposition period  
356 begins. Even during this period, sublimation still acts on ice particles, albeit slowly. Figure 8b shows the sublimation time  
357 expected for monodisperse spheroids during the deposition period. The DS was considered 5 °C hotter than air. As shown,  
358 during the deposition period the relative humidity of air also plays a role, even if secondary. In summer ( $T_{\text{air}}=-30^{\circ}\text{C}$ ),  
359 sublimation can take less than a minute for particles smaller than 100  $\mu\text{m}$  and ten minutes for 300  $\mu\text{m}$  particles. During winter  
360 ( $T_{\text{air}} = -70^{\circ}\text{C}$ ), all particles are expected to survive through the deposition period. As a rule-of-the-thumb, simulation showed

361 that working with  $dT=+5^{\circ}\text{C}$  resulted in an increase of the rate of sublimation by a factor 2-3 compared with a DS in thermal  
362 equilibrium with ambient air ( $dT=0$ ) for the whole range of air temperatures and RHair shown in Fig. 8b.  
363 Results of Fig.8b show that the effective lower limit of ICE-CAMERA particle detection is not limited solely to the resolution  
364 of the optical system and/or image processing software. In summer, particles smaller than  $100\text{ }\mu\text{m}$  may be decimated during  
365 deposition, unless they fall just before scanning. Such small particles dominate diamond dust events. As a result, ICE-  
366 CAMERA, during the summer period, is best suited to the study of cloud precipitation. Nevertheless, visual screening of ICE-  
367 CAMERA images showed only a limited number of small particles revealing signs of partial sublimation, such as rounded  
368 corners, smooth edges, or a spheroidal appearance. Some small plates (observed mainly in winter, when sublimation during  
369 the deposition period is very slow) showed smoothed corners, but it is not clear if this was induced by sublimation or is a  
370 natural feature of these ice grains. Also, even in summer, small DD particles such as plates (with no signs of edge smoothing)  
371 were normally observed (Sec.5.2). It is probable that most particles (other than, probably, pristine plates) never achieve thermal  
372 equilibrium with the DS glass, and that the results of Fig.8 should be considered as the worst case. Also, the sublimation of  
373 the high-curvature parts of the particle prior to assuming the spheroidal form (Jambon-Puillet et al.,2018) could take much  
374 more time than the sublimation time calculated here for the spheroid. A series of consecutive DS scans at fixed air temperatures  
375 is needed to measure the effective sublimation rate of small particles in deposition conditions ( $dT=+5^{\circ}\text{C}$ ).  
376 When a polydisperse particle population is deposited on the DS instead of monodisperse particles, a more complicated  
377 sublimation picture arises, because small spheroidal particles, shrinking, are continuously replaced in the size distribution by  
378 sublimating, initially bigger ones. An initial uniform particle size distribution (PSD) of the oblate spheroids (AR=5) was  
379 assumed with diameters between  $D=1$  and  $2000\text{ }\mu\text{m}$  for the simulations. The evolution over time (1 sec resolution) of the PSD  
380 was calculated (Fig. 9) in terms of particle survival (the ratio between the actual number of particles in a certain size bin and  
381 the initial number in the same bin). No vapour competition between ice particles was taken into consideration in the  
382 simulations. Results are similar to those of monodisperse particles (Fig. 8), with a slightly longer time of sublimation for  
383 polydisperse particles compared to monodisperse particles of the same size. Results for an air temperature of  $-70^{\circ}\text{C}$  confirm  
384 that most particles larger than  $500\text{ }\mu\text{m}$  survive, throughout the DS sublimation period ( $dT = 20^{\circ}\text{C}$ ), longer than 30 minutes.  
385 This means that sublimating by heating the glass is quite inefficient for large particles in winter. During the deposition period,  
386 at  $-70^{\circ}\text{C}$  losses for sublimation are scarce and limited to particles smaller than  $200\text{ }\mu\text{m}$ . Consequently, double counting of the  
387 same particle ( $D>500\text{ }\mu\text{m}$ ) is possible in two consecutive ICE-CAMERA scans in the cold DC winter. At  $-30^{\circ}\text{C}$  air temperature  
388 (summer) the heating of the DS with  $dT=20^{\circ}\text{C}$  leads to the sublimation of most particles up to  $2\text{ mm}$  diameter within 5 minutes.  
389 On the other side, during the deposition period particles smaller than  $500\text{ }\mu\text{m}$  can undergo sublimation over a period of just 10  
390 minutes in summer, thus limiting the effective period of deposition before a scan. As with monodisperse particles, this  
391 introduces bias in the summer because many small particles (typical of DD) can be removed before they are measured.

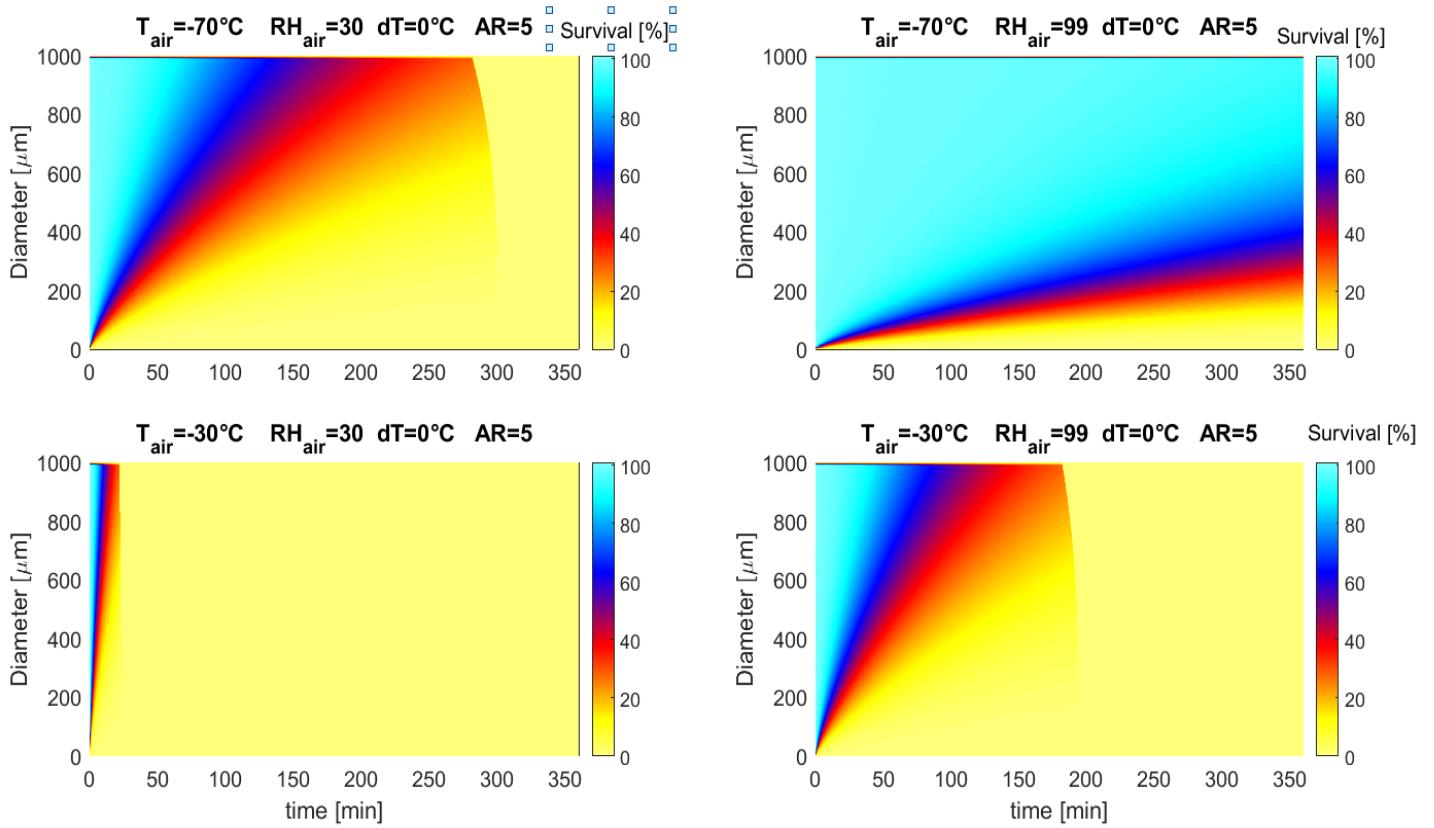


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393 **Fig.9: Evolution of a originally uniform PSD of ice spheroids (D=0-2000  $\mu\text{m}$ , AR=5) under different atmospheric**  
 394 **conditions. (RH<sub>air</sub> is a secondary factor affecting the results, shown here for RH=70%).**  
 395 **Left: sublimation period; right: deposition period. Top: winter, bottom: summer.**

396

397 Even if these results could be disappointing for interpreting ICE-CAMERA data, the same problems affect the actual method  
 398 of observing precipitation in DC: collecting and observing (every 24 hours) the ice particles deposited on flat surfaces  
 399 ('benches') is affected by the same problem as collecting particles on the ICE-CAMERA DS with  $dT=0$ . Fluctuations in  
 400 relative humidity over 24 hours result in sublimation and regrowth of particles on the "benches" in an almost unpredictable  
 401 manner. Fig. 10 shows the expected sublimation time for particles (with the same PSD of Fig. 9) placed on 'benches' (or  
 402 ICE-CAMERA DS) in equilibrium with air ( $dT=0$ ) for extreme, sub-saturated conditions: winter  $T_{\text{air}}=-70^\circ$  ( $\text{RH}_{\text{air}}=30\%$  and  
 403 99%), and summer  $T_{\text{air}}=-30$  ( $\text{RH}_{\text{air}}=30\%$  and 99%). The PSD evolution is computed with a resolution of 1 sec for a total  
 404 period of 6 hours. The results show that sublimation also works in winter and with almost saturated air (99%  $\text{RH}_{\text{air}}$ ), leading  
 405 to a complete loss of small particles ( $D<200 \mu\text{m}$ ) in a few hours. In summer conditions and 30%  $\text{RH}_{\text{air}}$  sublimation happens  
 406 much more quickly, with the disappearance of all particles up to 2000  $\mu\text{m}$  in 30 minutes. With  $\text{RH}_{\text{air}}=99\%$ , sublimation  
 407 removes all particles in just a few hours in summer. In presence of wind and dry air, sublimation rate could even increase, as  
 408 observed by Grazioli et al., 2017 in coastal areas. These simulations all refer to sub-saturated conditions: in the case of a  
 409 'bench' in thermal equilibrium with super-saturated air, hoar form on the surface, with a possible confusion with precipitation.



410

411 **Fig.10: Evolution of a originally uniform PSD of ice spheroids ( $D=0-2000 \mu\text{m}$ , AR=5) under different atmospheric**  
 412 **conditions. The DS (or ‘bench’) is in thermal equilibrium with air ( $dT=0$ ). Top: winter, bottom: summer**

413

414

#### 415 **4. Data processing.**

416

##### 417 **4.1 Image processing.**

418 ICE-CAMERA is not just designed to take photographs of ice particles, but to provide automatic morphometry and  
 419 classification of polar precipitation. This was accomplished through the use of image processing and machine learning  
 420 techniques. The process is divided into two parts: segmentation and measuring, and classification of ice crystals.

421

###### 422 **4.1.1 Image segmentation and measurement of ice particles.**

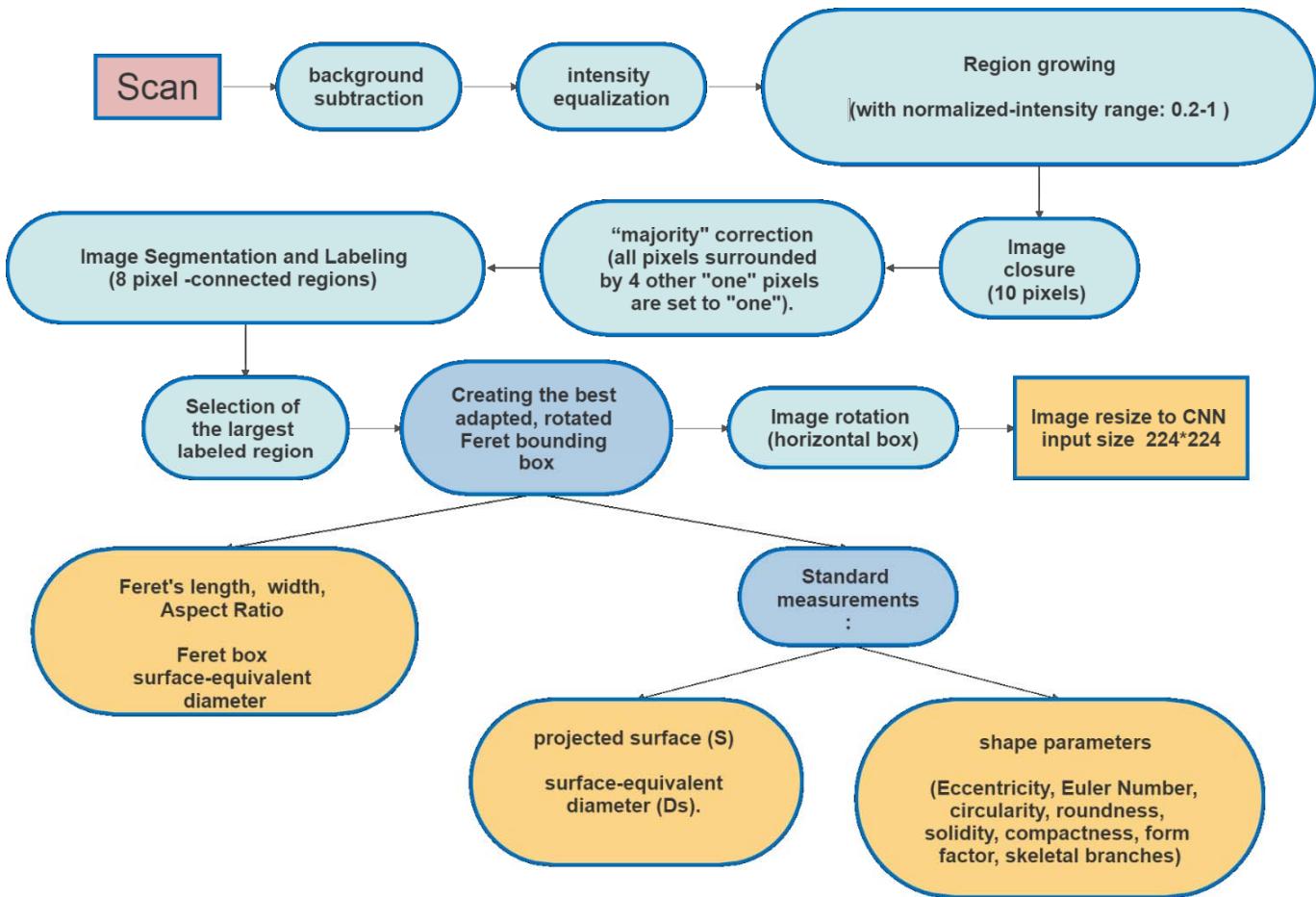
423 After acquisition, the raw ICE-CAMERA scans are segmented, using MATLAB software, to isolate all detected particles. The  
 424 process follows the workflow of Fig. 11. Refer to Pratt (2007) for image-processing nomenclature, to Walton, 1948 for Feret  
 425 measurement, to Russ and Brent Neal (2017) for the nomenclature of standard shape parameters such as Eccentricity, Euler  
 426 Number, circularity, roundness, solidity, compactness, form factor, and number of skeletal branches. The normalized central  
 427 moments  $f1 \dots f7$  were also computed as described by Hu, (1962).

428 The Aspect Ratio (AR) is defined as Feret’s length/ Feret’s width. The Feret-box surface-equivalent diameter ( $D_f$ ) is defined  
 429 as the diameter of the circle of the same area as the Feret bounding box, while the surface-equivalent diameter ( $D_s$ ) is  
 430 defined as the diameter of the circle having the same area as the segmented ice grain. The main steps of Fig. 11 are visually  
 431 summarized in Fig.12 for a rimed, columnar particle.

432

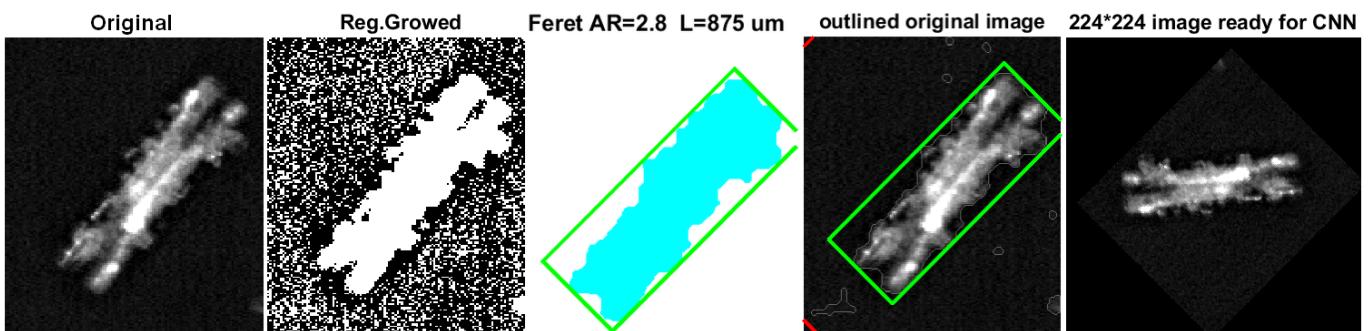
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Fig. 11: The image-processing flow chart.



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Fig. 12: The original image (in this case a rimmed column) is segmented using ‘region-growing’. The projected particle area (clear blue) is calculated. The bounding box is determined (green) and the Feret length and width measured.

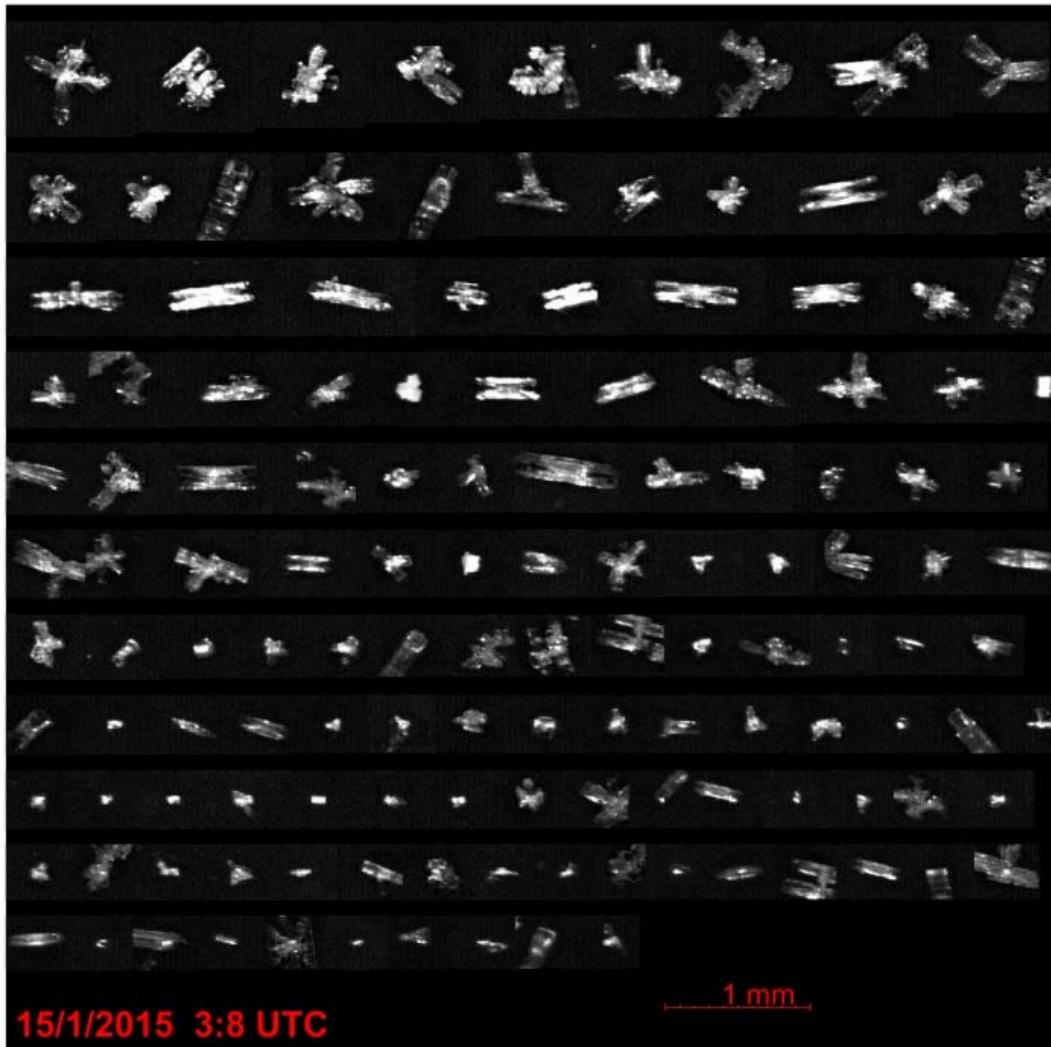
The image is finally rotated to have the major axis horizontal, re-scaled, and resized to the CNN input size.

452

453 **4.1.2 Summary-image of detected particles.**

454 The bounding boxes of all individual ice particles detected in a scan are sorted by Feret length, and reassembled in a summary-  
 455 image collecting all segmented particles (Fig 13). Each particle is also associated with a numerical record containing the  
 456 coordinates of its bounding rectangle on the summary-image, shape parameters, time of acquisition and local weather data. In  
 457 this way, the re-analysis of the summary-image is possible instead of re-processing the original, large image. The original  
 458 image is ultimately removed.

459

460  
461 **Fig. 13: Example of a summary-image for a single scan.**  
462  
463464 **4.1.3 Limitations and uncertainties in detecting and sizing ice particles.**

465 1) The total number of particles measured is actually limited to 2000 per scan, as a result of MATLAB memory  
 466 limitations. Extra particles are not treated.

467 2) Particles below  $3600 \mu\text{m}^2$  in bounding-box surface, (equivalent to approximately  $60 \mu\text{m}$  in diameter for a spherical  
 468 particle) are not preprocessed (smaller particles could be detected, but most have a seemingly circular shape due to  
 469 low pixelation or poor focus).

470 3) The segmentation becomes difficult when overlapping particles or aggregates of particles are present. In such  
 471 situations, double counting of the same particle may occur in up to 12% in a scan in the presence of an intense  
 472 precipitation event. The same particles can in fact fall inside different segmented areas of the image, because of the

473 lack, on the original image, of defined boundaries between particles. The process of "region growing" which leads  
474 to segmented particles can actually start, independently, from several bright ("seed") regions located in different parts  
475 of the image of the overlapping particles. The 'region growing' processes can then propagate through the overlapping  
476 particles leading to several 'copies' of the same, segmented image. This unwanted effect could be prevented by  
477 looking for similar copies of the same segmented image, but this method was not implemented at DC due to limited  
478 PC resources. Overlapping particles are normally classed by the CNN algorithm as "clusters". A few occasional  
479 arrangements of three or more overlapping columns are sometimes mistaken for single plates. The Feret measurement  
480 of these particles is meaningless. At DC this situation occurs only after heavy cloud precipitation, a relatively rare  
481 event.

482 4) Multiple counts of the same particle also occur for non overlapping particles when multiple bright spots exist within  
483 the same particle. As in 3), the region-growing process can start independently from several 'hotspots', leading to  
484 false copies of the same segmented particle. This effect could potentially be avoided by comparing segmented images  
485 and deleting copies, but this method was not implemented at DC.

486 5) Particles close each other in the original image could be segmented into a single particle by region-growing and thus  
487 misclassified.

488 6) In the case of defocused images, the particle shapes are all close to a fuzzy, round or elliptical shape, which can cause  
489 a misclassification into irregular particles, spheroidal particles or plates. ICE-CAMERA images dominated by this  
490 type of particles are normally eliminated during a preliminary manual screening. Also, a few big particles in summer  
491 resulted rounded by partial sublimation. A few images containing only rounded or "spheroidal" particles of 500  $\mu\text{m}$   
492 diameter or greater were collected during the warmest part of summer, and were manually discarded before the  
493 statistical data analysis.

494 7) Needles and hexagonal plates (typically small, see Fig.23) may be very bright in ICE-CAMERA images due to  
495 enhanced light diffusion at preferred angles. For the same reason, hollow columns sometimes have a shiny spot in the  
496 middle. In the case of needles, this effect can reduce the apparent aspect ratio, as the width is apparently increased  
497 by the scattered light saturating the camera. For plates, the bright specular reflection blurs sometimes the polygonal  
498 contour, especially in the case of small plates.

499

#### 500 **4.2 Automated classification of ice particles.**

501 An initial attempt at automatic classification of ICE-CAMERA segmented images was made in 2014 using shape factors. This  
502 kind of technique has also been used by others (e.g. Lindqvist et al., 2012) for attempting the classification of ice particles. In  
503 the case of ICE-CAMERA this approach resulted extremely unreliable. A much more promising approach was offered after  
504 2015 by the rapid development of transfer learning and convolutional neural networks (CNN) (Le Cun et al., 2015;  
505 Schmidhuber, 2014). Xiao et al. (2019) successfully applied deep transfer learning to ice particle images obtained with airborne  
506 Cloud Particle Imagers (CPI). The CNN approach has added much value to ICE-CAMERA because a reliable classification  
507 of ice particles into simplified classes became possible. The CNN used for the ICE-CAMERA particle classification is  
508 "GoogleNet" (Szegedy et al. 2015), a variant of the Inception network, a deep convolutional neuronal network developed by  
509 Google scientists. GoogleNet is a type of convolutional neural network based on the Inception architecture. It utilises Inception  
510 modules, which allow the network to choose between multiple convolutional filter sizes in each block. The GoogleNet  
511 architecture consists of 22 layers (27 layers including pooling layers), and part of these layers are a total of 9 inception modules.  
512 In this work, GoogleNet was used in MATLAB R2020b environment. The GoogleNet CNN, pretrained on the ImageNet  
513 dataset (Deng et al. 2009), was used, with its final, fully connected layer changed to size 14. The input layer of the GoogleNet  
514 architecture requires images of size 224 x 224.

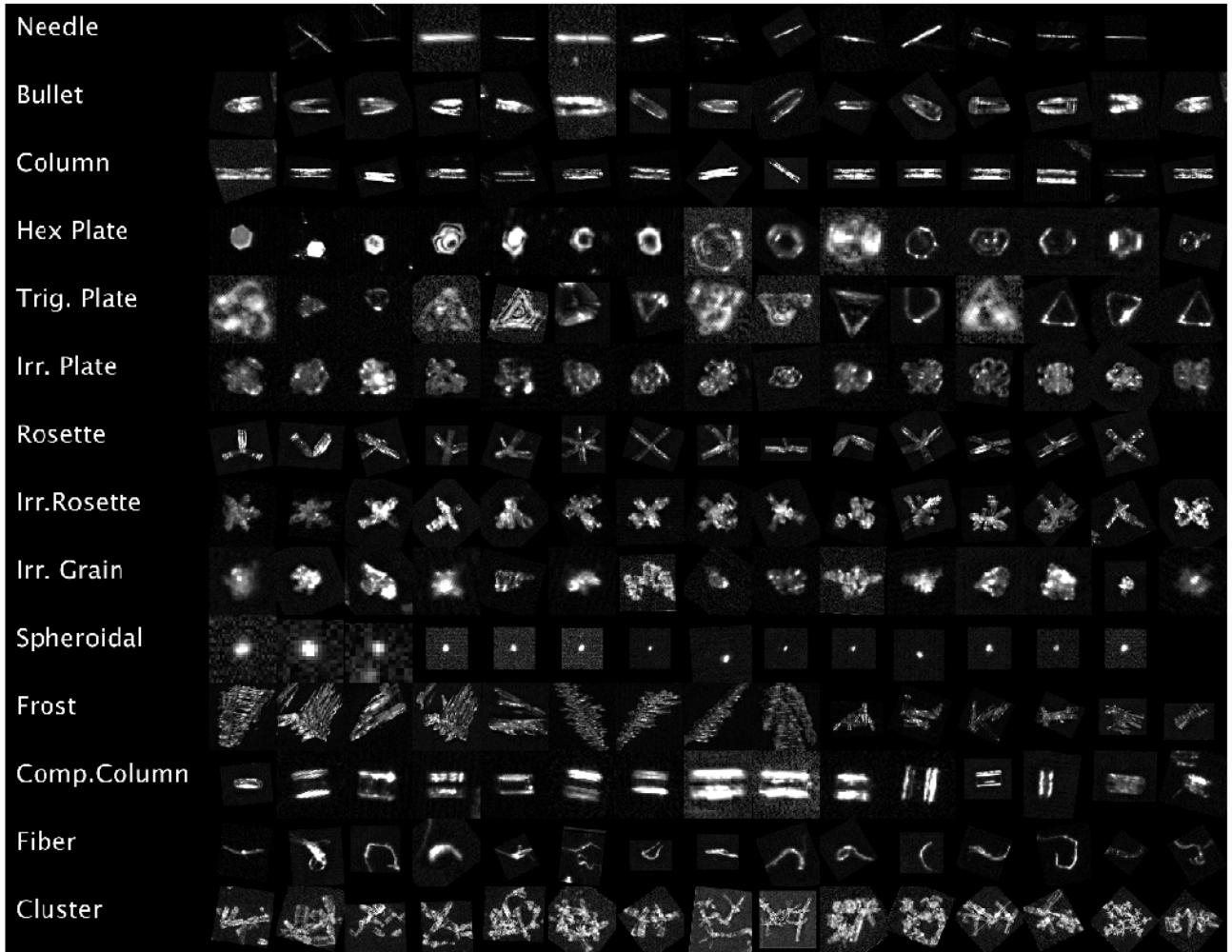
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516 **4.2.1 The CNN classification classes.**

517 Low temperatures and humidity on the high Antarctic plateau reduce the diversity of ice particle shapes. This is observed on  
 518 the field at DC, at South Pole station (Lawson et al., 2006), and suggested by review works such as Bailey and Hallett (2009).  
 519 Following an initial survey of the ICE-CAMERA image database, a set of 14 types of particles was selected, as shown in Fig.  
 520 14. When choosing the 14 classes, I assumed that shapes easily recognizable by a human operator could also be easily  
 521 recognizable by a CNN.

522

523



524

525 **Fig. 14: A sample of ICE-CAMERA images of the 14 classes of ice particles used to train the CNN.**

526

527 In the following scheme I tried to fit the classes chosen for ICE-CAMERA with the classification scheme of the ice particles  
 528 of Kikuchi et al. (2013), an updated version of the original classification of Magono and Lee (1966).

529

530 -**Needles:** covering the classes C1a,C1b,C3d (Kikuchi et al. 2013)531 -**Bullets:** covering the C4b-C4c classes.532 -**Columns:** columns covering classes C2a, R2b, C3a, C3b.533 -**Hexagonal plates:** covering classes P1a, P1b, P1c, P4f, G2a, G3a, CP3f, CP3d.534 -**Trigonal plates:** covering the class G2b.

535 **-Irregular plates:** plate-like particles with irregularities, riming, overgrowing plates, etc. But keeping a basic hexagonal  
536 shape, covering P6a, P6b, P7a, CP6d, R1b, R2b, R2c, R3a, G4b.

537 **-Rosettes:** bullet-rosettes or column-rosettes, with a minimum of two branches, covering C2c, C3e, C4d

538 **-Irregular rosettes:** rosettes with irregularities, riming, but preserving the typical stellar outline of rosettes. Covering  
539 classes P7a,P7b,CP2d ,CP4c,CP5a,CP6e,CP6f,CP6g,R1d

540 **-Irregular grains:** covering CP3e, CP5a, CP6d, G4c,G4a,I3a,I2a,I1a,H1a,H1b

541 **-Spheroidal:** particles with spheroidal or spherical appearance, covering H1a, H1c. (Large particles with D>600μm)  
542 detected as 'spheroidal' in DC are usually artifacts caused by defocused images and are not considered in the statistical  
543 analysis.

544 **-Compact columns:** short columns covering classes G1a, C3a

545 **-Clusters of particles:** covering A1a, A3a, H2a, H1b, P8b, CP3e, CP5a, CP6h

546 **-Frost:** frost formed on the DS CP7,CP8,CP9

547 **-Fibers:** non-volatile fibrous material (from local human activities, Styrofoam particles, textile particles, dust, etc)

548

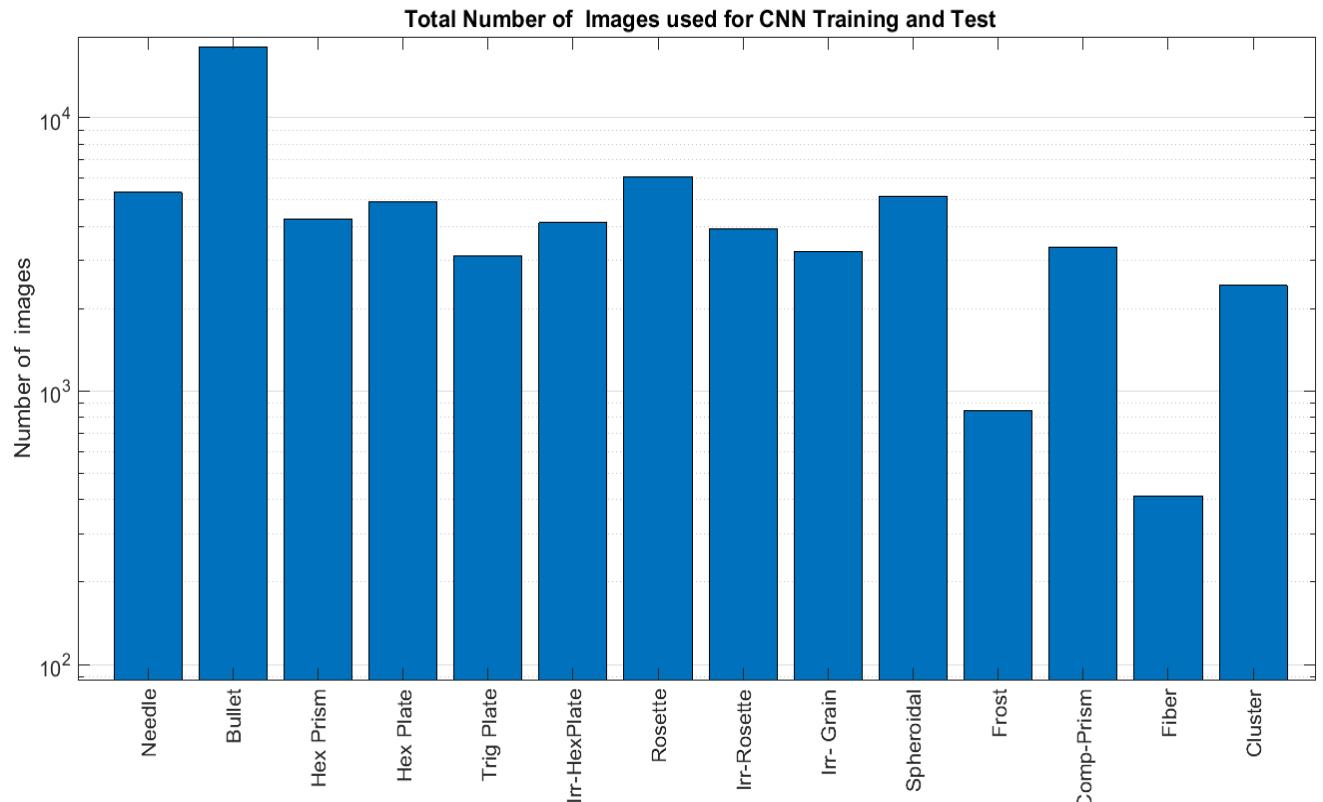
549 The last two classes are not considered in the statistical analysis of ICE-CAMERA data: they are just used to detect occasional  
550 frost formed on the DS in case of super-saturation, and man-made, non-evaporable (thus persisting on the DS) materials.  
551 Uncommon ice particle typologies present at Concordia were not considered in the present work. Trigonal plates have been  
552 included, although they are rare, simply because they are seemingly easy to detect with CNN.

553

554 **4.2.2 The training dataset.**

555 For the training of a first CNN, a set of 5500 ICE-CAMERA segmented images of single particles, sampled randomly from  
556 the 2014-2017 ICE-CAMERA database, have been manually sorted into 14 image data stores, corresponding to the 14 classes.  
557 Fourteen of the computer keyboard keys were marked with the symbols of the 14 classes in order to expedite the manual  
558 classification of the initial training dataset. These images were used for a first CNN training. 10% of the images were dedicated  
559 to validation, 10% for testing, and the remaining 80% for training. The first CNN was used for the classification of the ICE-  
560 CAMERA dataset for the years 2014 to 2017. In addition to the classification, the individual crystal images were also sorted  
561 and stored in 14 folders, according to the CNN classification. Selected images from these folders were manually reclassified  
562 into all 14 classes (when misclassified by CNN) and added to a second CNN training dataset. Also misclassified images were  
563 thus re-labeled and used in the training dataset for the new CNN. In this way, a potential positive bias of the confusion matrices  
564 due to the exclusion of misclassified images in the new training dataset was avoided. In the selection of the new training  
565 images, care was taken to ensure a balanced number of training images in the 14 classes. The updated image dataset was  
566 finally divided into validation (10%), test (10%) and training (80%) datasets for training a second CNN. This process was  
567 repeated three times to expand the training database and thus improve the overall precision of the CNN classifier.  
568 Figure 15 shows the final number of training and test images selected for each class. The total number of images used for the  
569 training was 81800. Trigonal plates were rare, and their number in the training dataset was thus artificially augmented by  
570 duplicating the training images, in order to avoid their absence in the small (64-images) training mini-batches.

571



572

573

**Fig. 15: The final number of images used for CNN training + validation + test.**

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#### 577 **4.2.3 CNN training details.**

578 To meet Google's input requirements, all images of single particles were resized to 224\*224 pixels. In the training process,  
 579 'data augmentation' was applied to the original dataset. Artificially 'augmenting' the image dataset has been shown to be  
 580 effective in CNN training (Shorten and Khoshgoftaar, 2019). Images inside each mini-batch are automatically, randomly  
 581 'augmented' in order to reduce CNN overfitting. The following transformations were used in augmentation:

582     • X, Y reflection

583     • random X, Y translations  $\pm 30$  pixels

584     • Random scaling 80-120%

585       Other changes such as rotation have not been introduced since the ICE-CAMERA images to be classified are  
 586 typically oriented horizontally by the image processing procedure (e.g. Fig. 13)

587 The following learning options were utilized in GoogleNet training:

588

589 Solver: stochastic gradient descent with momentum (SGDM)

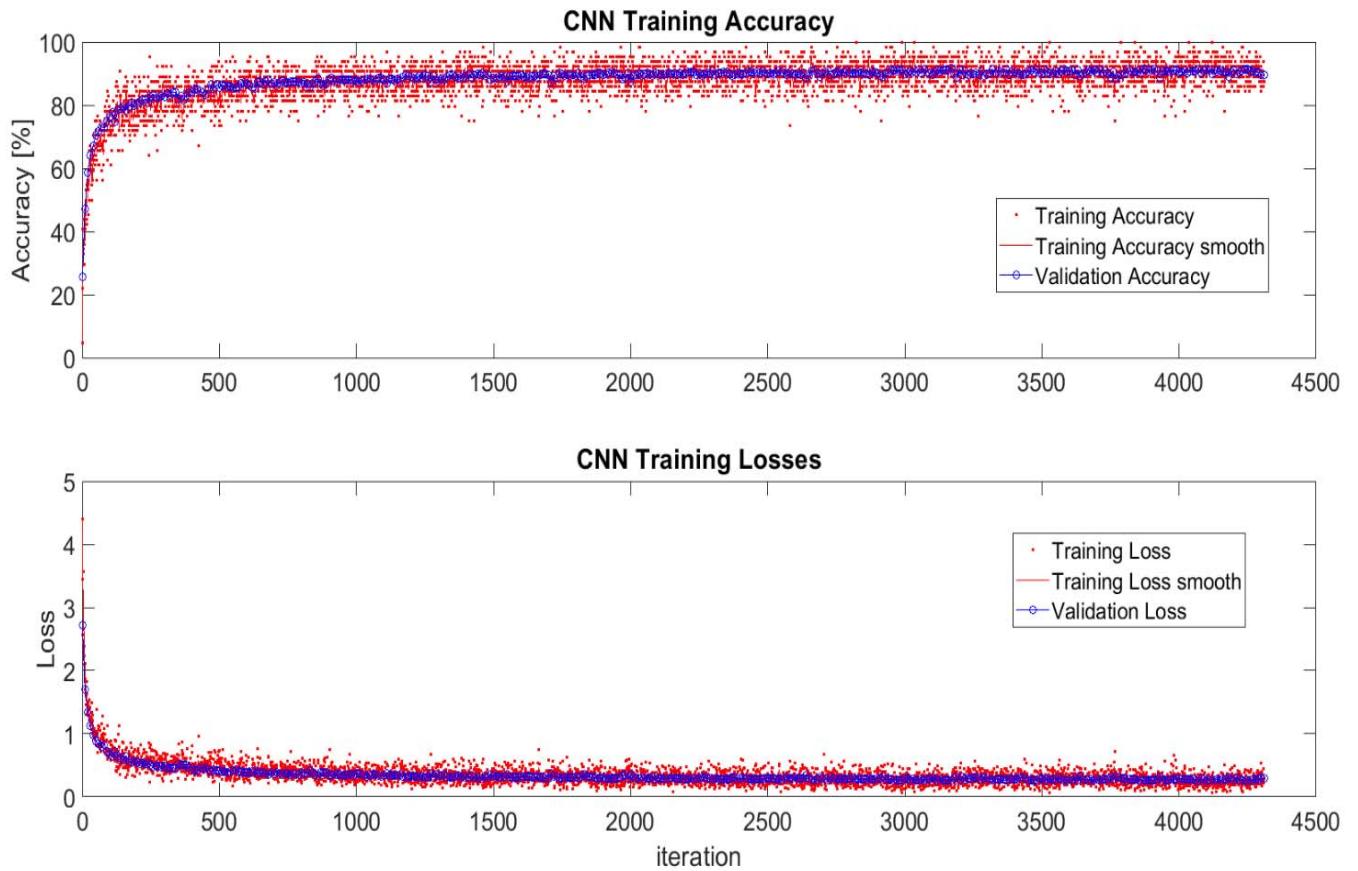
590 activation: softmax

591 Number of Epochs=5

592 Learn Rate=0.001

593 Batch Size=64

594 L2 weight regularization factor=0.005  
 595 Validation frequency= every 30 iterations  
 596 Shuffle of the dataset at every epoch  
  
 597 The evolution of the CNN training in terms of accuracy and losses is presented in Fig. 16. The validation line closely tracks  
 598 the training line, showing the absence of overfitting.



599  
 600 **Fig. 16: Evolution of the CNN training.**

601  
 602 **4.2.4. Testing the CNN classifier.**  
 603 CNN's performance test results are summarized in confusing matrix graphs like Fig. 17a. Each row corresponds to a  
 604 predicted class (Output Class) and each column corresponds to a true class (Target Class). Diagonal cells refer to correctly  
 605 classified observations. Off-diagonal cells are improperly classified observations (red color markings increasing  
 606 misclassification). The column on the far right of the plot shows the percentages of all the examples predicted to belong to  
 607 each class that are correctly and incorrectly classified (positive predictive value and false discovery rates, respectively). The  
 608 row at the bottom of the plot shows the percentages of all the examples belonging to each class that are correctly and  
 609 incorrectly classified (true positive rate and false negative rate, respectively).

610

611

612 **4.2.5 Accuracy of the classifier.**

613 In the column-normalized summary (Fig. 17a), the percentages along the i-th column shows the probability (P) of a "true"

614 particle in class i-th being classified in each of the 14 output classes.

Predicted Class	Spheroidal	99.8%			0.1%								1.2%		99.1%	0.9%	
	Bullet		98.7%	0.3%		1.4%	3.3%		0.2%		5.7%		7.7%	14.6%	0.2%	95.1%	4.9%
	Needle			98.6%		0.1%								1.9%		99.7%	0.3%
	Trig. Plate				96.4%	0.3%	0.5%		0.5%			0.2%	1.9%		0.2%	97.4%	2.6%
	Rosette		0.0%	0.2%		89.5%	0.2%	0.3%		5.5%	0.9%		1.9%	2.5%	0.5%	93.4%	6.6%
	Comp.column		0.5%	0.2%		0.1%	89.3%	0.7%	0.7%			0.2%		4.5%	1.0%	88.4%	11.6%
	Cluster					0.3%		89.1%	0.3%	2.2%	6.6%	1.4%		0.2%	0.7%	89.1%	10.9%
	Hex Plate	0.2%			1.8%	0.1%	2.1%	1.6%	89.1%			9.7%	1.9%		6.2%	84.7%	15.3%
	Irr.Rosette					2.9%	0.2%	1.0%	0.2%	85.7%	0.9%	2.9%	3.8%	0.2%	0.7%	89.6%	10.4%
	Frost					0.1%				0.4%	84.0%			0.4%	1.2%	89.9%	10.1%
	Irr. Plate				1.5%		0.2%	6.9%	8.3%	5.3%	0.9%	83.9%		0.2%	10.4%	74.4%	25.6%
	Fiber					1.6%			0.2%				80.8%			76.4%	23.6%
	Column		0.7%	0.8%		2.6%	3.6%				0.9%			77.5%		87.8%	12.2%
	Irr. Grain				0.3%	0.8%	0.5%	0.3%	0.7%	0.8%		1.7%			77.5%		92.1%

99.8%	98.7%	98.6%	96.4%	89.5%	89.3%	89.1%	89.1%	85.7%	84.0%	83.9%	80.8%	77.5%	77.5%	
0.2%	1.3%	1.4%	3.6%	10.5%	10.7%	10.9%	10.9%	14.3%	16.0%	16.1%	19.2%	22.5%	22.5%	

a)

Spheroidal Bullet Needle Trig. Plate Rosette Comp.column Cluster Hex Plate Irr.Rosette Frost Irr. Plate Fiber Column Irr. Grain True Class

Predicted Class	Needle	99.7%			0.2%								0.2%		99.7%	0.3%	
	Spheroidal		99.1%			0.2%	0.8%								99.1%	0.9%	
	Trig. Plate			97.4%		0.5%	0.3%			0.5%		0.8%	0.3%	0.3%	97.4%	2.6%	
	Bullet	0.1%			95.1%	0.5%	0.0%	0.3%		0.6%	3.3%	0.0%	0.2%		95.1%	4.9%	
	Rosette	0.1%			0.1%	93.4%	0.3%	0.1%	3.7%	0.1%	0.1%	1.8%		0.1%		93.4%	6.6%
	Irr. Grain		0.3%			1.8%	92.1%		1.2%	0.3%	0.6%		1.2%		2.6%	92.1%	7.9%
	Frost					1.0%	5.1%	89.9%	2.0%			2.0%				89.9%	10.1%
	Irr.Rosette					4.7%	0.6%	0.2%	89.6%	0.6%	0.2%	0.2%	0.2%	0.4%	3.2%	89.6%	10.4%
	Cluster					0.7%	1.0%	2.3%	3.6%	89.1%		0.3%	0.7%		2.3%	89.1%	10.9%
	Comp.column	0.2%			2.8%	0.2%	0.9%			0.5%	88.4%	5.7%	0.9%		0.2%	88.4%	11.6%
	Column	1.1%			3.4%	4.3%		0.2%			3.2%	87.8%				87.8%	12.2%
	Hex Plate		0.2%	1.1%		0.2%	3.9%			0.8%	1.4%		84.7%	0.2%	7.7%	84.7%	15.3%
	Fiber					21.8%						1.8%	76.4%			76.4%	23.6%
	Irr. Plate			1.0%			7.2%	0.2%	4.5%	3.6%	0.2%	0.2%	8.8%		74.4%		74.4%

b)

98.6%	99.8%	96.4%	98.7%	89.5%	77.5%	84.0%	85.7%	89.1%	89.3%	77.5%	89.1%	80.8%	83.9%	
1.4%	0.2%	3.6%	1.3%	10.5%	22.5%	16.0%	14.3%	10.9%	10.7%	22.5%	10.9%	19.2%	16.1%	

615

616

Fig. 17: Confusion plot of the CNN : a) column-normalized, b) row-normalized

617 Reading the columns of Fig 17a from left to right, the accuracy of the CNN in properly classifying a particle belonging to the  
618 i-th true class (bottom row) can be assessed. The results are summarized below:

619

620 -Good accuracy ( $P > 90\%$ ) in identifying needles, spheroidal, bullets, trigonal plates.

621 -Compact columns are misclassified into columns (3% of the time) and bullets (3% of the time).

622 -Hexagonal and irregular plates are confused approximately 10% of the time. This is expected since the edges of the plates  
623 (usually small) are sometimes blurred in the image.

624 -Irregular rosettes are misclassified in 5% of cases as pristine rosettes and in 5% of cases as irregular plates.

625 -Irregular plates are confused with hexagonal plates 10% of times.

626 -Irregular grains are sometimes mistaken with irregular plates (10%) and hex plates (6%).

627 -Columns are misclassified as bullets 15% of the times.

628 The three-dimensional structure of the ice particles is lost in the ICE-CAMERA images, so that some thick ice forms such as  
629 C4a, P1b, G3b, CP1a, etc. (Kikuchi et al, 2013), if any, are likely to be misclassified by this CNN.

630 A different view to read the CNN test is the row-normalized summary of the confusion matrix (Fig. 17b).

631 Percentages along the i-th row now show the probability for a particle classified into the i-th class to effectively belong to  
632 each of the 14 true classes. Reading the rows of Fig. 17b from top to bottom, results are:

633 -Particles classified as needles, spheroidal, trigonal plates, bullets, pristine rosettes and irregular grains effectively ( $P > 90\%$ )  
634 belong to their class.

635 -Particles classified as irregular rosettes have a 5% chance of being regular rosettes

636 -Particles classified as compact columns have a 6% chance of being columns.

637 -Particles classified as columns have a 4% chance of being a 2-branch rosette and 3% of being bullets or compact columns.

638 -Particles classified as pristine plates have a 4% chance of being irregular grains.

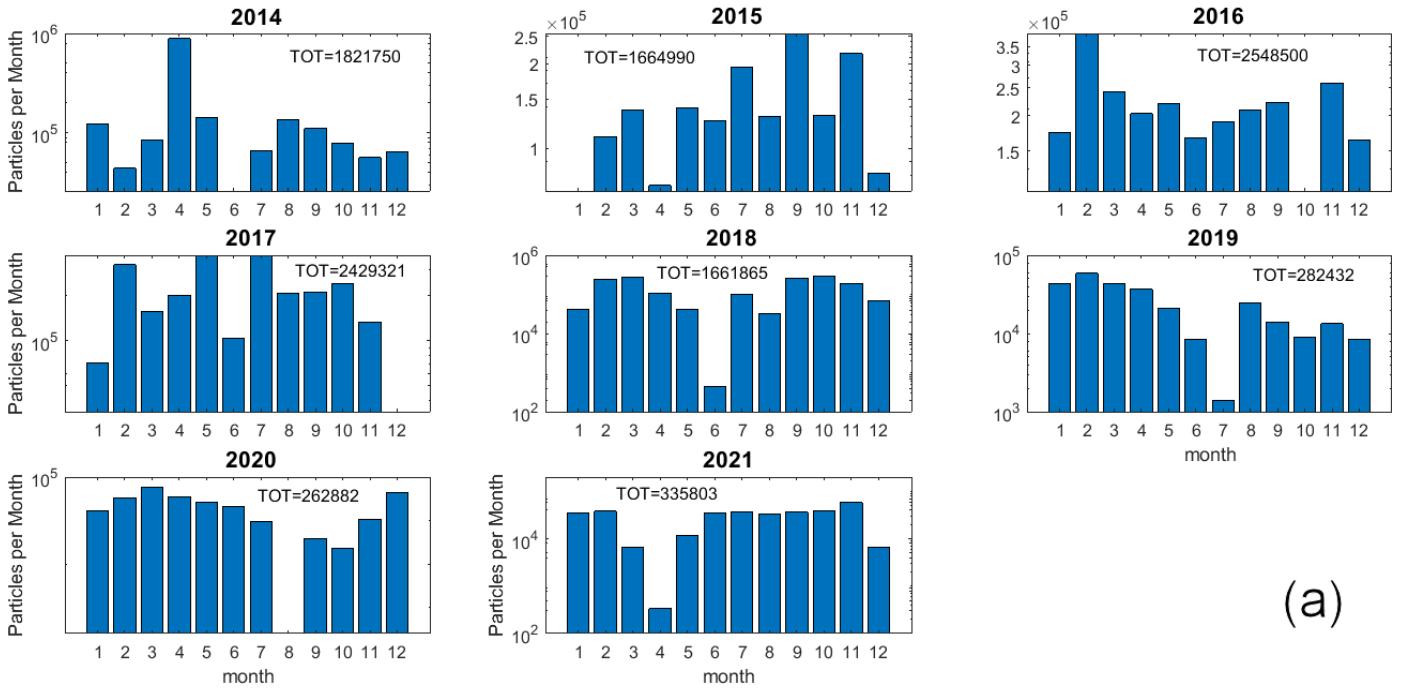
639 -Particles classified as irregular plates have a 7% chance of being irregular grains, 9% hex. plates and 5% of being irregular  
640 rosettes

641 **5. Results.**

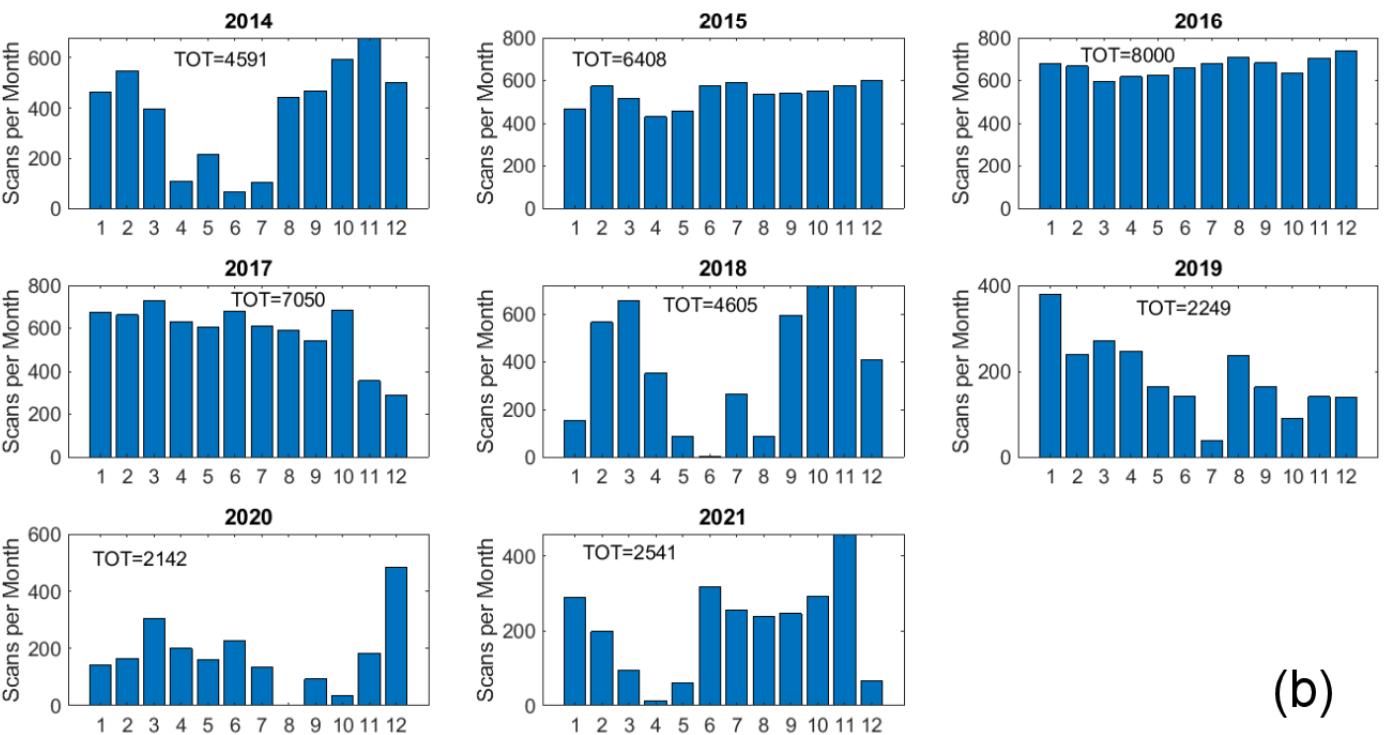
642 **5.1 Overview of ICE-CAMERA dataset.**

643 From January 2014 to December 2021, ICE-CAMERA has segmented a total of 11.007.543 particles. This gross count includes  
644 particulates successively rejected for the statistical analysis. Some whole scans were eventually ignored because of poor focus,  
645 sledge motor failures, or the presence of layers of snow or frost. Individual particles were omitted from the analysis due to  
646 their small size or defocus. The distribution of the number of particles observed during the months is shown on Fig. 18a. The  
647 number of scans per month is shown in Fig. 18b. Under optimal conditions, one scan per hour is planned, with a typical total  
648 of 740 scans per month. Some months, problems with ICE-CAMERA, focusing, or processing software resulted in the small  
649 number of scans or particles observed. In most other cases, scans were not recorded when fewer than ten particles were detected  
650 on the DS.

651



(a)



(b)

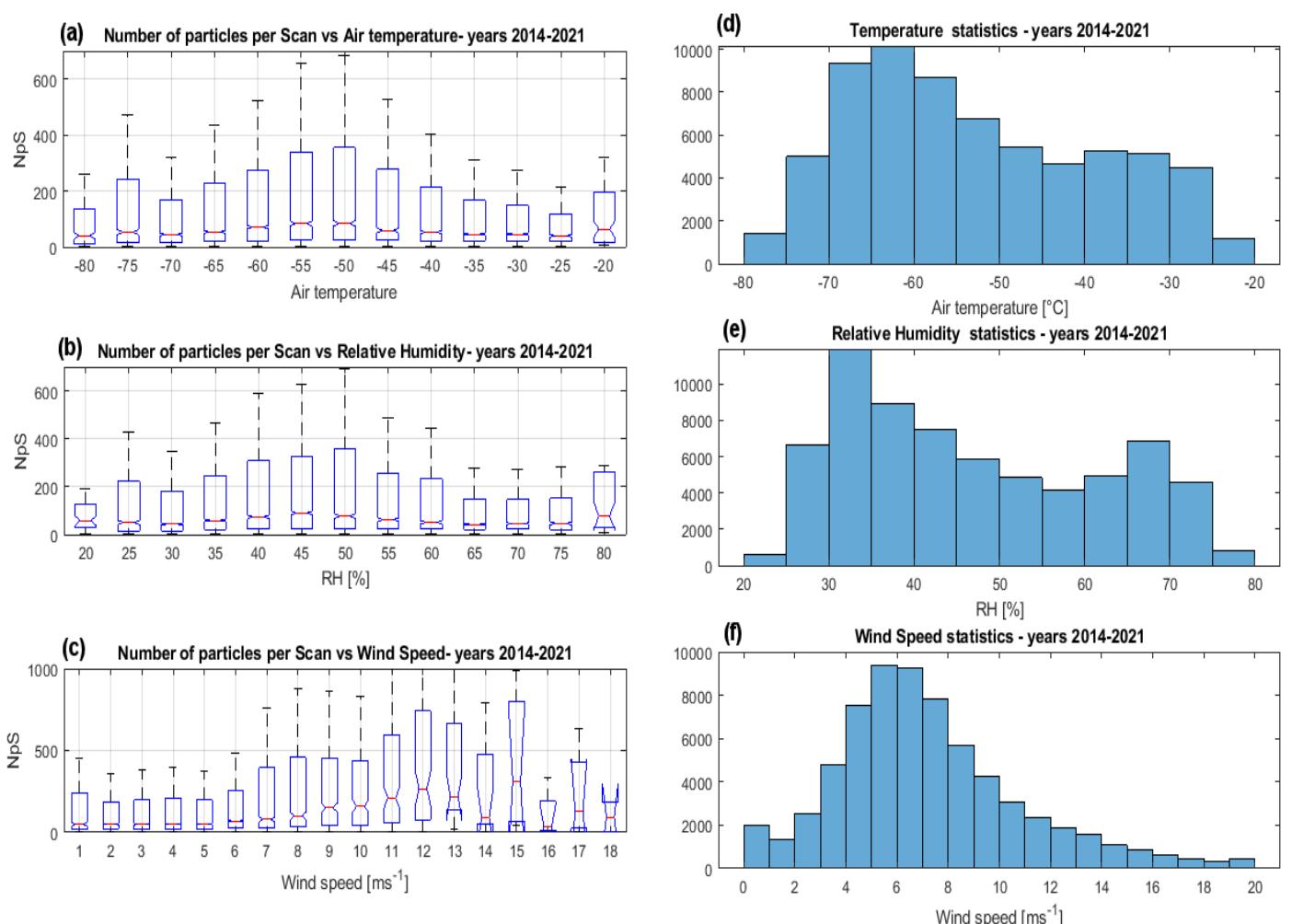
652

653 **Fig. 18: Statistics per month for the years 2014 to 2021. a) Ice particle counts per month (total counts per year are also**  
 654 **reported). (b) Number of scans per month (total number per year is also reported)**

655

656 The number of particles per scan (NpS) is a rough indicator of the intensity of the collected precipitation, but it could be  
 657 affected by sublimation, because in condition of ‘warm’ air the smallest particles could disappear from the DS before being  
 658 detected (sec.3.2). Figure 19a shows the NpS in relation to the air temperature for the whole period 2014-2021, in box and  
 659 whisker format. On each box, the middle mark indicates the median, and the lower and upper edges indicate the 25th and 75th  
 660 percentiles, respectively. The lower and upper whiskers indicate an interquartile below the 25th percentile and an interquartile  
 661 above the 75th percentile.

662 Most ice particles were detected at temperatures between  $-60^{\circ}\text{C}$  and  $-45^{\circ}\text{C}$ , characteristic temperatures in spring and autumn.  
 663 The NpS at  $-70^{\circ}\text{C}$  is not statistically different from the NpS at  $-30^{\circ}\text{C}$ . This observation shows that a statistically important  
 664 number of particles is measured also at the highest DC temperatures, when sublimation is expected (Sect.3.2) to rapidly deplete  
 665 the number of collected ice particles. According with the DC air temperature statistics (fig.19d), most particles were detected  
 666 at temperatures above the median DC temperature.  
 667 Looking at NpS statistics with relative humidity (Fig. 19b), most particles were detected with relative humidity ranging from  
 668 40% to 50%.  
 669 Figure 19c shows NpS in relation to wind velocity: ice particles were collected by ICE-CAMERA under all wind conditions  
 670 encountered in DC. Ice particles were numerically more abundant when the wind was between  $7\text{ ms}^{-1}$  and  $15\text{ ms}^{-1}$ . As the  
 671 average surface wind speed at DC resulted around  $\approx 6\text{ m s}^{-1}$  for the measurement period (Fig.19f), particles were collected on  
 672 the DS preferentially with winds stronger than the average, a condition typically encountered in winter in coincidence with  
 673 warming events (Argentini et al.,2014). These winds exceed the threshold value of  $5\text{ m s}^{-1}$  for blowing snow at ICE-CAMERA  
 674 altitude, and may ultimately contain some drifting snow. The drop of NpS for wind speeds above  $15\text{ m s}^{-1}$  (very rare in DC)  
 675 is probably due to the limited attachment of snow to the DS with strong winds.  
 676



677

678

Fig. 19: NpS statistics in relation to a)Tair, b) RH , c) wind speed.

679

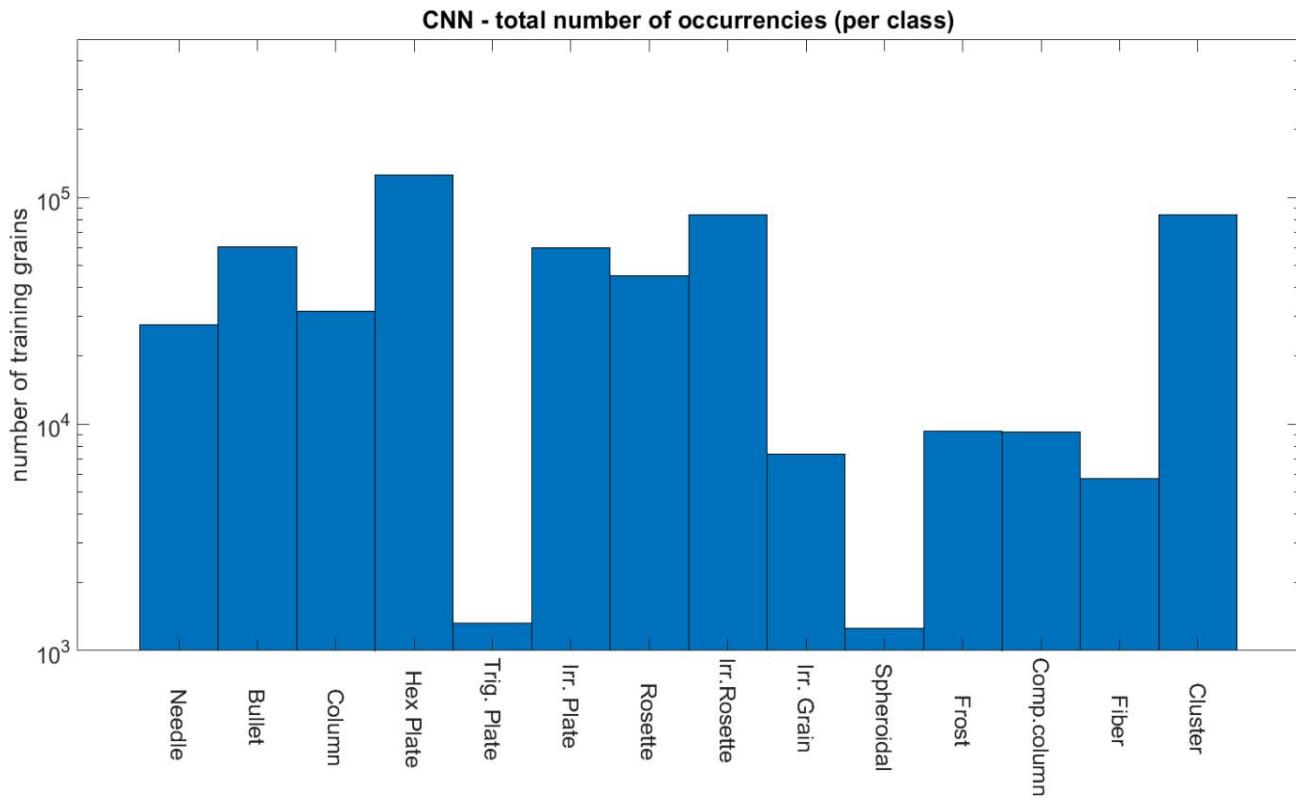
680

For comparison, the statistics for d)Tair, e) RH, and f) wind speed are shown for the same period (2014-2021).

681

682 **5.2 Image processing and CNN used on ICE-CAMERA data.**

683 MATLAB post-processing software, including the CNN classifier (Sect.4.2) and measurement tools (Sect.4.1) has been  
684 applied to the 2014-2017 ICE-CAMERA dataset. Even if the detailed analysis of these data is the task of a separate paper, a  
685 sample of the capacity of the instrument is presented in this section for the first two years of measurement (2014-2015). The  
686 total particles analyzed resulted in N=553.358. The number of particles classified in the 14 classes is reported in Fig. 20. The  
687 relative rarity of trigonal plates and spheroidal particles is evident.



688

689

**Fig. 20: Total numbers of particles classified in the 14 classes for years 2014-2015 .**

690

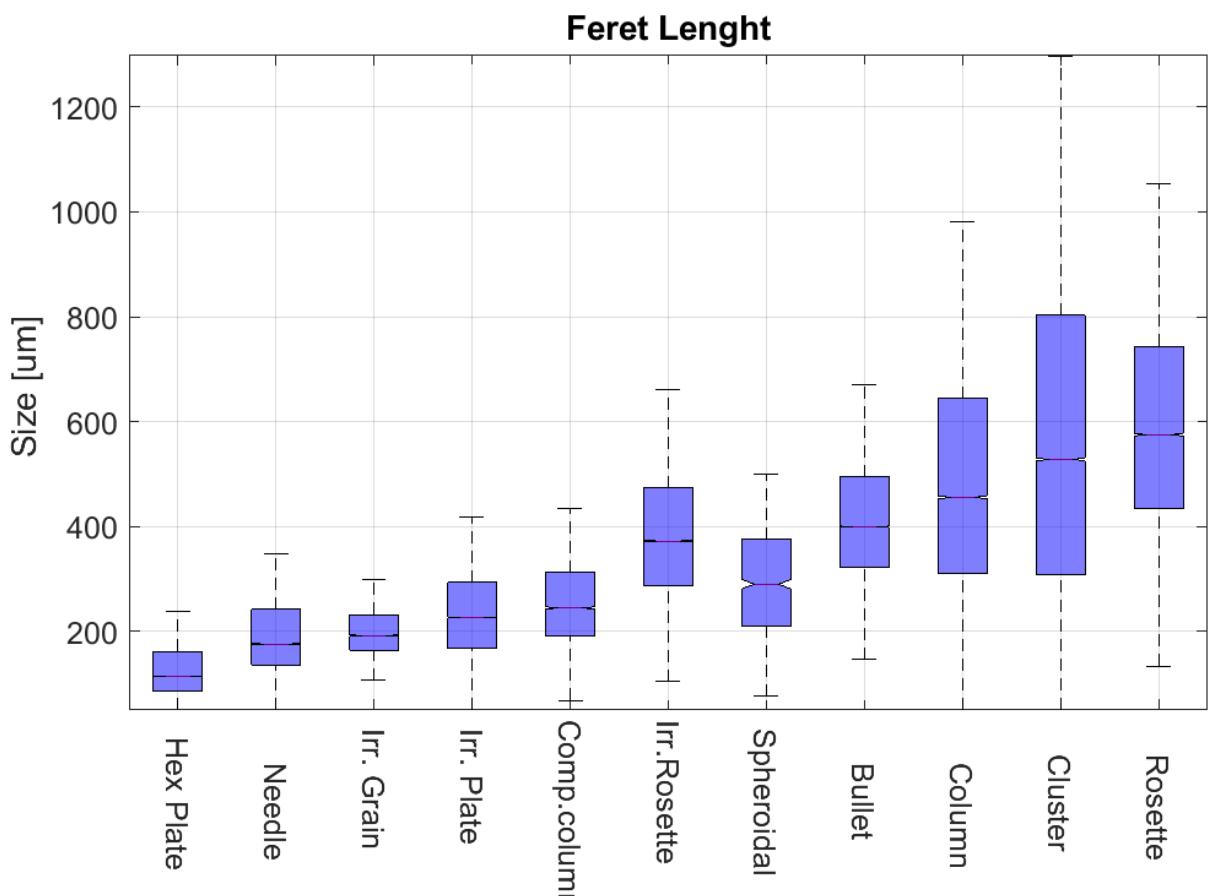
691 Figure 21 shows the Feret length statistics in box and whisker format. Particles classified as plates, needles, compact columns,  
692 spheroidal and irregular grains gave an average length lower than 300  $\mu\text{m}$ . Bullets and columns mean length resulted in the  
693 400-500  $\mu\text{m}$  range, while for rosettes and irregular rosettes was in the 350-550  $\mu\text{m}$  range.

694

695

696

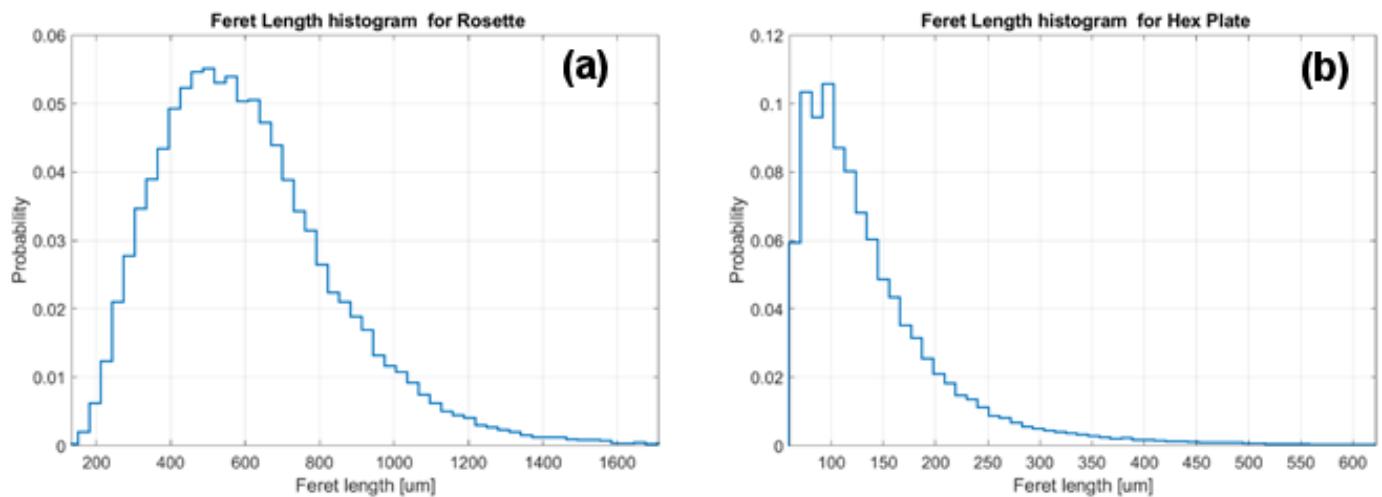
697



**Fig. 21: Feret length statistics for the years 2014-2015**

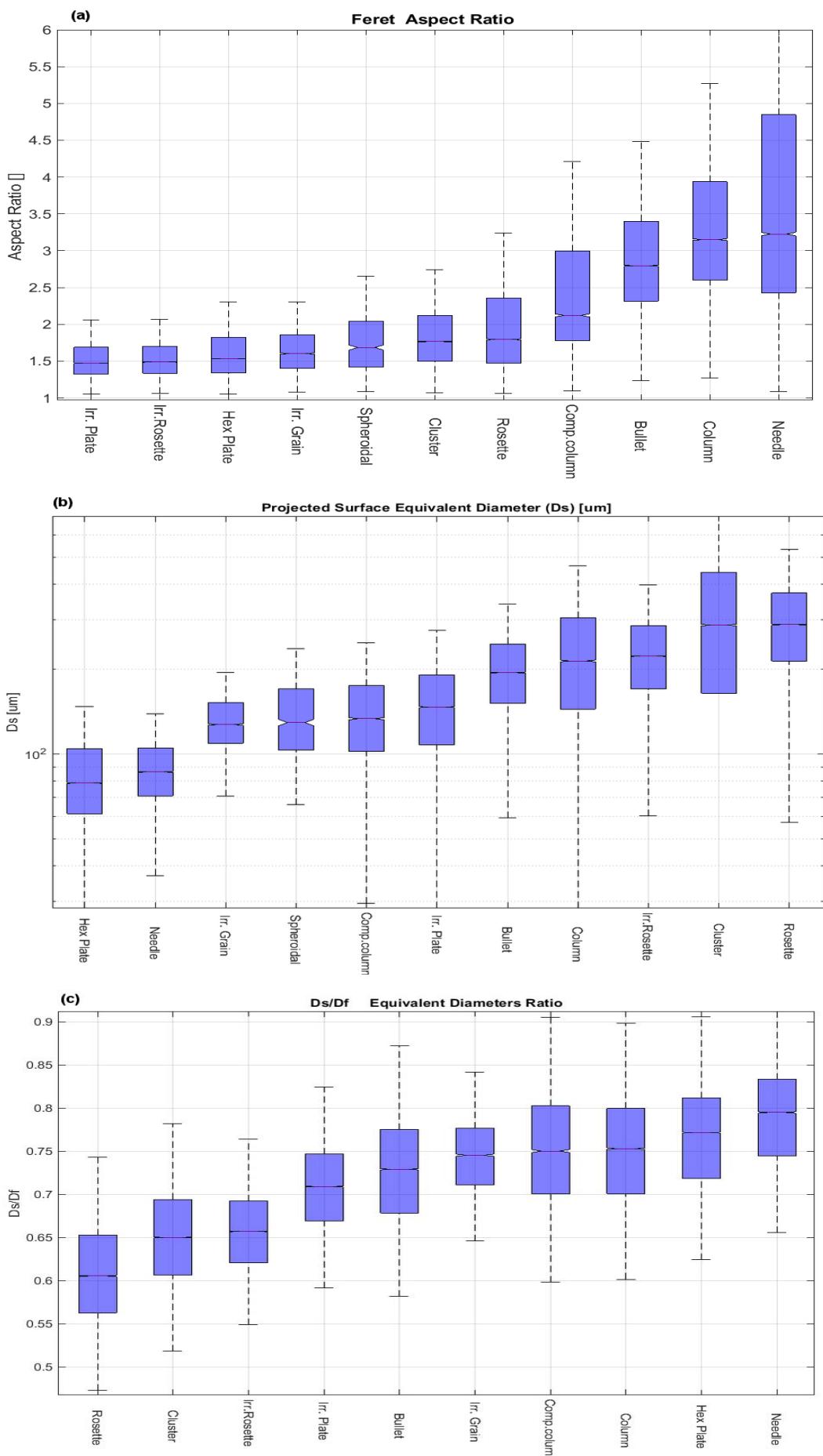
701 Figure 22 shows in detail the probability distribution of the Feret length for plates and rosettes. For plates (Fig22b), the peak  
 702 of the distribution is for  $L_{\text{Feret}}=100 \mu\text{m}$ , similar to the peak of the diamond dust (maximum) size distribution measured by  
 703 Lawson et al. (2006) at SPS in summer (it must be pointed out that Lawson et al. (2006) measured also particles as small as  
 704  $30 \mu\text{m}$ , while particles below  $60 \mu\text{m}$  are not processed by the ICE-CAMERA software, and are therefore missing from the  
 705 probability distribution of Fig22b). This finding suggests that sublimation of particles less than  $100-200 \mu\text{m}$  in diameter during  
 706 the deposition period (Sect. 3.2) is not relevant for shaping the final particle size statistics. The loss of particles in the lowest  
 707 size range for sublimation cannot be quantitatively assessed from these data. Nevertheless, the first effects of sublimation are  
 708 expected to be evident in the ICE-CAMERA images of small particles in the form of loss of sharp edges, eventually leading  
 709 to spheroidal shapes (sect.3.2). An overview of the images collected in the 2014-2015 summers (where sublimation is most  
 710 likely to occur) indicates that this effect is rarely observed. Either sublimation is slower than expected from the simulations  
 711 of Sect.3.2, or what is observed in the summer images of ICE-CAMERA is just the result of the crystals felt just before the  
 712 scan, with the majority of previously fallen small particles definitely sublimated and not detected by image segmentation.  
 713 This ambiguity will be resolved in DC by taking a continuous serie of acquisitions of DD in summer conditions. The results  
 714 obtained from ICE-CAMERA for pristine rosettes (Fig22a), differ considerably from those of Lawson et al (2006), because  
 715 the peak of the probability distribution resulted  $L=480 \mu\text{m}$ , to be compared with  $L=120 \mu\text{m}$  of Lawson et al (2006). This  
 716 difference is not explicable with the eventual sublimation of the smallest rosettes on the DS. Instead, this result is a realistic  
 717 feature, sustained by the direct visual observation of rosettes in DC precipitation. A much greater amount of rosettes is actually  
 718 observed in DC during precipitation from clouds than during diamond dust events. Even if rosettes in diamond dust are much

719 smaller than rosettes from clouds, the numerical dominance of cloud rosettes explains the large median value of their Feret  
720 length.



722 **Fig. 22: Feret length probability distribution for a) rosettes and b) hex. plates. The relevant presence of small plates**  
723 **(D<200 $\mu$ m) suggests that sublimation on the DS is not relevant. years 2014-2015**

724 Figure 23a shows Feret's aspect ratio per class. Not surprisingly, many "rounded" classes (plates, rosettes, etc.) have an AR<2.  
725 Compact columns show a median AR close to 2.4, while columns and bullets are close to 3. The average AR for the needles  
726 was 3.2, which is lower than expected for the reasons outlined in section 4.1.3.



**Fig. 23: Statistics of (a) Aspect Ratio, (b) Projected surface-equivalent diameter (Ds), (c) ratio between surface-equivalent (Ds) and Feret box-equivalent (Df) diameters for the years 2014-2015.**

731 Figure 23b shows the surface-equivalent diameter  $D_s$  of the particles. Figure 23c shows the ratio between the surface  
 732 equivalent diameter ( $D_s$ ) and the Feret-box equivalent diameter ( $D_f$ ). The difference between the two diameters is relevant for  
 733 "fluffy" particles like rosettes and clusters. For those particles,  $D_s/D_f$  gave values of 0.6 to 0.65. For comparison, a round  
 734 particle is expected to have a ratio of  $D_s/D_f=0.89$ .

736 **6. Conclusions.**

737 ICE-CAMERA, although very similar to a simple flatbed scanner in its basic design, has represented a technical challenge for  
 738 its implementation at DC. Hardware and software have been continuously and extensively modified at DC over the past five  
 739 summer campaigns. The result is now a reliable instrument, running throughout the year on an hourly basis, for the statistical  
 740 study of precipitation in internal polar areas. Particle size and morphology are automatically obtained, and some semi-  
 741 quantitative precipitation estimates can be derived. The collected data are automatically pre-analyzed, but they can be post-  
 742 processed at any time, in order to follow the continuous improvements of the image processing and machine learning  
 743 algorithms. The GoogleNet CNN, trained specifically for this instrument, has succeeded in classifying ICE-CAMERA images  
 744 into 14 form classes, with an accuracy of more than 80% for most of them. The instrument is particularly useful for  
 745 automatically measuring the size of individual ice particles in precipitation, a process virtually impossible manually, and  
 746 certainly impossible on the field in DC and elsewhere on the Antarctic plateau in winter. ICE-CAMERA scans are carried out  
 747 every hour. Keeping the surface of the instrument free of frost all the time and cleaning it by heating the deposition surface  
 748 after each scan is paid with the possible loss of small ice particles. Particles less than 100-200  $\mu\text{m}$  can disappear by sublimation  
 749 before being recorded, especially in summer. This problem is complementary to the problem encountered when observing  
 750 precipitation manually: when observing precipitation manually every 24 hours, (as is the case of DC) the reprocessing of  
 751 particles, or the formation of ice and hoar artifacts cannot be prevented. In ICE-CAMERA, frost and ice regrowth are  
 752 suppressed, but small particles may disappear for sublimation. The effect of sublimation on particles observed with ICE-  
 753 CAMERA cannot be easily quantified, also given the broad range of atmospheric conditions encountered by the DS throughout  
 754 the year. Images of small particles (100-200  $\mu\text{m}$ ) (such as plates) collected during the 2014-2015 summers rarely show  
 755 evidence of early sublimation such as edge smoothing or rounding. While encouraging specific experiments with ICE-  
 756 CAMERA, this observation suggests that sublimation could be slower than predicted by simulations. ICE-CAMERA data,  
 757 collected since 2014, have already been statistically processed and the results will be described in a specialized paper. Results  
 758 from a subset of data (years from 2014 to 2015), was presented in this work. As preliminary results, in DC the rosettes were  
 759 found to be significantly larger (480  $\mu\text{m}$ ) than those observed at SPS by Lawson et al (2006), while the plates were of similar  
 760 size (120  $\mu\text{m}$ ). These results demonstrated the capability of the instrument to classify and size individual ice particles in DC  
 761 precipitation. Unfortunately, only non-polluted, very cold, low humidity, low precipitation environments (like high mountain  
 762 tops, dry polar environments) could house a similar instrument. In the presence of pollution, marine aerosols or dust, manual  
 763 cleaning of the DS would be required to remove solid particles and salts escaping sublimation. For coastal zones, the  
 764 temperature is generally close to zero, making the thermal cleaning of the DS by sublimation problematic. In these  
 765 environments, if an instrument like ICE-CAMERA were installed, a mechanical wiper would replace the heated glass of the  
 766 current instrument. Furthermore, the CNN presented in this paper should be re-trained with different classes of ice crystals.

771 **7. Technical issues.**

772 Using ICE-CAMERA at DC, as well as other automated instruments, was difficult. The instrument had several failures along  
773 years, and each one was difficult to fix, at least in winter, when the instrument had to be dismounted from the roof of the  
774 shelter at -70°C and eventually fixed in the local lab by the winter-over crew, with remote assistance from Europe. Until a few  
775 years ago, communicating with DC was limited to email with small attachments, making remote assistance a lengthy task.  
776 Even today, connecting the rest of the world remotely with the ICE-CAMERA PC, to operate with the instrument software, is  
777 virtually impossible. Most hardware failures in DC were due to software bugs or computer failures. Rather than having trouble  
778 with low temperatures, operating in DC means dealing with limited heat-dissipation of PC parts such as power supply and  
779 hard disks, electrostatic discharge issues in low-humidity, heated environments, lack of spare parts for most of the year, a  
780 varied skill-ness of winter-over personnel. Failures in the thermal control of ICE-CAMERA caused some mechanical stress  
781 and failures in the focusing sledge, while water condensation eventually rusted the bearings of the stepper motors (all bearing  
782 were de-greased for a better low temperature operation). The CNN used to classify ICE-CAMERA images is continually  
783 changing and improving and the CNN training dataset increases with time, as new images collected by ICE-CAMERA are  
784 used as new training ones.

785

786

787 **8. Code and Data availability.**

788 The CNN developed as part of this work (under Mathworks MATLAB R2020B), along with the image dataset (224\*224  
789 images for the 14 classes of particles) used for training, validation and testing the CNN are available in the ZENODO  
790 repository (Del Guasta, 2022)

791

792

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797 hard to permit our scientific activity. I am also grateful to Francesco Castagnoli (INO CNR) for the initial design of the  
798 instrument.

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