Aerosol properties derived from COCCON ground-based Fourier Transform spectra

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Abstract. Fourier Transform Infrared (FTIR) spectroscopy is particularly relevant for climate studies due to its ability to provide information on both fine absorption structures (i.e. trace gases) and broadband continuum signatures (i.e. aerosols or water continuum) across the entire infrared (IR) domain. In this context, this study assesses the capability of the portable and compact EM27/SUN spectrometer, used within the research infrastructure COCCON (COllaborative Carbon Column Observing Network), to retrieve spectral aerosol properties from low-resolution FTIR solar absorption spectra (0.5 cm^{-1}). The study focuses on the retrieval of Aerosol Optical Depth (AOD) and its spectral dependence in the 873-2314 nm spectral range from COCCON measurements at the subtropical high-mountain Izaña Observatory (IZO, Tenerife, Spain), which were coincidentally carried out with standard sun photometry within the Aerosol Robotic Network (AERONET) in the 3-year period from December 2019 to September 2022. The co-located AERONET-COCCON database ~~in the 2021-2022 period (post-COVID-19 lockdown)~~ was used to cross-validate these two independent techniques in the common spectral range (870-1640 nm), demonstrating an excellent agreement at the near-coincident spectral bands (mean AOD differences limited to 0.005, standard deviations up to ~~0.019~~ 0.021 and Pearson regression coefficients up to ~~0.99~~ 0.97). This indicates that the low-resolution COCCON instruments are suitable for detecting the aerosol broadband signal contained in the IR spectra in addition to the retrieval of precise trace gas concentrations provided ~~that its calibration is performed frequently enough~~ a robust calibration procedure (Langley-based or absolute calibration procedures) is used to compensate for the optical degradation of the external system ($\sim$~~0.6~~0.72% per month). The study also assesses the capability of the EM27/SUN to simultaneously infer aerosols and trace gases, and relate their common emission sources in two case study events: a volcanic plume from the La Palma eruption in 2021 and a nearby forest fire in Tenerife in 2022. Overall, our results demonstrate the potential of the portable low-resolution COCCON instruments to enhance the multi-parameter capability of the FTIR technique for atmospheric monitoring.

1 Introduction

Atmospheric aerosols play a significant role in climate. Despite their short lifetime in the troposphere, aerosols are the second-largest contributor to climate change (Li et al., 2022). They are recognized as an "essential climate variable" by the Global Climate Observing System (GCOS) World Meteorological Organization (WMO) program (Bojinski et al., 2014) and a key focus area for the WMO Global Atmosphere Watching (GAW) program. Over the years, the scientific community has increasingly focused on atmospheric aerosols, and the most recent assessment report by the Intergovernmental Panel on Climate Change (IPCC) estimated the total effective radiative forcing (ERF) due to aerosols over the industrial era to be negative (-1.1 Wm^{-2}), with a reduced level of uncertainty compared to the previous IPCC report (Forster et al., 2021). This enhanced understanding of aerosol processes has been made possible by a combination of improved observational analyses and data-assimilated reanalyses (Ma et al., 2014; Benedetti et al., 2018; Rémy et al., 2018; Kinne, 2019; Bellouin et al., 2020; Chen et al., 2022) together with improved modelling approaches (Forster et al., 2021, and references therein), and can be considered relevant in the context of near-term and long-term climate mitigation and decarbonization strategies (Dreyfus et al., 2022).

To further advance our understanding of atmospheric aerosols, WMO considers the development of new, reliable, and ~~SI-traceable~~ International System of Units (SI)-traceable measurement techniques, and non-conventional measurement methods with open availability of validation data, as a core activity (WMO, 2010, 2017). Ground-based remote sensing has made recent advancements in this regard, allowing us to monitor aerosol micro-physical and optical properties from observations of atmospheric spectral transmission with excellent space and temporal (long-term) coverage (Holben et al., 2001; Smirnov et al., 2002; Torres et al., 2017; Cuevas et al., 2019b; Nakajima et al., 2020; Karanikolas et al., 2022, and among others).

Spectrometers and spectroradiometers provide excellent spectral performance across a wide range of the solar irradiance spectrum (Gröbner et al., 2017; Gröbner and Kouremeti, 2019; García-Cabrera et al., 2020; Egli et al., 2022; García et al., 2023), enabling radiometric information to be extended to other spectral regions far from the common ultraviolet (UV) and visible (VIS) regions (Barreto et al., 2020; Frausto-Vicencio et al., 2023). For instance, recent works have shown that high-resolution solar spectroscopic measurements from Fourier Transform Infrared (FTIR) spectrometers can be used to retrieve the spectral aerosol optical depth (AOD) and their spectral dependence in the infrared (IR) region, despite the instrument not being designed for absolute radiometric stability required for aerosol monitoring (Barreto et al., 2020; Frausto-Vicencio et al., 2023). Monitoring aerosols in the IR spectral range can provide useful information for a better characterization of atmospheric aerosols, especially for large aerosols like mineral dust, which can effectively interact with both solar and terrestrial radiation (Clarisse et al., 2013; Barreto et al., 2020). The significant impact of scattering and absorption processes in the longwave region, caused by these large particles (Dufresne et al., 2002; Does et al., 2018), may introduce a potential bias towards a radiative cooling effect in existing dust models (Ryder et al., 2019).

This paper focuses on the use of solar spectroscopic ground-based measurements from an EM27/SUN FTIR spectrometer (Gisi et al., 2012; Hase et al., 2016) to retrieve AOD at eight different bands in the near-infrared (NIR) and short-wave infrared (SWIR) spectral regions. The near-infrared (NIR) and short-wave infrared (SWIR) spectral regions are portions of the electromagnetic spectrum that extend beyond the visible light range. NIR refers to wavelengths between approximately 700

to 900 nanometers, while SWIR refers to wavelengths between approximately 900 to 2500 nanometers. These regions have
unique properties that make them valuable for various applications, including remote sensing, spectroscopy, and imaging. The
EM27/SUN is a portable, compact, and low-resolution FTIR (0.5 cm^{-1}) spectrometer designed to measure the main green-
house gases (GHG) in the atmospheric column and is used for this purpose by the COllaborative Carbon Column Observing
Network (COCCON) (Frey et al., 2019; Alberti et al., 2022). The simultaneous retrieval of fine absorption structures (i.e.
trace gases) and broadband continuum signatures (i.e. aerosols) across the entire NIR and SWIR domains with the same de-
vice, within an extended network like COCCON, is a promising opportunity to develop integral GHG-aerosol infrastructures
to target current global challenges in atmospheric composition monitoring. Additionally, the wealth of spectral information
available from the EM27/SUN instruments opens up new perspectives related to satellite validation and improves the spectral
characterization of aerosol particles. This, in turn, will help to improve the accuracy in estimating the aerosol radiative effects
on climate, especially for larger particles (Ryder et al., 2019; Clarisse et al., 2013).

The EM27/SUN aerosol properties in the 873–2314 nm spectral range have been determined at the subtropical high-
mountain Izaña Observatory (IZO, Tenerife, Spain), using COCCON measurements in a 3-year period from December 2019 to
September 2022. Firstly, the consistency and quality assessment of this new AOD database is addressed by comparing them to
high-precision aerosol observations performed within the Aerosol Robotic Network (AERONET). Secondly, the potential of
the EM27/SUN spectrometer to simultaneously monitor aerosols and trace gases for two case studies (forest fire and volcanic
emissions) has been evaluated. To address all these tasks, this paper is structured as follows: Sects. 2 and 3 describe the test site
and the instrumentation used (i.e. the EM27/SUN FTIR spectrometer and Cimel CE318-T radiometer), respectively. Section
4 introduces the methodology developed in this study to retrieve the EM27/SUN AOD products and to ensure their quality,
while Section 5 addresses the consistency and intercomparison study, as well as discusses the multi-parameter capability of the
EM27/SUN spectrometer to monitor aerosols and trace gases. Finally, Section 6 summarizes the most significant results and
conclusions drawn from this work.

2 Test Site

The Izaña Observatory (IZO) is a high-altitude observatory situated in Tenerife (Canary Islands, Spain) at 2373 meters above
sea level (at 28.3°N, 16.5°W). Managed by the Izaña Atmospheric Research Center (IARC), which is part of the State Me-
teorological Agency of Spain (AEMET), IZO is located in the North Atlantic subtropical zone that is well-known for its free
troposphere conditions, with a dry atmosphere and pristine skies. These conditions are a result of its elevation as well as its lo-
cation in the descending branch of the Northern Hadley's circulation cell (Cuevas et al., 2017, 2019b, a, and references therein).
Long-term atmospheric monitoring has been recorded under a comprehensive measurement program at IZO contributing to
the WMO GAW program since 1984. In addition, IZO was designated as a Commission for Instruments and Methods of
Observation (CIMO) test bed for aerosols and water vapour remote sensing instruments by the WMO in 2014 (WMO, 2014).

Many international networks for atmospheric monitoring activities have been developed at IZO. Among them, the FTIR
program was established in 1999 in the framework of a collaboration between the AEMET and the KIT (Karlsruhe Institute of

Technology), with the main goals being the long-term monitoring of atmospheric gas composition and the validation of satellite remote sensing measurements and climate models (Schneider et al., 2005; García et al., 2021). Based on high-resolution FTIR measurements (an IFS 125HR spectrometer), these activities have routinely contributed to two of the most prestigious

90 networks for atmospheric gas composition monitoring: NDACC (Network for the Detection of Atmospheric Composition Change, https://www.ndaccdemo.org, last access: 1st March 2023) since 1999, and TCCON (Total Carbon Column Observing Network, https://tccon-wiki.caltech.edu, last access: 1st March 2023) since 2007. Since 2018 NDACC and TCCON activities have been complemented by a portable, low-resolution FTIR spectrometer (a Bruker EM27/SUN), which operates within the COCCON research infrastructure (Frey et al., 2019; Alberti et al., 2022). More information about IZO and its atmospheric

95 monitoring programs can be found in Cuevas et al. (2019a).

IZO is a widely-used site for optimal application of the Langley-Plot technique (Shaw, 1982, 1983) for a variety of activities, such as zero airmass extrapolations for direct sun/moon calibrations or reference radiometric observations (Barreto et al., 2016; Toledano et al., 2018; Cuevas et al., 2019b; García-Cabrera et al., 2020). The observatory serves as reference calibration site for two of the most important ground-based photometer aerosol networks: National Aeronautics and Space Administra-

100 tion - Aerosol Robotic Network (NASA-AERONET; https://aeronet.gsfc.nasa.gov/; last access: 30th Abril 2023) and GAW precision-filter radiometers (GAW-PFR; http://www.pmodwrc.ch/worcc/; last access: 30th April 2023). Additionally, IZO is used for retrieving the reference top-of-atmosphere (TOA) solar spectrum from the ground at different spectral ranges (Bolsee et al., 2014; Gröbner et al., 2017). Other relevant aspect of IZO is its proximity to the African Continent (approximately 300 km), where the Earth's largest sources of mineral dust are located (Rodríguez et al., 2011; Prospero and Mayol-Bracero, 2013).

105 This fact makes IZO a unique and strategic location for studying dust transport and dust physical-chemical properties (e.g. Rodríguez et al., 2011; Cuevas et al., 2017; Barreto et al., 2022a, b).

Aerosol characterization studies have shown that summer months are related to high AOD conditions at IZO due to the prevalence of Saharan dust transport, with AOD at 500 nm above 0.1, which can be associated with the presence of large particles (Angstrom Exponent, AE) at 440-870 nm below 0.25 for pure mineral dust (Guirado, 2015; Berjón et al., 2019;

110 Barreto et al., 2022a, b). Cleaner conditions can be found at this station for the rest of the year, with AOD at 500 nm below 0.1 and AE above 0.6. Overall, mineral dust is the predominant aerosol in the North Atlantic free troposphere, where IZO is located, ensuring that the IR spectral AOD characterization performed in this study is mainly focused on mineral dust particles. Exceptionally, the 2021 La Palma eruption gave us the extraordinary opportunity to study volcanic aerosols during our period of study (Milford et al., 2023; García et al., 2023).

115 **3 Instrumentation**

3.1 Bruker EM27/SUN

The Bruker EM27/SUN is a robust and portable FTIR spectrometer, which has been developed by KIT in collaboration with Bruker Optics GmbH (Gisi et al., 2012; Hase et al., 2016). This instrument, based on a RockSolid^{TM} pendulum interferometer, acquires solar absorption spectra in the near-infrared region from 4000 to ~~11000~~ 11500 cm^{-1} with a spectral resolution of 0.5

120 cm^{-1} (maximum optical path difference, OPD$_{max}$, of 1.8 cm), using a CaF$_2$ beamsplitter and two InGaAs photodetectors. The primary detector covers the spectral section between 5500 and 11500 cm^{-1}, while the secondary detector covers the 4000–5500 cm^{-1} region (hereafter referred to as InGaAs-1 and InGaAs-2, respectively) (Hase et al., 2016). The fieldstop of 0.6 mm diameter limits the field-of-view (FOV), leading to a narrow semi FOV angle of only 2.96 mrad. While taking solar spectra, the EM27/SUN FOV is automatically centred at the solar disc by means of the
125 solar Camtracker system (Gisi et al., 2011).

In order to increase the signal-to-noise ratio, five forward and backward scans are routinely added in each measurement, thereby the acquisition of one spectrum lasts about one minute. The detectors' signal is DC coupled, allowing corrections for solar intensity variations due, for example, to thin clouds (Keppel-Aleks et al., 2007). Then, spectra are generated from the raw interferograms (IFGs) using a preprocessor tool developed by KIT in the framework of the COCCON-PROCEEDS I - III
130 projects funded by the European Space Agency (ESA). This tool includes the so-called DC correction of the IFGs, a dedicated phase correction scheme for double-sided IFGs, and several quality control tests (Sha et al., 2020; Frey et al., 2021). These quality filters will be used in the quality-control procedure established in this work (see details in Section 4.3).

In this work, EM27/SUN solar spectra are used to retrieve aerosol properties. Nonetheless, this FTIR spectrometer is mainly devoted to monitoring different atmospheric trace gas concentrations (O$_2$, CO$_2$, CH$_4$, CO, H$_2$O, among others) under different
135 conditions (e.g. metropolitan areas, mining regions or volcanic eruptions, Hase et al., 2015; Butz et al., 2017; Frey et al., 2019; Tu et al., 2022, and reference therein). The IZO EM27/SUN is operated in accordance with COCCON requirements (Frey et al., 2019; Alberti et al., 2022). This guarantees strict common methods for ensuring the quality of measurements (evaluation of the optical alignment and instrumental line shape), proper calibration of all COCCON spectrometers with respect to the TCCON site Karlsruhe and the COCCON reference EM27/SUN spectrometer operated permanently at KIT (in terms
140 of the standard retrieved species), and adherence to the COCCON data analysis scheme ensures the generation of precise and accurate data products. The COCCON gas data processing and analysis tools are freely available at the COCCON website (https://www.imk-asf.kit.edu/english/COCCON.php, last access: 24 April 2023).

3.2 Cimel CE318-T Radiometer

The Cimel CE318-T radiometer, referred to as CE318-AERONET hereafter, is the reference instrument in AERONET for direct
145 solar and diffuse sky measurements. It utilizes a two-axis robot to automatically point to the sun and a four-quadrant sensor installed in the radiometer's head for each direct solar/lunar measurement. Solar/lunar and sky measurements are normally taken every ∼15 minutes or at fixed air mass intervals at specific wavelengths, with a FOV of ∼1.3° (Holben et al., 1998; Torres et al., 2013). In the case of photometric information used in this paper, Cimel solar observations have been retrieved with a higher frequency, between 2 and 6 minutes. The instrument is equipped with Silicon and InGaAs detectors and nine
150 narrow band-pass filters, allowing for the recording of wavelengths at 340, 380, 440, 500, 675, 870, 940, 1020, and 1640 nm. The bandwidth, characterized by the full-width-at-half-maximum (FWHM), ranges from 2-4 nm in the UV (340-380 nm) to 10 nm in the visible and near-infrared (440-1020 nm) and 25 nm in the shortwave infrared (1640 nm). Additionally, the radiometer has an integrated barometer for atmospheric pressure measurements.

AOD, with its spectral variations, can be calculated from the direct solar/lunar measurements, and the aerosol's microphys-
ical properties can be retrieved from multi-angular observations using operational inversion algorithms (Holben et al., 1998;
Giles et al., 2019; Dubovik and King, 2000). The AOD total uncertainty for CE318-AERONET, from UV to NIR, was esti-
mated to be between 0.002-0.009 and 0.01-0.02 for reference and field instruments, respectively, with higher errors expected
for UV spectral bands (Eck et al., 1999). A specific analysis is still required to determine the AOD uncertainty in the SWIR
(1640 nm). However, the estimated uncertainty of 0.004, determined in Barreto et al. (2020) by comparing the AERONET
1640 product to high-resolution FTIR AOD retrievals, could be considered as a conservative uncertainty value for this band.

The current study utilizes the level 1.5 version 3 AOD products from the AERONET database, which provides cloud-
screened and quality-controlled AOD data.

4 Methodology

4.1 Micro-windows Selection

In order to estimate AOD from FTIR solar absorption spectra, it is necessary to use atmospheric windows with high solar
transmission, as this approach avoids strong absorption bands of gases and reduces the impact of gaseous signatures on the
EM27/SUN AOD retrievals.

EM27/SUN solar spectrum for the 870-2500 nm spectral region, acquired at a spectral resolution of 0.5 cm^{-1} with the
detectors InGaAs-1 and InGaAs-2, displayed as light grey and grey curves, respectively. Note that both detectors have different
gains (greater for InGaAs-2, i.e., B6-B8 micro-windows in SWIR), therefore the observed spectral behaviour is not the one
expected for the solar radiance: the higher the wavelength, the lower the measured radiation. Dashed coloured vertical arrows
show the position of the eight narrow spectral bands selected for the EM27/SUN AOD retrievals (B1-B8).

In this study, we followed a similar procedure for the AOD calculation as described by Barreto et al. (2020). We used low-
resolution (0.5cm^{-1}) EM27/SUN solar absorption spectra, acquired at IZO in the NIR and SWIR spectral ranges, between
4000 to $\frac{11000}{}$11500 cm^{-1} (corresponding to 870-2500 nm, Figure ??1). As can be read off from Figure 1, the water vapour
absorption bands are clearly distinguished from the atmospheric transparent windows normally used in remote sensing studies.
Within these atmospheric windows, we selected eight narrow micro-windows for the EM27/SUN AOD retrievals at the central
wavelengths of 872.55, 1020.90, 1238.25, 1558.25, 1636.00, 2133.40, 2192.00, and 2314.20 nm (hereafter referred to as B1,
B2, B3, B4, B5, B6, B7, and B8 bands, respectively, Table 1). In this study, Seven of the presented spectral bands (B2-B8)
were selected with respect to those presented in Barreto et al. (2020), while an additional channel (B1) has been incorporated
at 872.55 nm with respect to those presented in Barreto et al. (2020), for the purposes of this study due to the wider coverage
range of the EM27/SUN InGaAs detector. Thus, micro-windows B1-B5 are recorded by the first InGaAs detector and micro-
windows B6-B8 by the second one(hereafter referred to as InGaAs-1 and . Note that both detectors have different gains
(greater for InGaAs-2, respectively)i.e., B6-B8 micro-windows in SWIR), therefore the observed spectral behaviour is not the
one expected for the solar radiance: the higher the wavelength, the lower the measured radiation. It is also worth highlighting

that the solar spectrum displayed in Figure ~~??~~ 1 corresponds directly to the raw radiation signal as observed for both detectors, as in this study the EM27/SUN solar spectra were neither calibrated nor referenced to any traceable lamp.
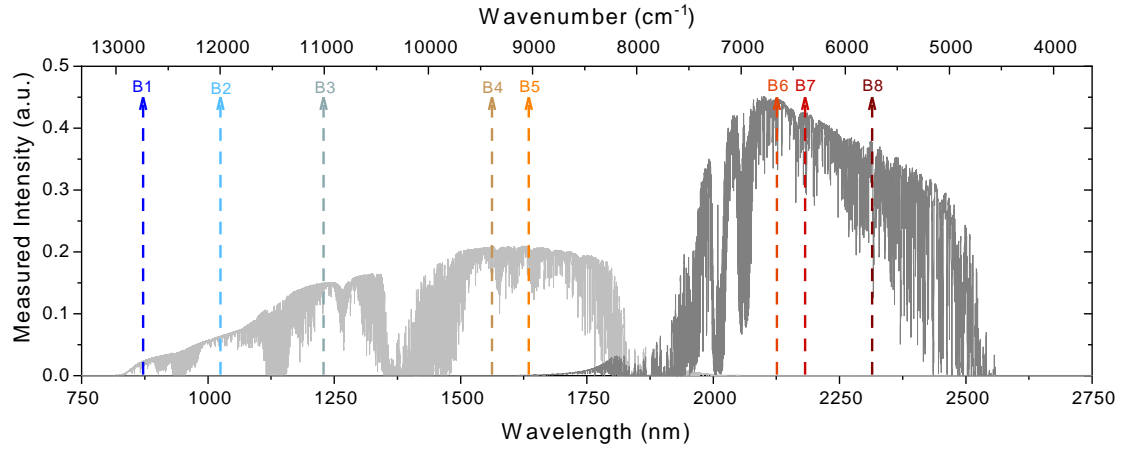
Figure 1. EM27/SUN solar spectrum for the 11500-4000 cm^{-1} spectral region, acquired at a spectral resolution of 0.5 cm^{-1} with the detectors InGaAs-1 and InGaAs-2, displayed as light grey and grey curves, respectively. Dashed-coloured vertical arrows show the position of the eight narrow spectral bands selected for the EM27/SUN AOD retrievals (B1-B8). The covered spectral range corresponds to 870-2500 nm, with a spectral resolution ranging between 0.04 and 0.27 nm for the B1 and B8 bands, respectively.

Table 1. Description of the eight spectral micro-windows used for the EM27/SUN AOD retrievals: spectral range and bandwidth (in nm). It also includes the mean (M) and the standard deviation (σ) of the intra-band coefficient of variation (CV, in %) time series considering all available EM27/SUN solar measurements between December 2019 and September 2021 (N = 56190), and those acquired for optical air masses (m) smaller than 7 (N = 55664). CV is estimated as the ratio between the standard deviation of the raw radiation signal and the mean value in each band. Micro-windows in bold character represent those bands coincident with CE318-AERONET.

Spectral range	Bandwidth	Intra-band CV [All]	Intra-band CV [m<7]
(nm)	(nm)	M, σ (%)	M, σ (%)
B1: 872.20 – 872.90	0.7	0.36, 0.02	0.36, 0.02
B2: 1020.55 – 1021.25	0.7	0.25, 0.01	0.25, 0.01
B3: 1237.75 – 1238.75	1	0.25, 0.15	0.24, 0.12
B4: 1557.75 – 1558.75	1	0.30, 0.07	0.30, 0.05
B5: 1635.50 – 1636.50	1	0.67, 0.19	0.66, 0.13
B6: 2132.90 – 2133.90	1	0.27, 0.04	0.27, 0.03
B7: 2191.50 – 2192.50	1	0.29, 0.05	0.29, 0.03
B8: 2313.80 – 2314.60	0.8	0.19, 0.08	0.19, 0.06

When selecting the optimal micro-windows for AOD retrievals, the temporal stability of the solar radiation and the EM27/SUN instrumental noise should be analyzed in detail for each proposed band. To achieve this, we estimated the intra-band coefficient of variation (CV), which is calculated as the standard deviation of radiation counts divided by the mean value in each band. This parameter can be a good indicator of instrumental issues, similar to the signal-to-noise ratio, or the effect of gas absorption interferences within each micro-window (García et al., 2016; Barreto et al., 2020). The EM27/SUN database used for this analysis was filtered according to the criteria defined in Section 4.3 about quality-control procedures, resulting in a total number of 187 days between December 2019 and September 2021 (56190 instantaneous AOD values for each spectral band).

As presented in Table 1, the mean and scatter of the intra-band CV values for the eight EM27/SUN micro-windows (B1-B8) and for the ~~3-year~~ 3 years period are overall consistent among the different bands and with those reported for the high-resolution IFS 125HR spectrometer at IZO (Barreto et al., 2020). Consistently to these authors, larger variability is expected to be observed in B3 (0.15%) and B5 (0.19%) bands. B3 seems to be somewhat sensitive to the water vapour content, whereas B5 contains CO_2 and CH_4 absorption lines, leading to an increase of the mean and scatter intra-band CV values provided they are spectrally resolved by the spectrometer. Nonetheless, unlike the IZO IFS 125HR, the EM27/SUN shows considerably lower CV for the B7 (0.05% compared to 0.27%), which could be attributed in part to the impact of the different spectral resolutions of the low- and the high-resolution FTIR systems (originally at 0.5 cm^{-1} in the EM27/SUN and truncated low-resolution 0.5 cm^{-1} in the case of the IFS 125HR). For the remaining bands, variations are expected to be lower than 0.10%. Another factor that could account for the observed variability is the range of optical air masses (m) spanned by the measured EM27/SUN solar spectra. This factor may be particularly important at higher solar zenith angles, where airmass changes occur faster than for lower air masses and measured IFGs are therefore more susceptible to intensity variations. As also listed in Table 1, intra-band CV retrieved for m below 7 (which is the normal operating range of the common photometers) shows a significant decrease compared to those obtained for all air masses, particularly for bands B3 and B5.

Based on these findings, the proposed EM27/SUN micro-windows could be highly effective in estimating AOD estimation under a wide range of atmospheric conditions.

4.2 Zero airmass Extrapolation and AOD Calculation

Solar transmittance methods have been used for many years to study aerosol columnar concentration from direct or reflected sunlight measurements. The Beer-Lambert-Bouguer attenuation law is applied to narrow spectral bands (Schmid et al., 1998, 2001), and the retrieval of $\tau_{a,\lambda}$ (or AOD_λ) from direct solar measurements is based on this attenuation law, which can be written as follows:

$$V_\lambda = V_{0,\lambda} \cdot d^{-2} \cdot exp^{-m \cdot \tau_\lambda} \tag{1}$$

This equation describes the exponential decay of direct sun measurements at a spectral band ~~centered~~ centred at the nominal wavelength λ (V_λ). The $V_{0,\lambda}$ term represents the instrument's signal measured at TOA ~~, where d is the Earth-Sun distance in~~

astronomical units and m is the optical air mass. In Equation 1, the terms τ_λ and m refer to different attenuators or absorbers in the atmosphere, such as aerosols (τ_a or AOD and m_a), Rayleigh scattering by molecules (τ_R and m_R), or absorption by gases such as O_3, NO_2, or H_2O . We calculated these (τ_{abs} and m_{abs}). Distance correction factor has not been included in this study considering the reduced FOV of the EM27, which is much smaller than the solar diameter. We calculated the previous variables following AERONET procedures (Holben et al., 1998; Giles et al., 2019) and the methodology proposed by Barreto et al. (2020). The Langley-Plot technique (Shaw, 1982, 1983) was used to determine $V_{0,\lambda}$, following the criteria defined by Toledano et al. (2018).

The final equation for AOD_λ or $\tau_{a,\lambda}$ determined at the central wavelength λ is retrieved by taking logarithms on both sides of Equation 1 as follows:

$$\tau_{a,\lambda} = \frac{1}{m_a} \cdot [ln(V_{0,\lambda}) - ln(V_\lambda) - \tau_{R,\lambda} \cdot m_R - \tau_{abs,\lambda} \cdot m_{abs}] \qquad (2)$$

4.3 Cloud Mask and Quality-Control Procedure

Three different filters have been implemented in the EM27/SUN data set to avoid cloud-contaminated data and instrumental issues, following the methodology proposed by Barreto et al. (2020). These filters are:

1. Cloud Mask: 1-minute measurements of co-located global and diffuse short-wave downward radiation were used to screen out cloud-contaminated data (García et al., 2014).

2. Quality-Control: this process is based on the analysis of the EM27/SUN IFGs and the Langley zero airmass extrapolation:

    (a) COCCON Quality Filters: As established by COCCON (e.g. Frey et al., 2021), several quality filters are applied on the raw EM27/SUN IFGs, such as requiring a minimum DC level of the maximum detector signal level (5%), restricting the tolerable DC variation of the measured level (10%), checking the centerburst location in the IFG and the centerburst amplitudes of forward and backward scans and the relative amplitude of out-of-band artifacts. The only difference introduced in this work compared to COCCON's standard quality filters is that we applied a more stringent filtering DC criterion, removing IFGs with intensity fluctuations above 1% and intensities below 1% of the maximal signal range.

    (b) Flag for non-valid Langley plots: this flag was manually implemented to avoid relevant changes in $V_{0,\lambda}$ resulting from variations in the instrument's calibration (e.g. mineral dust events or necessary cleanings of external EM27/SUN components).

As a result of implementing these flags, the original data set between December 2019 and September 2022, containing 59494 measurements, was filtered up to 9.4% resulting in 56190 quality-controlled observations.

5 Results

5.1 Top-of-Atmosphere Reference Irradiance Determination

250 The Langley-Plot method was used in this study to retrieve the $V_{0,\lambda}$ term and, therefore, to evaluate the radiometric stability of the EM27/SUN system and to calculate spectrally-resolved AOD. The developed Langley analysis considers that the derived calibration constant $V_{0,\lambda}$ is sensitive to any change in atmospheric conditions, which are assumed to be ~~constant~~ stable during the measurements.

In contrast to the morning calibration approach used for the IZO high-resolution IFS 125HR (Barreto et al., 2020), the Lan-
255 gley coefficients have been computed using both morning and evening data, with a higher frequency of Langley measurements taken at sunrise. Thus, the Langley days selected for this study, following the criteria published by Toledano et al. (2018), amounted to 31 high-quality Langley plots at the eight EM27/SUN spectral bands between December 2019 and September 2022, as displayed in Figure ~~??~~2(a). All these cases present similar and high Pearson correlation coefficients (R between 0.97 and 1.00) and standard deviations of the fitting analysis reaching Toledano's criterion (σ_{fit} below 0.006). It should be noted
260 that each EM27/SUN measurement ~~day~~ is calibrated using a daily $V_{0,\lambda}$, which is estimated from the smoothing spline functions determined ~~throughout the~~ with the information of the 31 Langleys performed over the entire period. This is a robust calibration procedure aimed at compensating for the continuous optical degradation of such an exposed optical system. An example of a morning Langley–Plot for the three ~~AERONET~~ CE318-AERONET coincident spectral bands (B1, B2 and B5), performed on 5th May 2022, is represented in Figure ~~??~~2(c)-(e).

265 Temporal degradation and continuous drift ~~, but also some improvements in~~ of the EM27/SUN calibration coefficients ~~,~~ are observed in Figures ~~??~~2(a)-(b). The latter depicts the relative variation in $V_{0,\lambda}$ with respect to the first valid Langley day for a better identification of the change-points presented in the calibration evolution. In both figures, important operational procedures and natural events that could affect the calibration are emphasized with ~~backward~~ black arrows. During the first year of the study, several logistical issues prevented the continuous performance of EM27/SUN Langley-Plot calibrations, resulting
270 in temporal gaps in the calibration time series. Specifically, due to the COVID-19 lockdown and subsequent restrictions, the EM27/SUN was not calibrated from March to September 2020. However, the external mirrors were manually cleaned on 1st August 2020, which improved the overall EM27/SUN signal by $\sim$~~6~~2% (black arrow (1) in Figures ~~??~~2(a)-(b)).
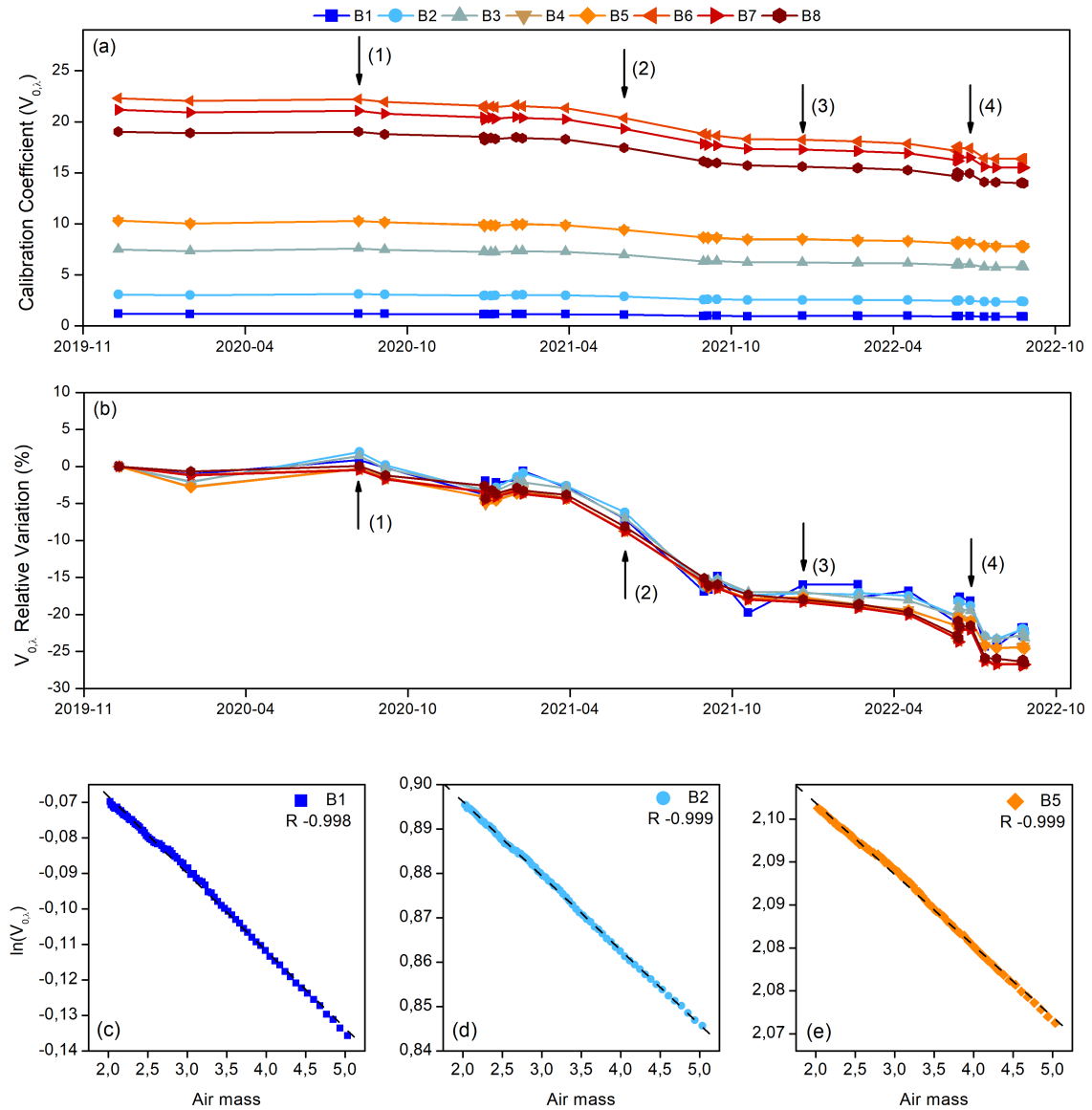
Figure 2. Top panel: (a) Langley-Plot calibration curve for the EM27/SUN coefficient $V_{0,\lambda}$ (in a.u.) at the eight EM27/SUN spectral bands from December 2019 to September 2022: solid lines are the smoothing splines of the $V_{0,\lambda}$ throughout the whole period. Black arrows mark important changes in the calibration coefficients: (1) 2nd September 2020, (2) 24th June 2021, (3) 8th January 2022, and (4) 13rd July 2022. (b) Relative variation in $V_{0,\lambda}$ with respect to the first valid Langley day (in %). Bottom panel: morning Langley-Plot for the three ~~AERONET~~ CE318-AERONET coincident spectral bands performed on 5th May 2022 for optical air masses between 2 and 5 are presented for (c) B1 (870 nm), (d) B2 (1020 nm), and (e) B5 (1640 nm). R is the Pearson correlation coefficient.

The most significant loss of EM27/SUN signal over the study period occurred from June 2021 to January 2022, when a $V_{0,\lambda}$ relative drop between 15% and ~~17~~18% was recorded for the eight bands (black arrows (2) and (3) in Figure ~~??~~2(a)-(b)). During these six months, several dust events were observed at IZO, and the volcanic activity from the La Palma eruption, which occurred 140 km away from IZO, led to volcanic ash deposition at the site (Milford et al., 2023). Other natural events, such as a forest fire that occurred at the end of July 2022 very close to the observatory, could also cause a loss of signal (event (4) in Figures ~~??~~2(a)-(b)). As a result, a significant drop in the Langley-Plot calibration curve was documented between two nearby calibrations performed on 13th July and 29th July 2022 (~~4% and 8% for~~ between 3% in B5 and 6% in B1~~bands, respectively,~~ with an all-band mean of ~~6~~4.2% and a σ of ~~1~~0.9%).

Considering the entire study period, the most significant changes in EM27/SUN's $V_{0,\lambda}$ coefficients were observed in the SWIR bands recorded by the InGaAs-2 detector. The overall decrease in $V_{0,\lambda}$, from December 2019 to September 2022, ranged from ~~19~~22% (B1-B2), ~~20-21~~23-24% (B3-B5) and ~~22~~26-27% (B6-B8), with a mean of ~~21~~24% across all bands (σ of ~~1.5~~1.8%). Consequently, the $V_{0,\lambda}$ degradation resulted in a drop between ~~0.56~~0.66%month^{-1} (B1-B2) and ~~0.66~~0.78%month^{-1} (B6-B8), with an all-band mean of ~~0.62~~0.72%month^{-1} (σ of ~~0.04~~0.27%month^{-1}). These values are relatively low compared to that of the high-resolution IFS 125HR system at the same station, which ranged between about 1.61%month^{-1} (B8) and 1.75%month^{-1} (B2), reaching a total decrease of 14.5% (B8) and 15.8% (B2) from May 2019 to February 2020 (Barreto et al., 2020).

5.2 AOD Retrievals

The EM27/SUN AOD database was created over the course of 187 days (comprising 56190 instantaneous AOD values for each spectral band) during the study period of December 2019 to September 2022. This database was constructed using the cloud screening and quality-controlled procedure, as described in Section 4.3. Figure ~~??~~3(a) shows the daily-mean spectral AOD from EM27/SUN for each defined band. It should be noted that there was a lack of calibration during the 6-month period from March to September 2020 (black arrow (1)), as mentioned in Section 5.1. This fact causes ~~that~~ the AOD values ~~were~~ to be calculated by taking into account smoothing spline functions ~~that were poorly conditioned. This could lead to overfitting and less confident results in this period resulting in, for example, negative AOD retrievals as observed in spring and summer of 2020 (black arrow (1) in Figure ??(a)).~~ described in Section 5.1.

Throughout the entire EM27/SUN measurement period, two significant dust events were recorded at IZO on 31st July 2020 and 20th June 2022 (black arrows (2) and (4), respectively, in Figure ~~??~~3(a)), with the maximum AOD recorded by the EM27/SUN exceeding 0.40 in the NIR B1 band. Another noteworthy dust event occurred in September 2020 with AOD values reaching ~0.30 in the lowest wavelength. Additionally, several minor dust events occurred during the study period. Together with the predominant presence of this aerosol, high AOD values were also observed during the volcanic eruption in La Palma from September to December 2021 (black arrow (3)), as well as during a significant forest fire that occurred in Tenerife in July 2022 (black arrow (5)). Note that the λ-AOD dependence is stronger under high-turbidity conditions, as observed by Barreto et al. (2020) using the high-resolution FTIR spectrometer in the same site.

All these events were simultaneously and consistently recorded by both the EM27/SUN and AERONET Cimel photometer, as presented in Figure ~~??~~3(b) for the coincident B2/1020 nm band. As an example, a zoomed-in view of a 14-day period with

continuous measurements is shown in Figure ??3(c), demonstrating the excellent temporal coherence between the measurements obtained by both techniques. This period was marked by the arrival of volcanic aerosols and gases from the La Palma eruption (Sicard et al., 2022; Bedoya-Velásquez et al., 2022; Córdoba-Jabonero et al., 2023; García et al., 2023; Milford et al., 2023) during the first part of the period, a Saharan dust outbreak from 30th September to 5th October (AOD_{1020nm} up to 0.15) (García et al., 2023; Milford et al., 2023), and background conditions (AOD_{1020nm} less than 0.05) for the rest of the time.
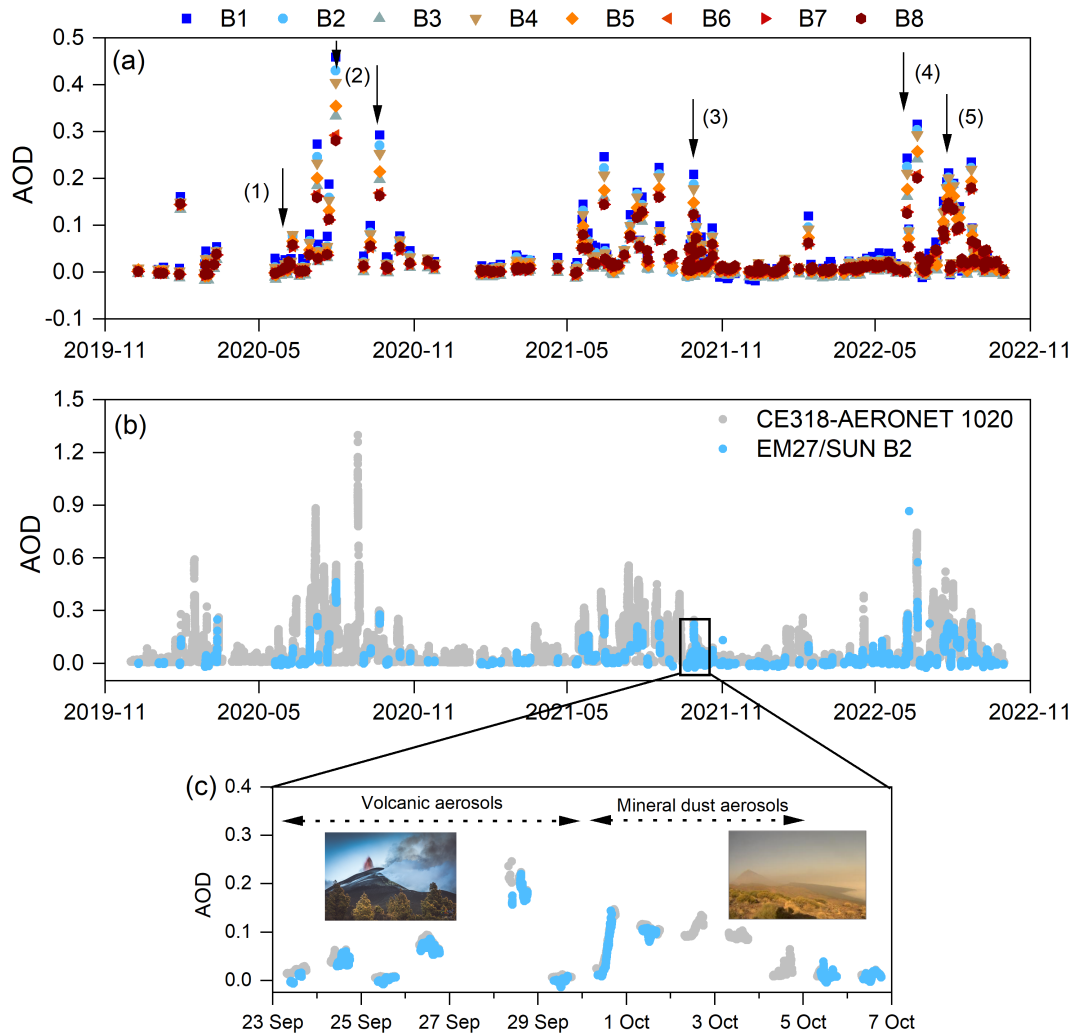
Figure 3. Time series, from December 2019 to September 2022, of (a) the daily-mean AOD data for the eight EM27/SUN spectral bands and (b) the ~~AERONET~~ CE318-AERONET and EM27/SUN AOD values measured at the coincident B2 and 1020 nm band. (c) AOD evolution in a 14-day period in September 2021. The images on (c) correspond to the volcano on La Palma (LuzLux/AEMET) and a Saharan dust event at IZO. Black arrows in (a) mark important AOD events: (1) spring/summer 2020, (2) mineral dust outbreaks in July and September 2020, (3) La Palma volcanic eruption between September and December 2021, (4) mineral dust outbreak in June 2022, and (5) forest fire in July 2022.

5.3 AOD uncertainty analysis

The uncertainty analysis of the EM27/SUN AOD retrievals has been addressed in this section through the utilization of uncertainty propagation techniques, specifically employing the guidelines outlined in the "Guide to the Expression of Uncertainty

315 in Measurement" (GUM) and the application of the Monte-Carlo method (MCM), as detailed in the work by GUM (2008). The input parameters along with their corresponding relative uncertainties are presented in Table 2. Following the procedure outlined in the publication by García et al. (2021), the number of iterations (M) has been set at 10^6 in order to achieve a coverage probability of 95%. To conduct the Monte-Carlo calculations, the Python module MetroloPy (available at https://pypi.org/project/r accessed on July 16, 2023) was employed.

320 The standard uncertainty estimated with the MCM has been computed considering that our model is given by the equations 3, 4 and 5, retrieved by applying Equation 2 at the three EM27/SUN spectral bands coincident to the CE318-AERONET and using the AERONET standard corrections (Holben et al., 1998; Giles et al., 2019) as follows:

$$\tau_{B1} = \frac{1}{m_a} \cdot [ln(V_{0,B1)}) - ln(V_{B1}) - \tau_R \cdot m_R] \tag{3}$$

$$\tau_{B2} = \frac{1}{m_a} \cdot [ln(V_{0,B2)}) - ln(V_{B2}) - \tau_R \cdot m_R - m_a \cdot (a_1 \cdot PWV + a_2)] \tag{4}$$

325 $$\tau_{B5} = \frac{1}{m_a} \cdot [ln(V_{0,B5)}) - ln(V_{B5}) - \tau_R \cdot m_R - m_a \cdot (a_1 \cdot PWV + a_2) - m_a \cdot M_1 - m_a \cdot M_2] \tag{5}$$

In these equations, m_a represents the air mass of atmospheric aerosols (Kasten and Young, 1989), m_R is the Rayleigh air mass (computed using the same equation as for aerosols), τ_R is the Rayleigh optical depth (Bodhaine et al., 1999) and PWV is the precipitable water vapour. Following Giles et al. (2019), the impact of carbon dioxide and methane in the 1640 nm spectral band are estimated with the M_1 and M_2 terms, including their estimated concentrations adjusted to the station elevation.

330 Some uncertainties presented in Table 2 have been sourced from relevant bibliography, while others have been estimated through empirical processes. The assessment of EM27/SUN instrumental uncertainties (V_λ) incorporates the mean voltage ratios between co-located EM27/SUN and CE318-AERONET measurements. This ratio offers insight into potential instrumental issues, such as possible tracking problems or temperature dependence. The average voltage ratios were observed to range between 1.7% for 870 nm and 0.9% for 1640 nm (detailed in Table 2). These values are consistent with the fact that the

335 spectral band centred at 872 nm is the closest band to the detector cut-off, which consequently introduces more noise to the EM27/SUN measurements.

The uncertainty linked to the determination of the V_0 term has been conservatively approximated at 1.06%, derived from the maximum standard error of the intercept within the Langley-Plot linear regression conducted across the entire duration and the three distinct spectral bands encompassing the 31 Langley plots.

Table 2. Relative uncertainties of the input parameters in the AOD retrieval and their corresponding references.

Parameter	Uncertainty	Reference
V_{B1}	1.7%	empirical
V_{B2}	1.2%	empirical
V_{B5}	0.9%	empirical
V_0	1.06%	empirical
τ_R	0.7%	Fröhlich and Shaw (1980)
m_a	0.065%	Kasten and Young (1989)
PWV	10%	Alexandrov et al. (2009); Giles et al. (2019)
M_1, M_2	4.5%	Smirnov et al. (2004)
a_1	2% (1020nm) - 5% (1640nm)	Smirnov et al. (2004)
a_2	5% (1020nm) - 2% (1640nm)	Smirnov et al. (2004)
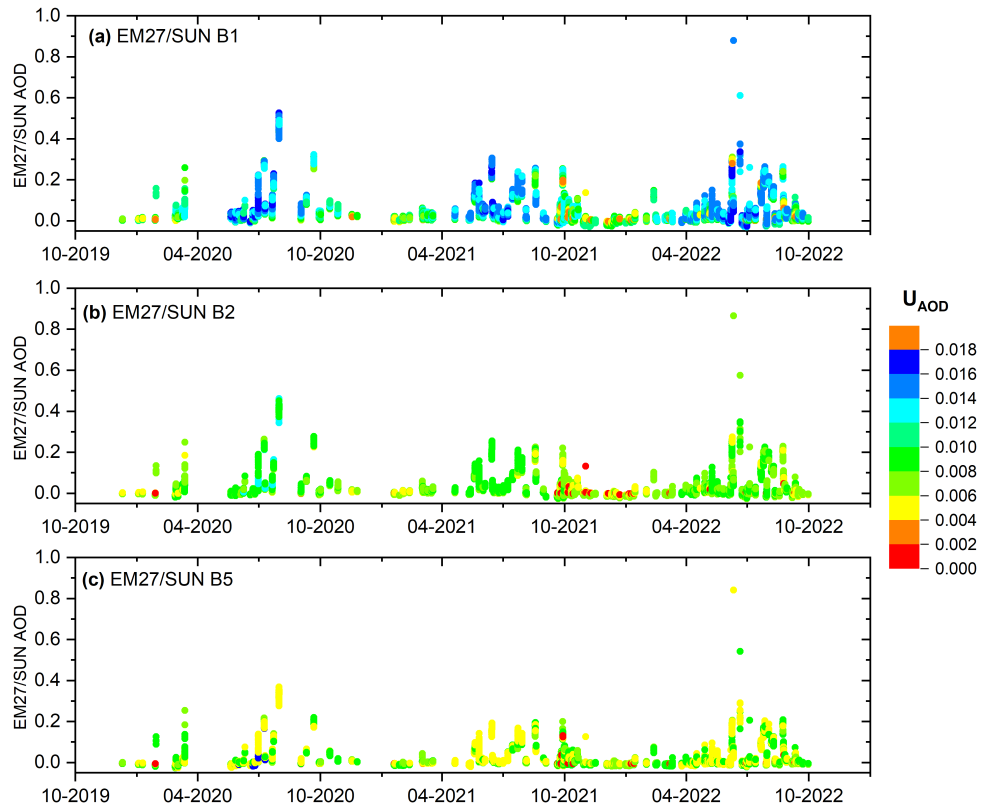
Figure 4. Time series, from December 2019 to September 2022, of EM27/SUN AOD (a) B1, (b) B2 and (c) B5. The colour scale indicates the absolute uncertainty U_{AOD} calculated by means of the Monte-Carlo method.

340   Figure 4 illustrates the temporal evolution of AOD retrieved using the EM27/SUN instrument spanning the period from December 2019 to September 2022. The graph incorporates the AOD absolute uncertainty (U_{AOD}), which has been computed using the Monte-Carlo method (MCM). The U_{AOD} values exhibit a dependence on both wavelength and AOD magnitude, with elevated uncertainties observed within the 870 nm spectral band for higher AOD values. The calculated U_{AOD} from the 870 nm spectral band varies from 0.013 (for AOD values < 0.05) to 0.281 (for AOD values > 0.20). In the case of the 1020 nm

345   band, U_{AOD} range spans from 0.011 (AOD < 0.05) to 0.275 (AOD > 0.20). For the 1640 nm spectral band, the uncertainties range from < 0.001 (AOD < 0.05) to 0.301 (AOD > 0.20), with this last case encompassing only 1% of the dataset. Average uncertainties are 0.012, 0.013, and 0.009 for 870, 1020, and 1640 nm spectral bands, respectively.

5.4 AOD Quality Assessment

The consistency of EM27/SUN retrievals was evaluated over the 3-year period by comparing synchronous AOD measurements

350   (within 1 minute) of the three coincident spectral bands from EM27/SUN and ~~AERONET~~ CE318-AERONET instruments (870, 1020, and 1640 nm). The scatterplot in Figure~~??~~ 5 and the main statistics presented in Table 3 show the results of these synchronous AOD data sets, considering separately the two time periods with different Langley-Plot calibration frequencies: (1) from December 2019 to December 2020 ~~, where the lack of correct calibration frequency could cause more uncertain results~~ ~~(open circles in Figure??)~~(Figure 5(a)), characterized by a significant calibration gap resulting from the COVID-19 lockdown

355   restrictions; and (2) after January 2021, ~~when~~ marked by a notable increase in the calibration frequency ~~was adapted to meet the instrumental needs~~(Figure 5(b)).
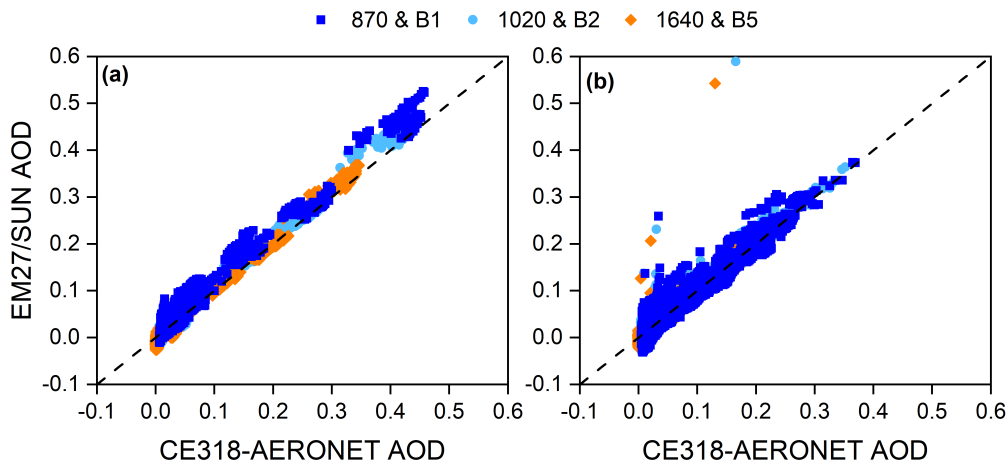
Figure 5. Scatterplot for the coincident EM27/~~SUN-AERONET~~ SUN versus CE318-AERONET AOD values from (a) December 2019 to December 2020 and (b) January 2021 to September 2022 considering the EM27/SUN B1 (870 nm), B2 (1020 nm) and B5 (1640 nm) micro-windows. ~~Open circles correspond to data in the December 2019 - December 2020 time period.~~ The number of coincidences is 14575 and 2863 in the period ~~December 2019~~ January 2021 - September 2022 and December 2019 - December 2020, respectively.

Our ~~results~~ findings confirm the consistency between these two independent techniques ~~, especially from January 2021 onward. The main statistics revealed excellent comparability in~~ during two distinct periods marked by varying calibration frequencies. The principal statistical outcomes unveil a notable consistency across the three spectral bands, with mean ~~differences~~ discrepancies limited to 0.005~~, which are similar to the values reported~~. These differences align with values akin to those documented in Barreto et al. (2020). While B5 ~~has the lowest~~ showcases the most favourable results, evidenced by its minimal σ (0.011) and RMSE values (root-mean-squared error of ~~the differences)~~ differences, 0.012), as well as ~~the~~ its highest R values ~~among~~ across the spectral bands (0.967), B1 ~~exhibits the worst performance~~with demonstrates comparatively inferior performance. B1 is characterized by an R ~~of 0.947~~ value of 0.938 and a σ ~~of 0.019. This~~ value of 0.021. This performance ~~discrepancy~~ is attributed to the fact that B1 is the closest band to the detector cut-off, which consequently introduces more noise to the EM27/SUN solar spectra in this particular spectral range. ~~These results can serve as an estimation of the maximum uncertainty (air mass equals to 1) of the AOD EM27/SUN product, ranging between 0.038 (for B1) and 0.014 (for B3) with a 95% level of statistical significance. Considering a mean air mass of 1.60 calculated for the 2021-2022 period, these uncertainties decrease to 0.023 (for B1) and 0.009 (for B3). Our results also emphasize the importance of maintaining a regular calibration frequency, ideally at least once per month, to compensate for the optical degradation of the external system, as described in Section 5.1~~

The observed bias within this intercomparison analysis, described by means of the mean AOD differences, falls within the bounds of uncertainty determined via the Monte-Carlo method outlined in Section 5.3. These results confirm the uncertainty estimation with respect to the input values performed in our study.

When comparing data from two instruments with different spectral resolutions, like the EM27/SUN and the Cimel CE318-T, a common intercomparison procedure is to degrade the higher-resolution spectra considering the response functions of the lower-resolution instrument. However, as found by ~~Barreto (2020)~~Barreto et al. (2020), both techniques provide a straightforward AOD intercomparison, and the convolution approach yielded very consistent and equivalent results.

Table 3. Main statistics of the ~~AERONET~~ CE318-AERONET and EM27/SUN AOD comparison for the coincident spectral bands in the January 2021 - September 2022 period, including the results for the entire period (2019-2022) in parenthesis. MD stands for the mean AOD difference (~~AERONET-EM27~~CE318-AERONET minus EM27/SUN), σ is the standard deviation and RMSE is the root-mean-squared error of the differences. The fit parameters are also included: Pearson correlation coefficient (R), slope, and intercept. The number of coincidences is 14575 and 2863 in the period December 2019 - September 2022 and January 2021 - September 2022, respectively.

Spectral Band	B1/870 nm	B2/1020 nm	B5/1640 nm
MD	~~-0.002 (<-0.001~~-0.003 (0.014)	~~0.001 (0.003<~~-0.001 (-0.004)	0.005 (~~0.011~~0.005)
σ	~~0.019 (0.020~~0.021 (0.018)	~~0.012 (0.017~~0.015 (0.013)	~~0.007 (0.018~~0.011 (0.010)
RMSE	~~0.019 (0.020~~0.021 (0.023)	~~0.011 (0.017~~0.015 (0.013)	~~0.008 (0.021~~0.012 (0.010)
R	~~0.947 (0.953~~0.938 (0.985)	~~0.974 (0.961~~0.962 (0.992)	~~0.987 (0.934~~0.967 (0.994)
Slope	~~0.982 (0.974~~0.997 (1.078)	~~0.975 (0.951~~0.989 (1.063)	~~0.990 (0.940~~1.008 (1.064)
Intercept	~~0.002 (0.001~~0.003 (0.010)	0.001 (~~-0.001~~0.001)	-0.005 (~~-0.010~~-0.007)

~~Figure ??~~ Figure 6 investigates the relationship between the AOD differences (~~AERONET-EM27~~CE318-AERONET minus EM27/SUN) and important factors such as aerosol loading (i.e. ~~AERONET~~ CE318-AERONET AOD at 1640 nm, AOD_{1640nm}, as proxy), precipitable water vapour content (PWV) from AERONET, and optical path (i.e. optical air mass), as well as their temporal variations. The comparison with the air mass provides valuable information about the traceability of the two AOD data sets. According to the WMO traceability criteria (WMO, 2005), AOD measurements are considered traceable if 95% or more of the AOD differences fall within a specified acceptance limit, ~~$U95$~~U_{95}, defined in terms of the aerosol optical air mass as:

$$U_{95} = \pm(0.005 + \frac{0.010}{m_a} \frac{0.010}{m})$$
(6)

In this equation, the first term represents instrumental and algorithmic uncertainties, while the second term is related to the calibration uncertainty, estimated to be less than 1% at maximum (for air mass equals ~~to~~ 1). The optical air mass for U_{95} has been calculated by using the equation provided by Kasten and Young (1989).

The temporal evolution of the AOD differences is indicated in Figure~~??~~ 6(a) ~~, where AOD differences in~~ for the three coincident bands of ~~AERONET~~ CE318-AERONET and EM27/SUN (870, 1020, and 1640 nm)~~are higher during the period affected~~. We noted that these differences remained consistent even throughout the period impacted by the absence of Langley-Plot calibrations (~~spanning~~ a 6-month period from March to September 2020) ~~due to~~ as a result of COVID-19 lockdown and subsequent restrictions. ~~The number of Langley-Plot calibrations increased during 2021 and especially during 2022, resulting in a considerably better agreement between the two instruments. Figure??~~Figure 6(b) shows the behaviour of AOD differences with the air mass for the entire period of this study, with the September 2019 - December 2020 period marked as open circles. This analysis serves to calculate traceability limits (~~$U95$) of 48%, 64~~U_{95}) of 51.4%, 70.4%, and ~~72~~82.2% for the 870, 1020, and 1640 nm coincident spectral bands, respectively. ~~An~~ A slight improvement in traceability has been observed since January 2021, when it increases to values up to ~~50%, 71~~52.0%, 71.1%, and ~~84~~83.5% for the respective bands. It should be noted that even for the 2021-2022 period, the proportion of data meeting the WMO traceability requirements is low. However, it is important to consider that the EM27/SUN instrument was not specifically designed to offer the absolute photometric stability necessary for aerosol monitoring~~, with a maximum estimated uncertainty in AOD up to 0.038 (air mass equals 1) and average uncertainties (for an average air mass of 1.6)of 0.023. Additionally,~~. The low traceability identified in our study stands in contrast to the remarkable traceability established between CE318-AERONET and the IFS 125HR, as reported by Barreto et al. (2020). This disparity, evident despite employing identical methodology and spectral resolution, might indicate the existence of mechanisms introducing a variable spectral ordinate calibration in the ~~uncertainty term (0.005) mentioned in Equation (??) was estimated for the UV-VIS range and is assumed to be wavelength-independent~~case of the EM27/SUN.

~~AOD comparison between AERONET and EM27/SUN (2021-2022) in the three coincident spectral bands for AERONET (870, 1020, and 1640 nm) and EM27/SUN (B1, B2, and B5) according to (a) time, (b) optical air mass, (c) PWV (cm) and (d) AOD at 1640 nm from AERONET (AOD_{1640nm}). Open circles in (a) and (b) correspond to the September 2019 - December 2020 period. The solid curves in (b) represent the $U95$ uncertainty limit. The AOD differences in (c) and (d) are displayed as~~

box plots, where the lower and upper boundaries for each box are the 25th and 75th percentiles, the solid line is the median value, the hyphens are the maximum and minimum values and the asterisks indicate the 1th and 99th percentiles. The number of cases in each box was 2778, 3450, 4021, 1193 and 271 for PWV <0.2, 0.2-0.4, 0.4-0.8, 0.8-1.0 and >0.1 cm, respectively. In the case of AOD differences with respect to AERONET AOD, the number of cases was 9850, 929, 541, 364 and 29 for AOD_{1640nm} <0.05, 0.05-0.1, 0.1-0.15, 0.15-0.2 and >0.2, respectively.

Figure??Figure 6(c) and (d) indicate that there is no correlation between AOD differences and reveal an absence of correlation between differences in AOD and both PWV and AOD values at 1640 nm measured by AERONET in the period , as measured by CE318-AERONET, during the period spanning from January 2021 to September 2022, in agreement with the results found 2022. This observation holds true for the 1020 and 1640 nm spectral bands, aligning with the findings documented in Barreto et al. (2020). Therefore, the In contrast, a dependence on these variables is observed for the 870 nm spectral band. This correlation is attributed to the anticipated higher uncertainty in the AOD estimation within this specific spectral band, particularly under conditions of elevated AOD. Such conditions are frequently encountered at the Izaña site due to the presence of the Saharan Air Layer, which manifests as a distinct, well-stratified and humid layer when contrasted with the standard free-troposphere conditions at the location (Barreto et al., 2022a).

Consequently, the EM27/SUN AOD retrievals presented here can be well-suited to be applicable under a wide range of atmospheric conditions(i.e. humid environments , different kinds of aerosols, etc.). in this study exhibit strong potential for broad applicability across diverse atmospheric conditions, encompassing scenarios such as humid environments and varying aerosol compositions.
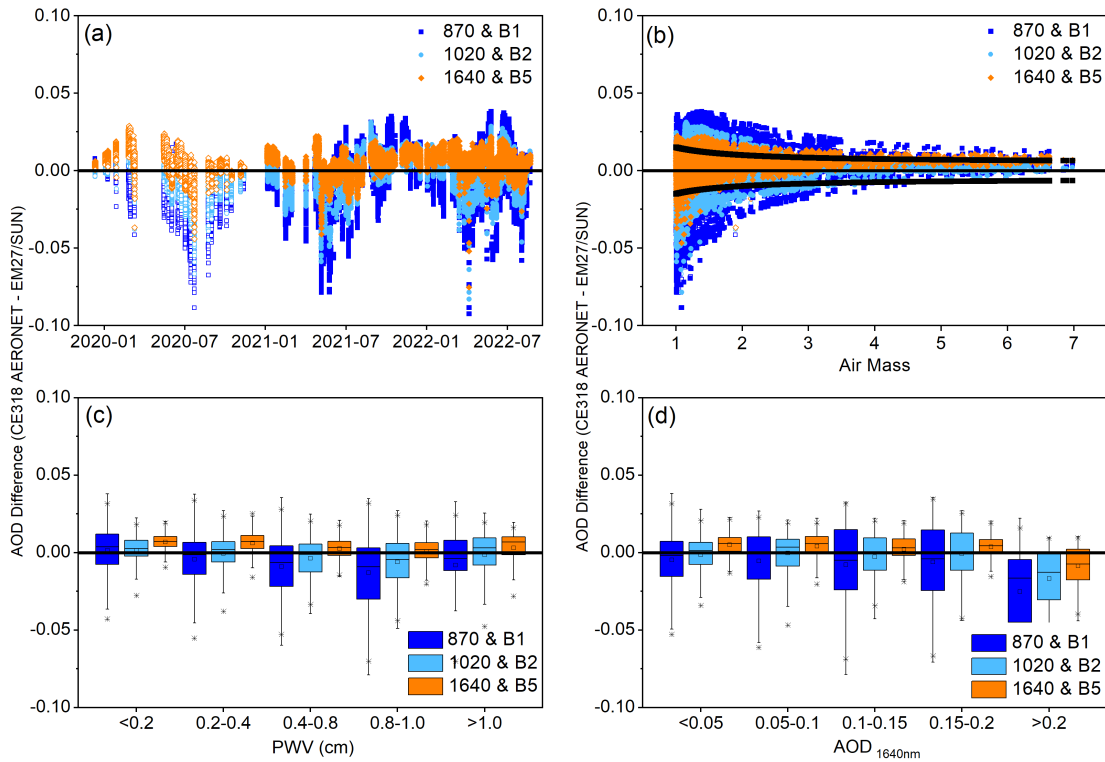
Figure 6. AOD comparison between CE318-AERONET and EM27/SUN (2021-2022) in the three coincident spectral bands for CE318-AERONET (870, 1020, and 1640 nm) and EM27/SUN (B1, B2, and B5) according to (a) time, (b) optical air mass, (c) PWV (cm) and (d) AOD at 1640 nm from CE318-AERONET (AOD_{1640nm}). Open circles in (a) and (b) correspond to the September 2019 - December 2020 period. The solid curves in (b) represent the U_{95} uncertainty limit. The AOD differences in (c) and (d) are displayed as box plots, where the lower and upper boundaries for each box are the 25th and 75th percentiles, the solid line is the median value, the hyphens are the maximum and minimum values and the asterisks indicate the 1th and 99th percentiles. The number of cases in each box was 3191, 3729, 3690, 773 and 238 for PWV <0.2, 0.2-0.4, 0.4-0.8, 0.8-1.0 and >0.1 cm, respectively. In the case of AOD differences with respect to CE318-AERONET AOD, the number of cases was 9757, 929, 542, 364 and 29 for AOD_{1640nm} <0.05, 0.05-0.1, 0.1-0.15, 0.15-0.2 and >0.2, respectively.

430 5.5 Spectral Consistency

~~Section5.3~~ Section 5.4 compared and evaluated EM27/SUN AOD retrievals using AERONET data for the three coincident spectral bands (870, 1020, and 1640 nm). To indirectly assess the spectral consistency of the EM27/SUN AOD database in the entire 873-2314 nm spectral range (B1-B8), we verified the λ-AOD dependence by performing a correlation matrix analysis. This information served as a cross-validation of the EM27/SUN AOD observations across different spectral bands. 435 The assessment has been limited to the 2021-2022 period of reliable EM27/SUN AOD records due to the high-frequency ~~calibrationof~~ Langley-Plot calibrations.

As shown in previous works (e.g. Toledano et al., 2019; Barreto et al., 2020, and references therein), the spectral dependence of AOD does not follow the Ångström power law (Junge, 1955), especially in the shortwave infrared for mineral dust aerosols. As the linear λ-AOD behaviour (in logarithmic space) is not the expected one, the Kendall rank correlation coefficient (R_K) has been used here as an indicator of the degree of correlation between all the EM27/SUN AOD observations. This non-parametric statistical test checks whether two variables are statistically dependent without making any assumptions about the distributions of the two variables being compared (Kendall, 1938).

~~Kendall rank correlation coefficient between EM27/SUN AOD retrievals for the eight bands in the 2021-2022 period for different aerosol regimes: (a) clean conditions (AOD$_{1640nm}$ < 0.074), (b) dusty conditions (AOD$_{1640nm}$ > 0.074) and (c) all aerosol regimes.~~

~~Figure ??~~ Figure 7 shows the Kendall rank correlation coefficient between the AOD retrieved in the eight EM27/SUN spectral bands for three different aerosol regimes at IZO, according to Barreto et al. (2022a, b). Following these authors, background conditions at IZO can be characterized by $AOD_{1640nm} < 0.074$. We subsequently calculated matrix correlation coefficients for clean conditions ($AOD_{1640nm} < 0.074$) in Figure~~??~~ 7(a), for dusty conditions ($AOD_{1640nm} > 0.074$) in Figure~~??~~ 7(b), and for all conditions in Figure~~??~~ 7(c). bands, particularly under high-AOD regimes (dusty conditions). Three correlation blocks are clearly distinguished, where we found spectral consistency between nearby spectral bands, particularly under high-AOD regimes (dusty conditions). These results are in agreement with the results published by Barreto et al. (2020). B1 and B2 exhibited R_K values up to 0.85, B3-B5 had the highest correlations in the NIR, with R_K values > 0.90, and B6-B8 (SWIR) had the highest R_K values of the entire spectral range (> 0.90). It is particularly relevant that a weaker correlation was observed between B3 and B4, which are only separated by 320 nm, while the correlation between B2 and B5, with a λ separation of 398 nm, is higher. We attribute these results to the presence of gas absorption bands, such as H_2O in the case of B3, which were not considered in this analysis. The weakest relationship between the different AOD bands is documented under clean conditions. In these cases, the retrieved EM27/SUN AOD values are susceptible to including ~~artefacts~~ artifacts due to the rather low AOD values, especially close to the edge-cutoff limits of the EM27/SUN detectors. Therefore, our results demonstrate the agreement between EM27/SUN AOD observations in the 873-1021, 1238-1636, and 2133-2314 nm wavelength ranges, which we attribute to the spectral AOD coherence in each range.

5.6 Combination of Aerosols and Gases Retrievals

As demonstrated in this study, the EM27/SUN's multi-parameter capability enables simultaneous retrievals of spectrally-resolved aerosol and trace gas information, opening up new opportunities to enhance atmospheric chemistry monitoring (Paton-Walsh et al., 2005, 2010; Frausto-Vicencio et al., 2023). Providing such combined measurements is crucial for monitoring and understanding in-plume aerosol-gas interactions or assessing the impact of aerosols on remote gas observations. The latter is especially important for validating satellite missions devoted to GHG observations, which are strongly affected by artificial distortions due to the presence of aerosols, particularly in complex emission scenes such as urban/industrial, biomass burning, or mining emissions (e.g. Jung et al., 2016; Ke et al., 2022). Here, two case studies have been selected to ~~analyse~~ analyze
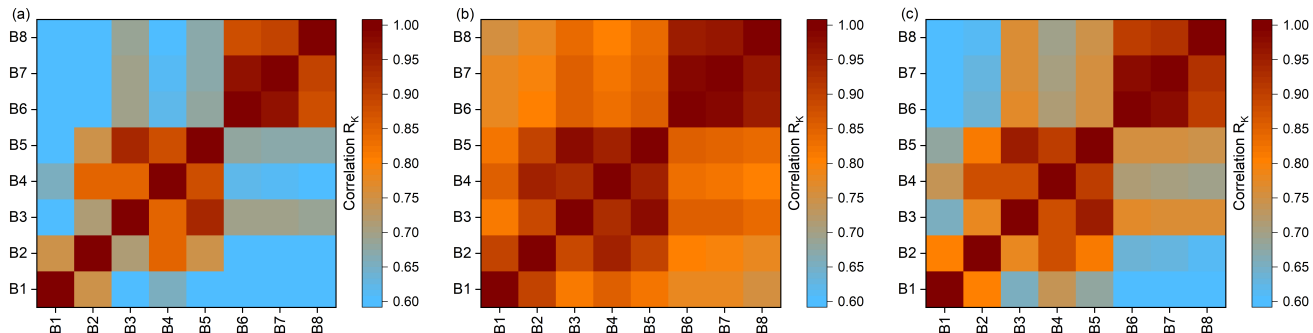
Figure 7. Kendall rank correlation coefficient between EM27/SUN AOD retrievals for the eight bands in the 2021-2022 period for different aerosol regimes: (a) clean conditions ($AOD_{1640nm} < 0.074$), (b) dusty conditions ($AOD_{1640nm} > 0.074$) and (c) all aerosol regimes.

470   the relationship between co-emitted tracers (aerosols, CO_2 and CO molecules), when these atmospheric compounds share a similar emission source.

    The first case study corresponds to a volcanic event that impacted IZO on 11st October 2021 (Figure ??8(a)), with a peak surface ~~sulphur~~ sulfur dioxide (SO_2) of 377.1 μgm^{-3} compared to the very low background level of 0.2 μgm^{-3} at the site (Cuevas et al., 2022; Milford et al., 2023). SO_2 total column amounts up to 11 DU (Dobson Units) were also recorded by the

475 high-resolution IFS 125HR spectrometer at IZO, when typical background values are lower than 1 DU (García et al., 2022). The volcanic influence of this event from the La Palma eruption was also confirmed using Sentinel-5P TROPOMI SO_2 product (ESA, 2021a) (Figure ??8(b)) and backtrajectory analysis (not shown here).

    The second case study analyses biomass burning emissions from a forest fire that started on 11st July 2022 in the municipality of Los Realejos, at ~10 km distance from IZO, on the northern side of Tenerife (Figure ??8(c)). This fire affected an area

480 of more than 2700 hectares and was considered stabilized on 25th July 2022. We included EM27/SUN data affected by the biomass burning emissions in the afternoon of 21st July 2022. TROPOMI XCO product corresponding to this event is presented in Figure ??8(d) (ESA, 2021b).

    The EM27/SUN gas observations have been retrieved at the same time as AOD values. For this purpose, the measured EM27/SUN absorption spectra were ~~analysed~~ analyzed using the non-linear least-squares fitting algorithm PROFFAST, ob-

485 taining the total columns of different trace gases (O_2, CO_2, CH_4, CO, and H_2O) (Sha et al., 2020; Frey et al., 2021, and reference therein). Then, the total column-averaged dry-air mole fractions of these gases (XGas) were computed by using the co-observed O_2 total columns. The XGas mole fractions were calibrated onto the WMO's gas standard scale maintained by NOAA (National Oceanic and Atmospheric Administration,https://www.noaa.gov/, last access: 3rd May 2023) and provided as standard COCCON products (Hase et al., 2016; Frey et al., 2019; Sha et al., 2020; Frey et al., 2021).

490 For both case studies, the tracer-tracer anomaly correlation technique has been used to ~~analyse~~ analyze the relationship between aerosol and trace gases enhancements (ΔAOD and $\Delta XGas$, respectively) (e.g. Wong et al., 2016; García et al., 2021). To quantify the enhancements, we followed the approach of Frausto-Vicencio et al. (2023) and subtracted background from instan-
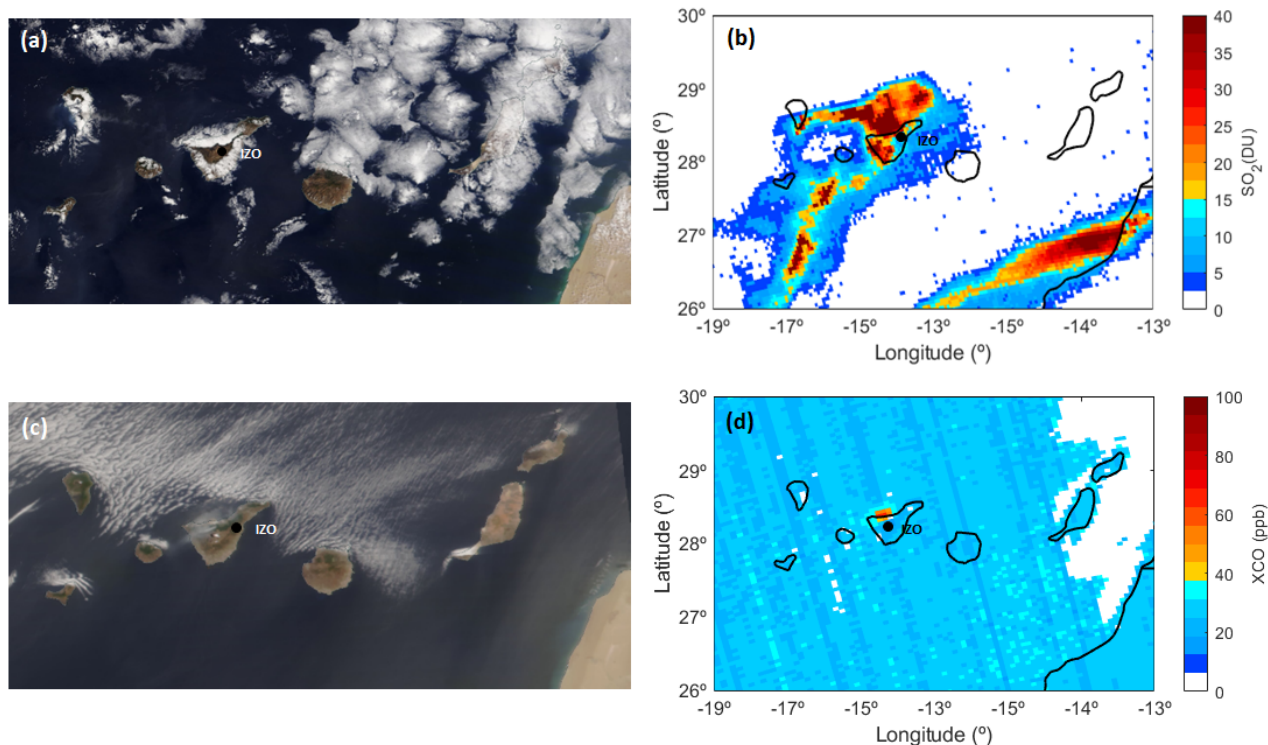
Figure 8. Images captured by the Moderate Resolution Imaging Spectroradiometer (MODIS) sensor onboard the NASA's Terra satellite (https://worldview.earthdata.nasa.gov) on (a) 11st October 2021 (volcanic plume) and (c) 21st July 2022 (forest fire plume). Satellite column-integrated (b) ~~sulphur~~ sulfur dioxide (SO_2, in DU) on 11st October 2021 and (d) monoxide carbon mole fraction (XCO, in ppb) between 20th and 22nd July 2022 from Copernicus Sentinel-5P TROPOMI (https://maps.s5p-pal.com/, last access: 28th April 2023). The location of the Izaña Observatory (IZO) is indicated with a black dot.

taneous values. Background gas values were calculated using the 2nd percentile of daily measurements, whereas for ΔAOD the climatology background value at IZO of 0.001 at 1020 nm was taken as reference (Toledano et al., 2018). In Figure 7,

495 which summarizes the two case studies, each panel depicts the instantaneous EM27/SUN and ~~AERONET~~ CE318-AERONET AOD values at B2/1020 nm, EM27/SUN XCO_2 and XCO mole fractions, and correlation plots for ΔXCO - ΔXCO_2, ΔXCO - ΔAOD, and ΔXCO_2 - ΔAOD. The time periods selected for the tracer-tracer anomaly assessment were identified based on observed increases, especially, in AOD and XCO records (shaded areas in Figure ~~7~~9(a), (b), (f), and (g)). It should be noted that although only the analysis for AOD at B2/1020 nm is discussed as a reference, similar conclusions were drawn from the

500 other AOD bands (data not shown).

For the volcanic event, significant ΔXCO_2 and ΔXCO of up to 1 ppm and 3 ppb were recorded, which represent mean increases of $\sim$0.15% and 2% with respect to the IZO background values, respectively. These ΔXCO_2 values at IZO, which is located 140 km from the La Palma volcano crater, are comparable to those observed, for example, for the Kilauea eruption
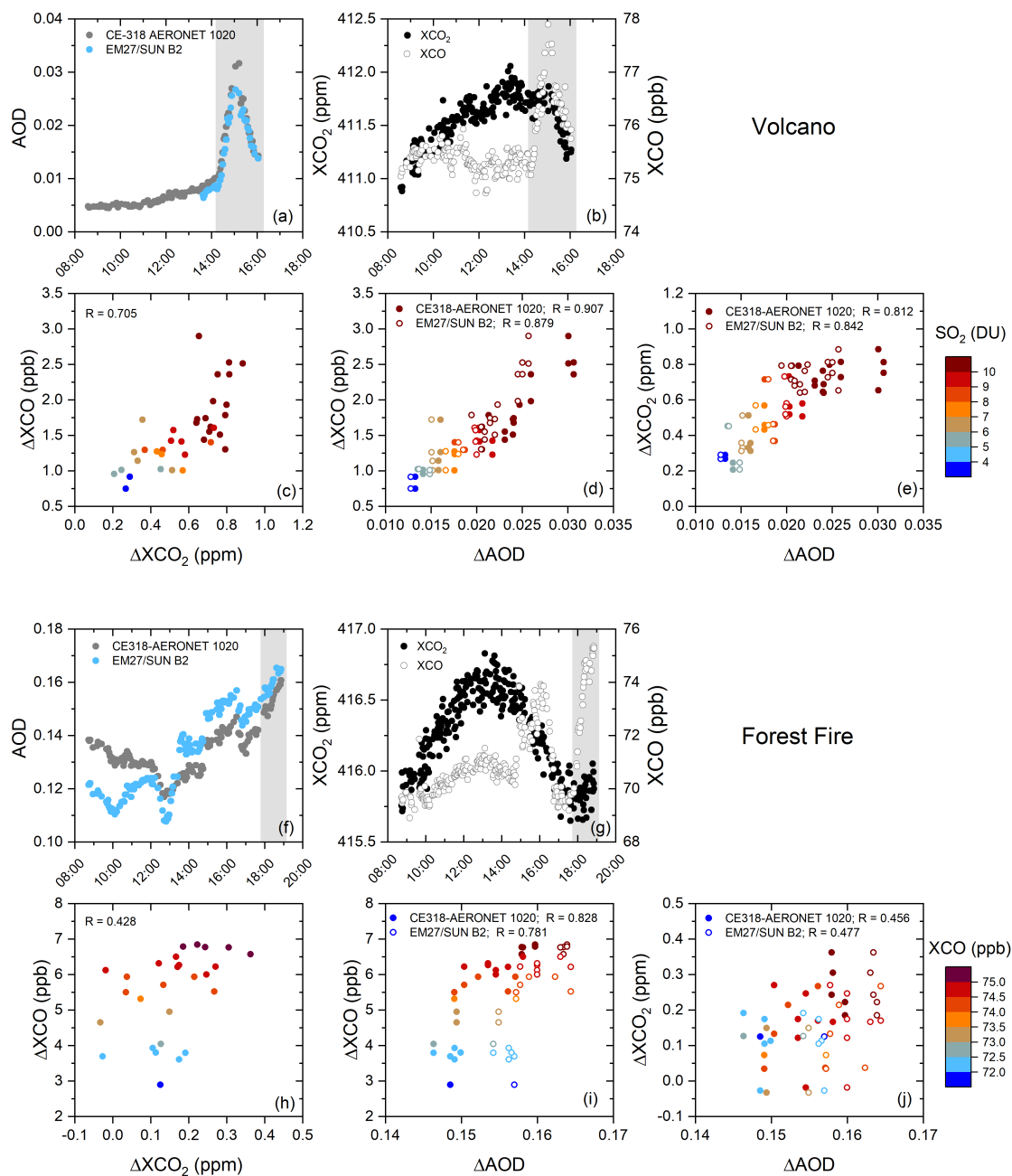
Figure 9. Top panel: volcanic event on 11st October 2021 (a-e). Bottom panel: forest fire on 21st July 2022 (f-j). Each panel displays the instantaneous EM27/SUN and ~~AERONET~~ CE318-AERONET AOD values for B2 and 1020 nm band (a and f), EM27/SUN mole fractions of XCO_2 (ppm) and XCO (ppb) (b and g), scatterplots of ΔXCO (ppb) versus ΔXCO_2 (ppm) (c and h), ΔXCO (ppb) versus ΔAOD at B2/1020 nm (d and i), and ΔXCO_2 (ppm) versus ΔAOD at B2/1020 nm (e and j). The time periods selected for the tracer-tracer anomaly assessment displayed in (c)-(e) and (h)-(j) correspond to the shaded areas in (a)-(b) and (f)-(g), respectively. R stands for the Pearson correlation coefficient.

(Hawaii) in 2018 at a distance of $\sim$200 km from the volcano ($\sim$1-2 ppm, Johnson et al., 2020) or for passive degassing in Mt. Etna's plume (Italy) only at 5-10 km distance (up to 1.5 ppm, Butz et al., 2017) and in the Popocatepetl's plume at a distance of $\sim$12 km ($\sim$ 1 ppm, Stremme et al., 2023). Despite the minor aerosol abundance observed during this event, both ΔXCO_2 and ΔXCO show significant agreements with aerosol enhancements (R larger than 0.80 for ΔXCO_2 and 0.85 for ΔXCO, Figure ~~7~~9(d) and (e)), suggesting that aerosol and CO_2/CO emission sources, as well as the vertical mixing of emissions and their transport patterns, are similar. We found mean $\Delta XCO_2/\Delta AOD$ ratios of 30.8$\pm$3.8 ppm/ΔAOD and 43.0$\pm$4.7 ppm/ΔAOD for ~~AERONET~~ CE318-AERONET and EM27/SUN, respectively, whereas the mean $\Delta XCO/\Delta AOD$ ratios were 90.5$\pm$7.2 ppb/ΔAOD and 118.2$\pm$11.0 ppb/ΔAOD for ~~AERONET~~ CE318-AERONET and EM27/SUN, respectively. The clear linear relationship also observed with the SO_2 total columns further confirms the volcanic origin of aerosol and trace gases emissions (R ranging between 0.80 and 0.90, with mean $\Delta CO_2/SO_2$ ratio of 48.4$\pm$4.8 and $\Delta CO/SO_2$ of 0.10$\pm$0.02).

These results are also corroborated by the significant correlation found between CO_2 and CO variations (R of 0.70), with a mean $\Delta XCO/\Delta XCO_2$ ratio of 1.86$\pm$0.32 ppb/ppm. This ratio is in the lower range of that reported for other volcanoes (1 to 120 ppb/ppm) (e.g. Oppenheimer et al., 2002; Wardell et al., 2004; Mori and Notsu, 2008; Oppenheimer et al., 2018, and references therein). From a volcanological perspective, $\Delta XCO/\Delta XCO_2$ ratios in volcanic plumes are likely to provide valuable information on the redox conditions of magma, required for degassing models (Oppenheimer et Kyle, 2008; Moussallam et al., 2014; Ilanko et al., 2014). Following Moussallam et al. (2014) and considering a magma temperature of 1140 °C at La Palma (e.g. Day et al., 2022), a $\Delta XCO/\Delta XCO_2$ ratio of 1.86 ppb/ppm would correspond to a log(fO_2) of -6.4 equivalent to FMQ +2.6. It is in agreement with the vanadium-in-olivine oximetry results reported by Day et al. (2022) for the 2021 La Palma eruption.

For the forest fire event, the estimated ΔXCO values were twice as high as those from volcanic emissions, with values reaching up to 7 ppb (i.e. $\sim$7% with respect to IZO background levels). This highlights the significant contribution of CO from combustion processes during biomass burning emissions. In fact, the CO variations inferred in this study are similar to those associated to urban/industrial environments (e.g. Madrid and St. Peterburg, García et al. (2021); Makarova et al. (2021)). Conversely, the CO_2 signal was relatively weak with enhancements smaller than 0.5 ppm (i.e. $\sim$0.04% with respect to IZO background levels). The subtle CO_2 signature during this event was also corroborated by the surface CO_2 records taken at IZO in the framework of the WMO GAW program (data not shown). As a result, the correlation between ΔXCO_2 and ΔXCO, and ΔAOD is found to be moderate (R between 0.40 and 0.50, Figure ~~7~~9(h) and (j)), with mean $\Delta XCO/\Delta XCO_2$ ratio of 5.2$\pm$2.2 ppb/ppm and $\Delta XCO_2/\Delta AOD$ ratios of 10.7$\pm$4.8 ppm/ΔAOD and 13.8$\pm$5.1 ppm/ΔAOD for ~~AERONET~~ CE318-AERONET and EM27/SUN, respectively.

Although the $\Delta XCO/\Delta XCO_2$ ratio strongly depends on the type of vegetation, the ratio values obtained in this study are significantly lower than those found in the literature. For instance, previous studies have reported $\Delta XCO/\Delta XCO_2$ ratios of $\sim$60 ppb/ppm for biomass burning emissions from tropical forests (Yokelson et al., 2007), whereas ratios ranging from 80 to $\sim$150 ppb/ppm have been attributed to boreal forest emissions (Akagi et al., 2014; Vasileva et al., 2017). It is worth highlighting that, despite ~~of~~ observing marked variability in the $\Delta XCO/\Delta XCO_2$ distribution, the $\Delta XCO/\Delta XCO_2$ ratio estimated here are comparable to those found for more diluted plumes from urban/industrial emissions using ground- (i.e. 5.9 - 6.2 in St.

Peterburg, 5.68 - 8.44 in Paris, 1.92 - 6.6 in London, 6 - 9 in Indianapolis, Makarova et al., 2021, and references therein) and satellite-based observations (ranging from 3.7 in Dallas to 23.9 in Tehran as reported by MacDonald et al., 2023).

The relationship between CO and aerosol biomass burning emissions is more evident (R of ∼0.80), with mean $\Delta XCO/\Delta AOD$ ratios of 233±31 ppb/ΔAOD and 271±43 ppb/ΔAOD for ~~AERONET~~ CE318-AERONET and EM27/SUN, respectively. Considering CO total column amounts, the $\Delta CO/\Delta AOD$ ratios for ~~AERONET~~ CE318-AERONET and EM27/SUN are calculated to be 1.57±0.05 molec cm^{-2}/ΔAOD and 1.85±0.06 molec cm^{-2}/ΔAOD, respectively. These values are in good agreement with the ratio of 1.50±0.10 molec cm^{-2}/AOD (at 500 nm) reported by Paton-Walsh et al. (2005) for Australian forest fires.

The moderate agreement found for the forest fire event compared to volcanic emissions could be attributed, in part, to the proximity of the biomass burning source to IZO, resulting in more temporally and spatially variable plumes. Additionally, these events coincided with the presence of mineral dust aerosols at IZO, which contributed to the AOD records recorded at the observatory (see Figure ~~7~~9(f)). Therefore, this complex mixture of scenes could account for part of the observed variability.

6 Summary and Conclusions

This study utilized the low-resolution EM27/SUN FTIR ~~(Fourier Transform Infrared)~~ spectrometer to obtain columnar spectral aerosol information in eight micro-windows in the infrared spectral range (873-2314 nm) under different conditions at the Izaña Observatory (IZO, Tenerife, Spain). The Langley-Plot ~~calibration~~ procedure was applied to calibrate the instrument, not only to provide ~~aerosol optical depth (AOD)~~ AOD retrievals but also to evaluate the radiometric stability of the EM27/SUN system. It is important to note that this Langley-Plot calibration is dispensable for standard trace gas retrievals since high-resolution solar absorption spectra are self-calibrating in the sense that the absorption signature is referenced to the surrounding continuum. The temporal degradation observed in the calibration coefficients $V_{0,\lambda}$ was relatively low compared to that of the high-resolution IFS 125HR system at the same station, ranging between ~~0.56~~0.66%month^{-1} (B1-B2) and ~~0.66~~0.78%month^{-1} (B6-B8), with an all-band mean of ~~0.62~~0.72%month^{-1} (σ of ~~0.04~~0.27%month^{-1}).

The ~~study~~ assessment of uncertainty in the EM27/SUN AOD retrievals (U_{AOD}) was conducted in this study through the use of the Monte-Carlo method. This approach encompassed the propagation of uncertainties linked to the parameters and measurements integral to the AOD retrieval process. Our results revealed expected uncertainties of 0.012, 0.013, and 0.009 for 870, 1020, and 1640 nm spectral bands, respectively.

The study successfully retrieved EM27/SUN spectral AOD in the infrared region, which was validated against the Cimel 318-T radiometer (as part of AERONET). ~~Although a~~ A 3-year period of synchronous ~~AERONET~~ CE318-AERONET and EM27/SUN measurements was available to verify the coherence between the two data sets~~, a final time period of two years (2021-2022) was used to overcome the lack of calibration due to the lockdown associated with the~~ . The complete time span covered both an initial phase (2019-2020) characterized by a significant Langley-Plot calibration gap resulting from the COVID-19 ~~restrictions in 2020.~~ lockdown restrictions, in addition to a subsequent two-year interval (2021-2022) marked by a notable increase in the calibration frequency. In this regard, our results ~~emphasize the importance of maintaining a regular calibration frequency, ideally at least once per month,~~ demonstrate that the calibration approach used in this paper based on

Langley-plot regular calibrations and smoothing spline functions to cover the calibration gaps is adequate to compensate for the optical degradation of the system. Other possible solutions to address this issue could involve the design of protective domes to prevent system degradation during operation, or other absolute radiometric calibration procedures, such as using
575 high-intensity calibration sources or robust calibration transfers, as already implemented during sporadic field campaigns by Gardiner et al. (2012); Menang et al. (2013); Elsey et al. (2017). The AERONET-COCCON cross-validation showed average AOD differences limited to 0.005, with Pearson regression coefficients reaching up to ~~0.99~~0.967. Based on the World Meteorological Organization (WMO) traceability criteria (~~$U95$~~U_{95}), EM27/SUN AOD was traceable to ~~AERONET in the 50%, 71~~CE318-AERONET in the 52.0%, 71.1%, and ~~84~~83.5% for the 870, 1020, and 1640 nm coincident spectral bands, respec-
580 tively. ~~AOD uncertainty (for average air mass of 1.6) of the EM27/SUN product was estimated to range between 0.023 (for B1) and 0.009 (for B3) with a 95% level of statistical significance.~~

The study also checked the spectral coherence between the EM27/SUN AOD retrievals by using a correlation matrix analysis, which showed a better spectral consistency in high-turbidity conditions between the bands within the 873-1021 nm (Kendall correlation coefficient, $R_K \sim 0.90$), 1238-1636 nm ($R_K = 0.92$-0.98), and 2133-2314 nm ($R_K = 0.93$-0.98) spectral ranges.

585 Two case studies were selected to demonstrate the potential of the EM27/SUN in improving the monitoring and understanding of atmospheric processes, due to its ability to simultaneously retrieve column-integrated aerosol and trace gas information. These events were chosen when aerosols and trace gases shared similar emission sources, allowing AOD information to be used as a proxy to trace gas concentrations and vice-versa, and providing a comprehensive perspective of events. One case study was a volcanic plume that impacted IZO on 11st October 2021 from the La Palma eruption, and the other was a forest
590 fire that also affected the station on 21st July 2022. The analysis revealed a strong correlation between the enhancements of aerosols and trace gases (AOD, CO_2, and CO), particularly for the volcanic case, which may be attributed to the well-mixed volcanic plume reaching IZO.

In conclusion, EM27/SUN observations can enhance our understanding of atmospheric chemistry during events that generate high concentrations of particles or gases, such as dust events, forest fires, urban/industrial or volcanic emissions. Simultaneous
595 measurements of gases and aerosols are crucial, for example, to understand degassing processes in volcanoes, monitor in-plume chemistry, and assess the impact of aerosols on remote gas measurements. The capability of the EM27/SUN to provide aerosol information in the NIR and SWIR spectral ranges can also be used to improve our knowledge of the radiative effect of large aerosols on climate. This portable instrument is highly versatile and can be deployed at numerous stations worldwide to meet specific measurement needs. Therefore, it has the potential to serve as a crucial tool for densifying current ground-based
600 networks for observing aerosols and gases, as well as for validating satellite-based gas and aerosol products.

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Competing interests. At least one of the (co-)authors is a member of the editorial board of Atmospheric Measurement Techniques/Atmo-
615 spheric Chemistry and Physics.

Data availability. The AERONET–Cimel data from the Izaña Observatory ("Izaña") are available on the AERONET website: https://aeronet.
gsfc.nasa.gov/cgi-bin/data_display_aod_v3?site=Izana&nachal=2&level=3&place_code=10 (Holben et al., 1998). The EM27/SUN and IFS
125HR data might be available upon request to the corresponding authors.

Appendix A:

620 Abbreviations
ACTRIS Aerosols, Clouds, and Trace gases Research Infrastructure
AE Angström Exponent
AEMET State Meteorological Agency of Spain
AERONET Aerosol Robotic Network
625 AOD Aerosol Optical Depth
CIMO Commission for Instruments and Methods of Observation
COCCON COllaborative Carbon Column Observing Network
CV Coefficient of Variation
DU Dobson Units
630 EMPIR European Metrology Program for Innovation and Research
ERF Effective Radiative Forcing

ESA European Space Agency

EURAMET EURopean Association of national METrology institutes

FWHM Full Width at Half Maximum

635   GAW Global Atmospheric Watch

GAW-PFR Precision Filter Radiometer

GCOS Global Climate Observing System

GHG GreenHouse Gases

GUM Guide to the Expression of Uncertainty in Measurement

640   FOV Field Of View

FTIR Fourier Transform Infrared

IARC Izaña Atmospheric Research Center

InGaAs Indium Gallium Arsenide

IPCC Intergovernmental Panel on Climate Change

645   IR InfraRed

IZO Izaña Observatory

KIT Karlsruhe Institute of Technology

MAPP Metrology for Aerosol Properties

MD Mean Deviation

650   MCM Monte-Carlo method

NASA National Aeronautics and Space Administration

NDACC Network for the Detection of Atmospheric Composite Change

NIR Near-Infrared spectral Region

NOAA National Oceanic and Atmospheric Administration

655   OPD Optical Path Difference

PWV Precipitable Water Vapour

R_K Kendall rank correlation coefficient

RMSE Root Mean Squared Error

SI International System of Units

660   SWIR Short-Wave Infrared spectral Region

TCCON Total Carbon Column Observing Network

TOA Top Of Atmosphere

TROPOMI Tropospheric Monitoring Instrument

UV UltraViolet spectral region

665   VIS Visible region

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