



1 Performance evaluation of an integrated path 2 differential absorption LIDAR model for surface 3 pressure from low-Earth orbit

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10 **Abstract.** Remote sensing of surface pressure from space is critical; differential absorption
11 LIDAR and differential absorption radar are only two kinds of remote sensing instruments
12 with this potential. The differential absorption LIDAR works in integral path mode from the
13 satellite in low-Earth orbit. It measures the differential optical depth of the Oxygen A-band,
14 and the surface pressure is thereafter obtained by performing circle-iterative calculation.
15 Performance evaluation of the differential absorption LIDAR model was conducted with
16 respect to the advanced system parameters of the space instrument, Low echo pulse energy at
17 ocean surface and the challenging calculation of repetitive cumulative average of echo on
18 uneven land surface yielded random errors in surface pressure measurement. On the other
19 hand, uncertain atmospheric temperature and water vapor mixture profiles resulted in
20 systematic error of surface pressure. Consequently, controlling the error of surface pressure
21 within 0.1% proved challenging. Under a strict implementation of the error budget, the time
22 resolution is 6.25 s and along-orbit distance resolution is 44km, and the results showed that
23 765.6735/765.4637 nm is suitable as the working wavelength pair. Further, error could be
24 expected to within 0.2–0.3% for the cumulative average of 625 ocean surface laser pulse
25 echoes, cumulative average of more than 144 pulse echoes on land, and observation from the
26 400km orbit.

27 1 Introduction

28 Atmospheric pressure plays a vital role in several atmospheric processes related to
29 atmospheric dynamics. Low/high pressure, low pressure troughs, high pressure ridges, and
30 anticyclones and other related information have been introduced into the atmospheric model.
31 Hurricanes are profound low-pressure systems that originate from low-pressure cyclones in
32 the tropical or subtropical oceanic regions. Accurate prediction of their formation, landing
33 direction, and movement trajectory requires atmospheric pressure gradient distribution data.
34 Fundamentally, the density of the atmosphere in high latitudes increase during winter, causing
35 the air to shrink and sink, thereby increasing the pressure and gradually resulting in the
36 formation of a powerful, deep, and broad air mass. Upon the accumulation of a sufficient cold
37 high-pressure force, a cold wave is formed, which rolls out and pours down. Meanwhile,
38 airspace for the release of radio sounding balloons is restricted; thus, continuous detection
39 during the entire day is not possible. Brown et al.(1986) reported that the accuracy of the
40 weather models is primarily limited by the regional sparsity of the input data. Specifically,
41 atmospheric pressure data is very sparse in large areas of the ocean, desert, plateaus, and polar
42 regions. Consequently, the International Meteorological Organization aims to achieve remote
43 sensing of surface pressure at an accuracy of 0.1–0.3%(Korb et al., 1995)(WMO-ICSU, 1973),
44 which however, remains a big challenge.



45 In 1983, Korb, C. L. et al., (1983) of Laboratory for Atmospheres, NASA Goddard Space
46 Flight Center, proposed a method of detecting atmospheric pressure using differential
47 absorption LIDAR and the trough between oxygen absorption lines. In 1987, Schwemmer et
48 al. (1987) structured a novel differential absorption LIDAR system. It employs a flash-pump
49 alexandrite laser to emit a beam of two wavelengths of approximately 13160cm^{-1} , coupled with
50 an oxygen photoacoustic absorption cell and a high-precision wavelength meter to stabilize
51 the emission wavelength. Moreover, the seed source is a continuous wave from either a
52 Ti:sapphire single longitudinal mode laser or a diode laser (Schwemmer et al., 1987). In June
53 and July 1989, a series of flight measurement tests were conducted on the east coast of the
54 United States (Korb et al., 1989). In 1999, Flamant, C. N., Schwemmer, G. K., Korb, C. L.,
55 Evans, K. D. and Palm, S. P. published their report "Pressure measurements Using and
56 Airborne Differential absorption LIDAR. Part I: Analysis of the systematic error
57 sources," (Flamant et al., 1999) where in the instrumental and systematic error sources of
58 differential absorption LIDAR was analyzed when measuring atmospheric pressure profile.

59 In the ASCENDS (Active Sensing of CO₂ Emission over Nights, Days, and Seasons)
60 program, the surface pressure was determined to accurately measure the CO₂ dry mixing
61 ratio (Zaccheo et al., 2014; Crowell, et al., 2015). Between 2007 and 2013 Stephen, M.
62 Krainak, M. Riris, H. and others of NASA Goddard Space Flight Center and Allan, G. R. of
63 Sigma Space Corporation reported on the use of an aircraft as a platform and transmitter to
64 continuously send out pulse trains of multiple wavelengths of approximately 764.7nm with
65 the receiver receiving the return echoes. (Stephen et al., 2007-2008; Riris et al., 2012-
66 2013, 2017) Thus, multiple pulse train return signals were accumulated, using which the
67 oxygen absorption spectrum curve of the 764.5–764.9nm trough segment was plotted.
68 Subsequently, the differential optical depth of oxygen was calculated from the transmittance
69 curve.

70 Dual-wavelength (detection/reference wavelengths) laser pulses are launched downwards
71 from the space platform (Millán et al., 2014); consequently, the reflected laser pulses energy
72 from the earth's surface or the top of a cloud are received. Subsequently, the atmospheric
73 optical depth and flight time of the laser pulses passing through the air column are measured.
74 Thus, the atmospheric pressure and altitude of the surface/cloud top can be simultaneously
75 obtained, and the top of the cloud ground can be distinguished from the ground. Such data is
76 meaningful for various meteorological applications. By obtaining the pressure values on the
77 surface and cloud tops and combining the results with a vertical temperature profile obtained
78 from other sensors or weather models and utilizing statistical equations, the vertical profile of
79 atmospheric pressure can be obtained. The differential absorption LIDAR is installed on a
80 sun-synchronous orbit, and it makes a polar orbit around the earth from south to north. It
81 allows much denser surface/cloud top atmospheric pressure data than ground meteorological
82 stations to be obtained.

83 This paper is structured as follows. Section 1 presents the Introduction, and Section 2, the
84 mechanism of differential absorption LIDAR for detecting surface pressure was introduced.
85 whereas Section 3 evaluates the performance of a differential absorption LIDAR model,
86 Finally, Section 4 presents the summary.

87 2 Mechanism of Differential Absorption LIDAR to Detect Atmospheric Pressure

88 Differential absorption LIDAR selects two wavelengths in the A absorption band of oxygen
89 (759–770nm). The laser beam with one wavelength value passes through the atmosphere
90 twice; its absorption coefficient, although insensitive to changes in atmospheric temperature,
91 is sensitive to variations in atmospheric pressure. This wavelength is referred to as the
92 detection wavelength (online). Further, the absorption coefficient of another wavelength from
93 the laser beam passing through the atmosphere twice is relatively smaller, and it is referred to
94 as the reference wavelength (offline), with its value being close to the detection wavelength.



95 Let the atmospheric pressure at altitude R_0 , where the LIDAR is located, be $p(R_0)$; the
 96 atmospheric pressure at altitude R be $p(R)$; and $g(z)$ be the gravitational acceleration at
 97 altitude z . The difference in the atmospheric pressure between altitude R_0 and R is equal to the
 98 weight of the air column between R_0 and R per unit area, where the dry air molecular mass
 99 $m_{dry}=28.9644$ g/mol and water vapor molecular mass $m_{wv}=18$ g/mol.
 100 Atmospheric quasi-static equation:

$$101 \quad dp = -n_{dry}(z) \cdot (m_{dry} + m_{wv}\chi_{wv}(z)) \cdot g(z) dz \quad (1)$$

102 Gas state equation:

$$103 \quad p(z) = n_{dry}(z) \cdot (1 + \chi_{wv}(z)) \cdot kT(z) \quad (2)$$

$$104 \quad p(R) = p_{surface} \cdot \exp^{-\int_0^R \frac{(m_{dry} + m_{wv}\chi_{wv}(z))g(z)}{kT(z)(1 + \chi_{wv}(z))} dz} \quad (3)$$

105 This integration is performed at an altitude z , $n_{dry}(z)$ is the density of dry air molecules, $\chi_{wv}(z)$
 106 is the water vapor volume mixing ratio, $p_{surface}$ is the surface pressure, and k is the Boltzmann
 107 constant. Thus, by remote sensing the weight or mass per unit area of a vertical air column
 108 between two altitudes, the difference in atmospheric pressure between these two altitudes can
 109 be obtained.

110 Oxygen is among the most stable components in the atmosphere in terms of space and
 111 time. $n_{O_2}(z)$ is the number density of oxygen molecules at altitude z . The number of oxygen
 112 molecules accounts for a fixed proportion of 20.948% of the number of dry air atmospheric
 113 molecules. Further, the optical depth of the atmosphere between R_0 and R is the integral of its
 114 extinction coefficient with respect to the beam path, which can be expressed as

$$115 \quad OD(R_0, R) = \int_{R_0}^R [\alpha_a(v, z) + \alpha_m(v, z) + n_{O_2}(z)\sigma(v, p(z), T(z))] dz \quad (4)$$

116 where OD is the optical depth in Beer's theorem, σ is the absorption cross-section of the
 117 oxygen molecule to the A-band λ wavelength, and $\Delta\sigma$ is the difference $\sigma(\lambda_{on}, p(z), T(z)) - \sigma(\lambda_{off},$
 118 $p(z), T(z))$. Further, $\alpha_a(\lambda, z)$ and $\alpha_m(\lambda, z)$ are the aerosol extinction coefficient and the extinction
 119 coefficient of atmospheric molecules except for oxygen absorption, respectively,
 120 and $n_{O_2}(z)\sigma(\lambda, p(z), T(z))$ is the oxygen absorption coefficient of the corresponding
 121 wavelength. The difference in the single-pass optical depth compared to the dual-wavelength
 122 between R_0 and R is referred to as the differential optical depth $dOD(R_0, R)$. Although the
 123 weight of the atmospheric column between R_0 and R per unit area is unknown, the differential
 124 optical depth can be expressed as

$$125 \quad dOD(R_0, R) = \int_{R_0}^R n_{O_2}(z)(\sigma(v_{on}, z) - \sigma(v_{off}, z)) dz \quad (5)$$

$$126 \quad n_{O_2}(z) = \frac{0.20948p(z)}{kT(z) \cdot (1 + \chi(z))} \quad (6)$$

127 where $N_{s,on}(R)/N_{s,off}(R)$ represents the online/offline dual-wavelength echo pulse energy
 128 (number of photons) received by the LIDAR, which is expressed using the LIDAR equations
 129 as follows:

$$130 \quad N_{s,on}(R) = \frac{c}{2} \frac{A_r \rho}{R^2 \pi} E_{on} \eta_r \eta_d \times \exp[-2OD(R_0, R)] \quad (7)$$

$$131 \quad N_{s,off}(R) = \frac{c}{2} \frac{A_r \rho}{R^2 \pi} E_{off} \eta_r \eta_d \times \exp[-2OD(R_0, R)] \quad (8)$$

132 Where E_{on}/E_{off} is the energy of a single shot emitted laser for both online/offline. Further, η_r is
 133 the receiving efficiency of light beam, η_d is the quantum efficiency of the detector, and A_r is
 134 the effective receiving area of the telescope. In the space-to-earth observation, IPDA, receives
 135 return echo from hard targets on the ground, and ρ represents the reflectivity of ground targets.



136 The beam of dual-wavelength has the same path, receiving/sending time, footprint, and
 137 random process in the atmosphere. Further, except for $\sigma(\lambda, p(z), T(z))$, all other parameters are
 138 considered to be similar (but not equal). Dividing Eq. (7) by Eq. (8), the differential optical
 139 depth can be calculated by measuring the energy of the pulse emitted and the energy of
 140 received return echo by the LIDAR as follows:

$$141 \quad dOD(R_0, R) = -\frac{1}{2} \ln \left\{ \left[\frac{N_{on}(R)}{N_{off}(R)} \right] \left(\frac{E_{off}}{E_{on}} \right) \right\} + C \quad (9)$$

$$142 \quad C = \int_{R_0}^R [\alpha_a(v_{on}, z) - \alpha_a(v_{off}, z)] dz + \int_{R_0}^R [\alpha_m(v_{on}, z) - \alpha_m(v_{off}, z)] dz \quad (10)$$

143 where C represents a systematic error between the $\frac{1}{2} \ln \left(\frac{N_{off}(R) E_{on}}{N_{on}(R) E_{off}} \right)$ value calculated from
 144 the LIDAR data and the differential optical depth. Thus, laser shots of these two wavelengths
 145 are simultaneously emitted and the reflections from the surface/cloud tops are received.
 146 However, owing to the difference in oxygen absorption, the atmospheric transmittance of the
 147 two wavelengths is different. The logarithm of the ratio of $\frac{N_{off}(R) E_{on}}{N_{on}(R) E_{off}}$ can be used to obtain
 148 the atmospheric differential optical depth from the satellite to the surface/cloud top. IPDA
 149 launches several laser pulses from the space platform to the ground, and it detects the surface
 150 pressure, with point R_0 representing the satellite location. Further, there is almost no air
 151 pressure $p(R_0)=0.0$, and $p(R)$ represents the surface pressure $p_{surface}$.

152 In the path of the laser beam, only the section from the altitude of 71km to the ground has
 153 a significant effect on the optical depth, whereas the effect of atmosphere above 71km can be
 154 ignored. Further, the gravitational acceleration $g(z)$ can be regarded as a constant 9.80616
 155 N/m^2 at atmospheric altitude below 71 km.

156 On transforming the elevation z coordinates in Eq. (1) into atmospheric pressure p
 157 coordinates, the following is obtained

$$158 \quad n_{dry}(z) dz = \frac{dp}{(m_{dry} + m_{wv} \chi_{wv}(p)) g(p)} \quad (11)$$

159 Thus, the absorption cross-section $\sigma(\lambda, p(z), T(z))$ is related to atmospheric temperature and
 160 pressure, and thus it can be rewritten as $\sigma(\lambda, p, T(p))$ in pressure p coordinates. Combining Eq.
 161 (11), we can transform Eq. (5) from the elevation z coordinate to the pressure p coordinate.
 162 The differential optical depth dOD associated with pressure p coordinates is expressed by

$$163 \quad dOD(p_{ground}, p_{top}) = 0.20948 \int_{p_{top}}^{p_{ground}} \frac{\Delta\sigma(v, p, T(p))}{(m_{dry} + m_{wv} \chi_{wv}(p)) g(p)} dp \quad (12)$$

164 Here we assume that the pressure at the top of the atmosphere is $p_{top}=0.0$, and the atmospheric
 165 pressure at the surface (or cloud top) $p_{ground}=p_{surface}$.

$$166 \quad dOD(p_{surface}) = \frac{0.20948}{g} \int_0^{p_{surface}} \frac{\sigma(v_{on}, p, T(p)) - \sigma(v_{off}, p, T(p))}{m_{dry} + m_{wv} \chi_{wv}(p)} dp \quad (13)$$

167 Equation (13) establishes the implicit expression of the differential optical depth of the entire
 168 aerosphere with respect to the surface pressure $p_{surface}$. Theoretically, the true value of the
 169 differential optical depth is the state of the atmosphere and is not related to the LIDAR
 170 parameters. Further, it is independent of the measurement method. However, the
 171 measurement error of the differential optical depth is closely related to the LIDAR parameters.
 172 In the pressure p coordinate, the differential optical depth $dOD(p_{surface})$ is expressed through
 173 the integral Eq. (14) as follows:

$$174 \quad dOD(p_{surface}) = \int_0^{p_{surface}} \frac{\sigma(v_{on}, p, T(p)) - \sigma(v_{off}, p, T(p))}{2.251667 \times 10^{-24} \times (1 + 0.6214 \chi_{wv}(p))} dp \quad (14)$$



175 On differentiating both sides of Eq. (13) with respect to $p_{surface}$ and considering the derivative
 176 function of $dOD(p_{surface})$ with respect to $p_{surface}$, we obtain

$$177 \quad \frac{\partial(dOD(p_{surface}))}{\partial p_{surface}} = \frac{1}{46.8119} \times \frac{\sigma(v_{on} p_{surface} T(p_{surface})) - \sigma(v_{off} p_{surface} T(p_{surface}))}{m_{dry} + m_{wv} \chi_{wv}(p_{surface})} \quad (15)$$

178 Subsequently, the relationship between the errors of the surface pressure and the differential
 179 optical depth of the entire atmosphere is obtained as

$$180 \quad \delta p_{surface} = \frac{46.8119 \times [m_{dry} + m_{wv} \chi_{wv}(p_{surface})]}{\sigma(v_{on} p_{surface} T(p_{surface})) - \sigma(v_{off} p_{surface} T(p_{surface}))} \delta [dOD(p_{surface})] \quad (16)$$

181 Assuming that the vertical profile of atmospheric temperature $T(R)$ and the vertical profile of
 182 water vapor mixing ratio $\chi_{wv}(R)$ are known from data obtained from other sensors or weather
 183 models, the surface pressure can be inverted from the differential optical depth of the entire
 184 atmosphere. The steps are shown in Fig. 1.

185 a. The differential optical depth measurement value $(dOD)_m$ of the atmosphere from the
 186 echo signal N_s and emission energy E of the differential absorption LIDAR is calculated.

187 b. Utilizing the atmospheric temperature profile $T(R)$ coupled with the pressure profile and
 188 the surface pressure in the standard atmosphere mode as the initial value of the
 189 atmospheric pressure profile $p_1(R)$ and the initial value of the surface pressure $p_{surface,1}$,
 190 respectively, and using the oxygen HITRAN database, the initial value of the absorption
 191 coefficient profile of the entire atmosphere is calculated. Thereafter, the initial value of the
 192 differential optical depth $(dOD)_{c,1}$ of the entire atmosphere is calculated.

193 c. If the differential optical depth $(dOD)_{c,i}$ of the entire atmosphere is numerically
 194 calculated in i -th cycle and $(dOD)_{c,i}$ is not equal to the differential optical depth $(dOD)_m$
 195 measured by the LIDAR, then the surface pressure $p_{surface,i}$ calculated using the numerical
 196 value is not equal to the true value $p_{surface}$ of the pressure at the footprint, and thus, $(dOD)_m$
 197 is subtracted from $(dOD)_{c,i}$.

198 d. The surface pressure varies with the differential optical depth. In the i -th cycle, the
 199 difference between $(dOD)_{c,i}$ and $(dOD)_m$ is multiplied by a coefficient
 200 $\frac{46.8119 \times [m_{dry} + m_{wv} \chi_{wv}(surface,i)]}{\sigma(v_{on} p_{surface,i} T_{surface,i}) - \sigma(v_{off} p_{surface,i} T_{surface,i})}$ in Eq. (16) as the compensation amount
 201 and added to the calculated value of the surface pressure $p_{surface,i}$. Consequently, the
 202 resulting sum is used as the new surface pressure $p_{surface,i+1}$; $p_{surface,i+1} = p_{surface,i} +$
 203 $\frac{46.8119 \times [m_{dry} + m_{wv} \chi_{wv}(surface,i)]}{\sigma(v_{on} p_{surface,i} T_{surface,i}) - \sigma(v_{off} p_{surface,i} T_{surface,i})} (dOD_{c,i} - dOD_m)$.

204 e. Subsequently, with atmospheric temperature profile $T(R)$ and water vapor mixing ratio
 205 $\chi_{wv}(R)$ provided by other sensors or numerical weather models, coupled with the surface
 206 pressure result $p_{surface,i}$ obtained in the $i+1$ -th cycle, the atmospheric pressure profile $p_{i+1}(R)$

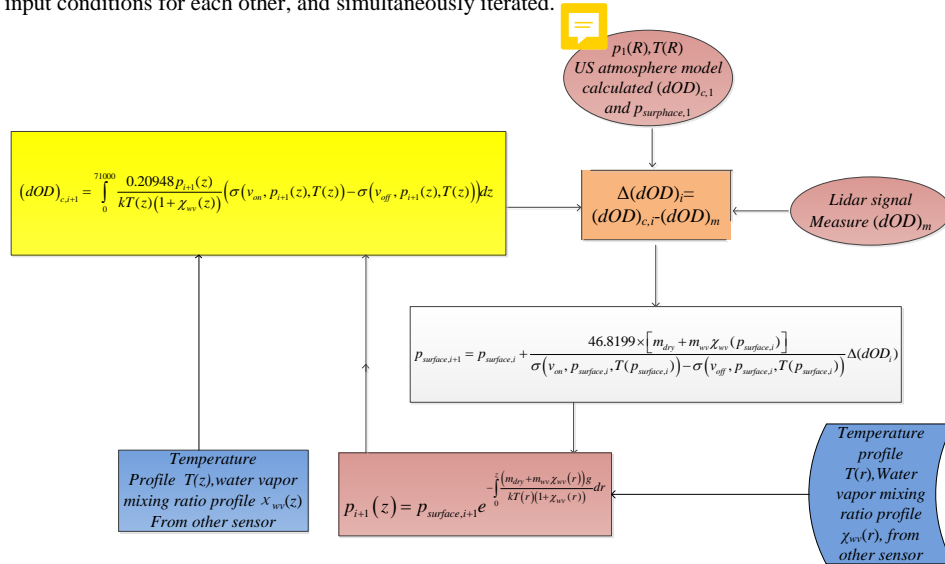
$$207 \quad \text{is calculated, } p_{i+1}(R) = p_{surface,i} \cdot \exp^{-\int_0^R \frac{(m_{dry} + m_{wv} \chi_{wv}(z))g(z)}{kT(z)(1 + \chi_{wv}(z))} dz}$$

208 f. Further, with the atmospheric temperature profile $T(R)$, profile for water vapor mixing
 209 ratio $\chi_{wv}(R)$, and the atmospheric pressure profile $p_{i+1}(R)$, based on the HITRAN database,
 210 differential optical depth $(dOD)_{c,i+1}$ calculations are repeated.

211 g. Repeat steps c–f. In the case of the above iterative process, with increase in i , the
 212 difference between $(dOD)_{c,i}$ and $(dOD)_m$ decreases till $i=M$, $p_{surface, M+1} - p_{surface, M}$ is
 213 comparable to the error. If that happens, the iterative loop stops. Herein, the output surface
 214 pressure $p_{surface, M}$ calculation result is considered to be sufficiently close to the true value
 215 $p_{surface}$.



216 The above calculation steps also suggest that the parameters related to the temperature,
 217 pressure, and humidity of the atmosphere should be detected synchronously in the future, as
 218 input conditions for each other, and simultaneously iterated.



219
 220

221 **Figure 1:** Iterative calculation process of the atmospheric pressure on the surface of the
 222 differential optical depth measured by the differential absorption LIDAR.

223 3 Performance evaluation of an integrated path differential absorption LIDAR model

224 3.1 A-band absorption spectrum of Oxygen

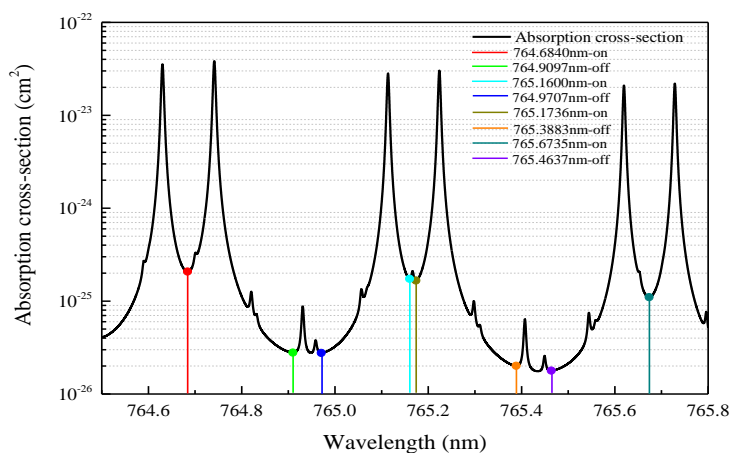
225 The absorption line of oxygen molecules is broadened in the atmosphere via collision and
 226 Doppler broadening. They are expressed via the famous Lorentz and Gauss line shapes,
 227 respectively. Below 15km in the atmosphere, collision broadening is dominant, with “n”
 228 representing the sensitivity factor of collision broadening with respect to air temperature, that
 229 is, the average value of its own broadening and nitrogen broadening sensitivity factors. We
 230 consider $n=0.73-0.59$ from HITRAN database, normal pressure $p_0=1013.25\text{hPa}$, normal
 231 temperature $T_0=296\text{K}$, γ_0 is the pressure broadening under normal temperature and normal
 232 pressure, S_0 is the intensity of the absorption line at room temperature and pressure, σ_0 is the
 233 peak absorption cross section of the absorption line at room temperature and pressure, c is the
 234 speed of light, h is the Planck constant, m is the molecular mass of oxygen, and $\nu_0(\text{cm}^{-1})$
 235 represents the position of the center wave number (light frequency) of the absorption line.
 236 Further, E'' is the energy of the low-energy state of the electron. Moreover, in the application
 237 of differential absorption LIDAR, the absorption line shape of the oxygen molecule can be
 238 represented using the Voigt line shape, which is a form of the convolution Gauss line shape
 239 with Lorentz line shape. The arbitrary real number t is the variable of the Voigt linear integral.
 240 The absorption cross-section $\sigma(\nu)$ at the light wave number ν is written as $\sigma(\nu) =$
 241 $\sigma_0 \frac{y}{\pi} \int_{-\infty}^{\infty} \frac{\exp(-t^2)}{y^2 + (x-t)^2} dt$, where $x = \left(\frac{\nu - \nu_0}{\nu_0}\right) \left(\frac{m}{2kT}\right)^{0.5} c$ and $y = \nu_0 \gamma_0 \left(\frac{p}{p_0}\right) \left(\frac{T_0}{T}\right)^n \left(\frac{m}{2kT}\right)^{0.5} c$, and
 242 $\sigma_0 = \frac{S_0 c}{\nu_0} \left(\frac{2\pi kT}{m}\right)^{0.5} \left(\frac{T_0}{T}\right)^{1.5} \exp\left[\frac{E'' c}{k} \left(\frac{1}{T_0} - \frac{1}{T}\right)\right]$. The wavelength in the trough between the oxygen
 243 absorption line P13Q12 and P13P13, between the oxygen absorption line P15P15 and
 244 P15Q14, and even absorption lines P17Q16 and P17P17 can be selected as the detection



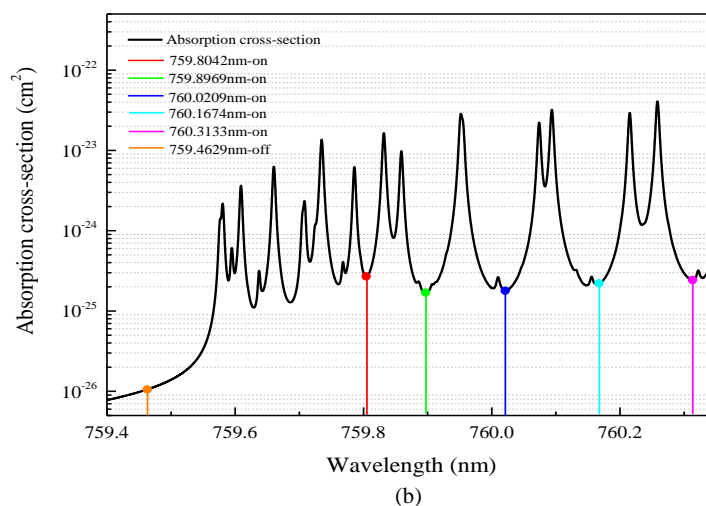
245 wavelength. Table 1 lists certain parameters of the six absorption lines: for example, the
 246 linear function $\sigma_q(\nu-\nu_{01})$ of the absorption cross-section of the oxygen absorption line P13Q12
 247 with respect to the wave number ν , and the line function $\sigma_p(\nu-\nu_{02})$ of the absorption cross
 248 section of the P13P13 with respect to the wave number ν . Further, the wavelength λ_{on}
 249 (wavenumber ν_{on}) we selected is located at the minimum of the absorption cross-section
 250 between the two spectral lines, that is, its absorption cross-section is the superposition of the
 251 values of the extension lines of two adjacent Voigt linear functions at ν_{on} . Moreover, its
 252 absorption cross-section $\sigma_{on}(\nu) = \sigma_p(\nu_{on}) + \sigma_q(\nu_{on})$ —the wings of the two Voigt lineshape
 253 functions—is the manifestation of their pressure expansion.

254 **Table 1 Parameters of the three groups of absorption lines of Oxygen A(296K) (Brown and Plymate, 2014)**

Assignment	Line center ν_0 cm ⁻¹	Intensity S_0 cm mole ⁻¹	Low energy E'' cm ⁻¹	Half Widths		Pressure-introduced shift Average(δ) cm ⁻¹ /atm	Temperature dependence n
				γ_0			
				γ_{air} cm ⁻¹ /atm	γ_{self} cm ⁻¹ /atm		
P13Q12	13078.2275	5.61×10^{-24}	260.6824	0.0466 (0.6)	0.0461 (1.1)	-0.0061	0.73
P13P13	13076.3273	6.13×10^{-24}	262.5827	0.0467 (0.8)	0.0460 (1.1)	-0.0068	0.73
P15Q14	13069.9619	4.33×10^{-24}	343.9694	0.0457 (1.2)	0.0449 (1.6)	-0.0051	0.73
P15P15	13068.0818	4.68×10^{-24}	345.8495	0.0455 (1.2)	0.0452 (1.6)	-0.0064	0.73
P17Q16	13061.3273	3.09×10^{-24}	438.7010	0.044	0.045	-0.00898	0.73
P17P17	13059.4665	3.31×10^{-24}	440.5618	0.0452	0.045	-0.00902	0.59



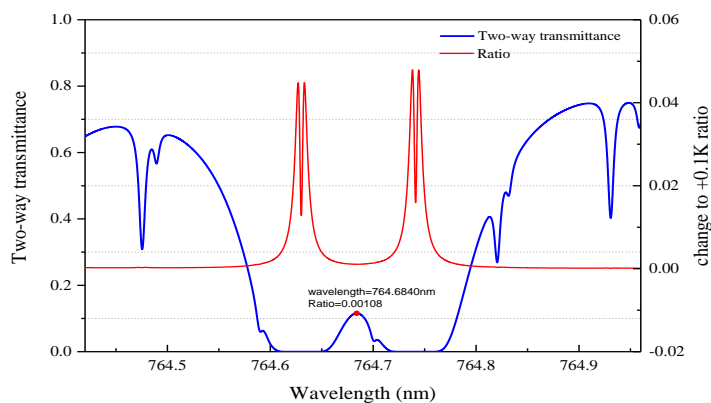
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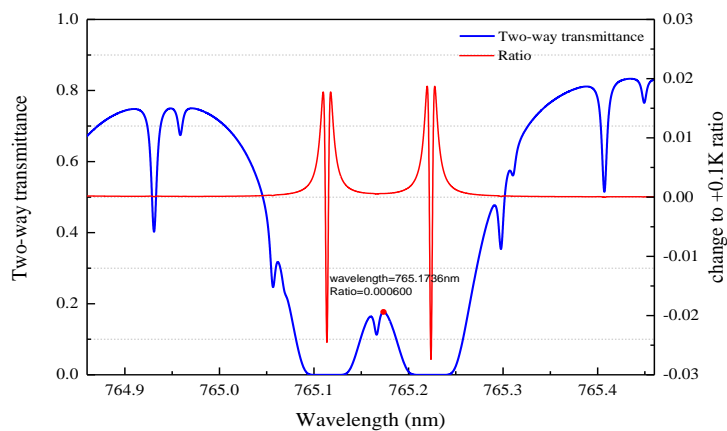
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Figure 2: (a)Near-ground absorption cross-section near 765nm,(b)Near-ground absorption cross-section near 760nm

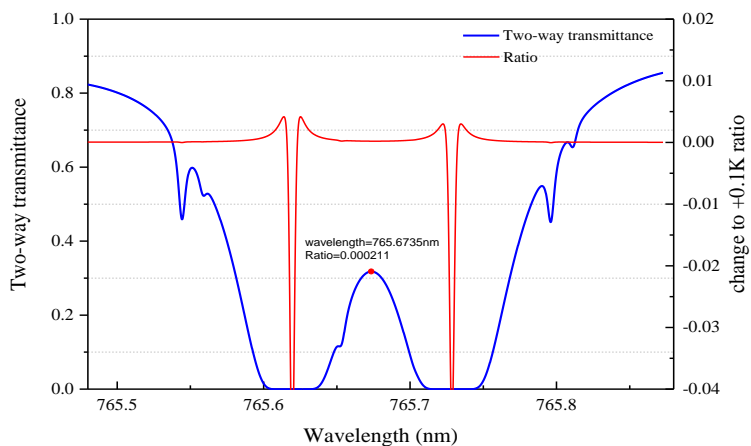
262 Within the A absorption band of oxygen (759–770nm), the spectral transmittances in the
263 vicinity of 760 and 765nm were relatively insensitive to temperature, this band is the low
264 interference of water vapor and carbon dioxide molecules. Further,the atmospheric
265 transmittance at 760nm is lower than the atmospheric transmittance at 765nm, and when
266 ground-based LIDAR (RR-DIAL) detects tropospheric backscattering owing to the round-trip
267 optical path being shorter, two adjacent lines near 760nm can be selected. In addition, the
268 plots of absorption spectra shown in Figure 2, the Oxygen absorption features with a number
269 of smaller and sharper absorption spikes, it contains isotopologues of Oxygen molecules,
270 showing some subtle differences, the wavelength in the middle of the trough area between
271 lines is more suitable as the ~~detection wavelength~~ online and the reference wavelength offline
272 (Korb et al., 1983,1989; Schwemmer, et al., 1987). When IPDA shoots lasers
273 from the satellite and receives the echo from the ground surface, the laser beam that passes through the
274 entire atmosphere twice results in the path being longer. Thus, 765nm is relatively more
275 suitable for remote sensing of surface pressure from satellites(Riris et al., 2017).



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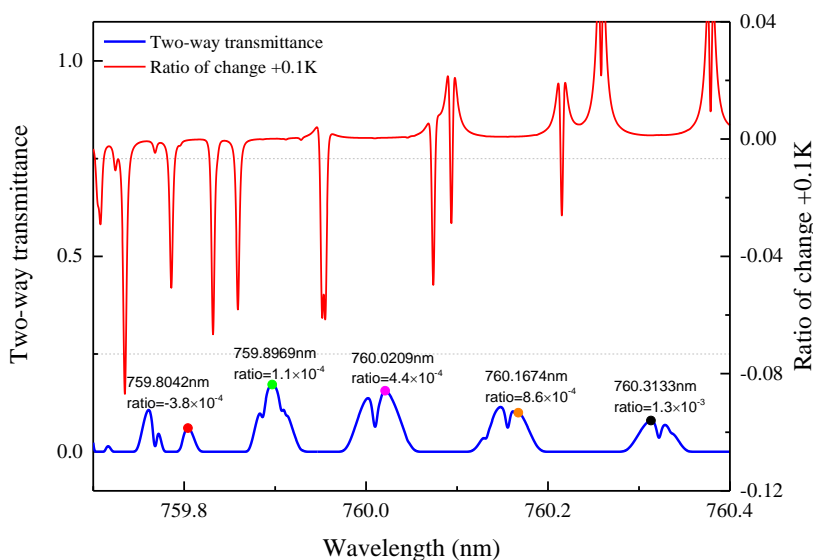


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Figure 3: Transmittance spectra of three intervals around 765nm and their ratio to 0.1K change in temperature



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Figure 4: Transmittance spectra around 760nm and their ratio to 0.1K change in temperature

284 The absorptive optical depth of oxygen $dOD(p_{surface})$ with respect to the path 0–71km
 285 corresponds to the optical transmittance e^{-2dOD} of this path with respect to the ν wavenumber.
 286 Figure 3 shows the transmittance spectra of three intervals around 765nm and their sensitivity
 287 to temperature changes of +0.1K. Similarly, Figure 4 shows the transmittance spectra of
 288 around 760nm and their sensitivity to temperature changes of +0.1K.
 289

290
 291

Table 2 differential optical depth and differential absorption cross section of 9 pairs of wavelengths in standard atmospheric mode

Wavelength(nm)	$\int_0^{71} (\alpha_m + \alpha_a) dz$	$\int_0^{71} (n_{O_2} \sigma) dz$	$OD(0,71km)$	$\sigma_{O_2} (cm^2)$	$dOD(0,71km)$	$\Delta\sigma_{O_2} (cm^2)$
on 764.6840	0.1852	0.493	0.678	2.08×10^{-25}	0.428	1.80×10^{-25}
off 764.9097	0.1851	0.0653	0.251	2.80×10^{-26}		
on 765.1600	0.1851	0.389	0.574	1.74×10^{-25}	0.325	1.46×10^{-25}
off 764.9707	0.1851	0.0638	0.249	2.77×10^{-26}		
on 765.1736	0.1849	0.373	0.558	1.67×10^{-25}	0.328	1.47×10^{-25}
off 765.3883	0.1849	0.0448	0.230	2.01×10^{-26}		



on	765.6735	0.1848	0.231	0.416	1.10×10^{-25}	0.192	9.25×10^{-26}
off	765.4637	0.1849	0.0391	0.224	1.78×10^{-26}		
on	759.8042	0.1872	0.515	0.702	2.60×10^{-25}	0.493	2.49×10^{-25}
off	759.4629	0.1873	0.0216	0.209	1.46×10^{-26}		
on	759.8969	0.1871	0.343	0.530	1.66×10^{-25}	0.321	1.55×10^{-25}
off	759.4629	0.1873	0.0216	0.209	1.46×10^{-26}		
on	760.0209	0.1871	0.385	0.573	1.77×10^{-25}	0.364	1.67×10^{-25}
off	759.4629	0.1873	0.0216	0.209	1.46×10^{-26}		
on	760.1674	0.1870	0.506	0.693	2.22×10^{-25}	0.484	2.12×10^{-25}
off	759.4629	0.1873	0.0216	0.209	1.46×10^{-26}		
on	760.3133	0.1870	0.576	0.763	2.46×10^{-25}	0.554	2.36×10^{-25}
off	759.4629	0.1873	0.0216	0.209	1.46×10^{-26}		

292

293 Equation (16) clearly indicates that in various factors that result in $\delta[dOD(p_{\text{surface}})]$, the error
 294 of differential optical depth conditionally causes conditionally causes the error $\delta p_{\text{surface}}$ of the
 295 surface pressure. Additionally, the detection wavelength absorption cross-section difference
 296 $\Delta\sigma(p_{\text{surface}})$ near the ground is inversely proportional to the surface pressure error. Evidently, a
 297 key factor affecting DIAL sensitivity is the online and offline wavelength positions. However,
 298 for the candidate wavelengths marked in Fig. 2 used as detection wavelengths, each would
 299 offer its own advantages and disadvantages, and consequently, comprehensive evaluation is
 300 required.

301 3.2 Differential absorption LIDAR system model

302 The research results reported by Coney, *et al.*(Munk et al., 2016-2019; Coney et al., 2021;
 303 Thomas et al., 2016; Strotkamp et al., 2019)along with those reported by Wulfmeyer
 304 and ösenberg et al.(1996), refer to the ADM-Aeolus in orbit ALADIN system parameters of
 305 the Aeolus mission (Lemmerz et al., 2017); the receiver is based on the GLAS-Mission-
 306 1064nm receiver, the orbit altitude is 400km, and diameter of the telescope is 1.5m.
 307 Consequently, the model parameters of the differential absorption LIDAR has been proposed,
 308 as shown in Table 3. The transmitter model parameters, with the exception of the pulse
 309 energy of 100mJ, have been separately reported in different documents(Coney et al.,
 310 2021).However, these indicators have been achieved in the same laser, and thus, more
 311 research is required.

312

Table 3 System parameters of differential absorption LIDAR

Transmitter	
Laser pulse energy	100mJ



Laser pulse Width	88ns
Pulse repetition rate	100Hz
Laser Divergence Angle	90 μ rad for $\pm 3\sigma$
Spectral purity	99.99%
Pointing stability	< 10 μ rad
Receiver	
Telescope Diameter (A_r)	1.5m(SiC)
Receiver Field-of-view (full)	100 μ rad
Optical Filter Bandwidth (FW)	0.8nm(FWHM)
Fabry-Perot Elton (thickness=2mm)	25pm(free spectral range \approx 0.1nm)
Receiver Efficiency	50%
Combined filter width	0.025nm
Detector and amplifier	
Detector(Laser Components DG, Inc)	Si-APD(SAR1500/C30956/S3884-04)
APD Quantum Efficiency (η_d)	75%
Detector Diameter	Φ 1.5mm
Electronic system bandwidth (BW)	3MHz
APD dark current (I_d)	1nA type
APD gain(M)	100
APD excess noise factor(F)	2.4
APD capacitance (C_d)	4pF
trans-impedance amplifier gain (R_f)	20kV/A
trans-impedance amplifier input current noise (I_{nA})	2.5pA/Hz ^{1/2}
trans-impedance amplifier input voltage noise (V_{nA})	20nV/Hz ^{1/2}
operate temperature	293 K
Platform and environment	
Orbit altitude and velocity	400 km, 7 km/s
Orbit type	Polar, sun synchronous, dawn/dusk
Along-track resolution	44 km
Simulation top altitude	71 km
Viewing geometry	Nadir
Atmosphere model	US standard atmosphere
Aerosol model	Median aerosol profile
(765nm)the surface albedo over ocean/land	0.1575/0.314
Pointing stability	< 50 μ rad
Spectroscopic data base	HITRAN 2012

313 As reported in reference (Lancaster et al., 2005), the equivalent Lambertian reflection
 314 coefficient of the sub-satellite point laser on the ocean surface has an empirical relationship
 315 $\rho_{eff} = \frac{\rho}{4(S^2)}$, where the Fresnel reflection coefficient is $\rho = 0.02$, and $\langle S^2 \rangle$ is the variance of
 316 the wave steepness distribution. Further, Bufton et al.(1983)and Menzies et
 317 al.(1998)individually adopted relationship as follows:



$$\langle S^2 \rangle = \begin{cases} (\ln U_{10} + 1.2) \times 10^{-2} \\ (0.85 \ln U_{10} - 1.45) \times 10^{-1} \end{cases} \quad U_{10} \leq 7.0 \text{ m/s} / U_{10} > 7.0 \text{ m/s} \quad (17)$$

where U_{10} is the wind speed of segment 10m above the ocean surface. The general ocean surface wind speed is taken as 8m/s, whereas $\rho_{\text{eff}} = 0.1575$ and $\rho_{\text{eff}}\pi = 0.025 \text{ sr}^{-1}$. In addition, the reflectivity of terrestrial lasers is generally 0.314.

3.3 Performance evaluation of A-band DIAL system

3.3.1 Random error of differential optical depth caused by noise

The number of received return echo photons $N_{s,\text{on}}$ and $N_{s,\text{off}}$ is obtained using the LIDAR Eq. (7) and Eq.(8). Equations(18)–(25) are commonly used for the on and off channels.

$$N_s(R_0, R) = \frac{\lambda \cdot E}{h \cdot c} \cdot A_r \cdot \eta_d \cdot \eta_r \cdot \left(\frac{\rho}{\pi}\right) \cdot \frac{\exp[-2OD(R_0, R)]}{(R_0 - R)^2} \quad (18)$$

Here the working and reference wavelengths λ_{on} and λ_{off} , the Planck constant is h , ρ is the surface reflectivity, and c is the speed of light. Further, the effective pulse width τ_w of the echo signal is a combination of the emitted laser pulse width τ_L , and the detection electronic system bandwidth BW (unit Hz), effective target altitude within the laser footprint ΔH , $R_0=400\text{km}, \Delta H=2 \text{ m}$, can be expressed as (Ehret, et al., 2008):

$$\tau_w = \sqrt{\tau_L^2 + \left(\frac{1}{3} \cdot BW\right)^2 + \left(\frac{2 \cdot \Delta H}{c}\right)^2} \quad (19)$$

The background signal N_{BG} (photoelectrons), assuming a Lambertian surface and zenith sun, is calculated as

$$N_{BG}(\lambda) = \frac{\lambda \cdot S_{BG}}{h \cdot c} \tau_w \cdot A_r \cdot \eta_d \cdot \eta_r \cdot \left(\frac{\rho}{\pi}\right) \cdot \left(\frac{FOV}{2}\right)^2 \pi \cdot FW \cdot \exp[-2OD(R_0, R)] \quad (20)$$

where S_{BG} is the exo-atmospheric solar irradiance value ($1.221 \text{ W m}^{-2} \text{ nm}^{-1}$) (ASTM international, United States, 2019) at 765nm. FW is the bandwidth of the optical filter ($0.025 \text{ nm} \times 4$) and the field of view (FOV) (unit rad) of the FOV receiving telescope. The bandwidth of the Fabry–Perot etalon, free spectral range, and width of the narrowband filter were 25pm, 0.1nm, and 0.8nm, respectively. There are 8 longitudinal modes of Fabry–Perot etalon that can pass through. However, the transmittance of each longitudinal mode is different, and thus, the equivalent solar window width is $25 \text{ pm} \times 4 = 0.1 \text{ nm}$. Further, the backscattering coefficient (ρ/π) on the surface of land such as ocean and vegetation during the daytime (Thomas et al., 2016), and ρ were calculated as 0.1575 and 0.314, respectively.

Here the q electrons charge is $1.6 \times 10^{-19} \text{ C}$, M is the gain of silicon avalanche diode (APD). The total noise associated with the detection signal is divided into fixed circuit noise and signal-dependent shot noise. The total circuit noise current spectral density (unit $\text{A}/\text{Hz}^{1/2}$) I_n , can be expressed as (Refaat, et al., 2013)

$$I_n = \sqrt{2 \cdot q \cdot I_d \cdot M \cdot F + I_{nA}^2 + \frac{V_{nA}^2}{R_f^2} + \frac{4 \cdot k \cdot T}{R_f + \frac{(2 \cdot \pi \cdot V_{nA} \cdot C_d \cdot BW)^2}{3}}} \quad (21)$$

where I_d and F are the dark current and excess noise factors of the detector, respectively; I_{nA} and V_{nA} are the preamplifier integrated input current and input voltage noise spectral density, respectively; R_f is the feedback resistance of the preamplifier; and C_d is the equivalent input capacitance of the amplifier and the detector. The circuit noise is often limited by the shot noise of the dark current of the detector or the noise of the preamplifier. In this analysis, all circuit noises refer to the detector input and the equivalent circuit noise-generated photoelectrons, and $N_{n,C}$, is calculated as



357
$$N_{n,C} = \frac{I_n \cdot \tau_w \cdot \sqrt{BW}}{q \cdot M} \quad (22)$$

358 Similarly, the equivalent shot noise-generated photoelectrons, $N_{n,S}$, are calculated as

359
$$N_{n,S} = \sqrt{2 \cdot N_S \cdot F \cdot \tau_w \cdot BW} \quad (23)$$

360 Further, the photoelectron $N_{n,BG}$, equivalent to the equivalent shot noise associated with the
 361 background radiation can be calculated as

362
$$N_{n,BG} = \sqrt{2 \cdot N_{BG} \cdot F \cdot \tau_w \cdot BW} \quad (24)$$

363 These noises are regarded as the equivalent photoelectron number generated in the detector
 364 (before the multiplication process), and are proportional to the actual detected photoelectron
 365 number. The total signal-to-noise ratio is expressed as follows(Ehret, *et al.*, 2008):

366
$$SNR_{on/off} = \frac{N_{s,on/off}}{\sqrt{N_{n,C}^2 + N_{n,S,on/off}^2 + N_{n,BG,on/off}^2}} \quad (25)$$

367 Where s is the number of echo signal pulses accumulated and averaged by the LIDAR. The
 368 error ε_R caused by the noise of the LIDAR receiving a single echo is a random error.
 369 Moreover, it is necessary to calibrate the LIDAR echo signal detection channel and the laser
 370 emission pulse energy monitoring channel to remove nonlinear and nonzero biased
 371 background voltage. Further, the calibration error and AD conversion error comprise the
 372 systematic error. In addition, the error ε_A caused by the uncertainty of atmospheric
 373 environment (atmospheric temperature profile, atmospheric water vapor mixing ratio profile),
 374 and the associated error ε_T of the laser emission characteristics (jitter of the center wavelength
 375 of the emitted beam, the emission spectrum width, and the purity of the emission spectrum)
 376 are all systematic errors. The total error of the differential optical depth can be expressed as

377
$$\delta[dOD] = \frac{\varepsilon_R}{\sqrt{s}} + \sqrt{\varepsilon_A^2 + \varepsilon_T^2} \quad (26)$$

378 Equation(9) indicates that the random noise of the echo signals $N_{on}(R)$ and $N_{off}(R)$ and the
 379 random measurement error of the pulse energies E_{on} and E_{off} result in the random error of the
 380 differential optical depth $\varepsilon_R = \delta[dOD(p_{surface})]_R$ as follows:

381
$$\varepsilon_R = \delta[dOD(p_{surface})]_R = \frac{1}{2} \sqrt{\left(\frac{\delta N_{s,on}(R)}{N_{s,on}(R)}\right)^2 + \left(\frac{\delta N_{s,off}(R)}{N_{s,off}(R)}\right)^2 + \left(\frac{\Delta E_{on}}{E_{on}}\right)^2 + \left(\frac{\Delta E_{off}}{E_{off}}\right)^2} \quad (27)$$

382 The signal-to-noise ratio of LIDAR can be calculated using Eq.(18)–(25). Simultaneously, it
 383 is considered that the measurement error of pulse energy $\Delta E_{on}/E_{on} \approx \Delta E_{off}/E_{off}$ is very small
 384 and can thus be ignored. The random error of LIDAR echo (noise) measurement is calculated
 385 as

386
$$\varepsilon_R = \frac{1}{2} \sqrt{SNR_{on}^{-2} + SNR_{off}^{-2}} \quad (28)$$

387 When the laser irradiates the ocean surface (for example, the average wind speed is 8m/s),
 388 0.1575 represents median for the ocean surface reflectivity(Ehret, *et al.*, 2008), and the
 389 random error ε_R of the differential optical depth is calculated considering the single pulse echo;
 390 with a time resolution of at least 6.25 s ,the distance resolution along the track of 44km and S
 391 $= 625$ laser pulse echoes are taken as a group for cumulative average, the random error of the
 392 atmospheric differential optical depth above the ocean surface is $\frac{\varepsilon_R}{\sqrt{625}}$.

393 The ocean surface possesses low laser reflectivity and weak echoes. The averaging
 394 method employed is as follows: first add up 625 echoes; thereafter subtract B and normalize,



395 where B is the level background baseline of the LIDAR output; and finally Eq. (25) provides
 396 the differential optical depth.

397
$$E_{on} = \sum_{i=1}^{625} E_{i,on} \quad (29-1)$$

398
$$E_{off} = \sum_{i=1}^{625} E_{i,off} \quad (29-2)$$

399
$$N_{s,on} = \sum_{i=1}^{625} N_{i,s,on} \quad (29-3)$$

400
$$N_{s,off} = \sum_{i=1}^{625} N_{i,s,off} \quad (29-4)$$

401
$$dOD = OD_{on} - OD_{off} = -\frac{1}{2} \ln\left(\frac{N_{s,on}-B}{E_{on}}\right) + \frac{1}{2} \ln\left(\frac{N_{s,off}-B}{E_{off}}\right) \quad (30)$$

402

403

Table 4 Error from noise

Wavelength(nm)	SNR		Single shot Noise Error		average Noise Error	
	ocean	land	ocean	land	Ocean/25	Land/12
	0.1575	0.314	0.1575	0.314	0.1575	0.314
764.6840	97.74	139.68	0.0061	0.0043	2.44×10^{-4}	3.55×10^{-4}
764.9097	152.1	215.8				
765.1600	109.0	155.5	0.0056	0.0040	2.24×10^{-4}	3.30×10^{-4}
764.9707	152.3	216.2				
765.1736	110.9	158.0	0.0055	0.0039	2.20×10^{-4}	3.24×10^{-4}
765.3883	155.4	220.5				
765.6735	128.4	182.6	0.0050	0.0035	2.00×10^{-4}	2.96×10^{-4}
765.4637	156.3	221.8				
759.8042	95.06	135.9	0.00614	0.00430	2.46×10^{-4}	3.58×10^{-4}
759.4629	158.1	224.3				
759.8969	113.7	161.9	0.00542	0.00381	2.17×10^{-4}	3.17×10^{-4}
759.4629	158.1	224.3				
760.0209	108.8	155.1	0.00558	0.00392	2.23×10^{-4}	3.27×10^{-4}
759.4629	158.1	224.3				
760.1674	95.98	137.2	0.00609	0.00427	2.44×10^{-4}	3.56×10^{-4}
759.4629	158.1	224.3				



760.3133	89.15	127.7				
759.4629	158.1	224.3	0.00644	0.00451	2.58×10^{-4}	3.76×10^{-4}

404

405 It is believed that the reflectance value of 0.314 is typical and representative of the surface
 406 reflectance for most features on land (vegetation, sand, and soil). When the laser irradiates the
 407 land, the noise of the single pulse echo causes a random error in the optical depth. The time
 408 resolution of 6.25 s and resolution along track of 44 km are maintained. First, the very high
 409 footprint is highlighted, and the very low footprint points is removed. Subsequently, the
 410 average altitude of most of the remaining footprints is calculated. Considering this average
 411 altitude, all the echo pulses whose footprint altitude and the average altitude are within 2 m as
 412 a group are acquired, and the number of pulses obtained is no less than 144 pulses. Further,
 413 they are accumulated and averaged to decrease the random measurement error of the
 414 differential optical depth.

415 The averaging method involves first subtracting the background baseline B_i of a single
 416 echo, normalizing, and thereafter performing the cumulative average, where

417
$$s \geq 144, \quad \frac{N_{s,on/off}}{E_{on/off}} \approx \sum_{i=1}^M \frac{N_{i,s,on/off} - B_i}{E_{i,on/off}} \quad (31)$$

418
$$dOD = OD_{on} - OD_{off} = -\frac{1}{2} \ln \left(\frac{N_{s,on}}{E_{on}} \right) + \frac{1}{2} \ln \left(\frac{N_{s,off}}{E_{off}} \right) \quad (32)$$

419 When the laser irradiates the plain area, it is easy to confidently pick out the footprint altitude
 420 of more than 144 shots from 625 pulses. However, if uneven terrain is encountered, there are
 421 less than 144 echo pulses, difference between the footprint altitude and the average altitude is
 422 within 2 m, and average of multiple laser pulse echoes becomes unreasonable. Moreover, if
 423 greater than 144 pulses with similar footprints are still not found, this set of data is discarded.

424 In conclusion, when laser irradiates the ocean surface, its single echo signal is relatively
 425 weak, whereas its footprint altitude is relatively consistent; thus, more pulse echoes can be
 426 accumulated and averaged. Further, when the laser irradiates the land surface, the altitude
 427 consistency of the landing footprint is poor and accumulated average echo pulse is less,
 428 although the signal-to-noise ratio of the land single echo is relatively high.
 429

430 3.3.2 Uncertainty of vertical distribution profile of atmospheric temperature

431 Within a certain time resolution (distance resolution), the uncertainty of the vertical profile of
 432 the atmospheric temperature results in an absolute systematic error with respect to the
 433 differential optical depth (Refaat, et al., 2013).

434 When the temperature change $\pm 1K$, Oxygen number density is

435
$$n_{O_2}(T \pm 1K) = \frac{0.20948p(z)}{k(T(z) \pm 1) \cdot (1 + \chi(z))} \quad (33)$$

436
$$\Delta[dOD(0, surface)]_{\pm 1K} = \int_0^{71} n_{O_2}(T \pm 1K) \Delta\sigma(p(z), T(z) \pm 1) dz$$

$$- \int_0^{71} n_{O_2}(T) \Delta\sigma(p(z), T(z)) dz \quad (34)$$

437 **Table 5 Temperature sensitivity of differential optical depth in case of $\pm 1K$ uncertainty**

Regimentation	$\Delta dOD(p, T - 1) $	$\Delta dOD(p, T + 1) $	$Max\{\Delta dc$ $- 1) , \Delta d$ $- +1) \}$
---------------	-------------------------	-------------------------	---



764.6840/ 764.9097	0.00216	0.00215	0.00216
765.1600/ 764.9707	0.000929	0.000933	0.000933
765.1736/ 765.3883	0.00102	0.00102	0.00102
765.6735/ 765.4637	0.000139	0.000146	0.000146
759.8042/ 759.4629	0.00141	0.00139	0.00141
759.8969/ 759.4629	0.000173	0.000162	0.000173
760.0209/ 759.4629	0.000624	0.000631	0.000631
760.1674/ 759.4629	0.00170	0.00170	0.00170
760.3133/ 759.4629	0.00285	0.00284	0.00285

438

439 **3.3.3 Error of differential optical depth $[dOD(0, p_{surface})]_{wv}$ caused uncertainty of the**
 440 **mixture ratio of water vapor**

441 The mixture ratio of near-ground water vapor in standard atmospheric mode is 1.247% higher
 442 than $\chi_{wv}(p_{surface})$.

443 The 20% uncertainty of profile for water vapor mixture ratio introduces uncertainty in
 444 differential optical depth $[dOD(0, p_{surface})]_{wv}$.

$$\Delta[dOD(p_{surface})]_{\pm 20\%wv} \approx 0.20948 \int_0^{71km} \frac{p(z)}{kT(z)(1 + (1 \pm 20\%) \chi(z))} (\sigma(\lambda_{on}, z) - \sigma(\lambda_{off}, z)) dz$$

$$445 - 0.20948 \int_0^{71km} \frac{p(z)}{kT(z)(1 + \chi(z))} (\sigma(\lambda_{on}, z) - \sigma(\lambda_{off}, z)) dz \quad (35)$$

446 **Table 6 Error differential optical depth $[dOD(0, p_{surface})]_{wv}$ caused uncertainty of the mixture of water vapor**

Wavelength(nm)	Differential optical depth error(20%)
764.6840	5.29×10^{-4}
764.9097	
765.1600	4.15×10^{-4}
764.9707	
765.1736	4.18×10^{-4}
765.3883	



447	$\frac{765.6735}{765.4637}$	2.53×10^{-4}
448	$\frac{759.8042}{759.4629}$	6.71×10^{-4}
449	$\frac{759.8969}{759.4629}$	4.27×10^{-4}
450	$\frac{760.0209}{759.4629}$	4.71×10^{-4}
451	$\frac{760.1674}{759.4629}$	6.11×10^{-4}
452	$\frac{760.3133}{759.4629}$	6.88×10^{-4}
453		
454		
455		
456		
457		
458		
459		

3.3.4 Error of the differential optical depth caused the difference in the altitude of the inner surface and between the land footprints.

The largest oxygen density is near the ground, and thus, differential optical depth is sensitive to high uncertainty near ground; $\Delta H=2\text{m}$.

$$\begin{aligned}
 \Delta[dOD(P_{surface})]_{\Delta H} &\approx \int_0^{71000} n_{O_2}(z)(\sigma(\lambda_{on}, z) - \sigma(\lambda_{off}, z))dz \\
 &\quad - \int_{\pm 2}^{71000} n_{O_2}(z)(\sigma(\lambda_{on}, z) - \sigma(\lambda_{off}, z))dz \\
 &\approx \int_0^{\pm 2} n_{O_2}(z)(\sigma(\lambda_{on}, z) - \sigma(\lambda_{off}, z))dz
 \end{aligned} \tag{36}$$

Table 7 Differential optical depth error caused by the 2 m altitude difference

Wavelength	Error [dOD]
764.6840	0.4578×10^{-4}
764.9097	
765.1600	0.3723×10^{-4}
764.9707	
765.1736	0.3737×10^{-4}
765.3883	
765.6735	0.2356×10^{-4}
765.4637	
759.8042	0.139×10^{-4}
759.4629	
759.8969	0.0905×10^{-4}
759.4629	
760.0209	
759.4629	0.102×10^{-4}



467	760.1674	0.136 × 10 ⁻⁴
	759.4629	
468	760.3133	0.156 × 10 ⁻⁴
469	759.4629	
470		

471 **3.3.5 Relative error in calibration for the echo and energy monitoring channels.**

472 The absolute error in the differential optical depth is $dOD \times 0.025\%$ (Ehret, et al., 2008),
 473 which also belongs to the systematic error.

474 **Table 8 Calibration error for echo detection channels and energy monitoring channels**
 475

Wavelength(nm)	$dOD(0, 71\text{km})$	error of dOD
764.6840	0.428	1.07 × 10 ⁻⁴
764.9097		
765.1600	0.325	0.81 × 10 ⁻⁴
764.9707		
765.1736	0.328	0.82 × 10 ⁻⁴
765.3883		
765.6735	0.192	0.48 × 10 ⁻⁴
765.4637		
759.8042	0.493	1.23 × 10 ⁻⁴
759.4629		
759.8969	0.321	0.803 × 10 ⁻⁴
759.4629		
760.0209	0.364	0.910 × 10 ⁻⁴
759.4629		
760.1674	0.484	1.21 × 10 ⁻⁴
759.4629		
760.3133	0.554	1.39 × 10 ⁻⁴
759.4629		

476

477 **3.3.6 Error in optical depth due to the wavelength dependence of aerosol scattering**

478 In the standard atmosphere mode, Mie and Rayleigh scatterings of the 765nm or 760nm dual-
 479 wavelength are similar but not equal. The coefficient C in Eq. (10) expresses the systematic
 480 error caused by the difference as follows:

481
$$C = \int_0^{71\text{km}} [\alpha_a(\lambda_{on}, z) + \alpha_m(\lambda_{on}, z)] dz - \int_0^{71\text{km}} [\alpha_a(\lambda_{off}, z) + \alpha_m(\lambda_{off}, z)] dz \quad (37)$$

482 **Table 9 Optical depth error caused by the wavelength dependence of the aerosol scattering**
 483
 484



	Wavelength(nm)	C
485	on 764.6840	0.90×10^{-4}
486	off 764.9097	
487	on 765.1600	0.76×10^{-4}
488	off 764.9707	
489	on 765.1736	0.86×10^{-4}
490	off 765.3883	
491	on 765.6735	0.84×10^{-4}
492	off 765.4637	
493	on 759.8042	1.39×10^{-4}
494	off 759.4629	
495	on 759.8969	1.77×10^{-4}
496	off 759.4629	
497	on 760.0209	2.28×10^{-4}
498	off 759.4629	
499	on 760.1674	2.88×10^{-4}
500	off 759.4629	
501	on 760.3133	3.47×10^{-4}
502	off 759.4629	
503	on 760.3133	3.47×10^{-4}
504	off 759.4629	
505	on 760.3133	3.47×10^{-4}
506	off 759.4629	

507 The errors in aerosol Mie scattering with the wavelength dependence error, spectral width
 508 error, and spectral purity error, can be eliminated by correction. Wavelength dependence of
 509 extinction coefficient of aerosol, $0.0 < k < 2.0$, k is uncertain and varies with the particle size,
 510 shape and concentration of aerosol, so it will bring systematic error of differential optical
 511 depth.
 512

513 3.3.7 Error in differential optical depth from the spectral purity of the on/off laser

514 The spectral purity ζ of the spaceborne IPDA LIDAR is 99.99% (Wulfmeyer and ösenberg et
 515 al.1996), which results in an increase in the on-channel echo and the absolute error of the
 516 optical depth. For a spectral purity of 100%, the relationship between the two on/off channel
 517 signals is considered to be $N_{s,on}' = N_{s,off}' e^{-2dOD}$ and $dOD = -\frac{1}{2} \ln \left(\frac{N_{s,on}'}{N_{s,off}'} \right)$. However,
 518 because the spectral purity is not 100%, but it only is ζ , the relationship between the two
 519 channel signals is approximately $N_{s,on} = N_{s,off} [(1 - \zeta) + \zeta e^{-2dOD}]$, $\ln \left(\frac{N_{s,on}}{N_{s,off}} \right) =$
 520 $\ln[(1 - \zeta) + \zeta e^{-2dOD}]$, and the spectral purity yield the following error in optical depth:

$$\begin{aligned}
 \varepsilon_{\xi} &= \frac{1}{2} \ln \left(\frac{N_{s,on}}{N_{s,off}} \right) - \frac{1}{2} \ln \left(\frac{N_{s,on}'}{N_{s,off}'} \right) \\
 521 \quad &\approx dOD + \frac{1}{2} \ln[1 - \zeta(1 - e^{-2dOD})] \quad (38)
 \end{aligned}$$



522

Table 10 Error in differential optical depth from spectral purity of 99.99%

Wavelength	$dOD(0, 71\text{km})$	Error (spectral purity of 99.99%)
764.6840	0.428	6.79×10^{-5}
764.9097		
765.1600	0.325	4.60×10^{-5}
764.9707		
765.1736	0.328	4.66×10^{-5}
765.3883		
765.6735	0.192	2.36×10^{-5}
765.4637		
759.8042	0.493	8.40×10^{-5}
759.4629		
759.8969	0.322	4.51×10^{-5}
759.4629		
760.0209	0.364	5.35×10^{-5}
759.4629		
760.1674	0.484	8.16×10^{-5}
759.4629		
760.3133	0.554	1.01×10^{-4}
759.4629		

523

524 3.3.8 Error of differential optical depth caused jitter of ν in central optical frequency

525 The laser source of the spaceborne IPDA has a jitter of ν central optical frequency, whereas
 526 the online/offline emission laser spectrum has a central frequency jitter of 10MHz(Strotkamp
 527 et al., 2019). $\Delta\nu_{on} = \pm 10\text{MHz}$ introduces uncertainty in the Oxygen molecule absorption cross-
 528 section $\sigma(\nu_{on})$, resulting in a systematic error in the optical depth. However, for solid-state
 529 lasers, stabilizing the center wavelength within $\pm 10\text{MHz}$ is quite easy compared to within
 530 $\pm 1\text{MHz}$. For example, if the center wavelength is offset by $\pm 10\text{MHz}$, the jitter of the emitted
 531 laser frequency that affects the swing of the optical depth can be expressed as

532

$$\begin{aligned}
 [dOD(0, surface)]_{10\text{MHz}} &= \int_0^{71\text{km}} \frac{0.20948p(z)}{kT(z)(1+\chi(z))} (\sigma(\nu_{on} \pm 10\text{MHz}, z) - \sigma(\nu_{off}, z)) dz \\
 &- \int_0^{71\text{km}} \frac{0.20948p(z)}{kT(z)(1+\chi(z))} (\sigma(\nu_{on}, z) - \sigma(\nu_{off}, z)) dz \quad (39)
 \end{aligned}$$

534 It is evident from the curves of the two bands in Fig.5 that because the former is next to an
 535 absorption line $13069.062119\text{cm}^{-1}$ of the oxygen isotope $^{16}\text{O}^{18}\text{O}$, the curvature at the former
 536 position protrudes, and the abscissa of the curve is the laser frequency (Wavelength).

537 **Table 11 Jitter in the emission laser frequency ν causes a 10MHz change in the differential optical depth**



538

539

540

$\lambda(\text{nm})$	$dOD(v \pm 10\text{MHz}, p) - dOD(v, p)$		$\text{Max}\{ \Delta(+10\text{MHz}) , \Delta(-10\text{MHz}) \}$
	$ \Delta(+10\text{MHz}) $	$ \Delta(-10\text{MHz}) $	
764.6840	6.18×10^{-7}	9.62×10^{-7}	9.62×10^{-7}
764.9097			
765.1600	2.23×10^{-5}	2.38×10^{-5}	2.38×10^{-5}
764.9707			
765.1736	2.00×10^{-6}	2.81×10^{-6}	2.81×10^{-6}
765.3883			
765.6735	1.04×10^{-6}	8.53×10^{-7}	1.04×10^{-6}
765.4637			
759.8042	4.28×10^{-3}	2.52×10^{-3}	4.28×10^{-3}
759.4629			
759.8969	1.56×10^{-3}	1.19×10^{-3}	1.56×10^{-3}
759.4629			
760.0209	7.71×10^{-4}	7.76×10^{-4}	7.76×10^{-4}
759.4629			
760.1674	6.84×10^{-4}	1.60×10^{-3}	1.60×10^{-4}
759.4629			
760.3133	3.83×10^{-4}	9.49×10^{-4}	9.49×10^{-4}
759.4629			

541 According to the comprehensive analysis of the above differential optical depth errors
 542 induced in many factors, the comprehensive evaluation index items of the wavelength to be
 543 selected are shown in Figure 12 and Figure 13.
 544

545

Table 12 Comprehensive of various errors at near 765nm

Wavelength		764.6840	765.1600	765.1736	765.6735
		/764.9097	/764.9707	/765.3883	/765.4637
Random error	ocean	2.44×10^{-4}	2.24×10^{-4}	2.20×10^{-4}	2.00×10^{-4}
	land	3.55×10^{-4}	3.30×10^{-4}	3.24×10^{-4}	2.96×10^{-4}
Temperature	1K	21.54×10^{-4}	9.31×10^{-4}	10.16×10^{-4}	1.43×10^{-4}
Vapor mixing ratio	20%	5.29×10^{-4}	4.15×10^{-4}	4.18×10^{-4}	2.53×10^{-4}
Energy monitor channel calibration		1.07×10^{-4}	0.813×10^{-4}	0.820×10^{-4}	0.480×10^{-4}
Echo channel calibration		1.07×10^{-4}	0.813×10^{-4}	0.820×10^{-4}	0.480×10^{-4}



Elevation 2m error	0.458×10^{-4}	0.372×10^{-4}	0.374×10^{-4}	0.236×10^{-4}
Aerosol Mie scattering	0.903×10^{-4}	0.756×10^{-4}	0.857×10^{-4}	0.837×10^{-4}
99.99% spectral purity	0.679×10^{-4}	0.460×10^{-4}	0.466×10^{-4}	0.236×10^{-4}
Frequency jitter(10MHZ)	9.62×10^{-7}	0.238×10^{-4}	2.81×10^{-6}	1.04×10^{-6}
Differential absorption cross section (m^2)	1.80×10^{-29}	1.46×10^{-29}	1.47×10^{-29}	9.25×10^{-30}
Geometrically added	25.8×10^{-4}	19.6×10^{-4}	14.3×10^{-4}	6.08×10^{-4}
Absolute error(hPa)	3.29	3.08	2.23	1.51
Relative error(%)	0.324	0.304	0.220	0.150

546

547

Table 13 Comprehensive of various errors at near 760nm

Wavelength		759.8042 /759.4629	759.8969 /759.4629	760.0209 /759.4629	760.1674 /759.4629	760.3133 /759.4629
Random error	ocean	2.59×10^{-4}	2.09×10^{-4}	2.22×10^{-4}	2.64×10^{-4}	2.96×10^{-4}
	land	3.78×10^{-4}	3.06×10^{-4}	3.25×10^{-4}	3.84×10^{-4}	4.29×10^{-4}
Temperature	1K	26.2×10^{-4}	4.64×10^{-4}	9.10×10^{-4}	27.5×10^{-4}	47.8×10^{-4}
Vapor mixing ratio	20%	7.68×10^{-4}	4.86×10^{-4}	5.38×10^{-4}	7.03×10^{-4}	7.95×10^{-4}
Energy monitor channel calibration		1.91×10^{-4}	1.25×10^{-4}	1.40×10^{-4}	1.96×10^{-4}	2.28×10^{-4}
Echo channel calibration		1.91×10^{-4}	1.25×10^{-4}	1.40×10^{-4}	1.96×10^{-4}	2.28×10^{-4}
Elevation 2m error		0.22×10^{-4}	0.14×10^{-4}	0.16×10^{-4}	0.22×10^{-4}	0.26×10^{-4}
Aerosol Mie scattering		1.21×10^{-4}	1.59×10^{-4}	2.09×10^{-4}	2.69×10^{-4}	3.29×10^{-4}
99.99% spectral purity		0.840×10^{-4}	0.451×10^{-4}	0.535×10^{-4}	0.816×10^{-4}	1.01×10^{-4}
Frequency jitter(10MHZ)		42.8×10^{-4}	15.6×10^{-4}	7.76×10^{-4}	16.0×10^{-4}	9.49×10^{-4}
Differential absorption cross section (m^2)		2.45×10^{-29}	1.51×10^{-29}	1.63×10^{-29}	2.08×10^{-29}	2.32×10^{-29}
Geometrically added		54.6×10^{-4}	20.2×10^{-4}	16.7×10^{-4}	36.7×10^{-4}	53.9×10^{-4}
Absolute error(hPa)		5.12	3.07	2.35	4.05	5.34
Relative error(%)		0.505	0.302	0.232	0.400	0.527

548

549 Consequently, according to Table 12 and Table 13, the pulse energy, pulse repetition, time
 550 resolution, and distance along track are 100mJ, 100Hz, 6.25 s, and 44km, respectively. The
 551 765.6735/765.4637 wavelength pairs are used as detection wavelength and reference
 552 wavelength. In the 1K error temperature profile, 20% error vapor mixing ratio result in $6.08 \times$
 553 10^{-4} error in differential optical depth, which corresponds to an absolute error in surface
 554 pressure of 1.51hPa, therefore, it is a desirable result that the relative error of surface pressure
 555 could be considered as 0.150%.



556 **4 Summary**

557 The calculation process of retrieving the surface pressure from the atmospheric differential
558 optical depth was also discussed. The performance of the differential absorption LIDAR
559 model was evaluated. Owing to the influence of temperature on Oxygen absorption
560 coefficient and the uncertainty of atmospheric mixing ratio, maintaining the relative error of
561 surface atmospheric pressure below 0.1% is a challenging task. The main factors affecting the
562 random error of surface pressure are the low sea reflectivity, random error of low signal-to-
563 noise ratio, and the uneven ground, which renders the multi pulse echo unable to be
564 accumulated and averaged directly. Further, 765.6735/765.4637nm was selected as the
565 working wavelength, the pulse energy, pulse repetition, time resolution, distance along the
566 track resolution are 100mJ, 100Hz, 6.25 s, and 44km, respectively, while the relative error of
567 surface atmospheric pressure was controlled in the range of 0.2–0.3%.

568

569 **Author contribution**

570 Guanglei Hong developed the model and Yu Dong performed the simulations, was
571 responsible for data processing and software code. Guanglei Hong provided part of the
572 manuscript and analyzed the method respectively. Huige Di supervised and modified the
573 manuscript.

574

575 **Competing interests**

576 I declare that there is no conflict of interest.

577

577 **References**

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