



On the temperature stability requirements of free-running Nd:YAG lasers for atmospheric temperature profiling through the rotational Raman technique

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Abstract. We assess the temperature stability requirements of unseeded Nd:YAG lasers in lidar systems for atmospheric temperature profiling through the rotational Raman technique. Taking as a reference a system using a seeded laser assumed
15 to emit pulses of negligible spectral width and wavelength-drift free, we estimate first the effect of the pulse spectral widening of the unseeded laser on the output of the interference filters, then we derive the limits of the allowable wavelength drift for a given bias in the temperature measurement that would add to the noise-induced uncertainty. Finally, using spectroscopic data, we relate the allowable wavelength drift to allowable temperature variations of the YAG rod. We find that, in order to keep the bias affecting atmospheric temperature measurements smaller than 1 K, the Nd:YAG rod
20 temperature should also be kept within 1 K.

1 Introduction

Profiles of humidity and temperature throughout the troposphere are of the utmost importance to weather forecast(Statement of Guidance for Global Numerical Weather Forecasting (GNWP), 2023). Radiosondes are the staple way to provide such profiles, but they cannot do it in a continuous or quasi-continuous basis, except for short periods of time during campaigns.
25 Raman lidars offer an alternative to retrieve these profiles with relatively high temporal resolution without spending balloons and their payloads. With respect to temperature profiles, the Raman lidar technique is based on the dependence on temperature of the intensity of the atmospheric N₂ and O₂ rotational Raman lines (Cooney, 1972). In the observed range of atmosphere temperatures, the backscatter cross section of the Raman lines with low rotational quantum numbers tends to decrease when the temperature increases, while that of lines with high quantum numbers has the opposite behavior.
30 Therefore, the power ratio of the backscattered radiation in two spectral regions provides information on the temperature of the backscattering volume.



In the past, the separation of the backscatter of different regions of the pure rotational Raman spectrum has been implemented through the use of diffraction gratings (Arshinov et al., 1983). The current state of the art permits the use of interference filters to do that separation (Vaughan et al., 1993; Behrendt and Reichardt, 2000).

35 The power ratio of the filter outputs corresponding to the different regions of the spectrum needs to be calibrated to retrieve the atmospheric temperature. This is usually done by comparison with a radiosonde profile (Hammann et al., 2015), although in-situ calibrations using local light sources (Vaughan et al., 1993), or by constructing a model of the power ratio between the outputs of the filters for high- and low-quantum numbers (Mahagammulla Gamage et al., 2019) are also described in the literature. For the calibration to be valid over long periods, the stability of the system to ensure that the
40 filters' response keeps constant in time is of paramount importance. Modern interference filters can have temperature dependences between 2 pm/°C and 5 pm/°C (Johansen et al., 2017) and, being usually kept in a controlled environment, where temperature can be easily stabilized, their response stability is usually not an issue. Of more concern regarding system stability is the laser wavelength. Although some systems use free-running lasers (Di Girolamo et al., 2004; Mahagammulla Gamage et al., 2019), laser frequency stability is achieved in many systems through the use of injection-seeded lasers, where
45 the stability and spectral purity of the seeder is transferred to the host laser (Behrendt and Reichardt, 2000; Behrendt et al., 2002; Behrendt et al., 2004; Lange et al., 2019), which is forced to operate in a single longitudinal mode.

In this paper we assume that the receiver is in a well-controlled environment, such that the wavelength drift of the filters can be neglected, and focus on the wavelength stability requirements for a non-seeded laser to be used in a lidar measuring atmospheric temperature profiles using the pure rotational Raman spectra of N₂ and O₂ under the excitation by the third
50 harmonic of a Nd:YAG laser. In section 2, the short-term spectral widening of a free-running frequency tripled Nd:YAG laser is taken into account. Section 3 considers the far more important effect of temperature variations in the YAG rod. Section 4 presents the conclusions of this work.

2 Effect of short-term emission spectral widening

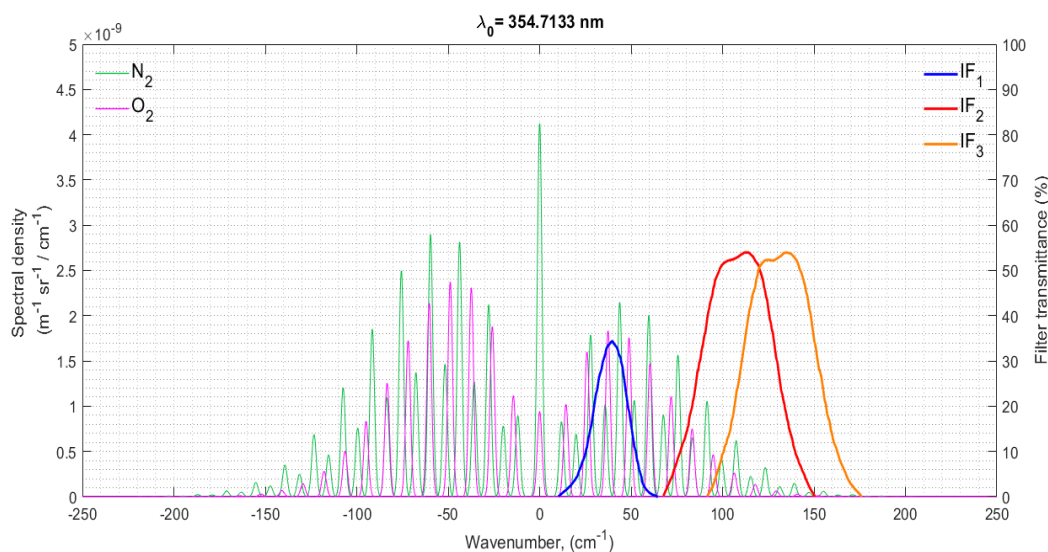
We assume that a free-running Nd:YAG laser oscillates in the most intense line, i.e. the R2 → Y3 line (Kushida, 1969) at
55 1064 nm. This line has a gain full width at half maximum of approximately 5 cm⁻¹ (Verdeyen, 1989; Sato and Taira, 2012). As only the modes closest to the maximum gain can oscillate, the emission linewidth is considerably narrower. We will assume a typical emission linewidth of the free-running laser of 1 cm⁻¹ at 1064 nm (in fact, references (Armandillo et al., 1997; Lumibird, 2018) give a somewhat lower figure of approximately 0.7 cm⁻¹), which would result in 3 cm⁻¹ at the third harmonic wavelength. Although this linewidth is made of many competing modes, we will assume that its average over a
60 sufficient number of pulses can be modeled by a Gaussian curve (Armandillo et al., 1997). Under the excitation of a widened emission, the Raman backscatter spectrum lines are widened accordingly, therefore we will consider that each line of backscatter differential cross-section σ_i will show a profile as a function of frequency given by



$$f_{\sigma_i}(\nu) = \frac{2\sqrt{\ln 2}\sigma_i}{\sqrt{\pi}\Delta\nu} \exp\left[-\frac{4\ln 2(\nu - \nu_{oi})^2}{(\Delta\nu)^2}\right], \quad (1)$$

with ν_{oi} the central emission-line frequency and $\Delta\nu$ the full width at half maximum of the excitation spectrum.

65 For the purpose of comparison, we will consider the filter responses like those in (Hamann et al., 2015). Figure 1 shows the widened pure rotational Raman backscatter spectrum (in terms of spectral density) of N_2 and O_2 at sea level (NASA, 1976) with the filters of (Hamann et al., 2015) and assuming a central wavelength of the excitation radiation of 354.7133 nm, the 3rd harmonic wavelength of a Nd:YAG laser at a rod temperature of 38 °C (see section 3). Table 1 shows the wavelengths and wavenumbers considered. The calculation is based on the spectrum frequencies and intensities used in
 70 (Zenteno-Hernández et al., 2021), which in turn rely on expressions and spectroscopic data found in (Zenteno-Hernández et al., 2021; Murphy et al., 2016; Alms et al., 2008; Buldakov et al., 1996); Long, 2002).



75 **Figure 1. Pure rotational backscatter Raman spectrum of N_2 (green lines) and O_2 (magenta lines) at sea level under the assumption of a Gaussian-widened excitation at 354.7133 nm. The spectrum is given as wavenumber shift from the excitation wavenumber. The transmission curves of the filters and combination of filters in (Hamann et al., 2015) are superimposed, transposing the wavelength scale in that reference to the wavenumber scale here. The filter denoted as IF2 (red transmission curve) is used in high background-radiation situations, whereas the IF3 filter (orange line) is optimized for low background ones. The Cabannes-line amplitudes (zero shift) are divided by 500 to fit them into the graph.**

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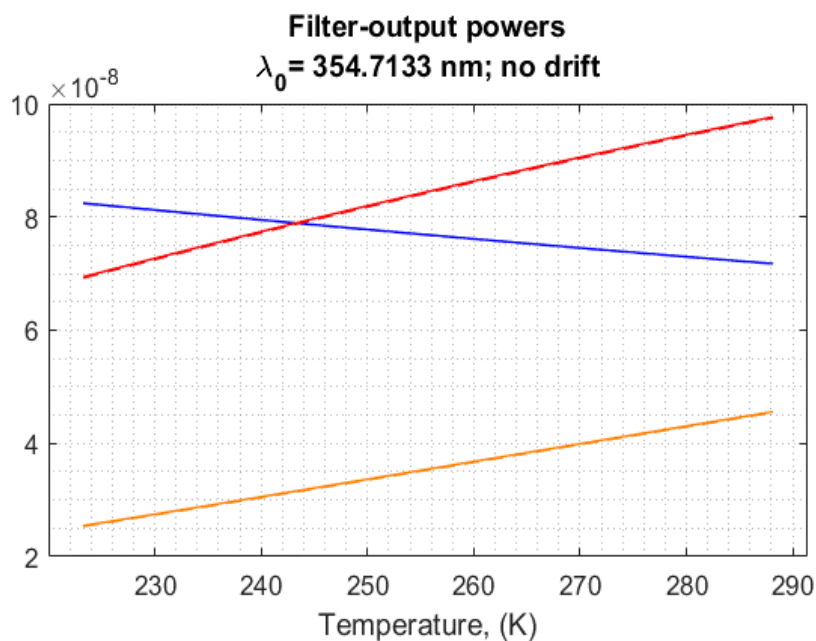


85 **Table 1. Wavelengths and wavenumbers of lasers and filter passbands considered**

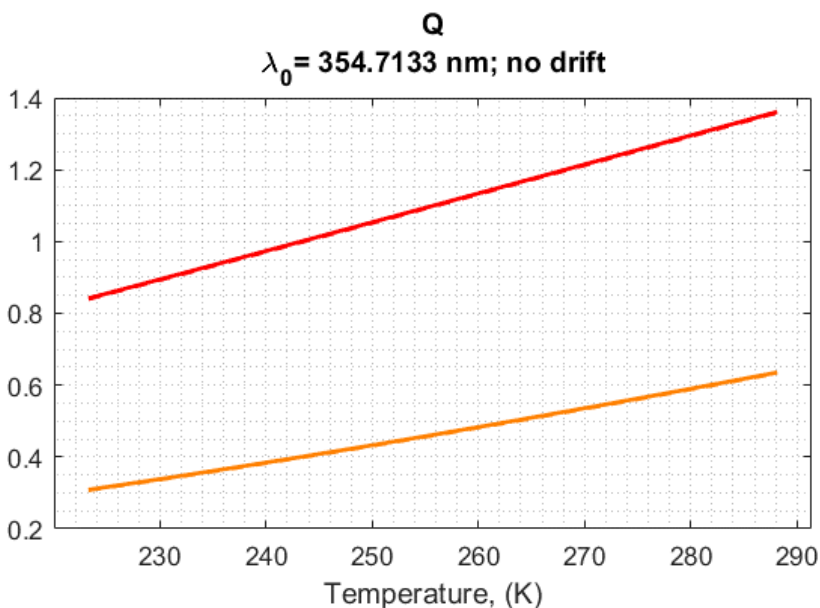
	Wavelength (nm)		Wavenumber (cm ⁻¹)	
3 rd harmonic of Nd:YAG laser fundamental wavelength	354.7133		28191.78	
	Non-shifted filters		Shifted filters*	
	HM wavelengths (nm)	HM wavenumbers (cm ⁻¹)	HM wavelengths (nm)	HM wavenumbers (cm ⁻¹)
Low quantum number filter IF1	354.0926 – 354.3488	28220.78 – 28241.20	353.9852 – 354.2412	28229.35 – 28249.77
High quantum number filter IF2	353.0767 – 353.6324	28277.95 – 28322.46	352.9699 – 353.5253	28286.52 – 28331.03
High quantum number filter IF3	352.7945 – 353.3415	28301.23 – 28345.11	352.7945 – 353.3415	28309.80 – 28353.68

* See section 3. A shift of 8.57 cm⁻¹ is considered in the filter passbands with respect to those non-shifted.

The power at the filter outputs is shown in Fig. 2, under the assumption of a widened spectrum and of an ideal line spectrum. The effect of the Gaussian spectral widening is so small that the pairs of curves are almost undistinguishable. Likewise, the spectral widening effect is very small in the curves (Fig. 3) obtained as the power ratio between the high quantum number filter outputs (IF2 and IF3 filters in Fig. 1) and the low quantum number filter output (IF1 filter in Fig. 1), henceforth called Q-curves, which, after calibration, give the atmosphere temperature. We conclude that the emission spectral widening of a free-running laser with respect to that of a seeded laser does not have a significant impact on the system performance under the assumption of a Gaussian widening smaller than 1 cm⁻¹ at the fundamental wavelength (Armandillo et al., 1997). No leakage of the Cabannes lines into the filters (especially into the low quantum number one) is noticed in these conditions, in spite their higher cross-section (around two orders of magnitude) with respect to the next more intense lines of the rotational Raman spectrum. One should make sure that this hypothesis is satisfied in particular cases and that the presence of aerosols does not increase the return at the excitation wavelength so as to produce leakage.



100 Figure 2. Power (in arbitrary units) at the filter outputs for the widened spectrum and for an unwidened spectrum as a function of the atmosphere temperature. The colors of the curves indicate the corresponding filter according to the colors of transmission curves in Fig. 1. There are two curves for each color, but the effect of the Gaussian widening of the spectrum lines is so small that they are almost undistinguishable.



105 Figure 3. Ratio (Q) between the outputs of the “red” filter and the “blue” filter of Fig. 1 (red lines) and between the outputs of the “orange” filter and the “blue” filter (orange lines) as a function of the atmosphere temperature. The lines corresponding to the widened spectrum and those corresponding to the unwidened spectrum are almost undistinguishable.



3 Effect of central wavelength drift

110 We now turn to the effect of drift of the laser central wavelength. In an unseeded Nd:YAG laser, the central wavelength coincides with the center of the gain curve, which in turn depends on the rod temperature. We have used the approximate analytical models in (Sato and Taira, 2012) that reproduce very well experimental results of the central frequency temperature dependence of various Nd:YAG emission lines, among which the one corresponding to the R2 → Y3 transition, in the temperature range 15 °C – 350 °C. For the benefit of the reader, we reproduce here Eq. (6) of (Sato and Taira, 2012):

$$115 \quad \nu_{if}(T) = \nu_{if}(0) - c_{if} \left(\frac{T}{\Theta_D} \right)^4 \int_0^{\frac{\Theta_D}{T}} \frac{x^3}{e^x - 1} dx, \quad (2)$$

where ν_{if} is the central wavenumber, T is the temperature, Θ_D is the Debye temperature, and c_{if} is a fitting parameter.

In particular we have taken the transition wavenumber at 0 K as 9403.15 cm⁻¹, the temperature-dependence fitting parameter c_{if} as 130.6 cm⁻¹, and the Debye temperature as 795 K (Sato and Taira, 2012) (see also Table 1 of this reference).

120 Figure 4 shows the dependence of the center of the 3rd harmonic wavelength emission with the temperature according to that model. The dependence is nearly linear with a slope of approximately 1.5 pm/°C.

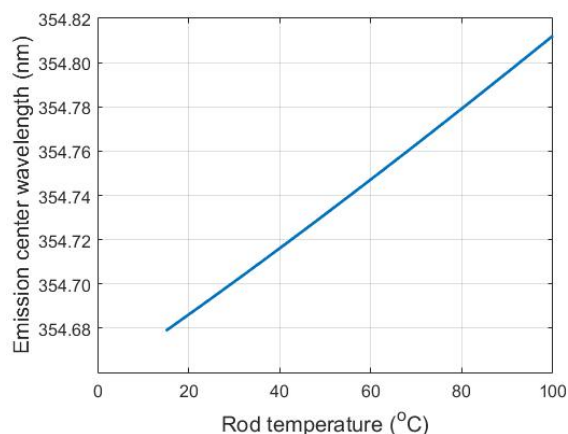


Figure 4. Dependence of the 3rd harmonic emission center wavelength on the rod temperature.

125 To assess the variations of the Q-curve with the laser central wavelength, we assume the central wavelength of 354.7133 nm and the widened spectrum of the unseeded laser already considered in the previous section, and let it drift ±1.5 pm, which would correspond to a temperature variation in the Nd:YAG rod of slightly less than ±1 °C around 38 °C. The effect of the wavelength drift on the power at the filter outputs is shown in Fig. 5, where small differences can be seen (more noticeable in the “red” filter output). However small, these differences impact the Q-curves, as shown in Fig. 6, and entrain an uncertainty in the temperature retrieval. For example, let us consider Fig. 7, which is a zoom of the right-hand panel of

130 Fig. 6 around the temperature of 260 K (with the horizontal and vertical axes swapped), where we assume that the Q-curve



has been calibrated for a laser rod temperature at 38 °C (solid curve). If the laser central wavelength drifts subsequently by ± 1.5 pm, roughly corresponding to a temperature variation of the laser rod of ± 1 °C, the uncertainty in the retrieved temperature for the approximately 0.514 value of Q is the atmosphere temperature interval encompassed by the double-headed arrowed line, i.e. slightly more than 1 K.

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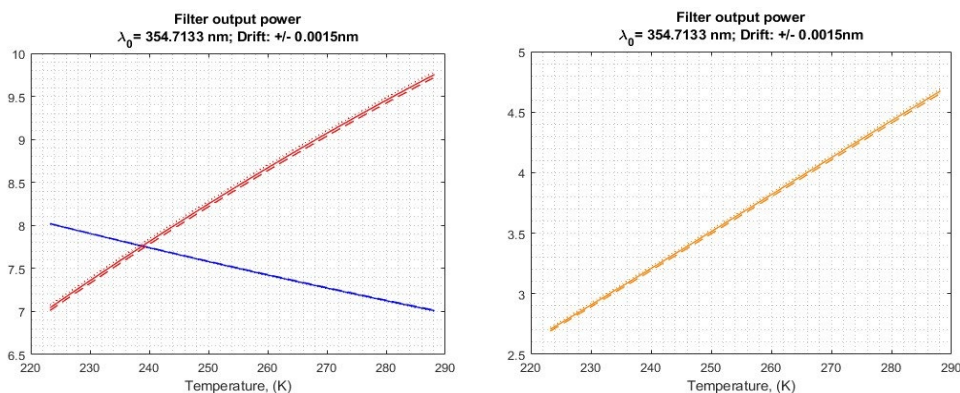


Figure 5. Effect of the wavelength drift on the power (in relative units) at the filter outputs. Left: IF1 (blue) and IF2 (red) filters of Fig. 1. Right: IF3 (orange) filter of Fig. 1. Solid lines correspond to the filter outputs at the 354.7133 nm wavelength. The dashed line corresponds to the +1.5 pm drift and the dotted one to the -1.5 pm one.

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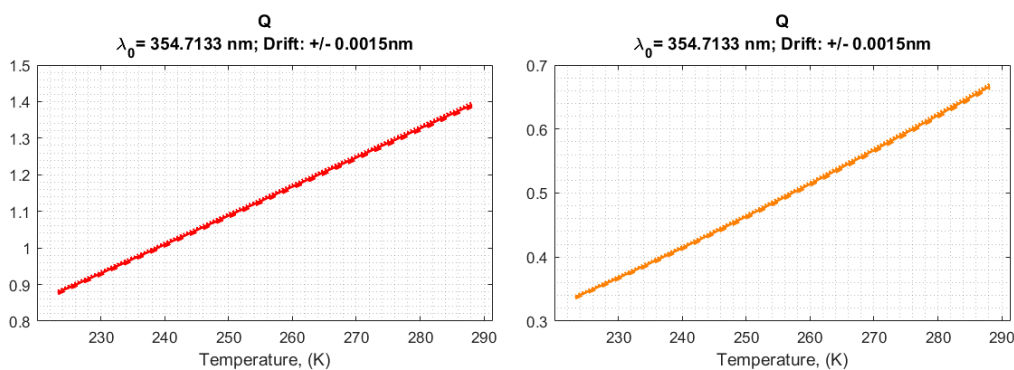
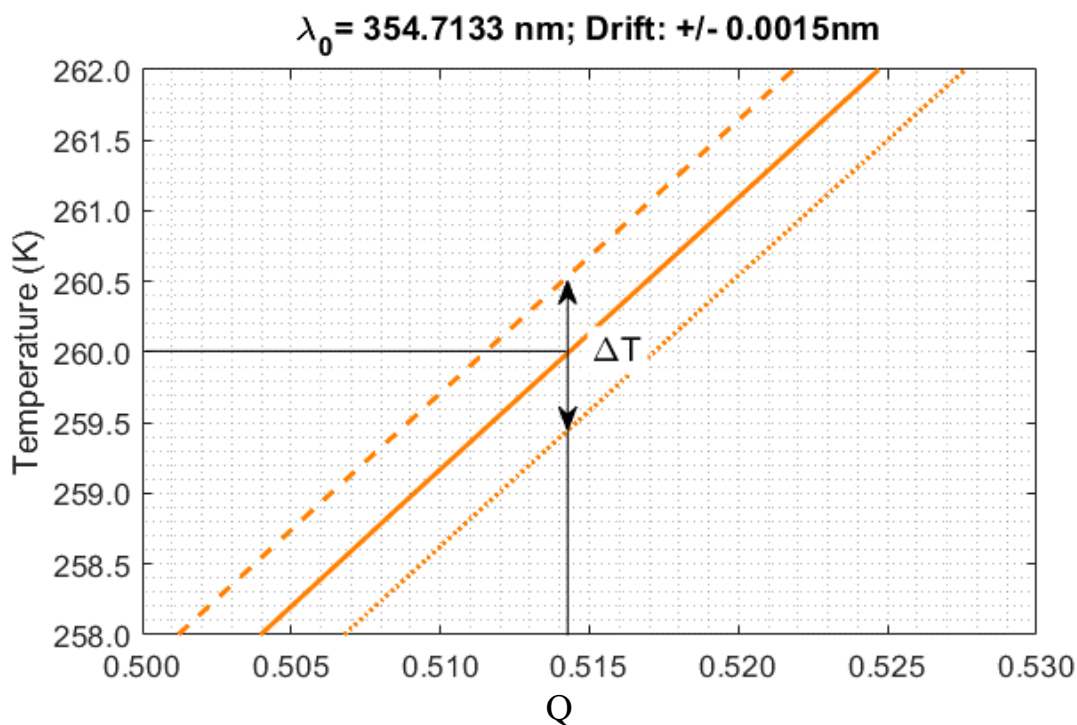


Figure 6. Effect of the wavelength drift on the Q ratio. Left: ratio between the output power of the IF2 (red) filter and the IF1 (blue) filter of Fig. 1. Right: ratio between the output power of the IF3 (orange) filter and the IF1 (blue) filter of Fig. 1. Solid lines correspond to the filter outputs at the 354.7133 nm wavelength. The dashed line corresponds to the +1.5 pm drift and the dotted one to the -1.5 pm one.

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150 **Figure 7. Example of uncertainty in the temperature retrieval (ΔT) due to the uncertainty in the laser central wavelength. Solid curve: assumed calibrated Q obtained as the ratio between the output powers of the IF3 (orange) filter and the IF1 (blue) filter of Fig. 1 at the 354.7133 nm wavelength; dashed and dotted curves: Q obtained for wavelength deviations with respect to the calibration wavelength of +1.5 pm and -1.5 pm respectively. Note that the axes are swapped with respect to those in Fig. 6.**

155 This assessment has been done for the laser central wavelength of 354.7133 nm wavelength and a drift of ± 1.5 pm around it. However, in (Hammann et al., 2015), the laser wavelength was 354.83 nm, fixed by the seeder wavelength. If we assume that the filter responses in Fig. 1 can be shifted by 116.7 pm (i.e. around 9 cm^{-1}) towards shorter wavelengths while preserving their shape, in order to be in a similar filter passbands situation with respect to the Raman spectrum as in (Hammann et al., 2015), we obtain, for the ratio of IF3 (orange) filter output power to the IF1 (blue) filter output power, the

160 curves of Fig. 8, where the uncertainty in the retrieval of the temperature around 260 K because of a non-controlled ± 1.5 pm laser wavelength drift from the calibration wavelength is smaller than 1 K. Note that we have implicitly assumed trailing zeros in the wavelength of the laser in (Hammann et al., 2015), which is given only to the hundredth of nm; what is important however is that by shifting the filters by around 9 cm^{-1} we operate in a region where the sensitivity of the Q -curve to wavelength drifts is smaller.

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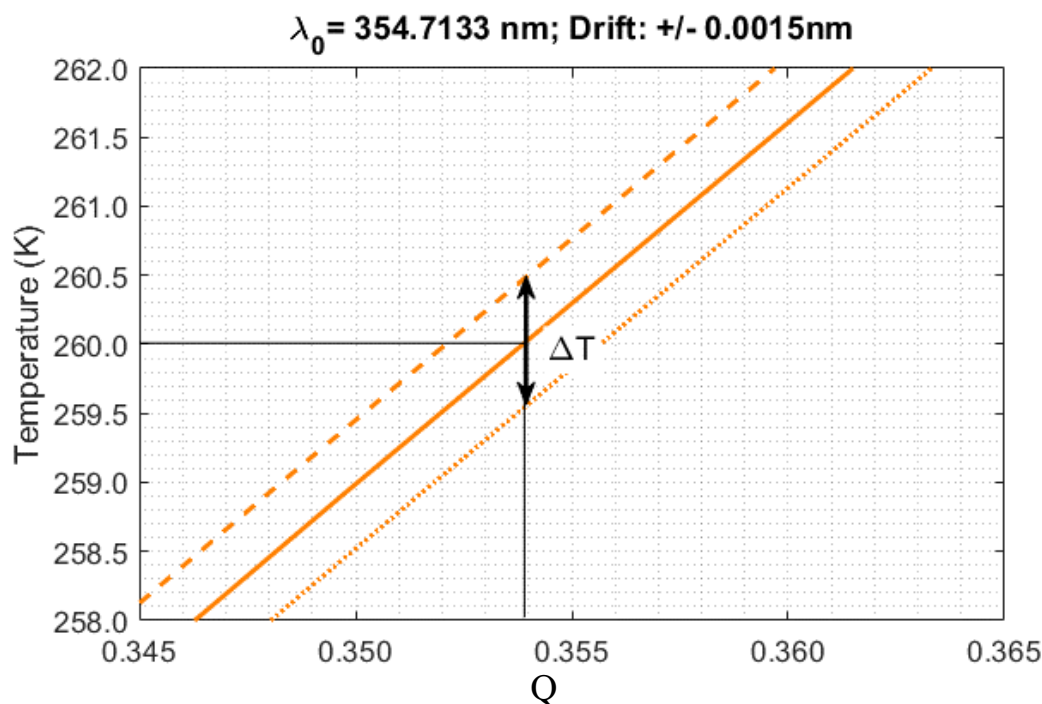


Figure 8. Like Fig. 7 with the filters of Fig. 2 shifted 8.57 cm^{-1} towards shorter wavelengths.

170 The uncertainty in the temperature retrieval due to a $\pm 1.5 \text{ pm}$ wavelength drift with respect to the calibration wavelength has been estimated as in the examples above illustrated in Figs. 7 and 8 for all the considered temperature range, and is shown in Fig. 9 for the case of the filters shifted by 8.57 cm^{-1} . The uncertainty varies between approximately 0.87 K for a 220 K temperature and 1.22 K at 290 K considering that the calibrated Q-curve has been obtained as the ratio of the output power of the IF2 filter of Fig. 1 to the output power of the IF1 filter. For the Q-curve given by the ratio of the output power of the IF3 filter of Fig. 1 and the IF1 filter, the uncertainty ranges between 0.77 K at 220 K and 1.07 K at 290 K .

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4 Conclusions

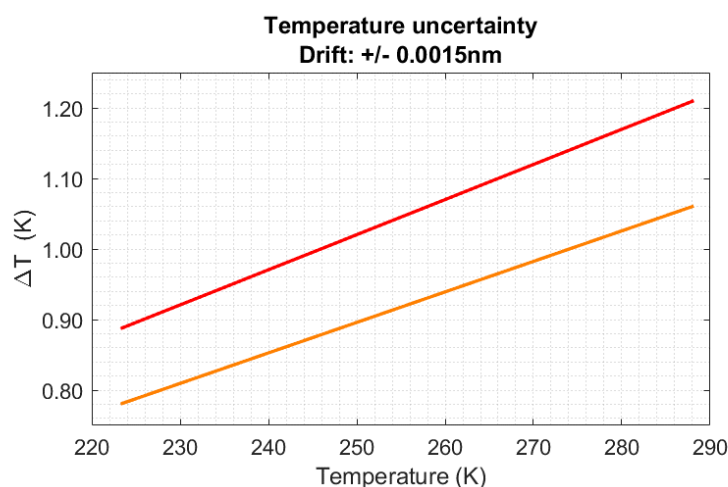
We have estimated, using a real filter configuration, that, in lidar systems employing a Nd:YAG laser to measure the atmosphere temperature profiles through the pure rotational Raman technique, the short-term spectral widening of a commercial-grade unseeded laser virtually does not affect the ratio of the output power of the high quantum-number filter to the output power of the low quantum-number filter (Q-curves), at least under the assumption that the widening is Gaussian (Armandillo et al., 1997) with a FWHM smaller than 1 cm^{-1} at the fundamental wavelength. One should pay attention that aerosols do not increase the elastic return to the point that leakage into the filters can happen. However, slight drifts on the central wavelength of the laser emitted spectrum entail small changes in the Q-curves that impair the calibration and cause an uncertainty in the retrieved atmosphere temperature. The drifts can be related to temperature changes of the YAG rod. For

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185 the 3rd harmonic of the laser, the drift is around $1.5 \text{ pm}/^\circ\text{C}$. The allowable uncertainty in the temperature retrieval depends of course on the application, but if we take as a reference value $\pm 1 \text{ K}$, it turns out that the YAG rod temperature needs also to be kept within $\pm 1 \text{ K}$. Note “en passant” that, in a seeded laser, changes of temperature in the seeder could also cause wavelength drifts, hence uncontrolled biases in the atmosphere temperature measurements that would add to their uncertainty.

190 It must be stressed that the analysis above takes only into account the temperature uncertainty due to the spectral behavior of the laser radiation. To this, the uncertainty due to the limited signal-to-noise ratio in the detected power should be added (Behrendt and Reichardt, 2000; Di Girolamo et al., 2004).



195 **Figure 9.** Retrieved atmosphere temperature uncertainty due to a $\pm 1.5 \text{ pm}$ wavelength drift over all the considered temperature range for Q obtained as the ratio between the output powers of the IF2 (red) filter and the IF1 (blue) filter of Fig. 1 (red curve) and as the ratio between the outputs of the IF3 (orange) filter and the IF1 (blue) filter in the same figure (orange curve). The filter responses have been shifted by 8.57 cm^{-1} .

200 Author contribution

JAZ: formal analysis, software; AC: conceptualization, writing; FD: methodology, formal analysis; AR: visualization, supervision; CM: software, visualization; MS: visualization; NF: validation; AB: supervision, visualization; PDG: supervision, validation.

Competing interests

205 The authors declare that they have no competing interests.



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